

Falsifying Baryogenesis Mechanisms through Observation of Lepton Number and Flavour Violation

Julia Harz

ILP - LPTHE - UPMC - CNRS Paris

in collaboration with

F. Deppisch, L. Graf, W. Huang, M. Hirsch, H. Päs

27/09/2016

DESY Theory Workshop

F. Deppisch, JH, M. Hirsch, PRL 112 (2014) 221601

F. Deppisch, JH, W. Huang, M. Hirsch, H. Päs, Phys. Rev. D92 (2015) 036005

F. Deppisch, L. Graf, JH, W. Huang, work in progress

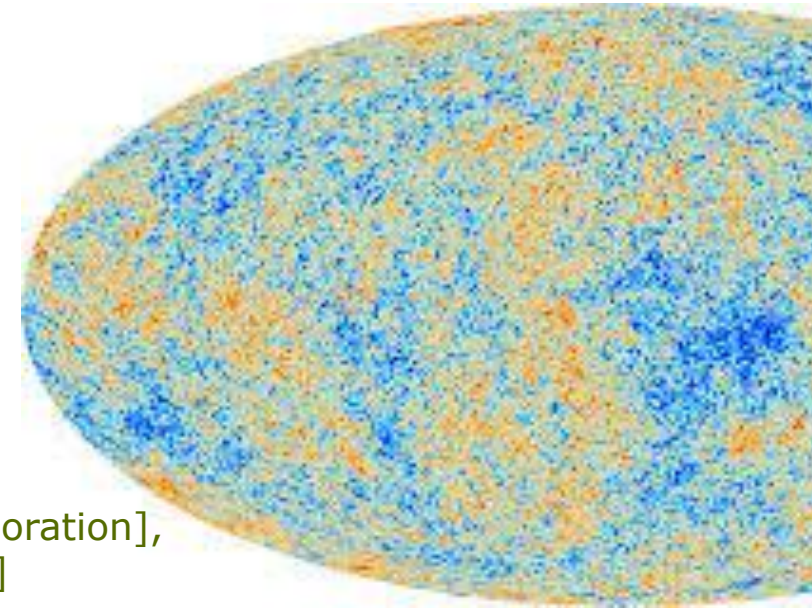


Baryon Asymmetry

- observation of a baryon asymmetry of the Universe (BAU)

$$\eta_B^{\text{obs}} = \frac{n_B - n_{\bar{B}}}{n_\gamma} = (6.09 \pm 0.06) \times 10^{-10}$$

P. A. R. Ade et al. [Planck Collaboration],
arXiv:1502.01589 [astro-ph.CO]



- theoretical requirements for generating a baryon asymmetry: 3 Sakharov conditions

A. D. Sakharov, JETP Lett. 5, 24 (1967)

- C and CP violation
- departure from thermal equilibrium
- $(B-L)$ -violation



not fully fulfilled within the Standard Model



physics beyond the Standard Model

- popular scenarios for explaining baryon asymmetry:
 - electroweak baryogenesis, leptogenesis, etc. ...

How can we shed light on the underlying mechanism that generated the baryon asymmetry with current experiments?

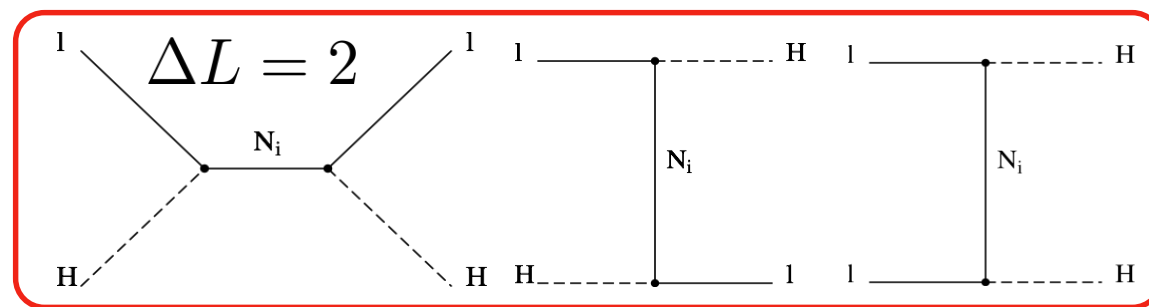
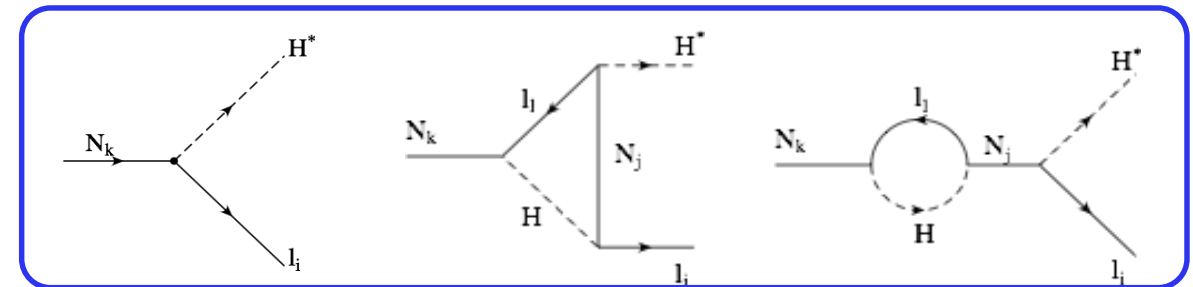
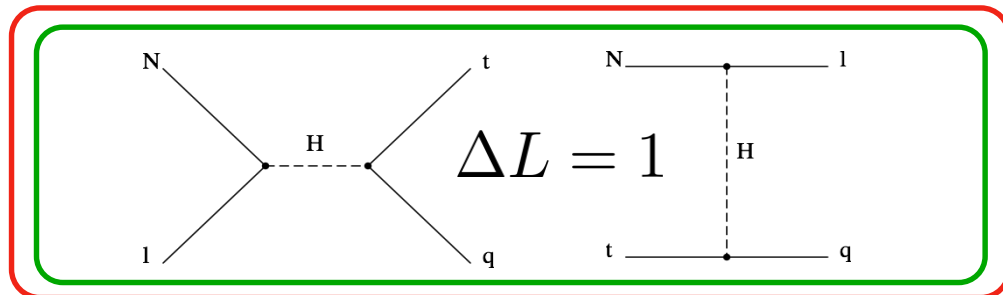
Example: Leptogenesis

