



**Universidad**  
Zaragoza



**LSC**

*Laboratorio Subterráneo de Canfranc*

# Status of the ANAIS dark matter project at the Canfranc underground laboratory

Miguel Ángel Oliván  
on behalf of the ANAIS team

Universidad de Zaragoza  
Laboratorio Subterráneo de Canfranc

# Outline

- ANAIS Experiment
- ANAIS-25
- ANAIS-37
- ANAIS status and prospects

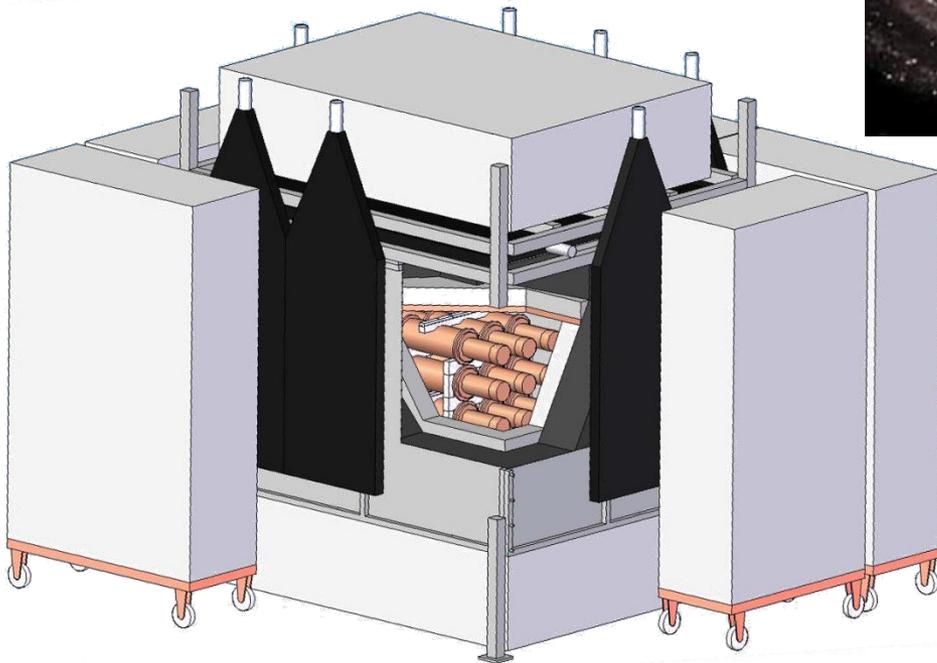
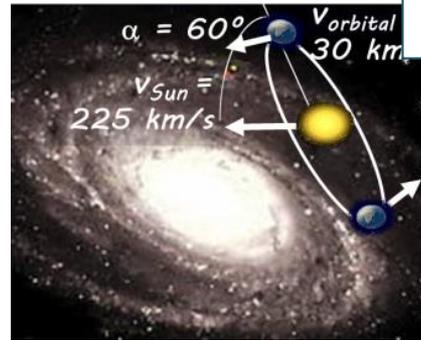
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# ANAIS Experiment

Project for 3×3 matrix of NaI(Tl) scintillators 12.5 kg each to look for Dark Matter (DM) annual modulation at Canfranc Underground Laboratory (LSC)

LSC, Spain  
2450mwe



Goal:  
Confirmation of DAMA/LIBRA  
positive signal with same target and  
technique

Experimental requirements:

- Energy threshold  $< 2 \text{ keVee}$
- Background few counts/(keV kg day)

# ANAIS Experiment

Long effort by the University of Zaragoza through the operation of different NaI detectors in Canfranc

- Using hexagonal, 10.7 kg NaI crystals from BICRON:

- Pioneer modulation analysis at **NaI32 experiment**

M. L. Sarsa *et al.*, *Phys. Lett. B* **386**, 458 (1996)

M. L. Sarsa *et al.*, *Phys. Rev. D* **56**, 185 (1997)

- First ANAIS prototypes

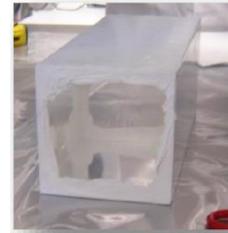
- Using a 9.6 kg NaI crystal from St. Gobain:  
**ANAIS-0 prototype**

- Bulk NaI(Tl) scintillation low energy events selection with the ANAIS-0 module. C. Cuesta *et al.*, *EPJ C* **74** (2014) 3150.

- Analysis of the 40K contamination in NaI(Tl) crystals from different providers in the frame of the ANAIS project. C. Cuesta *et al.*, *Int. J. Mod. Phys. A*. **29** (2014) 1443010.

- Slow scintillation time constants in NaI(Tl) for different interacting particles. C. Cuesta *et al.*, *Opt. Mat.* **36** (2013) 316.

- Background model for a NaI(Tl) detector devoted to dark matter searches. S. Cebrián *et al.*, *Astrop. Phys.* (2012) 60.



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*A.* **29** (2014) 144

- Slow scintillation

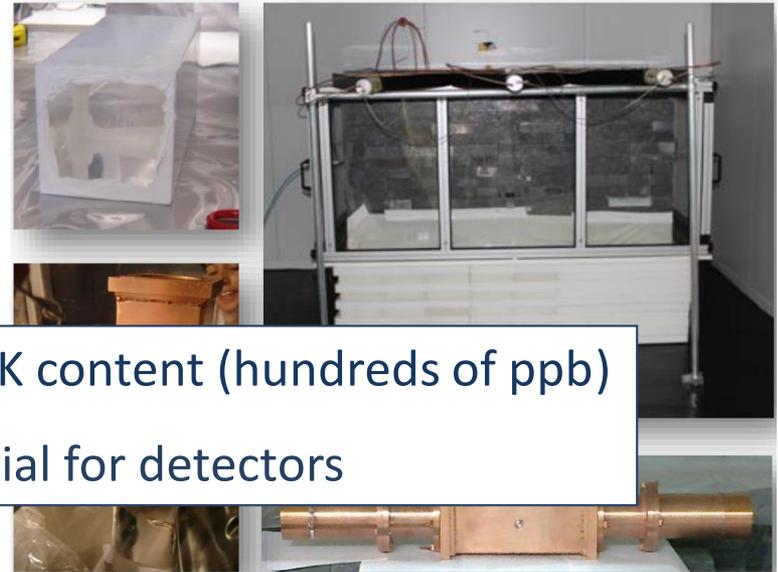
C. Cuesta *et al.*, *C*

- Background m

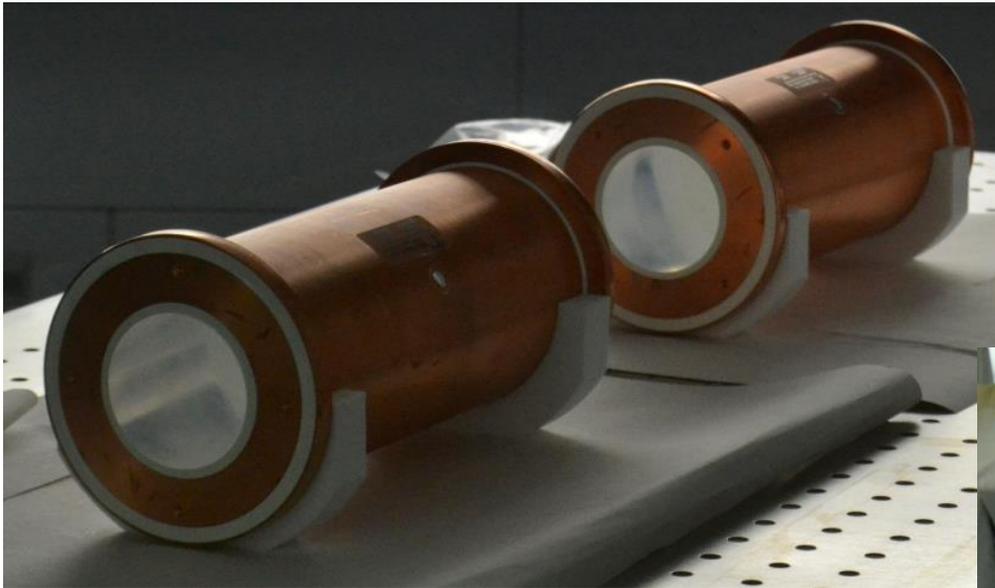
S. Cebrián *et al.*,

Disregarded due to an unacceptable K content (hundreds of ppb)

Search for cleaner material for detectors

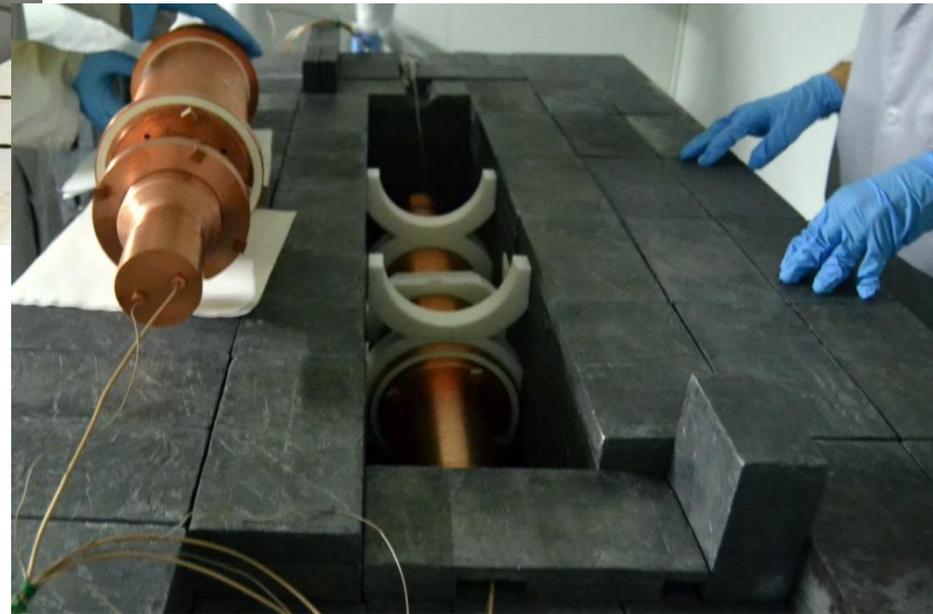


# ANAIS-25



## Two Alpha Spectra modules:

- Clean starting NaI powder (< 90 ppb K)
- 12.5 kg, cylindrical
- Mylar window for low energy calibration
- PMT Ham (R12669SEL2 & R11065SEL)



## Goals:

- Measure internal contamination  $^{40}\text{K}$  and  $^{238}\text{U}$  and  $^{232}\text{Th}$  chains
- Determine light collection, fine tuning of DAQ, filtering and analysis protocols, general background assessment

