

# Detection of axion dark matter



**Erika Garutti**  
Uni. HH

on behalf of the  
MADMAX  
collaboration

Towards the

# Detection of axion dark matter



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# Strong CP problem

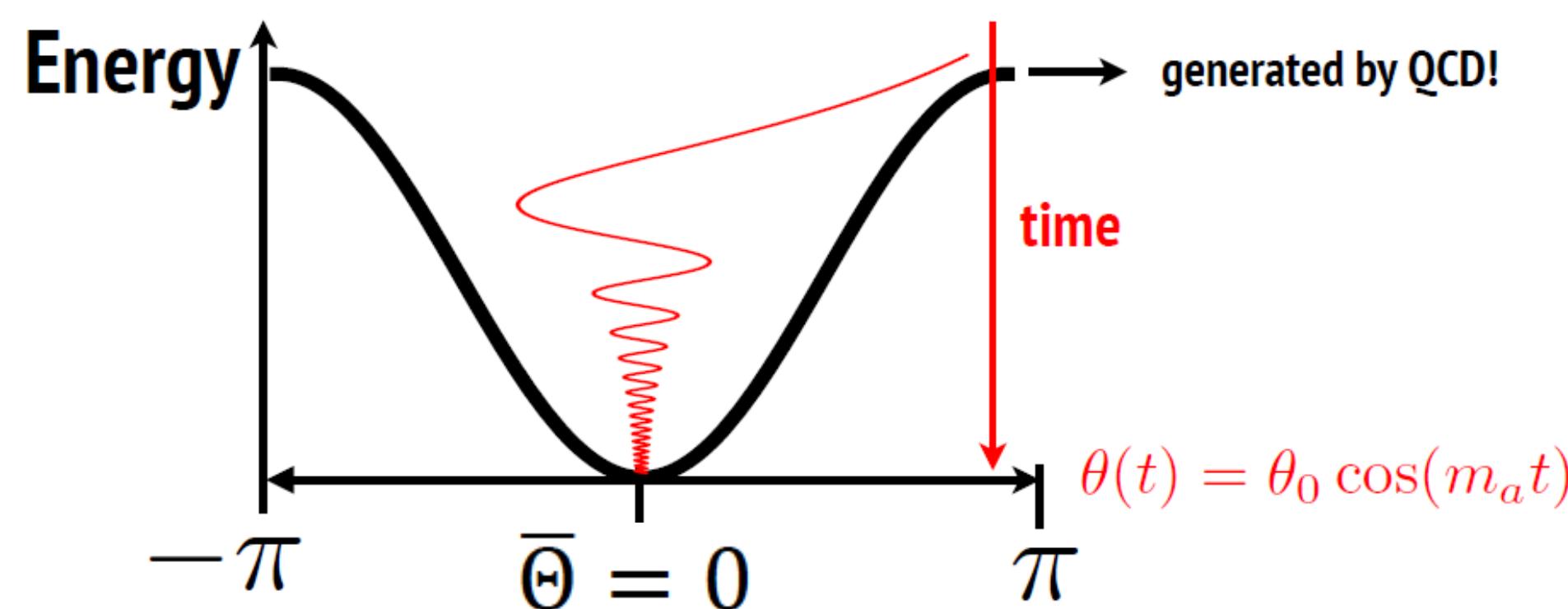


Strong force invariant under CP

$$\mathcal{L}_{\text{QCD}} = \sum_q \bar{\psi}_q (iD - m_q) \psi_q - \frac{1}{4} G_{\mu\nu a} G_a^{\mu\nu} - \overline{\Theta} \frac{\alpha_s}{8\pi} G_{\mu\nu a} \tilde{G}_a^{\mu\nu}$$

→ violates T reversal AND Parity  
→ CP violating term induces **electric dipole moment** of neutron (EDM):  
 $d \sim \overline{\Theta} \cdot 10^{-16} \text{ e cm}$

Peccei-Quinn symmetry breaking @  $T \sim f_a$  (very early universe,  $f_a > 10^9 \text{ GeV}$ )



# Strong CP problem



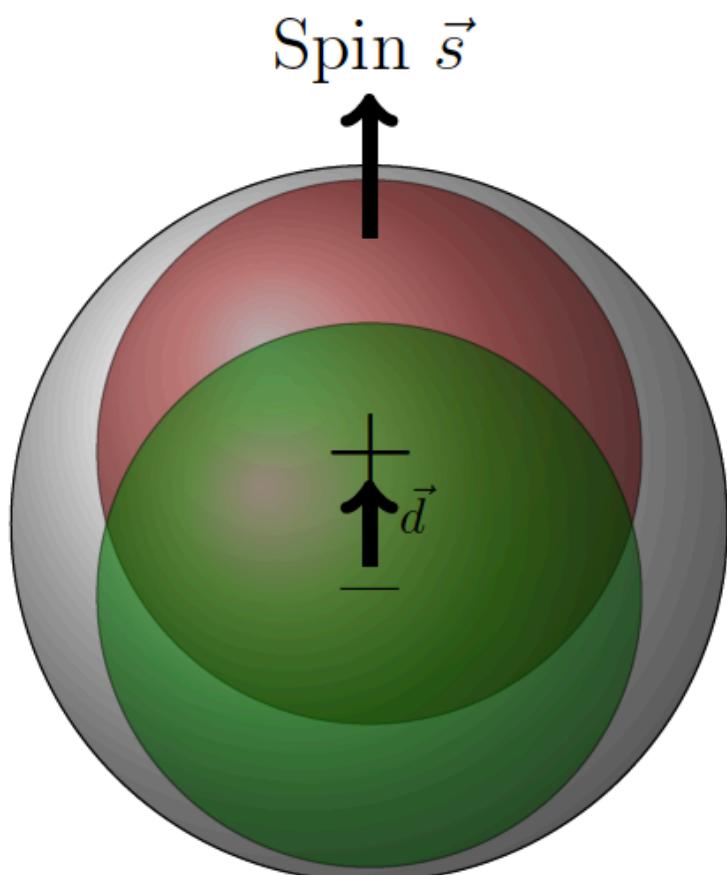
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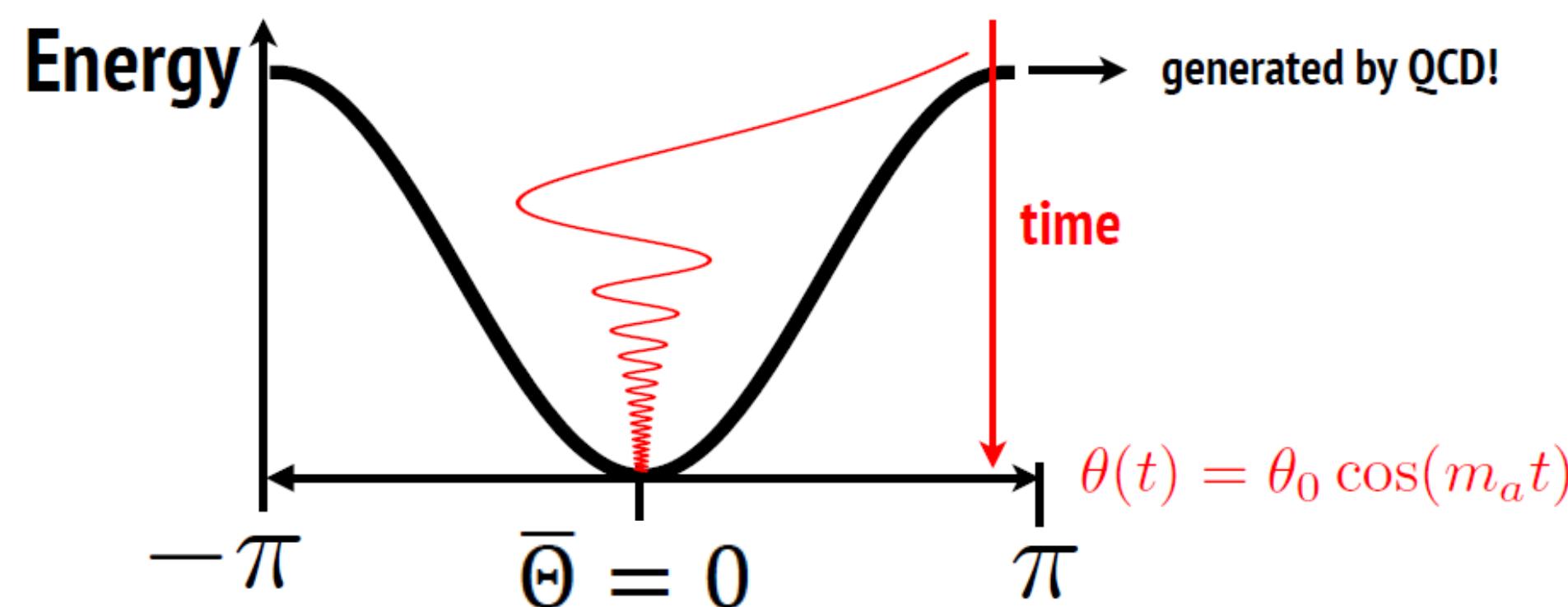
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## EDM:

- permanent separation of positive and negative charge
- fundamental property of particles (like magnetic moment, mass, charge)
- existence of EDM only possible via violation of time reversal  $T = CP$  symmetry
- has nothing to do with electric dipole moments observed in some molecules (e.g. water molecule)
- close connection to “matter-antimatter” asymmetry
- axion/ALP field leads to oscillating EDM

# Strong CP problem



Strong force invariant under CP

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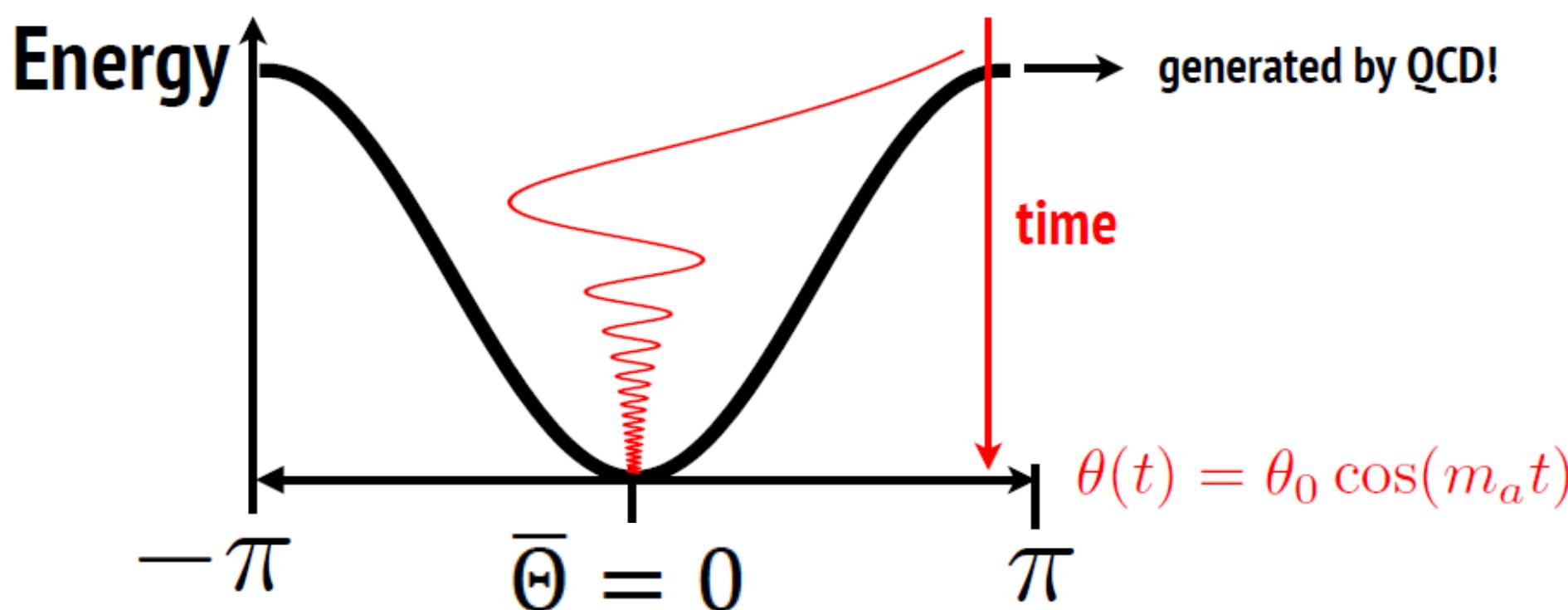
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→ CP violating term induces **electric dipole moment** of neutron (EDM):

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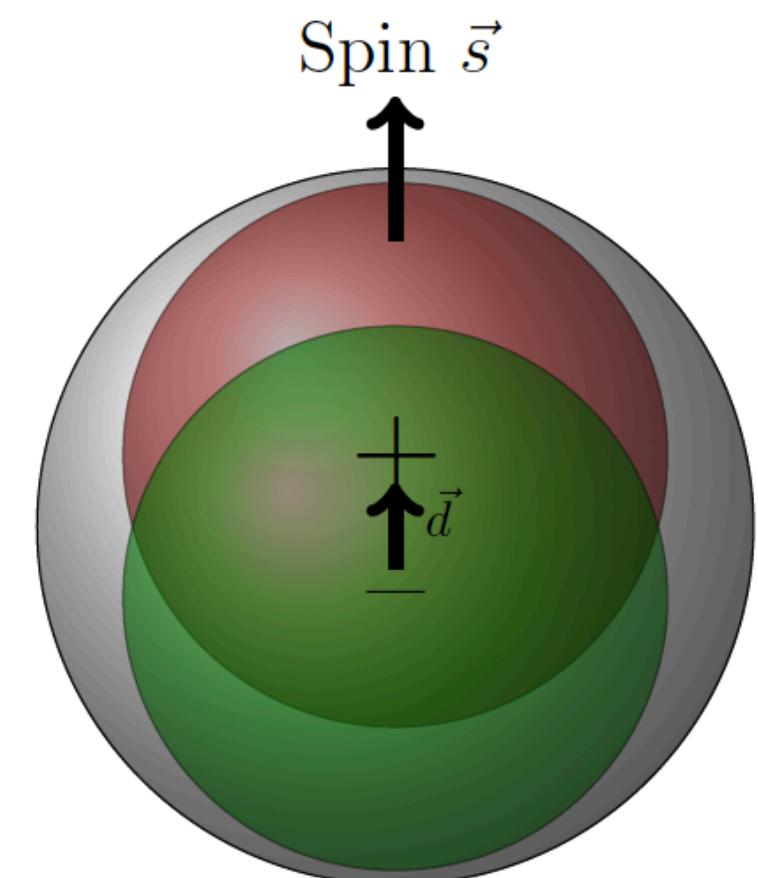
But experimentally:  $d < 10^{-26} \text{ e cm}$  →  $\overline{\Theta} < 10^{-10}$  WHY SO SMALL !?!