- generation of lepton asymmetry via **heavy neutrino decays**
- competition with lepton number violating (LNV) **washout processes**
- conversion to baryon asymmetry via **sphaleron processes**

$$Hz \frac{dN_{N_1}}{dz} = -(\Gamma_D + \Gamma_S)(N_{N_1} - N_{N_1}^{\text{eq}})$$

$$Hz \frac{dN_L}{dz} = \epsilon_1 \Gamma_D (N_{N_1} - N_{N_1}^{\text{eq}}) - \Gamma_W N_L$$

source of CP-asymmetry



washout processes

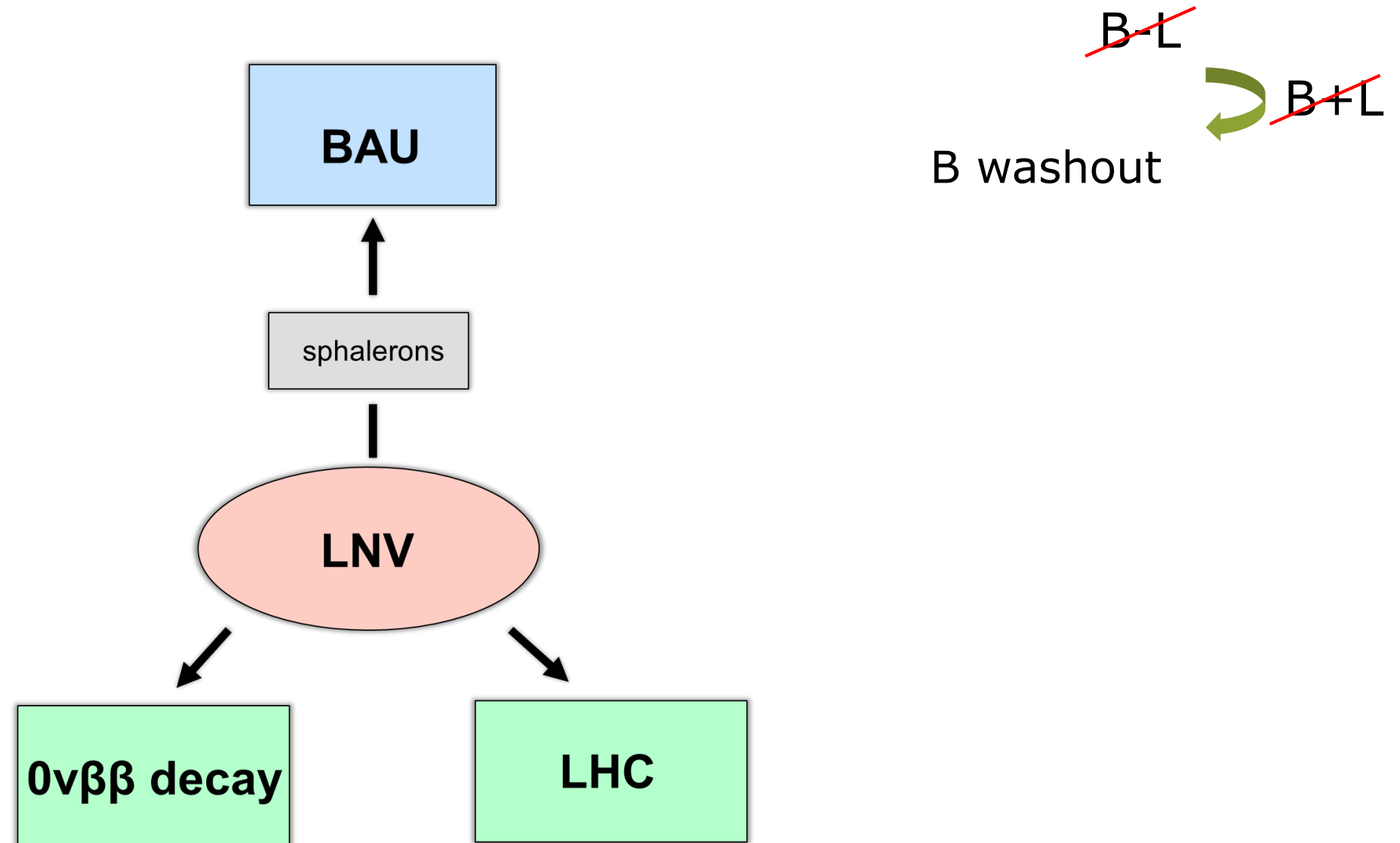
washout highly efficient if: $\frac{\Gamma_W}{H} > 1$

B asymmetry

~~B-L~~ \rightarrow ~~B+L~~

Lepton Number Violation (LNV)

- experimental observation of LNV implies existence of washout processes in the Early Universe
- due to sphaleron processes this allows for a measure of the corresponding baryon asymmetry washout



observation of low energy LNV will have far-reaching consequences on mechanisms of baryogenesis

Neutrinoless Double Beta Decay (0vbb)

- $0\nu\beta\beta$ ($2n \rightarrow 2p + 2e^-$) is a sensitive probe of low energy LNV

- current experimental limits on the half life of $0\nu\beta\beta$:

$$T_{1/2}^{76\text{Ge}} > (1.1 - 1.9) \times 10^{25} \text{ y} \quad (\text{EXO-200, KamLAND-Zen})$$

$$T_{1/2}^{136\text{Xe}} > 2.1 \times 10^{25} \text{ y} \quad (\text{GERDA})$$

- general Lagrangian describing different **non-SM contributions to $0\nu\beta\beta$** can be written in terms of **effective couplings ϵ_α^β** , e.g. for the **long range contribution**:

$$\mathcal{L} = \frac{G_F}{\sqrt{2}} \{ j_{V-A}^\mu J_{V-A,\mu}^\dagger + \sum_{\alpha,\beta} \epsilon_\alpha^\beta j_\beta J_\alpha^\dagger \}$$

$$T_{1/2}^{-1} = |\epsilon_\alpha^\beta|^2 G_i |M_i|^2$$

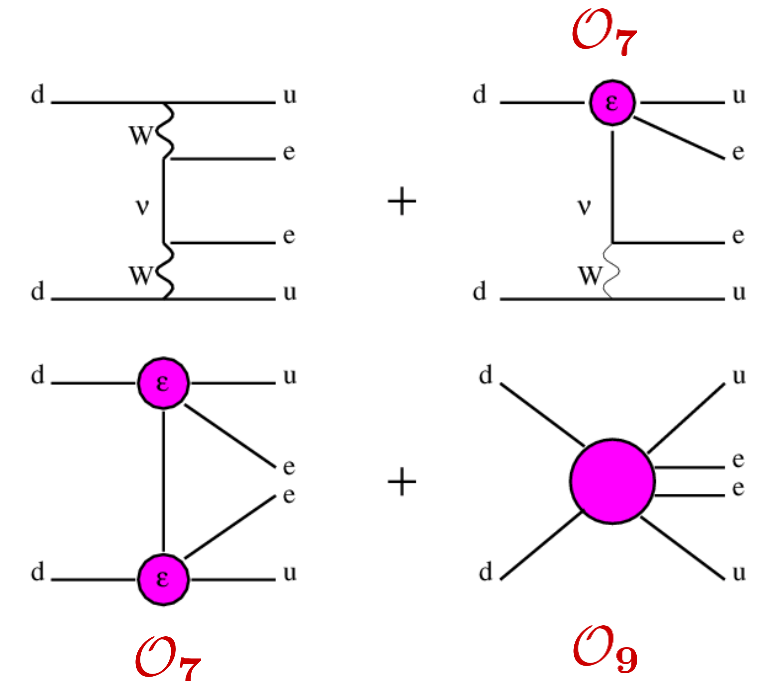
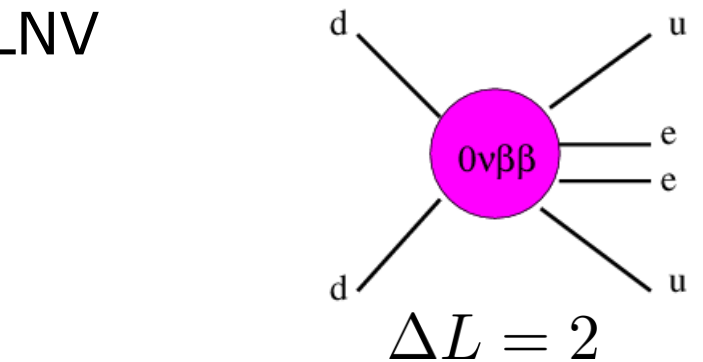
$$j_\beta = \bar{e} \mathcal{O}_\beta \nu$$

