

31 October 2016 |। Fellows Day, DESY Hamburg


## The Flavour Problem: masses



- Quark masses generically hierarchical
- Charged lepton masses generically hierarchical
- Absolute neutrino mass not yet known, only mass-squared differences up to a sign


## The Flavour Problem: mixings

- Mismatch between flavour and mass bases leads to $3 \times 3$ unitary mixing matrices

Quarks


## Leptons

- Non-hierarchical, but:
- 'bi-maximal' (?)
- 'tri-maximal' (?)
- Non-zero 13 element

$$
\left|U_{\text {PMNS }}\right|\left(\begin{array}{ll}
\left(\begin{array}{l}
0.845 \\
0.791
\end{array}\right. \\
\left(\begin{array}{l}
0.521 \\
0.254
\end{array}\right. & \left(\begin{array}{l}
0.592 \\
0.512 \\
0.698 \\
0.521 \\
0.521 \\
0.254
\end{array}\right)
\end{array}\left(\begin{array}{l}
\binom{0.172}{0.133} \\
0.698 \\
0.455
\end{array}\right)\left(\begin{array}{l}
\left(\begin{array}{l}
0.782 \\
0.604 \\
0.782 \\
0.604
\end{array}\right)
\end{array}\right)\right.
$$

## Symmetries as solutions


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DISCRETE 2014

## Discrete flavour symmetries

- The data (arguably) indicate some ordering to flavoured parameters-new flavour symmetries can provide for such organization.
- Discrete symmetries (imposed via finite groups) have been favored candidates, especially in the leptonic sector.
- Such discrete symmetries can quantize precise mixing patterns and provide interesting relations amongst masses.
- Furthermore, breaking discrete symmetries does not necessitate goldstone modes that could spoil phenomenology, and vacuum alignment can also be achieved.
- Discrete symmetries can also be embedded into Grand Unified Theories, and could have origins in extra dimensions, e.g. heterotic orbifold compactifications, thus naturally connecting them to UV complete theories


## Model-independent symmetry searches



## Status of discrete flavour symmetries?

- Multiple symmetries predict the same mixing patterns, and the same symmetry can predict multiple patterns
- In the absence of an exact symmetry, sub-leading corrections become important for phenomenology.
- It is not presently clear that any discrete symmetry can, without special modeling, successfully describe all fermionic structure.
- Vacuum alignment mechanisms are often involved, and additional symmetries often needed.
- It is also not yet clear how such models should be completed / realized in the UV.

Input is needed from UV physics. Guideposts could come from:
Renormalization Group Evolution
Anomaly cancellation constraints
Higher dimensional theories

## Projects, ideas, and interests

## Generalized anomaly constraints w/ Sven Krippendorf (Oxford)

## Indirect model for quarks and leptons -

 w/ GG Ross (Oxford)Can the RGE for mass and mixing parameters be generalized with an EFT approach?
Are there alternative mechanisms/ constraints for flavoured vacuum alignment?
What are the connections between flavour and cosmology?


## Atmospheric charm production

Cosmic ray


## Prompt neutrinos @ terrestrial detectors



- Our central result is just below the most recent IceCube bound, indicating that a prompt component of the incoming flux should be observed soon....
- Our central result is consistent with the recent BERSS collaboration, though with better estimates of the uncertainties, which also encompass the 2008 ERS result and the most recent GMS calculation.



## SCET, an effective theory of QCD

- SCET permits the derivation of all-order factorization theorems:

$$
d \sigma \sim H \cdot \mathcal{J} \otimes \mathcal{J} \otimes \mathcal{S}
$$



- Once factorized, we resum logs via RG Equations:

$$
\frac{d H\left(Q^{2}, \mu\right)}{d \ln \mu}=\left[2 \Gamma_{c u s p} \ln \left(\frac{Q^{2}}{\mu^{2}}\right)+4 \gamma_{H}\left(\alpha_{s}\right)\right] H\left(Q^{2}, \mu\right)
$$

- To increase the accuracy of the resummations one needs the anomalous dimensions and the matching corrections to higher orders.


# Automated calculation of dijet soft functions 

| Soft function | $\gamma_{0}^{S} / C_{F}$ | $c_{1}^{S} / C_{F}$ | $\gamma_{1}^{C_{A}}$ | $\gamma_{1}^{n_{f}}$ | $c_{2}^{C_{A}}$ | $c_{2}^{n_{f}}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Thrust [168,169] | 0 | $-\pi^{2}$ | 15.7945 <br> $(15.7945)$ | 3.90981 <br> $(3.90981)$ | -56.4992 <br> $(-56.4990)$ | 43.3902 <br> $(43.3905)$ |
| C-parameter [142] | 0 | $-\pi^{2} / 3$ | 15.7947 <br> $(15.7945)$ | 3.90980 <br> $(3.90981)$ | -57.9754 <br> $(-)$ | 43.8179 <br> $(-)$ |
| Thresh. Drell-Yan [167] | 0 | $\pi^{2} / 3$ | 15.7946 <br> $(15.7945)$ | 3.90982 <br> $(3.90981)$ | 6.81281 <br> $(6.81287)$ | -10.6857 <br> $(-10.6857)$ |
| W@large $p_{T}[172]$ | 0 | $\pi^{2}$ | 15.88 <br> $(15.7945)$ | 3.905 <br> $(3.90981)$ | -2.65034 <br> $(-2.65010)$ | -25.3073 <br> $(-25.3073)$ |

Table 3.3: Anomalous dimensions and finite terms of the renormalized soft function for sample $\mathrm{SCET}_{1}$ observables. The upper numbers are the numerical results that we obtain with the SecDec implementation of our algorithm, and the lower ones correspond to the known analytic expressions.

## NNLL resummation of angularities

$$
\tau_{a}(X)=\frac{1}{Q} \sum_{i \in X} E_{i}\left|\sin \theta_{i}\right|^{a}\left(1-\left|\cos \theta_{i}\right|\right)^{1-a}
$$



Figure 3.6: LEFT: The central values of the NNLL' resummed and $\mathcal{O}\left(\alpha_{s}^{2}\right)$ matched angularity distributions at all 7 values of the parameter $a$. RIGHT: Theory bands demonstrating the convergence between NLL $\rightarrow$ NNLL' resummations. The plot is for $a=.25 . Q=91.2 \mathrm{GeV}$ in both plots.

## NNLL resummation of angularities



Figure 3.7: NNLL' resummed and $\mathcal{O}\left(\alpha_{s}^{2}\right)$ matched angularity distributions at four values of the parameter $a, a \in\{-.5,-.25, .25, .5\}$. The blue (PT) curves represent the purely perturbative cross-section, whereas the green (NP) curves are shifted according to (3.126). $Q=91.2 \mathrm{GeV}$ in all four plots.

## Projects, ideas, and interests

Finalizing automated calculation of NNLO soft functions w/ Guido Bell (Siegen) and Rudi Rahn (Bern)

NNLL resummation of angularities

## w/ Chris Lee (LANL), Andrew Hornig, and Guido Bell

What's the value of the strong coupling constant at M_\{Z\}?

Are there any other systematic uncertainties in the prompt atmospheric neutrino flux?

What can SCET say about the (forward) production of heavy mesons?

