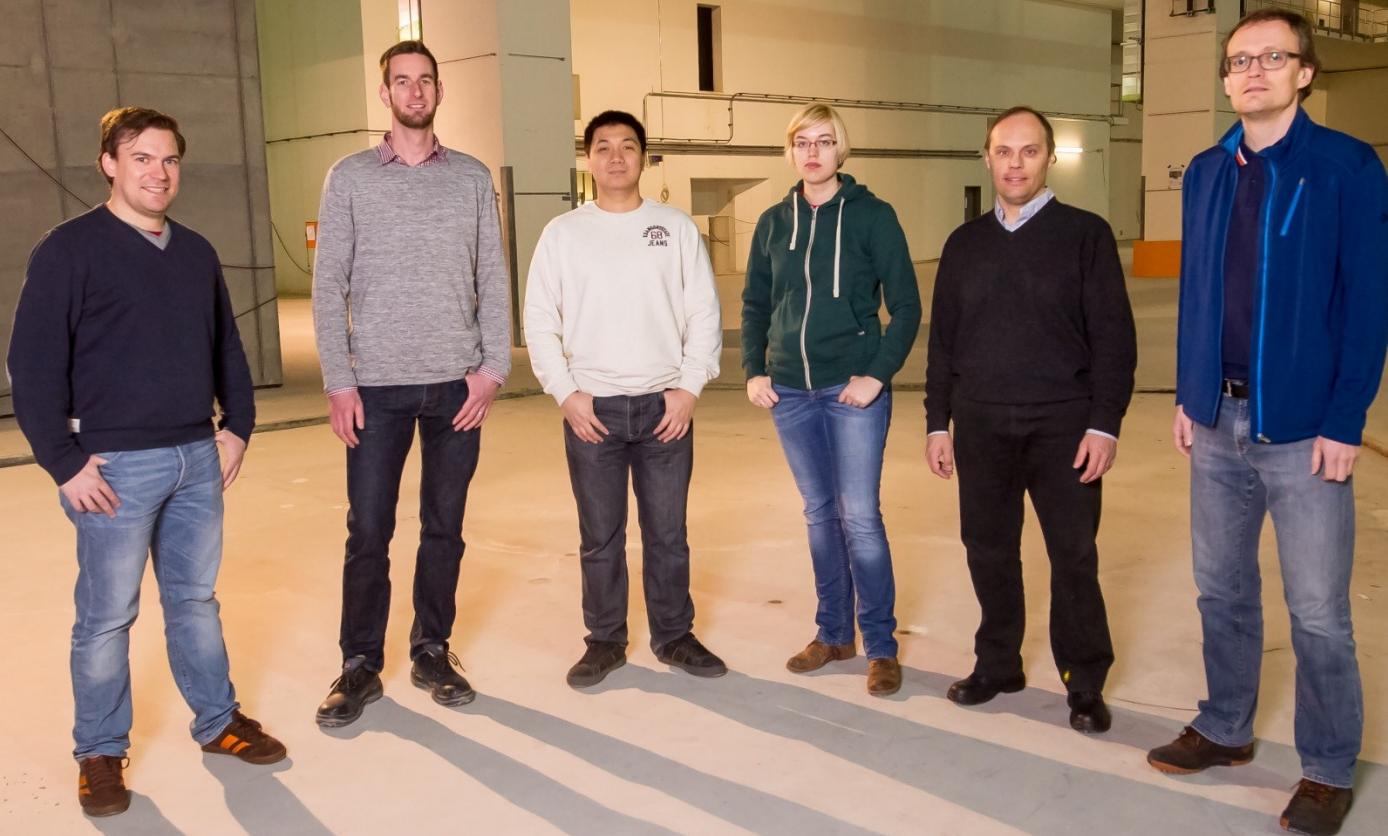




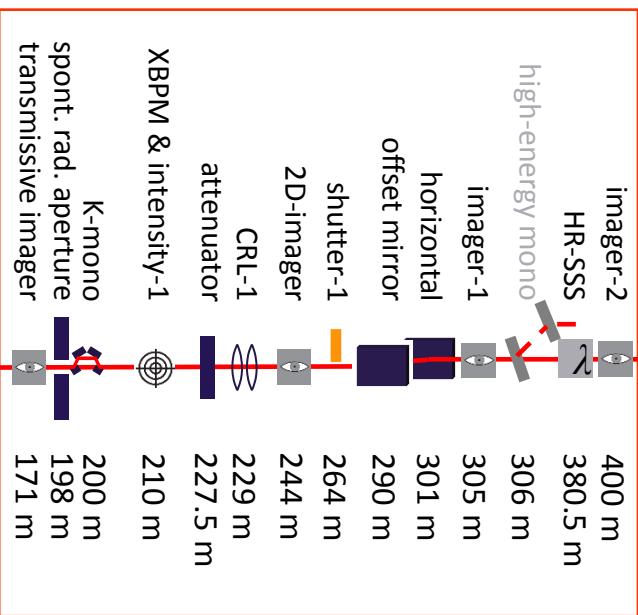
First experiments and commissioning of the MID instrument (SASE-2)

A. Madsen,
European XFEL
anders.madsen@xfel.eu

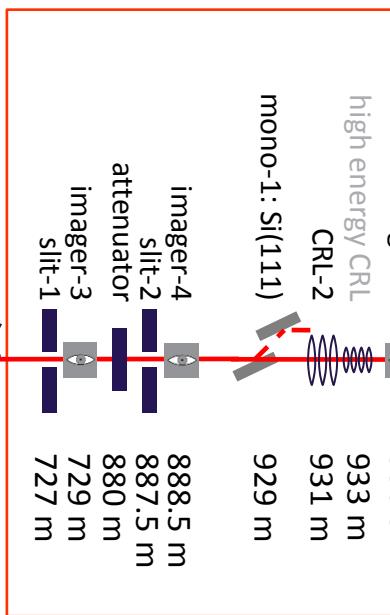


MID beamline overview

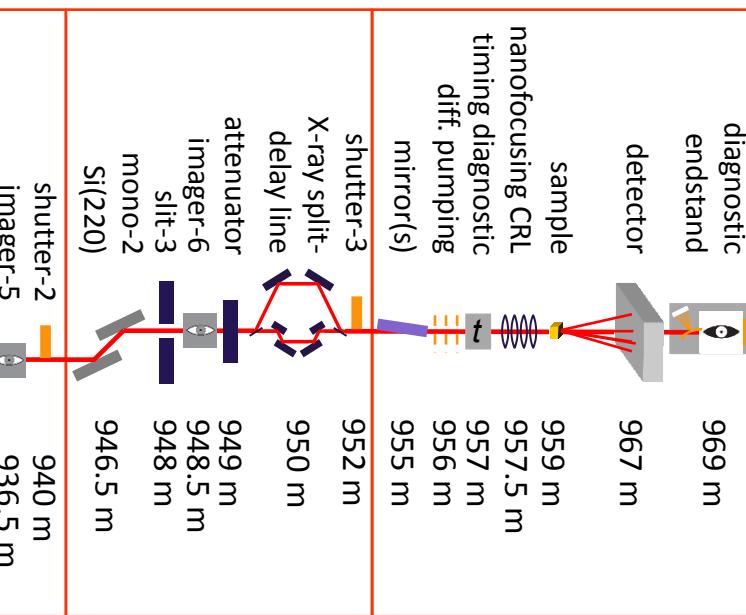
common SASE-2 beamline (MID/HED)



MID photon beamline



MID optics hutch



Not shown:

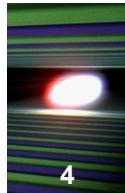
MCP at 303m (fine tuning of SASE)

Distribution mirror(s) at 390m and 395m (MID on central branch)