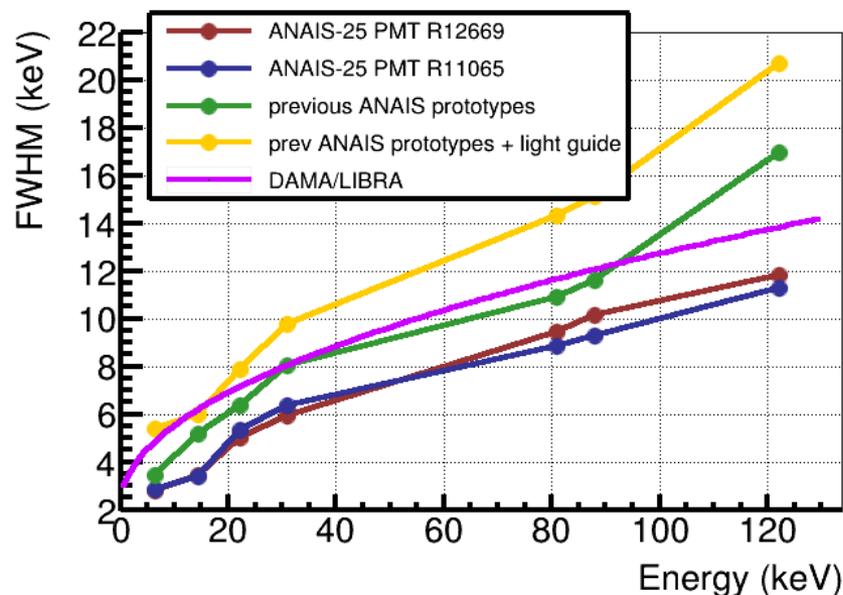
Taking data at LSC: Dec 2012 - Mar 2015

# ANAIS-25 – Light collection

Detector	PMT model	Phe <sup>-</sup> /keV
D0	Ham R12669SEL2	<b>16.13 ± 0.66</b>
D1	Ham R12669SEL2	<b>15.19 ± 0.09</b>

Excellent light collection (~15 phe<sup>-</sup>/keV), better with high quantum efficiency PMTs (R12669SEL2)

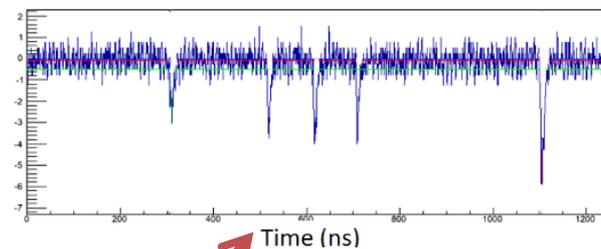
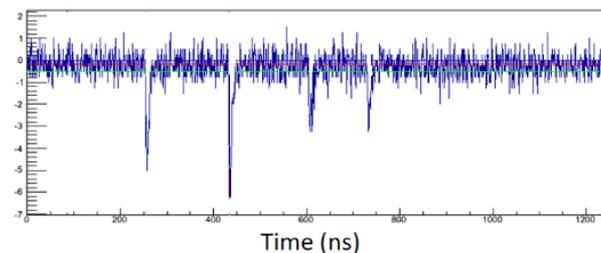
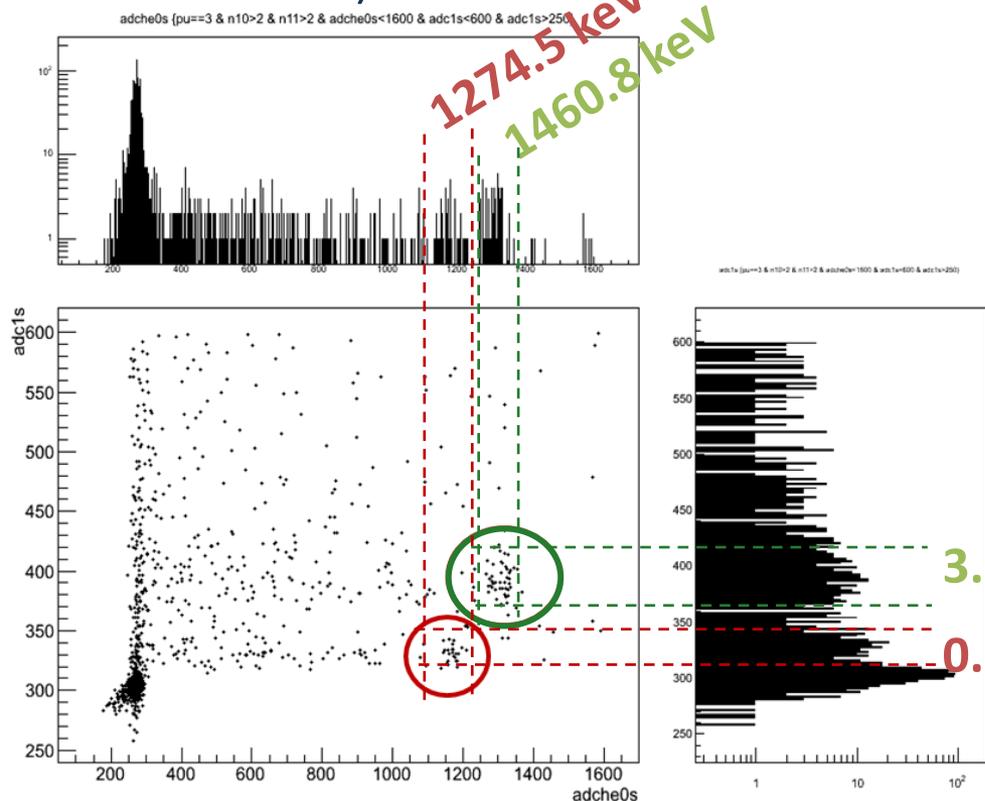
(See “Light collection in the prototypes of the ANAIS mark matter project” poster for details)



Resolution improvement in all low energy calibration lines

# ANAIS-25 – Threshold

Coincident events → two energy coincident lines ( $^{40}\text{K}$  and  $^{22}\text{Na}$  internal contamination)



1 keVee energy threshold seems to be achievable. Trigger efficiency at this energy is under investigation

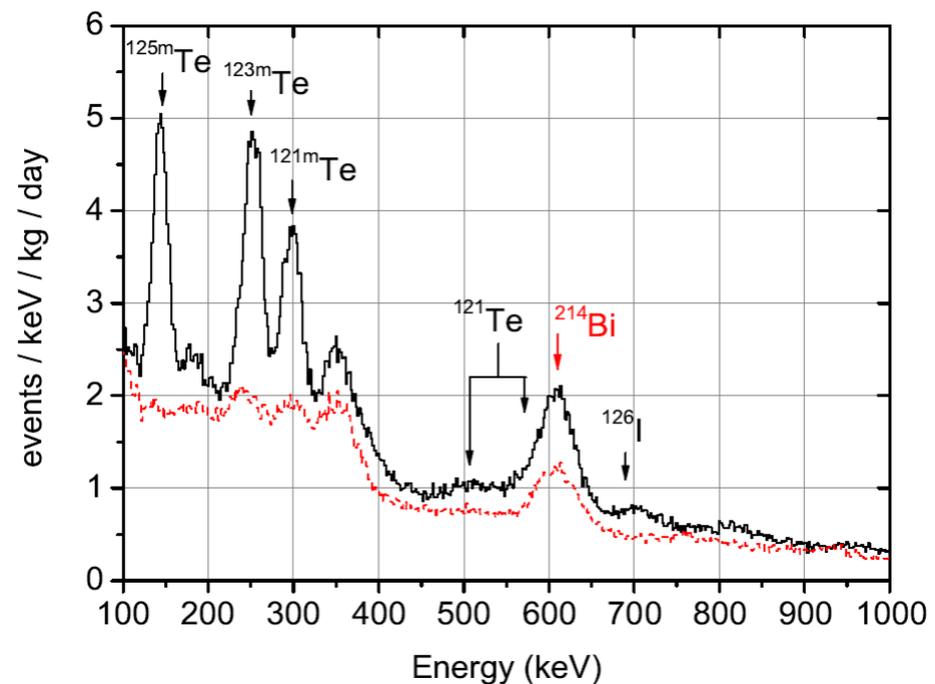
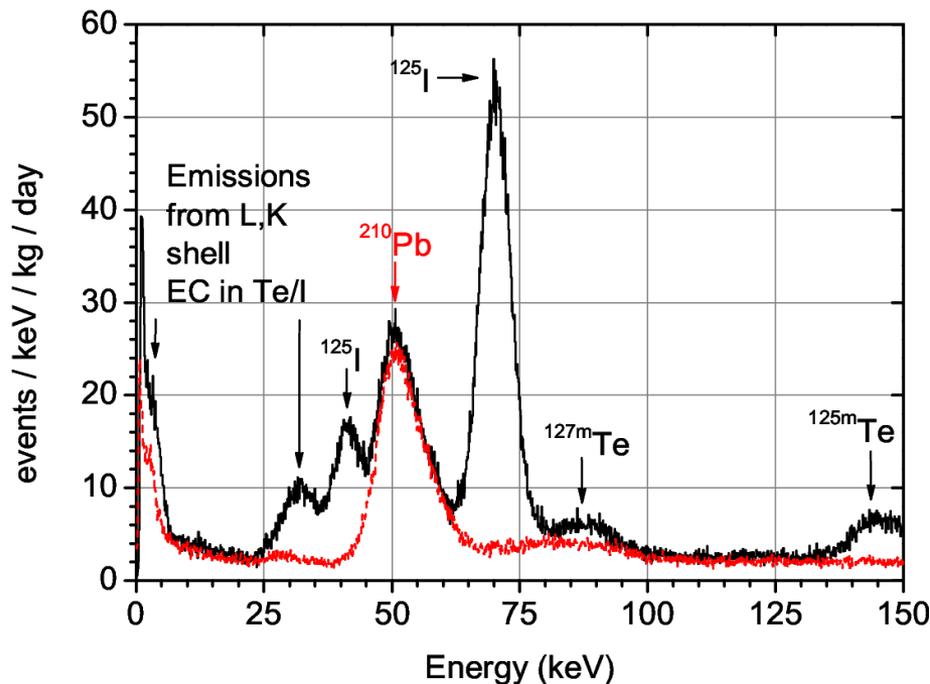
# ANAIS-25 – Background

Contribution from cosmogenic isotopes induced in NaI crystals:

Determined for the first time the cosmogenic production of some isotopes in NaI (I, Te, Na)

J. Amaré et al, JCAP 02 (2015) 046

- First month of data-taking
- 15 months later

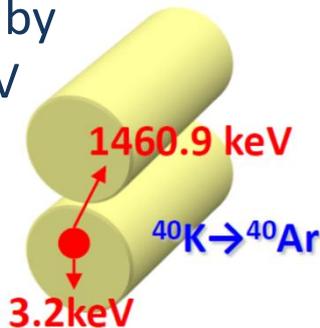


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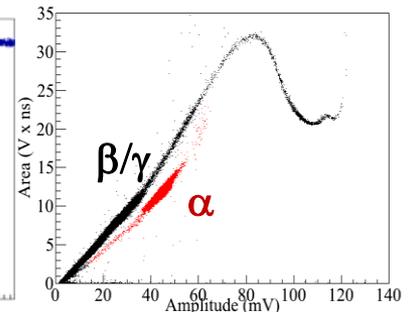
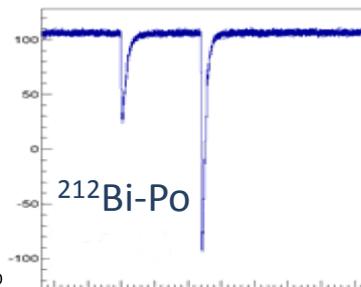
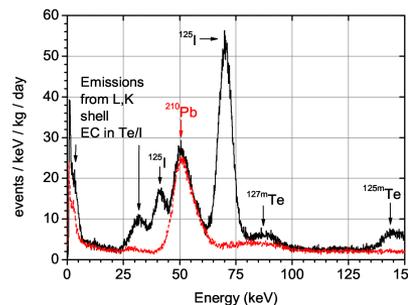
Bulk Contamination:

$^{40}\text{K}$	$^{238}\text{U}$	$^{210}\text{Pb}$	$^{232}\text{Th}$
1.25 mBq/kg (41 ppb K)	10 $\mu\text{Bq/kg}$	<b>3.15 mBq/kg</b>	3 $\mu\text{Bq/kg}$

Determined by  
1461/3.2 keV  
coincidence  
analysis



Determined by alpha rate and Bi/Po sequences.  
Verified by simulations (see later).



