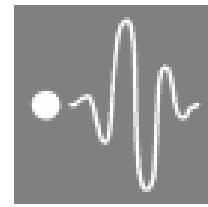


# Threshold corrections and vacuum stability

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Loops and Legs in Quantum Field Theory  
Leipzig, 24–29 April 2016



In collaboration w/ A. Bednyakov, F. Bezrukov, M. Yu. Kalmykov,  
A. Pikelner, M. Shaposhnikov, and O. Veretin.

Introduction  
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Running & Matching  
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EW vacuum stability  
oooooooooooo

Cosmological implications  
oooo

Outlook  
ooo

# Outline

## 1 Introduction

## 2 Running & Matching

## 3 EW vacuum stability

## 4 Cosmological implications

## 5 Outlook

# Introduction



The Nobel Prize in Physics 2013 was awarded jointly to François Englert and Peter W. Higgs "for the theoretical discovery of a mechanism that contributes to our understanding of the origin of mass of subatomic particles, and which recently was confirmed through the discovery of the predicted fundamental particle, by the ATLAS and CMS experiments at CERN's Large Hadron Collider."

# Higgs potential

## SM Higgs sector: complex scalar doublet $\Phi$

$$\mathcal{L} = \partial_\mu \phi^\dagger \partial^\mu \phi - V(|\phi|^2), \quad V = \mu^2 |\phi|^2 + \lambda |\phi|^4$$

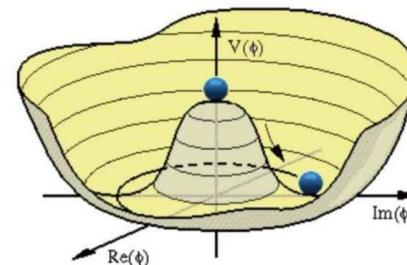
Unbroken phase:  $\mu^2 > 0$

Broken phase:  $\mu^2 < 0$

## After SSB and Higgs mechanism:

$$\Phi = \begin{pmatrix} 0 \\ \frac{v+H}{\sqrt{2}} \end{pmatrix}$$

$$\mathcal{L} = \frac{1}{2} \partial_\mu H \partial^\mu H - V(H), \quad V = -\frac{\lambda v^4}{4} + \lambda v^2 H^2 + \lambda v H^3 + \frac{\lambda}{4} H^4$$



- $\frac{\partial V}{\partial H} \Big|_{H=v} = 0 \quad \rightsquigarrow \quad -\mu^2 = \lambda v^2 \equiv \frac{m_H^2}{2}$
  - $\frac{gv}{2} \equiv m_W \quad \rightsquigarrow \quad v = 2^{-1/4} G_F^{-1/2} = 246.220 \text{ GeV}$
  - $m_H$  is free parameter.

So far, **bare** fields and parameters.

# Properties of the Higgs boson

$$\mathcal{L}_H = \left( m_W^2 W_\mu^+ W^{-\mu} + \frac{m_Z^2}{2} Z_\mu Z^\mu \right) \left( 1 + \frac{H}{V} \right)^2 - \sum_f m_f \bar{f} f \left( 1 + \frac{H}{V} \right) - \frac{m_H^2}{2} H^2 \left( 1 + \frac{H}{2V} \right)^2$$

Quantum numbers	$Q = 0$ $J^{PC} = 0^{++}$
VEV	$v = 2^{-1/4} G_F^{-1/2} \approx 246.22 \text{ GeV}$
Couplings	$g_{VVH} = 2^{5/4} G_F^{1/2} m_V^2 \quad V = W, Z$ $g_{VVHH} = 2^{3/2} G_F m_V^2$ $g_{ffH} = 2^{3/4} G_F^{1/2} m_f$ $\lambda = 2^{-1/2} G_F m_H^2$ $g_{HHH} = 6v\lambda$ $g_{HHHH} = 6\lambda$
Mass	$M_H$ is free parameter

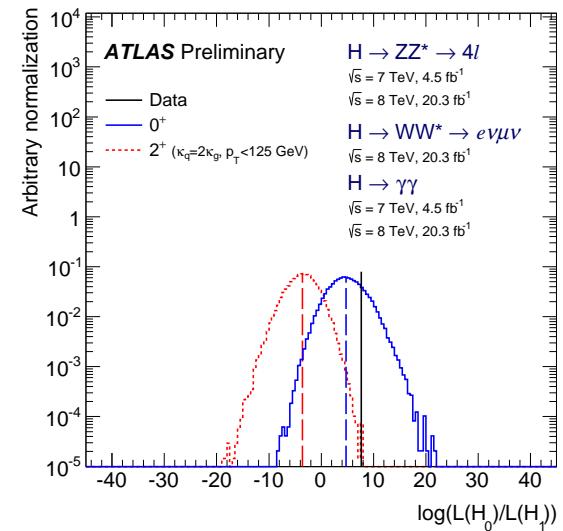
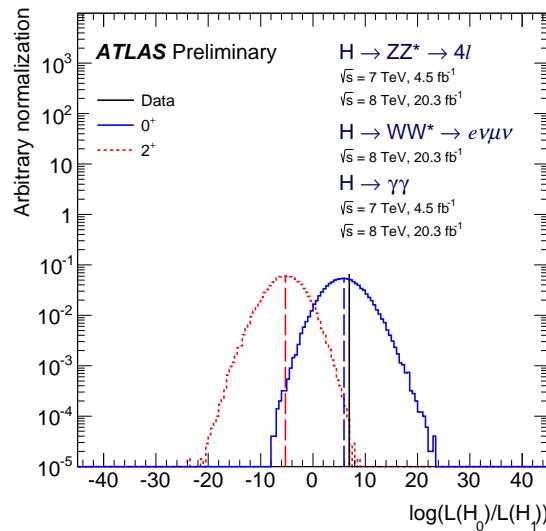
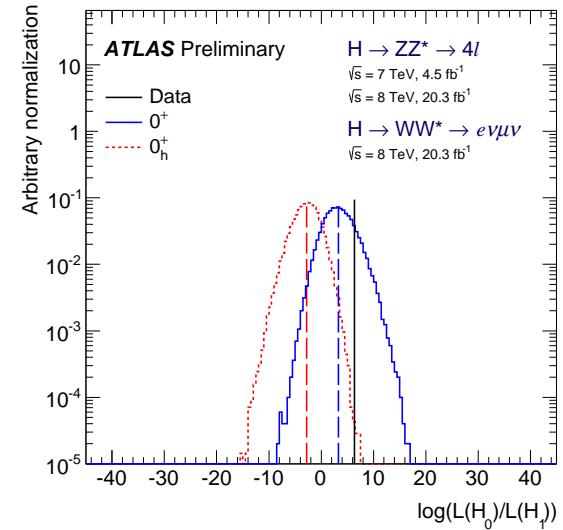
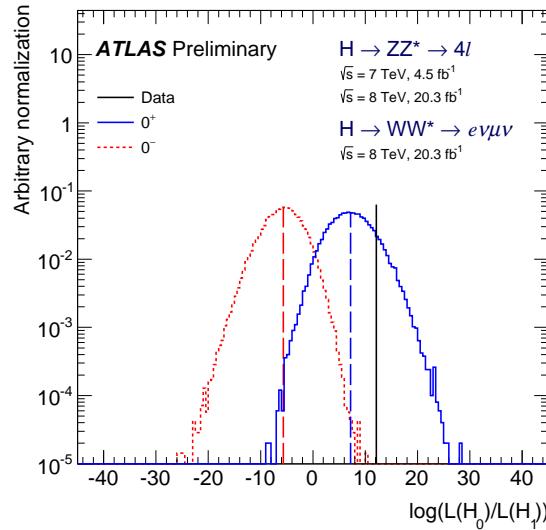
# $J^P$ & Lagrangian tensor structure

$0^+:$   $H V_\mu V^\mu$

$$0^-: \frac{1}{\Lambda} H V_{\mu\nu} \tilde{V}^{\mu\nu}$$

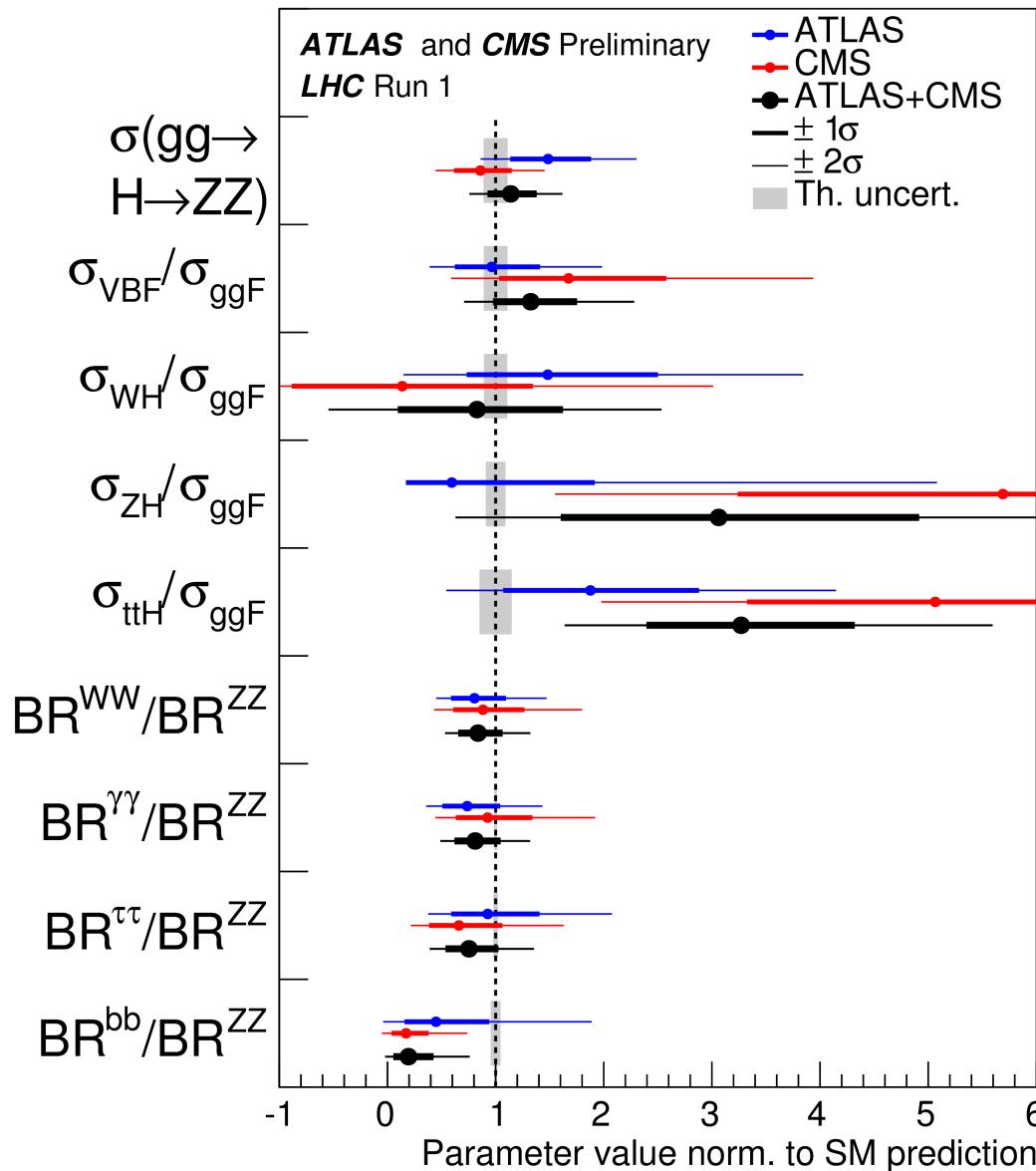
$$O_h^+ \equiv \frac{1}{\Lambda} H V_{\mu\nu} V^{\mu\nu}$$

$$2^-: \frac{1}{\Lambda} H_{\mu\nu} \\ \times [\kappa_g \mathcal{T}_g^{\mu\nu} \\ + \kappa_q \mathcal{T}_q^{\mu\nu}]$$



ATLAS-CONF-2015-008 (13 Mar 2015)  $4.5 \text{ fb}^{-1}$  @  $\sqrt{s} = 7 \text{ TeV}$ ,  $20.3 \text{ fb}^{-1}$  @  $\sqrt{s} = 8 \text{ TeV}$