$$J_\alpha^\dagger = \bar{u} \mathcal{O}_\alpha d$$

$$\mathcal{O}_{V\pm A} = \gamma^\mu (1 \pm \gamma_5)$$

$$\mathcal{O}_{S\pm P} = (1 \pm \gamma_5)$$

$$\mathcal{O}_{TR,L} = \frac{i}{2} [\gamma_\mu, \gamma_\nu] (1 \pm \gamma_5)$$



Isotope	$ \epsilon_{V-A}^{V+A} $	$ \epsilon_{V+A}^{V+A} $	$ \epsilon_{S-P}^{S+P} $	$ \epsilon_{S+P}^{S+P} $	$ \epsilon_{TL}^{TR} $	$ \epsilon_{TR}^{TR} $
^{76}Ge	$3.3 \cdot 10^{-9}$	$5.9 \cdot 10^{-7}$	$1.0 \cdot 10^{-8}$	$1.0 \cdot 10^{-8}$	$6.4 \cdot 10^{-10}$	$1.0 \cdot 10^{-9}$
^{136}Xe	$2.6 \cdot 10^{-9}$	$5.1 \cdot 10^{-7}$	$6.2 \cdot 10^{-9}$	$6.2 \cdot 10^{-9}$	$4.4 \cdot 10^{-10}$	$7.4 \cdot 10^{-10}$

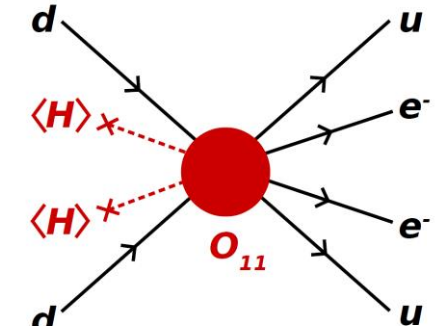
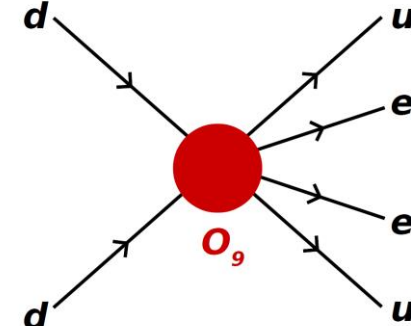
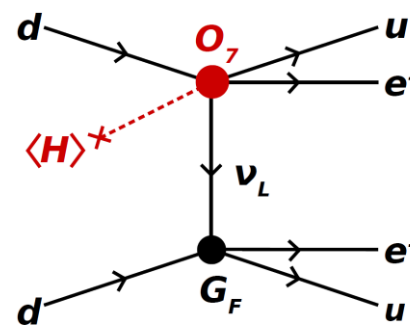
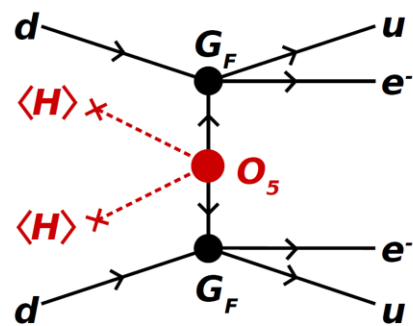
F. Deppisch, M. Hirsch, H. Päs, J. Phys. G 39 (2012) 124007, arXiv:1208.0727 [hep-ph], updated

$0\nu\beta\beta$ half life sets constraints on effective couplings ϵ_α^β

Possible underlying LNV Operators

- four examples from the complete list of all possible LNV $\Delta L = 2$ effective operators

K. S. Babu, C. N. Leung, Nucl. Phys. B 619 (2001), arxiv:0106054 [hep-ph]
A. de Gouvea, J. Jenkins, PRD 77 (2008), arXiv:0708.1344 [hep-ph]



$$\mathcal{O}_5 = (L^i L^j) H^k H^l \epsilon_{ik} \epsilon_{jl}$$

$$\mathcal{O}_7 = (L^i d^c) (\bar{e}^c \bar{u}^c) H^j \epsilon_{ij}$$

$$\mathcal{O}_9 = (L^i L^j) (\bar{Q}_i \bar{u}^c) (\bar{Q}_j \bar{u}^c)$$

$$\mathcal{O}_{11} = (L^i L^j) (Q_k d^c) (Q_l d^c) H_m \bar{H}_i \epsilon_{jk} \epsilon_{lm}$$

If $0\nu\beta\beta$ was observed, the scale of the underlying operator can be determined

$$m_e \epsilon_5 = \frac{g^2 v^2}{\Lambda_5}$$

$$\frac{G_F \epsilon_7}{\sqrt{2}} = \frac{g^3 v}{2 \Lambda_7^3}$$

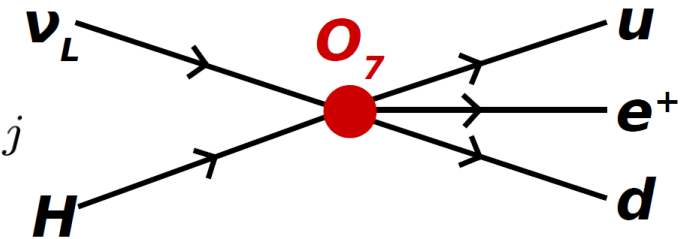
$$\frac{G_F^2 \epsilon_{\{9,11\}}}{2m_p} = \left\{ \frac{g^4}{\Lambda_9^5}, \frac{g^6 v^2}{\Lambda_{11}^7} \right\}$$

\mathcal{O}_D	Λ_D^0 [GeV]
\mathcal{O}_5	9.1×10^{13}
\mathcal{O}_7	2.6×10^4
\mathcal{O}_9	2.1×10^3
\mathcal{O}_{11}	1.0×10^3

Lepton Asymmetry Washout

- LNV operator would cause washout of pre-existing net lepton asymmetry in the Early Universe – we take now as an example \mathcal{O}_7

$$\mathcal{O}_7 = (L^i d^c)(\bar{e}^c \bar{u}^c) H^j \epsilon_{ij}$$



$$zHn_\gamma \frac{d\eta_N}{dz} = - \sum_{a,i,j,\dots} \left(\frac{n_N n_a \dots}{n_N^{\text{eq}} n_a^{\text{eq}} \dots} - \frac{n_i n_j \dots}{n_i^{\text{eq}} n_j^{\text{eq}} \dots} \right) \gamma^{\text{eq}} (Na \dots \leftrightarrow ij \dots)$$

$$n_\gamma H T \frac{d\eta_L}{dT} = c_D \frac{T^{2D-4}}{\Lambda_D^{2D-8}} \eta_L$$

