



Long-lived particles and other unusual signatures

Loïc Quertenmont

Terascale school @ DESY
07 April 2016



WARNING

- This talk is
 - Not exhaustive
 - Not about theory motivation for LL
 - Biased toward what I know (CMS)
 - About experimental techniques used @LHC
 - Not exhaustive at all !

- Predicted by many BSM models:
 - Split SUSY, GMSB, AMSB, pMSSM, Hidden Valley, RPV SUSY, UED, ...
 - Long lifetime is a consequence of :
 - a new symmetry conservation (R-Parity, KK-Parity)
 - a small decay phase space
- Constraints from prompt particles searches generally don't apply.
 - Lifetime may explain why SUSY was not yet discovered at the LHC ☺
- Who knows what BSM looks like
 - Search in every single corner
 - Model independent searches when possible
 - Agnostic searches when possible

Non-exhaustive list of models predicting LL charged particles

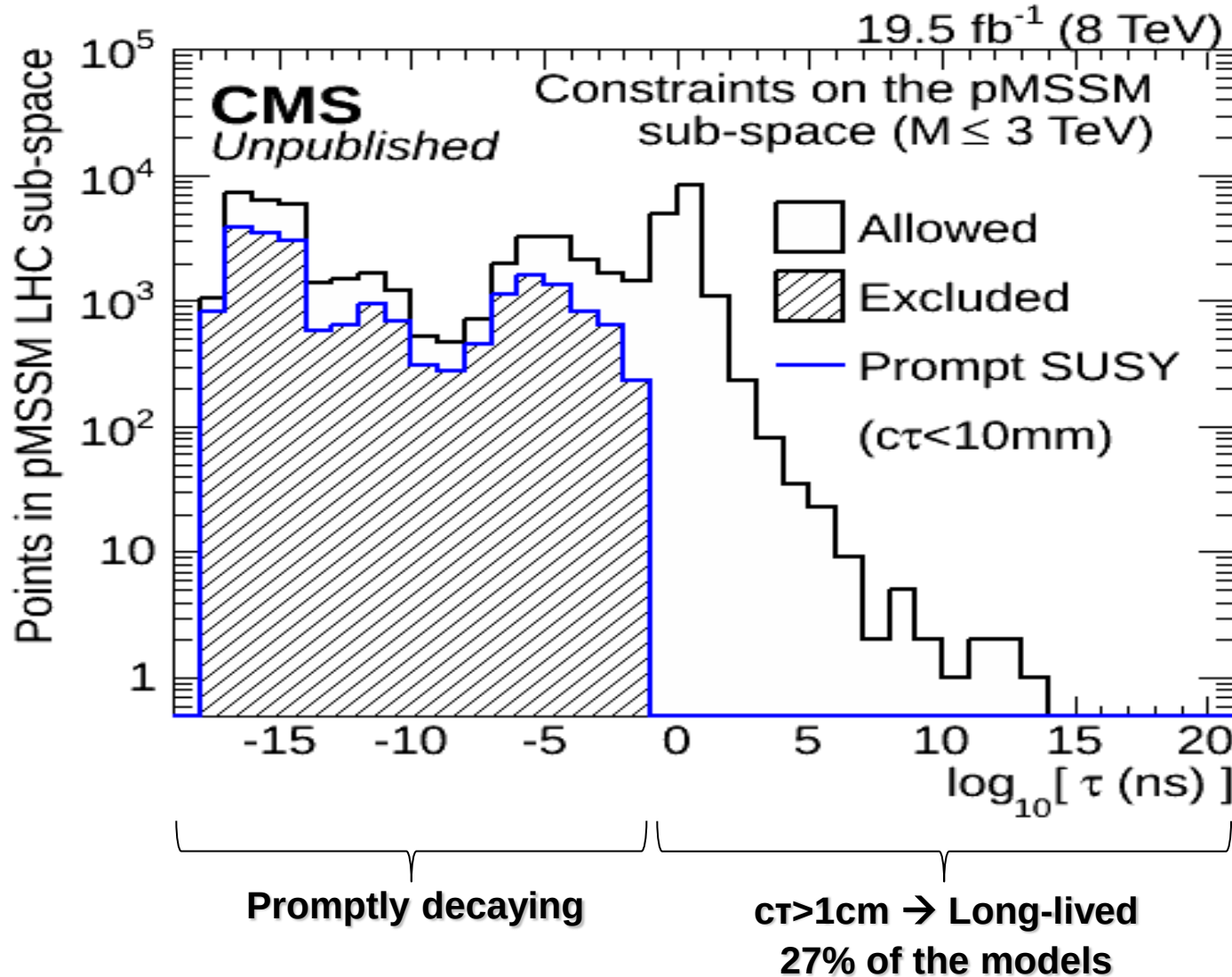
SMP	LSP	Scenario	Conditions	SUSY
$\tilde{\tau}_1$	$\tilde{\chi}_1^0$	MSSM	$\tilde{\tau}_1$ mass (determined by $m_{\tilde{\tau}_{L,R}}^2$, μ , $\tan\beta$, and A_τ) close to $\tilde{\chi}_1^0$ mass.	
	$\tilde{G}$	GMSB	Large N , small M , and/or large $\tan\beta$.	
		$\tilde{g}$ MSB	No detailed phenomenology studies, see [23].	
		SUGRA	Supergravity with a gravitino LSP, see [24].	
$\tilde{\tau}_1$		MSSM	Small $m_{\tilde{\tau}_{L,R}}$ and/or large $\tan\beta$ and/or very large A_τ .	
		AMSB	Small m_0 , large $\tan\beta$.	
		$\tilde{g}$ MSB	Generic in minimal models.	
$\tilde{t}_{41}$	$\tilde{G}$	GMSB	$\tilde{\tau}_1$ NLSP (see above). $\tilde{e}_1$ and $\tilde{\mu}_1$ co-NLSP and also SMP for small $\tan\beta$ and μ .	
	$\tilde{\tau}_1$	$\tilde{g}$ MSB	$\tilde{e}_1$ and $\tilde{\mu}_1$ co-LSP and also SMP when stau mixing small.	
$\tilde{\chi}_1^+$	$\tilde{\chi}_1^0$	MSSM	$m_{\tilde{\chi}_1^+} - m_{\tilde{\chi}_1^0} \lesssim m_{\pi^+}$. Very large $M_{1,2} \gtrsim 2 \text{ TeV} \gg \mu $ (Higgsino region) or non-universal gaugino masses $M_1 \gtrsim 4M_2$, with the latter condition relaxed to $M_1 \gtrsim M_2$ for $M_2 \ll \mu $. Natural in O-II models, where simultaneously also the $\tilde{g}$ can be long-lived near $\delta_{\text{GS}} = -3$.	
		AMSB	$M_1 > M_2$ natural. m_0 not too small. See MSSM above.	
$\tilde{g}$	$\tilde{\chi}_1^0$	MSSM	Very large $m_{\tilde{q}}^2 \gg M_3$, e.g. split SUSY.	
	$\tilde{G}$	GMSB	SUSY GUT extensions [25–27].	
	$\tilde{g}$	MSSM	Very small $M_3 \ll M_{1,2}$, O-II models near $\delta_{\text{GS}} = -3$.	
		GMSB	SUSY GUT extensions [25–29].	
$\tilde{t}_1$	$\tilde{\chi}_1^0$	MSSM	Non-universal squark and gaugino masses. Small $m_{\tilde{q}}^2$ and M_3 , small $\tan\beta$, large A_t .	
$\tilde{b}_1$			Small $m_{\tilde{q}}^2$ and M_3 , large $\tan\beta$ and/or large $A_b \gg A_t$.	

Q_{em}	C_{QCD}	S	Model(s)	EXOTIC
0	8	1	Universal Extra Dimensions (KK gluon)	
± 1	1	$\frac{1}{2}$	Universal Extra Dimensions (KK lepton)	
			Fat Higgs with a fat top (ψ fermions)	
			4th generation (chiral) fermions	
			Mirror and/or vector-like fermions	
		0	Fat Higgs with a fat top (ψ scalars)	
$\pm \frac{4}{3}$	3	$\frac{1}{2}$	Warped Extra Dimensions with GUT parity (XY gaugino)	
		0	5D Dynamical SUSY-breaking (xyon)	
$-\frac{1}{3}, \frac{2}{3}$	3	$\frac{1}{2}$	Universal Extra Dimensions (KK down, KK up)	
			4th generation (chiral) fermions	
			Mirror and/or vector-like fermions	
			Warped Extra Dimensions with GUT parity (XY gaugino)	
$\epsilon < 1$	1	$\frac{1}{2}$	GUT with $U(1) - U(1)'$ mixing	
			Extra singlets with hypercharge $Y = 2\epsilon$	
			Millicharged neutrinos	
?	?	$0/\frac{1}{2}/1$	“Technibaryons”	

+++ many models predicting neutral LL
 +++ other more exotic models/particles
 (monopoles, quirks, dyons, etc.)

[arXiv:hep-ph/0611040](https://arxiv.org/abs/hep-ph/0611040) Fairbairn et al.

Example with the pMSSM model (generic R-parity conserving SUSY)

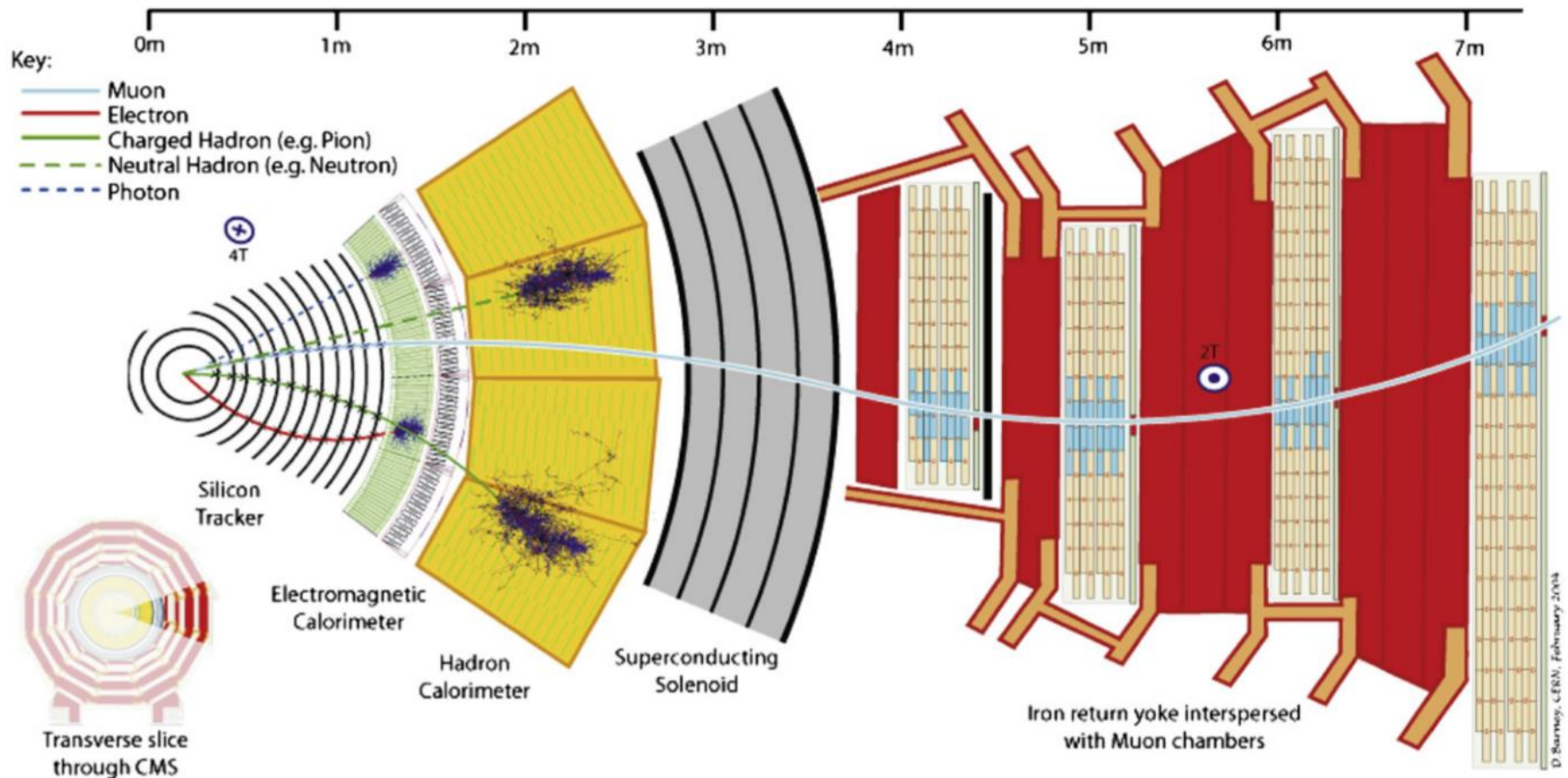


Each entry correspond to a point in the 19D pMSSM parameter space

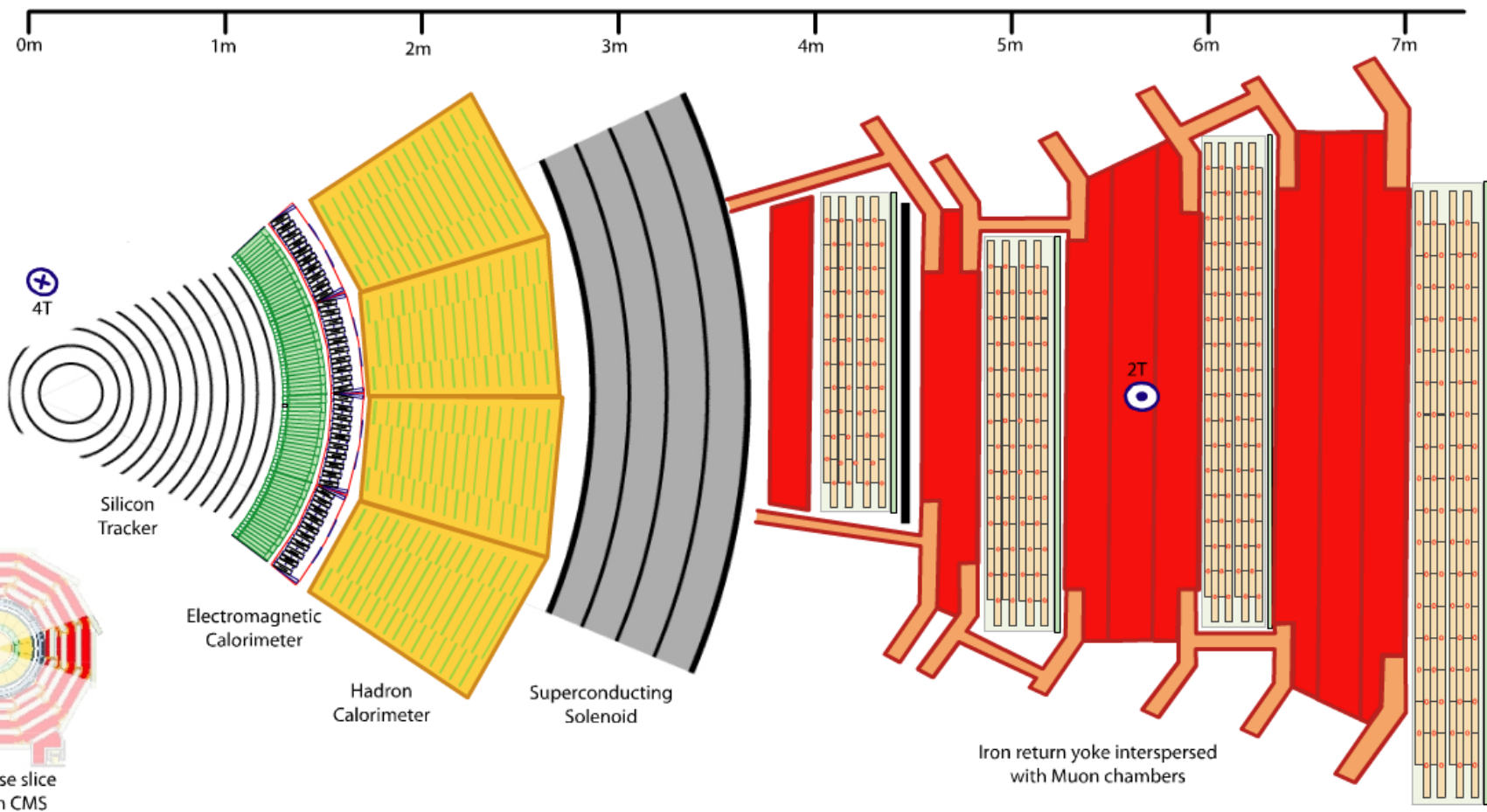
$\rightarrow$ each point is a different SUSY model

NLSP (chargino) lifetime

- **Modern detectors are**
 - Made of layers optimized to reconstruct and identify SM particles
 - Photons, Electrons, Hadrons, Muons, Neutrino (through MET)



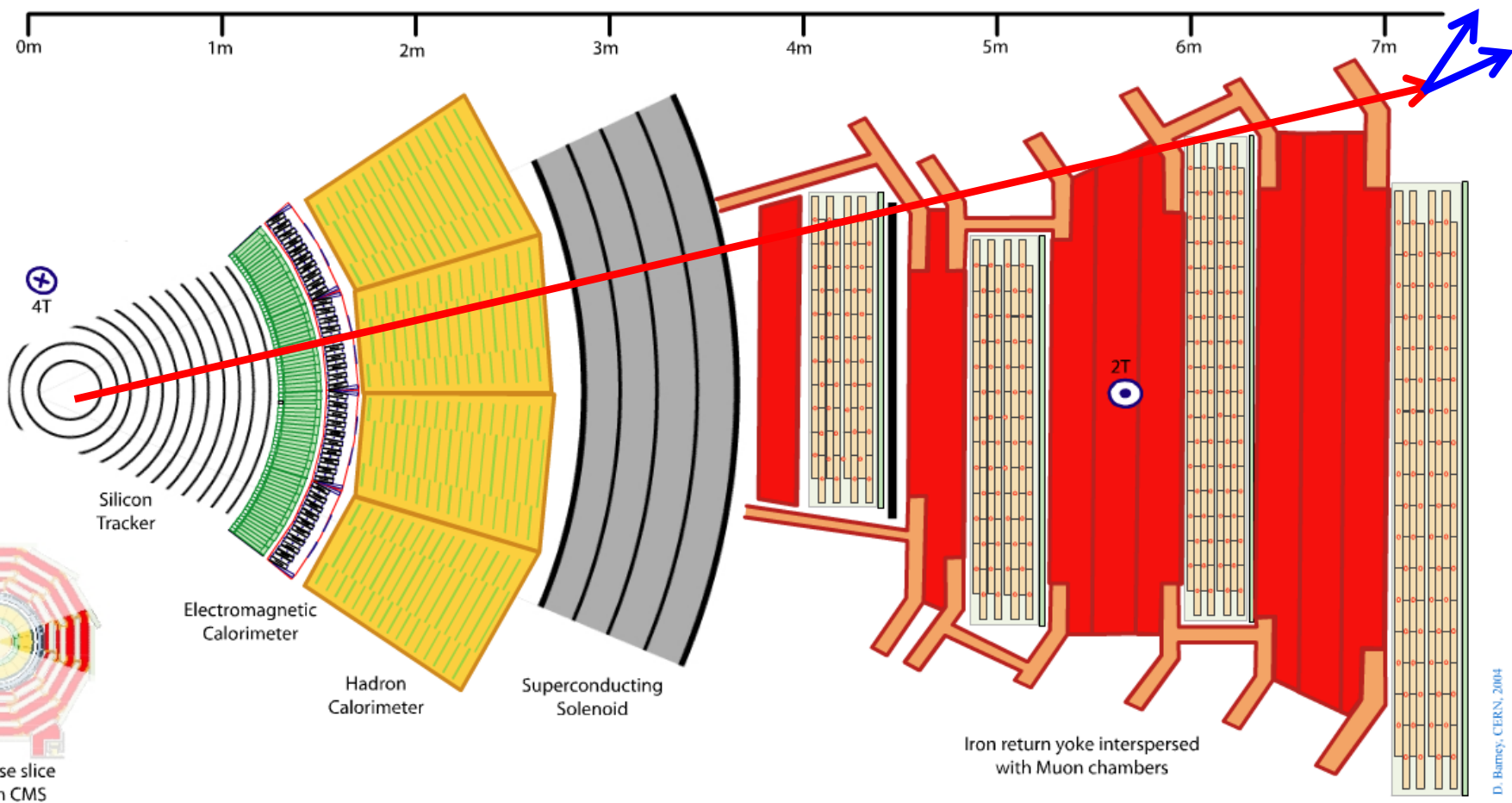
- **Modern detectors are**
 - Made of layers optimized to reconstruct and identify SM particles
 - Photons, Electrons, Hadrons, Muons, Neutrino (through MET)
 - Very efficient for SM physics measurements
 - Very efficient to search for BSM particles
 - Promptly decaying to SM particles (leptons or jets)
 - Being invisible (but detectable via MET like SM neutrinos)
- **Searching for Long-Lived BSM particles is more challenging...**
 - **Strategy depends on the particle type and lifetime**



$C\tau \geq O(m)$

Quasi-stable

Searches based on
ionization/velocity/calor



$C\tau \leq O(m)$

$C\tau \geq O(m)$

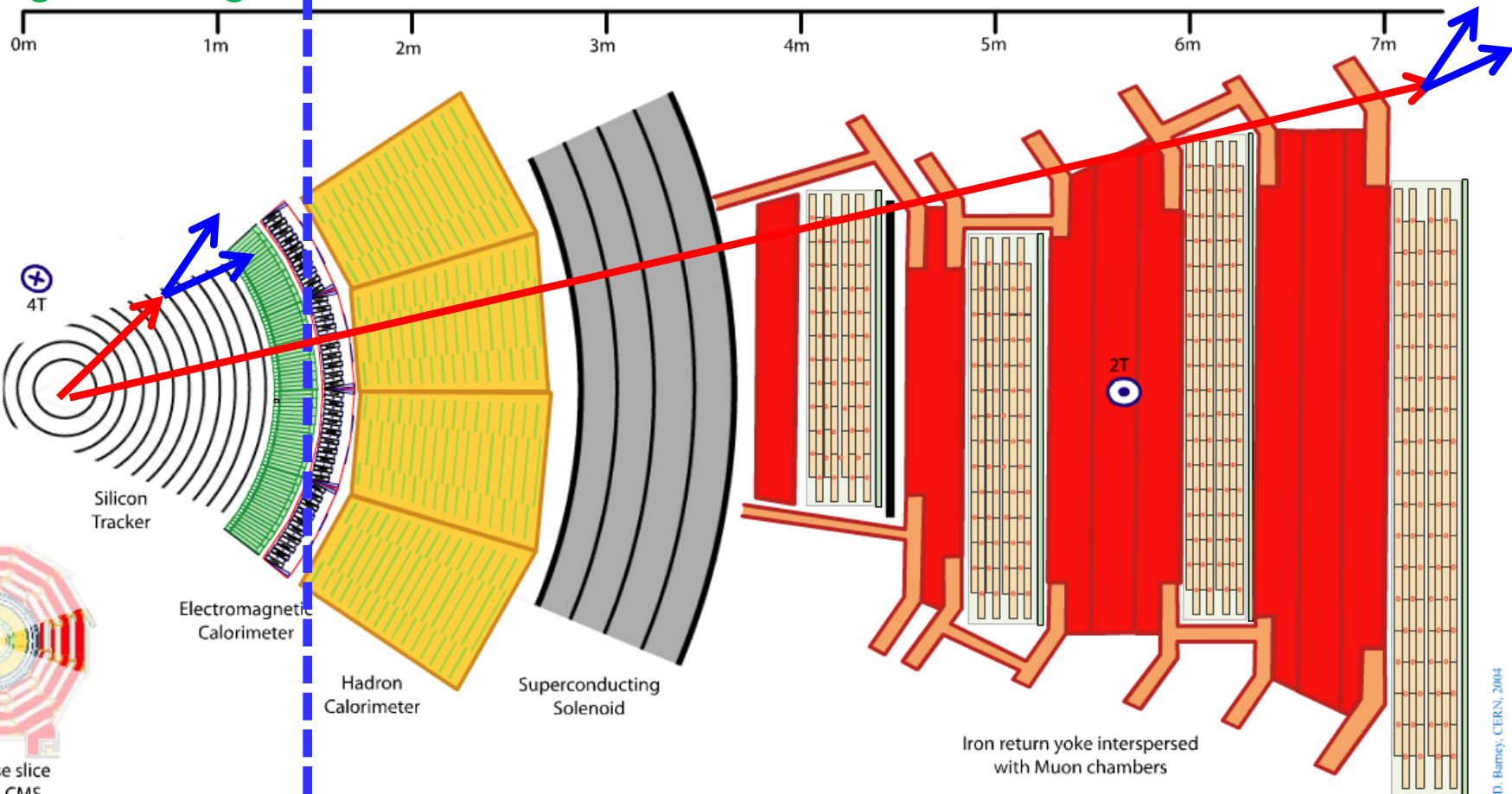
Decay in flight

Quasi-stable

Searches based on
tracking/vertexing

Searches based on
ionization/velocity/calor

Not strict boundary

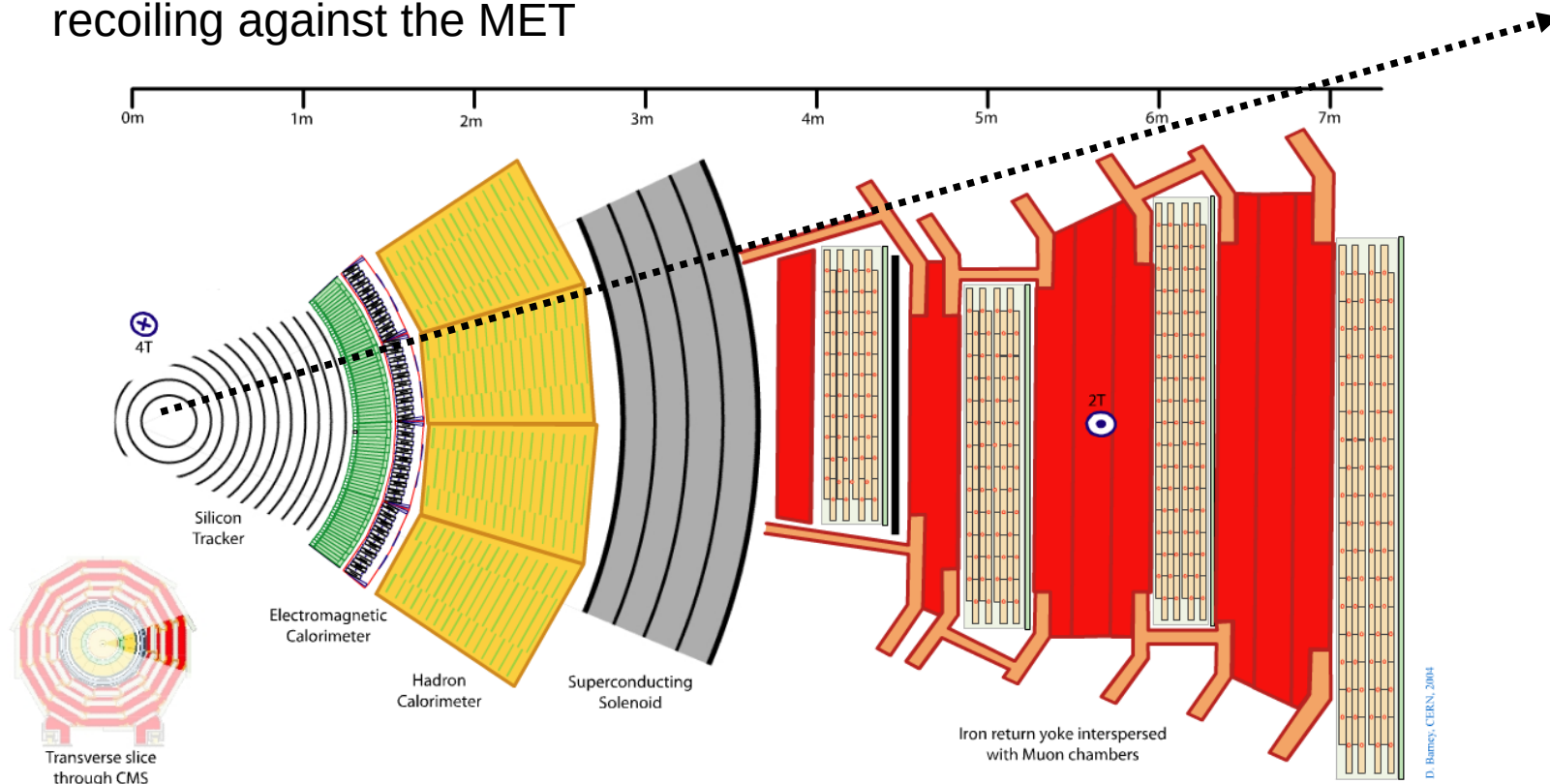


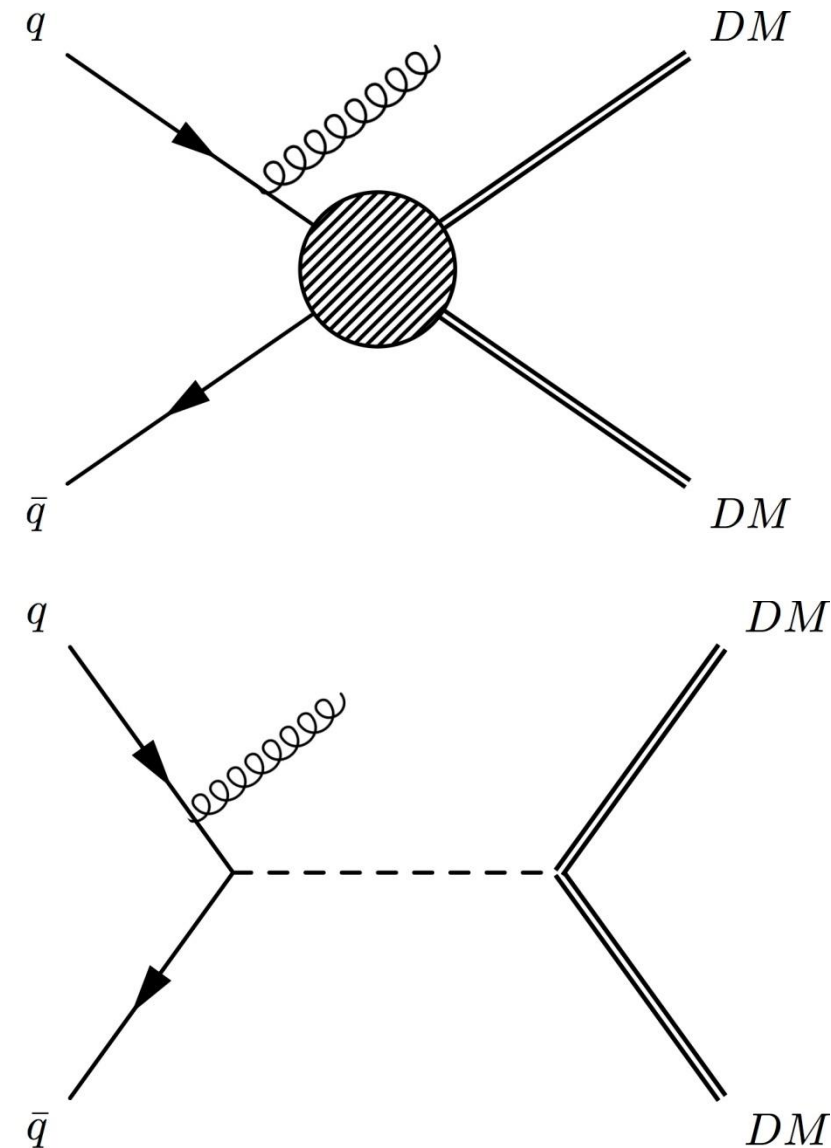


- **Introduction to LL physics (Why ? How ?)** ✓
- **Searches for LL particles decaying OUTSIDE the detector**
 - **Neutral**
 - **Charged**
- **Searches for LL particles decaying INSIDE the detector**
 - **Charged**
 - **Neutral**
- **Super exotic particles**
- **Summary**

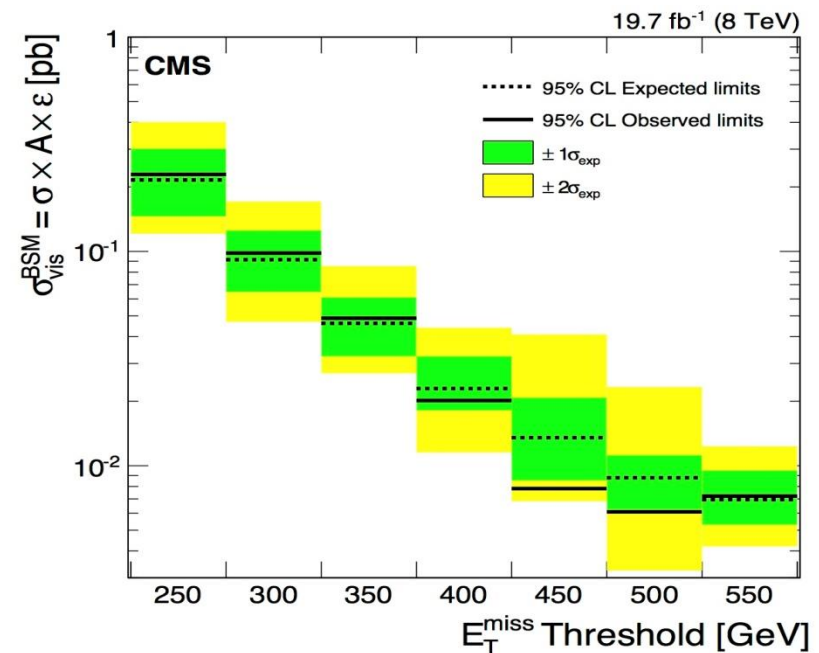
- Introduction to LL physics (Why ? How ?) ✓
- Searches for LL particles decaying OUTSIDE the detector
 - Neutral ←
 - Charged
- Searches for LL particles decaying INSIDE the detector
 - Charged
 - Neutral
- Super exotic particles
- Summary

- Direct detection is not possible
 - Similar to SM neutrino
- Indirect observation via MET
 - Covered by many SUSY/DM searches looking for mono-X objects recoiling against the MET





- The CMS monojet search (arXiv:1408.3583) explicitly limits pair production of dark matter particles accompanied by ISR jet.
- The same limits apply to neutral LL particles that decay outside CMS ($R \gtrsim 10$ m).



- Introduction to LL physics (Why ? How ?) ✓
- Searches for LL particles decaying OUTSIDE the detector
 - Neutral ✓
 - Charged ←
- Searches for LL particles decaying INSIDE the detector
 - Charged
 - Neutral
- Super exotic particles
- Summary

What are we talking about ?

A rich phenomenology, dictating the experimental signature

- **Heavy**
- **Stable**
- **Charged**

What are we talking about ?

A rich phenomenology, dictating the experimental signature

- **Heavy**

Slow particle $\rightarrow \beta < 1$

- **Stable**

- **Charged**

A rich phenomenology, dictating the experimental signature

- **Heavy**

Slow particle $\rightarrow \beta < 1$

- **Stable $\rightarrow$ Long-lived**

1) Escape the tracker

$\rightarrow c\tau > O(1)m$

2) Escape the detector

$\rightarrow c\tau > O(10)m$

3) Stop in $\rightarrow$ Asynchronous decay

$\rightarrow c\tau > O(100)m$

- **Charged**

A rich phenomenology, dictating the experimental signature

- **Heavy**

Slow particle $\rightarrow \beta < 1$

- **Stable $\rightarrow$ Long-lived**

1) Escape the tracker

$\rightarrow c\tau > O(1)m$

2) Escape the detector

$\rightarrow c\tau > O(10)m$

3) Stop in $\rightarrow$ Asynchronous decay

$\rightarrow c\tau > O(100)m$

- **Charged**

Electric $\rightarrow$ lepton-like

$|Q| \ll e, |Q| < e, |Q| = e, |Q| > e, |Q| \gg e$

Color $\rightarrow$ hadron-like $\rightarrow$ R-hadron properties ?!
(gluinoball, charge flip, interaction w/ matter, etc.)

A rich phenomenology, dictating the experimental signature

- **Heavy**

Slow particle $\rightarrow \beta < 1$

- **Stable $\rightarrow$ Long-lived**

1) Escape the tracker

$\rightarrow c\tau > O(1)m$

2) Escape the detector

$\rightarrow c\tau > O(10)m$

3) Stop in $\rightarrow$ Asynchronous decay

$\rightarrow c\tau > O(100)m$

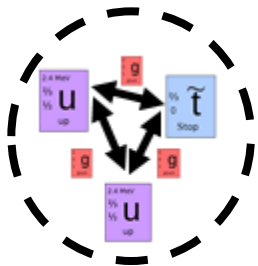
- **Charged**

IONIZATION

{	Electric	$\rightarrow$ lepton-like	
	Color	$\rightarrow$ hadron-like	$\rightarrow$ R-hadron properties ?!

(gluinoball, charge flip, interaction w/ matter, etc.)