# ANAIS-25 – Background

Bulk Contamination:

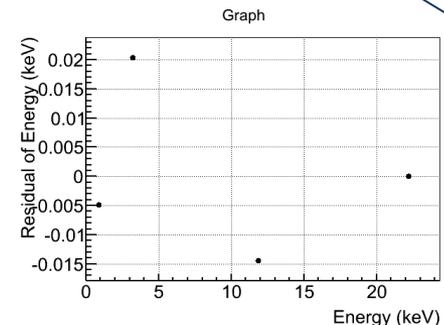
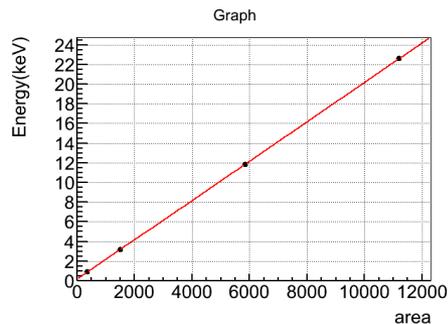
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- Radiopurity goals are fulfilled for  $^{40}\text{K}$  and  $^{232}\text{Th}$  and  $^{238}\text{U}$  chains, but a  $^{210}\text{Pb}$  contamination out-of-equilibrium is present in ANAIS-25 crystals.
- Origin of the  $^{210}\text{Pb}$  contamination identified (crystal growing) and being solved by Alpha Spectra.
- **New material prepared at Alpha Spectra using improved protocols:  
new detector under test → ANAIS-37**

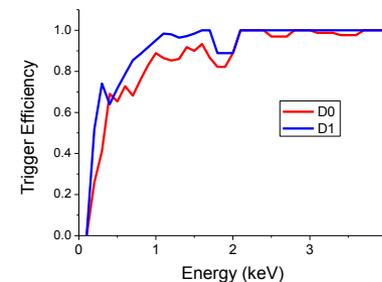
# ANAIS-25 – Data analysis

PRELIMINARY

Calibration with  $^{109}\text{Cd}$  and coincident lines ( $^{40}\text{K}$  and  $^{22}\text{Na}$ ): accurate calibration at 1keVee

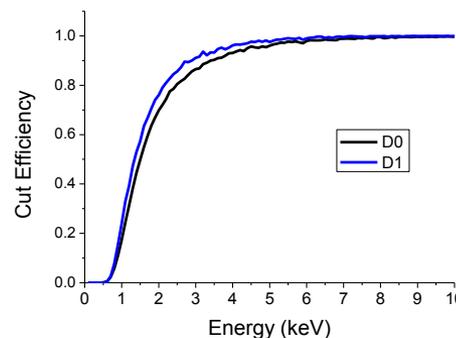


At 1 keV the trigger efficiency is larger than 80% (estimated with coincident events,  $^{22}\text{Na}$  and  $^{40}\text{K}$ )



Cuts for removing PMT events:

- 1) Cut in the number of peaks in the pulse ( $n > 2$  in each PMT)
- 2) Cut in temporal parameters of the pulse
- 3) Cut in asymmetry in the light sharing

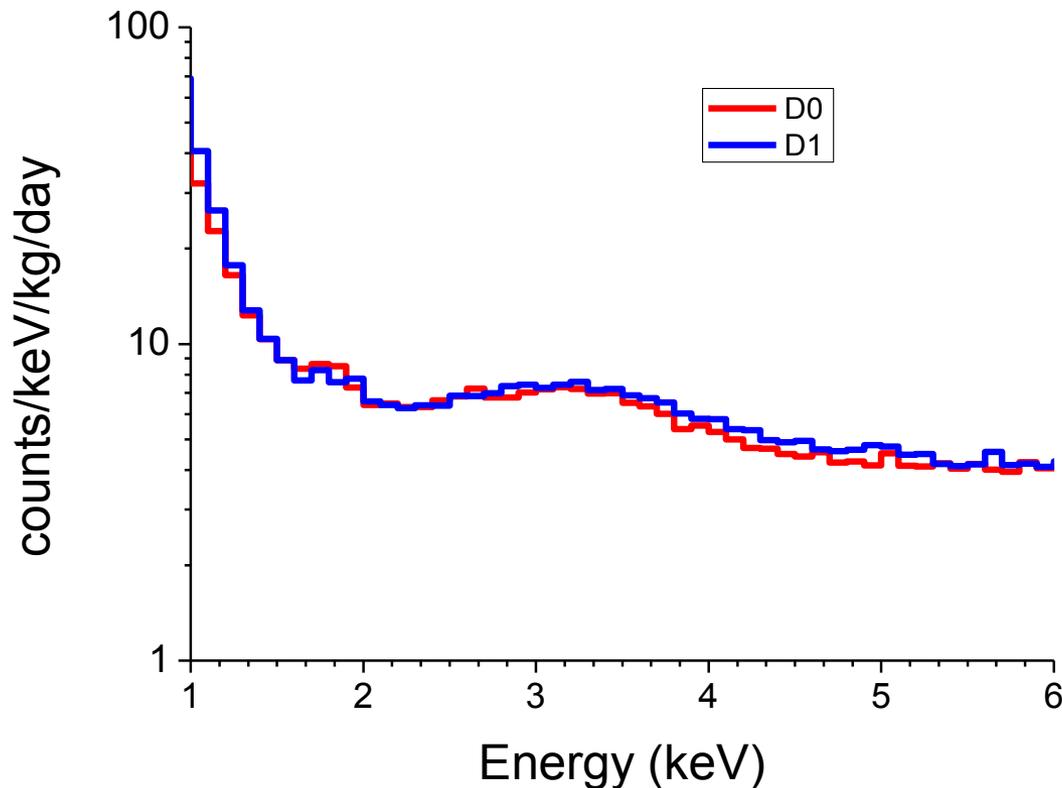


Cut efficiency at 1 keVee 20%  
Very fast increasing 1-2 keVee  
(calculated with  $^{109}\text{Cd}$  events)

# ANAIS-25 – Data analysis

PRELIMINARY

Background after applying the cuts, efficiency corrected:

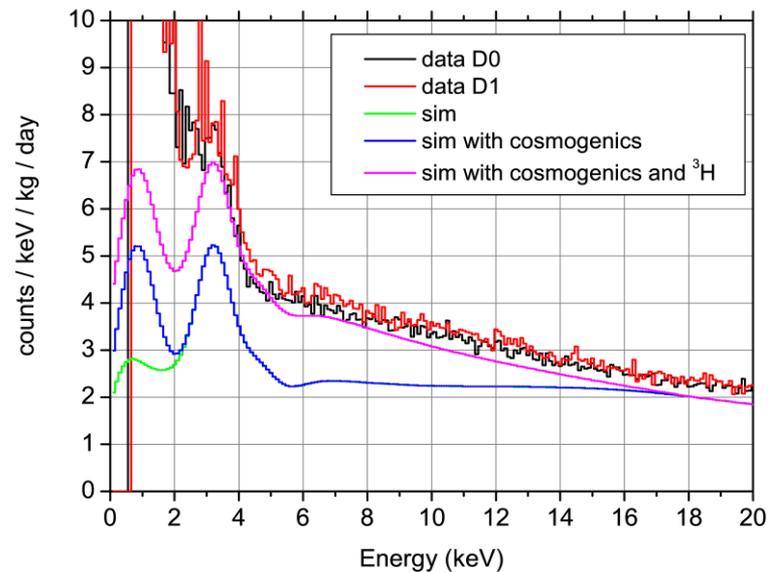


- Working on more aggressive cuts to reduce background below 2 keVee
- Working on a *blank module* (without crystal) to study the PMT coincident events

# ANAIS-25 – Background model

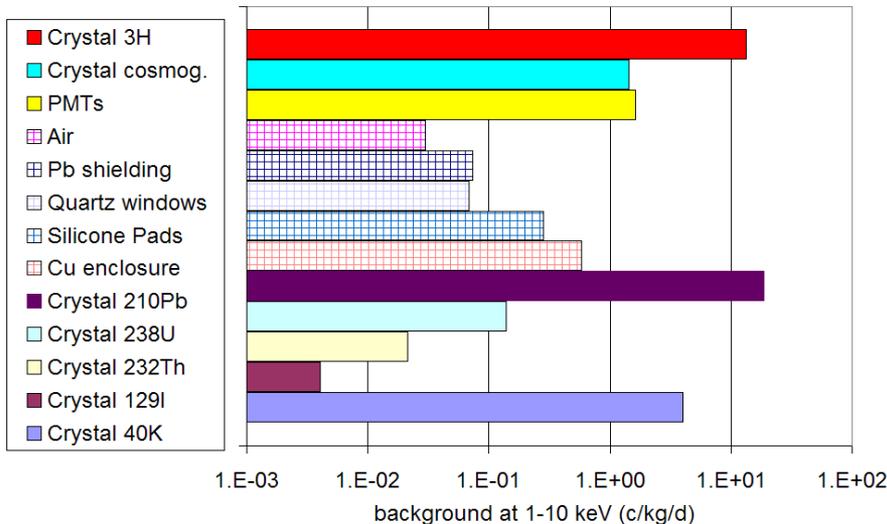
PRELIMINARY

- Measured activity of external components (PMTs, copper, ...)
- Measured activity in NaI crystals ( $^{40}\text{K}$ ,  $^{238}\text{U}$ ,  $^{232}\text{Th}$ ,  $^{210}\text{Pb}$ )
- Cosmogenics  
J. Amaré et al, JCAP 02 (2015) 046
- $^3\text{H}$  activity



Good agreement with filtered background, coincident and anticoincident spectra down to 3 keVee.