# Cross sections & branching ratios



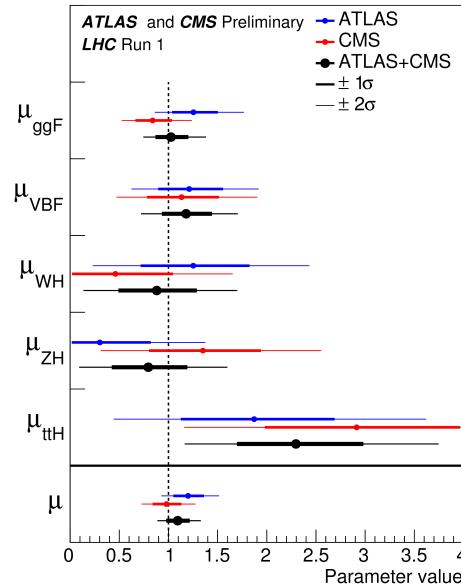
ATLAS-CONF-2015-044  
CMS-PAS-HIG-15-002  
(15 September 2015)  
5  $fb^{-1}$  @  $\sqrt{s} = 7$  TeV  
20  $fb^{-1}$  @  $\sqrt{s} = 8$  TeV  
per experiment

# Signal strengths

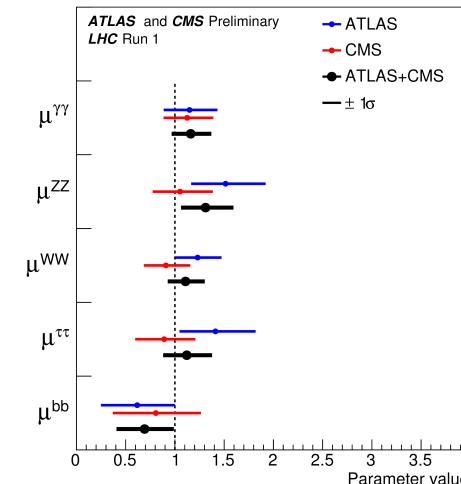
Production and decay channel  $i \rightarrow H \rightarrow f$

Narrow-width approximation:  $\sigma(i \rightarrow H \rightarrow f) = \sigma_i \times \text{BR}^f = \sigma_i \times \frac{\Gamma^f}{\Gamma_H}$

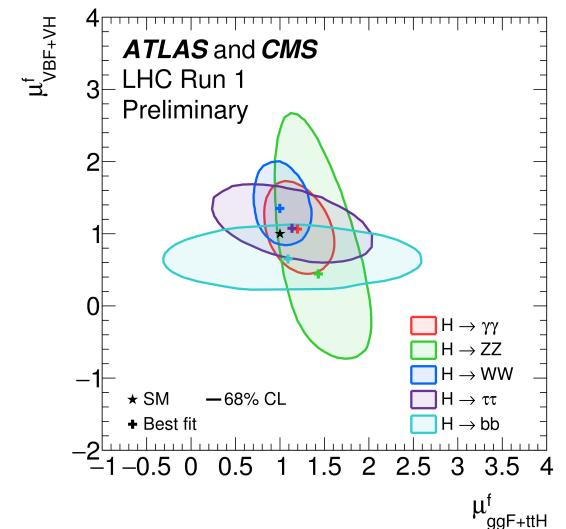
$$\mu_i = \frac{\sigma_i}{(\sigma_i)_{\text{SM}}}$$



$$\mu_f = \frac{\text{BR}^f}{(\text{BR}^f)_{\text{SM}}}$$

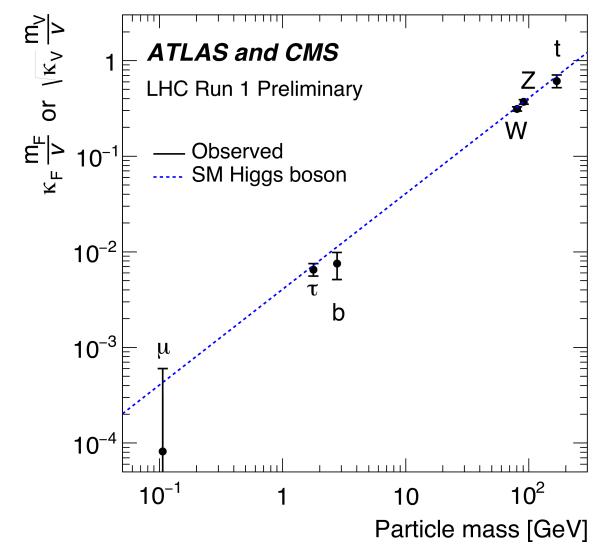
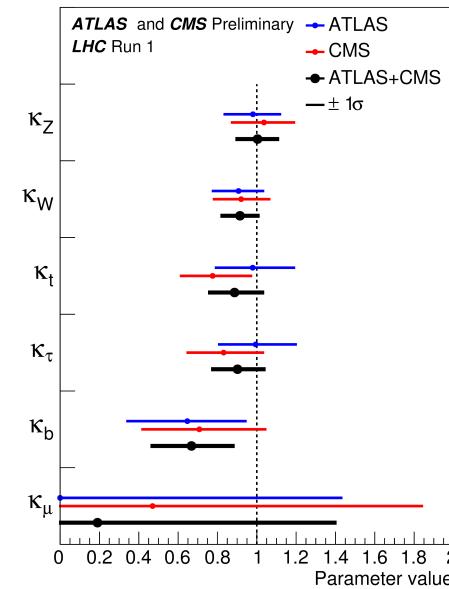
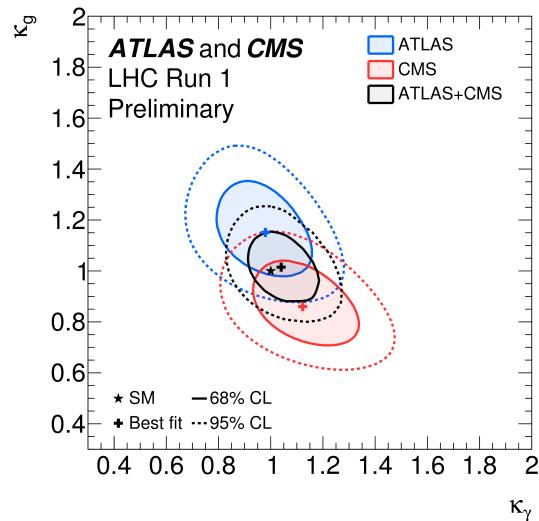


$$\mu_i^f = \mu_i \times \mu_f$$

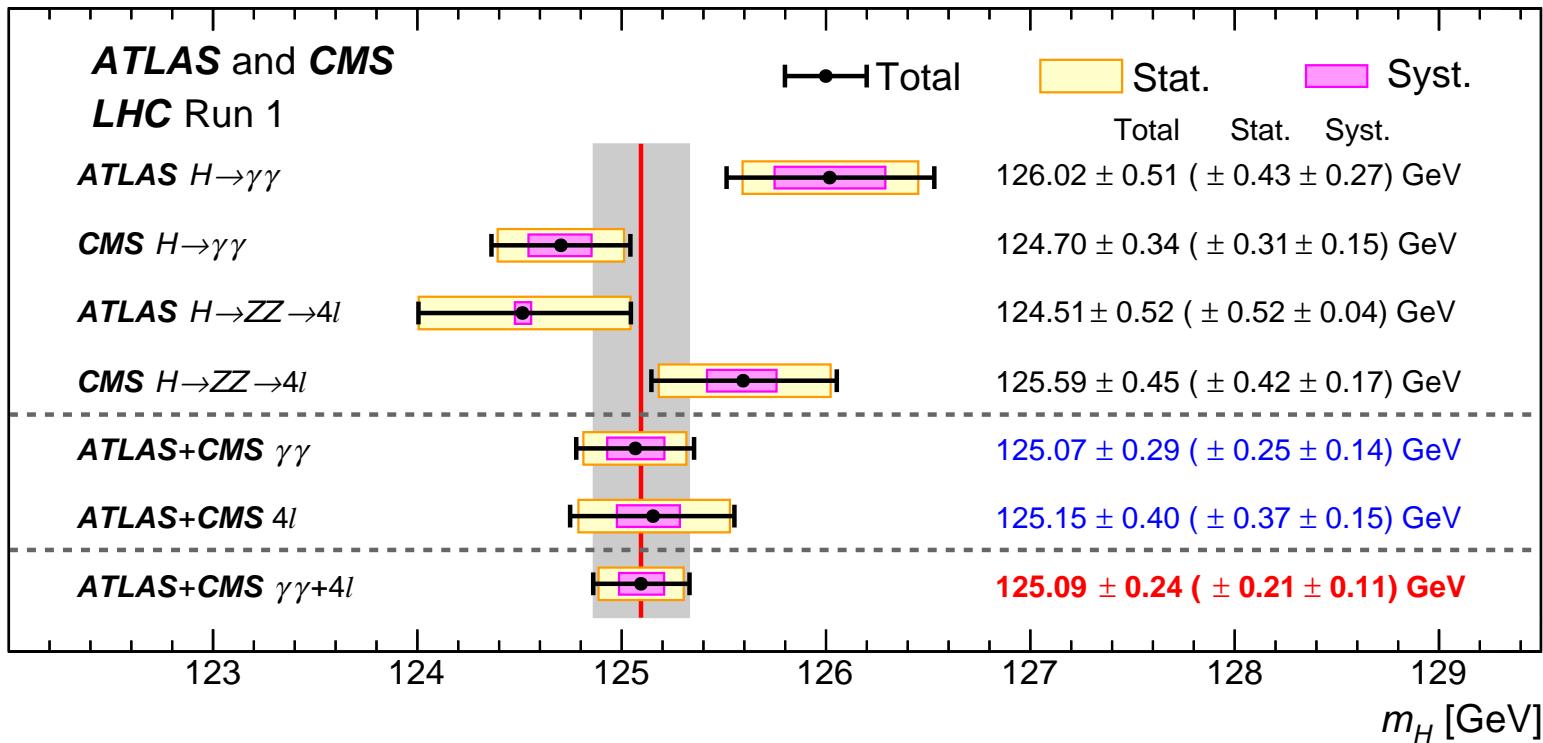


# Coupling modifiers

$\kappa_i^2 = \frac{\sigma_i}{(\sigma_i)_{\text{SM}}}$  except for  $\kappa_g$   
 $\kappa_f^2 = \frac{\Gamma^f}{(\Gamma^f)_{\text{SM}}}$  except for  $\kappa_\gamma$



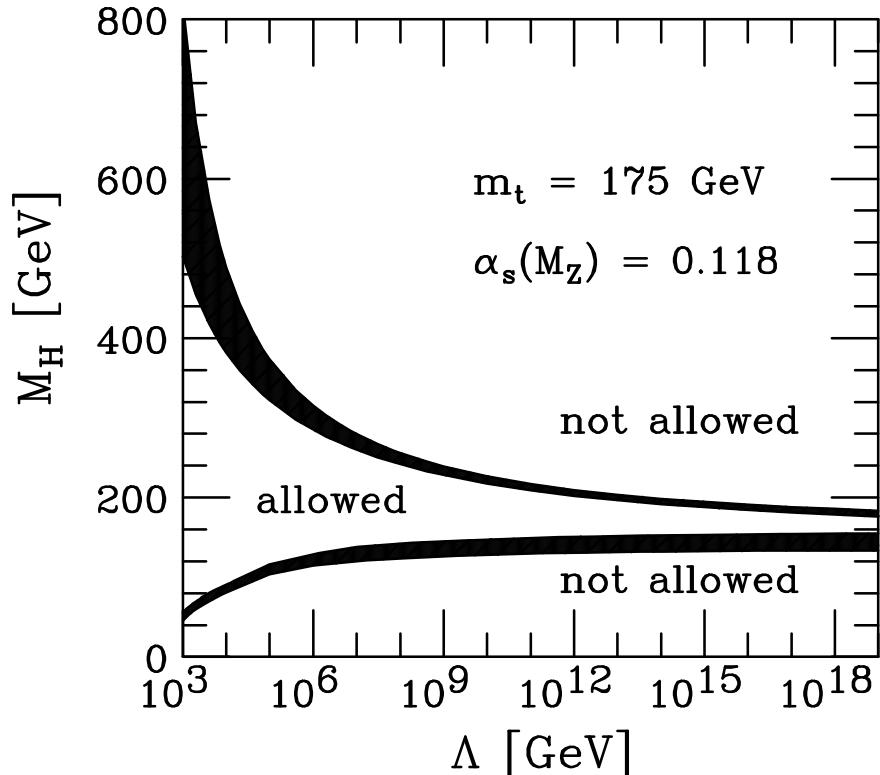
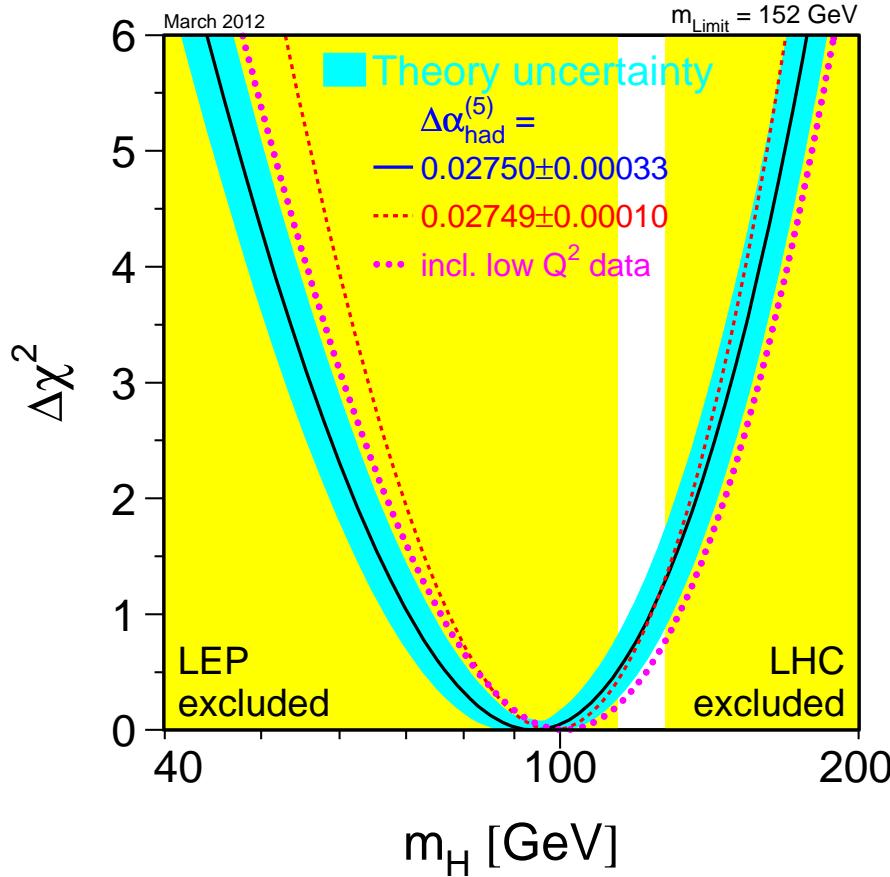
# Mass measurements