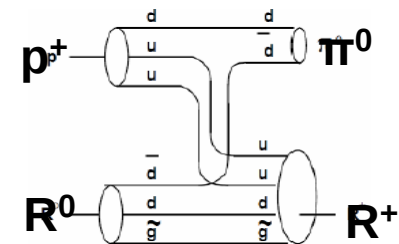
Terminology : HSCP=SMP=ChaMP=CmLLP=...

$$Q = 2/3 + 2/3 - 1/3 = +1$$

$$Q = 2/3 + 2/3 - 1/3 = +1$$
[illegible]

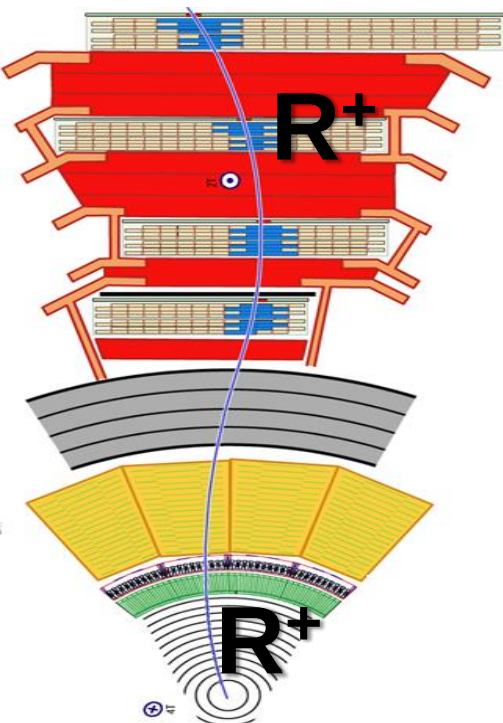
Detector Material

$$Q = \frac{2}{3} - \frac{1}{3} - \frac{1}{3} = 0$$
$$Q = 2/3 + 2/3 + 2/3 = +2$$

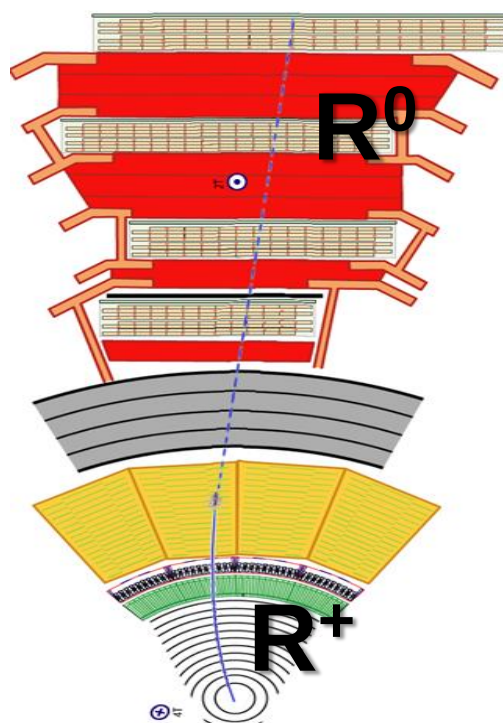
- R-Hadron charge can flip due to nuclear interaction
- Many possible R-Hadron interaction models
→ some predicts that flipping to neutral is more likely
- Gluino ball fraction (f) is also unknown (even if 10% in pythia)



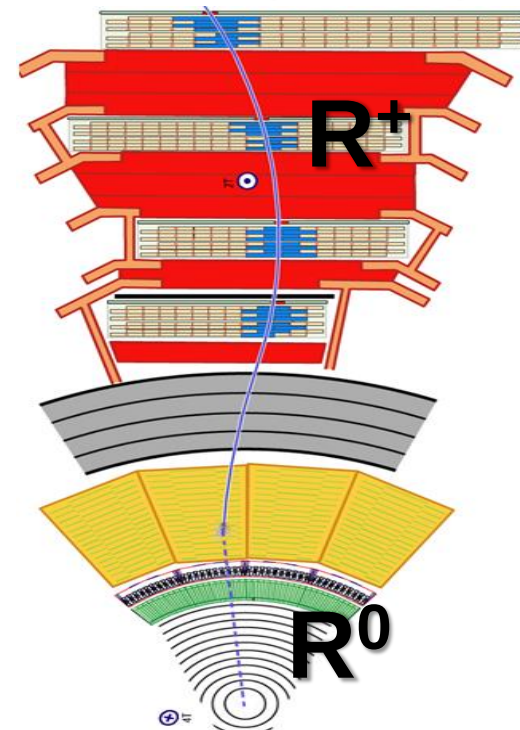
**Stay charge
all along**



**Charged in
tracker only**



**Charged in
muon syst. only**

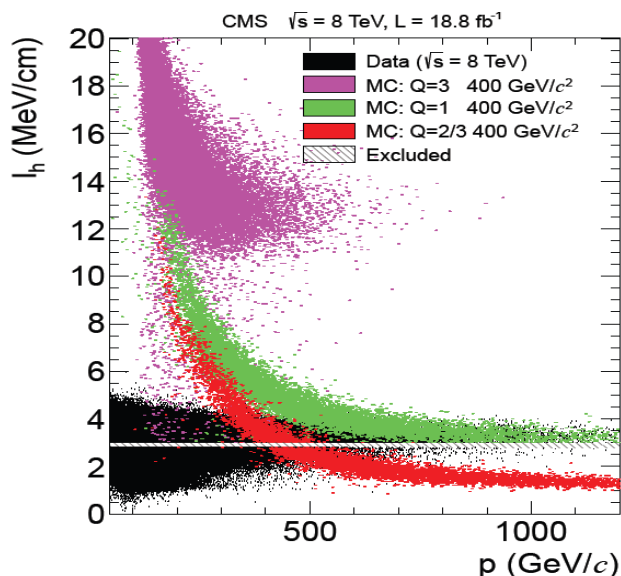


IONIZING



dE/dx

$$\frac{dE}{dx} \approx Q^2 \left(\frac{A}{\beta^2} + B \right) \text{ for } \beta \ll 1$$



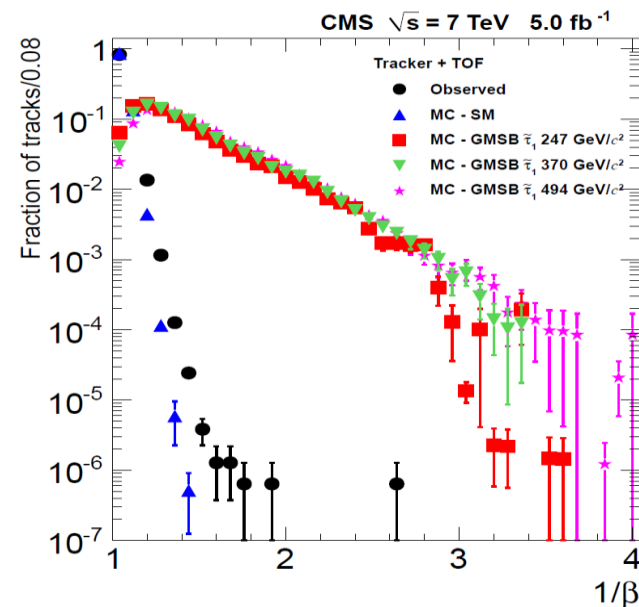
dE/dx resolution improves
with #layers, sensor thickness
and readout electronics

SLOW



Time Of Flight

$$1/\beta = 1 + \frac{c\delta_t}{L} \text{ with } \delta_t = \text{delay}$$



β resolution improves
with time resolution (δt)
and with lever arm (L)

IONIZING



dE/dx

$$\frac{dE}{dx} \approx Q^2 \left(\frac{A}{\beta^2} + B \right) \text{ for } \beta \ll 1$$

SLOW



Time Of Flight

$$1/\beta = 1 + \frac{c\delta_t}{L} \text{ with } \delta_t = \text{delay}$$

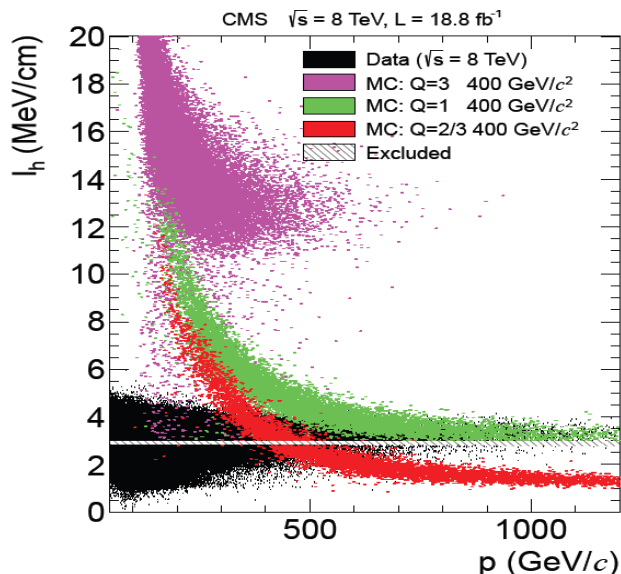
STOPPED



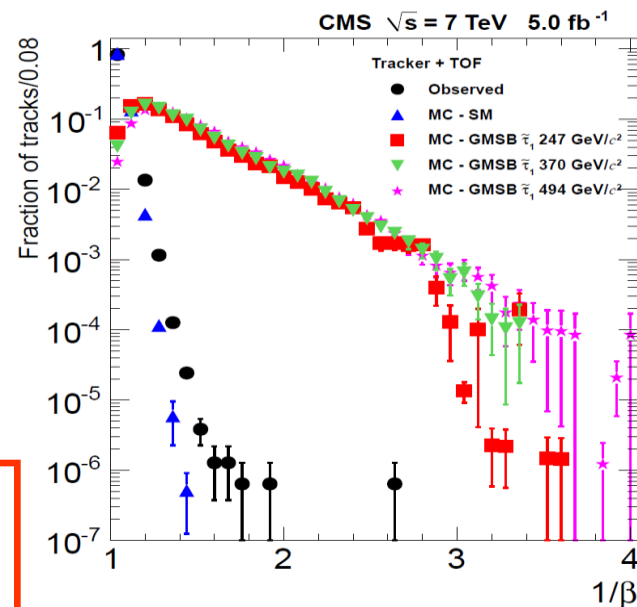
**Decay
asynchronously
w.r.t LHC bunch**



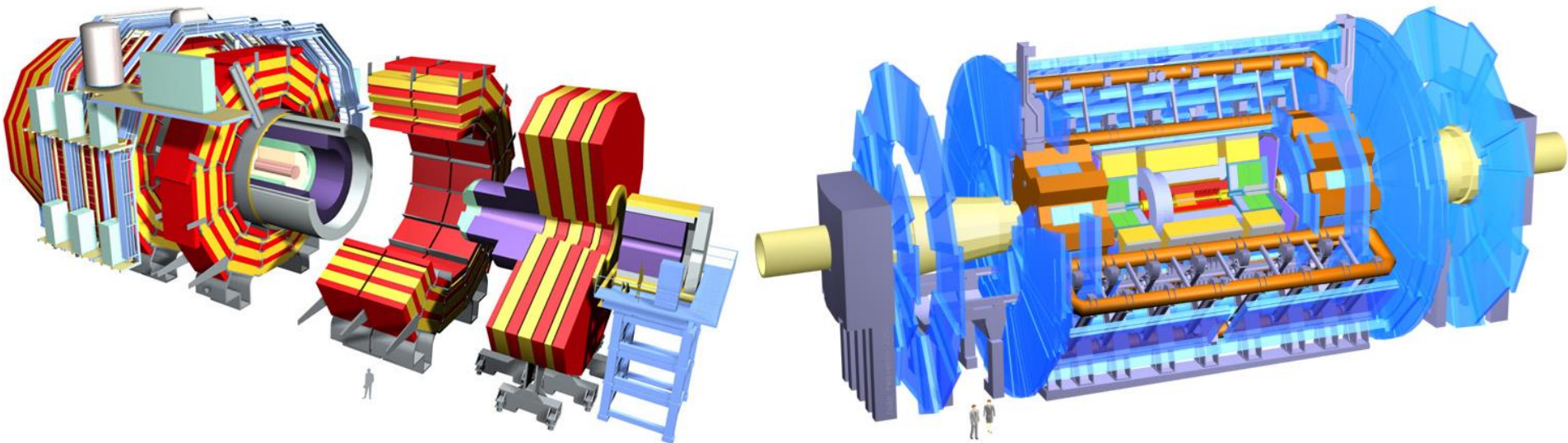
If dE/dx is (high)
comparable to the
kinetic energy
The particle may
stop in the detector
and decay later



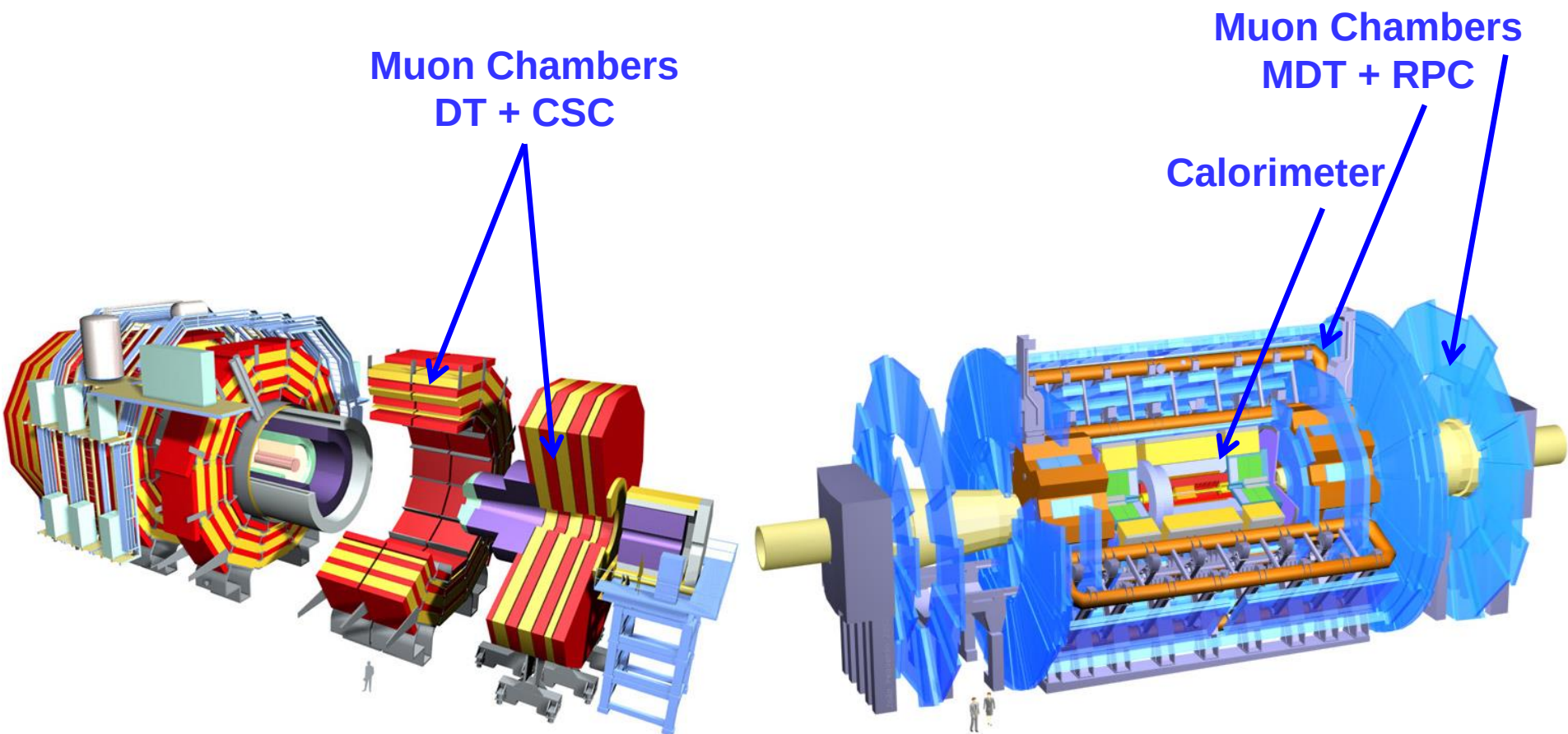
dE/dx resolution improves
with #layers, sensor thickness
and readout electronics



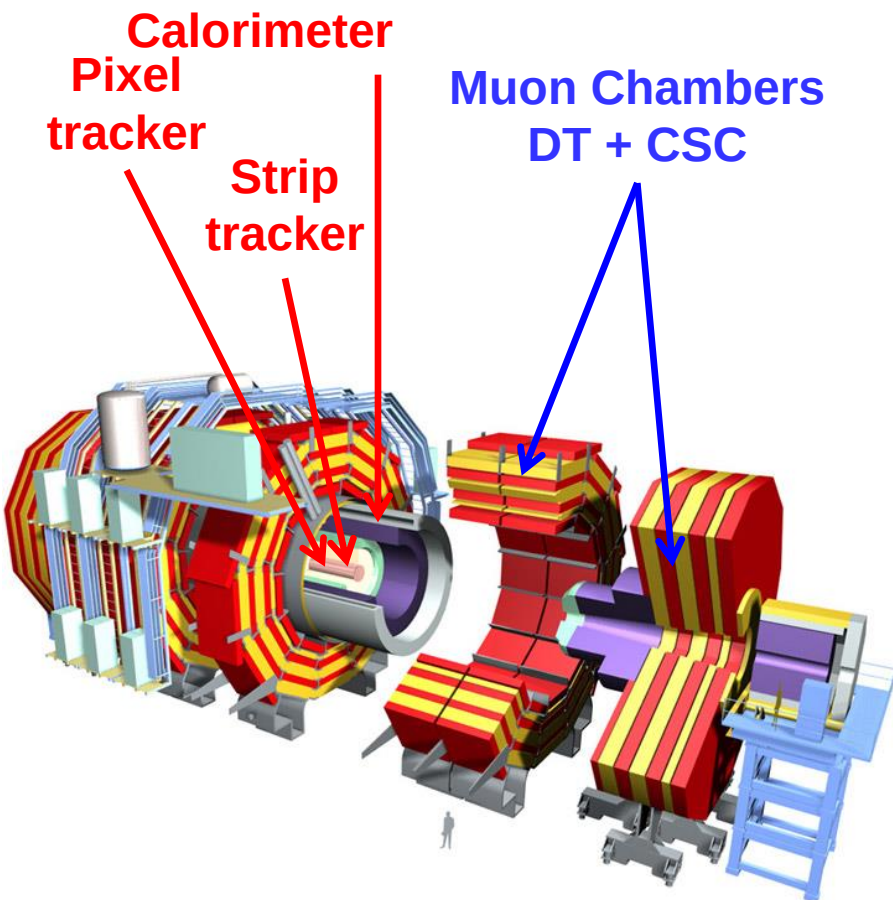
β resolution improves
with time resolution (δt)
and with lever arm (L)



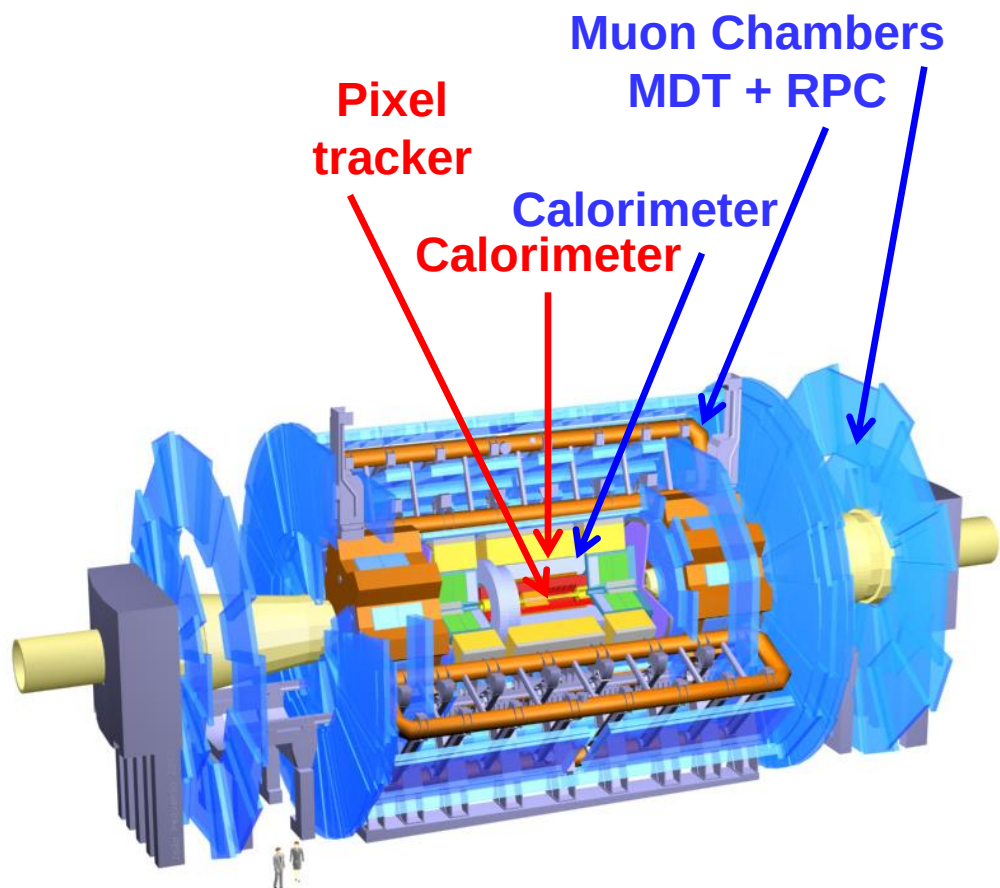
Time of Flight



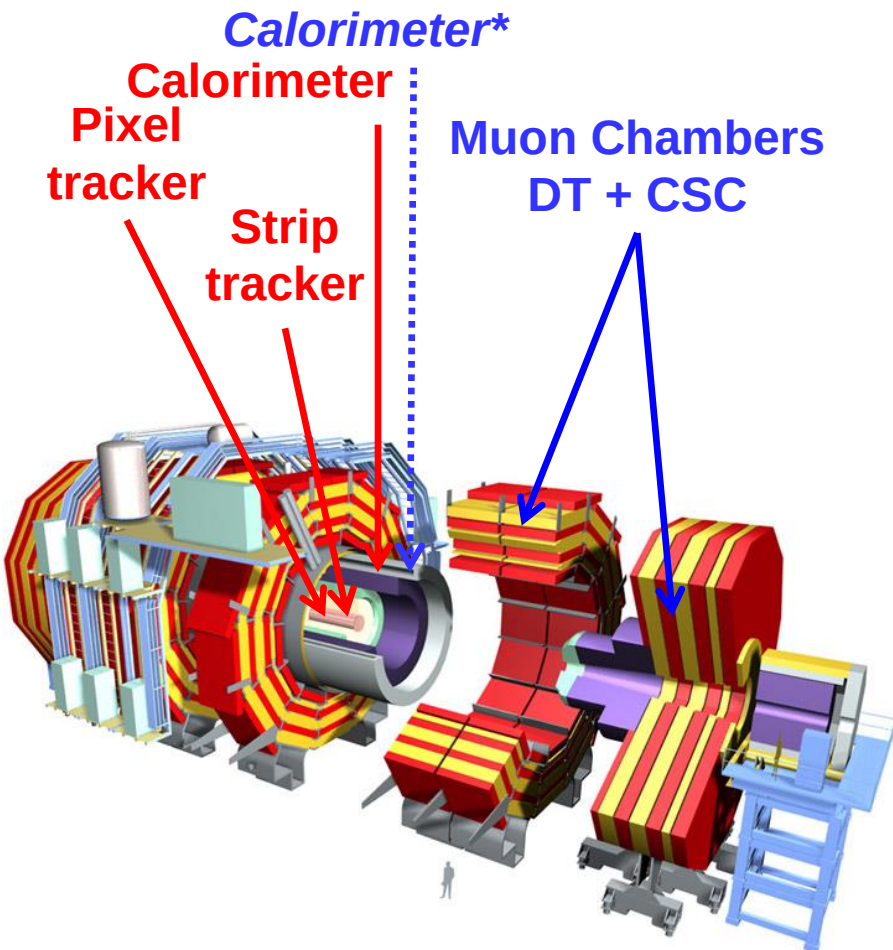
dE/dx



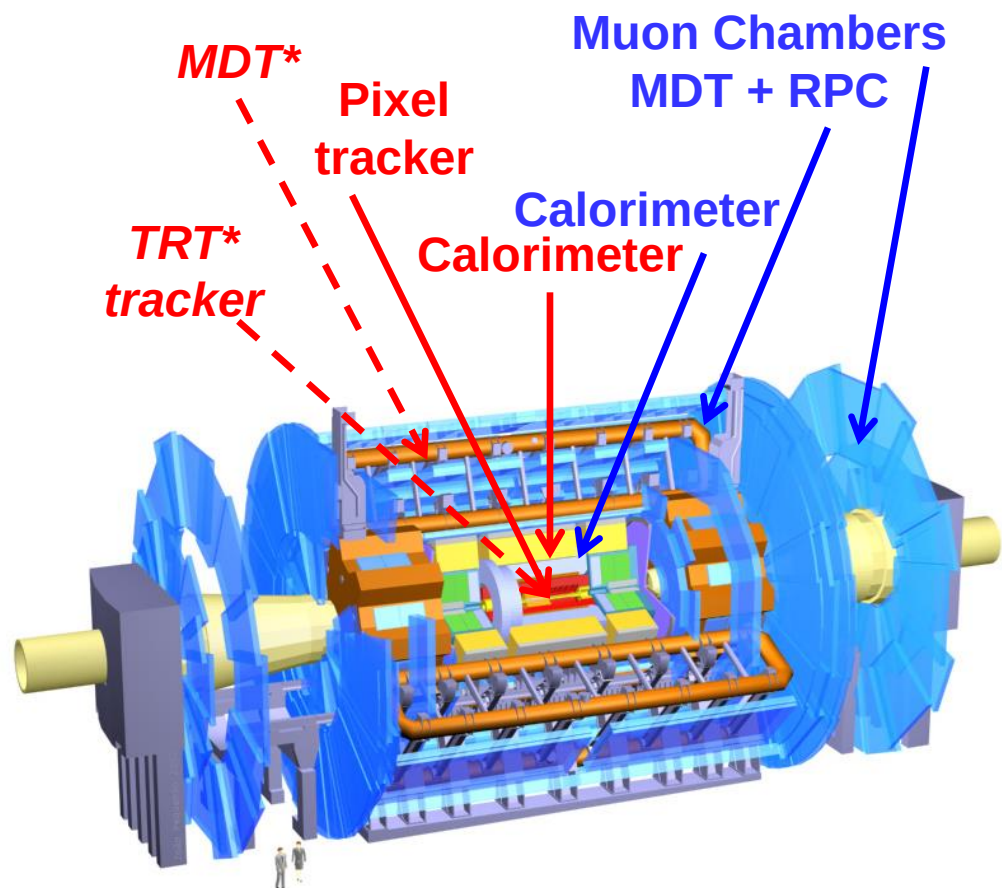
Time of Flight



dE/dx

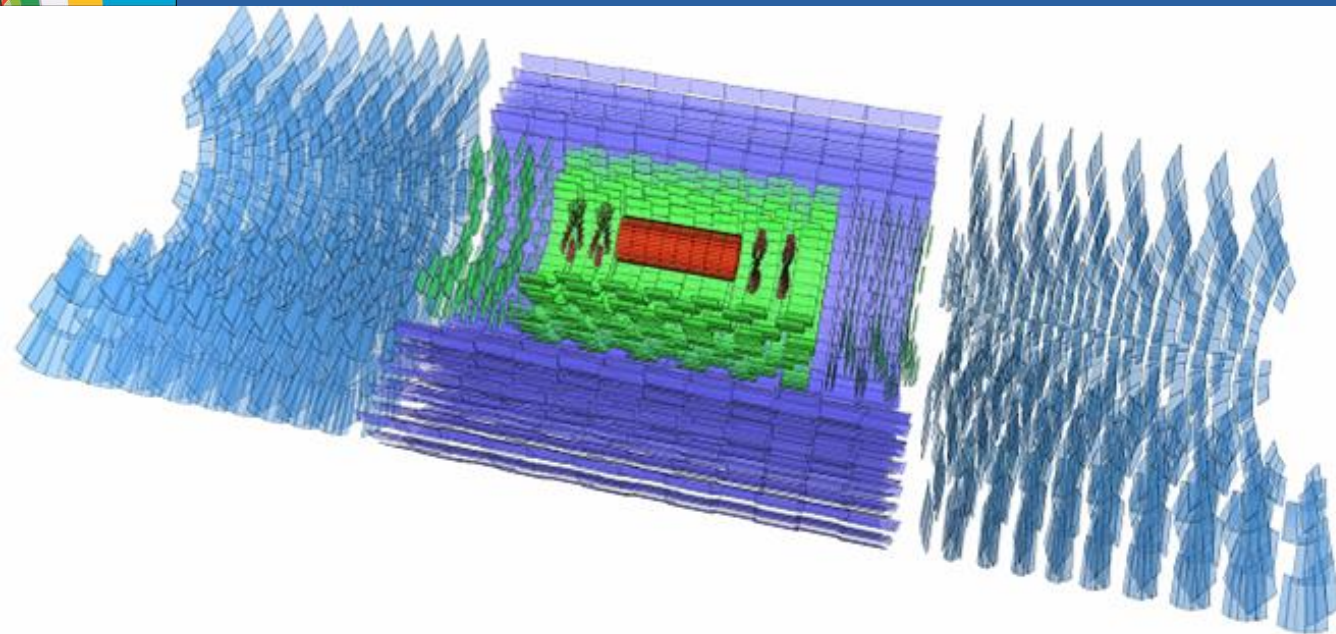


Time of Flight



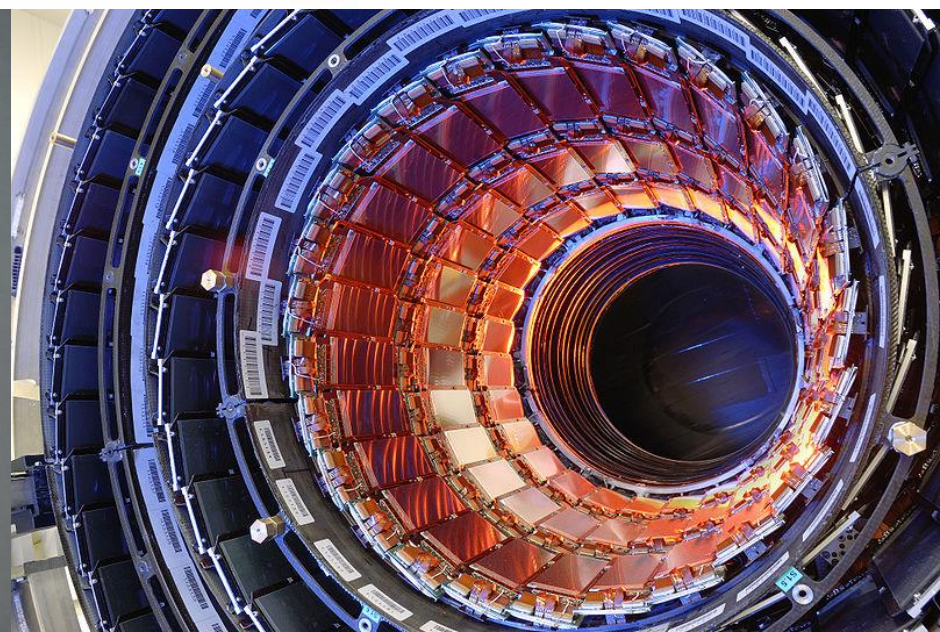
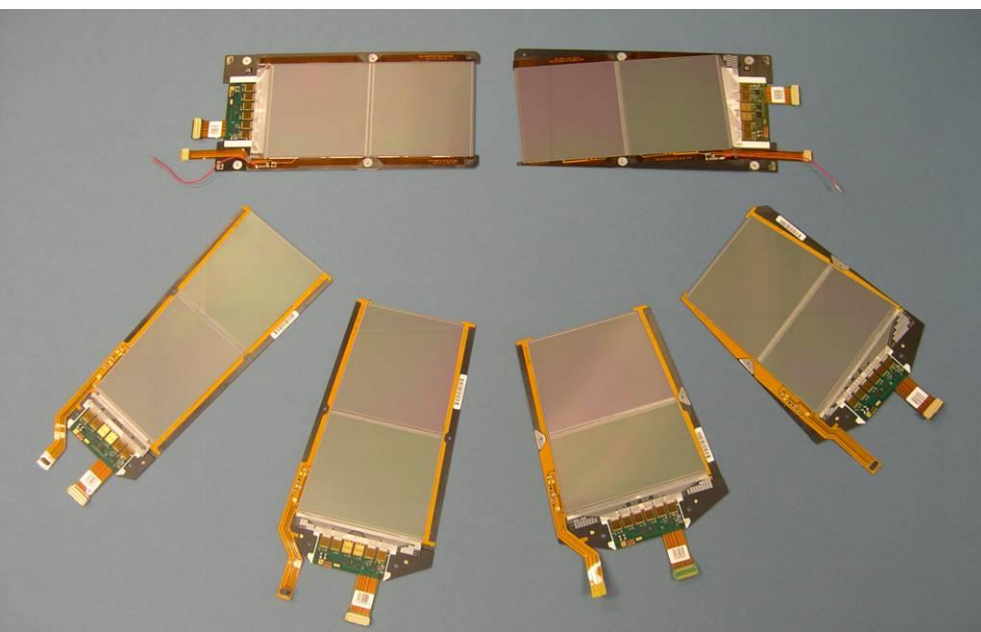
*** = only if $dE/dx \gg O(1)$ GeV**

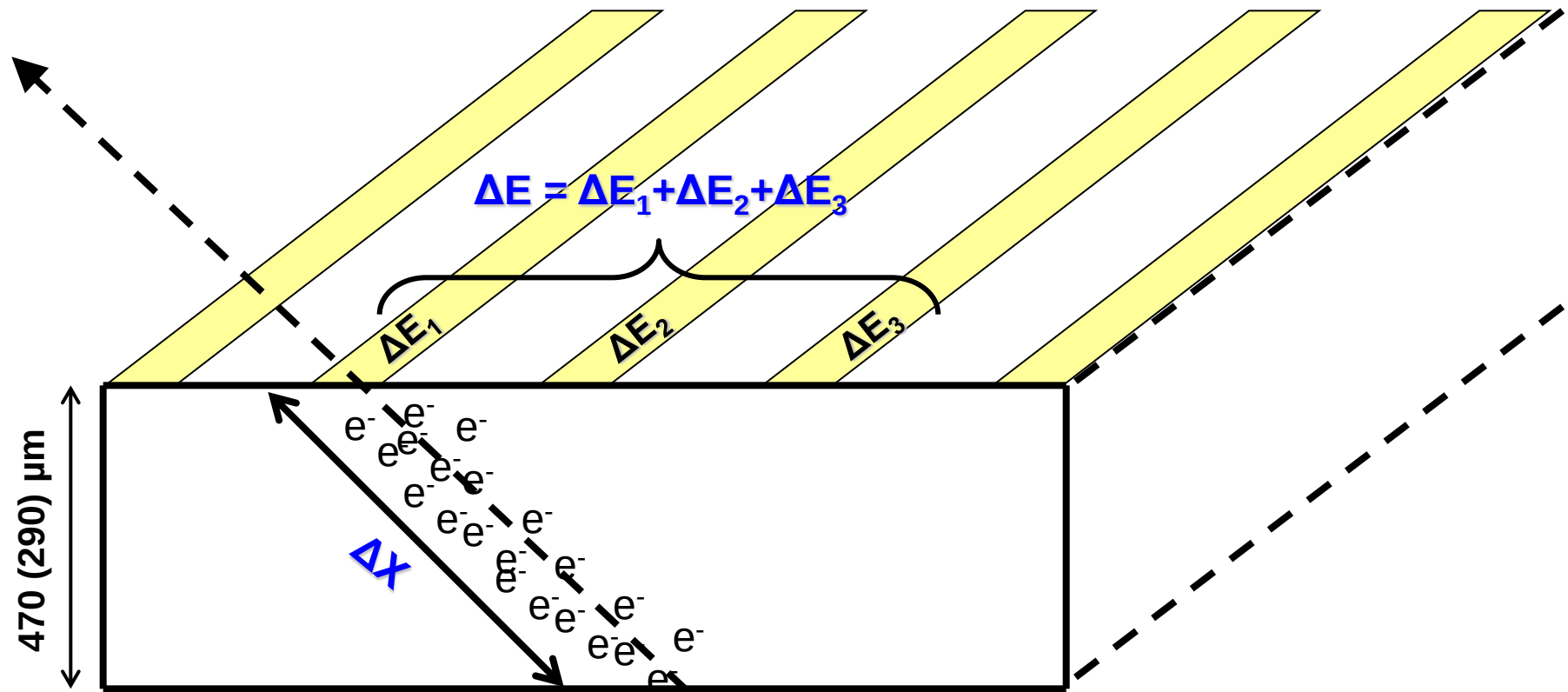
*** = only in special stream**



Strip Tracker
15148 Modules
9.7M Channels

**A particle goes
through ~15
modules**





$$\frac{\Delta E}{\Delta x} \cong K \frac{m^2}{p^2} + C \left[\frac{1}{2} \ln \frac{2 m^2 v_{\text{max}}^2}{I^2 \beta^2} - \frac{0.6009}{\beta^2} \right]$$

Slow particles lose more energy

- **dEdx discriminator - Asymmetric Smirnov (I_{as}) :**

- Main variables to **discriminate signal from backgrounds**
- Probability that a track as a higher energy loss than a MIP
- P_i are hit-level probability extracted from templates of dEdx vs dx for MIPs in minimum bias samples.

$$I_{as} = \frac{3}{N} \times \left(\frac{1}{12N} + \sum_{i=1}^N \left[P_i \times \left(P_i - \frac{2i-1}{2N} \right)^2 \right] \right)$$

N in [10, 30]
<N> = ~17

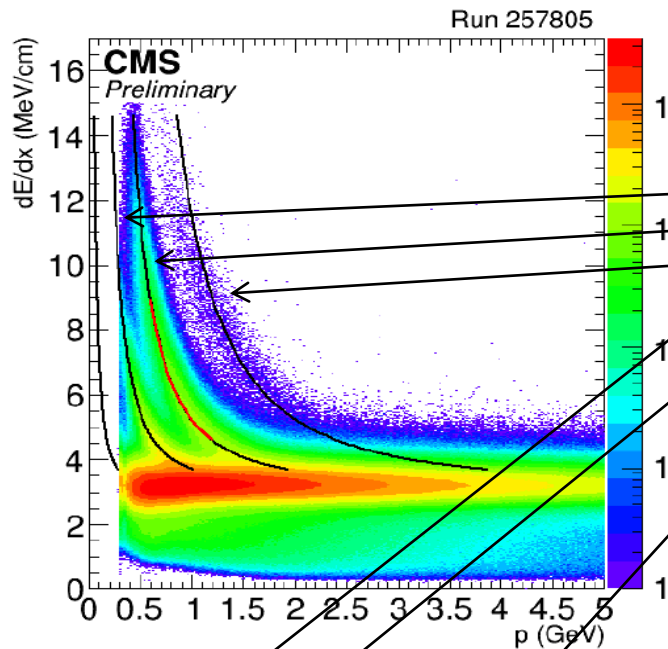
- **dEdx estimator - Harmonic-2 :**

- **only used to reconstruct mass** of HSCP candidates and separate S from Bkg.

$$I_h = \left(\frac{1}{N} \sum_i c_i^{-2} \right)^{-1/2} = K \frac{m^2}{p^2} + C.$$

mass
 momentum
 K,C= constants

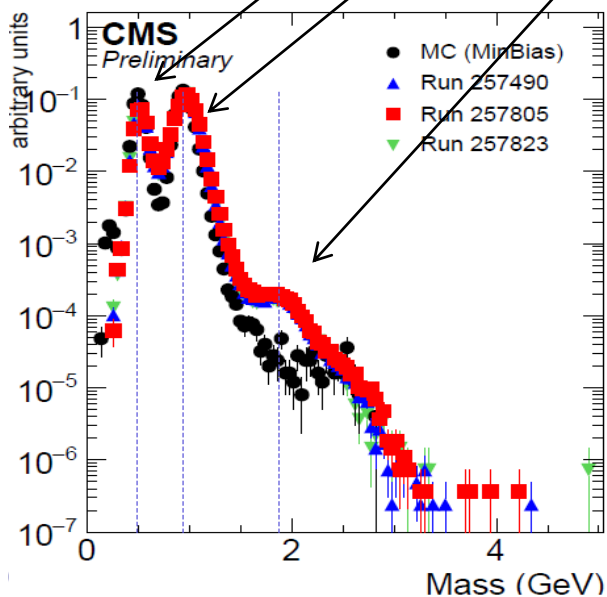
- hamonic-2 is very **robust** against **upward fluctuation** of C_i (i.e. Landau tail)
- hamonic-2 is very much **affected** by **downward fluctuation** (i.e. noise)

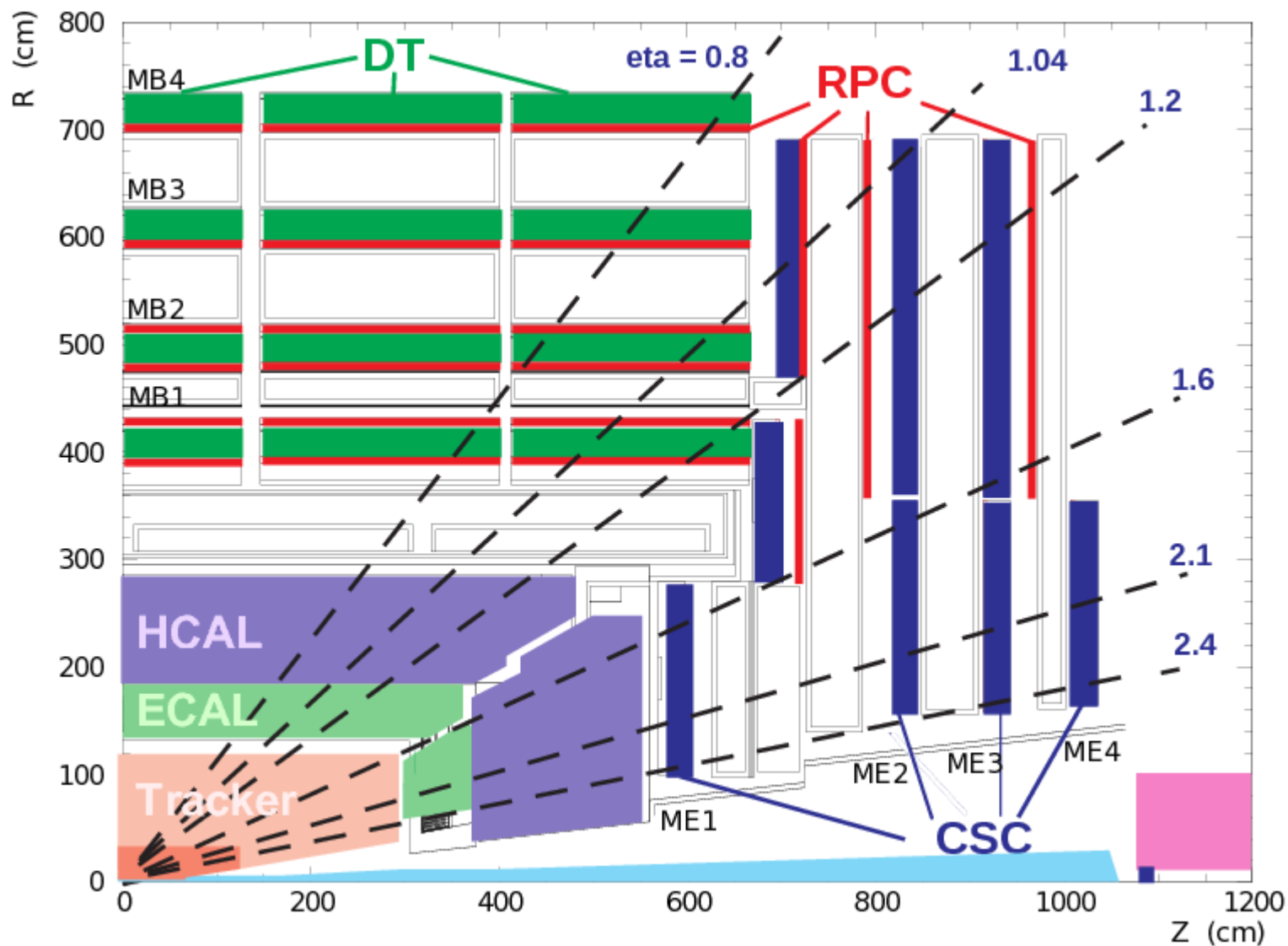


- Constant K & C are extracted from a fit to low momenta protons

- $K = 2.535 \text{ MeV/cm}$
- $C = 3.339 \text{ MeV/cm}$

- Mass of kaons, protons and deuterons can then be reconstructed in data.
- Deficit of deuterons in the simulation is due to pythia not producing them.
- Difference between reconstructed and true mass due to electronic saturation
- Protons are also used as a candle to infer systematic uncertainty due to both I_{as} and I_h data/MC difference.





