

Influence of the geomagnetic field on gamma/hadron separation in Cherenkov telescopes

Outline

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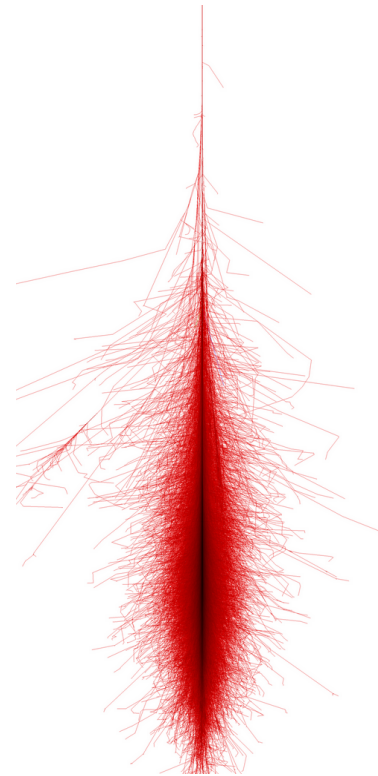
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- 4 ground based Cherenkov telescopes in Arizona (US)
- Maximum sensitivity: 100 GeV – 10 TeV
- For this project:
 - Background sample: VERITAS data
 - Signal sample: simulations

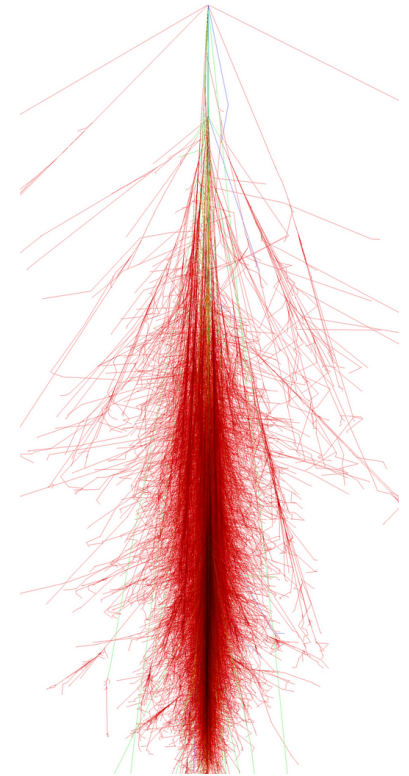


Gamma-ray and hadron showers

- Gamma rays produce an electromagnetic cascade: extensive air shower
- Gamma-ray showers are different from proton showers
- The geomagnetic field influences the shape of the showers