See “Background model for NaI(Tl) detectors for the ANAIS dark matter project” poster for details



# Outline

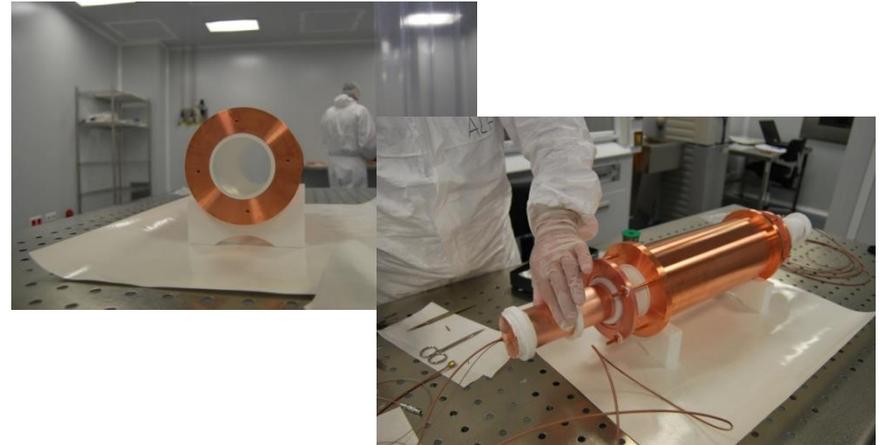
- ANAIS Experiment
- ANAIS-25
- **ANAIS-37**
- ANAIS status and prospects

# ANAIS-37

New module by Alpha Spectra (D2)

Goals, check:

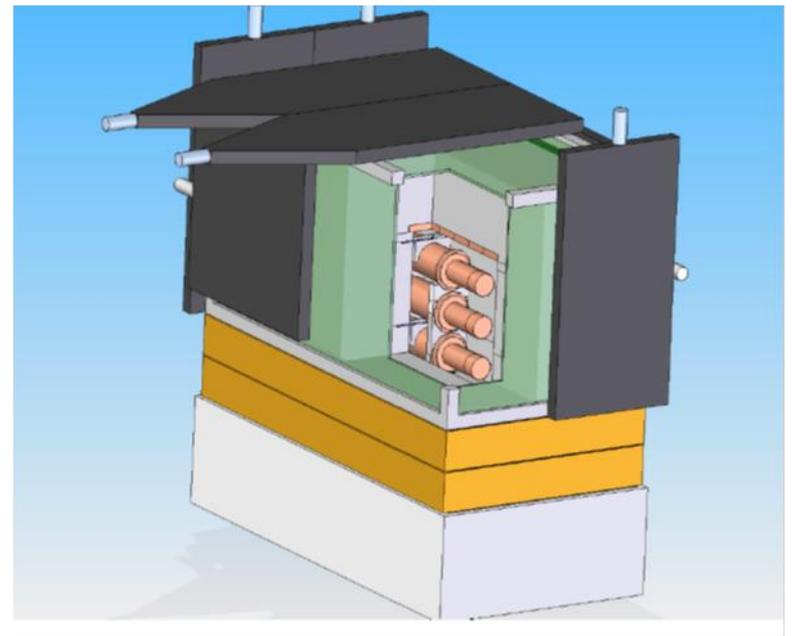
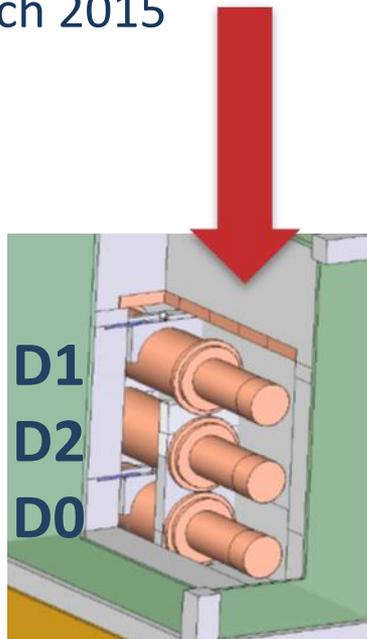
- $^{210}\text{Pb}$  reduction
- $^{40}\text{K}$  and  $^{238}\text{U}$  and  $^{232}\text{Th}$  chains
- Light collection



Taking data at LSC since March 2015



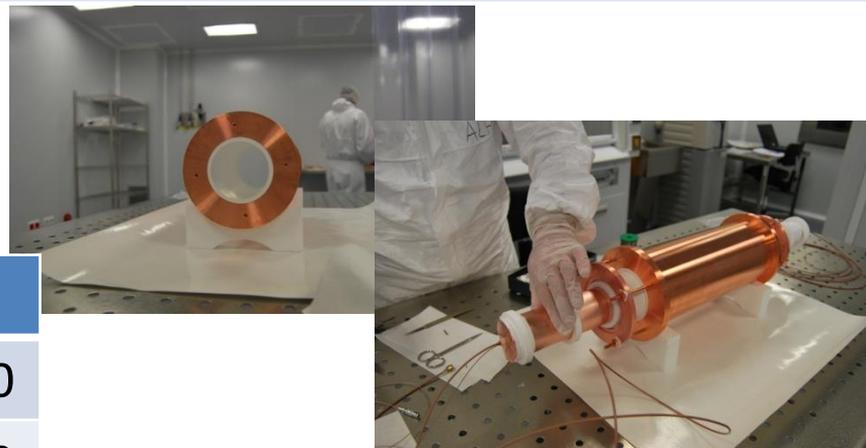
PATRAS Workshop 2015, Zaragoza



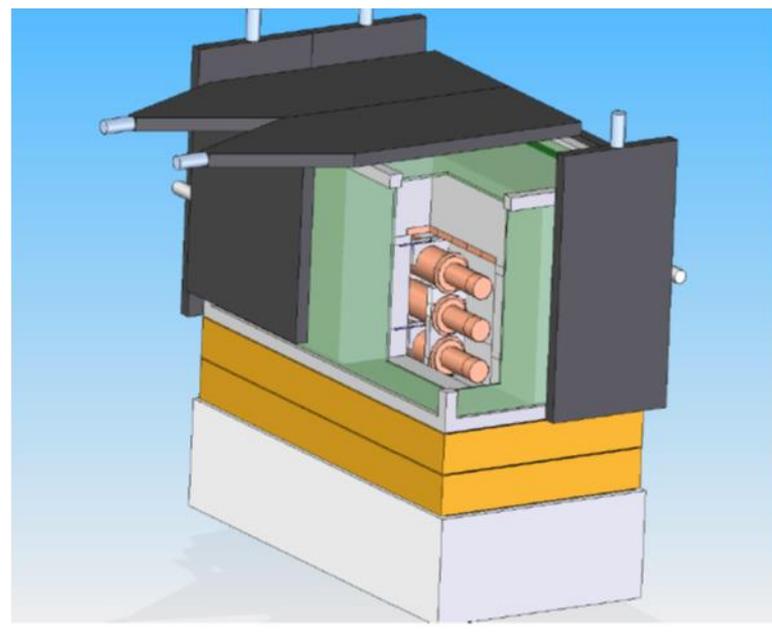
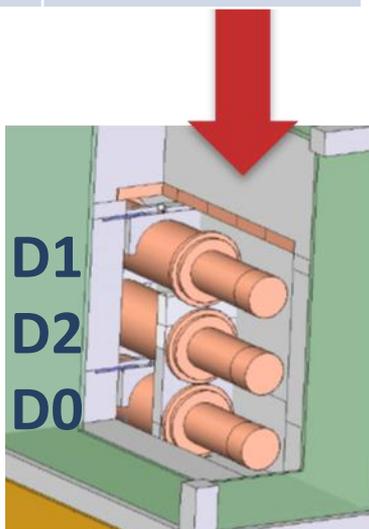
# ANAIS-37 – Light Collection

New module by Alpha Spectra (D2)  
~ 15 phe/keV

Detector	PMT model	Phe <sup>-</sup> /keV
D0	Ham R12669SEL2	15.26 ± 0.10
D1	Ham R12669SEL2	14.44 ± 0.09
D2	Ham R12669SEL2	<b>15.41 ± 0.05</b>



PATRAS Workshop 2015, Zaragoza

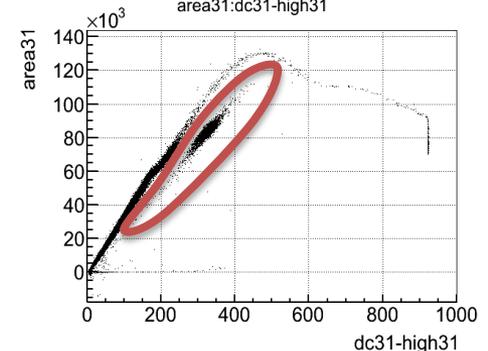
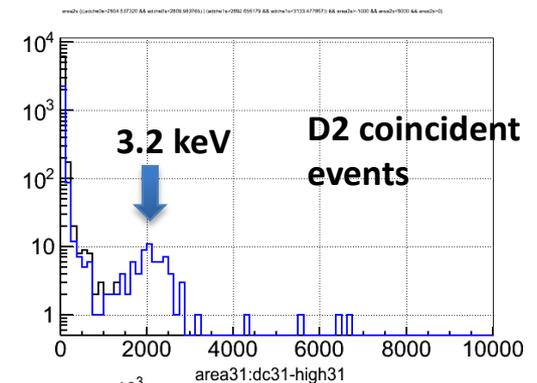
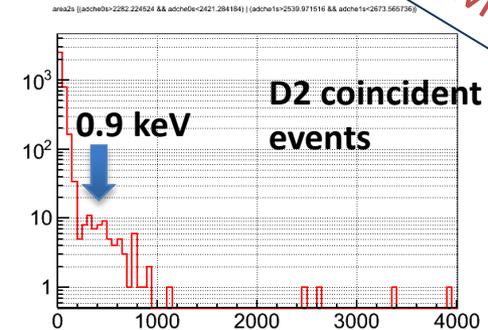


# ANAIS-37 – Background

PRELIMINARY

New detector (D2) background (Live Time: 23.7 days)

- Coincidence window at 1274 keV in D0 | D1  $^{22}\text{Na}$  coincident events as **cross-check of the threshold**
- Coincidence window at 1461 keV in D0 | D1  $^{40}\text{K}$  Potassium Content:  $44 \pm 4$  ppb
- Total alpha rate:  $0.58 \pm 0.01$  mBq/kg  
ANAIS-25 alpha rate: 3.15 mBq/kg (factor 0.18)

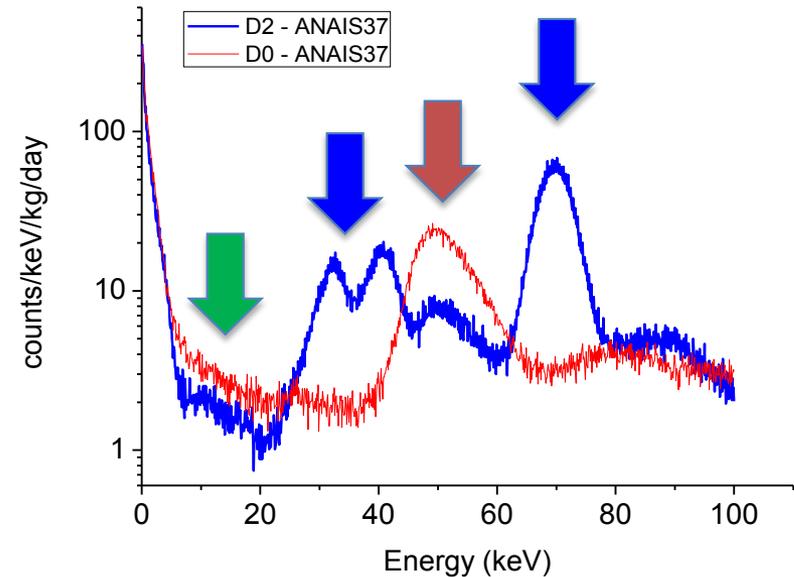


# ANAIS-37 – Background

Raw D2 background without filtering:

Very promising reduction under 20 keV  
and cosmogenics still decaying

$^{210}\text{Pb}$  clear reduction with respect to  
ANAIS-25 modules



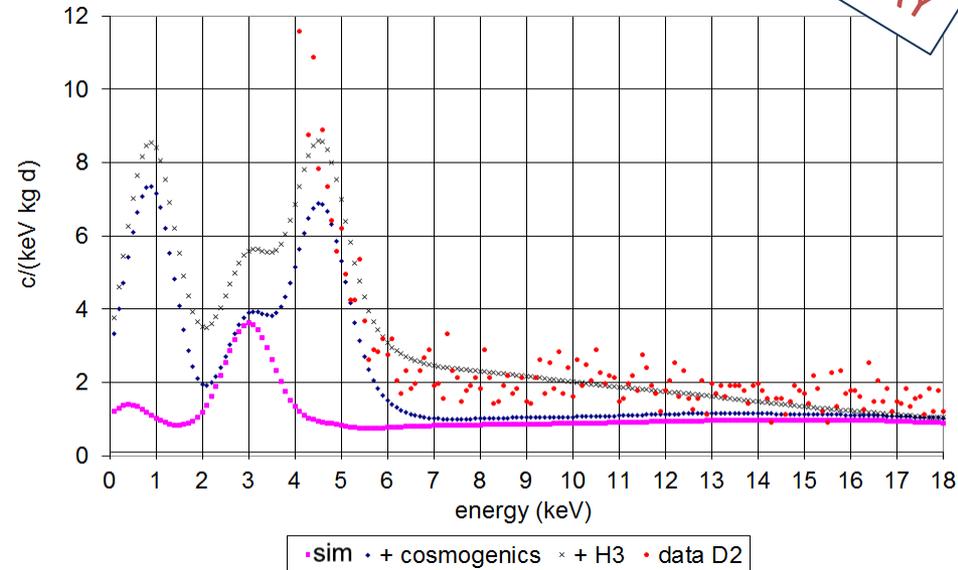
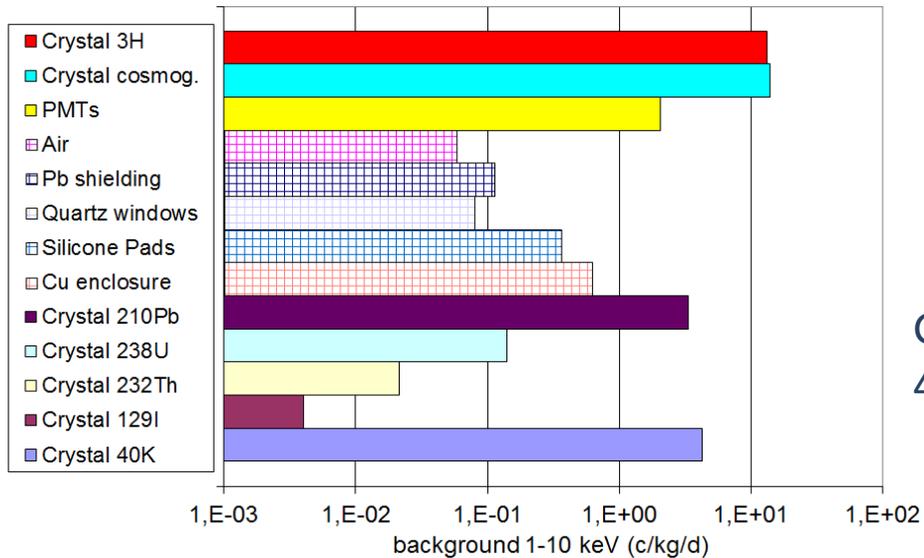
# ANAIS-37 – Background model

PRELIMINARY

Same ANAIS-25 contributions except:

- Less  $^{210}\text{Pb}$
- new detector  $\rightarrow$

Cosmogenic isotopes still decaying



Good agreement with **raw** background down to 4 keVee

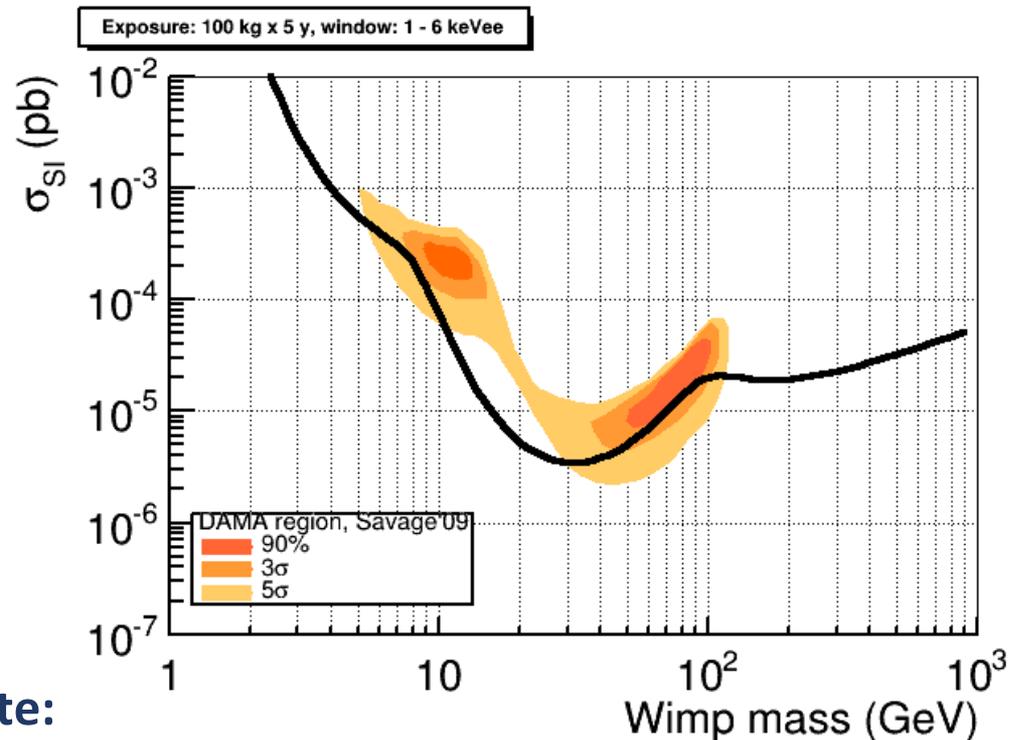
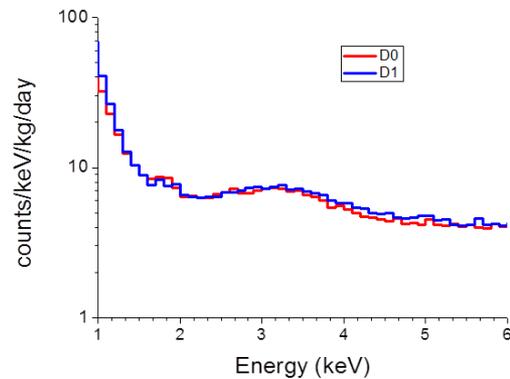
See “Background model for NaI(Tl) detectors for the ANAIS dark matter project” poster for details

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- **ANAIS status and prospects**

# ANAIS – Sensitivity prospects

Detection limit at 90% C.L. with a critical limit at 90% C.L.

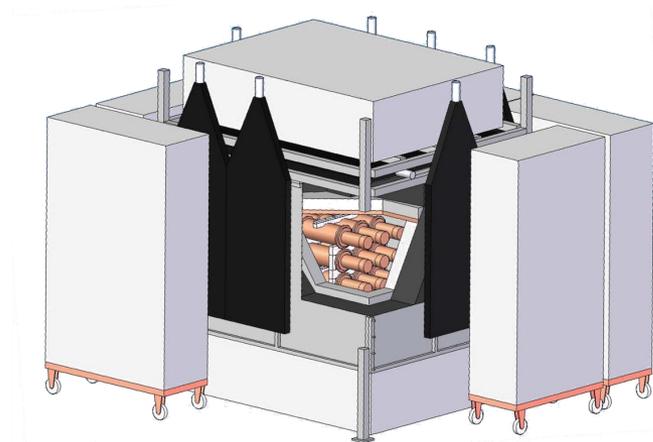


## Conservative background estimate:

Contribution of 2.57 mBq/kg of  $^{210}\text{Pb}$  subtracted to the background measured at ANAIS-25. Further improvement is expected in next modules, coincident background rejection not included.

