

Long Baseline: LBNF Events in THEIA

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THEIA Workshop - Hamburg
UC Davis
23/03/2017

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Contents

- Predicted event numbers from long baseline running. Dependent on:
 - Different beam modes.
 - o True cp-phase.
 - Neutrino mass hierarchy.
- Event display.
 - Producing event displays within the rat-pac framework, using GENIE simulated events.
- Reconstruction.
 - Separation of Cherenkov and scintillation light.
 - Cherenkov ring reconstruction.
- Light yield device.



Predicted Events

Predicted Events - Considerations



- In order to calculate the number and composition of events we must consider:
 - o Oscillation probabilities.
 - o Neutrino beam flux.
 - Cross-section measurements.
 - o Detector size.
 - o Detector composition.
 - o Baseline.
 - o Run period.

Oscillation Probabilities



 To calculate the oscillation probabilities we must consider matter effects and the cp phase:

$$\begin{split} P(\nu_{\mu} \rightarrow \nu_{e}) = & 4C_{13}^{2}S_{13}^{2}S_{23}^{2}\sin^{2}\Phi_{31}(1 + \frac{2a}{\Delta m_{31}^{2}}(1 - 2S_{13}^{2})) \\ +8C_{13}^{2}S_{12}S_{13}S_{23}(C_{12}C_{23}\cos\delta_{CP} - S_{12}S_{13}S_{23})\cos\Phi_{32}\sin\Phi_{31}\sin\Phi_{21} \\ -8C_{13}^{2}C_{12}C_{23}S_{12}S_{13}S_{23}\sin\delta_{CP}\sin\Phi_{32}\sin\Phi_{31}\sin\Phi_{21} \\ +4S_{12}^{2}C_{13}^{2}(C_{12}^{2}C_{23}^{2} + S_{12}^{2}S_{23}^{2}S_{13}^{2} - 2C_{12}C_{23}S_{12}S_{23}S_{13}\cos\delta_{CP})\sin^{2}\Phi_{21} \\ -8C_{13}^{2}S_{13}^{2}S_{23}^{2}(1 - 2S_{13}^{2})\frac{aL}{4E_{\nu}}\cos\Phi_{32}\sin\Phi_{31}, \end{split}$$

Where:

$$\phi_{ji} = \Delta m_{ji}^2 L/4E_{\nu}$$

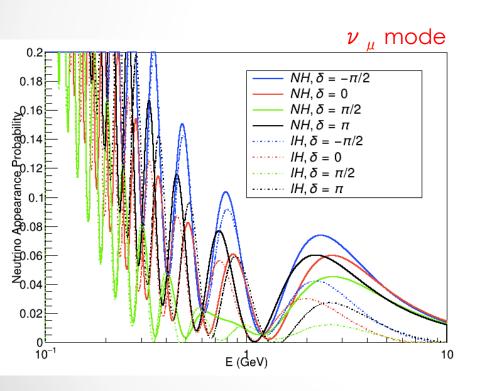
$$a \equiv 2\sqrt{2}G_F n_e E_{\nu} = 7.56 \times 10^{-5} [eV^2] (\frac{\rho}{[g/cm^3]}) (\frac{E_{\nu}}{[GeV]})$$

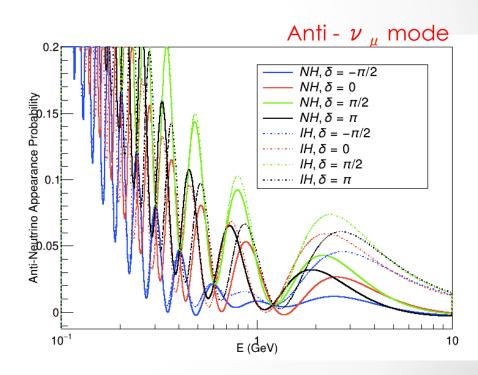
- Assumed a baseline of 1300 km.
- Constant crust density of 2.6 gcm⁻³.
- Used PDG 2016 values.

