

# Experimental Astroparticle Physics

**PRC open session, 22.10.2015, DESY, Hamburg**

## **Groups & People:**

**H.E.S.S.:** Klepser, Stegmann

**Magic:** Bernardini, Garczarczyk

**VERITAS:** Maier, Pohl, Schlenstedt

**CTA:** Bühler, Garczarczyk, Knapp,  
Maier, Schlenstedt, Stegmann,  
Wischnewski

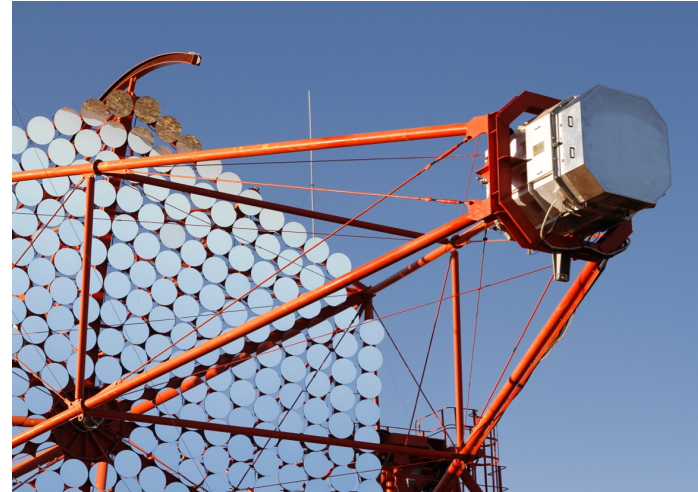
**FERMI:** Bühler, Ackermann

**IceCube:** Ackermann, Bernardini,  
Franckowiak (start Nov),  
Kowalski, Nanhauer

**Zwicky Transient Facility:**  
Franckowiak (start Nov), Kowalski

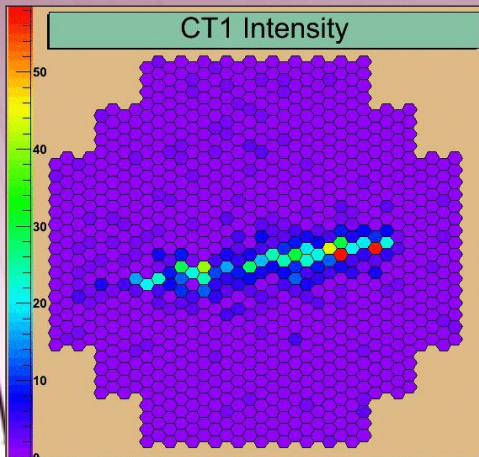
# H.E.S.S. I Camera Upgrade

- Redesign of Camera electronics for lower dead time and down time
- July 2015: Deployment of first camera
- Installation could be finished within the planned schedule of 3 weeks



