

# Classification of Maximally Supersymmetric Backgrounds in Supergravity Theories

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# Introduction


- ▶ Supersymmetric solutions in supergravity are well studied, many are classified.
- ▶ They are believed to uplift to solutions of the full string theory and are related to string solitons, i.e. non-perturbative string theory.
- ▶ Here: Maximally supersymmetric solutions for all gauged and deformed supergravity theories in  $D \geq 3$  space-time dimensions.

# Outline / Results

A) Set the stage to discuss solutions in a generic framework.

B) Solutions *without* fluxes:

- ▶ Only two possible cases:

$Mink_D$                       and                       $AdS_D$   
(only gauged / deformed SUGRAs) 


C) Solutions *with* (non-trivial) fluxes:

- ▶ Exist only for a small class of theories.
  - ▶ Solutions coincide with those of the corresponding ungauged theories.
  - ▶ For these theories all solutions are known and classified.
- Exhaustive list of solutions.

# Maximally supersymmetric backgrounds

- ▶ Classical solutions / backgrounds for which

$$\langle \delta_\epsilon B \rangle = \langle \delta_\epsilon F \rangle = 0.$$

 SUSY variations of  
bosonic / fermionic fields

- ▶ Here: bosonic solutions, i.e.  $\langle F \rangle = 0$

$$\Rightarrow \langle \delta_\epsilon B \rangle \sim \langle F \rangle = 0.$$

$\Rightarrow$  only remaining condition:  $\langle \delta_\epsilon F \rangle = 0$

- ▶ indep. supersymmetry parameters  $\epsilon \leftrightarrow$  preserved supercharges

# Supersymmetry variations

- ▶ gravitini:

$$\delta\psi_\mu^i = D_\mu\epsilon^i + (\mathcal{F}_{0\mu})_j^i \epsilon^j + A_{0j}^i \gamma_\mu \epsilon^j$$

fluxes from the gravity multiplet      gaugings / deformations

covariant derivative:  $D_\mu\epsilon^i = \nabla_\mu\epsilon^i + (Q_\mu)_j^i \epsilon^j$

- ▶ spin-1/2 fermions from the gravity multiplet (“dilatini”):

$$\delta\chi^a = (\mathcal{F}_1)_i^a \epsilon^i + A_{1i}^a \epsilon^i$$

- ▶ spin-1/2 fermions from other multiplets (“gaugini”, “hyperini”, ...):

$$\delta\lambda^s = (\mathcal{F}_2)_i^s \epsilon^i + A_{2i}^s \epsilon^i$$

fluxes from other multiplets

# Solutions without fluxes (1)

- ▶ Easiest case: all fluxes vanish:

$$\mathcal{F}_{0\mu} = \mathcal{F}_1 = \mathcal{F}_2 = 0$$

- ▶ Supersymmetric solutions of gauged supergravity without background fluxes need to satisfy the Killing spinor equation:

$$\delta\psi_\mu^i = \nabla_\mu \epsilon^i + A_{0j}^i \gamma_\mu \epsilon^j = 0$$

- ▶ Integrability condition:

$$\left( \frac{1}{4} R_{\mu\nu}{}^{\alpha\beta} \delta_k^i + 2A_{0j}^i A_{0k}^j \delta_\mu^\alpha \delta_\nu^\beta \right) \gamma_{\alpha\beta} \epsilon^k = 0$$

- ⇒ Unbroken supersymmetry (without fluxes):  
**Mink<sub>D</sub>** and **AdS<sub>D</sub>** are the only possible solutions.

## Solutions without fluxes (2)

- ▶ Spin-1/2 variations:

$$\delta\chi^a = A_{1i}^a \epsilon^i = 0, \quad \delta\lambda^s = A_{2i}^s \epsilon^i = 0$$

### Algebraic conditions

$$A_0^2 = -\frac{\Lambda}{2(D-1)(D-2)} \mathbb{1}, \quad A_1 = A_2 = 0$$

- ▶ Compare with the potential:

$$V = -c_0 \text{tr}(A_0^\dagger A_0) + c_1 \text{tr}(A_1^\dagger A_1) + c_2 \text{tr}(A_2^\dagger A_2),$$

- ▶ **Mink<sub>D</sub>** ( $\Lambda = 0$ ) solutions exist for all theories.
- ▶ For  $D \leq 7$  most gauged theories admit **AdS<sub>D</sub>** ( $\Lambda < 0$ ) solutions.

