



# MadMax: a new QCD Dark Matter Axion search

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- **Axion to photon conversion at surfaces**
- **The open structure idea & simulations of boost factor**
  - **Experimental idea & first measurements  
with Seed project setup**
    - **Design ideas for magnet**
      - **Further plans**

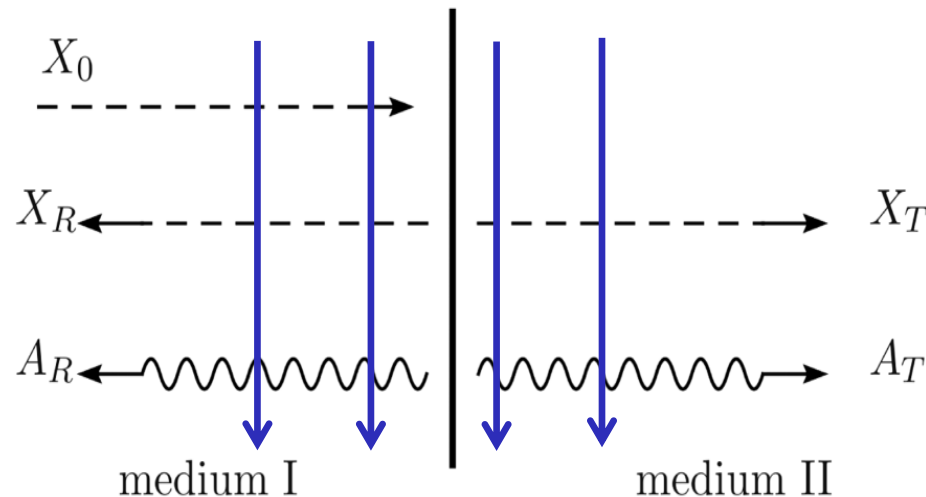


# Axion to photon conversion at surfaces:

Axion mixes with photon in a static B-field

At surfaces (reflecting or change in refractive index): emission of photons in both directions

D. Horns, J. Jaeckel, A. Lindner, A. Lobanov, J. Redondo and A. Ringwald JCAP 1304 (2013) 016 [[arXiv:1212.2970](https://arxiv.org/abs/1212.2970)].



$$(P/A)_{\text{single surface}} \sim 2 \cdot 10^{-27} \text{ W/m}^2 \cdot (B_{\parallel}/10\text{T})^2 \cdot c_V^2 \cdot f(\epsilon_{m1}, \epsilon_{m2})$$

$\epsilon_{m1}, \epsilon_{m2}$  dielectric constant of media I and II

$f(\epsilon_{m1}, \epsilon_{m2})$  increases with increasing  $\Delta(\epsilon_{m1}, \epsilon_{m2})$

Vacuum-mirror transission:  $f(0, \infty)=1$

Axion to photon conversion appears as source term in Maxwell equations

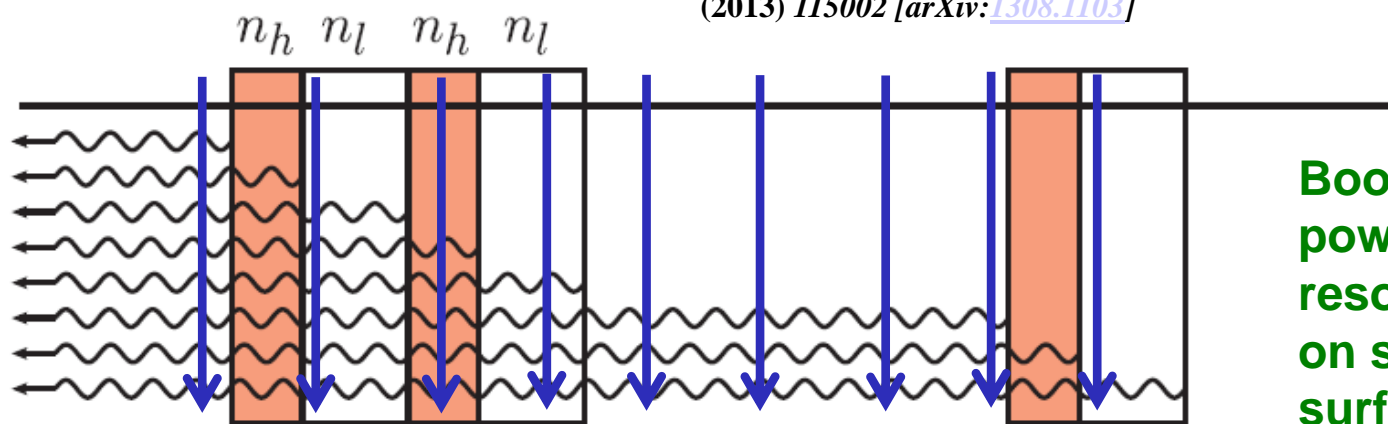
# Axion to photon conversion at surfaces:

Many surfaces:

→ constructive interference of photon part of wave function

→ “photon boost”

J. Jaeckel and J. Redondo, Phys. Rev. D 88  
(2013) 115002 [[arXiv:1308.1103](https://arxiv.org/abs/1308.1103)]



Boost factor  $\beta$ :  
power emitted by  
resonator/power emitted  
on single metallic ( $\epsilon_r = \infty$ )  
surface

$$(P/A)_{\text{resonant cavity}} \sim 2 \cdot 10^{-27} \text{ W/m}^2 \cdot (B_{\parallel}/10\text{T})^2 \cdot c_V^2 \cdot f(\epsilon_{m1}, \epsilon_{m2}) \cdot \beta$$

$\beta$  : Boost factor, depends on:

frequency (axion mass),  $\epsilon$  of materials, number of surfaces,  
displacement between surfaces, etc.

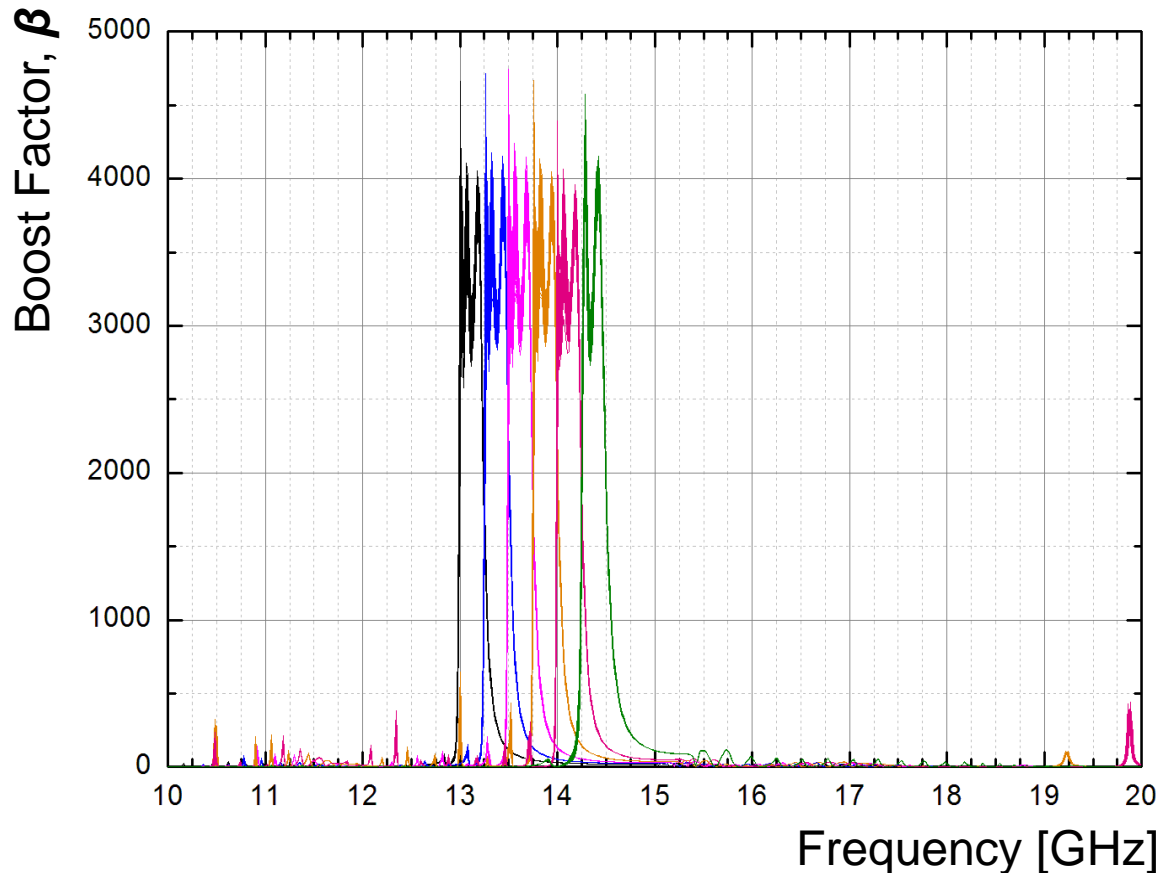
Axion to photon conversion appears as source term in Maxwell equations

→ Can treat problem classically!

