

The GERDA Experiment: Search for the Neutrinoless Double Beta Decay

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Magellan Workshop 2016
DESY Hamburg, 17-18 March 2016



**Universität
Zürich**^{UZH}

1. Matter antimatter asymmetry

- ▶ Origin and possible explanation

2. Neutrinos Beyond the Standard Model

- ▶ Is the neutrino a **Majorana particle**?
- ▶ Absolute neutrino mass scale and hierarchy
- ▶ Baryogenesis via leptogenesis

3. Neutrinoless double beta decay

- ▶ Prime avenue of neutrino research

4. The GERDA Experiment

- ▶ Set-up and detection method
- ▶ Phase I results
- ▶ Upgrade to Phase II



Ettore Majorana

All galaxies and stars consist of baryons.

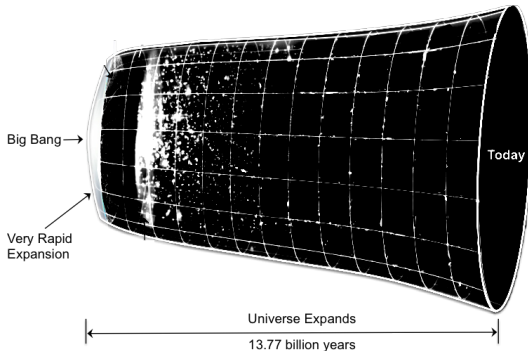
- ▶ Universe completely **matter dominated**.
- ▶ Antimatter exists only as a product of high-energy particle collisions.
- ▶ Antimatter regions of space would be detectable by gamma rays from annihilation reactions along its boundary.



Large Magellan Cloud. Source: nasa.gov

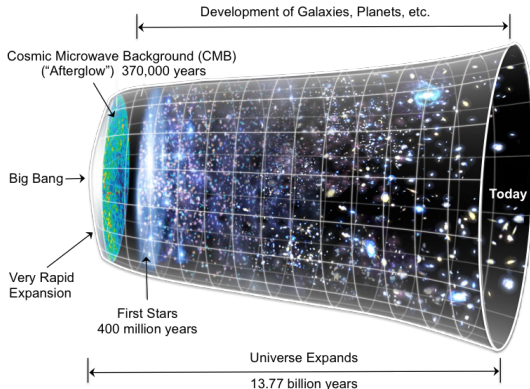
According to the Standard Model of Particle Physics:

- ▶ Early Universe after Big Bang completely symmetrical.
- ▶ **Equal** amounts of matter and antimatter.
- ▶ Quark and antiquarks annihilate in pairs, producing photons: $n_q, n_{\bar{q}} \approx n_\gamma$.
- ▶ Leading to the complete **annihilation** of a perfectly symmetrical early universe.
- ▶ Only radiation is left.



In reality:

- ▶ Slight **asymmetry** introduced in the early stages of the universe.
- ▶ **Surplus** of matter in comparison to antimatter.
- ▶ Quark and antiquarks annihilate in pairs, producing photons: $n_q, n_{\bar{q}} \approx n_\gamma$.
- ▶ After annihilation, low fraction of matter **survives**.
- ▶ All the galaxies and stars descend from this small **excess of matter**.



Origin of asymmetry could lie in neutrino nature.

- ▶ In 1937, **Ettore Majorana** suggested that neutrino could be its own antiparticle.
- ▶ Recent discovery of neutrino **oscillations** establish non-zero **mass** of neutrinos.