# August 15: First light!

- First cosmic gamma-ray showers

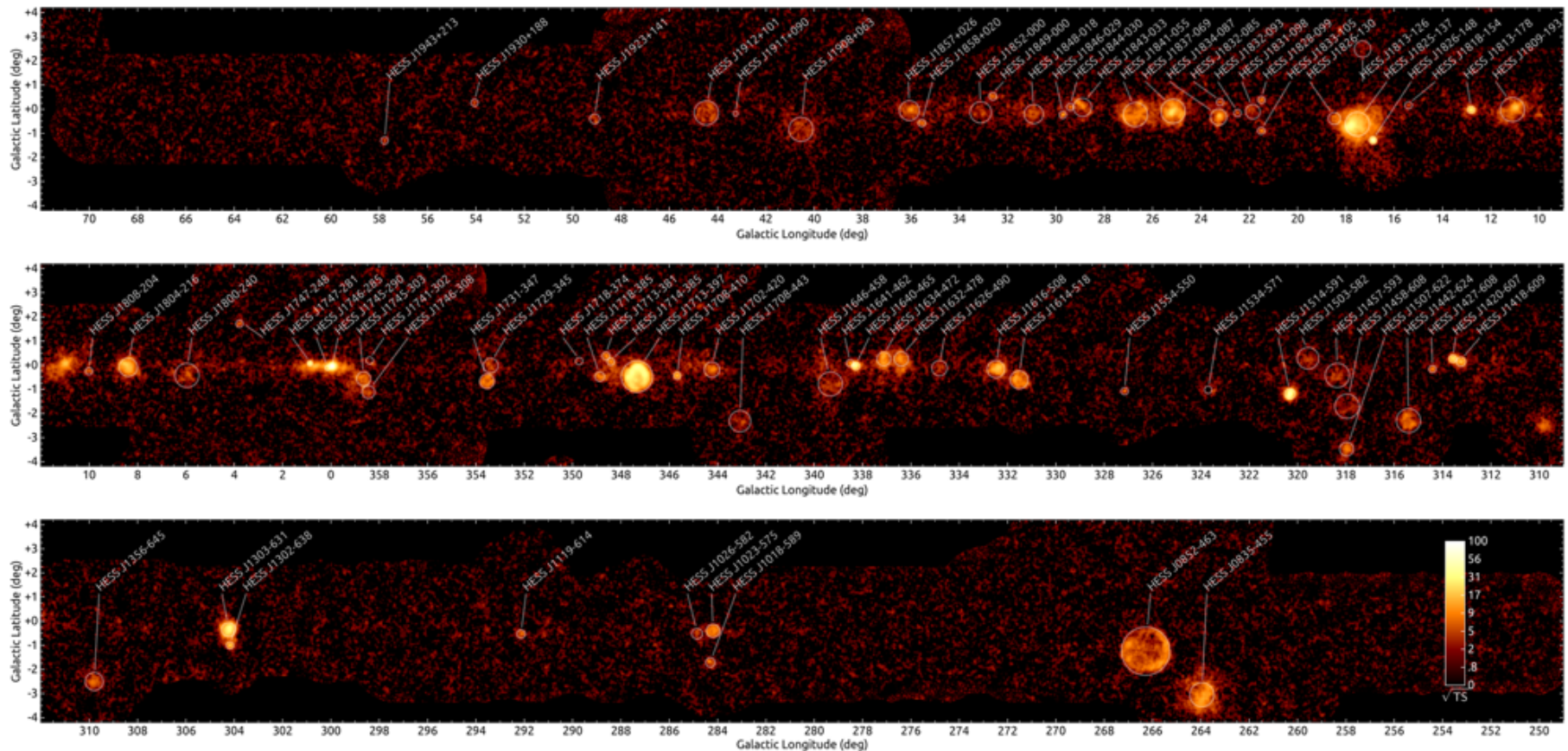


# Plan for Cameras 2-4

- Production has already begun
- Tentative deployment in September 2016

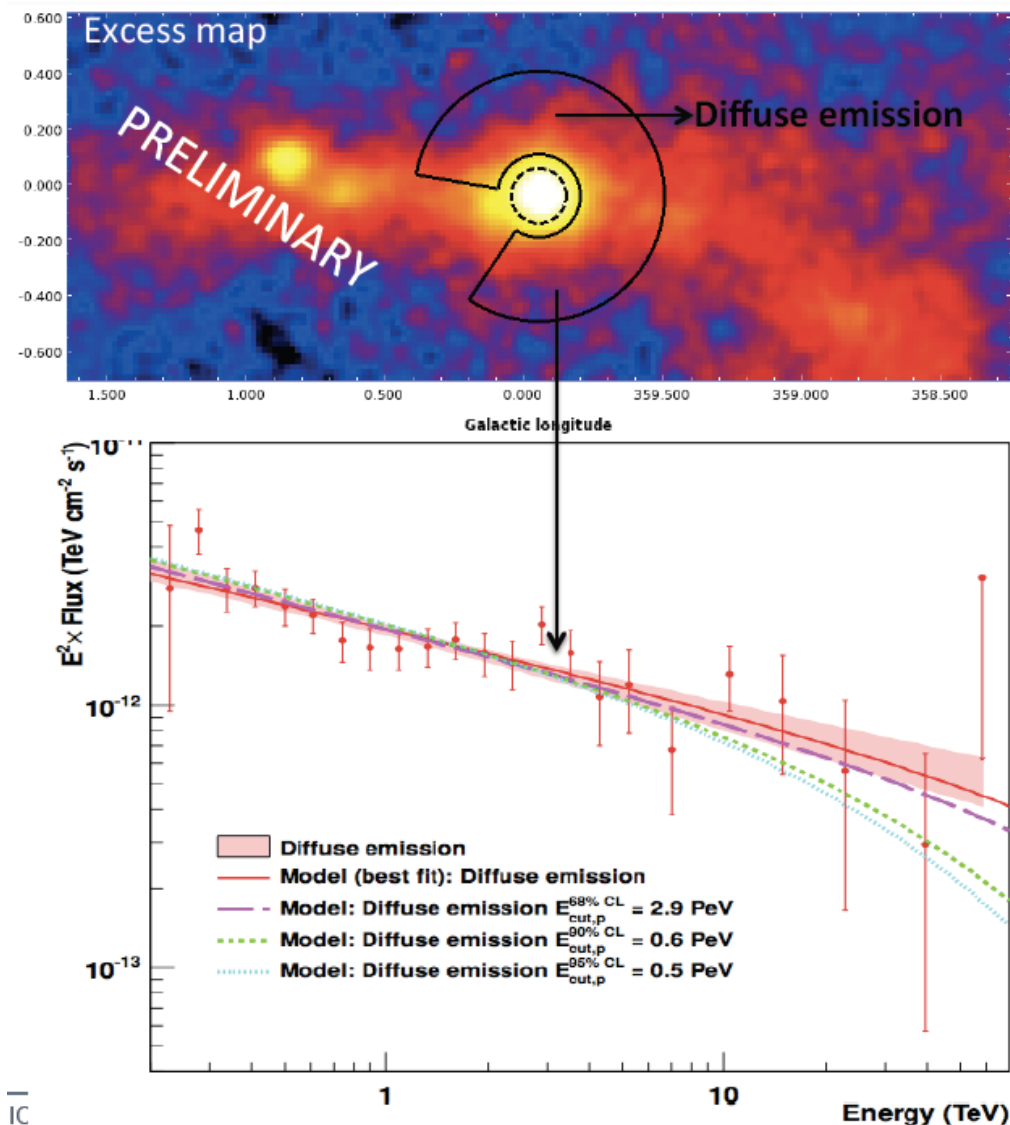


- Work on A&A special issue around galactic plane scan
- Some results shown on ICRC 2015



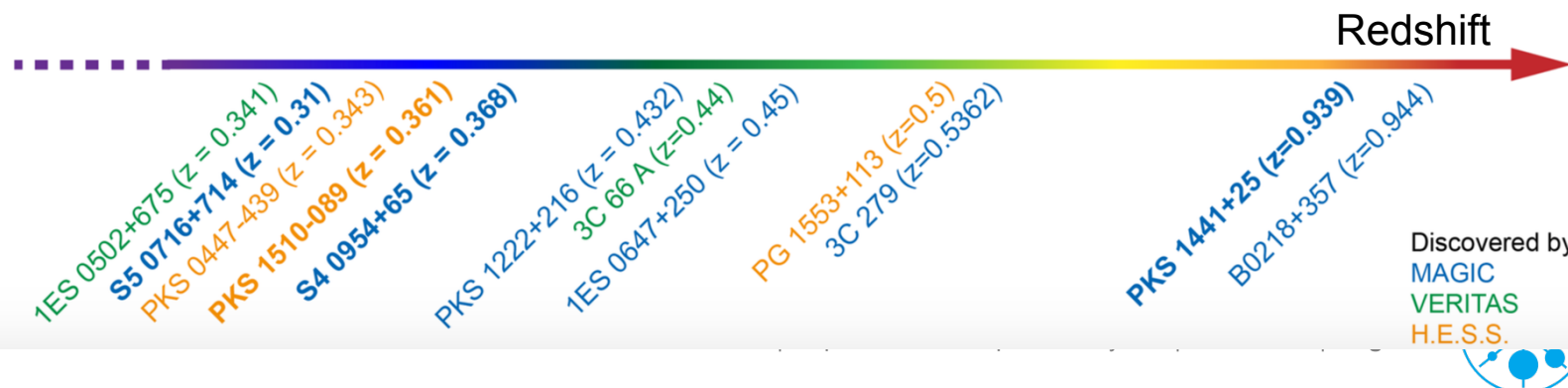
# A Pevatron in the Galactic Centre

- Very hard spectrum of diffuse emission
- $1/R$  profile suggests continuous injection and diffusive propagation
- Conclusion: Central PeV accelerator injecting particles in quasi-continuous fashion
- Detection of first cosmic PeVatron!



## Astronomical Telegram – ATEL!

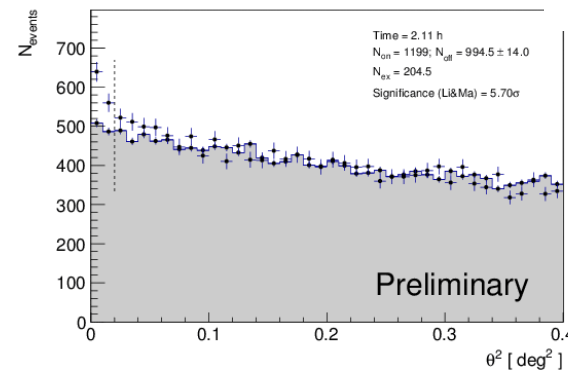
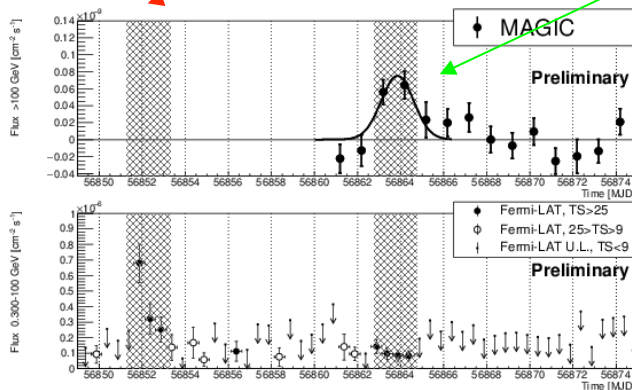
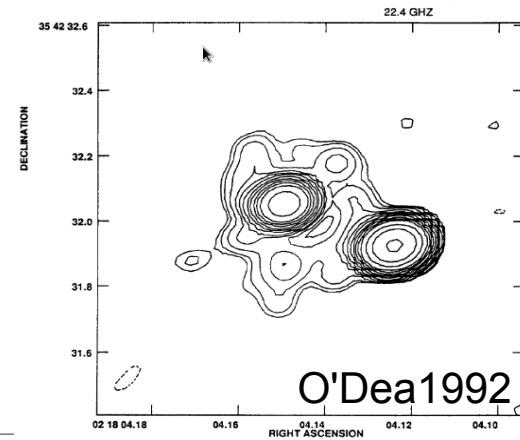
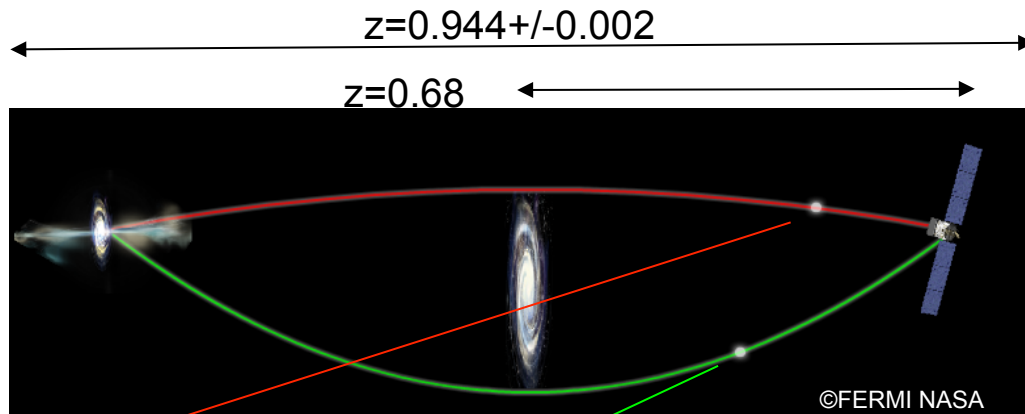
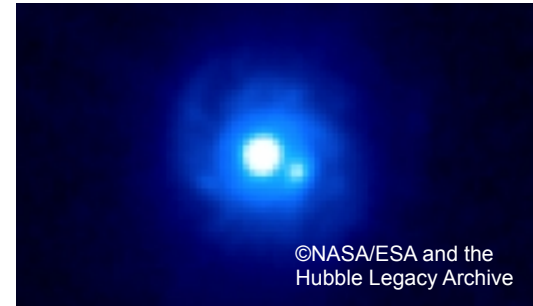
- ATEL #6999 “MAGIC detects Very High Energy gamma-rays from S5 0716+714” G.Pedaletti @DESY analyzer
- ATEL #7080 “Discovery of Very High Energy Gamma-Ray Emission from the FSRQ S4 0954+65 with the MAGIC telescopes” → G.Pedaletti @DESY analyzer and leading the publication
- ATEL #7416 “Discovery of Very High Energy Gamma-Ray Emission from the distant FSRQ PKS 1441+25 with the MAGIC telescopes” ( $z=0.939$ )
- ATEL #7542 “MAGIC detects an increased activity from PKS 1510-089 at very high energy gamma-rays” G.Pedaletti @DESY analyzer
- ATEL #7660 “MAGIC detects an increased activity from BL Lacertae at very high energy gamma-rays” “Archetipical” BL Lac celebrates the MAGIC collaboration meeting held at DESY Zeuthen
- ATEL #7844 “Discovery of Very High Energy Gamma-Ray Emission from the intermediate BL Lac S2 0109+22 with the MAGIC telescopes” G.Pedaletti @DESY taking data in La Palma