$$\gamma^{\text{eq}} \propto \frac{T^{2D-4}}{\Lambda_D^{2D-8}}$$

c_D operator specific factor

η_L lepton density

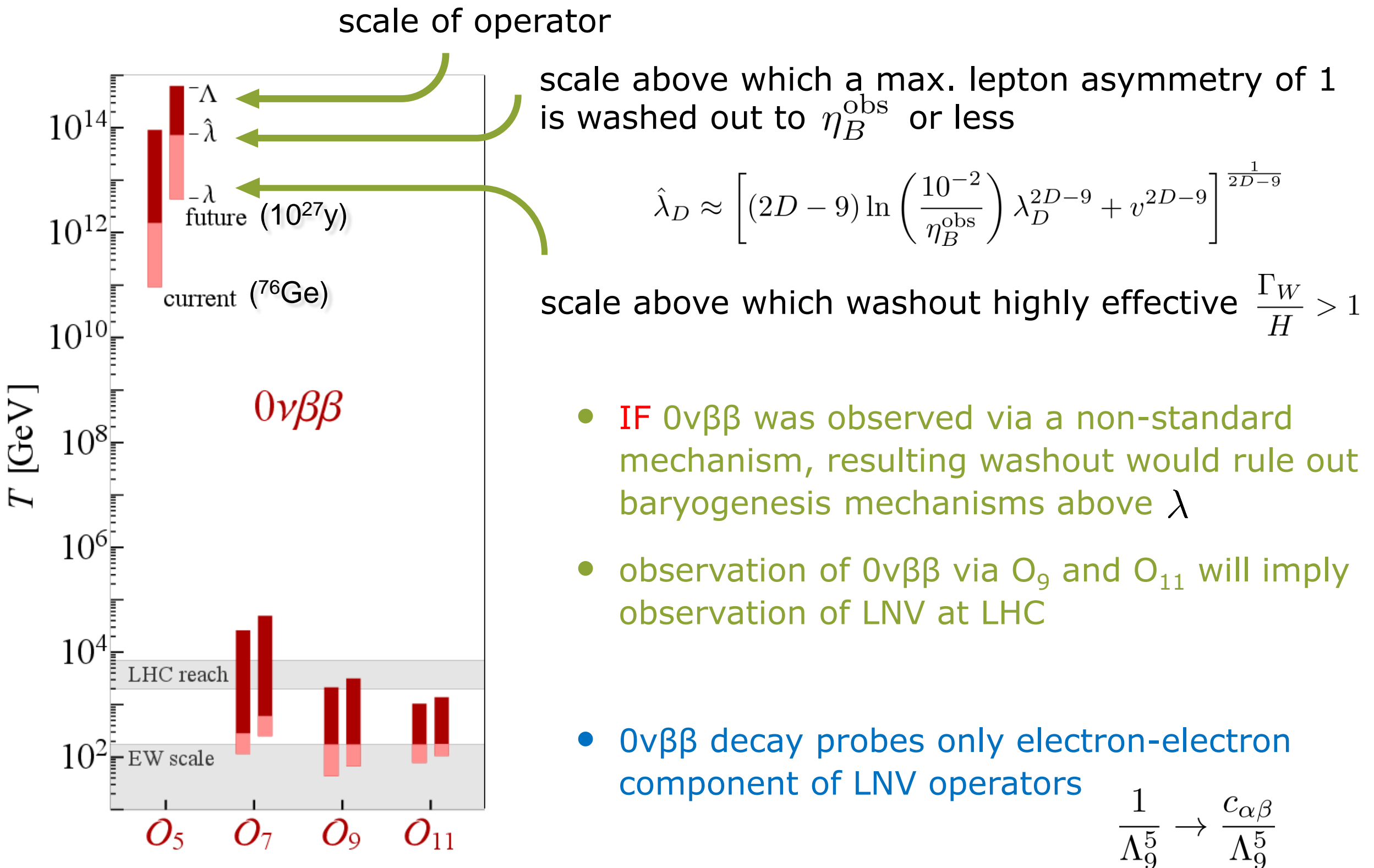
- washout effective if

$$\frac{\Gamma_W}{H} \equiv \frac{c_D}{n_\gamma H} \frac{T^{2D-4}}{\Lambda_D^{2D-8}} = c'_D \frac{\Lambda_{\text{Pl}}}{\Lambda_D} \left(\frac{T}{\Lambda_D} \right)^{2D-9} > 1$$

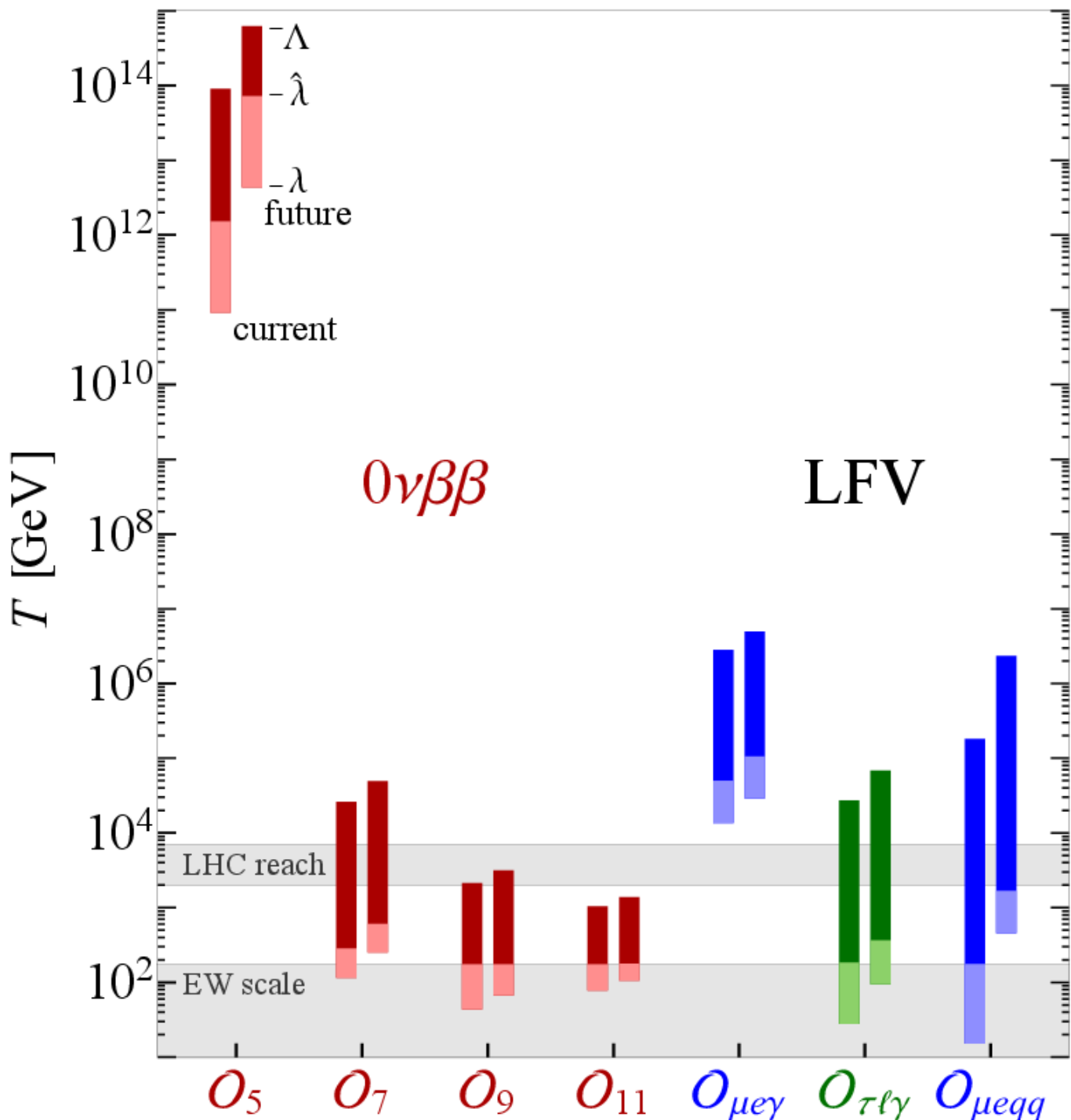
- **If** $0\nu\beta\beta$ is observed, washout effective in the temperature interval

$$\Lambda_D \left(\frac{\Lambda_D}{c'_D \Lambda_{\text{Pl}}} \right)^{\frac{1}{2D-9}} \equiv \lambda_D < T < \Lambda_D$$

