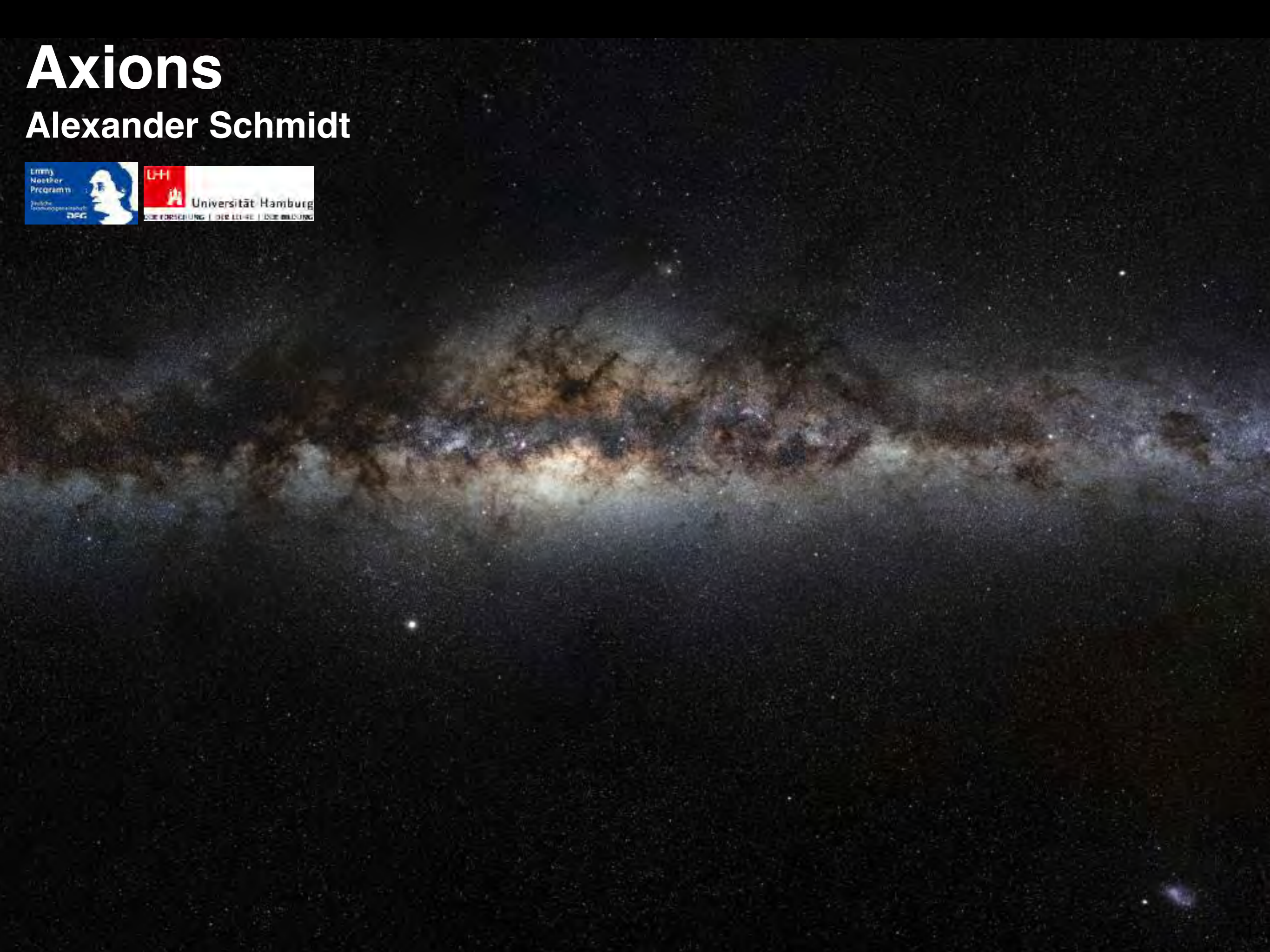
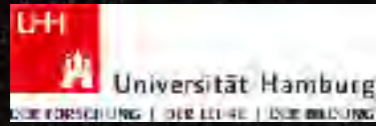


Axions

Alexander Schmidt

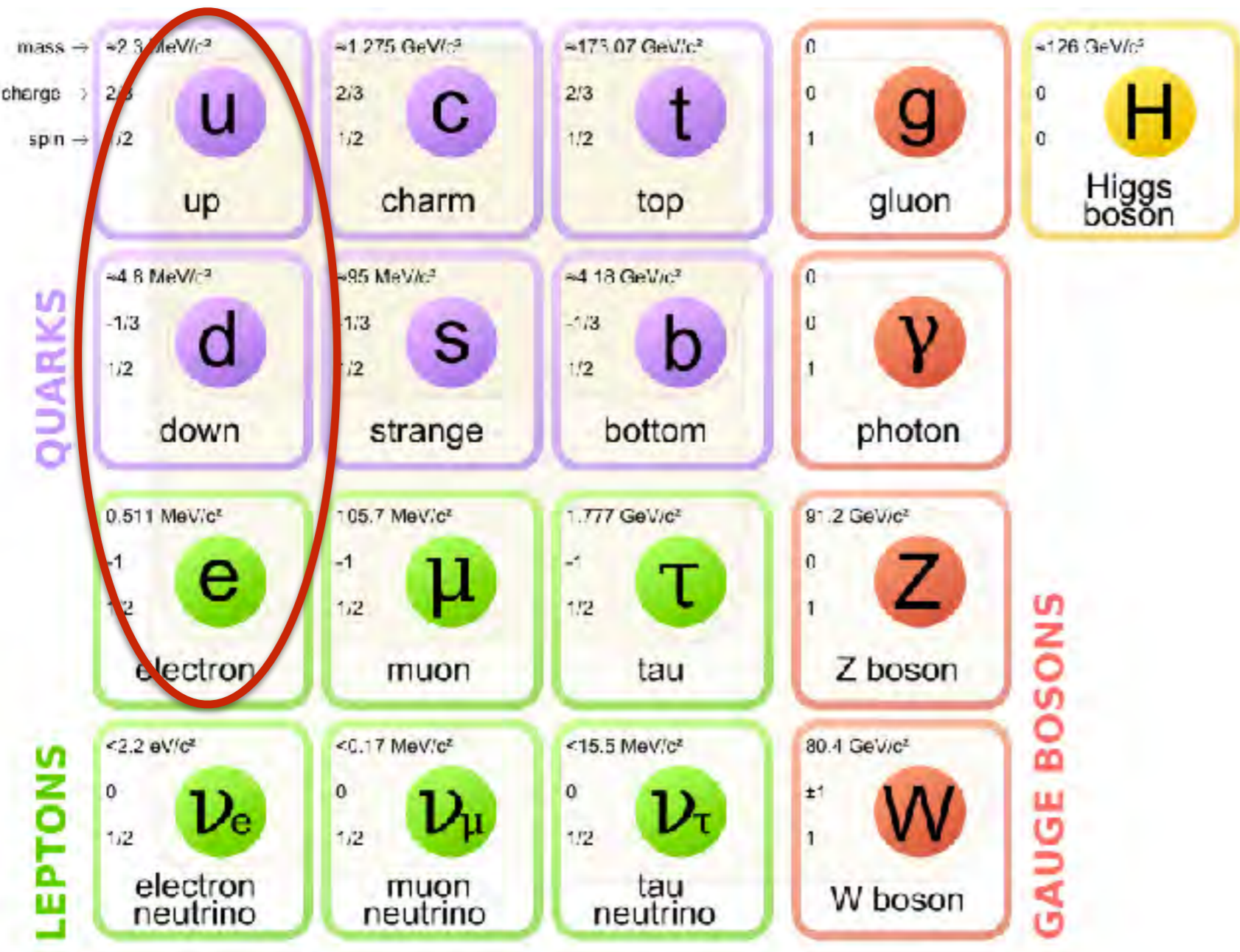


axions?

