

CP-conservation in QCD and why only “invisible” axions work

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From a fundamental point of view, presumably the most fundamental parameters are given at a defining scale. In this sense,

Natural mass scales are the Planck mass and/or some scales suppressed by small coupling times that, i.e. GUT scale.

All smaller mass scales are better to come from symmetry arguments. Fermions by chiral symmetry and scalars by the Goldstone theorem.

Among Goldstone bosons, “Invisible” axion is the most interesting one.

1. Strong CP problem



- Because of instanton solutions of QCD, there exists an effective interaction term $G \tilde{G}$ -dual. It is the flavor singlet and the source solving the U(1) problem of QCD: 't Hooft, Phys. Rep. (1986).
- This term is physical, but leads to
- The strong CP problem, “Why is the $n\text{EDM}$ so small?”
- The remaining ‘natural solution’ is “invisible” axion as given in my title.



- The gluon interaction.

$$\mathcal{L} = \bar{\theta}\{G\tilde{G}\} \equiv \frac{\bar{\theta}}{64\pi^2} \epsilon^{\mu\nu\rho\sigma} G_{\mu\nu}^a G_{\rho\sigma}^a,$$

- The neutron mass term.

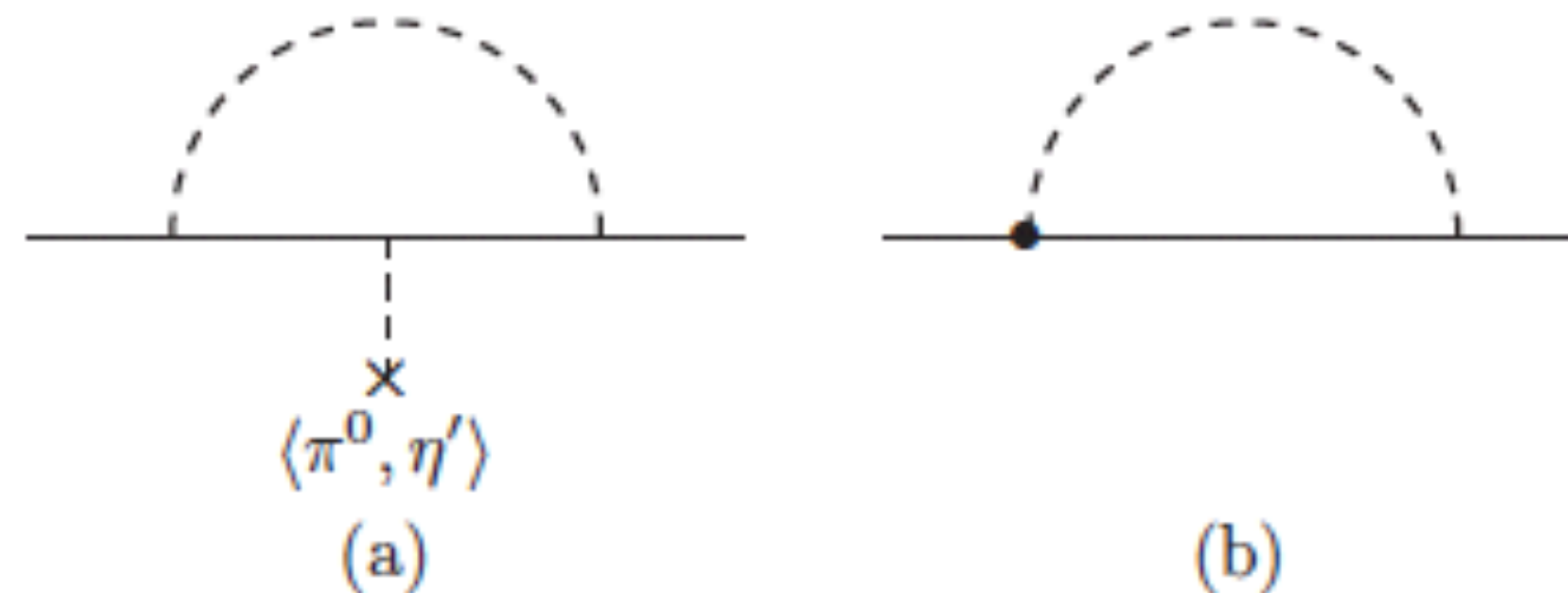
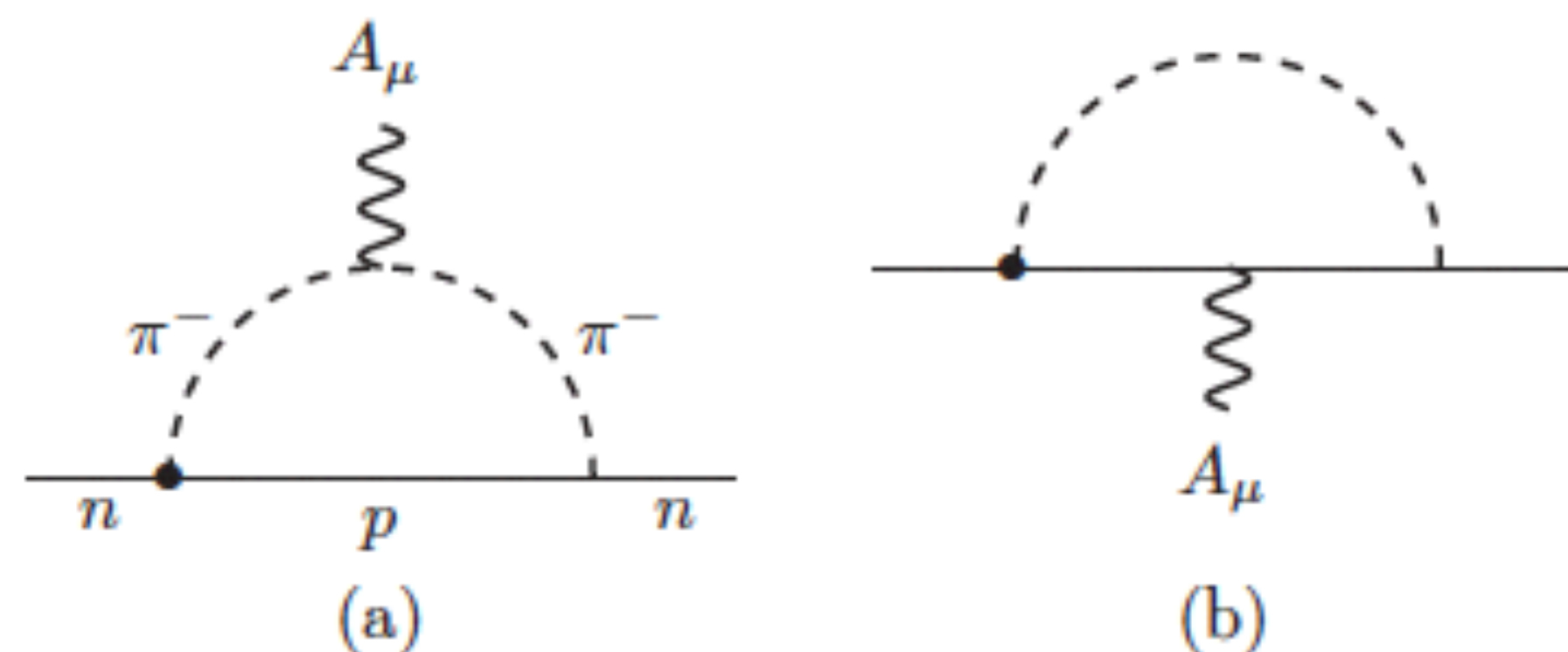


FIG. 4. Loop corrections for $\bar{n}n$ -meson coupling. Insertion of the CP violation effect by VEVs of π^0 and η' in (a). They can be transferred to one vertex shown as a bullet in (b). With this bullet, CP violation is present because of a mismatch between the CP -conserving RHS vertex and CP -violating LHS vertex.



- The neutron EDM term.



$$\frac{d_n}{e} = \frac{\overline{g_{\pi NN} g_{\pi NN}}}{4\pi^2 m_N} \ln\left(\frac{m_N}{m_\pi}\right),$$

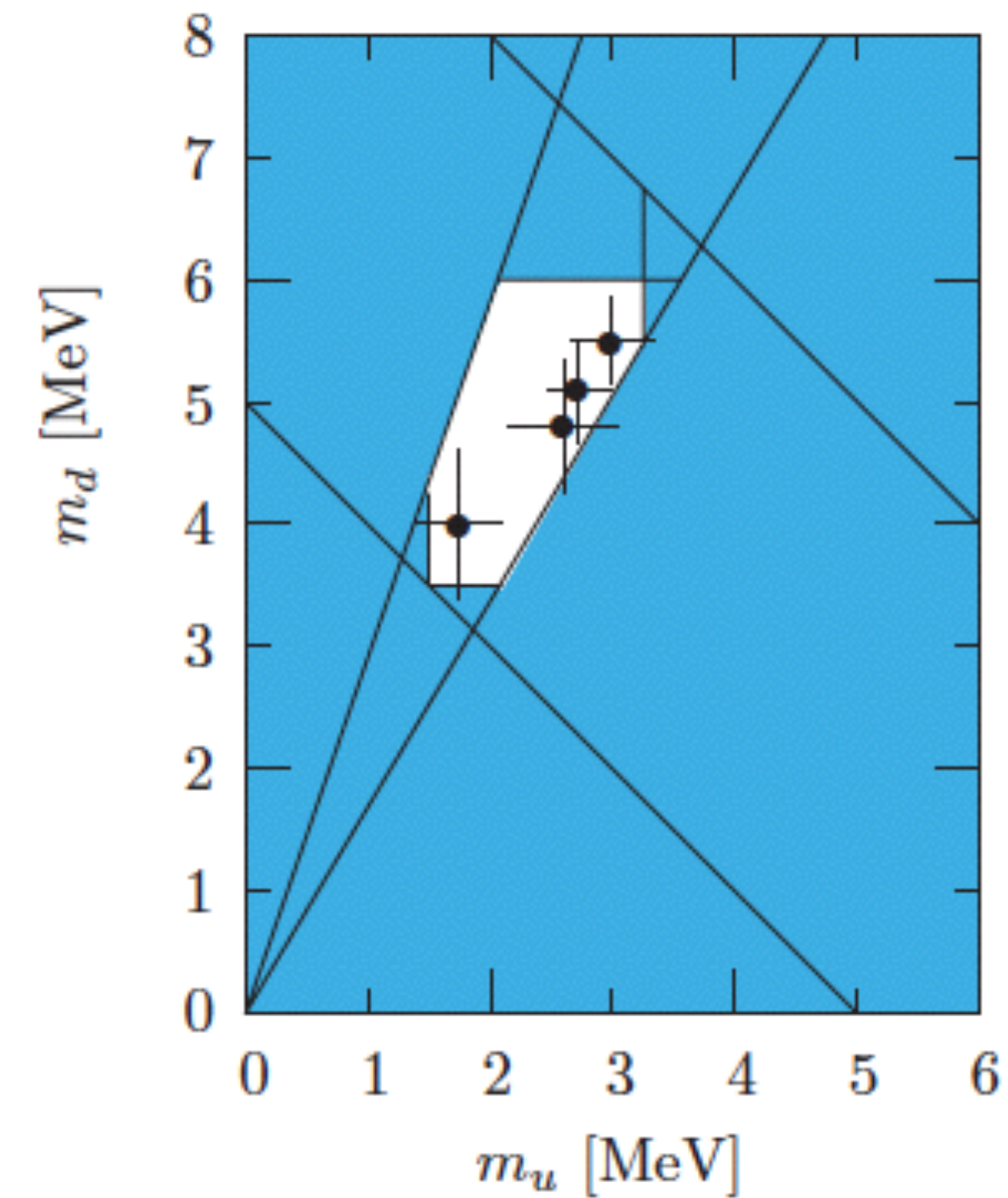
FIG. 5. Diagrams contributing to the NEDM with the bullet representing the CP violation effect. (a) is the physically observable contribution.