Explanation through Standard Model extension:

- ▶ See-saw mechanism introduces **right-handed neutrinos**.
- ▶ CP violating decays of these particles spontaneously generate leptons resulting in a **lepton asymmetry** in the early universe.
- ▶ **Baryogenesis via leptogenesis** could lead to the observed matter antimatter asymmetry.
[Phys. Lett. B 174: 45]

Investigation of neutrino nature:

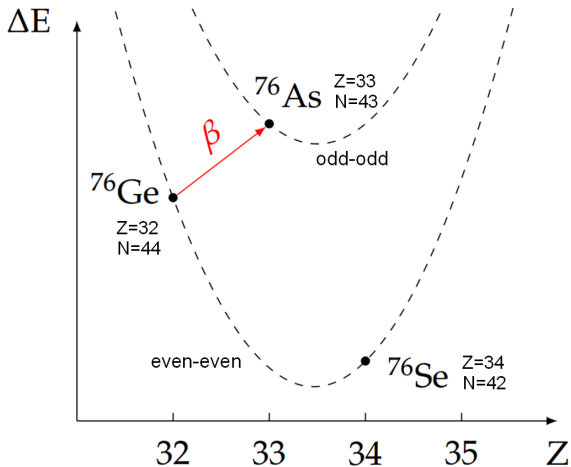
- ▶ Prime avenue is **neutrinoless double beta decay**.



Ettore Majorana

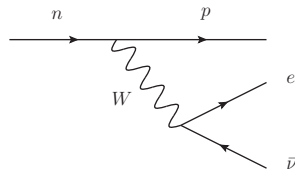
Decay channels of ^{76}Ge :

- Beta decay β^- to ^{76}As energetically **forbidden**.

Beta decay β^-

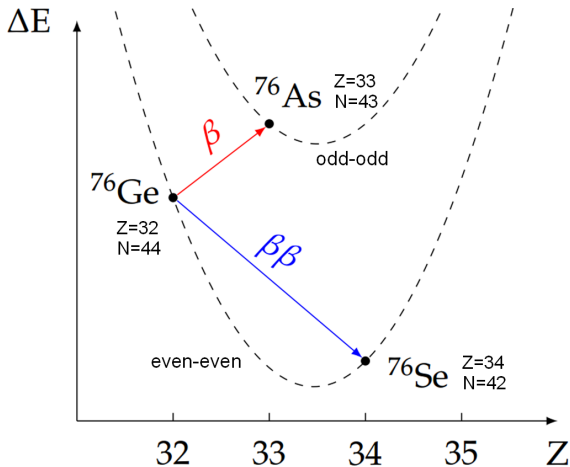
- Neutron decays to

$$\Rightarrow \begin{cases} \text{proton} \\ \text{electron} \\ \text{antineutrino} \end{cases}$$



Decay channels of ^{76}Ge :

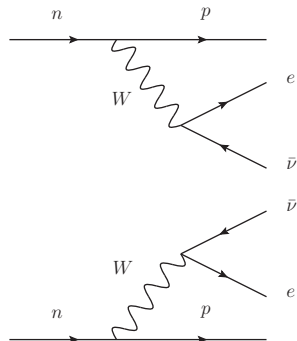
- ▶ Double beta decay $\beta\beta$ to ^{76}Se **allowed**.



Double beta decay $\beta\beta$

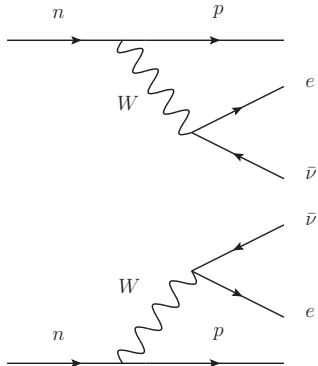
- ▶ 2 neutrons decay to

$$\Rightarrow \begin{cases} 2 \text{ protons} \\ 2 \text{ electrons} \\ 2 \text{ antineutrinos} \end{cases}$$



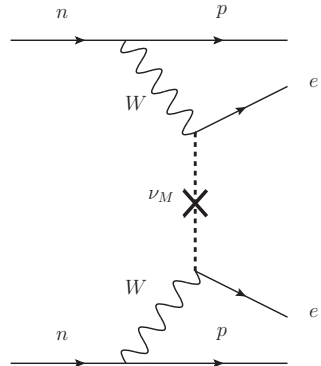
Double beta decay $2\nu\beta\beta$:

- ▶ **Rarest** observed decay.
- ▶ Possible for Dirac and Majorana neutrinos, SM $\Delta L=0$.
- ▶ **2 neutrinos** in the final state.