$$M_H = (125.09 \pm 0.24) \text{ GeV}$$

ATLAS & CMS, PRL114(2015)191803

# EW precision tests, triviality & vacuum stability



- $m_H = (125.09 \pm 0.24) \text{ GeV}$  agrees w/ EW precision data.
- Triviality bound satisfied.
- How about vacuum stability bound?

# Renormalization: RG evolution

Cosmological applications require reliable predictions over very large range of scales:  $v \lesssim \mu \lesssim M_P$

Use  $\overline{\text{MS}}$  renormalization scheme: running couplings

$\lambda(\mu), y_t(\mu), g_s(\mu), \dots$

Two-step procedure: 1. RG evolution:

$$\mu^2 \frac{d\lambda(\mu)}{d\mu^2} = \beta_\lambda = \frac{1}{16\pi^2} (12\lambda^2 + 6\lambda y_t^2 - 3y_t^4) + \dots$$

$$\mu^2 \frac{dy_t(\mu)}{d\mu^2} = \beta_{y_t} = \frac{1}{16\pi^2} y_t \left( \frac{9}{4} y_t^2 - 4g_s^2 \right) + \dots$$

$$\mu^2 \frac{dg_s(\mu)}{d\mu^2} = \beta_{g_s} = \frac{1}{16\pi^2} g_s^3 \left( -\frac{11}{2} + \frac{n_f}{3} \right) + \dots$$

$$\beta_\lambda^{(3)}, \beta_{y_t}^{(3)}$$

Chetyrkin, Zoller, JHEP06(2012)033; 04(2013)091

Bednyakov *et al.*, JHEP01(2013)017; PLB722(2013)336; NPB875(2013)552

$$\beta_{\dots}^{(3)}, \beta_{g_s, y_t}^{(3)}$$

Mihaila *et al.*, PRL108(2012)151602; PRD86(2012)096008

$$\beta_{g_s}^{(3)}$$

Tarasov *et al.*, PLB93(1980)429

# Threshold corrections

## 2. Matching at $\mu = \mathcal{O}(v)$ :

$$\lambda(\mu) = 2^{-1/2} G_F m_H^2 [1 + \delta_H^{(1)}(\mu) + \dots]$$

$$\delta_H^{(1)}(\mu) = \frac{G_F m_H^2}{8\pi^2 \sqrt{2}} \left[ 6 \ln \frac{\mu^2}{m_H^2} + \frac{25}{2} - \frac{3}{2} \pi \sqrt{3} + \mathcal{O}\left(\frac{m_Z^2}{m_H^2} \ln \frac{m_H^2}{m_Z^2}\right) \right]$$

Sirlin, Zucchini, NPB266(1986)389

$$y_t(\mu) = 2^{3/4} G_F^{1/2} m_t [1 + \delta_t^{(1)}(\mu) + \dots]$$

$$\delta_t^{(1)}(\mu) = \frac{Q_t^2 \alpha + C_F \alpha_s(\mu)}{4\pi} \left( -3 \ln \frac{\mu^2}{m_t^2} - 4 \right)$$

$$+ \frac{G_F m_t^2}{8\pi^2 \sqrt{2}} \left[ \frac{9}{2} \ln \frac{\mu^2}{m_t^2} + \frac{11}{2} - 2\pi \frac{m_H}{m_t} + \mathcal{O}\left(\frac{m_H^2}{m_t^2} \ln \frac{m_t^2}{m_H^2}\right) \right]$$

Hempfling, BK, PRD51(1995)1386

$\delta_H^{(\alpha\alpha_s)}, \delta_t^{(\alpha\alpha_s)}$  Bezrukov, Kalmykov, BK, Shaposhnikov, JHEP10(2012)140

$\delta_H^{(y_t^4)}, \delta_t^{(y_t^4)}$  Degrassi *et al.*, JHEP08(2012)098; BK, Veretin, NPB885(2014)459

$\delta_H^{(\alpha^2)}, \delta_t^{(\alpha^2)}$  Buttazzo *et al.*, JHEP12(2013)089

$\delta_x^{(\alpha^2)}$  for all  $x$  BK, Veretin, Pikelner, NPB896(2015)19

# $\overline{\text{MS}}$ renormalization scheme

Parameters of the symmetric phase:  $g, g', \lambda, m_\phi, y_f$

Parameters of the broken phase:  $e, m_W, m_Z, m_H, m_f$

Tree-level relationships:

$$\frac{1}{e^2} = \frac{1}{g^2} + \frac{1}{g'^2}$$

$$\frac{4m_W^2}{v^2} = g^2 \quad \frac{4m_Z^2}{v^2} = g^2 + g'^2 \quad \frac{m_H^2}{2v^2} = \lambda \quad \frac{2m_f^2}{v^2} = y_f^2$$

$$\frac{1}{v^2} = \frac{\lambda}{-m_\phi^2} = \frac{e^2}{4m_W^2(1 - m_W^2/m_Z^2)}$$

Treat as exact in the  $\overline{\text{MS}}$  renormalization scheme.

# On-shell renormalization scheme

- Pole masses:

$$p^2 = M_B^2 : 0 = p^2 - m_{B,0}^2 - \Pi_{BB}(p^2) \quad (B = H, W)$$

$$p^2 = M_Z^2 : 0 = p^2 - m_{Z,0}^2 - \Pi_{ZZ,T}(p^2) - \frac{\Pi_{\gamma Z,T}^2(p^2)}{p^2 - \Pi_{\gamma\gamma,T}(p^2)}$$

$$\not{p} = m_f : 0 = \not{p} - m_{f,0} - \Sigma_f(\not{p})$$

- Fine-structure constant:  $\alpha_{\text{Th}}$  absorbs radiative corrections to Thomson scattering.

Induces large corrections  $\propto \alpha \ln(q^2/m_\ell^2)$  and hadronic uncertainties!  $\rightsquigarrow$  Use instead Sirlin, PRD22(1980)971

$$G_F = \frac{\pi \alpha_{\text{Th}}}{\sqrt{2} M_W^2 (1 - M_W^2/M_Z^2)} (1 + \Delta r)$$

# Matching

- Masses:

$$m_0^2 = M^2 - \Pi(M^2) = m^2(\mu) \left( 1 + \frac{Z^{(1)}}{\varepsilon} + \frac{Z^{(2)}}{\varepsilon^2} + \dots \right)$$

$$Z^{(j)} = \frac{g^2}{16\pi^2} Z_\alpha^{(j)} + \frac{g^2}{16\pi^2} \frac{g_s^2}{16\pi^2} Z_{\alpha\alpha_s}^{(j)} + \left( \frac{g^2}{16\pi^2} \right)^2 Z_{\alpha^2}^{(j)} + \dots$$

- Couplings:

$$2^{1/2} G_F = \frac{1 + \Delta \bar{r}(\mu)}{v^2(\mu)}$$

$$\frac{e^2}{8m_W^2(1 - m_W^2/m_Z^2)} (1 + \Delta \bar{r}) = \left[ \sqrt{Z_{2,e} Z_{2,\nu_e} Z_{2,\mu} Z_{2,\nu_\mu}} A(e + \nu_e \rightarrow \mu + \nu_\mu) \right]_{\text{hard}}$$

**hard:** Nullify external four-momenta and light-fermion masses before loop integration. Awramik *et al.*, PRD68(2003)053004

# Threshold corrections

- Couplings:

$$g^2(\mu) = 2^{5/2} G_F M_W^2 [1 + \delta_W(\mu)]$$

$$g^2(\mu) + g'^2(\mu) = 2^{5/2} G_F M_Z^2 [1 + \delta_Z(\mu)]$$

$$e^2(\mu) = 2^{5/2} G_F M_W^2 [1 + \delta_W(\mu)] \left[ 1 - \frac{M_W^2}{M_Z^2} \frac{1 + \delta_W(\mu)}{1 + \delta_Z(\mu)} \right]$$

$$\lambda(\mu) = 2^{-1/2} G_F M_H^2 [1 + \delta_H(\mu)]$$

$$y_f(\mu) = 2^{3/4} G_F^{1/2} M_f [1 + \delta_f(\mu)]$$

$$g_s^2(\mu) = 4\pi\alpha_s^{(5)}(\mu) [1 + \delta_{\alpha_s}(\mu)]$$

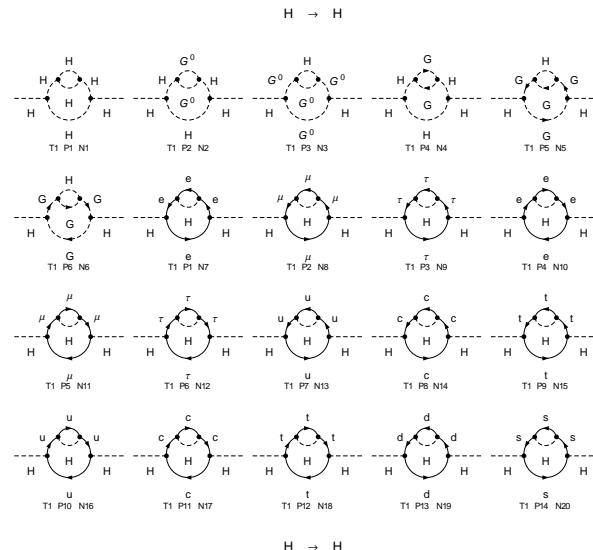
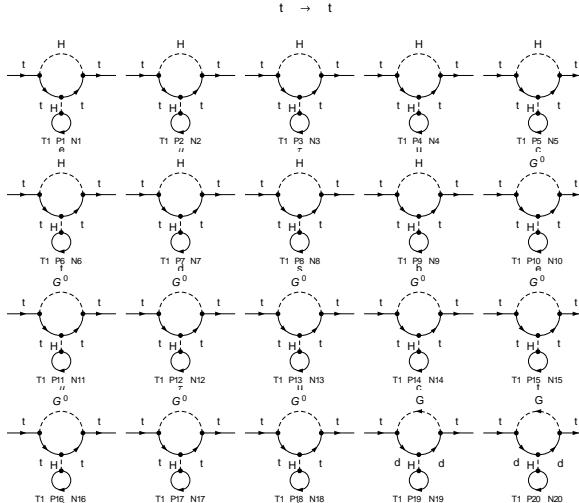
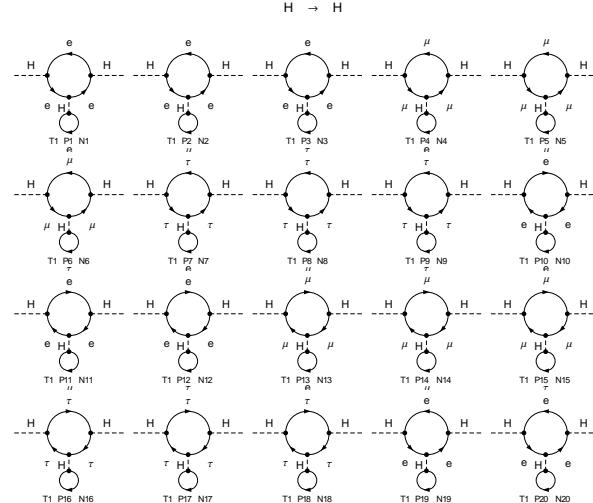
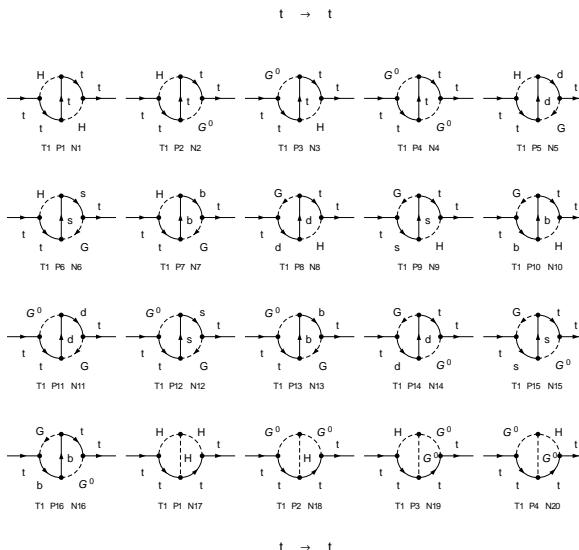
- Masses:

$$m_B^2(\mu) = M_B^2 [1 + \Delta\bar{r}(\mu)] [1 + \delta_B(\mu)] \quad B = W, Z, H$$

$$m_f(\mu) = M_f [1 + \Delta\bar{r}(\mu)]^{1/2} [1 + \delta_f(\mu)] \quad f = t, b$$

Exact two-loop results. BK, Veretin, Pikelner, NPB896(2015)19

# Typical Feynman diagrams

 $\delta_H(\mu)$ 

 $\delta_t(\mu)$ 


# Tools

## Packages used:

- Generation of diagrams: **QGRAF**, **DIANA** Nogueira, Tentyukov
- Reduction: **TARCER** (*Mathematica*) Mertig
  - ~~> Gauge invariance upon inclusion of all tadpoles ✓
- Numerical evaluation of master integrals: **TSIL** (*C++*) Martin

## Program library created: **mr** for matching and running (*C++*)

BK, Pikelner, Veretin, 1601.08143 [hep-ph]

- Matching @ 2-loop EW & 4-loop QCD level
- RG evolution @ 3-loop EW & 4-loop QCD level

Available for download from URL: <http://apik.github.io/mr/>

# Numerical results

- Corrections to  $\delta_H(M_t)$  in  $10^{-4}$

$M_H$ [GeV]	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-114.8	-107.5	-26.6 (-29.1)	-248.7
125	-114.5	-105.2	-26.4 (-29.2)	-246.1
126	-114.1	-103.1	-26.3 (-29.3)	-243.5

- Corrections to  $\delta_t(M_t)$  in  $10^{-4}$

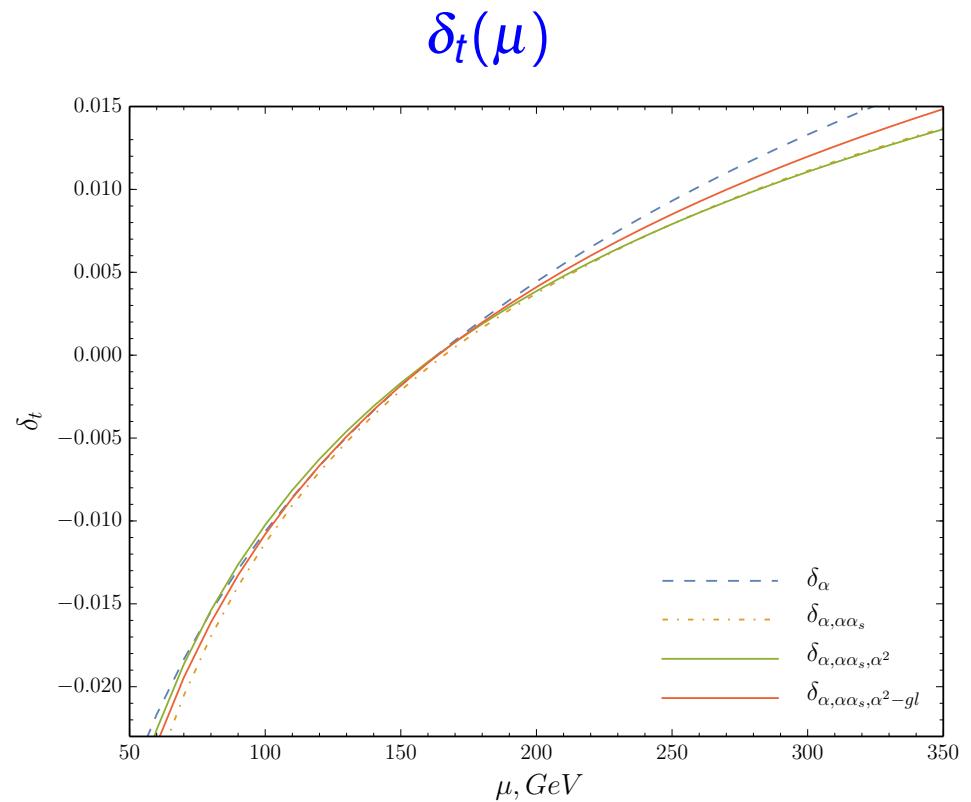
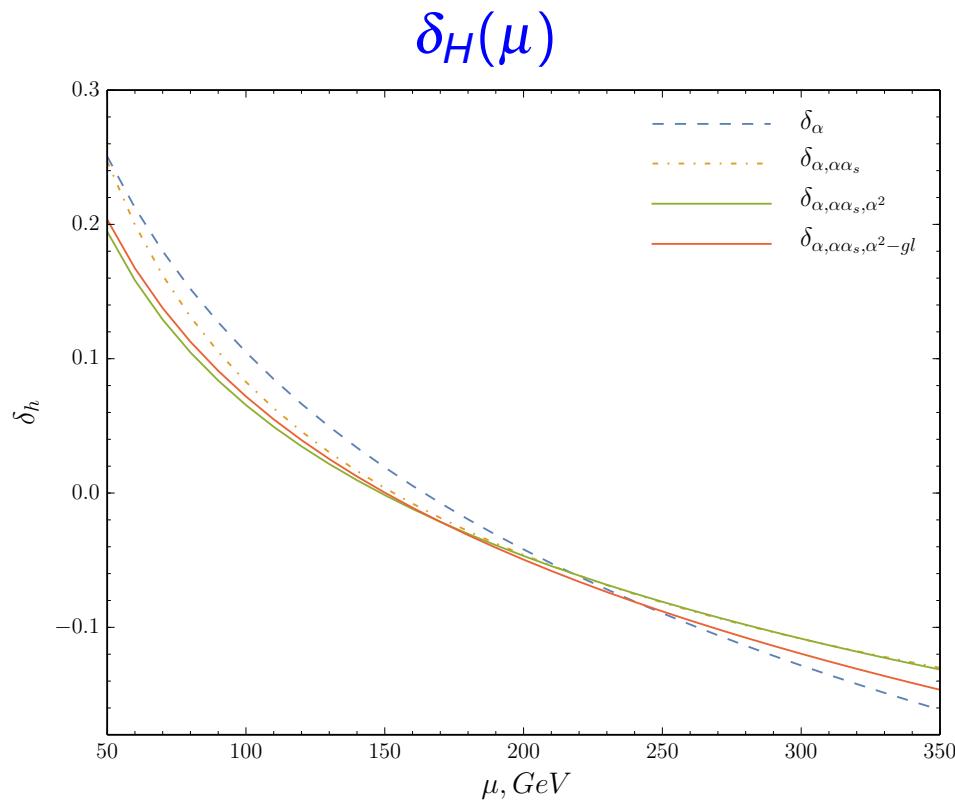
$M_H$ [GeV]	QCD	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-599.3	13.5	-4.4	2.7 (3.1)	-587.4
125	-599.3	13.2	-4.3	2.7 (3.1)	-587.7
126	-599.3	12.9	-4.2	2.7 (3.1)	-587.9

- Corrections to  $\delta_b(M_b)$

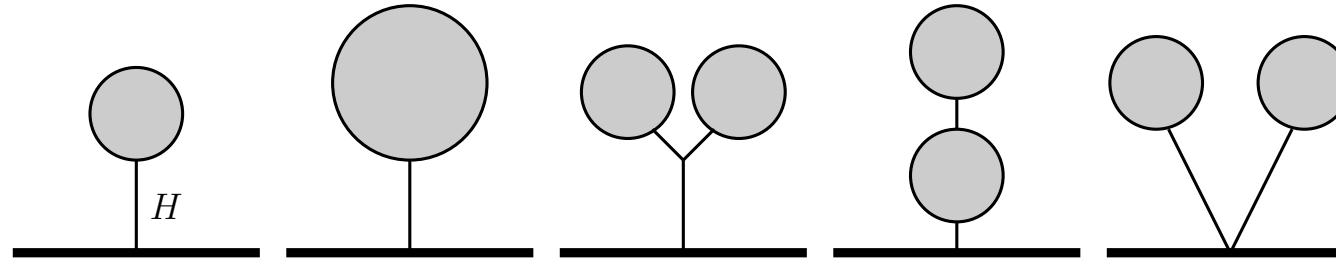
$$\begin{aligned} \{1 + \delta_b(M_b)\}_{\text{QCD}, O(\alpha), O(\alpha\alpha_s), O(\alpha^2)} = \\ 1 - 0.1728 - 0.0190 - 0.0112 + 0.0032(0.0033) \end{aligned}$$

BK, Pikelner, Veretin, NPB896(2015)19

# Numerical results



# Role of tadpoles



- Tadpole is gauge dependent and  $\propto 1/M_H^{2n}$  for  $M_H \rightarrow 0$ .
- Adjust vev  $v^0 = \sqrt{-(m_\phi^0)^2/\lambda^0}$  to eliminate term  $\propto H$  in bare  $\mathcal{L}$ .  
*Hempfling, BK, PRD51(1995)1386*
- No tadpole counterterm.
- Include tadpoles order by order to ensure finiteness and gauge independence.
- $\Delta\bar{r}(\mu)$  and  $\delta_x(\mu)$  are gauge independent through  $\mathcal{O}(\alpha^2)$ .
- At  $\mathcal{O}(\alpha^2)$ ,  $\delta_x(\mu) \propto M_H^0$  for  $x = W, Z, f$ ;  $\delta_H(\mu) \propto M_H^{-2}$ ;  $\Delta\bar{r}(\mu) \propto M_H^{-4}$  for  $M_H \rightarrow 0$ .
- $m_f(\mu)$  gauge independent, but receive large EW corrections.  $\rightsquigarrow$   
Use instead *Jegerlehner, Kalmykov, BK, PLB722(2013)123*  

$$m_f^Y(\mu) = 2^{-3/4} G_F^{-1/2} y_f(\mu) = M_f [1 + \delta_f(\mu)] = m_f(\mu) [1 + \Delta\bar{r}(\mu)]^{-1/2}.$$

# Tadpole cancellation

- Consider  $m_f(\mu)$  and  $y_f(\mu)$  at  $\mathcal{O}(\alpha)$  Hempfling, BK, PRD51(1995)1386

$$m_f(\mu) = M_f(1 + \delta M_f/M_f)_{\overline{\text{MS}}}$$

$$y_f(\mu) = 2^{3/4} G_F^{1/2} M_f(1 + \delta M_f/M_f - \delta v/v)_{\overline{\text{MS}}}$$

$$\delta M_f/M_f = \text{Re}[\Sigma_V^f(M_f^2) + \Sigma_S^f(M_f^2)] - 2^{1/4} G_F^{1/2} T/M_H^2$$

$$\delta v/v = [\Pi_W(0)/M_W^2 + E]/2 - 2^{1/4} G_F^{1/2} T/M_H^2$$

- Exact tadpole cancellation also in  $\mathcal{O}(\alpha\alpha_s)$ . Jegerlehner, Kalmykov, NPB676(2004)365; BK, Piclum, Steinhauser, NPB695(2004)199
- Incomplete tadpole cancellation in  $\mathcal{O}(\alpha^2)$  BK, Veretin, NPB885(2014)459; BK, Pikelner, Veretin, NPB896(2015)19
- Similar for  $\lambda(\mu)$ . Sirlin, Zucchini, NPB266(1986)389; Bezrukov *et al.*, JHEP01(2012)140