Pion VEV (CP violating)
in the meson L can give

$$\overline{g_{\pi NN}} = -\bar{\theta} \frac{Z}{(1+Z)} \simeq -\frac{\bar{\theta}}{3}.$$

Masses up quark:

PDG book,
Manohar-Sachrajda



Symmetry solution (natural solution, calculable solution):


Beg-Tsao, Georgi, MS, ..., Nelson, Barr:
But no compelling model with very small θ -bar



2. PQ symmetry



After Cronin et al paper, “Need for a theory of weak CP violation”: KM+...

- (1) by light colored scalar,
 - (2) by right-handed current(s),
 - (3) by three left-handed families,
 - (4) by propagators of light color-singlet scalars, and
 - (5) by an extra $U(1)$ gauge interaction.
- 
- By Kobayashi-Maskawa

CP violation by many Higgs doublets

At the time when the third quark family was not discovered, Weinberg tried to introduce the weak CP violation in the Higgs potential. Due to the GW theorem, he introduced two Higgs doublets, one coupling to the up quarks and the other coupling to the down quarks. His Higgs potential with multi Higgs fields is

$$V = \frac{1}{2} \sum_I m_I^2 \phi_I^\dagger \phi_I + \frac{1}{4} \sum_{IJ} \left\{ a_{IJ} \phi_I^\dagger \phi_I \phi_J^\dagger \phi_J + b_{IJ} \phi_I^\dagger \phi_J \phi_J^\dagger \phi_I + (c_{IJ} \phi_I^\dagger \phi_J \phi_I^\dagger \phi_J + \text{H.c.}) \right\}$$

Removed by
Peccei-Quinn

CP violation by Weinberg,
including Yukawa of H_u
and H_d couplings.
A ha, then there appears a
 $U(1)$ global symmetry. The
Peccei-Quinn symmetry.

1. PQWW axion reported by Weinberg and Wilczek at Ben Lee Memorial, Oct 1977
2. Calculable models(no axion), 1978
3. Invisible axion, 1979
4. Invisible axion as CDM, 1983
5. Axion detection, 1983 [2013]
6. Model-Ind. axion, 1985
7. Anomalous U(1) gauge symmetry, 1986
8. Axion-photon coupling from string compactification, 1988, 2014, 2016



Fine-tuning problem in invisible axion

KSVZ: $f \bar{Q}_R Q_L \sigma + \text{H.c.}$

VEV of sigma is f_a . No fine-tuning problem.

DFSZ: $-\mu_1^2 H_u^* H_u - \mu_2^2 H_d^* H_d + \lambda_1 (H_u^* H_u)^2 + \lambda_2 (H_d^* H_d)^2 + \dots$
 $(+M H_u H_d \sigma') + \lambda_{h\sigma} H_u H_d \sigma^2 + \text{H.c.}$

VEV of sigma for f_a and electroweak scale v needs some fine-tuning.

(Mixing term)/ $\lambda_{(1,2)}$ needs a fine-tuning of order 10^{-18} .

A similar issue for WarmDM axino was studied at Bonn: Dreiner-Staub-Ubaldi, 1402.5977 [hep-ph]

Supersymmetry

KN term: $\frac{1}{M} H_u H_d \sigma^2$

M is determined
from the theory.

This term is the definition
of the PQ symmetry.

2. “Invisible” axions



SU(2)xU(1) singlet houses the invisible axion.

KSVZ : $\mathcal{L}_Y = \bar{Q}_L Q_R \sigma + \text{H.c.}; \quad \langle \sigma \rangle = \frac{f_a}{\sqrt{2}},$

DFSZ : $\mathcal{L}_Y = \bar{q}_L u_R H_u + \bar{q}_L d_R H_d + \text{H.c.},$

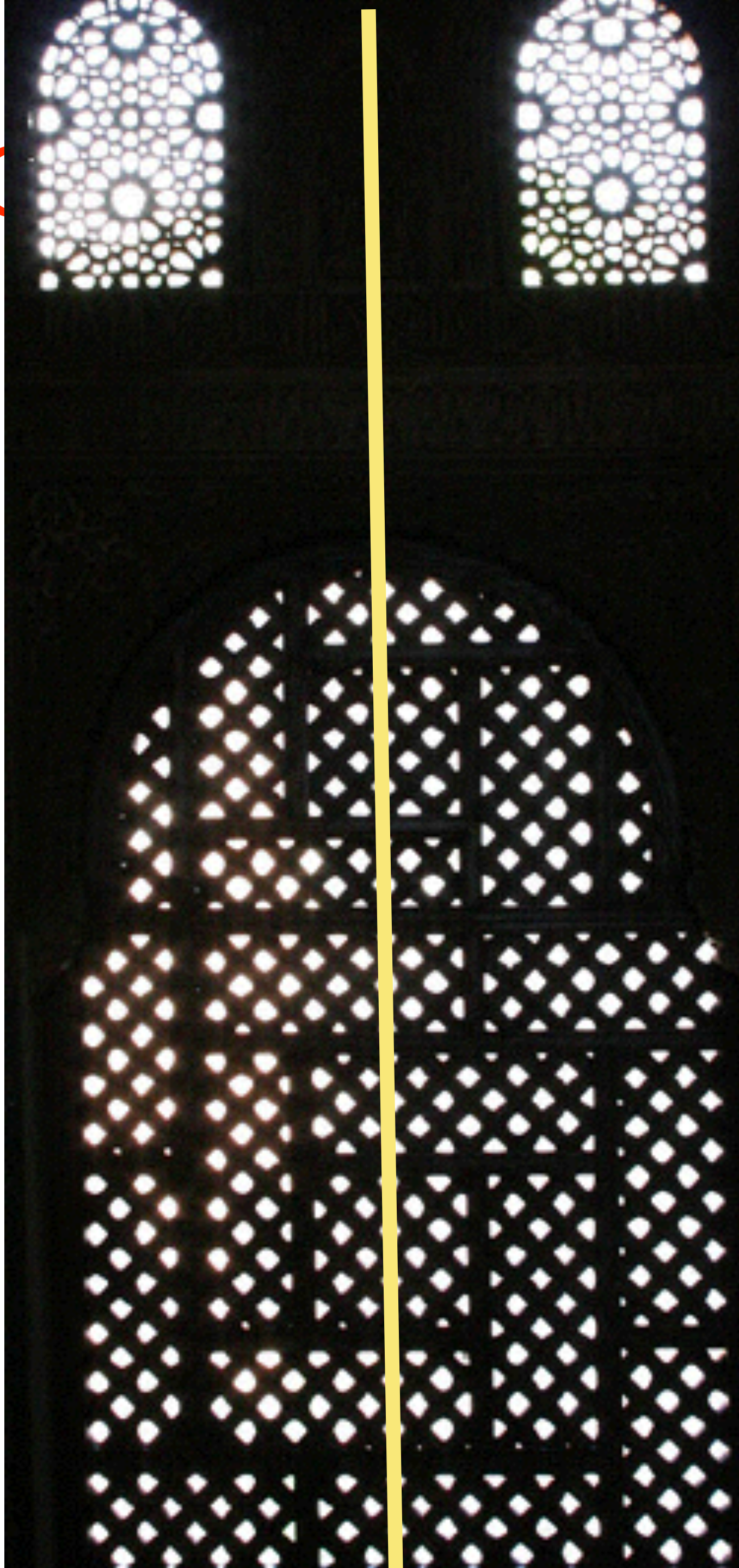
$V = H_u H_d \sigma^2 + \dots + \text{H.c.}; \quad \langle \sigma \rangle = \frac{f_a}{\sqrt{2}}$

$$\mathcal{L}_{\text{int}}^{\text{eff}} = c_1 \frac{(\partial_\mu a)}{f_a} \sum_q \bar{q} \gamma^\mu \gamma_5 q - \sum_q (\bar{q}_L m q_R e^{i c_2 a / f_a} + \text{h.c.}) + \frac{c_3}{32\pi^2 f_a} a G \tilde{G}$$

$$+ \frac{C_{aWW}}{32\pi^2 f_a} a W \tilde{W} + \frac{C_{aYY}}{32\pi^2 f_a} a Y \tilde{Y} + \mathcal{L}_{\text{leptons}},$$

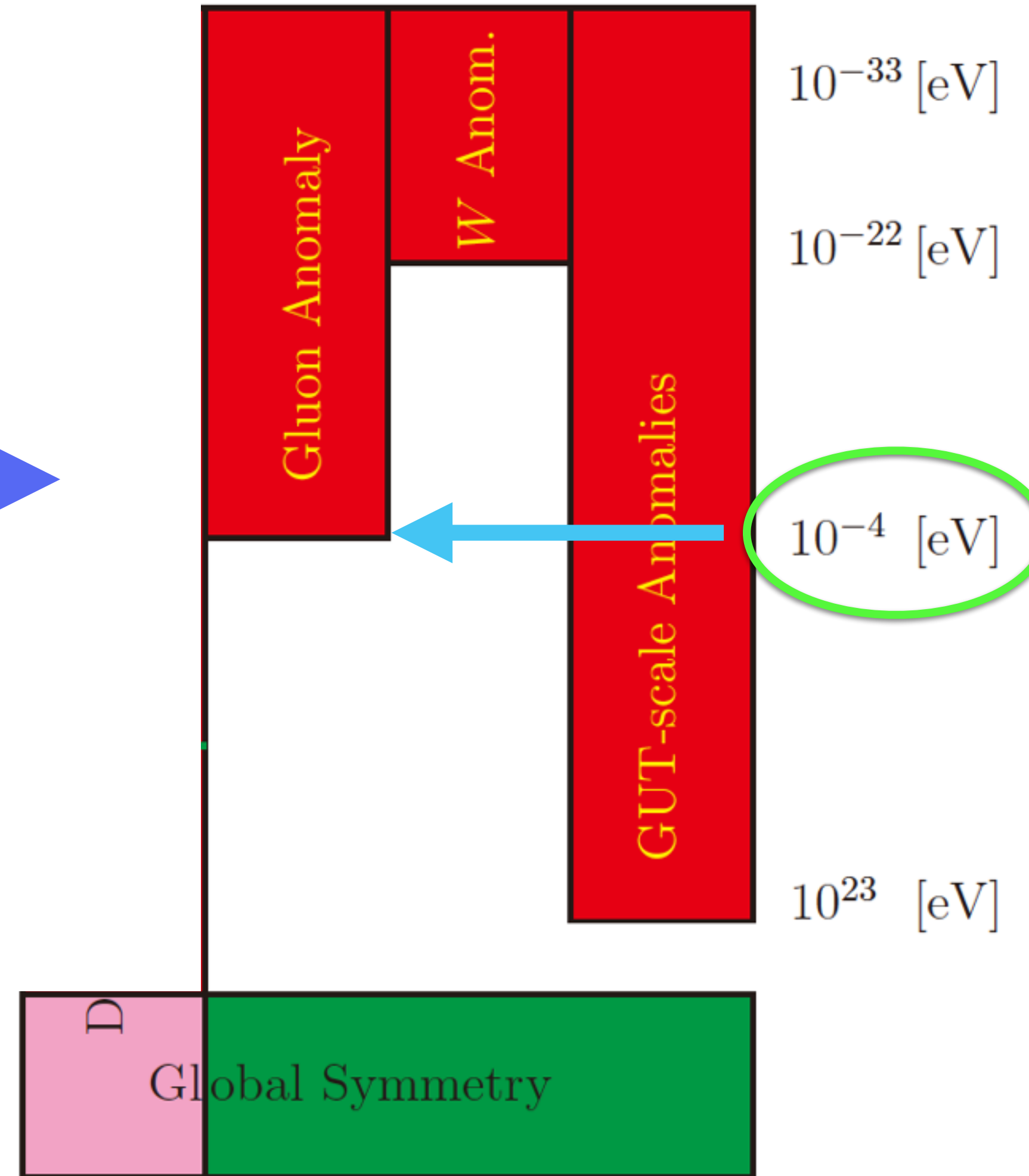
Light bosons are pseudo-Compton.