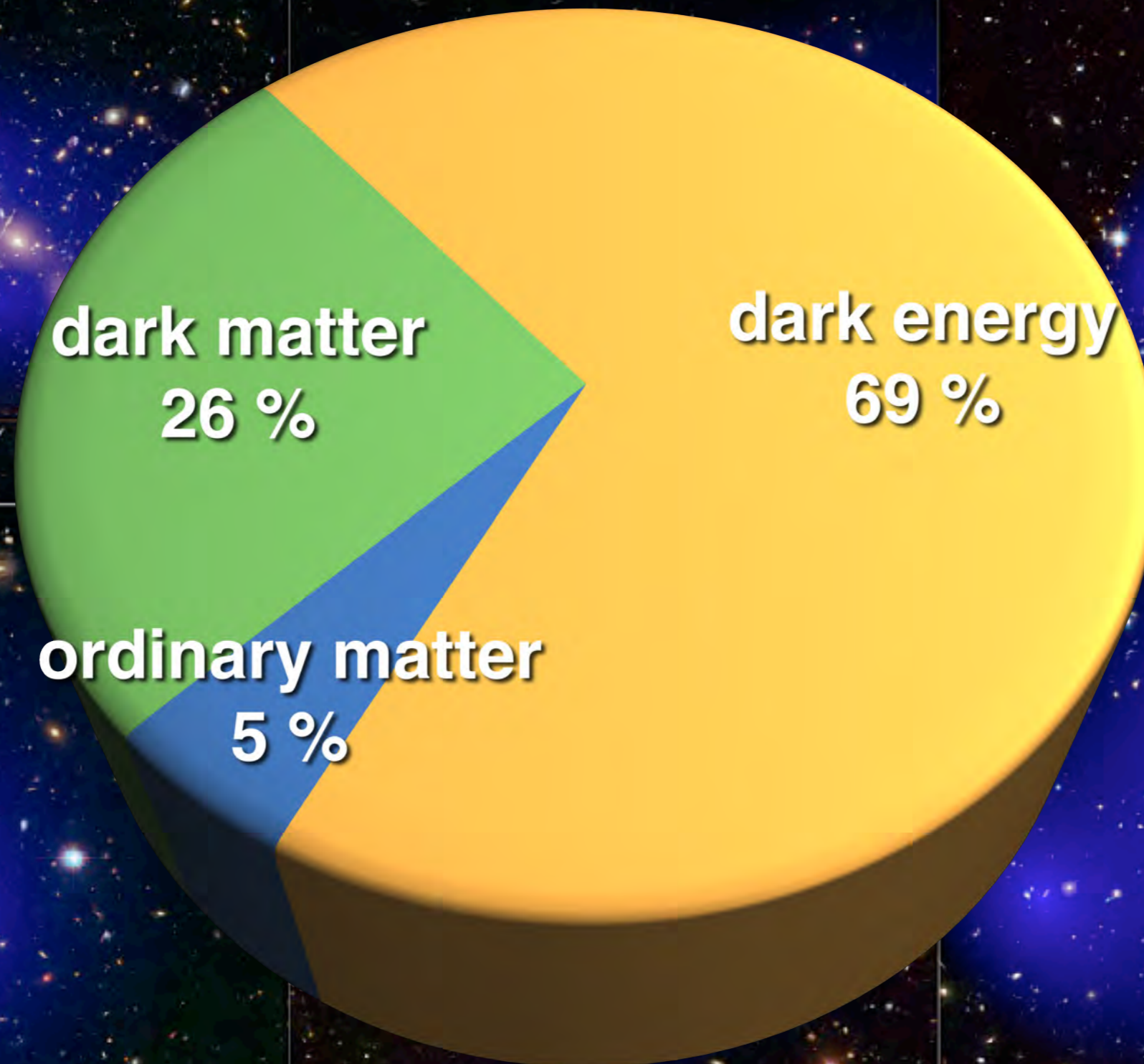


the big picture

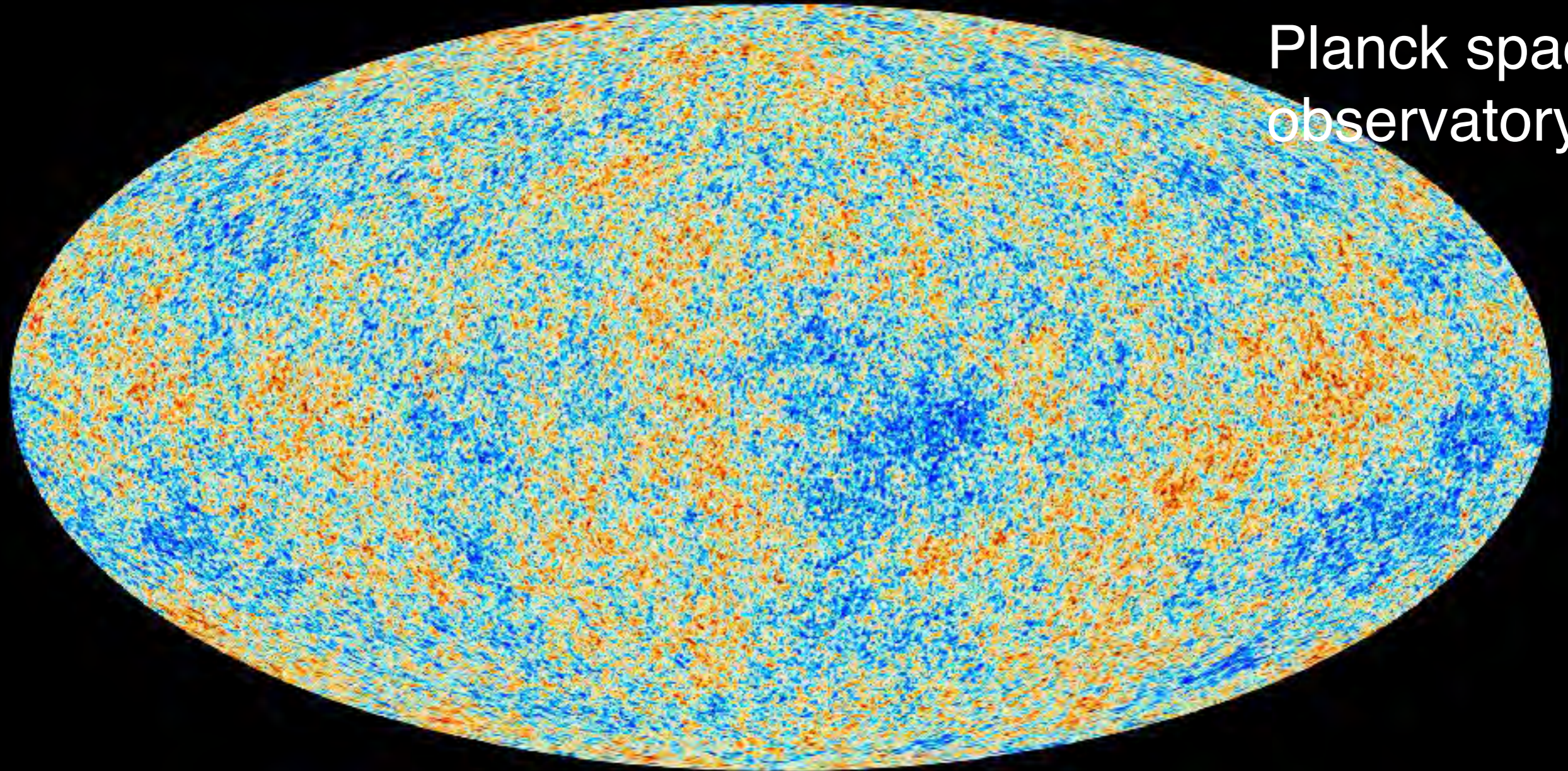
• confirmed again and again by all experiments:



- makes **incredibly** precise predictions about:
 - properties of particles
 - interactions between particles
- ➔ production cross sections
- ➔ decay rates of particles



cosmic microwave background

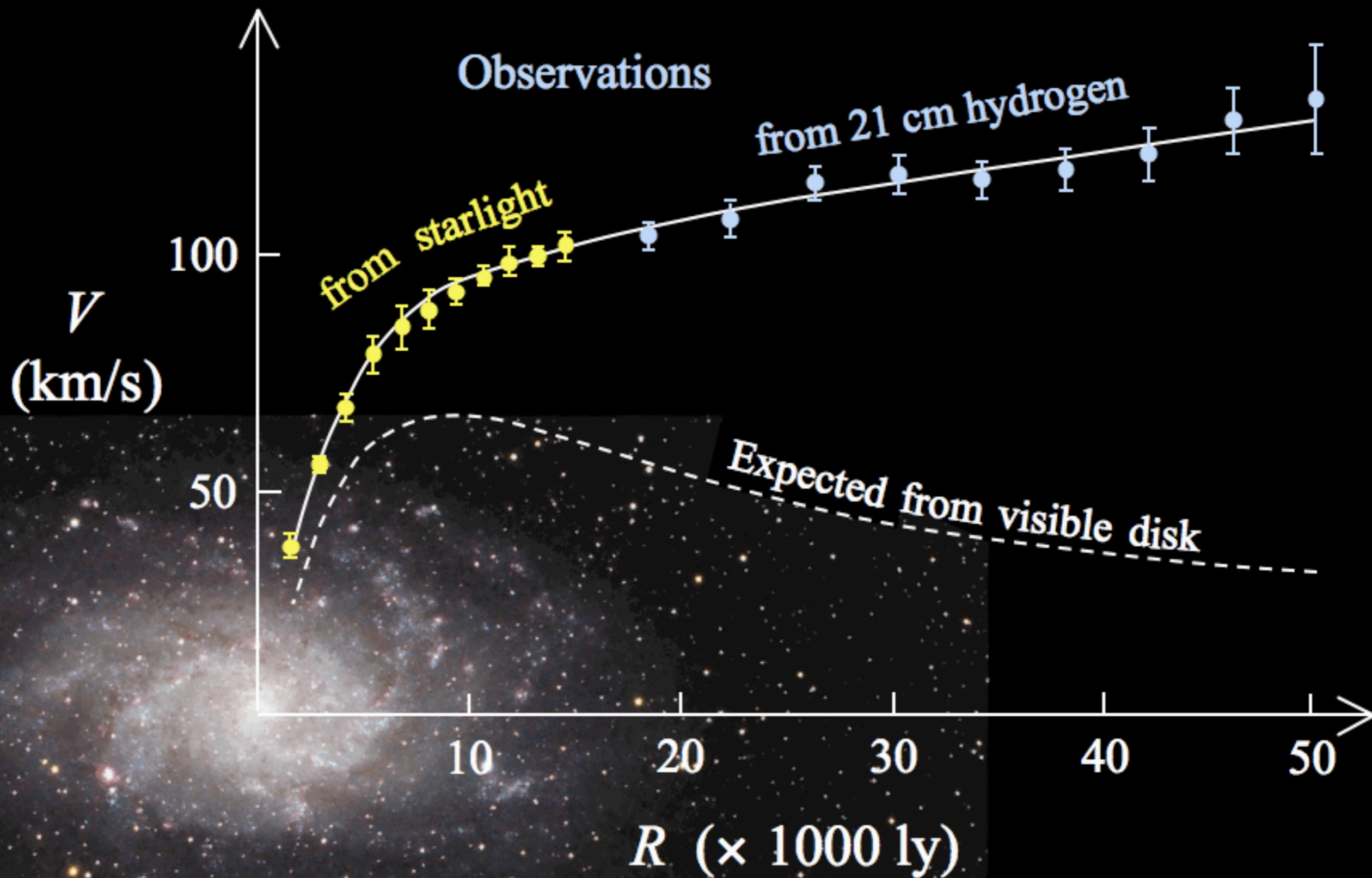


Planck space
observatory

**anisotropy distribution constrains
cosmological models:**

- expansion of the universe
- dark matter
- structure formation

indications for dark matter



dark matter candidates

WIMPs

- weakly interacting **massive** particles
- predicted by supersymmetry
- very large experimental efforts

subjective
plausibility: 50%

Axions

- ultra-**light** particles
- predicted by Peccei-Quinn theory
- large experimental efforts