Oscillation Probabilities **UCDAVIS**



- Deduce the oscillation probabilities for different combinations of true cp phase, mass hierarchy structure and neutrino mode.
- Of particular interest for long-baseline studies is the 1-10 GeV range.

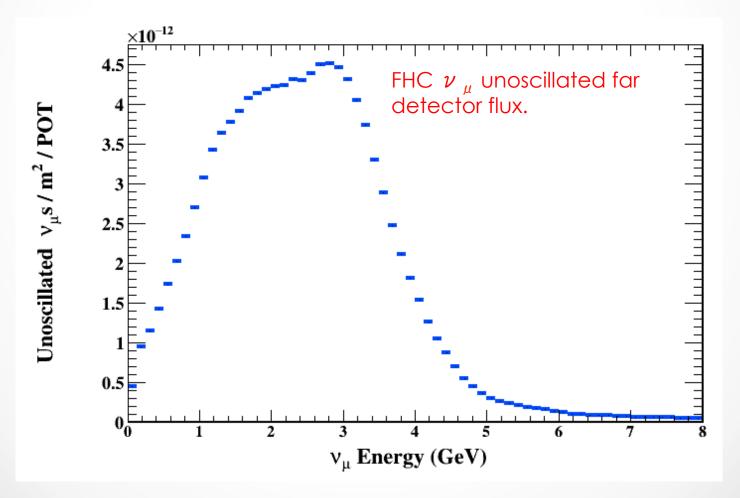




Neutrino Flux



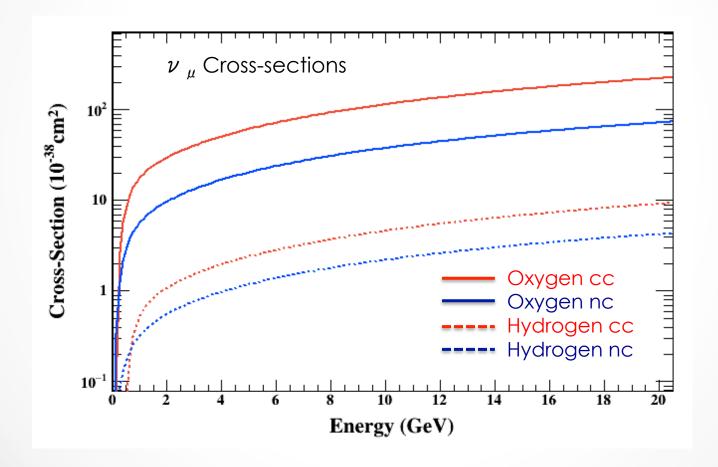
- Used the LBNE optimised neutrino flux files.
 - Gives increased sensitivity to cp violation and mass hierarchy determination measurements.
 - Full details of the configuration can be found in arXiv.1512.06148.



Cross-Sections

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- Using GENIE cross-sections.
- Considered interactions on H and O only.



Number of Events - Calculation



- In the example given, the following parameters were used:
 - o A 40kT FV.
 - o 3.5 years FHC neutrino mode run.
 - \circ $\delta_{cp} = 0$.
 - o Normal mass hierarchy.
 - \circ Not included intrinsic u $_{
 m e}$ background in the beam.
 - Only considered cross-sections of water and hydrogen (no carbon).
- 20 GENIE simulations.
- Average gives the expected value and the RMS of this gives the uncertainty.

Event Breakdown



Simulation of the following number of events...

Neutrino Flavour	Number of Events
Anti - $ u_{\mu}$	898
Anti - $\nu_{ m e}$	13
$ u_{\mu}$	12545
$ u_{ m e}$	1544

Leading to the following breakdown...

