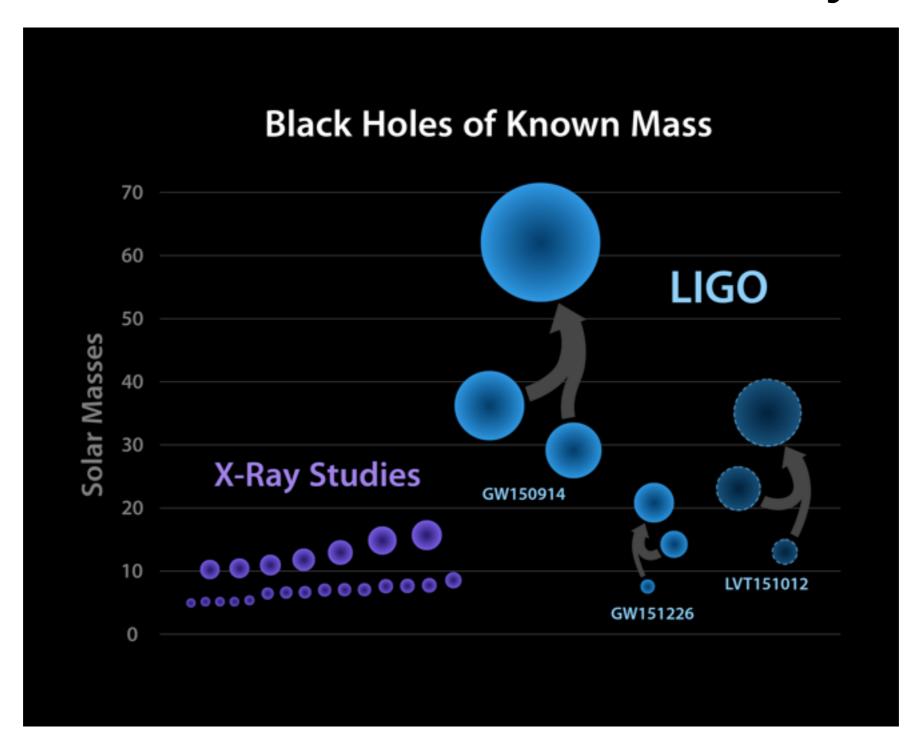
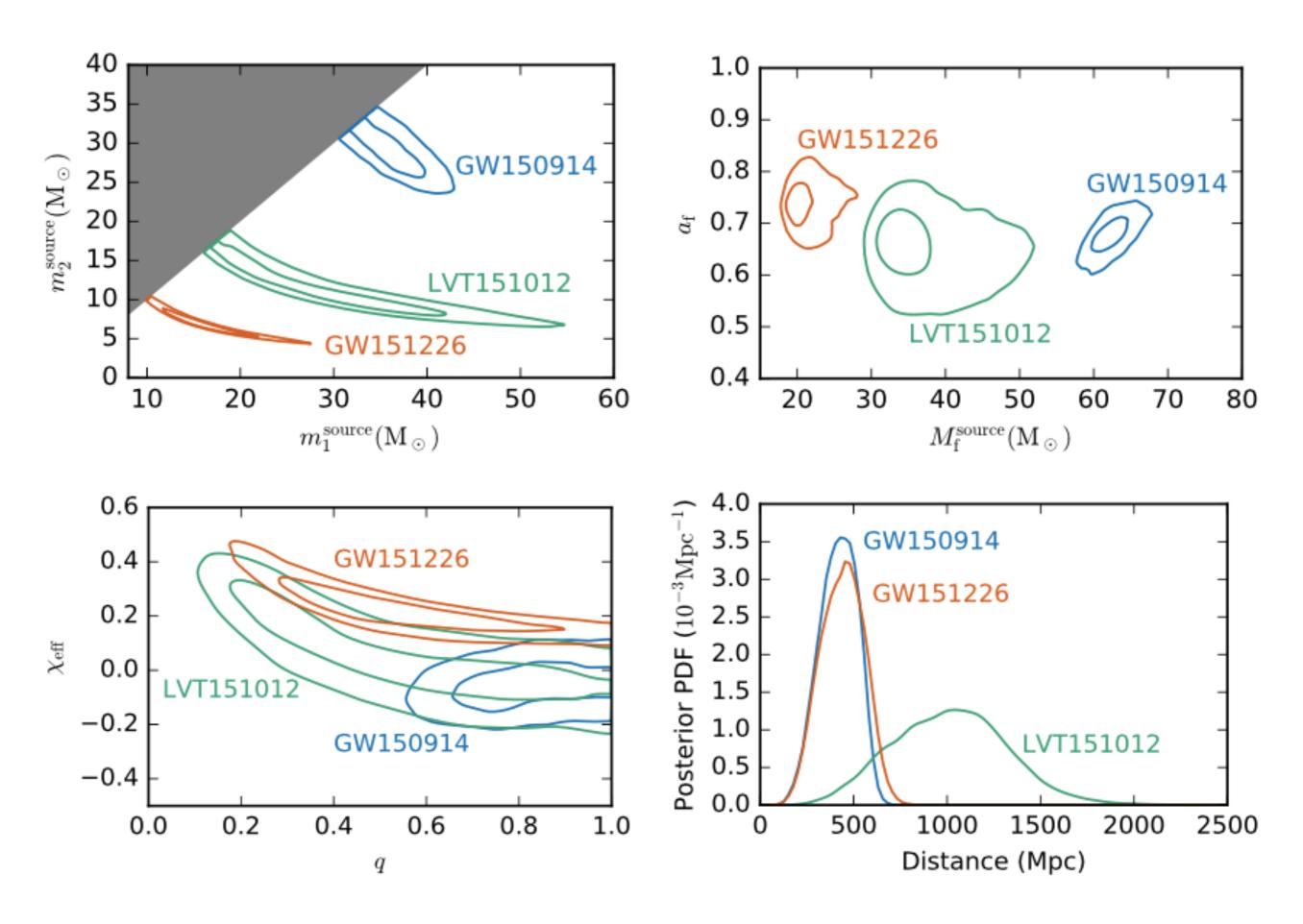
# Astrophysical consequences of the LIGO discovery



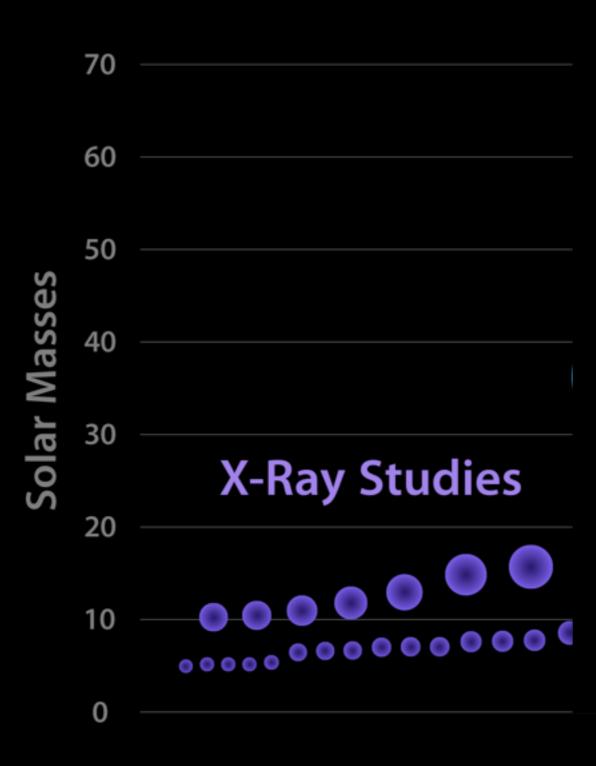
Ilya Mandel University of Birmingham

- BBHs exist
- Merging BBHs exist
- stellar-mass BHs with mass above 30 solar masses exist (and take part in mergers)
- Don't know formation channel from single event isolated binary with CE? chemically homogeneous evolution? dynamical formation?
- Likely formed in low metallicity environment (either locally, or at high z with long time delay)
- Primary has spin of <0.7 at 90% confidence; no evidence for spins being both large and strongly aligned