Pseudoscalars fit to this scheme. Somehow discrete symmetries (P) seem to work as the messengers to low and high energy scales. Mainly from the string compactification examples.



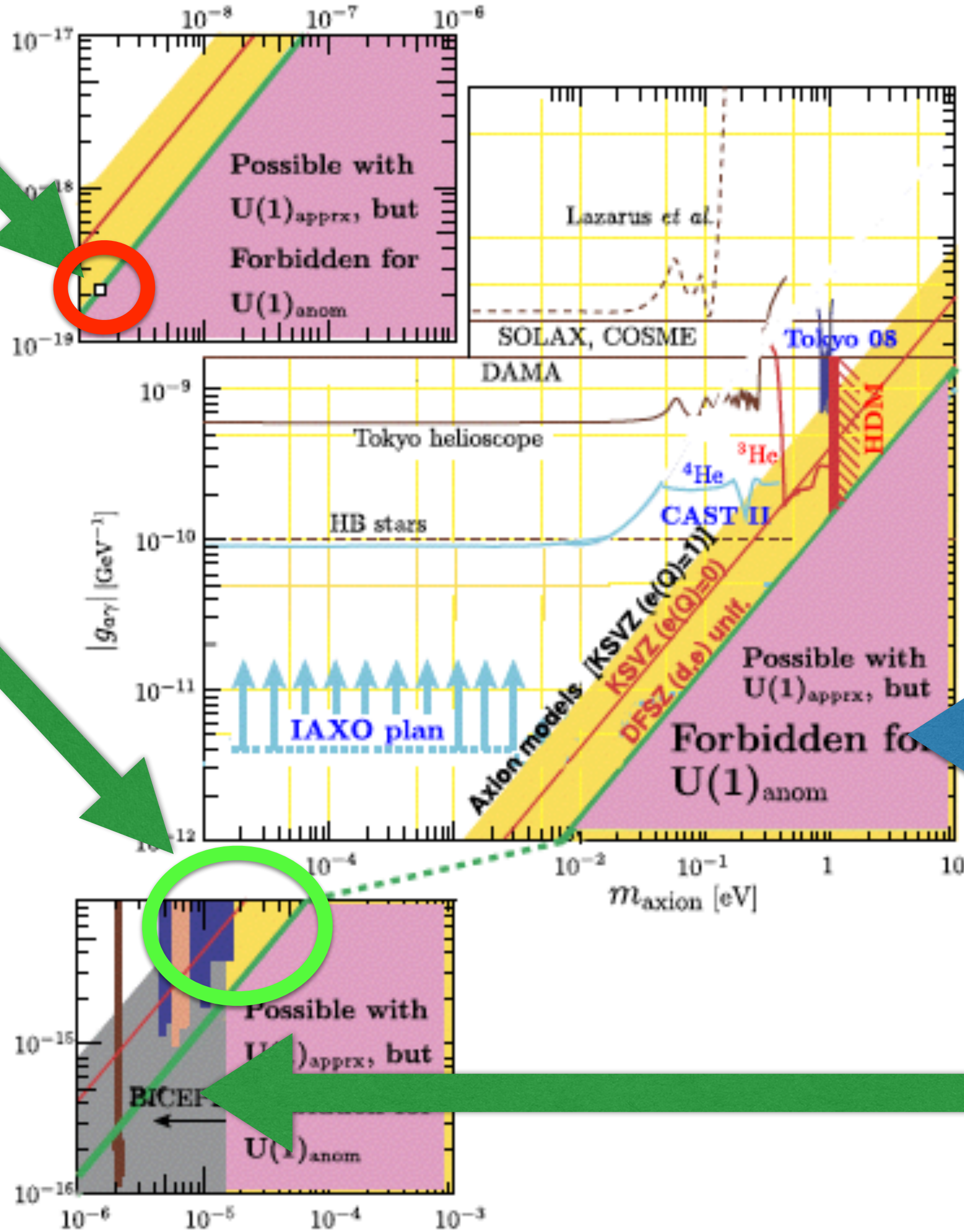
If this is absent,
it is called
axion. And
 $\theta=0$ is the
minimum.

Still some term in
 V is present with
discrete
symmetry, then
 $\theta=0$ is not
guaranteed to be
the minimum.



MI axion

A small allowed region by $U(1)_{anom}$



$g_{a\gamma}(= 1.57 \times 10^{-10} c_{a\gamma\gamma})$ vs. m_a plot

Kim-Semertzidis-Tsujikawa, Front. Phys. 2 (2014) 60

Kim-Nam, 1603.02145[hep-ph]

$U(1)_{anom}$ forbidden

If H_I is greater than f_a , there is the isocurvature constraint.

In SUSY, without extra small parameters, the following dim-4 W is the minimum example.

Summarized by
Weinberg operator:
[13.08.1979, Received]

$$\frac{1}{M} \ell \ell H_u H_u$$

gives ν mass

(Note: In the original image, a green box labeled 'L=2' has an arrow pointing to the $\ell \ell$ term, and a blue box labeled 'L=-2' has an arrow pointing to the $H_u H_u$ term.)

Kim-Nilles SUSY operator:
[24.11.1983, Received]

$$\frac{1}{M} S_1 S_2 H_u H_d$$

gives TeV scale μ term

(Note: In the original image, a green box labeled 'Q=2' has an arrow pointing to the $S_1 S_2$ term, and a blue box labeled 'Q=-2' has an arrow pointing to the $H_u H_d$ term.)

We neglected $S H_u H_d$ since the VEV of S at intermediate scale is too large for $SU(2)$ breaking.

Realized in seesaw:
Minkowski [13.04.1977, Published],
Yanagida [13-14 Feb 79, Conf. talk]

$$\ell_L H_u N_R$$

Realized in string comp.:
Many papers, ...

$$S_1 H_u X_{\text{doublet}}, S_2 H_d X'_{\text{doublet}}, \bar{Q}_L Q_R S_1, \dots$$

(Note: In the original image, yellow boxes with black borders contain the charges: '-1' below S_1 , '-1' below S_2 , '-1/2' above \bar{Q}_L , and '+1' above Q_R . Blue arrows point from these boxes to the corresponding terms in the equation.)

The Peccei-Quinn symmetry at low energy with SUSY is introduced in this way. Breaking the PQ symmetry in SUGRA is

$$W = Z \left(S_1 S_2 - \frac{f^2}{2} \right) \quad \text{Kim, PLB136 (1984) 378.}$$

Most SUGRA models, breaking a global symmetry use this superpotential. Sometimes used in hybrid inflation also.

But it is not a minimal model for “invisible” axion.

The minimal “invisible” axion model :

$$W_1 = \lambda ABC + mBD + \frac{1}{M} ADh_u h_d, \text{ Kim-Kyae-Nam-Nilles, 2016.}$$

$$W_2 = f_Q CQ_1\bar{Q}_2 + f_u^{ij} q_i u_j^c h_u + f_d^{ij} q_i d_j^c h_d,$$

This minimal model has a far reaching consequences:

$$\Gamma = Q_R + 2Q_{PQ}$$

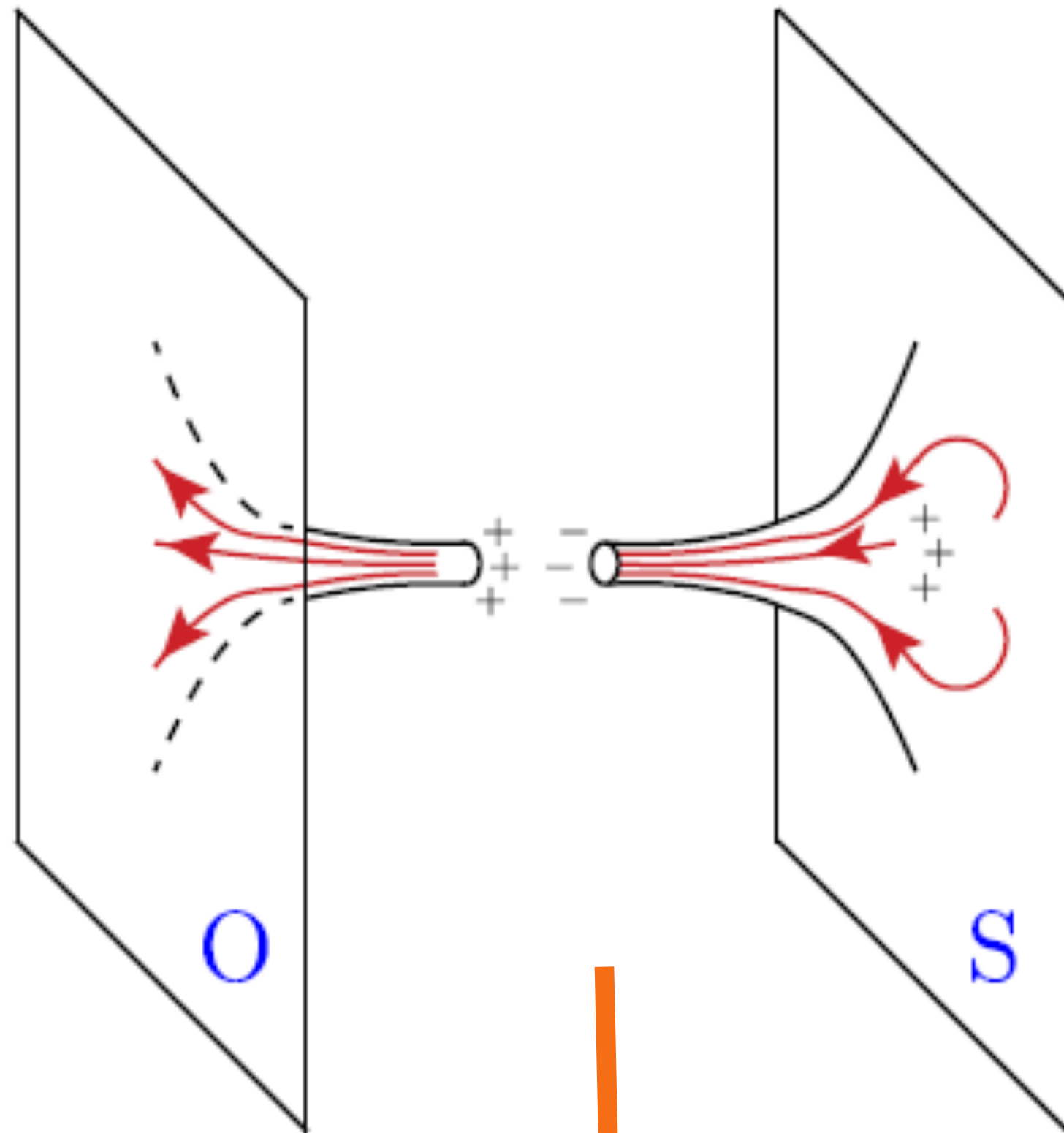
	<i>A</i>	<i>B</i>	<i>C</i>	<i>D</i>	h_u	h_d	Q_1	\bar{Q}_2	q_i	u_i^c	d_i^c	ℓ	e^c
Q_R	+4	0	-2	+2	-2	-2	+2	+2	+2	+2	+2	+2	+2
Q_{PQ}	-1	+1	0	-1	+1	+1	0	0	$\frac{-1}{2}$	$\frac{-1}{2}$	$\frac{-1}{2}$	$\frac{-1}{2}$	$\frac{-1}{2}$
Γ	+2	+2	-2	0	0	0	0	0	+1	+1	+1	+1	+1

This minimal ABCD model has a far reaching consequences:

(1) Heavy axino: Choi-Kim-Lee-Seto, 0801.0491; Kim-Huh, 0908.0152.