Beam loss monitors, PES

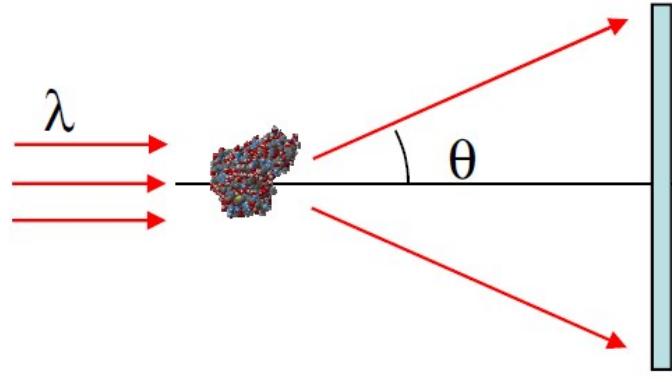
last 25 m in experimental hall

Coherent scattering and speckle

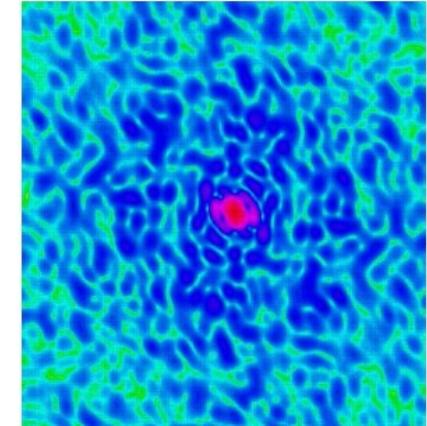


Diffraction microscopy:
Phase retrieval required
but no limiting optics

Isolated object

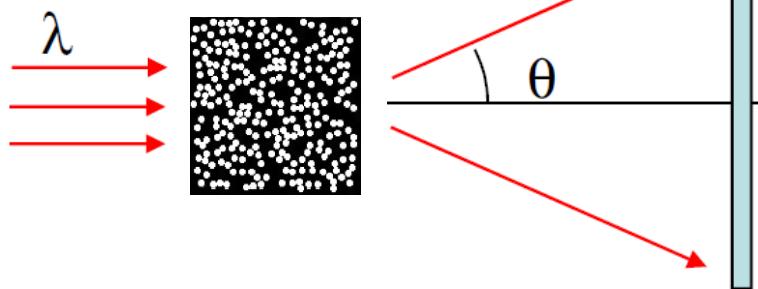


Static speckle

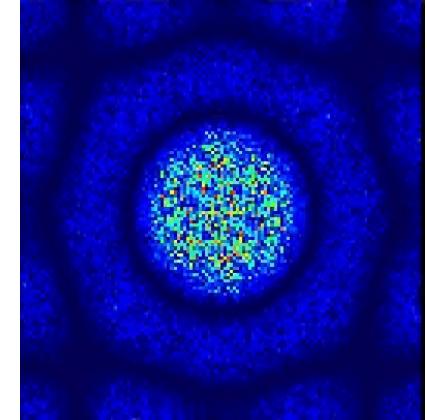


Correlation spectroscopy:
Temporal: XPCS
Spatial: XCCA

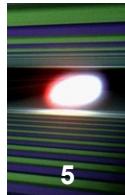
Ensemble of objects



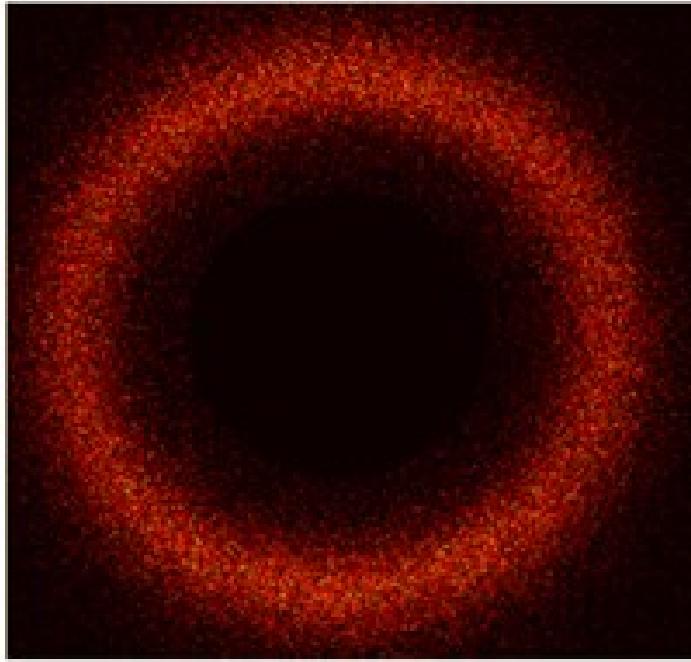
Dynamic speckle



Possible 1st experiment



High-Q coherent scattering from a glass former



Vitreous
 GeO_2

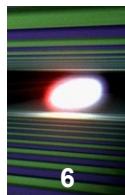
1st peak in $S(Q)$
at 15.5 nm^{-1}

Application: Study of structural dynamics in the supercooled liquid state

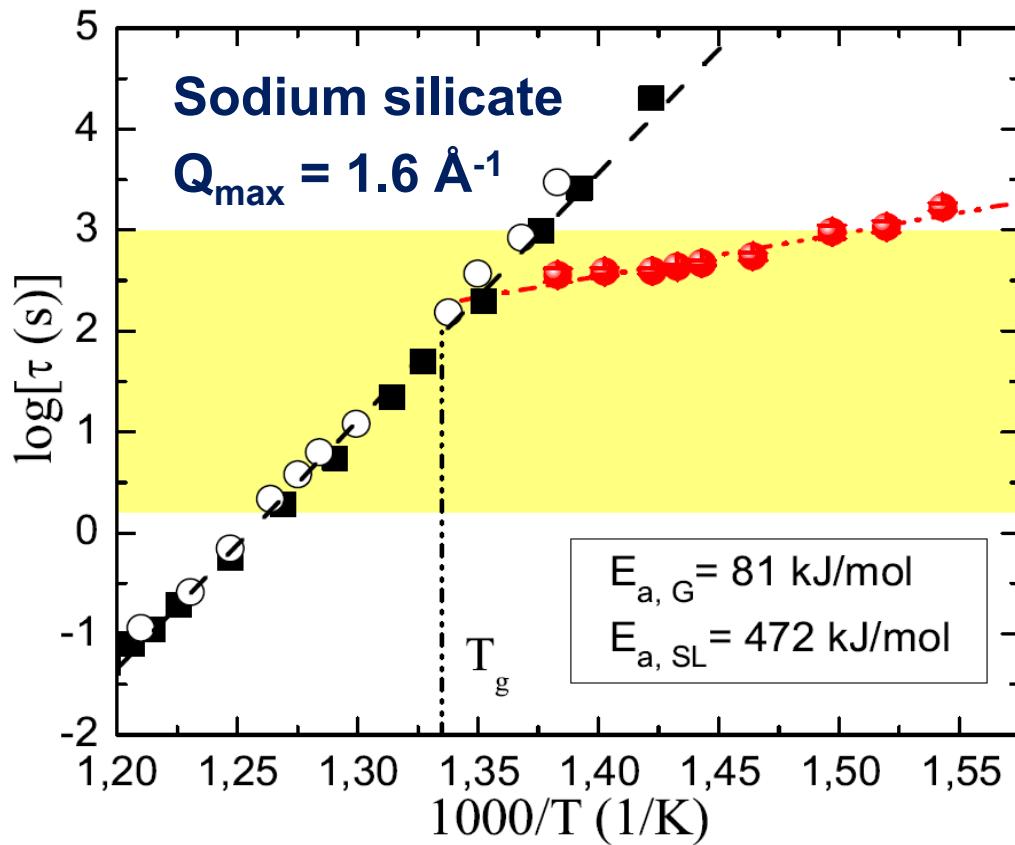
Techniques: X-ray Photon Correlation Spectroscopy (XPCS),

Time-resolved X-ray Cross Correlation Analysis (XCCA)

Possible 1st experiment

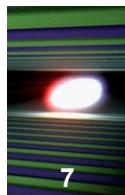


The deepest and most interesting unsolved problem in solid state theory is probably the nature of glass and the glass transition", Philip W. Anderson: "Viewpoint: The Future," Science, vol. 267, pp. 1615-1616 (1995)



- 1) What causes the dramatic increase of viscosity (~13 orders of magnitude) over 2 orders of magnitude in T, without any apparent change of structure?
- 2) Why is the structural dynamics only tracking the viscosity in a certain range of temperatures (or timescales)?
- 3) Can we understand the length-scale (Q) dependence of 2)

Assumed parameters for first experiment

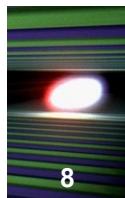


Electron energy	17.5 GeV
Photon energy	8.4 keV
Repetition rate	100 kHz (10 µs)
Max. number pulses per train	60
Undulator K-value	3.9
Undulator Gap	10 mm
Pulse energy	2 mJ (slightly oversaturated)
Divergence	2.2 urad
Pulse duration	43 fs
Saturation length	58 m

1 mJ ~ 1e12 ph @ 7 keV

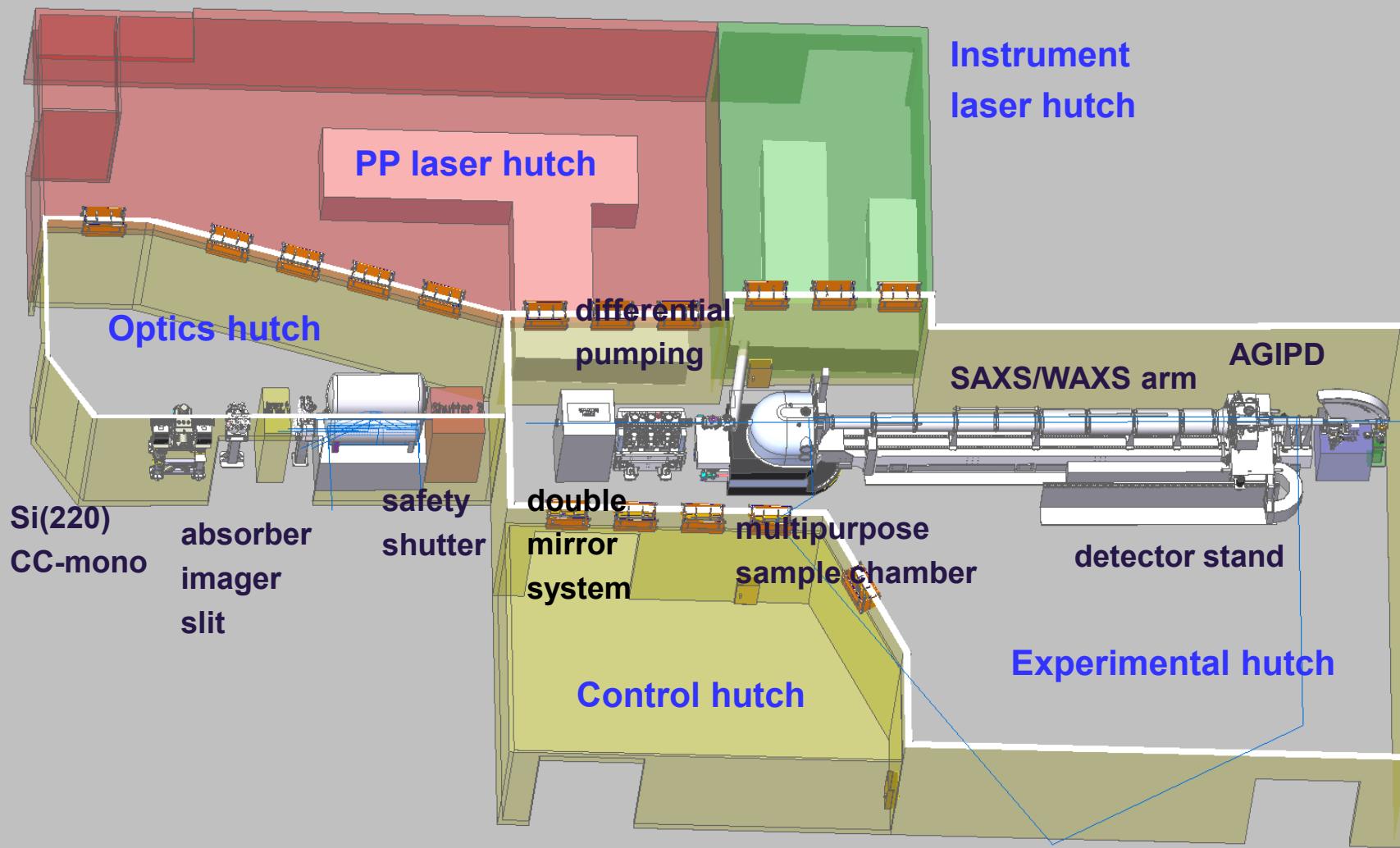
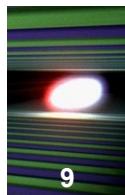
Courtesy of H. Sinn

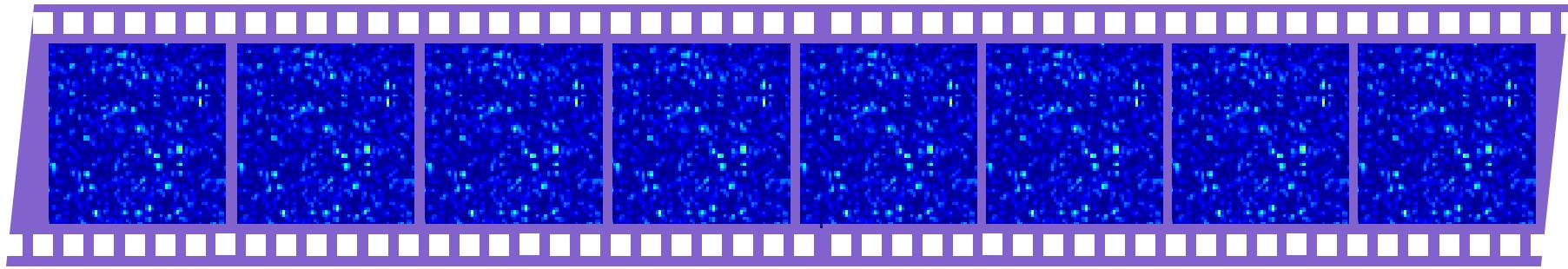
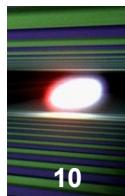
For the GeO₂ glass experiment we need



- 10¹⁰ or more photons per pulse
- 1e-4 bandwidth or less (SASE is only ~1e-3 so mono and/or seeding is needed)
- As high repetition rate as possible (100 kHz is fine to start with but...)
- As many pulses/train as possible (AGIPD detector can store ~350 images/train)
- A focused beam, ~10 x 10 micron (refractive lens system)
- A furnace (up to 1200 K)
- A good photon detection system, 2θ up to 25° (AGIPD and SAXS/WAXS arm...)
- Slits, attenuators, diagnostics
- ...

