



phenomenological MSSM interpretation of the 7 and 8 TeV CMS SUSY results

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on behalf of the CMS Collaoration

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DESY
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CMS pMSSM documentation

- Public 7 TeV results: CMS-SUS-12-030
twiki: <https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsSUS12030>
PAS: <http://cds.cern.ch/record/1552402>
- Public 7 + 8 TeV results: CMS-SUS-13-020
twiki: <https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResultsSUS13020>
PAS: <http://cds.cern.ch/record/1693148>

Many thanks to S. Bein, J. F. Gunion, S. Kraml, H. B. Prosper, L. Vanelderren and W. Waltenberger.

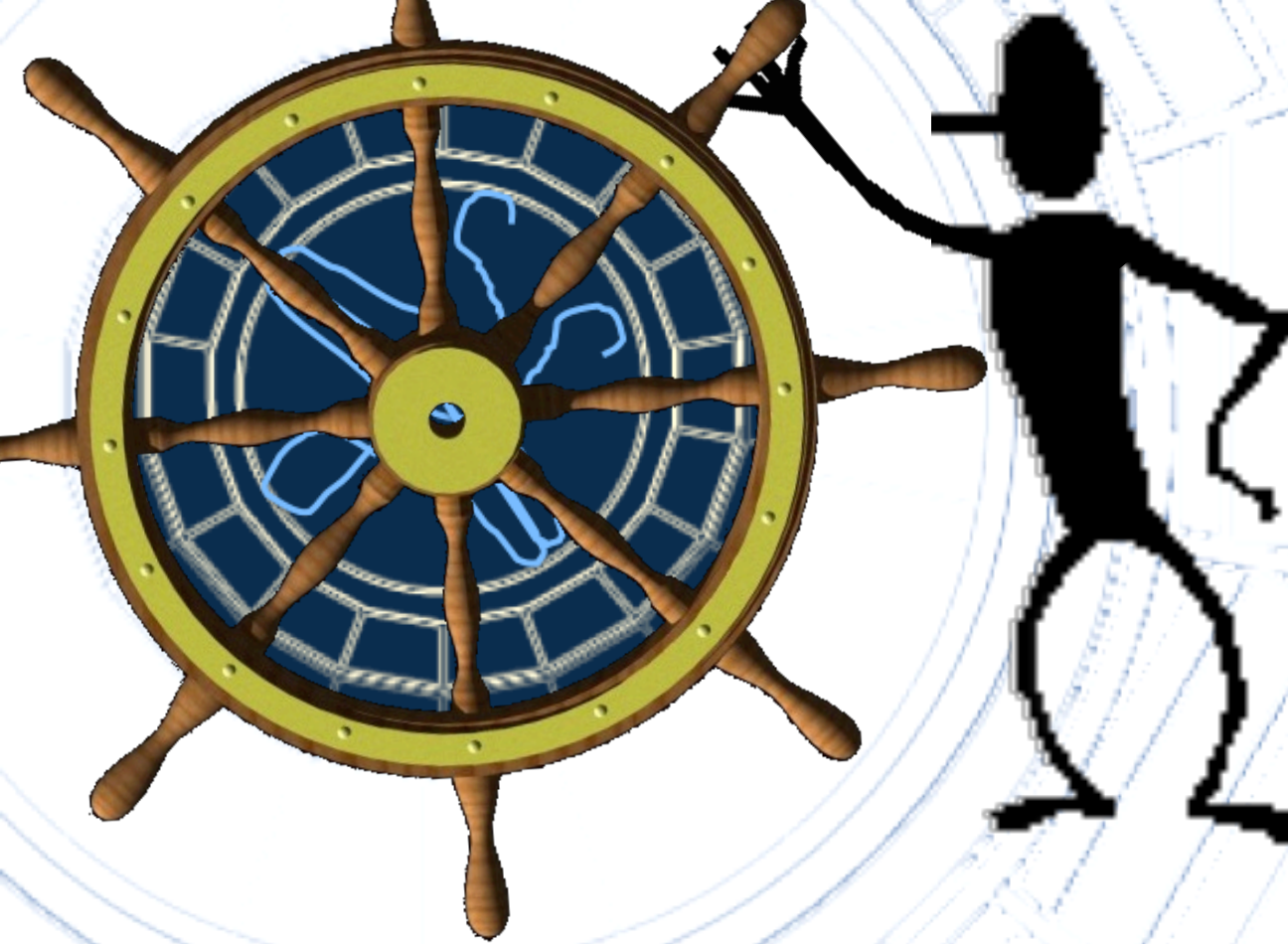
The model definition

p(henomenological)MSSM is a 19-dimensional parameterization of the 124-parameter MSSM at the SUSY scale $M_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$.

- the gaugino mass parameters M_1 , M_2 , and M_3 ;
- the ratio of the Higgs VEVs $\tan \beta = v_2/v_1$;
- the higgsino mass parameter μ and the pseudo-scalar Higgs mass m_A ;
- 10 sfermion mass parameters $m_{\tilde{F}}$, where $\tilde{F} = \tilde{Q}_1, \tilde{U}_1, \tilde{D}_1, \tilde{L}_1, \tilde{E}_1, \tilde{Q}_3, \tilde{U}_3, \tilde{D}_3, \tilde{L}_3, \tilde{E}_3$ (imposing $m_{\tilde{Q}_1} \equiv m_{\tilde{Q}_2}$, $m_{\tilde{L}_1} \equiv m_{\tilde{L}_2}$, etc.), and
- 3 trilinear couplings A_t , A_b and A_τ ,

Minimal assumptions: R parity conserved; no new CP phases; flavor-diagonal sfermion mass matrices and trilinear couplings; 1st/2nd generation degenerate; A-terms negligible; lightest neutralino is the LSP.

- A **well-founded theory**: Sufficiently **generic** 19-dimensional realization of the MSSM at the SUSY scale, which captures most of MSSM's phenomenological features
- **No assumptions** on the **nature of SUSY breaking mechanism**
- **No correlations** between the **sparticle masses**. Has a very rich phenomenology that covers **diverse** (and surprising) **LHC signatures**.
- Allows to **make very generic statements** on sparticle masses.
- Can **make diverse predictions**, on electroweak observables, branching ratios of rare decays, dark matter-related observables, etc.



Analysis

We perform a **global Bayesian analysis** and obtain **posterior probability distributions** of model parameters, masses and predicted observables.



Input parameters

pMSSM input parameters and parameter ranges:

$$-3 \text{ TeV} \leq M_1, M_2 \leq 3 \text{ TeV}$$

$$0 \leq M_3 \leq 3 \text{ TeV}$$

$$-3 \text{ TeV} \leq \mu \leq 3 \text{ TeV}$$

$$0 \leq m_A \leq 3 \text{ TeV}$$

$$2 \leq \tan \beta \leq 60$$

$$0 \leq \tilde{Q}_{1,2}, \tilde{U}_{1,2}, \tilde{D}_{1,2}, \tilde{L}_{1,2}, \tilde{E}_{1,2}, \tilde{Q}_3, \tilde{U}_3, \tilde{D}_3, \tilde{L}_3, \tilde{E}_3 \leq 3 \text{ TeV}$$

$$-7 \text{ TeV} \leq A_t, A_b, A_\tau \leq 7 \text{ TeV},$$

SM input parameters:

$$m_t, m_b(m_b) \text{ and } \alpha_s(M_Z)$$

(allowed to vary in an unlimited range, but constrained by a likelihood)



Building the non-DCS (direct CMS search) prior distribution

non-DCS prior

$$p^{\text{non-DCS}}(\theta) =$$

We obtain a discrete representation of this prior by sampling 20 000 000 points from the pMSSM space using MCMC.

We implement a prompt chargino requirement (since simulation of long-lived particle decays is not available in FastSim) afterwards.

non-DCS
likelihood

$$L(D^{\text{non-DCS}}|\theta)$$

theoretical
constraints

$$p(\text{theory}|\theta)$$

$$\cdot p(\text{prompt } \tilde{\chi}_1^\pm|\theta)$$

prompt
chargino

$$\cdot p_0(\theta)$$

initial
flat prior



non-DCS likelihood and measurements

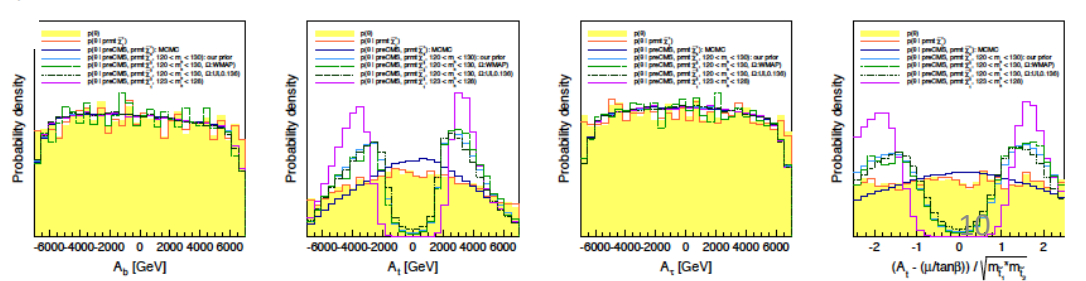
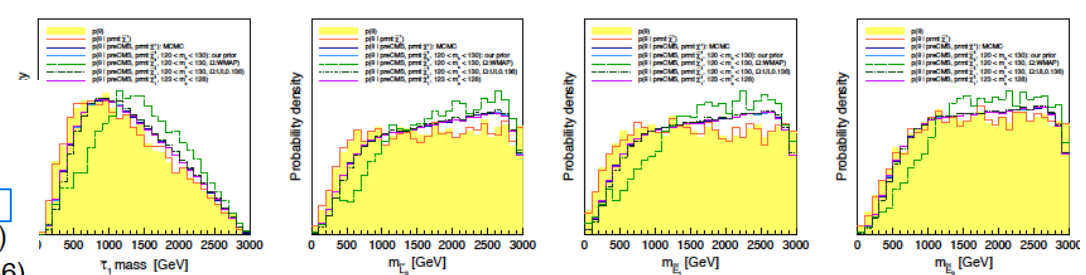
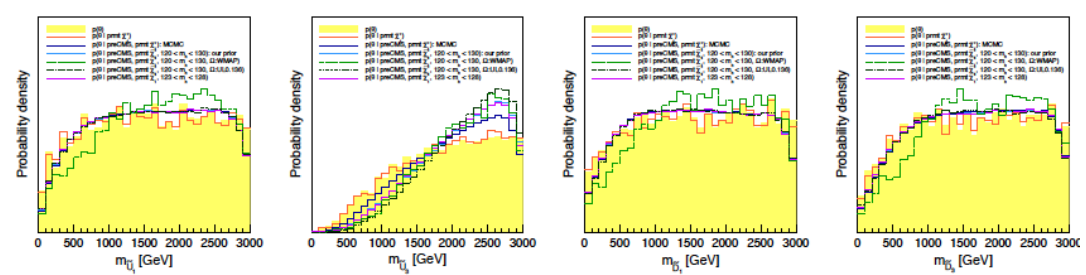
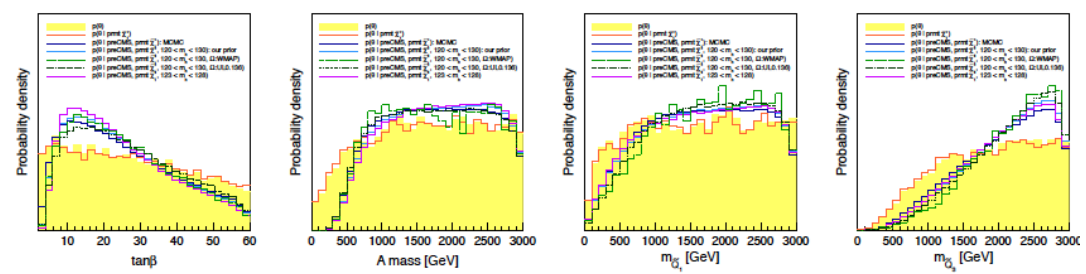
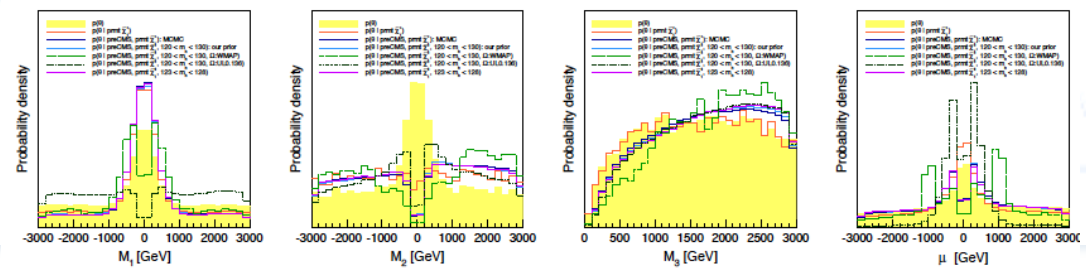
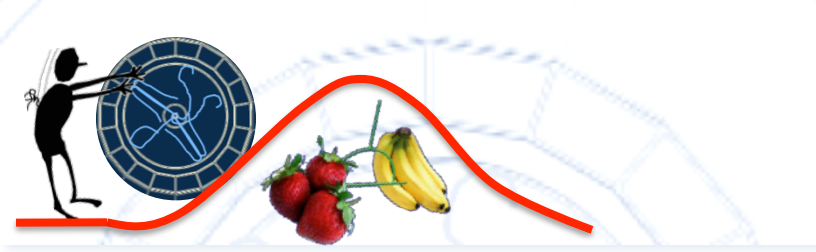
$$L(D^{\text{non-DCS}}|\theta) = \left\{ \prod_i L(D_i^{\text{non-DCS}}|\mu_i(\theta)) \right\}$$

