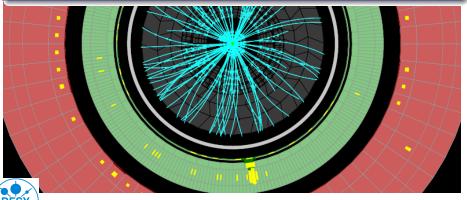
# Experimental particle physics at the LHC (3)

Kerstin Tackmann (DESY)



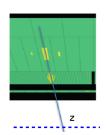
GRK1504/2: Autumn Block Course 2016

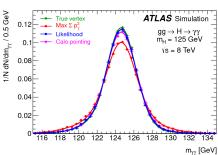
## Photon pointing and primary vertex selection (I)

$$m_{\gamma\gamma}^2 = 2E_1E_2(1-\cos\alpha)$$

Improve photon angle measurement using neural network (8 TeV) based on

- Photon pointing
  - Photon direction measured from calorimeter using longitudinal segmentation
  - Position of conversion vertex for converted photons (with Si hits)
- ullet  $\sum p_T^2$ ,  $\sum p_T$  (over tracks) and angular balance in  $\phi$  between tracks and diphoton system
- → Contribution of angle measurement to mass resolution negligible already without primary vertex information
- → Good primary vertex selection needed for selection of signal jets



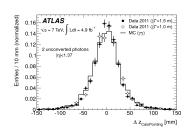


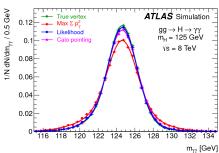
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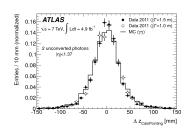


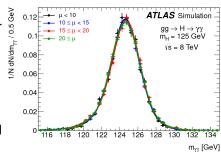
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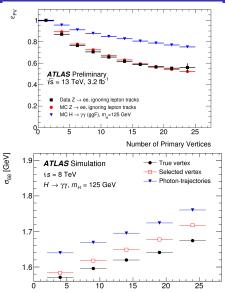
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# Photon pointing and primary vertex selection (II)

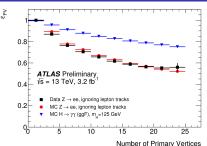
- ullet Efficiency of primary vertex selection can be measured with Z 
  ightarrow ee, disregarding the electron tracks
- Primary vertex selection more efficient in lower pileup
  - Fewer primary vertices with larger average spacing
- ullet Primary vertex selection more efficiency for events with higher Higgs  $p_T$ 
  - ★ More recoil track p<sub>T</sub>

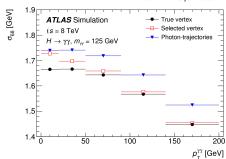


Number of primary vertices

### Photon pointing and primary vertex selection (II)

- ullet Efficiency of primary vertex selection can be measured with Z 
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  - $\star$  More recoil track  $p_T$





#### Invariant mass resolution – CMS vs ATLAS

#### Calorimeter resolution

 CMS crystal calorimeter with excellent intrinsic resolution

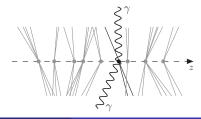
$$rac{\sigma_E}{E} = rac{2.8\%}{\sqrt{E}} \oplus rac{0.12}{E} \oplus 0.3\%$$
 vs ATLAS

$$rac{\sigma_E}{E} = rac{10\%}{\sqrt{E}} \oplus 0.7\%$$

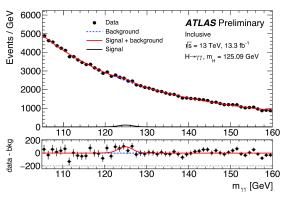
- ⇒ Narrower core of resolution function in CMS compared to ATLAS, e.g. best resolution event category
  - ★ CMS 1.18 GeV
  - \* ATLAS 1.39 GeV

#### Primary vertex selection

- ATLAS longitudinally segmented calorimeter allows for pileup-independent input to primary vertex selection
  - CMS primary vertex selection relies entirely on tracker
- ATLAS resolution function less affected by long non-Gaussian tails arising from wrong primary vertex choice