(2) Inflation, to appear.

(3) Correct SUGRA phenomenology, to appear.



Wormholes:
Gidding-Strominger,
Coleman, Cline

Exact global symmetries?
Kamionkowski-MarchRussel,
Holdom et al.
Exclude terms up to dim 8.
The example of acc symm.

Discrete gauge symmetry:
Krauss-Wilczek



After many years,
Model-independent axion (Green-Schwarz, Witten)
strikes back

With the anomalous $U(1)$
gauge symmetry in
string compactification



JEK, 1604.00716 [hep-ph]



Antisymmetric tensor fields: B_{MN}

They are gauge fields in 10D. Their couplings to matter fields are from compactification process, including the Green-Schwarz term. If some of them give color anomaly coupling without anomalous U(1), they must be necessarily the hadronic axion-type. Anyway, their decay constants are above the GUT scale. Without fine tuning, it is expected that f_a near the string scale.

MI axion: Choi-K, PLB 154 (1985) 393;

MD axion: Svrcek-Witten, JHEP 06 (2006) 051.

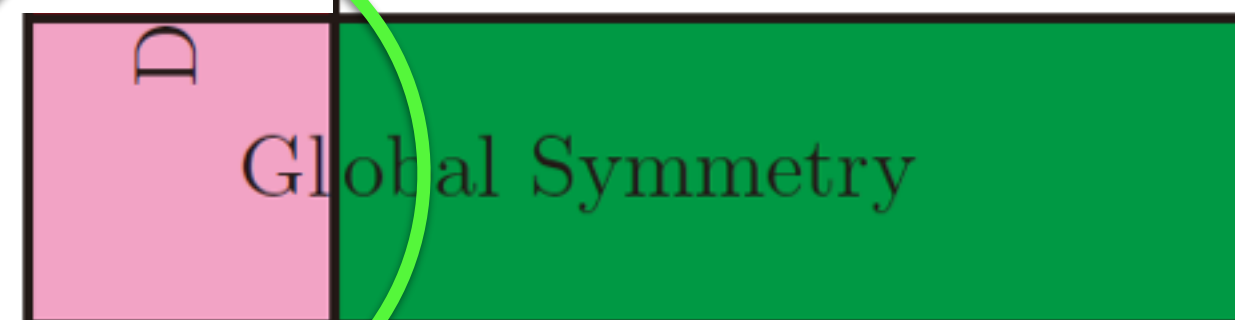
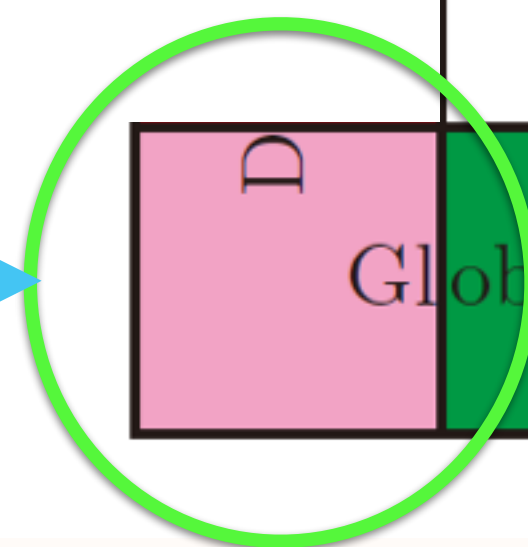
Gauge symmetry origin, but from compactification:

Anomalous U(1) gauge symmetry in string compactification:

becomes PQ global symmetry below 10^{15} GeV.

There can be terms in W , but the PQ symm breaking terms are absent. This can be achieved in compactification, with anomalous $U(1)$.

The KN term is the leading one here.



10^{-33} [eV]

10^{-22} [eV]

10^{-4} [eV]

10^{23} [eV]

Gluon Anomaly

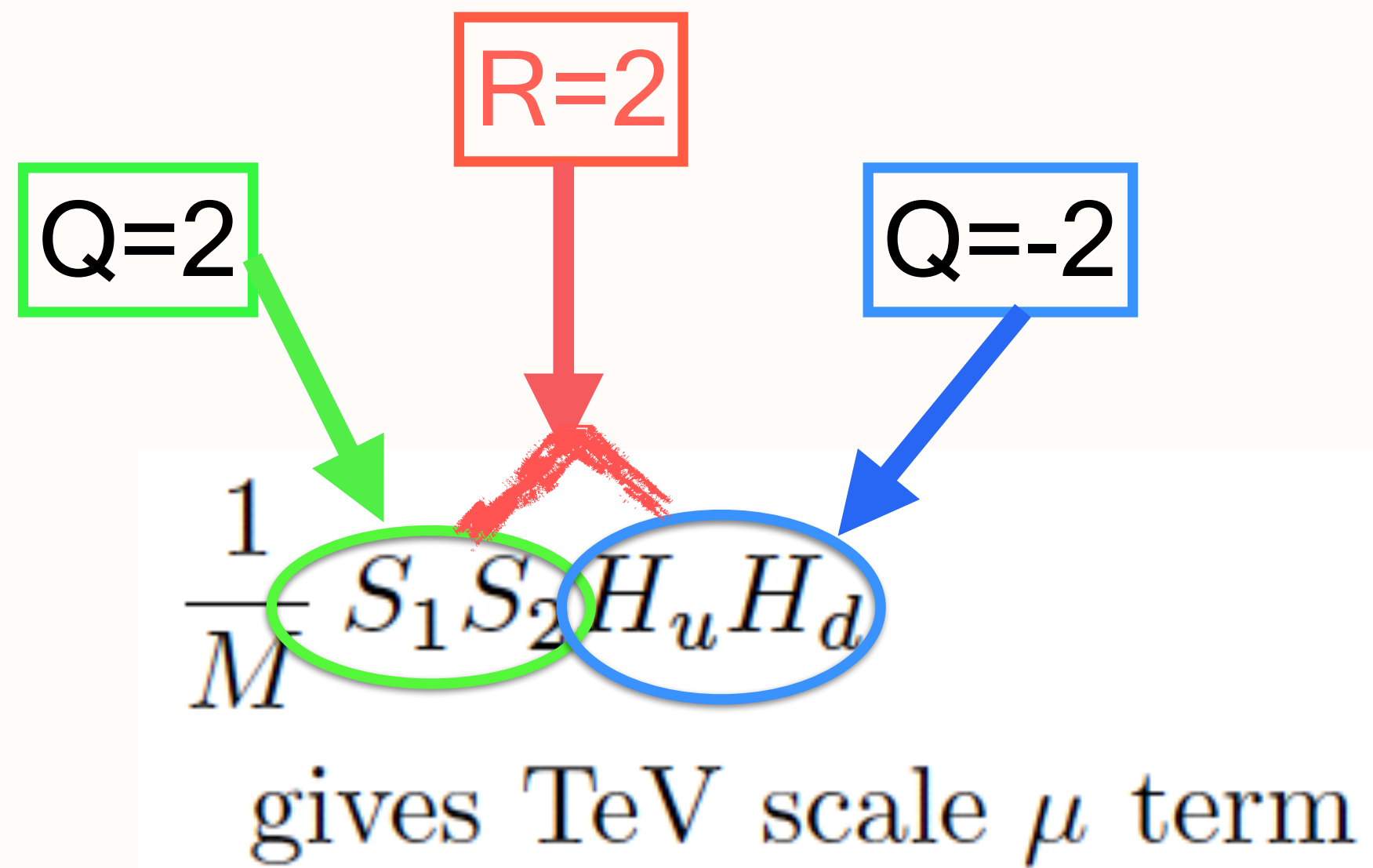
W Anom.

GUT-scale Anomalies

D
Global Symmetry



We have seen the KN SUSY operator:



The diagram shows the operator $\frac{1}{M} S_1 S_2 H_u H_d$ with three arrows pointing to it from boxes above. A green box labeled 'Q=2' has a green arrow pointing to the $S_1 S_2$ part, which is circled in green. A red box labeled 'R=2' has a red arrow pointing to the entire operator. A blue box labeled 'Q=-2' has a blue arrow pointing to the $H_u H_d$ part, which is circled in blue. Below the operator, the text 'gives TeV scale μ term' is written.

$$\frac{1}{M} S_1 S_2 H_u H_d$$

gives TeV scale μ term

It can be realized in string compactification. In the QCD axion coupling to photons, one has to work within a complete phenomenologically working model, taking into account all fermion spectra. Axion-photon-photon coupling:

JEK, PLB 207 (1988) 434;
JEK, PLB 735 (2014) 95, 741 (2014) 317;
JEK-Nam, PLB 759 (2016) 149; coupl.
JEK, PLB 759 (2016) 58. DW No.=1

The minimal model (Kim-Kyae-Nam-Nilles) satisfies both R and PQ global symmetries.

Here, DW number can be made 1 as in the non-SUSY KSVZ model.

But QCD axion and photon couplings are given phenomenologically in the BSM field theory.

Q_{em}	KSVZ $c_{a\gamma\gamma}$	x	$q^c - e_L$ pair	DFSZ $c_{a\gamma\gamma}$	$c_{a\gamma\gamma}$	Superstring $c_{a\gamma\gamma}$	Comments
0	-2	any x	(d^c, e)	$\frac{2}{3}$	$\frac{2}{3}$	arXiv:1405.6175	Anomalous U(1) as U(1) _{PQ}
$\pm\frac{1}{3}$	$-\frac{4}{3}$	any x	(u^c, e)	$-\frac{4}{3}$	$\frac{2}{3}$	hep-ph/0612107	Approximate U(1) _{PQ}
$\pm\frac{2}{3}$	$\frac{2}{3}$			Without	GUTs or	JEK-Nam, PLB 759,149 (16) $\geq \frac{2}{3}$	Anomalous U(1) as U(1) _{PQ}
± 1	4			SUSY	SUSY		$c_{a\gamma\gamma} = (1 - 2 \sin^2 \theta_W) / \sin^2 \theta_W$
(m, m)	$-\frac{1}{3}$			H_d or H_u^*	H_d or H_u		with $m_u/m_d = 0.5$.

