

Axion Monodromy and the Weak Gravity Conjecture

Fabrizio Rompineve

Institute for Theoretical Physics, Heidelberg

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Talk based on [A. Hebecker, FR, A. Westphal \[arXiv:1512.03768\]](#)



Large field Inflation and axions

→ Slow-roll parameters: $\epsilon \sim (V'/V)^2 \ll 1, \eta \sim V''/V \ll 1, M_{pl} \equiv 1$.

- CMB Observables:

$$n_s - 1 = 2\eta - 6\epsilon \simeq 0.04$$

$$r = 16\epsilon \leq 0.08$$

→ Natural choice, $\epsilon \lesssim \eta \sim 10^{-2}$, requires transplanckian field displacements $\Delta\phi \gtrsim 4(\epsilon/10^{-2})^{1/2} M_{pl}$.

- Shift symmetry needed to ensure flatness over large range → **Axions** with large decay constants!
- Axions with large periodicities are also important in the mechanism of *Cosmological Relaxation* of EW scale (see [Witkowski's talk](#)).
- However, stringy implementations of Axion Inflation exhibits a certain degree of tuning/complexity.

→ Why? [[Rudelius/Montero,Uranga,Valenzuela/...](#)]

The Weak Gravity Conjecture (WGC)

[Arkani-Hamed, Motl, Nicolis, Vafa '06]

→ Extremal ($Q/M = 1$) black holes should be allowed to decay.

The Electric WGC

Consider a $U(1)$ gauge theory with coupling g coupled to gravity in 4d ($M_{Pl} \equiv 1$). Then there exists an electrically charged particle of mass m_{el} and charge q s.t.:

$$\frac{qg}{m_{el}} \gtrsim 1$$

The conjecture can be generalised to any p -form gauge theory with coupling $e_{p,d}$ in d dimensions coupled to $p - 1$ -dimensional objects with tension T_{p-1} .

$$\frac{e_{p,d} q M_d^{d/2-1}}{T_{p-1}} \gtrsim 1$$

→ The case $p = 0$ constrains/rules out models of Axion Inflation where a transplanckian range arises in the field space of one or more axions [see [Arthur's talk](#)].

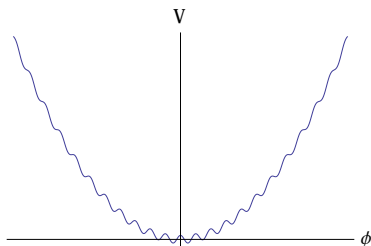
Axion Monodromy [Silverstein, Westphal '08/...

- However, transplanckian ranges can also be obtained by introducing *monodromies* to extend the compact field space of one axion.
- These models are *not* affected by the WGC for axions.

→ Phenomenological definition:

$$V = \underbrace{\frac{1}{2} m^2 \phi^2}_{\text{monodromy breaking shift symmetry}} + \alpha \cos(\phi/f)$$

→ For $\alpha/(m^2 f^2) > 1$, the potential exhibits local minima.



Domain walls in monodromy potential

→ The presence of local minima separated by “wiggles” can be effectively understood via a dual description:

$$\star F_4 \equiv \phi.$$

- The corresponding 3-form potential couples to 2-dimensional objects in $4d \Rightarrow$ Domain walls!

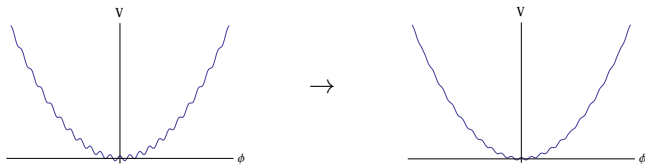
$$\mathcal{L} = \frac{1}{e^2} F_4^2 + \int_{DW} A_3.$$

- F_4 is quantized in units of e : $F_4 = ne, n \in \mathbb{Z}$.
- Across the membrane, the flux changes by one unit: $n \rightarrow n + 1$.
- The discretuum of points n corresponds to the values of the field ϕ in the cosine wells: $\phi \sim 2\pi n f$.
- Away from membranes, energy density: $\frac{1}{2} F_4^2 \sim \frac{1}{2} n^2 e^2 \rightarrow \frac{1}{2} m^2 \phi^2$, so $e = 2\pi m f$.

→ see Arthur's talk for Kaloper-Sorbo domain walls arising in axion monodromy

Constraints on Axion Monodromy, pt. 1

- Apply WGC for 3-form potential and domain walls.
- **WGC:** $T_{DW} \lesssim e \Rightarrow \alpha \lesssim m^2$.
- Tension of the heaviest domain walls: $T_{DW} \sim \sqrt{V} \Delta\Phi \sim \sqrt{\alpha} f$.
- The WGC requires a small tension, which can be obtained by lowering the height of the wiggles, i.e. α .
- **Crucial:** this is precisely what is needed to have slow roll inflation!
- ⇒ No tension between WGC and slow-roll ⇒ no constraint from the electric WGC!