• Time of flight $\rightarrow \beta^{-1}$:

$$\beta^{-1} = 1 + \frac{c\delta_t}{L}$$

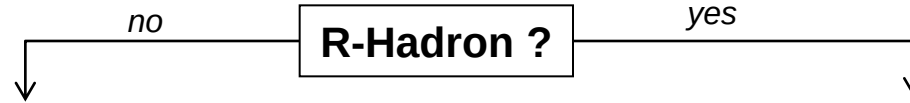
- β^{-1} measurements computed from:
 - flight distance L between IP and the muon chamber
 - δt delay with respect to a muon traveling at the speed-of-light.
- Combine β^{-1} is computed as the average of individual TOF measurement weighted by their individual resolution.

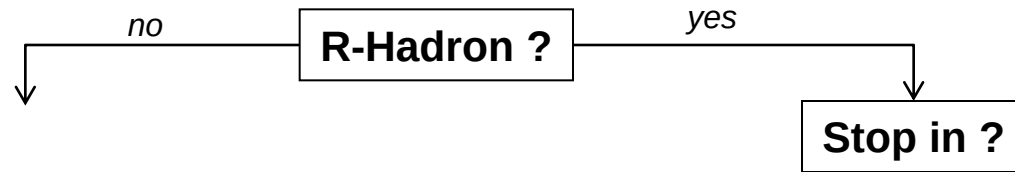
DT: $w_i = \frac{(n-2)}{n} \frac{L_i^2}{\sigma_{DT}^2}$

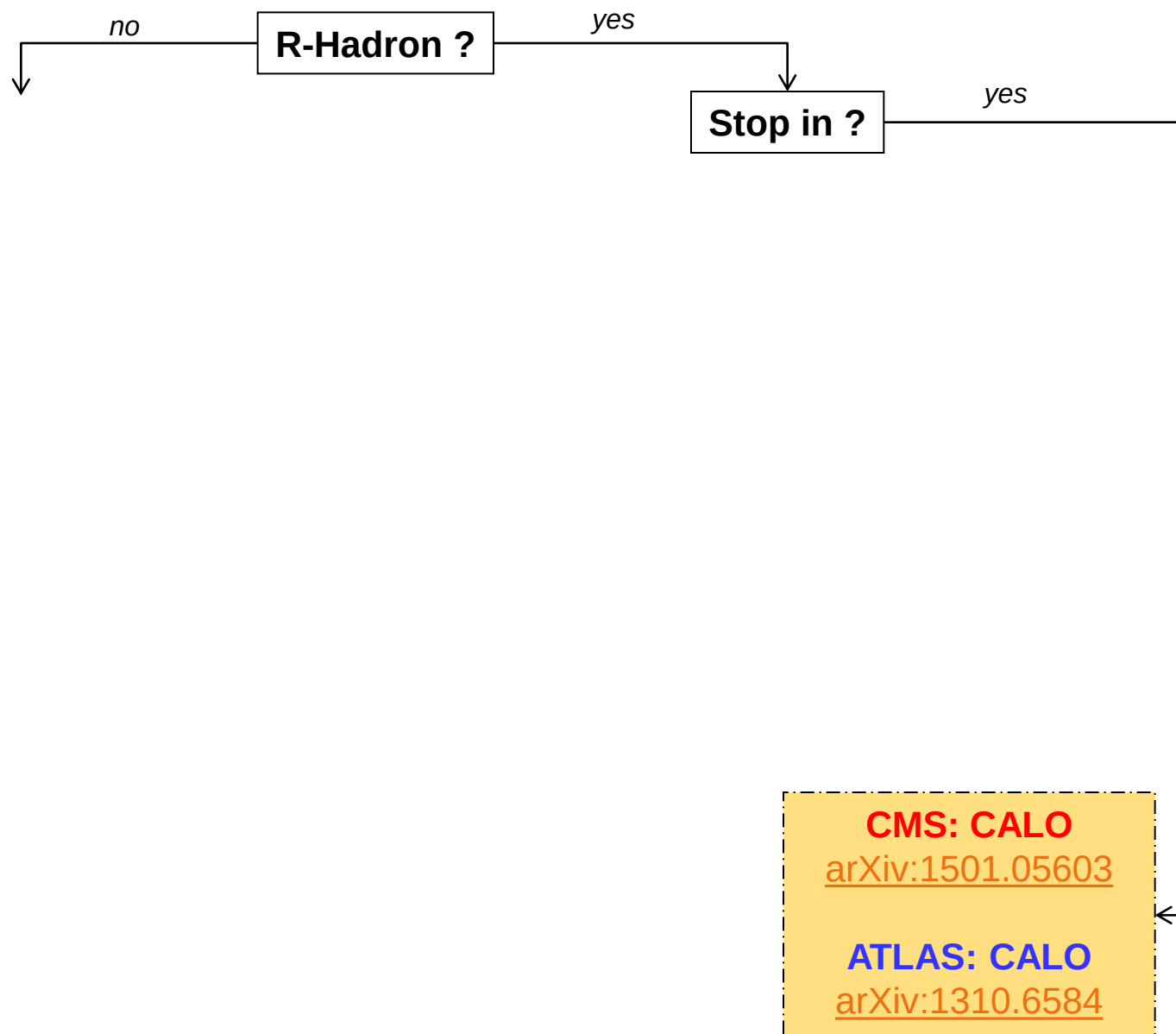
CSC: $w_i = \frac{L_i^2}{\sigma_i^2}$

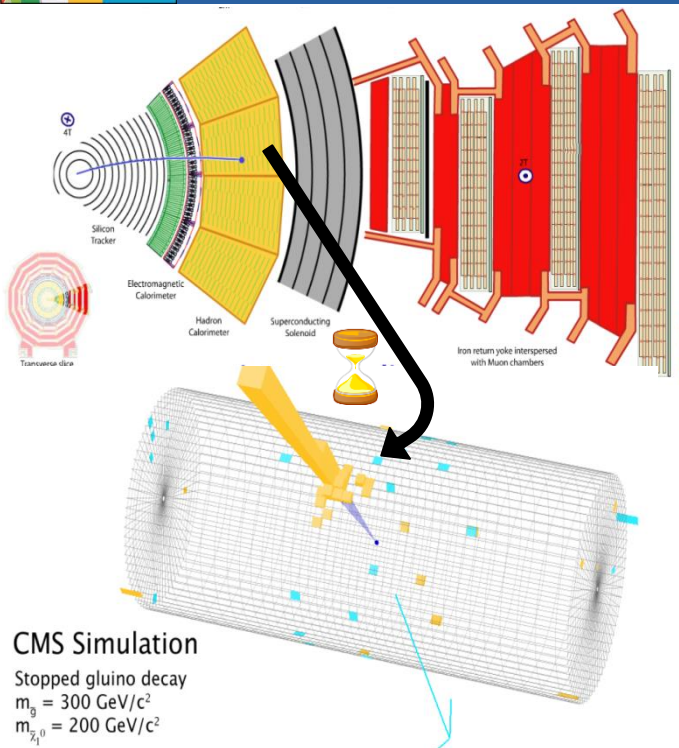
With $\sigma = 3\text{ns}$ for DT, 7ns for CSC cathode and 8.6ns for CSC anode.

$$\sigma_{1/\beta} = \sqrt{\sum_{i=1}^N \frac{(1/\beta_i - \overline{1/\beta})^2 \times w_i}{N-1}} = \sim 0.065 \text{ on average for muons}$$



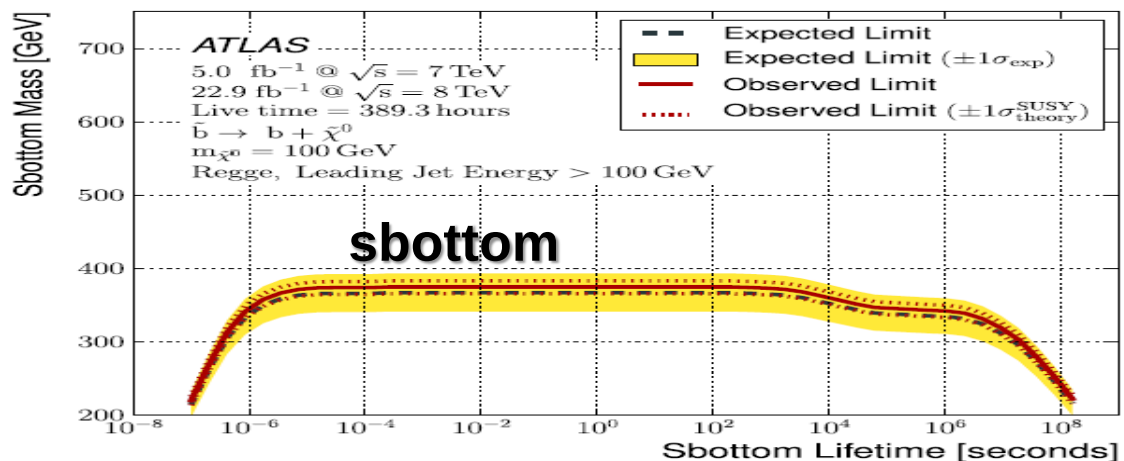
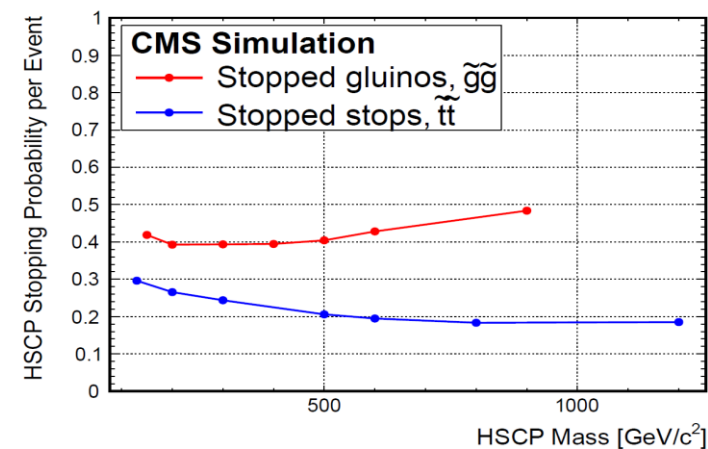
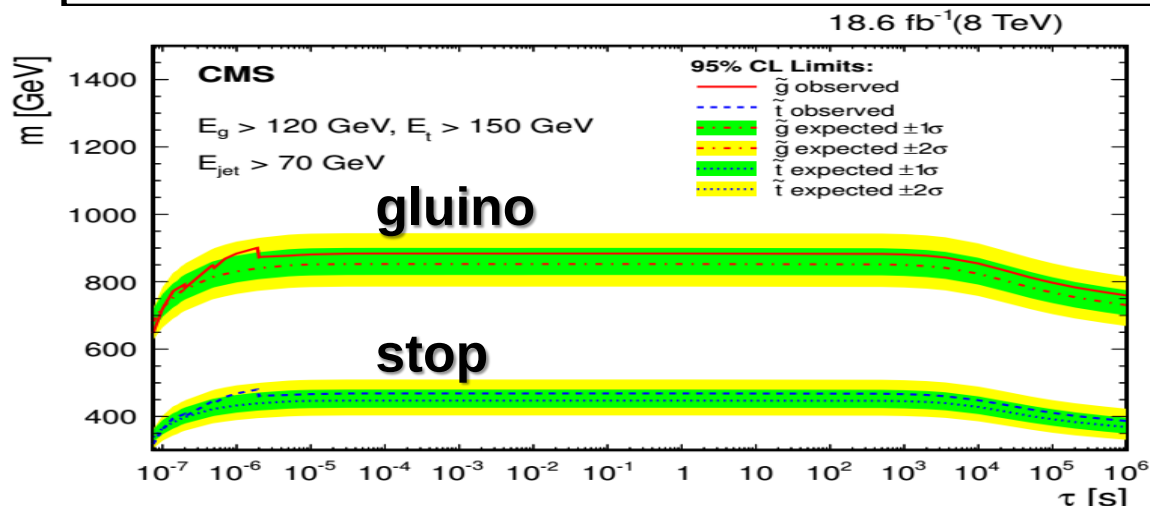


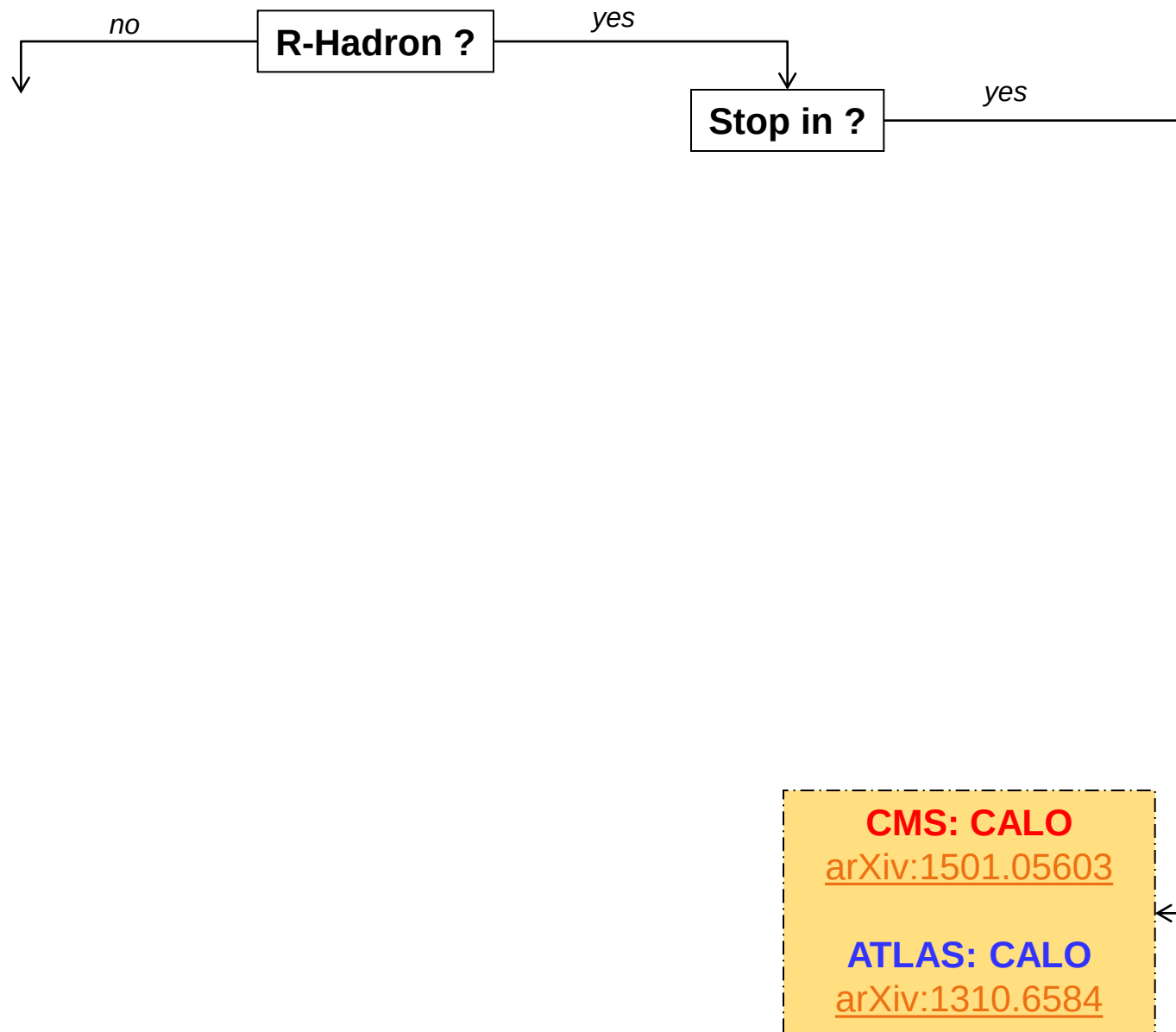


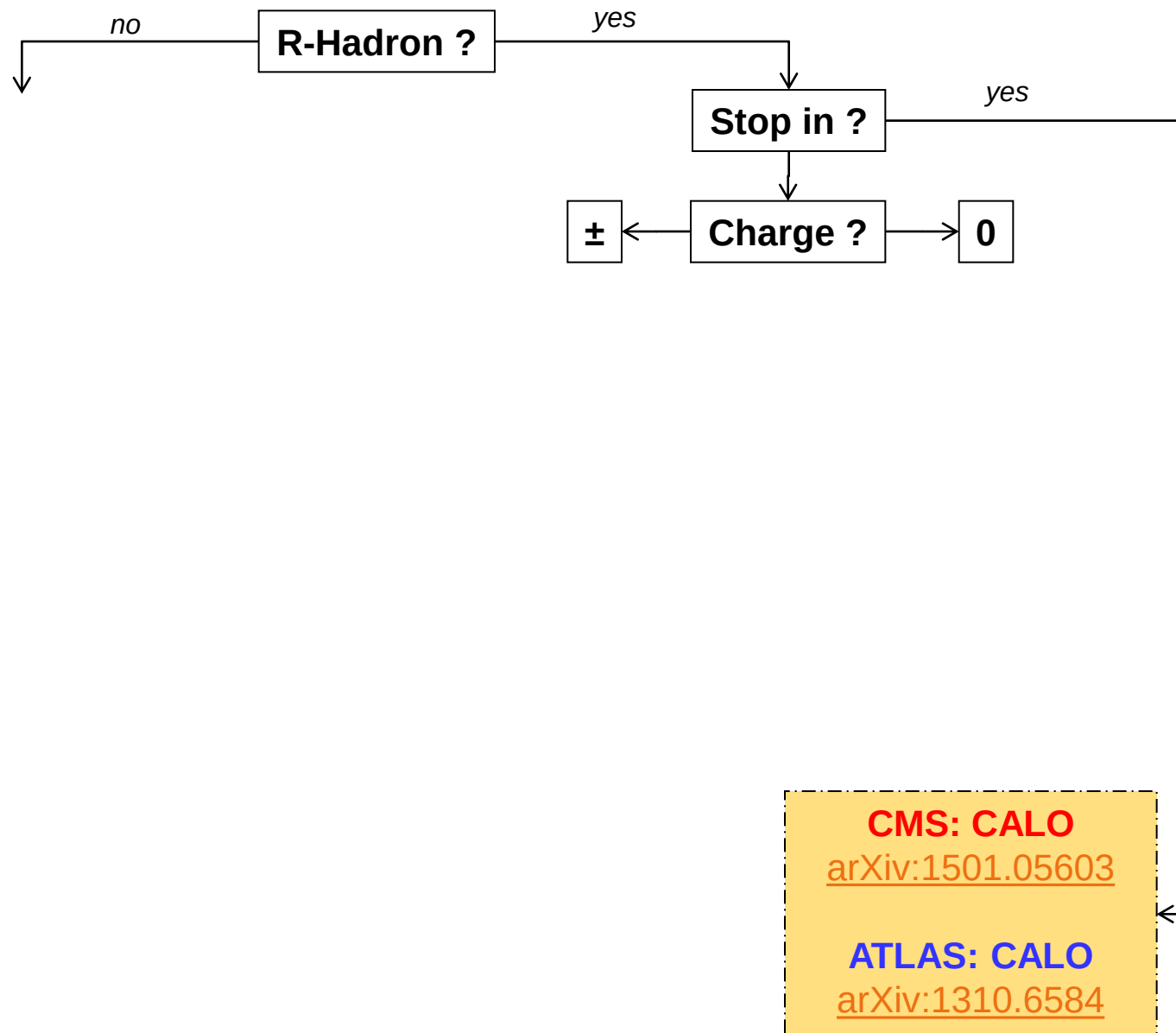


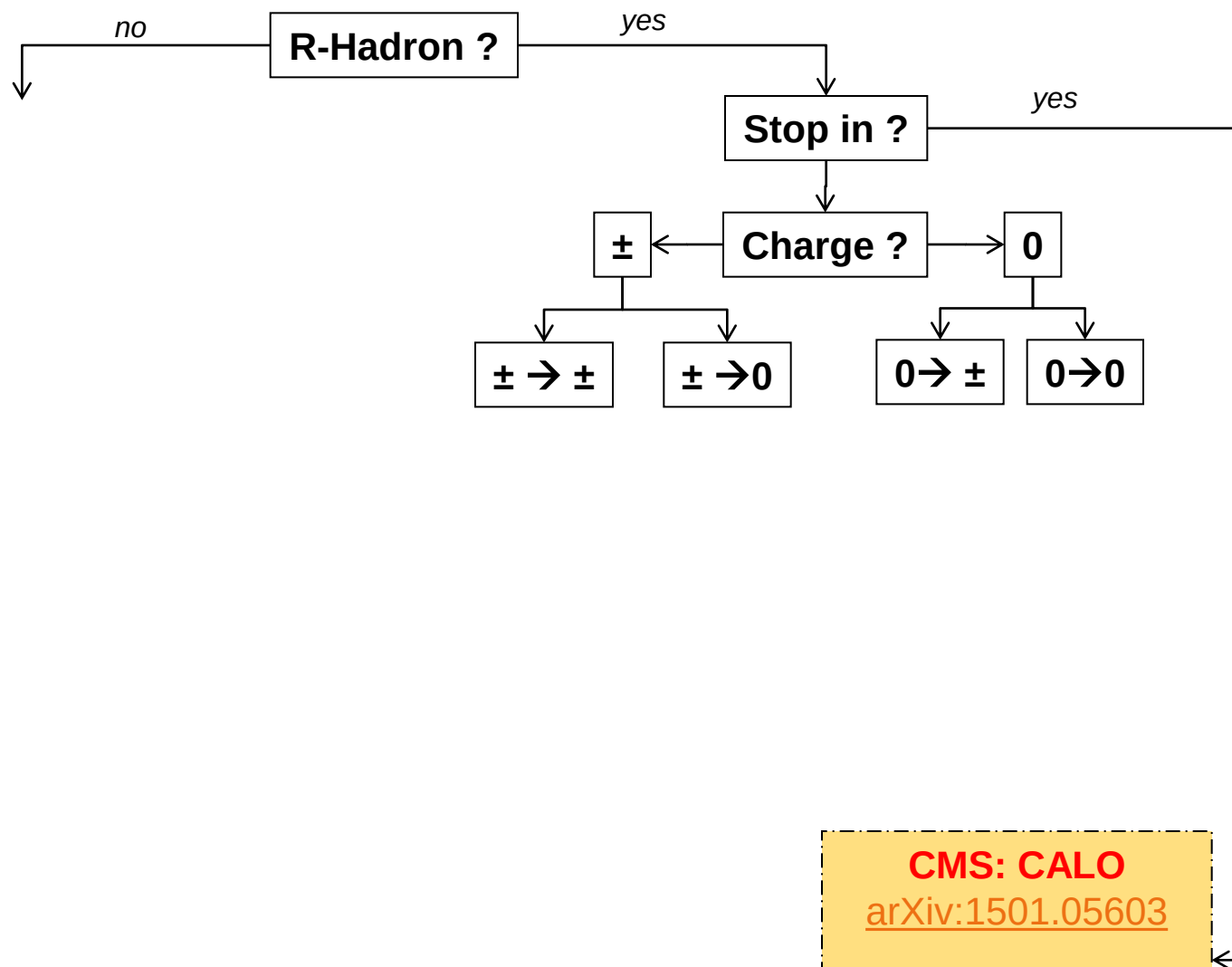
The search in a nutshell

- **Signature:** Energetic jet w/o pp collisions
- **Backgrounds:** cosmics, beam halo, det. noise







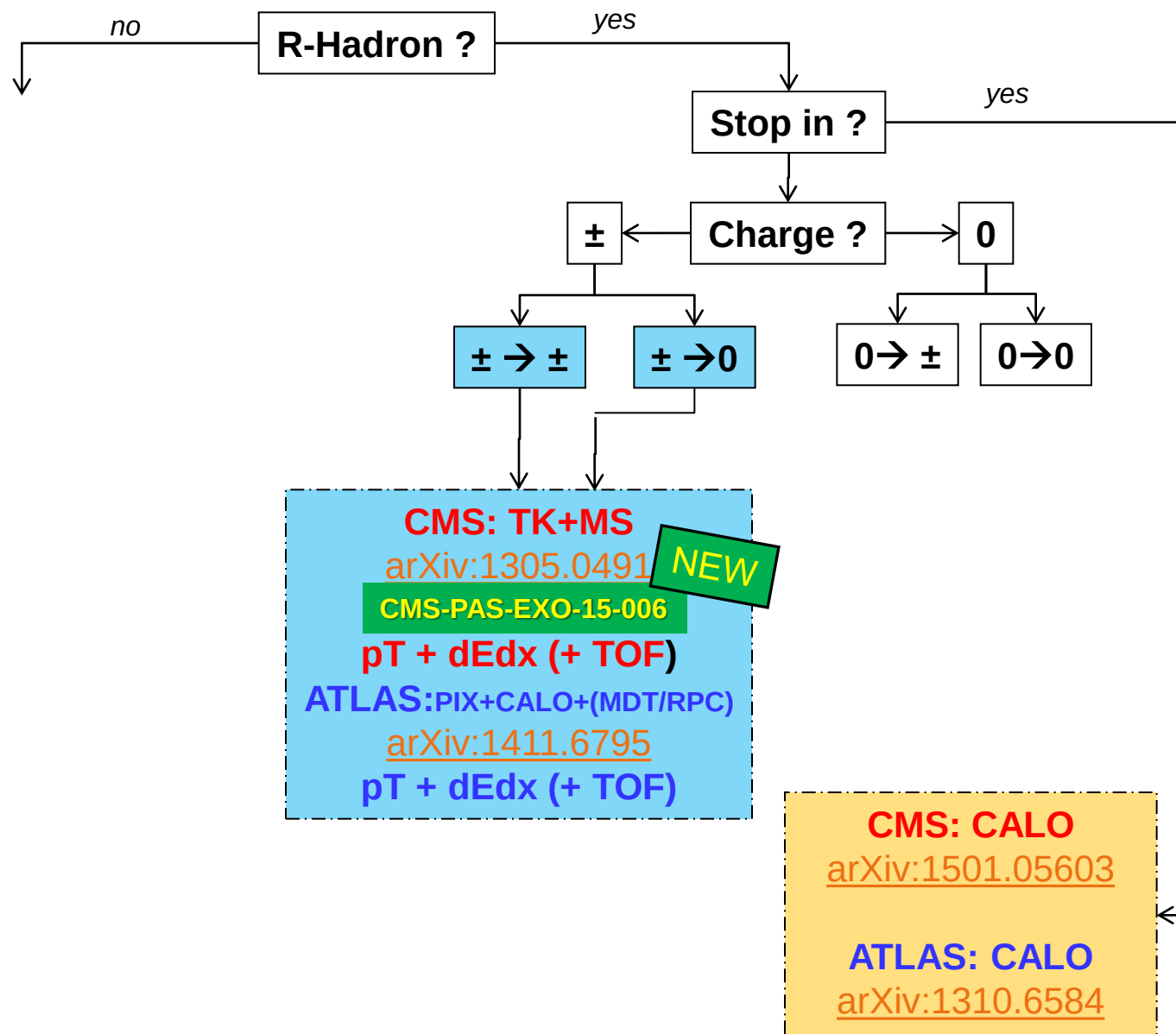


CMS: CALO

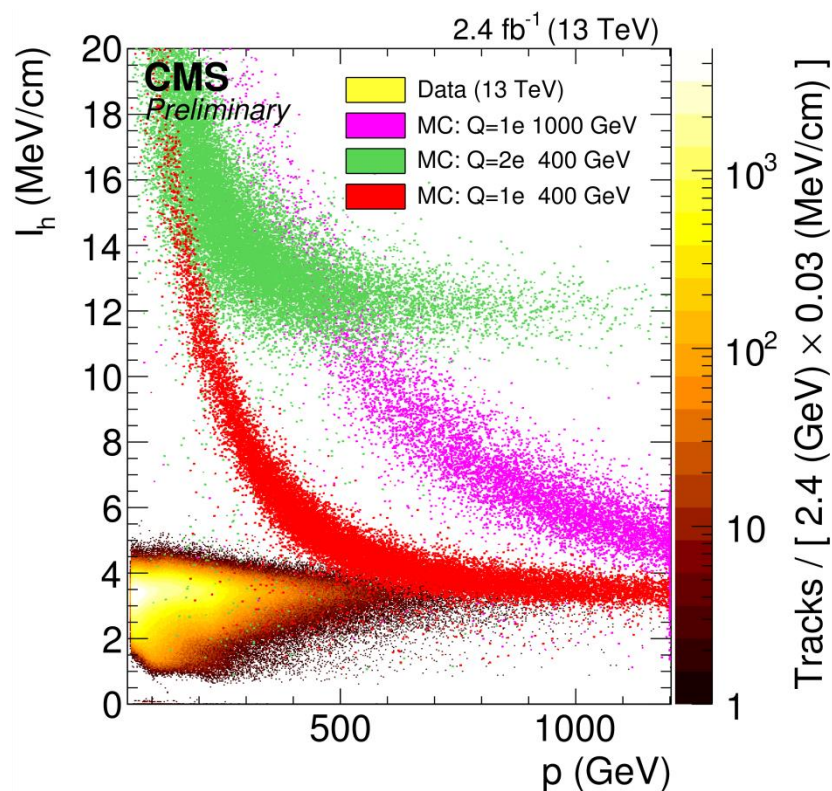
[arXiv:1501.05603](https://arxiv.org/abs/1501.05603)

ATLAS: CALO

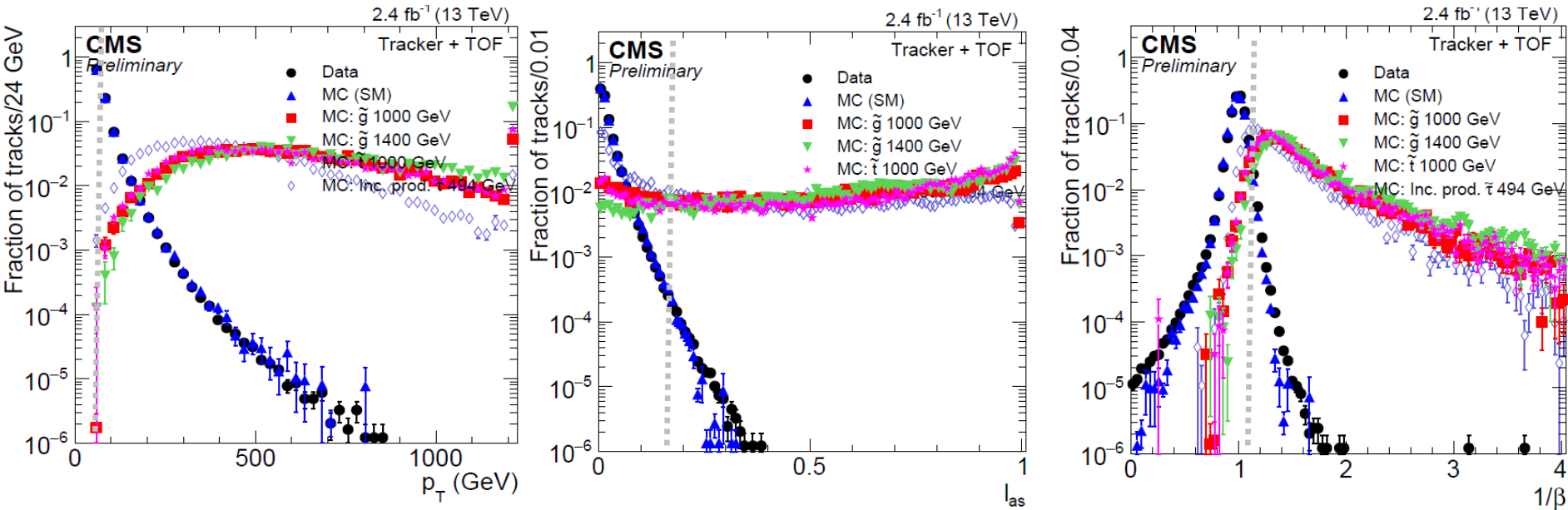
[arXiv:1310.6584](https://arxiv.org/abs/1310.6584)



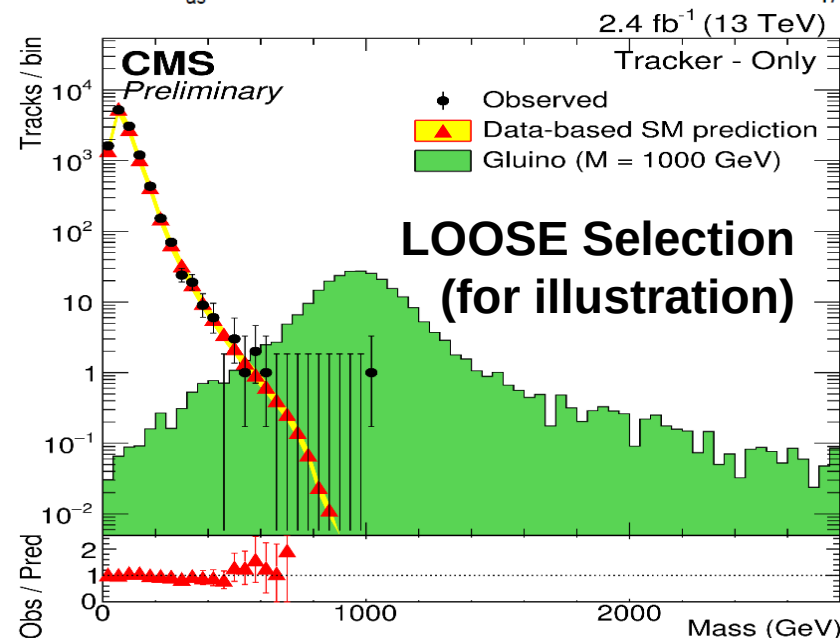
- Very similar approaches used by ATLAS and CMS
- Trigger : Muon OR MET
- Variables: p_T (tracker), dE/dx (inner detector) and TOF (Muon System)
- Background prediction based on ABCD and predicted mass sprectrum from pdf
- Main difference: ATLAS has a 2 HSCP/event category not considered by CMS