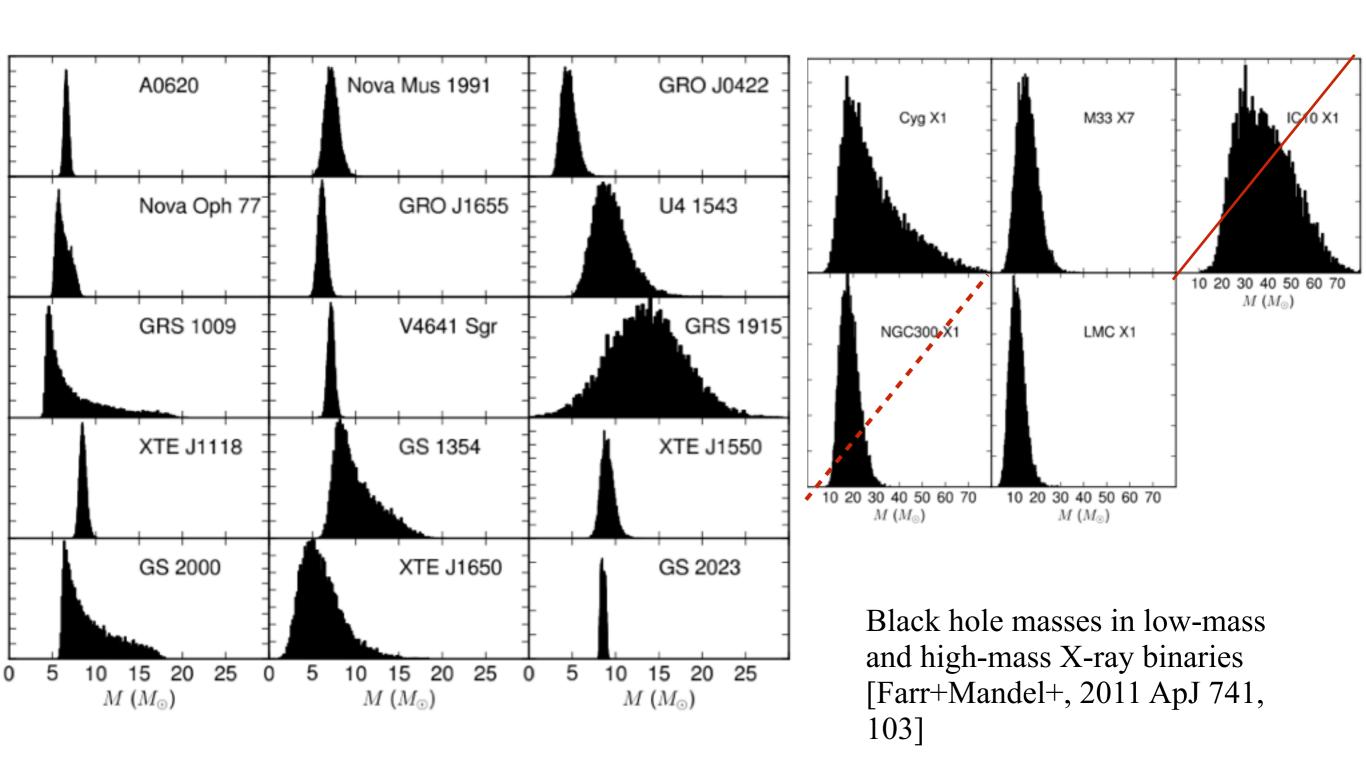


Abbott et al., 2016

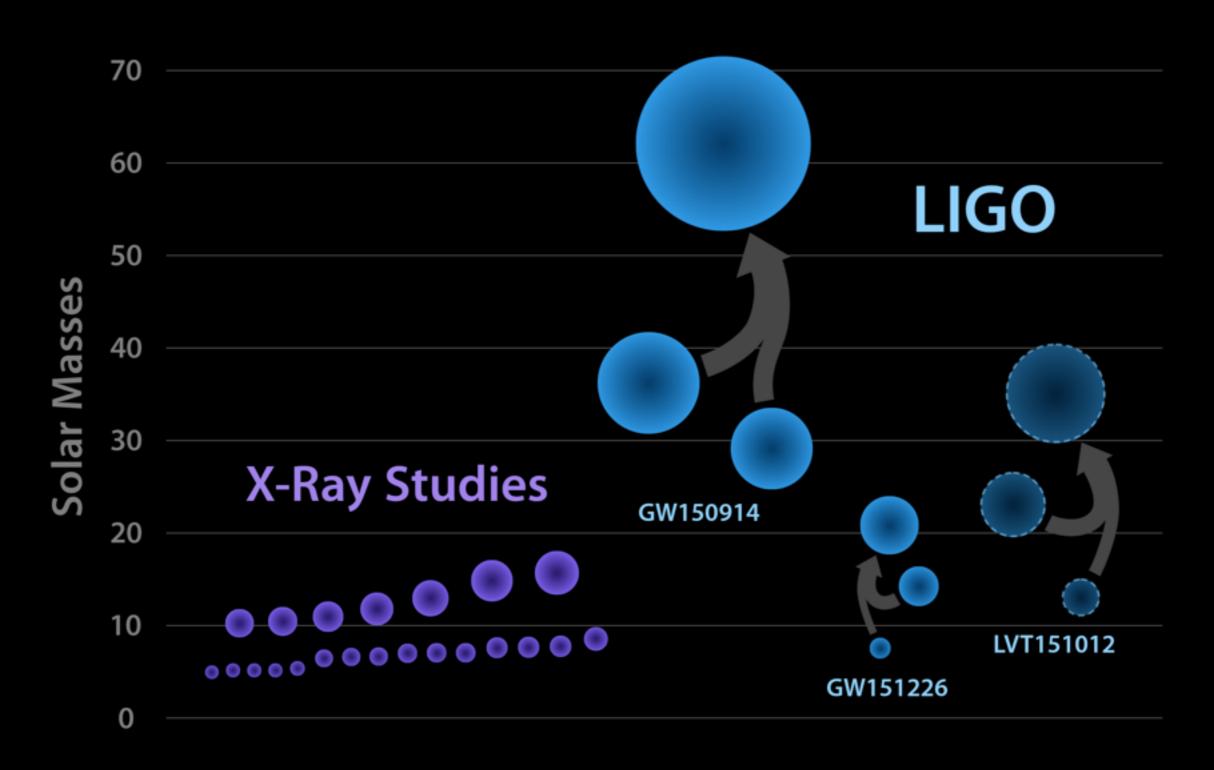
#### **Black Holes of Known Mass**



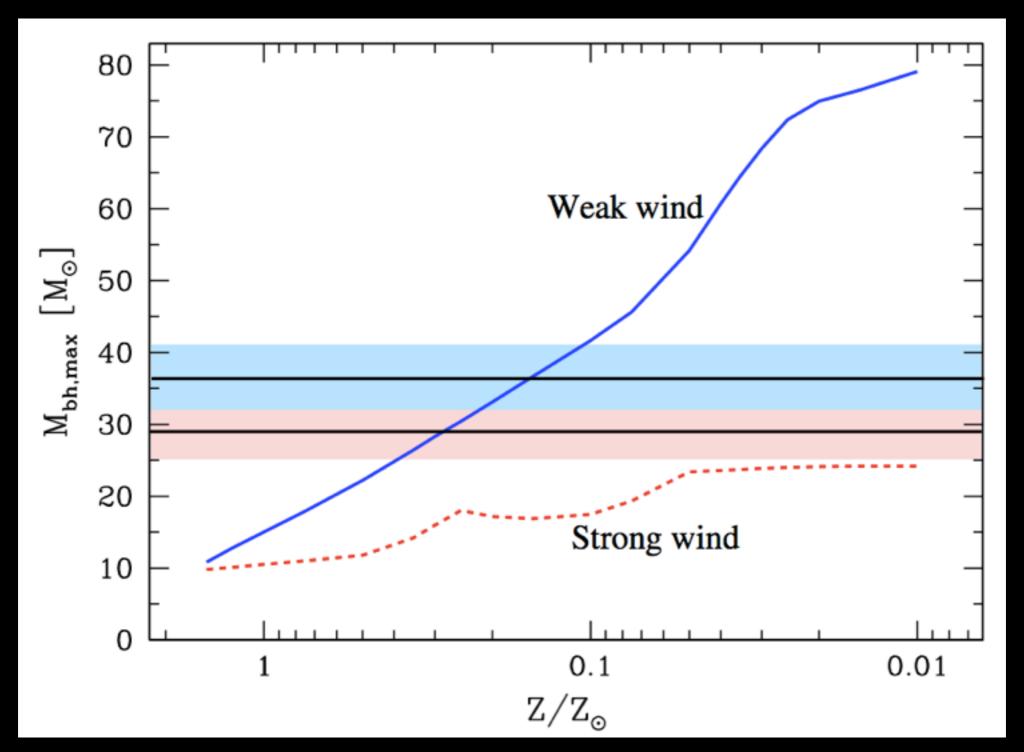
#### BH masses in X-ray binaries



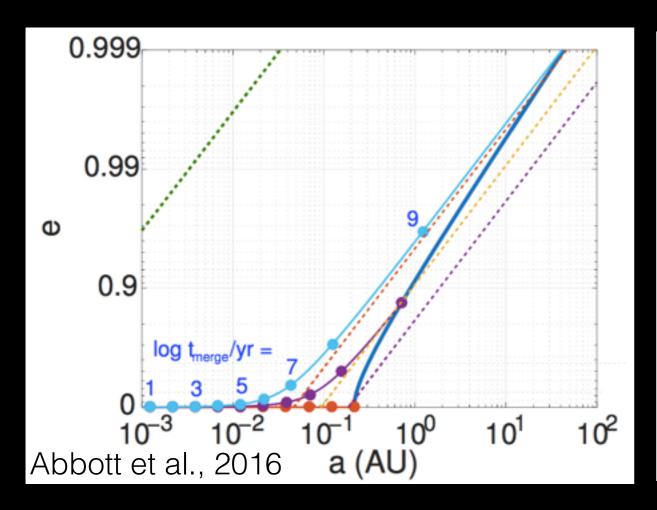
#### **Black Holes of Known Mass**

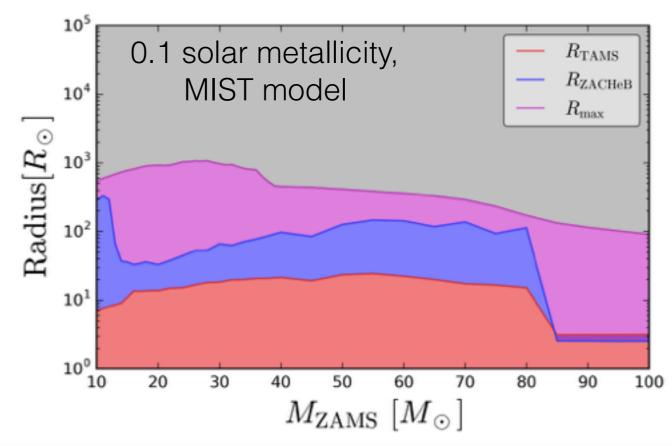


## The mass problem

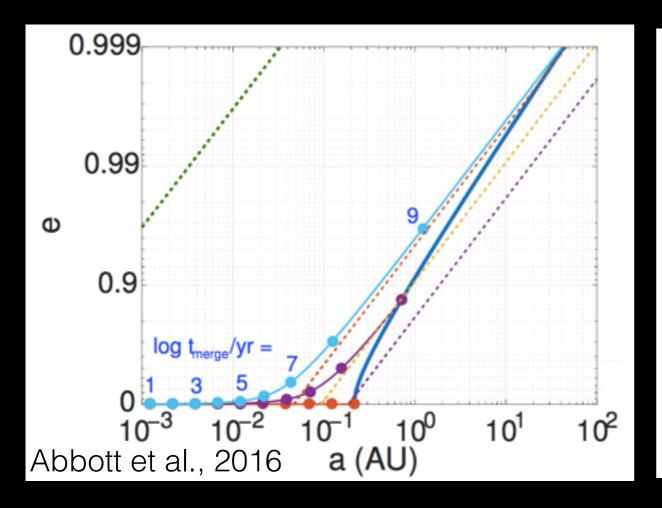


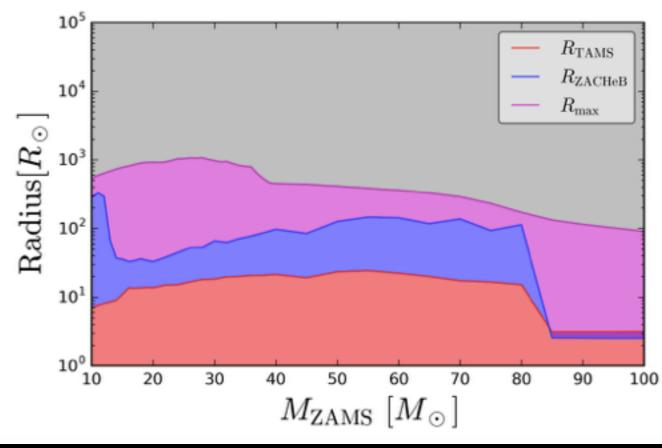
Belczynski+ (2010), adapted in Abbott+ (2016)



