

Probing TeV scale origin of neutrino mass R. N. Mohapatra





LAUNCH 2017: Lindner 60th

Neutrino mass- a signal of new physics beyond SM

- Understanding origin of nu mass will specify the direction of new BSM physics and thus a significant venue of research!
 Could this new physics be in the TeV scale, so it could be probed using current facilities ?
- This talk explores this question within the seesaw paradigm for neutrino masses and how we can look for its signals at LHC.

Left-right model for neutrino mass seesaw

Two key ingredients of seesaw i.e.
 (i) Right handed neutrinos
 (ii) Broken B-L symmetry

are both automatic in left-right models !

- A compelling model for type I seesaw is the left-right extension of standard model.
- Majorana mass of the RH neutrino is protected by a gauge symmetry- so seesaw scale is connected to the symmetry breaking scale.

Left-Right model: • Gauge group: $SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$

• Fermions $\begin{pmatrix} u_L \\ d_L \end{pmatrix} \stackrel{P}{\Leftrightarrow} \begin{pmatrix} u_R \\ d_R \end{pmatrix} \begin{pmatrix} v_L \\ e_L \end{pmatrix} \stackrel{P}{\Leftrightarrow} \begin{pmatrix} v_R \\ e_R \end{pmatrix}$

$$L = \frac{g}{2} [\vec{J}_{L}^{\ \mu} \cdot \vec{W}_{\mu L} + \vec{J}_{R}^{\ \mu} \cdot \vec{W}_{\mu R}]$$

Parity a spontaneously $M_{W_R} \gg M_{W_L}$ broken symmetry: (Pati, Salam'74; Mohapatra, Pati'74; '74; Senjanovic, Mohapatra'75)

Left-Right seesaw:

• Gauge group: $SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$ (needed for

• Fermions $(u_L) \xrightarrow{P} (u_R) (v_L) P$

$$\begin{pmatrix} u_L \\ d_L \end{pmatrix} \stackrel{P}{\Leftrightarrow} \begin{pmatrix} u_R \\ d_R \end{pmatrix} \begin{pmatrix} v_L \\ e_L \end{pmatrix} \stackrel{P}{\Leftrightarrow} \begin{pmatrix} v_R \\ e_R \end{pmatrix}$$

seesaw)

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Parity breaking as origin of Majorana Neutrino mass

Electric charge formula in LR

$$Q = I_{3L} + I_{3R} + \frac{B - L}{2}$$

Above EW scale,

$$\Delta Q = \Delta I_{3L} = 0 \rightarrow \Delta I_{3R} = -\frac{1}{2}\Delta L$$

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Parity breaking → Majorana nu (RNM, Marshak'80)
 Small neutrino mass is why Weak Int. V-A

Other attractive features LR models

- Simple picture of naturally stable WIMP DM: Add B-L=0 triplet fermions: No susy required (Heeck, Patra'15; Nagata, Olive, Zheng'15)
- Solution of strong CP without axion: Parity makes mass matrices hermitean $\rightarrow \theta = 0$

(Beg, Tsao'78; RNM, Senjanovic'78)

Both prefer W_R mass in multi-TeV range: a motivation for exploring TeV LR seesaw !!

How do neutrino masses work in TeV LR seesaw?

• Typically, $m_D = h_{\nu} v_{wk}$;

• So for $h_{\nu} \simeq 10^{-5.5} \sim h_e$; seesaw scale v_R could easily be in the TeV range and fit oscillation data;

 Hence W_R, Z' accessible to colliders. M_{WR} = g_Rv_R ~ fewTeV

 Plus low energy effects e.g. ββ_{0ν}, μ → e + γ...

$$\begin{array}{c} \textbf{Doublet breaking of LR and}\\ \textbf{Inverse seesaw alternative}\\ SU(2)_{L} \times SU(2)_{R} \times U(1)_{B-L} + \text{singlets S}\\ <\chi_{R}^{0} \\ \boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R} \quad (\Delta L=0) \quad \boldsymbol{M}_{N} = f \boldsymbol{\nabla}_{R}\\ (\nu \quad N \quad S)\\ SU(2)_{L} \times U(1)_{Y}\\ <\phi_{1}^{0} \\ \boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R} \quad (\boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R})\\ (\boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R})\\ (\boldsymbol{\nabla}_{R} \quad \boldsymbol{\nabla}_{R})\\ (\boldsymbol{$$

 $m_{
u} \simeq m_D (f v_R)^{-1} \mu (f v_R)^{-1} m_D^T$ (RNM'86; RNM, Valle'86)

• $\mu \sim \text{keV} \rightarrow \text{TeV LR}$

Left-right Inverse seesaw at TeV is GUT embeddable!