Observable $\mu_j(\theta)$	Constraint $D_j^{\text{non-DCS}}$	Likelihood function $L(D_j^{\text{non-DCS}} \mu_j(\theta))$	MCMC / post-MCMC
$BR(b \rightarrow s\gamma)$ [37, 38]	$(3.55 \pm 0.23^{\text{stat}} \pm 0.24^{\text{th}} \pm 0.09^{\text{sys}}) \times 10^{-4}$	Gaussian	MCMC
$BR(b \rightarrow s\gamma)$ [39]	$(3.43 \pm 0.21^{\text{stat}} \pm 0.24^{\text{th}} \pm 0.07^{\text{sys}}) \times 10^{-4}$	Gaussian	reweight
$BR(B_s \rightarrow \mu\mu)$ [40]	observed CLs curve from [40]	$d(1 - CLs)/d(BR(B_s \rightarrow \mu\mu))$	MCMC
$BR(B_s \rightarrow \mu\mu)$ [41]	$(2.9 \pm 0.7 \pm 0.29^{\text{th}}) \times 10^{-9}$	Gaussian	reweight
$R(B_u \rightarrow \tau\nu)$ [42]	1.63 ± 0.54	Gaussian	MCMC
$R(B_u \rightarrow \tau\nu)$ [39]	1.04 ± 0.34	Gaussian	reweight
Δa_μ [43]	$(26.1 \pm 6.3^{\text{exp}} \pm 4.9^{\text{SM}} \pm 10.0^{\text{SUSY}}) \times 10^{-10}$	Gaussian	MCMC
m_t [44]	$173.3 \pm 0.5^{\text{stat}} \pm 1.3^{\text{sys}}$ GeV	Gaussian	MCMC
m_t [45]	$173.20 \pm 0.87^{\text{stat}} \pm 1.3^{\text{sys}}$ GeV	Gaussian	reweight
$m_b(m_b)$ [42]	$4.19_{-0.06}^{+0.18}$ GeV	Two-sided Gaussian	MCMC
$\alpha_s(M_Z)$ [42]	0.1184 ± 0.0007	Gaussian	MCMC
m_h	pre-LHC: $m_h^{\text{low}} = 112$	1 if $m_h \geq m_h^{\text{low}}$ 0 if $m_h < m_h^{\text{low}}$	MCMC
m_h	LHC: $m_h^{\text{low}} = 120, m_h^{\text{up}} = 130$	1 if $m_h^{\text{low}} \leq m_h \leq m_h^{\text{up}}$ 0 if $m_h < m_h^{\text{low}}$ or $m_h > m_h^{\text{up}}$	reweight
sparticle masses	LEP [46] (via micrOMEGAs [31–33])	1 if allowed 0 if excluded	MCMC



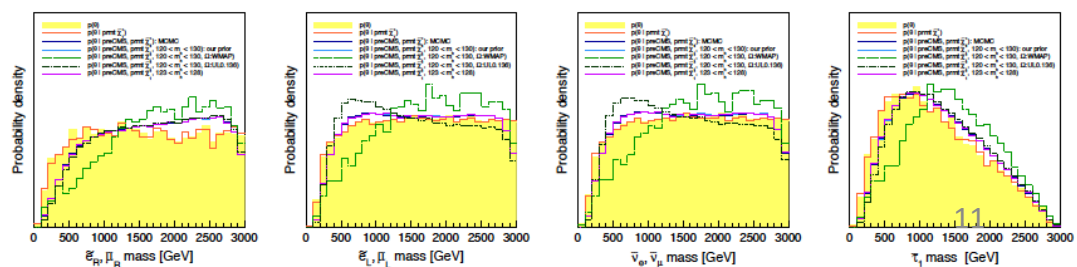
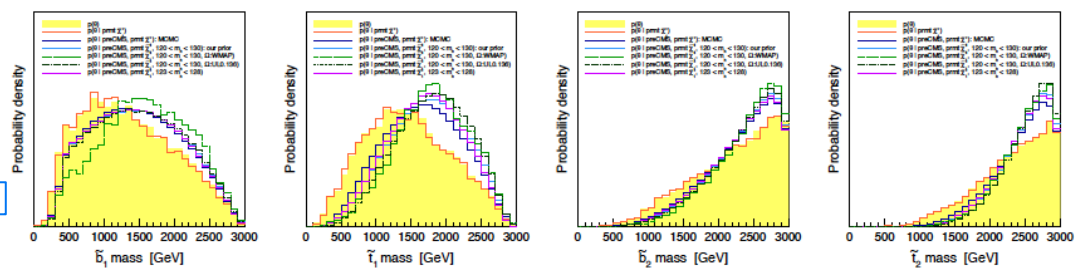
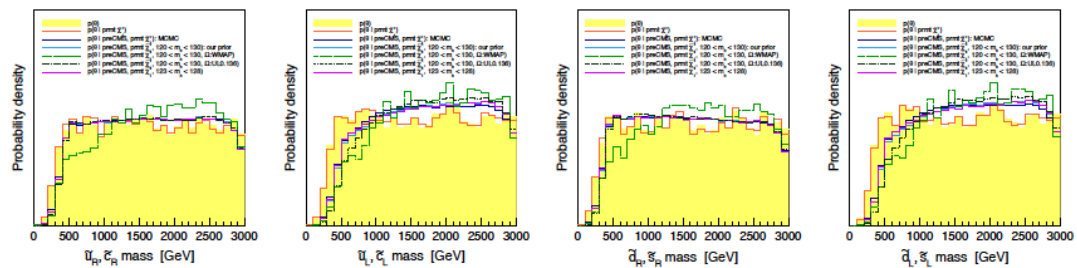
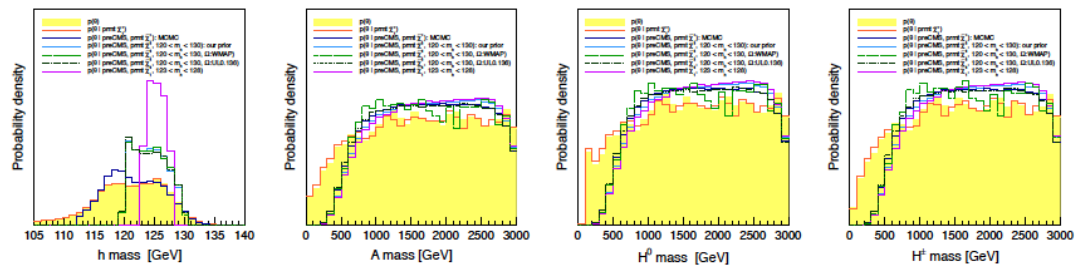
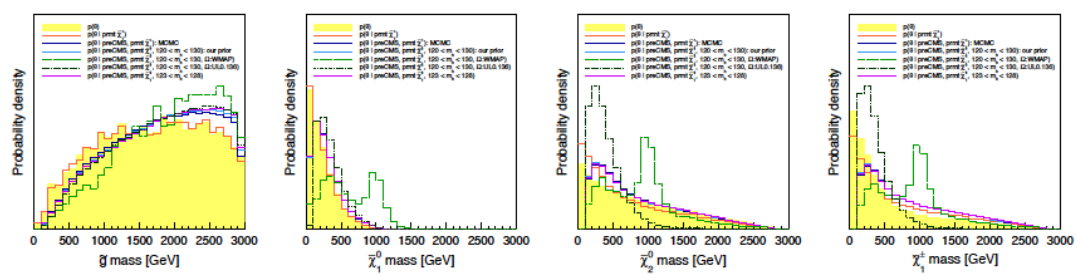
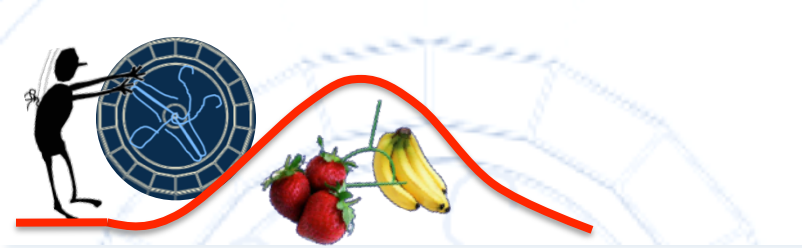
Tools

- SUSY spectrum calculation: **SOFTSUSY 3.3.1**
- EW and B-physics observables: **SuperIso 3.3**
- Dark matter relic density, DD cross section: **micrOMEGAs 2.4.5**
- SUSY mass limits: **micrOMEGAs 2.4.5**
- SUSY decays: **SUSYHIT SUSYHIT (SDECAY1.3b, HDECAY3.4)**
- SUSY event generation: **PYTHIA 6**
- Interfaces: **SUSY Les Houches Accord**
- Detector simulation: **CMSSW, FASTSIM**



1D preCMS posterior probability distributions for the pMSSM input parameters. Effects of the light Higgs mass and DM relic density constraints are shown.

- $p(\theta)$
- $p(\theta | \text{prmt } \tilde{\chi}_1^+)$
- $p(\theta | \text{preCMS, prmt } \tilde{\chi}_1^+)$ Our prior
- $p(\theta | \text{preCMS, prmt } \tilde{\chi}_1^+, 120 < m_h < 130)$: our prior
- $p(\theta | \text{preCMS, prmt } \tilde{\chi}_1^+, 120 < m_h < 130, \Omega: \text{WMAP})$
- $p(\theta | \text{preCMS, prmt } \tilde{\chi}_1^+, 120 < m_h < 130, \Omega: \text{ULO.136})$
- $p(\theta | \text{preCMS, prmt } \tilde{\chi}_1^+, 123 < m_h < 128)$

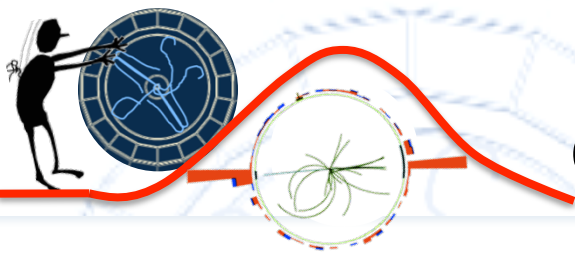


1D preCMS posterior probability distributions for selected sparticle masses. Effects of the light Higgs mass and DM relic density constraints are shown.

- $p(\theta)$
- $p(\theta \mid \text{prmt } \tilde{\chi}_1^+)$
- $p(\theta \mid \text{preCMS, prmt } \tilde{\chi}_1^+)$: MCMC
- $p(\theta \mid \text{preCMS, prmt } \tilde{\chi}_1^+, 120 < m_h < 130)$: our prior
- $p(\theta \mid \text{preCMS, prmt } \tilde{\chi}_1^+, 120 < m_h < 130, \Omega: \text{WMAF})$
- $p(\theta \mid \text{preCMS, prmt } \tilde{\chi}_1^+, 120 < m_h < 130, \Omega: \text{ULO.1})$
- $p(\theta \mid \text{preCMS, prmt } \tilde{\chi}_1^+, 123 < m_h < 128)$

Our prior





Building the CMS posterior distribution - I

CMS posterior distribution

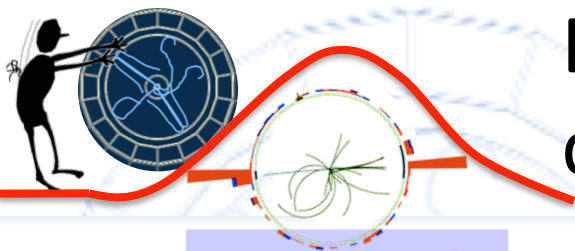
$$p(\theta|D^{\text{CMS}}) = L(D^{\text{CMS}}|\theta) p^{\text{non-DCS}}(\theta)$$

CMS likelihood

based on simulation of signal events:
per point 10000 events, `PYTHIA + FastSim`

analysis groups provide necessary (simple)
input to calculate an approximate likelihood

We select 7300 points out of the non-DCS point set with a prompt chargino and weight each by $L(D^{\text{CMS}}|\theta)$.