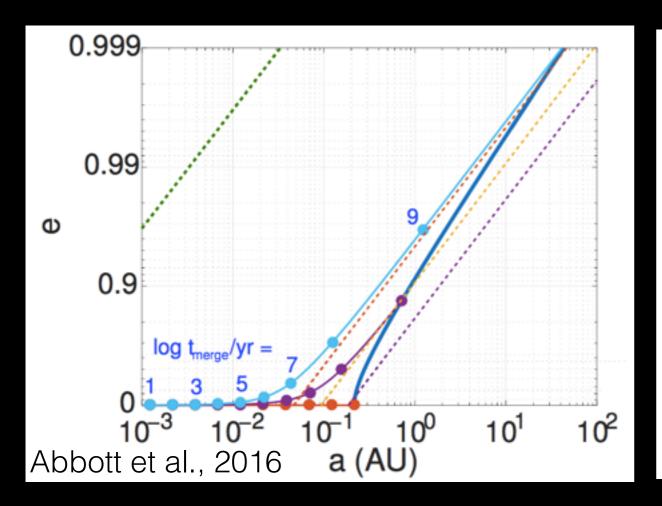


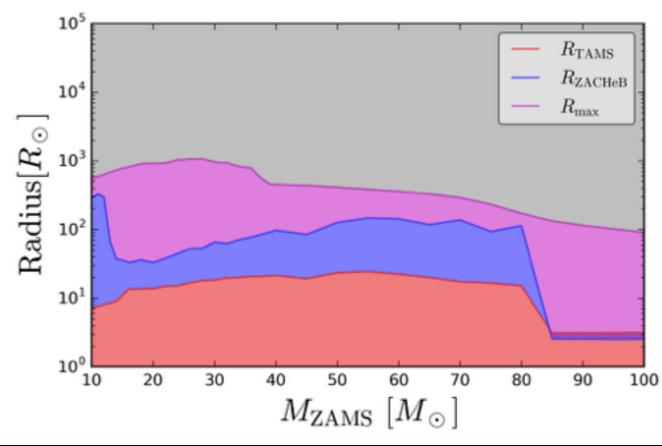
MIST: Dotter (2016), Choi+ (2016) based on MESA, Paxton+ (2011, 2013, 2015)





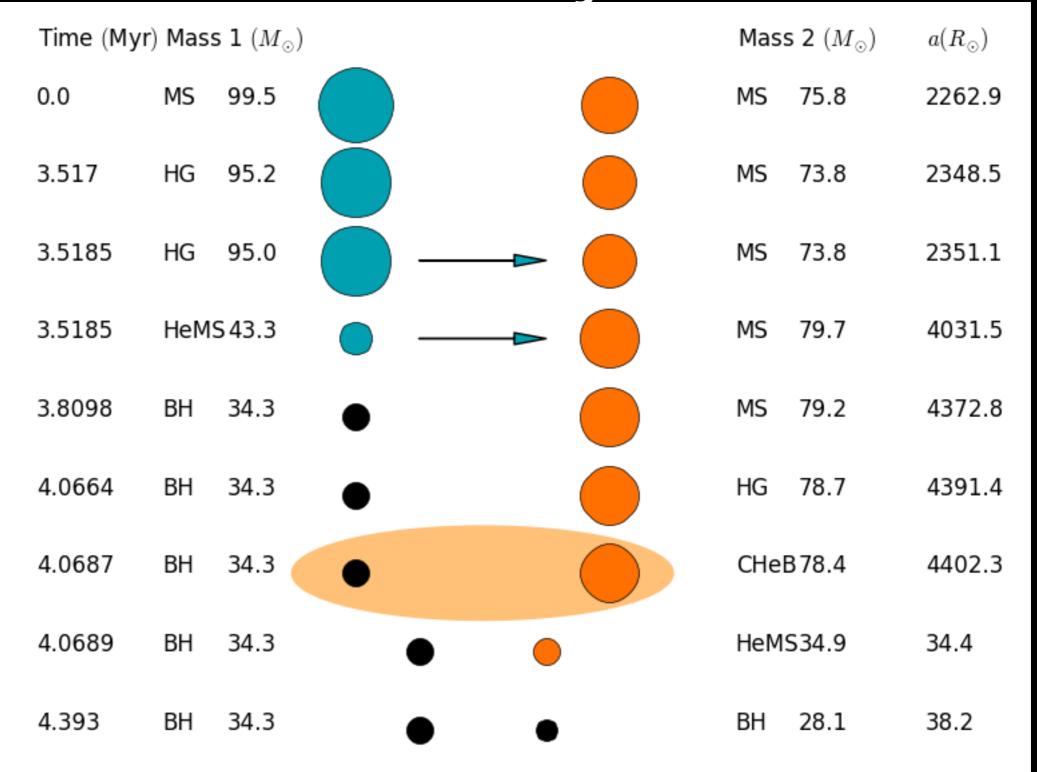
- Harden the binary after stars evolved
- Don't let stars expand
- Let stars evolve before they make a binary



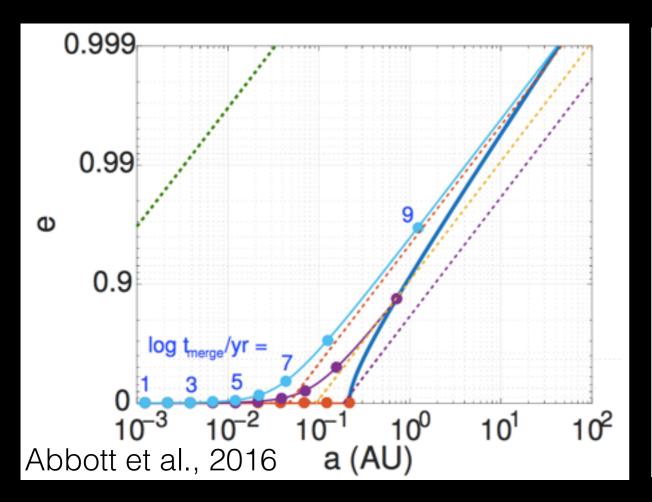


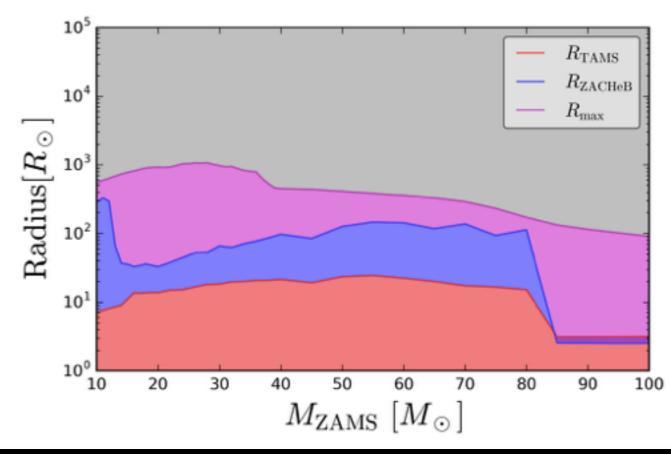
- Harden the binary after stars evolved: classical binary evolution via a common envelope
  - Tutukov & Yungelson 1973, 1993; Lipunov, Postnov & Prokhorov (1997), Bethe & Brown (1998), Bloom, Sigurdsson & Pols (1999), De Donder & Vanbeveren (2004), Grishchuk et al. (2001), Nelemans (2003), Voss & Tauris (2003), Pfahl, Podsiadlowski & Rappaport (2005), Dewi, Podsiadlowski & Sena (2006), Kalogera et al. 2007; O'Shaughnessy et al. (2008), Mennekens & Vanbeveren (2014), Dominik et al. (2012, 2013, 2015), Belczynski et al. 2016, Eldridge & Stanway (2016), Lipunov+ (2016), ...