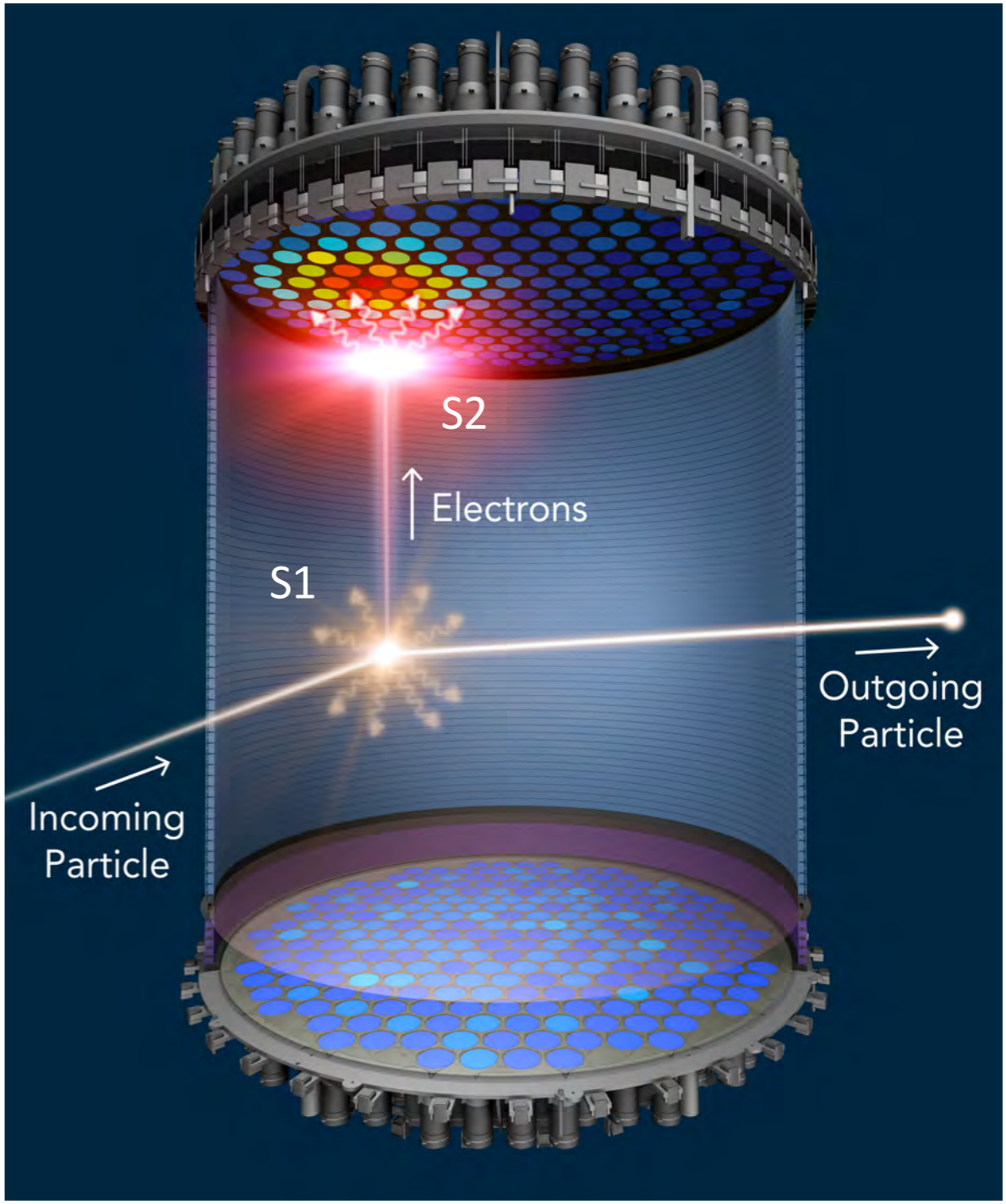
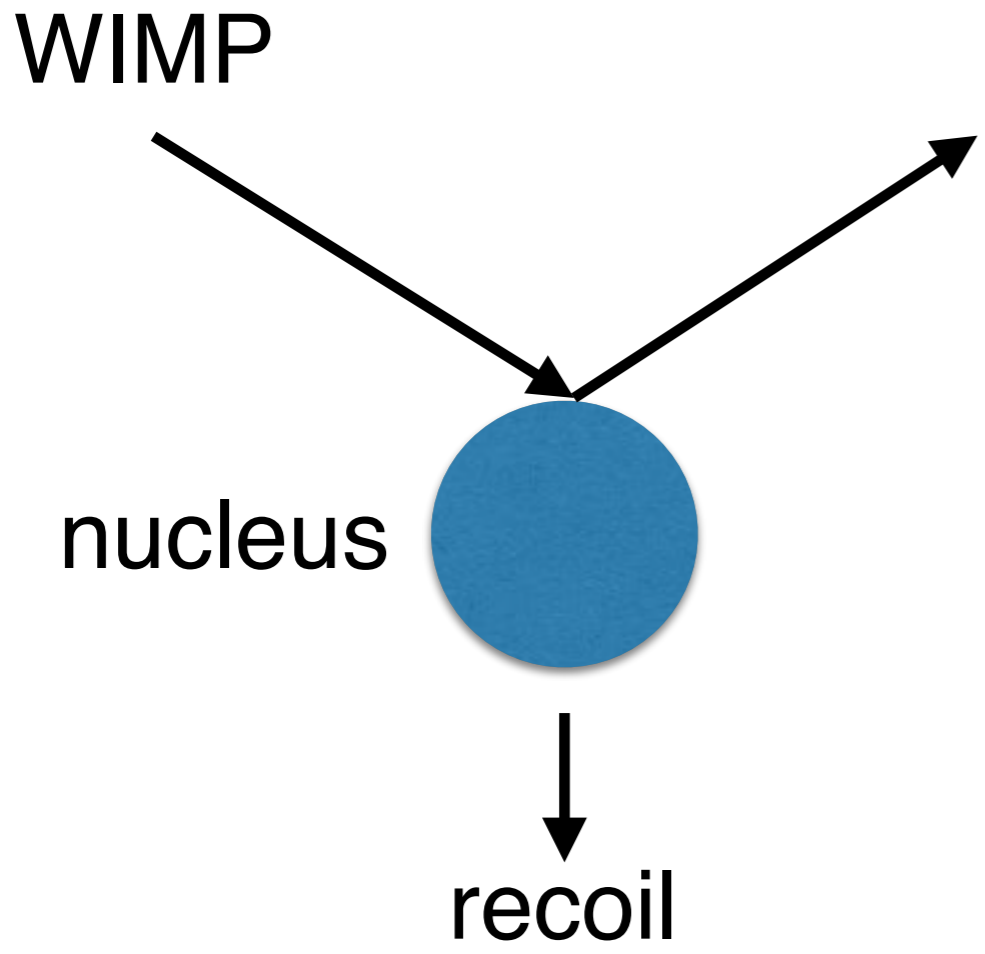
subjective
plausibility: 45%

other:

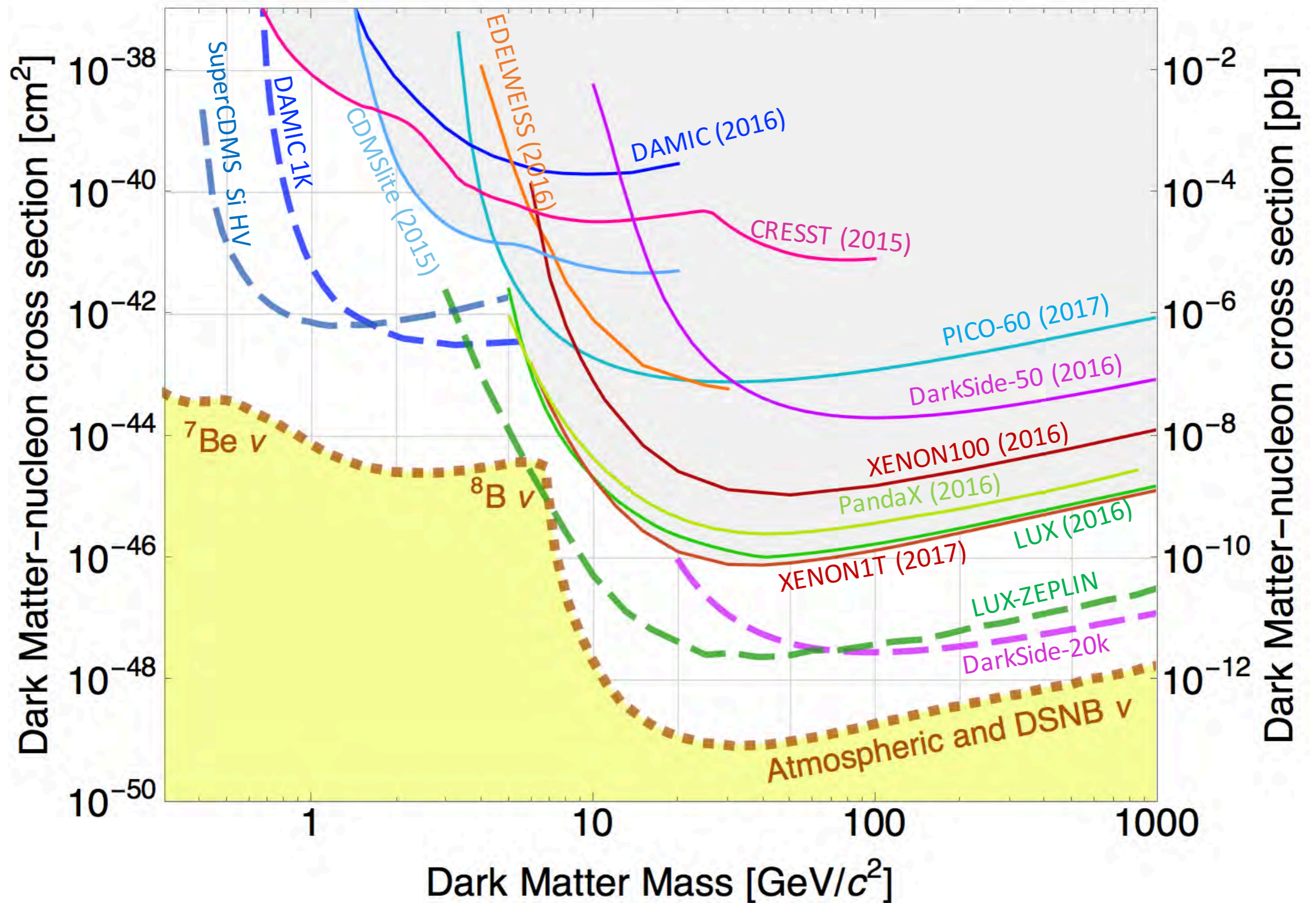
- modified Newtonian dynamics (MOND)
- massive compact halo objects (MACHO)