MID Experimental area @ SASE-2

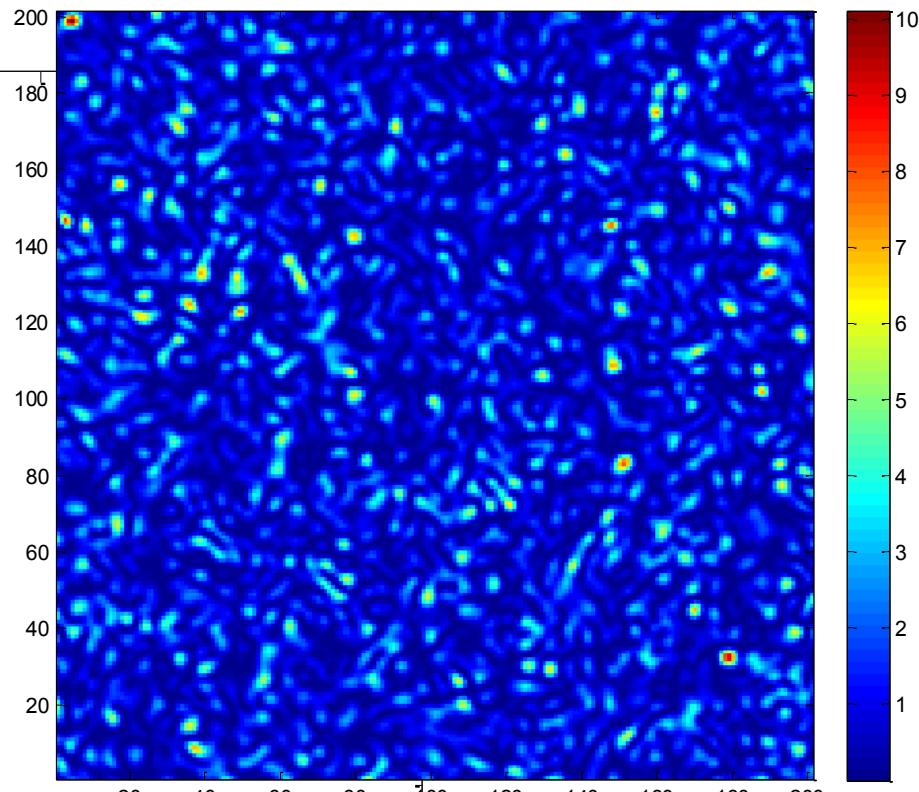
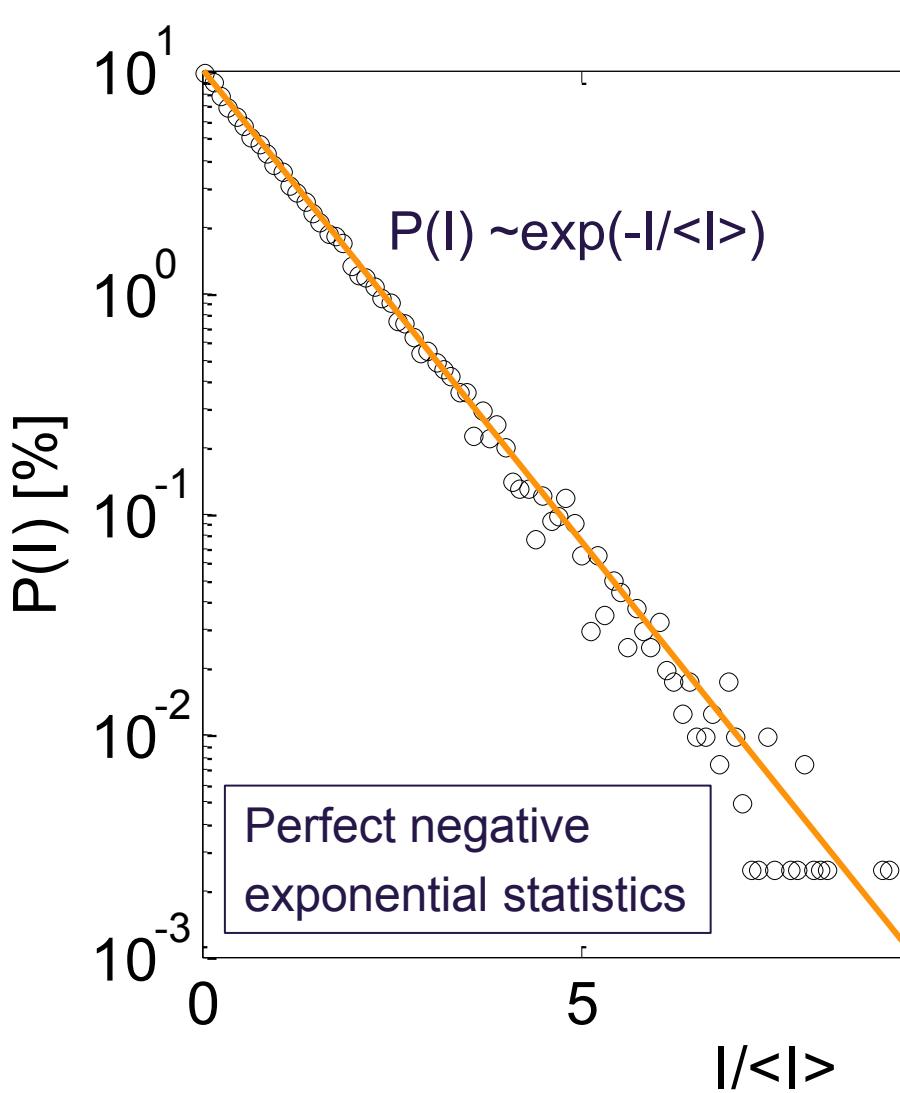




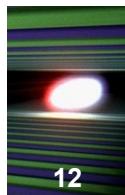
#1 #2 #3 #4 #5

- Recording movies of weak speckle patterns
- Average count rate: 1 ph/pixel/image, or less
- Up to 4.5 MHz and as many images per train as possible
- E range: 5 - 25 keV, typically 7-15 keV
- $\text{SNR} \propto 1/(\text{pixel area})$, smaller pixels are better
- $\text{SNR} \propto \text{number of pixels}$
- Important that image sequence (#1, #2, #3,...) is maintained
- Minimize image distortions....

100% contrast speckle pattern, mean Int = 1

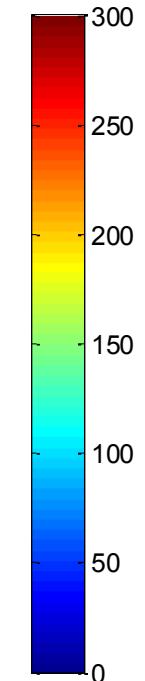
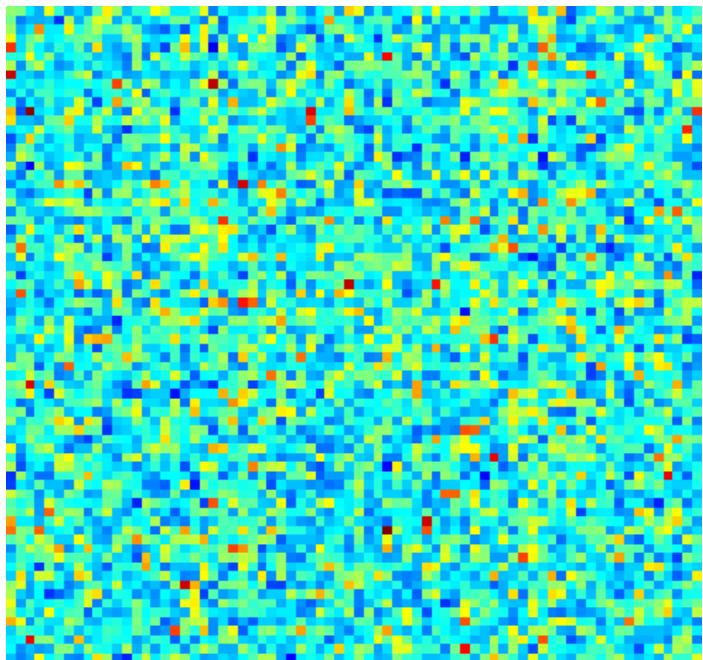


The effect of BIG pixels

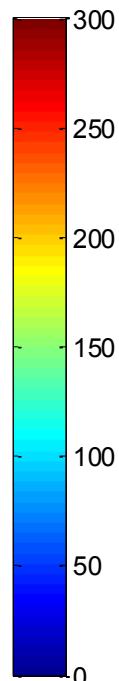
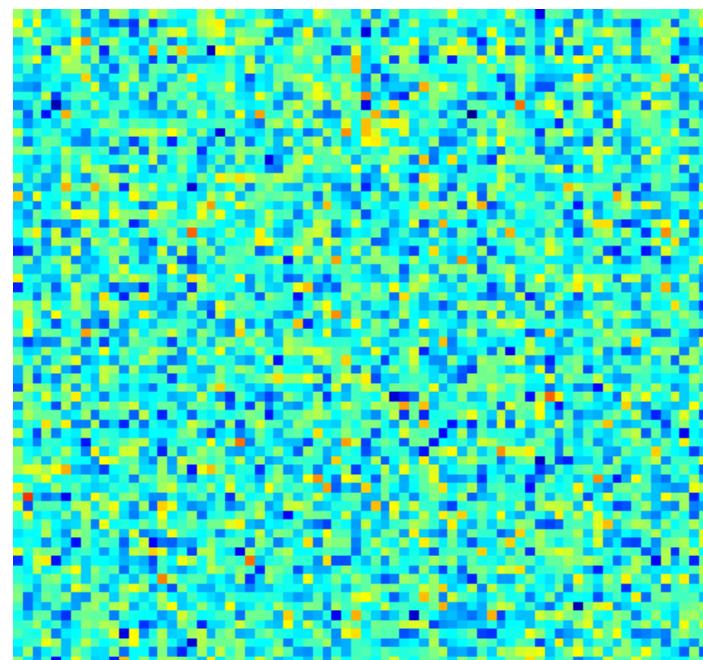


Challenge to distinguish between:

Undersampled speckle pattern
8% contrast (AGIPD case, 8m distance)