Results



Considering Lepton Flavour Violation (LFV)



- Most stringent limits on LFV set by 6-dim $\Delta L = 0$ operators

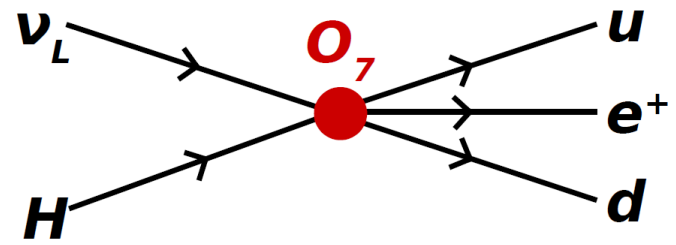
$$\mathcal{O}_{\ell\ell\gamma} = \mathcal{C}_{\ell\ell\gamma} \bar{L}_\ell \sigma^{\mu\nu} \bar{\ell}^c H F_{\mu\nu}$$

$$\mathcal{O}_{\ell\ell qq} = \mathcal{C}_{\ell\ell qq} (\bar{\ell} \Pi_1 \ell) (\bar{q} \Pi_2 q)$$

$$\mathcal{C}_{\ell\ell qq} = \frac{g^2}{\Lambda_{\ell\ell qq}^2} \quad \mathcal{C}_{\ell\ell\gamma} = \frac{eg^3}{16\pi^2 \Lambda_{\ell\ell\gamma}^2}$$

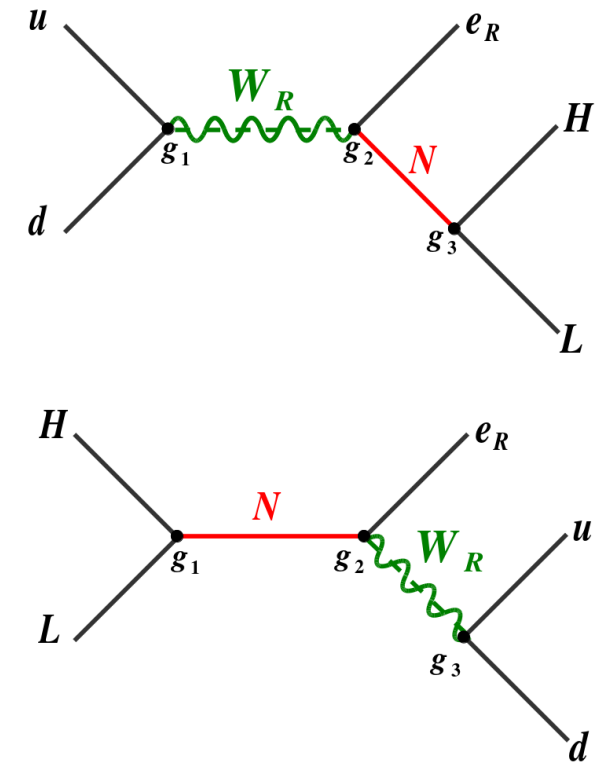
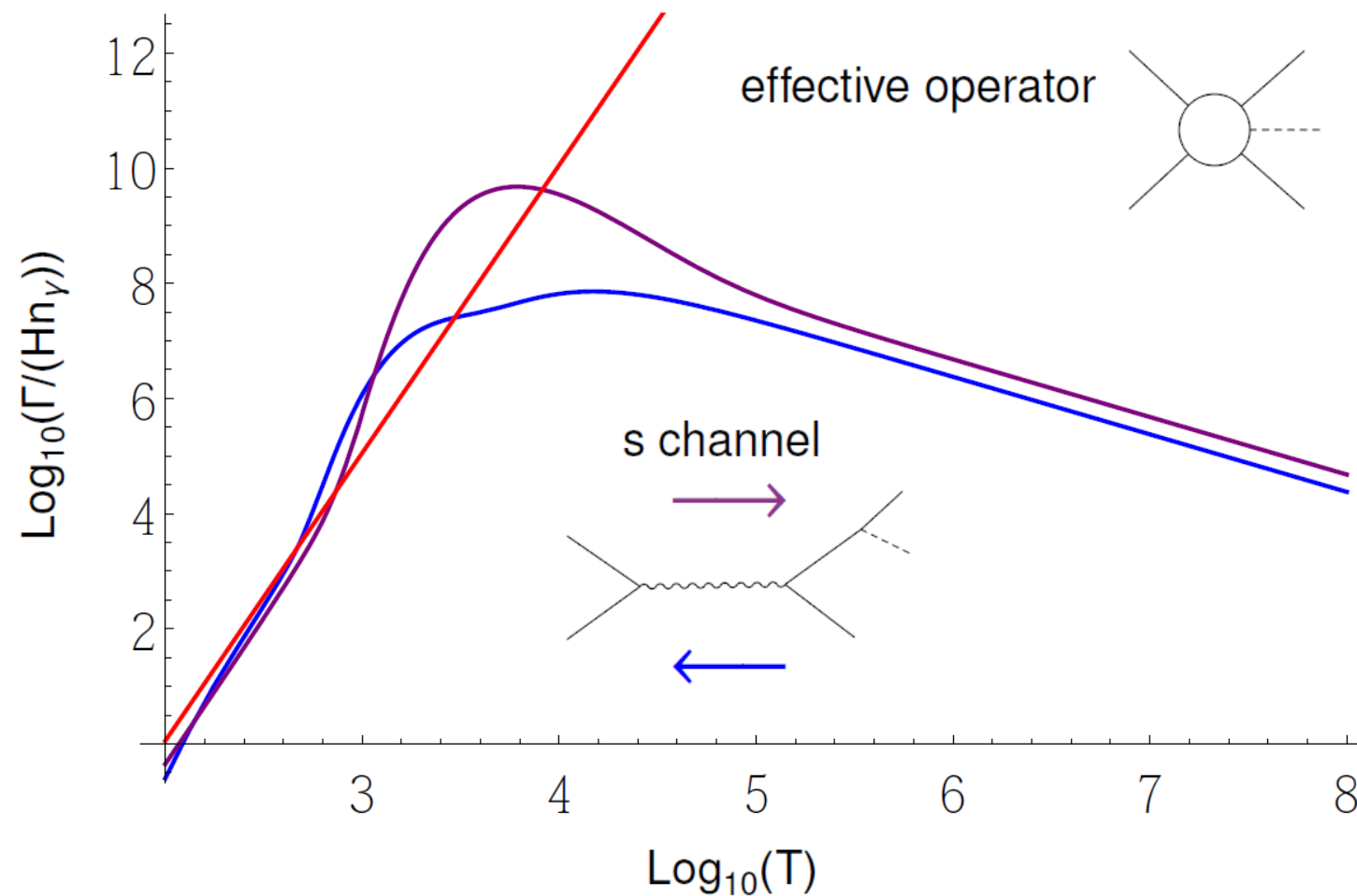
- determine temperature interval in which LFV process equilibrate pre-existing flavour asymmetry
- IF LFV processes are observed as well, loophole of asymmetry being stored in another flavour sector is ruled out