Given in JEK, PRD 58 (1998) 055006;
K-Carosi, RMP 82 (2010) 557.

$$\frac{1 - 2 \sin^2 \theta_W}{\sin^2 \theta_W} \geq \frac{2}{3}$$

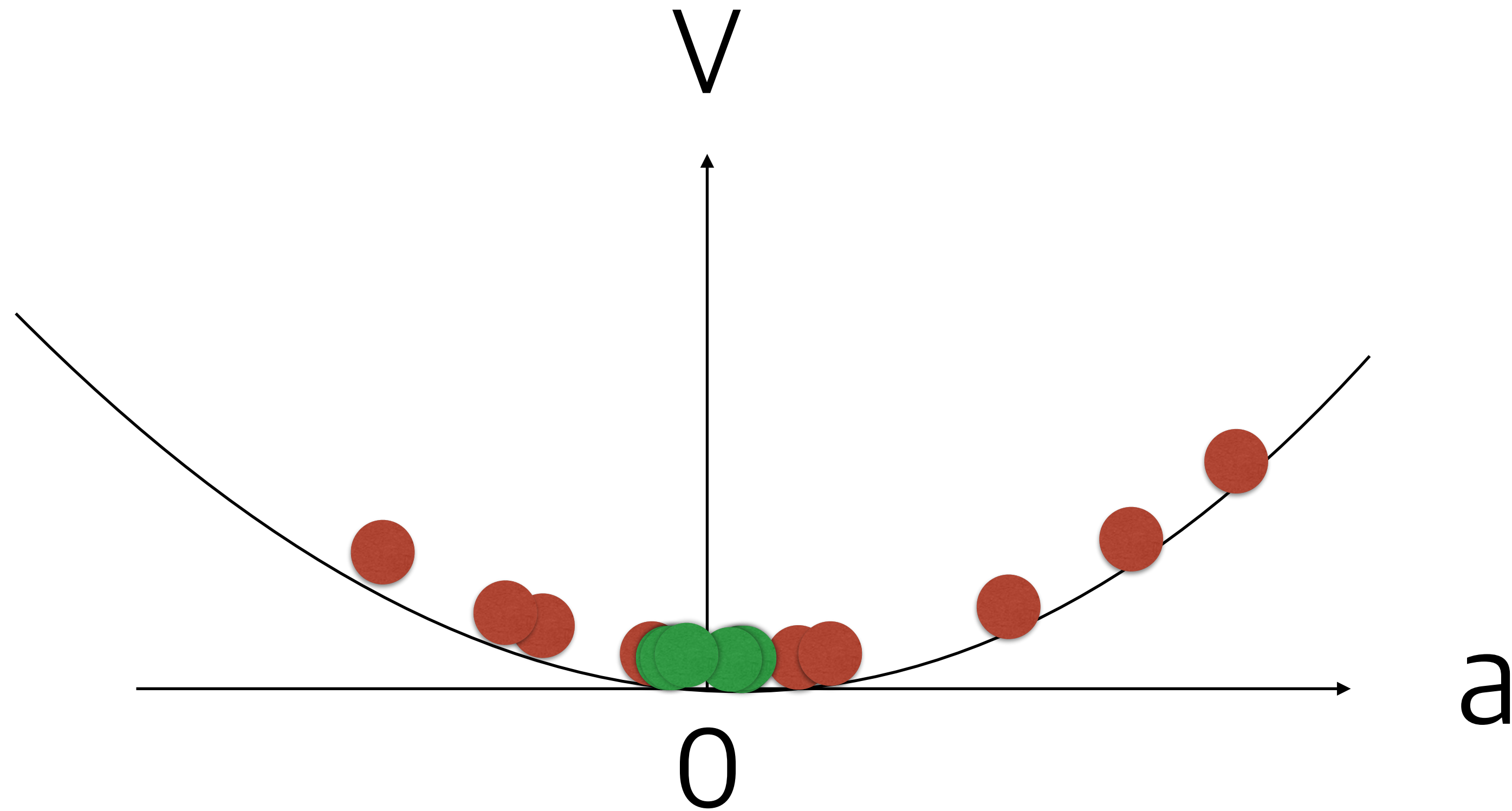
3. QCD axion in cosmology



Axion energy density



Axion solution = cosmological solution



for $\Lambda_{\text{QCD}} = 380 \text{ MeV}$,

$$\Omega_a = 0.025 \left(\frac{\theta_1^2 F(\theta_1)}{\gamma} \right) \left(\frac{0.68}{h} \right)^2 \left(\frac{f_{a, \text{GeV}}}{10^{11}} \right)^{1.184}$$

where $g_{*,\text{present}} \simeq 3.91$ and γ is the entropy production ratio,

Kim-Semertzidis-Tsujikawa, Front. Phys. 2, 60 (2014):
Bosonic coherent motion review
[Review in research topic BCM,
New one: Frontiers in Physics,
HE & astroparticle physics]

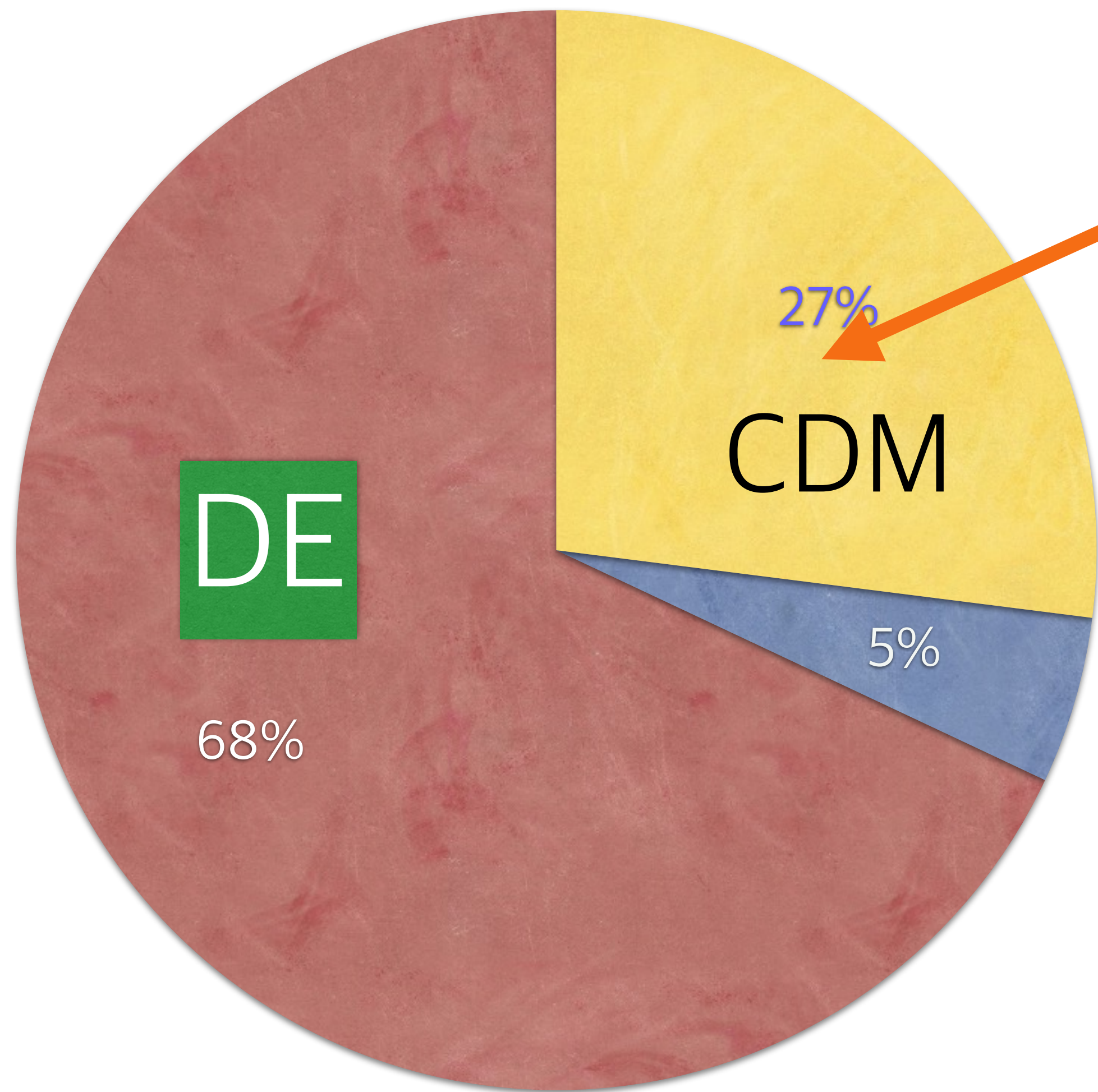
BCM initiated by three groups in 83.
(Preskill et al, Dine et al, Abbott et al.)

BCM:

BCM1 → Axion
BCM2 → Inflaton

CCtmp: ↓
Dark energy





For the invisible axion here, the present axion detection experiments on CDM axions are going on.

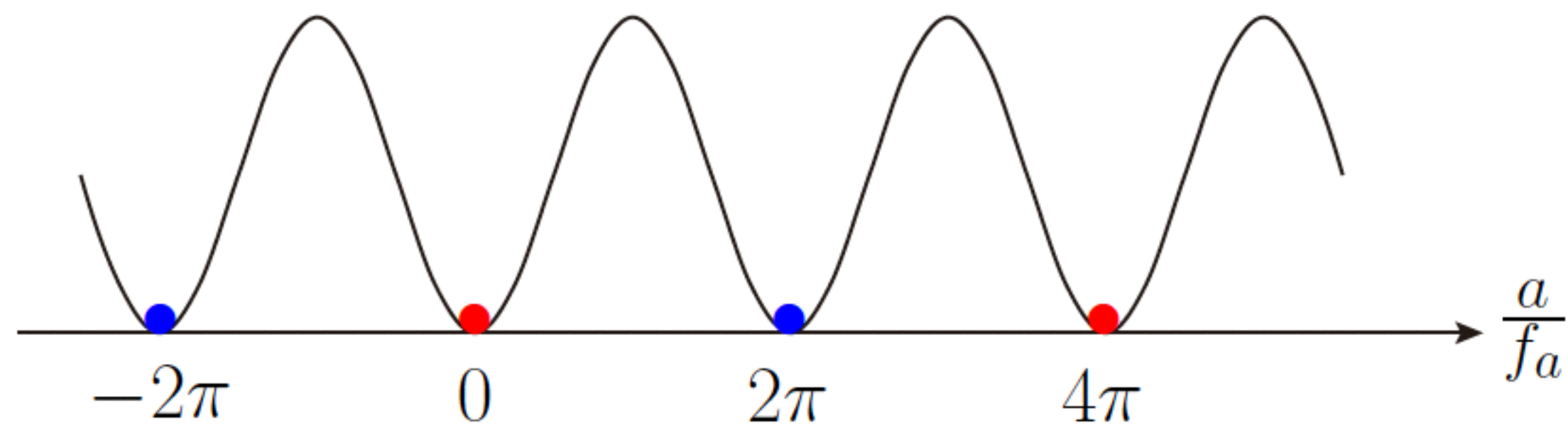
The QCD axion is studied for this region of parameter space. Any constraints are talked about with this CDM possibility.