Peccei-Quinn symmetry breaking @  $T \sim f_a$  (very early universe,  $f_a > 10^9 \text{ GeV}$ )



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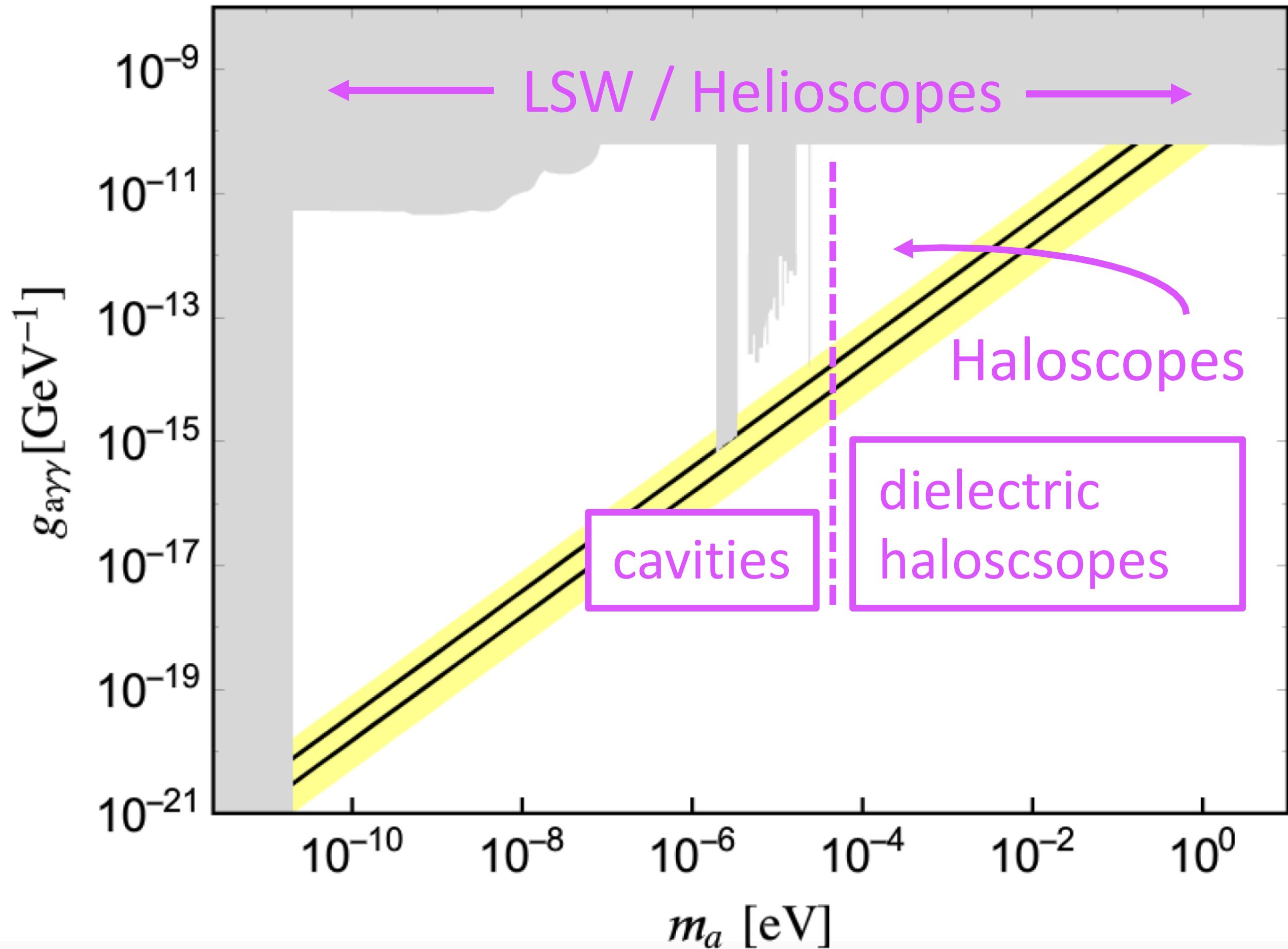


# Dark matter axion search



- Axion: arises from Strong CP problem via Peccei-Quinn mechanism
- Pseudo Nambu-Goldstone boson
- Axion can couple to two photons:  

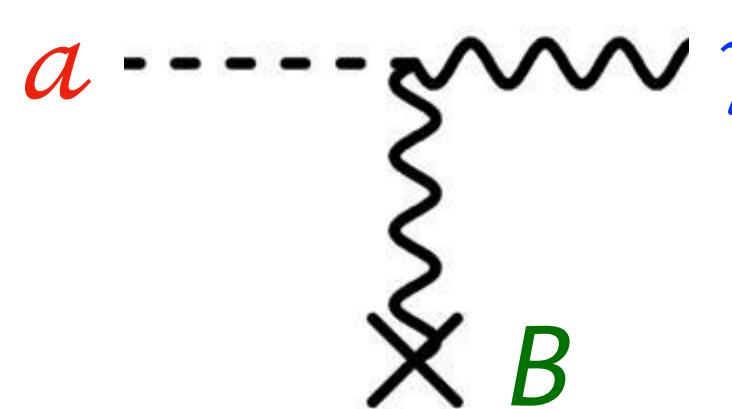
**“Sikivie process”**
- Axion model:  $m_a \propto g_{a\gamma\gamma}$  (axion-photon coupling)
- Axion can explain (part of) Cold Dark Matter



# Dielectric Haloscope



- In an external **B-field** the **axion** sources an oscillating **E-field**



$$E_a = -\frac{g_{a\gamma\gamma} B_e}{\epsilon} a$$

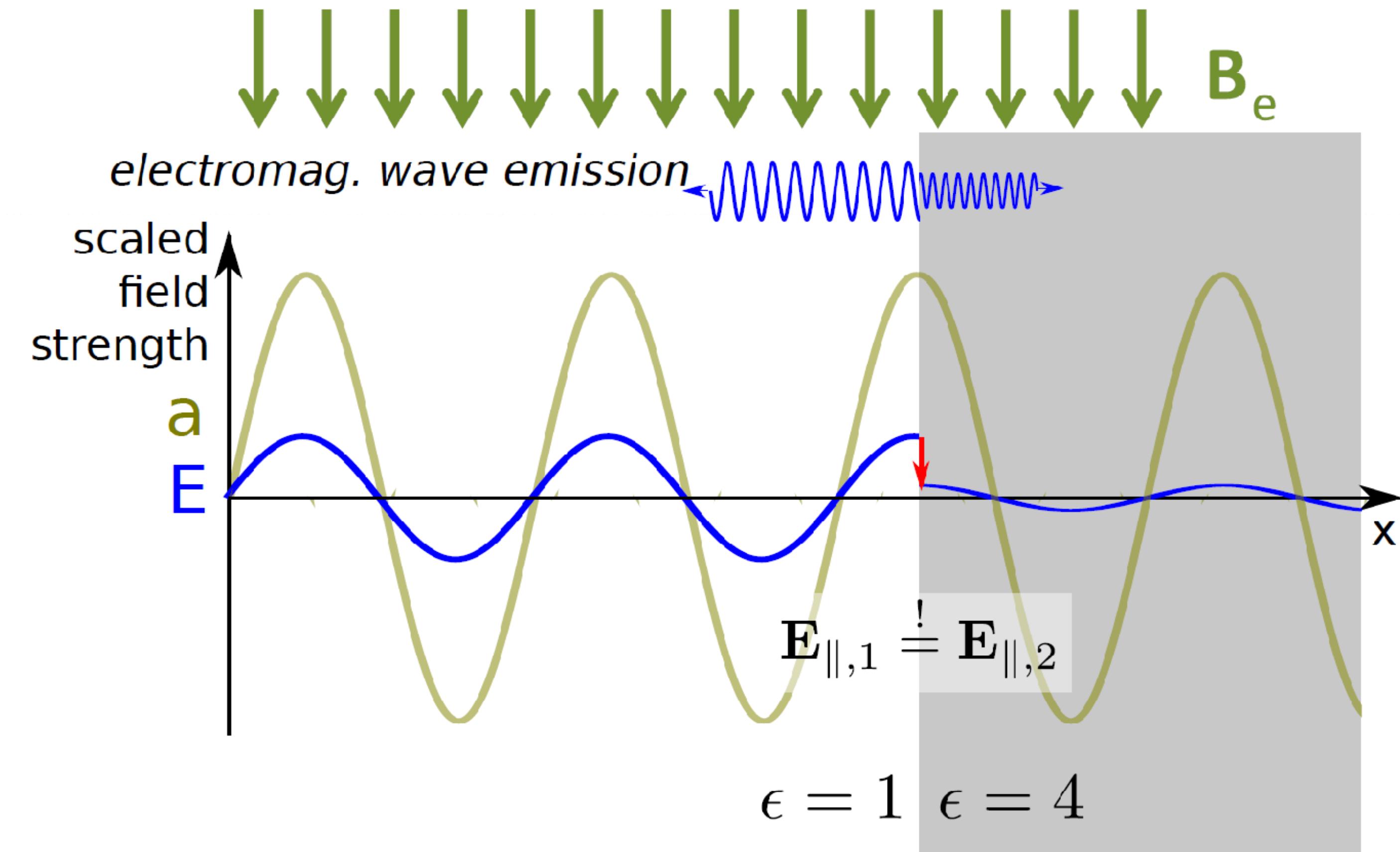
- At surfaces with transition of  $\epsilon_1 \neq \epsilon_2$ : E-field must be continuous  
→ **Emission of photons**

**Photon power :**

$$\frac{P}{A} = 2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} C_{a\gamma\gamma}^2 \left( \frac{B}{10\text{T}} \right)^2$$

$$C_{a\gamma\gamma}^2 \propto g_{a\gamma\gamma}^2$$

$$O(C_{a\gamma\gamma}^2) = 1$$



Based on the original idea of:  
D. Horns, J. Jaeckel, A. Lindner, A. Lobanov, J. Redondo and A. Ringwald  
JCAP 1304 (2013) 016 [arXiv:1212.2970].

# Dielectric Haloscope



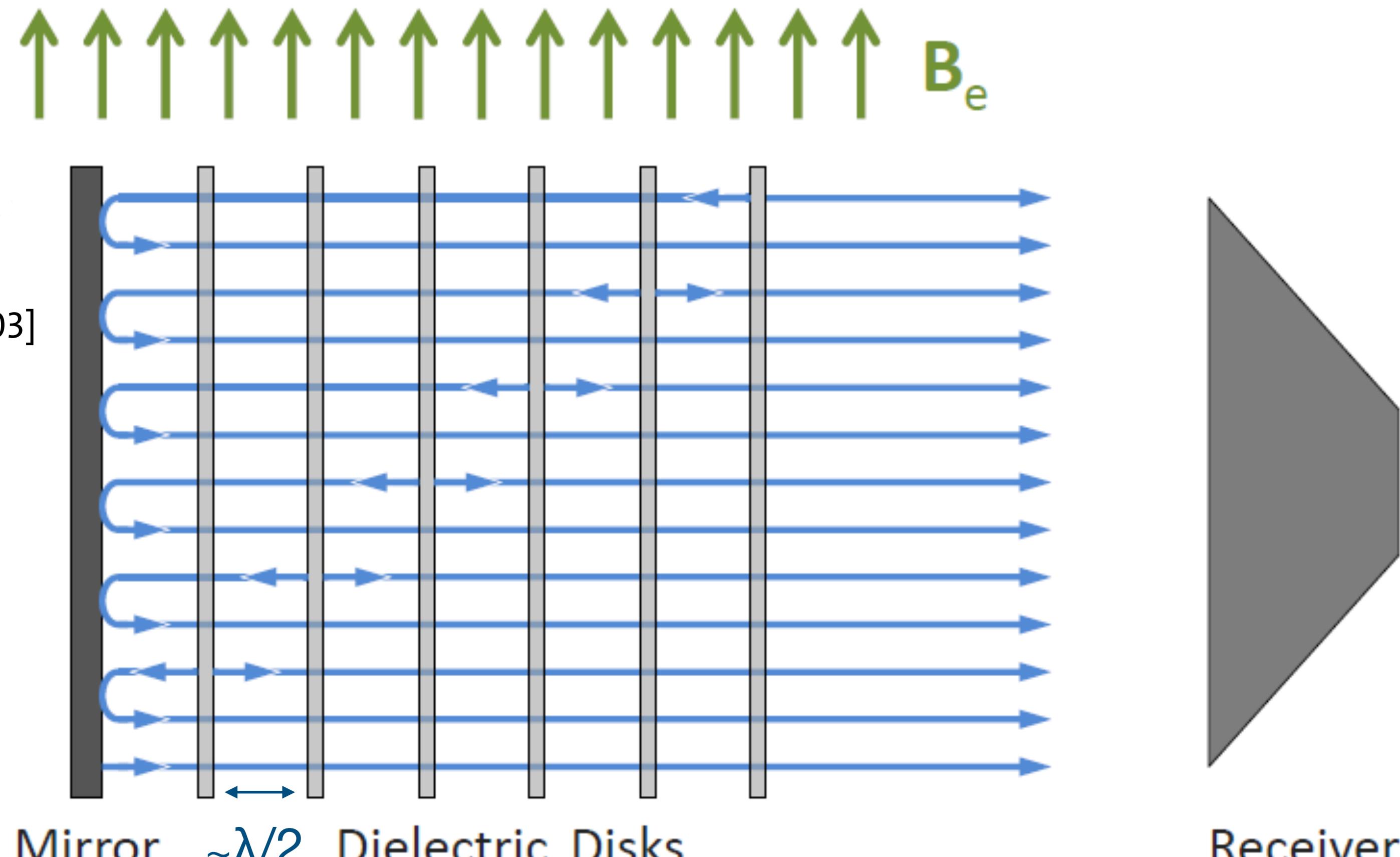
- Boost the power by coherent interference of photons generated on N discs plus resonance between discs

Jaeckel and J. Redondo, Phys. Rev. D 88, 115002, (2013) [arXiv:1308.1103]

**Photon power :**

$$\frac{P}{A} = 2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} C_{a\gamma\gamma}^2 \left( \frac{B}{10\text{T}} \right)^2 |\beta|^2$$

$$\beta^2 = \frac{P_{\text{Diel.Halosc.}}}{P_{\text{Mirror}}}$$



# Dielectric Haloscope



- Boost the power by coherent interference of photons generated on N discs plus resonance between discs

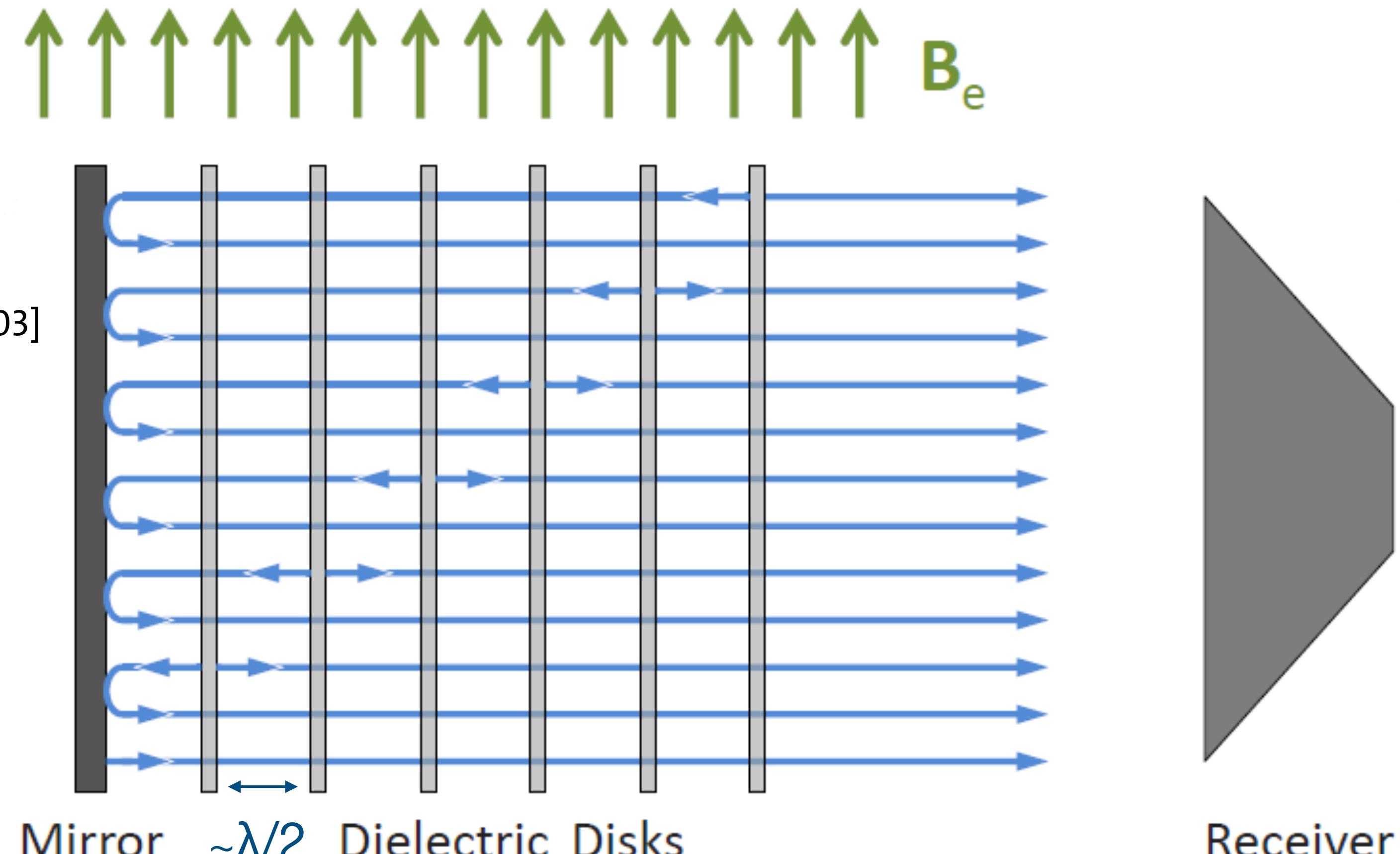
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**FoM =  $B^2 \text{ m}^2 = 100 \text{ T}^2 \text{ m}^2$**

$$\beta^2 = \frac{P_{\text{Diel.Halosc.}}}{P_{\text{Mirror}}}$$



# Dielectric Haloscope



- $|\beta|^2 > 10^4$  achievable with 80 discs of  $\text{LaAlO}_3$  ( $\epsilon = 24$ )