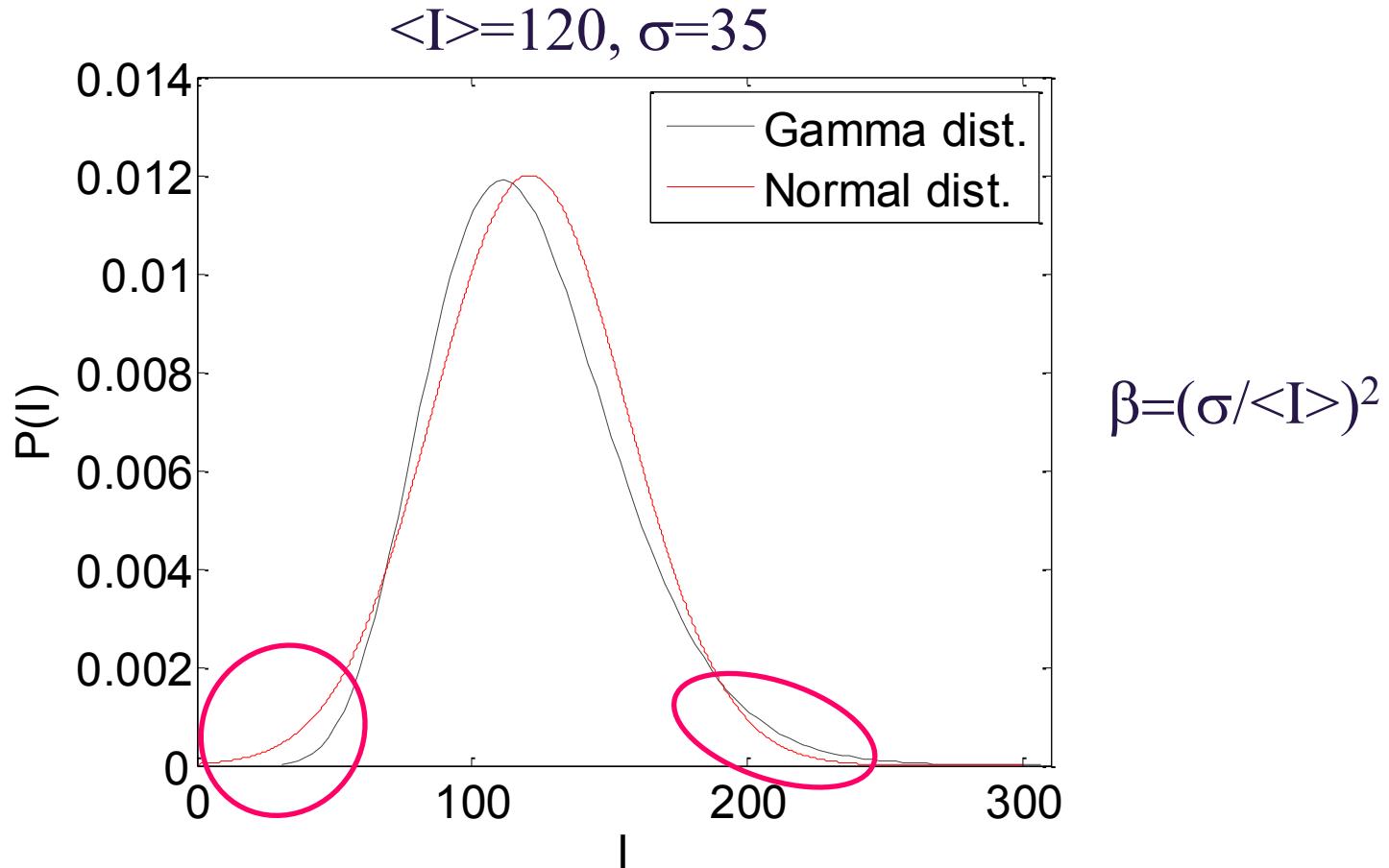
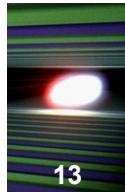


Random, normal-distributed
intensity pattern



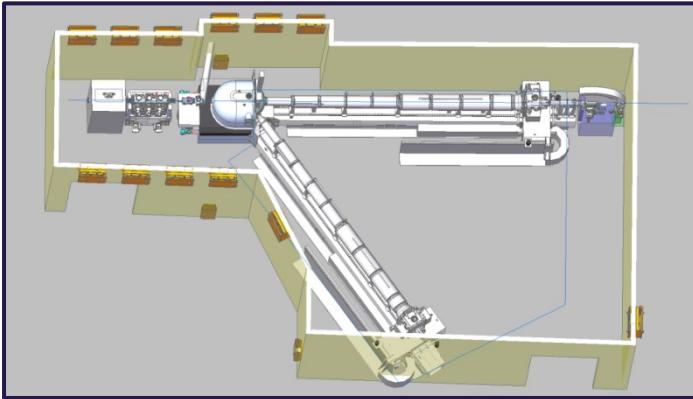
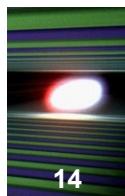
$$\langle I \rangle = 120, \sigma = 35$$

Photon Statistics of 8% Speckle Pattern

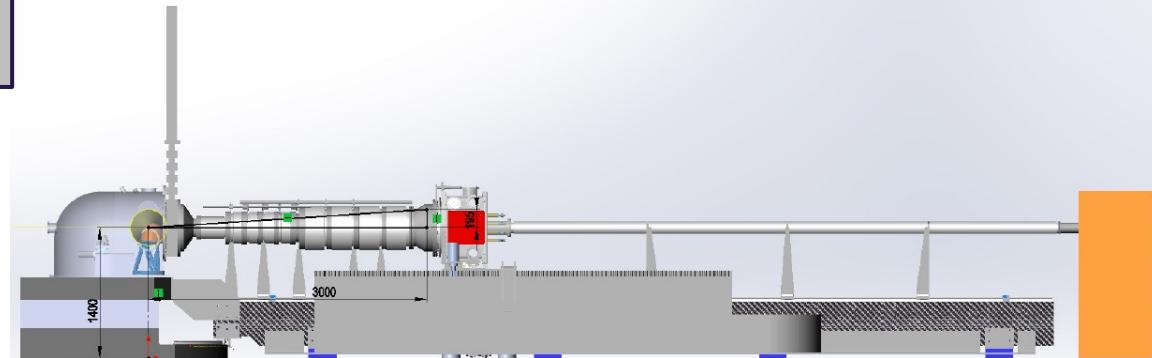


If detector isn't precise we cannot measure speckle patterns in low contrast conditions
Direct illumination CCDs can do it. Can AGIPD?

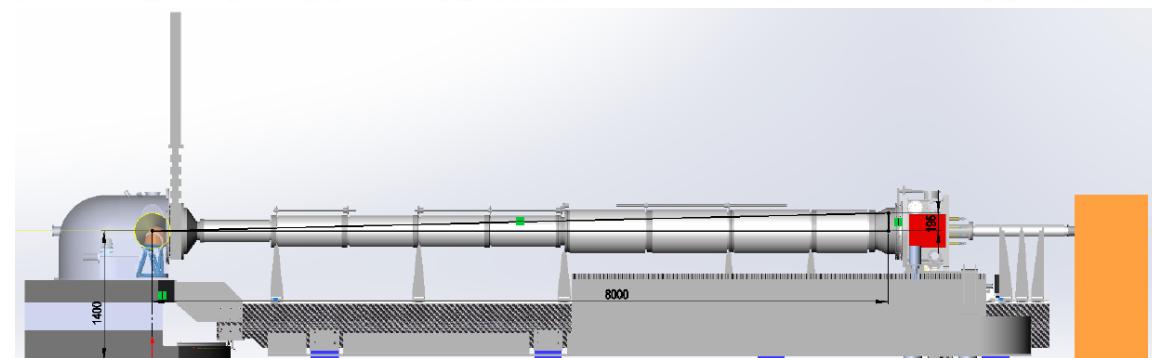
Long SAXS/WAXS arm



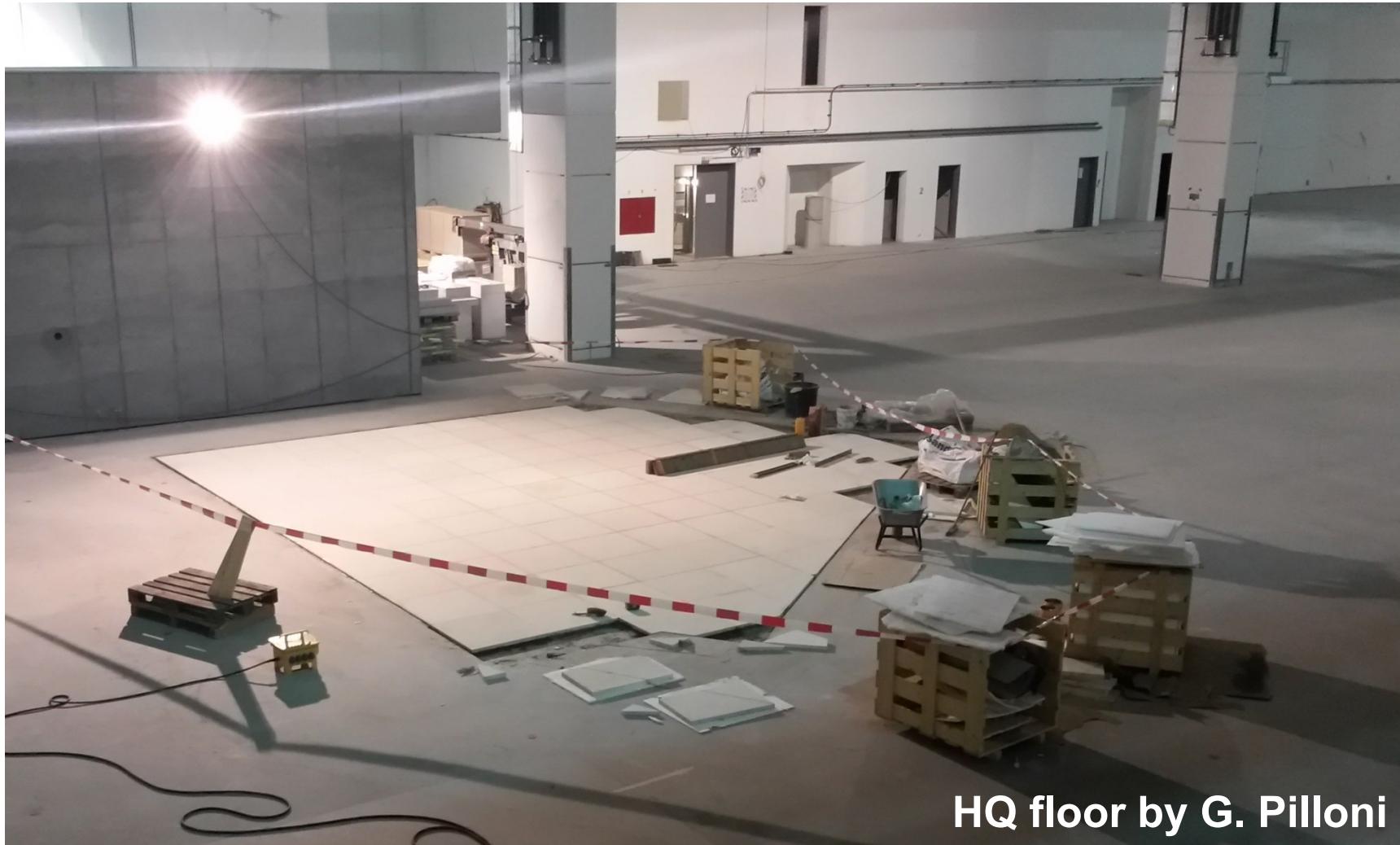
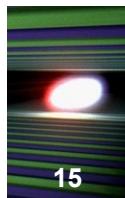
Under design in collaboration with
ESRFs engineering group