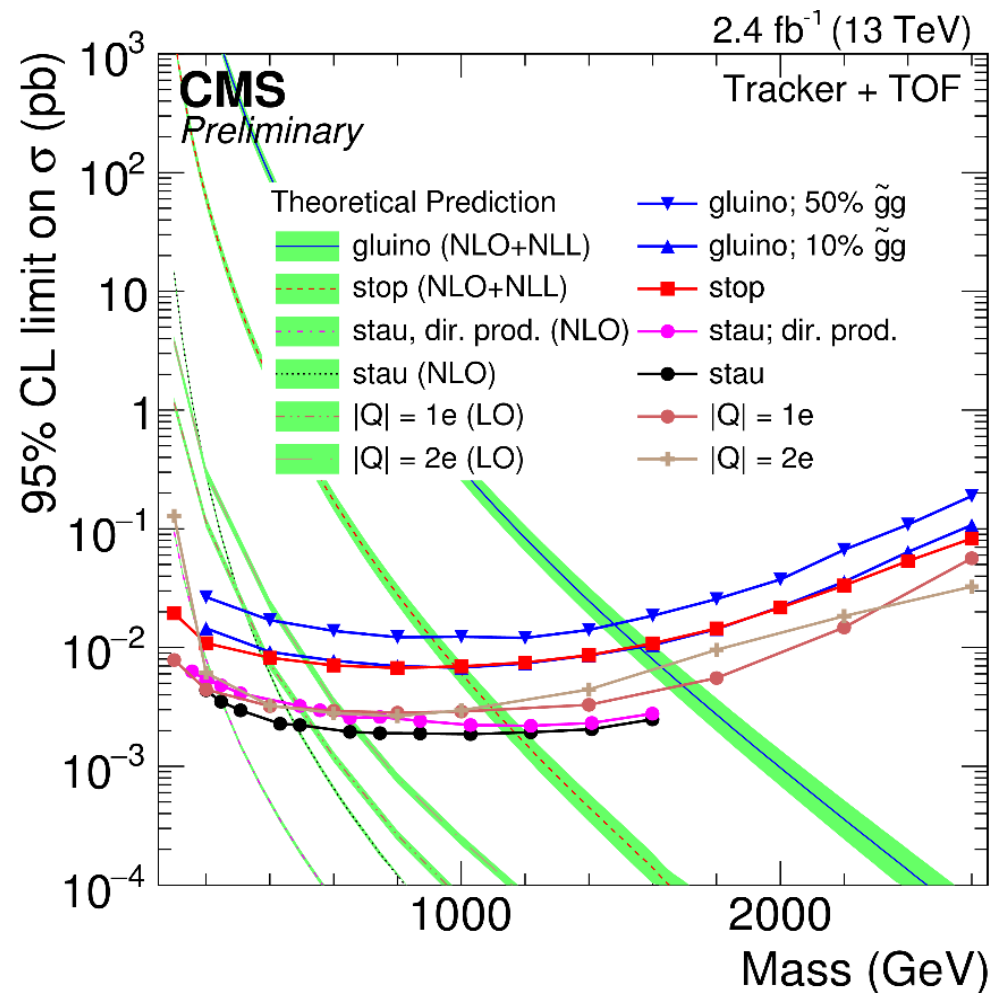


	Trk-only	Trk+TOF
Track Type	Inner Track	Global Muon
$ \eta $	< 2.1	
p_T (GeV)	> 55	
d_z and d_{xy} (cm)	< 0.5	
σ_{p_T}/p_T	< 0.25	
Track χ^2/dof	< 5	
# Pixel Hits	> 1	
# Tracker Hits	> 7	
Frac. Valid Hits	> 0.8	
# dE/dx Measurements	> 5	
dE/dx Strip Shape Test	yes	
$\Sigma p_T^{\text{trk}}(\Delta R < 0.3)$ (GeV)	< 50	
$E_{\text{cal}}(\Delta R < 0.3)/p$	< 0.3	
# TOF Measurements	/	> 7
$\sigma_{1/\beta}$	/	< 0.15
$1/\beta$	/	> 1



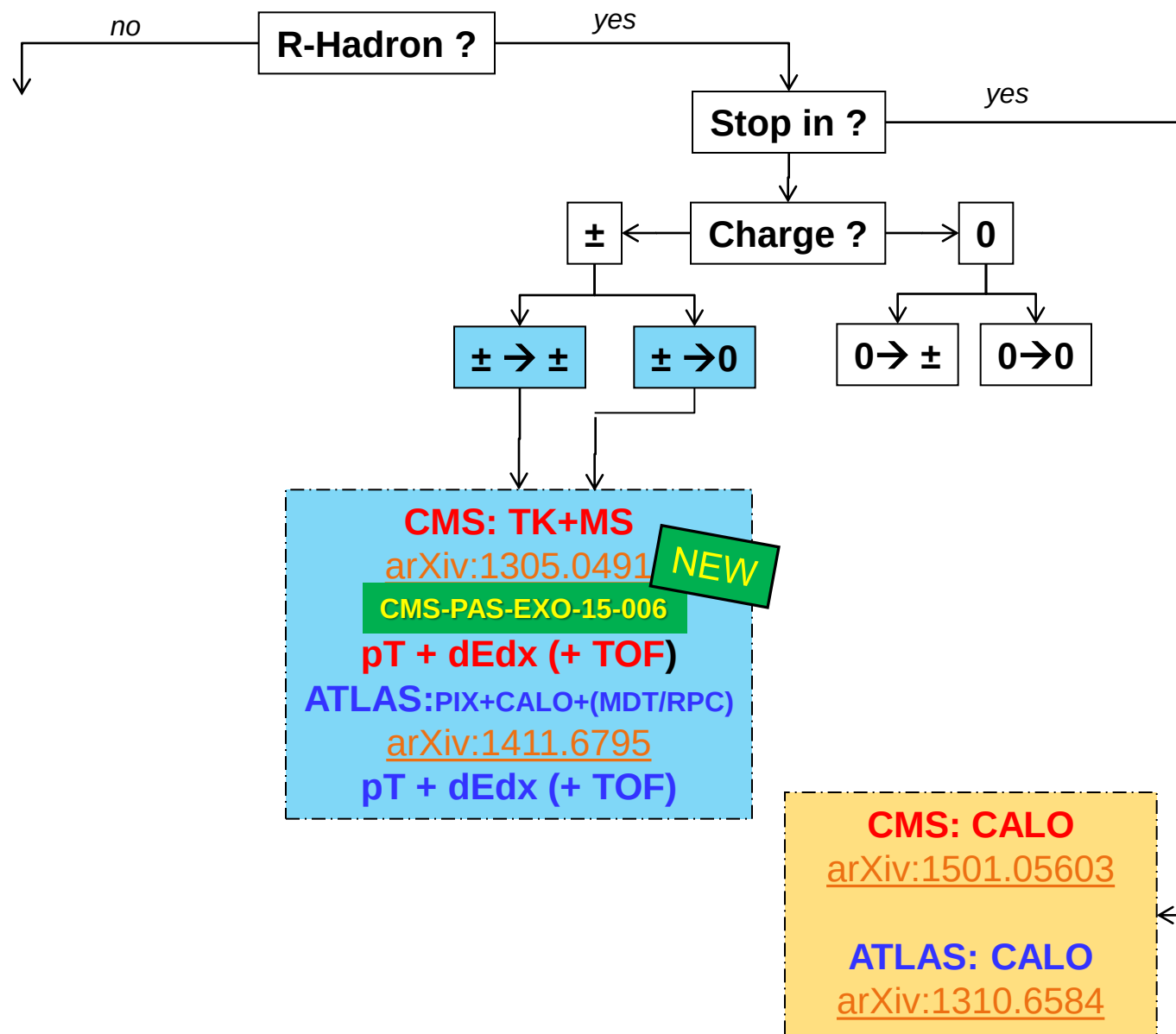
- The 3 variables are uncorrelated !
- The background rate is estimated from data using a 2D or 3D ABCD method.
- The background mass distribution is obtained from momentum and dEdx pdf in control regions.

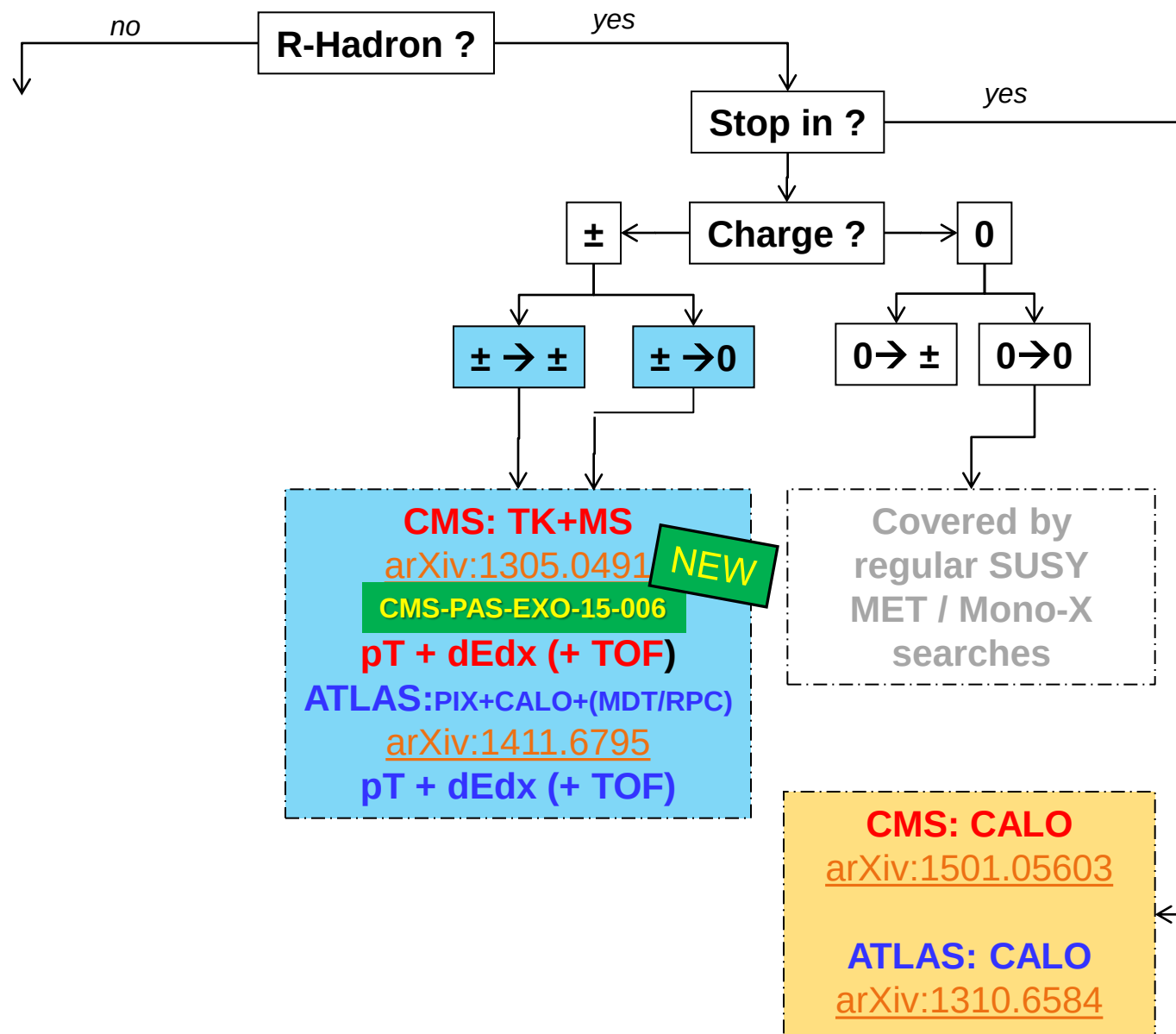


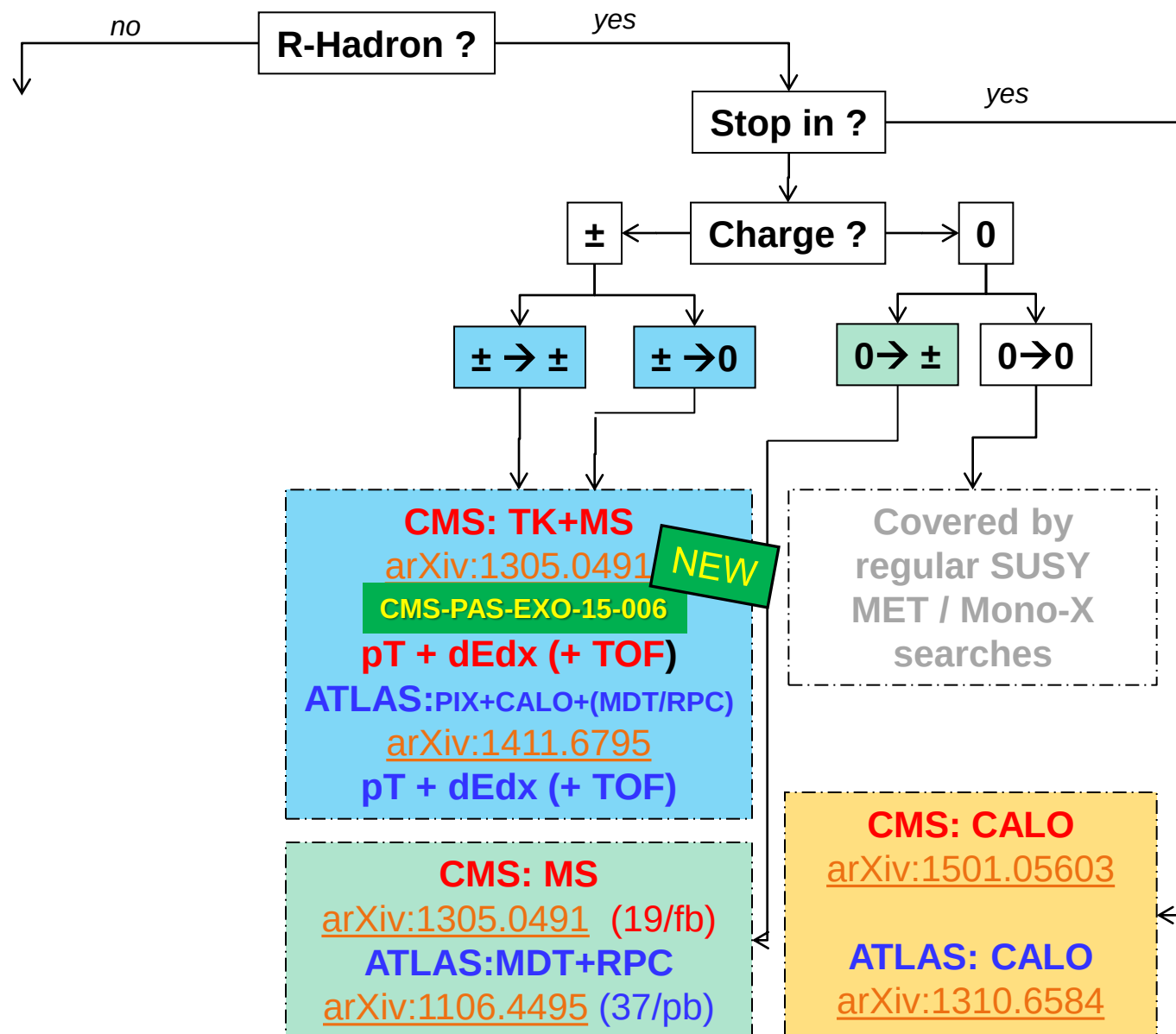


Model	analysis used	$\sqrt{s} = 7 + 8$ TeV	$\sqrt{s} = 13$ TeV
Gluino $f = 0.1$	tracker-only	$M > 1320$ GeV	$M > 1590$ GeV
	tracker+TOF	$M > 1290$ GeV	$M > 1560$ GeV
Gluino $f = 0.1$ CS	tracker-only	$M > 1230$ GeV	$M > 1570$ GeV
Gluino $f = 0.5$	tracker-only	$M > 1260$ GeV	$M > 1500$ GeV
	tracker+TOF	$M > 1220$ GeV	$M > 1480$ GeV
Gluino $f = 0.5$ CS	tracker-only	$M > 1150$ GeV	$M > 1530$ GeV
Stop	tracker-only	$M > 940$ GeV	$M > 1020$ GeV
	tracker+TOF	$M > 910$ GeV	$M > 980$ GeV
Stop CS	tracker-only	$M > 820$ GeV	$M > 970$ GeV
Stau inc. prod. (GMSB SPS7)	tracker+TOF	$M > 440$ GeV	$M > 480$ GeV
	tracker-only	$M > 390$ GeV	$M > 450$ GeV
Stau pair prod.	tracker+TOF	$M > 340$ GeV	$M > 230$ GeV
	tracker-only	$M > 190$ GeV	
DY $Q = 1e$	tracker-only	$M > 650$ GeV	$M > 510$ GeV
	tracker+TOF	$M > 650$ GeV	$M > 540$ GeV
DY $Q = 2e$	multiply charged tracker-only tracker+TOF	$M > 730$ GeV	-
		$M > 520$ GeV	$M > 600$ GeV
		$M > 520$ GeV	$M > 650$ GeV

**Gluino mass limits are at 1.6TeV
with just 2.4fb**





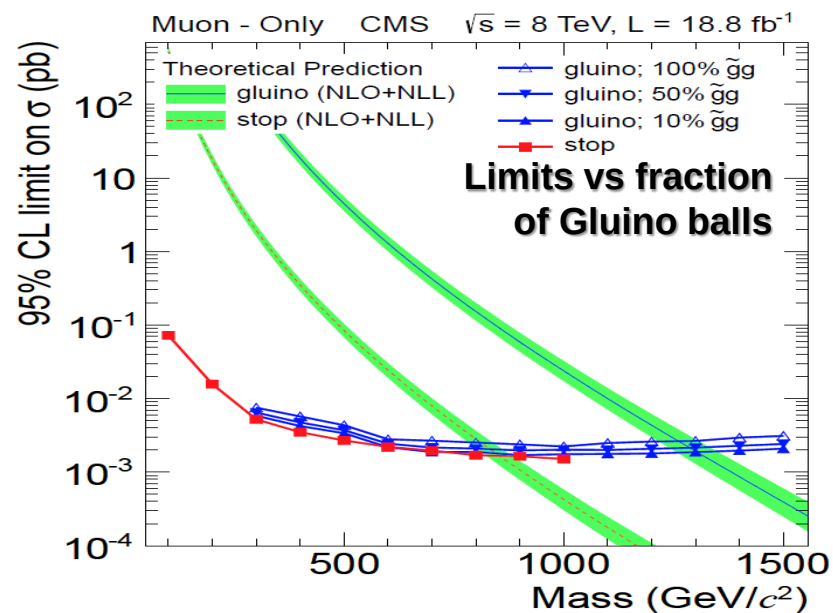
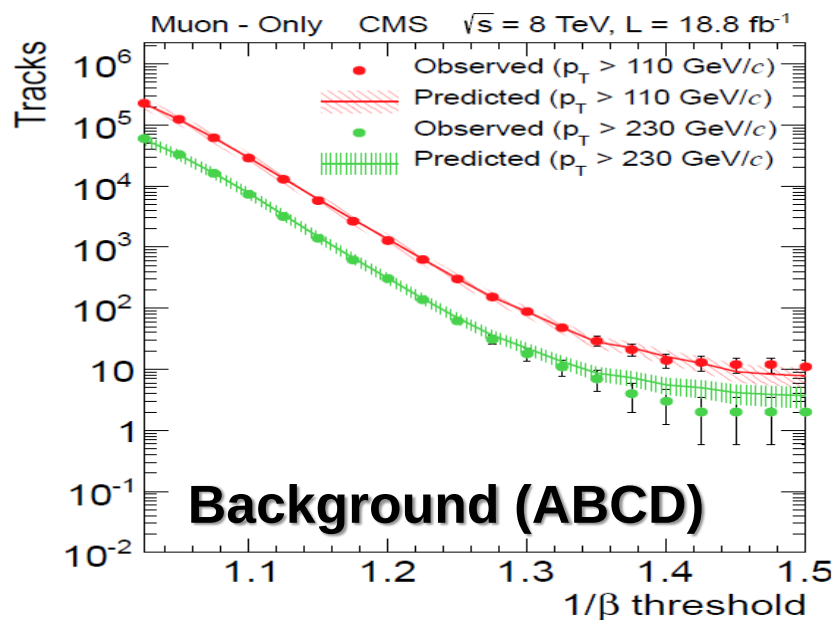


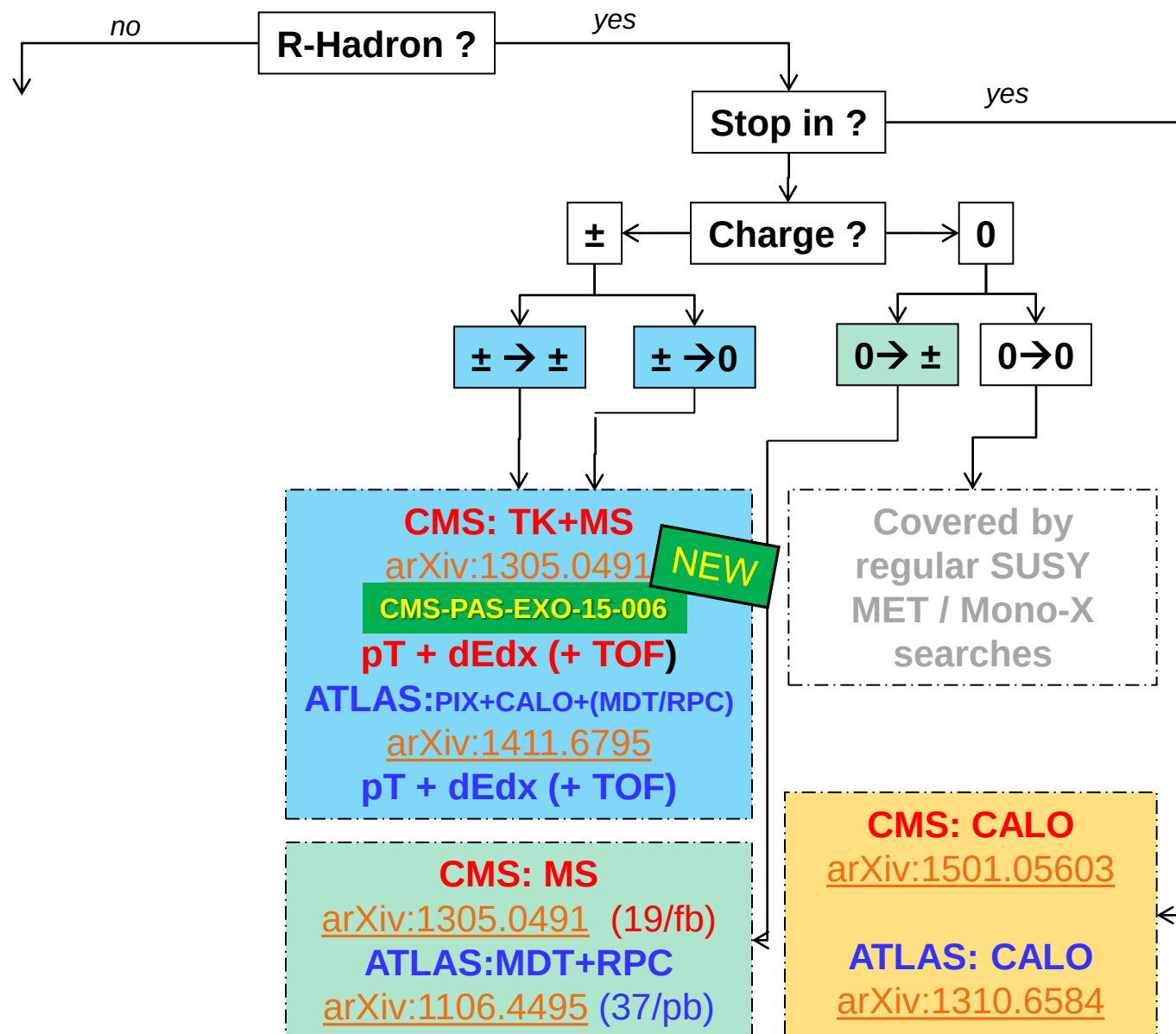
Challenges:

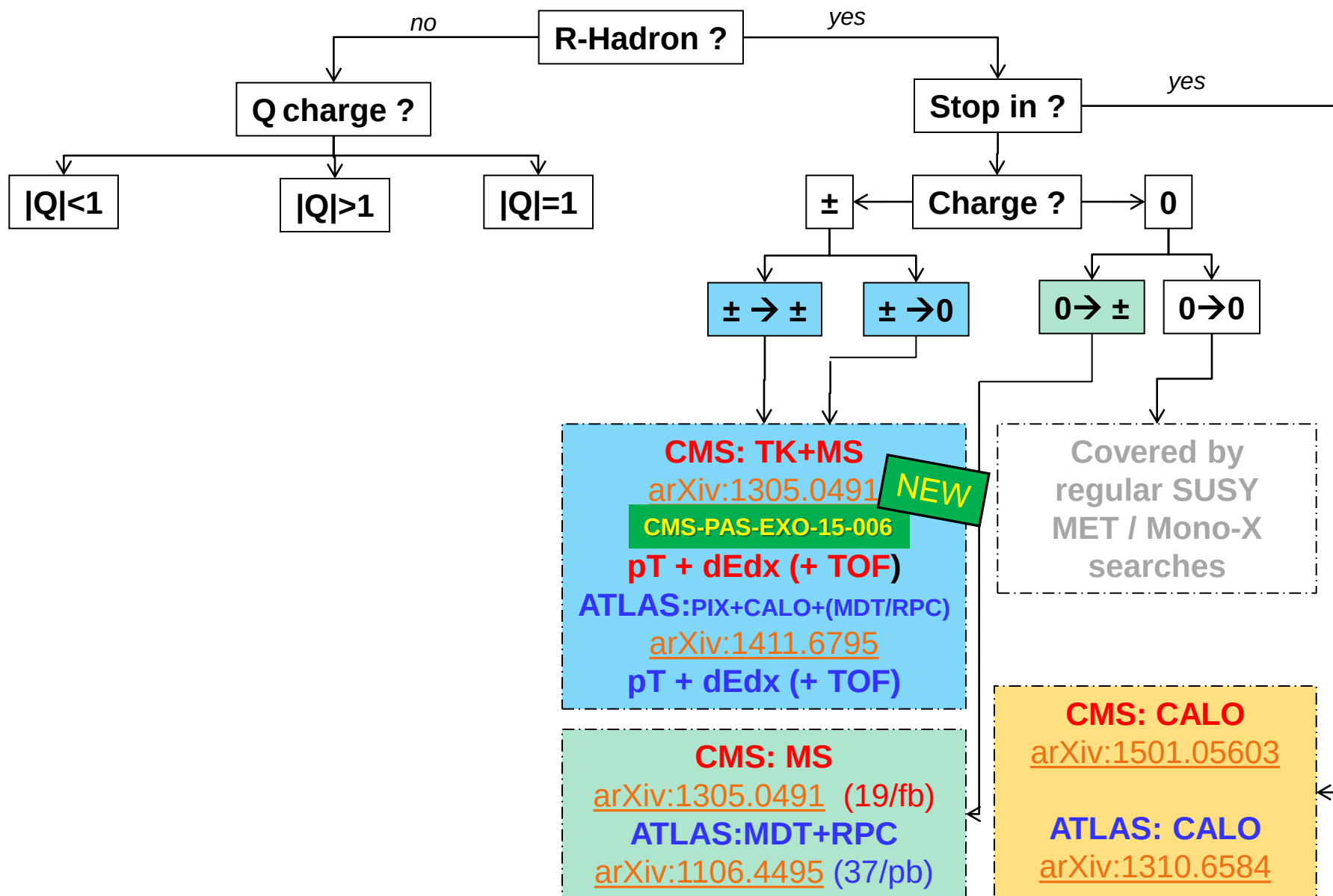
- Invisible in the tracker
- Worse p_T resolution (due to coarser hit position in the MS)
- High cosmic backgrounds (also out-of-time)

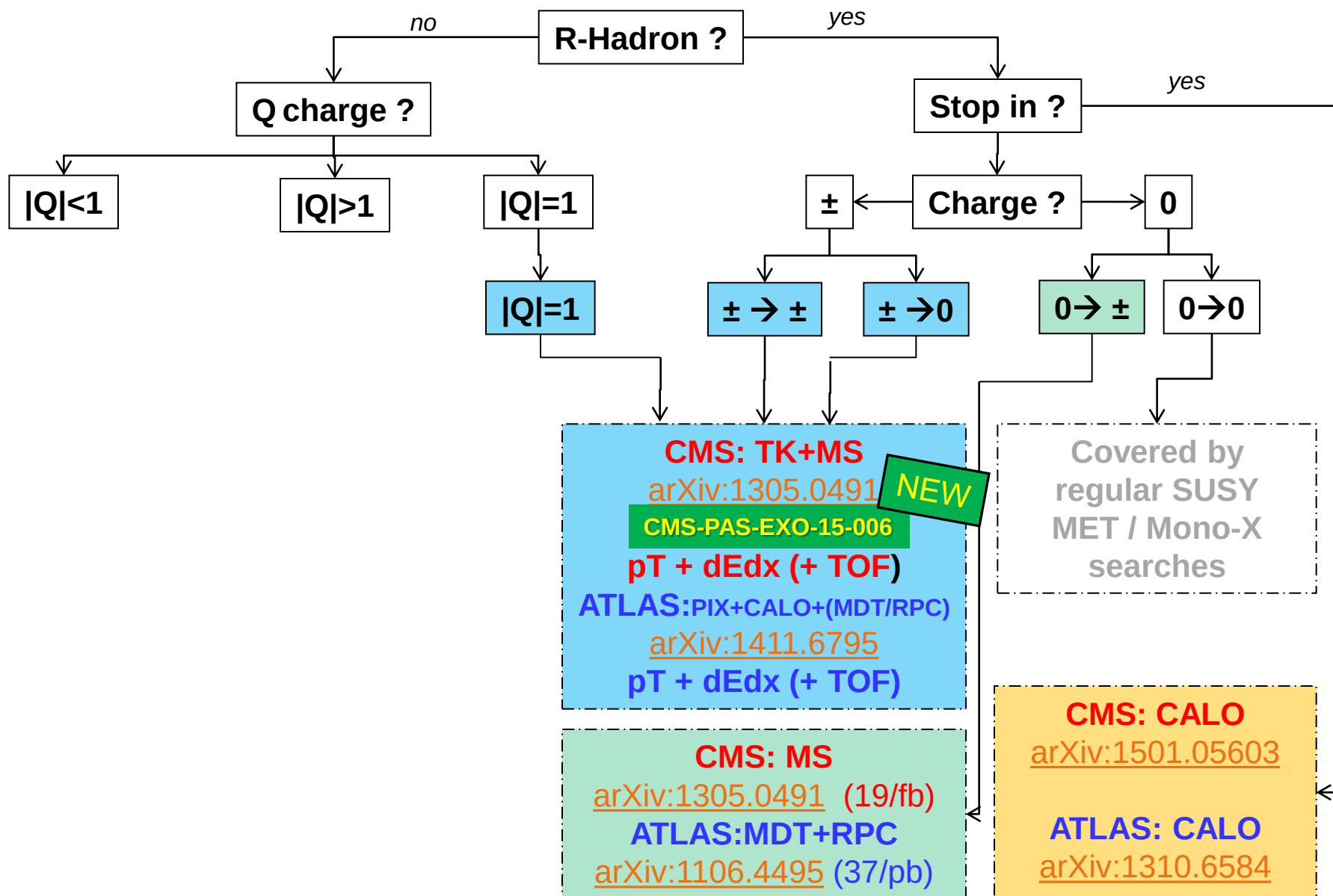
Solutions:

- Dedicated trigger (stand alone muons without inner track matching)
- Tight Cosmic Veto
- Muon TOF must be compatible with IP collisions

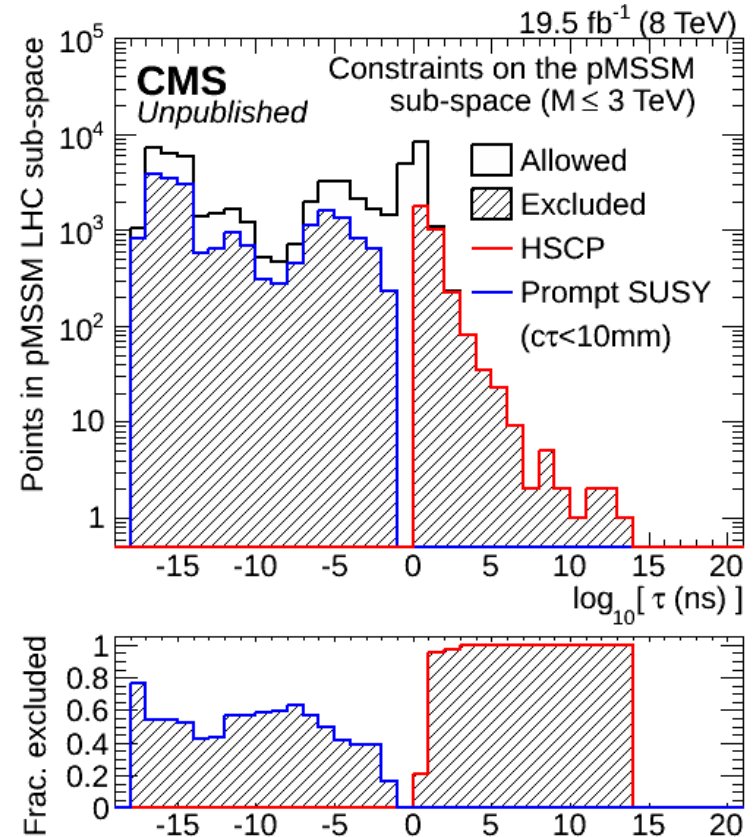
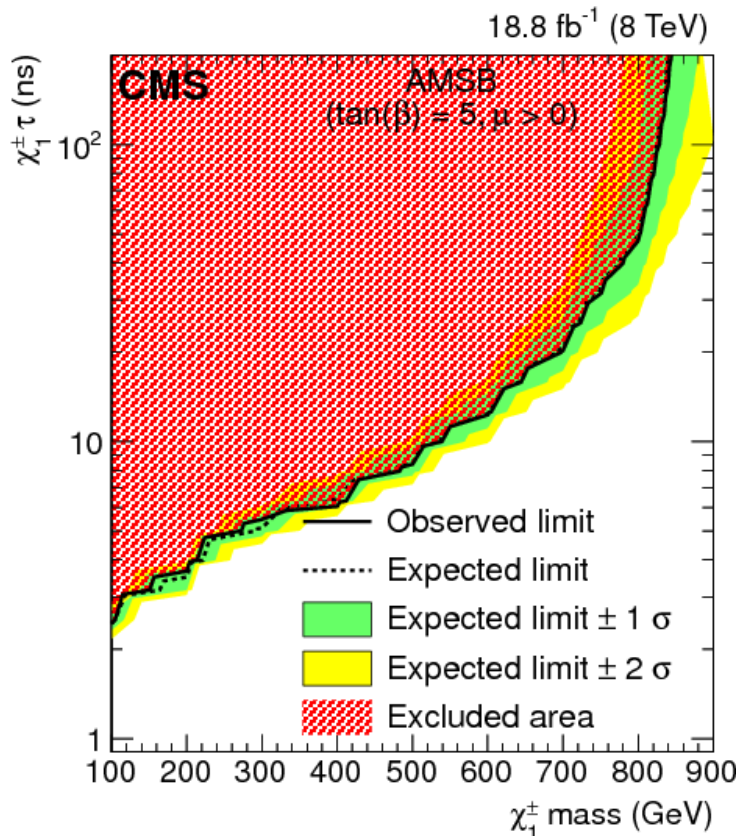


