### Test Beam Analysis with a Telescope and Irradiated Strip Sensors

# Combining an analog ALiBaVa strip DUT with a digital EUDET/AIDA pixel telescope





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### Outline

- > Test beam setup & measurements
  - Epitaxial strip sensors for the CMS Phase-II upgrade
  - ALiBaVa & telescope setup

Integrating an analog strip sensor into a digital pixel telescope

- ALiBaVa processors for EUTelescope
- Challenges and changes to EUTelescope

Sensor analysis and results



#### **Test Beam Setup and Measurements**



### **Measurement Goals & Program**

#### > Phase-II upgrade of the CMS tracker:

- Investigate possible future sensor materials
  - $\rightarrow$  Could epitaxial silicon be a HL-LHC radiation-hard sensor material?

#### Sensors / irradiations:

- 19 epitaxial, float-zone and magnetic czochralski silicon strip sensors from a CMS vendor run: 64 strips, 80µm pitch, active thickness 70µm 200µm
- Proton irradiations up to  $\Phi$  = 1.3e16 n<sub>eq</sub>/cm<sup>2</sup> at CERN (23 GeV) and Los Alamos (800 MeV)
- Both n- and p-type material, with p-stop and p-spray isolation
- 5 GeV e<sup>+</sup>/e<sup>-</sup> from beams 21 & 22 with DATURA and ACONITE telescopes

#### Measurement program:

- 3 beam incidence angles (0° to 51.3°) to investigate charge sharing
- How are signal/noise levels at different bias voltages (up to ±1000V)?
- Can the sensors still be used after extremely high irradiations?



### ALiBaVa Setup

A Liverpool Barcelona Valencia Read-out System

- DAQ in-a-box system based on LHCb Beetle chip to read out strip sensor data
- Motherboard for PC connection and data processing
- Sensor wire-bonded to Daughterboard with Beetle ROC
- Analog read-out of the 2 \* 128 channels every event
- Serial communication on two differential lines
- Rate limited to ~ 150 Hz
- Coldbox to mount setup in beam telescope

Signal cable connector

Beetle chip



Pt100 temp read-out

### ALiBaVa Setup in the Test Beam

### Difficult setup goals:

Reach -27 °C sensor temperature for highest irradiations

Limit access to beam area: use an Arduino for hardware resets

- Minimize scattering material in beam
- Movable setup for different beam incidence angles



Trigger scintillators Coolant pipes Beam N<sub>2</sub> supply Telescope HV, signal and plane temp cables 3D rotation stage

## **Trigger Synchronisation**



DESY

#### Integrating an analog strip sensor into a digital pixel telescope



#### **Reconstruction and Analysis Workflow**



### **Data Conversion and Pedestals**

Convert ALiBaVa data to LCIO

#### For off-beam runs: calculate pedestals of each channel

- Fill signals into a histogram and fit with a gaussian
- Mean → pedestal value
- Width → preliminary noise



Pedestals Channel 40



### **Channel Masking**

Not all channels are connected to the sensor

- The ALiBaVa has 2 \* 128 channels our sensor only has 64
- Some wire bonds might be faulty
- > Bad channels have to be masked  $\rightarrow$  look at preliminary noise
  - Flag outliers from here on out





### **Common Mode Subtraction**

Subtract pedestals, mask bad channels

#### > Calculate common mode

- Synchronous movement of all (good) channels in an event
- For each event: exclude outliers and fit a slope  $\rightarrow$  common mode



Common Mode Calculation



### **Data Correction and Noise**

- Correction values (pedestal + common mode) for each channel and event
  - Correct the 'real' on-beam data
- Get the noise values for each channel
  - Will be used later on for clustering, signal-to-noise, etc.
  - Gaussian fit on correction values
  - For unirradiated sensors or  $\Phi$  < 1.5e16 n<sub>eq</sub>/cm<sup>2</sup>  $\sim 4.5$  ADCs
  - After  $\Phi$  = 1.3e16 n<sub>eq</sub>/cm<sup>2</sup>:  $\sim$  7.1 ADCs for d = 100  $\mu$ m  $\sim 5.5$  ADCs for d = 200  $\mu$ m



Bias Voltage in |V

**Correction Values Channel 40** 

### Clustering

> On corrected data: search for channels with signal > 5 \* noise → 'seed'

- If neighbours have a signal > 2.5 \* noise  $\rightarrow$  add to cluster
- Cluster size: for 0° beam incidence: ~1.1, for 51° incidence ~ 1.8
- Write same format as telescope clusters → need an X-coordinate!
  - Look for cluster on preceding telescope plane nearest the relative Y-position
     → copy X CoG from here
- Control plots: η-distribution
  - =  $\eta = Q_{l} / (Q_{l} + Q_{r})$
  - Left / right is assigned to the seed and it's neighbour with the highest signal
    - $\rightarrow$  asymmetric!