#### Isolated binary evolution

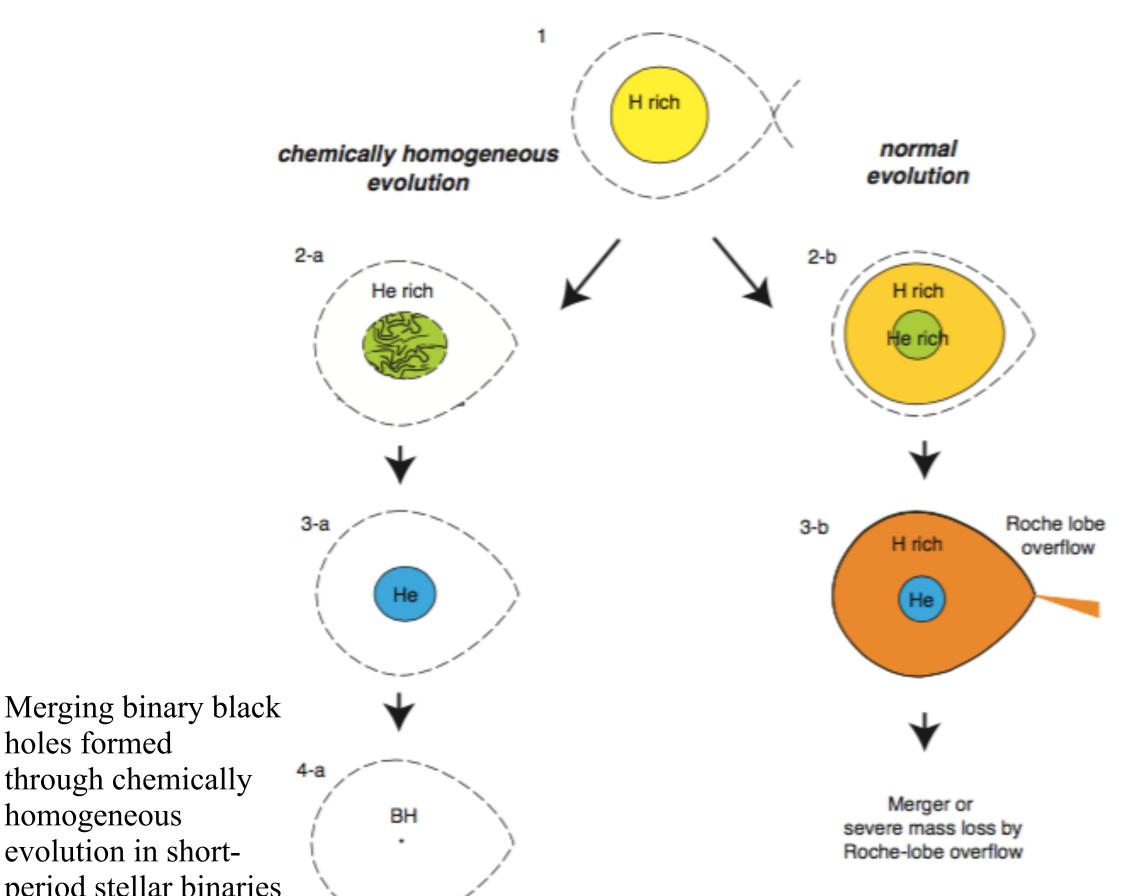


Stevenson, Vigna Gomez, Mandel, Perkins, Barrett, de Mink, 2016



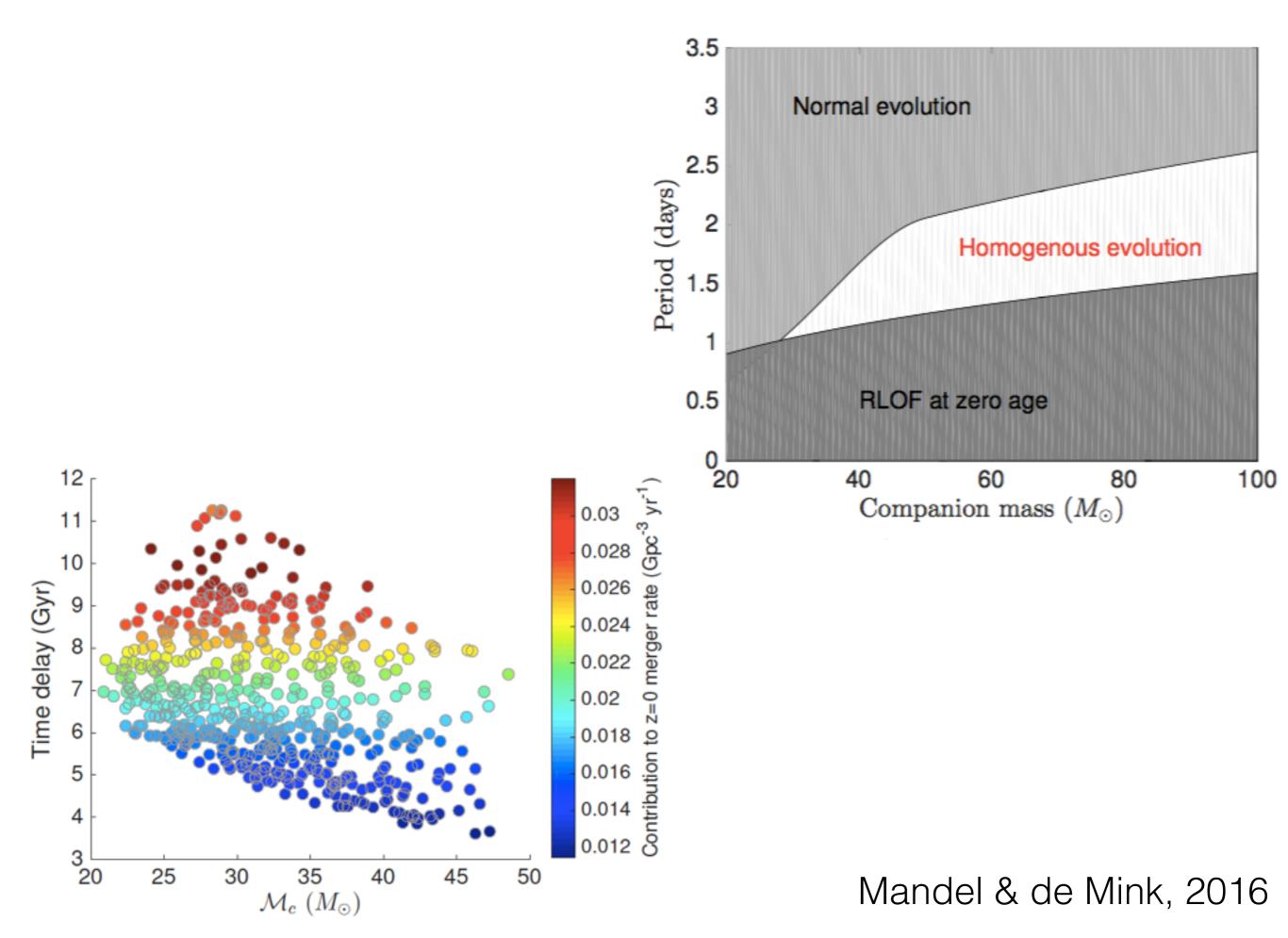


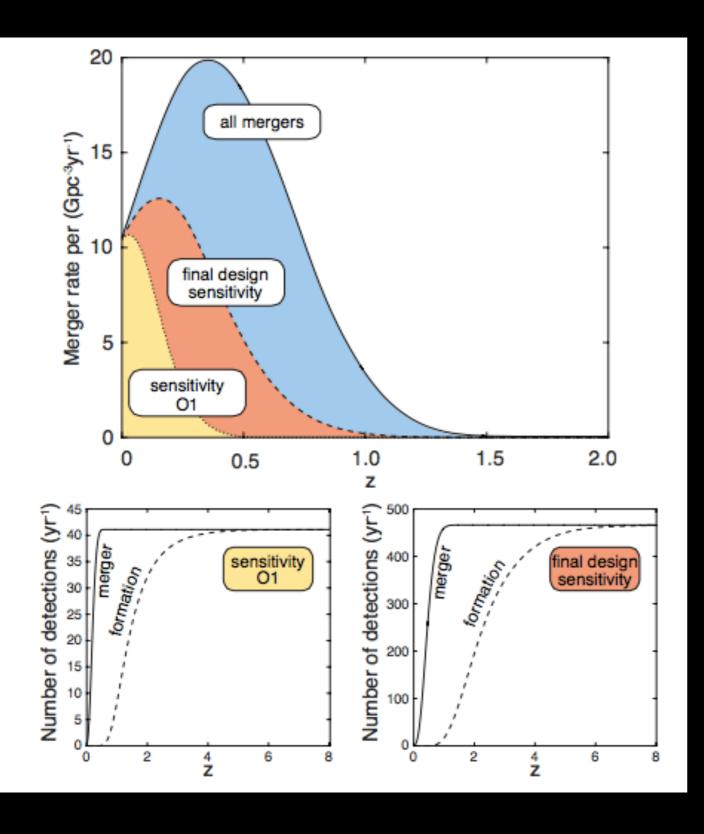
- Harden the binary after stars evolved
- Don't let stars expand: chemically homogeneous evolution
  - Mandel & de Mink, 2016; Marchant+, 2016; de Mink & Mandel, 2016
- Let stars evolve before they make a binary

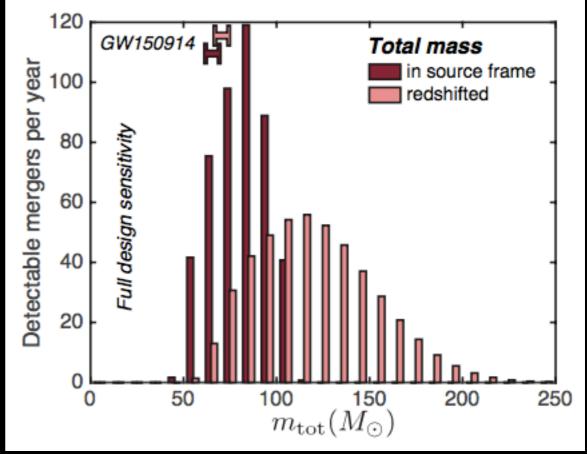


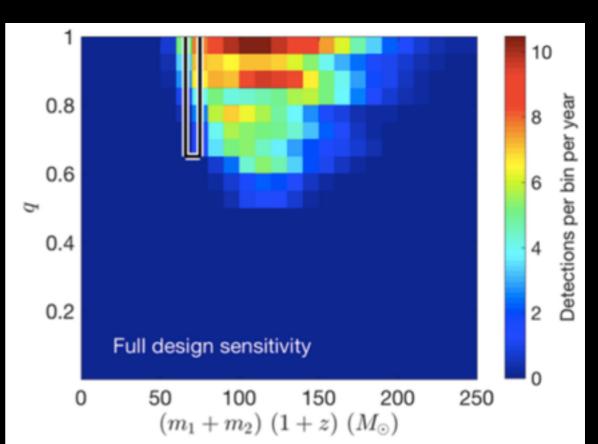
through chemically homogeneous evolution in shortperiod stellar binaries [Mandel & de Mink,2016, after de Mink+, 2009]