Interaction Type	Number of Events
Anti – ν_{μ} CC	650.2 ± 12.5
Anti – ν_{μ} NC	247.8 ± 12.5
Anti – v_e CC	9.1 ± 1.4
Anti – v_e NC	3.9 ± 1.4
v_{μ} CC	9255.7 ± 35.7
$v_{\mu}NC$	3289.3 ± 35.7
v_e CC	1138.9 ± 17.4
$v_{\rm e}$ NC	404.4 ± 17.4





· Which can be further subdivided into...

Number of Events
$1857.2 \pm 40.1 \ (109.6 \pm 9.9)$
$2770.9 \pm 45.2 \ (140.9 \pm 8.6)$
$4560.2 \pm 45.1 \ (390.9 \pm 17.3)$
$67.3 \pm 11.0 \ (8.7 \pm 3.5)$
$738.3 \pm 18.6 \ (41.0 \pm 5.4)$
$965.7 \pm 20.6 (57.8 \pm 7.8)$
$1546.8 \pm 35.9 (144.3 \pm 9.2)$
$38.6 \pm 4.4 \ (4.8 \pm 2.1)$
$224.9 \pm 10.3 \ (3.5 \pm 1.6)$
$405.9 \pm 12.8 \ (2.6 \pm 1.6)$
$497.6 \pm 21.6 \ (2.7 \pm 1.6)$
$10.5 \pm 3.4 \ (0.3 \pm 0.4)$
$86.7 \pm 9.2 (1.3 \pm 1.1)$
$144.2 \pm 13.0 \ (1.5 \pm 0.8)$
$168.7 \pm 8.5 \ (1.1 \pm 1.0)$
$4.8 \pm 2.5 \ (0.0 \pm 0.0)$

Event Display and Reconstruction



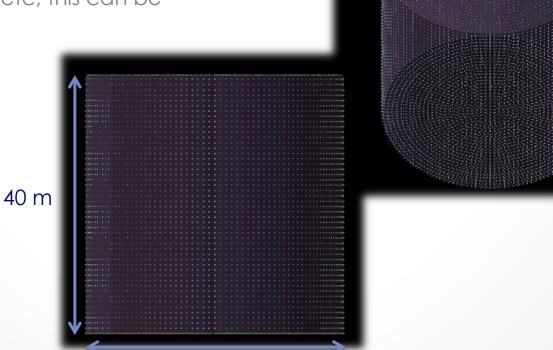
THEIA Simulations



- Able to simulate GENIE events in THEIA using rat-pac.
 - Version issues. Currently using Geant4 Version 10.01.p02.
 - When simulating Cherenkov events the simulation is slower than expected.

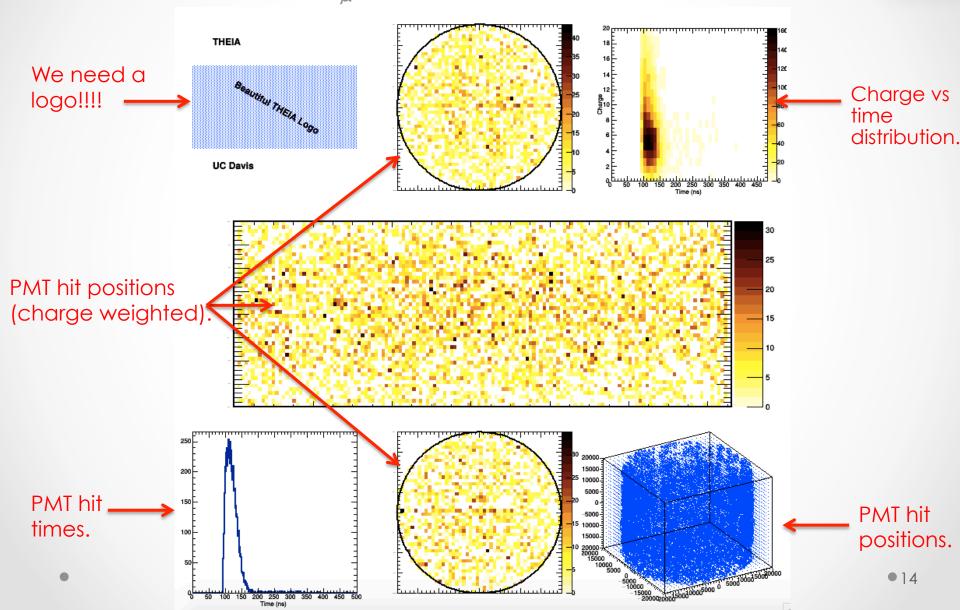
Current setup:

- 11471 R7081_hqe inner facing PMTS.
- o 26.9 % PMT coverage.
- Depending on final PMT choice, configuration etc, this can be increased.
- o WbLS 5%.
- o 40 kT volume.

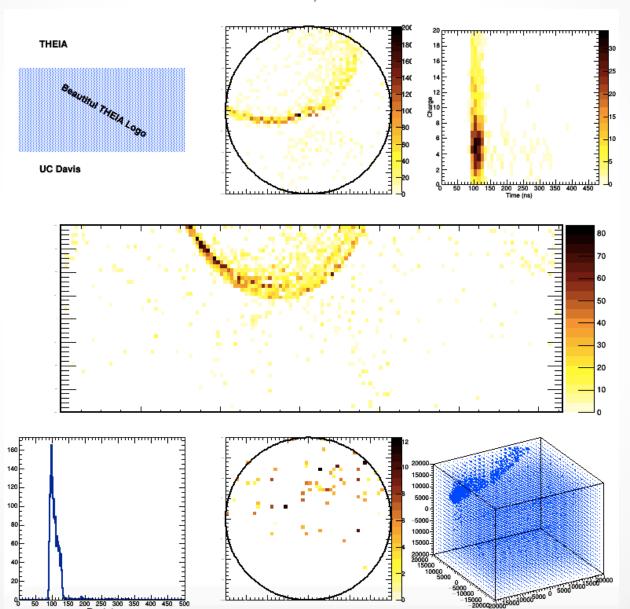


Event Display – Scintillation Only UCDAVIS

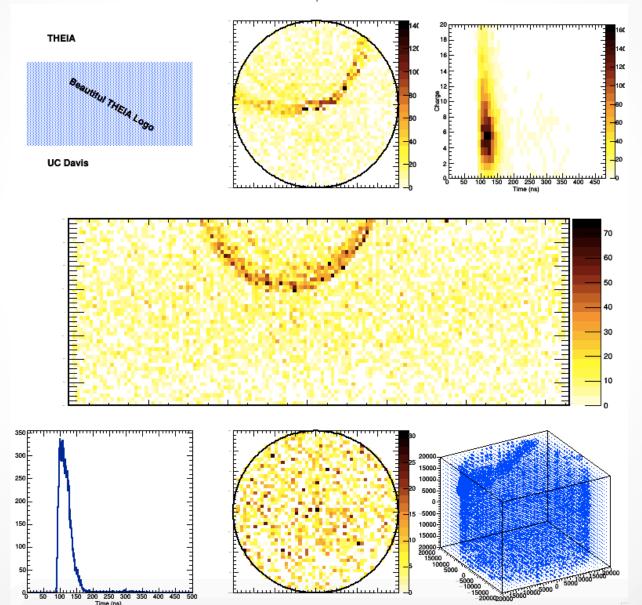
• Simulation of a ν_{μ} CC Quasi-Elastic Event.



Event Display – Cherenkov Only Simulation of the same ν_{μ} CC Quasi-Elastic Event.