seed cut

### Filtering

#### > Asymmetric cross-talk:

- Charge is spread to the next channel in the serial read out
- Compare charge distributions of seed neighbours:
- Channels on the right of a seed have ~6.4%, on the left only ~ 2.6% of the seed's charge
- Also seen in the ALiBaVa chip header and with ATLAS sensors  $\rightarrow$  no sensor effect

#### Finite Impulse Response filter to correct this

- Run through all events: Q<sub>new</sub>[i] = Q<sub>old</sub>[i] - a\*Q[i-1] – b\*Q[i-2]
- Calculate a,b from seed neighbour charge ratios
- Conserve charge: add the cross-talking charge back to where it originated
- After 2 filter iterations: recluster data
  - $\rightarrow$  symmetric  $\eta!$  :-)







### **More Filtering**

#### Feature of irradiated HPK n-bulk sensors: non-gaussian noise

- Show-stopper for a tracking detector
- Try to remove this by another filter
- Mask all signals > 150 ADCs
- p-bulk data has ~ 0.2 clusters/event → mask events with > 1 cluster
- If a seed candidate has a neighbour with a large negative signal → probably a discharge → mask event
- Finally recluster data
- Try to salvage some n-bulk runs



DUT Off-Beam Noise, Epi100N, F=1.0e+15, 400V, run 434

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### **Merging Data Streams**

- Both ALiBaVa and Telescope data streams are in sync on an event basis and have the same number of events
  - Create new TrackerPulse and TrackerData collections with clusters from both systems
- Check synchronisation over time / events:
  - In some runs severals telescope runs for one ALiBaVa run → does concatenation work?
  - Beam rate not always steady, sometimes hour-long gaps (e.g. user access)
  - Example: loss of sync after ~430k events





### Hitmaking

#### > First 'major' change:

- EUTelHitmaker.cc expects ZS binary data and disregards actual analog CoG position
- Fixed by an optional processor parameter to identify analog planes by their sensorid in the cluster
- Has this been fixed in newer versions?
- Setup rotations are applied:
  - Turns in -α direction
- Rough prealignment is performed







### **Alignment Strategy**

#### 5 iterations with DAF fitter:

- Narrow down residual cuts and resolutions
- Fix front and end plane, all Z coordinates

#### > Use DUT Residual in Y vs. track X for additional γ rotation

- Since no real DUT X information this can be a critical rotation
- Inclination gives γ
- Rotation not necessarily around plane center
- additional X/Y shift afterwards
- 2 iterations
- With final Z-alignment: 10 steps



DUT Residual in Y vs. X - Profile



### **Track Search**

#### > Rewrite of EUTelFitTuple.cc:

- Search tracks on upstream planes only, extrapolate to DUT
- Translate track impact point back into DUT channels → undo alignment & hitmaker steps
- Write tracks and reconstructed DUT sensor data to a ROOT Ntuple

#### Track resolutions:

- Assuming  $\sigma_{tele} = 5 \ \mu m$ , a DUT resolution of  $\sigma_{DUT} = 21.62 \ \mu m$  can be achieved ( $\sigma_{binary} \sim 22.55 \ \mu m$ )
- Improves for rotated data
- But: not for all runs :-(







### **Track Display**

#### CED to visualize tracks

Rebuild your EUTelescope installation with (more details in backup):

```
ilcsoft.install(CED(CED_version))
ilcsoft.module("CED").envcmake['CED_SERVER']='ON'
ilcsoft.install(CEDViewer(CEDViewer_version))
```





#### Sensor analysis with tracks from EUTelescope



### **Track Selection**

Many analysis steps done already by EUTelescope processors

- Run on track NTuple
- Telescope is much slower in read out (115µs) than ALiBaVa (25ns)
  - Many tracks in a (telescope) event  $\rightarrow$  much less in the ALiBaVa
    - $\rightarrow$  select the track that the ALiBaVa has most probably measured





### **Results: Charge Collection**

### Signal definition:

- Consider signal in 5 channels around extrapolated track impact
- Sum channel ADCs, fit Landau-Gaussian

### Charge collection

- Increase with beam incidence angles and bias voltage
- For Φ < 3e15 n<sub>eq</sub>/cm<sup>2</sup>:
   ~85% of unirr. signal collected
- For Φ = 1.3e16 n<sub>eq</sub>/cm<sup>2</sup>:
   ~70% of unirr. signal collected