Axionic domain walls



$$S = -\frac{\bar{\theta}}{32\pi^2} \int d^4x G_{\mu\nu}^a \tilde{G}^{a\mu\nu}.$$



But other matter fields
can give

$$\Phi \rightarrow e^{i\theta/N} \Phi, \quad \theta = \frac{a}{f_a}$$

Phi returns to its original value after
 $a \rightarrow a + 2\pi N f_a$. N is the domain-wall number.
[Sikivie (1982)]

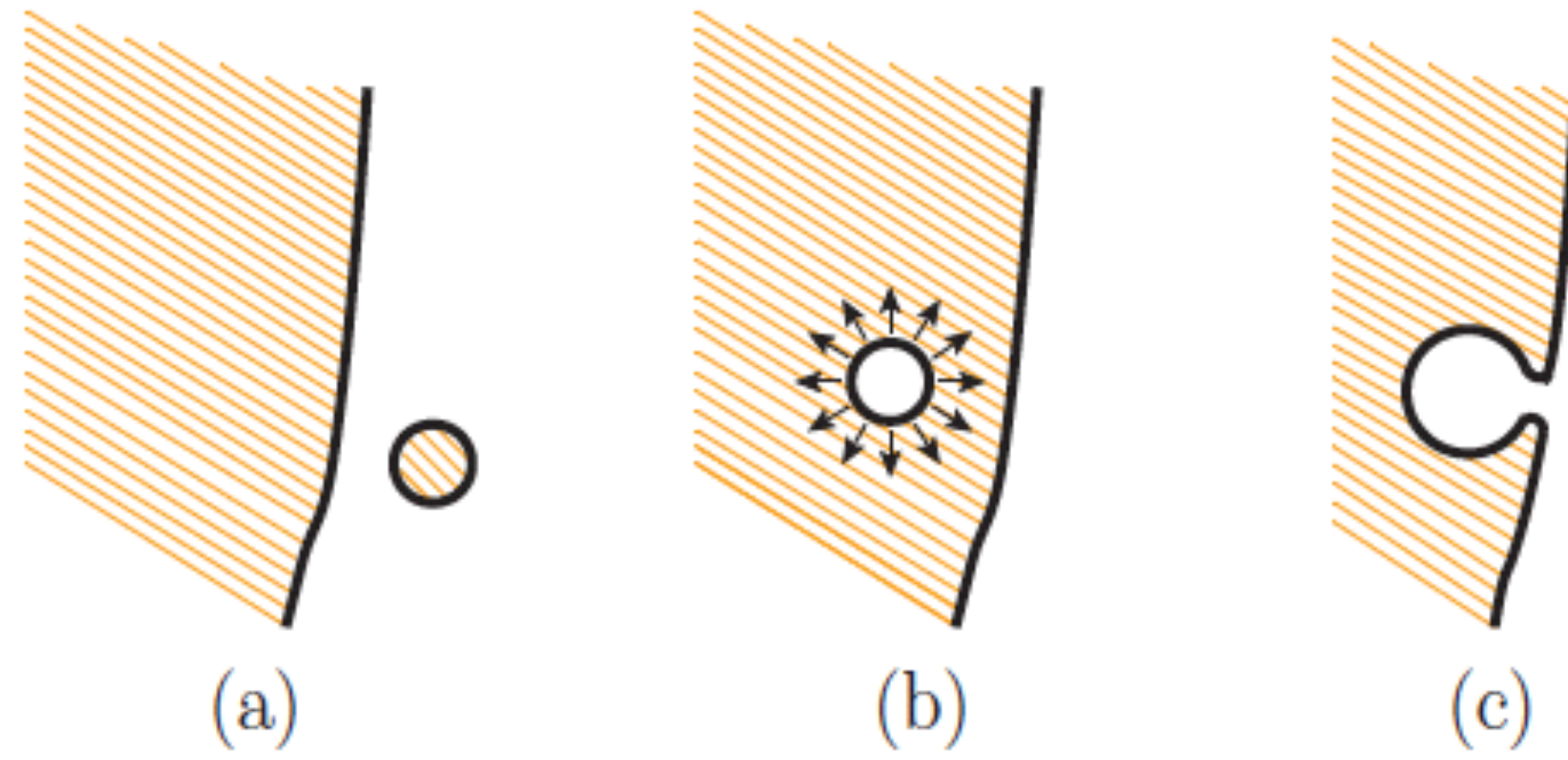
Small string-DW are not problematic.



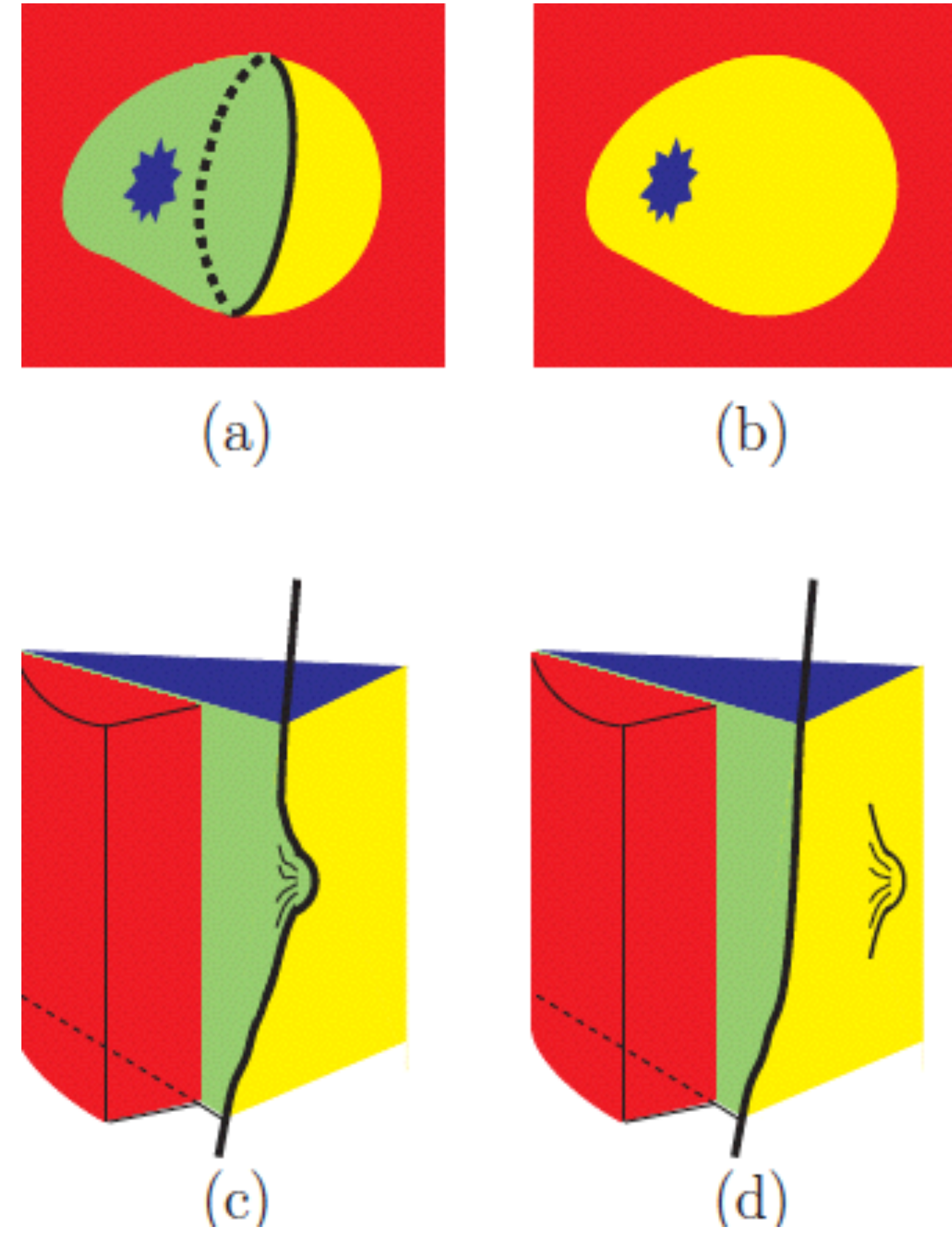
The horizon scale string-DW system is problematic.

Barr-Kim, PRL 113 (2014) 241301.





Vilenkin-Everett (1982);
Barr-Choi-Kim (1987)



Sikivie (1982)

Top-down approach, using string compactification

1. Anomalous $U(1)$ gauge symmetry.
2. Anomalous $U(1)$ becomes global $U(1)$ below the GUT scale.
3. The global $U(1)$ is broken at the axion window.
4. DW number here. By giving a VEV to $NPQ=1$ field, we obtain $NDW=1$.
5. Choi-Kim mechanism: with hidden sector force.



Better way: we need a
Goldstone boson direction

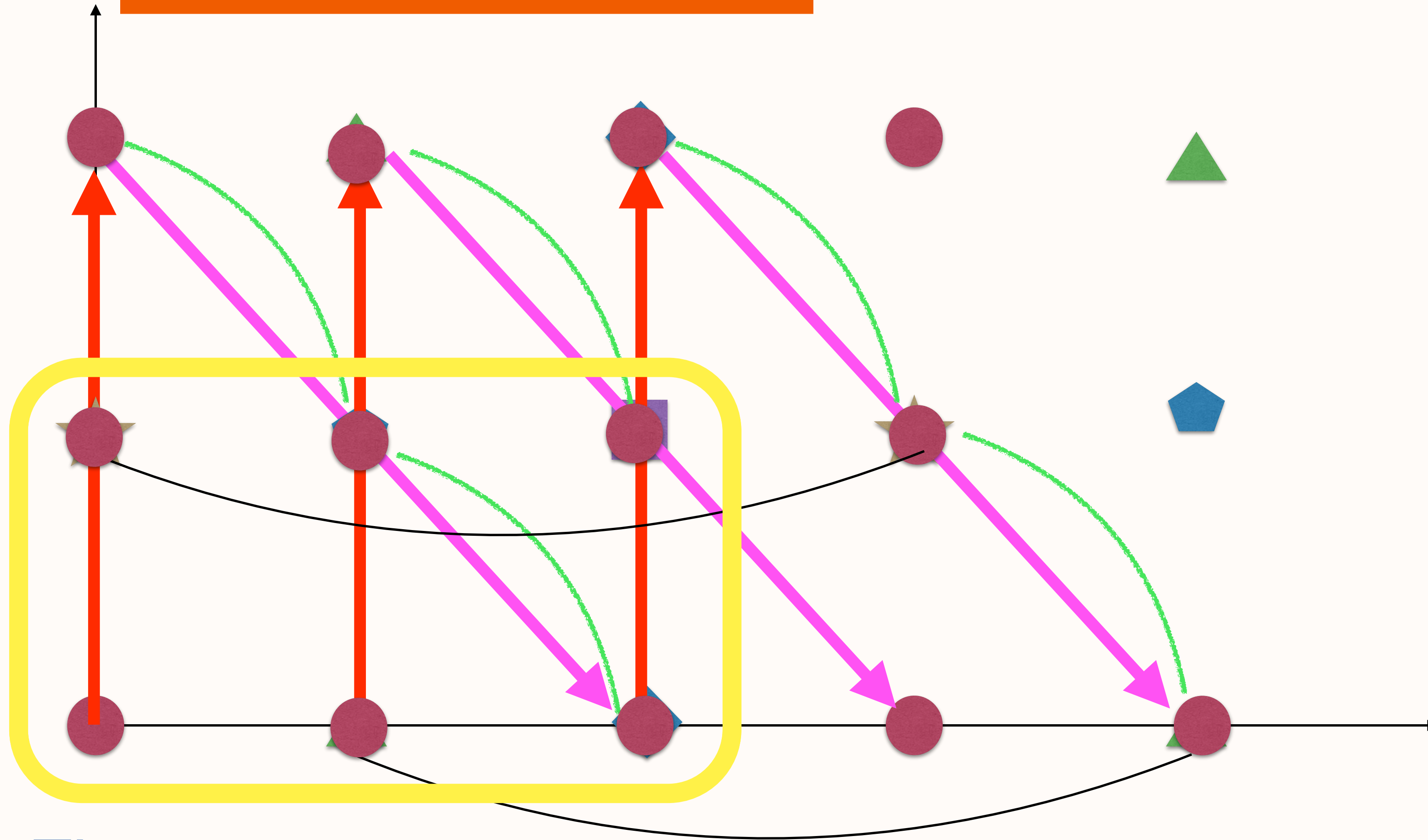


Choi-Kim, PRL55 (1985) 2637

with two confining forces

For the center of GUT group, Lazarides-Shafi (1982).
But, the following ideas are more widely applicable.