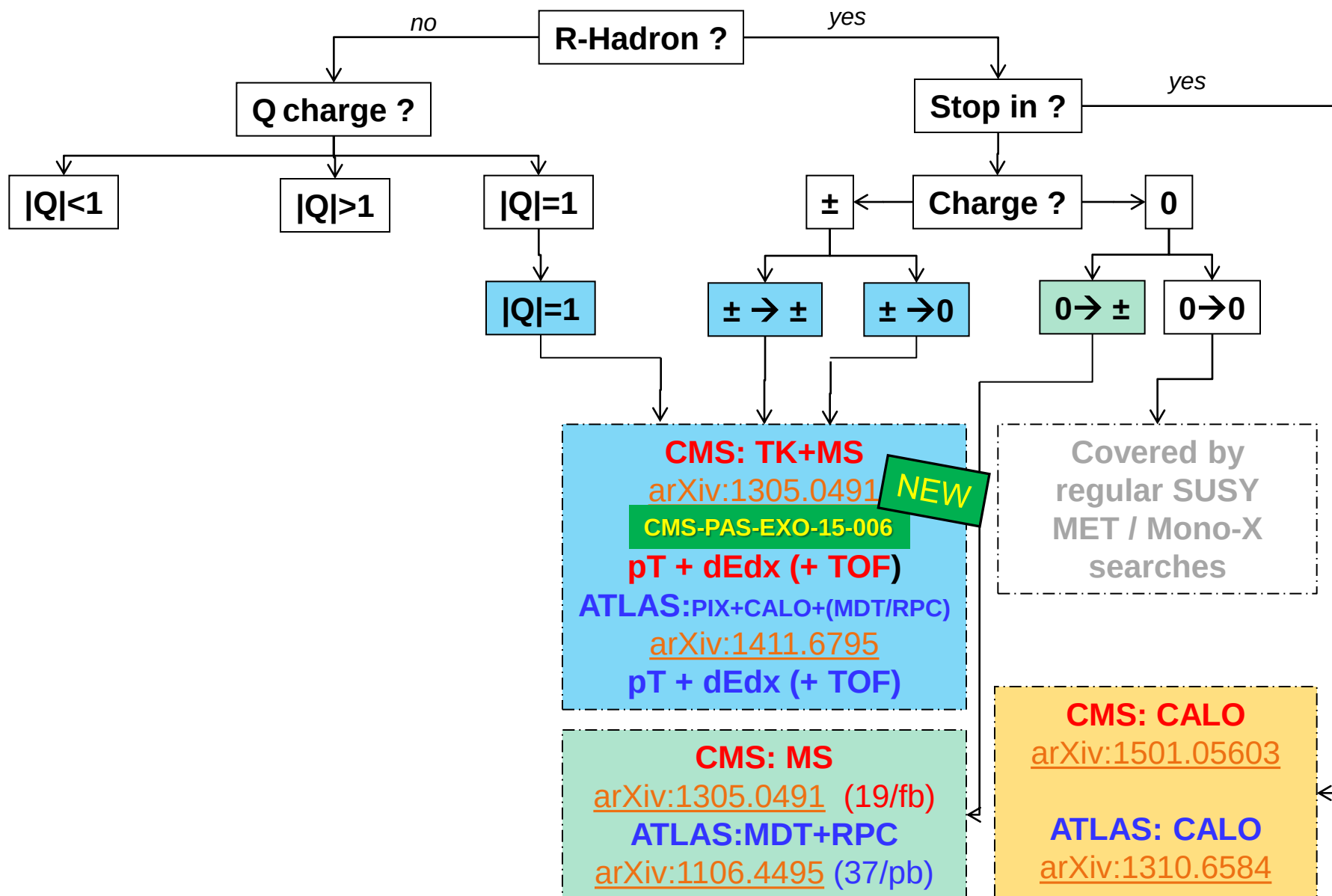


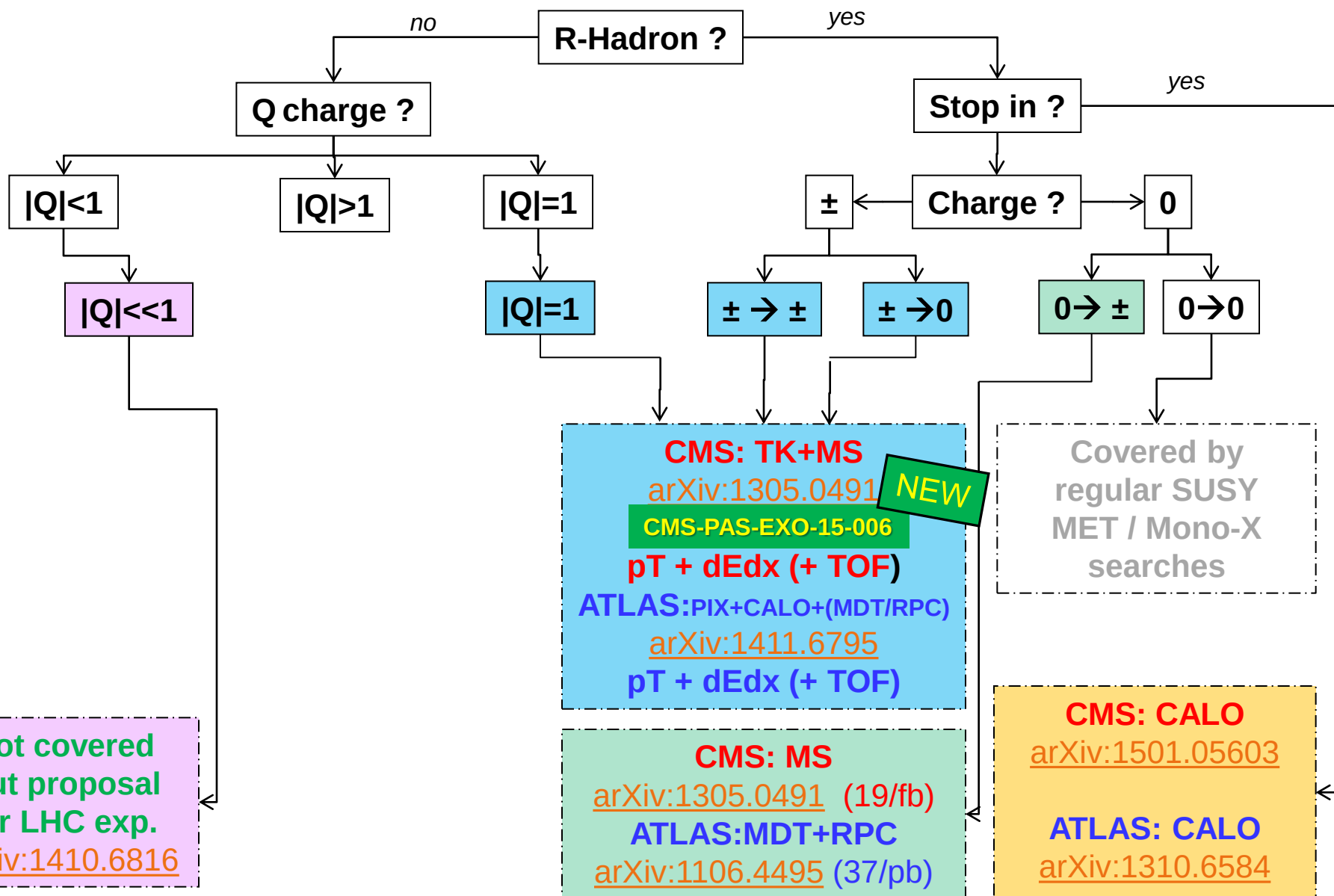


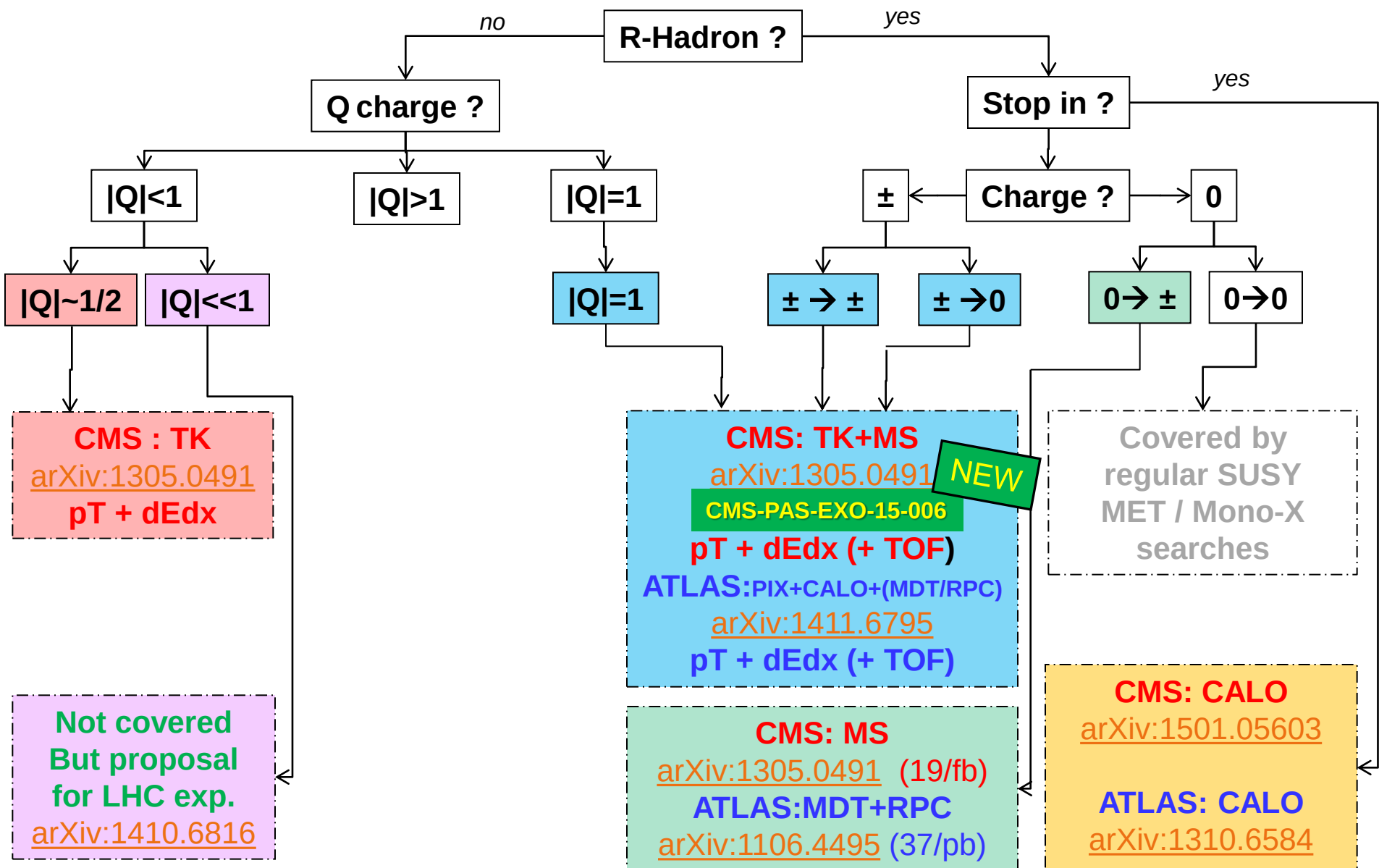


- Identical to the case regular ($\pm \rightarrow \pm$) R-Hadron search
- But constrains many additional models predicting « lepton-like » HSCP
 - GMSB (staus)
 - AMSB (charginos)
 - pMSSM (charginos)





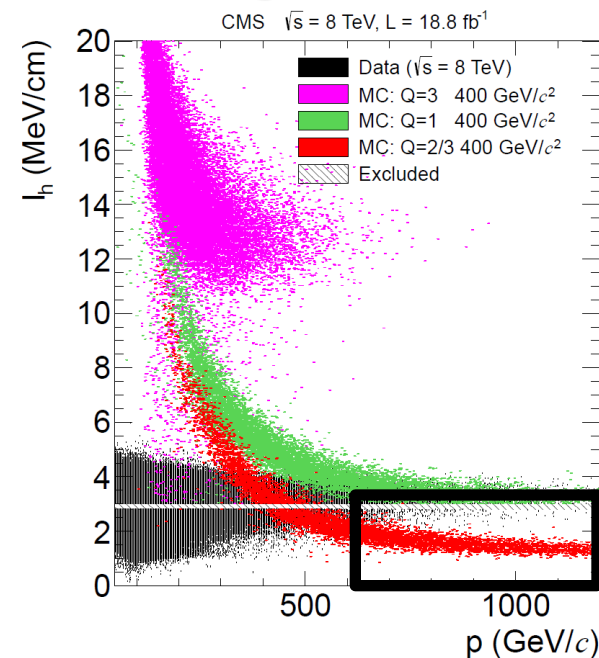




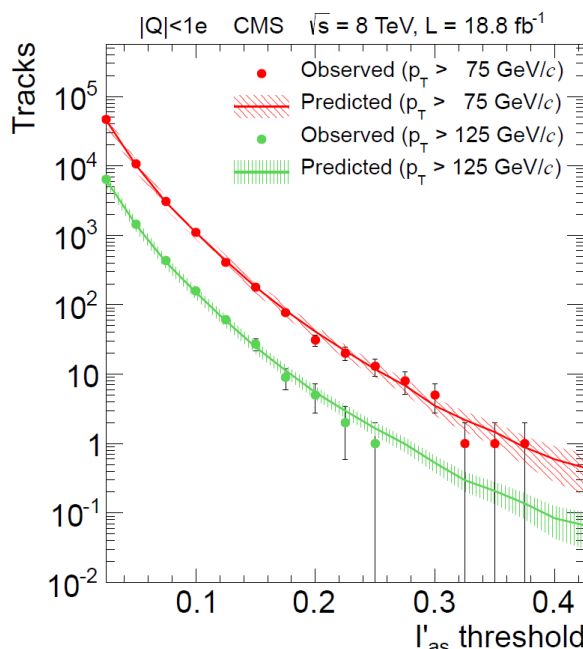
The search in a nutshell

- **Challenges:**
 - Small energy loss $\rightarrow$ LHC detectors and DAQ designed for $Q=1$
 - Out-of-time particles (cosmics/ghost) also lead to low dE/dx
- **Advantage:**
 - Tracking: high reco p_T ($=p_{T, \text{True}} / Q$)

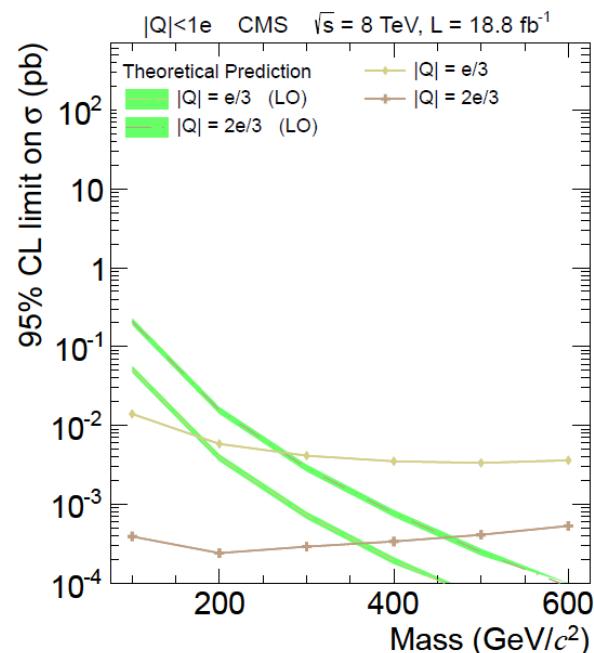
signature

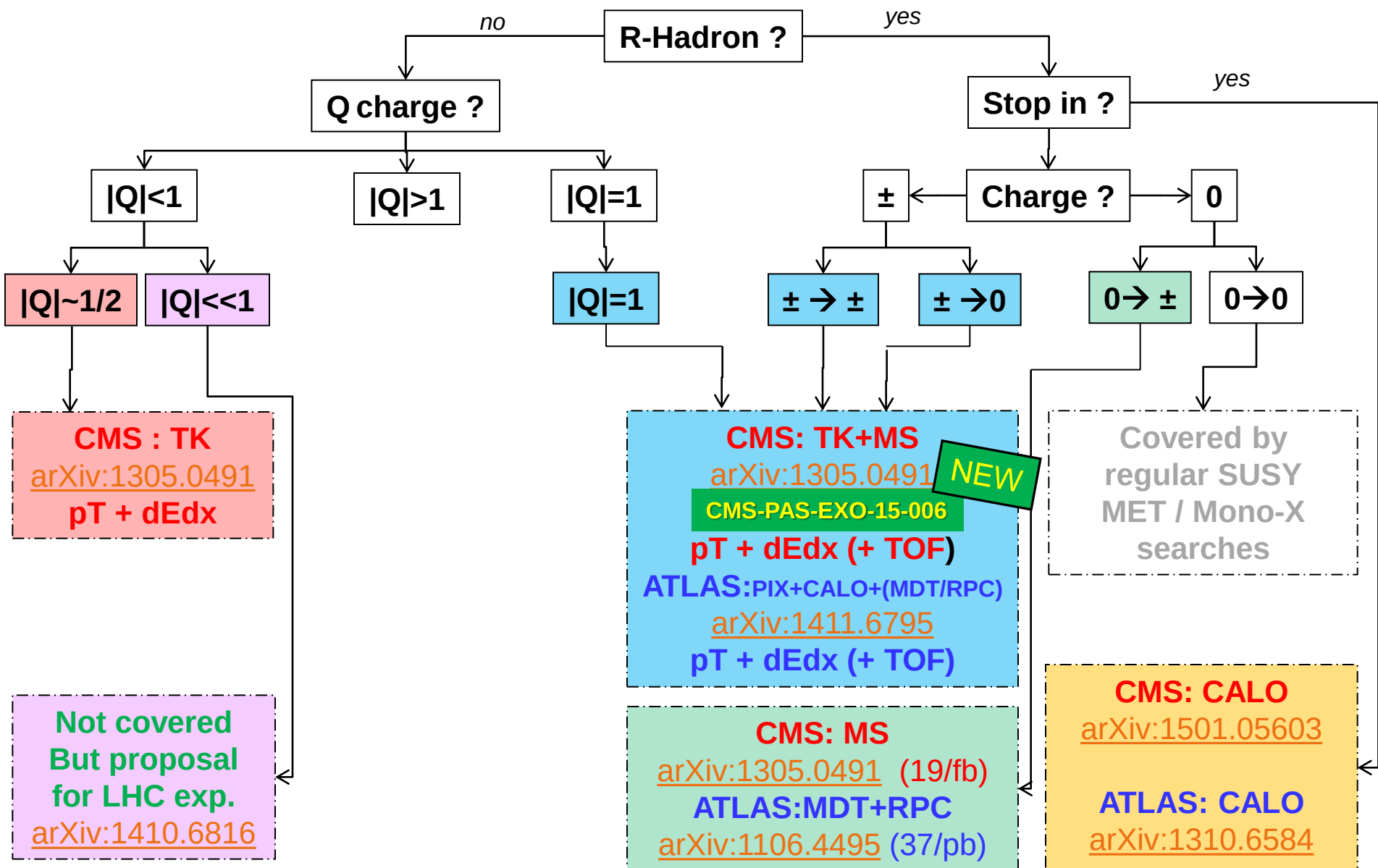


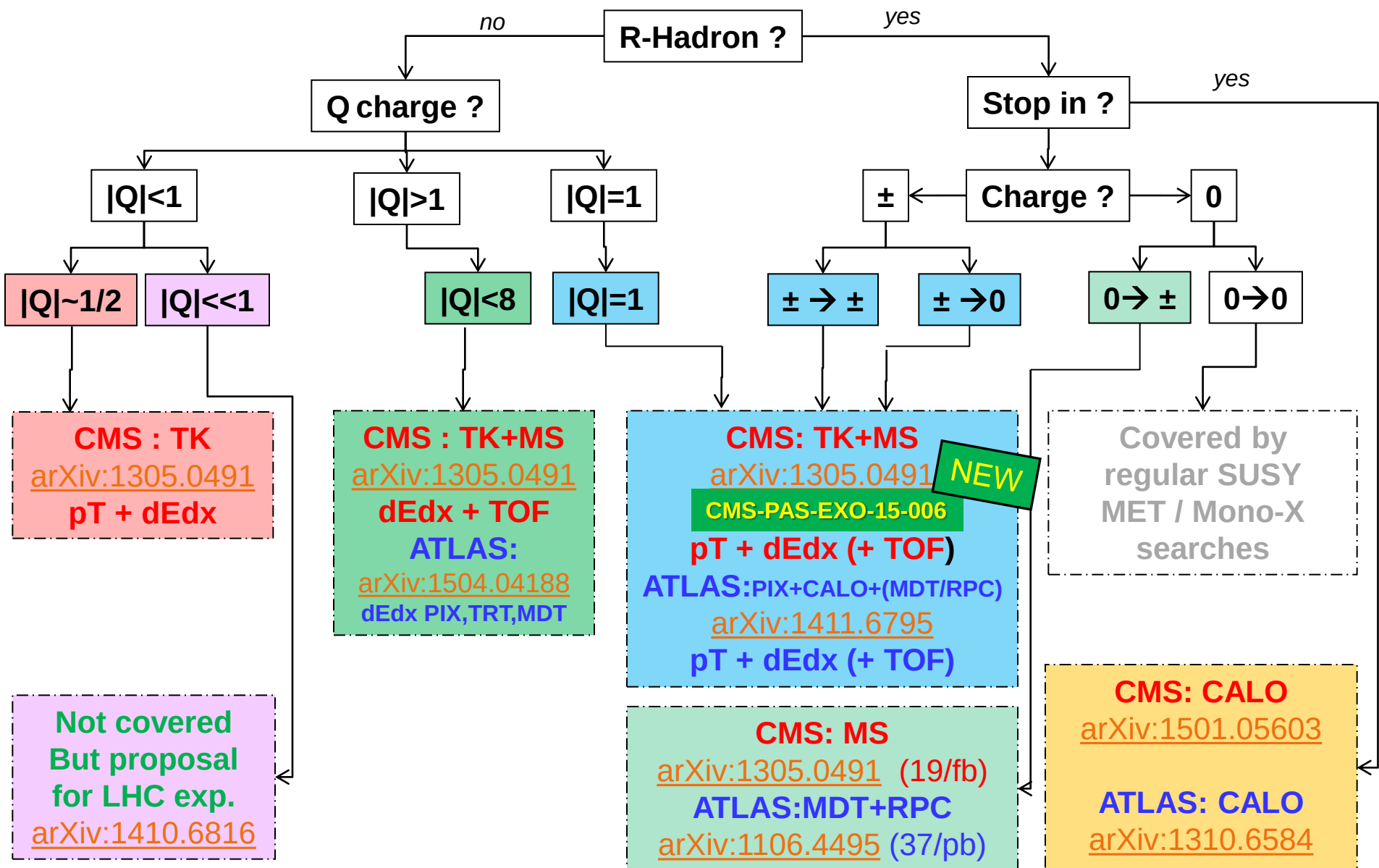
Background (ABCD)



Limits







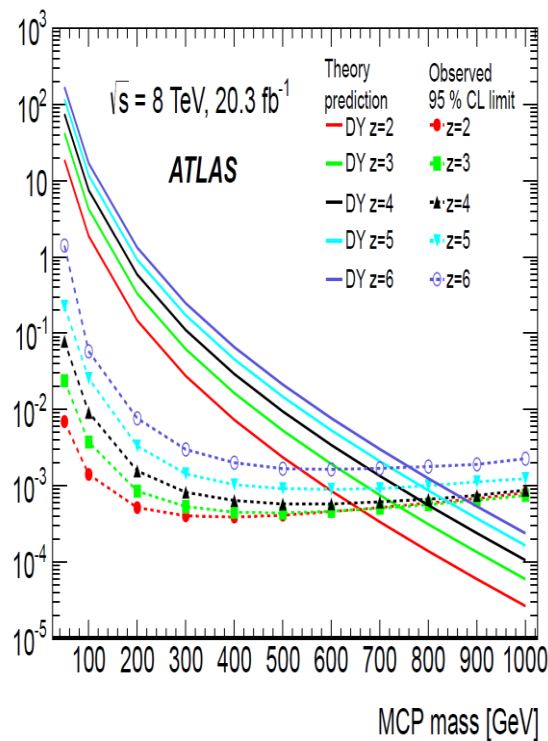
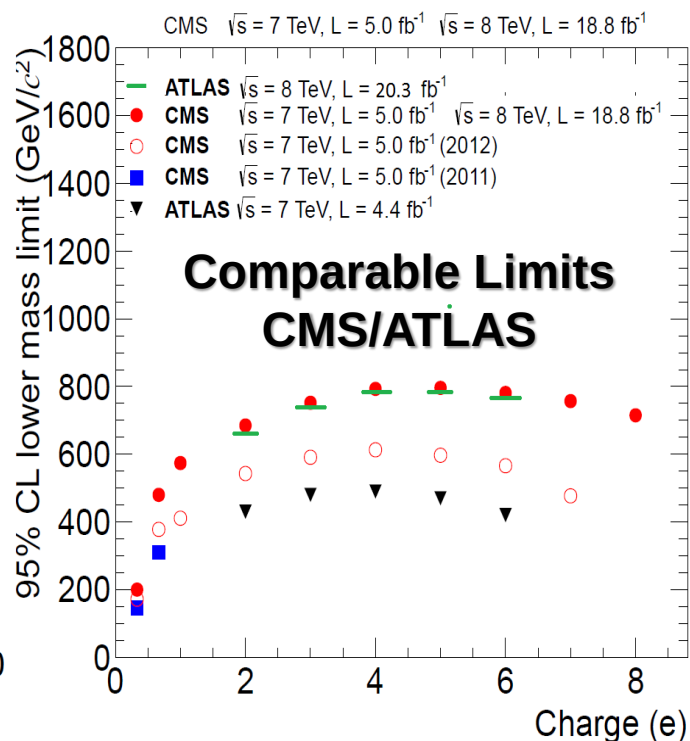
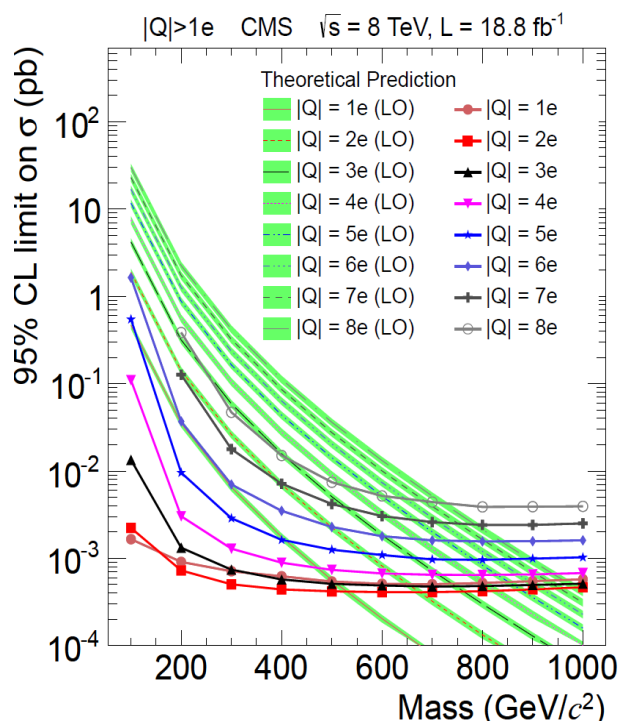
The search in a nutshell

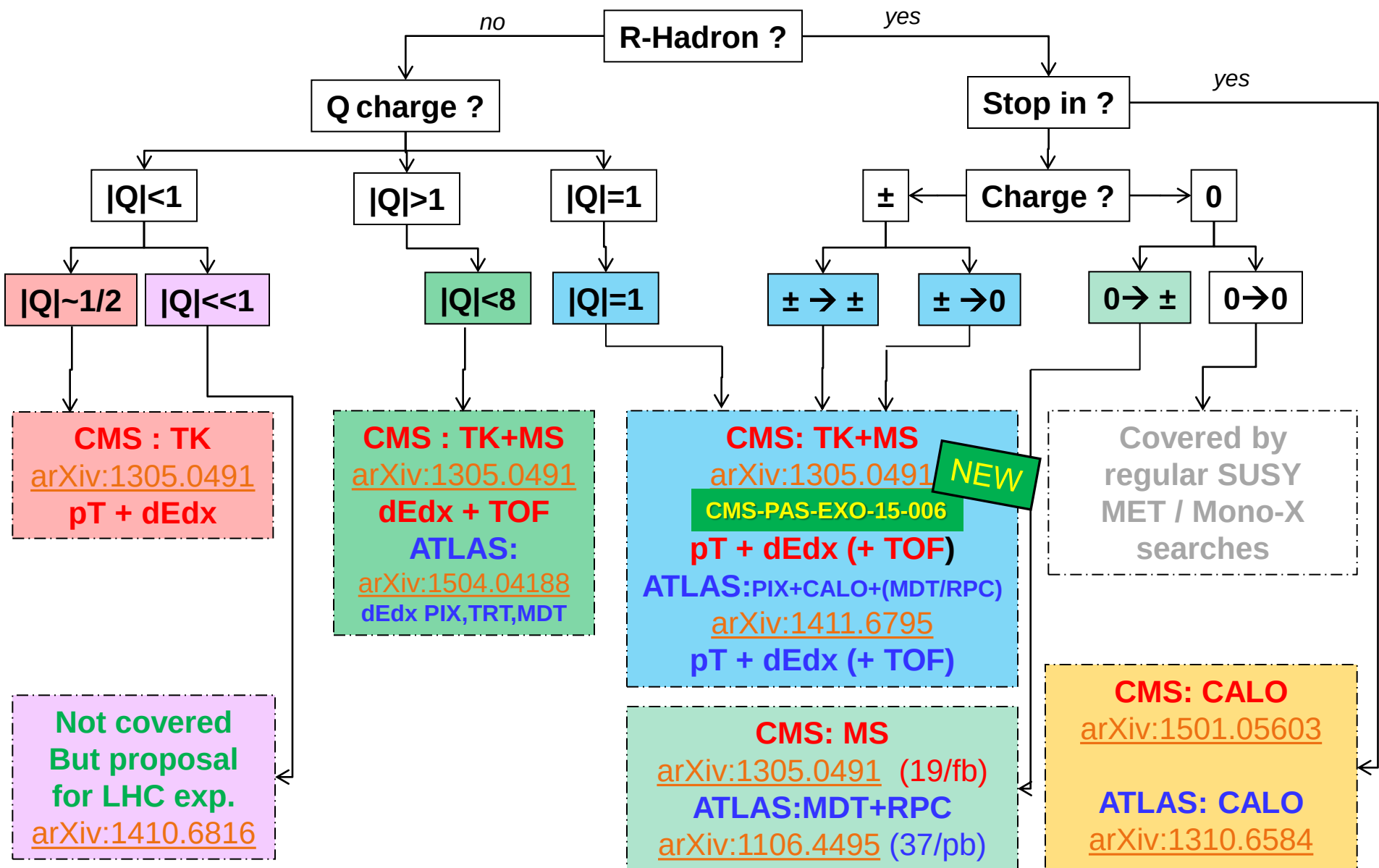
- Challenges:**

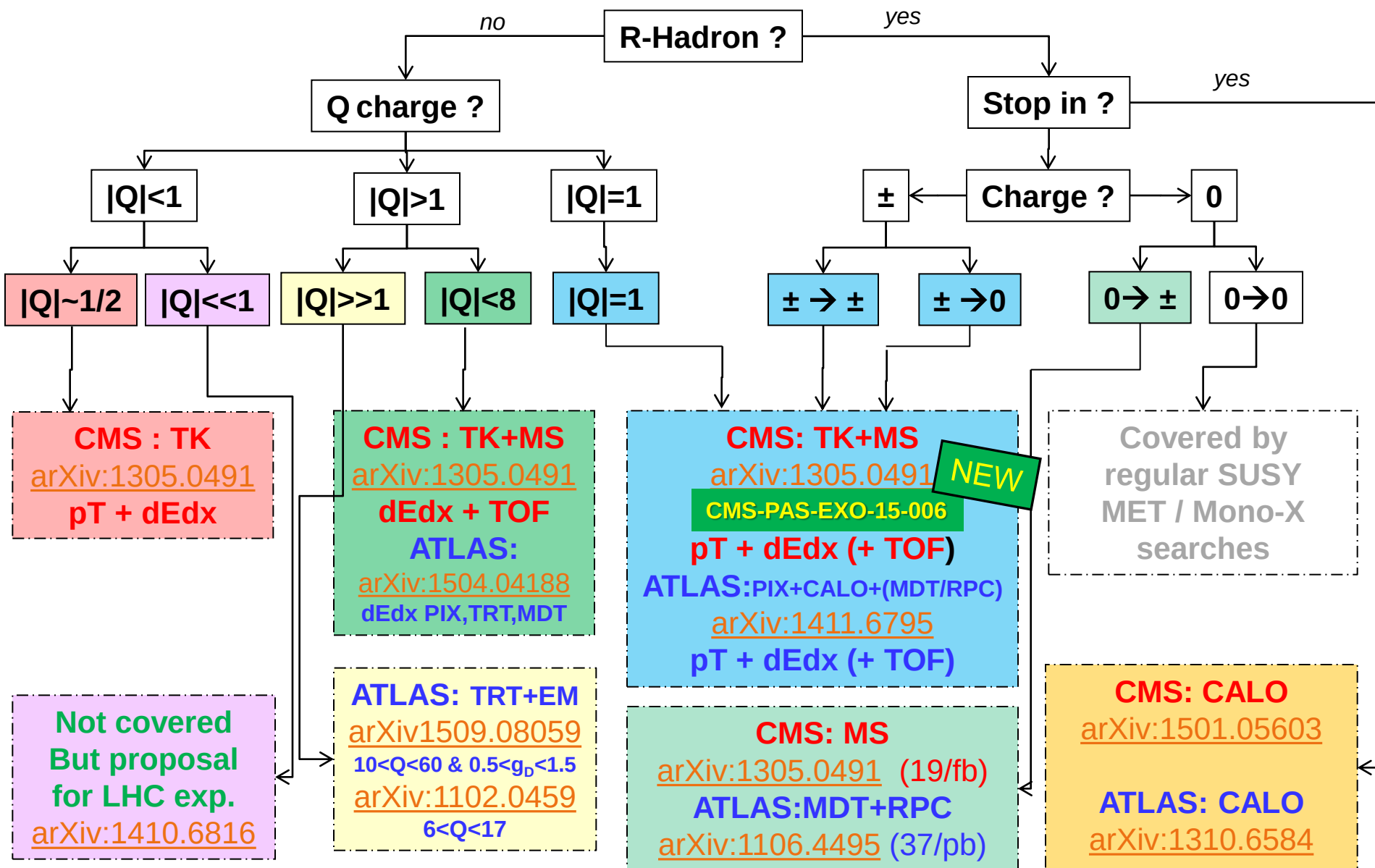
- Tracking: smaller reco p_T ($= p_{T, \text{True}} / Q$) $\rightarrow$ trigger & background rate

- Solutions:**

- Requires it to reach MS (many redundant measurements dE/dx or TOF)
- Keep track p_T threshold as low as possible (40 GeV)

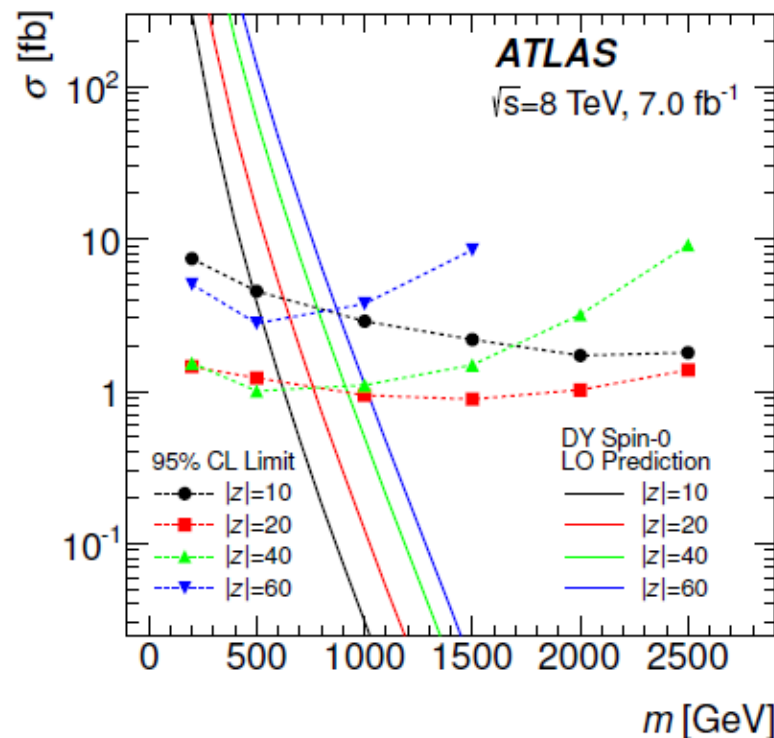
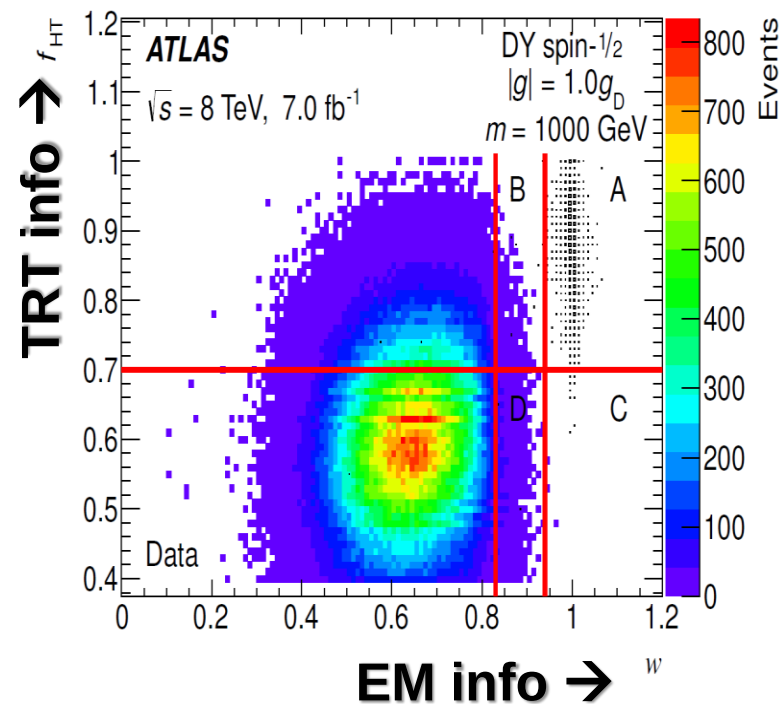


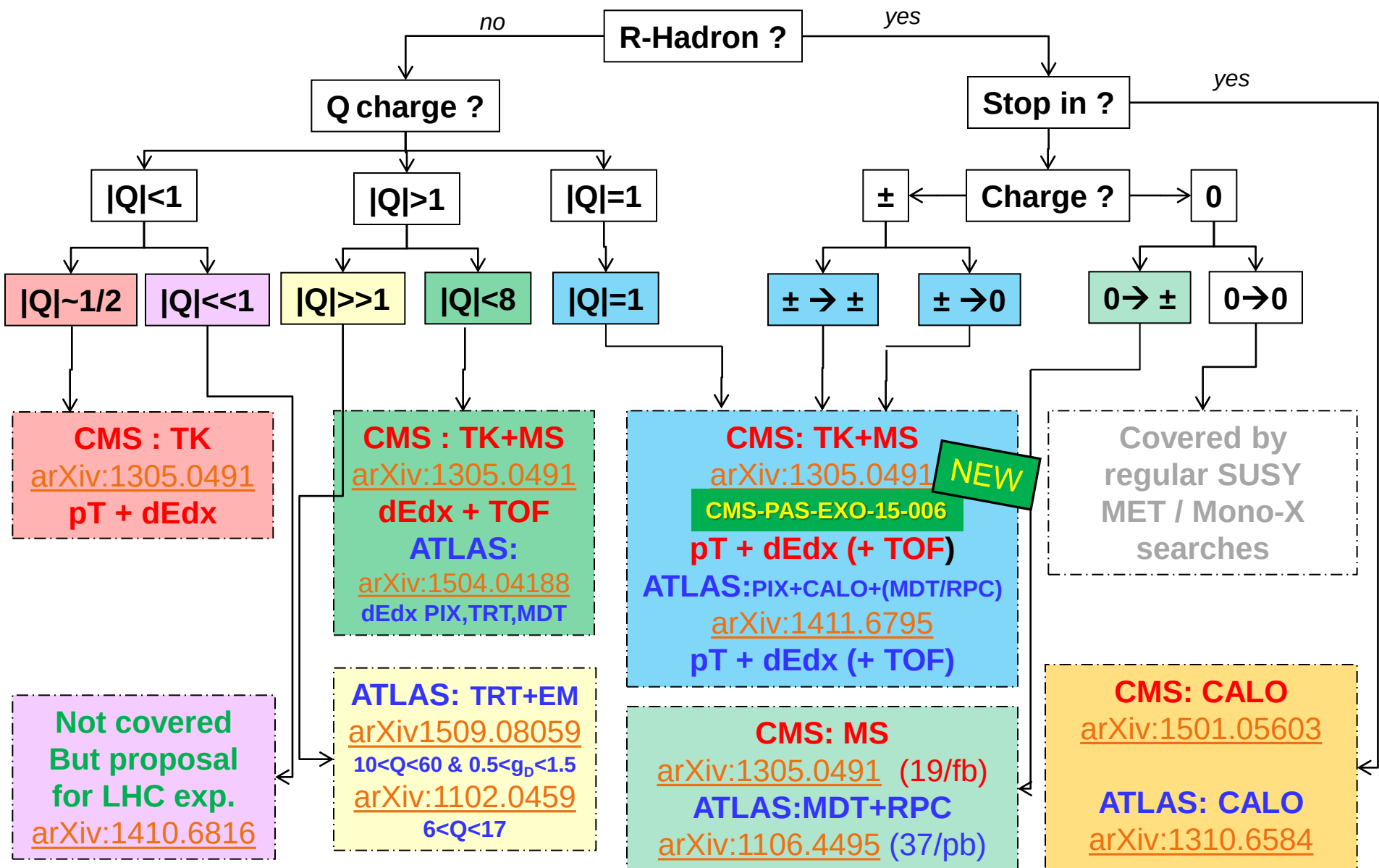




The search in a nutshell

- **Challenges:**
 - Tracking: small reco p_T ($=p_{T, \text{True}} / Q$)
 - Huge energy loss: not reach outer detector (+ possible DAQ issues)
- **Solutions:** Dedicated trigger + no tracking + only inner detectors (TRT and EM)