Air pads ensure the motion of
the arm on a high precision floor

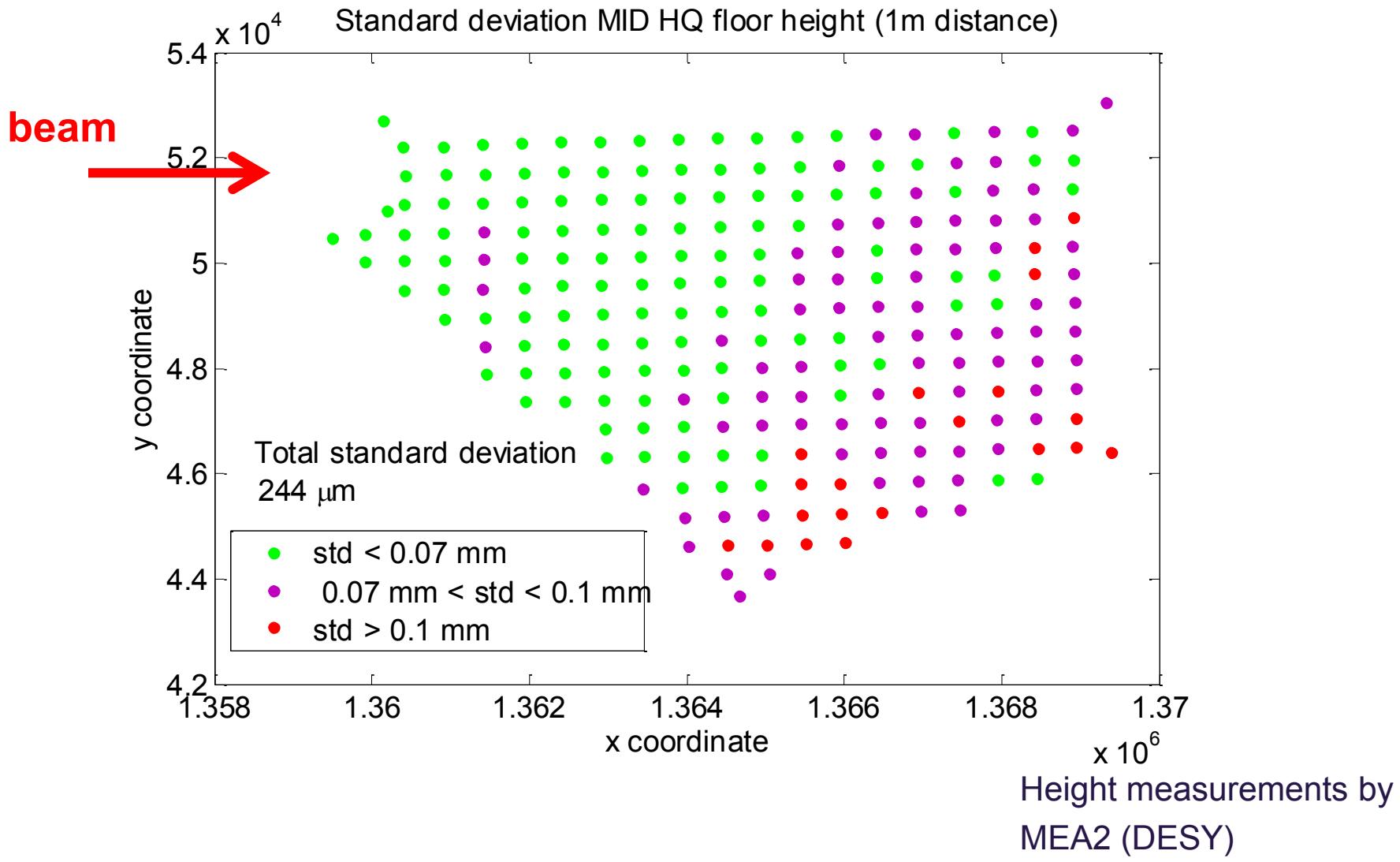
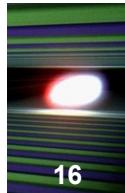


High precision floor is in a good state...

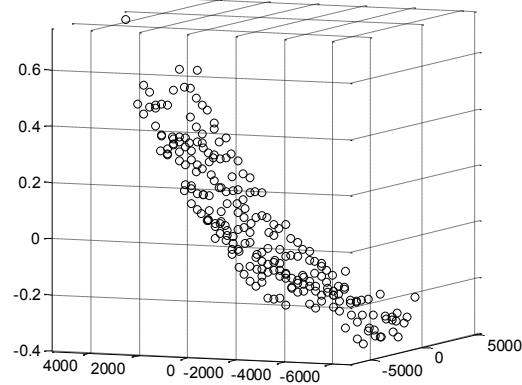
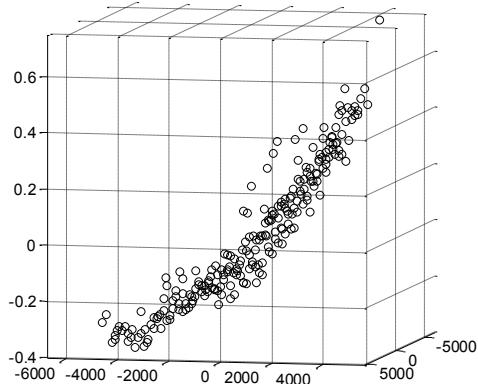
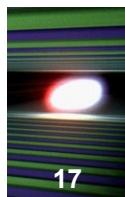


HQ floor by G. Pilloni

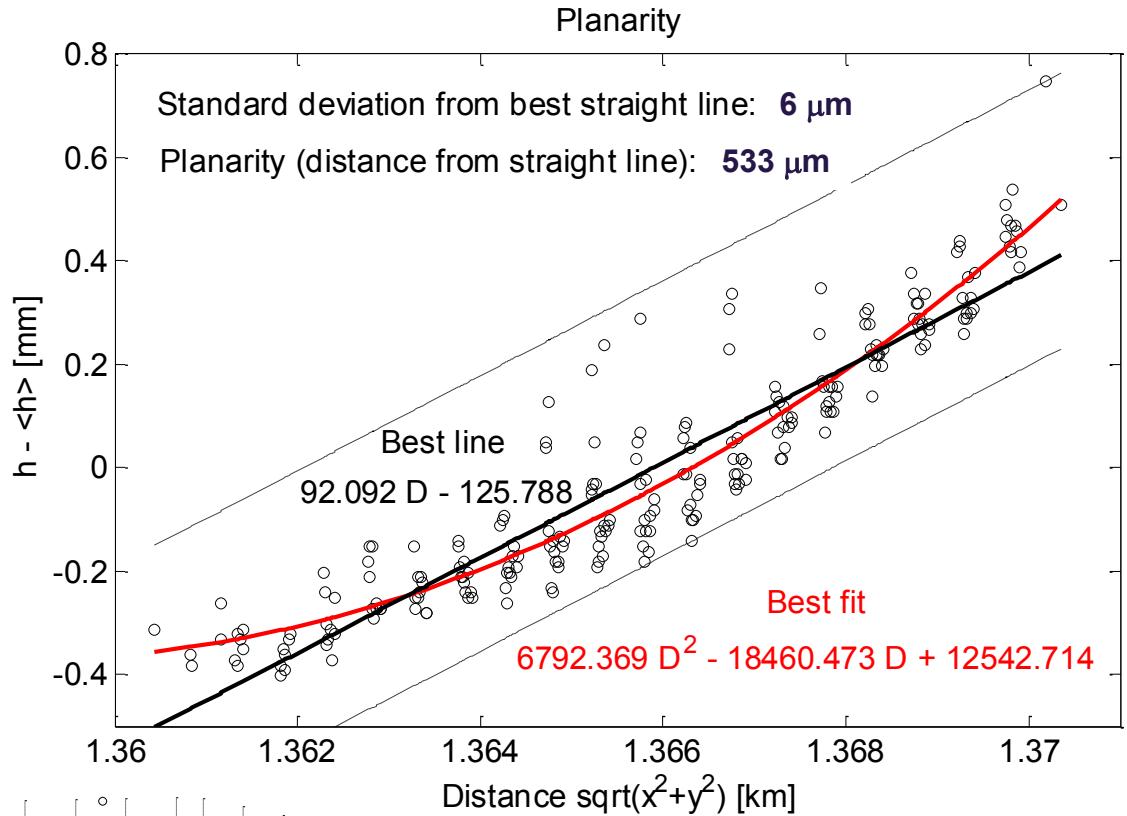
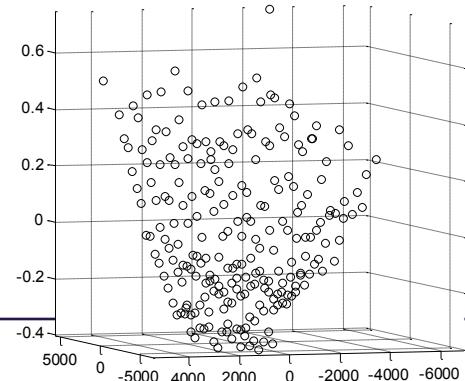
Floor quality