Goldstone boson direction



The same vacuum identification

Goldstone boson directions are in principle discussed here:

JEK, PLB 207 (1988) 434;

JEK, PLB 735 (2014) 95, 741 (2014) 317;

JEK-Nam, PLB 759 (2016) 149; coupl.

JEK, PLB 759 (2016) 58. DW No.=1

Axionic string contribution: It is important if strings are created after PQ symmetry breaking. With a high scale inflation, this string contribution has been considered.

Probably, this is the most significant implication of BICEP2 result on axion physics, with DW number 1:

Vissineli-Gondolo, 1403.4594.

Marsh et al, 1403.4216. 71 micro-eV???

But, it depends a calculation of axions from the system of string-domain walls.

1. Kamionkowski-March-Russel(1982)

2. Numerical estimates

Florida group: 1

Cambridge group: $O(100-1000)$

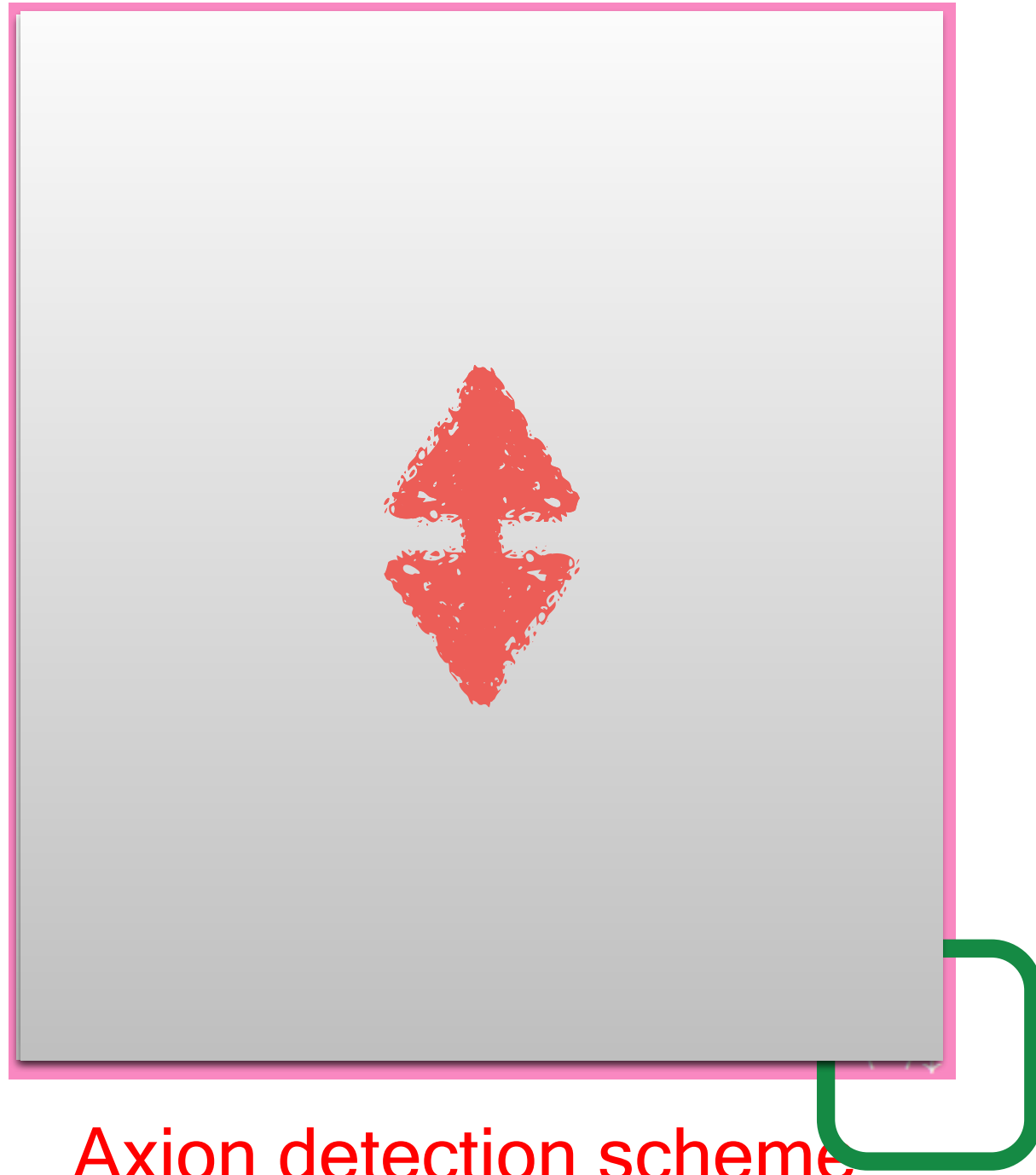
Tokyo group: 25

← This number is used



4. Axion detection



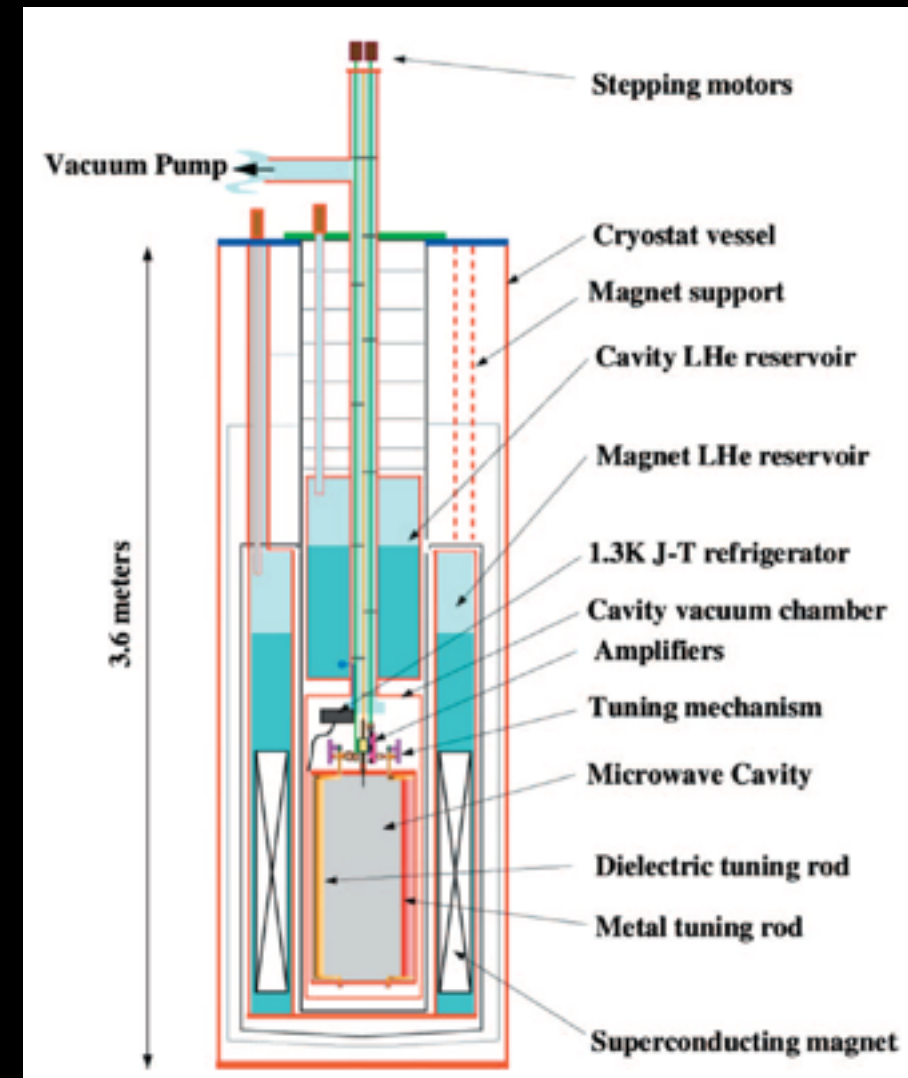


Axion detection scheme

Constant B field
E field follows the
axion oscillation



Very light axion : 1979
CDM candidate

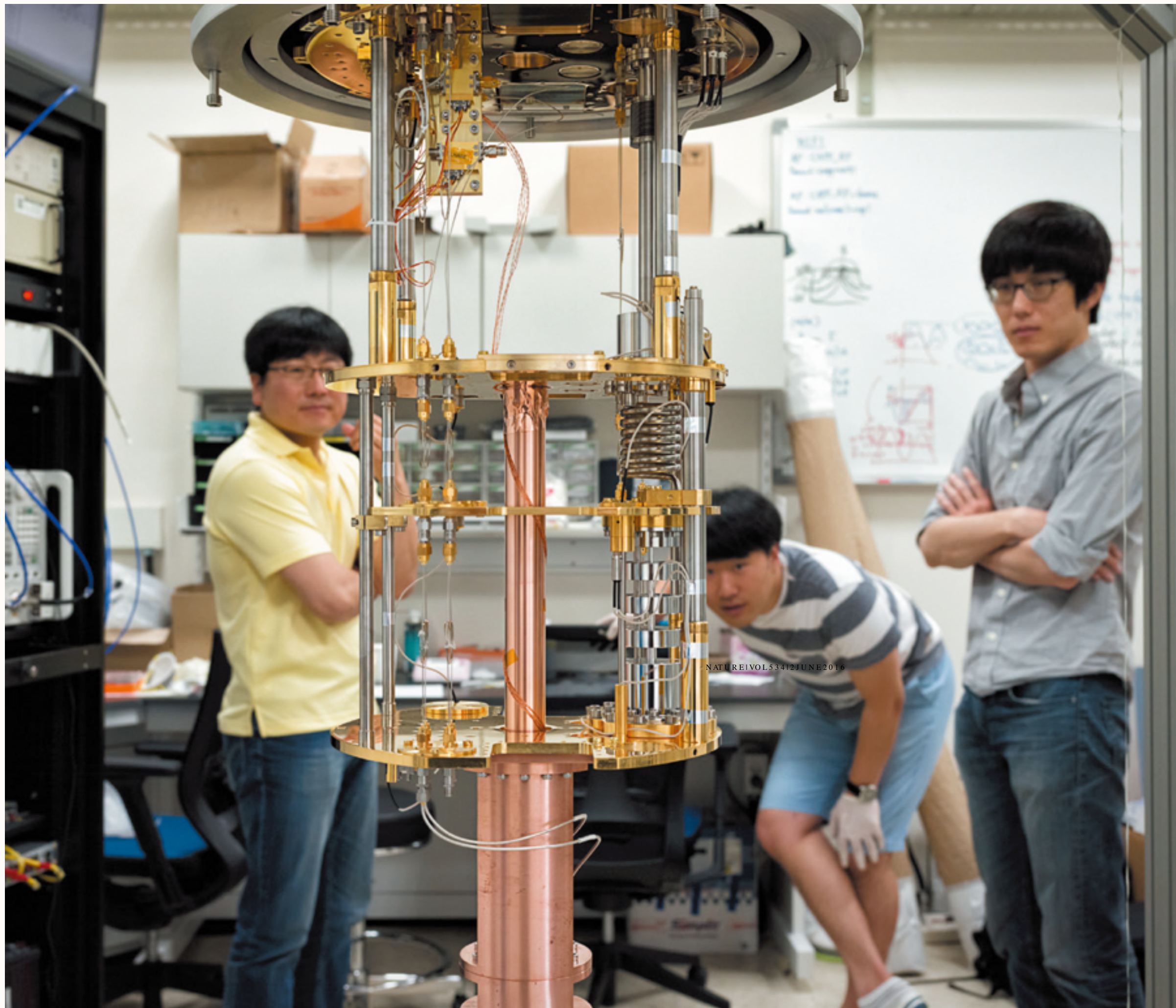


ADMX

CAPP: ???

Detection suggested : 1983
CAPP started : 2013

Sikivie's cavity detector



Nature Vol. 534 (June 2, 2016)



May 18, 2016

The design requires the solution in axio-electrodynamics, given in Hong-K-Semetrtzidis, 1403.1576 [hep-ph]: The record exists.