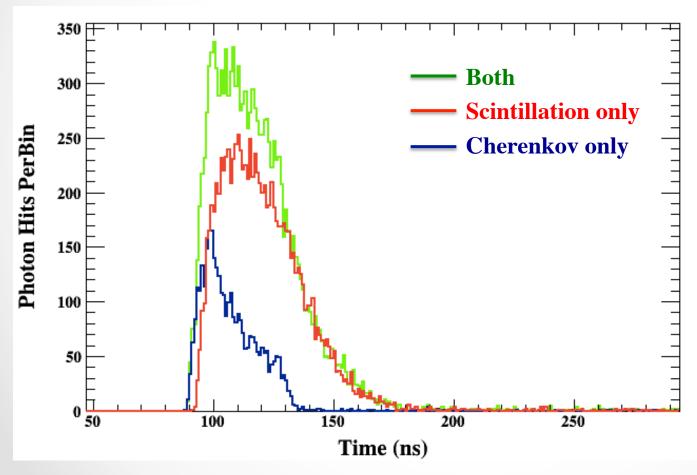
Event Display – Both • Simulation of the same ν_{μ} CC Quasi-Elastic Event.



PMT Hit Times



 PMT hit times of photons from Cherenkov emission, scintillation and both.

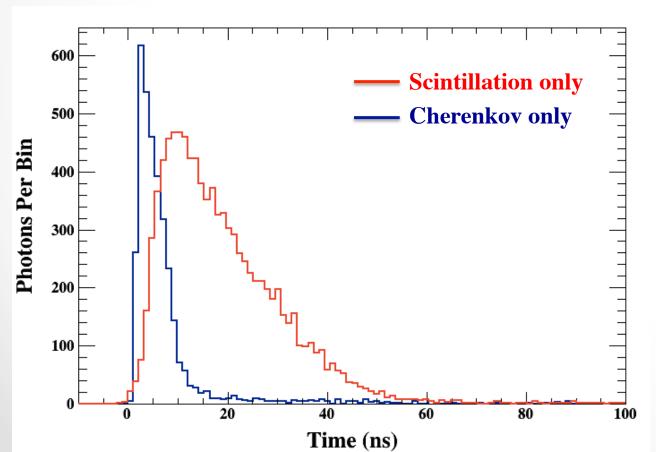


Although Cherenkov arrives sooner, there is still much overlap with the scintillation light. However, if we consider the time residuals...

Residual Timing



- Using muon truth information and knowledge of the vertex position, the following residuals are produced for Cherenkov and scintillation photons.
- This information can be used to locate the vertex.

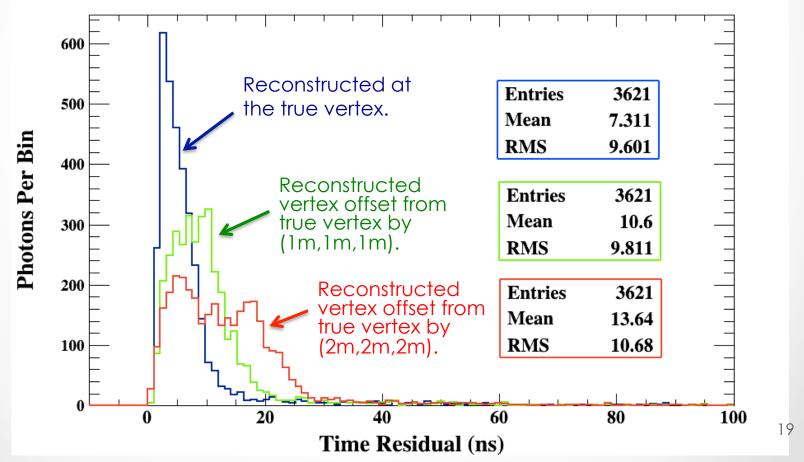


Greater separation of Cherenkov and scintillation light compared to the PMT hit timing presented on the previous slide.



Vertex Reconstruction – Cherenkov Light

- For an accurately reconstructed vertex, we would expect to see a narrower, more peaked, timing residual.
- For the Cherenkov radiation below, we see this is the case.

