Axion: Mass Dark Matter Abundance Relation

Guy Moore, TU Darmstadt

- Mystery 1: Dark Matter
- Mystery 2: T-symmetry of QED and QCD
- The Axion: a solution to both mysteries?
- Early Universe cosmology of the axion
- How to predict the axion mass if it's the dark matter

DESY Hamburg, 18 Jan. 2017 Folie 1 von 37

"Why move from Montréal to Darmstadt?"







DESY Hamburg, 18 Jan. 2017 Folie 2 von 37

Dark Matter: a Cosmic Mystery



Atoms: Standard Model.
Dark Energy: Cosmological Constant.
Strange value, but possible
Dark Matter: MYSTERY! NOT SM!

We only know 3 things about dark matter:

- It's Matter: gravitationally clumps.
- It's **Dark**: negligible electric charge, interactions too feeble to be detected except by gravity
- It's **Cold**: negligible pressure by redshift z = 3000

DESY Hamburg, 18 Jan. 2017 Folie 3 von 37

Another mystery: T-symmetry in QED and QCD

T symmetry: "when you run a movie backwards, the *microphysics* is correct."

Statistical mechanics breaks T.

But microphysics very nearly obeys it!

Weak physics breaks \mathbf{T} , but only through very small CKM effects. Observed in handful of experiments, all involving neutral meson oscillation.

No evidence for T viol in E&M or Strong interactions.

T in E&M

How do E, B fields change when you run movie backwards?



Q's unchanged, but J's flip. E same, but B flips.

DESY Hamburg, 18 Jan. 2017 Folie 5 von 37

Looking for T: Neutron EDM

Put neutron in \vec{B} field – spin lines up with \vec{B} .



Is there an Electric Dipole Moment (EDM) aligned with spin? If so: looks different when movie runs backwards, \mathbf{T} viol!

DESY Hamburg, 18 Jan. 2017 Folie 6 von 37

${f T}$ and the E&M Action

Action $S \Rightarrow$ all physics. Local field thy: $S = \int \mathcal{L} d^4 x$. \mathcal{L} a singlet (gauge symm) and spacetime scalar (Lorentz):

$$\mathcal{L} = \frac{\vec{B}^2 - \vec{E}^2}{2e^2} + \frac{\Theta}{4\pi^2}\vec{E}\cdot\vec{B} + (\text{electrons...})$$

T flip: $\vec{E} \to \vec{E}$ and $\vec{B} \to -\vec{B}$: $(B^2 - E^2) \to (B^2 - E^2)$ **BUT** $E \cdot B \to -E \cdot B$.

$$\mathcal{L} \xrightarrow{T} \frac{\vec{B}^2 - \vec{E}^2}{2e^2} - \frac{\Theta}{4\pi^2}\vec{E}\cdot\vec{B} + (\text{electrons...})$$

Nonvanishing Θ is a **T** violation!

DESY Hamburg, 18 Jan. 2017 Folie 7 von 37

E&M T violation is Illusory!

The $\Theta \vec{E} \cdot \vec{B}$ term has no *consequences*!

$$\vec{E} \cdot \vec{B} = \frac{1}{4} \epsilon_{\mu\nu\alpha\beta} F^{\mu\nu} F^{\alpha\beta} = \partial^{\mu} K_{\mu} , \quad K^{\mu} \equiv \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} A^{\nu} F^{\alpha\beta}$$

I can integrate $\vec{E} \cdot \vec{B}$ to a boundary term. Vanishes if $F^{\alpha\beta}$ vanishes on boundary. Alternately, EOM:

$$0 = \partial_{\mu} \left(\frac{1}{e^2} F^{\mu\nu} + \frac{\Theta}{8\pi^2} \epsilon_{\mu\nu\alpha\beta} \partial^{\alpha} A^{\beta} \right)$$

Second term zero by antisymmetry (if Θ constant)

DESY Hamburg, 18 Jan. 2017 Folie 8 von 37

QCD and its Lagrangian

QCD is like 8 copies of E&M, but with non-linearities:

 $\label{eq:Field strength} {\sf Field strength}: \quad F^{\mu\nu}_a = \partial^\mu A^\nu_a - \partial^\nu A^\mu_a + g f_{abc} A^\mu_b A^\nu_c \,,$

g: coupling. $a = 1 \dots 8$. f_{abc} "structure constants"

$$S = \int dt \int d^3x \, \left(\frac{\vec{E}_a^2 - \vec{B}_a^2}{2g^2} + \frac{\Theta}{8\pi^2} \vec{E}_a \cdot \vec{B}_a \right)$$

where $\vec{E}_a \cdot \vec{B}_a$ still a total derivative:

$$\vec{E}_a \cdot \vec{B}_a = \partial^{\mu} K_{\mu} , \qquad 2K_{\mu} = \epsilon_{\mu\nu\alpha\beta} \left(A^{\nu}_a F^{\alpha\beta}_a + \frac{gf_{abc}}{3} A^{\nu}_a A^{\alpha}_b A^{\beta}_c \right)$$

Last term need not vanish on boundary even if $\vec{E}_a = 0 = \vec{B}_a$ there! It's always $8\pi^2 N_I$ with N_I integer. So $\Theta \mod 2\pi$ has physical consequences G. 't Hooft, PRL 37, 8(1976); R. Jackiw and C. Rebbi, PRL 37, 172 (1976);

Callan Dashen and Gross, Phys Lett 63B, 334 (1976)

DESY Hamburg, 18 Jan. 2017 Folie 9 von 37

Theory: Neutron electric dipole moment should exist,

$$d_n = -3.8 \times 10^{-16} \, e \, \mathrm{cm} \times \Theta$$

SO long as Θ is not zero! Guo *et al*, arXiv:1502.02295, assumes Θ , modulo 2π , is small

Experiment: Consistent with zero! Baker et al (Grenoble), arXiv:hep-ex/0602020

 $|d_n| < 2.9 \times 10^{-26} \ e \ \mathrm{cm}$

Either $|\Theta| < 10^{-10}$ by (coincidence? accident?) or there is something deep going on here.

DESY Hamburg, 18 Jan. 2017 Folie 10 von 37

Why an accident seems unlikely

- **T** is not a fundamental symmetry. We have observed its violation (Kaon, B-meson, possibly D-meson physics)
- More **T** violation almost surely out there otherwise, tough to explain why Universe is filled with matter!
- Renormalization: T viol. one place finds its way into other places, including Θ. Θ does *not get smaller* as you go from high to low scales R.G. marginal feels T viol from all scales.

Axion: an explanation for $\Theta = 0$

Hypothesize an extra *complex scalar* field $\varphi = \phi e^{i\theta_A}$ Assume a *symmetry*: $\varphi \to e^{i\theta}\varphi \ [\theta_A = \theta_A + \theta]$. Lagrangian:

$$\mathcal{L} = g^{\mu\nu}\partial_{\mu}\varphi^*\partial_{\nu}\varphi + \lambda\left(\varphi^*\varphi - f_a^2/2\right)^2 (+\text{interactions})$$

Spontaneous symmetry breaking with VEV f_a . We will want $f_a \sim 10^{11} \text{GeV}$.



Peccei Quinn PRL 38, 1440 (1977);

J. E. Kim, PRL 43, 103 (1979); Shifman Vainshtein and Zakharov, NPB 166, 493 (1980)

DESY Hamburg, 18 Jan. 2017 Folie 12 von 37

Axion: Indirect QCD coupling

φ is singlet under QCD.

But add some heavy physics which couples it indirectly (such as heavy quarks which get mass due to $|\varphi|$)

Phase of $\varphi = f_a e^{i\theta_A}$ turns into part of $\overline{\Theta}$:

$$\mathcal{L} = \ldots + \frac{\Theta + \theta_A}{8\pi^2} \vec{E}_a \cdot \vec{B}_a$$

QCD cares about $(\Theta + \theta_A)$, not Θ alone.

DESY Hamburg, 18 Jan. 2017 Folie 13 von 37

Now the fun part

 θ_A is a *field* and can change value.

