# ATLAS Silicon strip upgrade: Results with the SLAC testbeam and AIDA telescope





#### John Keller (DESY) on behalf of many people



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### The team



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# ATLAS tracker upgrade

- Tracking detector forms innermost part of ATLAS.
- During "LS3" (2023-24), current ID will be replaced with allsilicon Integrated Tracker (ITk), to prepare for High-Lumi LHC.
- ITk consists of silicon pixels and strips surrounding the interaction point.





### This testbeam

- Had beam at SLAC End Station A from 2<sup>nd</sup> 16<sup>th</sup> July 2014.
- ~9 GeV, 5 Hz, O(100) electrons/bunch.
- Two measurements performed at once:
  - ABC 130: Measure the gain of new ATLAS binary chip.
  - Fast Cluster Finder: Demonstrate new readout technique for use in self-seeded trigger.
- Both being tested with beam for the first time.



### The setup



#### Integration with telescope





Telescope track x position [mm]

Both DUT's are in-sync with telescope. 

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# ABC 130

- New ATLAS Binary Chip ABC 130 intended to be used as default.
- First prototypes show lower gain than expected.
- However these measurements depend on internal calibration capacitor: want to use test beam as external calibration.



## Gain Measurement

- Given the known beam energy and sensor thickness, the most probable charge deposit due to ionization may be calculated: 3.6 fC.
- Cannot calculate voltage gained due to only binary readout: instead do a threshold scan counting events at each level.
- Differential of threshold scan should correspond to Landau curve of charge deposition.

The most probable ionization energy			
Thickness [um] E [MeV]		Theory[KeV]	Exp.[KeV]
1.1			
148	0.976	38.4	40.3
148	199	40.4	41.4
148	9 999	40.6	39.6
148	9 999	40.6	39.5
148	50 000	40.6	39.7
290	199	82.2	80.9
1 007	199	306	323

Moller et al. (1982)



## Gain results



- Preliminary result indicates maximum of threshold curve at around 314 mV.
- This would correspond to a gain of 88 mV/fC, in good agreement with the specification of 90 mV/fC.
- However there are many subtleties to the analysis: charge-sharing effects, DAC-mV conversion, etc.
- We are awaiting complementary measurements with a source before reaching a firm conclusion.

#### Fast Cluster Finder

## **Fast Cluster Finder**



- FCF readout finds coincidence between hits on two layers of silicon strips.
- Can be used as input for self-seeded track trigger algorithm.

#### FCF doublet



 No B field in test beam: instead slide sensors relative to one another to simulate effect of curved trajectory.

#### FCF correlations



Micrometer position [µm]

## FCF and tracks

- Integrating FCF with telescope is complicated by smaller telescope coverage, and slight mis-alignment with beam.
  - Partially overcome using 4-plane tracking.
- Applying isolation from multiple tracks suggests most of the "background" correlations are due to two nearby electrons.

#### filter 1: one track within 0.08 $\mu$ m. filter 2: also no other track within 0.3 $\mu$ m.



## Conclusions

- Two successful and exhilarating weeks of testbeam occurred at SLAC End Station A last July.
- External calibration of ABC 130 gain provided valuable input to the strip community, awaiting further corroboration from source tests.
- Fast Cluster Finder successfully demonstrated using data for the first time.
- AIDA telescope tracks correlated with DUT's and used to assist in the measurements.

#### Backup slides



### Track reconstruction details

- We mostly used the default settings in the datura\_noDET example of the EUTelescope package.
- DUT's not included in track reconstruction.
- Residual cut at alignment step set to 150.
- Use only plane 0 as fixed for ABC runs.
- Maximum chi2 of 1000.
- Maximum 300 clusters per plane for filter.
- For ABC runs, require only 5 of the 6 planes to have hits.
- For FCF runs, do tracking with either first or last 4 planes.

## Residuals

- A look at the fit residuals (i.e. track position hit position) from a typical run.
- The x axis is in mm.
- (Note, these are biased residuals).



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# Coverage issue

- The telescope was slightly rotated with respect to the incoming beam.
- This reduces the effective area of the sensors by about 5 mm (20 → 15)
- All of ABC chip is still inside coverage, but for FCF we miss two of the 3 big peaks.



reduced coverage

# Coverage issue

- Coverage in x direction is only an issue for the FCF; single-chip ABC is still totally covered.
- Mostly recovered by running track reconstruction twice, with either first or last 4 planes.
- Which set to use determined by address on FCF doublet.



Telescope track x position [mm]



- We can check to make sure the correlation is working as expected by plotting the track's distance from the expected line, as a function of time (or TLU ID number).
- This is an example of a run that is good the whole way through.

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- In some cases, the telescope gets "off" from the DUT, usually by one TLU ID.
- Simple correlation plot may not catch this.
- Usually can be recovered.

### Latency



When the TLU trigger arrives at HSIO, it's already a few BC cycles behind the actual event. A latency can be set to read out the right event.

Before data taking, a latency scan was done to find the right latency to use.