subjective
plausibility: 5%

direct search for WIMPs



search for WIMPs



dark matter candidates

WIMPs

- weakly interacting **massive** particles
- predicted by supersymmetry
- very large experimental efforts

subjective
plausibility: 25%

Axions

- ultra-**light** particles
- predicted by Peccei-Quinn theory
- large experimental efforts

subjective
plausibility: 70%

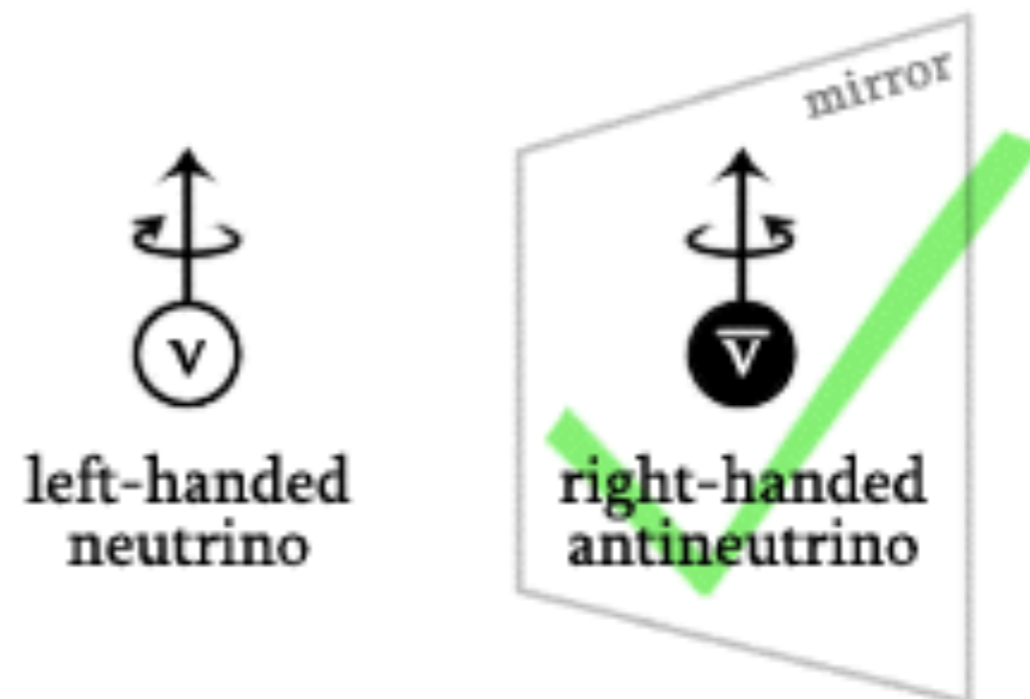
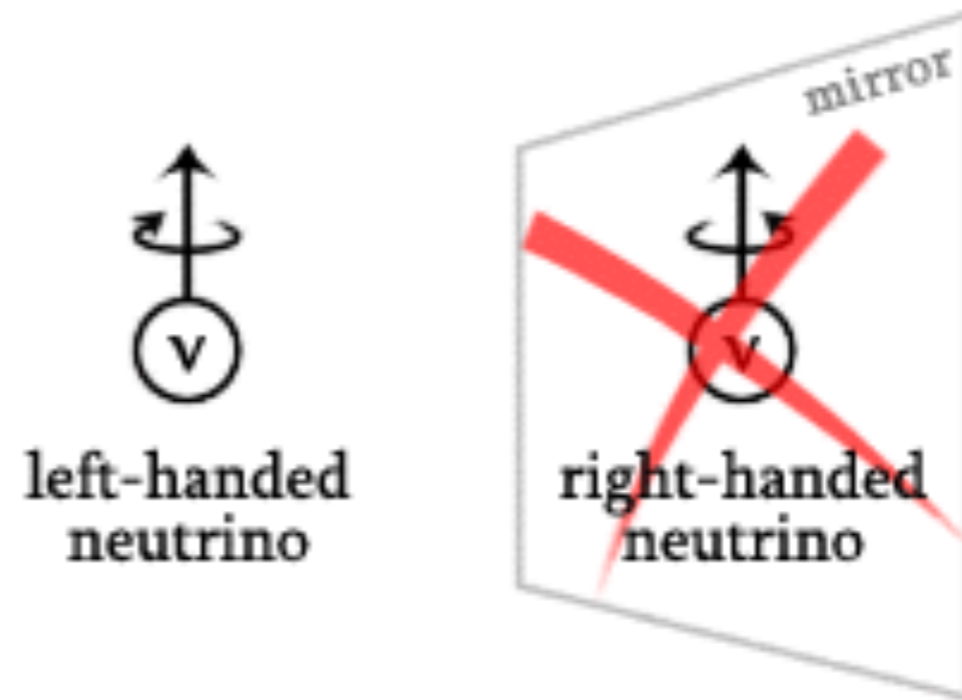
other:

- modified Newtonian dynamics (MOND)
- massive compact halo objects (MACHO)

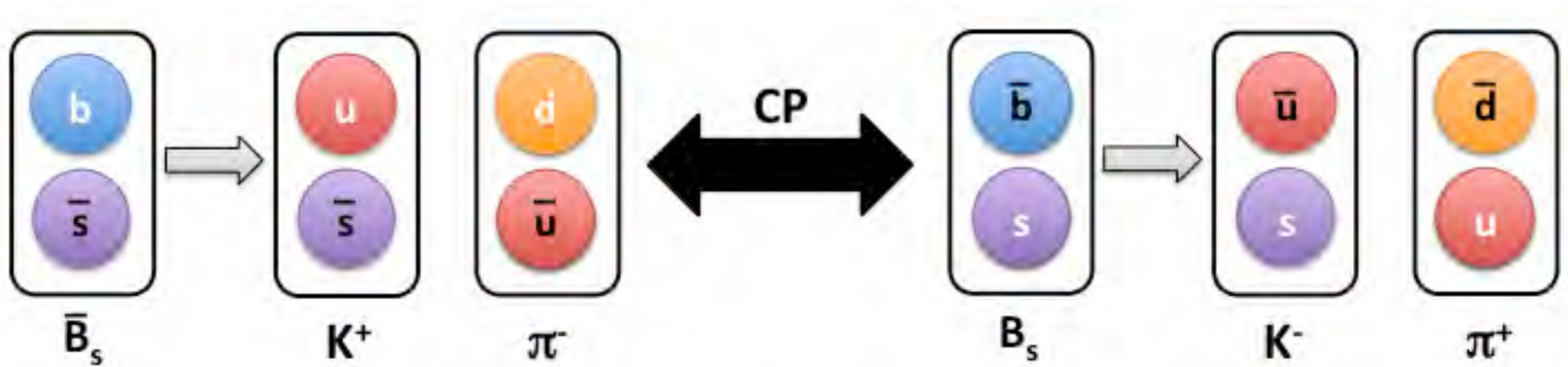
subjective
plausibility: 5%

P violation

- **maximum** parity (P) violation in weak interaction
- well established in theory
- CP is the **true symmetry**



CP violation



- decay rates **differ** at percent level
- CP symmetry is violated
- well established in EWK theory
- CPT is the **true symmetry**

strong CP problem

- CP violation
 - **not** established in strong interaction

$$\mathcal{L}_\theta = \theta \frac{g_s^2}{32\pi^2} \tilde{G}^{a\mu\nu} G_{\mu\nu}^a$$

- CP violating term
- θ is arbitrary angle

• gluon field

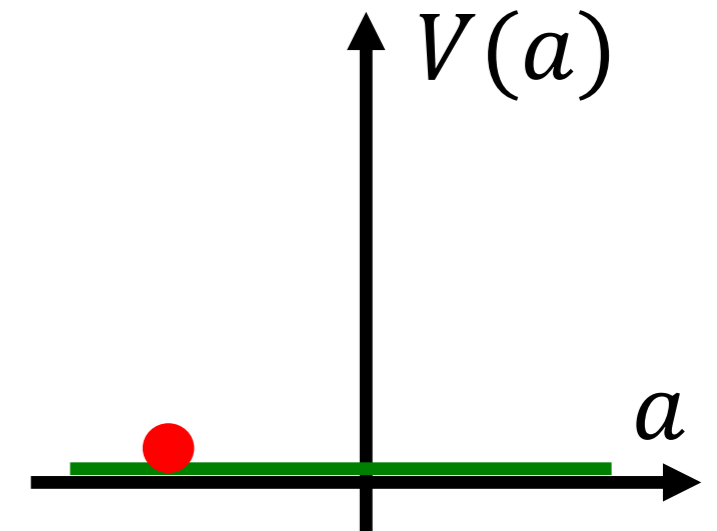
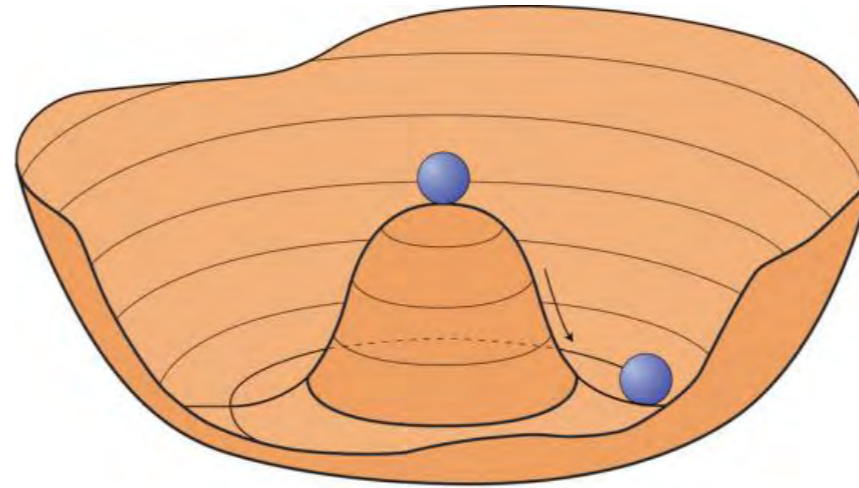
- why is $\theta == 0$?