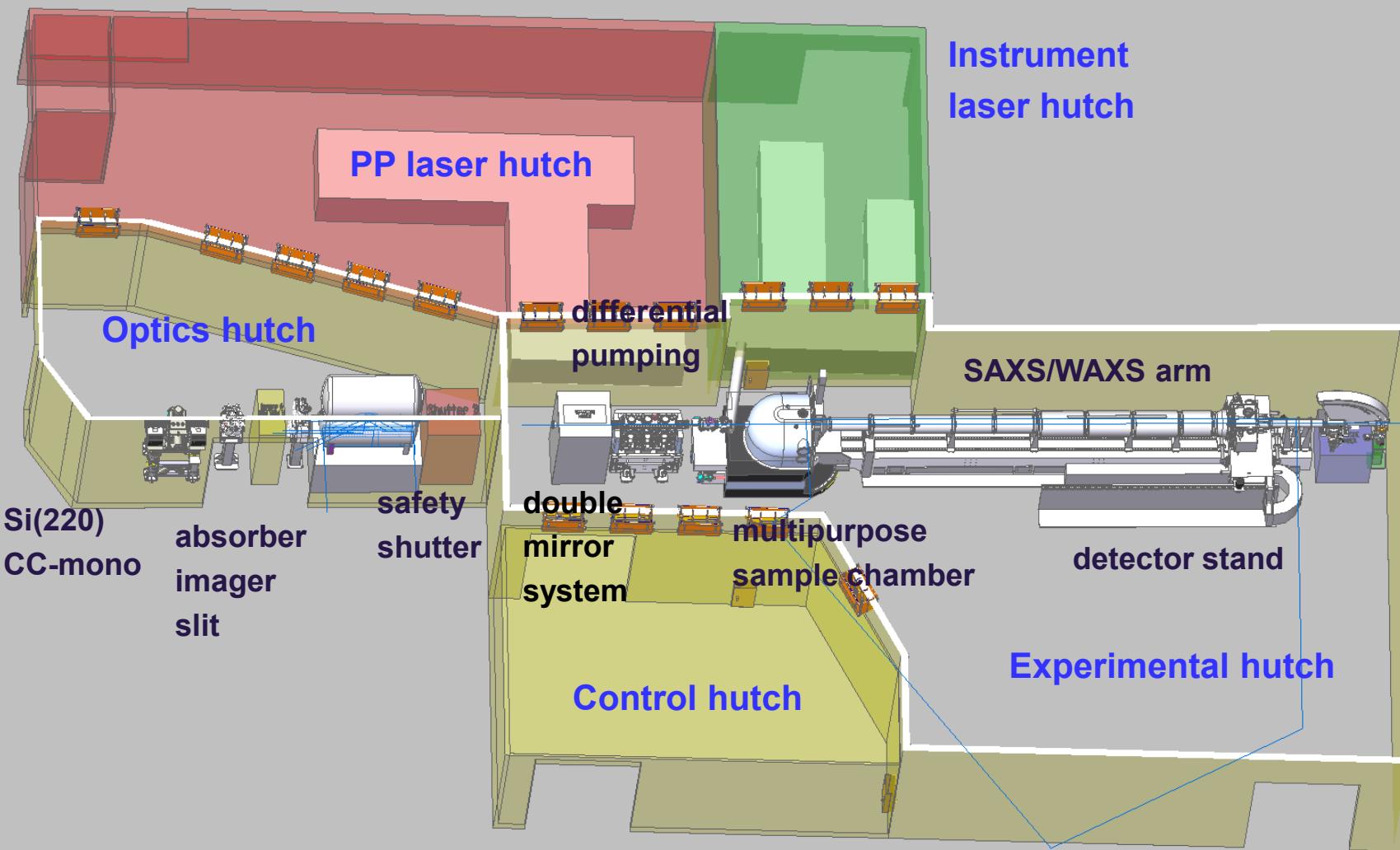
Floor quality



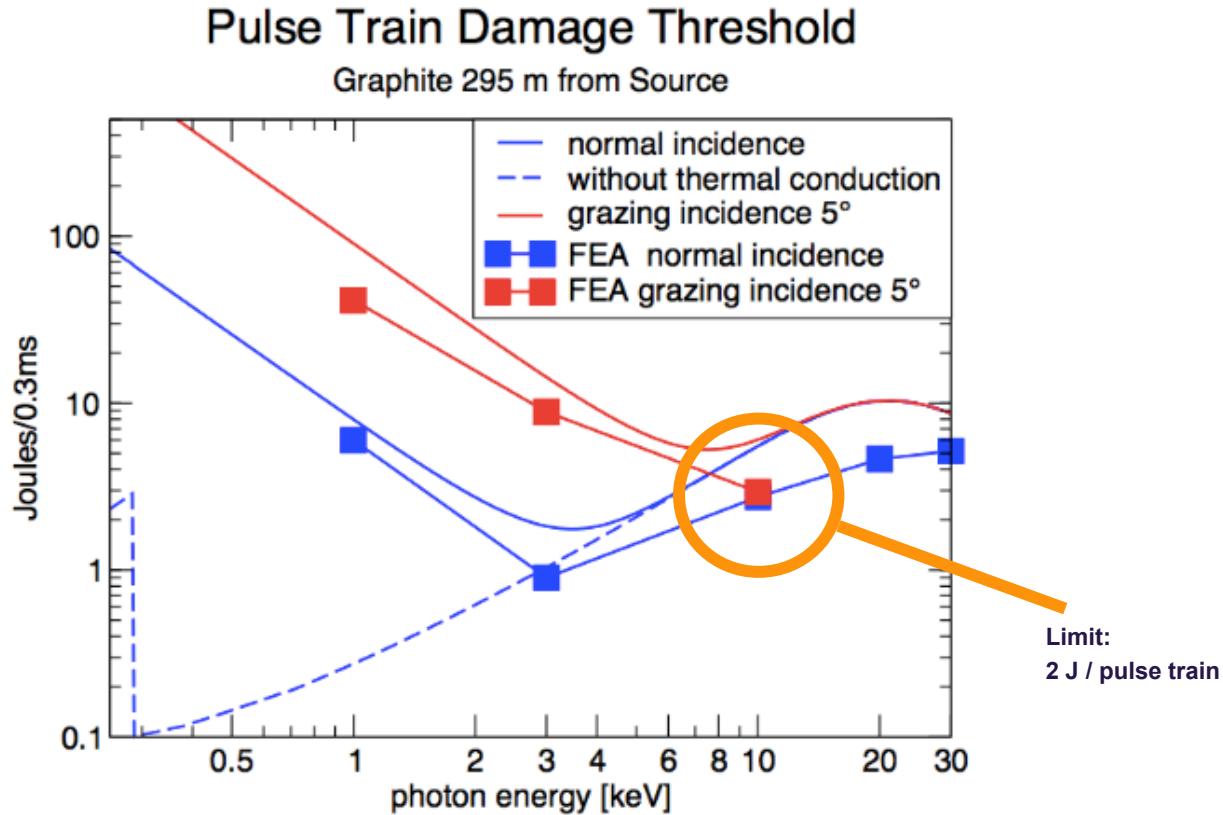
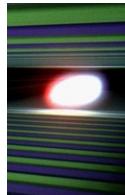
Height depends almost only
on distance along beam
(z coordinate).
Slope: 92 $\mu\text{m}/\text{m}$



Height measurements by
MEA2 (DESY)

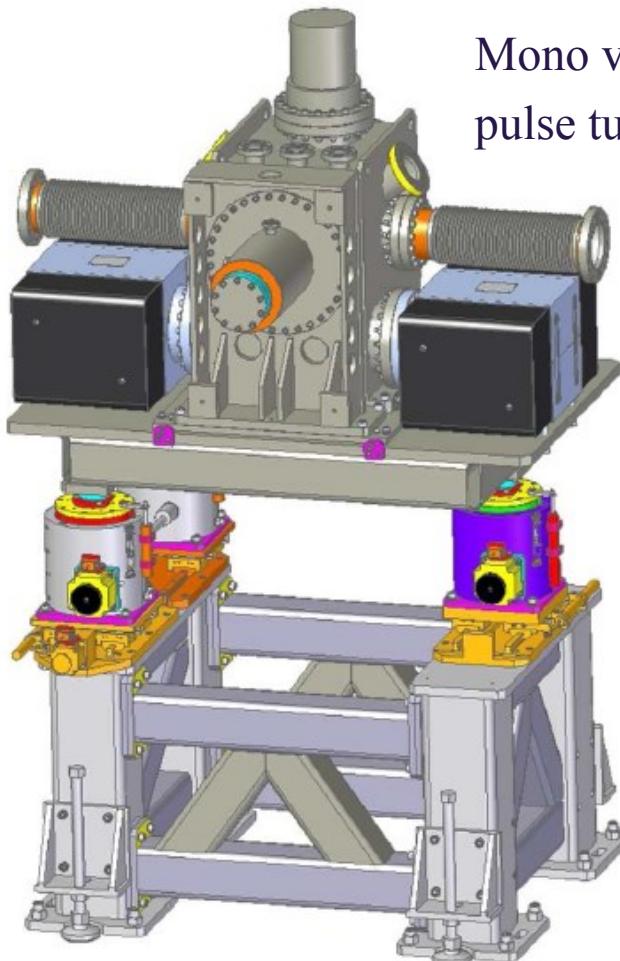
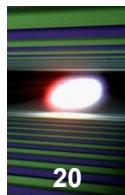


Damage on shutters

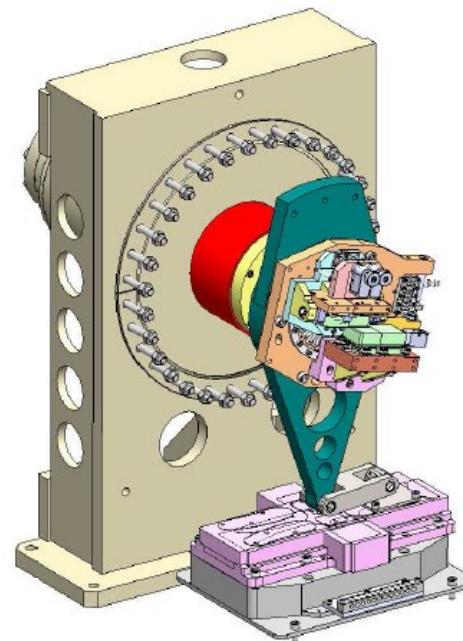


Actual load on shutter: $60 \times 2 \text{ mJ} = 0.12 \text{ J} / \text{pulse train}$
No problem with shutters, collimators etc.

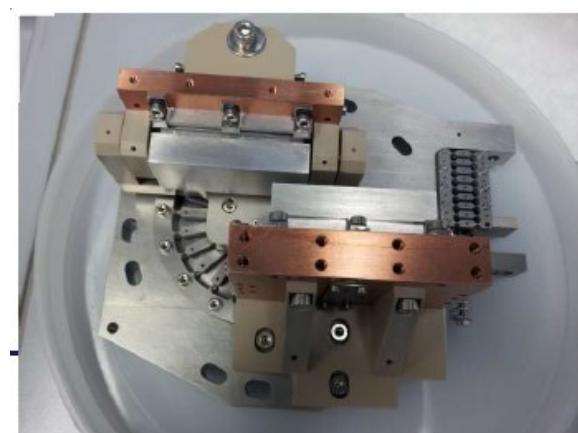
Monochromators for MID



Mono vessel with
pulse tube cryo cooler



D. Shu (APS), and X. Dong (WP-73, XFEL.EU)



Artificial channel-cut

Sine bar guided motion
for optimum stability

Silicon monochromator



1st mono in tunnel

929 m from source

Si (111)

Si (111)

Si (220)