Neutrinoless double beta decay $0\nu\beta\beta$:

- ▶ **Postulated** decay channel.
- ▶ Involved Majorana neutrinos **annihilate** off-shell, non-SM $\Delta L=2$.
- ▶ **0 neutrinos** in the final state.



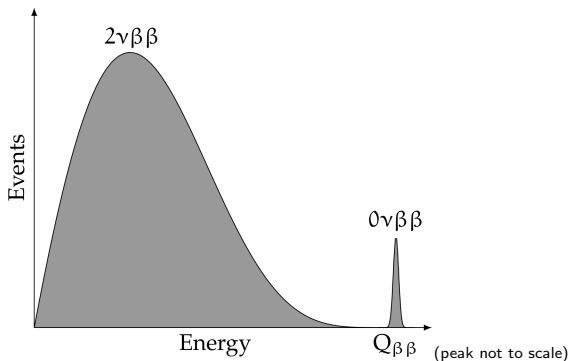
Summed **energy spectrum** of final state electrons:

Double beta decay $2\nu\beta\beta$:

▶ Continuum

Neutrinoless double beta decay $0\nu\beta\beta$:

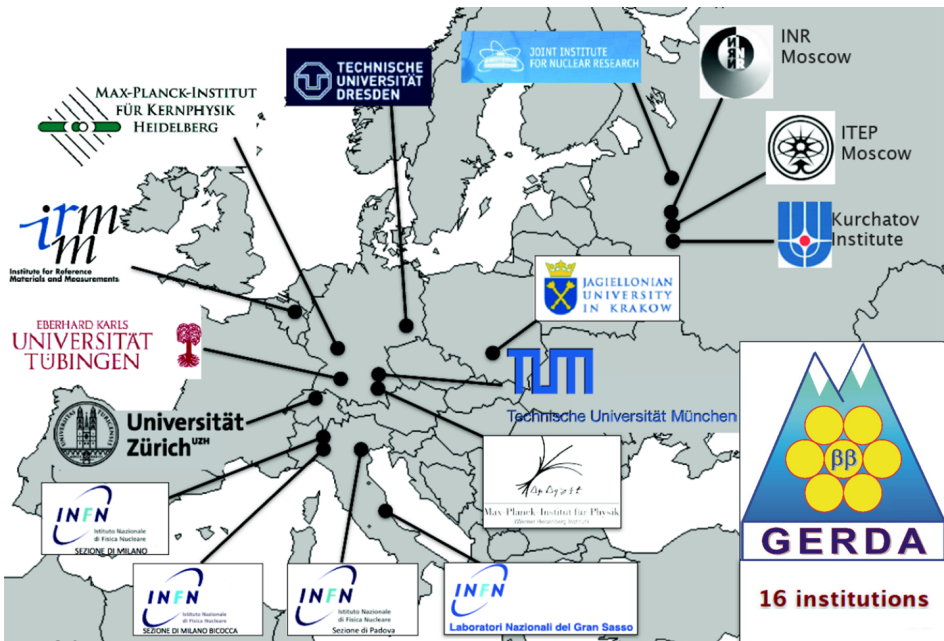
▶ Peak at $Q_{\beta\beta} = 2039$ keV



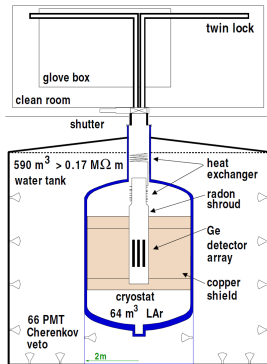
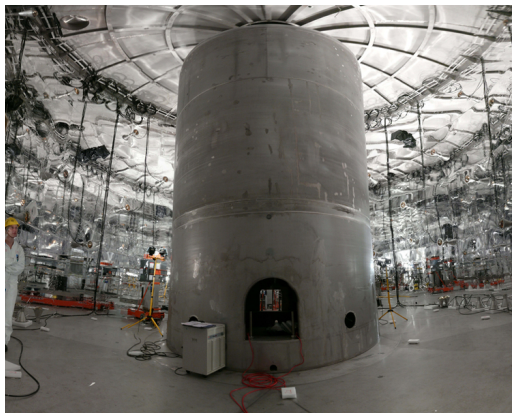
Consequences of discovery:

- ▶ Neutrinos have Majorana mass component.
- ▶ Rate of decay clue on absolute neutrino mass and hierarchy.
- ▶ Important step towards solving enigma of matter antimatter asymmetry.

The GERDA Collaboration



- ▶ **Germanium Detector Array** directly submerged in **liquid argon (LAr)** cryostat.
- ▶ Surrounded by **Cherenkov muon veto** water tank.
- ▶ Located underground in Hall A of **Laboratori Nazionali del Gran Sasso (LNGS)**.
- ▶ 1400 m overburden (μ flux $\sim 1 \text{ m}^{-2}\text{h}^{-1}$).

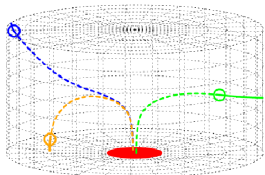


Low activity LAr, Water

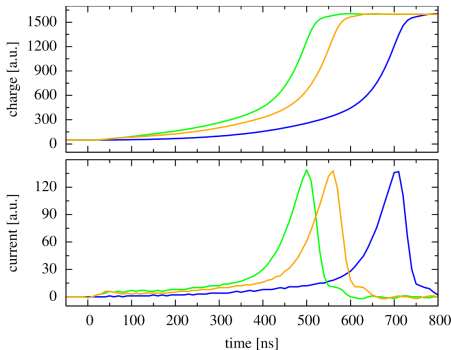
| Material | ^{208}Tl Activity [$\mu\text{Bq/kg}$] |
|---------------------|---|
| Rock, concrete | 3000000 |
| Stainless steel | ~ 5000 |
| Cu (NOSV), Pb | < 20 |
| Purified water | < 1 |
| LN_2 , LAr | ~ 0 |

Germanium detectors:

- ▶ Concept: Sources = Detectors. Isotopic enrichment to $\sim 86\%$.
- ▶ Reverse bias high voltage: Measurement of **ionisation** energy.
- ▶ Excellent **energy resolution**: $\sim 1.5\%$ FWHM at $Q_{\beta\beta}$.
- ▶ **Pulse Shape Discrimination (PSD)**
 - \Rightarrow $\left\{ \begin{array}{l} \text{single site events (e.g. } 0\nu\beta\beta) \\ \text{multi site events (e.g. Compton scattered } \gamma) \\ \text{surface events (e.g. } \alpha/\beta \text{ events)} \end{array} \right.$



- ⋯⋯⋯ anode
- cathode
- electrons
- - - holes
- ⊙ interaction point



Phase II String

Rare decay measurement requires low background.

- ▶ Operation of the experiment **underground**.
- ▶ Background suppression with **water** and **LAr shielding**.
- ▶ **Active veto** for cosmic muons and external radiation.
- ▶ **Minimise** radioactivity of materials close to detectors.
- ▶ **Pulse Shape Discrimination (PSD)**

Background limited scenario:

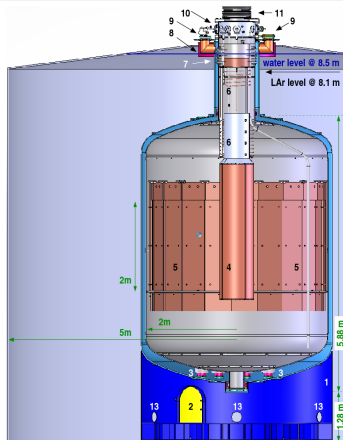
$$T_{1/2}^{0\nu} \propto \sqrt{\frac{M \cdot t}{\Delta E \cdot BI}}$$