Building the CMS posterior distribution - II

$$p(\theta | D^{\text{CMS}}) = L(D^{\text{CMS}} | \theta) p^{\text{non-DCS}}(\theta)$$

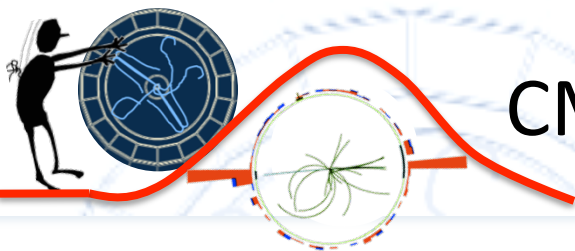
likelihood method 1: for pure count experiments

full treatment of statistical uncertainties bkg prediction,
correlation between systematic uncertainties ignored

likelihood method 2: for likelihood based experiments

full treatment of systematic uncertainties bkg prediction,
statistical uncertainties treated in asymptotic limit

- ▶ within each analysis signal regions are combined if they are statistically independent by multiplying the per signal region likelihoods
- ▶ for combinations of overlapping signal regions and analyses: see further

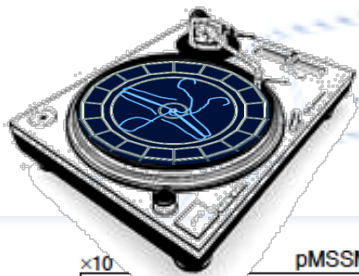


CMS analyses

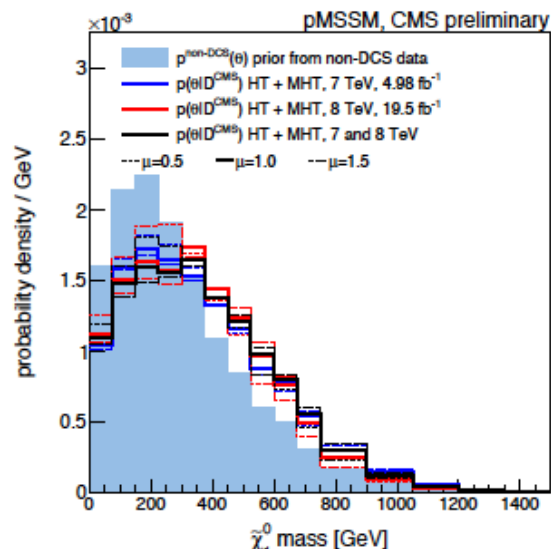
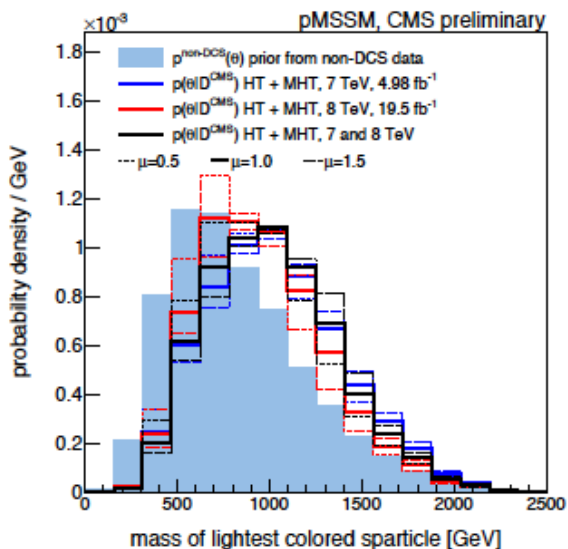
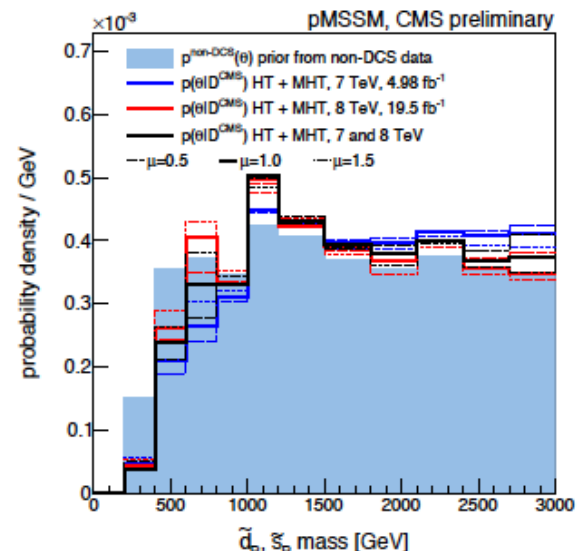
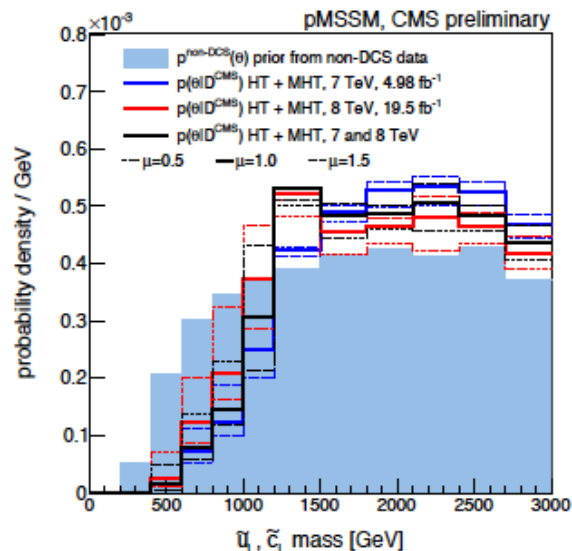
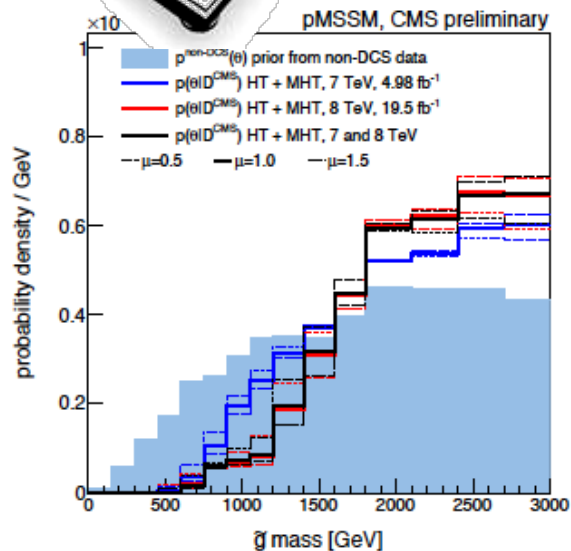
Analysis	\sqrt{s}	$\int dL$	CERN doc.
in the PAS			
$H_T + H_T^{miss}$ search	7 TeV	4.98 fb ⁻¹	SUS-12-011
$H_T + E_T^{miss} + b$ -jets search	7 TeV	4.98 fb ⁻¹	SUS-12-003
EW prod of $\tilde{\chi}^\pm, \tilde{\chi}^0, \tilde{l}$	7 TeV	4.98 fb ⁻¹	SUS-12-006
$H_T + H_T^{miss}$ search	8 TeV	19.5 fb ⁻¹	SUS-13-012
$H_T + E_T^{miss} + b$ -jets search	8 TeV	19.4 fb ⁻¹	SUS-12-024
EW prod of $\tilde{\chi}^\pm, \tilde{\chi}^0, \tilde{l}$	8 TeV	19.5 fb ⁻¹	SUS-12-006