holes formed

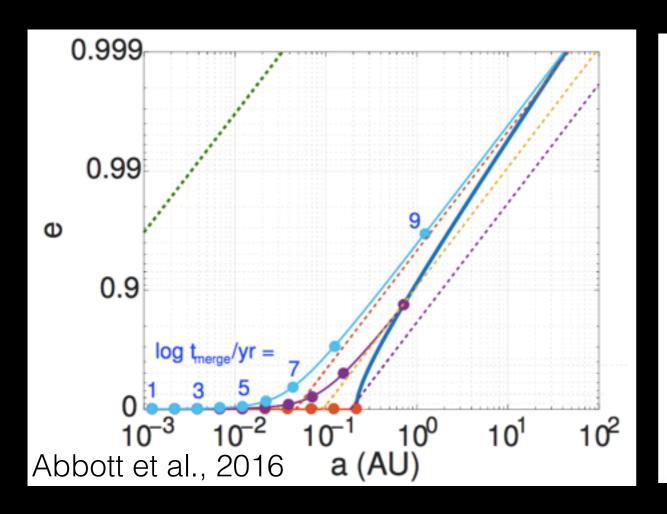


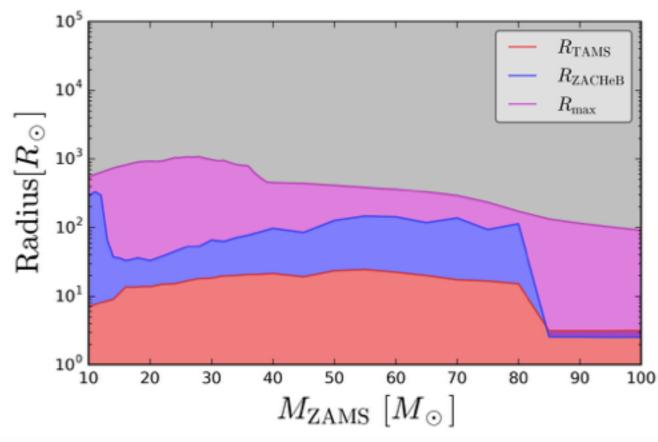




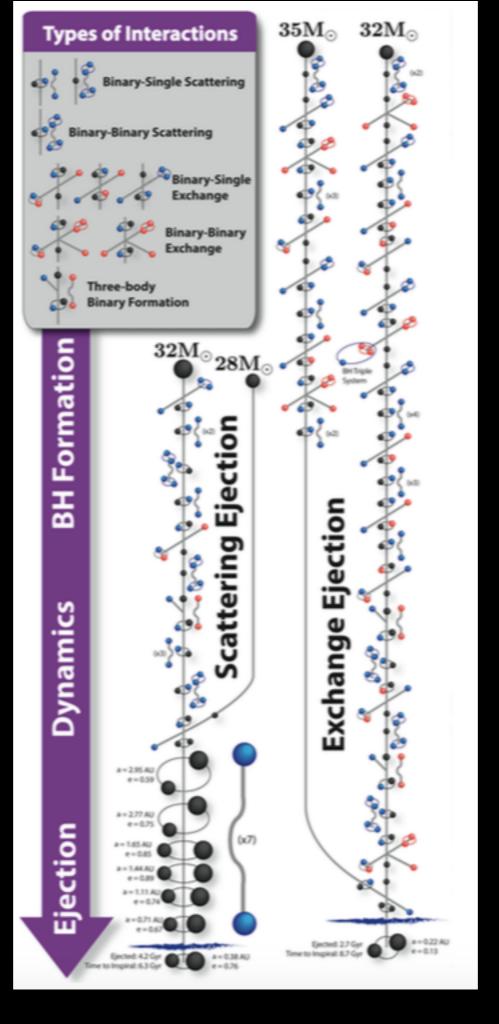


de Mink & Mandel, 2016





- Let stars evolve before they make a binary: dynamical format in clusters
  - Stellar and globular clusters: Sigurdsson & Hernquist (1993); Kulkarni+ (1993); Portegies Zwart & McMillan (2000); Gultekin+ (2004, 2006); Kocsis+ (2006); O'Leary+ (2006, 2007, 2016); Sadowski + (2008); Banerjee+ (2010); Downing+ (2010, 2011); Morscher+ (2013, 2015); Mapelli+ (2016); Rodriguez+ (2015, 2016)
  - Galactic nuclear clusters: Miller & Lauburg (2009); O'Leary+ (2009); Kocsis & Levin (2012); Tsang (2013); Bae+ (2014); Bartos+ (2016); Stone+ (2016)



Rodriguez, Haster+, (2016)



#### The inverse problem of gravitational-wave astrophysics:

how to go from a population of observed sources to understanding key uncertainties about binary evolution?







Alejandro Vigna Gómez



Simon Stevenson

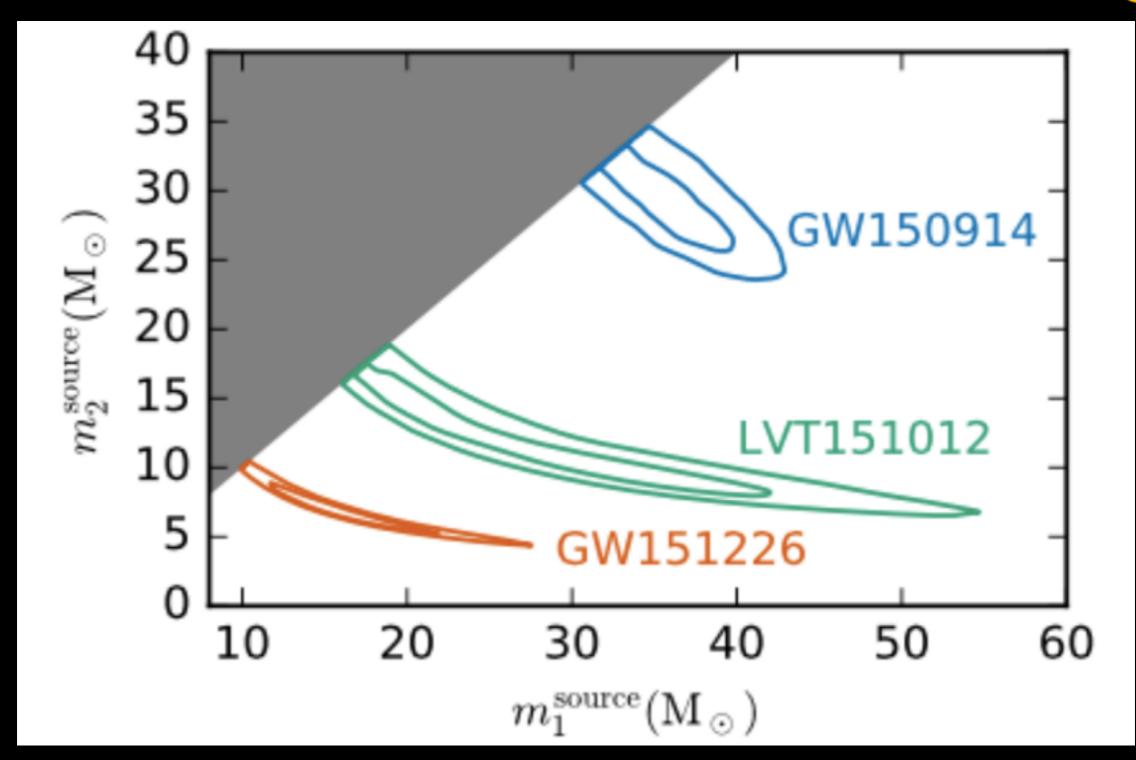


Jim Barrett

Key collaborators / advisors: Christopher Berry, Will Farr, Selma de Mink, Natasha Ivanova, Vicky Kalogera, Chris Belczynski, Gijs Nelemans,

Philipp Podsiadlowski... YOUR NAME COULD BE HERE!