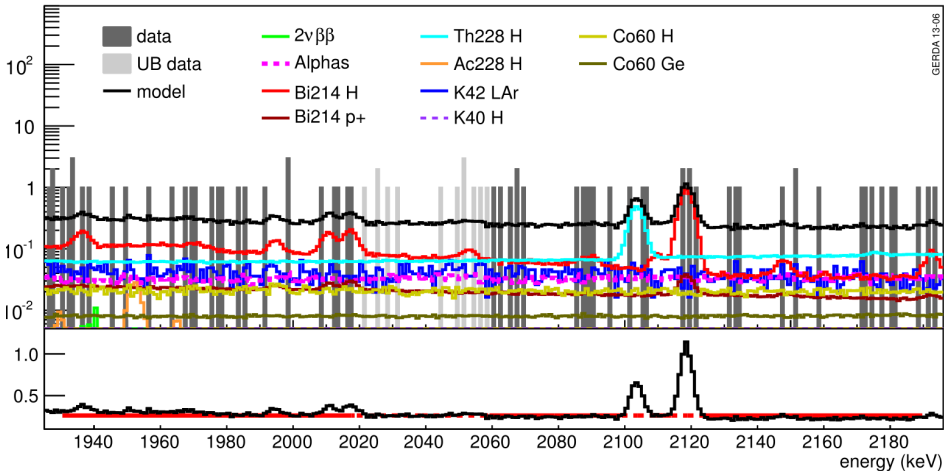
$M \cdot t$: exposure, ΔE : energy resolution, BI: background index.

Zero background regime:

$$T_{1/2}^{0\nu} \propto M \cdot t$$

- ▶ Goal: Achieve **zero background regime** to enable proportional scaling with exposure.





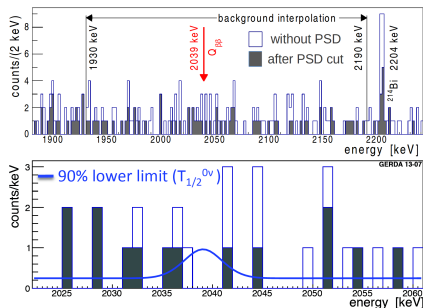
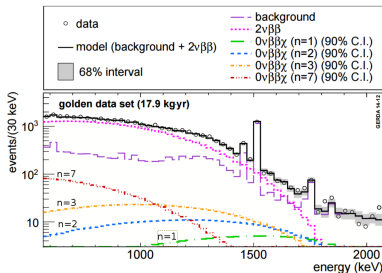
- ▶ Phase I background best fit **background model** and individual contributions.
- ▶ Fit performed **before unblinding** the 40 keV window around $Q_{\beta\beta}$.
- ▶ $BI = 1.0(1) \cdot 10^{-2}$ counts/(keV·kg·yr). Design goal fulfilled.
- ▶ Background **one order of magnitude better** than previous experiments.

Analysis of $0\nu\beta\beta$ decay [Phys. Rev. Lett. 111 (2013) 122503]

- ▶ 10 germanium detectors, 21.6 kg·yr exposure.
- ▶ Result: **No signal excess.**
(expected: 2.5, observed: 3 events)
- ▶ **Lower limit** on $T_{1/2}^{0\nu}$ of ^{76}Ge :
 $T_{1/2}^{0\nu} > 2.1 \cdot 10^{25}$ yr with 90% C.L.

Combined with HDM and IGEX:

$$T_{1/2}^{0\nu} > 3.0 \cdot 10^{25} \text{ yr with 90\% C.L.}$$

Analysis of $2\nu\beta\beta$ decay and $0\nu\beta\beta$ with Majoron Emission [Eur. Phys. J. C 75 (2015)]

- ▶ Measured half life:
 $T_{1/2}^{2\nu} = (1.926 \pm 0.095) \cdot 10^{21}$ yr
- ▶ Compatible with [J. Phys. G40 (2013) 035110].
- ▶ Search for $0\nu\beta\beta$ decay with Majoron emission performed for spectral index $n = 1, 2, 3, 7$.
- ▶ **No signal found**, limits of $O(10^{23})$ yr on half-lives.