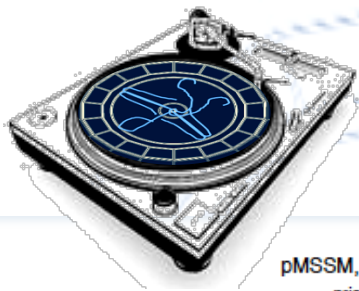


Results

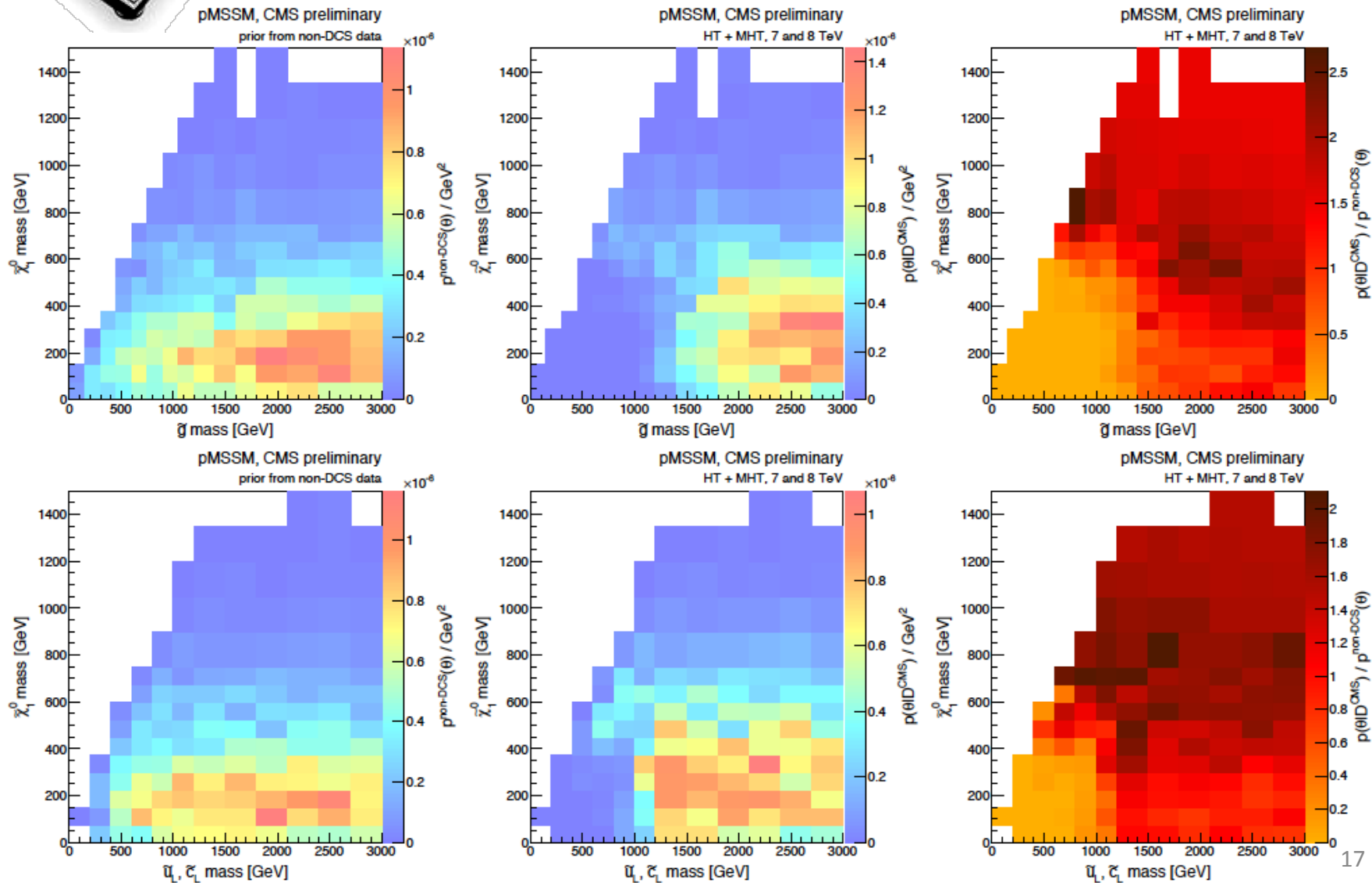


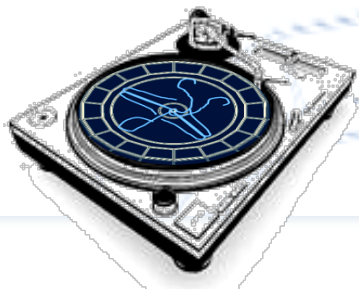
Impact of the HT + MHT search - I



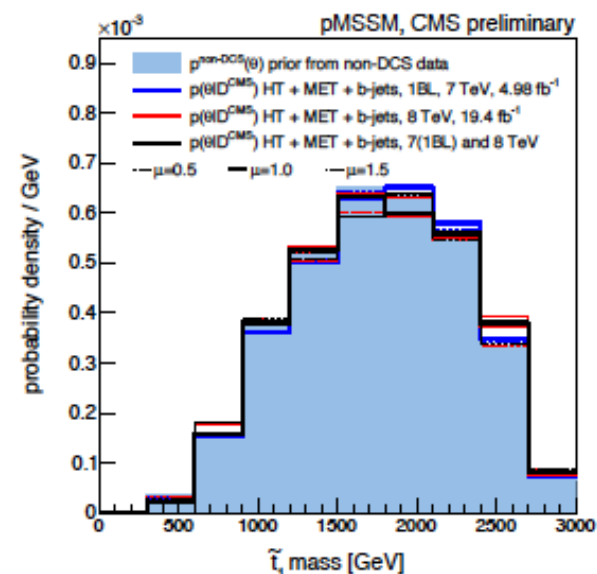
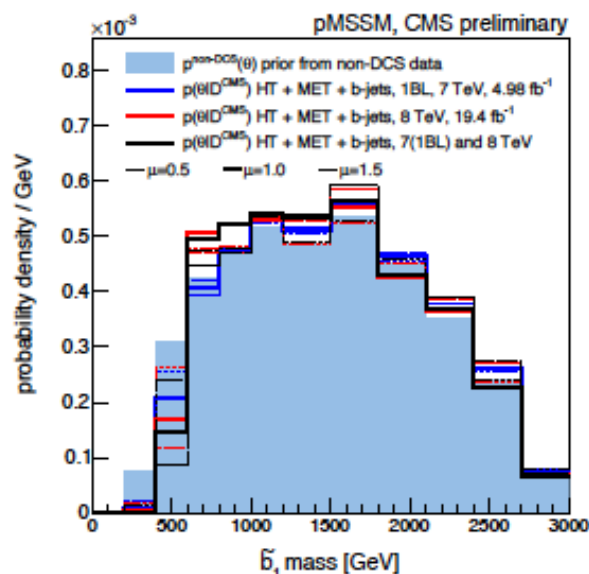
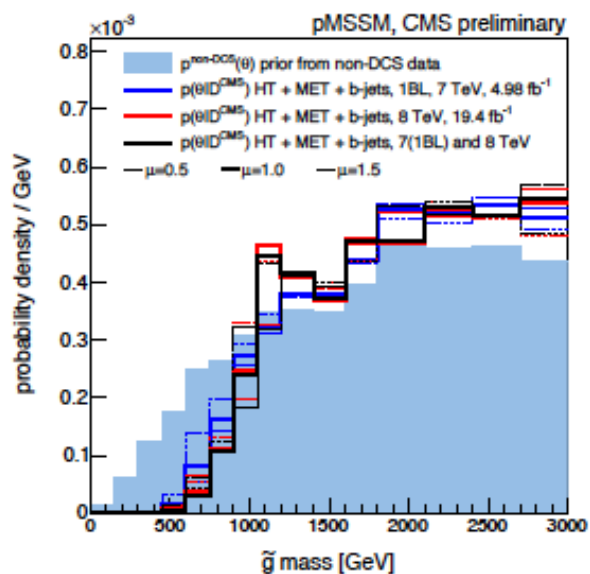


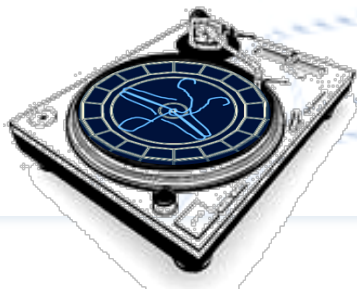
Impact of the HT + MHT search - II



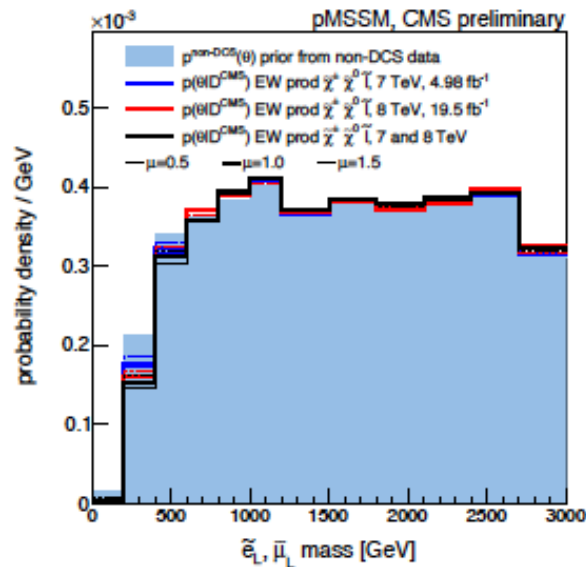
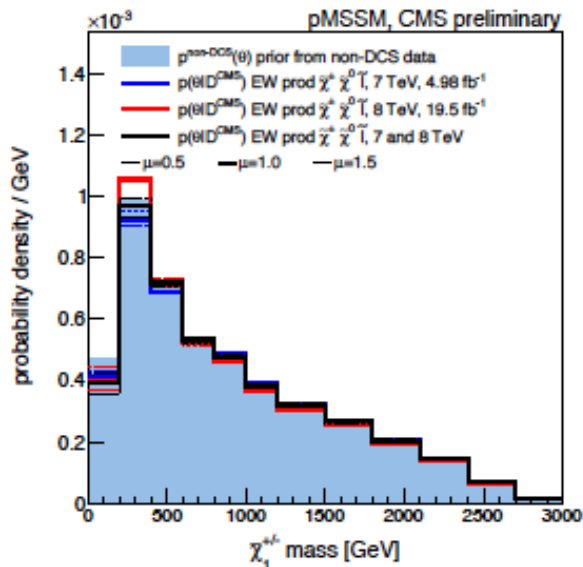
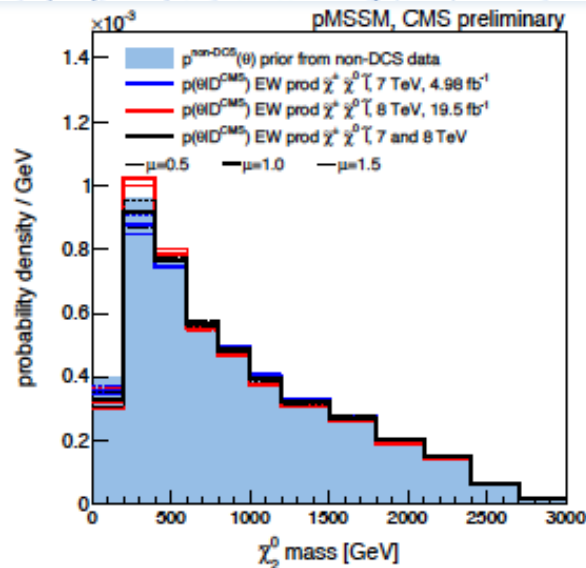
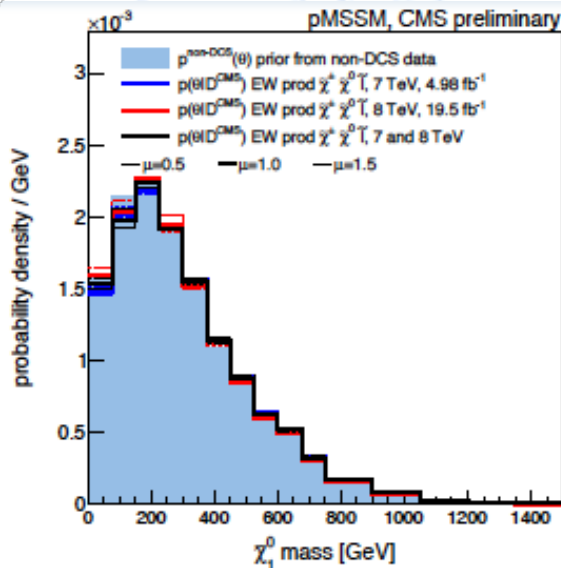


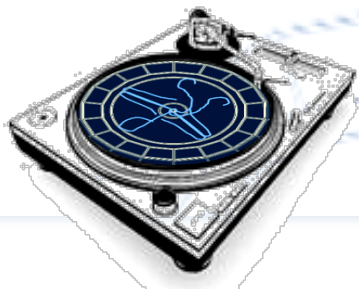
Impact of the HT + MET + jets search





Impact of the EW search





The Z significance

Bayesian analog of frequentist “ n -sigma’

$$Z(\theta) = \text{sign}(\ln B_{10}(\theta)) \sqrt{2 |\ln B_{10}(\theta)|}$$

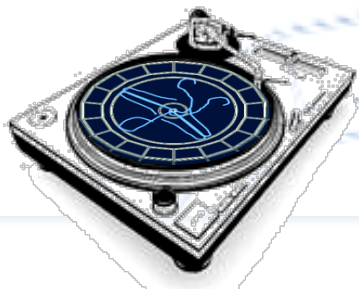
$$B_{10}(\theta) = L(D^{\text{CMS}}|\theta, H_1) / L(D^{\text{CMS}}|H_0)$$

$Z(\theta) < -1.64$ \approx signal θ excluded, 95% CL

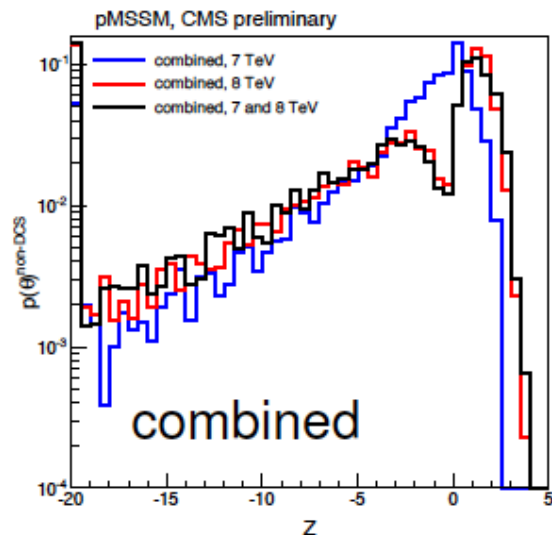
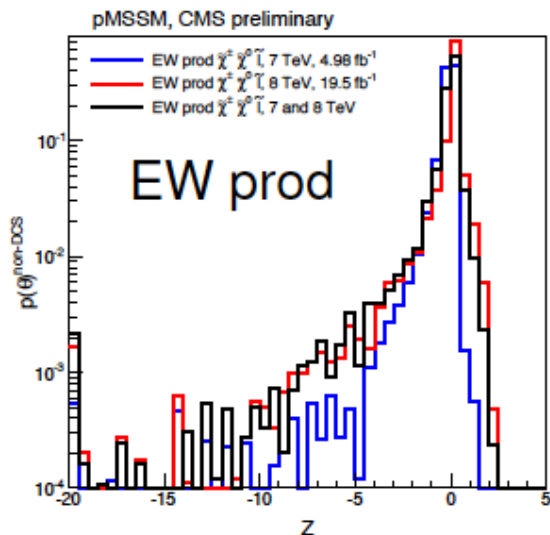
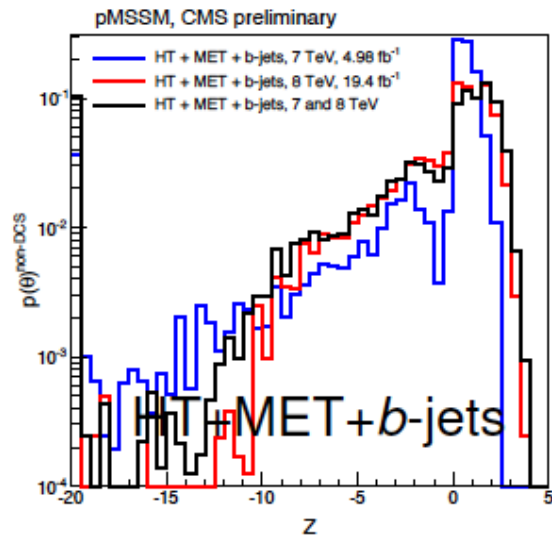
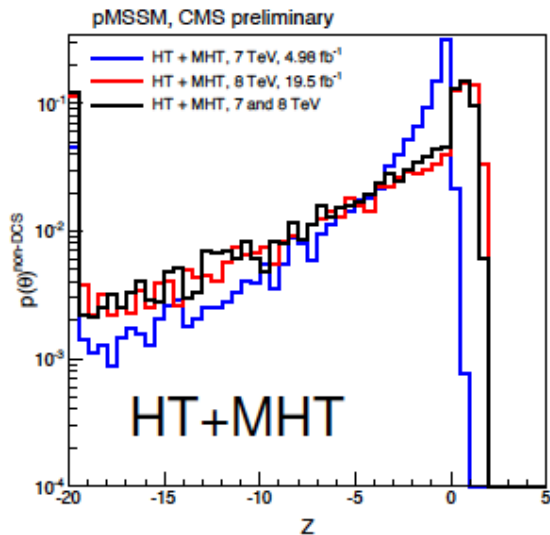
$Z(\theta) > 5$ \approx 5- σ discovery

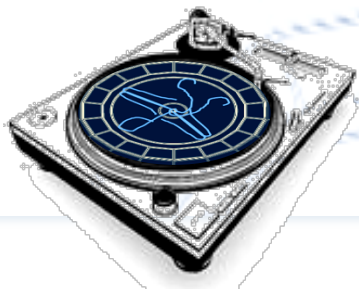
Combination:

Z : Z_i with $\max |Z_i|$



Z distributions

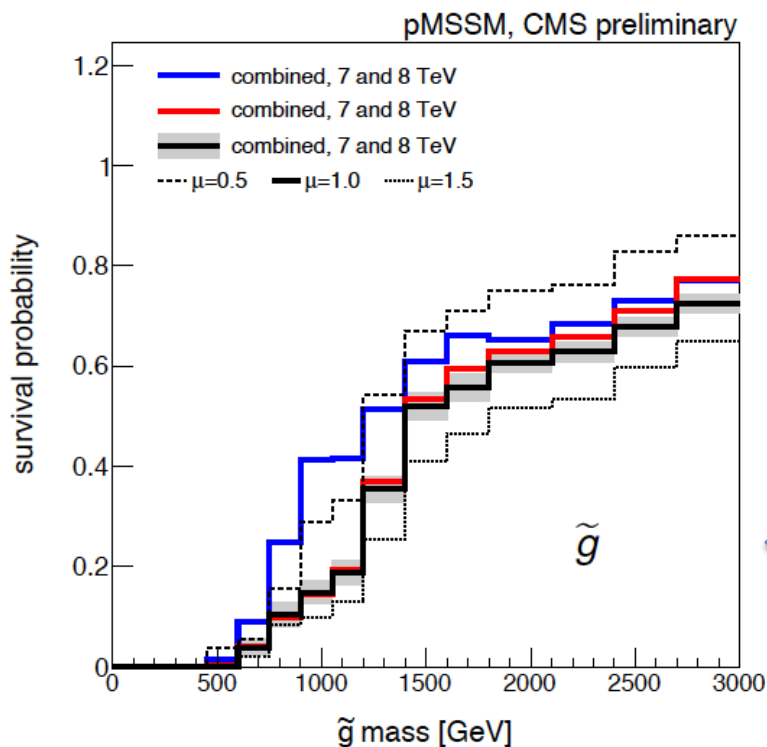
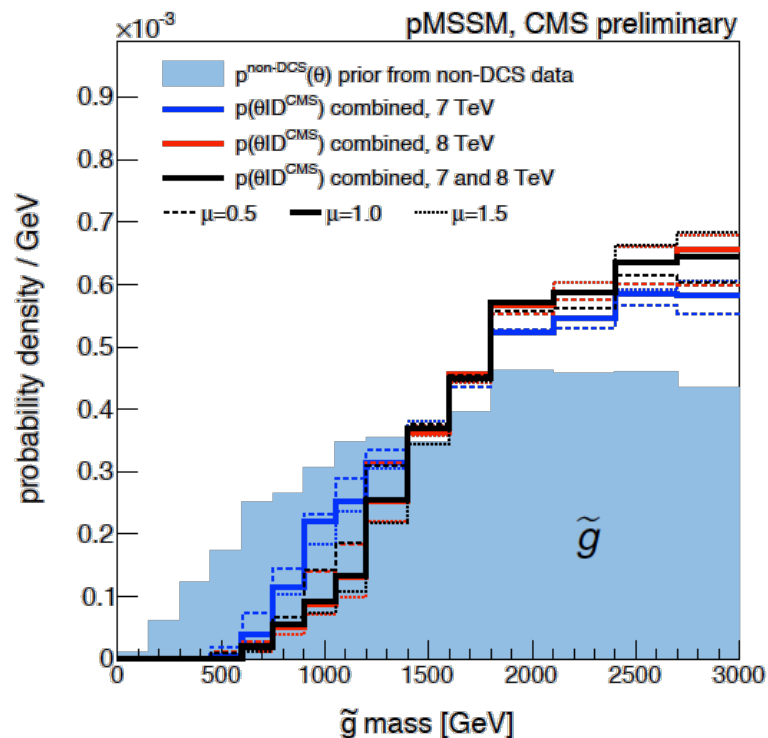




Z likelihood and survival probability

Z likelihood:

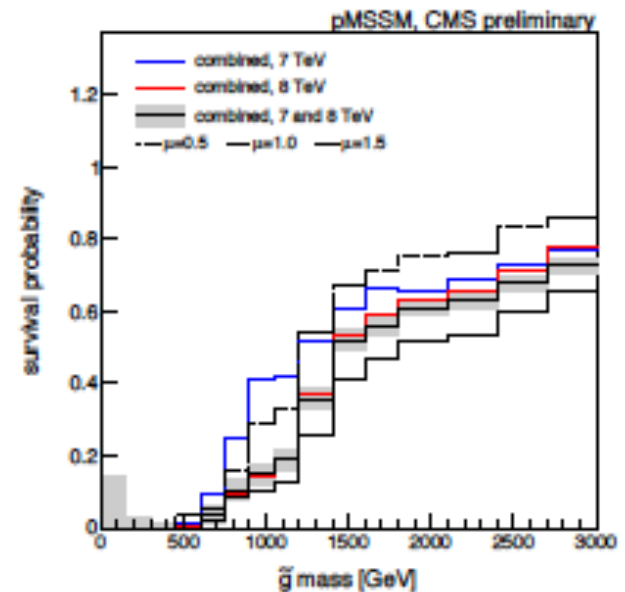
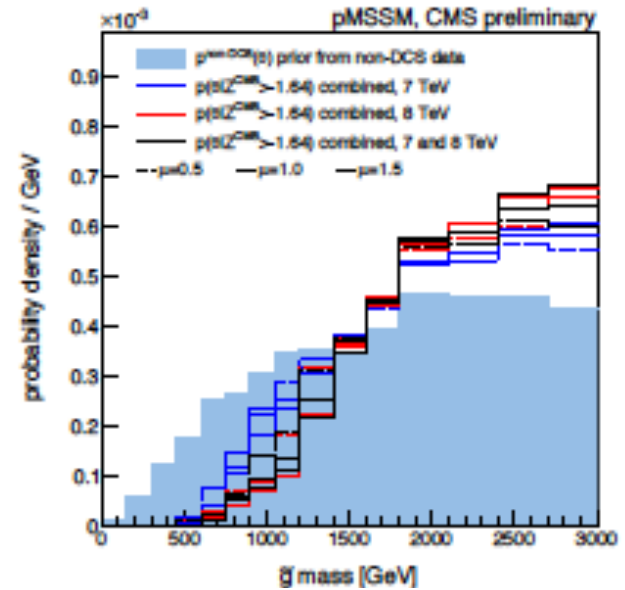
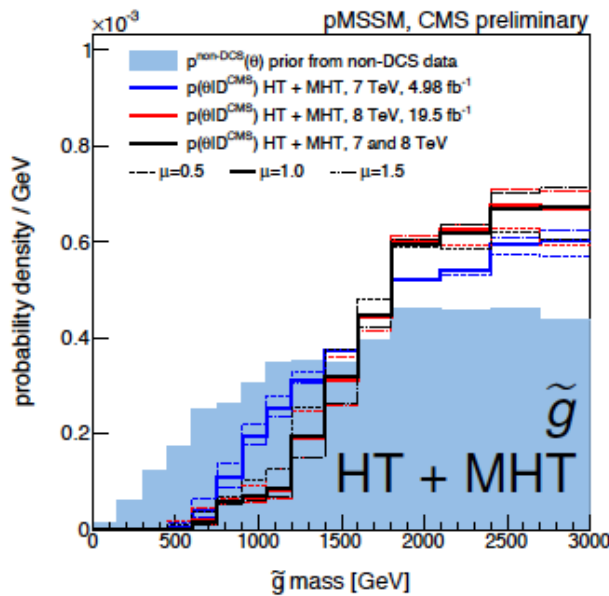
$$L(D^{\text{CMS}}|\theta) = \begin{cases} 0 & \text{if } Z_{\text{best}} < -1.64 \\ 1 & \text{if } Z_{\text{best}} \geq -1.64 \end{cases}$$



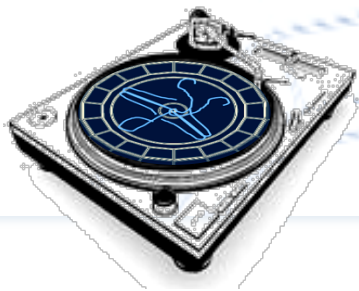
Survival probability is defined as the fraction of points per bin surviving the cut $Z > -1.64$.



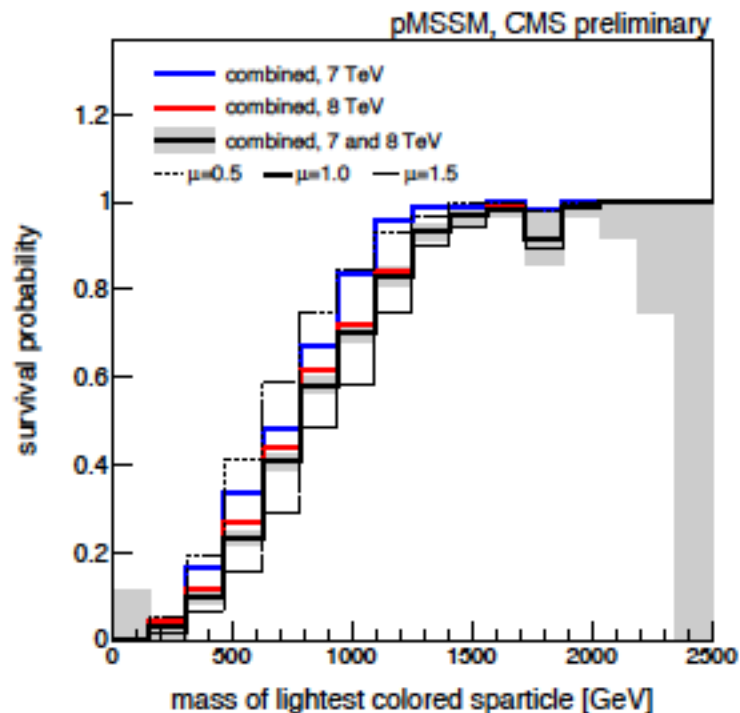
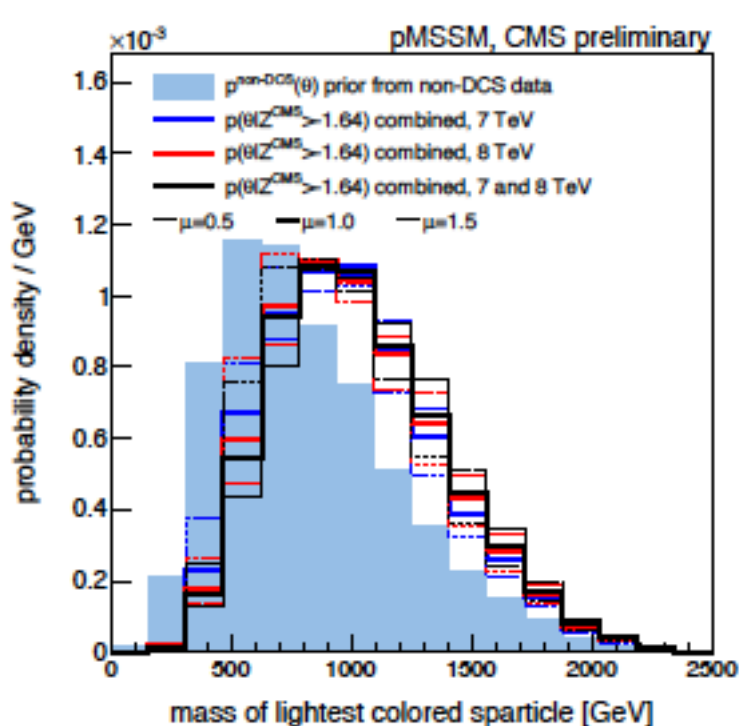
Combined impact: gluinos



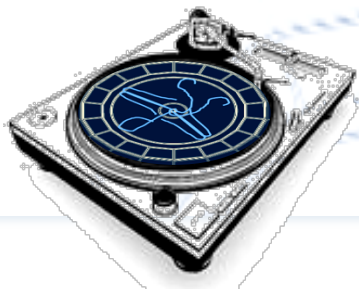
- HT + MHT search drives sensitivity in the colored sector.
- Gluino mass is strongly disfavored below 1200 GeV.