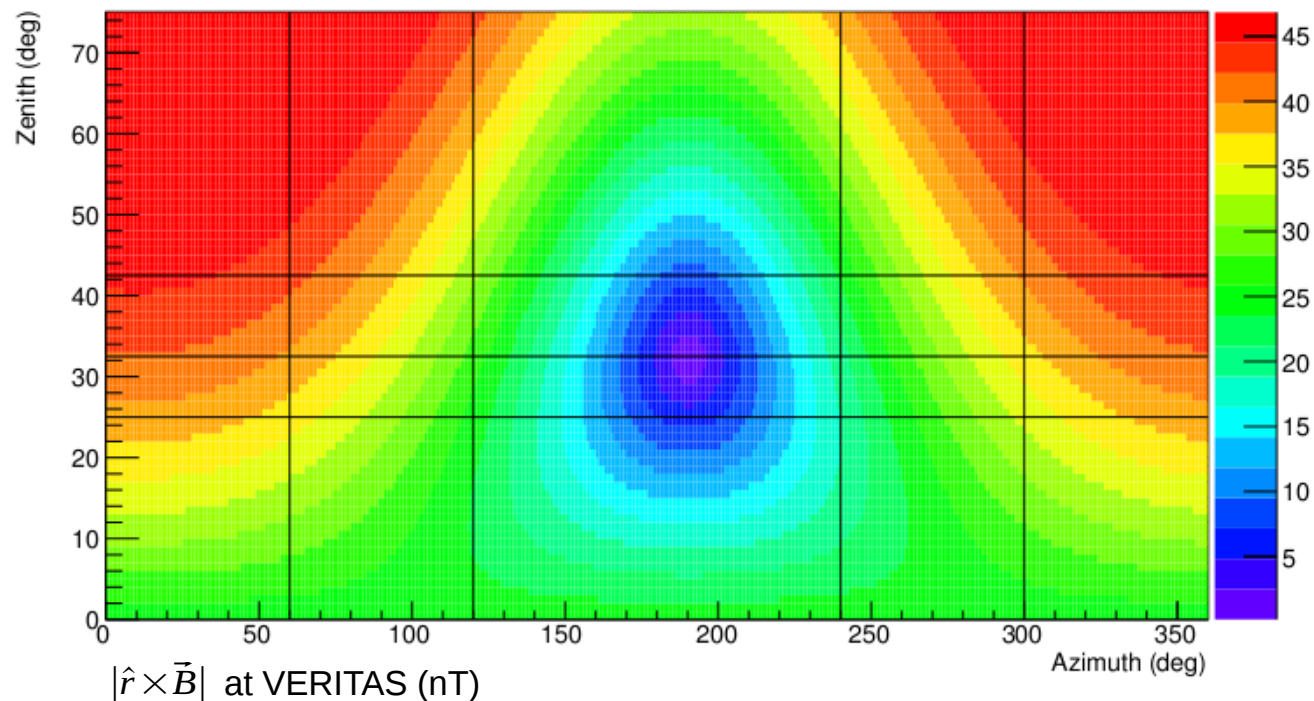
1 TeV gamma ray



1 TeV proton

The geomagnetic field

Lorentz force: $\vec{F} = q(\vec{v} \times \vec{B})$

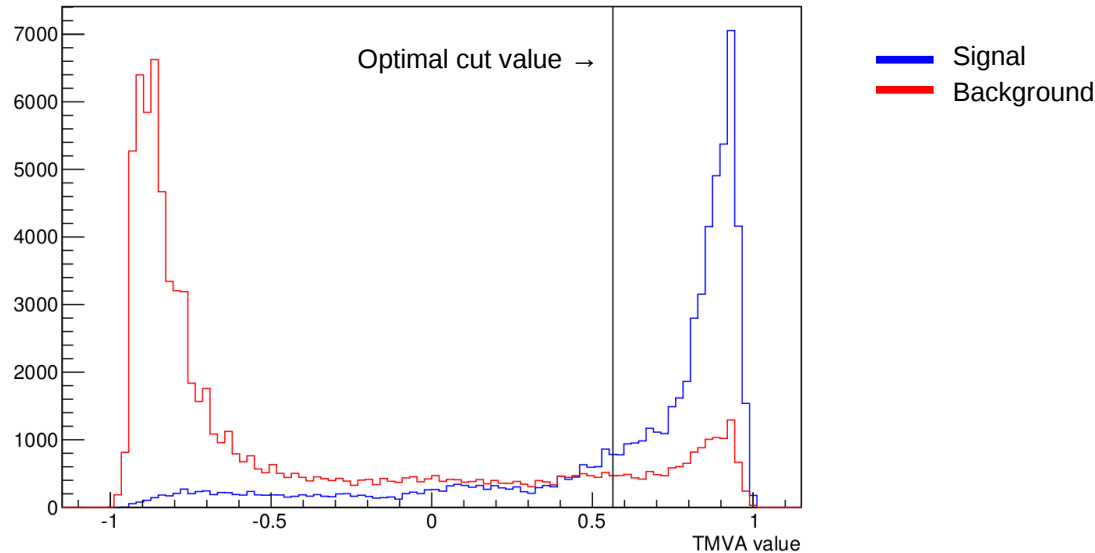


Project:

- Does this influence the separation of gamma-ray and proton showers?
- Can we improve gamma/hadron separation by taking it into account?