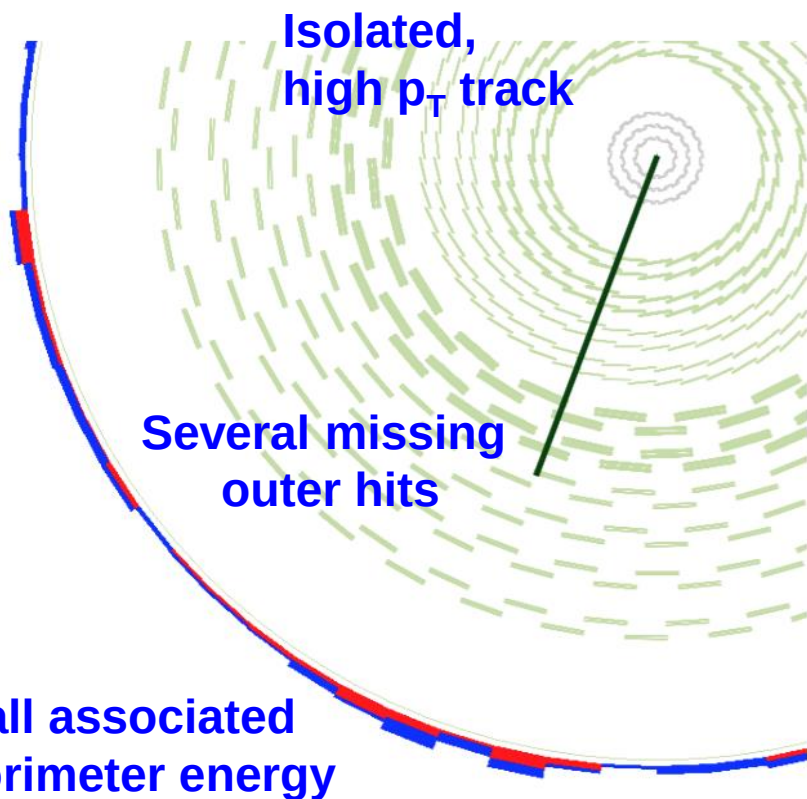
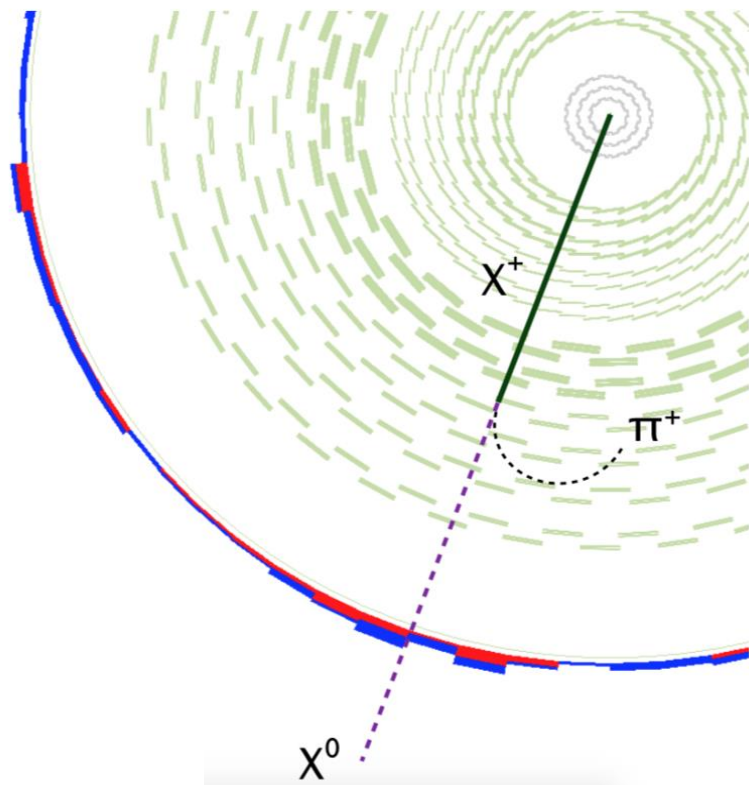


- Introduction to LL physics (Why ? How ?) ✓
- Searches for LL particles decaying OUTSIDE the detector
 - Neutral ✓
 - Charged ✓
- Searches for LL particles decaying INSIDE the detector
 - Charged ←
 - Neutral
- Super exotic particles
- Summary

AMSB SUSY predicts nearly degenerate neutralino and chargino: $\Delta M \approx 100\text{--}200\text{ MeV}$
→ Chargino is long-lived and decays to neutralino and a very soft pion

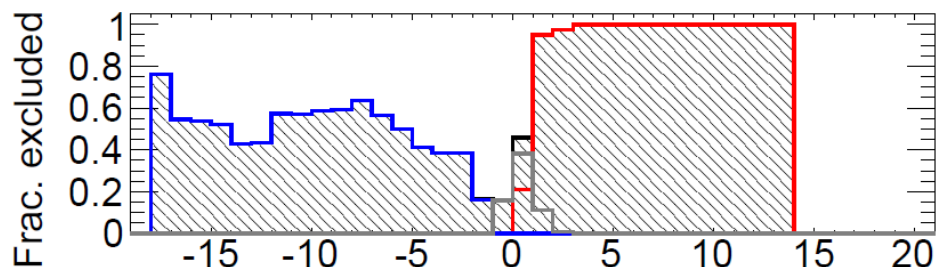
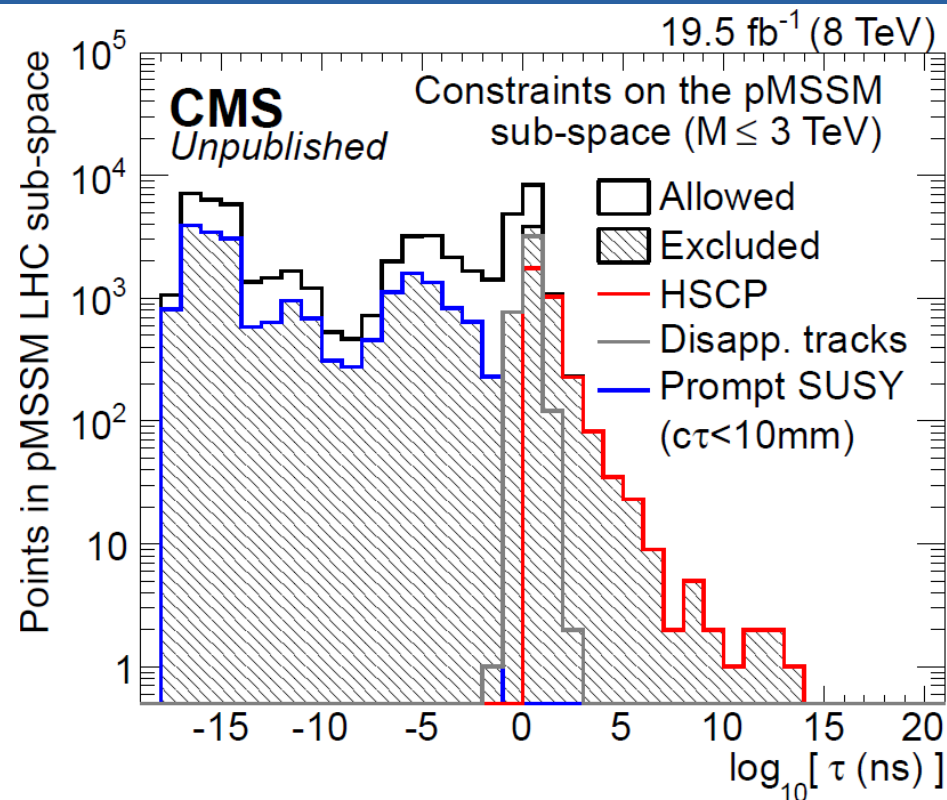
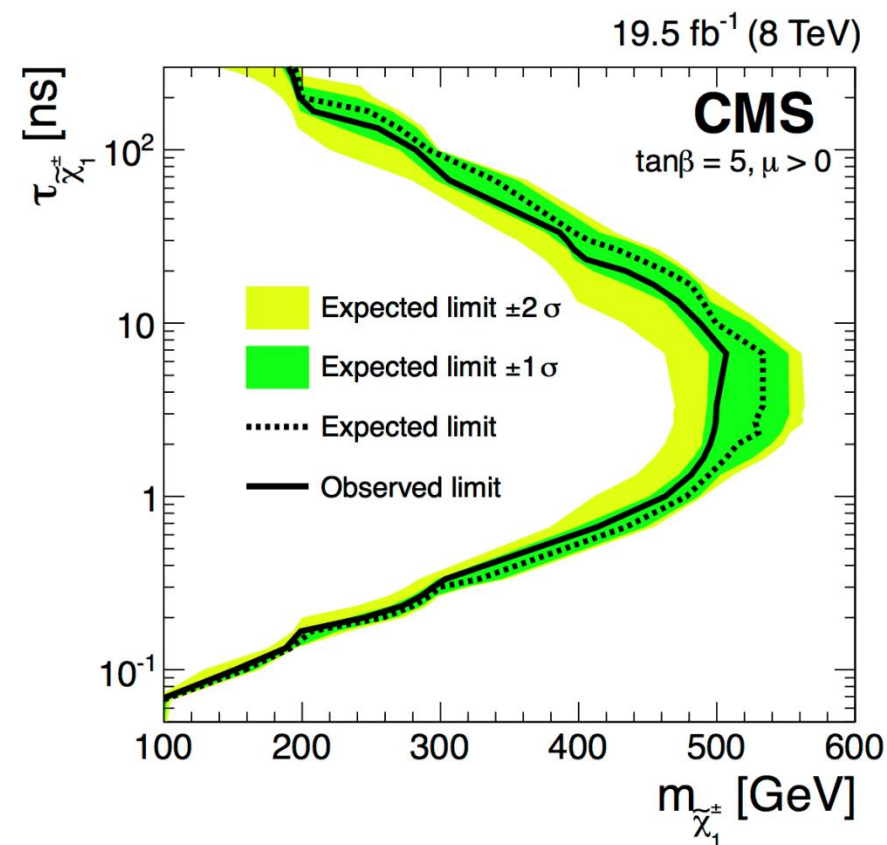
Disappearing High p_T Track

- **Signature** : tracker track with **high p_T** and **missing outer hits** and **small E_{calo}**
- **Backgrounds** : **fake tracks, kinked tracks (bremsstrahlung and nucl. int)**



AMSB models with a chargino mass less than 260 GeV, corresponding to a mean proper lifetime of 0.2 ns are excluded

Start filling the gap in pMSSM too

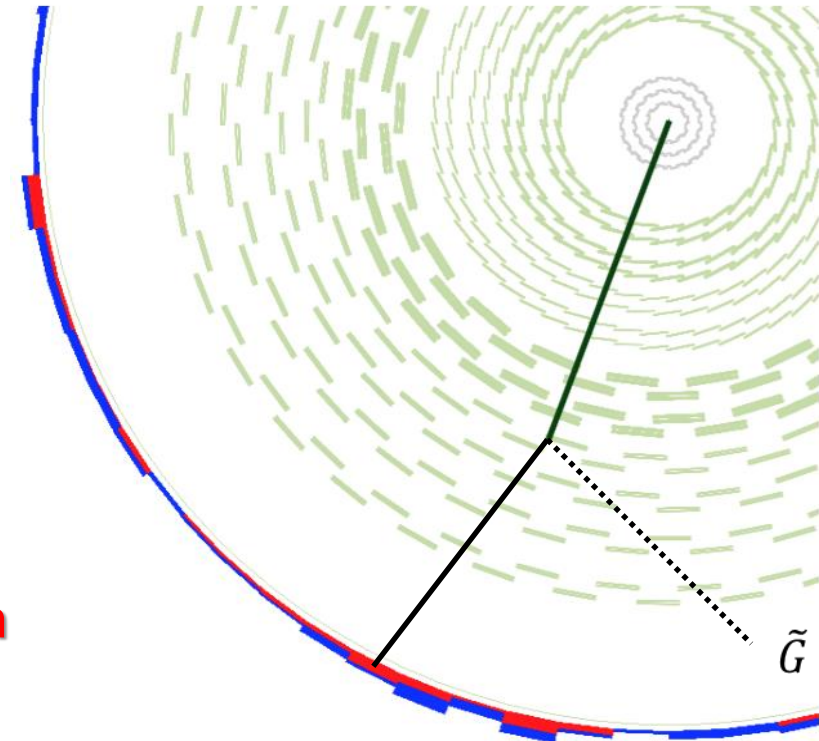
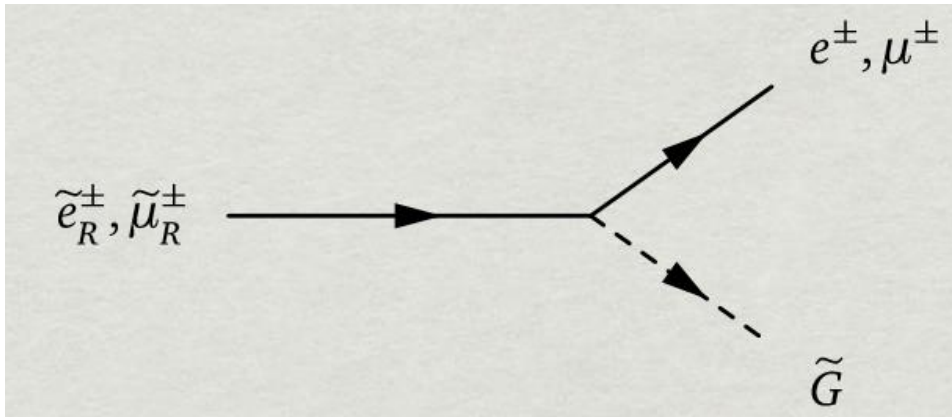


Still a gap to fill for intermediates lifetimes

Similar to the disappearing track search, except that the decay product has a high p_T charged candidate

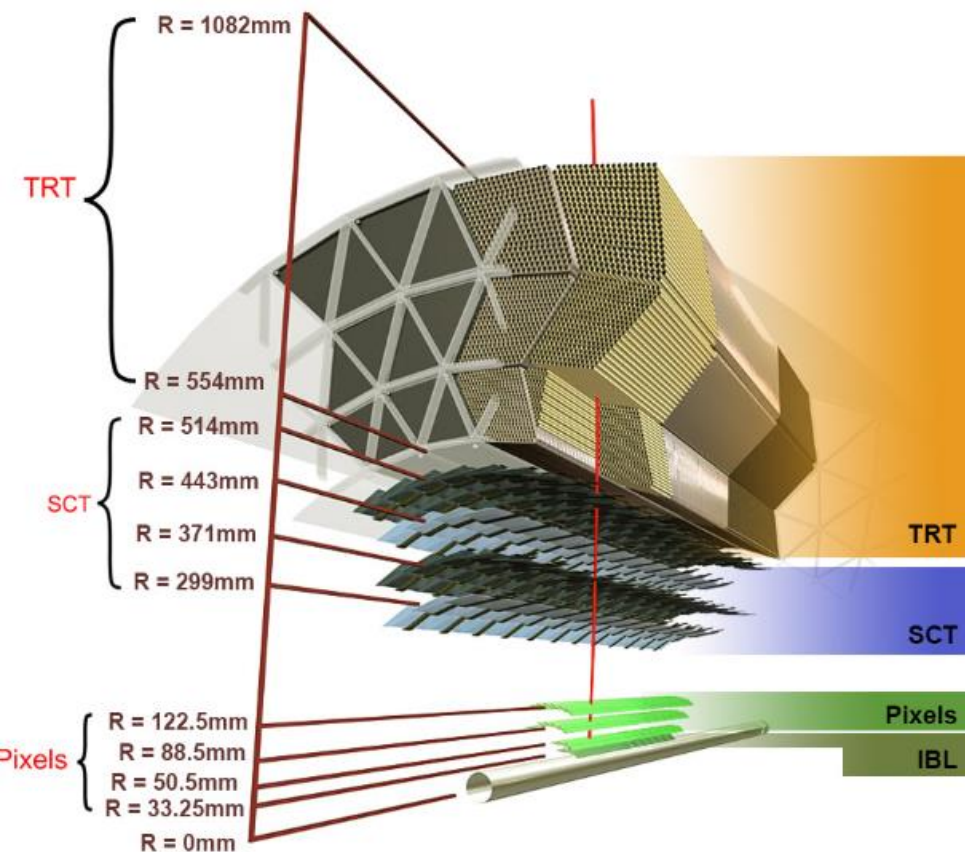
High p_T Kinked Track

- **Signature** : tracker track with **high p_T** and **a kink somewhere**
- **Backgrounds** : **fake tracks, bremsstrahlung and nucl. int**



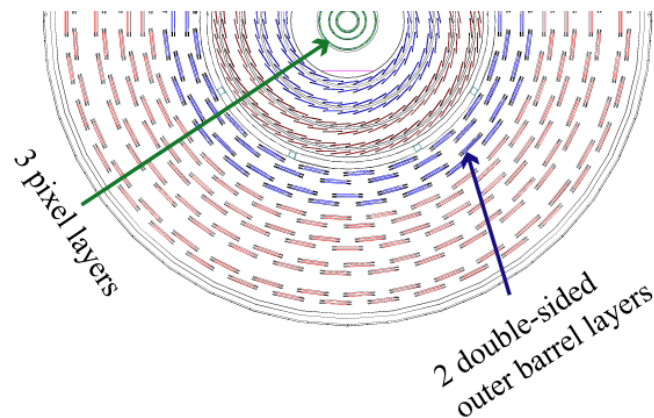
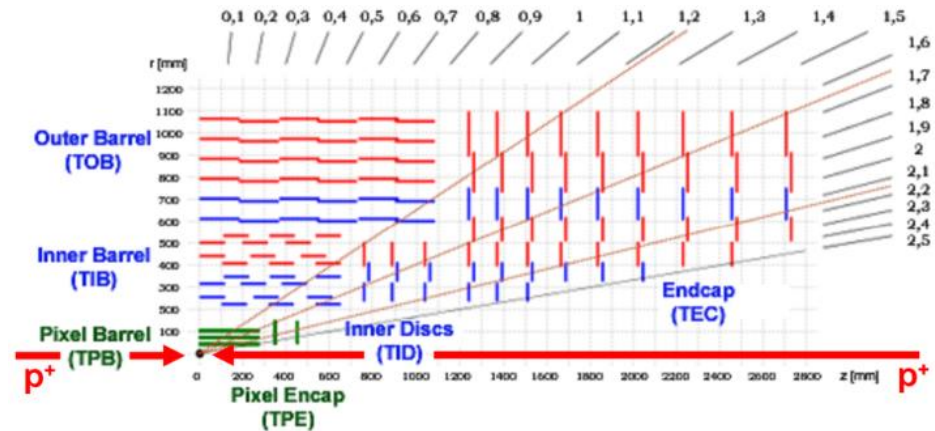
**No public results on this specific search
are currently public at the LHC.**

- Introduction to LL physics (Why ? How ?) ✓
- Searches for LL particles decaying OUTSIDE the detector
 - Neutral ✓
 - Charged ✓
- Searches for LL particles decaying INSIDE the detector
 - Charged ✓
 - Neutral ←
- Super exotic particles
- Summary



ATLAS tracker

- Inside a 2T solenoid
- Coverage up to $|\eta| \leq 2.5$
- $R \geq 50\text{mm}$ (run1), 33mm (run2)



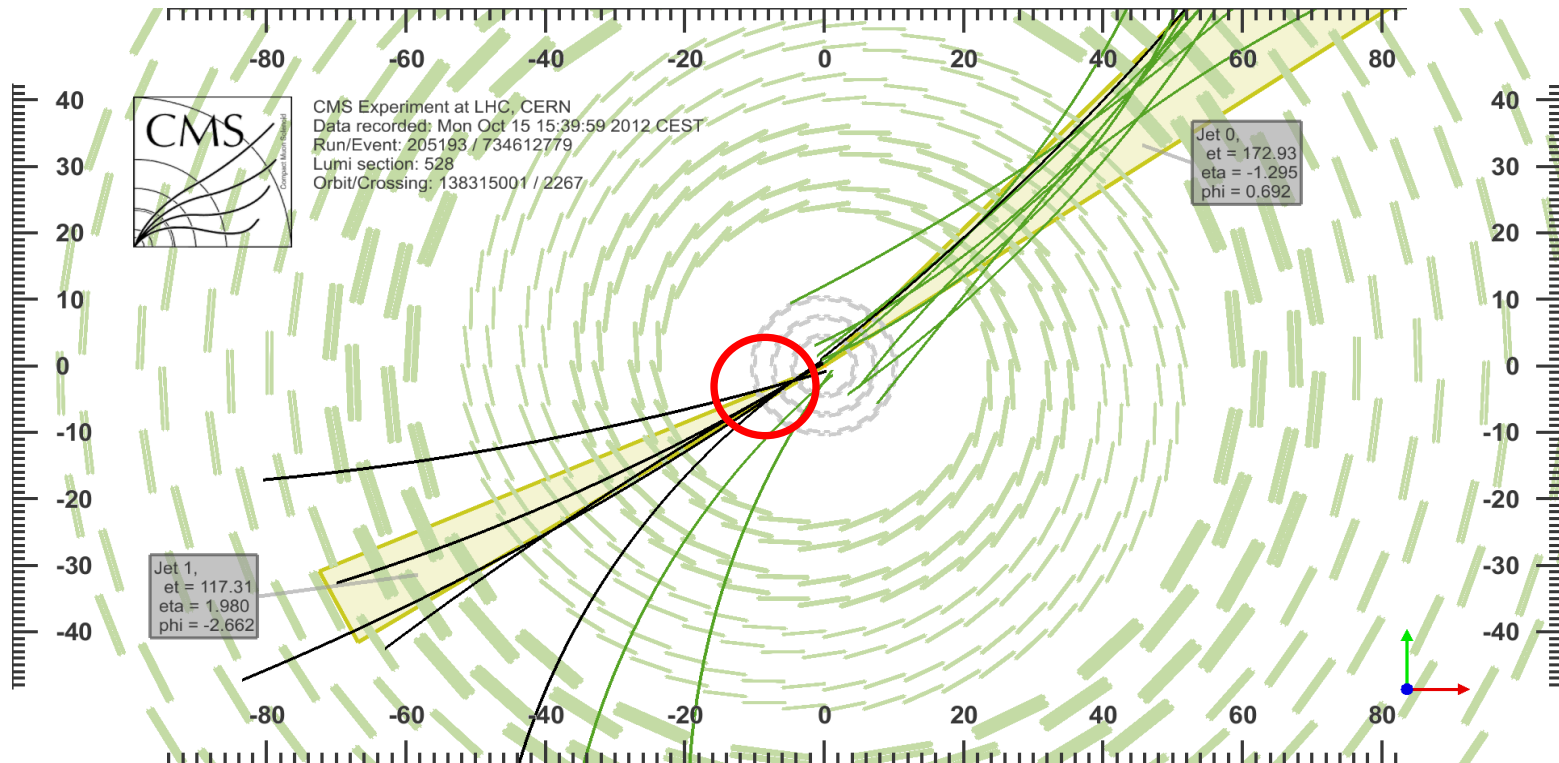
CMS tracker

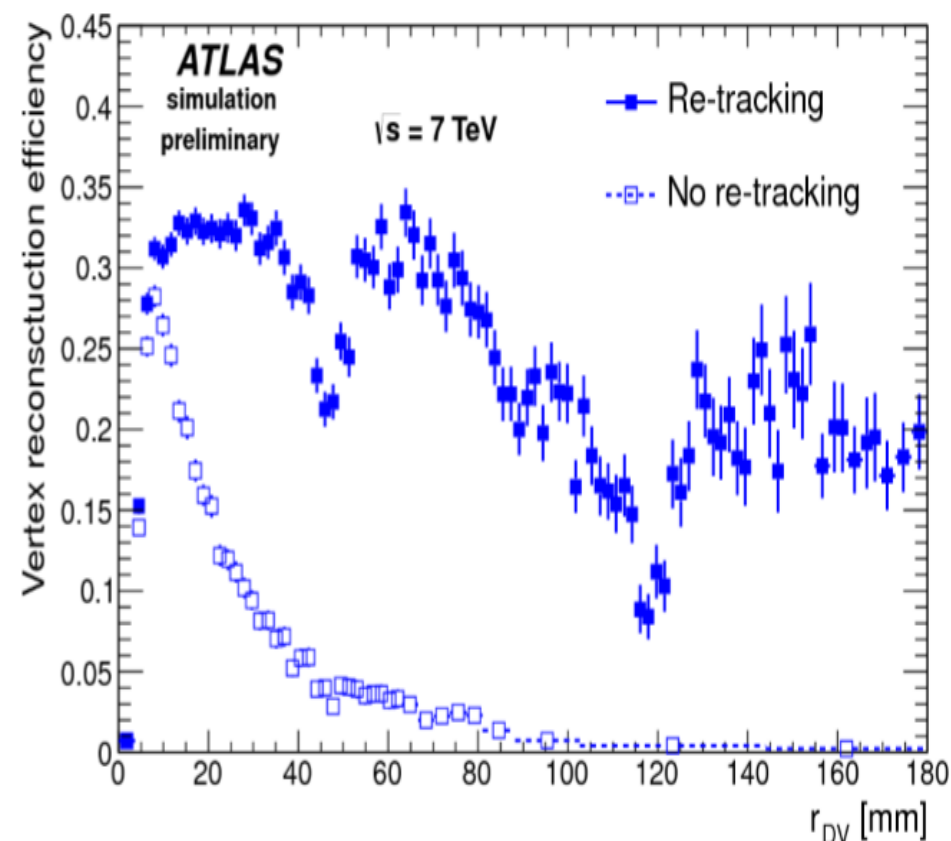
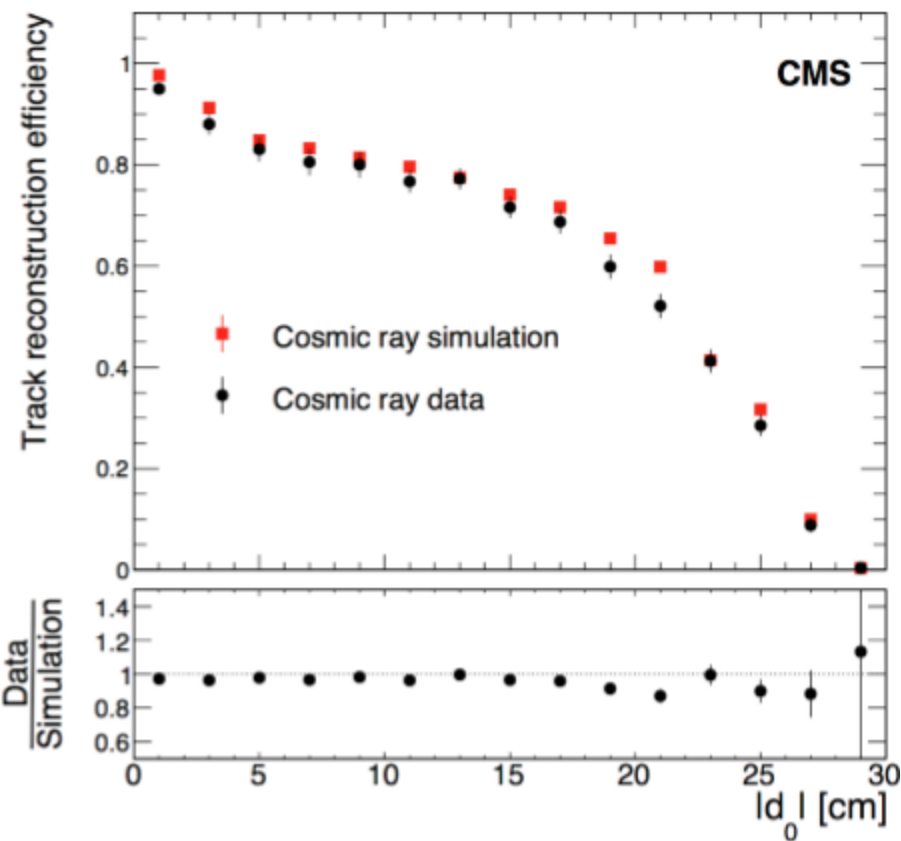
- Inside a 4T solenoid
- Coverage up to $|\eta| \leq 2.5$
- $R \geq 44\text{mm}$

- Vertexing is based on (high quality) track clustering
 - Clustering in CMS is done with Deterministic Annealing
 - Minimizes a free-energy-like function to determine the best number of vertices and their locations

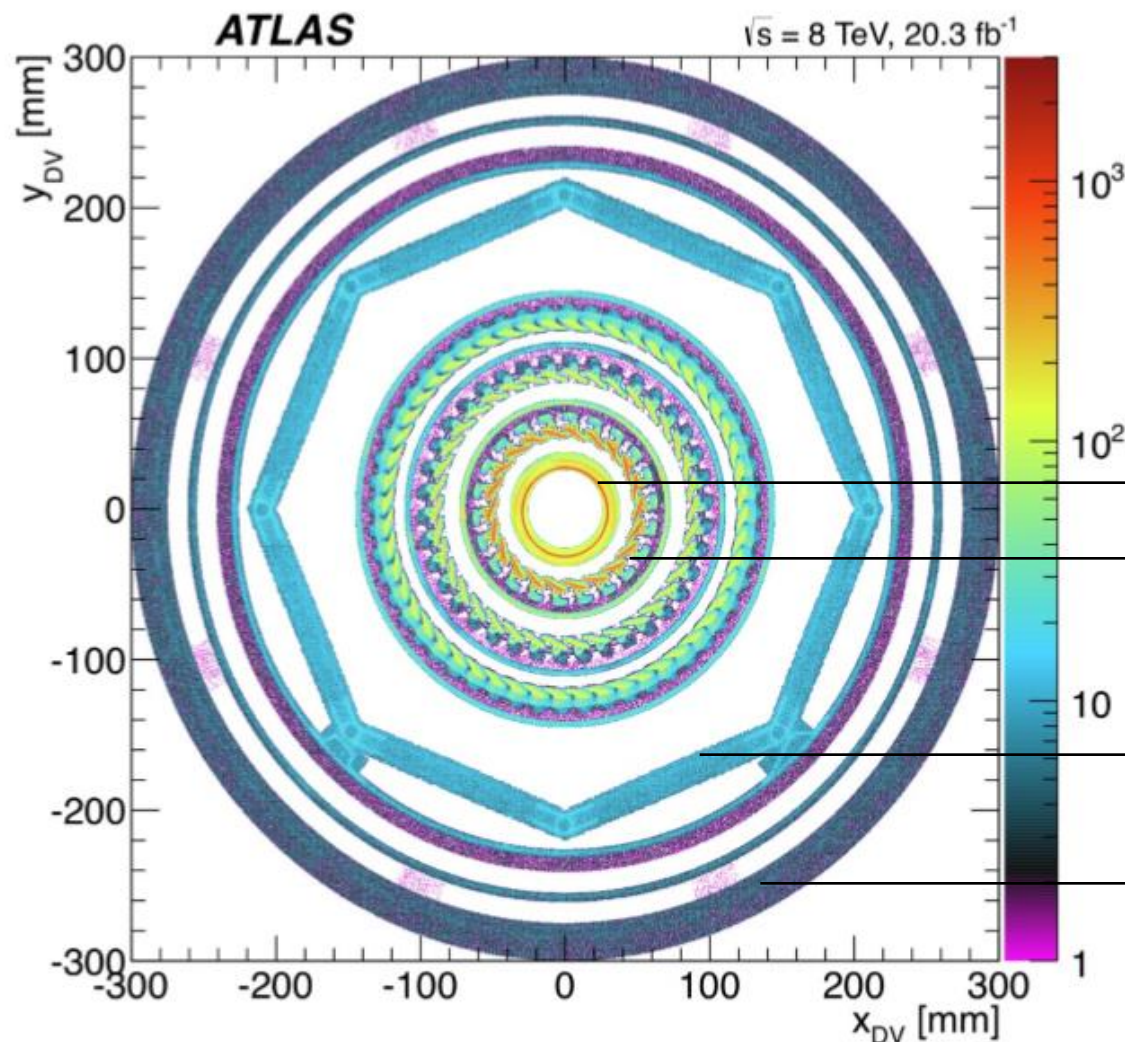
$$F = -T \sum_i^{\text{\# tracks}} p_i \log \sum_k^{\text{\# vertices}} \rho_k \exp \left[-\frac{1}{T} \frac{(z_i^T - z_k^V)^2}{\sigma_i^{z^2}} \right]$$

→ Track origin
 → Vertex position



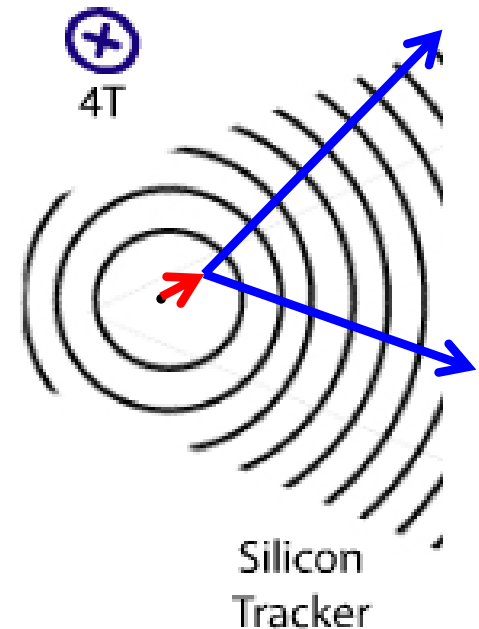


- For large displacement (longer life-time), we have Less tracking layers to reconstruct the tracks
 - Tracking efficiency drops for large displacement
 - Vertex efficiency also drops (consequently)



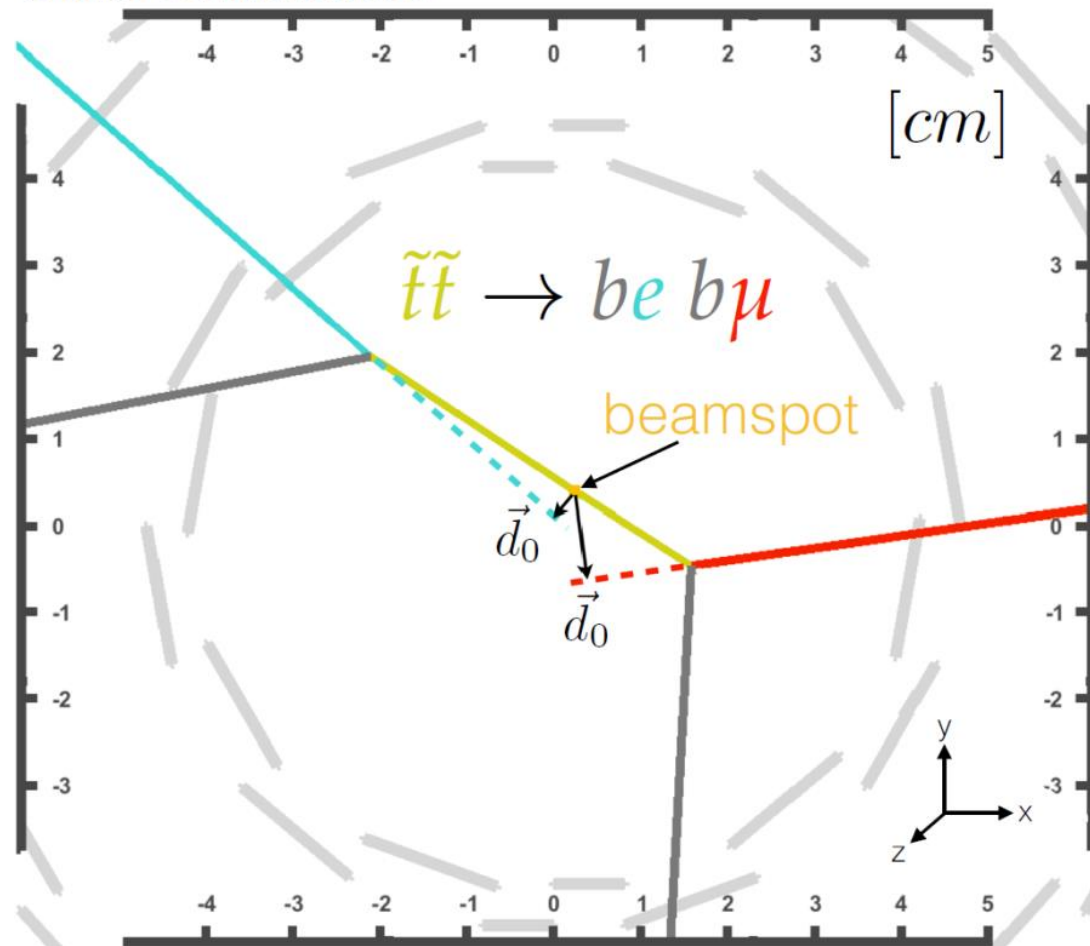
- Distribution of the position of reconstructed secondary vertex with ≥ 5 Tracks in ATLAS
 - Mechanical structures clearly appears
- Beam pipe
- Pixel layers (x3)
- Mechanical structures separating pixels and SCT

- **Challenges**
 - Many sources of backgrounds
 - Pileup
 - Tau leptons $\rightarrow c\tau = \sim 0.1\text{mm}$
 - B hadrons $\rightarrow c\tau = \sim 0.5\text{mm}$
 - Need good vertexing algorithms
 - Tracker (pixel) resolution is crucial
- **Advantages**
 - Decay product are almost prompt
 - Standard object reconstruction
 - Analysis similar to prompt search at the exception of the vertex requirement