Neutrino mass is determined by small mupper parameter $\rightarrow h_{\nu}$ can be >> 10^{-5.5} (could even be $\sim h_t$ allowing for quark lepton unification)



(Dev, RNM'10) SO(10) is the GUT theory for this. Collider signals for TeV scale LR type I seesaw

- Vector boson signal: W_R, Z' (M~TeVs)
- New fermion signal: N_{e,mu,tau} (M~GeV-TeV)
- Scalar boson signal: analog of SM Higgs

Vector boson signal: How light can W_R Be?

■ New interactions of quarks with W_R affects low energy observables e.g. K_L-K_S, ϵ , ϵ' , B_s-B_{s-bar}, \rightarrow M_{WR} > 2.5 TeV (g_R /g_L)

(Zhang,An,Ji,RNM; Maiezza, Nemevsek,Nesti,Senjanovic;Blanke, Buras,Gemmler,Hiedsieck)

• $g_R \neq g_L$ compatible with LR; $g_R/g_L > 0.6$

• Resolving ϵ'/ϵ puzzle \rightarrow TeV W_R (Cirigliano, Dekens, de Vries, Mereghetti'16)

Collider signals of LR seesaw (Type I) **Golden channel:** $W_R \rightarrow \ell_i \ell_k j j$ ■ W_R mediated graph(Keung, Senjanovic'82; Mitra, et al.'16; T. Han'17) $\sigma(W_R) \times BR(Ne) \sim 14.8$ fb $M_{MR} = 3 \text{ TeV}$ Predicts #

• Unlikely to be SM signal $A_{\ell\ell jj} \propto m_D M_N^{-1} m_D$

- What if $\#\ell^{\pm}\ell^{\pm} \neq \#\ell^{\pm}\ell^{\mp}$ General inverse seesaw $\begin{pmatrix} 0 & m_D & 0 \\ m_D^T & \mu_R & M_N \\ 0 & M_N^T & \mu_S \end{pmatrix} \begin{pmatrix} \nu \\ N \\ S \end{pmatrix}$
- N-S mix, α and split ΔM^2
- $\begin{array}{l} \bullet \quad \text{Coherence essential:} (\Delta M^2 \leq 2E\Gamma_{W_R} \text{; Akhmedov; Kayser}) \\ \frac{\ell^+ \ell^+}{\ell^+ \ell^-} \equiv R_{\ell \ell} = \quad \frac{\cos^2 2\alpha + \frac{4\Delta M^2}{\Gamma_0^2}}{1 + \sin^2 2\alpha + \frac{4\Delta M^2}{\Gamma_0^2}} \neq 1 \end{array}$

(Dev, RNM'15; Aniamati, Hirsch, Nardi'16; Das, Dev, RNM(to appear))

Type I case: CPV in N-decay: (Gluza, Jelinski)



Current limit from LHC

$\ell\ell jj$ final state



• Current limit: 2.8 TeV $(g_L = g_R)$

Limits on Z'

Minimal models:

Type I

$$M_{Z_2} = \sqrt{\frac{2\cos^2\theta_W}{\cos^2\theta_W}} M_{W_R}$$

Inverse seesaw

$$M_{Z_2} = \sqrt{\frac{\cos^2 \theta_W}{\cos 2\theta_W}} M_{W_R}$$

 $M_{Z'} > 5 \text{ TeV};$ $M_{Z'} > 3.7 \text{ TeV}$

In general LR models, where above relations do not hold, $M_{Z'} > 3.2$ TeV from 13 TeV data on $Z' \rightarrow \ell \ell$ (Lindner, Queiroz, Rodejohann'16)

Prospects for LR in future colliders

LHC14 reach 5-6 TeV; 100 TeV collider: 30 TeV signal: *pp→ lljj* LHeC: (~100 GeV e+7 TeV p) (M_{WR} < 6.5 TeV)



signal $pe^- \rightarrow e^+ jjj$

(Lindner, Queiroz, Rodejohann, Yaguna'16)



■ For low masses, new particles in seesaw are long lived → displaced vertices at LHC.