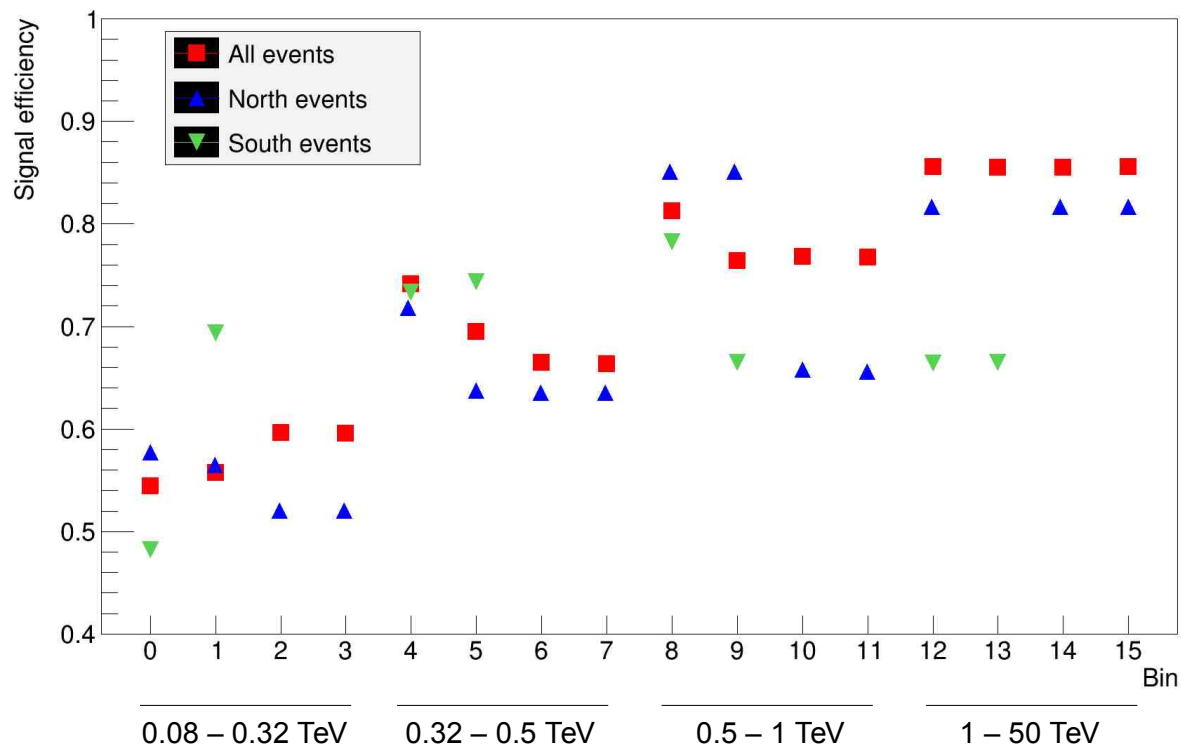
Boosted Decision Trees (BDTs)

- Multivariate analysis technique
- Built (trained) and tested with a sample of known signal and background events
- Each event \rightarrow value from -1 to 1
- Best cut value: highest significance
- To compare performance of BDT: signal and background efficiency



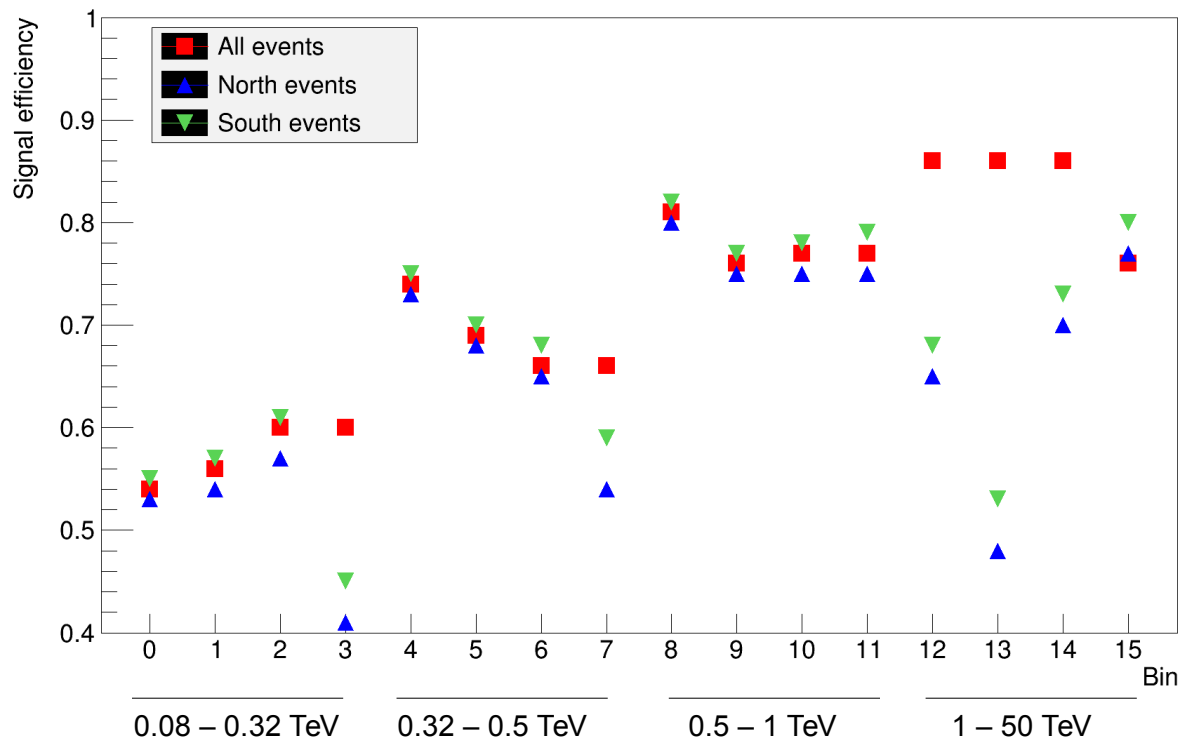
BDT training for north and south

- Two sets of BDTs trained with only N and S events (separately)
- Differences in signal efficiencies → direction influences training
- Not conclusive