#### Invariant mass spectrum



Background+signal fit, signal constrained to 125.09 GeV

#### Diphoton selection

Identified and isolated photons

$$p_T^{\gamma 1} > 0.35 m_{\gamma \gamma} \ p_T^{\gamma 2} > 0.25 m_{\gamma \gamma}$$

$$p_T^{\gamma 2} > 0.25 m_{\gamma \gamma}$$

#### Interlude

# CMS took some data with $B=0\,\mathrm{T}$ in 2015. How are the various steps affected?

Note: this data was not used in any  $H \to \gamma \gamma$  analysis, but was used in the search for high-mass diphoton resonances with 2015 data

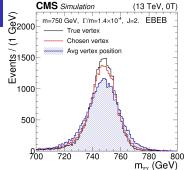
### Specifics for CMS 0 T dataset

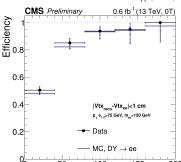
#### No energy spread from bremsstrahlung

- Dedicated calibration and photon identification
  - ★ 5-15% lower id efficiency
- Better intrinsic energy resolution
- Better representation of photons by electrons

#### No measurement of track momenta

- Track isolation relying on track counting instead of track momenta
- Primary vertex selection based on track counting (largest multiplicity)
  - ★ Reduced efficiency for identifying correct PV of 60% (instead of 90%)





I will discuss two measurements in the  $H \to \gamma \gamma$  decay channel in detail:

- Fiducial and differential cross section measurements
- "Coupling measurements": studies of the Higgs production processed

- I am showing plots from several different analyses, using different data sets, depending on which plot is available for which analysis/data set
  - ★ General principle is the same (details and results are of course different)

#### Cross section measurements

### Principle of a cross section measurement

#### Recall

$$\sigma = rac{N}{\int \mathcal{L} \mathrm{d}t} = rac{N_{\mathrm{meas}} - N_{\mathrm{bkgd}}}{\epsilon \cdot A \cdot \mathcal{B} \cdot \int \mathcal{L} \mathrm{d}t}$$

#### Experimental steps

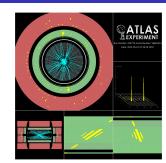
- Estimate and subtract the background(s)
- Correct for detector acceptance, and for efficiencies
- If needed/wanted, correct for branching ratio(s)
- Determine the luminosity (see earlier lecture)

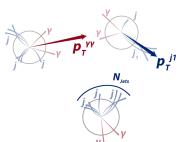
Differential cross section in variable x:  $\frac{d\sigma}{dx}$ 

- In practice: bin-averaged cross section  $\frac{\Delta \sigma}{\Delta x}$
- Background estimation and subtraction, efficiency and acceptance corrections performed for every bin
- Requires correction of resolution effects in x: unfolding

## Why cross section measurements for Higgs?

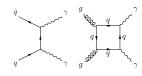
- Almost model-independent measurements of production and decay kinematics
- Measure kinematic distributions of Higgs, of associated jets, ...
- Sensitivity to Higgs production processes, QCD effects, CP, ...
- Measure inclusive cross section, and cross section in phase space enriched with VBF, and with a lepton
- $\bullet$  Differentially in  $p_T^{\gamma\gamma}, N_{\rm jet}, p_T^{\rm jet}, ...$
- $H o \gamma \gamma$  and  $H o 4\ell$  decays well suited thanks to good signal invariant mass resolution o comparably "simple" analyses



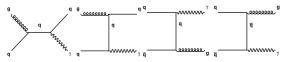


### Backgrounds

Irreducible backgrounds: events with two photons, e.g.



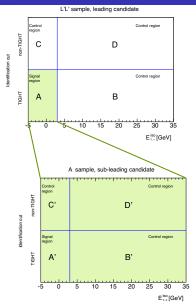
 Reducible backgrounds: events where at least one photon candidate is a misidentified jet, e.g.