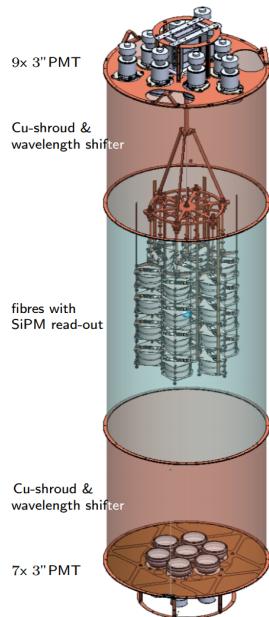
Germanium detector array upgrade:

- ▶ Addition of more **Broad Energy Germanium Detectors**
⇒ **Enhanced** PSD capabilities.
- ▶ Phase II in total **40 detectors** on **7 strings**.

LAr hybrid veto system:

- ▶ Installation of Photomultiplier Tubes (**PMTs**) and optical fibre curtain coupled to Silicon Photomultipliers (**SiPMs**).
- ▶ Detection of **LAr scintillation light** to reject external background events.

| Phase | Active Mass [kg] | BI [counts/(keV·kg·yr)] | $T_{1/2}^{0\nu}$ Sensitivity [yr] |
|---------------|---------------------|----------------------------|--------------------------------------|
| I (finished) | 15 | 10^{-2} | $2.1 \cdot 10^{25}$ |
| II (expected) | 35 | 10^{-3} | $1.4 \cdot 10^{26}$ |



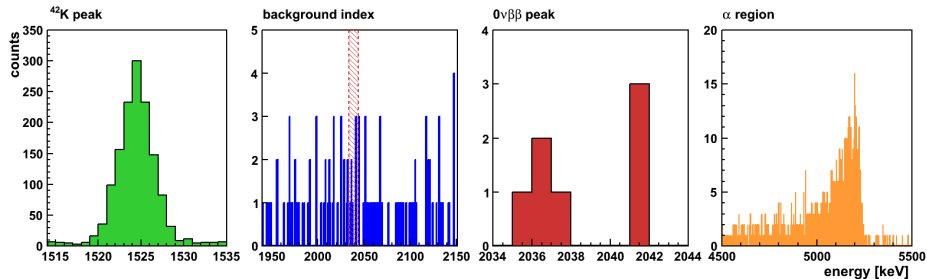
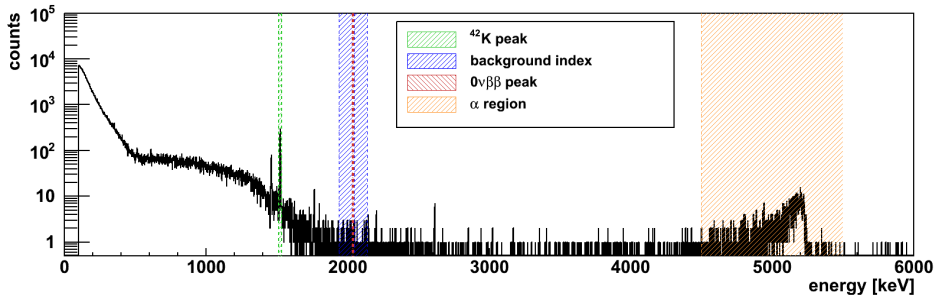
GERDA Phase I Results:

- ▶ Limit on the $0\nu\beta\beta$ decay half life: $T_{1/2}^{0\nu} > 2.1 \cdot 10^{25}$ yr (90% C.L.)
- ▶ Limit on the effective neutrino mass: $m_{\beta\beta} < 0.2-0.4$ eV (90% C.L.)