- see Ibanez, Montero, Uranga, Valenzuela '15 for constraints on KS domain walls for relaxion

The magnetic WGC

Weakly coupled gauge theories with cutoff Λ and magnetic monopole of charge q_m break down at a scale:

$$\Lambda \lesssim \frac{g}{q_m} \quad \text{parametrically smaller than } M_{Pl} \equiv 1!$$

- The minimally charged magnetic monopole should not be a black hole, i.e. $M_{mon} \sim \Lambda/e^2 \lesssim M_{BH} \sim R \sim \Lambda^{-1}$.
- **Strategy:** generalise this statement to any $(p+1)$ -form gauge theory with $d - (p+4)$ magnetic branes.

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2 Impose:

$$T_{d-(p+4)} \lesssim T_{BH} \sim M_d^{d-2} R^{p+1}.$$

Constraints on Monodromy Inflation, pt. 2

⇒ Specialise to domain walls in $4d$, i.e. $p = 2$:

$$\Lambda \lesssim e^{1/3}$$

→ **Caveat:** there are no (-2) -branes! Assume that the bound can be obtained by “analytical continuation” in p, d .

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Take home

- 1 **Electric** WGC does not constrain monodromy inflation [see also Ibanez, Montero, Uranga, Valenzuela '15].
- 2 **Magnetic** WGC limits the field range, but does not forbid phenomenological realisations.

How does ST fulfil the WGC?

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\rightarrow EFT is valid up to KK-scale: $M_{KK} \sim 1/R$.

- Tension of a q -dimensional object descending from a p -brane in $10D$:
 $T_q \sim M_s^{p+1} R^{p-q} / g_s$.

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\rightarrow **Take home:** the magnetic WGC as the more fundamental constraint from a ST point of view, therefore we assume that it is valid for any p -form.

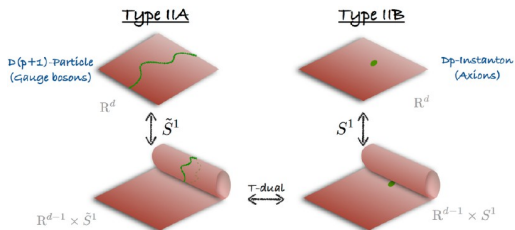
Problems with domain walls

General form of the WGC from extremality argument in d -dimensions:

$$\gamma_{p,d} T_p^2 \leq e_{p,d}^2 q^2 M_d^{d-2}$$

→ $\gamma_{p,d} = 0$ for axions ($p = 0$) and strings ($p = d - 2$). For domain walls ($p = d - 1$) the inequality cannot be solved.

→ Strategy for axions [Picture taken from Gary Shiu's talk at StringPheno 2015].



The WGC as a geometric constraint

- **Idea:** in string compactifications, the WGC translates into a constraint on the geometry of a chosen CY X
- **Our strategy:**
 - 1 Start with p -branes in 10D, compactify on p -cycle Σ of a CY X to get particles and gauge fields.

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$$\frac{V_X^{1/2} |q_\Sigma|}{V_\Sigma} \gtrsim 1,$$

where $|q_\Sigma|$ is the norm of the harmonic form related to Σ using the metric on X , V_X, V_Σ are volumes.

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- 3 Apply previously found constraint on the geometry to get the WGC for $q - p$ dim. objects in $4d$.

→ No need of string dualities! (see Brown, Cottrell, Shiu, Soler/Ibanez, Montero, Uranga, Valenzuela '15)

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1 Phenomenology:

- Models of axion monodromy (inflation/relaxation) exhibit low energy “wiggles”, which correspond to $4d$ domain walls.
- The WGC can be applied to these domain walls:
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2 String compactifications:

- In many cases, ST satisfies the WGC by lowering the cutoff of the EFT, i.e. the Kaluza-Klein scale.
 - Therefore, the magnetic WGC should be seen as the more fundamental constraint.
 - In this framework, the WGC can be formulated purely as a geometric constraint on the sizes of CY cycles.
- \rightarrow WGC can be extended to any p -dim object without need of string dualities!

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• Many thanks!

Application to Axion Inflation

- 1 For 0-form gauge fields, i.e. axions:

$$\mathcal{L}_{axions} \sim f^2 \int (\partial\phi)^2 + S_{inst}, \quad (1)$$

so the gauge coupling is $1/f$ and the tension is S_{inst} .

⇒ WGC for axions:

$$f \lesssim 1/S \lesssim 1, \quad (2)$$

because $S > 1$ in a calculable regime.

- Simplest model: single field natural inflation with $f > 1$ is ruled out!
- Models with two or more axions with subplanckian decay constants are also constrained.
- Monodromy models are not constrained, because they have a single axion with subplanckian decay constant.

WGC as a geometric constraint

w^i basis of $H^p(X, \mathbb{Z})$.

$$q_i^k = \int_{\Sigma_k} w^i = \int_X w^i \wedge w^k, \quad (3)$$

$$K_{ij} \equiv \int_X w_i \wedge \star w_j. \quad (4)$$

$$|q^\sigma|^2 \equiv K^{ij} q_i^\Sigma q_j^\Sigma \quad (5)$$