QCD cares about value of $\Theta + \theta_A$. Detailed **QCD** calculation *Topological Susceptibility of* **QCD**: Potential tilted



Energy lowest where $\Theta + \theta_A = 0$, where **T** conserved. Tilt is temperature dependent, smaller in very early Universe

Size of tilt: nonperturbative QCD. Solved for temp=0. Not for high temperature!

DESY Hamburg, 18 Jan. 2017 Folie 14 von 37

What is predicted:

- Neutron EDM: 10^{-5} smaller than current limit
- New particle, Axion, exists, with specific interactions
- Strength of axion interactions scale with its mass (Pseudo-Nambu-Goldstone Boson)

What is not predicted: Axion Mass!



The larger ϕ , the flatter the potential, and lighter the axion. ϕ not predicted by model.

DESY Hamburg, 18 Jan. 2017 Folie 15 von 37

Axion in cosmology

Assume first that φ starts homogeneous [inflation]

Classical axion field! Starts oscillating around $t = \pi m_a^{-1}$. Damped:

- Hubble drag
- effect of dm_a/dt

Acts Like Dark Matter!

Modern axion energy density larger if field starts oscillating later: $\varepsilon_{\rm axion} \propto f_a^{0.84} \propto m_a^{-0.84}$.

DESY Hamburg, 18 Jan. 2017 Folie 16 von 37

Initial state of φ field?

Maybe it's the same everywhere in space.

But most likely, it's randomly different in different places!

- Inflation stretches quantum fluctuations to classical ones: $\Delta \varphi \sim H_{\text{infl.}}$. If $N_{\text{efolds}}H^2 > f_a^2$, scambles field. If not: need $H < 10^{-5} f_a$ to avoid excess "isocurvature" fluctuations in axion field
- Gets scrambled *after* inflation if Universe was ever really hot $T > f_a \sim 10^{11}$ GeV.

Should be able to predict DM abundance from m_a . And that's valuable, also for the experimentalists!

DESY Hamburg, 18 Jan. 2017 Folie 17 von 37

L. Visinelli and P. Gondolo, PRL 103, 011802 (2014)

My Goal

Predict relation between **Dark Matter Density** and **Axion Mass** assuming space-random starting angle. Challenges:

- 1. $\chi(T):$ needs Lattice Gauge Thy. Open, subdominant problem
- 2. Axion field dynamics: classical but with large scale hierarchy $f_a/H\sim 10^{30}\gg 1$

I will explain 2., and show how to solve it.

Solving space-inhomogeneous case

Put the Lagrangian

$$\mathcal{L} = \partial_{\mu}\varphi^*\partial^{\mu}\varphi + \frac{\lambda}{8}(2-\varphi^*\varphi)^2 - \chi(t)\operatorname{Re}\varphi$$

as classical field thy. on **real-time lattice**, with random init. condit. Plus Hubble drag Nonperturbative approach. Easy enough to do!



DESY Hamburg, 18 Jan. 2017 Folie 19 von 37

What do you see?

Energy density has string-like maxima, complex dynamics:

 ε has stringlike structures: Annihilate when V tilts:

Axions and Topology I

φ is a complex number – plot as a 2D arrow. Axion field: a field of arrows. 2D slice for instance:

**************************** ************************* R 7 1 7 7 R KKKKKKKKKKKKKKKKK **ĸĸĸĸĸĸĸ**ĸ ********** KKKKKKKKKKKK KKKKKKKKKKKK ********* 661 2. アクトレストレイト *************** 11 11 KKFFFFFFFFFFFFFFFFFFFFFF アアア アクトクショ

Strings are vortices in the field.

DESY Hamburg, 18 Jan. 2017 Folie 21 von 37

Axions and Topology II

As you circle vortex, angle θ_A varies by 2π .



Continuity: angle θ_A must be undefined somewhere inside the circle. $\varphi = 0$ somewhere. Center of vortex.