12th PATRAS workshop on Axions, WIMPs and WISPs, 2016 June 20-24

# First simulations: the boost factor

20 plates with  $\epsilon_r = 24$  (LaAlO<sub>3</sub>) for bandwidth 250 MHz

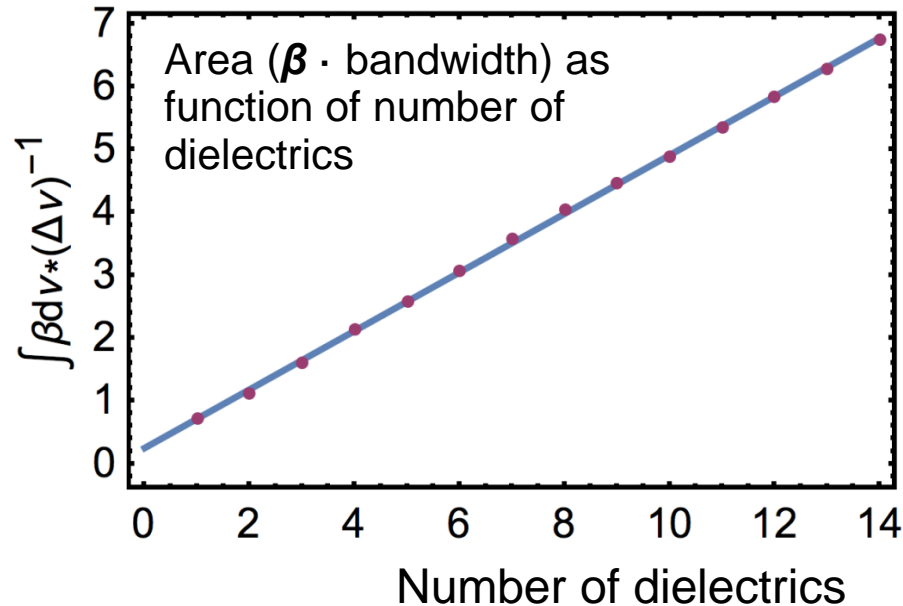


It is possible to adjust disc setting to reach sizeable  $\beta$  over broad bandwidth

Here: bandwidth per setting:  $\sim 250$  MHz

Estimated precision of placement of discs needed:  $\sim$  few  $\mu$ m

## First simulations: the boost factor

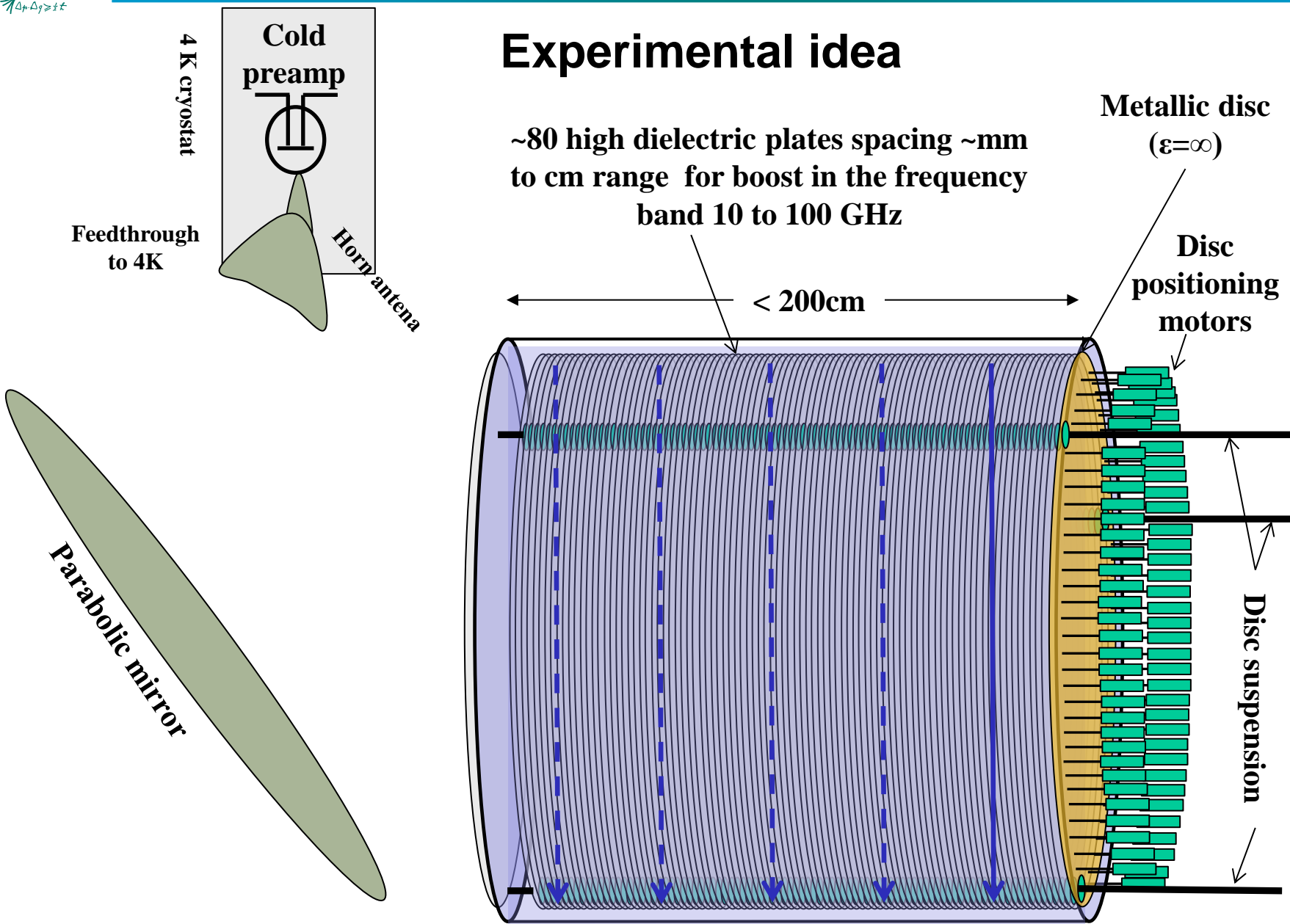


→ Area of boost peak scales ~linearly with number of discs

Simulations: disc placement (80 discs) with precision of  $\sim \mu\text{m}$   
enough for  $\beta \sim 10^5$  with bandwidth of tens of MHz

Boost factor can be probed by reflectance and  
transmittance measurements

# Experimental idea



## Experimental idea

### Chose dielectric material:

- High dielectric constant  $\epsilon$  (for large boost & conversion)
  - Low loss  $\rightarrow$  low  $\tan \delta$  (reduce photon losses)
    - Stable
    - Cheap