# ANAIS – Status and Schedule

- Acquisition of new 6 modules:
  - Start data taking with 4 modules January 2016, with final shielding
  - **Start data taking with 9 modules (3×3 matrix, 112.5 kg) along 2016**
- Ready to run:
  - Shielding available
  - Muon tagging system
  - Slow Control  
(temperatures, external Rn activity, N<sub>2</sub> flux, PMT voltage ... )
  - PMTs Ham R12669SEL2 under testing
  - Front-end modules available
  - DAQ software and analysis algorithms tested
- Improvement of filtering and selection protocols and efficiency estimation at low energy
- Simulation of 3×3 matrix and coincidence rejection factor
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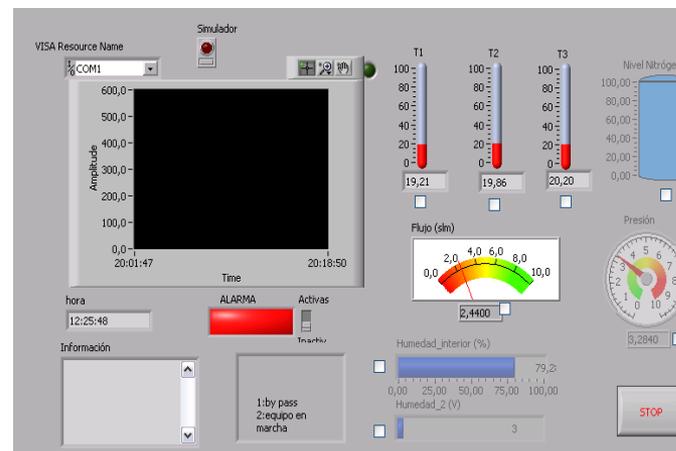
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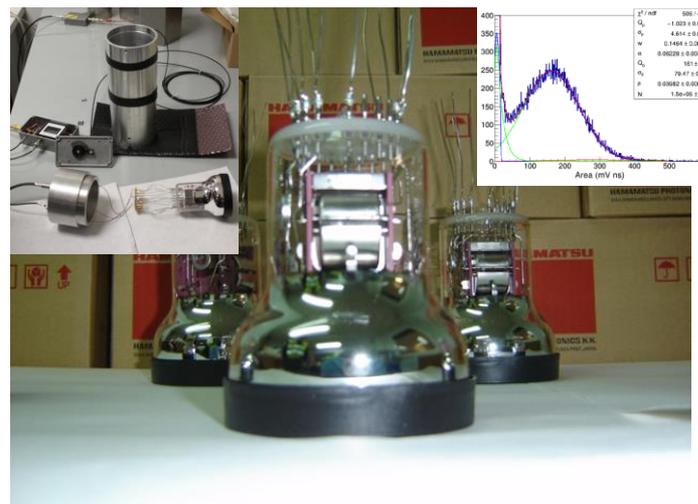
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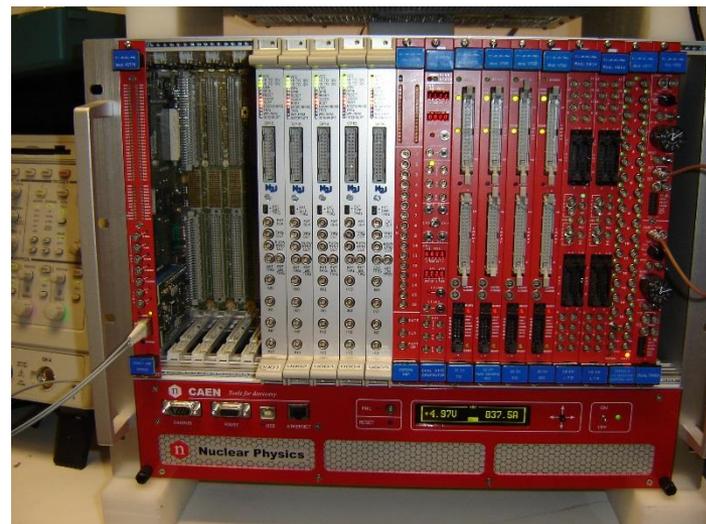
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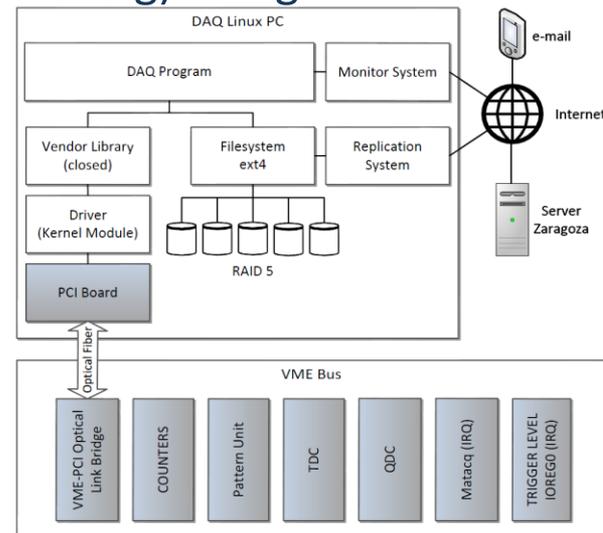
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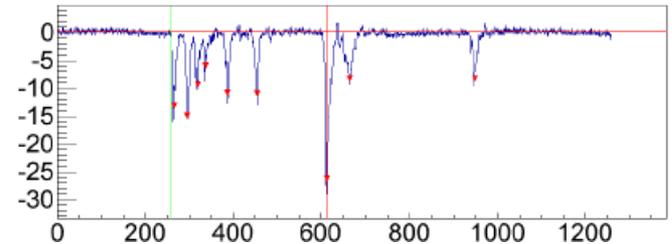
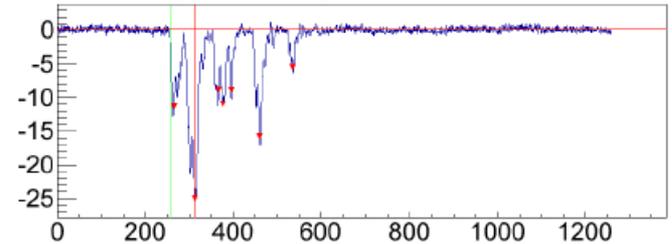
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- Simulation of 3×3 matrix and coincidence rejection factor
- Simulation of liquid scintillator veto



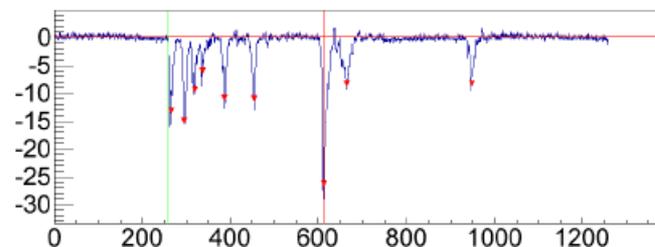
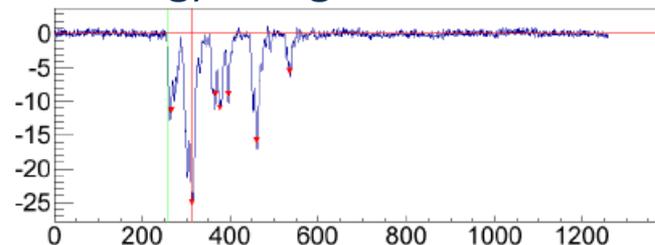
# ANAIS – Status and Schedule

- Acquisition of new 6 modules
  - Start data taking with 4 modules January 2016, with final shielding
  - Start data taking with 9 modules (3×3 matrix, 112.5 kg) along 2016
- Ready to run:
  - Shielding available
  - Muon tagging system
  - Slow Control  
(temperatures, external Rn activity, N<sub>2</sub> flux, PMT voltage ... )
  - PMTs Ham R12669SEL2 under testing
  - Front-end modules available
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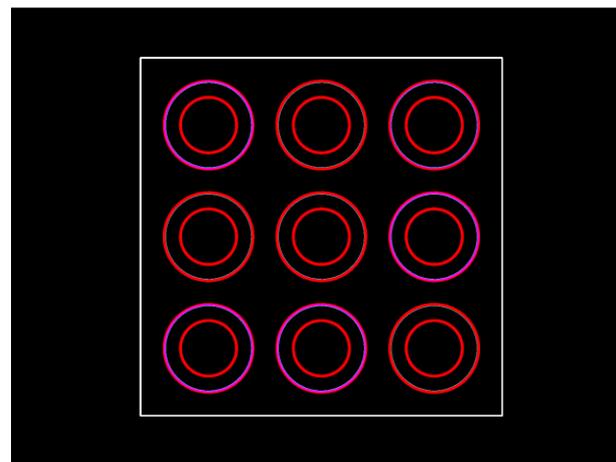
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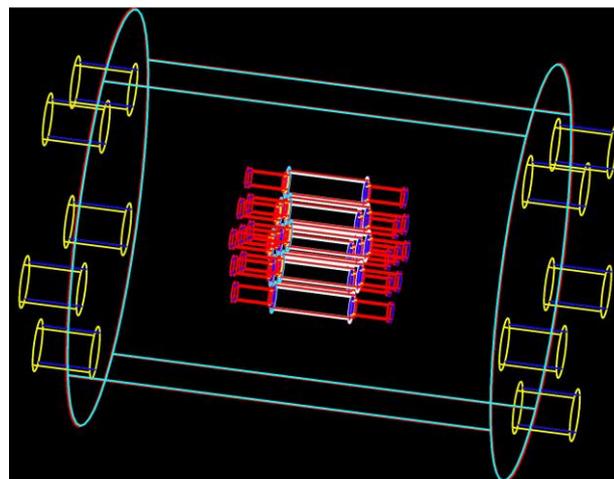
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# Summary

- The ANAIS experiment, intended to confirm the DAMA/LIBRA annual modulation positive signal, is under development at the Canfranc Underground Laboratory
  - 112.5 kg (3×3 crystal matrix) of NaI could be taking data along 2016
  - Shielding, electronic front-end, DAQ and software, slow control ready for the full experiment
- Good quality NaI detectors from Alpha Spectra have been fully characterized:
  - Outstanding light collection improving energy resolution and with 1 keVee energy threshold at reach
  - K content at the level of 40 ppb,  $^{210}\text{Pb}$  reduction 0.6 mBq/kg (reduced by a factor 5, further improvement is expected in next prototypes)
- Background measured in ANAIS-25 and ANAIS-37 set-ups is well understood down to 3 keVee
- Long-living cosmogenic isotopes  $^{22}\text{Na}$  and  $^3\text{H}$  are relevant in the region of interest
- Good sensitivity prospects for exploring the DAMA/LIBRA signal (even under very conservative assumptions) already achieved



**Universidad**  
Zaragoza



**LSC**

*Laboratorio Subterráneo de Canfranc*

# Status of the ANAIS dark matter project at the Canfranc underground laboratory

Miguel Ángel Oliván  
on behalf of the ANAIS team

Universidad de Zaragoza  
Laboratorio subterráneo de Canfranc

# ANAIS – Blank module

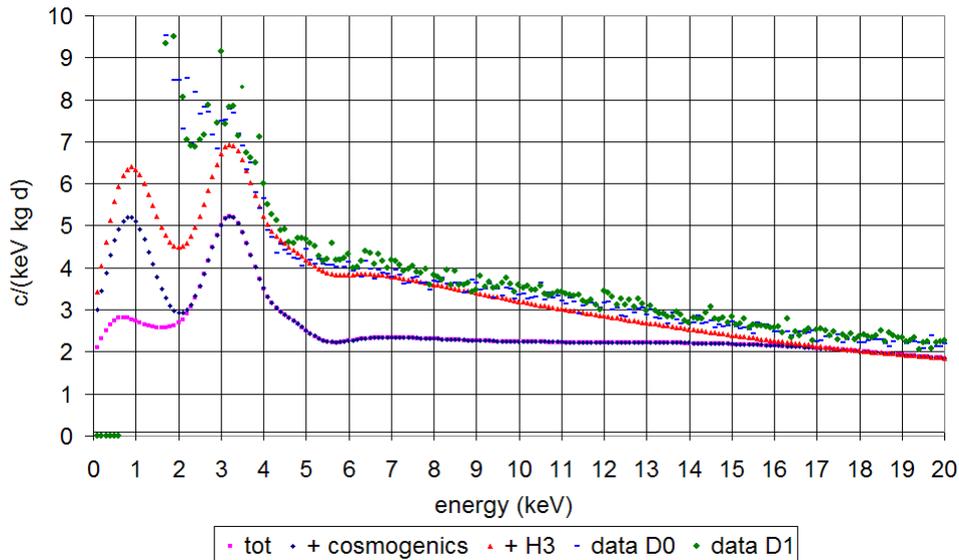
- A “detector” without crystal in order to characterize PMT events



# ANAIS-25 – Background model

PRELIMINARY

✓ Hypothesis:  $^3\text{H}$  contribution



An activity of 0.2 mBq/kg of  $^3\text{H}$  in the NaI crystals significantly improves the agreement with data