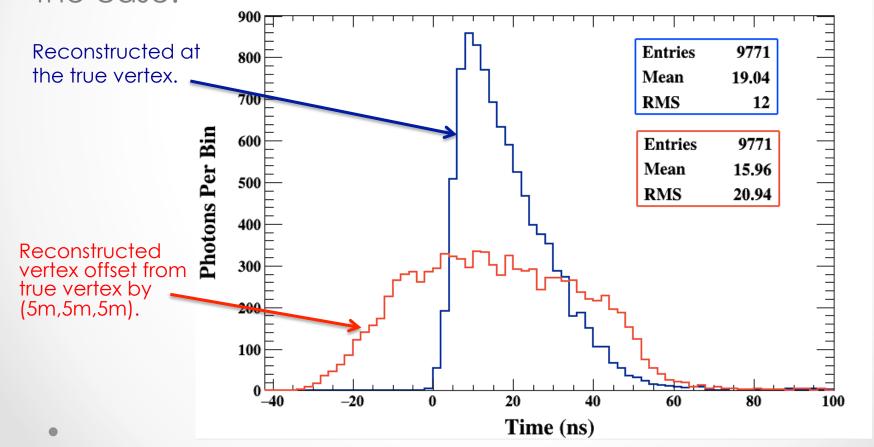


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Vertex Reconstruction – Scintillation Light

- For an accurately reconstructed vertex, we would expect to see a narrower, more peaked, timing residual.
- As with Cherenkov radiation, for scintillation light this is also the case.

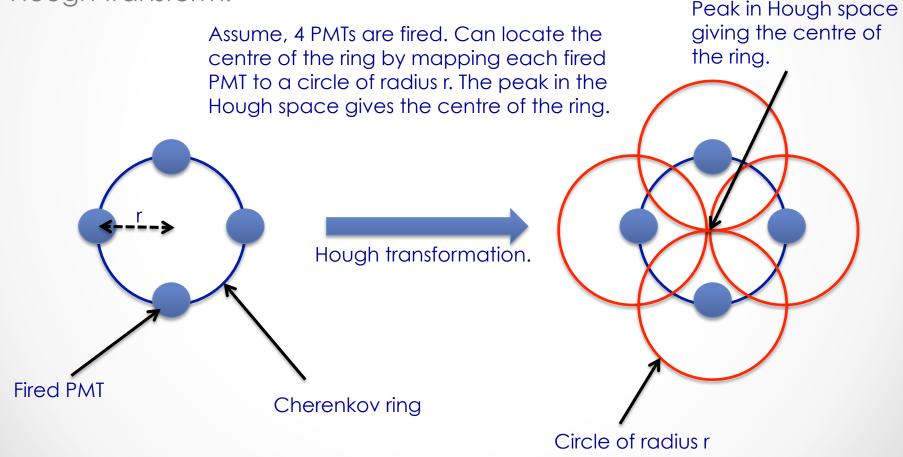


Near Future Reconstruction Work - Separation

- Originally, for Super-Kamiokande the first stage of reconstruction was fitting the vertex.
- As light is produced predominantly by Cherenkov radiation, the residual method is simple to use.
- However, we cannot calculate the time residuals in THEIA without knowing if the light is from scintillation or Cherenkov radiation.
- Perhaps make a PMT hit time cut in an attempt to isolate only scintillation?
- Perhaps search for Cherenkov rings first?

Near Future Reconstruction Work – Ring Fitting

 An easy to implement method for locating ring candidates is to use a Hough transform.