# Running top and bottom masses

- Corrections to  $m_t(M_t) - M_t$  in GeV

$M_H$ [GeV]	QCD	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-10.38	12.08	-0.39	-0.99 (-0.47)	0.32
125	-10.38	11.88	-0.39	-0.96 (-0.45)	0.14
126	-10.38	11.67	-0.38	-0.94 (-0.44)	-0.03

- Corrections to  $m_t^Y(M_t) - M_t$  in GeV

$M_H$ [GeV]	QCD	$\mathcal{O}(\alpha)$	$\mathcal{O}(\alpha\alpha_s)$	$\mathcal{O}(\alpha^2)$	total
124	-10.38	0.234	-0.076	0.047 (0.054)	-10.17
125	-10.38	0.229	-0.075	0.047 (0.054)	-10.18
126	-10.38	0.223	-0.073	0.047 (0.054)	-10.18

- Corrections to  $m_b(M_b) - M_b$

$$\begin{aligned} \{m_b(M_b) - M_b\}_{\text{QCD}, O(\alpha), O(\alpha\alpha_s), O(\alpha^2)} = \\ -0.85 - 1.90 - 1.53 + 1.75 \quad (1.80) \text{ GeV} \end{aligned}$$

- Corrections to  $m_b^Y(M_b) - M_b$

$$\begin{aligned} \{m_b^Y(M_b) - M_b\}_{\text{QCD}, O(\alpha), O(\alpha\alpha_s), O(\alpha^2)} = \\ -0.847 - 0.093 - 0.055 + 0.016(0.016) \text{ GeV} \end{aligned}$$

~~~  $m_q^Y$  is much more perturbatively stable than  $m_q(M_q)$ .

# Vaccum stability in a nutshell

Recall

$$\mu^2 \frac{d\lambda(\mu)}{d\mu^2} = \frac{1}{16\pi^2} (12\lambda^2 + 6\lambda y_t^2 - 3y_t^4) + \dots$$

$$\mu^2 \frac{dy_t(\mu)}{d\mu^2} = \frac{1}{16\pi^2} y_t \left( \frac{9}{4} y_t^2 - 4g_s^2 \right) + \dots$$

$$\lambda(m_H) = 0.130 \times \left( \frac{m_H}{125.7 \text{ GeV}} \right)^2, \quad y_t(m_t) = 0.995 \times \frac{m_t}{173.21 \text{ GeV}}, \quad g_s(m_Z) = 1.220$$

For  $m_t \gg m_H$ ,  $\lambda(\mu) = \lambda(v) - \frac{(3/8\pi^2)y_t^4(v)\ln(\mu/v)}{1 - (9/16\pi^2)y_t^2(v)\ln(\mu/v)}$

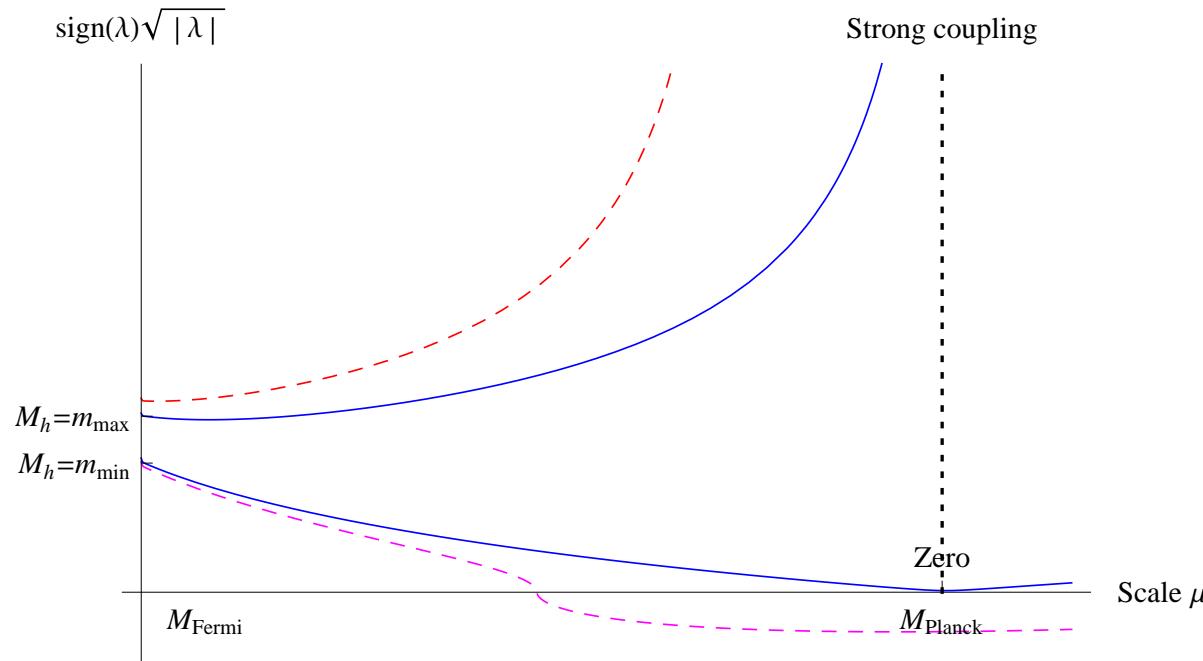
$$\rightsquigarrow \mu^{\text{cri}} \approx v \exp(4\pi^2 v^2 m_H^2 / 3m_t^4) = 8.0 \times 10^3 \text{ GeV}$$

Vacuum stability bound: Lindner, ZPC31(1986)295

If  $m_H < m_H^{\text{cri}}$ , then  $\lambda(\mu) < 0$  for  $\mu > \mu^{\text{cri}}$ .

$\rightsquigarrow$  Decay of universe

# Vaccum stability condition



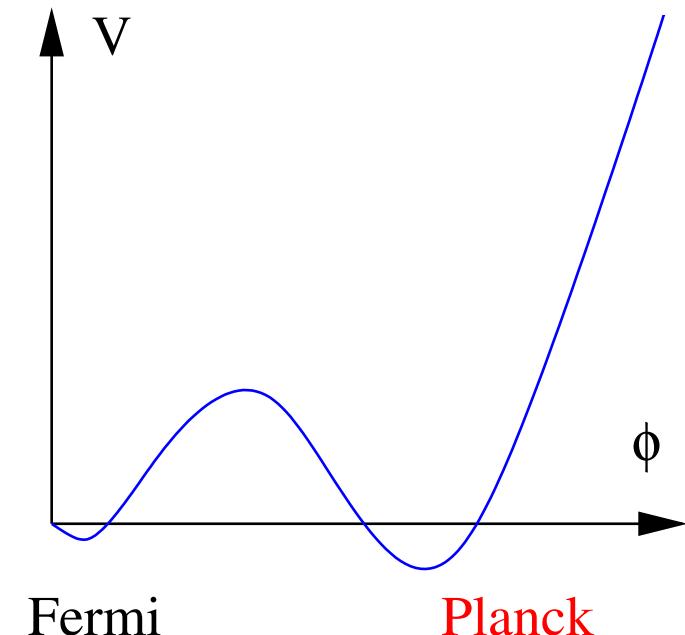
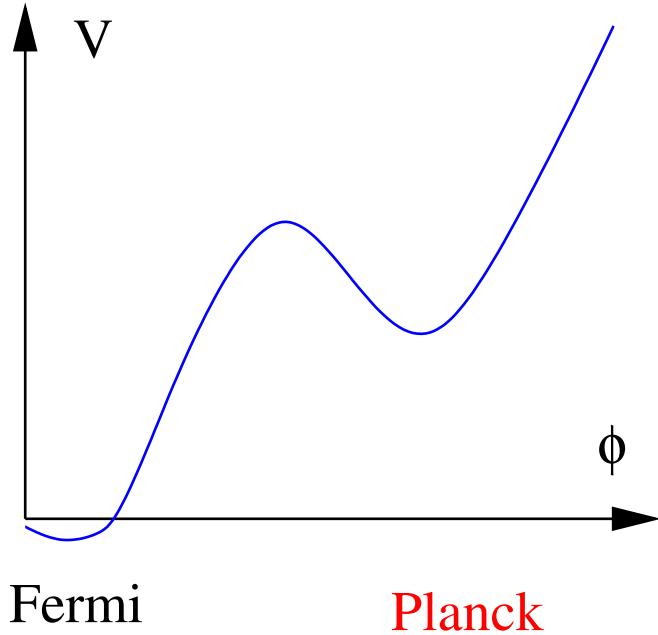
Determine  $\mu^{\text{cri}}$  and  $M_H^{\text{cri}}$  for given  $M_t$  (or  $M_t^{\text{cri}}$  for given  $M_H$ ) so that

$$\lambda(\mu^{\text{cri}}) = \beta_\lambda(\lambda(\mu^{\text{cri}})) = 0$$

↪ Vacuum is stable for  $M_H \geq M_H^{\text{cri}}$  (or  $M_t \leq M_t^{\text{cri}}$ ).

Caveat:  $\mu^{\text{cri}}$ ,  $M_H^{\text{cri}}$ ,  $M_t^{\text{cri}}$  are gauge independent, but (slightly) scheme dependent. ↪ theoretical uncertainty

# Effective potential



Determine  $\tilde{\mu}^{\text{cri}}$  and  $\tilde{M}_H^{\text{cri}}$  for given  $M_t$  (or  $\tilde{M}_t^{\text{cri}}$  for given  $M_H$ ) so that

$$V_{\text{eff}}(\tilde{\mu}^{\text{cri}}) = V_{\text{eff}}(v) \approx 0, \quad V'_{\text{eff}}(\tilde{\mu}^{\text{cri}}) = 0$$

↷ Vacuum is stable for  $M_H \geq \tilde{M}_H^{\text{cri}}$  (or  $M_t \leq \tilde{M}_t^{\text{cri}}$ ).

Caveat:  $\tilde{\mu}^{\text{cri}}, \tilde{M}_H^{\text{cri}}, \tilde{M}_t^{\text{cri}}$  are gauge dependent!

Degassi *et al.*, JHEP08(2012)098; Buttazzo *et al.*, JHEP12(2013)089

# Consistent approach to effective potential

- Reorganize  $V_{\text{eff}}(H)$  in powers of  $\hbar$  so that expansion coefficients are gauge independent at its extrema *Andreassen et al.*, PRL113(2014)241801