- possible solution:
 - θ is a new scalar field
 - minimum of QCD potential occurs at $\theta == 0$

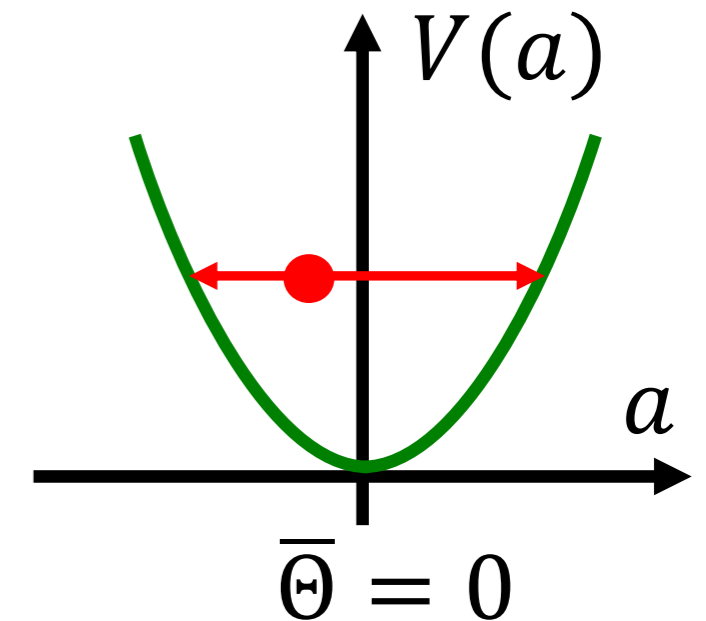
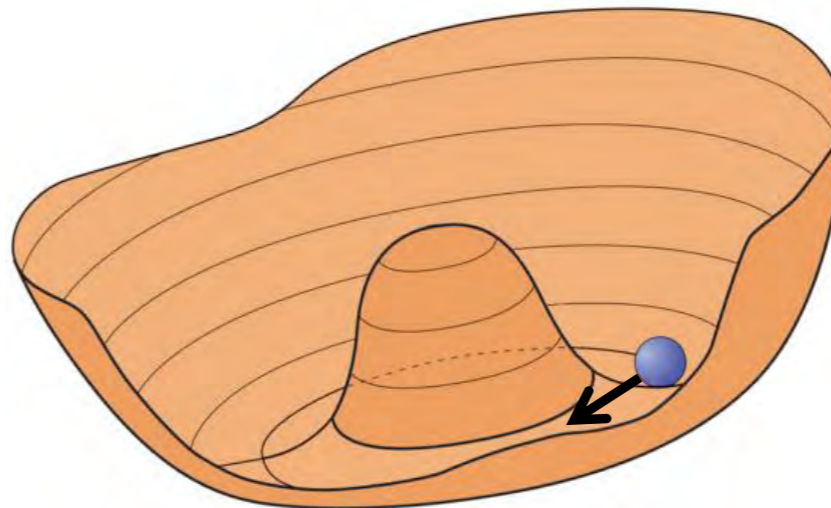
the axion in the early universe

pictures from G. Raffelt
(MADMAX workshop 2016)

- new $U(1)_{PQ}$ symmetry breaks at high scale f_a



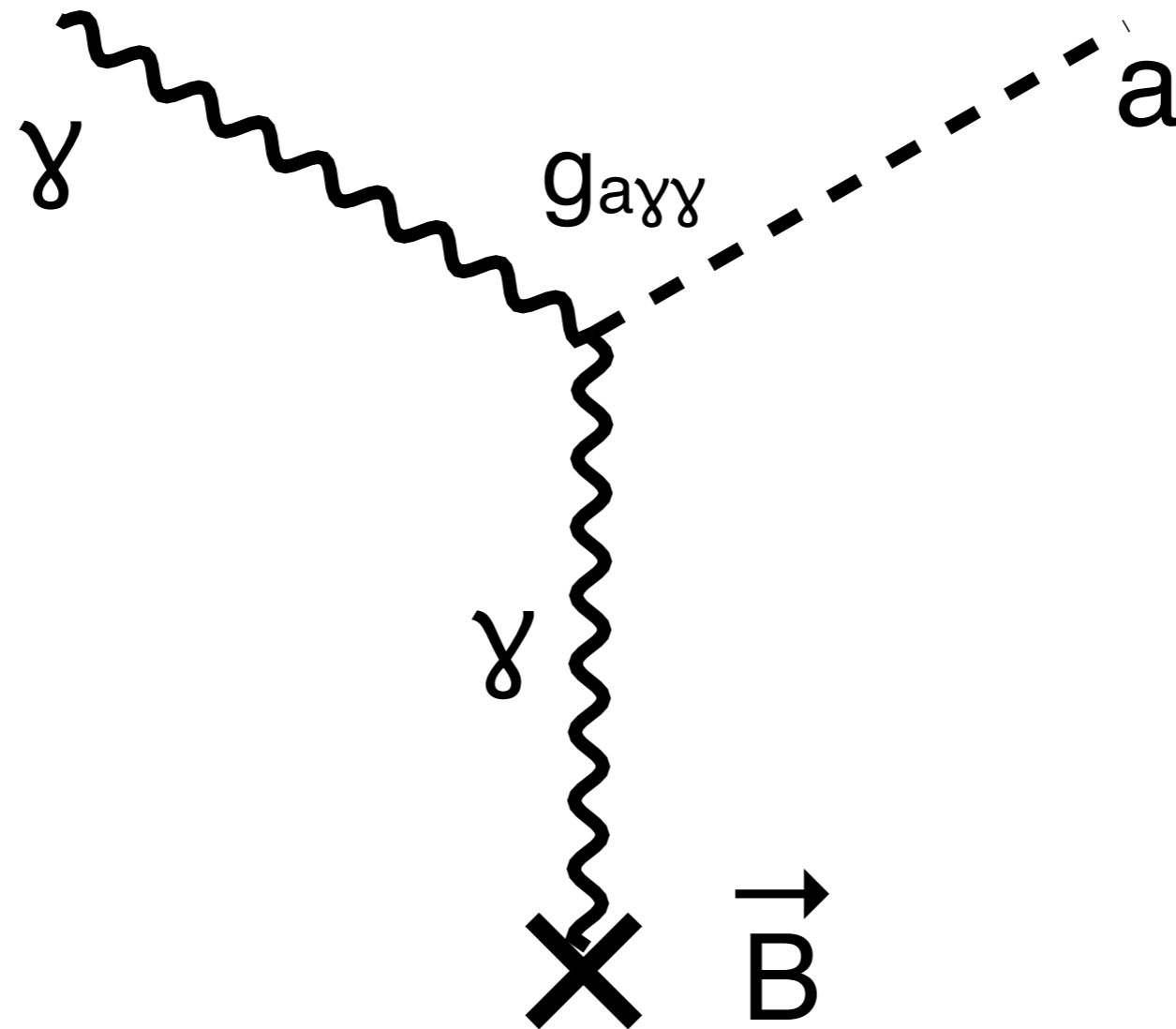
- potential changes shape when universe cools down ($T \sim 1 \text{ GeV}$)
- axion acquires mass
- field starts oscillating
- expected density compatible with DM