“Quasi-broadband” achieved by:

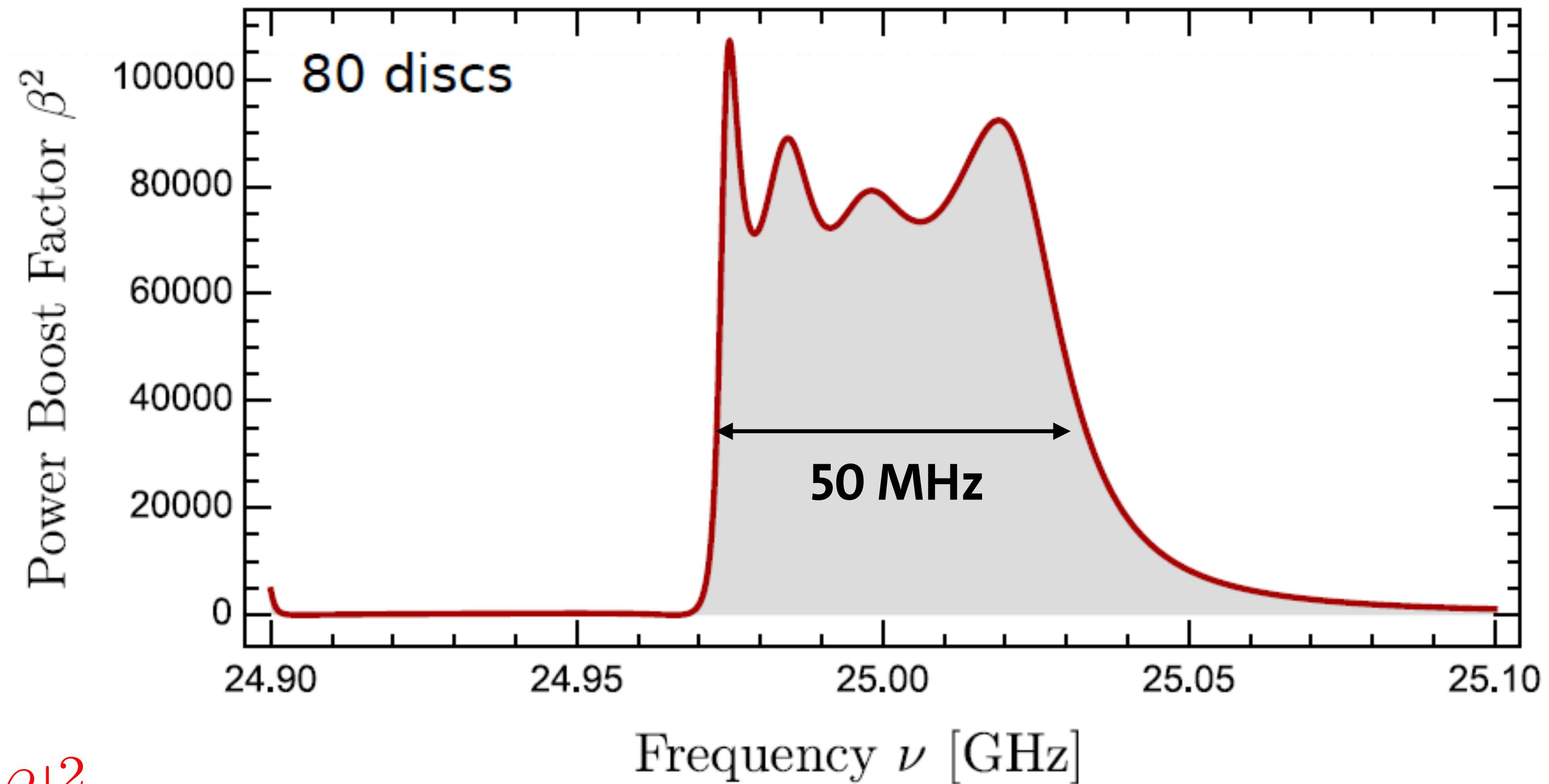
- positioning the discs with relative spacing of  $\sim \lambda/2$  according to simulation prediction
- with precision better than  $10 \mu\text{m}$

**Photon power :**

$$\frac{P}{A} = 2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} C_{a\gamma\gamma}^2 \left( \frac{B}{10\text{T}} \right)^2 |\beta|^2$$

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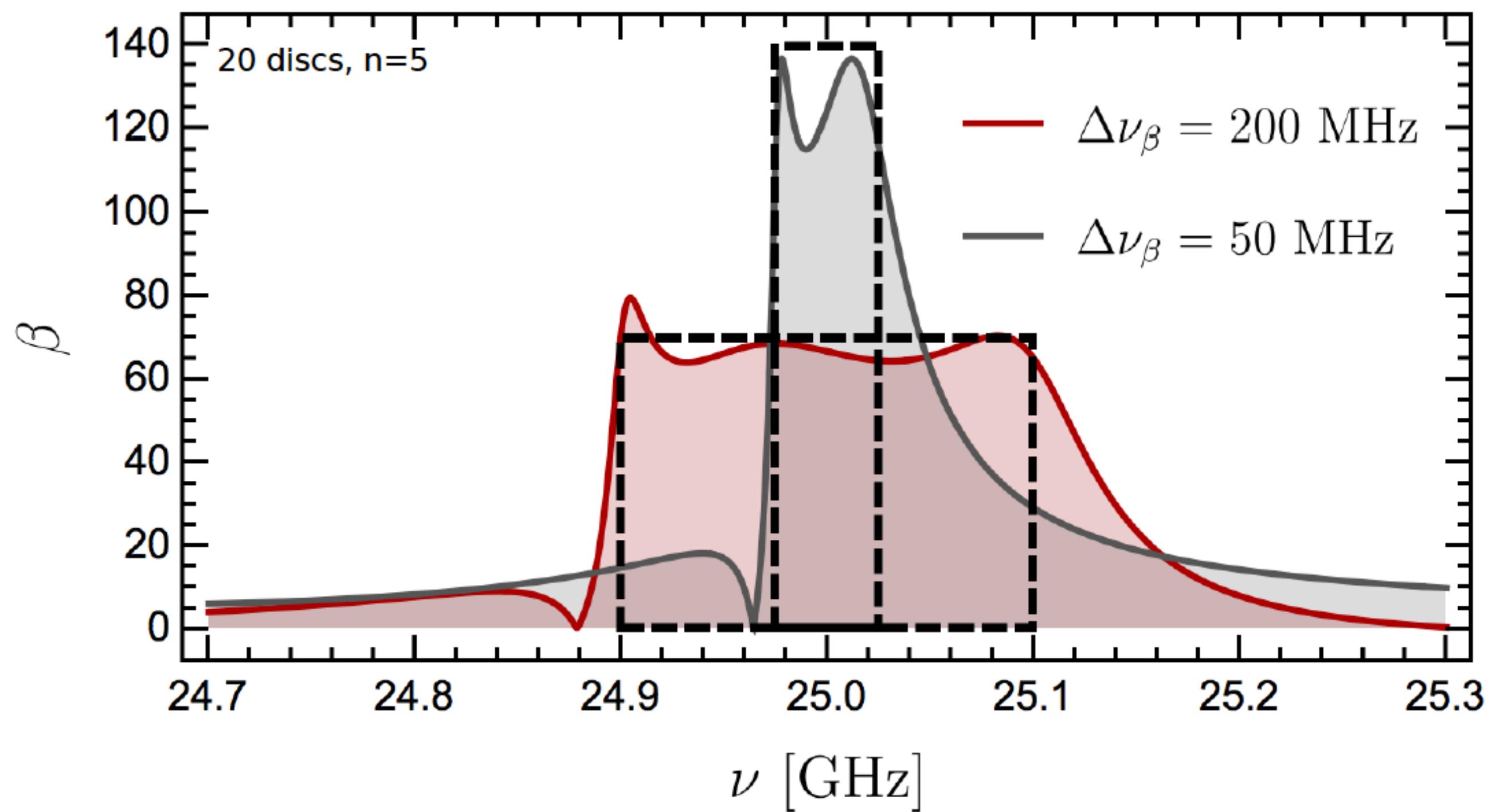
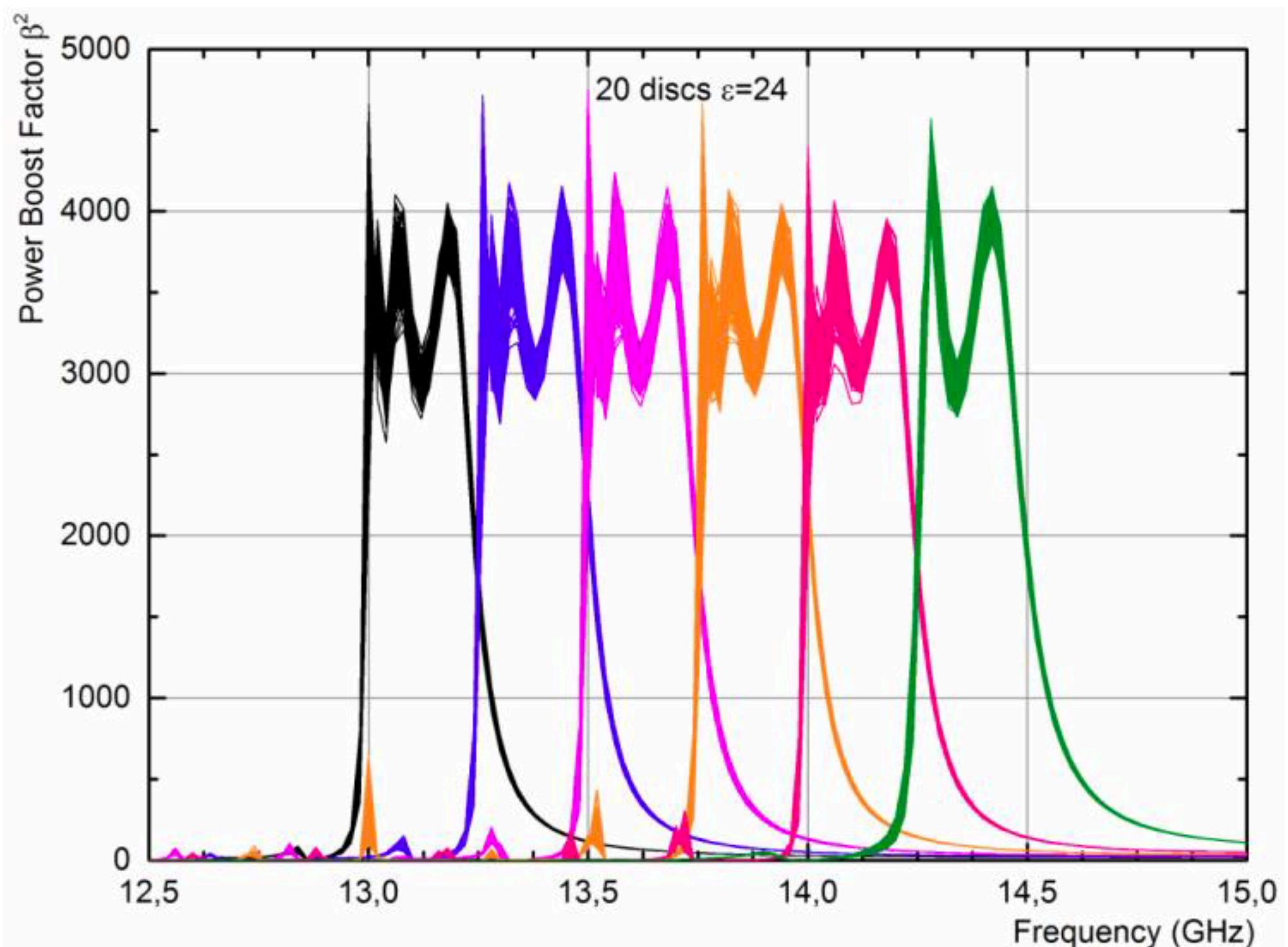


White paper: MADMAX Collaboration,  
Eur. Phys. J. C 79, 186 (2019), [arXiv:1901.07401]

# Frequency scan concept

- Tuning of sensitive frequency range by adjusting disc spacing

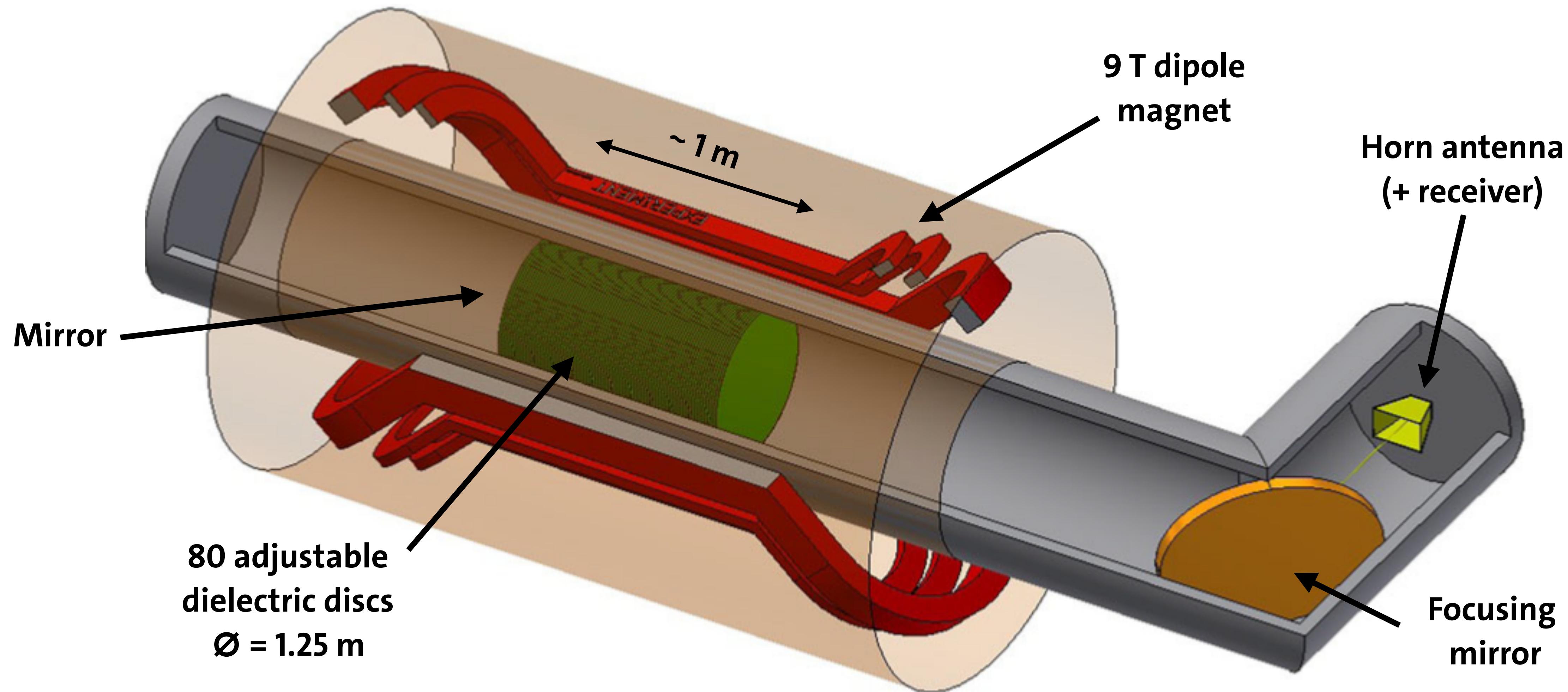
- Area law:  $\beta^2 \Delta\nu_\beta \sim \text{const.}$
- Broad-band scan for search
- Narrow-band to confirm possible signals