# Population Synthesis

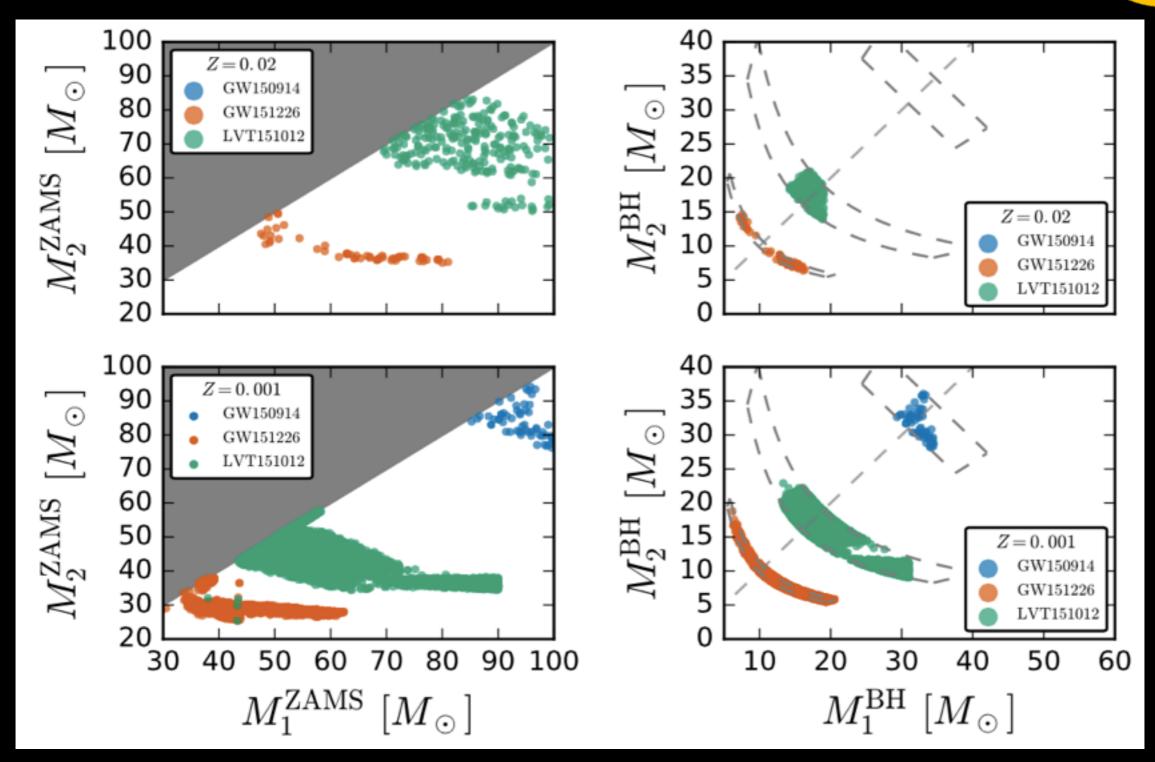


Abbott+ (LVC), 2016

# Population Synthesis

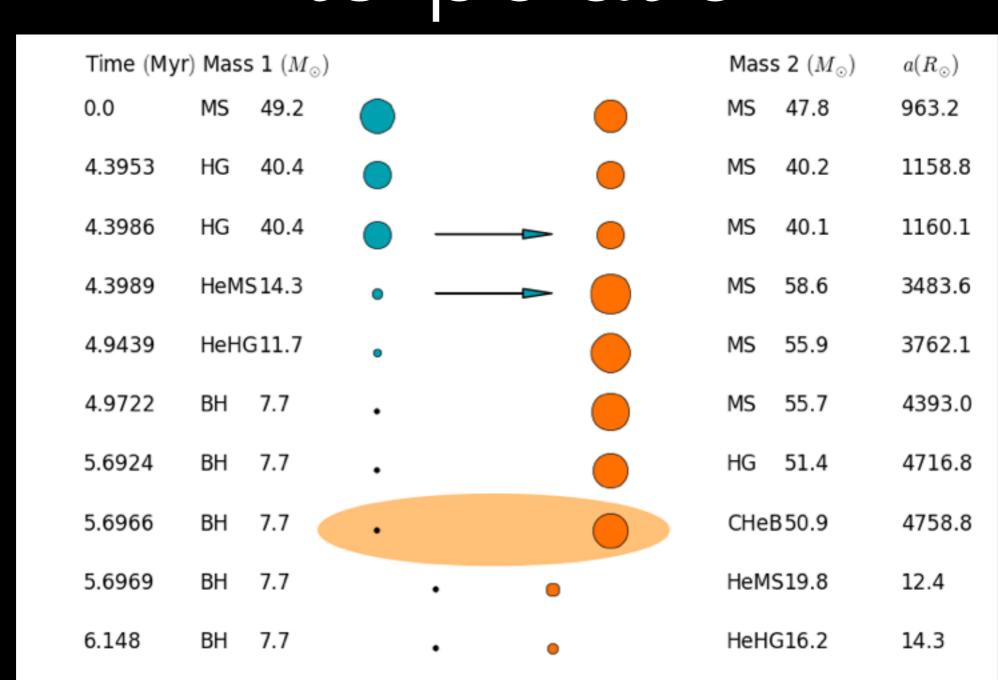


## Population Synthesis



Stevenson, Vigna Gomez, Mandel, Perkins, Barrett, de Mink, 2016

## Interpolation





6.1556

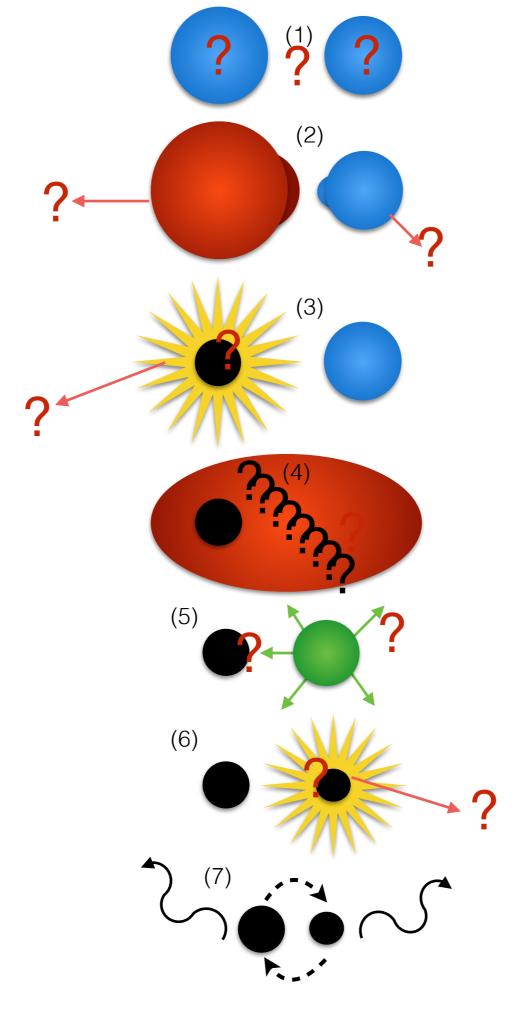
BH

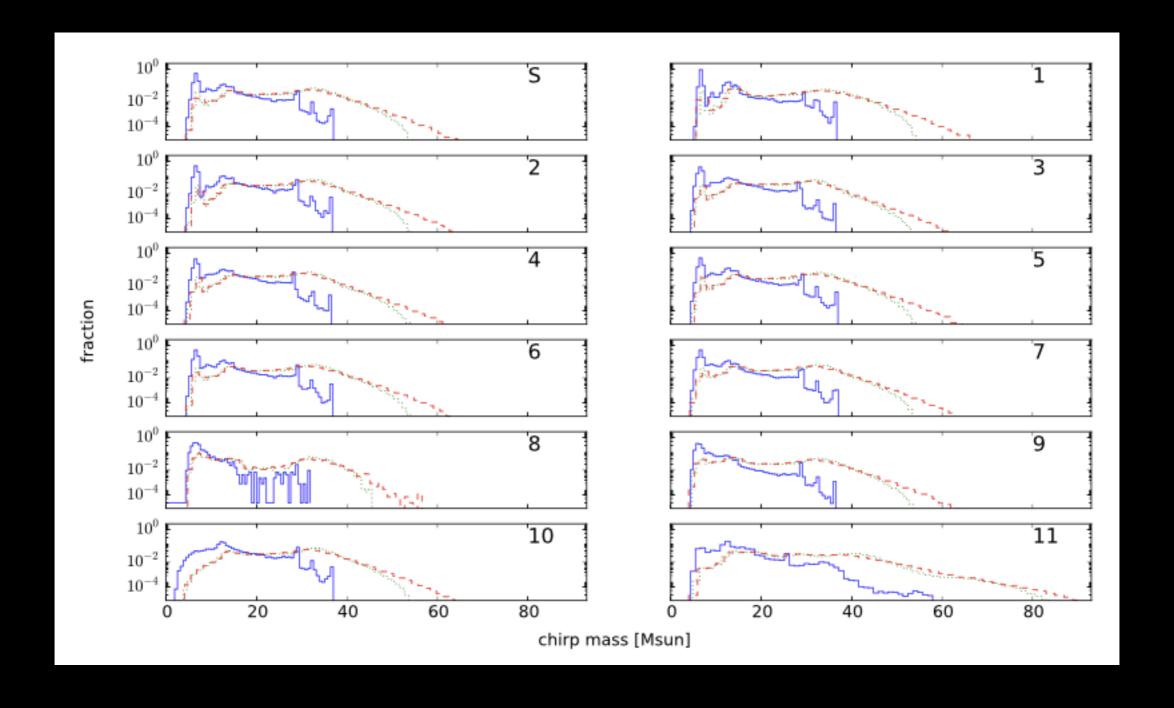
7.7

15.5

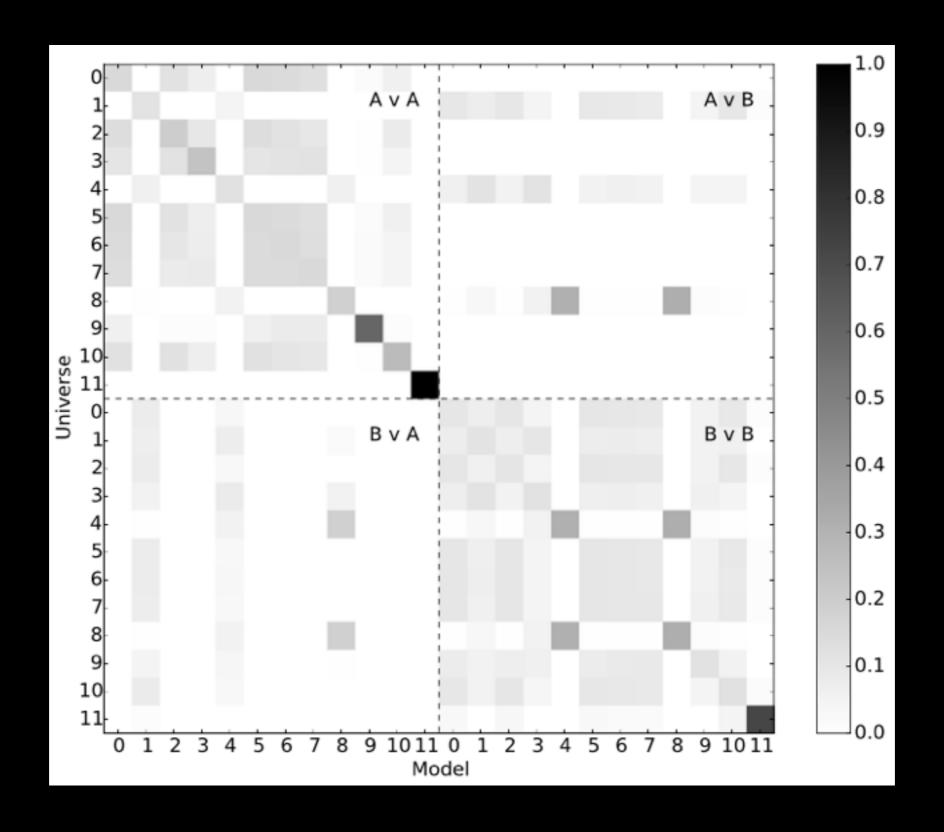
BH

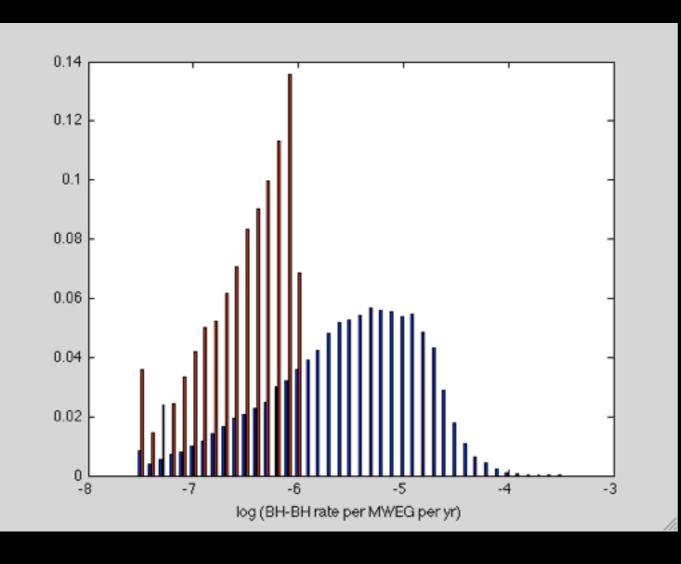
14.3

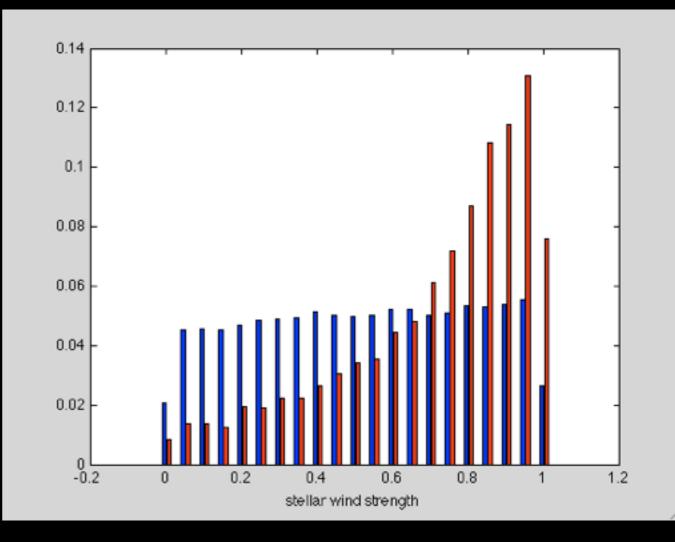




Stevenson+ 2015, from Dominik+ 2012

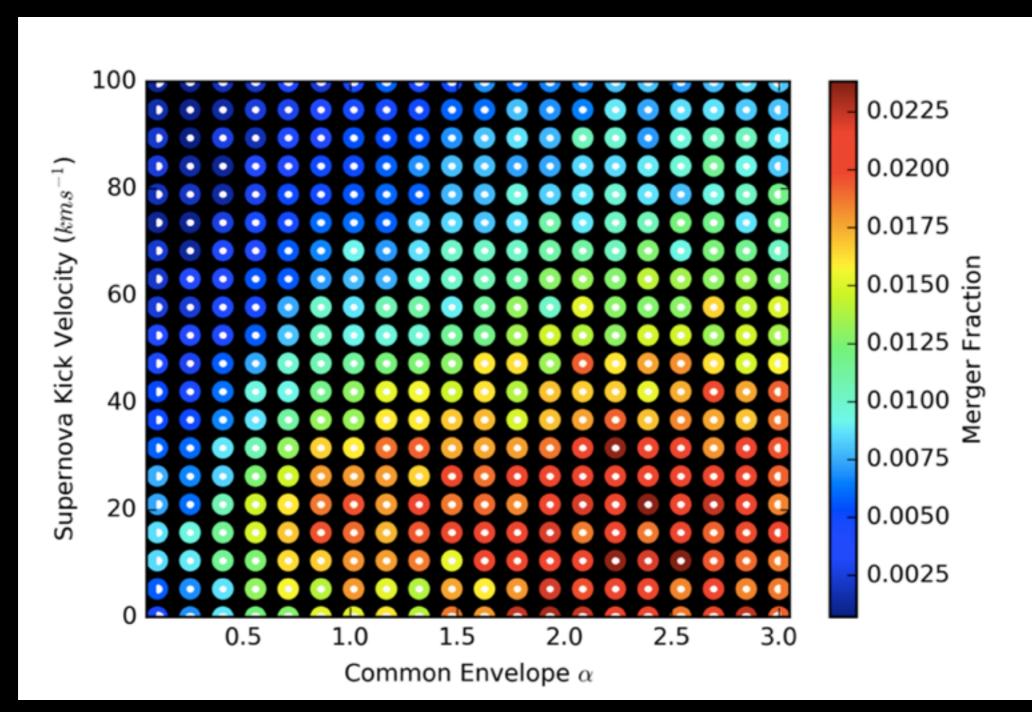






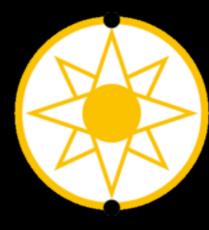
# Interpolation

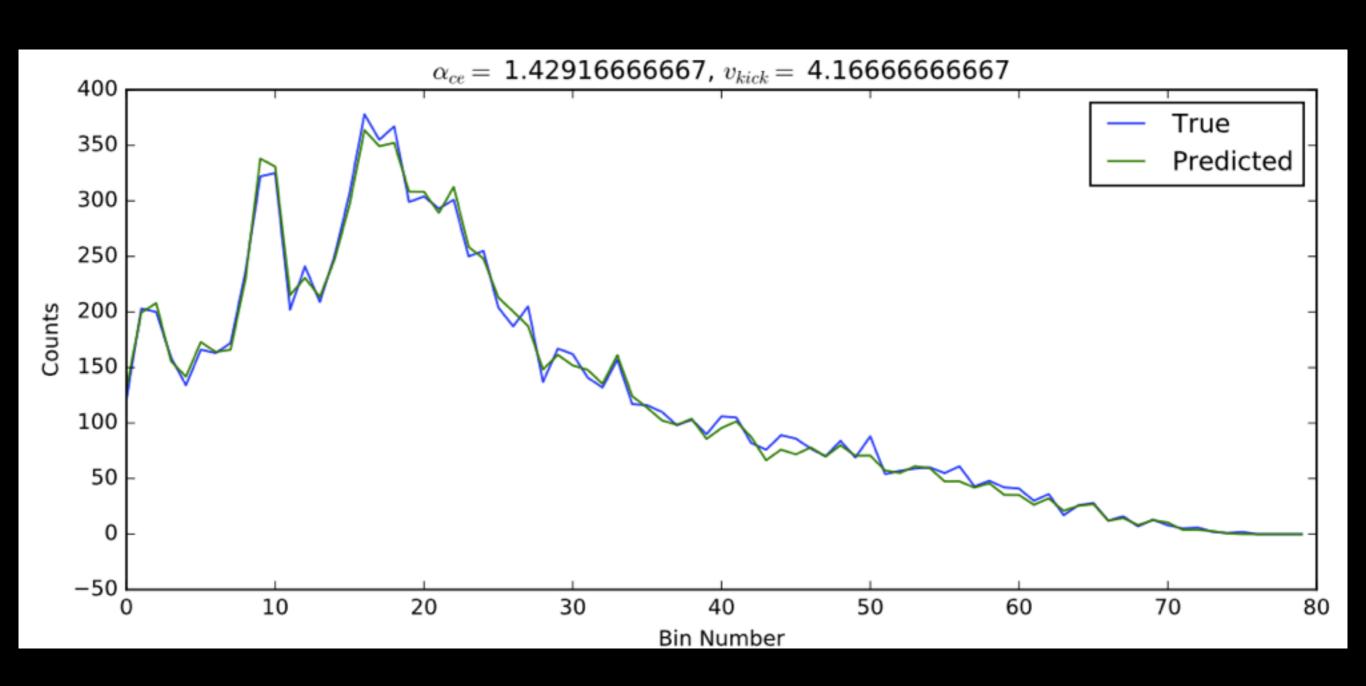




Barrett et al., in prep.

### Interpolation





Barrett et al., in prep.