Comparison with UV completed Model



$$\mathcal{O}_7 = (L^i d^c)(\bar{e}^c \bar{u}^c) H^j \epsilon_{ij}$$

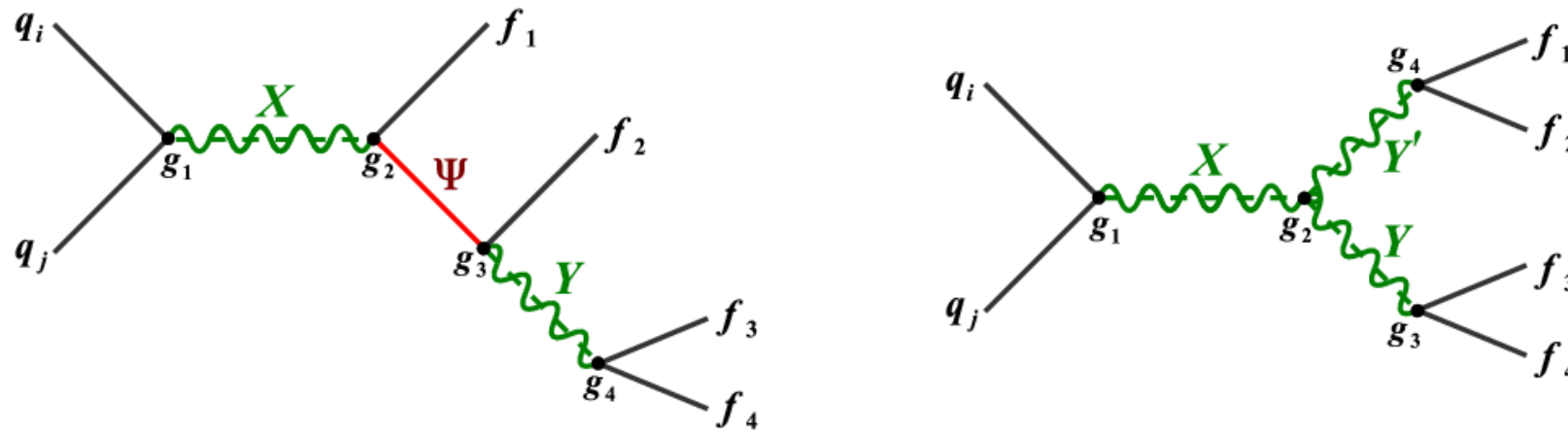
$$\Lambda_7^6 = m_{W_R}^4 m_N^2$$



effective operator approach is a conservative estimation of washout rate

LVN at the LHC

- **Signature:** $\Delta L = 2$ LVN at LHC through resonant process $pp \rightarrow l^\pm l^\pm + 2 \text{ jets}$ with two same-sign leptons and two jets without missing energy



$$\frac{\Gamma_W}{H} = \frac{1}{n_\gamma H} \frac{T}{32\pi^4} \int_0^\infty ds \, s^{3/2} \sigma(s) K_1 \left(\frac{\sqrt{s}}{T} \right)$$

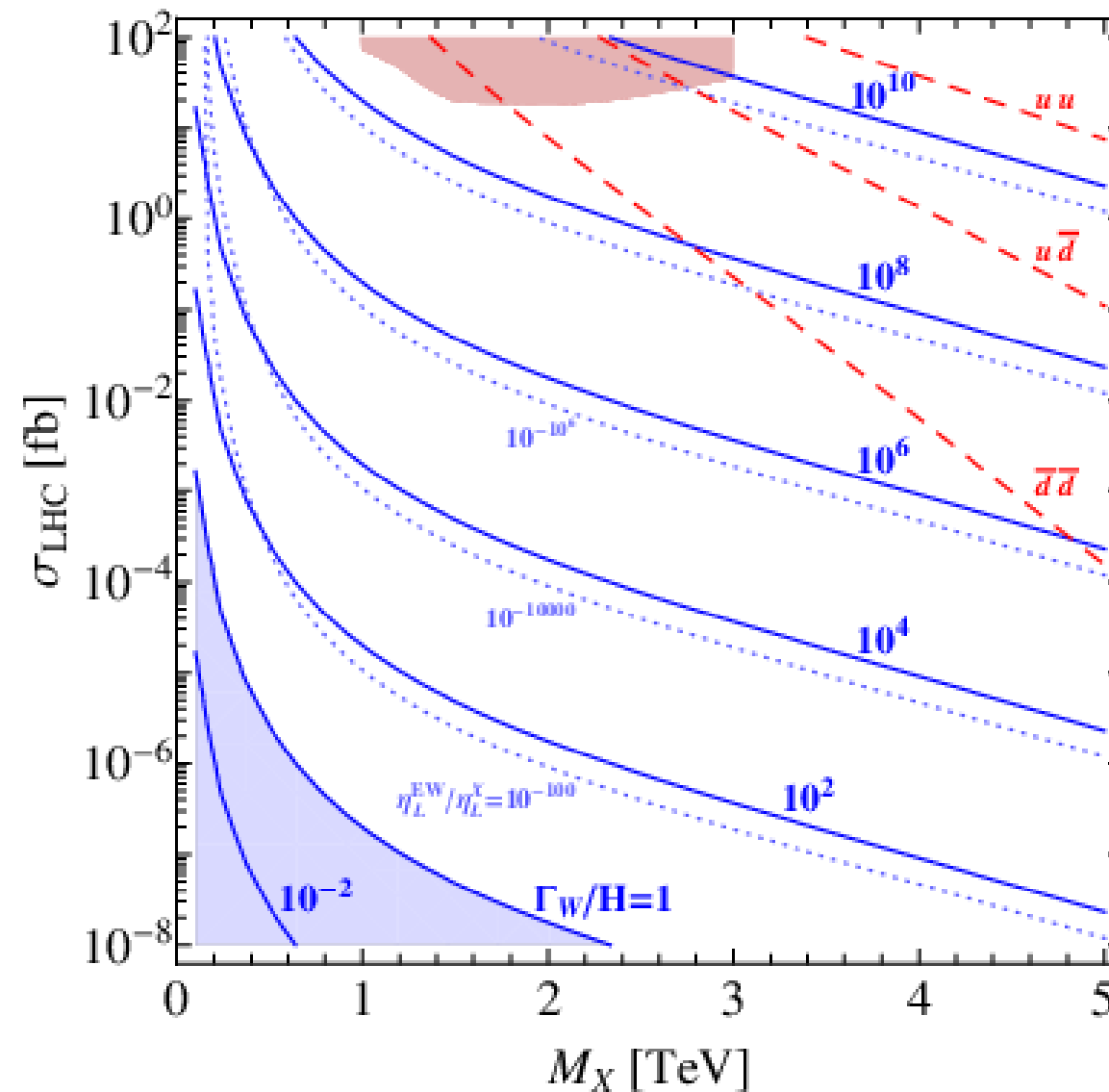
$$\sigma(s) = \frac{4 \cdot 9 \cdot s}{f_{q_1 q_2} (M_X / \sqrt{s})} \sigma_{\text{LHC}}$$

Measureable LVN signal at LHC and corresponding resonant mass sets lower limit on baryon asymmetry washout

$$\log_{10} \frac{\Gamma_W}{H} > 6.9 + 0.6 \left(\frac{M_X}{\text{TeV}} - 1 \right) + \log_{10} \frac{\sigma_{\text{LHC}}}{\text{fb}}$$