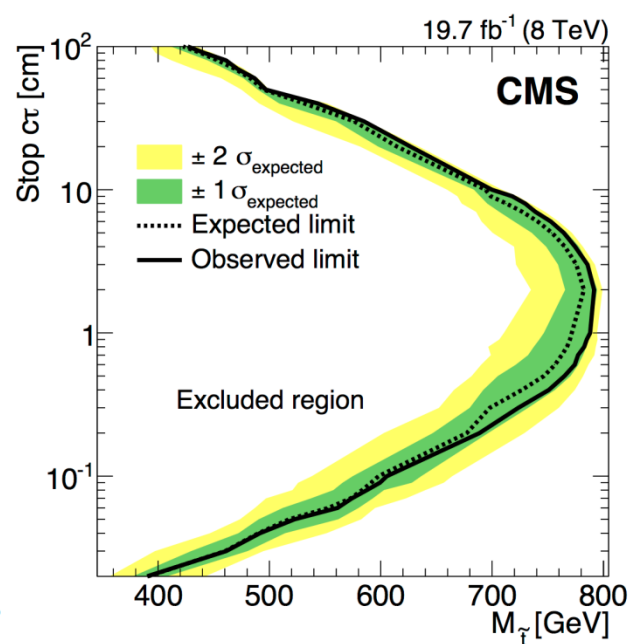
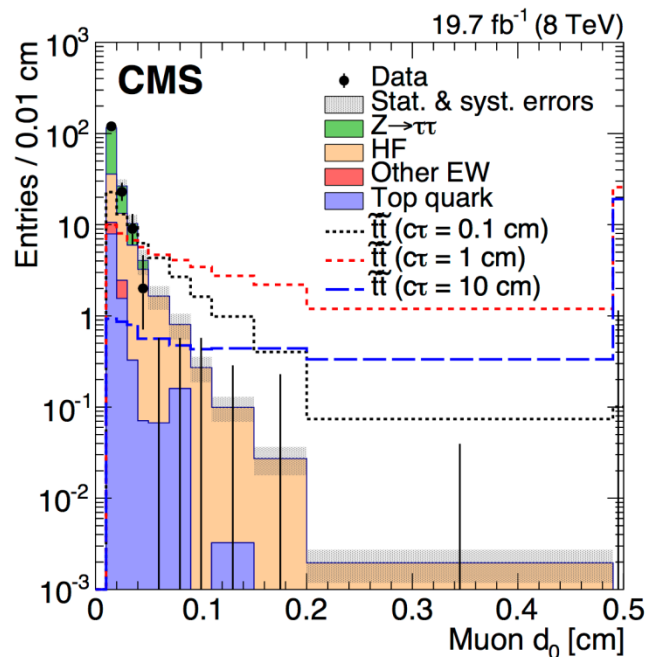
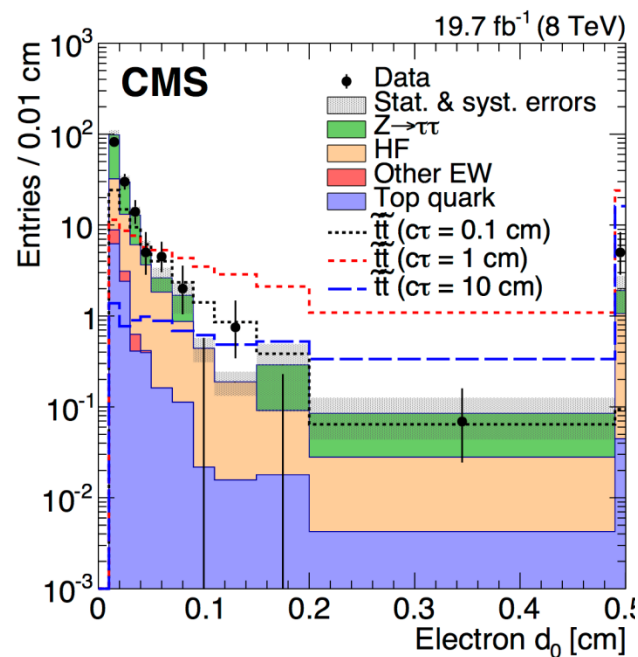


- **Search for Displaced SUSY in Dilepton Final States (arXiv:1409.4789)**
- **Signal model: $\tilde{t}\tilde{t} \rightarrow (be)(b\mu)$**
- **Require**
 - a displaced e
 - a displaced μ
 - (no vtx requirement)
- D_0 is the distance of closest approach in the transverse plane between lepton and the LHC beam axis

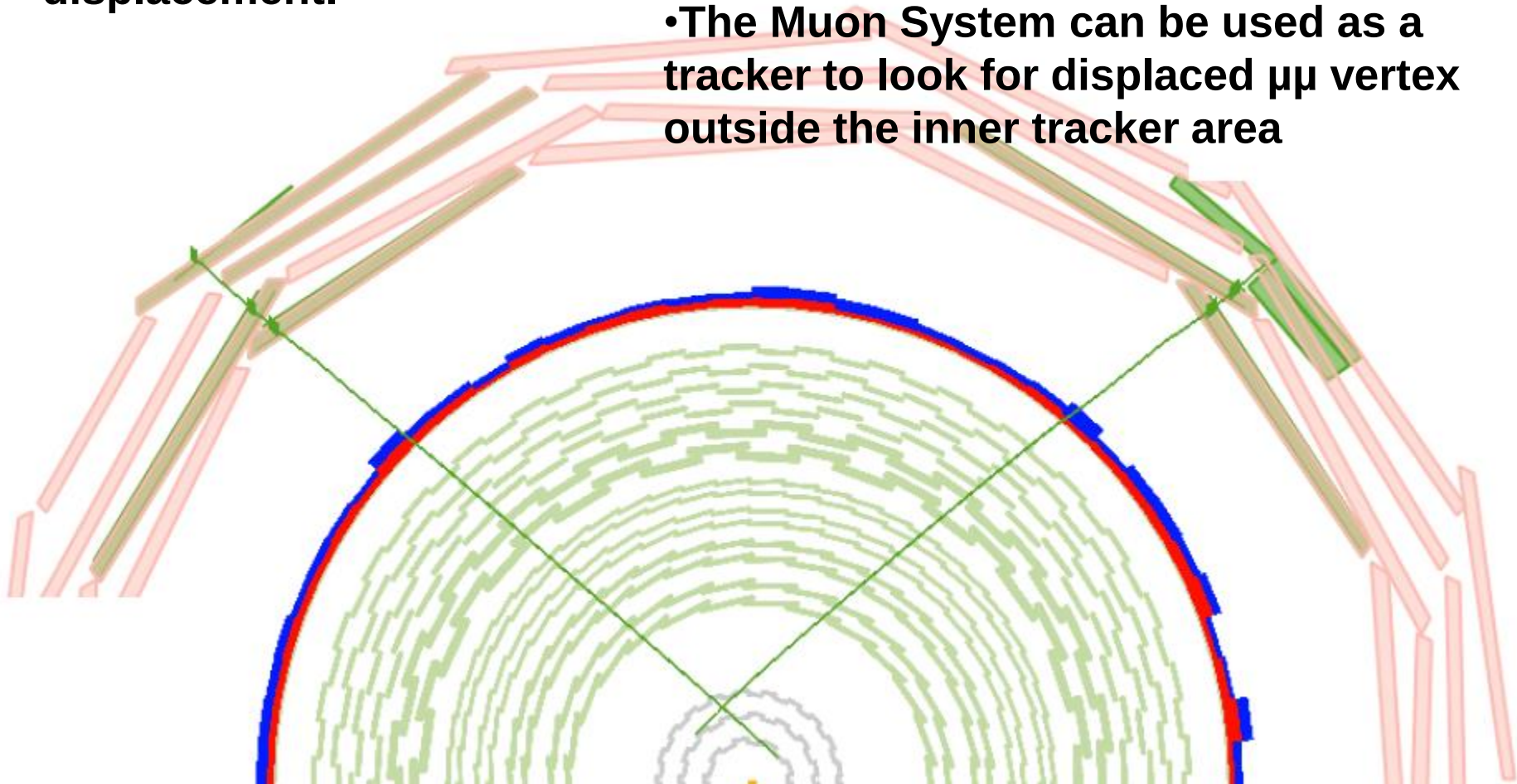
CMS Simulation



- **Search for Displaced SUSY in Dilepton Final States (arXiv:1409.4789)**
- Signal model: $\tilde{t}\tilde{t} \rightarrow (be)(b\mu)$
- **Require**
 - a displaced e
 - a displaced μ
 - (no vtx requirement)



- **Challenges**
 - Lepton trigger and reco efficiency is lower for large displacement.
- **Advantages**
 - Practically no backgrounds at the exception of fake / misreco tracks
 - The Muon System can be used as a tracker to look for displaced $\mu\mu$ vertex outside the inner tracker area



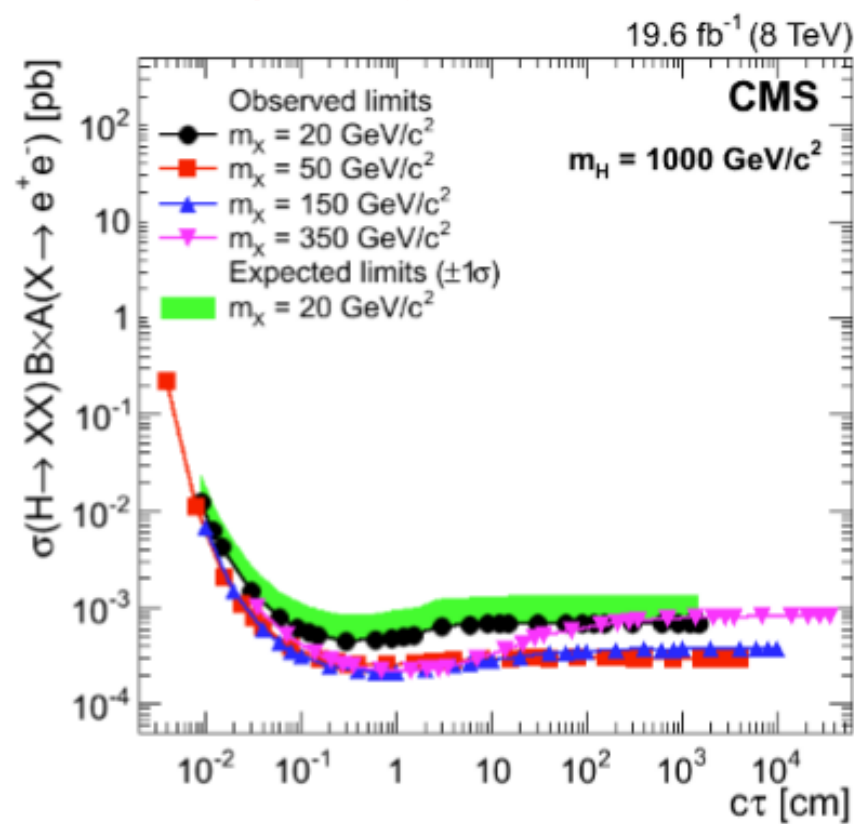
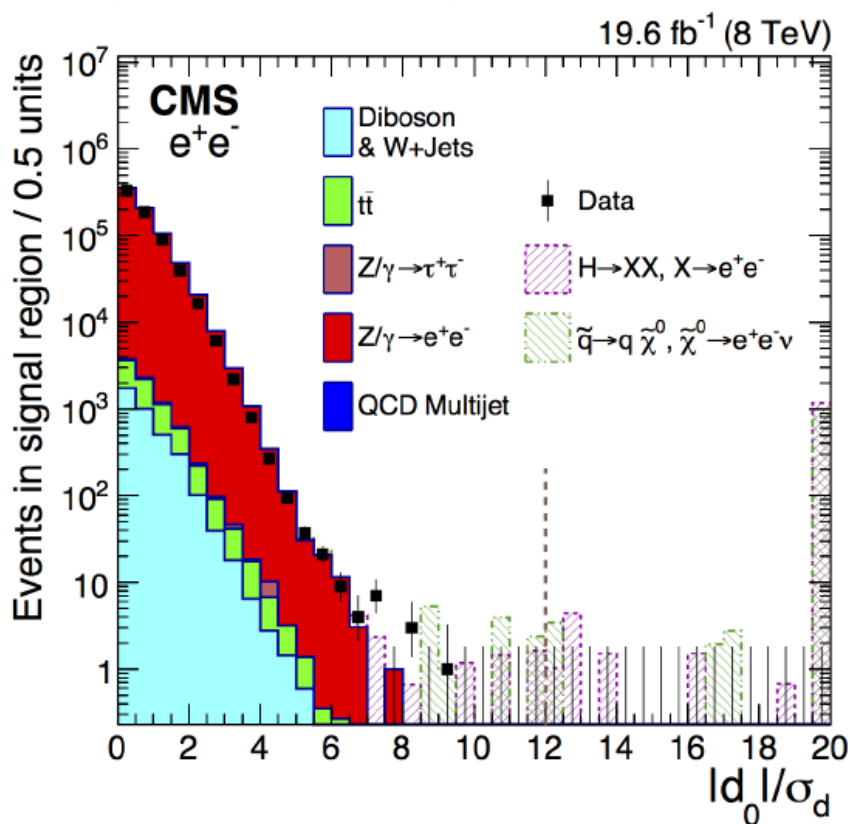
- **Search for Displaced Leptons (arXiv:1411.6977 + EXO-14-12)**

- **Signal model:**

- $H \rightarrow (X \rightarrow \ell\ell)(X \rightarrow \ell\ell)$ with X a long-lived neutral particle

- **RPV SUSY** $\tilde{q} \rightarrow (q)\tilde{\chi}^0 \rightarrow \ell\ell\nu$

- **Require the track pair to be incompatible with the primary vertex**

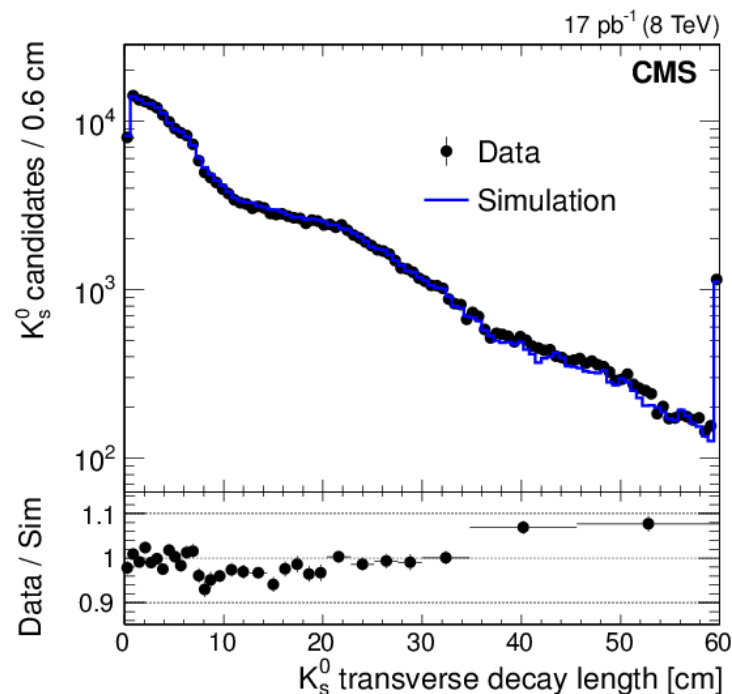
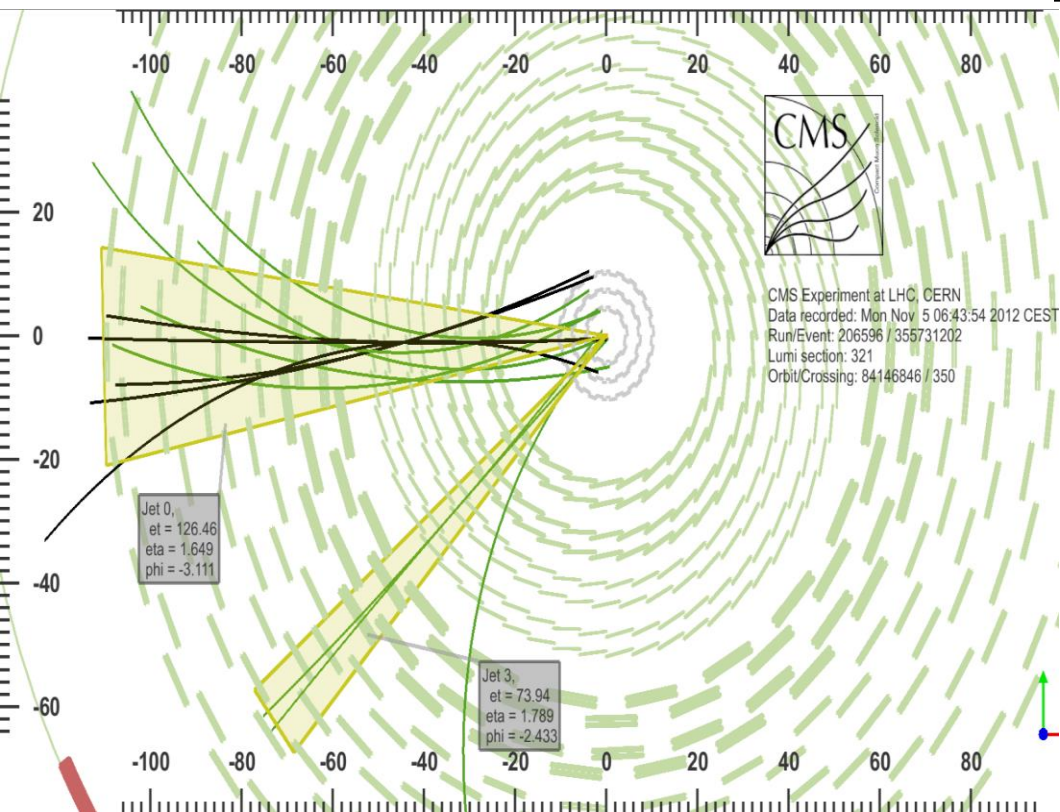


Challenges

- Triggering on such events
 - CPU time issues
 - High jet p_T thresholds
 - High HT thresholds

Advantages

- Generally easier to reconstruct a displaced vertex with more tracks. Hadronization help here
- Kaons can be used to verify data/MC for displaced vertex reconstruction



- **Search for Displaced Jets (arXiv1504.05162)**

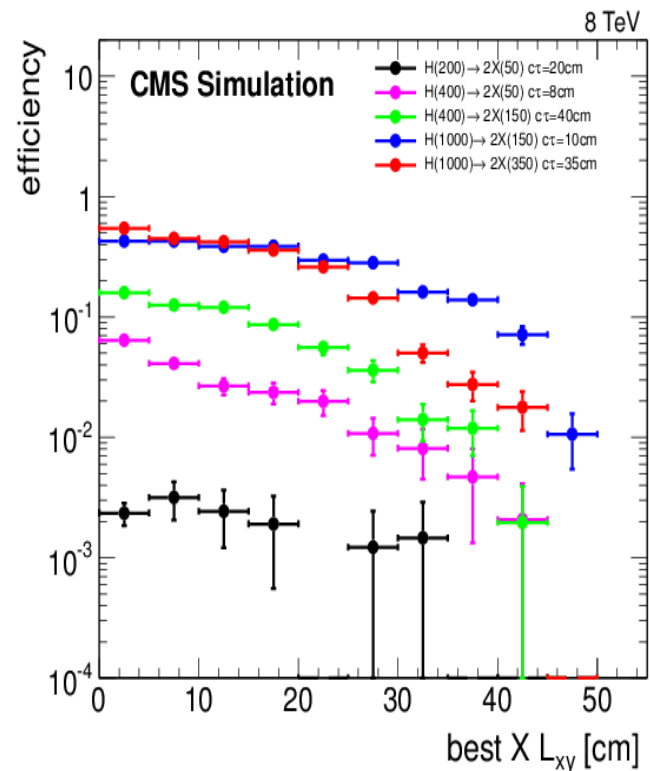
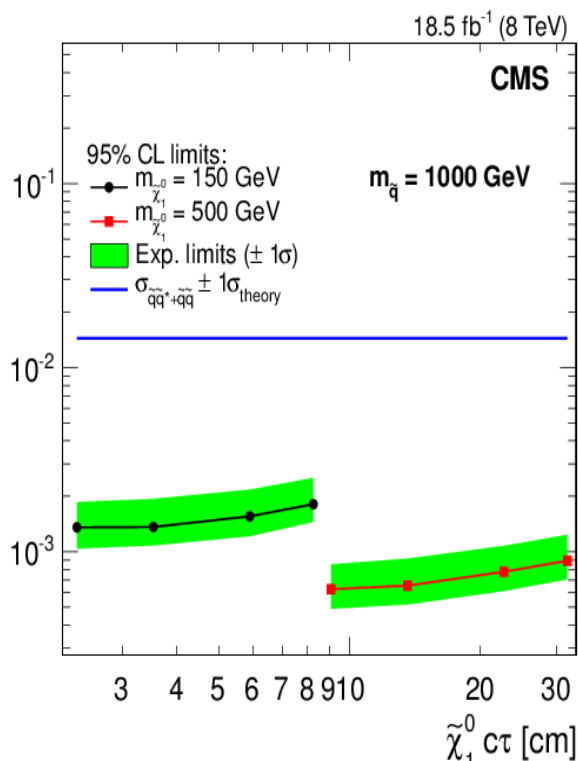
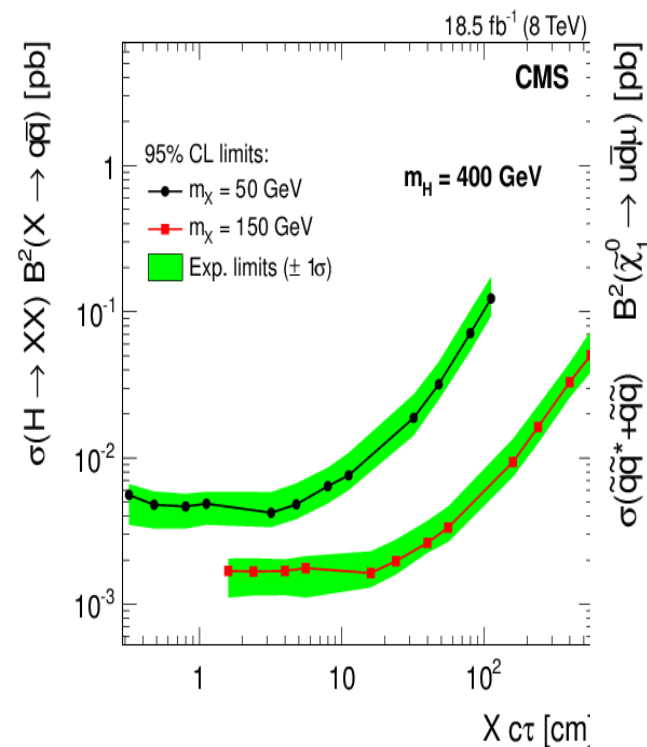
- **Signal model:**

- $H \rightarrow (X \rightarrow qq)(X \rightarrow qq)$ with X a long-lived neutral particle

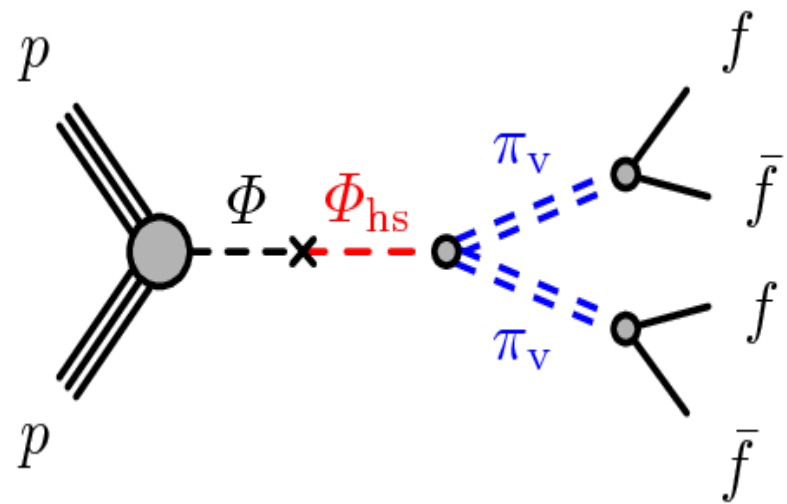
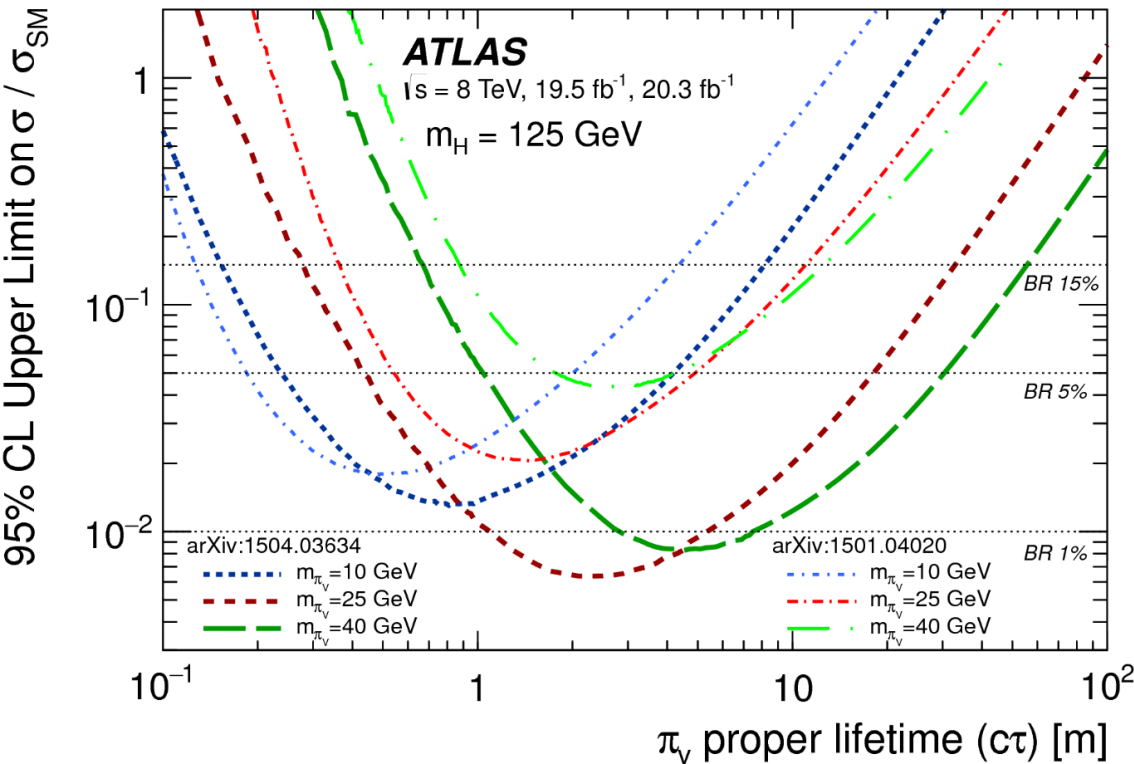
- **RPV SUSY**

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0 \rightarrow qq'\bar{q}''\mu$$

- **Require the track pair to be incompatible with the primary vertex**



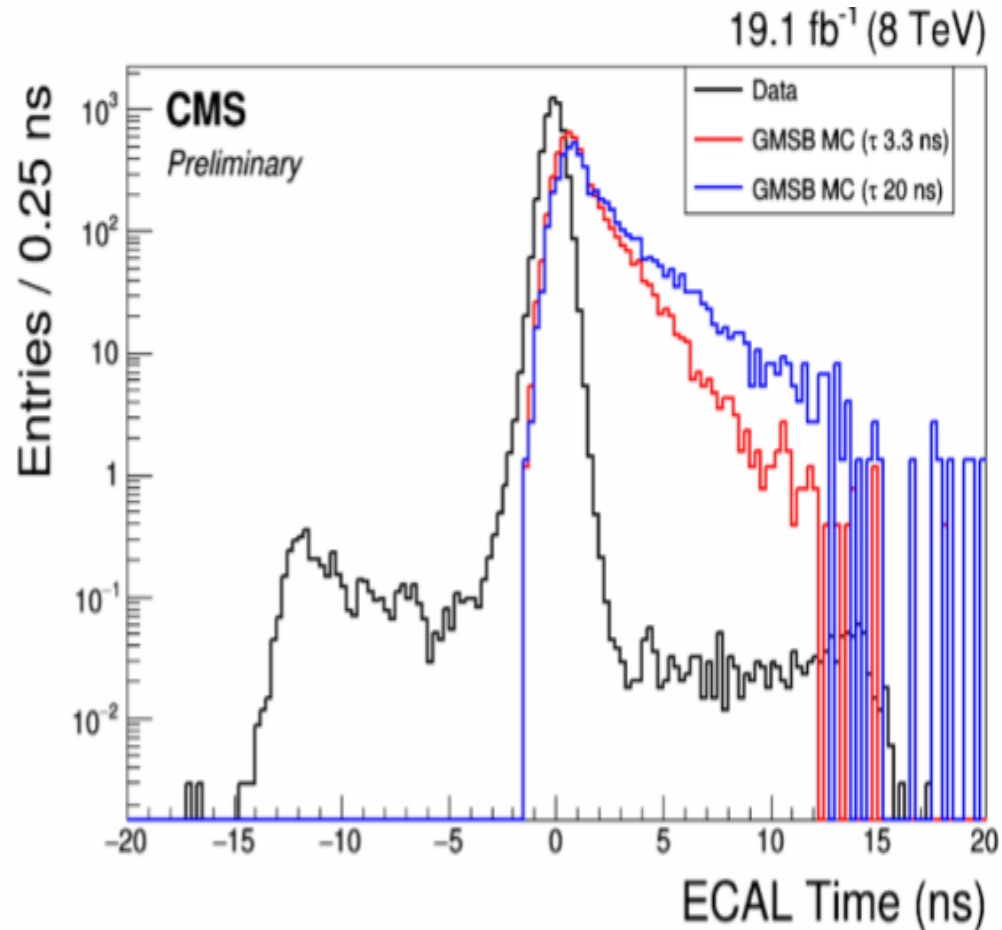
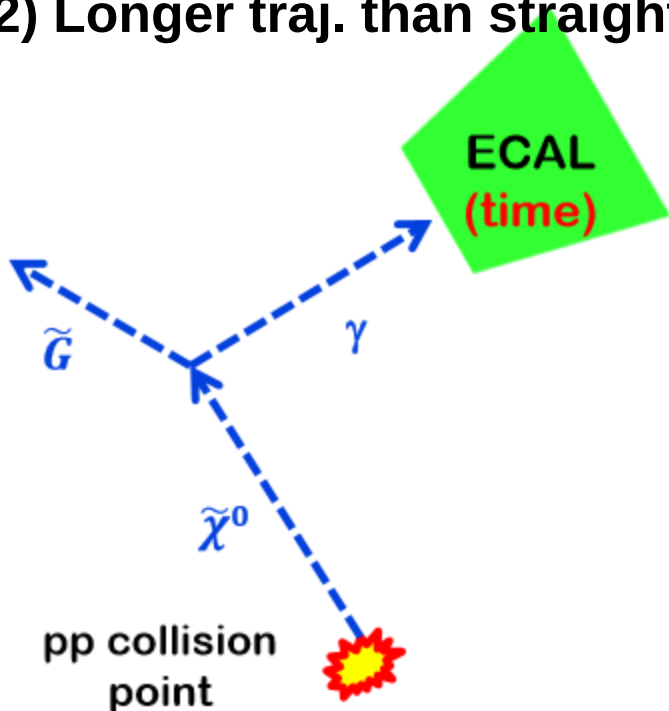
- **Search for Displaced Jets (arXiv1504.03634 & 1501.04020)**
- ***Sensitive to large displacement***
 - ***Outer layers***
 - ***ATLAS dimensions (larger lever arm)***
- ***Calorimeter segmentation & MS is used for vertexing***
- **Consider similar types of models**



Technique based on timing

Uses ECAL time ($\sigma=0.37\text{ns}$) to discriminate signal and bckg.

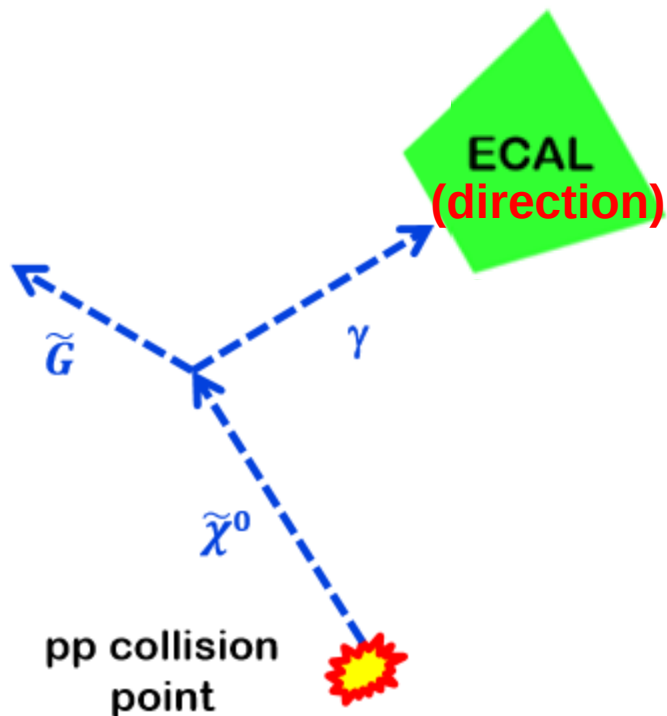
Delay is expected because of
1) Neutralino travelling at $v \ll c$
2) Longer traj. than straight line



Non-gaussian tails in data due to
bremsstrahlung photons from LHC
beam-halo or cosmic muons + ECAL
detector effects

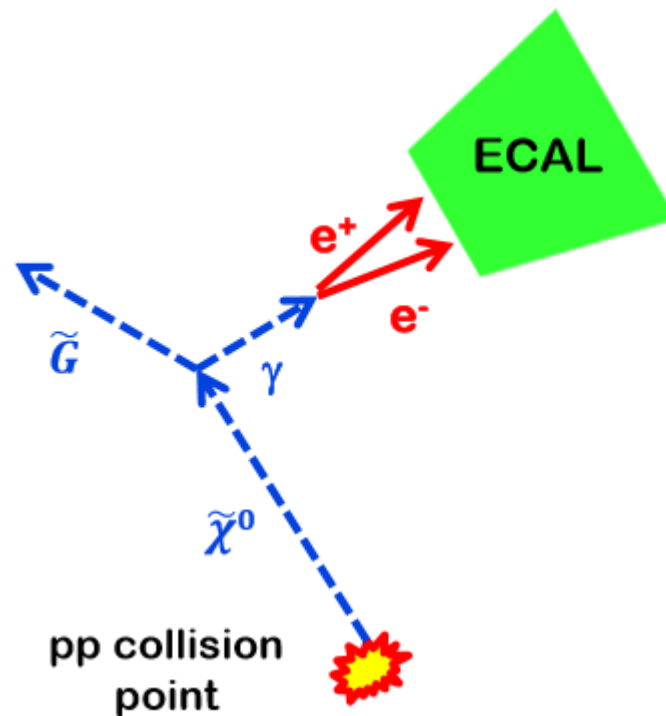
Technique based on shower profile

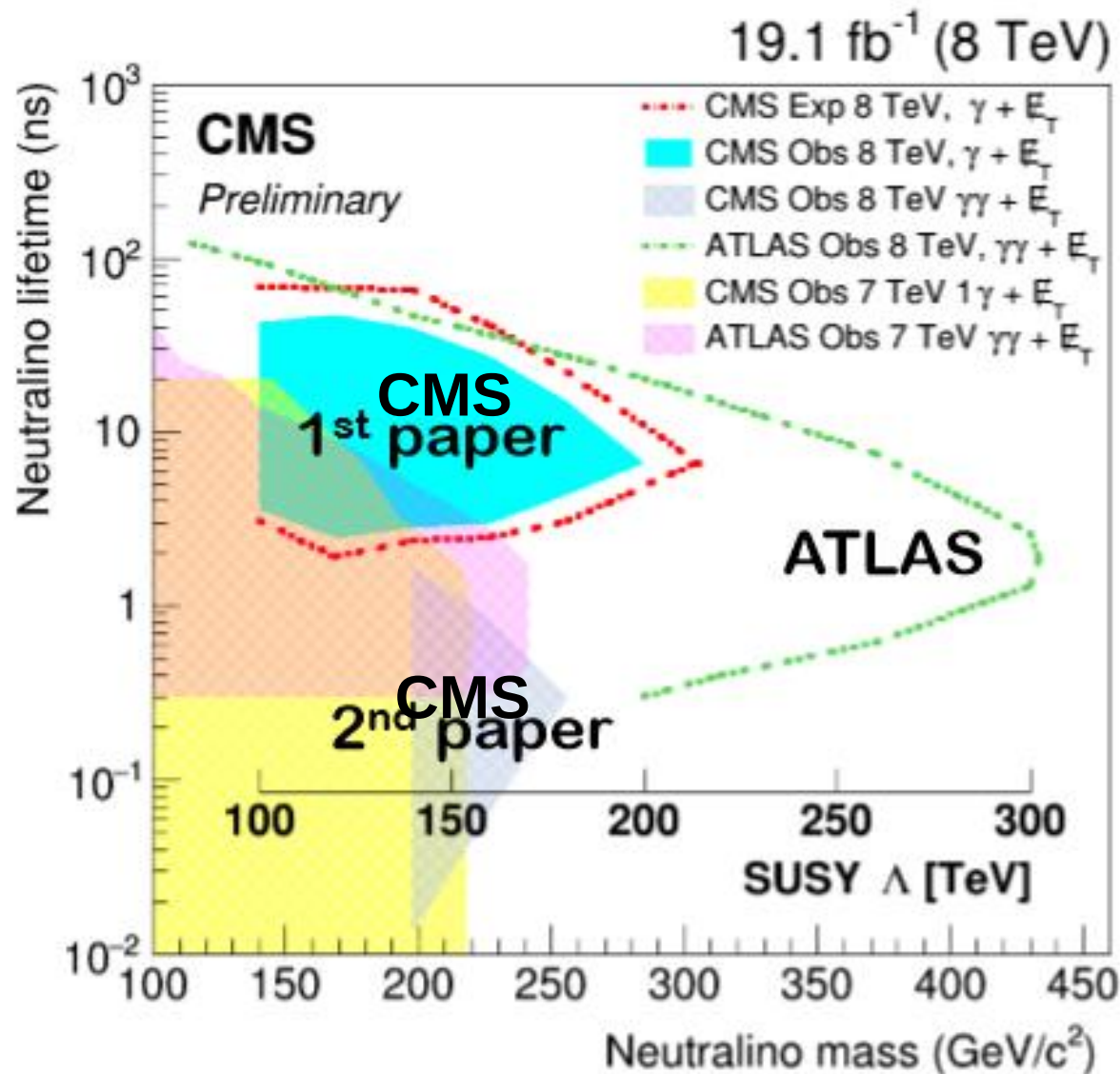
Take advantage of the longitudinal segmentation of the ATLAS calorimeter to verify if the shower is compatible from a photon originating from the IP.



Technique based on vertexing

Take advantage of the large amount of material in the tracker converting photons. Electron direction is used to show that photons are not originating from the IP.