# The experiment



## MAgnitized disc and Mirror Axion eXperiment



# Time scale

2017 -2019  
Design

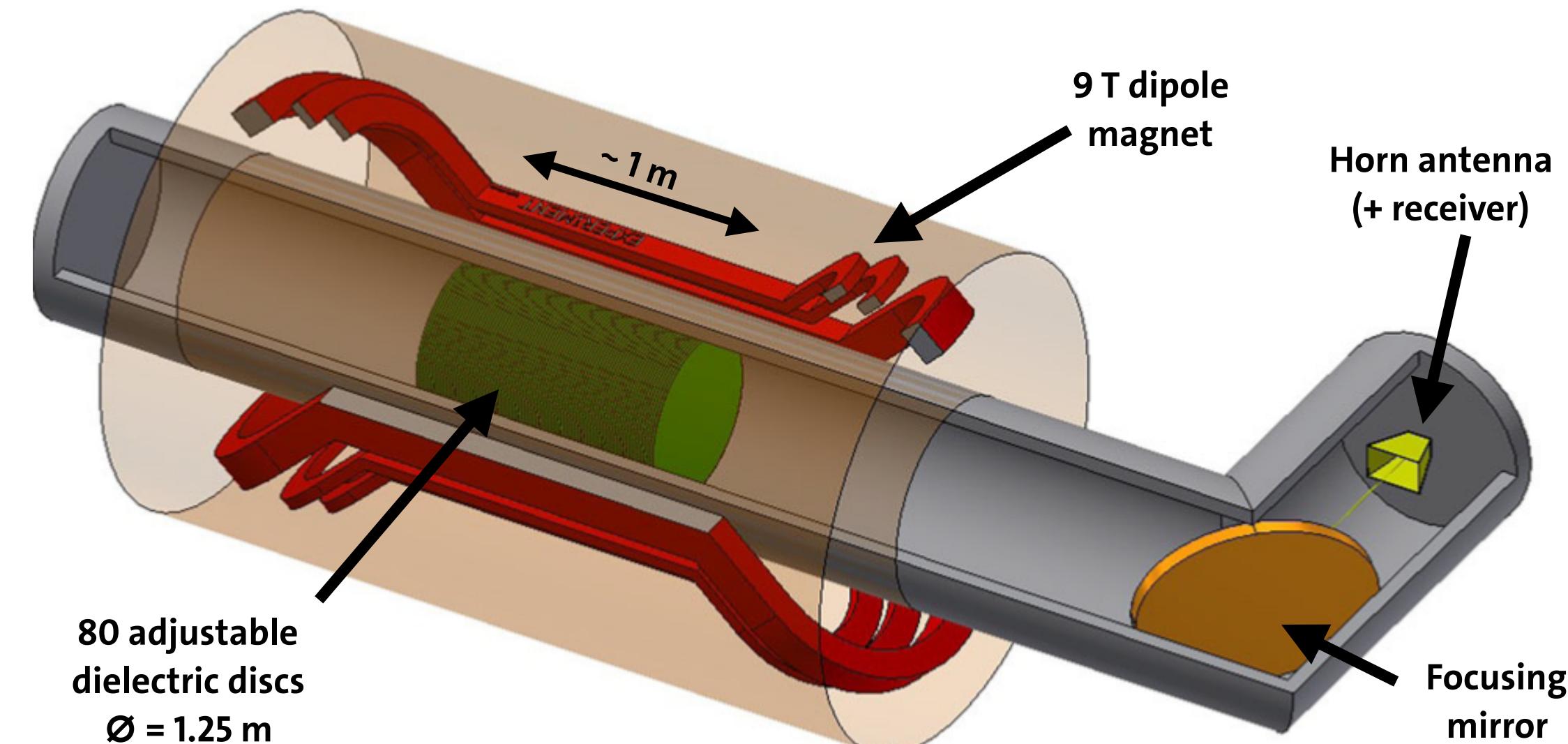
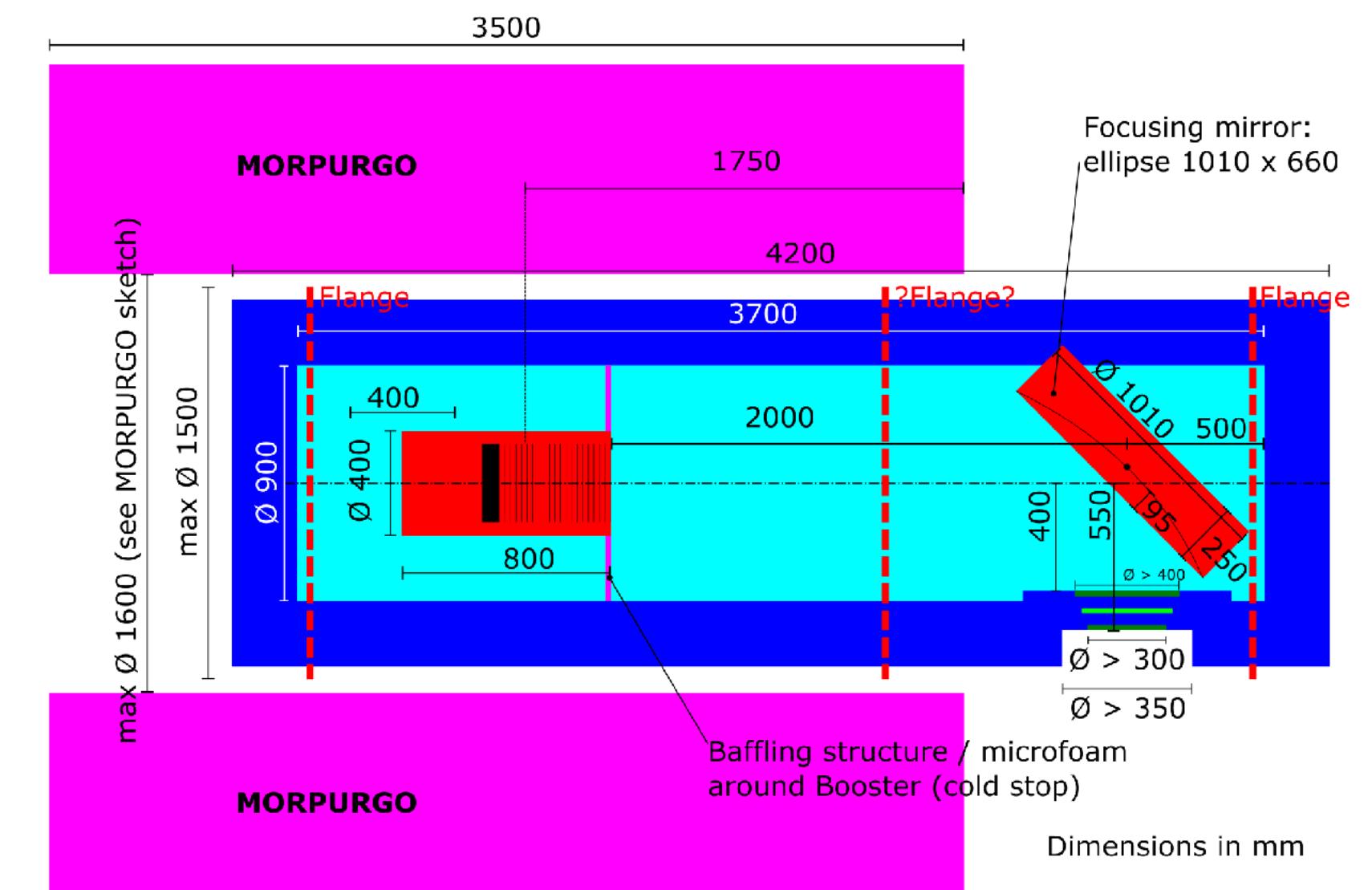
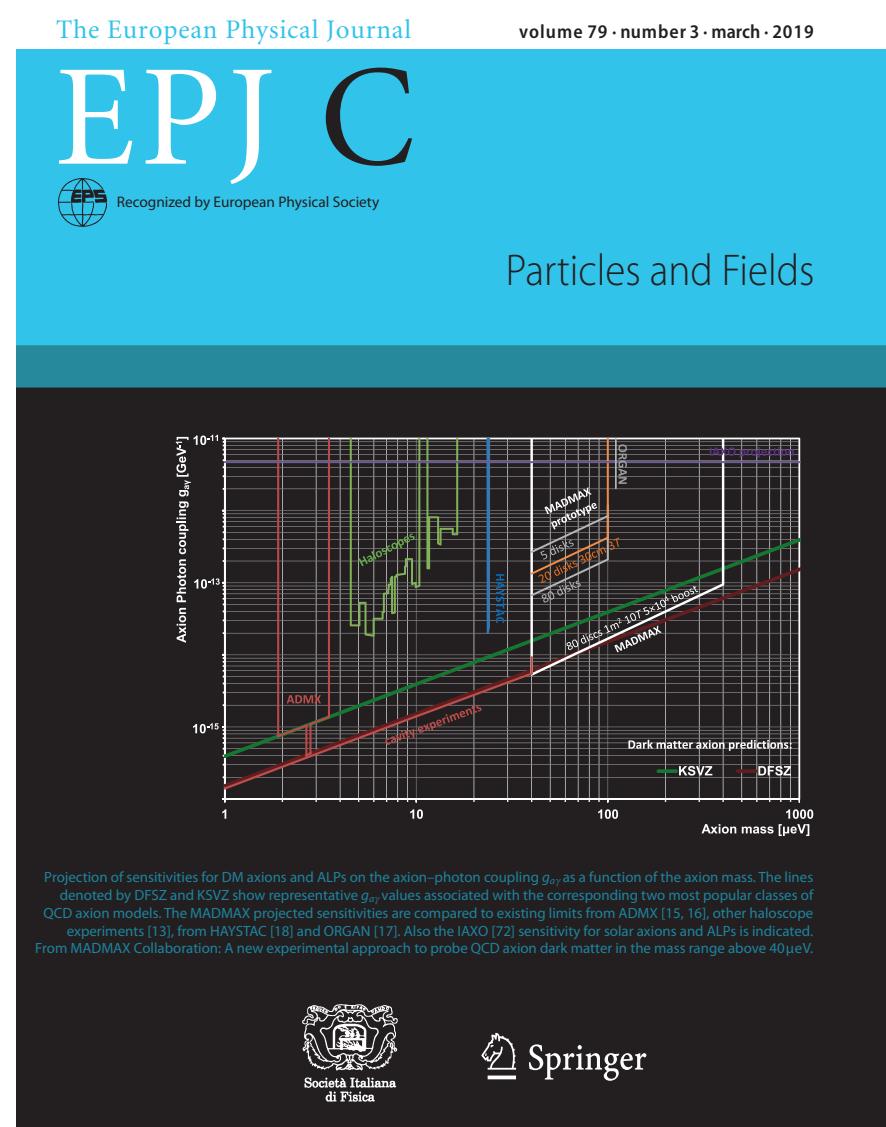
PRC 2019

2019 -2022  
Prototype

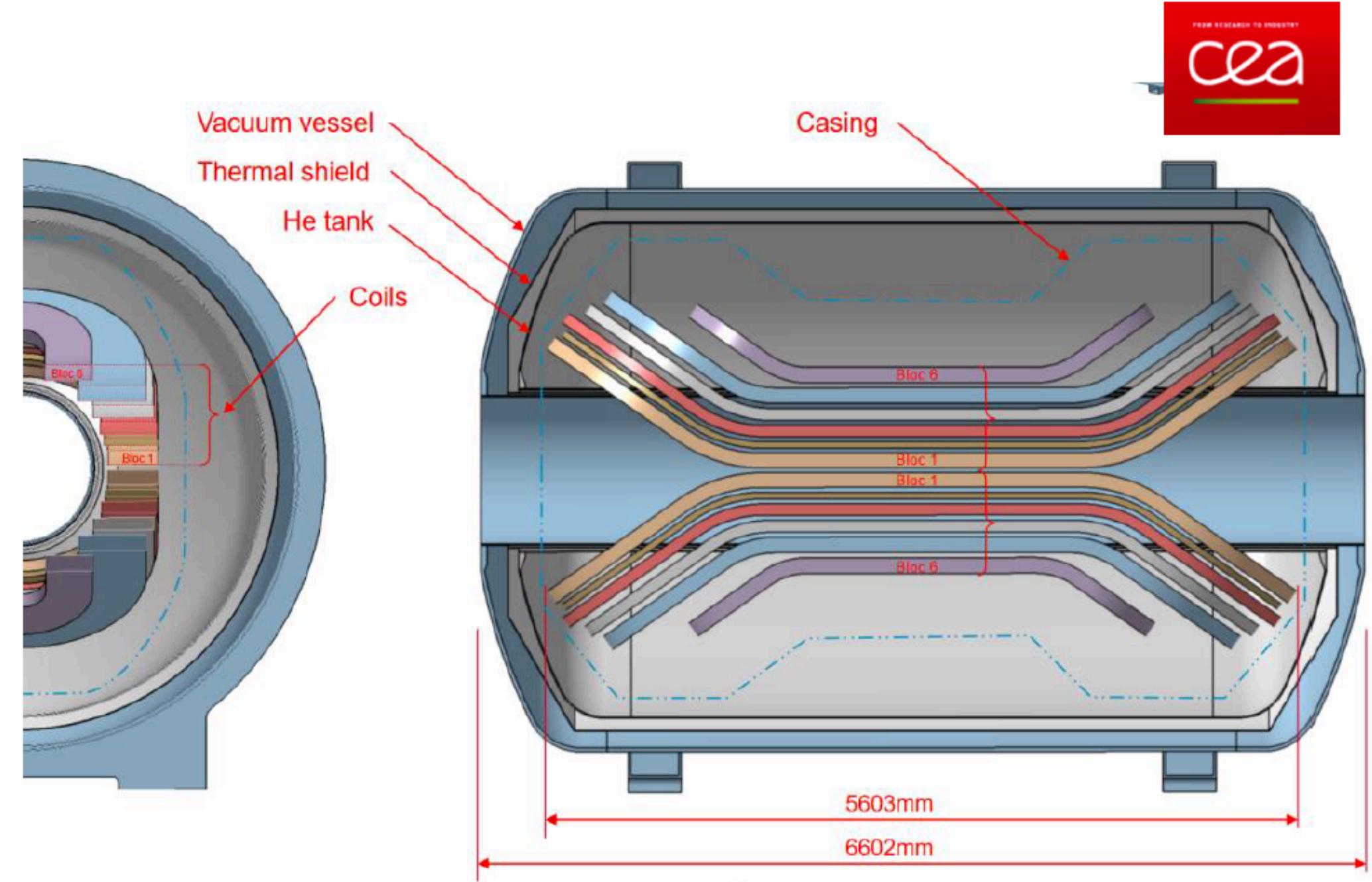
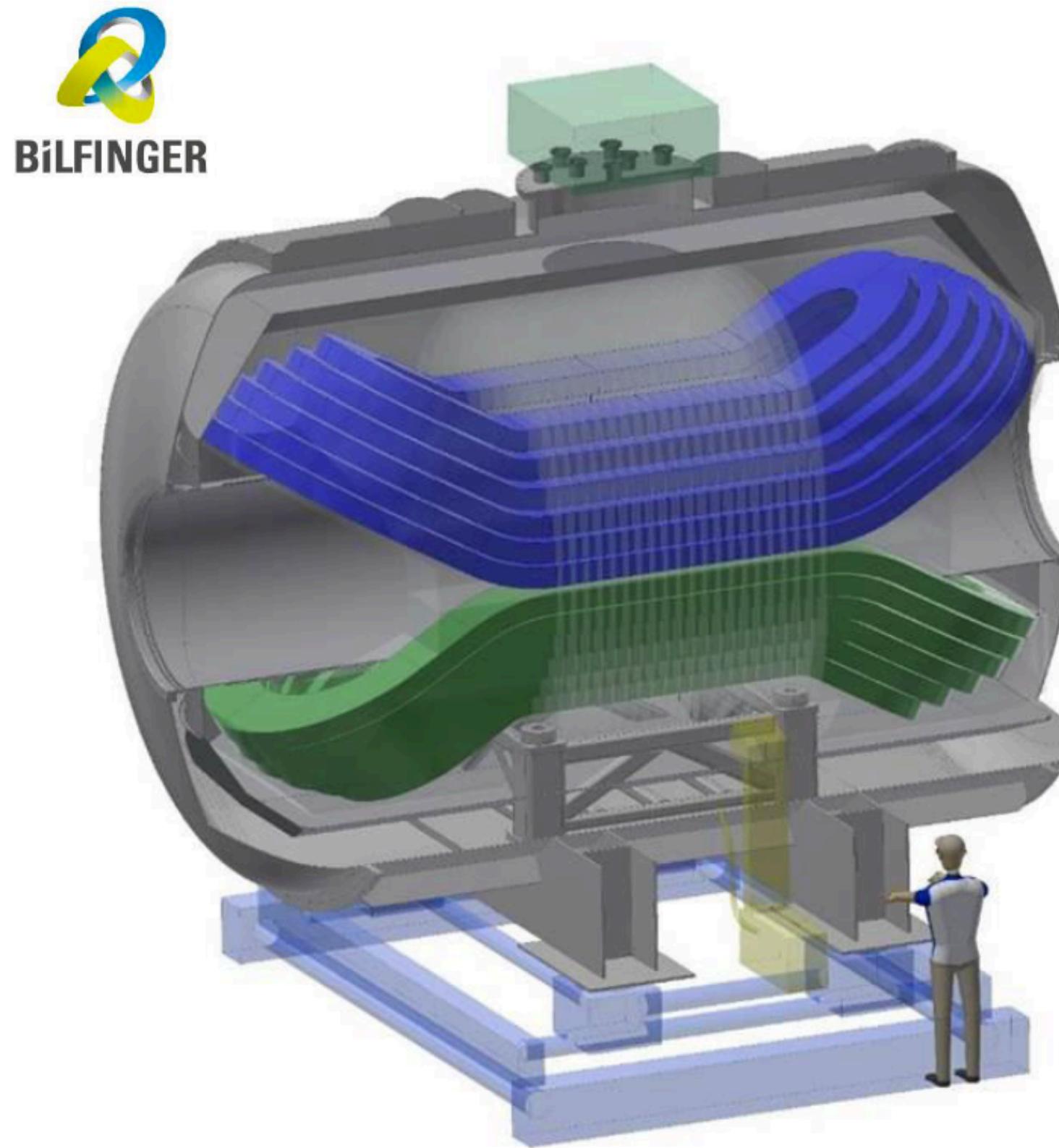
First physics run @ CERN