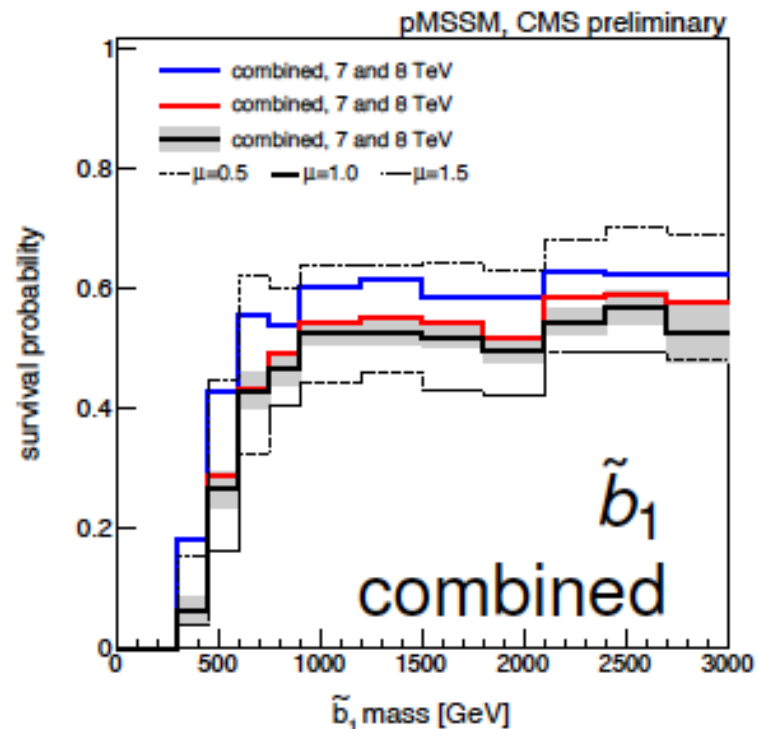
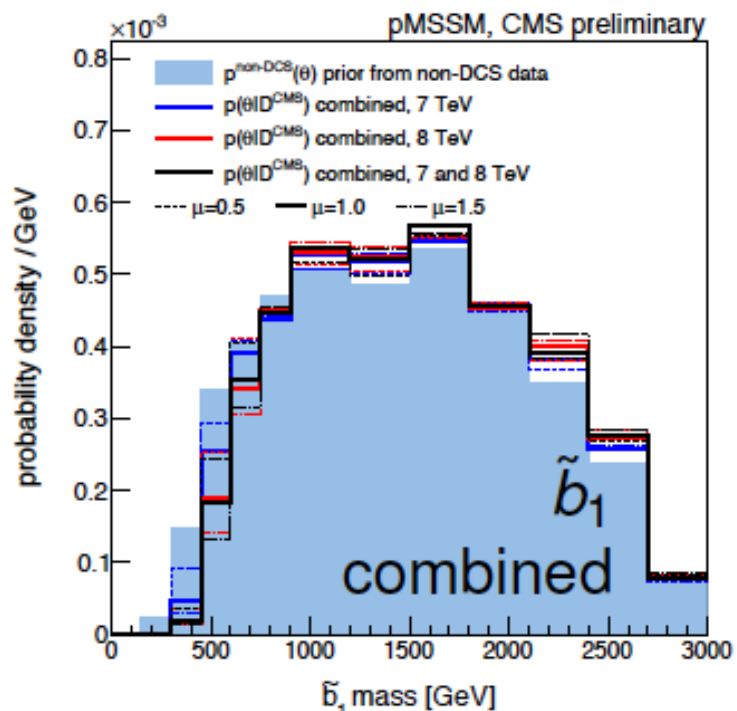
Combined impact: Colored particles



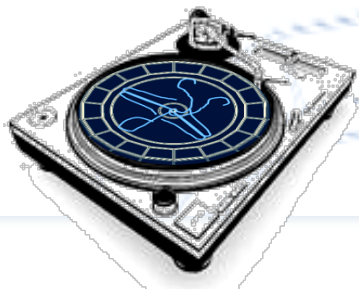
CMS sensitivity requires the model to have a colored sparticle with mass below 1300 GeV.



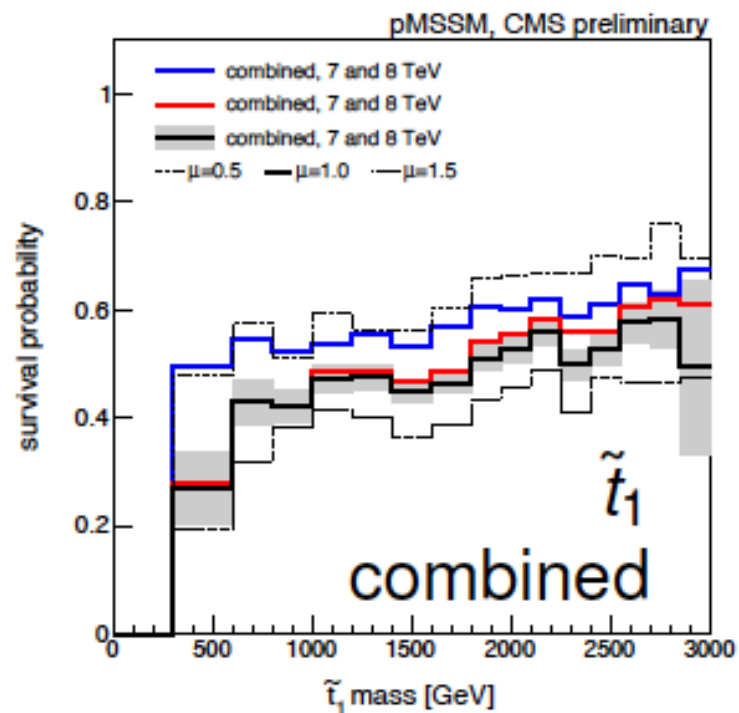
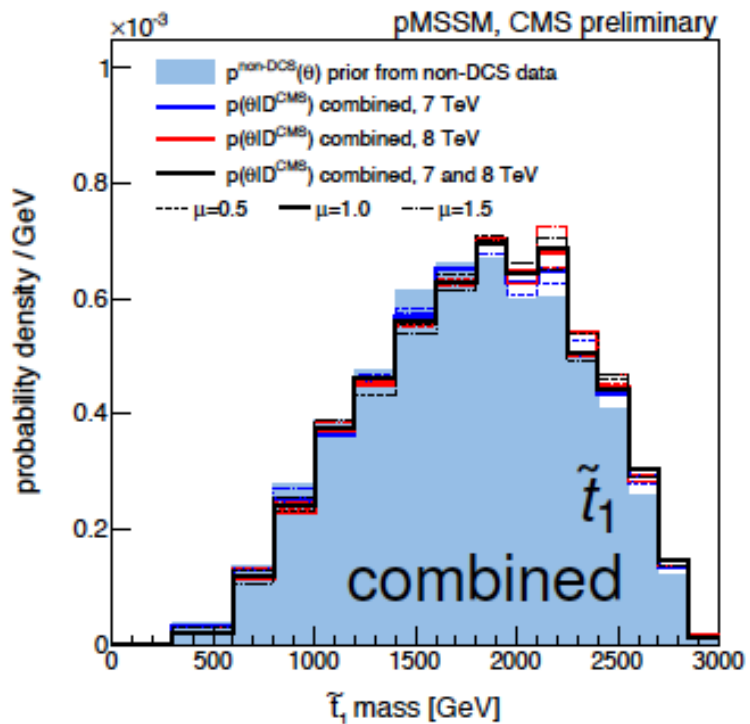
Combined impact: 3rd gen - I



CMS sensitivity to cases with sbottom mass below 600 GeV.



Combined impact: 3rd gen - II

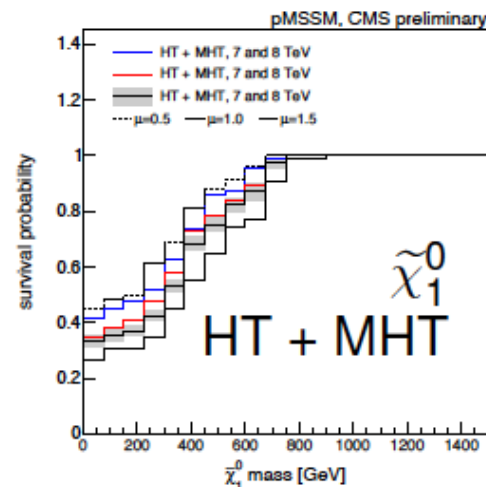
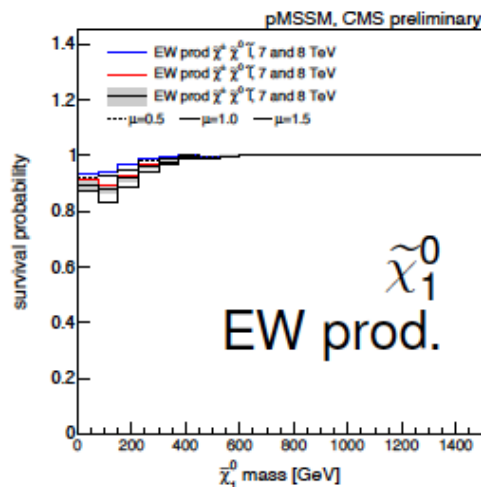
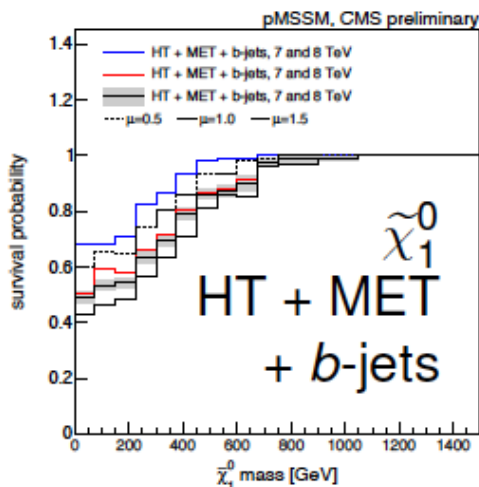
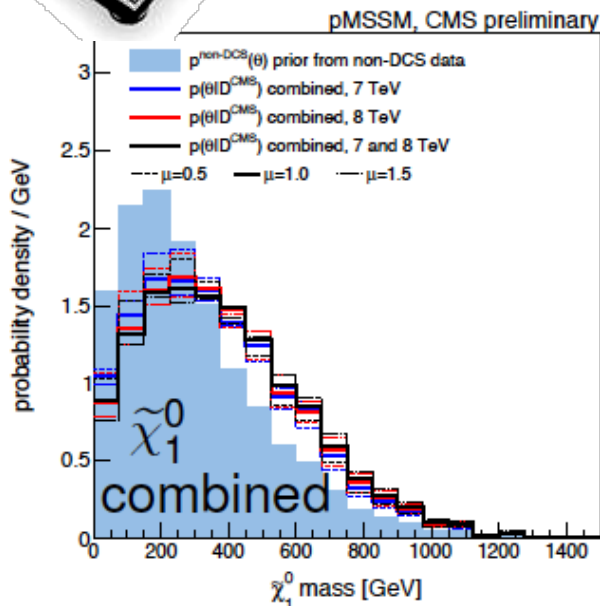


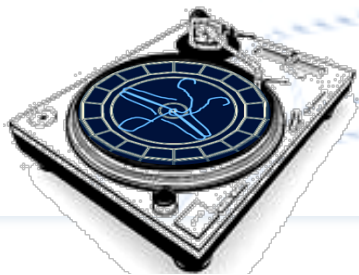
Mild CMS impact below 1 TeV



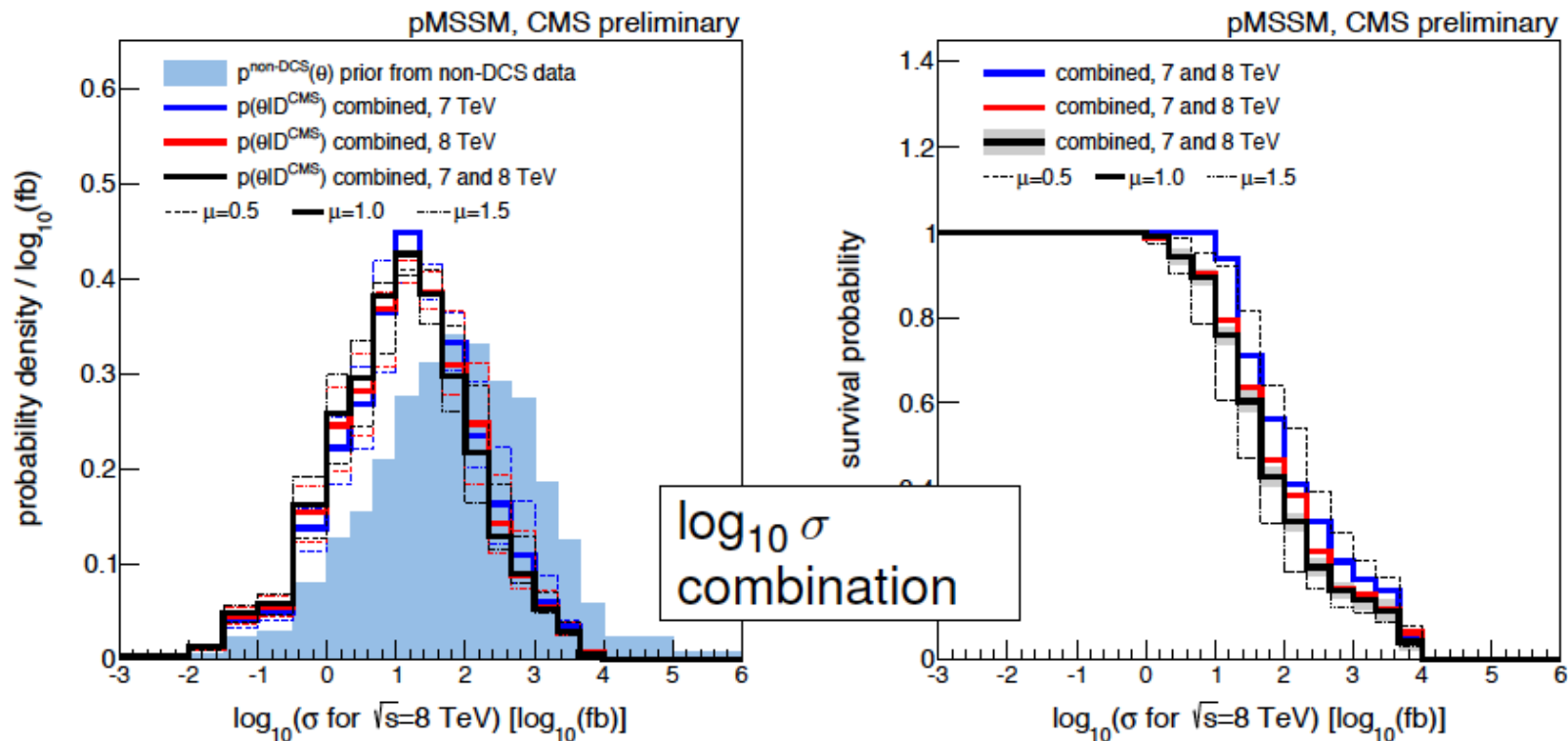
Combined impact: LSP

- There is CMS impact on the LSP mass.
- For the hadronic analyses, the impact arises indirectly. As colored sparticle masses are pushed up, LSPs with higher masses also become favorable.