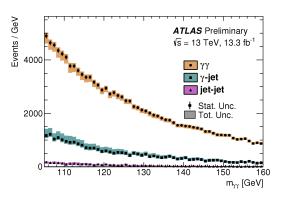
ullet Z 
ightarrow ee with the electrons misreconstructed as photons (mass tail reaches beyond  $m_Z=90\,{
m GeV})$ 

#### Understanding the backgrounds (I)

- Define control regions enriched in background
  - Photon candidates that fail a given set of the shower shape cuts and/or
  - \* Photon candidated that are less isolated
- Fit determines (given numbers of events in signal and control regions and photon identification and isolation efficiency)
  - Efficiencies for jet to pass photon identification and isolation for γjet and jetjet events, separately for higher and lower p<sub>T</sub> candidate
    - Correlation for both jets to pass isolation in jetjet events
  - $\star$  Number of  $\gamma\gamma$ ,  $\gamma$ jet, jet $\gamma$  and jetjet events
    - ightharpoonup Z 
      ightharpoonup ee included in  $\gamma\gamma$  as e look most like  $\gamma$  in id and isolation



### Understanding the backgrounds (II)



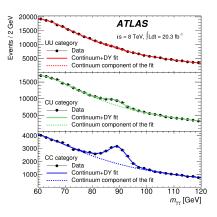
$$\begin{array}{ll} \gamma\gamma & (78.9 \pm 0.2^{+1.9}_{-4.0})\% \\ \gamma \text{jet} & (18.6 \pm 0.2^{+3.5}_{-1.7})\% \\ \text{jetjet} & (2.5 \pm 0.1^{+0.5}_{-0.4})\% \end{array}$$

Largest uncertainty: definition of control regions

- Study performed in every bin and every measured region of phase space
- Understanding of background composition not important directly to derive results, but for studies of background parametrization and photon identification

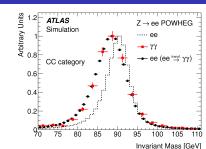
## Understanding the backgrounds (III)

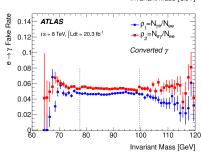
- Misreconstruction of electrons usually as single-track converted photons
- Suppressed by requiring no hit in innermost silicon layer for single-track conversions
- ullet Probability for an electron to be reconstructed as photon estimated using Z 
  ightarrow ee events
  - $\star$  Select " $Z o e \gamma$ " and Z o e e
- ullet For  $H o \gamma \gamma$ , only Z o ee tail relevant
- Not considered separately



## Understanding the backgrounds (III)

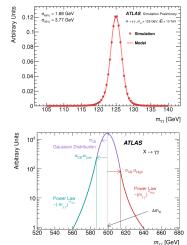
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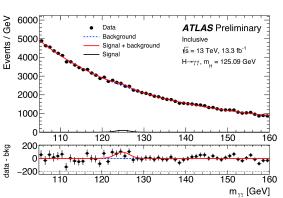


#### Parametrizing the signal

- ullet Signal is extracted by a signal+background fit to  $m_{\gamma\gamma}$  spectrum
- Signal is parametrized by a double-sided Crystal Ball function
  - \* Gaussian with exponential tail
- SM Higgs width  $4 \, \mathrm{MeV} \ (m_H = 125 \, \mathrm{GeV})$
- Parameters that determine the shape are determined on simulation
- Peak position (= Higgs mass) and Gaussian width (= detector resolution) constrained within uncertainties
  - $\star$  Energy scale and resolution, and  $m_H$
  - Peak position unconstrained for measurement of the Higgs mass
    - ➤ To be done with Run2 data once precision energy calibration achieved
  - \* Run1 Higgs mass measurement  $m_H = (125.09 \pm 0.21(\text{stat}) \pm 0.11(\text{syst})) \text{ GeV}$



# Parametrizing the backgrounds (I)



Background+signal fit, signal constrained to 125.09 GeV

Background modelled by smooth, monotonously falling function

- Polynomials (typically 3rd or 4th order)
- Exponentials of polynomials (typically 1st or 2nd order)