2022 -2025  
Construction

2025 -2035  
Data taking @ DESY



# The Magnet



Block design with NbTi as superconductor

- Magnet design and construction drives the time scale of the project
- Peak field **9 T**, homogeneity < 20%
- Magnet bore: Length ~ 1 m,  $\emptyset \sim 1.5$  m

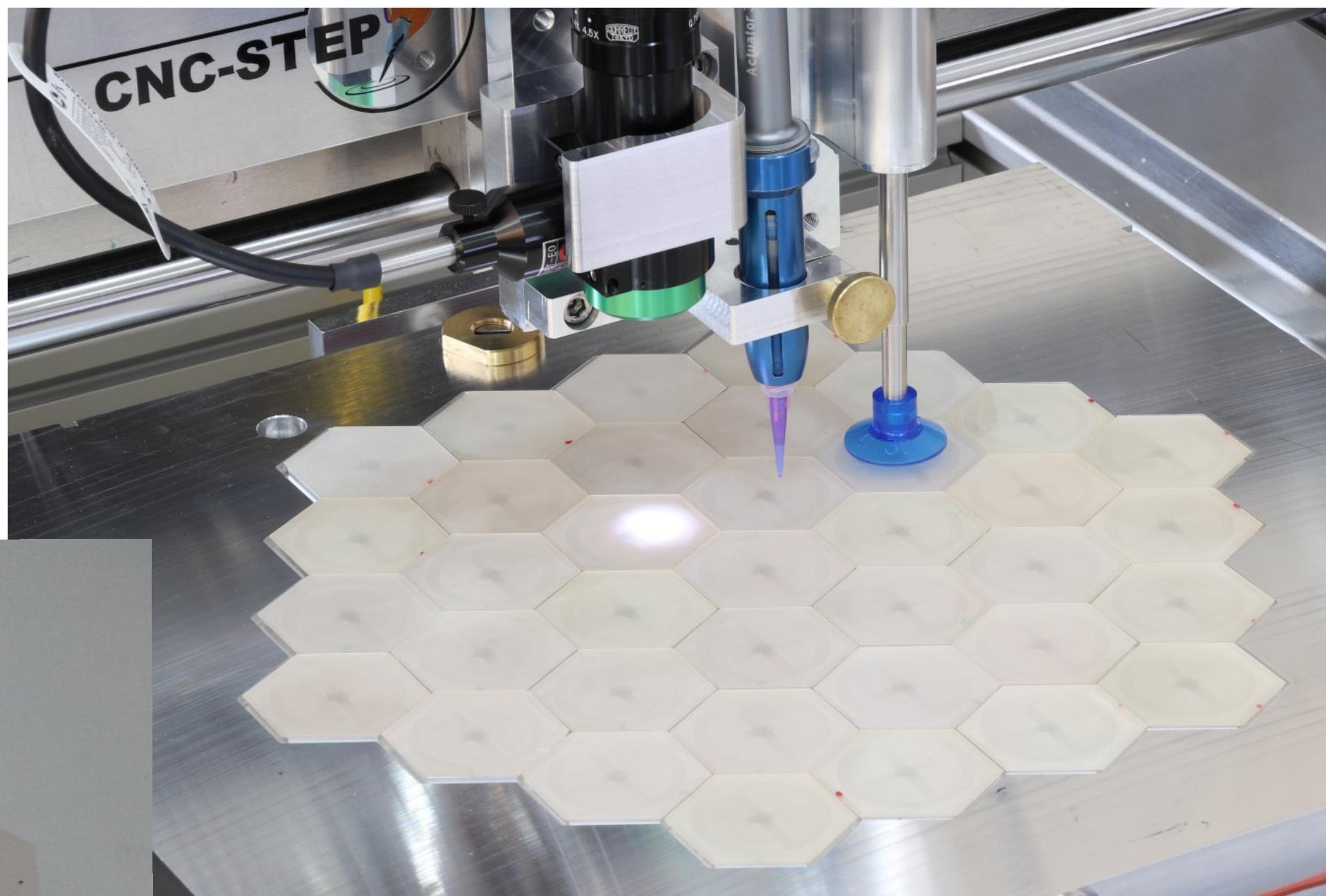
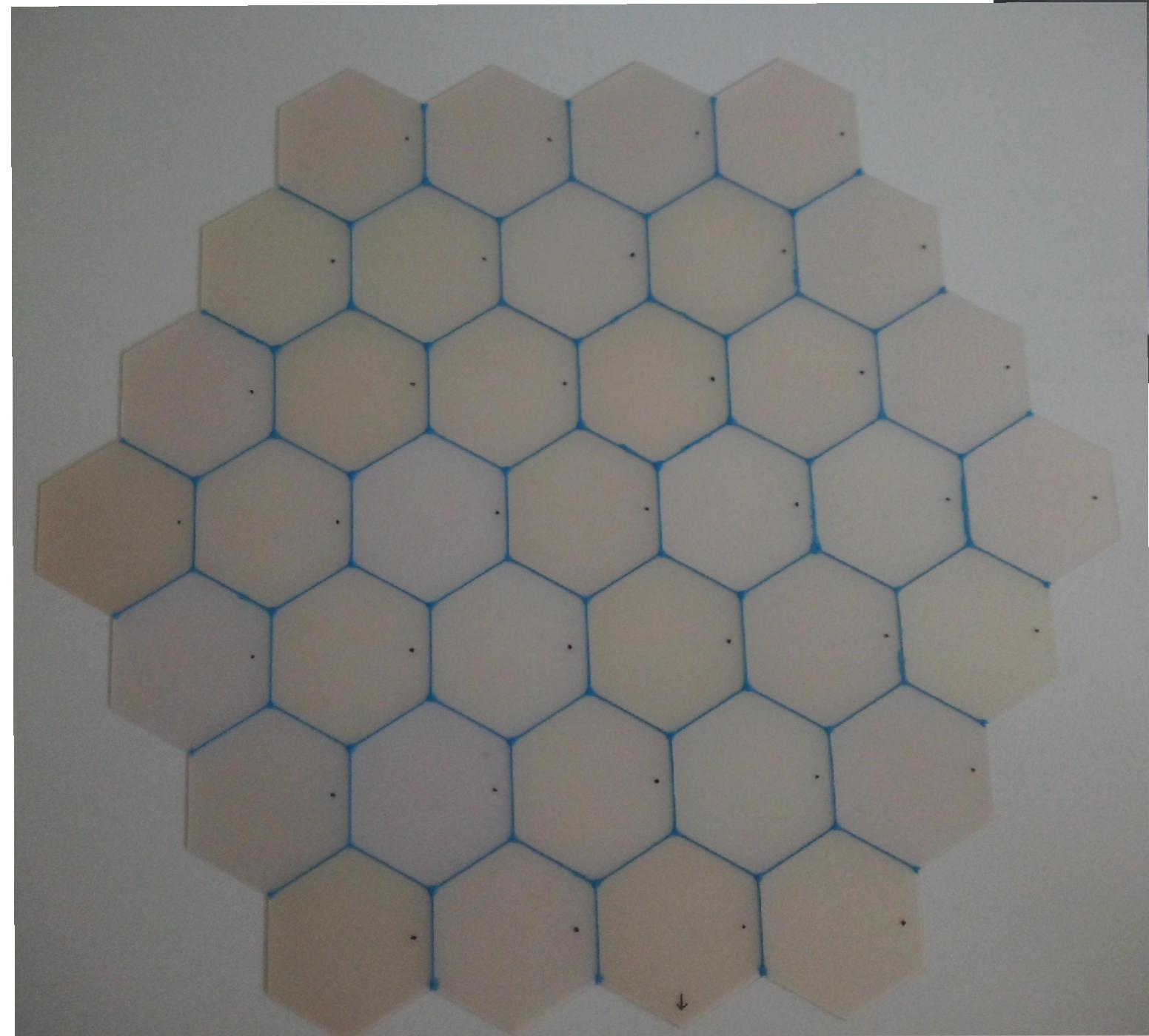
First of a kind!

$$\text{FoM} = B^2 m^2 = 100 T^2 m^2$$

# Discs

- Discs with  $\emptyset = 1.25 \text{ m}$  needed for
- Candidate materials:
  - LaAlO<sub>3</sub> ( $\epsilon \approx 24$ ,  $\tan\delta \approx \text{a few } 10^{-5}$ )**
  - Sapphire ( $\epsilon \approx 9$ ,  $\tan\delta \approx 10^{-5}$ )
- LaAlO<sub>3</sub> grown in 3" wafers max
- Tiling necessary
- Hexagonal tiles cut by laser cutter
- Glued with Stycast Blue
- Characterisation of dielectric properties @ 4 K,  $f = 10 - 15 \text{ GHz}$  ongoing

$$\text{FoM} = B^2 \text{ m}^2 = 100 \text{ T}^2 \text{ m}^2$$



**First tiled LaAlO<sub>3</sub> disc:**

$\emptyset = 30 \text{ cm}$

$d = 1 \text{ mm}$

Single wafer size 2"

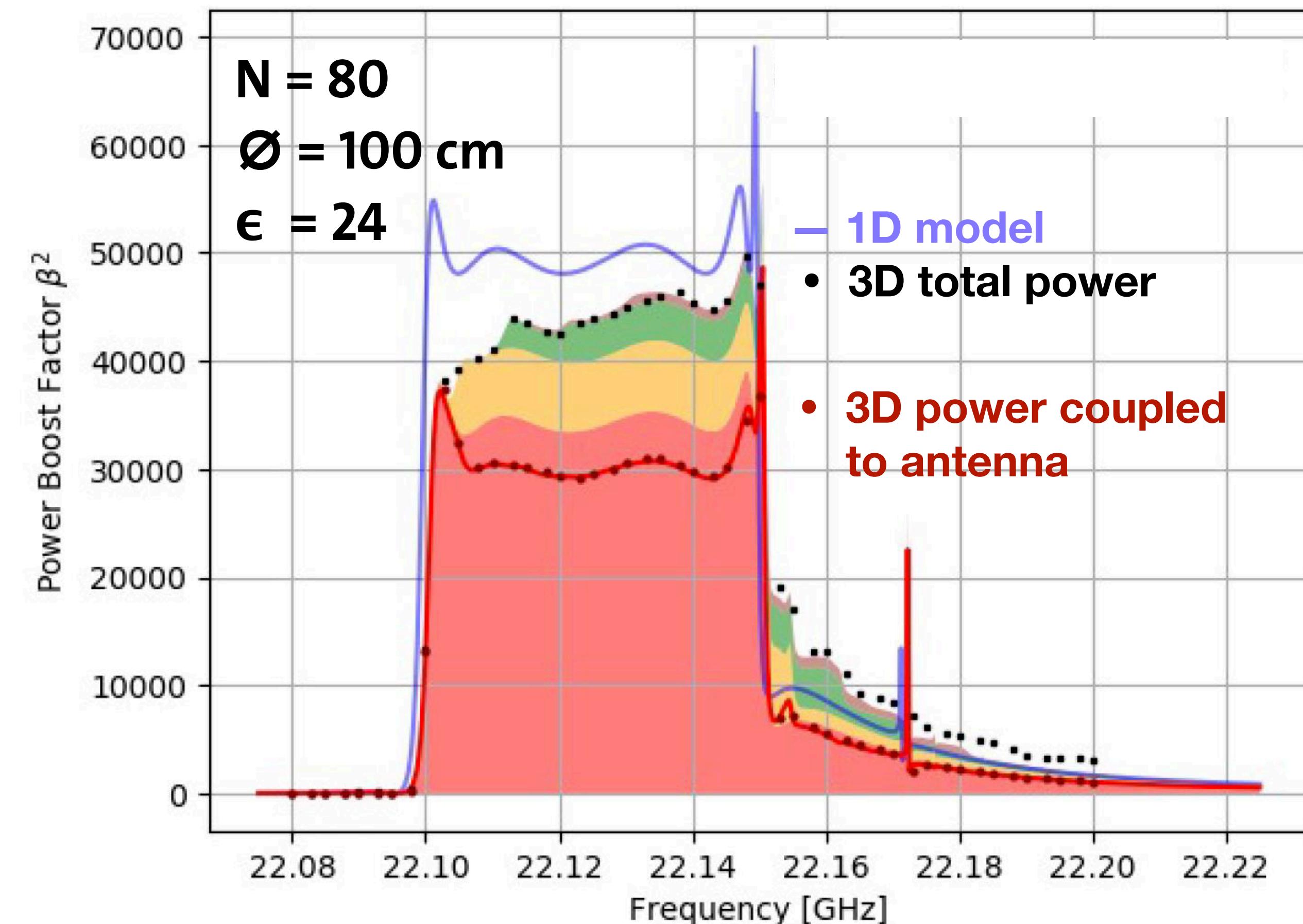
Scalability to  $\emptyset = 1.25 \text{ m}$  being investigated

# System design studies



## Detector feasibility study and design optimisation using simulation of achievable boost factor

- 3D effects (diffraction)
- Coupling to antenna (beam shape)  $\rightarrow \sim 10\text{-}20\% \text{ losses}$
- Dielectric loss
- Inaccuracy (position, roughness, tilt, thickness,...)  $\rightarrow \sim 10\text{-}20\% \text{ losses}$
- DM velocity dispersion
- Tiling of discs