## Data for $^3\text{H}$ in NaI:

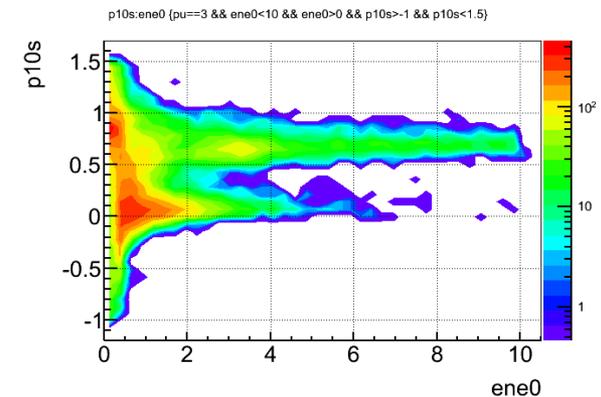
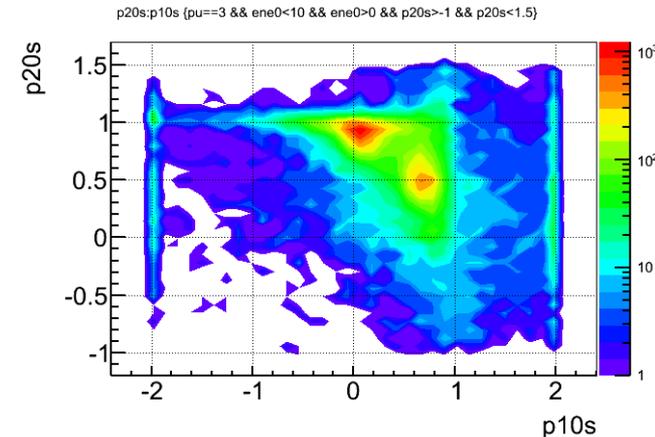
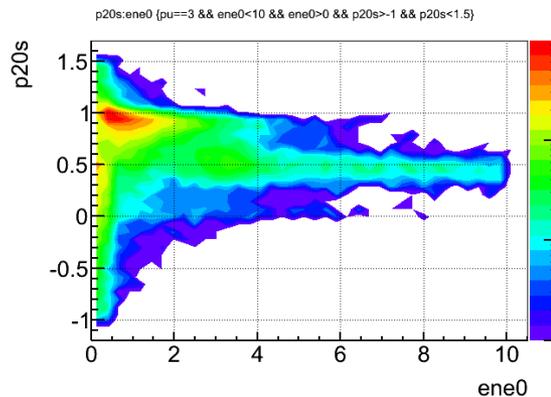
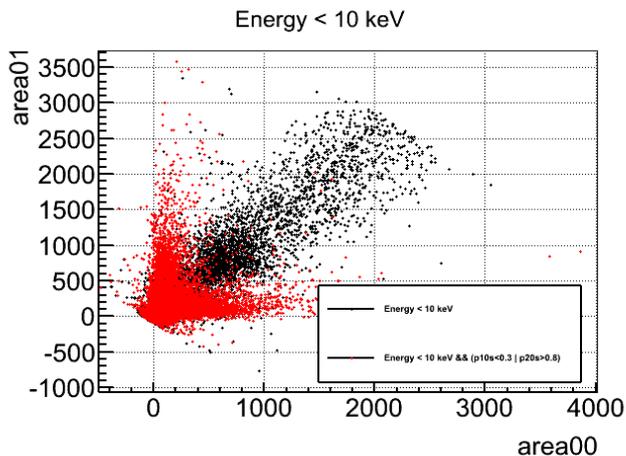
- DAMA/LIBRA:  $A < 0.09$  mBq/kg  
R. Bernabei et al., NIMA 592 (2008) 297
- Calculated production rate:  $R = 31.1$  kg $^{-1}$  d $^{-1}$   
D. M. Mei et al, Astropart. Phys. 31, 417–420 (2009).

For a production rate of 50 kg $^{-1}$ d $^{-1}$  an exposure of 1.9 y to the neutron flux at Grand Junction, Colorado, would produce the required  $^3\text{H}$  activity in ANAIS-25 crystals

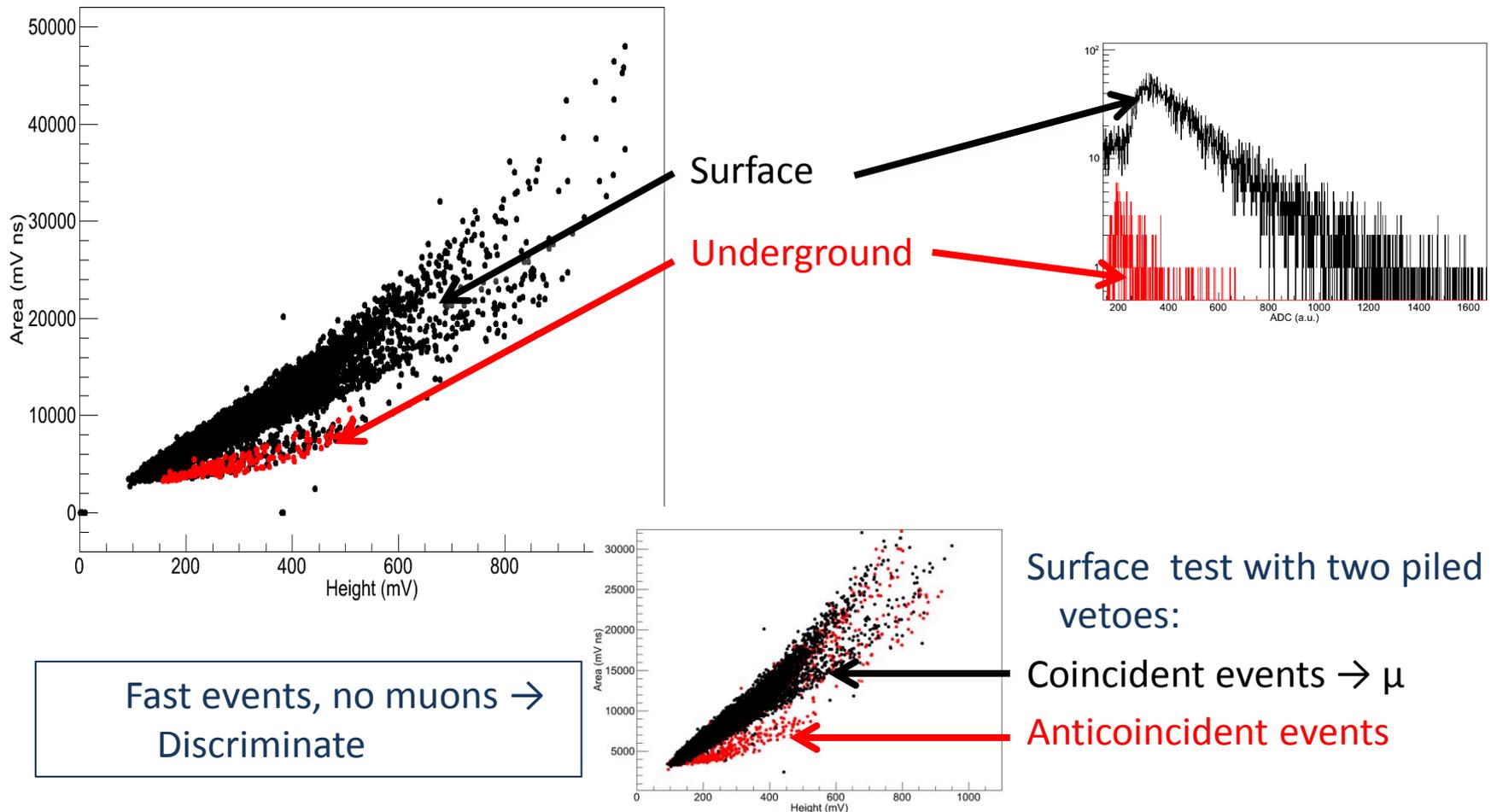
# ANAIS-25 – Data analysis

## Description of the cuts used to remove PMT events

- 1) Cut in the number of peaks in the pulse ( $n > 2$  in each PMT)
- 2) Cut in temporal parameters of the pulse:  $p1s$ ,  $p2s$
- 3) Cut in asymmetry in the light sharing



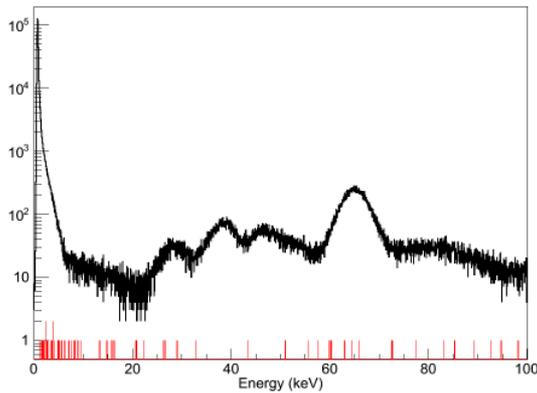
# ANAIS – Muon vs. Fast events



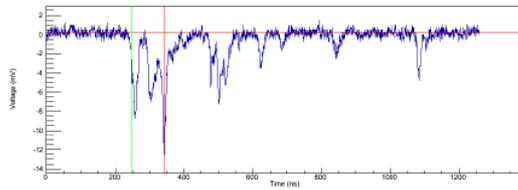
# ANAIS – Muon tagging

PRELIMINARY

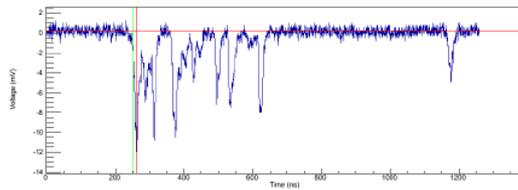
Muon event tagging by threshold and shape based on ADC values (fully compatible with digitized waveform)



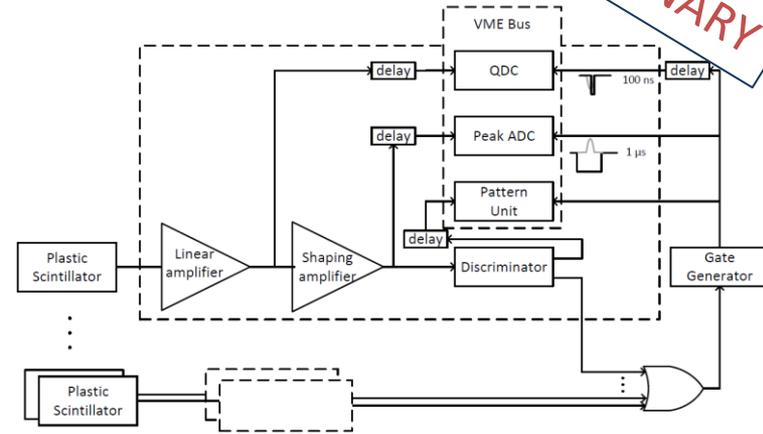
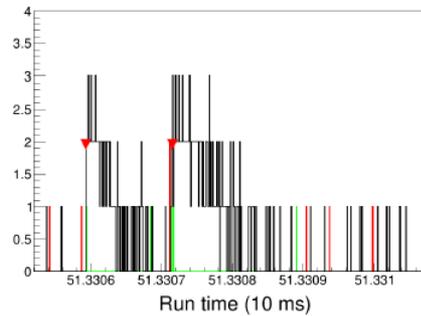
(a)



p21



(b)



Face	Plastic	Measured rate ( $\mu \times m^{-2} \times s^{-1}$ ) $\times 10^{-3}$	Measured rate ( $\mu \times m^{-2} \times s^{-1}$ ) $\times 10^{-3}$
South	#1	$5.27 \pm 0.10$	$5.43 \pm 0.07$
	#2	$5.53 \pm 0.10$	$5.43 \pm 0.07$
West	#3	$4.88 \pm 0.10$	$4.75 \pm 0.07$
	#4	$4.61 \pm 0.10$	$4.75 \pm 0.07$
East	#5	$4.53 \pm 0.10$	$4.54 \pm 0.07$
	#6	$4.54 \pm 0.10$	$4.54 \pm 0.07$
North	#7	$5.05 \pm 0.10$	$4.91 \pm 0.07$
	#8	$4.77 \pm 0.10$	$4.91 \pm 0.07$
Top	#9	$7.23 \pm 0.12$	
	#10	$7.54 \pm 0.12$	$7.36 \pm 0.07$
	#11	$7.32 \pm 0.12$	$7.36 \pm 0.07$