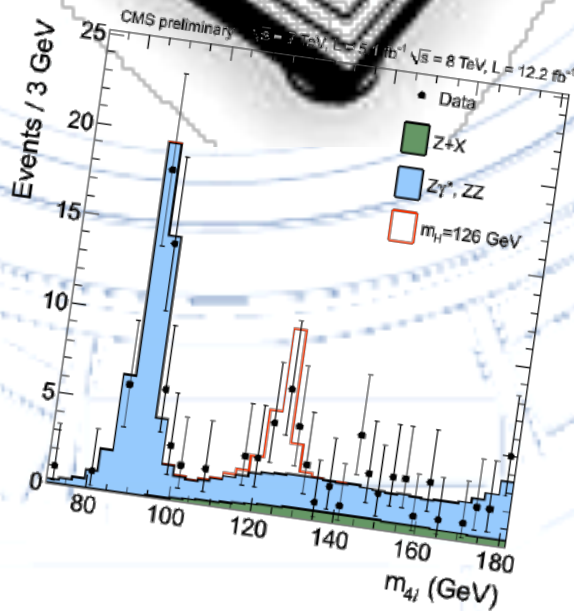
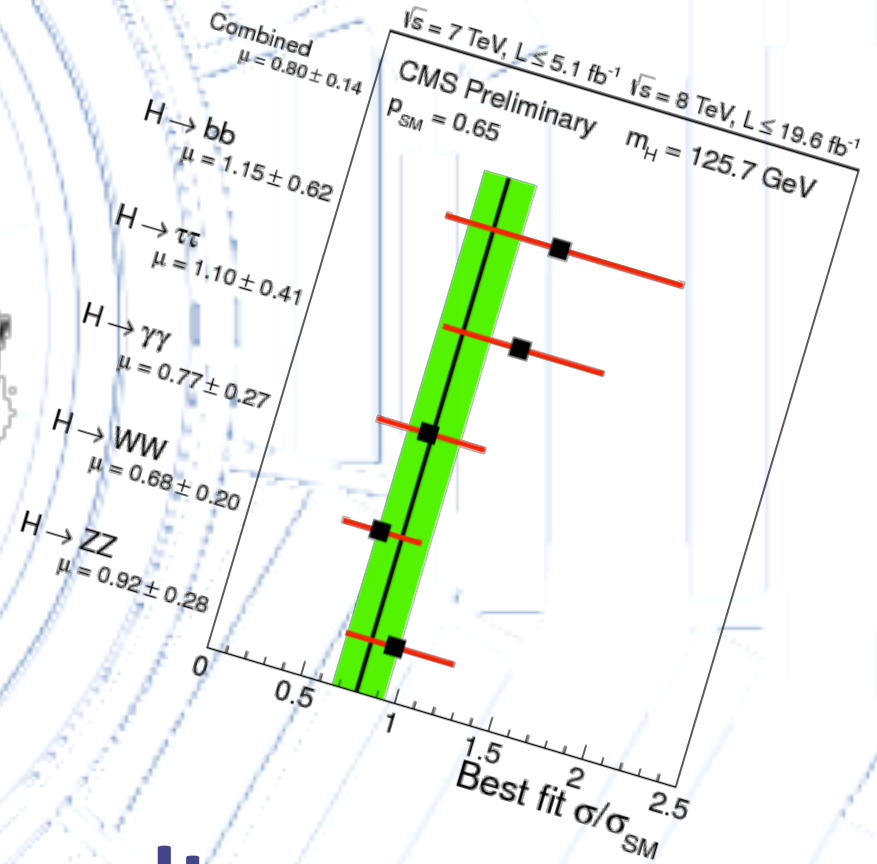
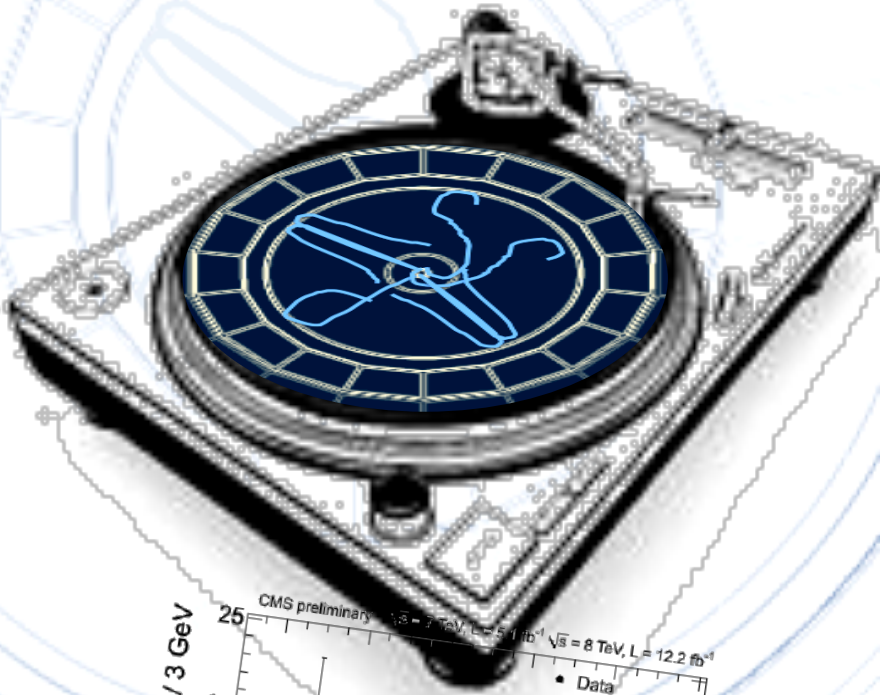




Combined impact: cross section



- Most likely cross section moved down from 100 to 10 fb.
- Scenarios with cross sections > 100 or 1000 fb are still possible.
- Probing scenarios with cross sections < 1 fb.



Results:
Consequences for
the Higgs bosons



Higgs production/decay rates

We take a closer look at the points within the $123 < m_h < 128$ window.

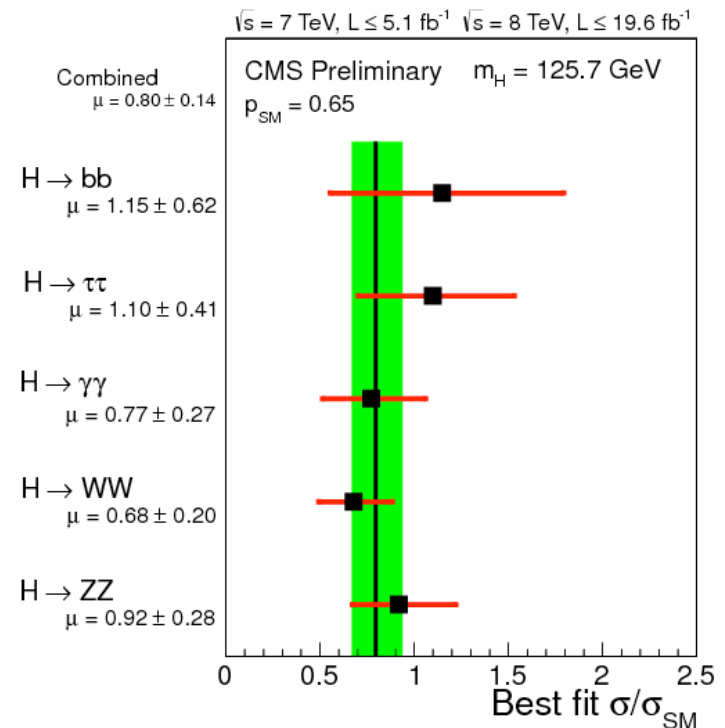
We examine the **production/decay rates** for the Higgs boson.

We use the **μ ratio** defined as

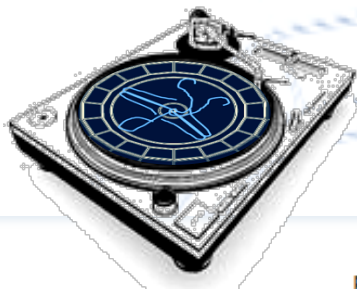
$$\mu_X(Y) = \frac{\sigma(X \rightarrow h) \text{BR}(h \rightarrow Y)}{\sigma_{\text{SM}}(X \rightarrow h_{\text{SM}}) \text{BR}(h_{\text{SM}} \rightarrow Y)}$$

$$X = gg, \text{VBF}, Vh$$

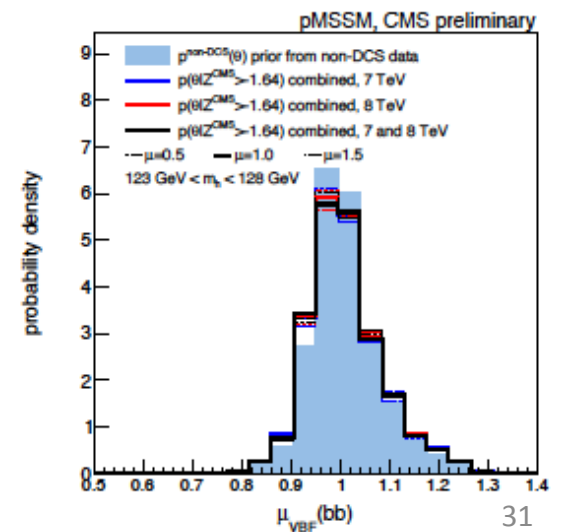
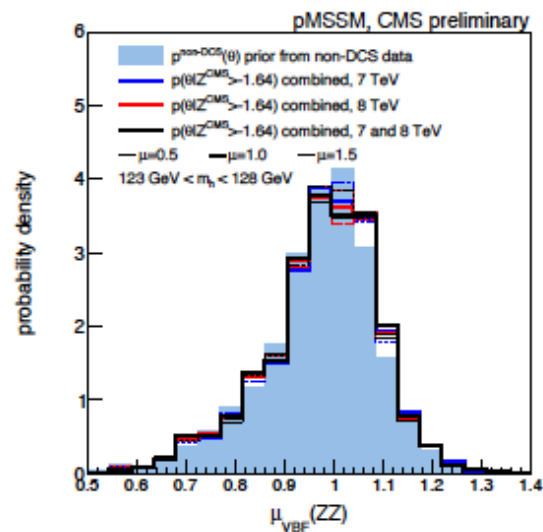
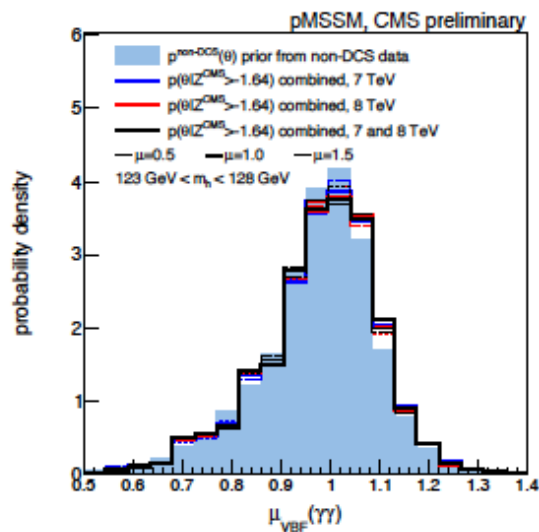
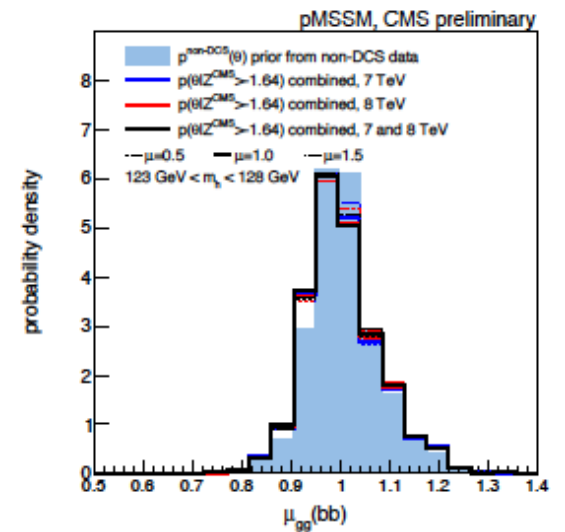
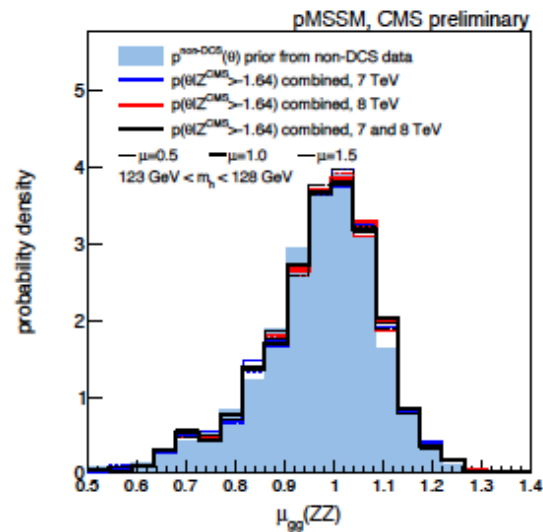
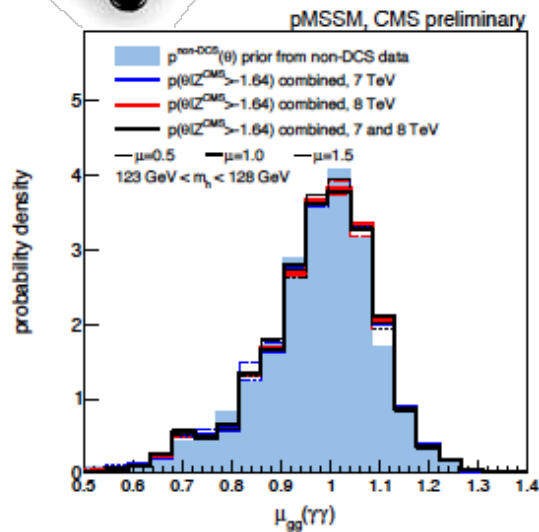
$$Y = \gamma\gamma, ZZ, WW, \tau\tau, bb$$