$\rightarrow$  Sapphire ( $\text{Al}_2\text{O}_3$ ) @ 300K, 10 GHz:

$$\epsilon \sim 10; \quad \tan \delta \sim \text{few} \cdot 10^{-5}$$

$\rightarrow$  Lanthanide Aluminate ( $\text{LaAlO}_3$ ) @ 77K

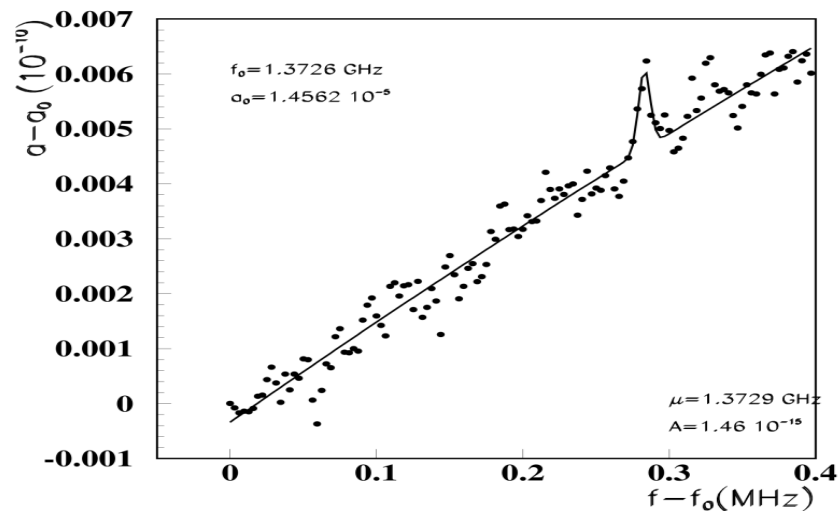
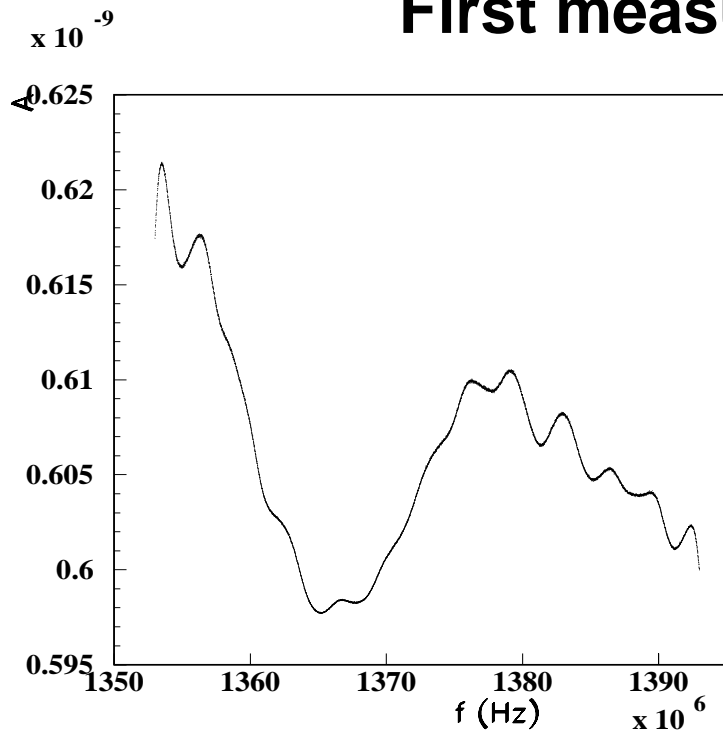
$$\epsilon \sim 24; \quad \tan \delta \sim 3 \cdot 10^{-5}$$

$\rightarrow$  Titanium dioxide – Rutil ( $\text{TiO}_2$ )

$$\epsilon \sim 100; \quad \tan \delta \sim \text{???}$$



# First measurements: sensitivity



- Inject fake axion signal with  $3 \cdot 10^{-21}$  W power
- Measurement for one week (integrate signal):  
Receiver at Room Temp.
  - Independent „blind“ analysis
  - found  $> 6\sigma$  signal successfully
  - At LHe: noise level factor 100 better
  - Sensitivity at the level of  $10^{-23}$  W expected