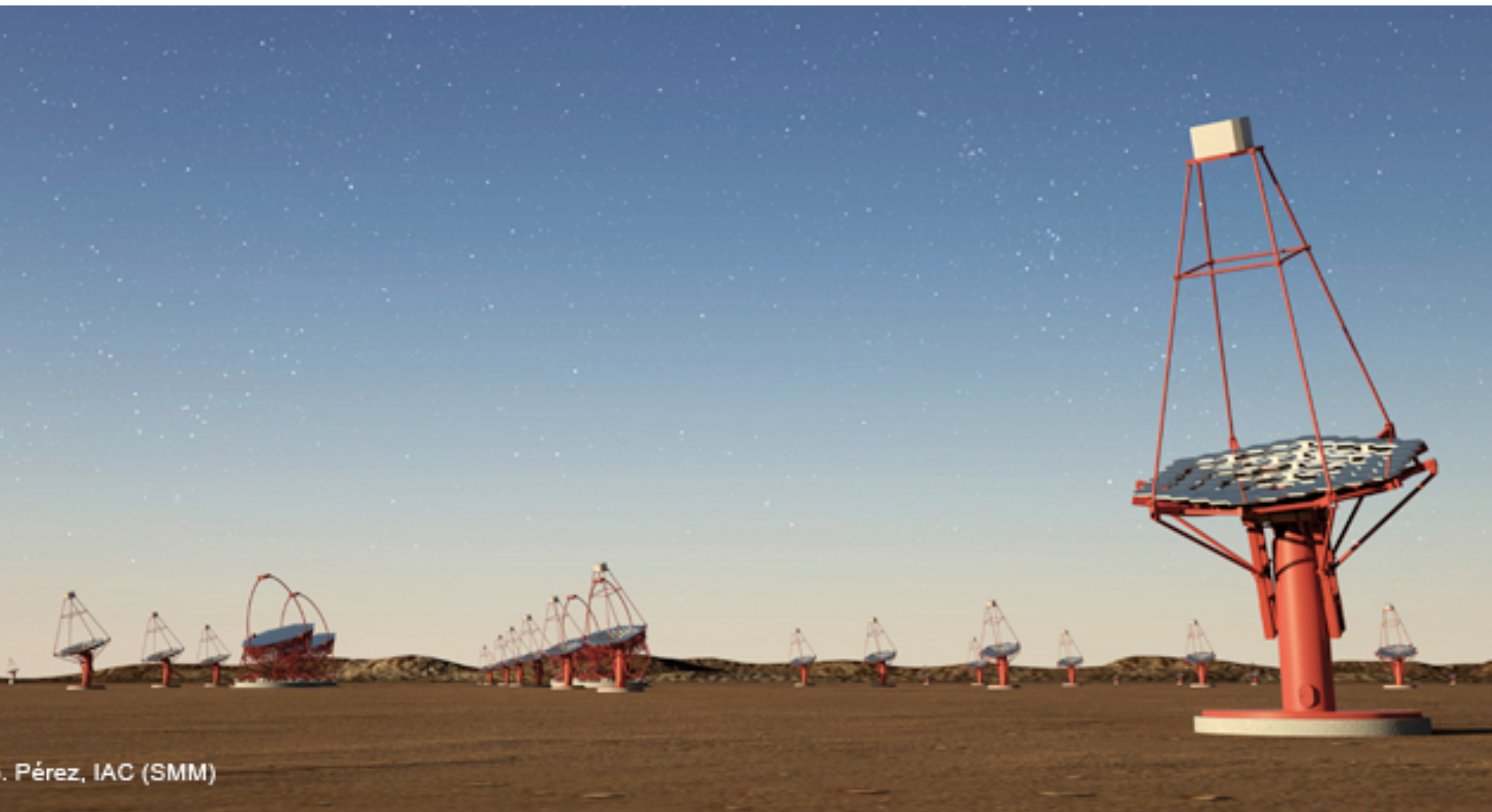
# Farther and farther: B0218+357



- A flare in Fermi from B0218+357 is detected on July 13/14 2014, full moon MAGIC
- Expected delay of  $\sim 11$  days  $\rightarrow$  MAGIC triggers, we get it  $\rightarrow$  at  $z \sim 1$  is an excellent probe of EBL