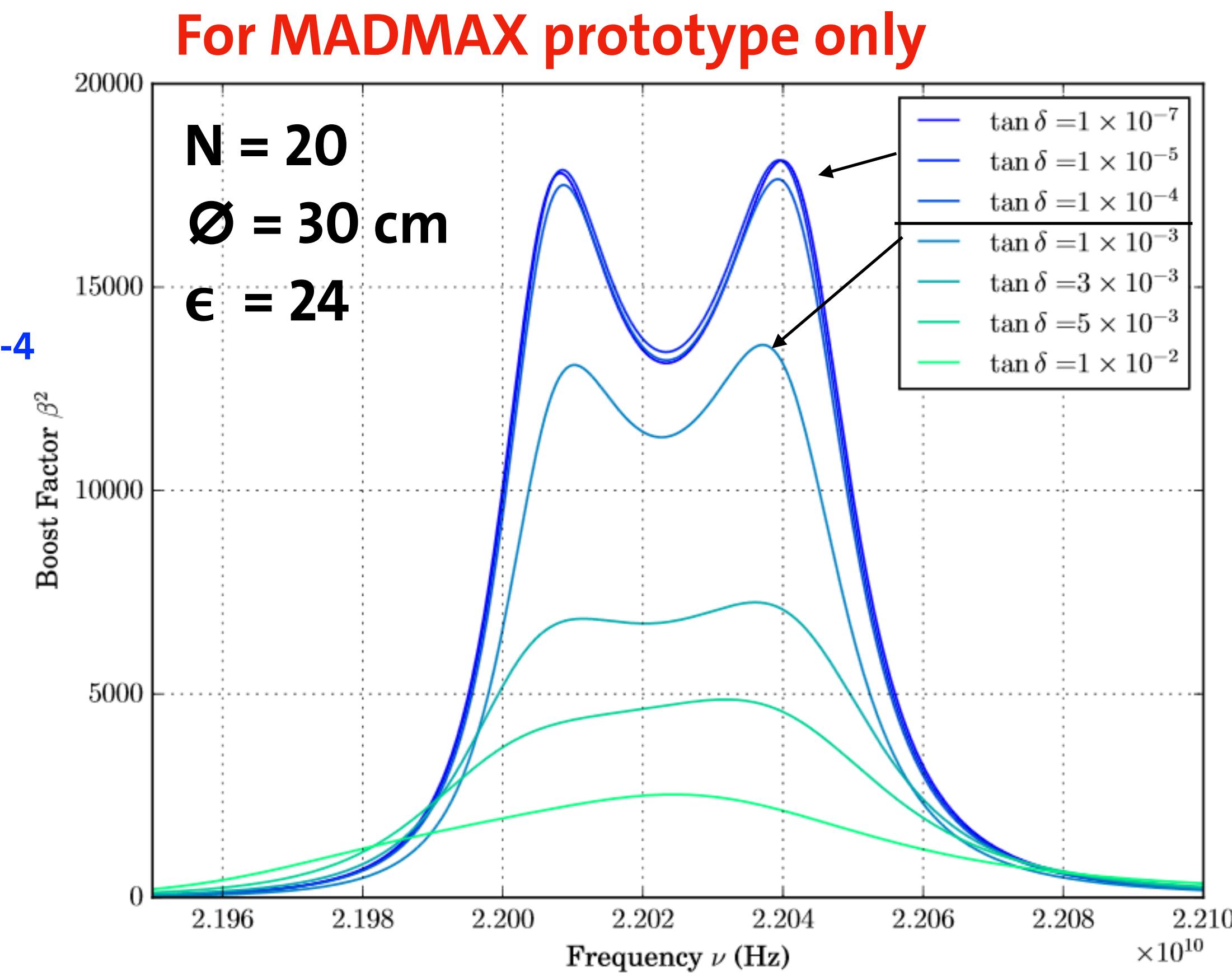


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- 3D effects (diffraction)
  - Coupling to antenna (beam shape)
  - **Dielectric loss**
  - Inaccuracy (position, roughness, tilt, thickness,...)
  - DM velocity dispersion
  - Tiling of discs
- ~10-20% losses
- ~10-20% losses
- **small losses for  $\tan\delta < 10^{-4}$**



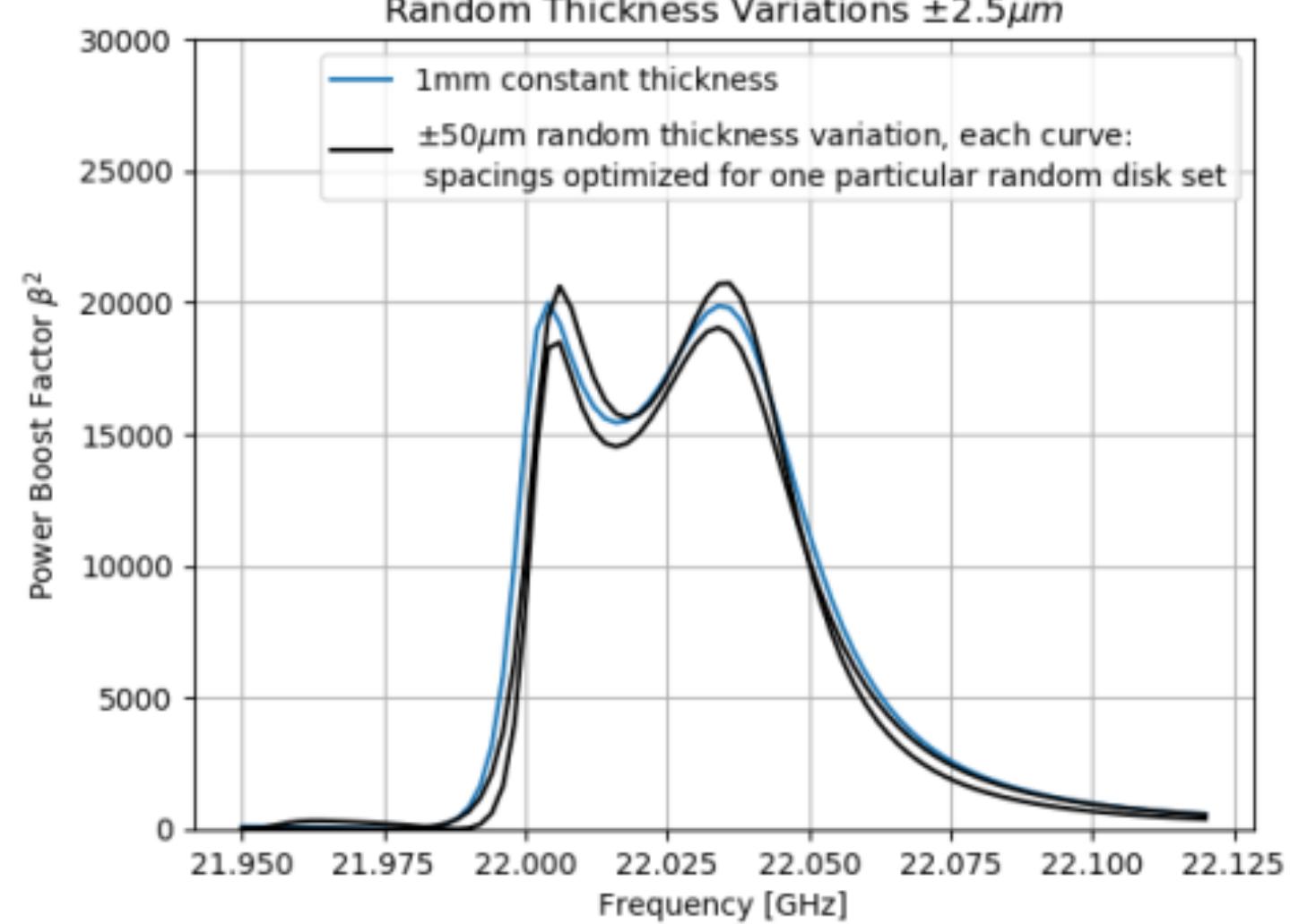
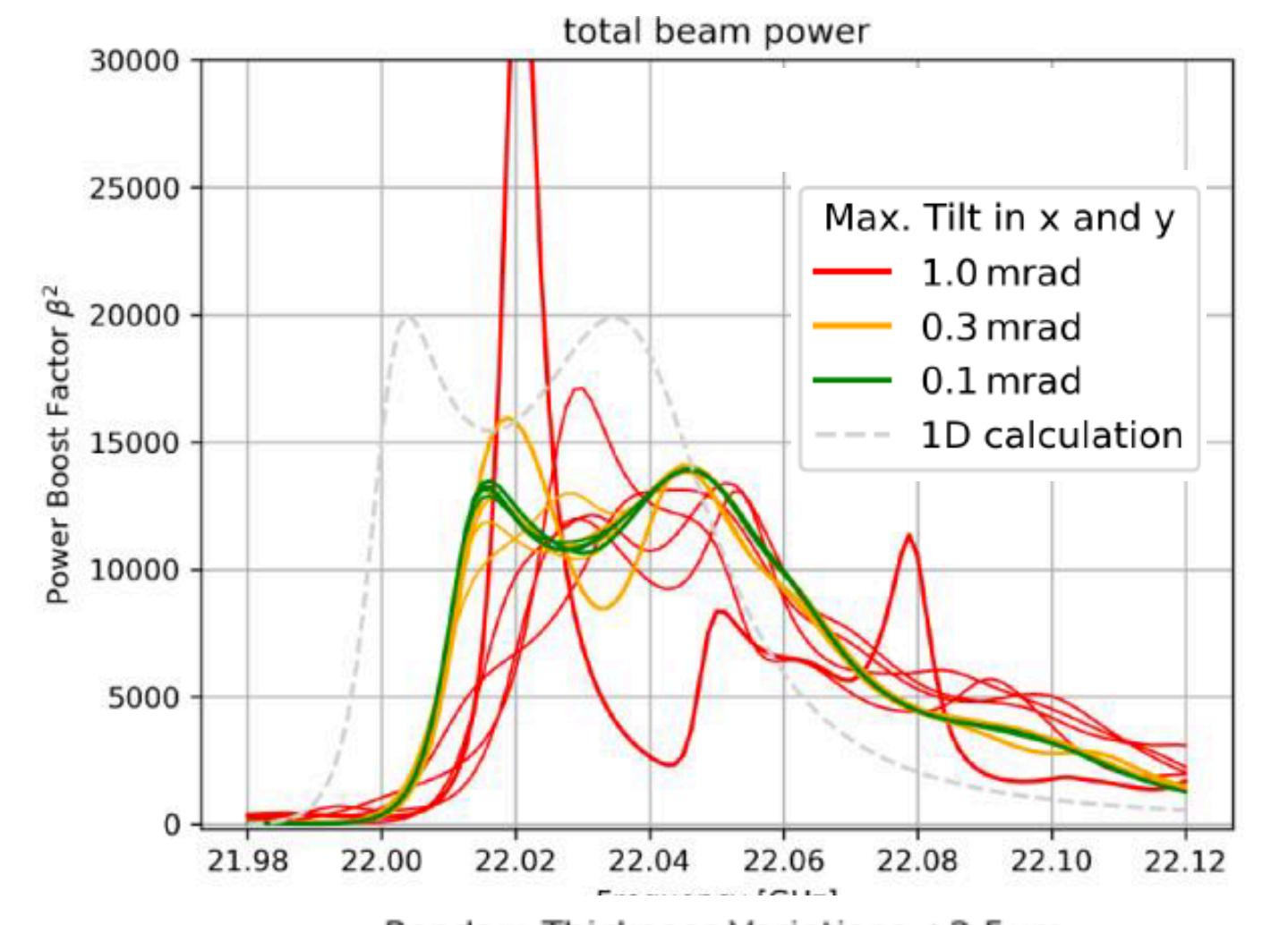
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- 3D effects (diffraction) → ~10-20% losses
- Coupling to antenna (beam shape) → ~10-20% losses
- Dielectric loss → small losses for  $\tan\delta < 10^{-4}$
- **Inaccuracy (position, roughness, tilt, thickness,...)** → **positioning precision < 10 μm**  
**roughness < 10 μm**  
**tilt < 0.1 mrad**  
**thickness measured to ± 5 μm**
- DM velocity dispersion
- Tiling of discs

$$\begin{aligned}N &= 20 \\ \varnothing &= 30 \text{ cm} \\ \epsilon &= 24\end{aligned}$$