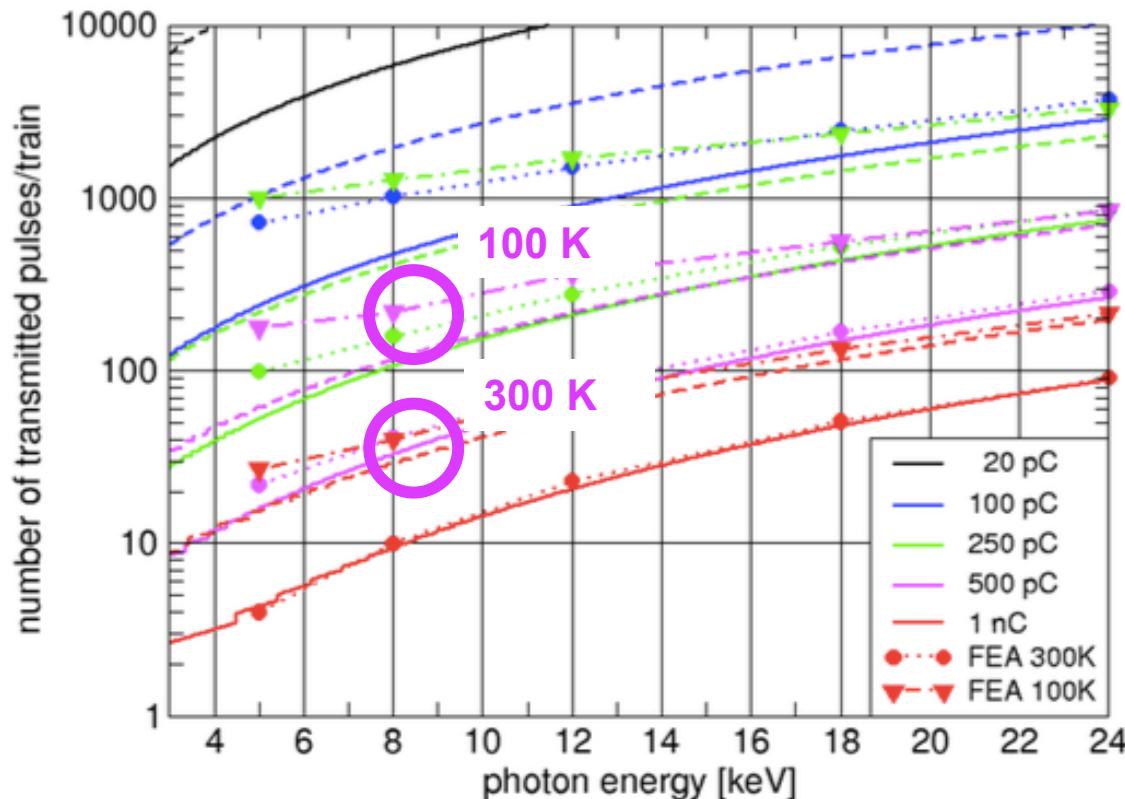
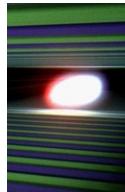
Si (220)

2nd mono in optics hutch
946 m from source

Table 13: Maximum temperature drifts for silicon and diamond monochromators

	Temperature [K]	Tolerable relative drift [10^{-4}]	Max. ΔT	W_{\max} /grams	W_{\max} /atom	W_{damage} /atom
Si (111)	300	1.4	50 K	36.3 J/g	10.5 meV	0.4 eV
Si (111)	100	1.4	144 K	64.3 J/g	18.7 meV	0.4 eV
C (111)	300	0.65	52 K	31.4 J/g	3.9 meV	1 eV
C (111)	100	0.65	168 K	29.0 J/g	3.6 meV	1 eV

Do we need cryo-cooling for the monos?



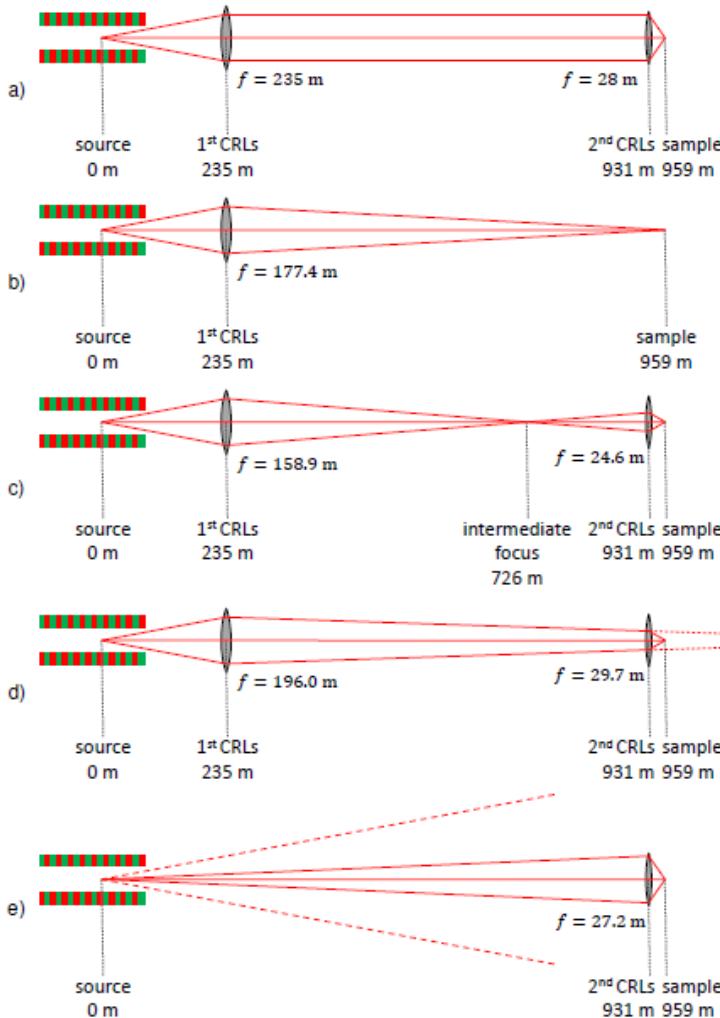
YES!
(on both 111 and 220)

~30 pulses/train will get through at 300K

~200 pulses/train will get through at 100K

Figure 46: Number of transmitted pulses through a Si (111) monochromator including heat flow perpendicular to the surface. Solid lines are 300 K, dashed lines 100 K. Dots are FEA results at 300K, triangles FEA at 100 K.

Focusing system at MID

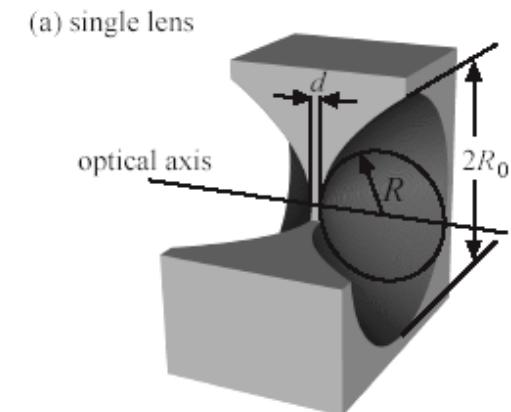


Best choice for Day-1

- CRL-1 @ 235m
- CRL-2 @ 931m



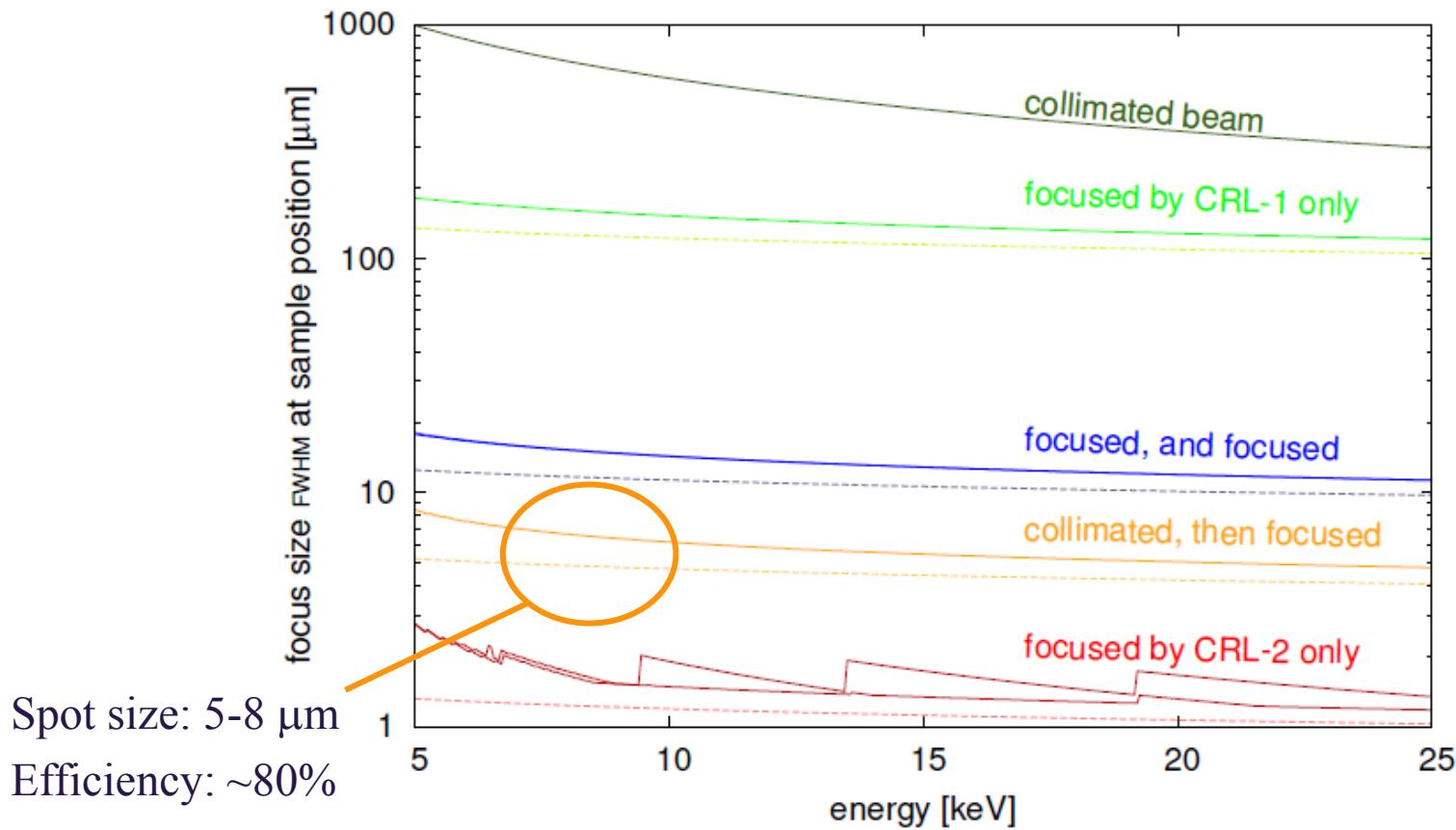
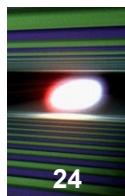
Too “dangerous“
for Day-1



Too small focus for the GeO₂ exp

Figure 5.2: Focusing schemes for the MID beamline: (a) Collimating then focusing scheme. (b) Direct focusing with the first CRL translocator (CRL-1). (c) Use of an intermediate focus. (d) Combined focusing scheme. (e) Brute focusing using only the second CRL translocator (CRL-2).