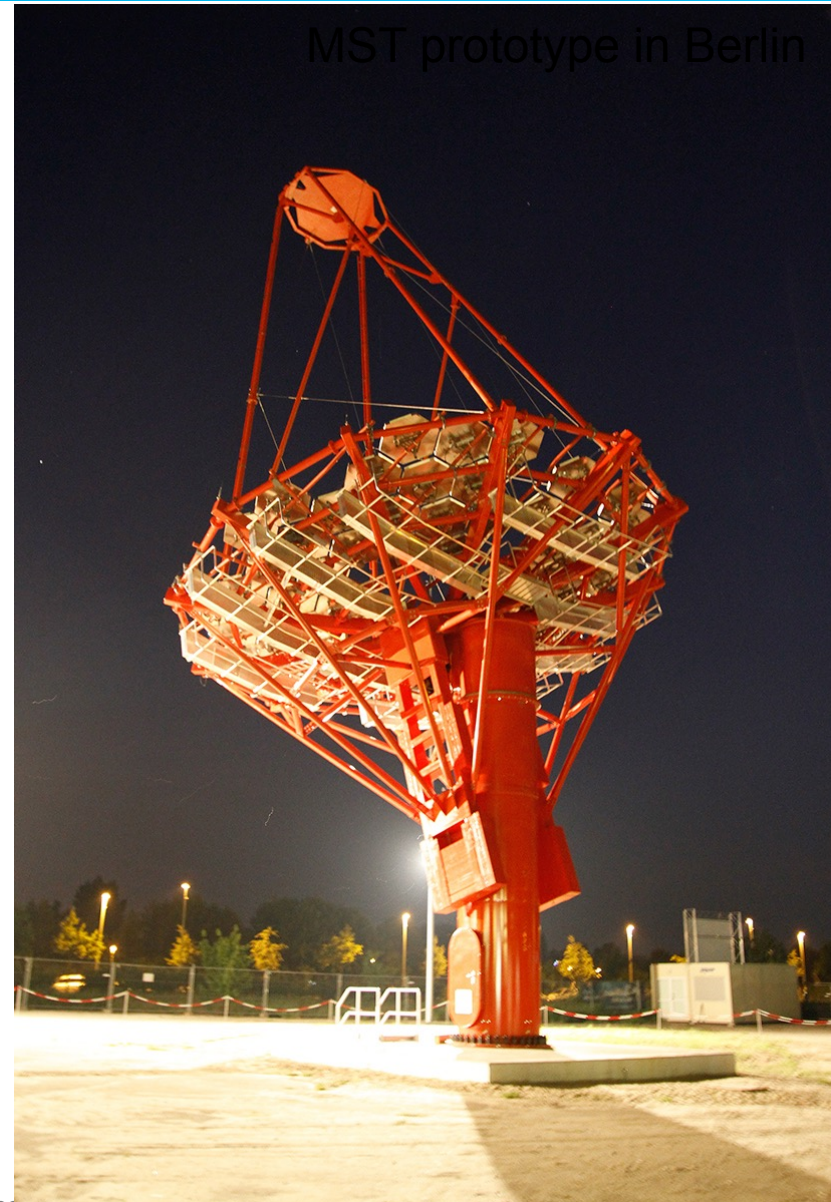
# CTA



. Pérez, IAC (SMM)

# CTA Involvement at DESY

- > Medium-Sized Telescopes MST
  - ❑ Design and build structure, drive and safety system
  - ❑ Project Management responsibility
- > Array Control
  - Work Package Array operation
  - Work Package On-site hardware
  - Timing System
  - Project Management responsibility, central activities
- > Sensitivity Studies, e.g. CTA layout
- > Data structures and pipelines
- > High level software (ctools)



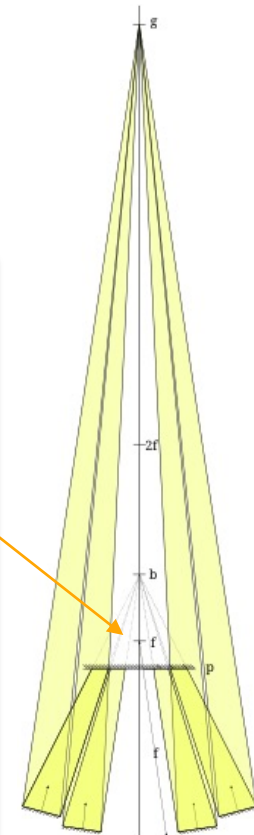
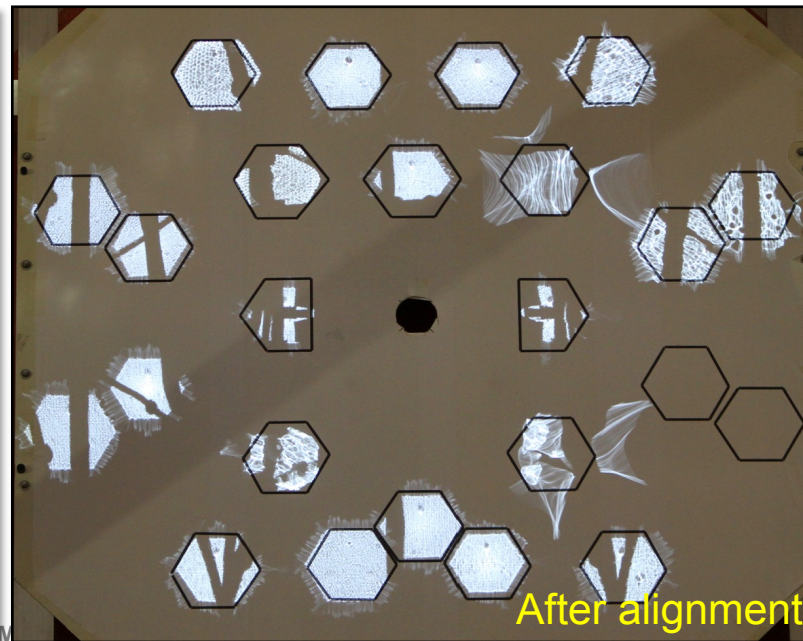
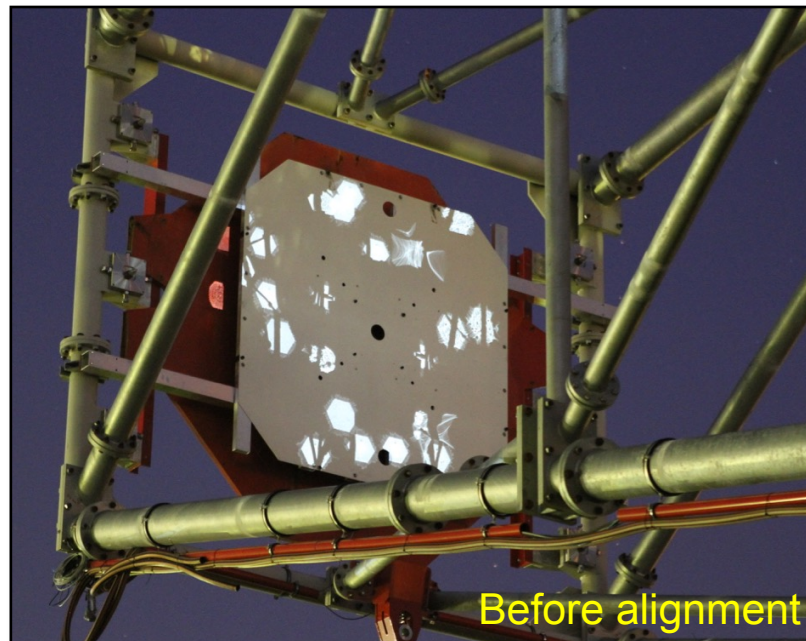
# MST: New Mirror Alignment Method