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Signal Distribution, Epi100P, F=1.3e+16, -800V, 0rot, run 225

### **Results: Charge Sharing**

#### > Charge sharing:

- Integrate η-distributions
- Consider entries > 0.2 and < 0.8 as 'shared charge'
- Sharing increases with rotation and fluence, declines at higher voltages
- Alternate definition:

count channels next to track impact over a threshold  $\rightarrow$  similar results, but introduces cuts







Epi100Y, F=1.5e+15, 25rot

### Summary

- > 3 runs in TB21, 2 parasitic runs in TB22 resulted in ~300h of data taking
  - Setup installed, connected and running within minutes → real plug and play :-)
- > Successful integration of an analog strip sensor a digital pixel telescope
- With custom EUTelescope processors many analysis steps can be performed during reconstruction
- > Results show epitaxial silicon is a promising material for future detectors



### Backup



### **Phase-II Upgrade Motivation**

### In 2023: upgrade of the LHC to a <u>High Luminosity-LHC</u>

- Increase the luminosity by a factor of 5
- Even harsher radiation environment



ATLAS  $t\bar{t}$  event simulation:  $\langle \mu \rangle = 40$ 



```
ATLAS tī event simulation: <µ> = 140
```

#### LHC experiment trackers need to be upgraded to maintain excellent current performance

- Improve granularity to keep low detector occupancy
- CMS: tracker to contribute to the Level-1 trigger  $\rightarrow$  new module concept
- Develop radiation-hard sensors → new sensor technologies



### **Expected Fluences**

Is epitaxially grown silicon a radiation-hard sensor material?

- Could it be used for a HL-LHC pixel sensor?
- > Current CMS tracker:
  - Design fluence: 4e14 n<sub>eq</sub>/cm<sup>2</sup> @ r = 20 cm

#### HL-LHC tracker:

- Expected fluence 2e15 n<sub>eq</sub>/cm<sup>2</sup> @ r = 20cm
- Fluences > 1e16 n<sub>eq</sub>/cm<sup>2</sup> for pixel sensors

#### > Desired signals:

- 6000 e<sup>-</sup> for strip regions
- 3000 e<sup>-</sup> could be sufficient for pixel





### **Sensor List**

List of successfully measured sensors

- Epi = epitaxial
- FTH = float-zone
- MCZ = magnetic czochralski

Fluence	Epi 100P	Epi 100Y	Epi 100N	Epi 70N	FTH 200Y	MCZ 200P	MCZ 200Y	MCZ 200N
0	$\checkmark$	$\checkmark$		$\checkmark$				
1e15 Los Alamos	$\checkmark$	$\checkmark$						
1.5e15 CERN	$\checkmark$	$\checkmark$						
3e15 CERN	$\checkmark$		$\checkmark$					
1.3e16 CERN		$\checkmark$	$\checkmark$		$\checkmark$		$\checkmark$	



### **DESY-II Testbeam Setup**

#### > Test beam from DESY-II synchrotron

- Carbon fibre generates a bremsstrahlung beam
- Metal plate converts photons to e<sup>+</sup>/e<sup>-</sup> pairs
- Dipole magnet spreads out beam
- End collimator cuts out final beam

#### > DESY beam telescopes

- Series of parallel Si sensor planes inserted into the particle test beam
- Reconstruct particle tracks through planes wrt. time and position
- Use gained information to evaluate device under test (DUT) performance
- Use telescope to investigate epitaxial sensors



### **Event Viewer**

How to install and use event-viewer:

1) Go to your ilcinstall directory. In examples/eutelescope, edit release-standalone.cfg

Uncomment or make sure these lines exist:

```
ilcsoft.install( CED( CED_version ))
ilcsoft.module("CED").envcmake['CED_SERVER']='ON'
ilcsoft.install( CEDViewer( CEDViewer_version ))
```

2) Rebuild ilcsoft: Make sure \$ILCSOFT is known, e.g. do "export ILCSOFT=/path/to/where/you/are", in the ilcinstall directory do: ./ilcinstall -i examples/eutelescope/release-standalone.cfg

3) Wait...

4) Source the environment: in ilcsoft/VERSION/Eutelescope/VERSION: source build\_env.sh

5) Start the display: glced &

6) In a new terminal, repeat 4) and run the eutelescope event-viewer: jobsub -c config.cfg -csv runlist.csv event-viewer RUNNUMBER

7) Enter moves to the next event, q quits.

Pro tip for the display window: zoom in and turn the background to black.

8) The event-viewer needs hits and tracks. The collections are specified in the steering file or in the config.cfg. Make sure these exist.