#### Population Reconstruction



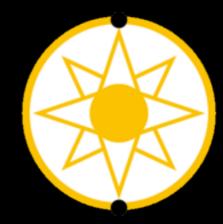
Selection effects and measurement uncertainty

$$p(\{\vec{d}^{(i)}\}|\vec{\lambda}) = \prod_{i=1}^{k} \frac{\int d\vec{\theta} p(\vec{d}^{(i)}|\vec{\theta}) p_{\text{pop}}(\vec{\theta}|\vec{\lambda})}{\int d\vec{\theta} p_{\text{det}}(\theta) p_{\text{pop}}(\vec{\theta}|\vec{\lambda})}$$

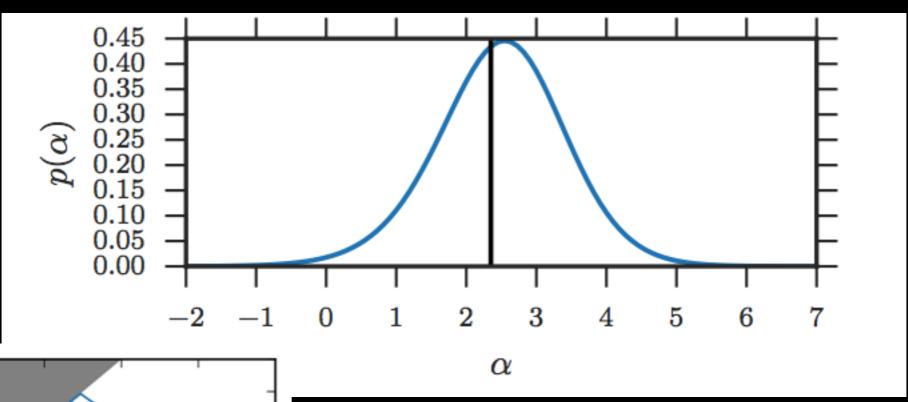
Mandel, Farr, Gair, in prep.

 [Counting and confusion — Farr, Gair, Mandel, Cutler, 2015]

#### Population Reconstruction



 $p(m_1) \propto m_1^{-\alpha}$  flat q



© 25 W) 20 

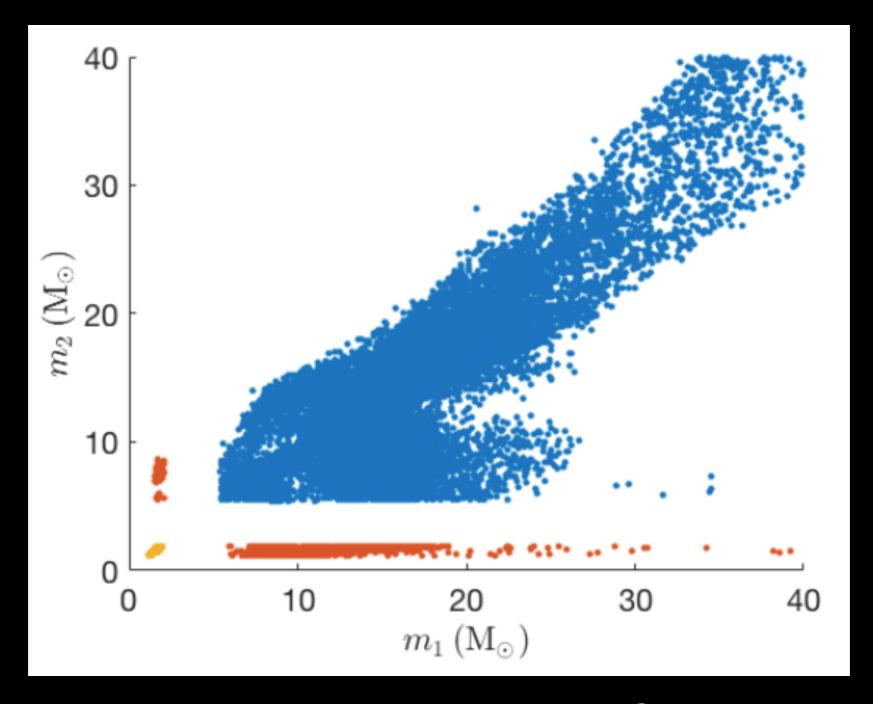
 $m_1^{
m source}({
m M}_{\odot})$ 

Abbott+ (LVC), 2016

$$h(f) \sim M^{5/6} => V \sim M^{2.5}$$

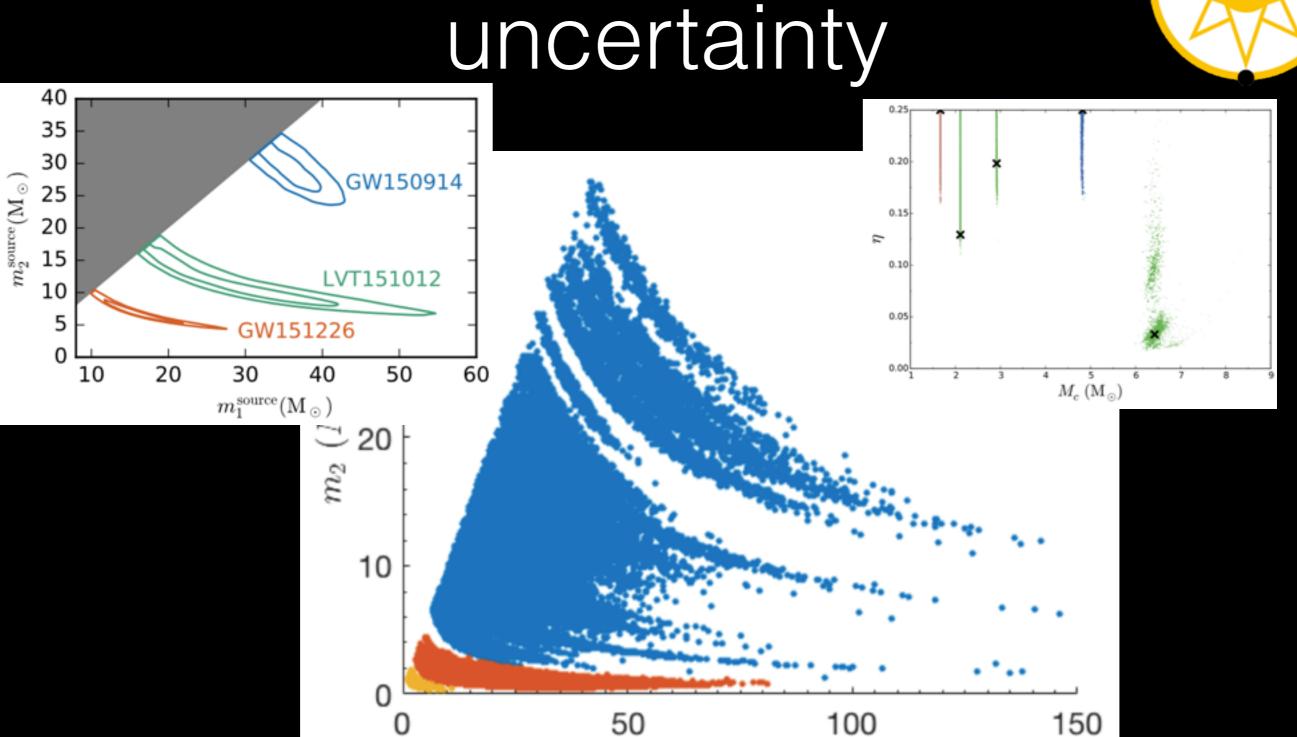
# Unmodeled Inference: Binary population clustering