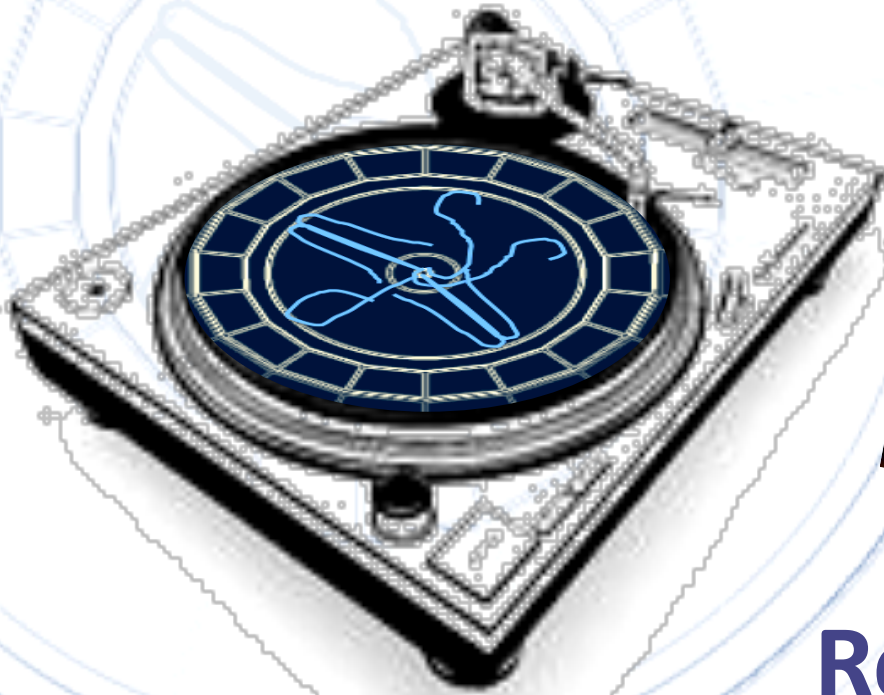


CMS-PAS-HIG-13-005



μ ratio distributions

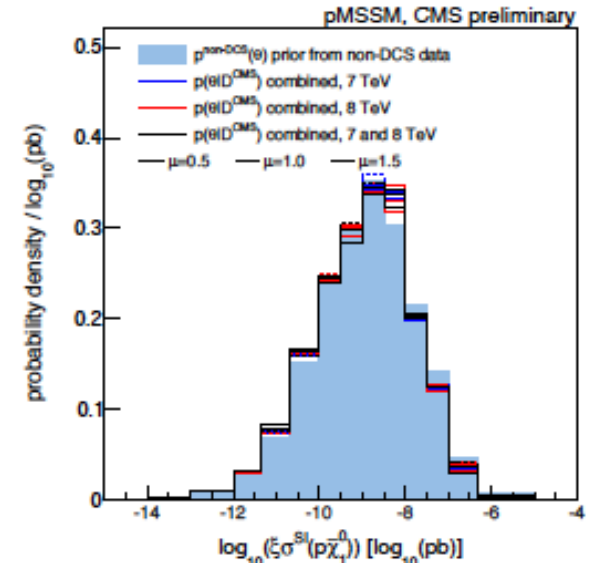
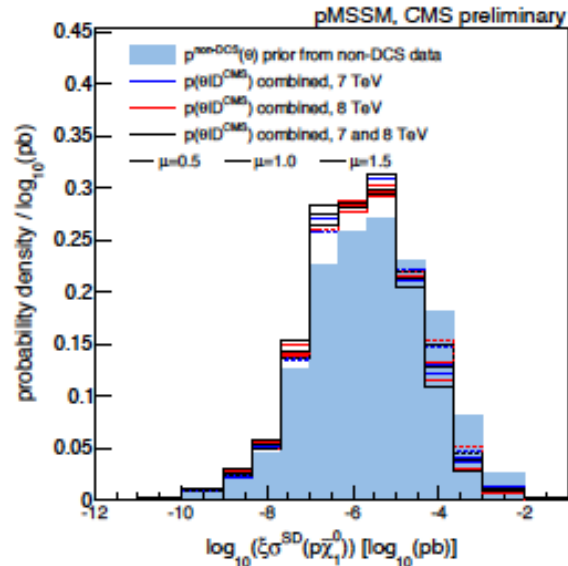
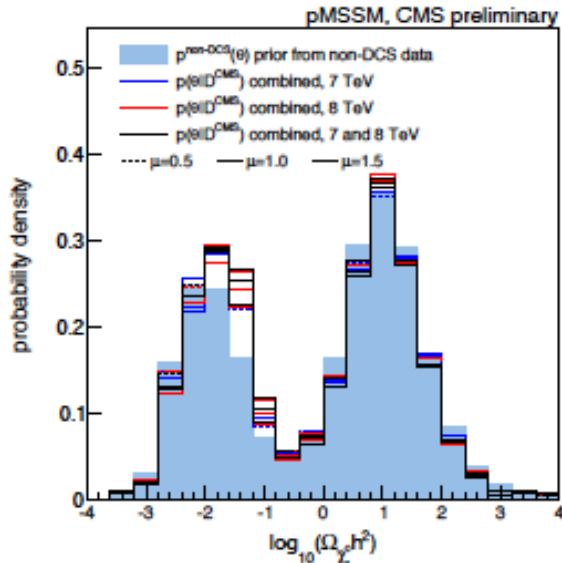




**Results:
Consequences for
dark matter
observables**



DM implications

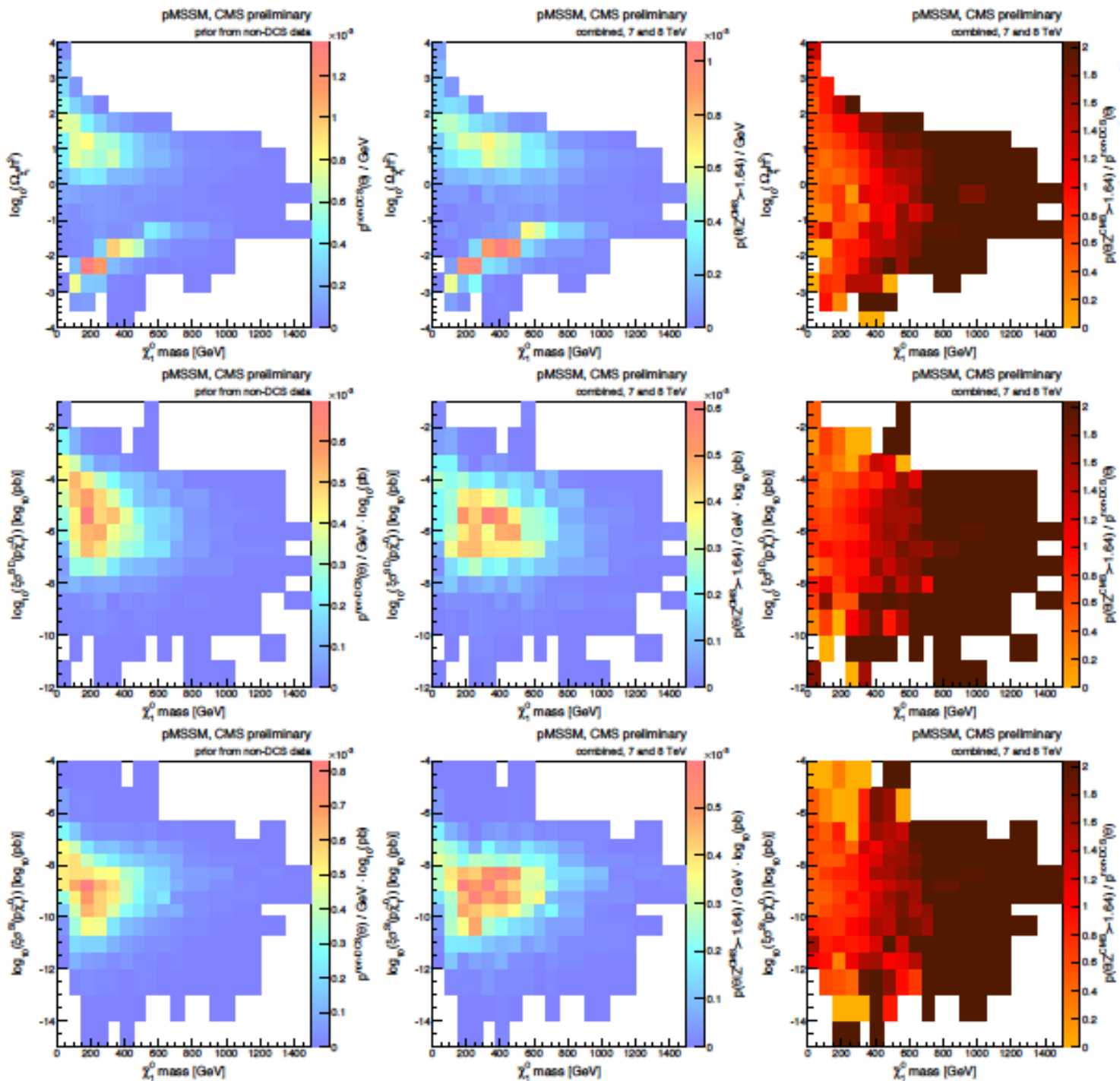


impact on DM relic density, spin-dependent and spin-independent DM direct detection cross sections.

- The Planck relic density window is at the dip.
- CMS data favors lower relic density.
- Mild effects on nucleon-LSP cross sections.



impact on DM relic density, spin-dependent and spin-independent DM direct detection cross sections vs. LSP mass. Mid CMS impact.





Conclusions

- We have investigated the impact of a subset of the 7 and 8 TeV CMS SUSY searches on a potentially accessible sub-space of the pMSSM.
- We investigated a subspace with sparticle masses up to 3 TeV.
- We interpreted the results of 3 searches: HT + MHT, HT + MET + b jets and a dedicated EW sparticle search.
- CMS impact mostly on the colored sector (gluino reach up to 1300 GeV), but also sensitive to EW sparticles up to 600 GeV.
- Also studied the effect of the SUSY searches on Higgs and DM variables. Only very mild impact.
- Currently extending the pMSSM interpretation by adding more CMS searches (inclusive, dedicated stop and EW searches).