GERDA Phase II Outlook:

- ▶ Expected background at $Q_{\beta\beta}$: 10^{-3} counts/(keV·kg·yr).
- ▶ Expected sensitivity on the $0\nu\beta\beta$ decay half life: $T_{1/2}^{0\nu} \sim 1.4 \cdot 10^{26}$ yr.
- ▶ Expected limit on the effective neutrino mass: $m_{\beta\beta} \sim 0.1$ eV.
- ▶ Upgrade of experiment has been **finished**.
- ▶ Phase II **data taking** has begun December 2015.



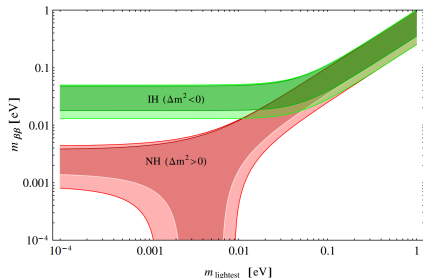


The mass mechanism

- ▶ For light Majorana ν exchange:

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q, Z) |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$

- ▶ $G^{0\nu}(Q, Z)$ = Phase Space integral
- ▶ $|M^{0\nu}|^2$ = nuclear matrix element
- ▶ $\langle m_{\beta\beta} \rangle^2 = \sum_i U_{ei}^2 m_i$ = effective ν mass
- ▶ U_{ei} = PMNS mixing matrix elements



Phys. Rev. D90 (2014) 033005

Experimental sensitivity:

- ▶ Number of signal events:

$$n_S = \frac{1}{T_{1/2}^{0\nu}} \cdot \frac{\ln 2 \cdot N_A}{m_A} \cdot f_{76} \cdot \varepsilon \cdot M \cdot t$$

- ▶ Number of background events:

$$n_B = BI \cdot \Delta E \cdot M \cdot t$$

where: f = enrichment fraction

N_A = Avogadro number

m_A = atomic mass

ε = total efficiency

M = detector mass

t = live time

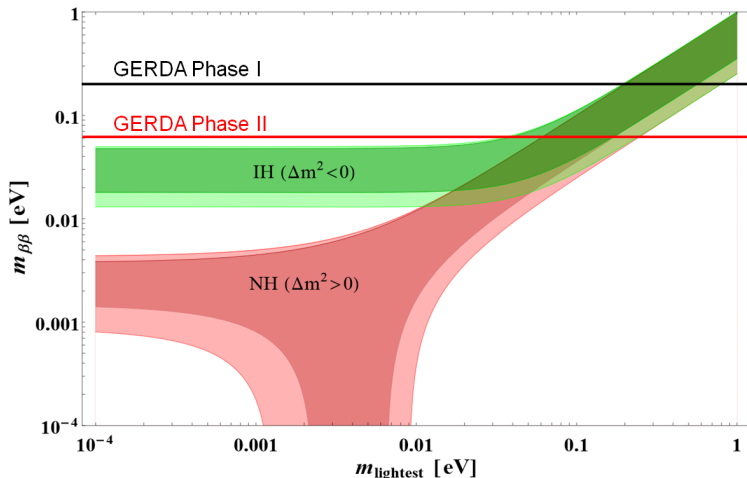
$M \cdot t$ = exposure

BI = Background Index

ΔE = energy resolution

Neutrino Mass Hierarchy Limits with GERDA:

- ▶ GERDA Phase I published.
- ▶ GERDA Phase II projected.

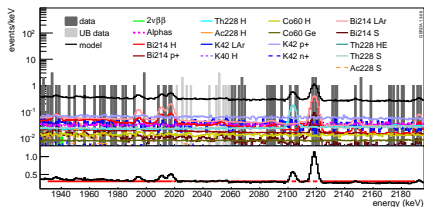
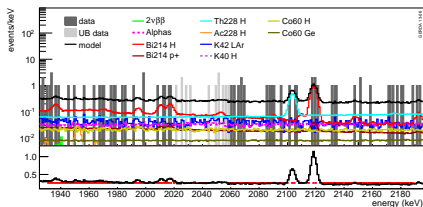


Background models

- ▶ Minimum model containing only known and visible background sources
- ▶ Alternative (maximum) model containing the same isotopes but more possible locations