LVN at the LHC

- assuming pre-existing lepton asymmetry generated at high scale

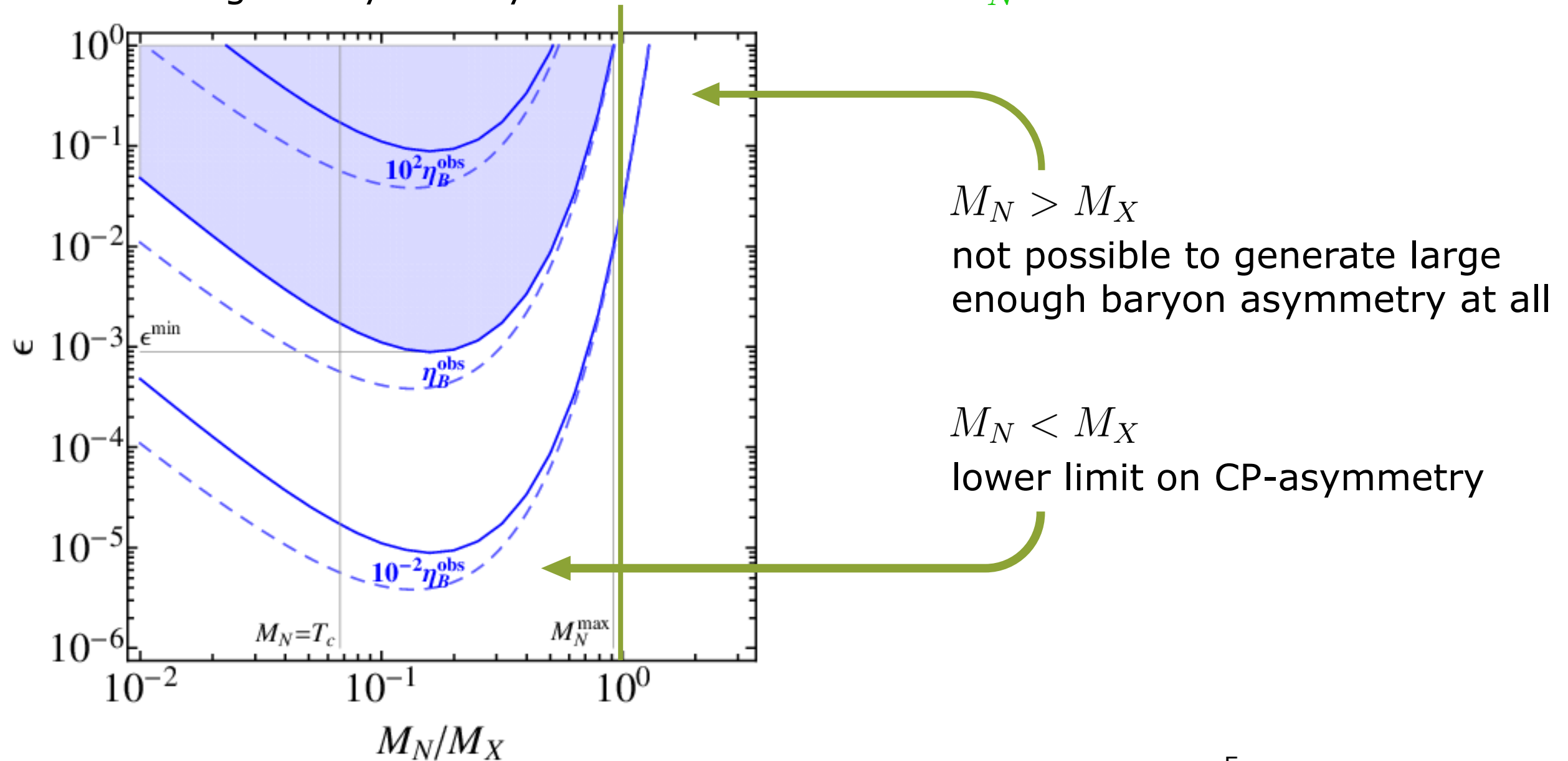


$$\log_{10} \frac{\Gamma_W}{H} > 6.9 + 0.6 \left(\frac{M_X}{\text{TeV}} - 1 \right) + \log_{10} \frac{\sigma_{\text{LHC}}}{\text{fb}}$$

observation of LVN process at the LHC implies very strong washout

LVN at the LHC

- assuming CP asymmetry ϵ is created at scale M_N



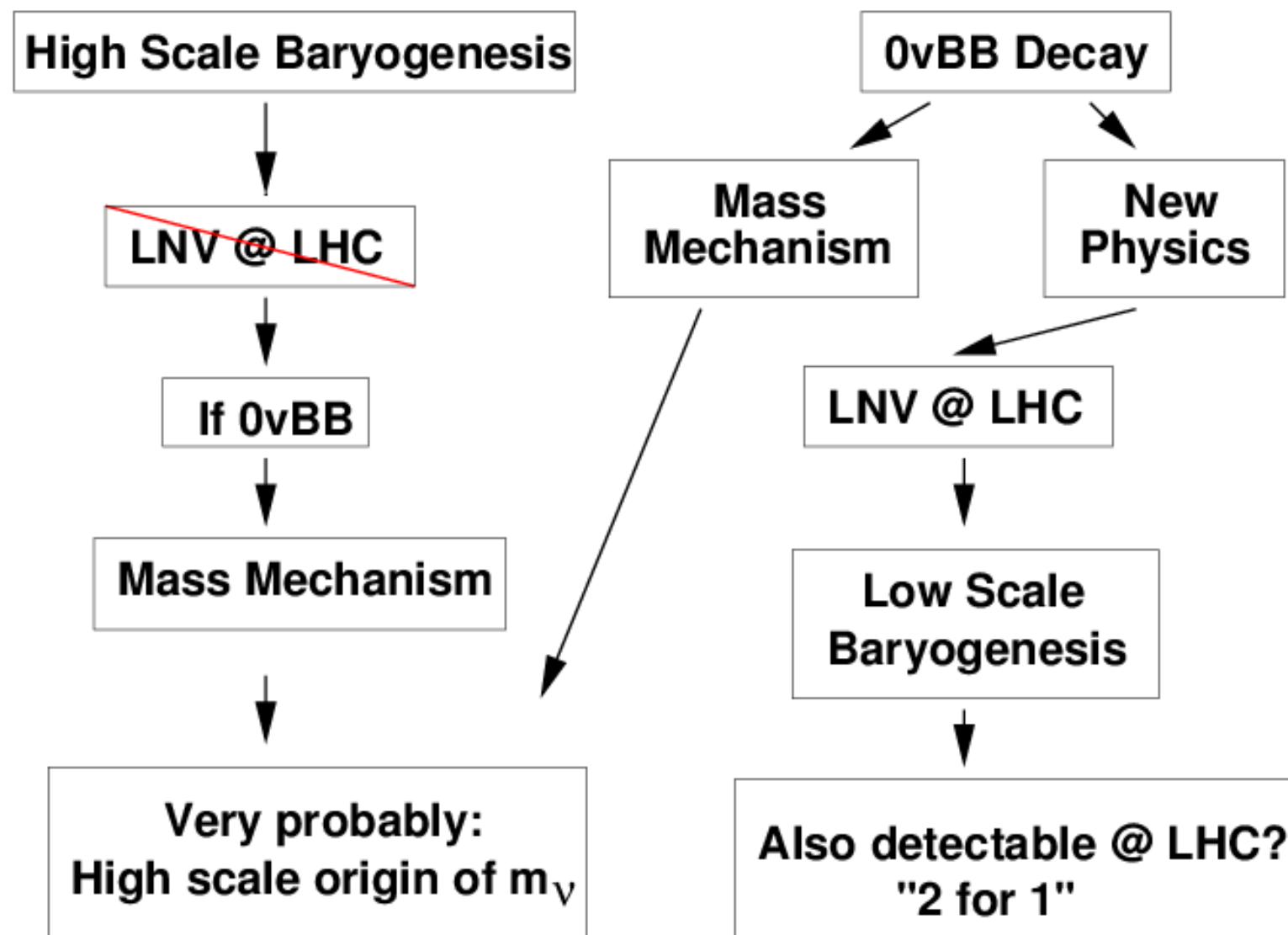
$$\sigma_{\text{LHC}} = 0.1 \text{ fb}$$