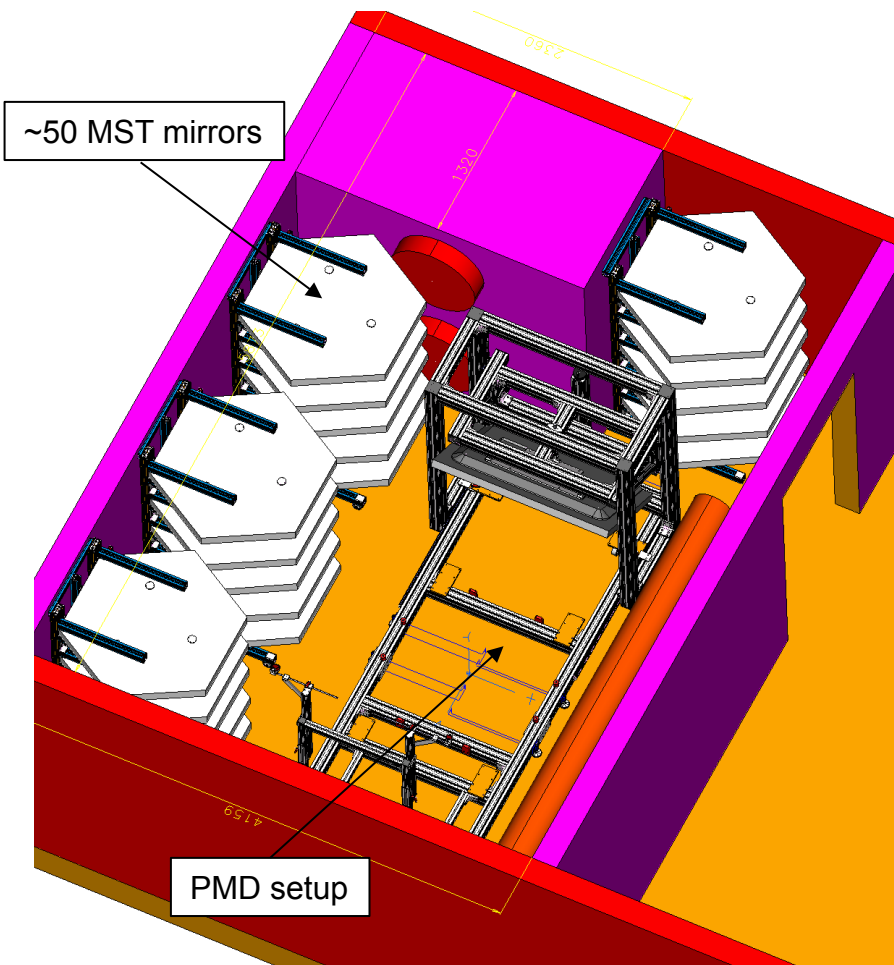
- > Difficult to align individual mirror facets under bright night sky conditions in Berlin
- > Bokeh (out of focus) method applied:

Nearby artificial light source

Ray-tracing of the image in focal plane



# MST: Mirror Evaluation at DESY

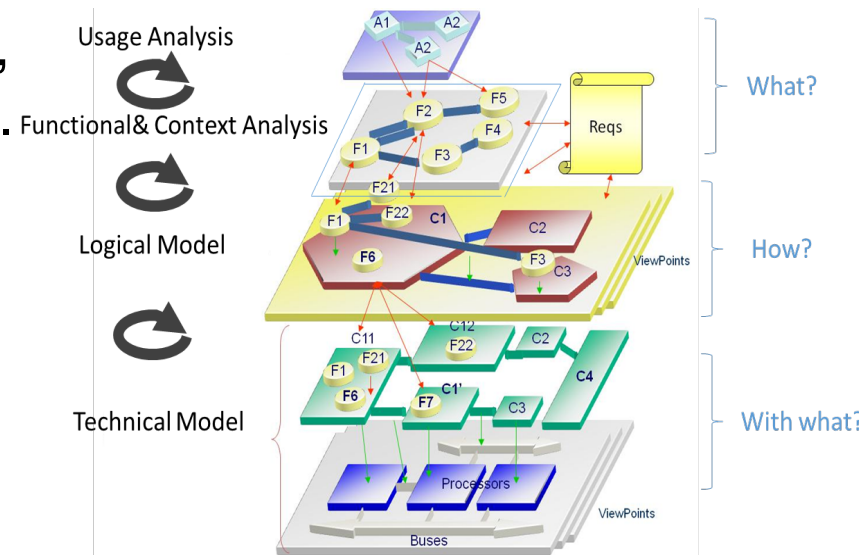


DESY climate chamber

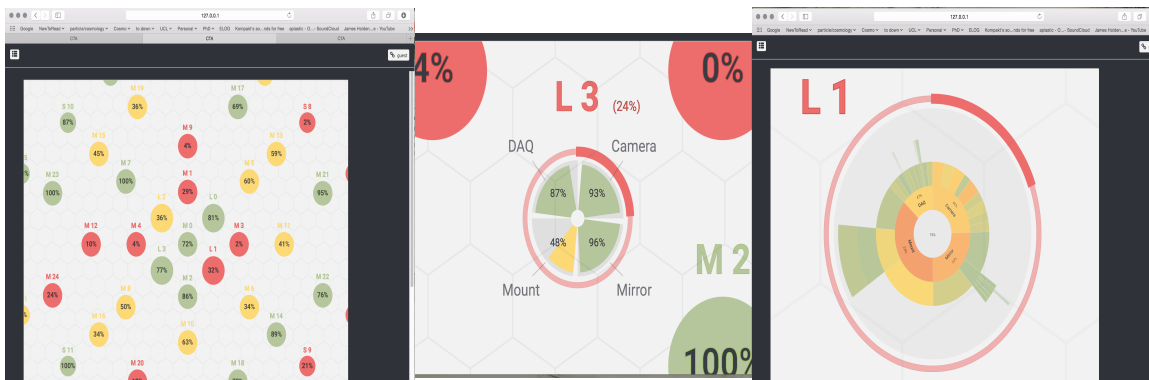
- > Three mirror stakeholders for MST: CEA, IFJ-PAN, MLT
- > New prototype mirrors in production, being delivered to DESY and CTA Mirror Test Facilities (MTF)
- > Tests performed at DESY:
  - > Point Spread Function, focal length and focused reflectivity – measurements using 2f setup and Phase Measuring Deflectometry
  - > Long term performance with varying temperature cycling ( $-20^{\circ}$  ...  $30^{\circ}\text{C}$ )

# Array Control (ACTL)

- Project leadership and management, Software architecture (see right Fig.).
- Design and implementation of the central control and operator GUI elements (see bottom).