- Solve  $V'_{\text{eff}}(H) = 0$  for  $H = \tilde{\mu}^{\text{cri}}$ :

$$\lambda = \frac{1}{256\pi^2} \left[ (g^2 + g'^2)^2 \left( 1 - 3 \ln \frac{g^2 + g'^2}{4} \right) + 2g'^4 \left( 1 - 3 \ln \frac{g'^2}{4} \right) - 48y_t^4 \left( 1 - \ln \frac{y_t^2}{4} \right) \right]$$

- Require that  $V_{\text{min}}^{\text{NLO}} = V_{\text{eff}}^{\text{NLO}}(\tilde{\mu}^{\text{cri}}) \geq 0$  for  $M_H \geq \tilde{M}_H^{\text{cri}}$  (or  $M_t \leq \tilde{M}_t^{\text{cri}}$ )  
e.g. in the Landau gauge
- **Caveat:**  $\tilde{\mu}^{\text{cri}} > M_P$ !

# Critical parameters

$$X = X_0 + \Delta X_{\alpha_s} \frac{\alpha_s^{(5)}(M_Z) - \alpha_s^{(5),\text{exp}}(M_Z)}{\Delta \alpha_s^{(5),\text{exp}}(M_Z)} + \Delta X_M \frac{M - M^{\text{exp}}}{\Delta M^{\text{exp}}} \pm \delta X_{\text{par}} + \delta X_\mu^\pm \pm \delta X_{\text{tru}}$$

| $X$                                    | $X_0$  | $\Delta X_{\alpha_s}$ | $\Delta X_M$ | $\delta X_{\text{par}}$ | $\delta X_\mu^+$ | $\delta X_\mu^-$ | $\delta X_{\text{tru}}$ |
|----------------------------------------|--------|-----------------------|--------------|-------------------------|------------------|------------------|-------------------------|
| $M_t^{\text{cri}}$                     | 171.44 | 0.23                  | 0.20         | 0.001                   | -0.36            | 0.17             | -0.02                   |
| $\log_{10} \mu_t^{\text{cri}}$         | 17.752 | -0.051                | 0.083        | 0.007                   | 0.007            | -0.006           | -0.002                  |
| $M_H^{\text{cri}}$                     | 129.30 | -0.49                 | 1.79         | 0.002                   | 0.72             | -0.33            | 0.04                    |
| $\log_{10} \mu_H^{\text{cri}}$         | 18.512 | -0.158                | 0.381        | 0.008                   | 0.173            | -0.082           | 0.008                   |
| $\tilde{M}_t^{\text{cri}}$             | 171.64 | 0.23                  | 0.20         | 0.001                   | -0.36            | 0.17             | -0.02                   |
| $\log_{10} \tilde{\mu}_t^{\text{cri}}$ | 21.442 | -0.059                | 0.094        | 0.005                   | -0.083           | 0.022            | 0.002                   |
| $\tilde{M}_H^{\text{cri}}$             | 128.90 | -0.49                 | 1.79         | 0.003                   | 0.73             | -0.34            | 0.04                    |
| $\log_{10} \tilde{\mu}_H^{\text{cri}}$ | 22.209 | -0.181                | 0.436        | 0.007                   | 0.092            | -0.062           | 0.013                   |

# Importance of higher orders

- $\mathcal{O}(\alpha^2)$  corrections to all  $\delta_i(\mu)$  BK, Pikelner, Veretin, NPB896(2015)19
- $\mathcal{O}(\alpha_s \alpha)$  and  $\mathcal{O}(\alpha_s^4)$  corrections to  $\delta_{\alpha_s}(\mu)$  Bednyakov, PLB741(2015)262; Schröder, Steinhauser, JHEP01(2006)051; Chetyrkin, Kühn, Sturm, NPB744(2006)121; BK *et al.*, PRL97(2006)042001
- $\mathcal{O}(\alpha_s^4)$  corrections to  $\delta_q(\mu)$  Marquard *et al.*, PRL114(2015)142002