[de Alwis, Louis, McAllister, Triendl, Westphal '13] [Louis, Triendl '14] [Louis, SL '15] [Louis, Triendl, Zagermann '15]

[Louis, Muranaka '16]

# Solutions with non-trivial flux (1)

- ▶ For more “interesting” solutions: Allow for non-vanishing flux.
- ▶ Firstly: spin-1/2 variations:

$$\delta\chi^a = (\mathcal{F}_1)_i^a \epsilon^i + A_{1i}^a \epsilon^i = 0, \quad \delta\lambda^s = (\mathcal{F}_2)_i^s \epsilon^i + A_{2i}^s \epsilon^i = 0$$

- ▶ Unbroken supersymmetry:

$$\mathcal{F}_1 = \mathcal{F}_2 = A_1 = A_2 = 0 \quad \Rightarrow \quad \mathcal{F}_{0\mu} = 0$$

- ▶ Generically no non-trivial fluxes possible!

## Two exceptions

1. theories without  $\chi$ 's in the gravitational multiplet.
2. chiral theories with selfdual fluxes.



## Solutions with non-trivial flux (2)

- ▶ Secondly: gravitino variations:

$$\delta\psi_{\mu}^i = \nabla_{\mu}\epsilon^i + (\mathcal{Q}_{\mu})_j^i + (\mathcal{F}_{0\mu})_j^i \epsilon^j + A_{0j}^i \gamma_{\mu}\epsilon^j = 0$$

- ▶ Integrability condition:

$$\left( \frac{1}{4} R_{\mu\nu\rho\sigma} \gamma^{\rho\sigma} \delta_j^i - (\mathcal{H}_{\mu\nu})_j^i + \dots \right) \epsilon^j = 0,$$

where  $\mathcal{H}_{\mu\nu}$  is the field strength corresponding to  $\mathcal{Q}_{\mu}$ .

- ▶ Unbroken supersymmetry:

$$\mathcal{H}_{\mu\nu} = 0 \quad \Rightarrow \quad \mathcal{Q}_{\mu} = A_0 = 0$$

(See also [Hristov,Looyestijn,Vandoren] [Gauntlett,Gutowski] [Akyol,Papadopoulos])

- ▶ The supersymmetry variations take the same form as in the ungauged / undeformed case.

## Solutions with non-trivial flux (3)

SUGRAs with non supersymmetry breaking flux:

dimension	supersymmetry	q	possible flux	classification
$D = 11$	$N = 1$	32	$F^{(4)}$	[Figueroa-O'Farrill,Papadopoulos]
$D = 10$	IIB	32	$F_+^{(5)}$	[Figueroa-O'Farrill,Papadopoulos]
$D = 6$	$N = (2, 0)$	16	$5 \times F_+^{(3)}$	[Chamseddine,Figueroa-O'Farrill,Sabra]
$D = 6$	$N = (1, 0)$	8	$F_+^{(3)}$	[Gutowski,Martelli,Reall]
$D = 5$	$N = 2$	8	$F^{(2)}$	[Gauntlett,Gutowski,Hull,Pakis,Reall]
$D = 4$	$N = 2$	8	$F^{(2)}$	[Tod]

The maximally supersymmetric solutions are classified:

- ▶  $AdS_p \times S^{(D-p)}$  and  $AdS_{(D-p)} \times S^p$ , for  $p$ -form flux.
- ▶ Hpp-wave as Penrose-limit of  $AdS \times S$  solutions.

[Penrose;Gueven;Blau,Figueroa-O'Farrill,Hull,Papadopoulos]

- ▶ Exceptional solutions in  $D = 5$ . [Gauntlett,Gutowski,Hull,Pakis,Reall]

# All maximally supersymmetric solutions

(with non-trivial flux)

dim.	SUSY	q	$AdS \times S$	Hpp-wave	others
$D = 11$	$\mathcal{N} = 1$	32	$AdS_4 \times S^7$ $AdS_7 \times S^4$	KG <sub>11</sub>	-
$D = 10$	IIB	32	$AdS_5 \times S^5$	KG <sub>10</sub>	-
$D = 6$	$\mathcal{N} = (2, 0)$	16	$AdS_3 \times S^3$	KG <sub>6</sub>	-
	$\mathcal{N} = (1, 0)$	8			
$D = 5$	$\mathcal{N} = 2$	8	$AdS_2 \times S^3$ $AdS_3 \times S^2$	KG <sub>5</sub>	Gödel-like, NH-BMPV*
$D = 4$	$\mathcal{N} = 2$	8	$AdS_2 \times S^2$	KG <sub>4</sub>	-

\* = near-horizon limit of the BMPV [Breckenridge, Myers, Peet, Vafa] black hole

# Summary

- ▶ Systematic classification of all maximally supersymmetric supergravity backgrounds:
- ▶ no fluxes: only flat space-time and anti-de Sitter.
- ▶ fluxes: - generically not possible
  - in the gauged case: same solutions as for ungauged theories
  - all solutions are known and classified

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Thank You!