BDT training for all directions

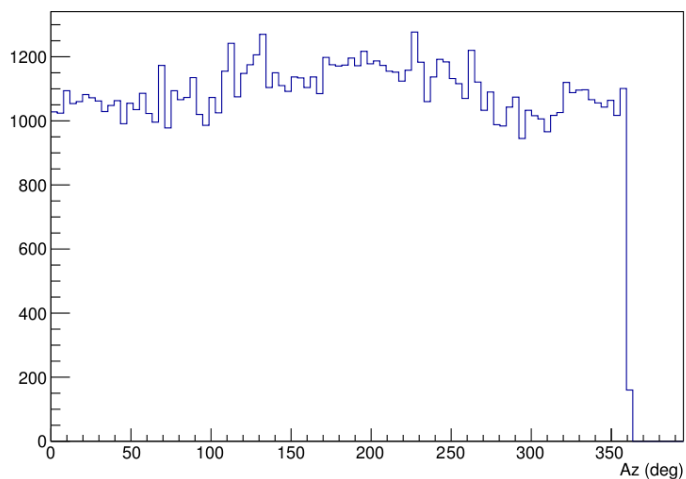
- BDTs trained with events from all directions
- Question: are events from a certain direction more accurately classified?
- Comparison of efficiency when looking only at N or S events
- South signal efficiency higher for all bins → agreement with initial expectations



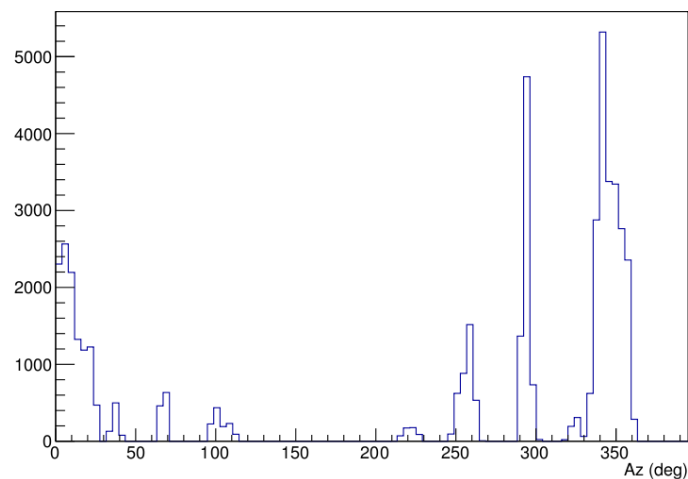
BDT training for all directions – Azimuth distribution

Surprise: background sample used for training is not evenly distributed (specially S).

- Limit in analysis
- Focus on N-trained BDTs



(a) All signal events



(b) All background events

$E = 0.5 - 1 \text{ TeV}$, $Z_e = 25 - 32.5^\circ$

- Analysis of a northern gamma ray source: binary system LSI +61 303
- 3 different methods:
 - Box cuts
 - Conventional BDTs
 - N-trained BDTs
- No improvement of significance for N-trained BDTs

Method	N_{on}	N_{off}	σ
Box cuts	1006	518	17.2
BDT cuts	792	336	19.1
North-trained BDT cuts	788	339	18.8



Conclusions

- **The direction of the showers influences their classification in BDT → the geomagnetic field affects BDT performance**
- In a conventional BDT, southern showers are better classified (as expected)
- Still unclear if training for different directions could improve BDT performance

Further studies:

- Training for N and S with larger and evenly distributed background sample
- Study of southern source with S-trained BDTs (improvement more likely)



A vibrant, multi-colored nebula with green, blue, and orange filaments against a starry black background. The nebula's structure is complex, with intricate patterns of gas and dust. The colors are bright and saturated, creating a striking contrast with the dark space. Numerous small, bright stars are scattered throughout the field of view, adding to the cosmic beauty of the scene.