DESY Hamburg, 18 Jan. 2017 Folie 22 von 37

Axions, Topology III

Singularity must exist somewhere inside such a circle



in 3D – true of *any* surface bounded by ring. Requires a *string* of singular points – a vortex line

DESY Hamburg, 18 Jan. 2017 Folie 23 von 37

Strings and Walls

Random starting conditions \Rightarrow network of these "strings" Network evolves, strings straighten out, find each other.

Potential tilts: $\theta_A \rightarrow 0$ where it can. Near string: explores all values. String has "domain wall" attached.



Pulls strings towards each other, speeds breakup of network. Complex string – domain wall network dynamics.

This dynamics produces Axions!

DESY Hamburg, 18 Jan. 2017 Folie 24 von 37

Conventional Wisdom

This dynamics produces axions! 3 sources:

- "misalignment" axions from $\theta_A \neq 0$ between strings
- "Wall" axions from breakup of axionic walls
- "String" axions radiated off of axionic strings

Axion production should be (much?) larger than angle-average of misalignment mechanism

Big complication

These strings are like onions (or scallions)



Layers! Innermost "core" has radius $r \sim f_a^{-1} \sim 10^{-27}$ m. Outer size $\sim H^{-1} \sim 10^2$ m. Energy stored logarithmically with radius $\varepsilon \propto \int \frac{rdr}{r^2}$. Name $\ln(f_a/H) \equiv \kappa \sim 70$.

DESY Hamburg, 18 Jan. 2017 Folie 26 von 37



Series of "sheaths" around string: equal energy in each $\times 2$ scale, 10^{30} scale range!

More "sheaths" \rightarrow outermost layer less important:

- inter-string interactions less important
- Radiation from strings less efficient

Both "thin out" string network. Big scale separation \rightarrow dense network

DESY Hamburg, 18 Jan. 2017 Folie 27 von 37

Does it matter?



String density varies strongly with # of layers. So does energy of string. Suggestive – axion production larger as more layers.

DESY Hamburg, 18 Jan. 2017 Folie 28 von 37

Wild extrapolation

Problem is – can I really believe an extrapolation over this large a range of scales, based on the information I have?



Need some way to capture string-core physics!

DESY Hamburg, 18 Jan. 2017 Folie 29 von 37

An effective description

The important axion-production physics is:

- Only long-range (light) degree of freedom is axion
- Axion strings: thin cores with high tension $T\simeq 70\times \pi f_a^2$
- Interaction between axion field and string is correct

Any modified UV (high-mass) physics which does this is OK!

Find massive fields which somehow increase string tension

DESY Hamburg, 18 Jan. 2017 Folie 30 von 37

Global vs Local Strings

Another theory with topological string objects: Abelian Higgs model

$$\mathcal{L}(\varphi, A_{\mu}) = \frac{1}{4} (\partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu})^2 + (D_{\mu} \varphi)^* (D^{\mu} \varphi) + \frac{\lambda}{8} \left(2\varphi^* \varphi - f_a^2 \right)^2$$

with $D_{\mu} = \partial_{\mu} - ieA_{\mu}$ covariant derivative

String where $\operatorname{Arg} \varphi$ winds by 2π as before. But now, \vec{B} -field flux in string core

$$A_{\phi} = rac{g(r)}{er}$$
 compensates so $D_{\mu}\varphi \rightarrow 0 (r \gg 1/ef_a)$

DESY Hamburg, 18 Jan. 2017 Folie 31 von 37

Abelian Higgs model

Strings carry energy in core: $\int B^2 + |D\varphi|^2 + V(\varphi)$ Energy density falls exponentially away from string core

Away from string core, only massive (heavy) fields:

$$m_{\varphi}^2 \sim \lambda f_a^2$$
, $m_A^2 = e^2 f_a^2$.

Network of strings with tension $T \simeq \pi f_a^2$.