$$\begin{aligned} \nabla \cdot \mathbf{E} &= \rho + g\nabla a \cdot \mathbf{B}, \\ \nabla \times \mathbf{B} - \partial_t \mathbf{E} &= \mathbf{j} - g\mathbf{B}\partial_t a - g\nabla a \times \mathbf{E}, \\ \nabla \cdot \mathbf{B} &= 0, \\ \nabla \times \mathbf{E} + \partial_t \mathbf{B} &= 0, \\ (\partial_t^2 - \nabla^2)a &= -V'(a) - g\mathbf{E} \cdot \mathbf{B} + \rho_a. \end{aligned}$$

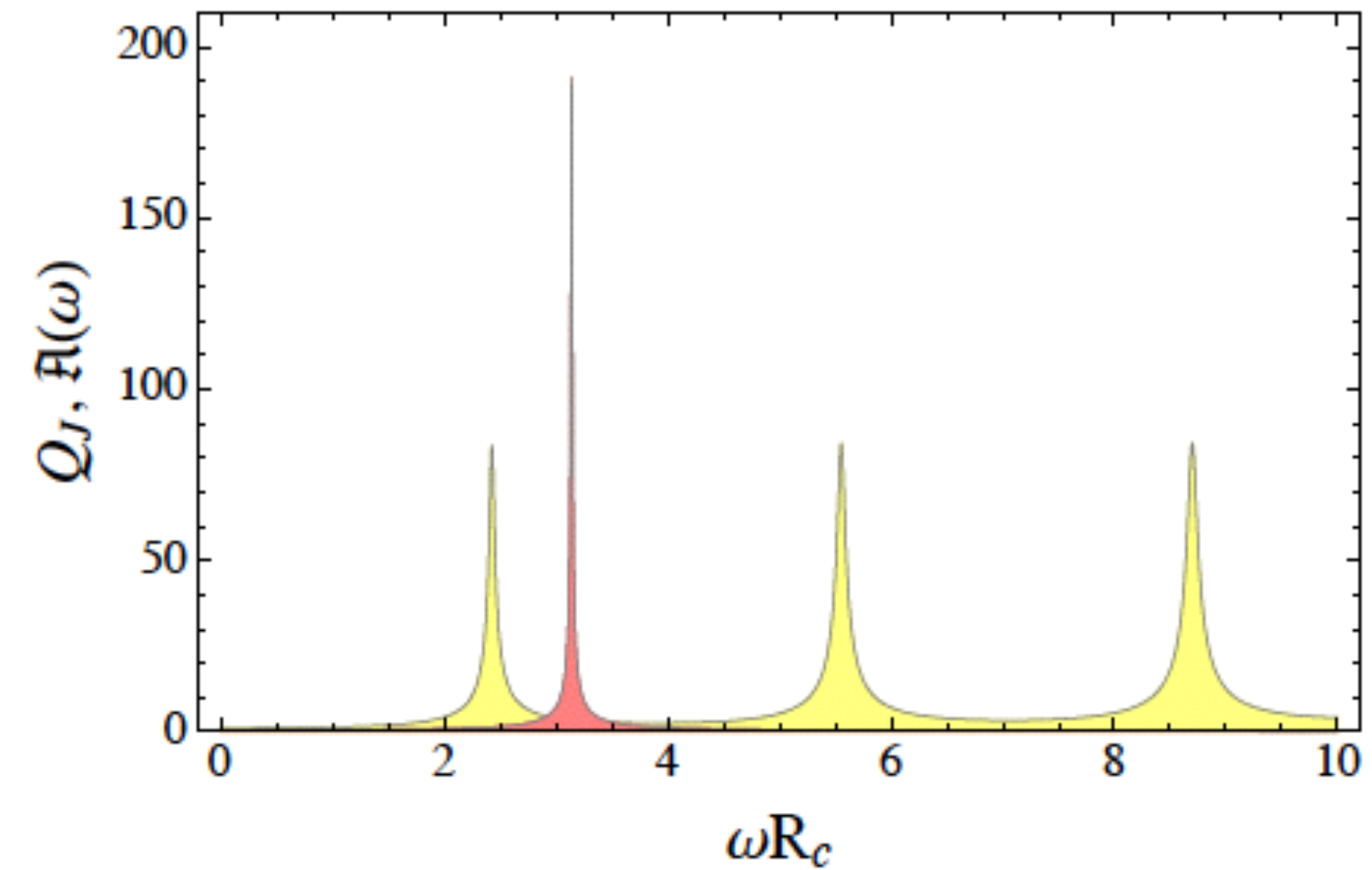
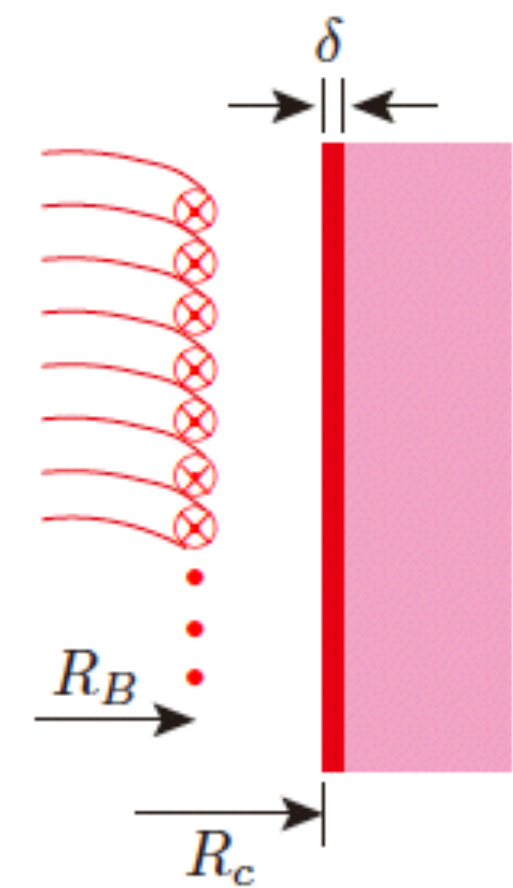


FIG. 2: An illustration of axion-tube resonance ($Q_l \sim 100$, $Q_a \sim 300$, $m_a \sim 1.3\omega_1$). The yellow curve is Q_J as a function of ωR_c and the reddish one is $|\mathcal{A}(\omega)|$ peaked at $m_a R_c$. A resonance occurs when the center of $|\mathcal{A}(\omega)|$ coincides with one of the peaks of Q_J , *i.e.* $\omega_l \simeq m_a$. The actual resonance curves would be much sharper.

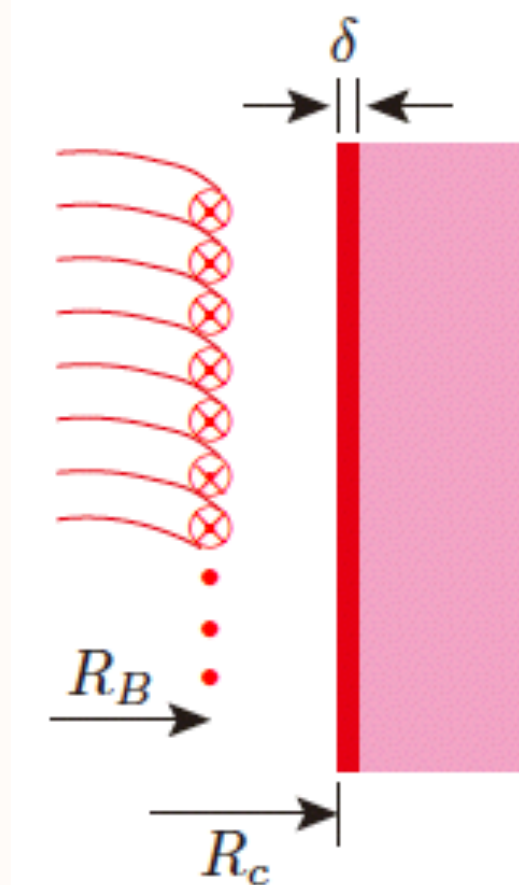
$$\nabla \cdot \mathbf{E} = \rho + g\nabla a \cdot \mathbf{B},$$

$$\nabla \times \mathbf{B} - \partial_t \mathbf{E} = \mathbf{j} - g\mathbf{B}\partial_t a - g\nabla a \times \mathbf{E},$$

$$\nabla \cdot \mathbf{B} = 0,$$

$$\nabla \times \mathbf{E} + \partial_t \mathbf{B} = 0,$$

$$(\partial_t^2 - \nabla^2)a = -V'(a) - g\mathbf{E} \cdot \mathbf{B} + \rho_a.$$



Cavity detectors:

Axion detection is going on at

ADMX: Seattle, USA

CAPP: Daejeon, Korea

(Haloscope: cosmic axions, f_a near 10^{11} GeV)

Planned at

IAXO: Spain

(Helioscope: Solar axions, f_a near 10^9 GeV)

5. Conclusion

1. The need: a solution of strong CP problem.
2. "Invisible" axions.
3. Prospective models from an ultra-violet completion, and presented couplings.
4. Commented on its implications in cosmology and possible detection.