[Mandel et al., 2015; Dominik et al., 2015; Stevenson et al., 2016]

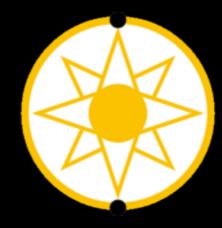
# Measurement uncertainty

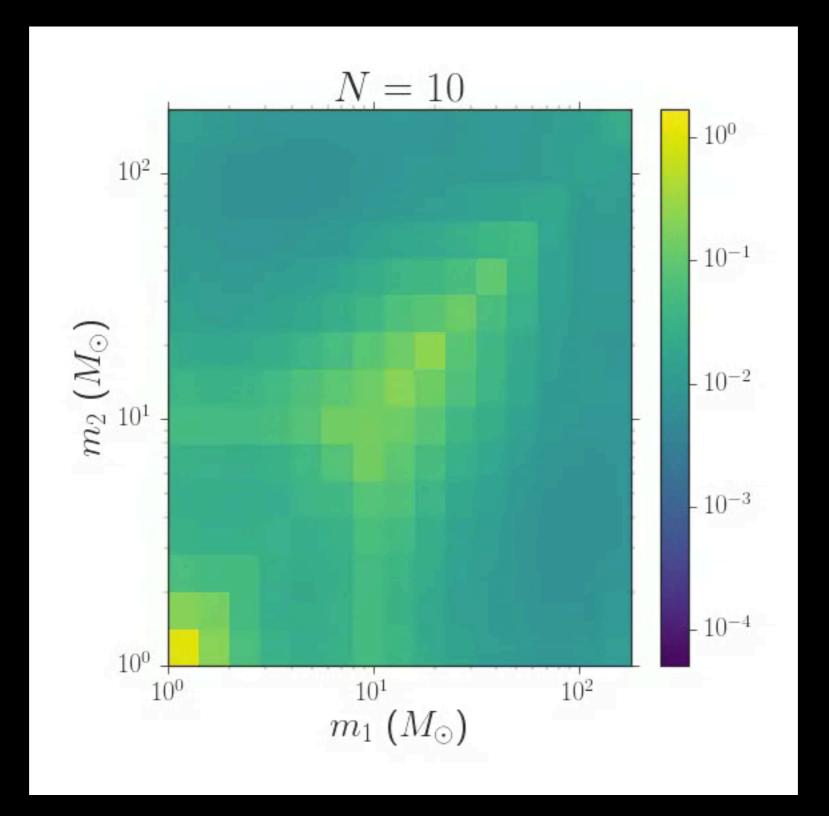


[Abbott et al., 2016; Mandel et al., 2015; see also Littenberg et al., 2015]

 $m_1 (M_{\odot})$ 

#### Unmodeled Inference: Binary population clustering

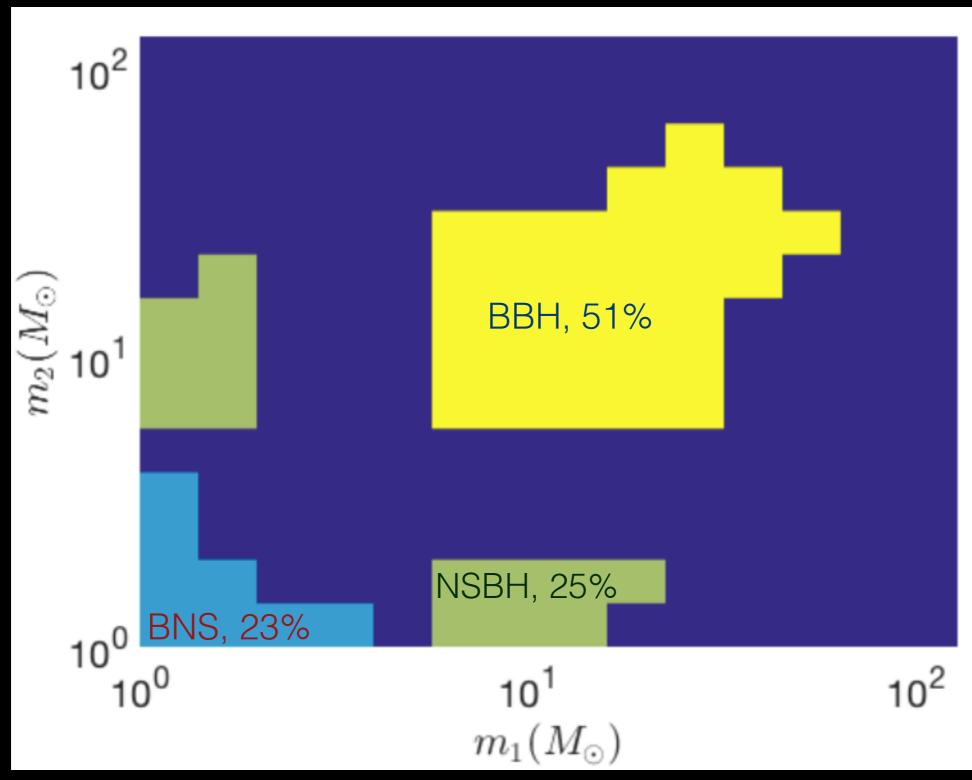




Mandel+, 2016

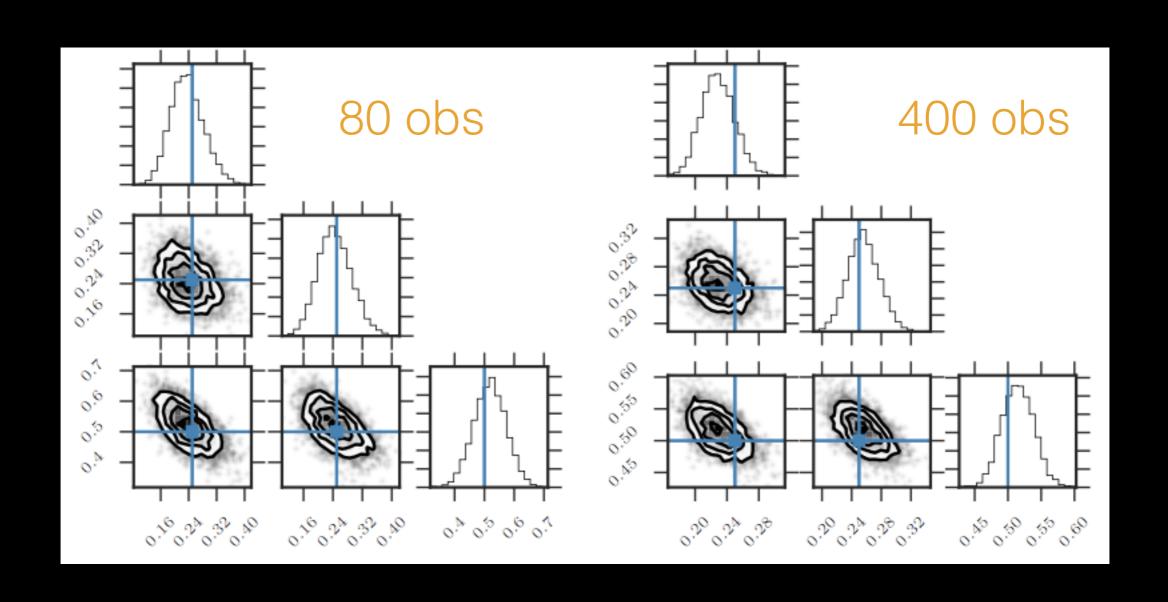
# Water filling on mean density





#### Unmodeled Inference





#### Future

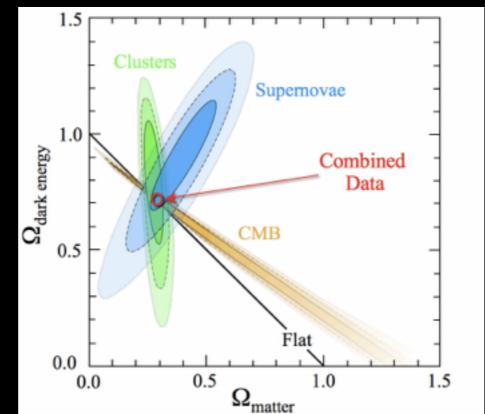


Bring together modelling and astrostatistics

 Figure out what questions we can realistically answer... and answer them!

Use full observation set — concordance binary

evolution?



THE GRAVITATIONAL WAVE
DETECTOR WORKS! FOR THE
FIRST TIME, WE CAN LISTEN
IN ON THE SIGNALS CARRIED
BY RIPPLES IN THE FABRIC
OF SPACE ITSELF!



EVENT: BLACK HOLE MERGER IN CARINA (30 M., 30 M.)

**EVENT:** ZORLAX THE MIGHTY WOULD LIKE TO CONNECT ON LINKEDIN

EVENT: BLACK HOLE MERGER IN ORION (20 Ma, 50 Ma)

**EVENT:** MORTGAGE OFFER FROM TRIANGULUM GALAXY

EVENT: ZORLAX THE MIGHTY WOULD LIKE TO CONNECT ON LINKEDIN

**EVENT:** MEET LONELY SINGLES IN THE LOCAL GROUP TONIGHT!



#### GW150914: our first Binary BH merger

-0.76s