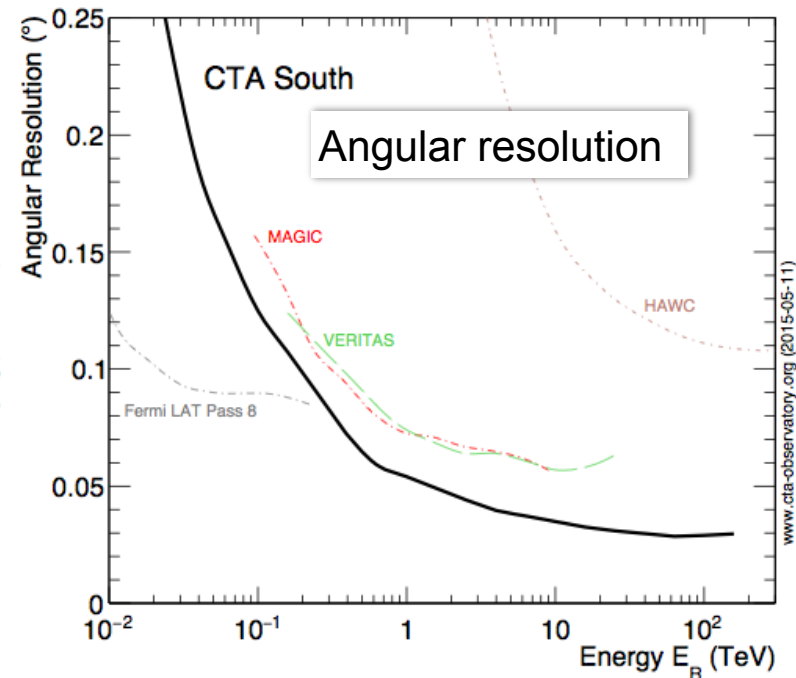
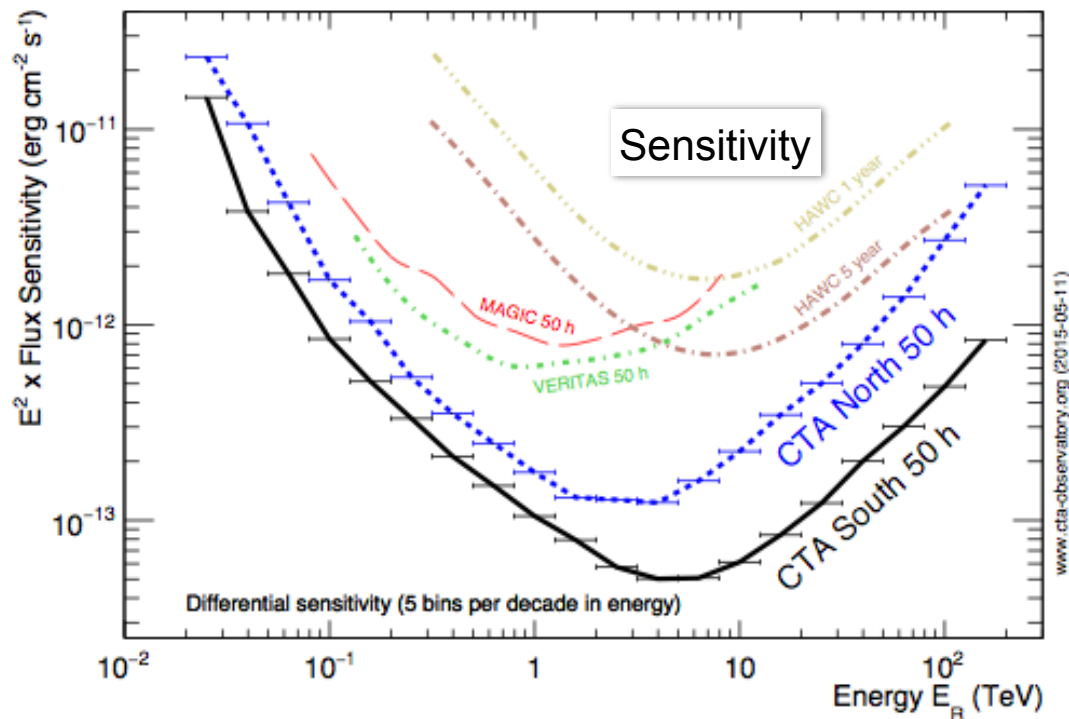
**Conceptual view of the adopted system architecture approach, based on SysML**



**Semantic zooming of CTA operator GUI prototype**

# CTA Performance

DESY group one of the key player in MC group: large input into site evaluation; next step is optimisation of array layout



CTA Public Performance Page:

[https://portal.cta-observatory.org/CTA\\_Observatory/performance/SitePages/Home.aspx](https://portal.cta-observatory.org/CTA_Observatory/performance/SitePages/Home.aspx)

# News from Cherenkov Telescope Array CTA

## > Critical Design Review June 2015

- ❑ Delivery of Detailed Technical Design Reports for all Work Packages
- ❑ Plans for pre-production and production for 70-100 telescopes on Southern Array and 20 telescopes on Northern Array
- ❑ Towards CTA Cost Book and Schedule

## > CTA Headquarters

Expression of Interest by Germany (DESY, Zeuthen), Italy (INAF, Bologna), Brazil (Rio de Janeiro)