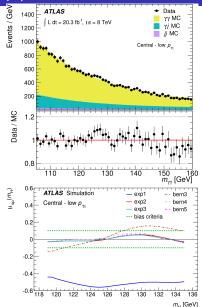
shape and normalization determined by the fit

Studied on high-statistics MC and chosen to give good statistical power while keeping potential biases acceptable

Potential bias accounted for as systematic uncertainty

### Parametrizing the backgrounds (II)

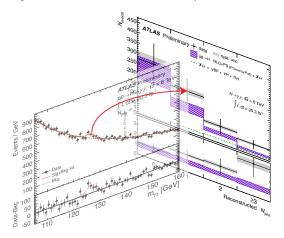
- S+B fits to high statistics background samples (→ "spurious signal")
- Choose the most performant parametrization that has small enough systematics:
  - $\begin{tabular}{ll} \bf O & Accept model if the spurious signal is \\ & < 10\% \ of expected signal or < 20\% \ of \\ & fitted signal uncertainty \\ \end{tabular}$
  - Among the models left, choose the one with the smallest number of free parameters
- Run1: γγ, γjet, jetjet simulation, detector effects included through weighting and smearing
- Run2: γjet, jetjet shapes from data control regions



# Signal+background fit

...carried out for...

- …all selected events → fiducial cross section
- ..after specific selections → fiducial cross section for that selection
- ullet ...in bins of a given variable o differential spectrum



## Signal+background fit

Likelihood function to be maximized

$$\mathcal{L} = \prod_{i} \left\{ \frac{\mathrm{e}^{-\nu_{i}}}{n_{i}!} \prod_{j}^{n_{i}} \left[ \nu_{i}^{\mathrm{sig}} \mathcal{F}_{i}^{\mathrm{sig}}(m_{\gamma\gamma}^{j}, \theta; m_{H}) + \nu_{i}^{\mathrm{bkg}} \mathcal{F}_{i}^{\mathrm{bkg}}(m_{\gamma\gamma}^{j}) \right] \right\} \times \prod_{l} G_{l}(\theta)$$

- for bin i and event j
- n<sub>i</sub> number of events in bin i
- $m{ ilde{
  u}}_i^{(
  m sig,bkg)}$  expected number of total/signal/background events
- $m{\circ}~ \mathcal{F}_i^{( ext{sig}, ext{bkg})}$  signal/background shape
- $m{ heta}$  nuisance parameters associated with systematic uncertainties, constraint via  $G_l(m{ heta})$
- ullet Energy scale and resolution uncertainties, and uncertainty on  $m_H$  correlated between all bins
  - → Nuisance parameters common between all bins

#### But wait a moment... is there a signal?

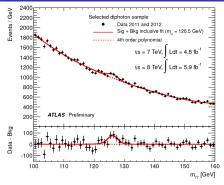
...back to summer 2012

★ Signal or statistical fluctuation of the background?

Compare compatibility of data with B-only and with S+B hypothesis with a signal scaling factor  $\mu$ 

Profile likelihood ratio

$$ilde{q}_{\mu} = -2 \ln rac{L( ext{data}|\mu,\hat{ heta}_{\mu})}{L( ext{data}|\hat{\mu},\hat{ heta})}$$



- Numerator and denominator are maximized independently
- $\hat{\theta}_{\mu}$  conditional maximum given  $\mu$ ;  $\hat{\mu}$ ,  $\hat{\theta}$  corresponding to global maximum of the likelihood
- ullet Large  $ilde{q}_{\mu}$  correspond to disagreement between data and hypothesis  $\mu$
- ullet  $ilde{q}_{\mu}$  behaves as  $\chi^2$  for large data samples and Gaussian heta
- Denominator is only normalization term, independent of  $\mu$