Table 4.3: Muon detection rate

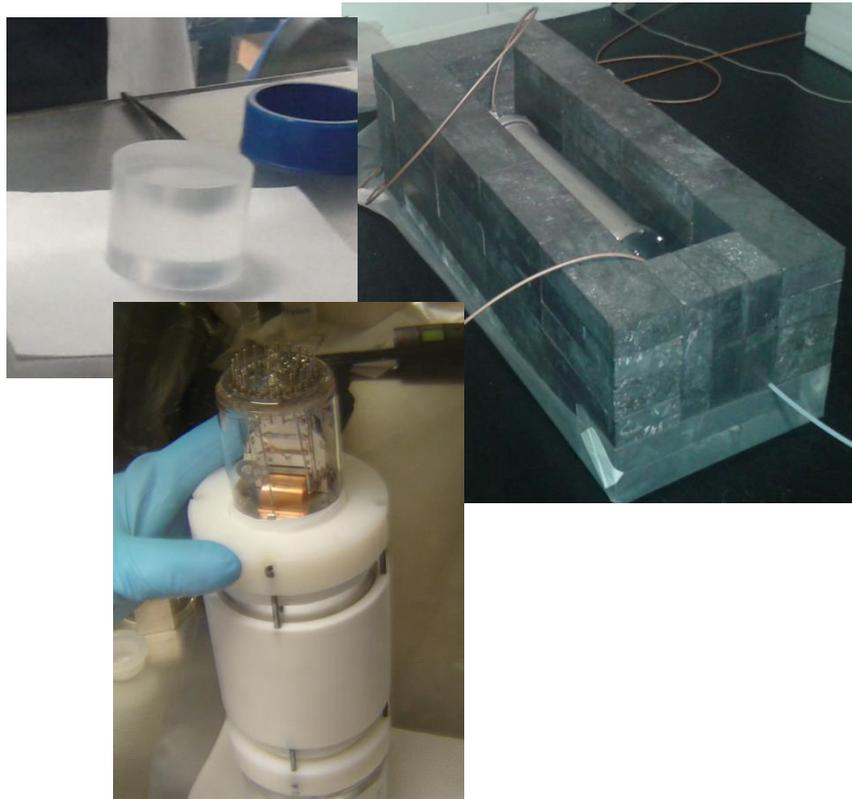
Faces	Coincident rate (mHz)
Top-North	$3.26 \pm 0.07$
Top-South	$1.05 \pm 0.04$
Top-East	$0.87 \pm 0.04$
Top-West	$0.95 \pm 0.04$
North-South	$0.37 \pm 0.02$
East-West	$0.14 \pm 0.01$
North-East	$0.48 \pm 0.03$
North-West	$0.34 \pm 0.02$
South-East	$0.45 \pm 0.03$
South-West	$0.47 \pm 0.03$

Table 4.4: Faces coincident rate

# AN AIS – Slow Control

- Slow control
  - External Rn
  - Temperatures
    - Environment
    - Frontend
    - Inside the shielding
  - N<sub>2</sub> flux
  - HV Power supply voltage & current

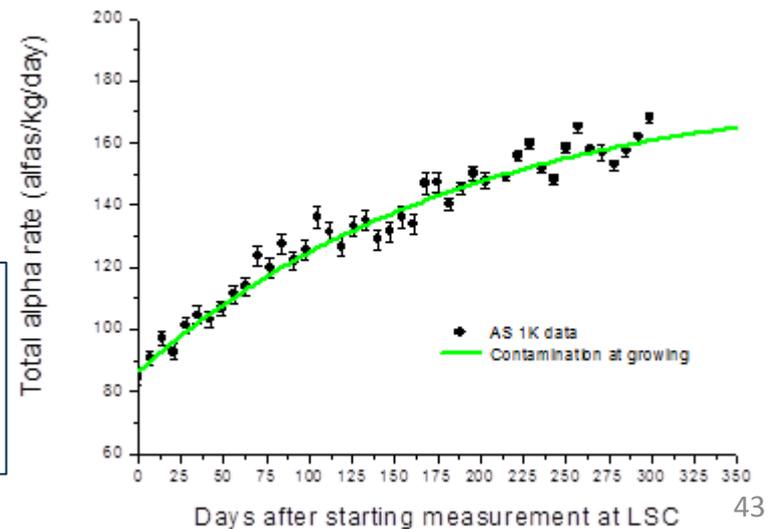
# AS1K – Alpha contamination



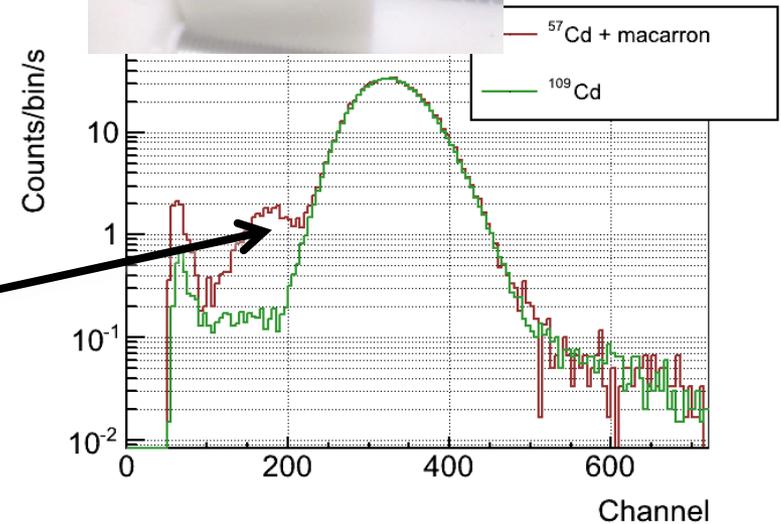
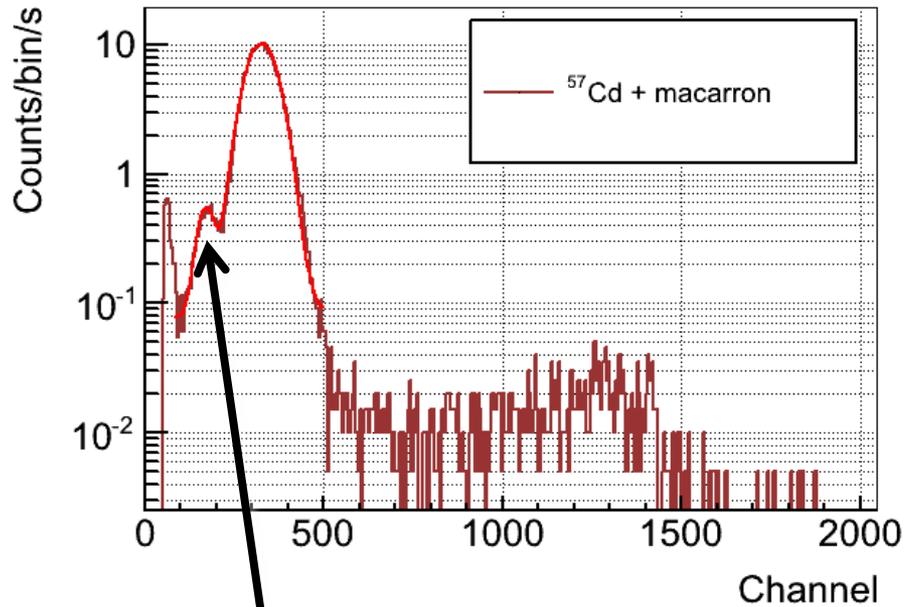
1 kg Alpha Spectra grown crystal, encapsulated at UZ to test at LSC for  $\alpha$  contamination .

$\alpha$  rate, determined by PSA. Compatible with broken chain at  $^{210}\text{Pb}$  and contamination at crystal growing.

Alpha Spectra has updated the purification and growing methods  $\rightarrow$  A new crystal available soon to test radiopurity

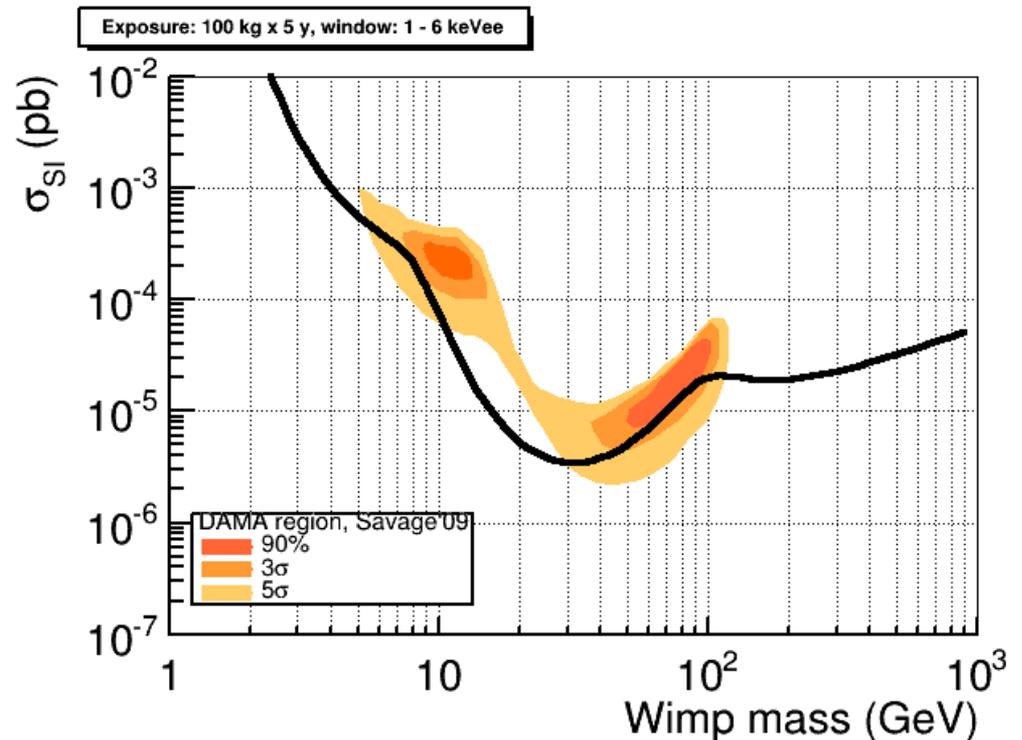
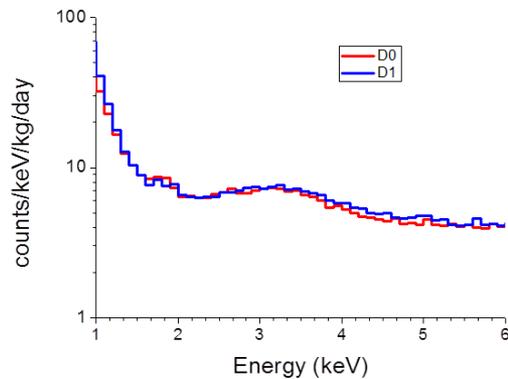


# ANAIS – Low energy calibration 12 keV line



# ANAIS – Prospects

Detection limit at 90% C.L. with a critical limit at 90% C.L.



Computed using Cebrian, S., et al. "Sensitivity plots for WIMP direct detection using the annual modulation signature." *Astroparticle Physics* 14.4 (2001): 339-350.