(SHiP; recent proposal:MATHUSLA detector)

(Curtin, Chou and Lubatti'16)



$$N_{obs} \sim N_X \cdot Br(X \to SM) \epsilon_{geom} \cdot \frac{L}{bc\tau}$$

- Fermion sector: X = light RHN
- Higgs sector: X=H₃ :seesaw Higgs

Light N and displaced vertices at LHC



New Higgs bosons in LR and displaced vertices

- Parity partner of SM Higgs $H_1^0, A^0; M_{H_1^0, A^0} \ge 10 \text{ TeV}$
- New : Seesaw Higgs $H_3^0 \equiv Re\Delta_R^0$ $M_{H_3^0} \simeq 1 - 1000 {\rm GeV}$
- Two domains of seesaw Higgs masses
- $M_{H3} \sim v_R >> M_h$
- $M_{H3} << M_h << v_R$ (~ few GeV to 100 MeV)



Light Seesaw Higgs H₃ (few GeV)

- No phenomenological limit; could it be light?
- If SU(2)_R x U(1)_{B-L} is broken radiatively, there is a good chance that $H_3^0 \equiv Re\Delta_R^0$ is light:

$$(m_{H_3}^2)^{\text{loop}} \simeq \frac{3}{2\pi^2} \left[\frac{1}{3} \alpha_3^2 + \frac{8}{3} \rho_2^2 - 8f^4 + \frac{1}{2} g_R^4 + (g_R^2 + g_{BL}^2)^2 \right] v_R^2$$

(Dev, RNM, Yongchao Zhang, PRD'17; 1703.02471)

Radiative breaking inspired by a conformal approach to gauge hierarchy

(Bardeen'95; Holthausen, Lindner, Schmidt'10)

How does light H₃ decay?

H_3 decay to quarks and leptons suppressed by FCNC constraints.



$\gamma\gamma$ Displaced vertices at LHC: signal of light H₃

Near GeV mass H_3^0 accessible at the LHC via displaced vertices ($\gamma\gamma$ mode 100% BR)

(Dev, RNM, Yongchao Zhang'16)

 M_{H3}- M_{WR} reach at LHC via this mode.

MATHUSLA Blue line





Take away lessons

- Observation of $\beta\beta_{0\nu}$ at the level of 20 to 30 meV does not mean inverse hierarchycould be W_R effect. Need supplementary information from LHC to pinpoint this.
- Suppose long base line → NH, any signal of $\beta\beta_{0\nu}$ at this level would strongly imply new particle effect e.g. W_R.
- Must find ways to disentangle heavy particle effects from light nu exchange

Implications of TeV WR for leptogenesis

- A'TeV W_R affects washout of lepton asymmetry
- Washout increases as M_{WR} decreases: →lower bound on M_{WR} > 10 TeV

(Frere, Hambye, Vertongen'09; Dev, Lee, RNM.'14)

Thus a discovery of W_R below 10 TeV will rule out leptogenesis as a way to understand the origin of matter

Summary

- Left-Right model with TeV scale seesaw: a compelling scenarios for neutrino masses:
 Rich set of predictions for colliders and low energies: (W_R, Z', N, Δ⁺⁺_R, ββ_{0ν} etc)
- Displaced vertex searches: a new window to seesaw physics.
- LHC can broaden our understanding of nu mass origin based on LR seesaw significantly!!



Happy birthday, Manfred!

Looking for this signal: Light H₃ production at LHC14





Light Higgs leads to displaced vertices at LHC

• H_3 is a linear combination of SM Higgs h and LR new Higgs $H_1(\theta_1, \theta_2)$ and has effective quark coupling of the form:

 $\begin{array}{ll} H_{3}\bar{u}u & \frac{1}{\sqrt{2}}\widehat{Y}_{U}\sin\tilde{\theta}_{1} - \frac{1}{\sqrt{2}}\left(V_{L}\widehat{Y}_{D}V_{R}^{\dagger}\right)\sin\tilde{\theta}_{2} \\ H_{3}\bar{d}d & \frac{1}{\sqrt{2}}\widehat{Y}_{D}\sin\tilde{\theta}_{1} - \frac{1}{\sqrt{2}}\left(V_{L}^{\dagger}\widehat{Y}_{U}V_{R}\right)\sin\tilde{\theta}_{2} \\ & \sin\tilde{\theta}_{2} = \sin\theta_{2} + \xi\sin\theta_{1} \end{array}$

 For light H₃, B and K-decays limit the value of mixing angles (barring cancellation)

Constraints on H₃ mixings from B-decays (Babar, Belle, LHCb)

Mixing with SM Higgs and heavy LR Higgs H⁰ (θ_1, θ_2) are strongly constrained for m~GeV;



Theoretical upper limit on M_N

• Like the top quark in the SM, very large RHN mass will destabilize the vacuum. This gives an upper limit on: $\Delta^0_R = \Delta^0_R$

$$\{\sum_{i} M_{N_{R},i}^{4}\}^{1/4} \leq 1.18 M_{W_{R}}$$



Minkowski'77; Gell-Mann, Ramond, Slansky; Yanagida; Glashow; Mohapatra, , Senjanovic; '79

 ν_R Majorana mass breaks B-L maximally !