#### Frequentist limit setting procedure

ullet Construct likelihood function  $L(\mu, heta)$ 

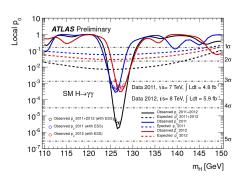
https://cds.cern.ch/record/1375842

- ullet Construct test statistics  $ilde{q}_{\mu}$
- ullet Perform fit to data and determine observed  $ilde{q}_{\mu,\mathrm{obs}}$  for hypothesis  $\mu$
- Generate pseudo MC to construct PDF  $p_{\mu}( ilde{q}_{\mu}|\mu,\hat{ heta}_{\mu,\mathrm{obs}})$  of  $ilde{q}_{\mu}$ 
  - $\star$  MC generation done with  $\hat{\theta}_{\mu, \mathrm{obs}}$ , but  $\hat{\theta}_{\mu}$  allowed to float in the fits
- Determine the observed p-value for hypothesis  $\mu$ :  $P(\mu) = \int_{\tilde{q}_{\mu} = 1}^{\infty} p_{\mu}(\tilde{q}_{\mu}|\mu, \hat{\theta}_{\mu, \text{obs}}) d\tilde{q}_{\mu}$
- Perform "discovery" test by computing  $P(\mu = 0)$
- ullet Find the 95% upper bound  $\mu=\mu_{95,
  m obs}$  for which  $P(\mu)=0.05$ 
  - \* To be conservative and to avoid that upward fluctuations of the background contribute to the p-value, LHC experiments compute upper limit from  $P_{\rm CL_s}(\mu) = P(\mu)/P(0) = 0.05$ 
    - CL<sub>s</sub> usually over-covers, so less than 5% of repeated experiments would lie outside the given bound

For complex fits pseudo-MC procedure can be very CPU intensive. Asymptotic formulae exist for cases with enough events. https://arxiv.org/abs/1007.1727

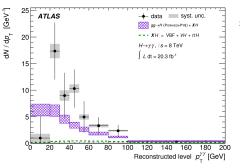
#### Testing background-only for $H o \gamma \gamma$ ICHEP 2012

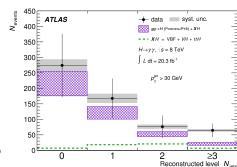
#### Maximum deviation from background-only expectation at $m_H=126.5\,{ m GeV}$



- Local significance 4.5  $\sigma$  (expected 2.4  $\sigma$ )
- ullet Global significance 3.6  $\sigma$
- Need to take into account "look-elsewhere effect": probability for a fluctuation somewhere in the studied mass range larger than for a given mass
- Require 5  $\sigma$  for discovery  $(p = 2.9 \cdot 10^{-7})$ 
  - $\star$  Reached at ICHEP in combination with  $H o ZZ^* o 4\ell$

#### Back to the measurements: Measured signal yield





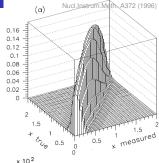
First part achieved:

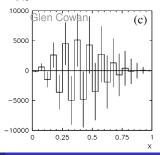
$$rac{\mathrm{d}\sigma}{\mathrm{d}x} = rac{N_{\mathrm{meas}} - N_{\mathrm{bkgd}}}{\epsilon \cdot A \cdot \mathcal{B} \cdot \mathrm{d}x \cdot \int \mathcal{L} \mathrm{d}t}$$

..although not quite...

#### Not mentioned so far: resolution corrections/unfolding

- $A_{ij}x_j = b_i$  (b measured, x true)
- Detector response matrix A encodes resolution (can also include efficiency and acceptance)
  - \*  $A_{ij}$  = Probability for event in true bin j to be reconstructed in reco bin i
  - \*  $A_{ij}$  is largely model independent, although there could be caveats in some cases
- "Naive" matrix inversion:  $x = A^{-1}b$ 
  - Unfolded spectrum x usually dominated by statistical fluctuations
    - Statistical fluctuations in measured spectrum get amplified
    - Nice explanation of this effect here
  - Unbiased estimator with smallest possible variance (typically see large negative correlations between adjacent bins)





#### Interlude

# **Unfolding**

Unfolding is a complicated business and one is well advised to ask in each problem if it can be avoided. (Glen Cowan)

Link to a (very readable) survey of unfolding methods from Glen Cowan here

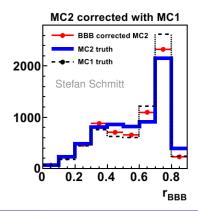
### Bin-by-bin correction (the simplest method)