No long-range interactions between strings. "String-only" Much (4× to 10×) higher string network density

DESY Hamburg, 18 Jan. 2017 Folie 32 von 37

Trick: global strings, local cores

Hybrid theory with A_{μ} and two scalars

$$\mathcal{L}(\varphi_{1},\varphi_{2},A_{\mu}) = \frac{1}{4}(\partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu})^{2} \\ + \frac{\lambda}{8}\left[(2\varphi_{1}^{*}\varphi_{1} - f_{a}^{2})^{2} + (2\varphi_{2}^{*}\varphi_{2} - f_{a}^{2})^{2}\right] \\ + |(\partial_{\mu} - iq_{1}eA_{\mu})\varphi_{1}|^{2} + |(\partial_{\mu} - iq_{2}eA_{\mu})\varphi_{2}|^{2}$$

Pick $q_1 \neq q_2$.

Two rotation symmetries, $\varphi_1 \rightarrow e^{i\theta_1}\varphi_1$, $\varphi_2 \rightarrow e^{i\theta_2}\varphi_2$ One combination gauged, one combination global (Axion)

DESY Hamburg, 18 Jan. 2017 Folie 33 von 37

For almost-equal charges $(q_1 = 3, q_2 = 4)$, initial $\varphi_1 = \varphi_2$, all strings have same φ_1 , φ_2 winding.

 φ_1 wants $B = 2\pi/eq_1$, φ_2 wants $B = 2\pi/eq_2$. Solution: $B = 2\pi(q_1 + q_2)/e(q_1^2 + q_2^2)$. String has large $2\pi f_a^2$ tension of gauge-string Plus long-range interactions $\propto (q_2 - q_1)^2/(q_2^2 + q_1^2)$

Behaves like axion-string with weak coupling to axion field. Or, with very large string-core tension

String cores modified. But far from cores – theory of axions only

DESY Hamburg, 18 Jan. 2017 Folie 34 von 37

String networks have larger tension

Strings denser, last longer, decay to more persistent loops

DESY Hamburg, 18 Jan. 2017 Folie 35 von 37

Results (preliminary)

- More, higher-tension, denser string network, but
- only 20% more axions than with axion-only simulation,
- Fewer axions than angle-averaged misalignment
- Usual argument **wrong**, involves **double-counting**
- energy in "walls" is *part* of misalignment energy
- Walls get "eaten" by strings, energy lost for axions

Conclusions

- Dark matter and ${\bf T}$ in QCD both mysteries
- Axion could explain both!
- Axion dynamics in early Universe: string defects
- Need auxiliary DOF to get high string tension.

Put in known dark matter density, vacuum $\chi(0)$ Recent lattice result for $\chi(T)$, these results: **Preliminary** result: $m_{\rm ax} \simeq () \mu eV$

DESY Hamburg, 18 Jan. 2017 Folie 37 von 37

Why is $\chi(T \sim 1 \text{GeV})$ Hard?

Lattice Monte-Carlo, find fraction of configs with instanton. But

- Instantons get rare fast, $\chi \sim T^{-7...-8}$. Statistics??
- Slowing-down of algorithm to change N_I
- Instanton counting must have very low false-positive rate!

Chiral limit *not* a problem: $\chi \propto (m_u m_d m_s/T^3)$, if $m \ll T$. Use $T \gg m \gg m_u$, multiply by $m_u m_d m_s/m^3$.

How do experimentalists look for Axions?

Most sensitive method: resonant cavity in magnetic field



Microwave cavity inside superconducting solenoid. \vec{E} field of cavity mode aligned with \vec{B} of solenoid. Cavity oscillation: oscillating $\vec{E} \cdot \vec{B}$. If cavity resonance matches m_A/\hbar : cavity resonance driven. Squid readout – tuneable cavity ...

What about Anthropic Principle?

Trendy Explanation for "coincidences" or "tunings"

Why is Cosmological Constant so small? If it were 100 times bigger, matter would fly apart or collapse before life could evolve. Nature plays dice, universes with all values occur, but only universes with life get observed.

Why does QCD respect T symmetry? If QCD violated T, something would go wrong with nuclear physics, which would make life impossible. Nature plays dice, only universes where life is possible get observed. Except that life is fine in a world where $\Theta = 10^{-2}$!

DESY Hamburg, 18 Jan. 2017 Folie 40 von 37