Focal spot size



Solid lines: Abbe's equation; Dashed lines: source size demagnification using geometrical optics

Source: MID TDR

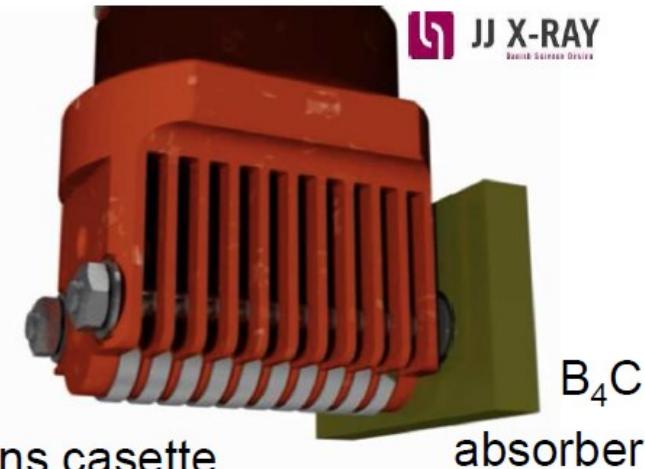
[doi:10.3204/XFEL.EU/TR-2013-005](https://doi.org/10.3204/XFEL.EU/TR-2013-005)

CRL transfocator

CRL 1	actuator	1	2	3	4	5	6	7	8
	filled with	1x5.8	1x4.9	1x4.0	1x3.3	2x5.8	3x4.0	7x4.0	7x2.0
CRL 2	actuator	1	2	3	4	5	6	7	8
	filled with	1x5.8	2x5.8	4x5.8	7x5.8	10x4.0	10x2.0	10x1.0	10x0.5

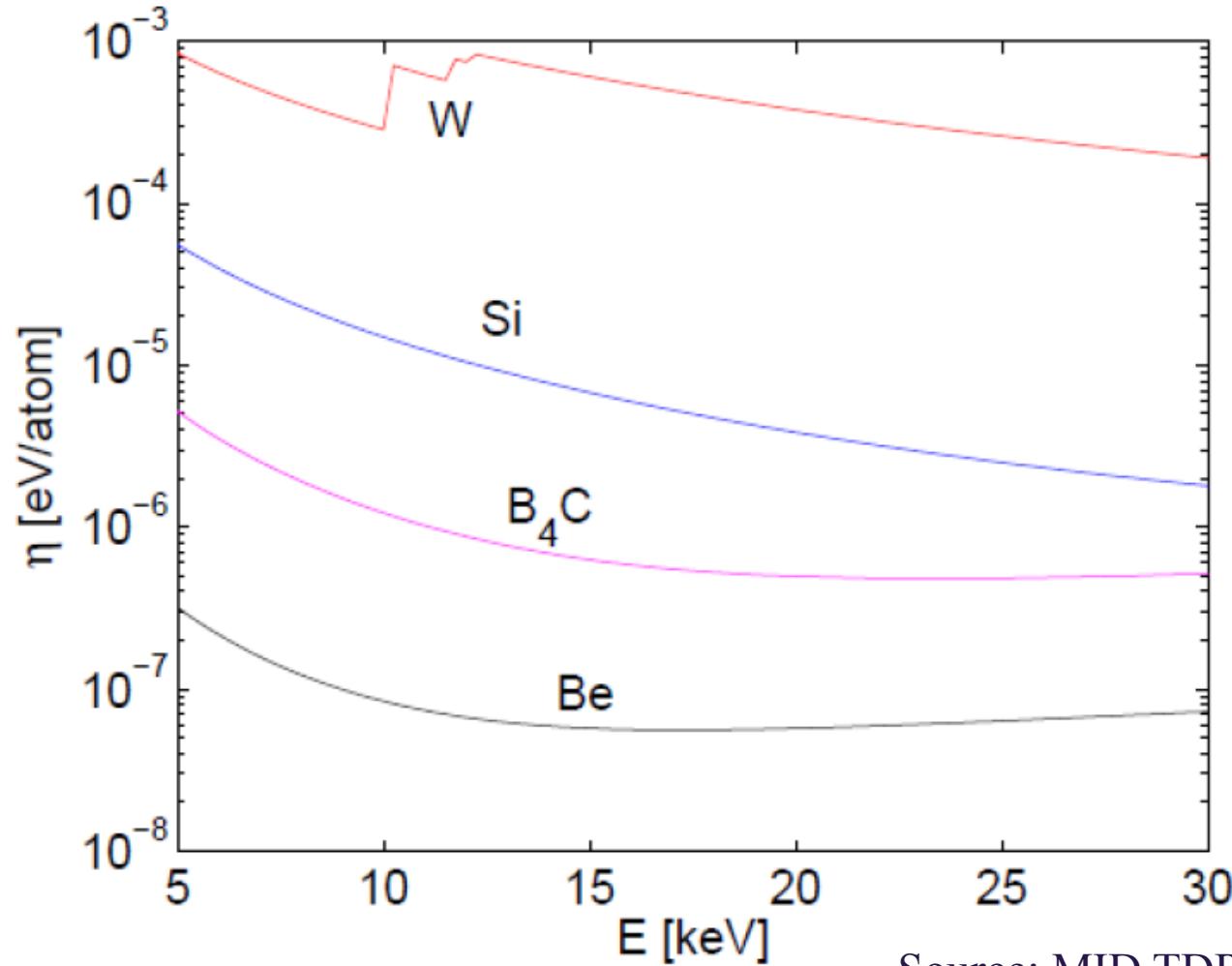


CRL 1: 23 lenses in total
CRL 2: 54 lenses in total



JJ X-ray & L. Batchelor

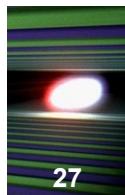
Example: 10^{10} photons in a $1 \times 1 \text{ mm}^2$ beam



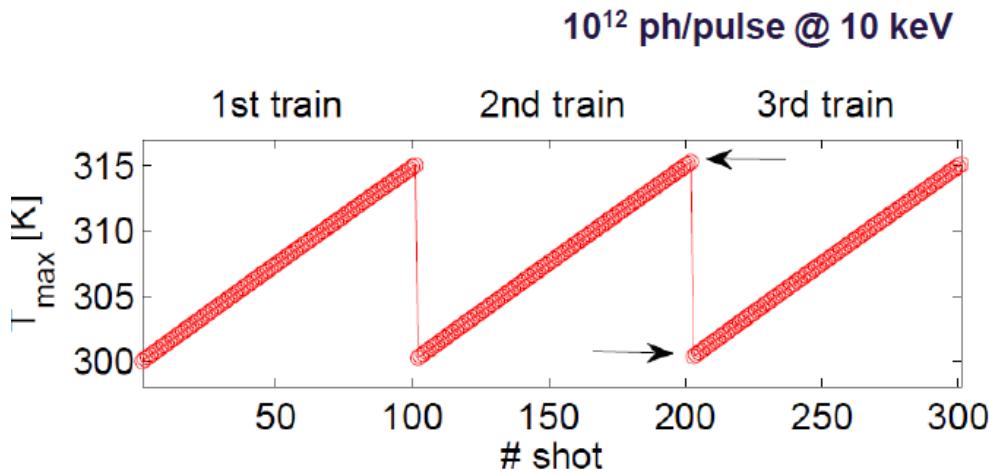
Source: MID TDR

[doi:10.3204/XFEL.EU/TR-2013-005](https://doi.org/10.3204/XFEL.EU/TR-2013-005)

CRL beam damage

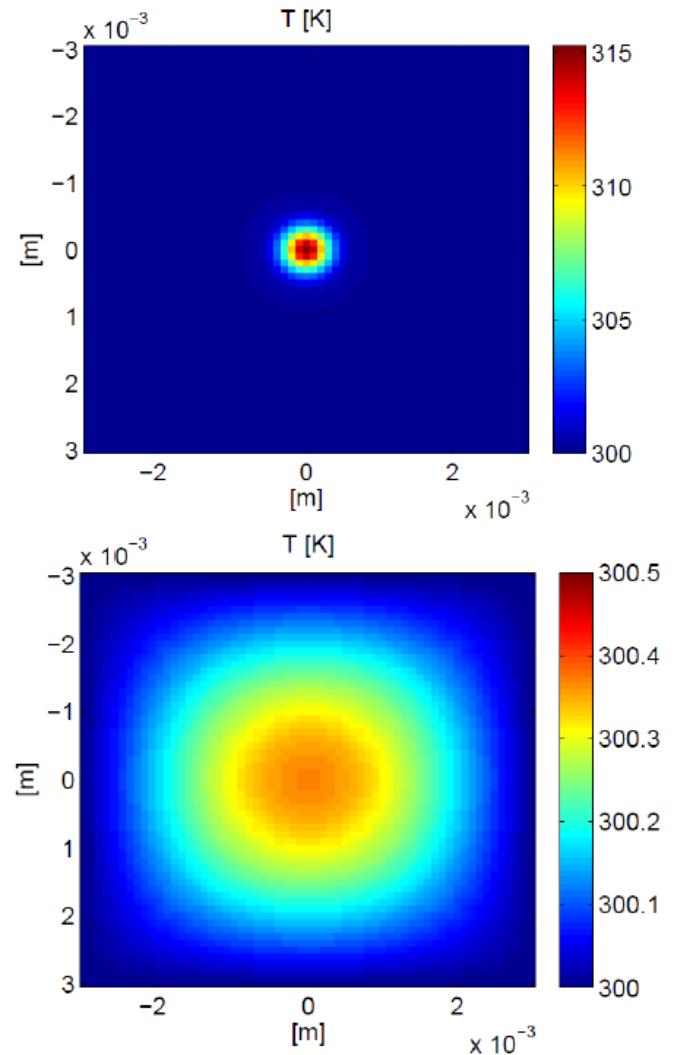


Adiabatic heating during trains,
cooling in between



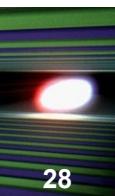
Limit on number of pulses per train.
Concern: repeated temperature cycles of Be

Shear stress limit?



Source: MID TDR
[doi:10.3204/XFEL.EU/TR-2013-005](https://doi.org/10.3204/XFEL.EU/TR-2013-005)

Commissioning toward 1st experiment



CRL-1 set for parallel beam at 8.4 keV (10 mm gap at 17.5 GeV)

CRL-2 to focus the beam on sample position (option a)

Diagnostics: screens, determination of focus size & Rayleigh length
(2 shifts)

Mono-1, Si(111) diffracting at 8.4 keV. Stability & heatload tests

Mono-2, Si(220) diffracting at 8.4 keV. Stability & heatload tests

Mono-1 & Mono-2 together
(3 shifts)

Coherence tests in SAXS geometry. Tuning of beamsize & intensity.

Ptychography and Speckle statistics. Scanning stage, Diagnostics &
Detector tune up

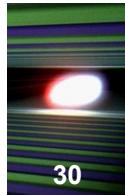
(4 shifts)

Data taking on GeO₂ at 100 kHz. Detector & diagnostics tweaking.
Large arm motion to change Q. Verification of mechanics precision.
On the fly-ish data analysis. Furnace regulation over a range of temperatures
(5 shifts)

TOTAL: 14 shifts (one week)

Very optimistic: Everything else has to work smoothly: Machine, Undulators,
WP-74 diagnostics, Offset mirrors, Safety, IT, Karabo & DAQ,....

Also: Enough time for installation and technical commissioning required
prior to commissioning with beam.



Hutches CfT

May 2015

Infrastructure CfT

Oct 2015

Hutches + infrastructure ready
(installation of sensitive components possible)

June 2016

1st SASE beam

April 2017