Call for proposals early October 2015

Deadline for applications Nov 22, 2015

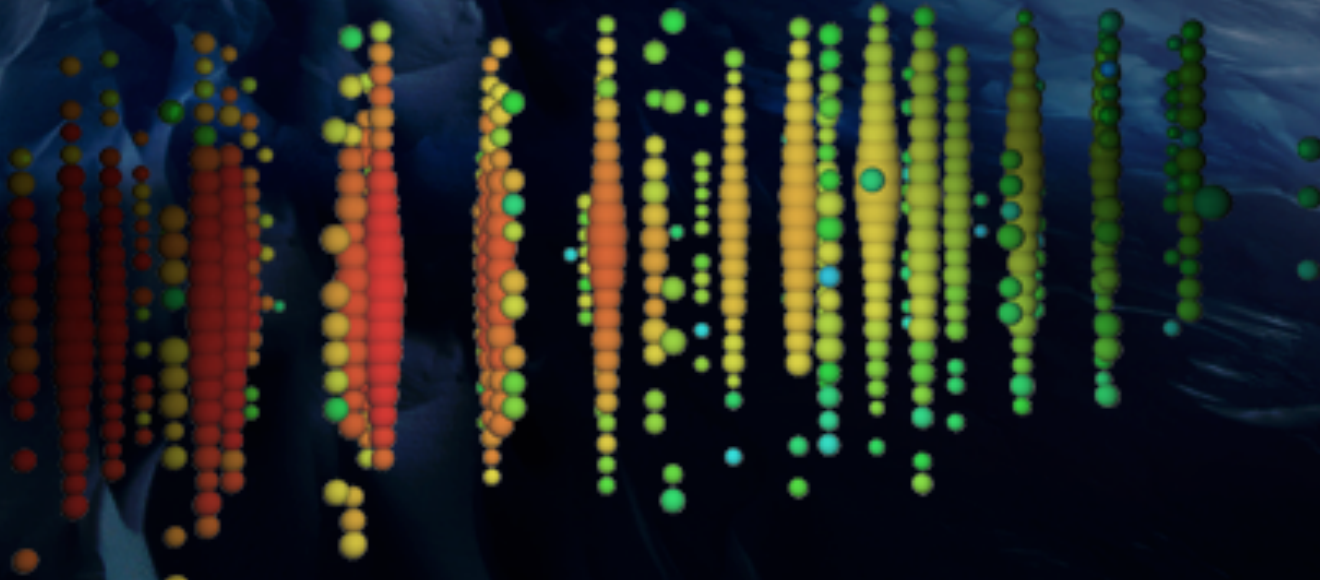
Decision spring 2016

## > Towards a CTA site decision: decision to start detailed negotiations in July with ESO (Chile) and Spain (La Palma)

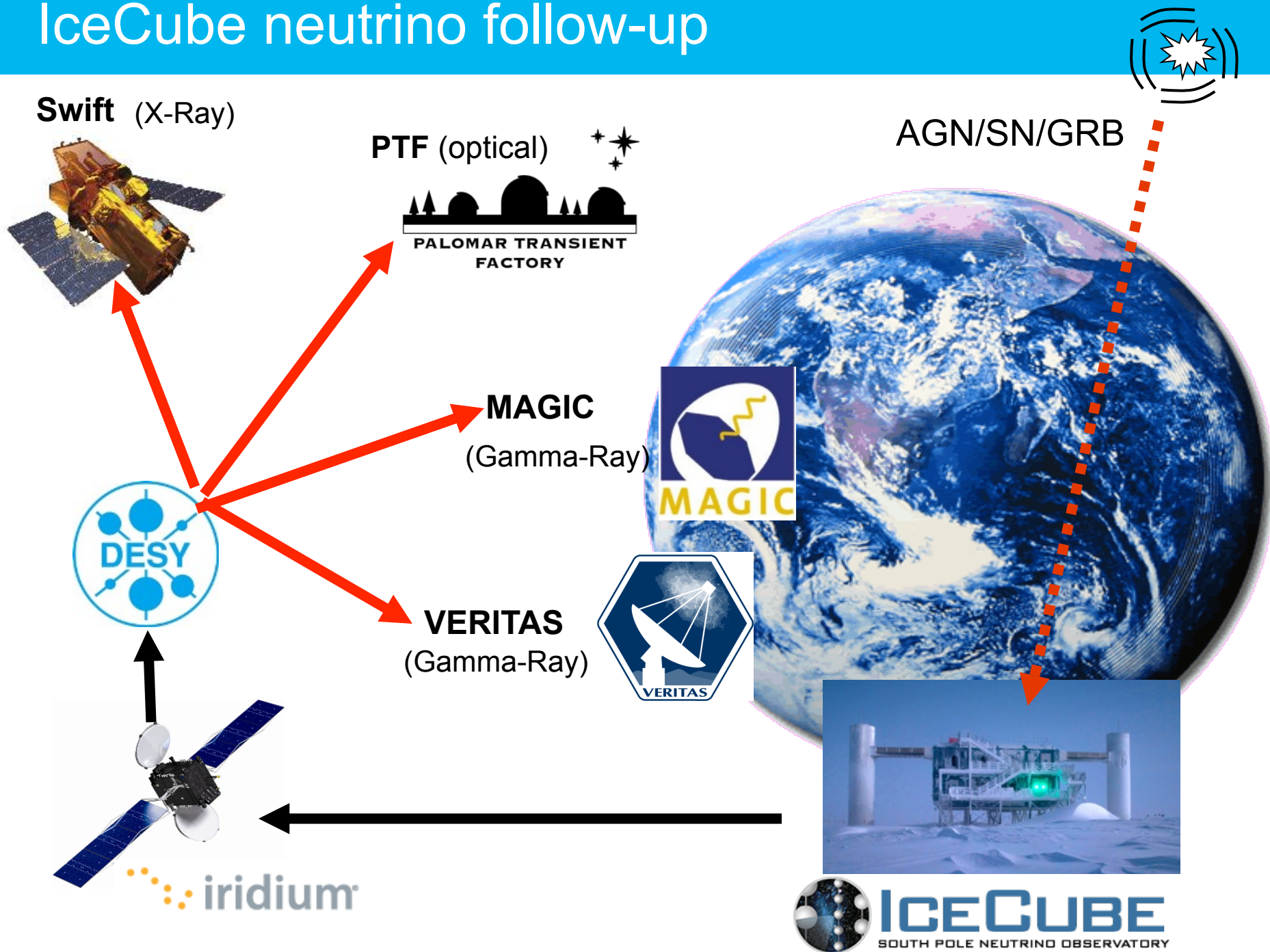


# IceCube

The highest energy event so far,  
deposited energy in IceCube:  $2.6 \pm 0.3$  PeV



# IceCube neutrino follow-up



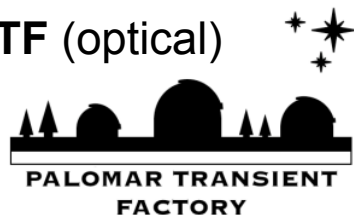
# IceCube neutrino follow-up



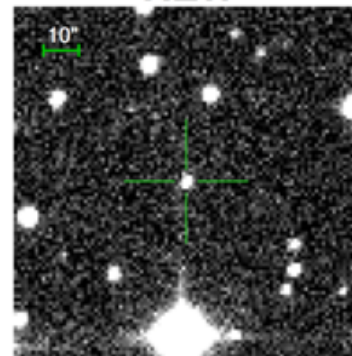
**Swift** (X-Ray)



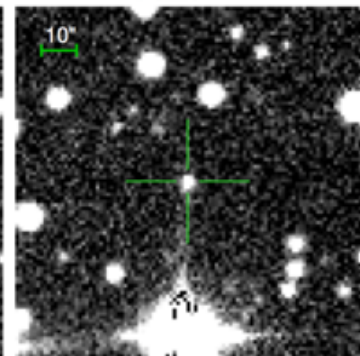
**PTF** (optical)



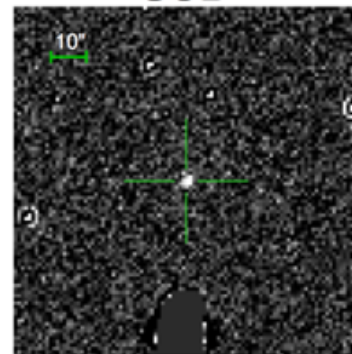
**NEW**



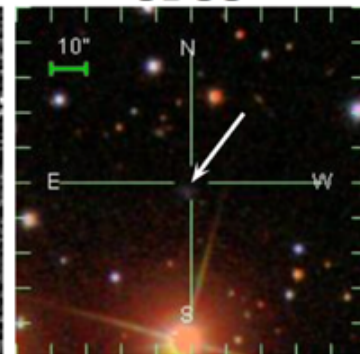
**REF**



**SUB**



**SDSS**



**Detection of Supernova IIN:**  
PTF12csy; chance coin. ~2%  
Astrophys. J. (2015)



**iridium**



**ICECUBE**  
SOUTH POLE NEUTRINO OBSERVATORY

# IceCube online data processing pipeline



Farm for online reconstruction

Upgrade by T. Kintscher  
and A. Stasik (DESY)