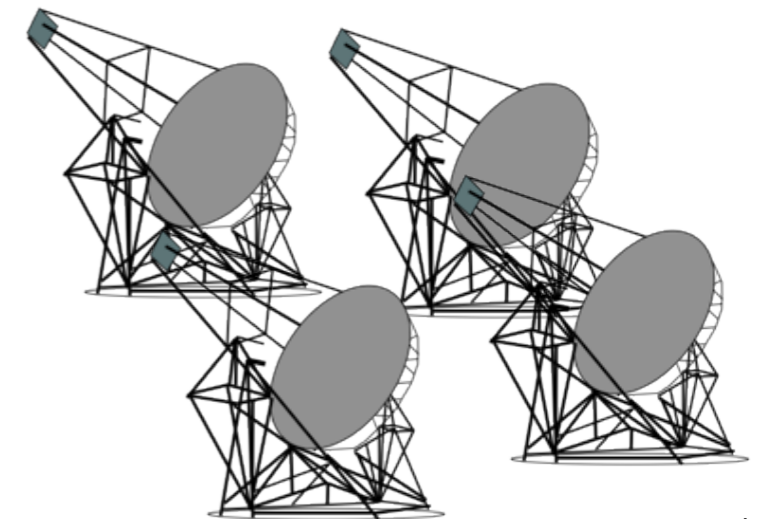
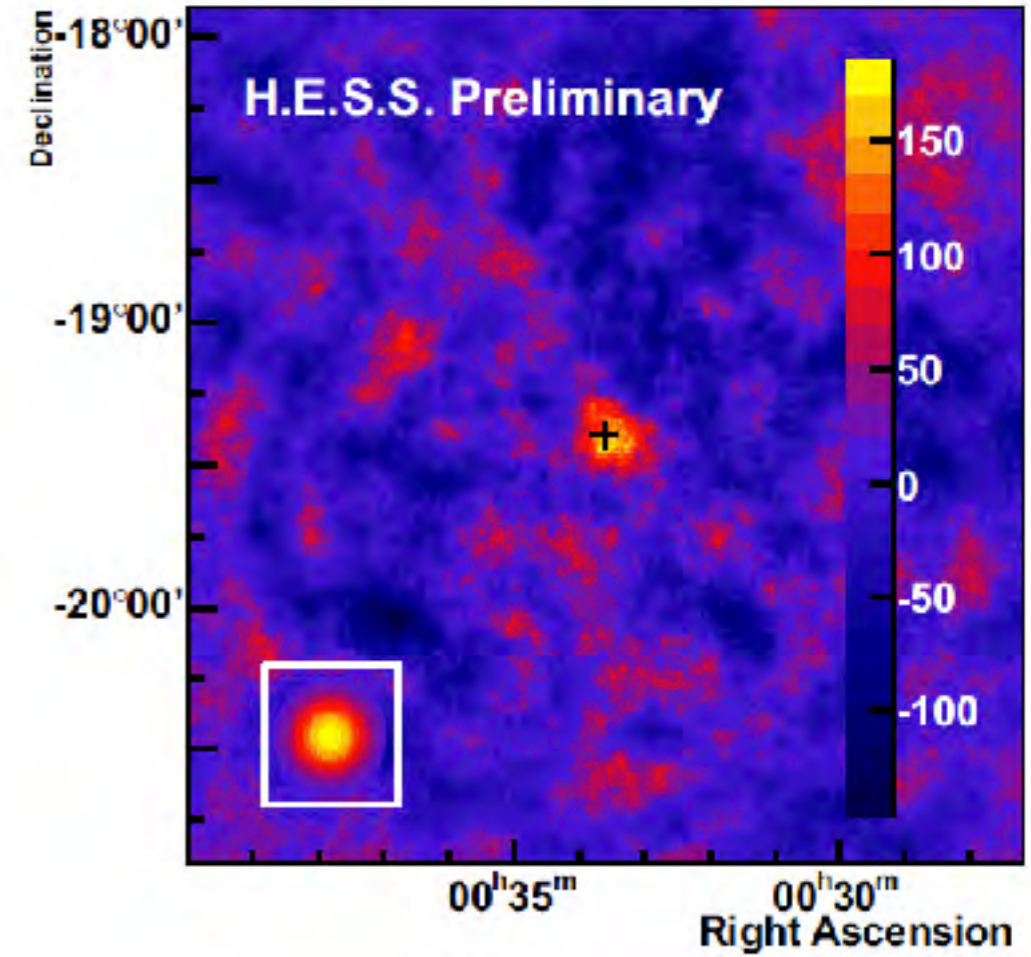
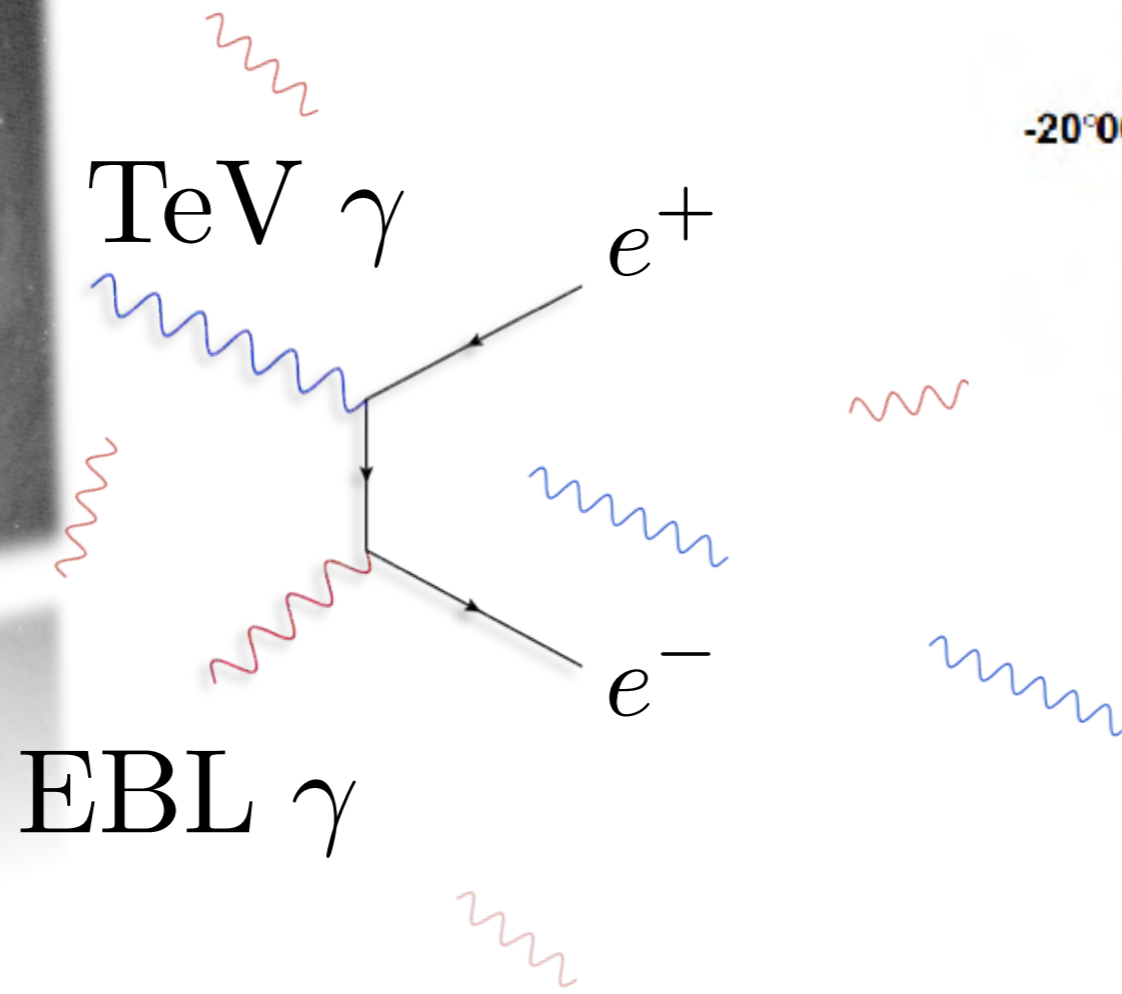
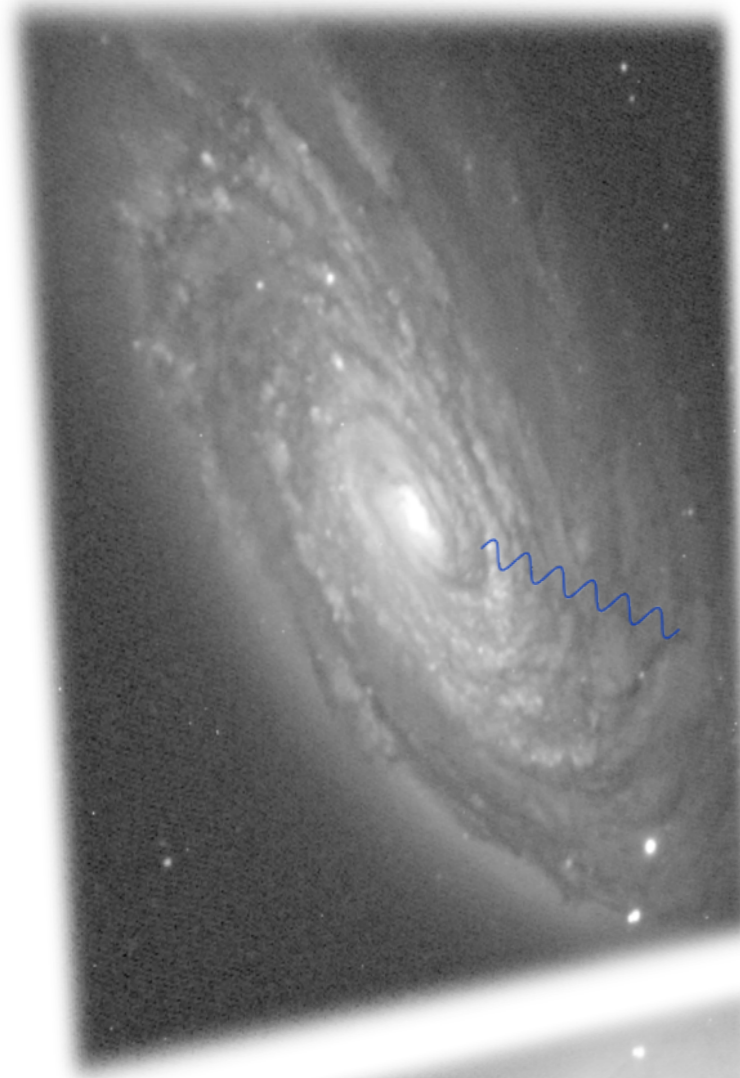
axions born from vacuum realignment