Both models predict a flat background at $Q_{\beta\beta}$



Analysis recipe:

- ▶ Fit with Gaussian peak and flat background in the 1930-2190 keV region, excluding known gamma peaks at 2104 (^{208}Tl SEP) and 2119 keV (^{214}Bi).

| PSD | Dataset | Obs. | Exp. bkg |
|-----|---------|------|----------|
| no | Golden | 5 | 3.3 |
| | Silver | 1 | 0.8 |
| | BEGe | 1 | 1.0 |
| yes | Golden | 2 | 2.0 |
| | Silver | 1 | 0.4 |
| | BEGe | 0 | 0.1 |

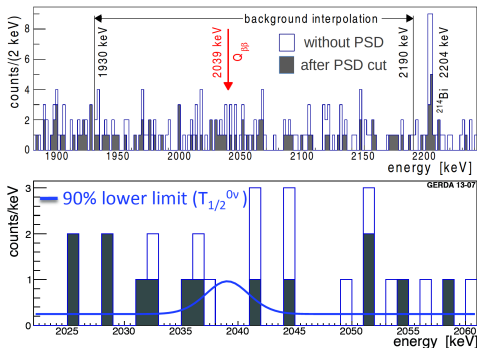
Profile Likelihood Method

- ▶ best fit $N^{0\nu} = 0$
- ▶ No excess of signal over bkg
- ▶ 90% C.L. lower limit:

$$T_{1/2}^{0\nu} > 2.1 \cdot 10^{25} \text{ yr}$$

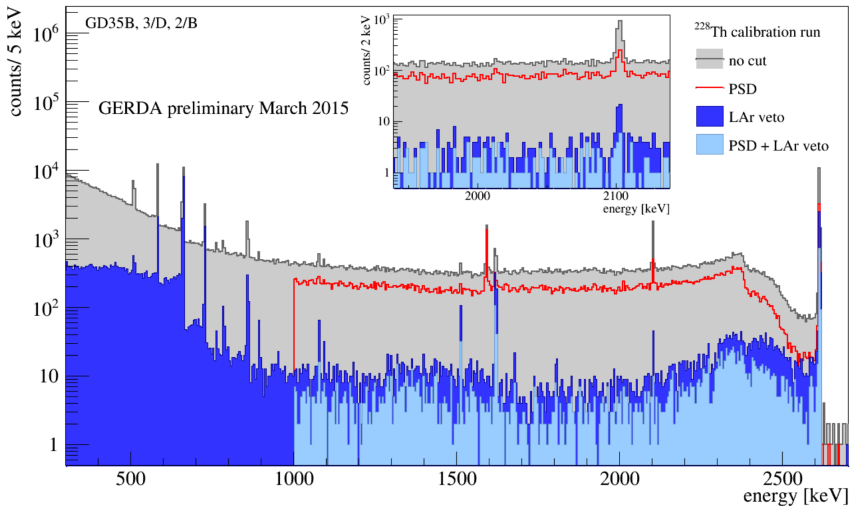
- ▶ Median sensitivity: $2.4 \cdot 10^{25} \text{ yr}$

GERDA Collaboration, Phys. Rev. Lett. 111 (2013) 122503



Bayesian Approach

- ▶ Flat prior for $1/T_{1/2}^{0\nu}$ in $[0; 10^{-24}] \text{ yr}^{-1}$
- ▶ best fit $N^{0\nu} = 0$
- ▶ 90% credibility interval:
 $T_{1/2}^{0\nu} > 1.9 \cdot 10^{25} \text{ yr}$
- ▶ Median sensitivity: $2.0 \cdot 10^{25} \text{ yr}$



- ▶ Spectrum from ^{228}Th calibration source
- ▶ 3 BEGe detectors (2 depleted, 1 enriched), 15 hours live time
- ▶ 15/16 PMTs and 7/16 SiPM working
- ▶ Background from ^{228}Th at $Q_{\beta\beta}$ suppressed by a factor ~ 100