$$M_X = 2 \text{ TeV}$$

$$\log_{10} \left| \frac{\eta_B}{\eta_B^{\text{obs}}} \right| < 2.4 \frac{M_X}{\text{TeV}} \left(1 - \frac{4}{3} \frac{M_N}{M_X} \right) + \log_{10} \left[|\epsilon| \left(\frac{\sigma_{\text{LHC}}}{\text{fb}} \right)^{-1} \left(\frac{4}{3} \frac{M_N}{M_X} \right)^2 \right]$$

observation of LVN process at the LHC excludes high-scale leptogenesis models and sets lower limit on the baryon asymmetry of an low-scale leptogenesis model

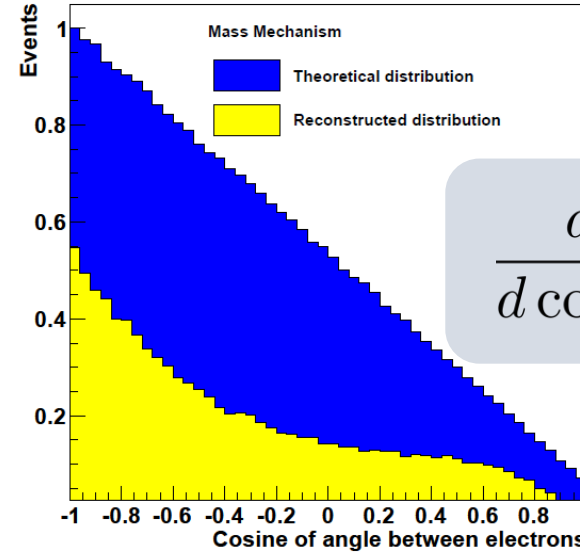
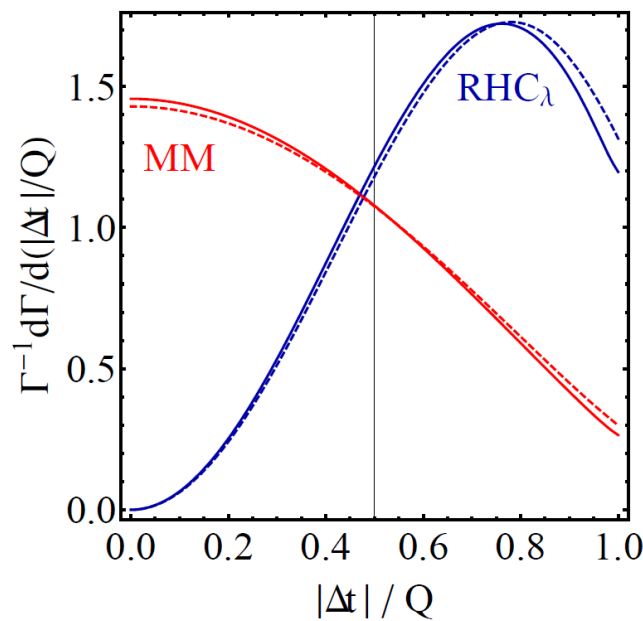
Conclusions



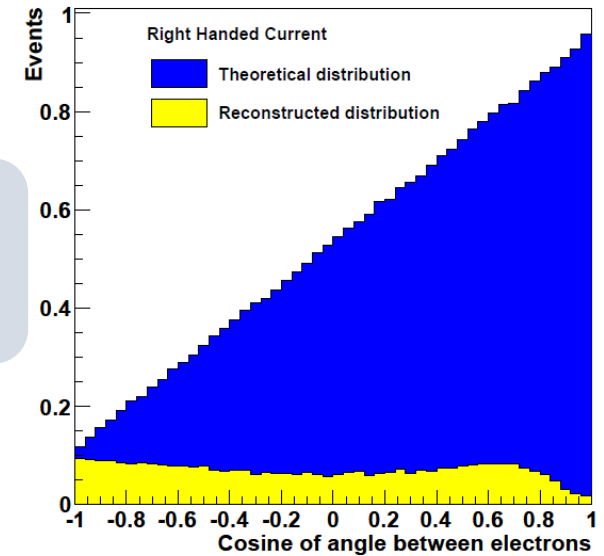
observation of low energy LNV processes (e.g. in 0vbb or LHC) indicates a washout of any pre-existing baryon asymmetry irrespective of the baryogenesis mechanism

Distinguishing between different Operators

- SuperNEMO can discriminate O_7 from other, due to e^-_R and e^-_L in final state

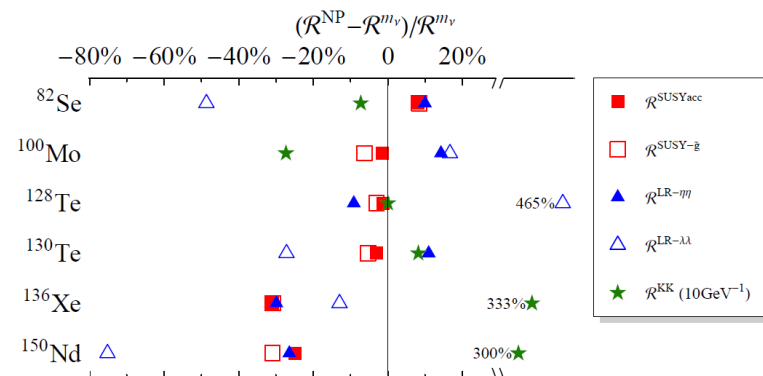


$$\frac{d\Gamma}{d \cos \theta_{12}} = \frac{\Gamma}{2} (1 - k_\theta \cos \theta_{12})$$



SuperNEMO collaboration, arXiv:1005.1241 [hep-ex]

- potential discrepancy between neutrino mass (cosmology) and $0\nu\beta\beta$ half live measurement could be an indication for $0\nu\beta\beta$ triggered by non-standard mechanism
- distinguishing between different mechanisms via measurements in different isotopes



$$\frac{T_{1/2}(^AX)}{T_{1/2}(^AX)} = \frac{|\mathcal{M}(^{76}\text{Ge})|^2 G(^{76}\text{Ge})}{|\mathcal{M}(^AX)|^2 G(^AX)}$$

Deppisch, Paes, PRL 98 (2007)
Gehmann, Elliott, J. Phys G 34 (2007)

- comparison of $0\nu\beta^-\beta^-$ with $0\nu\beta^+\beta^+$ Hirsch, Muto, Oda, Klapdor-Kleingrothaus, Z. Phys A347 (1994)
- observation of $0\nu\beta\beta$ via O_9 and O_{11} will imply observation of LNV at LHC