| $X$                                    | $X_0 + \delta X_\mu^\pm$   | w/o $\delta_i^{O(\alpha^2)}$ | w/o $\delta_{\alpha_s}^{O(\alpha \alpha_s, \alpha_s^4)}$ | w/o $\delta_q^{O(\alpha_s^4)}$ |
|----------------------------------------|----------------------------|------------------------------|----------------------------------------------------------|--------------------------------|
| $M_t^{\text{cri}}$                     | $171.44^{-0.36}_{+0.17}$   | $171.55^{-0.47}_{+1.04}$     | $171.43^{-0.36}_{+0.17}$                                 | $171.24^{-0.38}_{+0.19}$       |
| $\log_{10} \mu_t^{\text{cri}}$         | $17.752^{+0.007}_{-0.006}$ | $17.783^{+0.062}_{-0.008}$   | $17.754^{+0.007}_{-0.006}$                               | $17.751^{+0.007}_{-0.007}$     |
| $M_H^{\text{cri}}$                     | $129.30^{+0.72}_{-0.33}$   | $129.06^{+0.95}_{-2.14}$     | $129.32^{+0.73}_{-0.33}$                                 | $129.72^{+0.76}_{-0.38}$       |
| $\log_{10} \mu_H^{\text{cri}}$         | $18.512^{+0.173}_{-0.082}$ | $18.495^{+0.226}_{-0.531}$   | $18.518^{+0.174}_{-0.082}$                               | $18.602^{+0.184}_{-0.094}$     |
| $\tilde{M}_t^{\text{cri}}$             | $171.64^{-0.36}_{+0.17}$   | $171.74^{-0.46}_{+1.04}$     | $171.63^{-0.36}_{+0.17}$                                 | $171.43^{-0.37}_{+0.19}$       |
| $\log_{10} \tilde{\mu}_t^{\text{cri}}$ | $21.442^{-0.083}_{+0.022}$ | $21.485^{-0.085}_{+0.343}$   | $21.445^{-0.083}_{+0.022}$                               | $21.441^{-0.072}_{+0.014}$     |
| $\tilde{M}_H^{\text{cri}}$             | $128.90^{+0.73}_{-0.34}$   | $128.67^{+0.95}_{-2.15}$     | $128.92^{+0.73}_{-0.34}$                                 | $129.32^{+0.76}_{-0.38}$       |
| $\log_{10} \tilde{\mu}_H^{\text{cri}}$ | $22.209^{+0.092}_{-0.062}$ | $22.201^{+0.146}_{-0.171}$   | $22.217^{+0.094}_{-0.062}$                               | $22.312^{+0.113}_{-0.082}$     |

Introduction  
oooooooooooo

Running & Matching  
oooooooooooooooooooo

EW vacuum stability  
oooooooo●○○○

Cosmological implications  
oooo

Outlook  
ooo

# Combined results

PRL 115, 201802 (2015)

Selected for a Viewpoint in Physics  
PHYSICAL REVIEW LETTERS

week ending  
13 NOVEMBER 2015



## Stability of the Electroweak Vacuum: Gauge Independence and Advanced Precision

A. V. Bednyakov,<sup>1</sup> B. A. Kniehl,<sup>2</sup> A. F. Pikelner,<sup>2</sup> and O. L. Veretin<sup>2</sup>

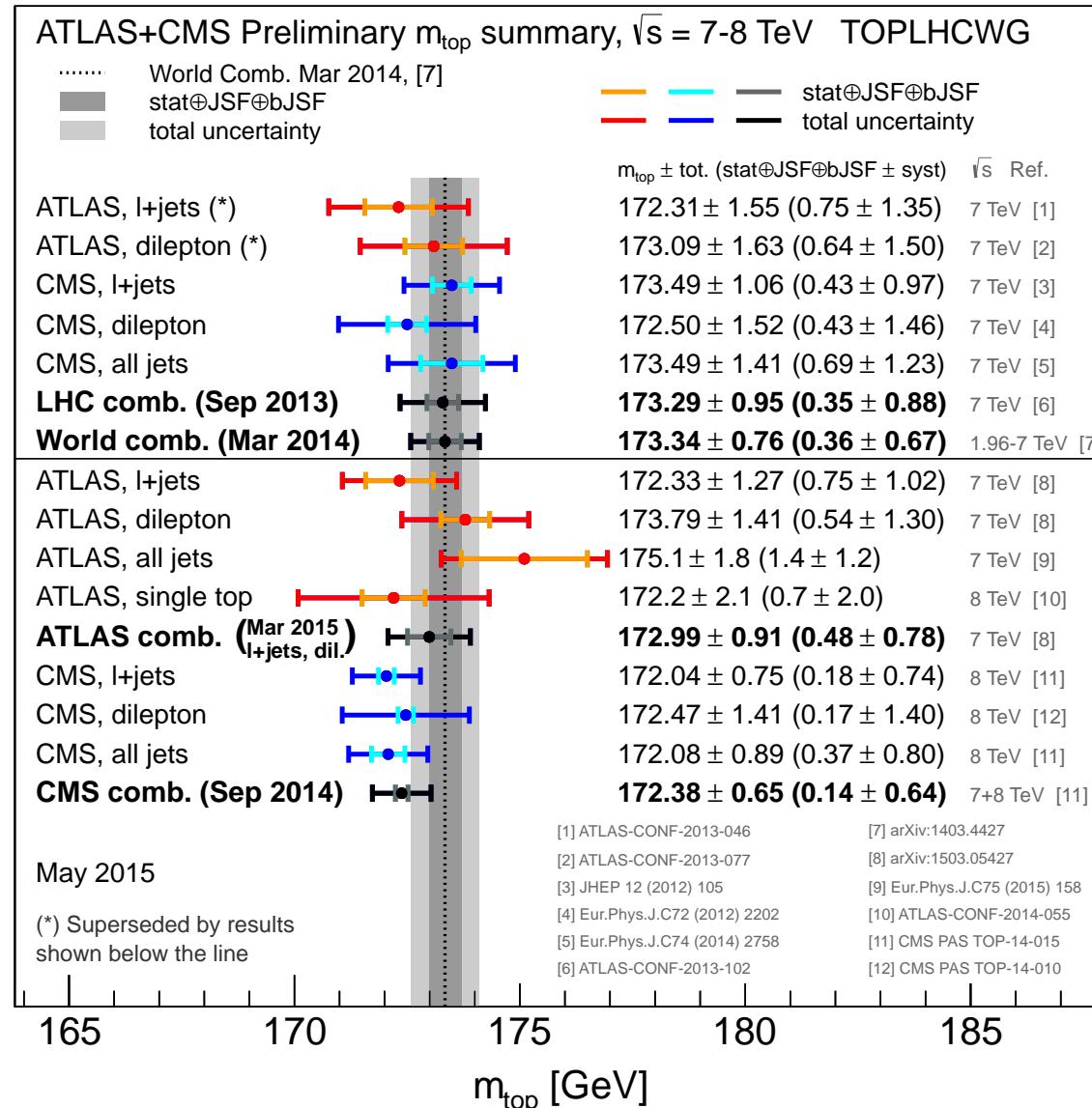
<sup>1</sup>Joint Institute for Nuclear Research, 141980 Dubna, Russia

<sup>2</sup>II. Institut für Theoretische Physik, Universität Hamburg, Luruper Chaussee 149, 22761 Hamburg, Germany

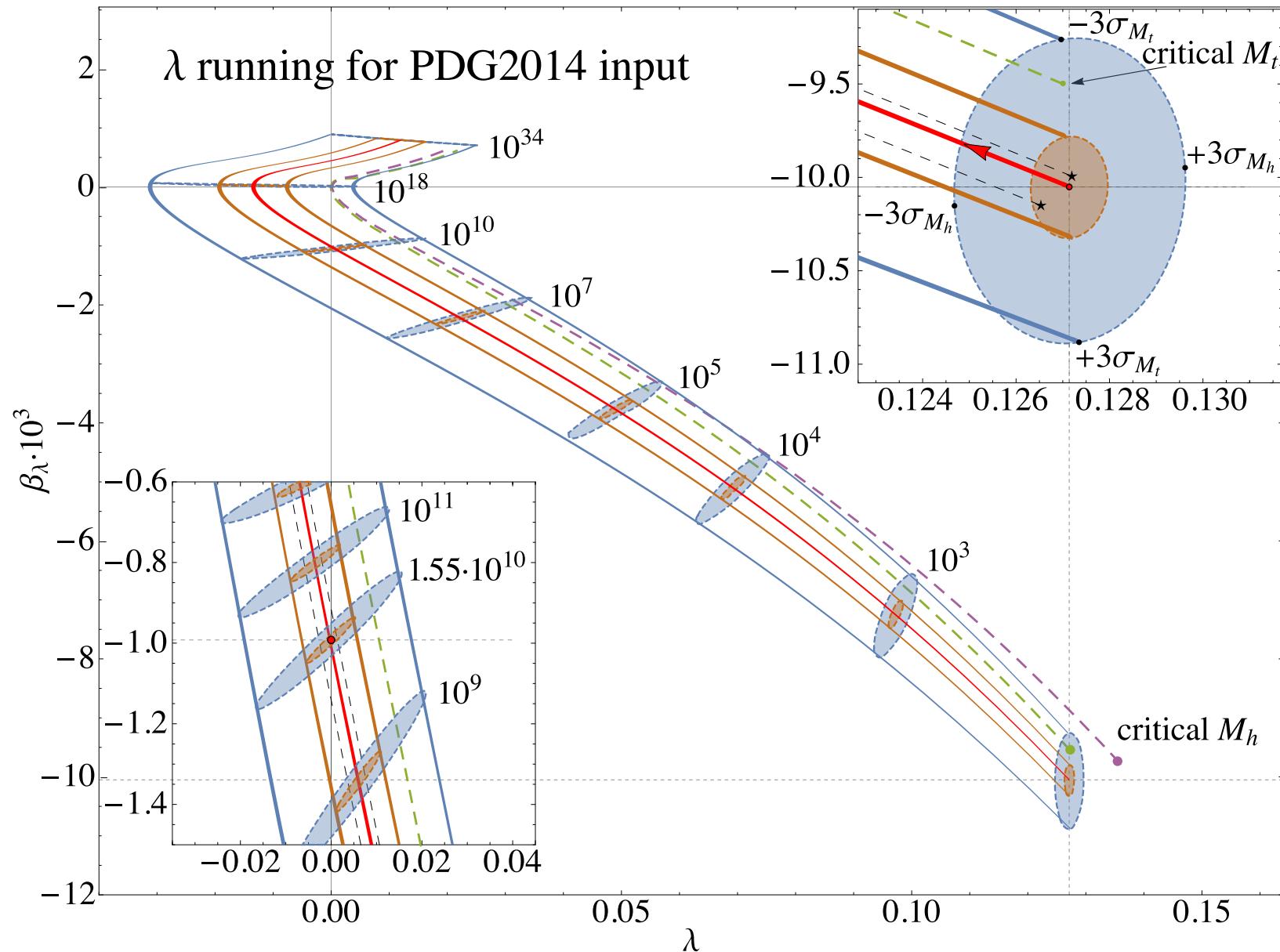
(Received 30 July 2015; revised manuscript received 24 August 2015; published 9 November 2015)

- From  $\lambda(\mu)$ :  $M_t^{\text{cri}} = (171.44 \pm 0.30)^{+0.17}_{-0.36}$  GeV
- From  $V_{\text{eff}}(H)$ :  $\tilde{M}_t^{\text{cri}} = (171.64 \pm 0.30)^{+0.17}_{-0.36}$  GeV
- Combination:  $\hat{M}_t^{\text{cri}} = (171.54 \pm 0.30)^{+0.26}_{-0.41}$  GeV
- Experiment:  $M_t^{\text{MC}} = (172.38 \pm 0.66)$  GeV ATLAS & CMS,  
[arXiv:1512.02244 \[hep-ex\]](https://arxiv.org/abs/1512.02244)

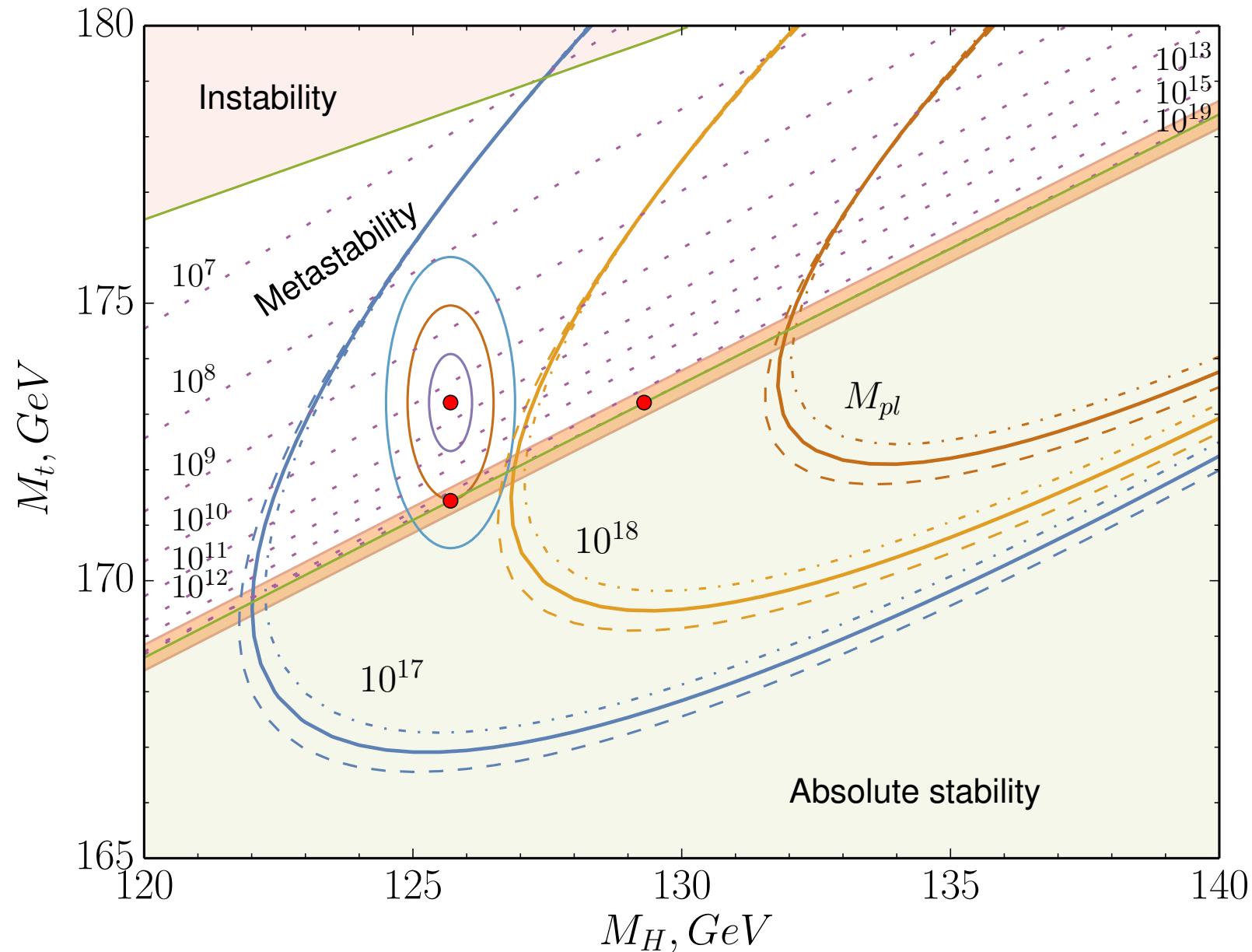