Light Yield Device

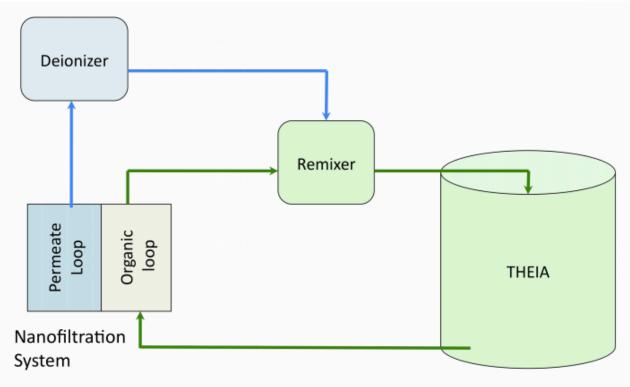


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Nanofiltration



- Testing a nanofiltration system at UC Davis, which will hopefully be scalable to use for THEIA.
- For an in-depth explanation see Prof. Svoboda's Frost Topical Workshop talk.²

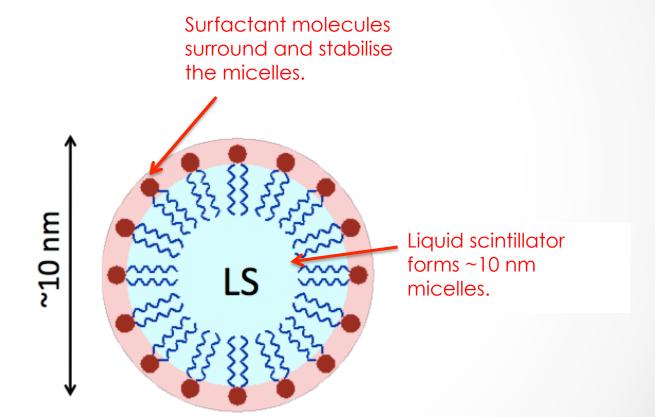


THEIA recirculation concept



Why Do We Need A Light Yield Device?

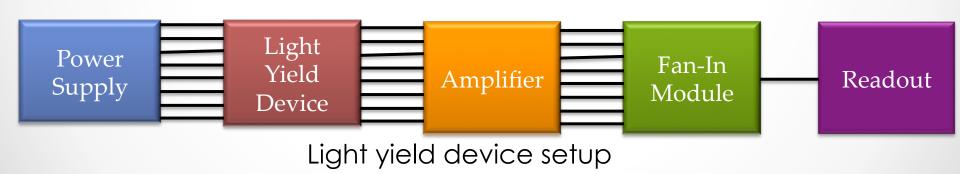
 One potential problem with nanofiltration, is that it might damage the micelles – affecting light yield.



Light Yield Device - Setup

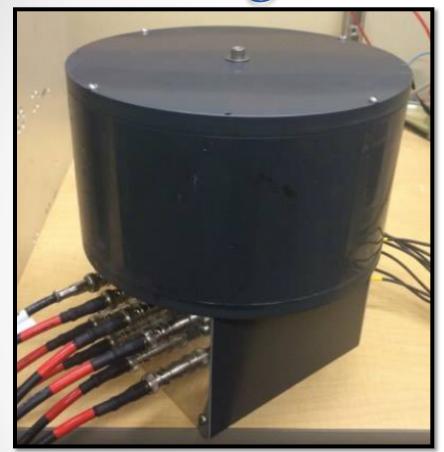


- Designed and built a light yield device that can hold a cuvette (dimensions: 45.2 x 12.5 x 12.5 mm) of WbLS.
- This is surrounded by 8 R928 Hamamatsu PMTs.
- Can measure the light yield when exposed to radiation from a ⁹⁰Sr source.
- 90 Sr undergoes β decay into 90 Y with a decay energy of 0.546 MeV.
- Source is placed above the cuvette, such that the emitted electron will be fully absorbed by the WbLS.
- Testing is underway and expect to take data soon...



Light Yield Device









Summary

- Infrastructure in place to predicted event numbers from long baseline running. Tunable depending on:
 - Different beam modes.
 - o True cp-phase.
 - Neutrino mass hierarchy.
- Event display produced.
 - Producing event displays within the rat-pac framework.
- Early stages of reconstruction have begun.
 - Separation of Cherenkov and scintillation light.
 - Cherenkov ring reconstruction.
- Light yield device.
 - Built and ready to begin taking data.