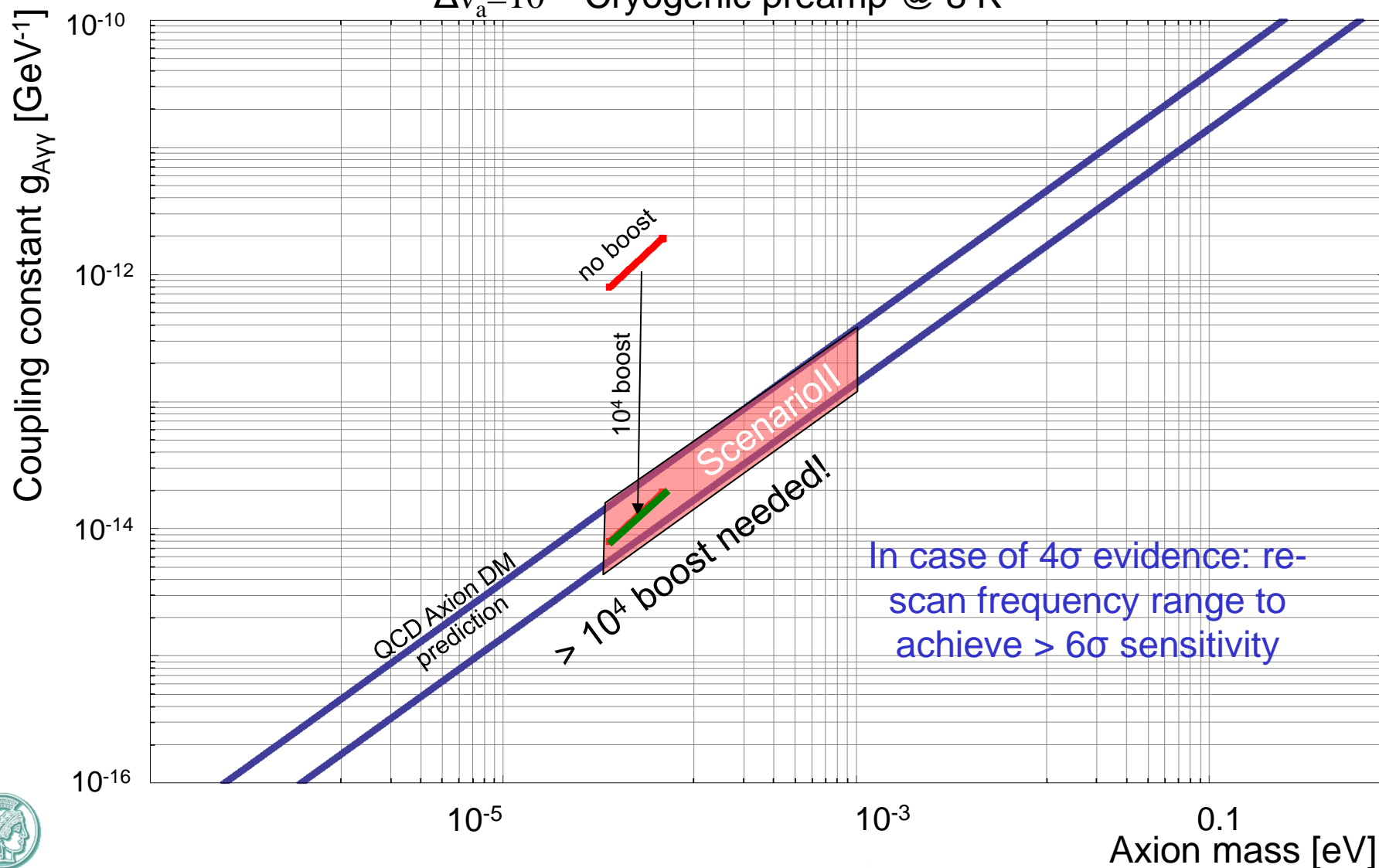




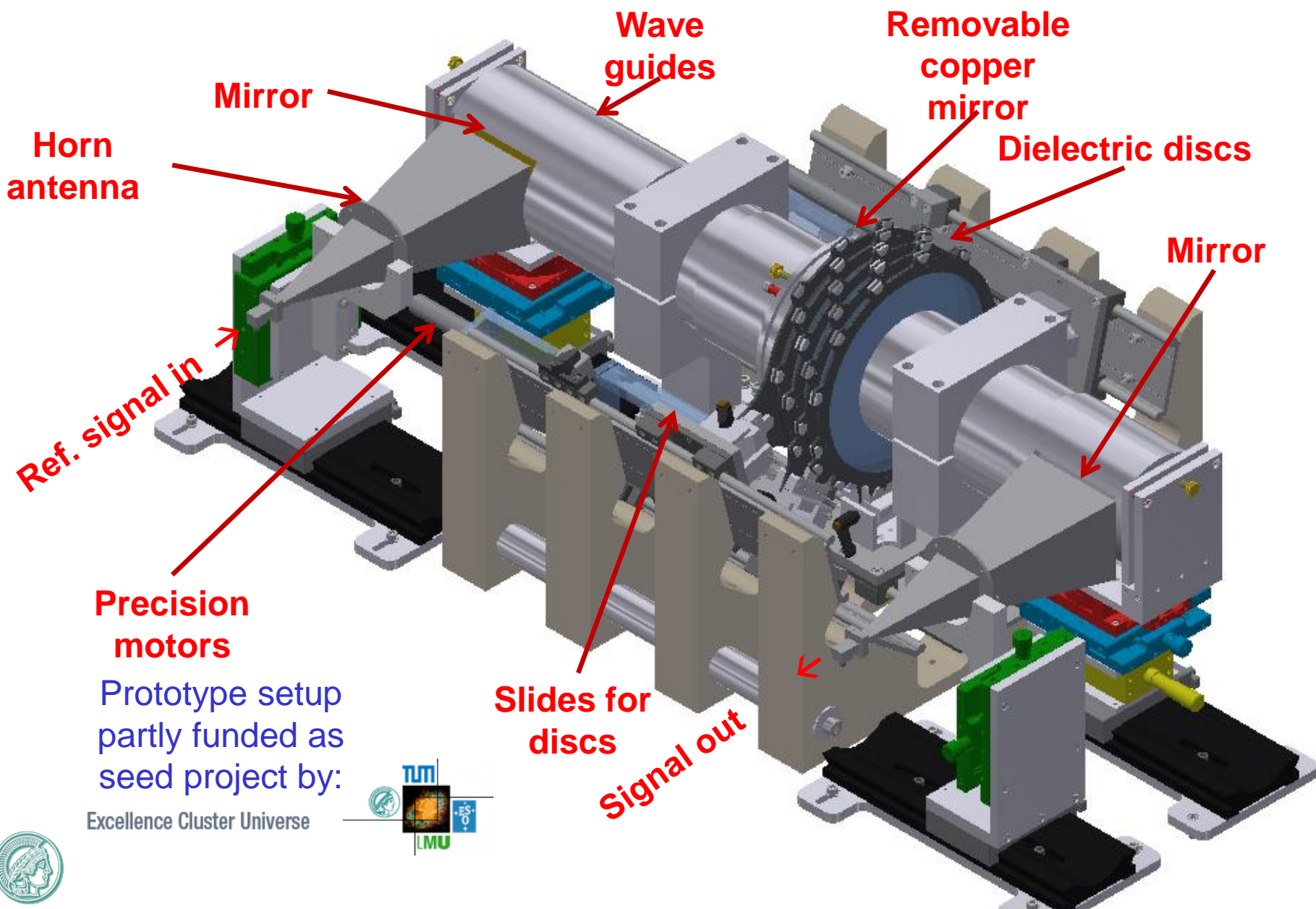
# First measurements: sensitivity

Expected  $4\sigma$  detection sensitivity **with** and **without** boost

for 80 discs,  $1\text{m}^2$  disc surface, 10T B-field,  $\tau=200\text{h}$ , 50MHz boost bandwidth,  
 $\Delta v_a=10^{-6}$ ; Cryogenic preamp @ 8 K



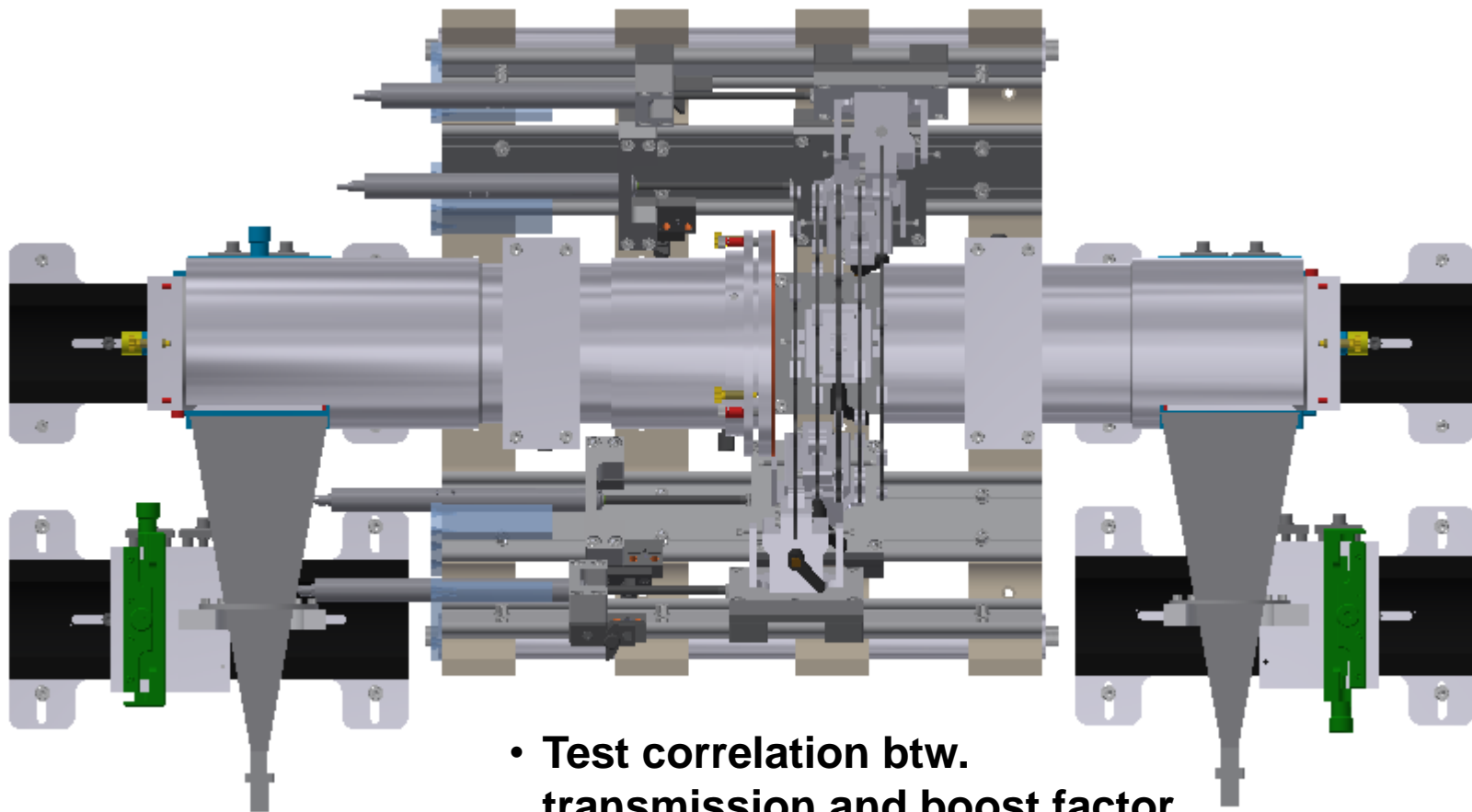
# First prototype setup at MPI



Prototype setup partly funded as seed project by:

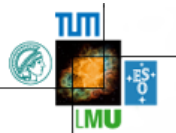


## First prototype setup at MPI



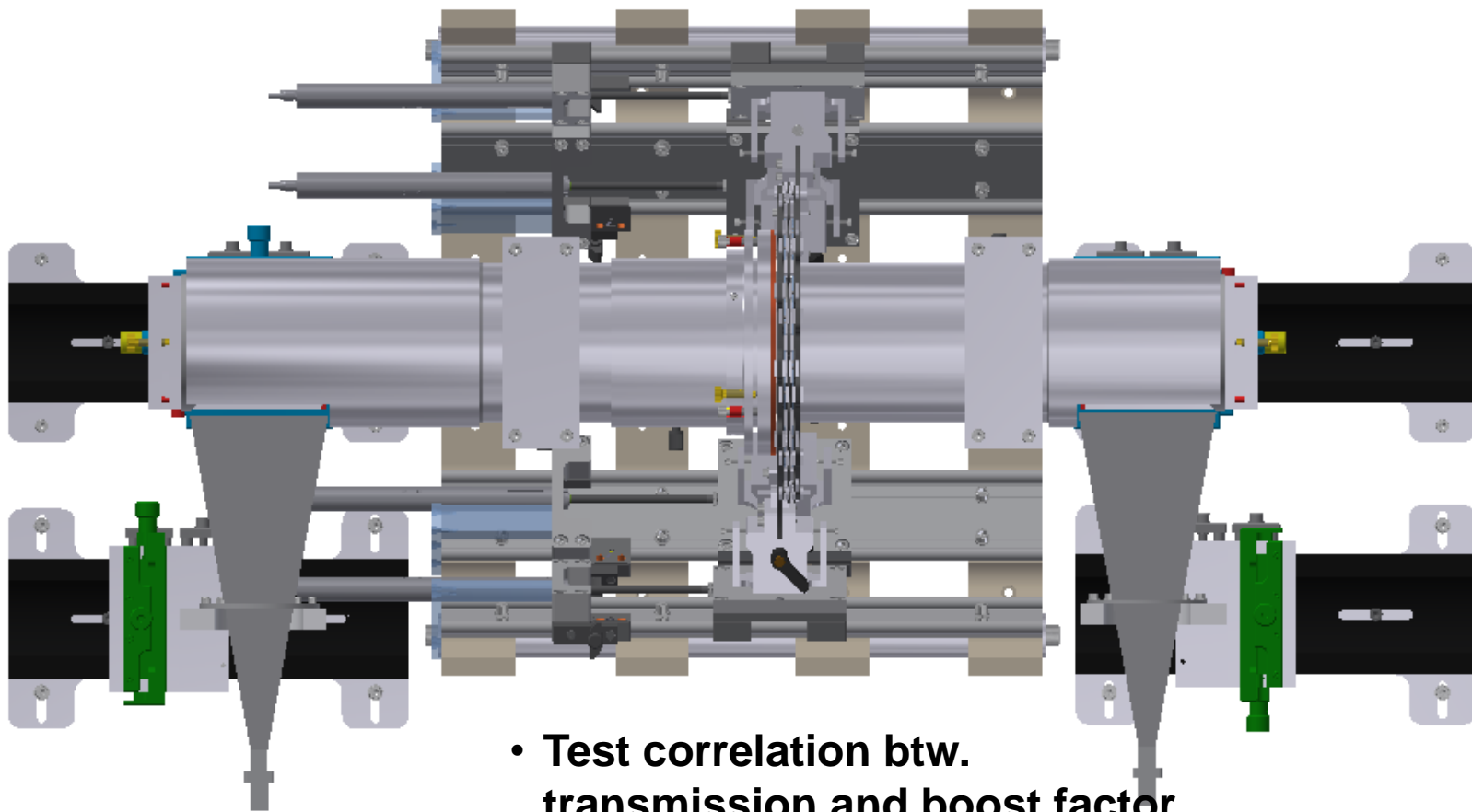
Prototype setup partly funded as seed project by:

Excellence Cluster Universe



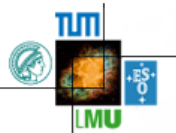
- Test correlation btw. transmission and boost factor
- Test needed disc precision
- Evaluate uncertainties
- R&D on tiling

## First prototype setup at MPI



Prototype setup partly funded as seed project by:

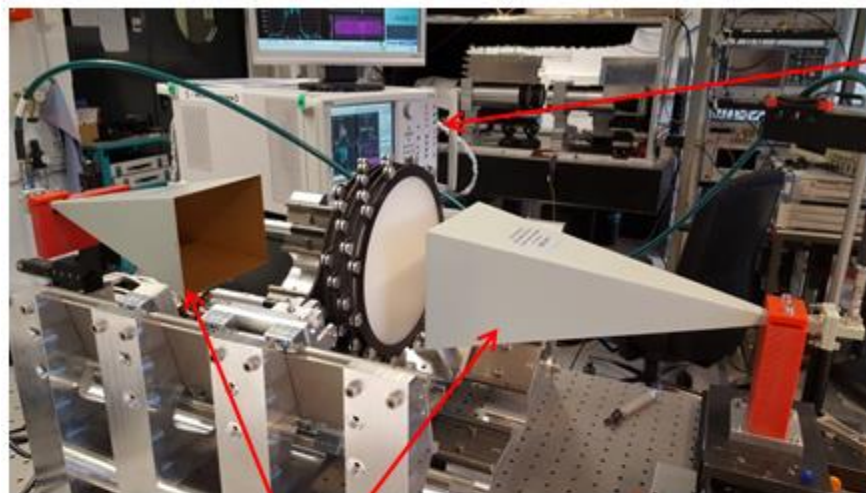
Excellence Cluster Universe



- Test correlation btw. transmission and boost factor
- Test needed disc precision
- Evaluate uncertainties
- R&D on tiling

# First prototype setup at MPI

## Transmission and reflection measurements



Network Analyzer

Antennas

Motors

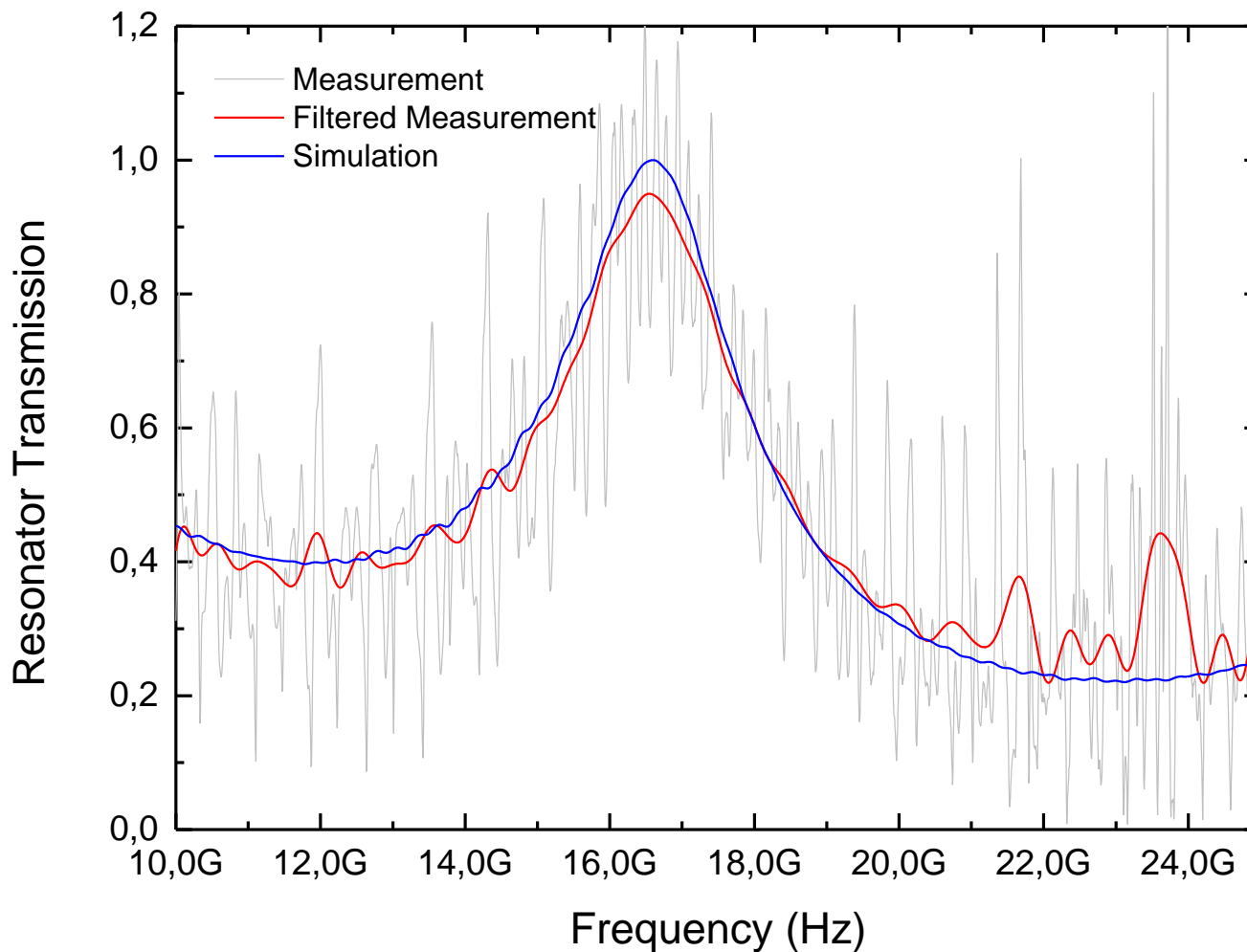


Sapphire discs

Prototype setup partly funded as seed project by:

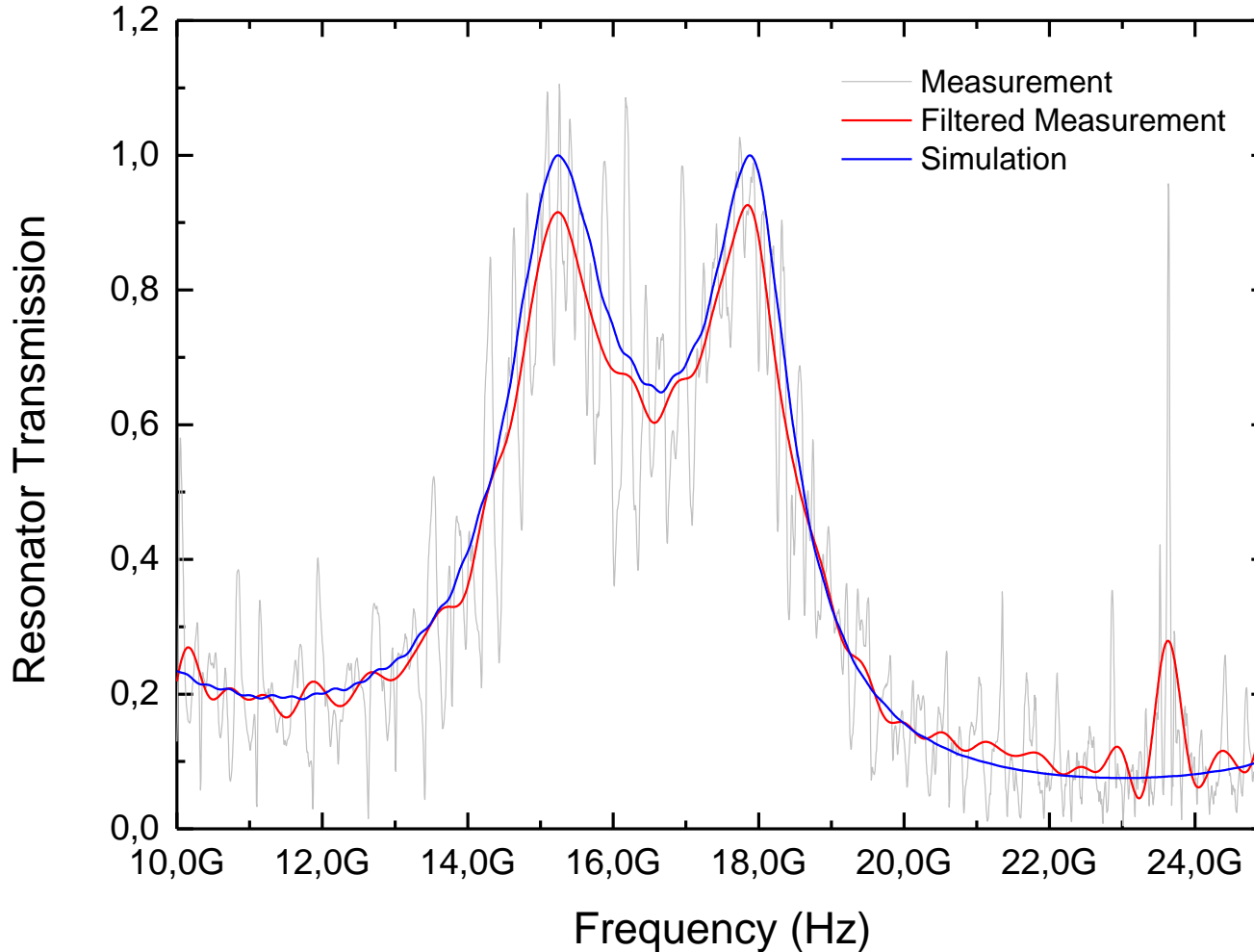


## Transmission measurements: 2 discs



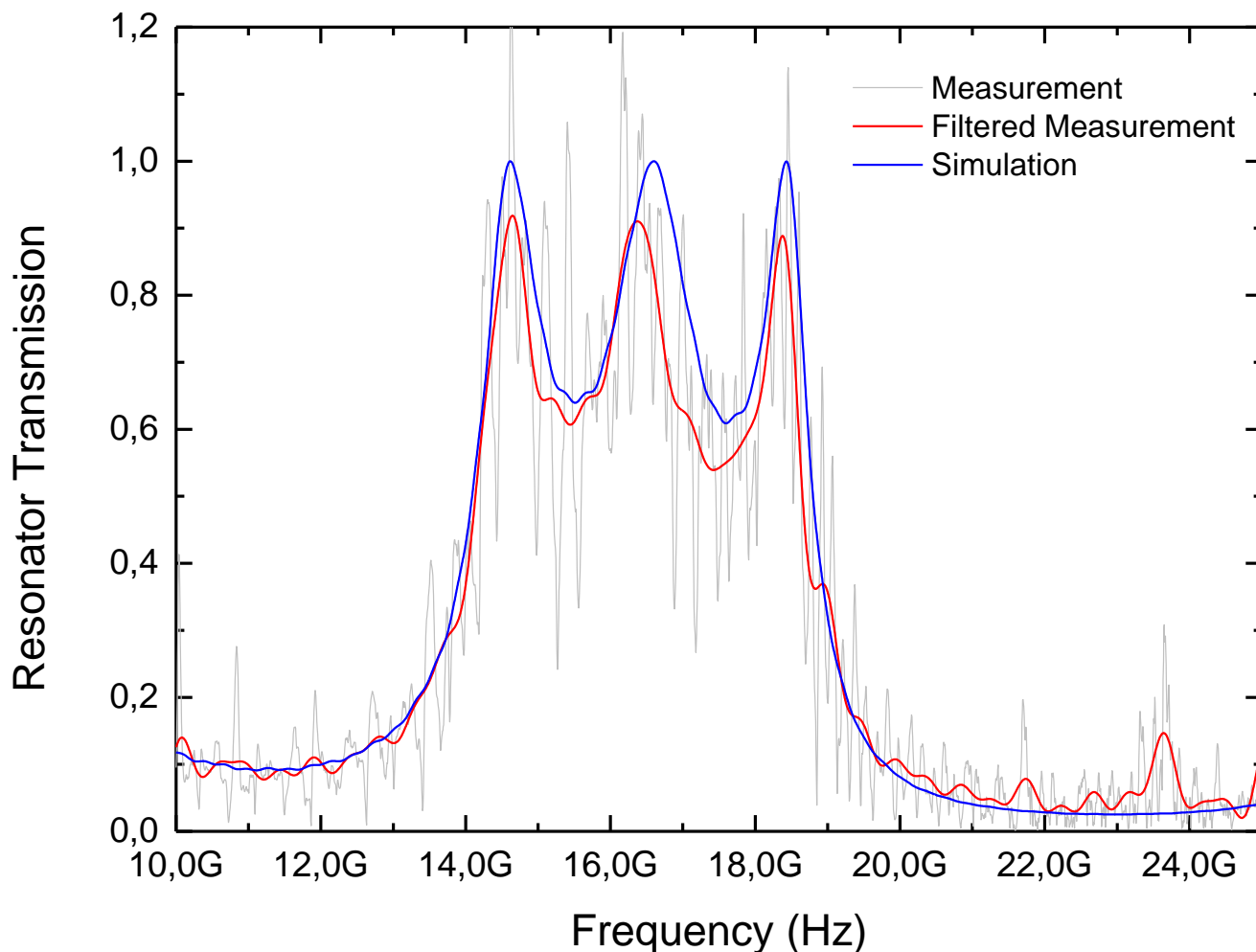
- **2**  $\text{AlO}_3$  discs with diameter 200mm Disc positions **transmission curve**
- Prediction (red) fits measurement (black) well.