# $M_t$ Measurements



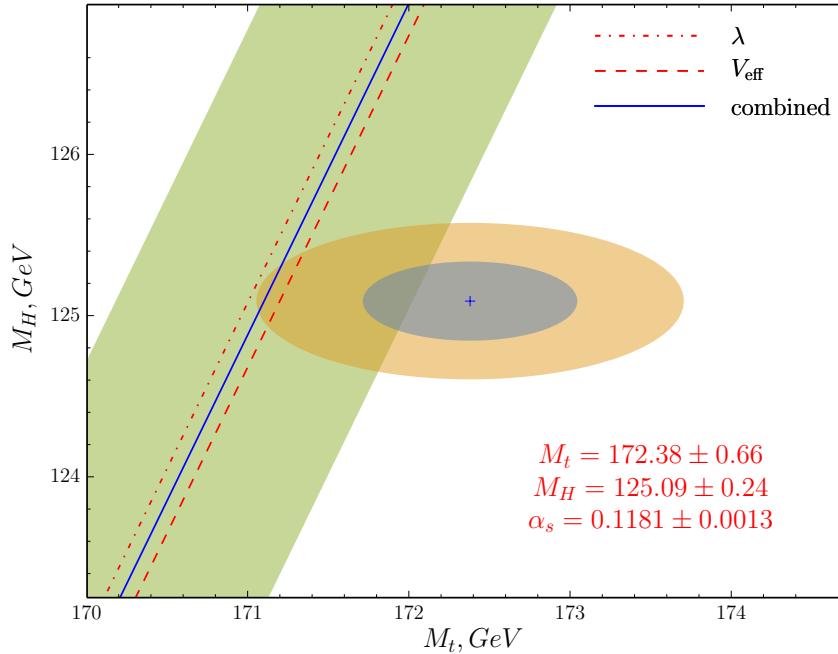
# RG flow



# Phase diagram



# SM stable all the way up to $M_P$ ?



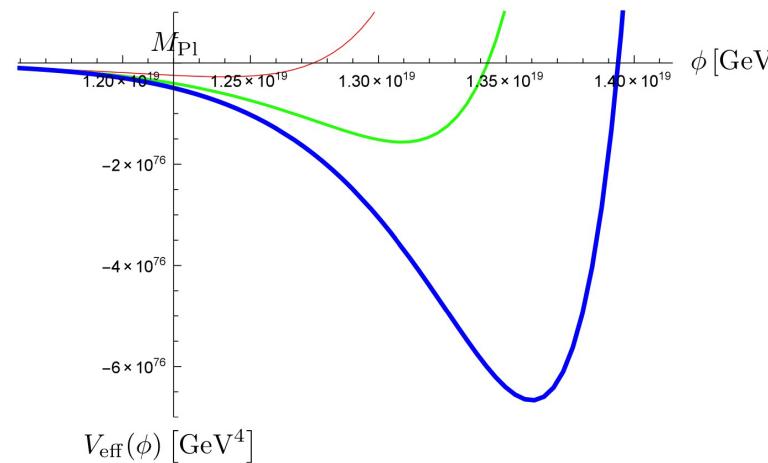
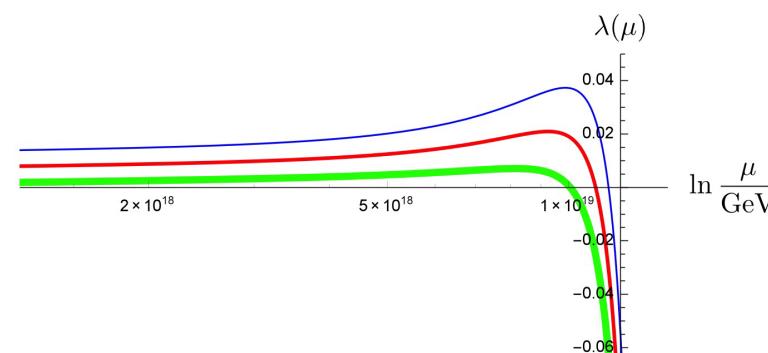
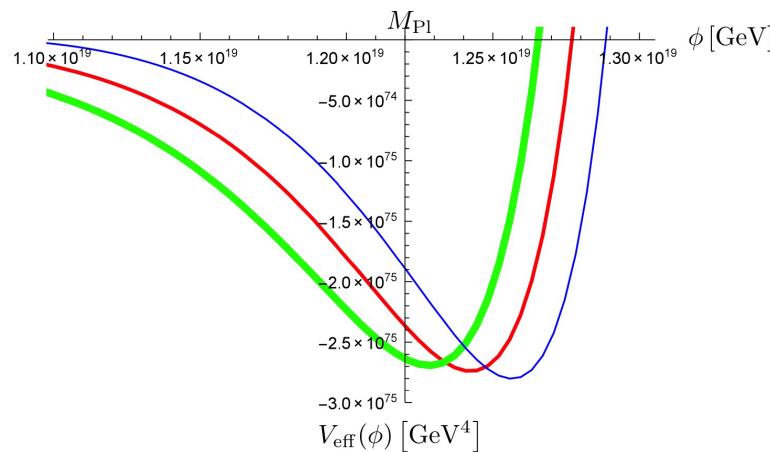
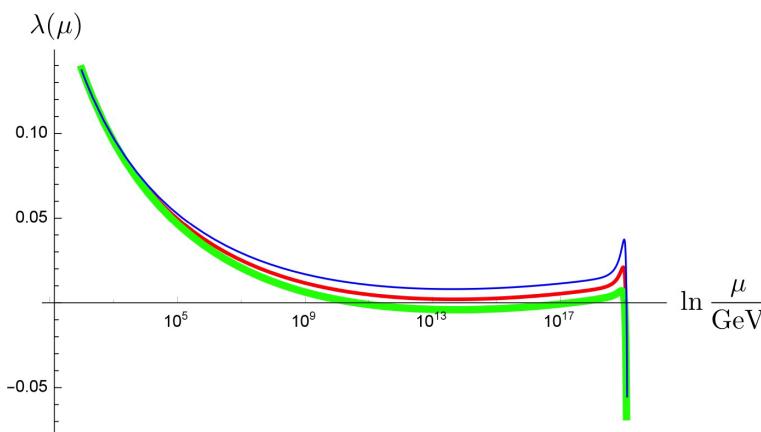
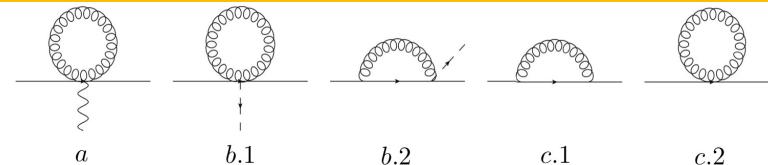
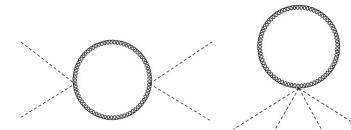
- Intriguing conspiracy of SM particle masses ↗  
 $\mu^{cri} \approx M_P = 1.22 \times 10^{18} \text{ GeV}$
- $\mu^{cri}$  stable w.r.t. parametric and higher-order uncertainties due to **asymptotic safety**
- Relationship between  $M_P$  and SM parameters?
- Electroweak scale determined by Planck scale physics?
- Implicit reduction of fundamental couplings?

# Einstein gravity loop corrections

A) 2-point functions



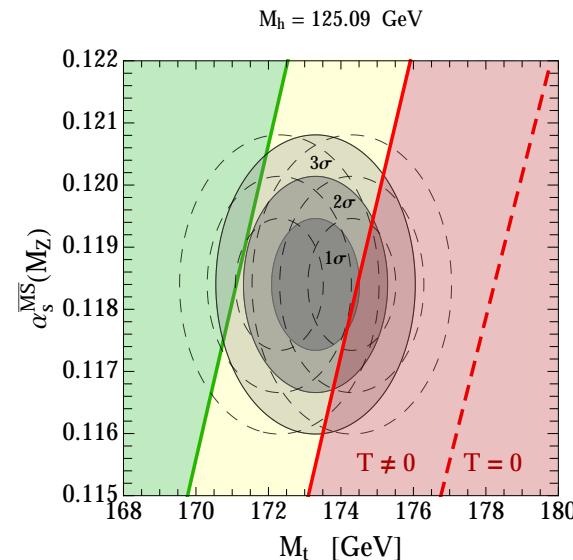
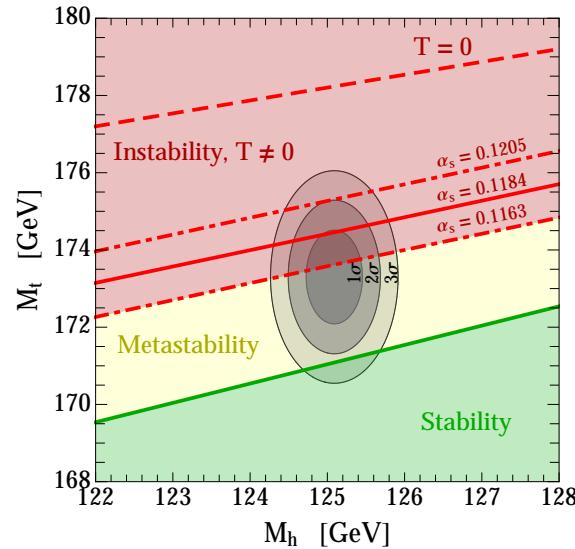
B) 4-point functions



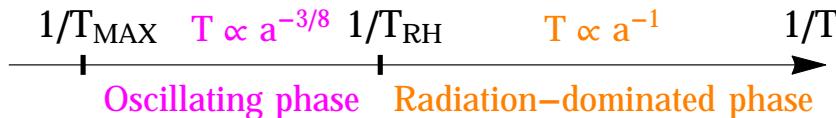
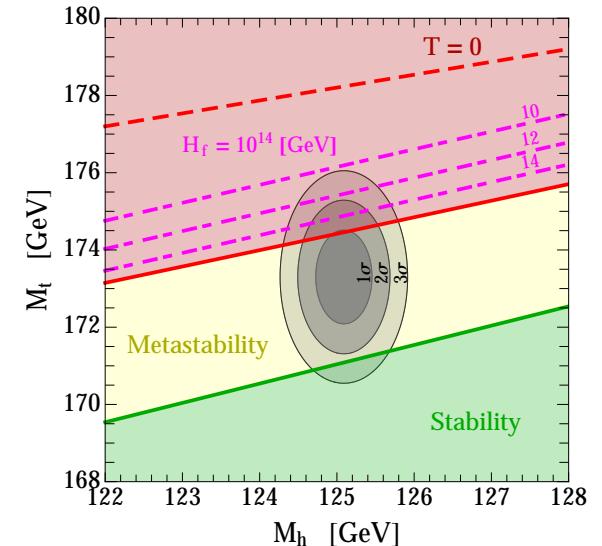
Quantum gravity corrections stabilize  $V_{\text{eff}}$ ! [Abe et al., 1602.03792 \[hep-ph\]](#)

# Thermal corrections

Cut-off scale  $\Lambda = 10^{19}$  GeV, i.e.  $T_{\text{cut}} \approx 10^{18}$  GeV



$T_{\text{RH}}$  dependence

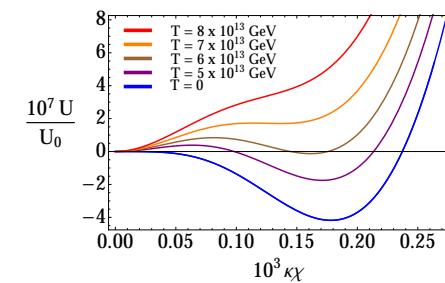
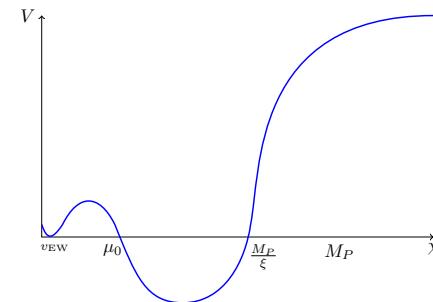
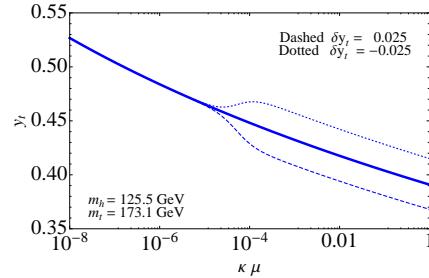
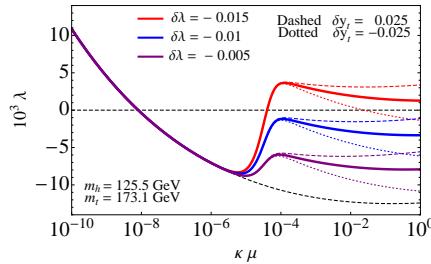


Finite- $T$  corrections increase tunneling probability and expand unstable phase!

Delle Rose *et al.*, 1507.06912 [hep-ph]

# Higgs inflaton

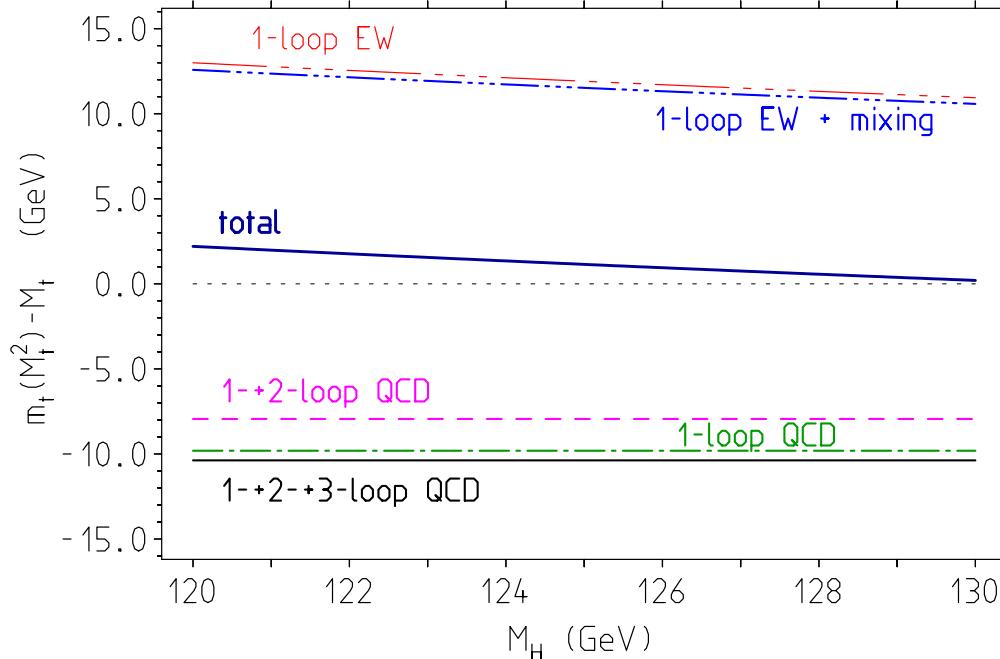
- Higgs field, nonminimally coupled to gravity with strength  $\xi$ , can be responsible for inflation
- Successful scenario possible even if EW vacuum is metastable
- Effective renormalization of SM couplings at scale  $M_P/\xi$
- Symmetry restoration after inflation due to high- $T$  effects temporarily eliminating vacuum at  $H \approx M_P$



Bezrukov *et al.*, PRD92(2015)083512

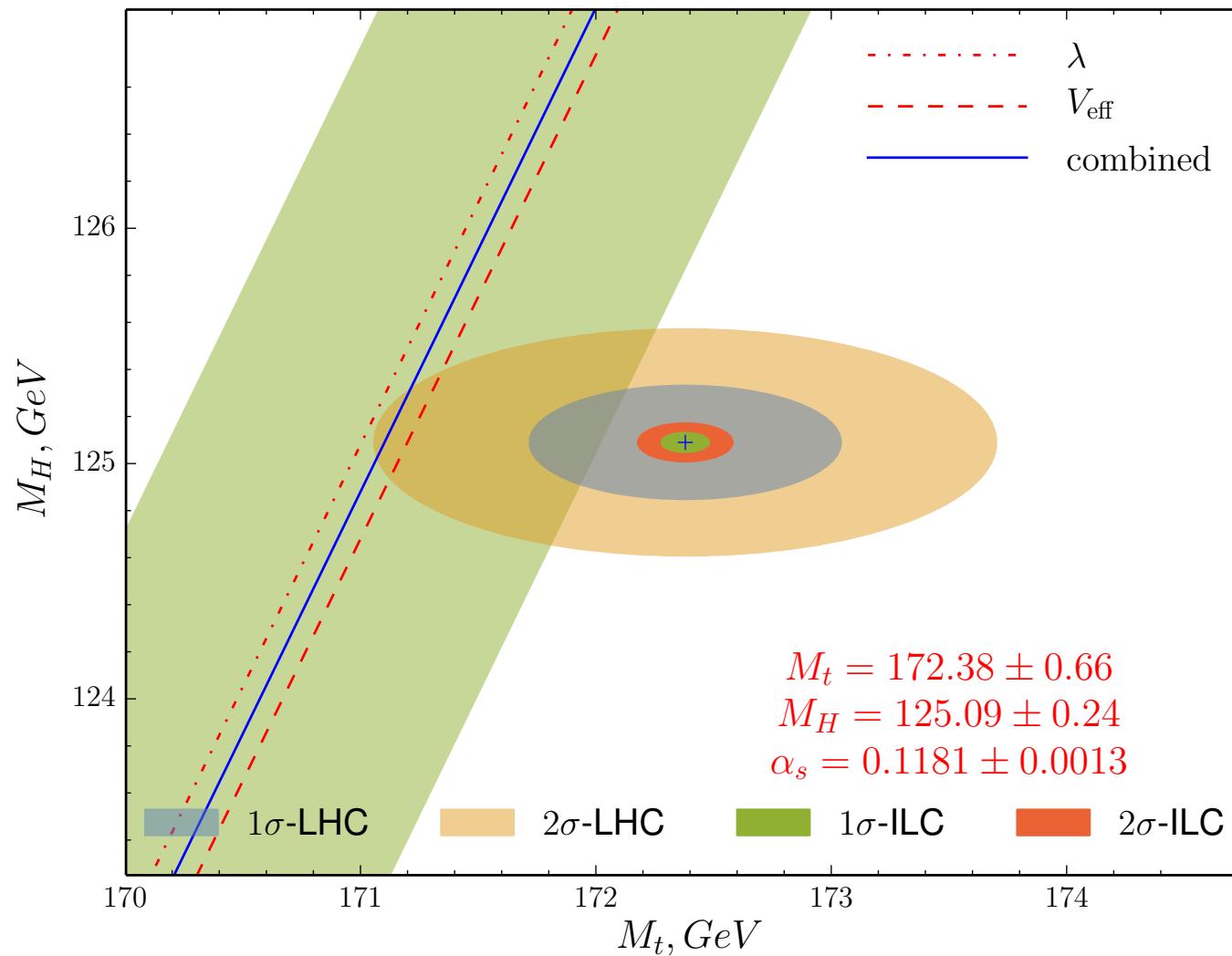
# Outlook: pole mass $M_t$

- PDG value  $M_X(t \rightarrow X) = (173.21 \pm 0.87)$  GeV is **not** pole mass  $M_t$ , but just parameter in MC programs w/o RC to partonic cross sections.
- Rigorous determination of  $\overline{\text{MS}}$  mass  $m_t(\mu)$  from  $\sigma_{\text{tot}}(p\bar{p}, pp \rightarrow t\bar{t} + X)$
- $m_t(\mu) - M_t$  receives large EW RC from tadpole contributions.



Jegerlehner, Kalmykov, BK, PLB722(2013)123

# ILC as top and Higgs factory



Anticipated errors  $\delta M_t = 100 \text{ MeV}$ ,  $\delta M_H = 40 \text{ MeV}$  Moortgat-Pick *et al.*,  
EPJC75(2015)371

# BSM physics

- Depending on future precision measurements of  $M_H$ ,  $M_t$ ,  $\alpha_s$  and higher-loop RC calculations, SM may be stable all the way up to  $M_P$ .
- Reduction of  $M_t \approx M_t^{\text{MC}} = 172.38 \text{ GeV}$  by  $0.84 \text{ GeV} \rightsquigarrow M_H^{\text{cri}} = M_H = 125.09 \text{ GeV}$
- Cf.  $m_t - m_{\bar{t}} = (-0.6 \pm 0.6) \text{ GeV}$ ,  $\Gamma_t = (2.0 \pm 0.5) \text{ GeV}$
- BSM physics still necessary to solve open problems, e.g.
  - smallness of neutrino masses
  - strong CP problem
  - dark matter
  - baryon asymmetry of universe
  - unification with gravity
- Higgs portals?