Thank you for your attention!

$$S = \sqrt{-2 \ln \lambda} = \sqrt{2} \left\{ N_{\text{on}} \ln \left[\frac{1 + \alpha}{\alpha} \left(\frac{N_{\text{on}}}{N_{\text{on}} + N_{\text{off}}} \right) \right] + N_{\text{off}} \ln \left[(1 + \alpha) \left(\frac{N_{\text{off}}}{N_{\text{on}} + N_{\text{off}}} \right) \right] \right\}^{1/2}$$

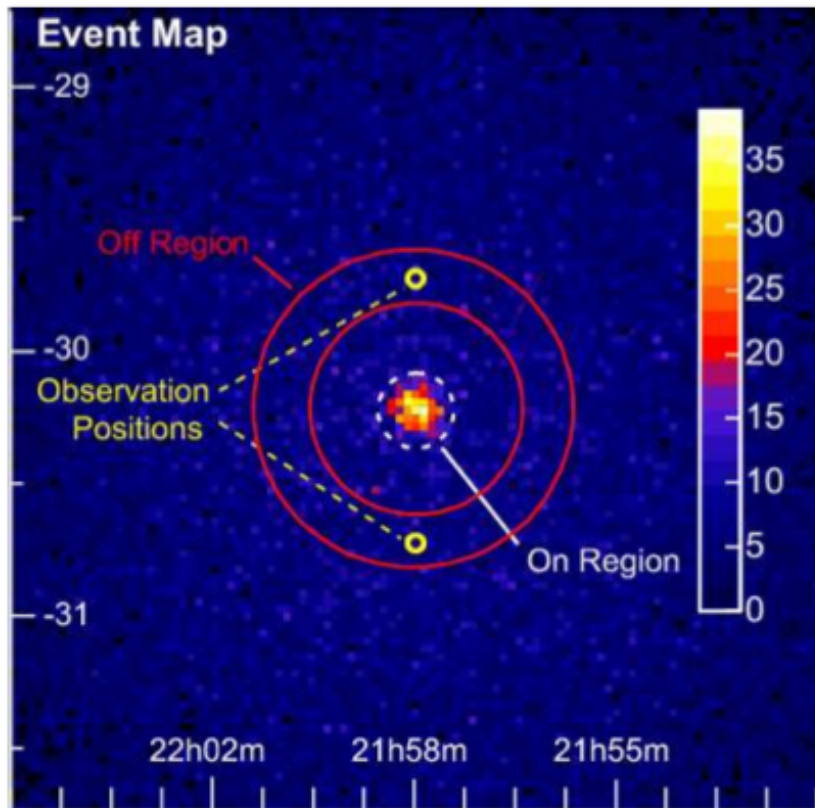
$$p = N(u = S; 0, 1)$$

Li & Ma, 1983

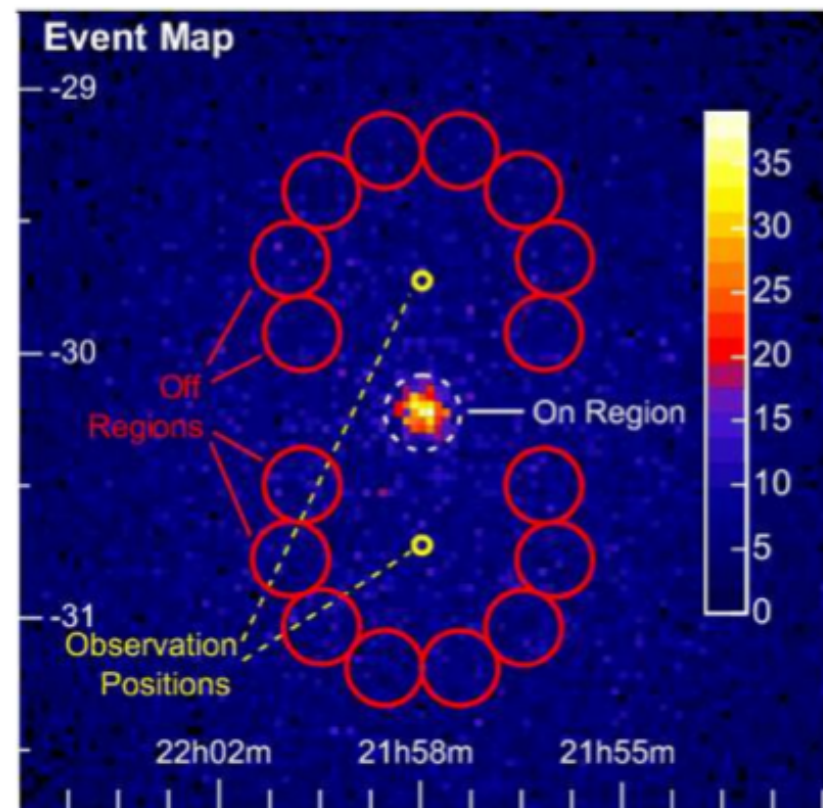


Extra slides

On and Off events



Ring background



Reflected region