## Transmission measurements: 3 discs



- **3**  $\text{AlO}_3$  discs with diameter 200mm Disc positions **transmission curve**
- Prediction (red) fits measurement (black) well.

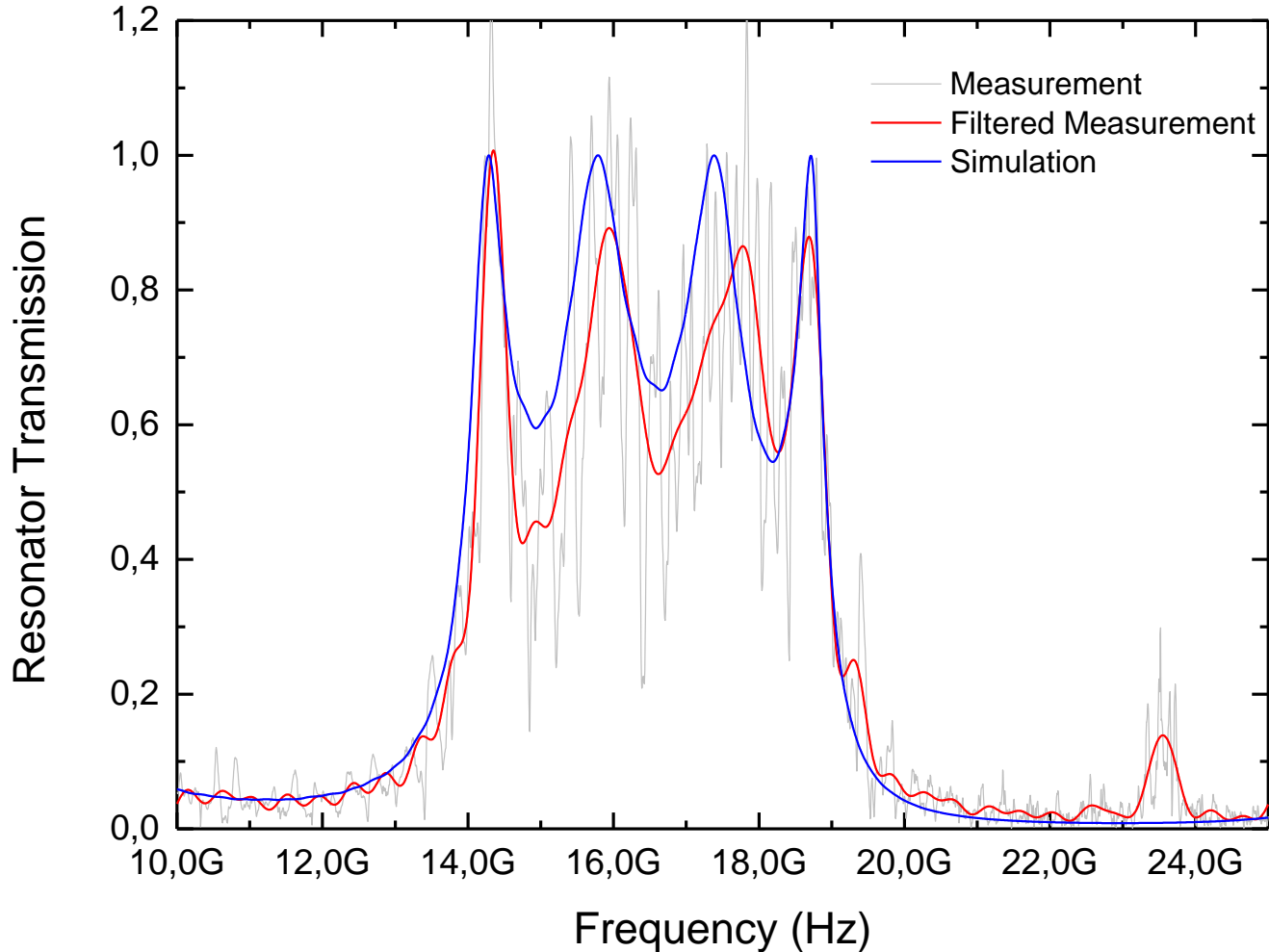
## Transmission measurements: 4 discs



- 4  $\text{AlO}_3$  discs with diameter 200mm Disc positions **transmission curve**
- Prediction (red) fits measurement (black) well.

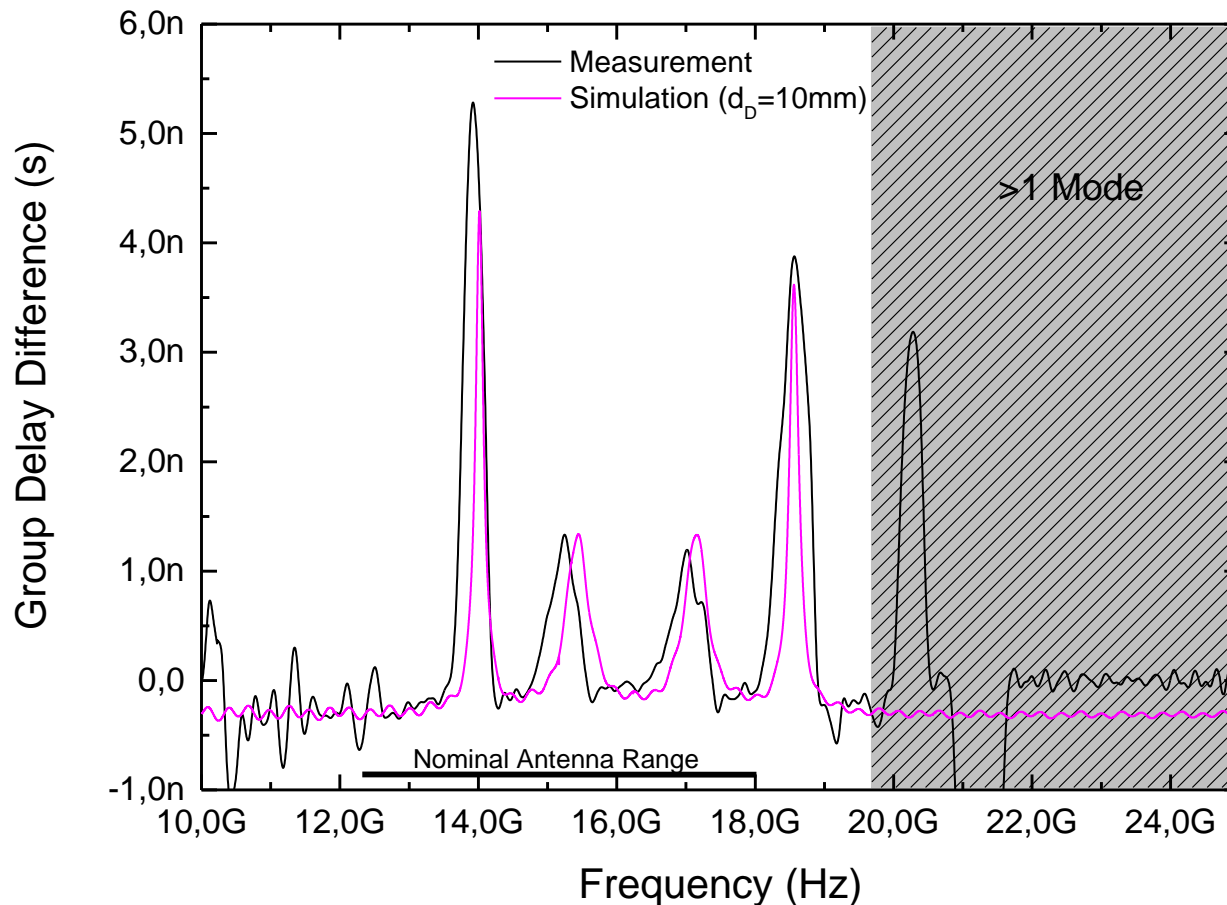


## Transmission measurements: 5 discs



- **5**  $\text{AlO}_3$  discs with diameter 200mm Disc positions **transmission** curve
- Prediction (red) fits measurement (black) well.

## Reflectivity measurement: 4 discs + mirror



- 4  $\text{AlO}_3$  discs with diameter 200mm plus mirror **reflection** curve
- Prediction (red) fits measurement (black) well.

