# Any Light Particle Search II

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# **Any Light Particle Search**

Any Light Particle Search experiment looks for Weakly Interacting Sub-eV Particles also called WISPs.

- Axion → Possible solution for the smallness of the CP violation in QCD
  - → Hints from astrophysics observations

$$\begin{array}{ccc} \gamma & \text{WISP} & \gamma \\ \hline & & & & \\ \hline \end{array} \\ \hline & & & \\ \hline \end{array} \end{array}$$

Light shining through the wall experiment: Photon-mixing + Additional light boson → Re-appearance of photons behind the barrier

1064 nm laser → 1.17 eV photons 1 photon every few hours

$$\mathcal{P}_{\gamma \to a} = \frac{\omega}{4k_a} \left( g_{a\gamma} BL \right)^2 |F|^2 = \mathcal{P}_{a \to \gamma} \qquad k_a^2 = \omega^2 - m_a^2$$

#### ALPS I



(Source: B. Döbrich)



Ehret et al., New ALPS results on hidden-sector lightweights , PLB (2010)

Laser	532 nm
Optic setup	10 m production cavity
Magnet	9m, 5T HERA dipole
Detector	CCD Camera (PIXIS)

#### From ALPS I to ALPS II



#### ALPS II

parameter	scaling	ALPS I	ALPS IIc	sens. gain
BL (total)	$g_{a\gamma} \propto (BL)^{-1}$	22 Tm	468 Tm	21
PC built up ( $P_{\text{laser,eff.}}$ )	$g_{ m a\gamma} \propto eta_{ m PC}^{-1/4}$	1 (kW)	150 (kW)	3.5
rel. photon flux $\dot{n}_{\text{prod}}$	$g_{a\gamma} \propto \dot{n}_{ m prod}^{-1/4}$	1 (532 nm)	2 (1064 nm)	1.2
RC built up $\beta_{\rm RC}$	$g_{a\gamma} \propto eta_{ m RC}^{-1/4}$	1	40,000	14
detector eff. DE	$g_{\mathrm{a}\gamma} \propto D E^{-1/4}$	0.9	0.75	0.96
detector noise DC	$g_{\mathrm{a}\gamma} \propto D C^{1/8}$	$1.8 \cdot 10^{-3}  s^{-1}$	$10^{-6}  \mathrm{s}^{-1}$	2.6
combined				3082

$$S(g_{a\mu}) \propto \frac{1}{BL} (\frac{DC}{T})^{1/8} \cdot \left(\frac{1}{\eta \dot{N}_{Pr} \beta_{PC} \beta_{RC}}\right)^{1/4}$$

#### **ALPS II Schedule**





#### ALPS IIa - setup



LASER	35 W, 1064 nm (Class 4 laser)	
LASERTYPE	Master Oscillator Power Amplifier (MOPA)	
РВРС	5000	
PB RC	40000	
VACUUM	10 <sup>-6</sup> mbar	

# Optics



# **Optical Layout for ALPS II**



Auxiliary green beam obtained via second harmonic generation (KTP crystals) from the IR production.  $\rightarrow$ Test resonance of the RC.

QPD7

(ovlp.

DC)

## Auto alignment



The PC and the RC needs to be in the same modal phase in order to reach the highest PB (aiming for 95% mode overlap).

M2 and M3  $\rightarrow$  Fixed position M1 and M4  $\rightarrow$  Piezo mount for compensation of ambient vibrations

**Beam alignment towards cavities** → Control beforehand the cavity

### **PB for production cavity**

Power modulation transfer function from upstream to downstream of the production cavity



The PB of the production cavity is not sufficient (< 5000).

Possible sources: • Coating of the mirrors • Cleanliness of the mirrors • Alignment

Clipping in the vacuum system

# **Mirrors quality testing**



• Optical loss of the high-reflectivity mirrors is a critical parameter.

• Use of a cavity ringdown technique to characterize them.

• Finesse values for each cavities are currently measured using this technique.

 $PB = \frac{F}{\pi}$ 

Reachable Power Build Up (PB) depends on the cavity finesse (F)

#### **Results and outlook**

Mirror 1	Mirror 2	Losses	Finesse
750 ppm	68 ppm	800 ± 30ppm	8600 ± 400
2 ppm	68 ppm	192 ± 30ppm	34000 ± 5000
68 ppm	68 ppm	213 ± 35ppm	30000 ± 5000

# ⇒ The mirrors are corresponding to the specs for small cavities and small laser beam radius.

All the mirrors are currently tested with a beam with a larger radius in order to enlarge the tested region on the surface of the mirrors.

# Coupling of the beam



### Setup



#### Latest results



The highest value is currently obtained for a lens with f= 35 mm (matching expectations).

The efficiency reached is lower than what was expected. The beam quality of the test laser is currently tested using a knife edge unit.

# Detector



#### **Technical challenges for the detector**

 $\Rightarrow$  Low energy (1.17 eV) and low rate (1 photon every few hours).

1) High efficiency
 2) Low dark count rate
 3) Long-term stability
 4) Good energy resolution
 5) Good time resolution

# **CCD in ALPS II**



**Von Seggern JE**, *Constraining Weakly Interacting Slim Particles with a Massive Star and in the Laboratory*, Dissertation, Univ. Hamburg, 2014

### **Transition Edge Sensor**

TES



Two channels module (3 cm \* 3 cm)



**Tungsten chip** (25 x 25 μm , 20 nm)

Tc ≈ 140 mK

A.E. Lita, A.J. Miller, S.W. Nam, *Counting nearinfrared single photons with* 95 % *efficiency*, Opt Expres. 2008

TES: SQUID Microcalorimeter measuring the temperature difference ΔT of the absorber material.



#### **NISTW-TES**

Efficiency (1064 nm)	95 %
Dark current	$10^{-4}  \mathrm{sec}^{-1}$
Long term stability	$\checkmark$
Good energy resolution	< 8%
Good time resolution	$\checkmark$

#### Photon absorption to signal output



#### **TES environment in ALPS II**



22

### Single photon events



### **Detection efficiency**

#### **Adiabatic Demagnetization Refrigerator**



# Magnets for ALPS II c



# Magnet



The power build up of the cavities depends highly on the aperture.

Hera magnets are bent → Small aperture (35 mm)

This aperture would allow us to have only 4+4 dipoles (10 + 10 forseen).

⇒ Straighten magnets

# Magnet



#### Unbending of the magnets

Straightening force applied by pressure screw in the middle of the magnet (cold mass).

Deformation successful, yielding 90% of maximum aperture. Magnet still working perfectly, quench current even slightly increased.

# Summary

- ALPS II experiment (DESY, Hamburg) follows the light-shining through the wall concept.
- ALPS II aims at an increase in sensitivity of 3000 compared to ALPS I, mainly by a regeneration cavity and more magnets.
- Basics of the optics setup have been demonstrated but the specifications are still to be reached.
- A tungsten Transition Edge Sensor operated below 100 mK has been successfully used to detect single-photons in the near-infrared.

#### ALPS II TDR: arXiv:1302.5647

Characterization, 1064 nm photon signals and background events of a tungsten TES detector for the ALPS experiment: arXiv:1502.07878v1



#### ⇒ First ALPS II data taking in 2016.



## Thank you for your attention !