- Used for  $H o \gamma \gamma$  differential cross section measurements
- Each bin is corrected with multiplicative factor,  $x_i = b_i/c_i$ , with:

$$c_i = rac{N_i^{
m reco}}{N_i^{
m true}} \quad \epsilon_i = rac{N_i^{
m true+reco}}{N_i^{
m true}} \quad p_i = rac{N_i^{
m true+reco}}{N_i^{
m reco}}$$

with  $\epsilon_i$  efficiency and  $p_i$  purity

- Pro: simple and robust
- Con: prone to produce biased results
  - $\star$  Data biased towards the MC used to compute the  $c_i$
  - → Carefully evaluate uncertainties!
- My personal take: ok to be used as long as biases are estimated properly and small compared to e.g. statistical uncertainties



### Regularization (I)

- Regularization dampens the oscillations by suppressing insignificant bins in the data distribution and detector response matrix
- In simplified form, can write unfolding problem as minimization of

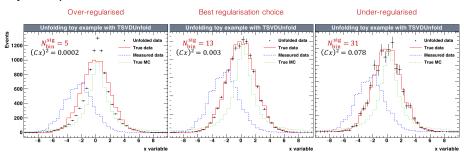
$$\boldsymbol{\chi}^{2}(\boldsymbol{x}^{\mathrm{data}}) = (\boldsymbol{A}^{\mathrm{MC}}\boldsymbol{x}^{\mathrm{data}} - \boldsymbol{b}^{\mathrm{data}})^{T}(\boldsymbol{A}^{\mathrm{MC}}\boldsymbol{x}^{\mathrm{data}} - \boldsymbol{b}^{\mathrm{data}}) + \boldsymbol{\tau}(\boldsymbol{C}\boldsymbol{x}^{\mathrm{data}})^{T}(\boldsymbol{C}\boldsymbol{x}^{\mathrm{data}})$$

where  $x^{\rm data}$  the unfolded data distribution, C is a matrix and  $Cx^{\rm data}$  e.g. the sum of squared of the 2nd derivative of  $x^{\rm data}$  (other choices possible)

- ullet au=0 corresponds to exact matrix inversion, au
  eq 0 regularizes inversion
  - $\star \tau$  too small  $\rightarrow$  oscillations
  - $\star$  au too large o unfolded spectrum biased towards MC used for unfolding
  - $\star$  Good size of au depends on statistics, binning, resolution, how well the MC describes the data, ... and needs to be carefully determined for every case

### Regularization (II)

#### Toy example



- ullet au=0 corresponds to exact matrix inversion, au
  eq 0 regularizes inversion
  - $\star$   $\tau$  too small  $\rightarrow$  oscillations
  - $\star$  au too large o unfolded spectrum biased towards MC used for unfolding
  - $\star$  Good size of au depends on statistics, binning, resolution, how well the MC describes the data, ... and needs to be carefully determined for every case

### Fit-based and SVD unfolding

- Solve unfolding problem by minimizing  $\chi^2$
- SVD: Based on singular value decomposition of detector response matrix (here)
  - Allows association of "physics" with large singular value contributions and "statistical fluctuations" with small singular value contributions
  - \* Regularization through suppression of small singular value constributions ("beyond *k*th singular value")

- Pro: in principle always possible to find decent working point (might require adaptation of discrete k choice)
- Con: curvature regularization iffy for spectra with sharp features (e.g. narrow peaks)

### Iterative unfolding

Iterative Bayesian unfolding (D'Agostini's method, here) is widely used

Employs Bayes theorem to infer the unfolding matrix

 $P( ext{true bin } j | ext{reco bin } i) \propto P( ext{reco bin } i | ext{true bin } j) P( ext{true bin } j)$  and iteratively determines

$$x_j = rac{1}{arepsilon_j} \sum_{i=1}^N P( ext{true bin } j | ext{reco bin } i) b_i$$

starting with an initial estimate for  $P(\operatorname{true} \operatorname{bin} j)$  (e.g. from MC)