Trigger  
2000 Hz

Level 1  
30 Hz

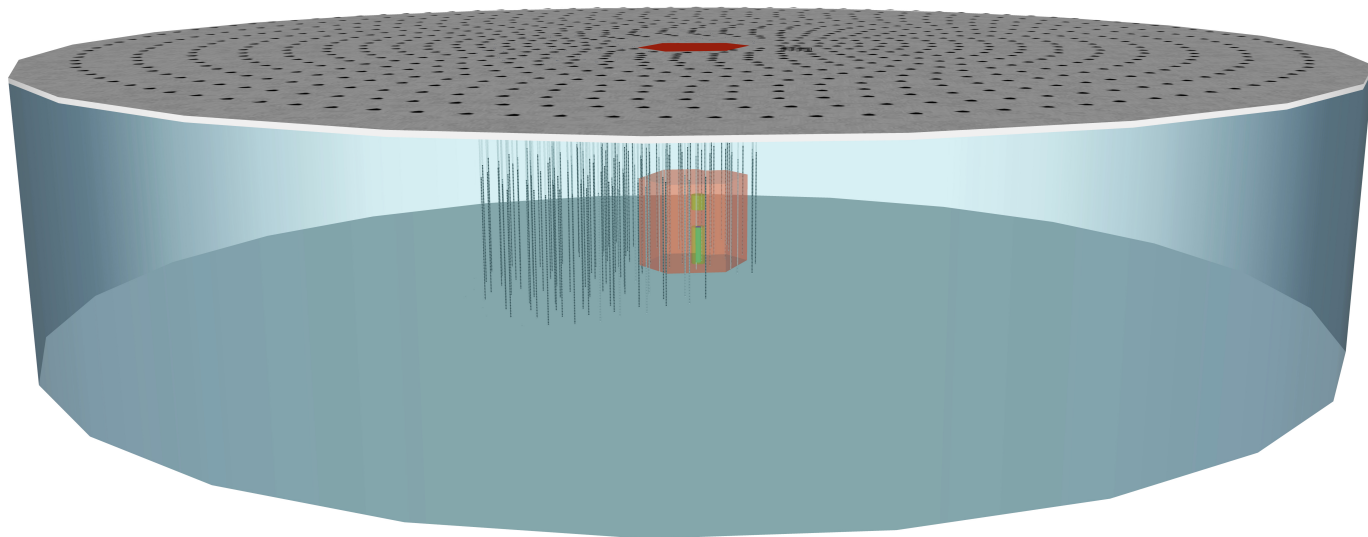
Neutrino-  
Level  
0.002 Hz

Single  
neutrinos  
send  
North

# Gen2: High Energy Baseline

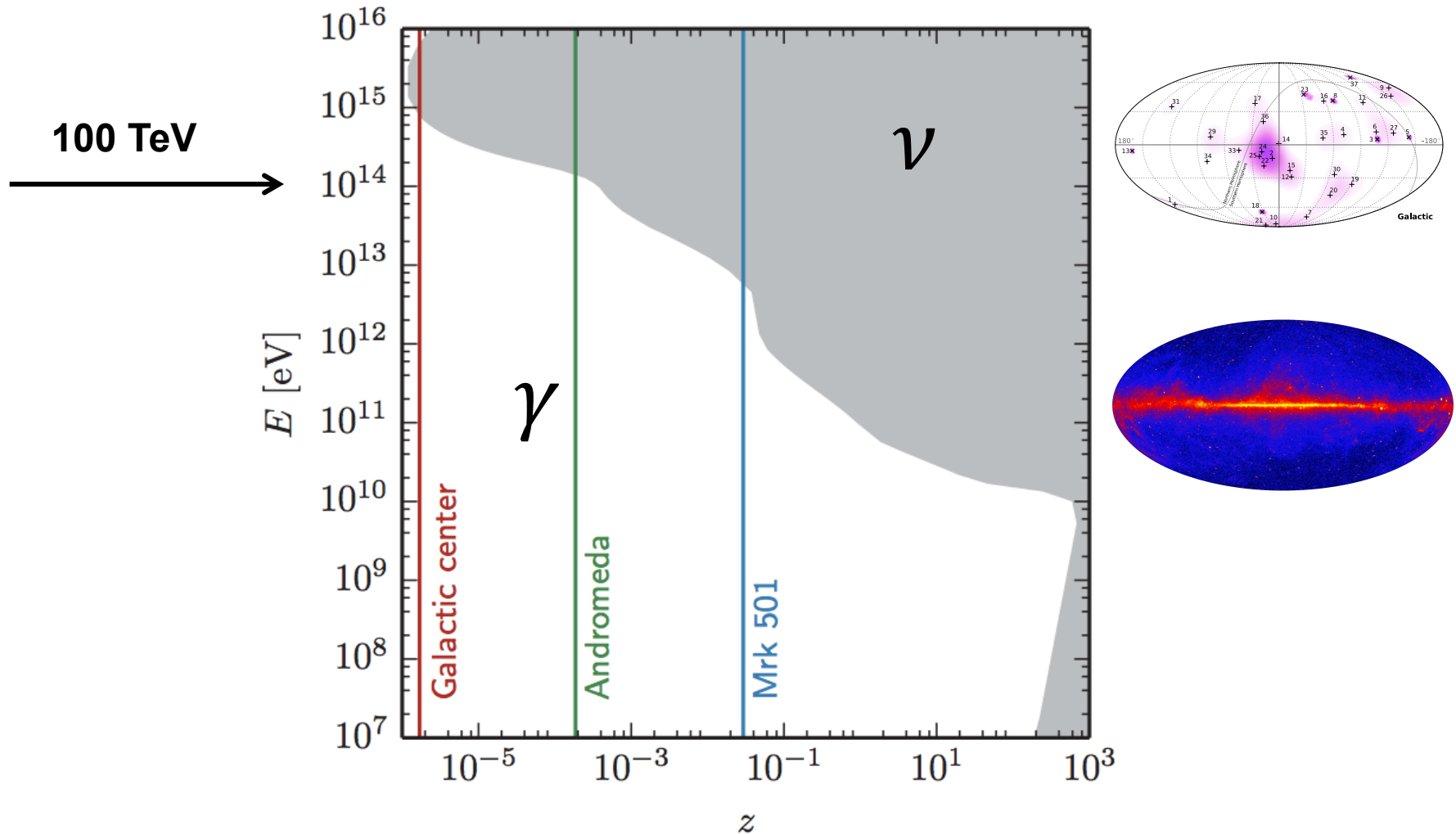
Conceptional design lead by DESY and U. Wisconsin.  
First full analysis pipeline available (J. van Santen, DESY)

Surface: 75 km<sup>2</sup> / 100 TeV threshold



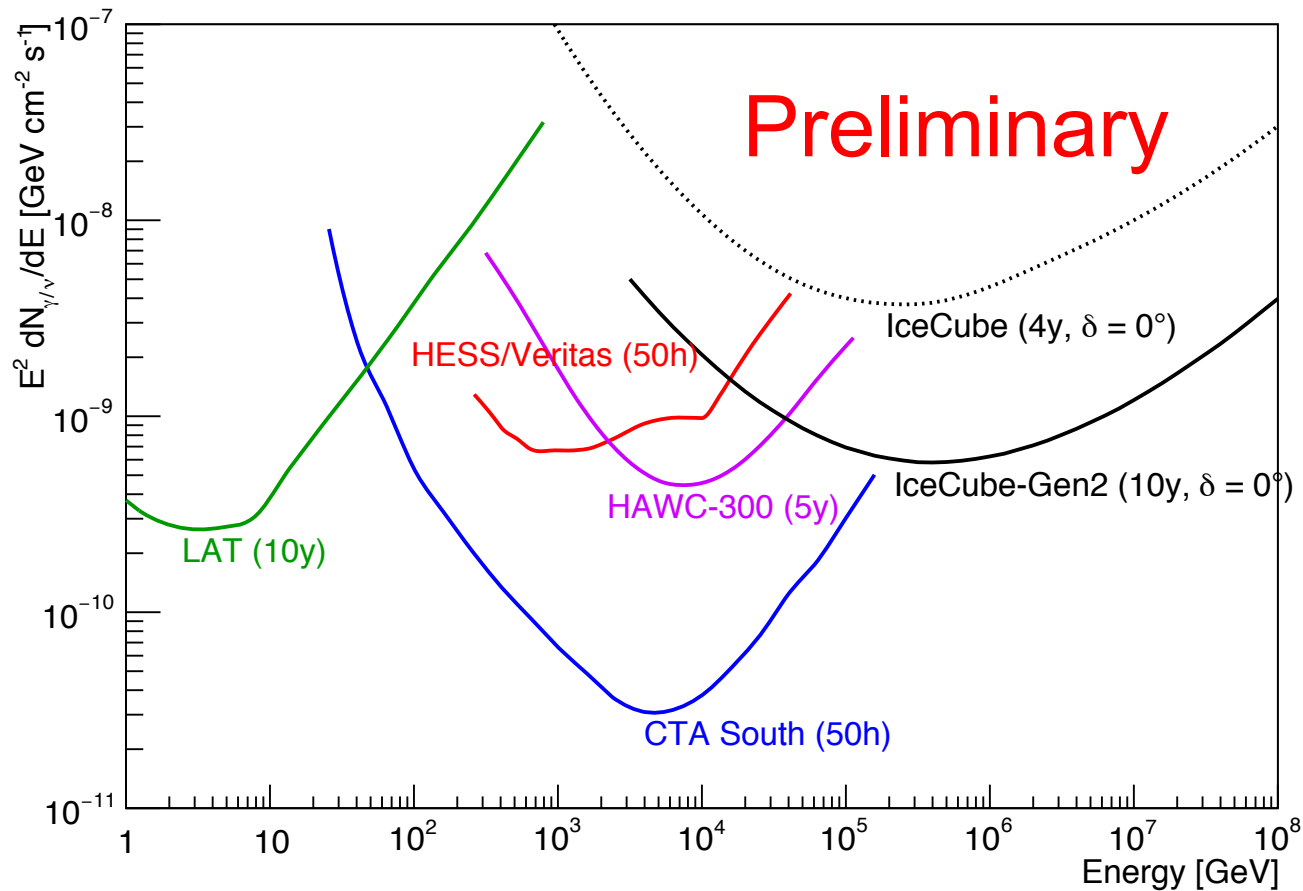
InIce: 120 strings, 240 m spacing

# Gen2: science case



**neutrino astronomy is THE window to the  
extragalactic Universe above 100 TeV**

# Point Source Sensitivity



# Conclusion

- Running experiments deliver top science
- CTA is ramping up well DESY key player
- IceCube-Gen2 starting to take shape