→ Needed precision of disc placing according to expectations

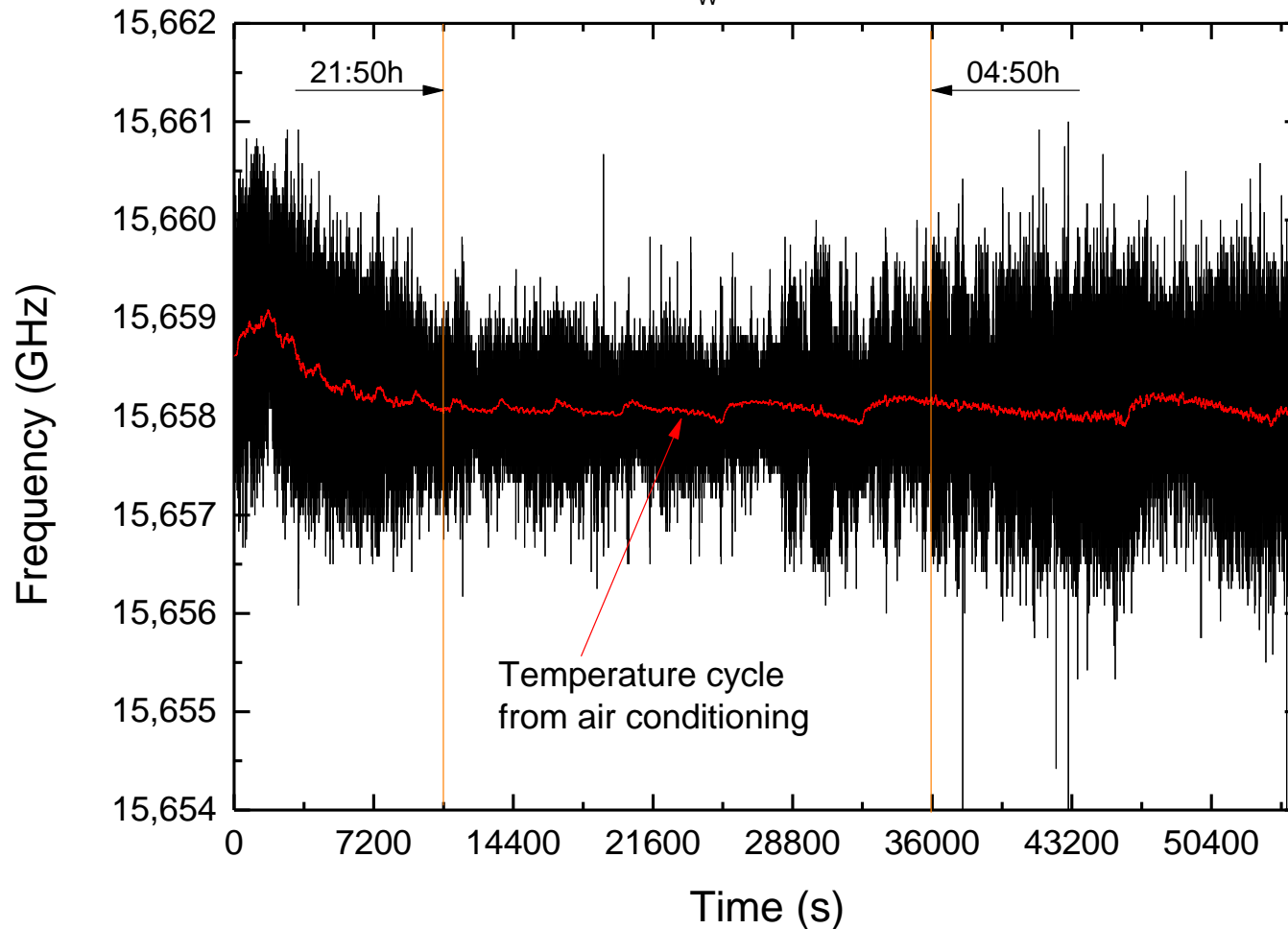
→ Gives confidence in boost factor calculation

12th PATRAS workshop on Axions, WIMPs and WISPs, 2016 June 20-24

## Reflectivity measurement: 4 discs + mirror

Center of group delay peak

$$B_W = 1\text{MHz}$$



Variation of group delay peak position  $\sim 1\text{MHz}$   
 $\rightarrow$  Long term stability of boost peak seems ok



# The 10T dipole magnet challenge:

Remember:

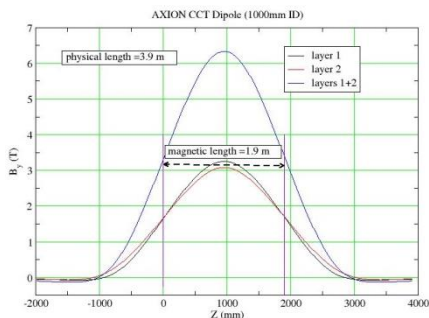
$$P \propto \mathbf{A}_{\text{mirror}} \cdot \mathbf{B}_{\parallel}^2$$

Sensitivity goal assumes use of 10T dipole magnet with 1m aperture!

Two design concepts available:

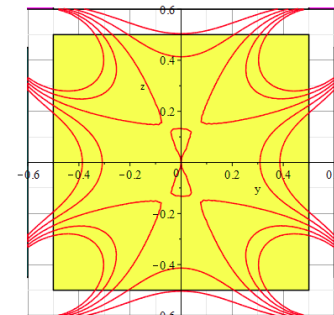
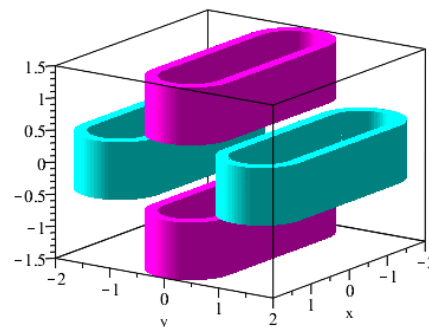
Berkeley superconducting magnet group:

Canted-Cosine-Theta magnet



CEA-IRFU Saclay magnet group:

4 racetracks actively shielded configuration



$\delta B = \pm 0.1 T$   
from 10 T

## Further plans

### 2016:

- Finish first test measurements at room temperature at MPI
- Test noise of preamplifier at LHe temperature
- Find additional collaborators for specific parts of project.  
presently in discussion with:  
**DESY, MPI-Radio astronomy, CEA-Saclay, LBNL-Berkeley**
- Order design study of 10T magnet
- Start to develop technique to cover frequencies above 40 GHz
- R&D on production of large diameter high- $\epsilon$  discs

### 2017-2020:

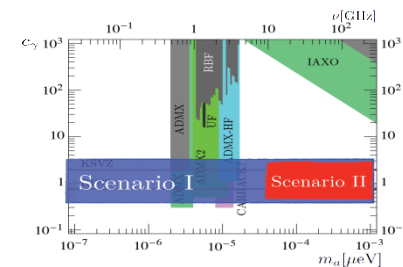
- R&D and Prototype experiments, tests in high B-field
- Design (and construction) of full scale experiment

### 2020 (?) :

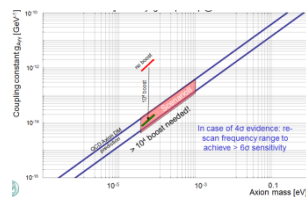
- Start full scale experiment

# CONCLUSIONS

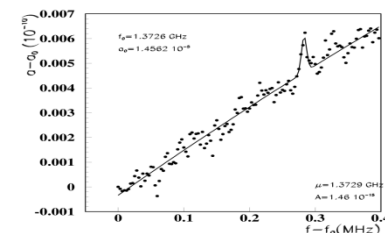
- Axions in mass range  $\sim 40\text{-}400 \mu\text{eV}$  could **solve strong CP & Dark Matter problem**



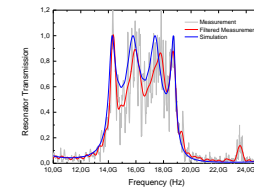
- Open dielectric structure with  $\sim 80$  discs expected to **boost axion detection rate by  $\sim 5$  orders of magnitude**



- First measurements with low noise preamp promising: 80 big enough discs in 10 T B-field: **sensitivity enough to probe models**



- First results from proof of principle setup are promising
- Magnet issue seems solvable (price?)



- We are ready to „go official“