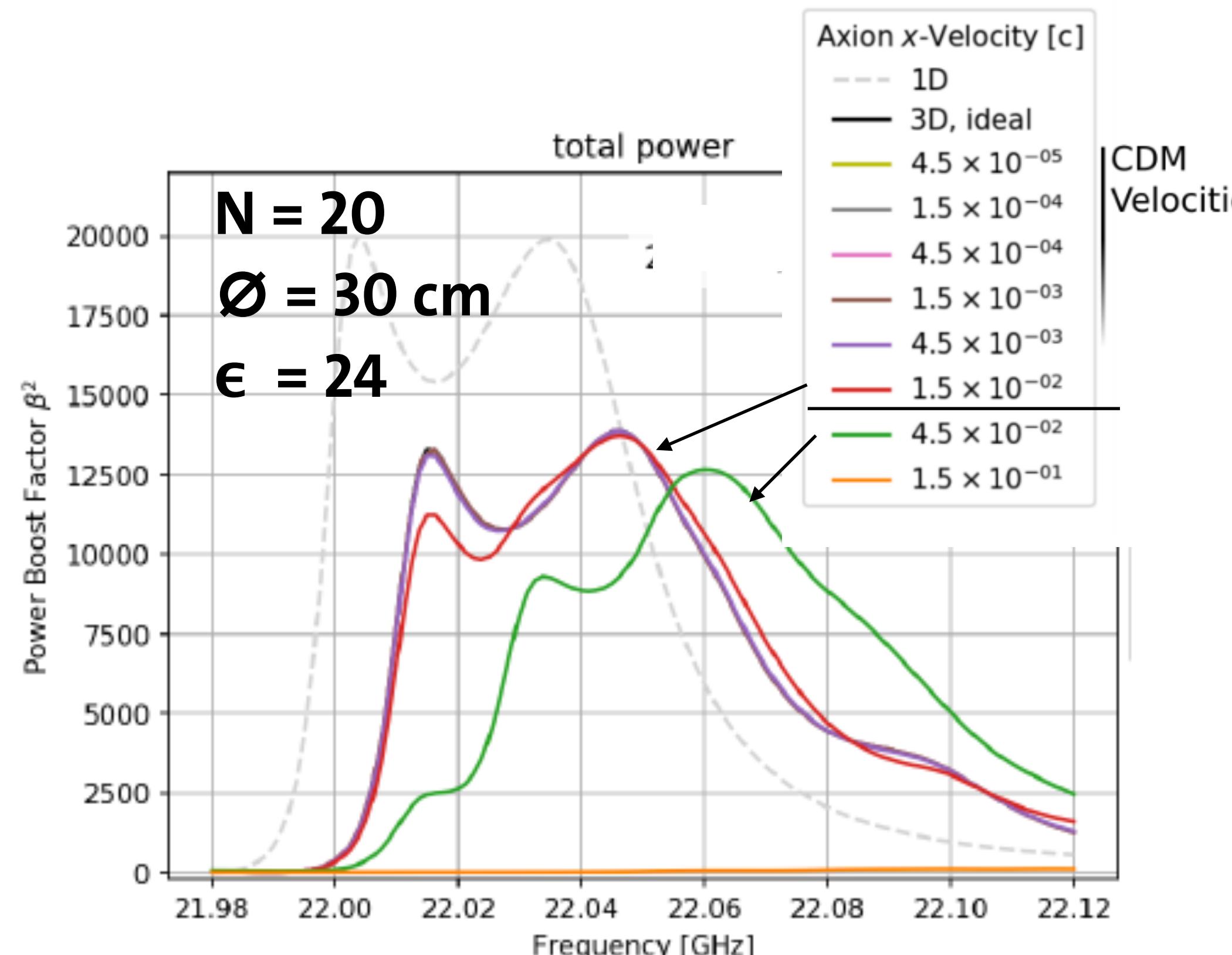


# System design studies



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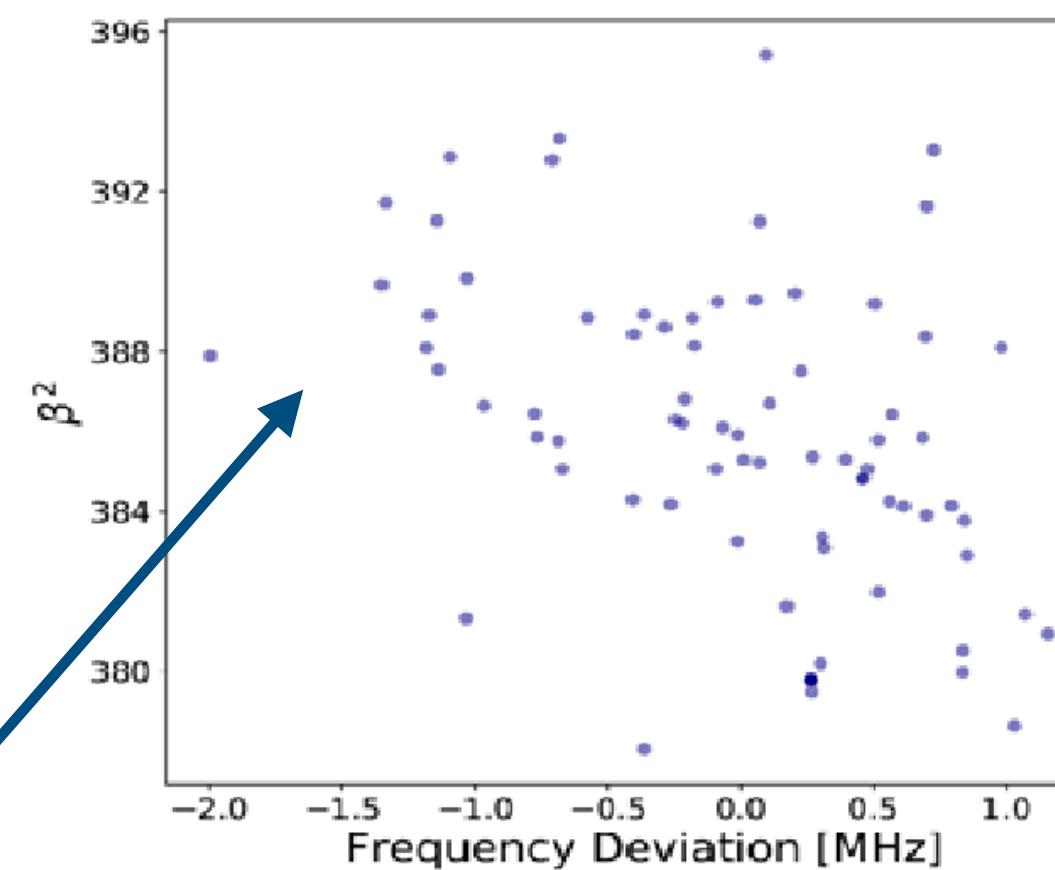
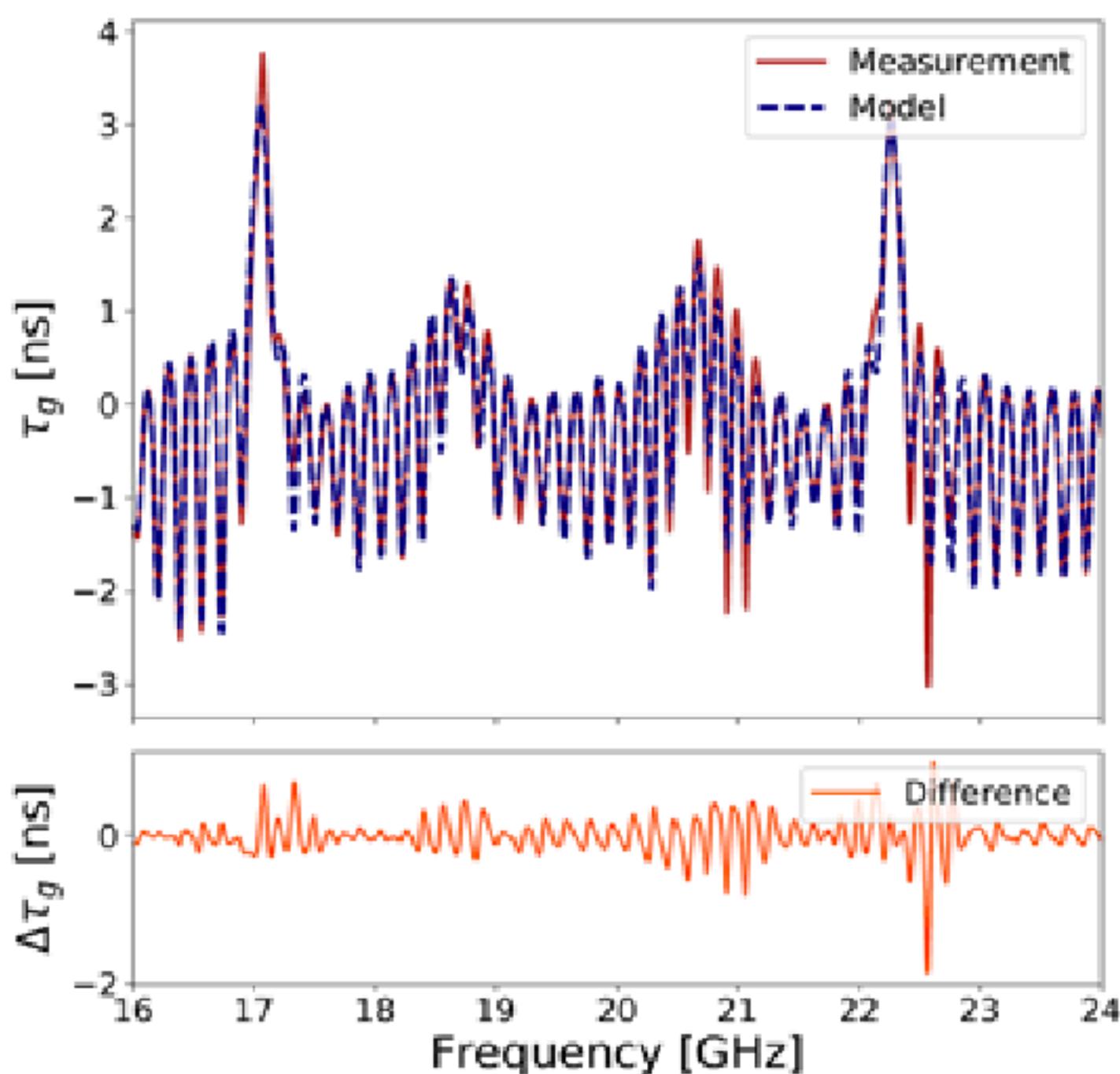
- 3D effects (diffraction) → ~10-20% losses
- Coupling to antenna (beam shape) → ~10-20% losses
- Dielectric loss → small losses for  $\tan\delta < 10^{-4}$
- Inaccuracy (position, roughness, tilt, thickness,...) → positioning precision < 10  $\mu\text{m}$
- **DM velocity dispersion** → **no significant loss if  $v < 10^{-2} c$**
- Tiling of discs



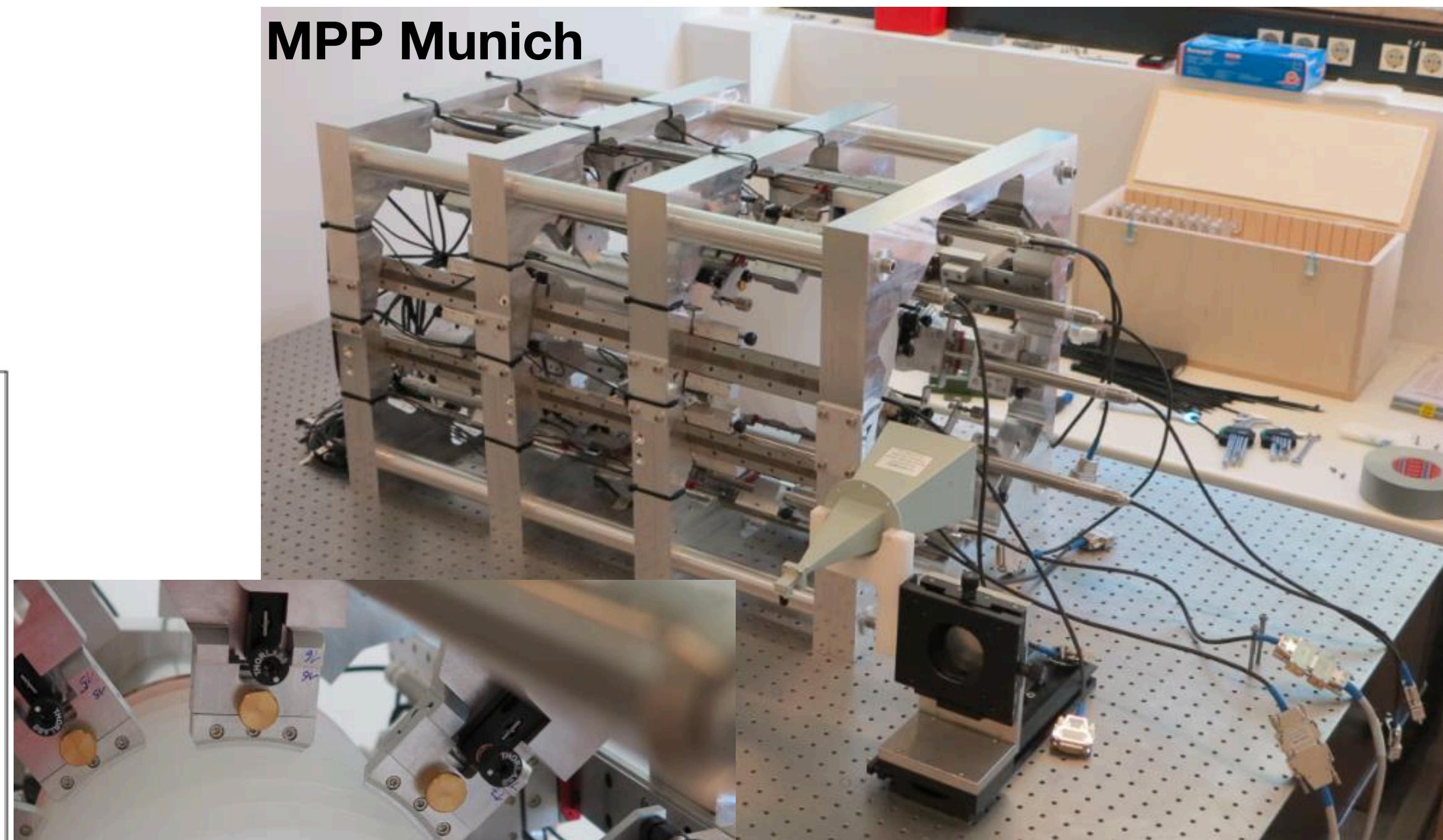
New on arxiv: 1906.02677  
by Jan Schütte-Engel

# Proof of principle

- Reflectivity measurements with up to 20 discs
- Reproducibility of discs **positioning few  $\mu\text{m}$**
- Achievable **boost factor** with 5 discs  
**reproducible within few MHz**

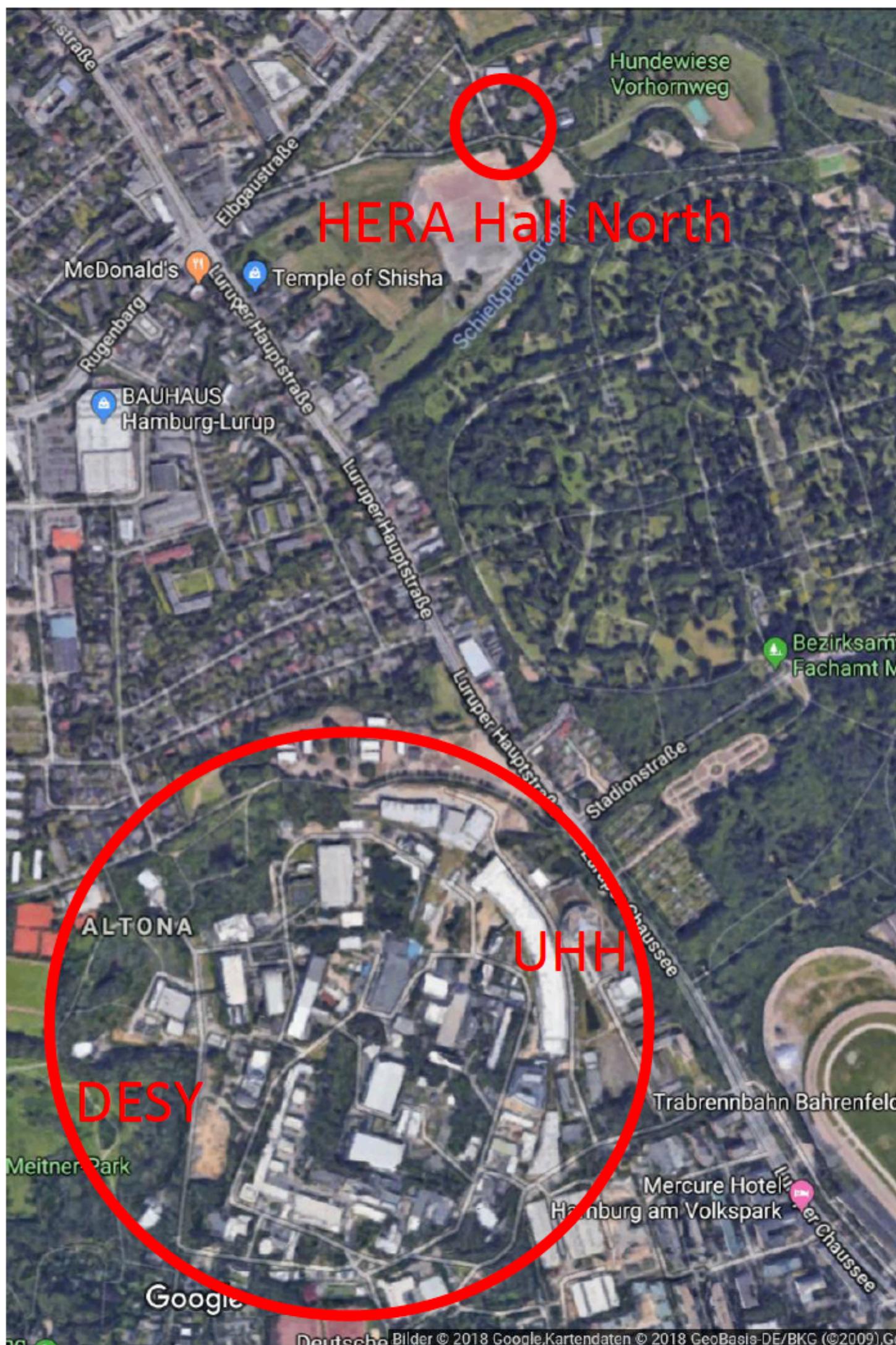


MPP Munich



Proof of principle booster  
 $N = 20$   
 $\emptyset = 20 \text{ cm}$   
 $\epsilon \approx 9$

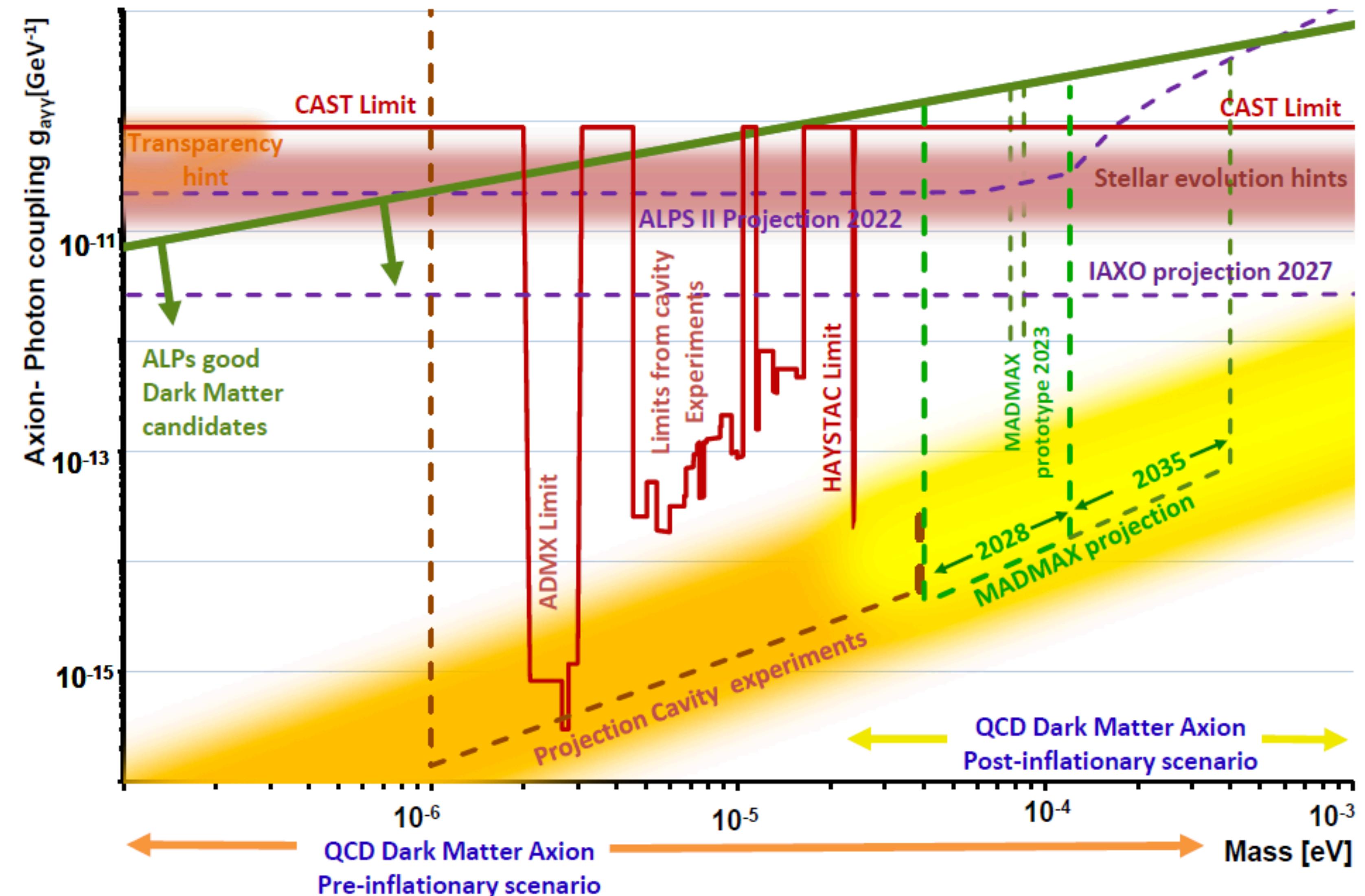
# Experimental site



- MADMAX to be built at Hera Hall North
- Make use of DESY infrastructure
- Benefit: re-use H1 yoke as magnetic shielding to reduce fringe field