- Effectively regularized by using a finite number of iterations
  - \* Small number of iterations → biased towards input MC
  - ★ Large number of iterations → large variance
- Pros: able to handle high-dimensional problems (e.g. picture deblurring), used in SM for 3d unfolding, some adaptation of "MC" to data (if enough iterations possible)
- Cons: iteration might not be possible with low statistics (seen in  $H \to \gamma \gamma$ ), can be problematic if there is a badly modelled quantity that is not measured but averaged over (no help from iteration)

### Back to the analysis: efficiency corrections

Efficiency of the reconstruction and selection

$$\epsilon = \frac{\text{Number of events reconstructed and selected}}{\text{Number of signal events in the kinematic range}}$$

Main contributions to inefficiencies in  $H o \gamma \gamma$ 

- photon identification
- photon isolation
- diphoton trigger

Efficiencies are measured in control samples

- Sometimes, efficiencies are determined from simulations
  - ★ Requires good simulation of detector and/or physics process

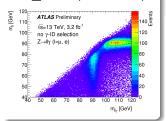
### Photon id efficiency measurements

Id efficiency for isolated photons:  $E_T^{iso}$  <4 GeV

Radiative Z decays:  $Z o \ell\ell\gamma$ 

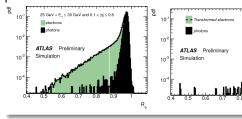
 $E_T^{\gamma}$  of 10-80 GeV Photon purity

- ~ 90% (10-15 GeV)
- > 98% (> 15 GeV)

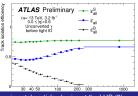


#### Z ightarrow ee tag-and-probe

+ transformation of electron showers to resemble photon showers



#### "Matrix method"



Purity determination from track isolation before and after id → id efficiency

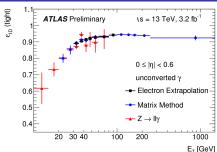
#### Photon id efficiency measurements

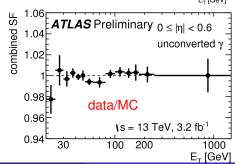
- ullet Partial overlap in  $E_T$  regions covered by the different methods
- Combination of measurements in overlap regions
  - ullet 1-2% uncertainties for  $E_T <$  40 GeV, 0.5-1% above 40 GeV

Uncertainty on  $H o \gamma \gamma$  signal yield

ICHEP 2012	10.8%
Dec 2012	5.3%
Moriond 2013	2.4%
ICHEP 2014	1%

Second-largest experimental uncertainty on  $H \rightarrow \gamma \gamma$  signal strength (final Run1 paper)





#### Acceptance corrections (I)

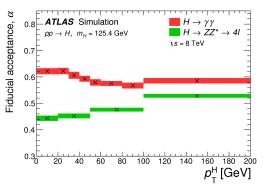
Acceptance of the kinematic selection

 $A = \frac{\text{Number of signal events in the kinematic range}}{\text{Number of all signal events}}$ 

- Experimentally accessible kinematic region is limited
  - $\star$  Small  $E_T$  photons not used due to large backgrounds
  - $\star$  Detector acceptance limited in  $\eta$
- Need to use theoretical predictions to extrapolate
  - ★ Usually in the form of simulations
  - ★ Introduced dependence on theoretical predictions and their uncertainties
- Unfold to a fiducial region defined by photons (and jets) to minimize acceptance corrections
  - \*  $p_T^{\gamma 1 (\gamma 2)} > 0.35 (0.25) m_{\gamma \gamma}, \quad |\eta^{\gamma 1, 2}| < 2.37$
  - \*  $p_T^{\rm iso} < 0.05~p_T^{\gamma}$  with  $p_T^{\rm iso} \sum p_T$  of all charged particles with  $p_T >$  1 GeV within  $\Delta R = 0.2$  around photon
  - $\star~p_T^j >$  30 GeV,  $|y^j| <$  4.4

#### Acceptance corrections (II)

Correcting from fiducial region to the full phase space would be a sizeable correction



- ...of course this means that theoretical predictions will have to be done for the same fiducial region
- where not available (yet), correction factors are derived from simulation