- **CMS 1st paper**
CMS-PAS-EXO-12-035
purely based on timing

- **CMS 2nd paper**
CMS-PAS-EXO-14-017
purely based on conversion

- **ATLAS paper**
[arXiv:1409.5542](https://arxiv.org/abs/1409.5542)
based on timing and direction

- Introduction to LL physics (Why ? How ?) ✓
- Searches for LL particles decaying OUTSIDE the detector
 - Neutral ✓
 - Charged ✓
- Searches for LL particles decaying INSIDE the detector
 - Charged ✓
 - Neutral ✓
- Super exotic particles ←
- Summary

- **Magnetic monopoles (from Dirac and others)**
 - Theoretically interesting → quantization of the electric charge:
 - Ionization energy loss:

For electric charge: $\left\langle -\frac{dE}{dx} \right\rangle = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[\frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 W_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$

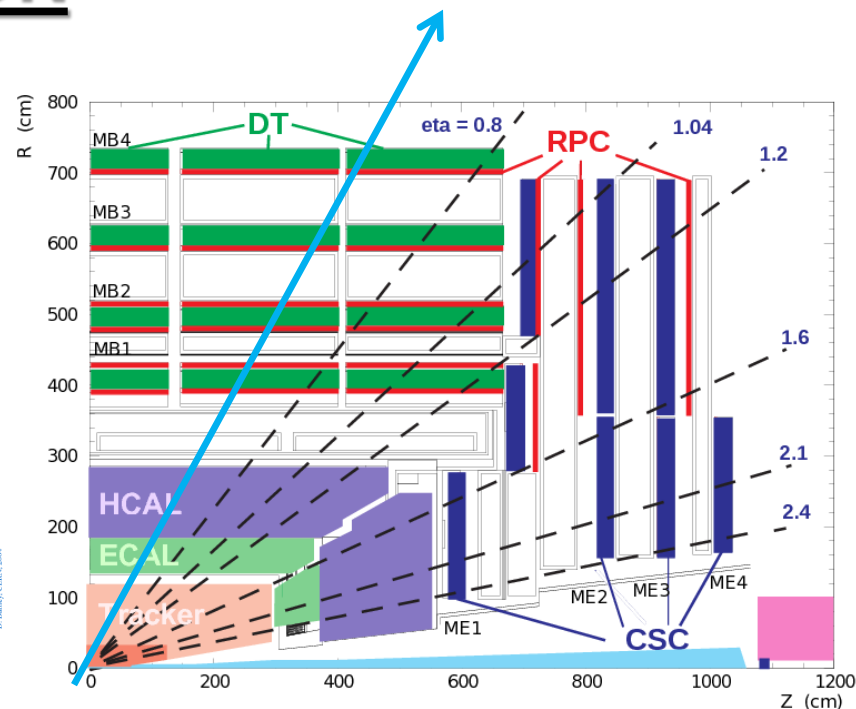
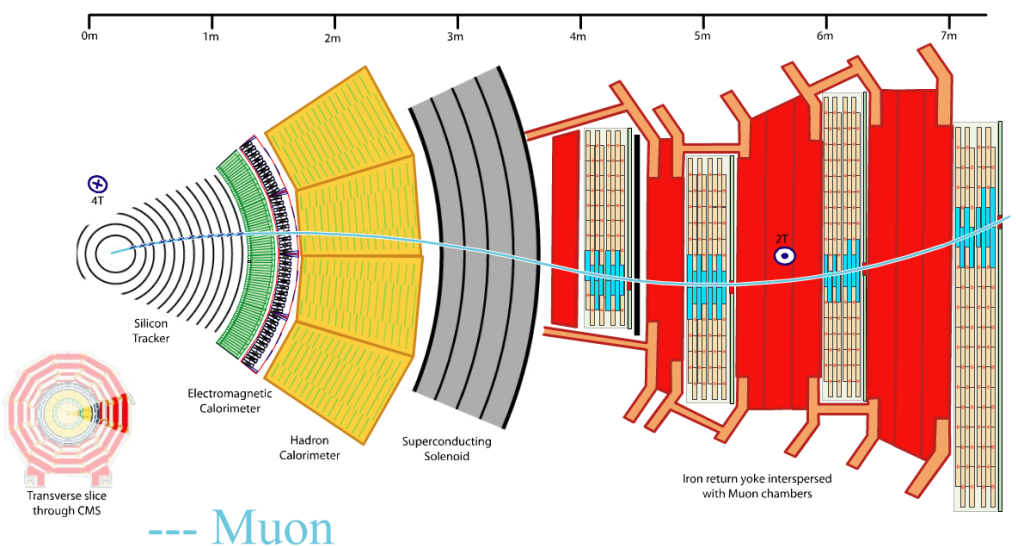
For magnetic charge: $\left\langle -\frac{dE}{dx} \right\rangle = (ng/e)^2 \frac{KZ}{A} \left[\frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 W_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$

- We replace ze with $ng\beta$: , dE/dx of a $n_D=1$ monopole is equivalent to a particle with charge $Q=68.5e$
 - → dE/dx 4500 higher than a proton
- **Dions** are particles which carry BOTH an electric and a magnetic charge (their dE/dx is therefore even larger)

• Trajectory:

- Due to the Lorentz force:
 - Electric particles $\rightarrow$ bending $\rightarrow$ perpendicular to B
 - Magnetic particles $\rightarrow$ bending $\rightarrow$ parallel to B
- Specific tracking might therefore be needed (depending on particle pT)

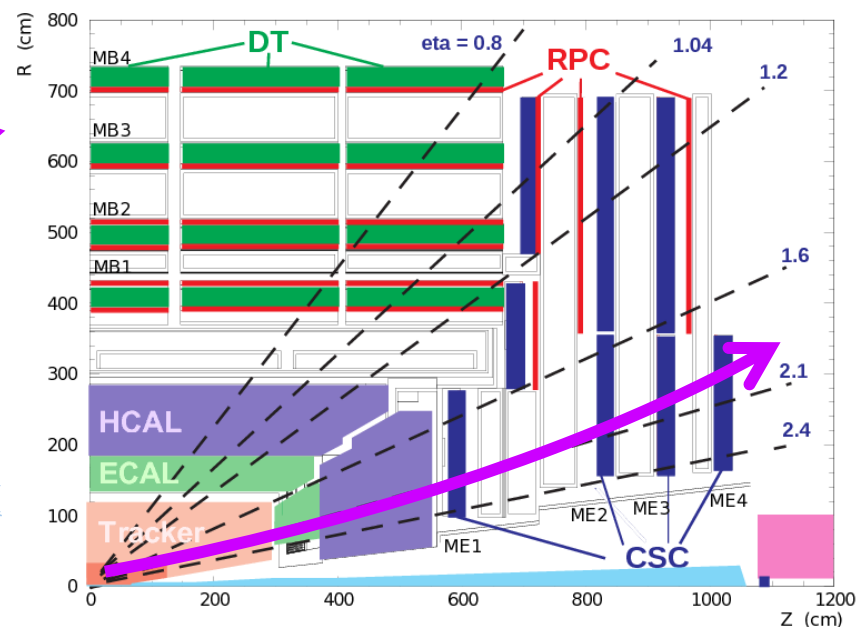
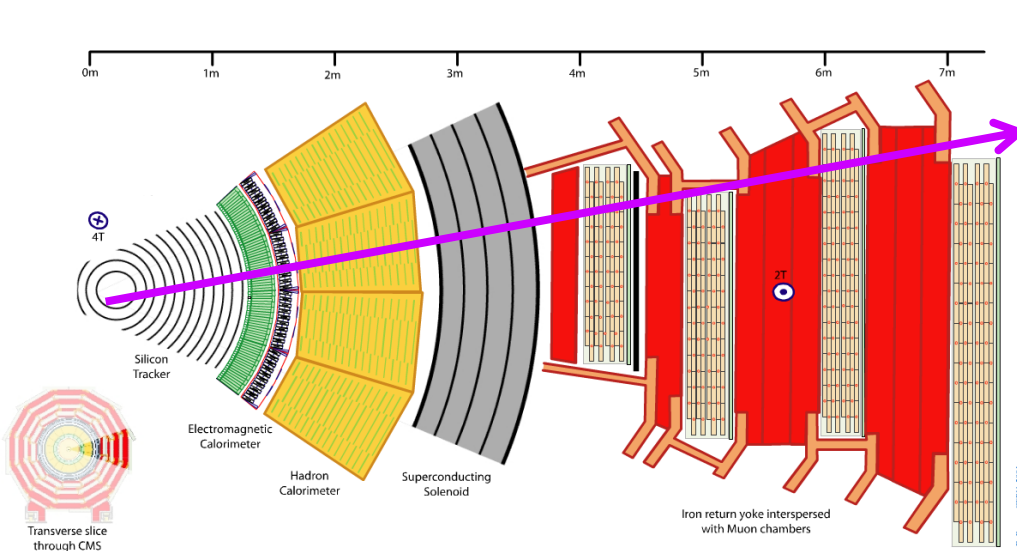
MUON



- **Trajectory:**

- Due to the Lorentz force:
 - Electric particles $\rightarrow$ bending $\rightarrow$ perpendicular to B
 - Magnetic particles $\rightarrow$ bending $\rightarrow$ parallel to B
- Specific tracking might therefore be needed (depending on particle pT)

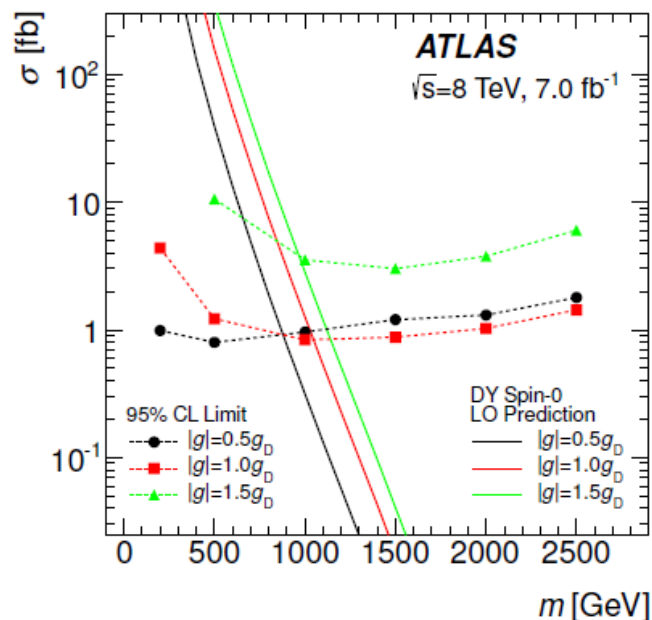
Magnetic MONOPOLE



ATLAS

Search similar to
HSCP with $Q \gg 1$
Trigger on ECAL energy

	$ g = 0.5g_D$	$ g = 1.0g_D$	$ g = 1.5g_D$
spin-1/2	1180	1340	1210
spin-0	890	1050	970

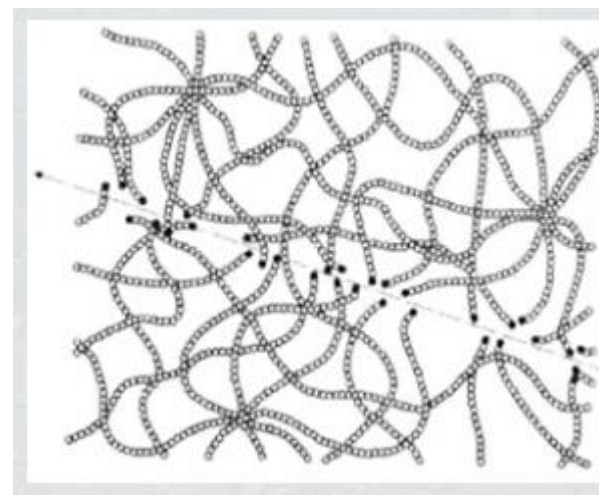
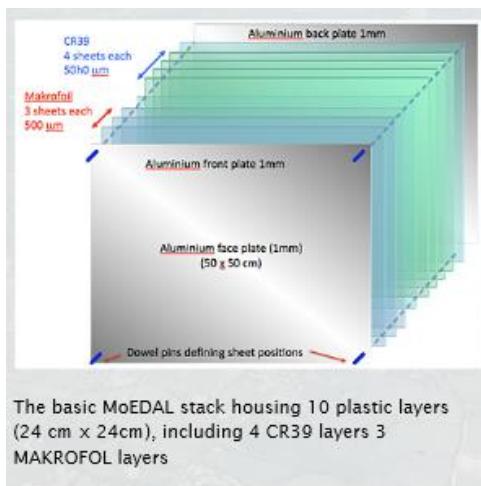


MoEDAL

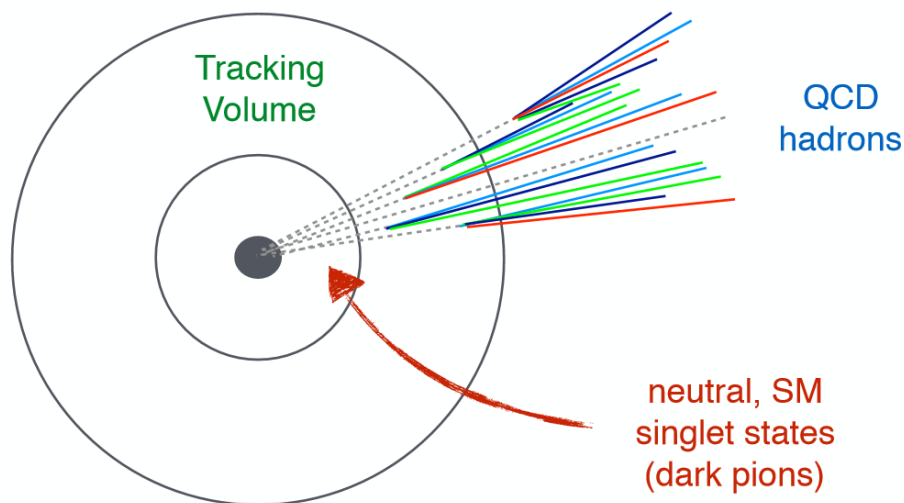
Monopole & Exotics Detector at the LHC



Search based on plastic
nuclear track detector

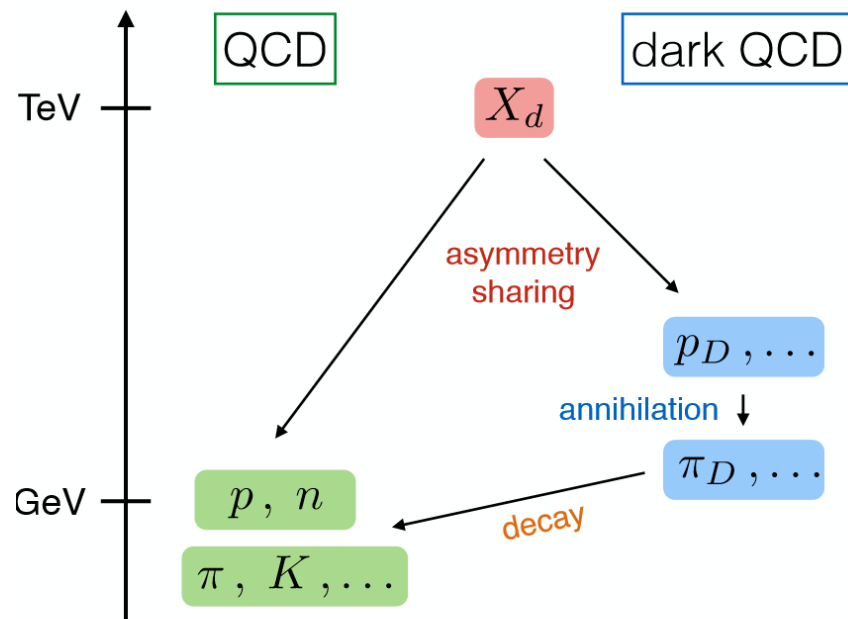


Possible origin for emerging Jets:
Dark QCD: hidden sector confinement (SU(N))

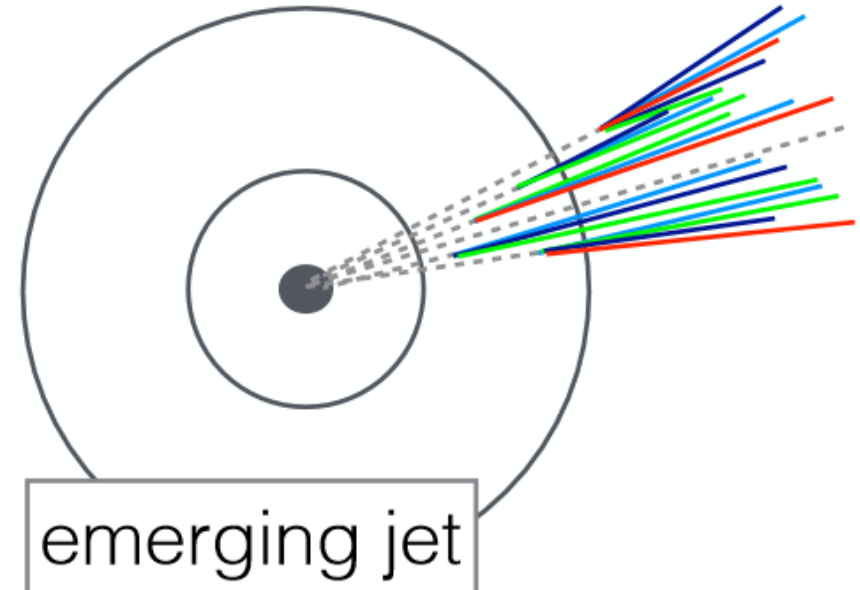
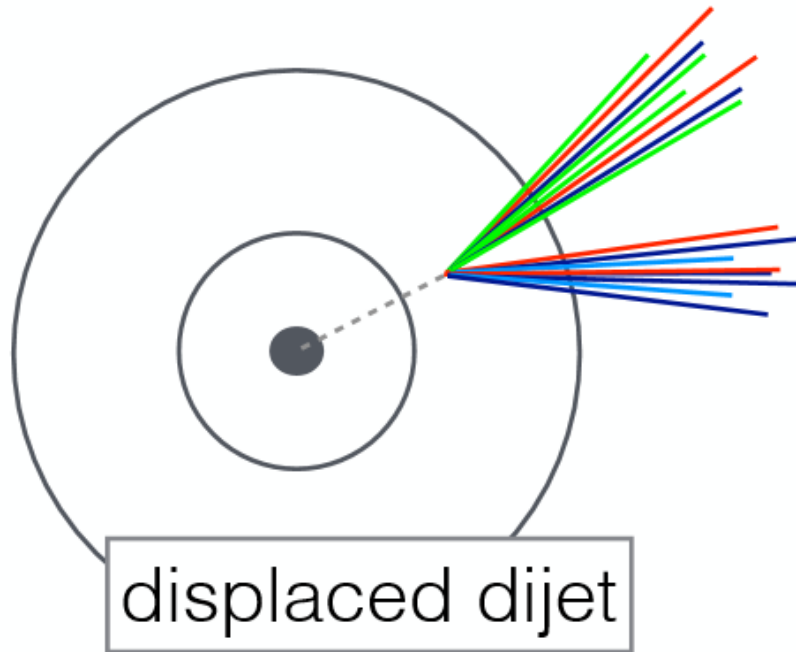


No specific search for emerging jets at the LHC, yet.

Slide inspired by
Pedro Schwaller
talk 11/13/15



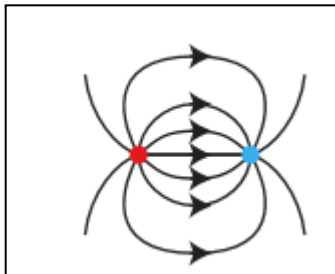
- SU(N) dark sector with neutral “dark quarks”
- DM is composite “dark proton”
- Confinement scale Λ_{darkQCD}
- “Dark pions” unstable, long lived



Displaced jet search have however some similarities
and results recasting is possible in some cases

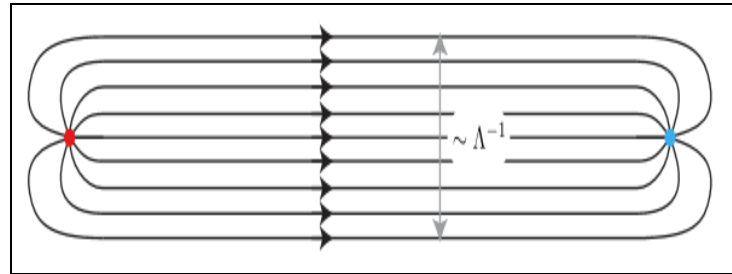
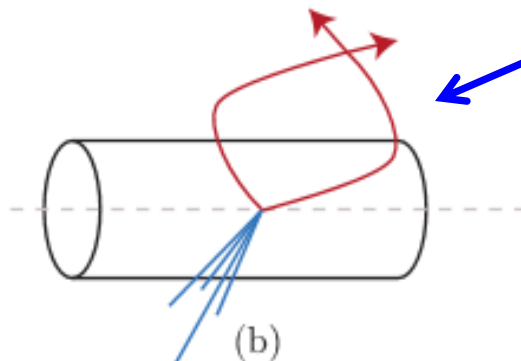
New gauge interaction (similar to QCD) with *infracolor* charge and a confinement scale that might be macroscopic.

Particle **phenomenology depends on string length** (confinement scale)...



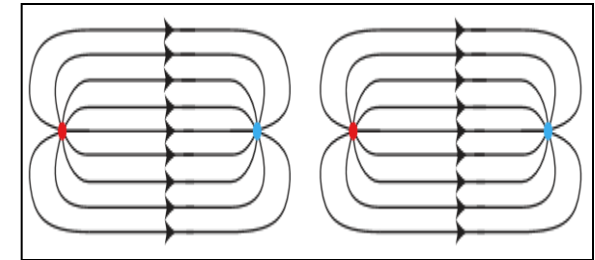
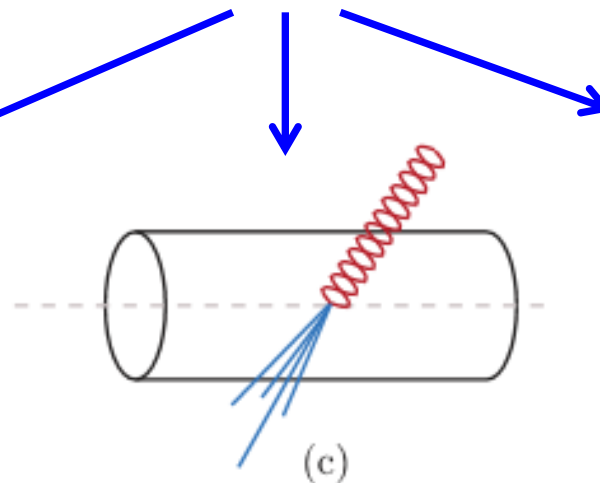
$$r \ll \Lambda^{-1}$$

Visible as
a single particle
with large dE/dx



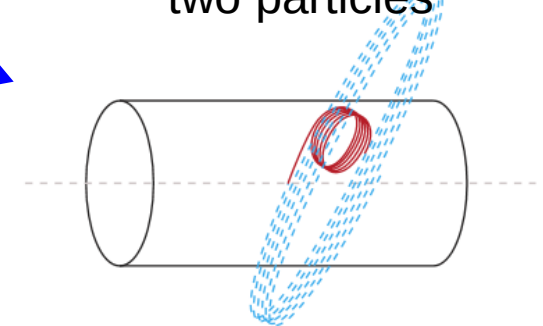
$$r \gg \Lambda^{-1}$$

Visible as two connected particles



String breaking $2m_Q \gg \Lambda$

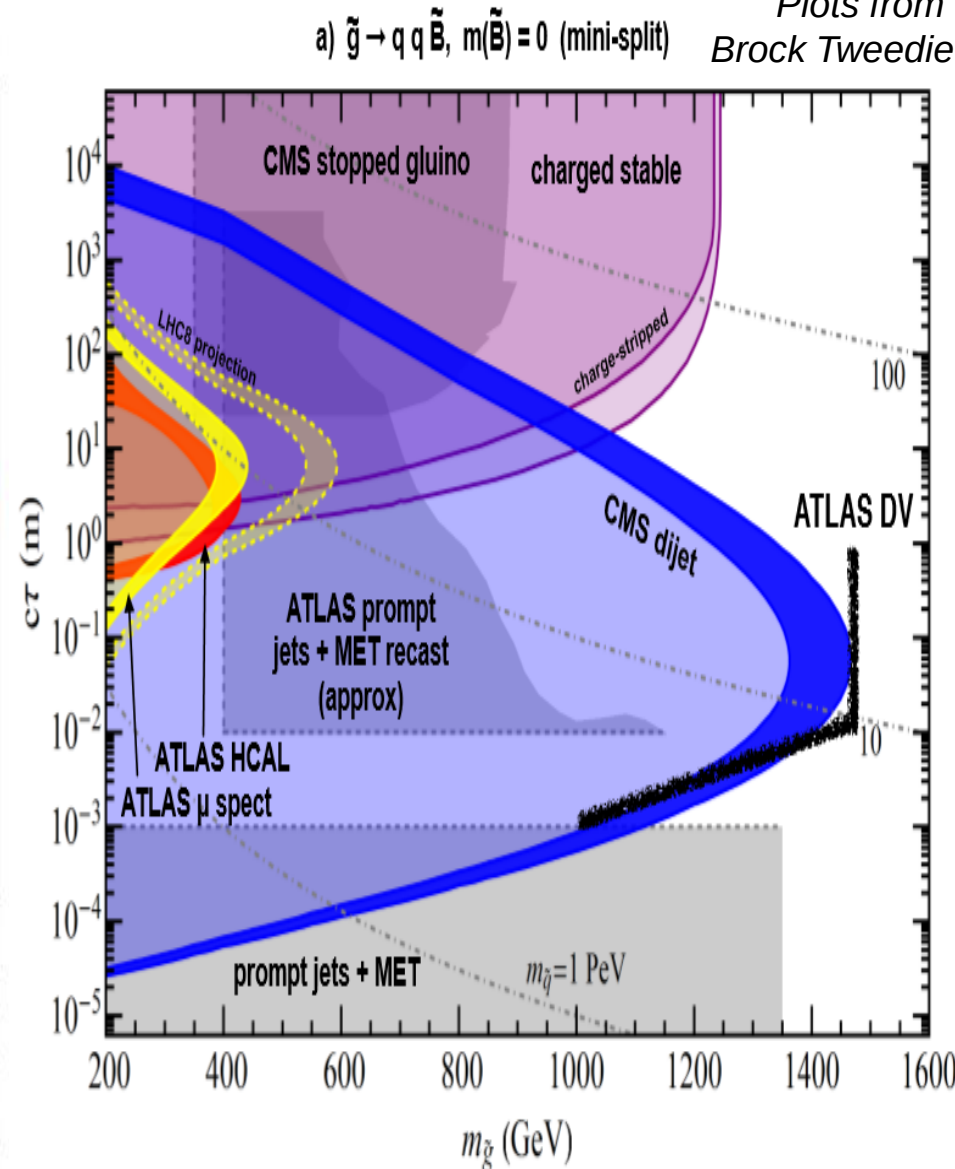
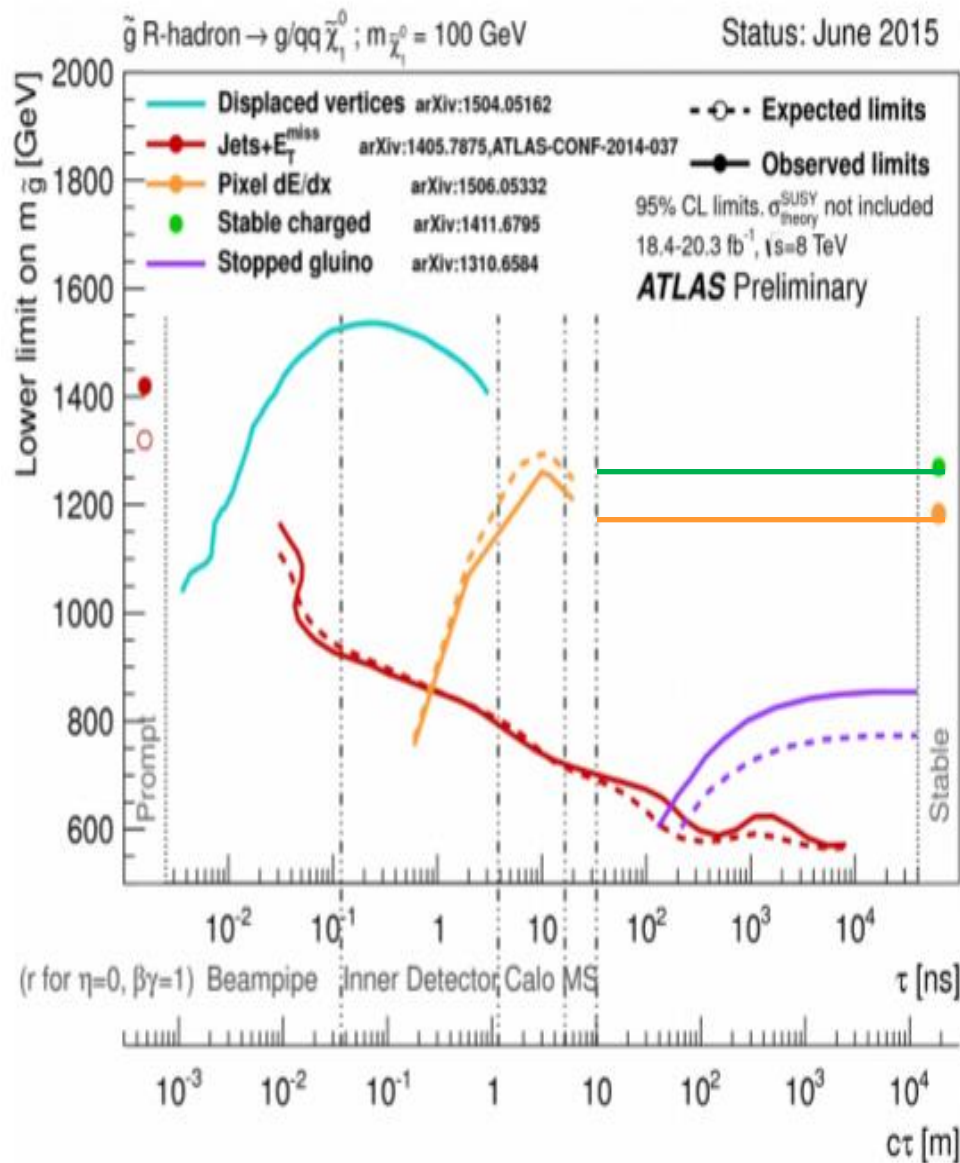
Visible as
two particles



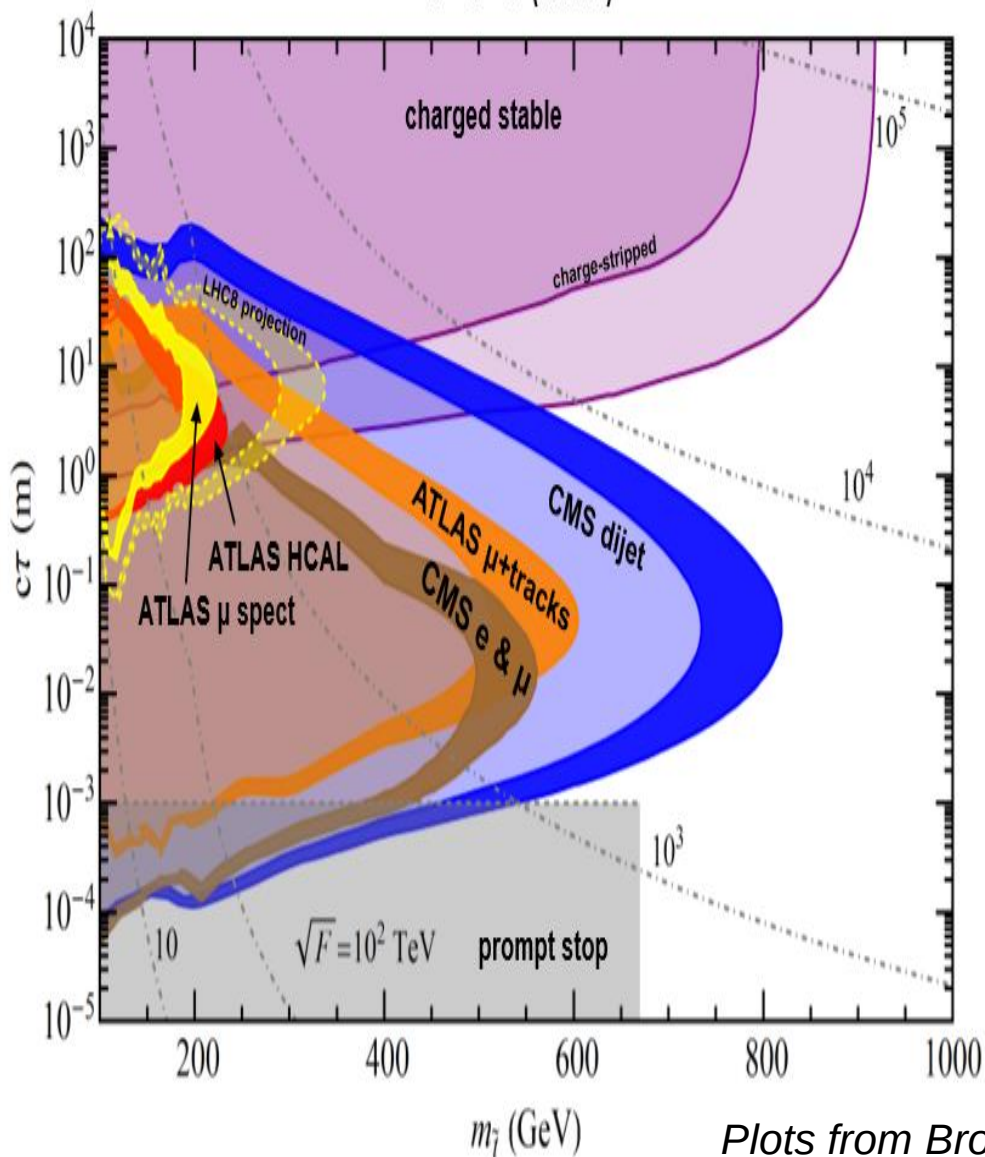
**No specific search for
quirks at the LHC, yet.**



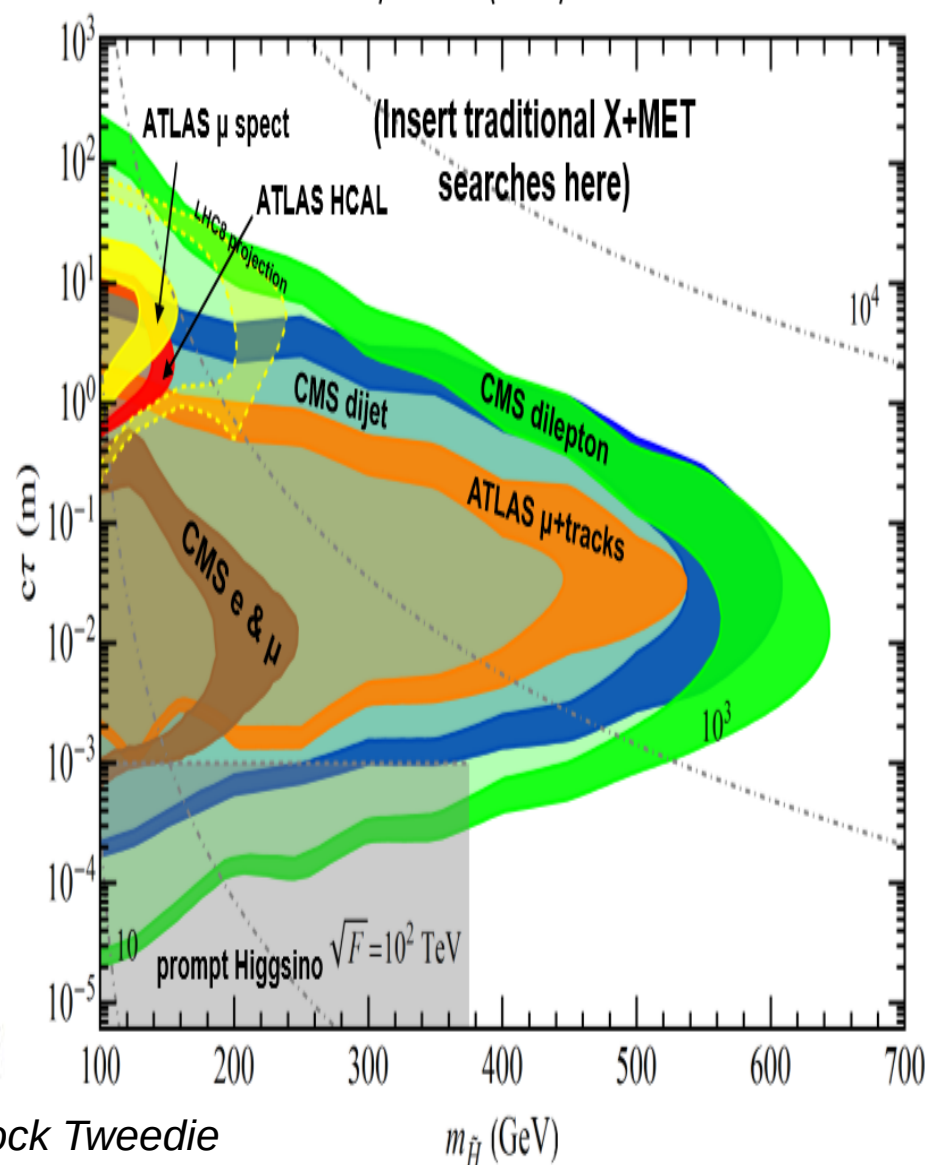
Summary (overview of LL SUSY constraints)



$\tilde{t} \rightarrow t^{(*)} \tilde{G}$ (GMSB)



a) $\tilde{H} \rightarrow Z \tilde{G}$ (GMSB)



Plots from Brock Tweedie

- Searches for LL particles are usually based on simple signatures that are sensitive to a wide range of models.
 - Massive, stable charged particles.
 - Displaced leptons, jets or photons.
 - Disappearing / Kinked tracks
 - Asynchronous (stopped) particle decay
- Searches / results are generally model independent as there are many models predicting LL particles and only a few people looking for them.
- Searches are generally quite challenging because of our detector/software being optimized for « prompt » decay.
 - Still, we set strong limits on (LL) SUSY and other BSM models...
- A diphoton excess at 750 GeV is not incompatible with long-lived particles
 - $\rightarrow \pi^0$ and π^+/π^- analogy





Backups

$-3 \leq$	M_1, M_2	$\leq 3 \text{ TeV}$
$0 \leq$	M_3	$\leq 3 \text{ TeV}$
$-3 \leq$	μ	$\leq 3 \text{ TeV}$
$0 \leq$	m_A	$\leq 3 \text{ TeV}$
$2 \leq$	$\tan \beta$	≤ 60
$0 \leq$	$\tilde{Q}_{1,2}, \tilde{Q}_3$	$\leq 3 \text{ TeV}$
$0 \leq$	$\tilde{U}_{1,2}, \tilde{U}_3$	$\leq 3 \text{ TeV}$
$0 \leq$	$\tilde{D}_{1,2}, \tilde{D}_3$	$\leq 3 \text{ TeV}$
$0 \leq$	$\tilde{L}_{1,2}, \tilde{L}_3$	$\leq 3 \text{ TeV}$
$0 \leq$	$\tilde{E}_{1,2}, \tilde{E}_3$	$\leq 3 \text{ TeV}$
$-7 \leq$	A_t, A_b, A_τ	$\leq 7 \text{ TeV}$

Exotic Objects ($pp \rightarrow XX$, $X \rightarrow o_1 o_2$ or $X \rightarrow o_1 o_2 o_3$)

$o_2 o_1$	j	b	γ	e	μ	τ	E_T
j	jj/dv	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv/em	jj/dv
b	X	jj/dv	jj/dv/dA	jj/dv	jj/dv	jj/dv/em	jj/dv
γ	X	X	dA	dA	dA	dA/em	dA
e	X	X	X	dv/	dv	dv/	em
μ	X	X	X	X	dv/	dv/	
τ	X	X	X	X	X	dv/ /em	

Search	arXiv	Symbol	Comments
CMS Displaced Dijets	1411.6530	jj	$m_X > 50$ GeV
ATLAS Displaced Vertex	1504.05162	dv	$m_X > 10$ GeV
CMS Displaced Dilepton	1411.6977		$m_X > 15$ GeV
ATLAS Delayed Photon	1409.5542	dA	$m_X \gtrsim 100$ GeV
CMS Displaced $e\mu$	1409.4789	em	Best at LFU