# Sensitivity to QCD dark matter axion



**MADMAX**  
 $A = 1 \text{ m}^2$   
 $B_{\parallel} = 10 \text{ T}$   
 $T_{\text{sys}} = 8 \text{ K}$   
 $\beta^2 = 5 \cdot 10^4$



# The MADMAX collaboration



Max-Planck-Institut  
für Radioastronomie



UH Universität Hamburg  
DER FORSCHUNG | DER LEHRE | DER BILDUNG



Max-Planck-Institut für Physik

Irfu - CEA Saclay  
Institut de recherche  
sur les lois fondamentales  
de l'Univers



EBERHARD KARLS  
UNIVERSITÄT  
TÜBINGEN



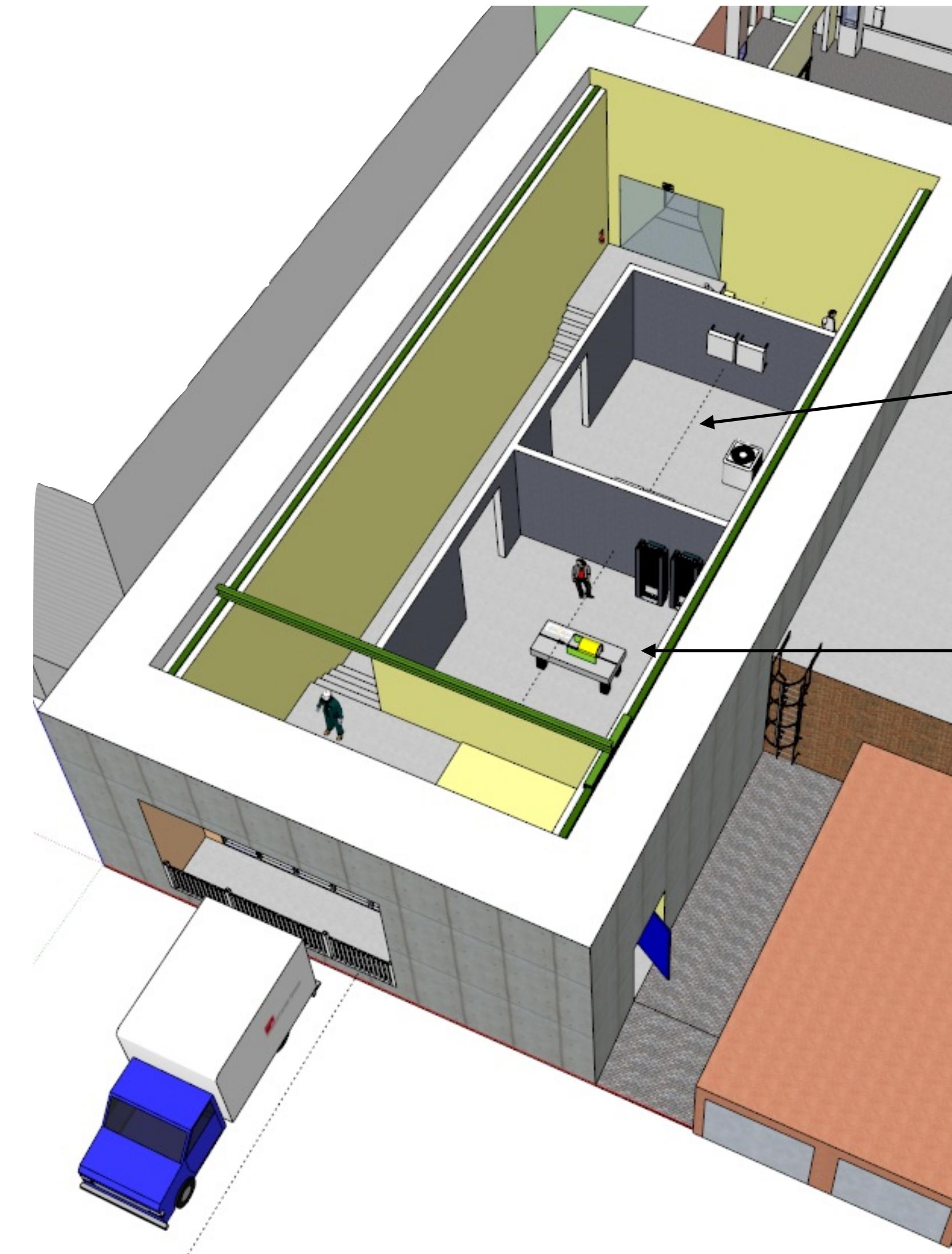
1542

Universidad  
Zaragoza

## associate members



# SHELL inauguration 08.07.19 @ 11:30



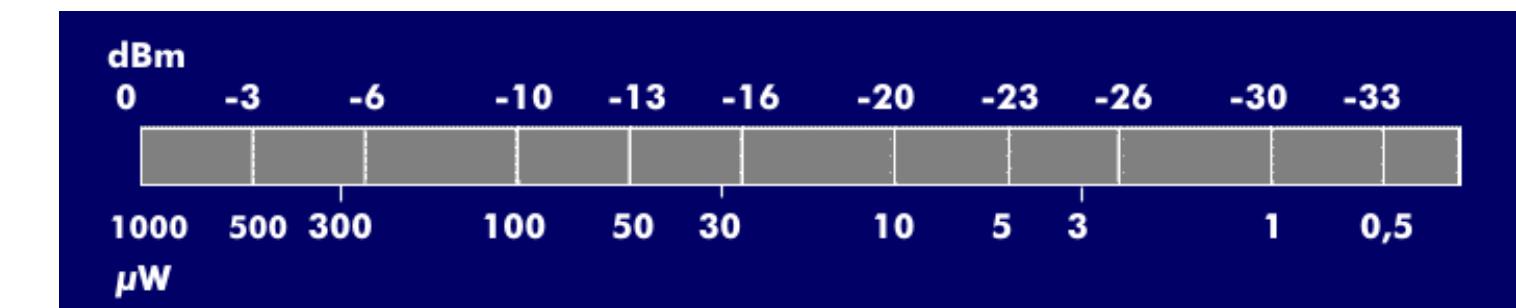
Ex-UHH synchrotron bunker  
Renewed thanks to QU funds  
Two RF shielded labs

< -50 dBm\*, for  $f < 10$  GHz  
< -100 dBm, for  $f > 10$  GHz

**BRASS**

**MADMAX**

\* dBm = Power in Decibel Milliwatt,  
 $p = 10\log(P/1 \text{ mW})$  [dBm]  
i.e. for  $P = p \mu\text{W}$ ,  $p = -90$  dBm

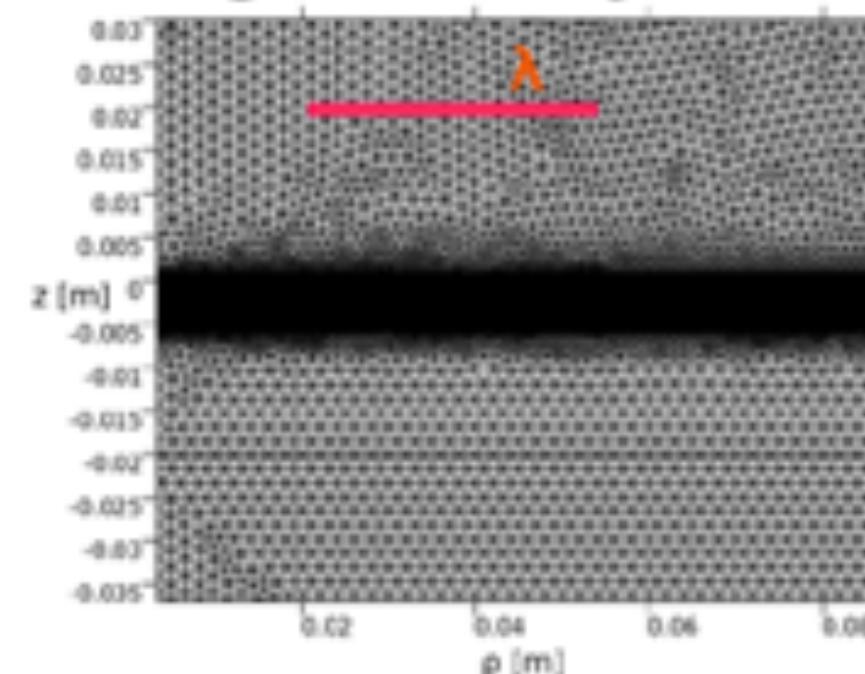


# Advanced simulation techniques



For large and many diks 3D FEM solution computationally not feasible.

## Using radial symmetry in FEM



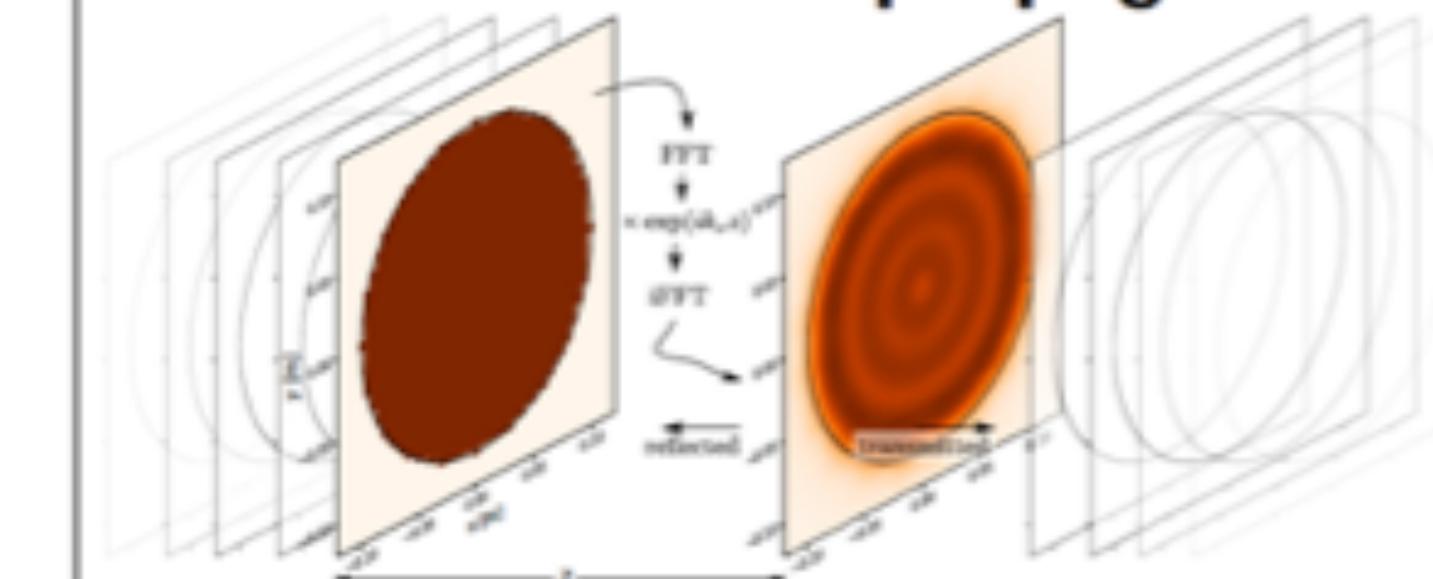
external  $B$ -field breaks symmetry  
⇒ restore symmetry by solving two circular polarized sources

$$\nabla \times (\nabla \times \mathbf{E}) - k_0^2 \epsilon \mathbf{E} = k_0^2 \mathbf{f}$$

$$\mathbf{f}(\rho, z) = \mathbf{f}^+(\rho, \phi, z) + \mathbf{f}^-(\rho, \phi, z)$$

approaches directly applicable to other open axion haloscopes

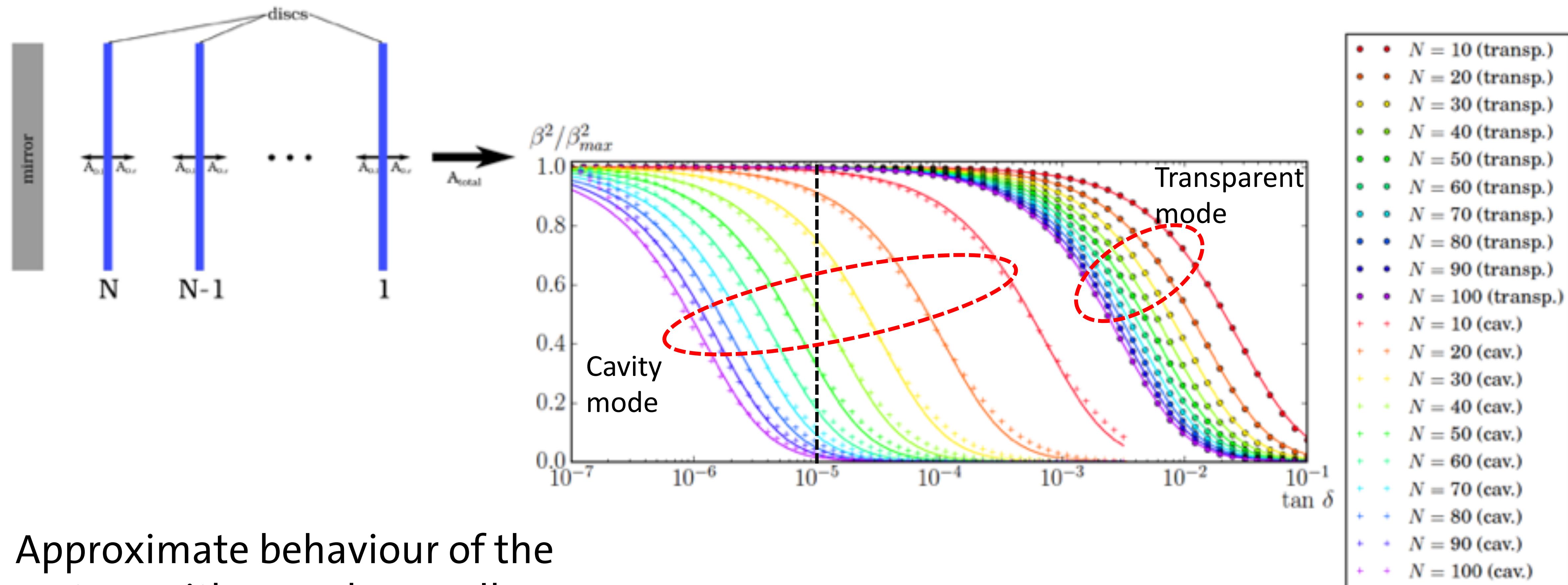
## Recursive Fourier propagation



For each propagation use a scalar diffraction theory:

$$E_i(\mathbf{x}) = \int_{\mathbb{R}^2} \frac{dk_x dk_y}{(2\pi)^2} \mathcal{F}(E_i)(k_x, k_y) \times e^{i|z-z_s|\sqrt{(\omega n)^2 - k_x^2 - k_y^2}} e^{ik_x x} e^{ik_y y}$$

# Transparent vs cavity mode



Approximate behaviour of the system with a random walk  
 $\rightarrow N_{\text{eff}} \sim N^2$

$\rightarrow$  cavity mode:  $N_{\text{eff}} = N^a$ ,  $a \approx 2.9 \pm 0.1$

# Antenna and receiver



## The challenges:

18-40 GHz: optimization of receiver-antenna system  
based on simulation results for beam  
shape/size (**MPIFR**)

50-100 GHz: develop a new concept for receiver  
(**MPIFR, NEEL**)