axion couplings predicted

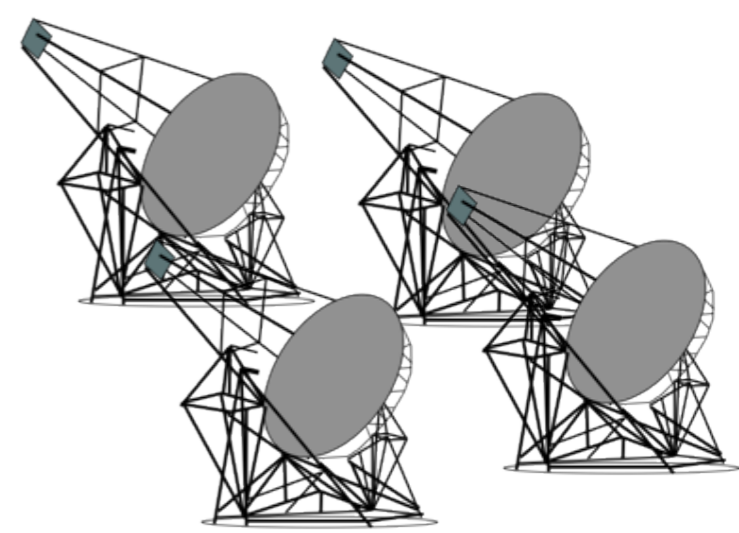
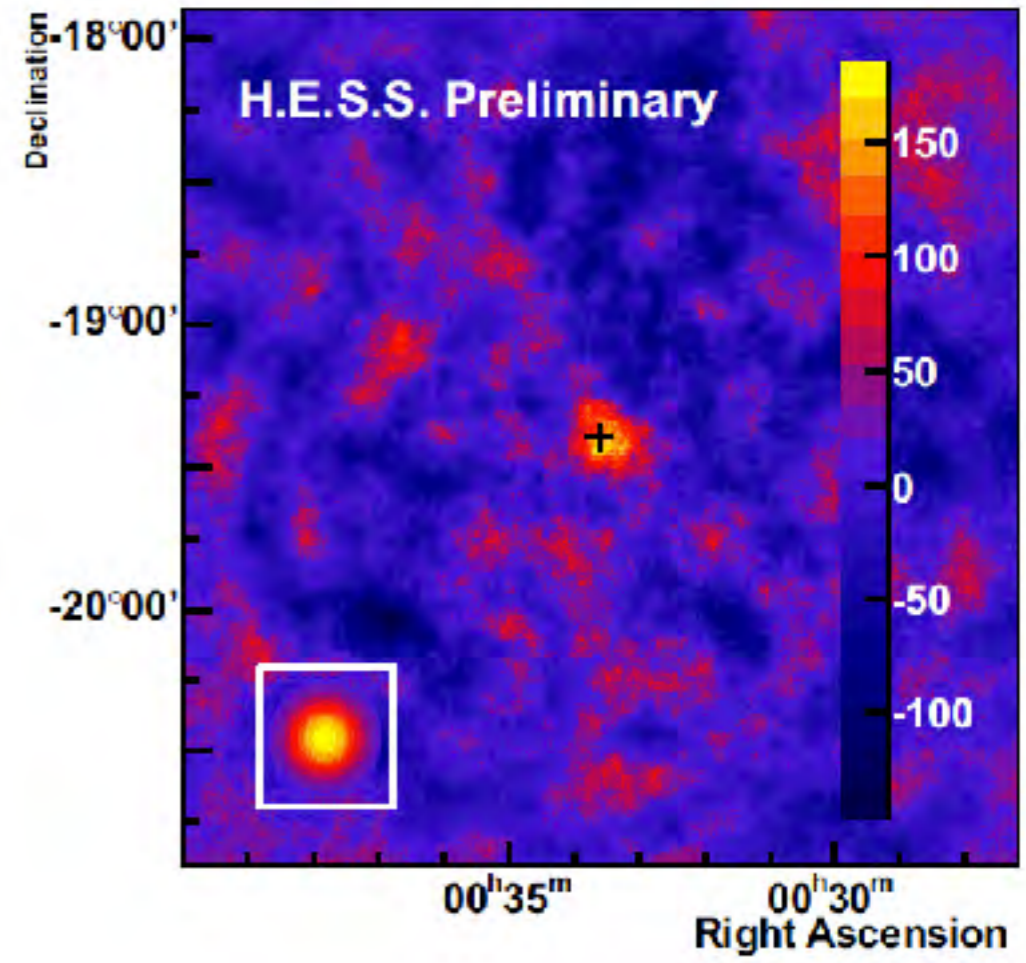
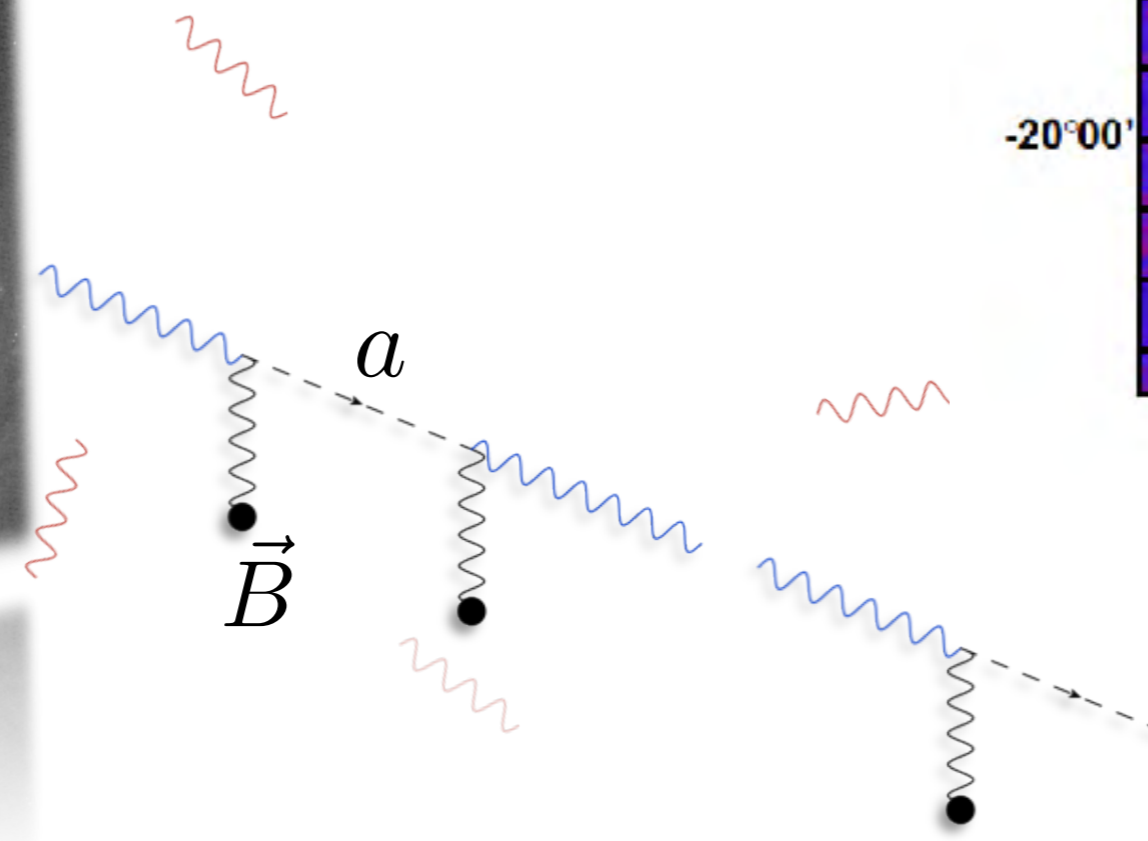
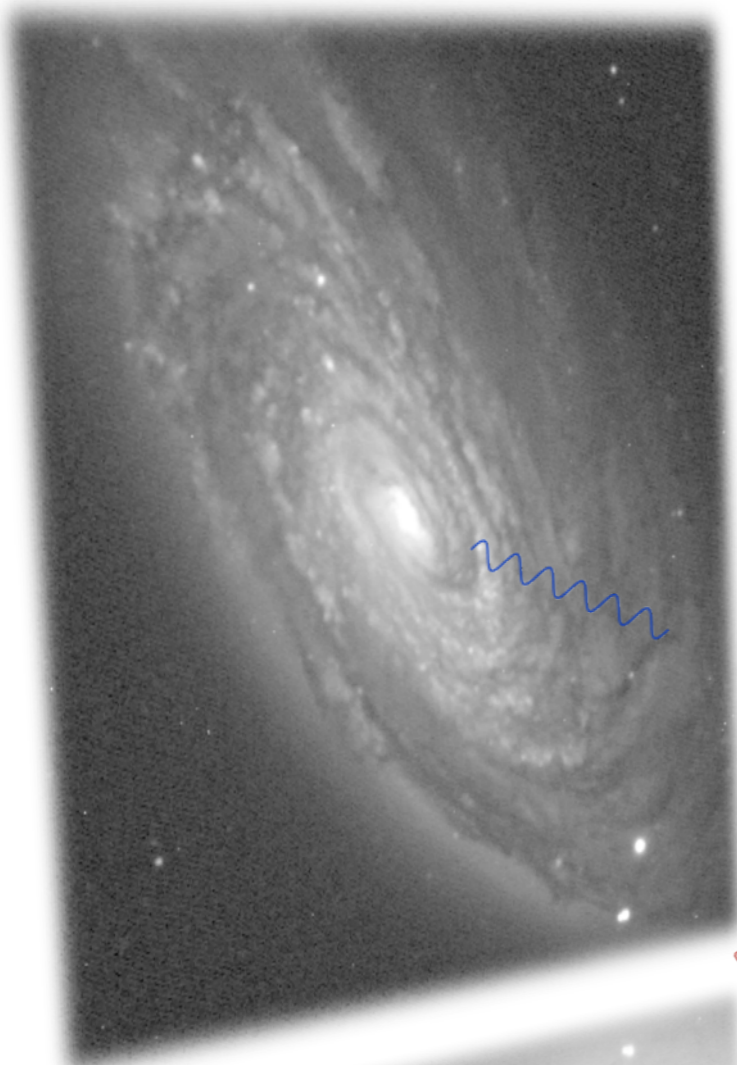
- Axion-photon coupling



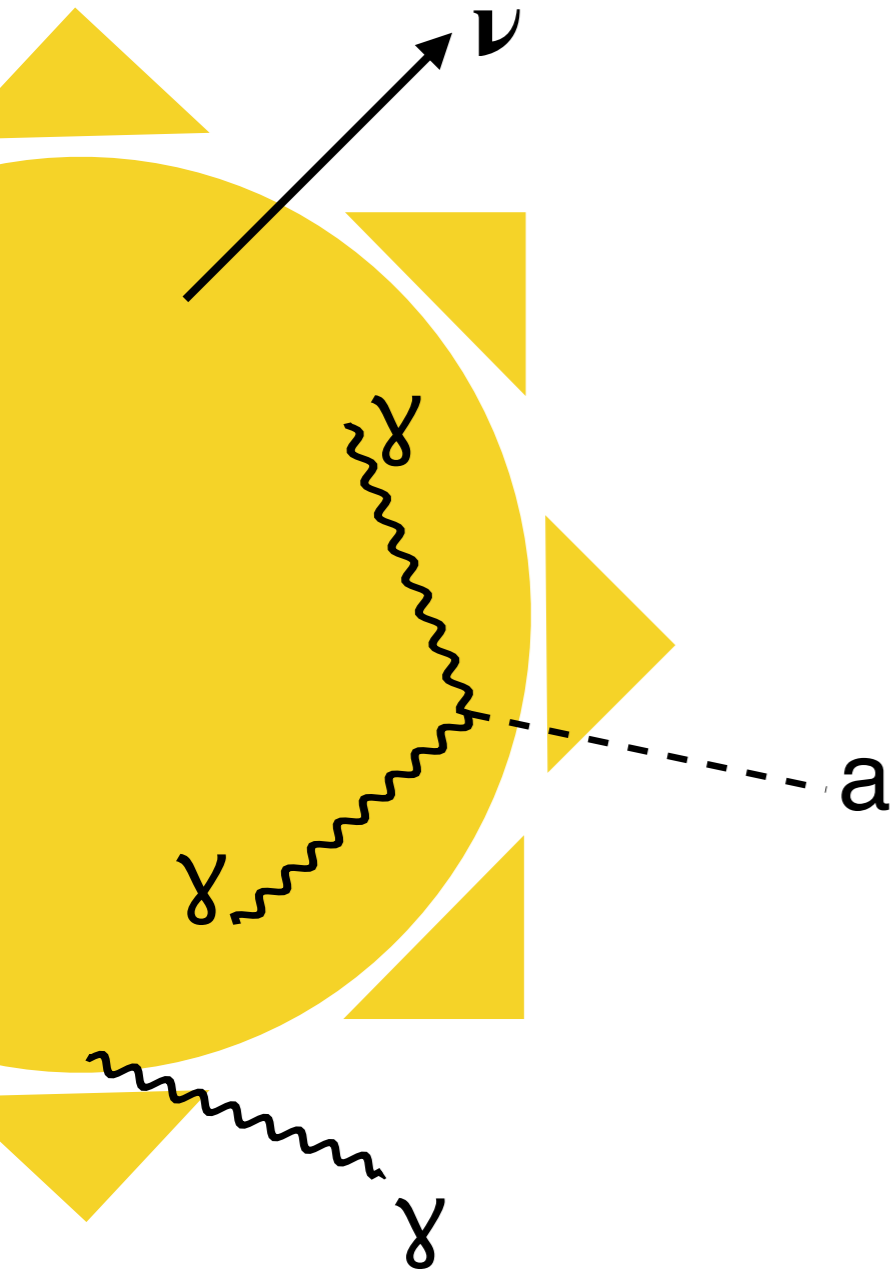
transparency of the universe



transparency of the universe



stellar cooling



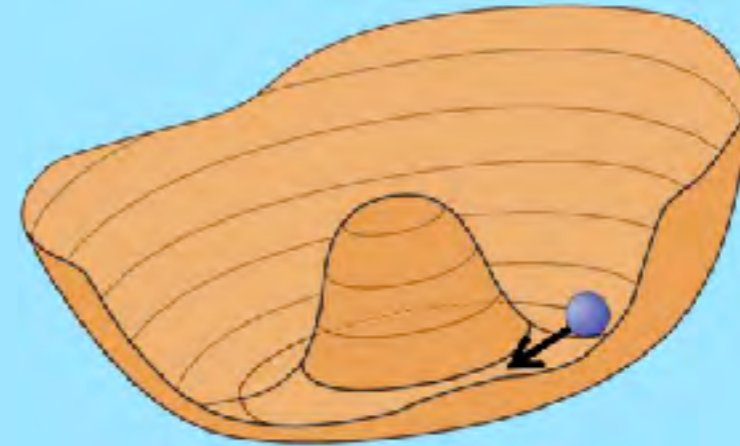
- excessive cooling in different types of stellar objects:
 - white dwarfs, red giants, ...
- single observation not statistically significant
 - combination is significant
- straight forward explanation with axions
 - hypothetical axion emission

axions solve many problems simultaneously

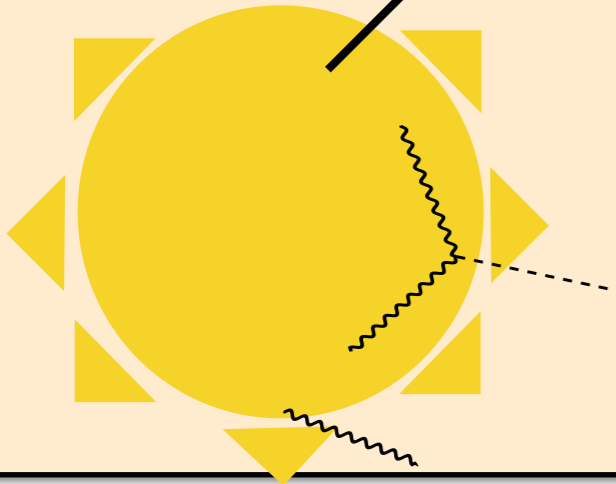
- strong CP problem

$$\mathcal{L}_\theta = \theta \frac{g_s^2}{32\pi^2} \tilde{G}^{a\mu\nu} G_{\mu\nu}^a$$

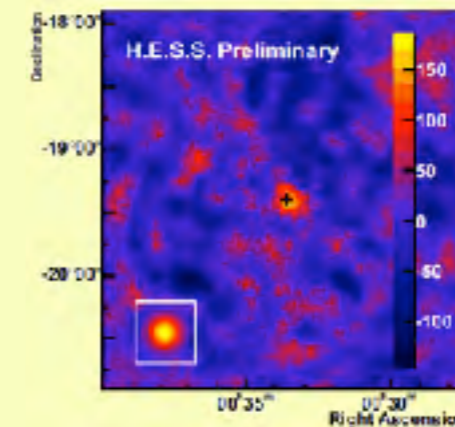
- dark matter



- stellar cooling

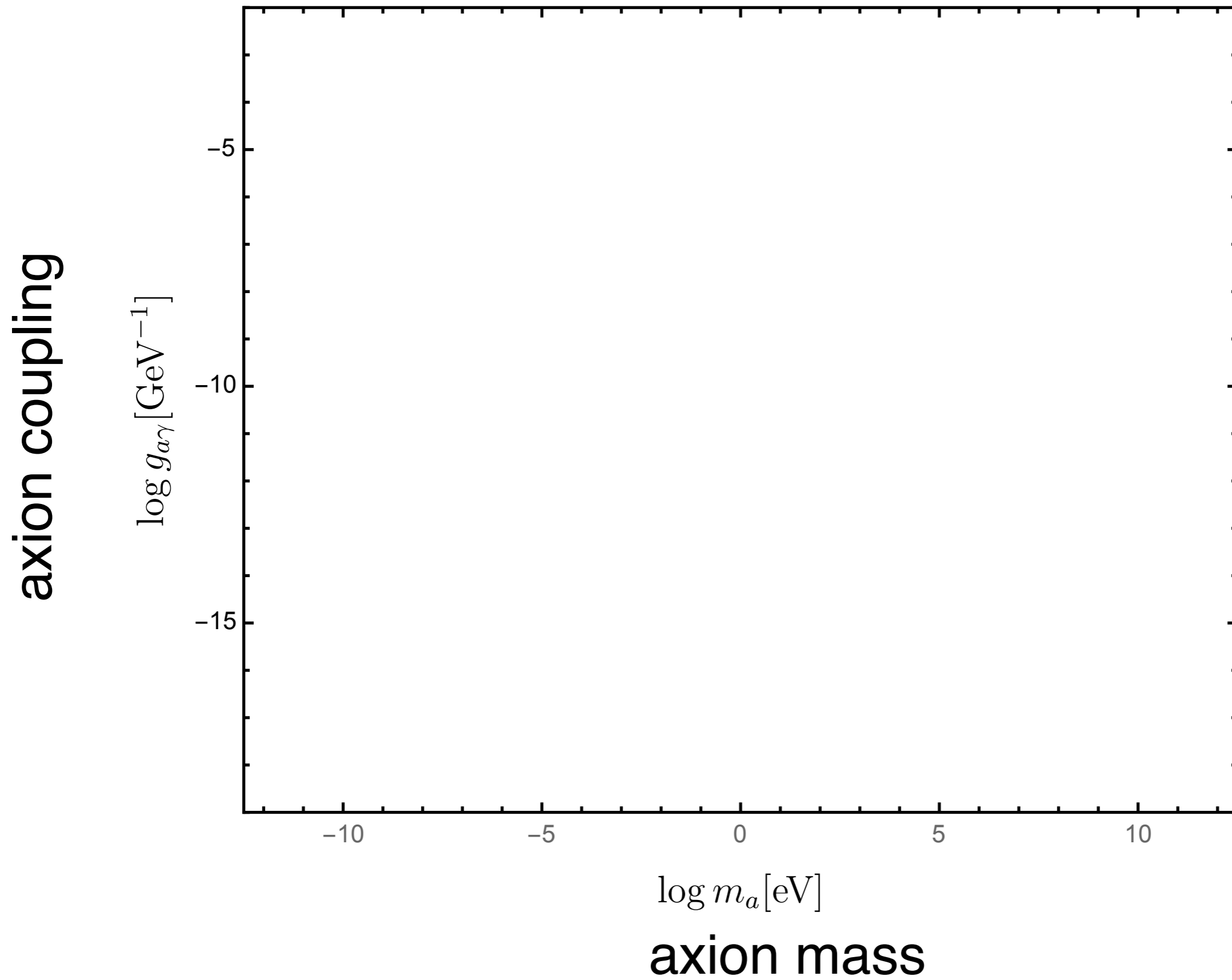


- transparency

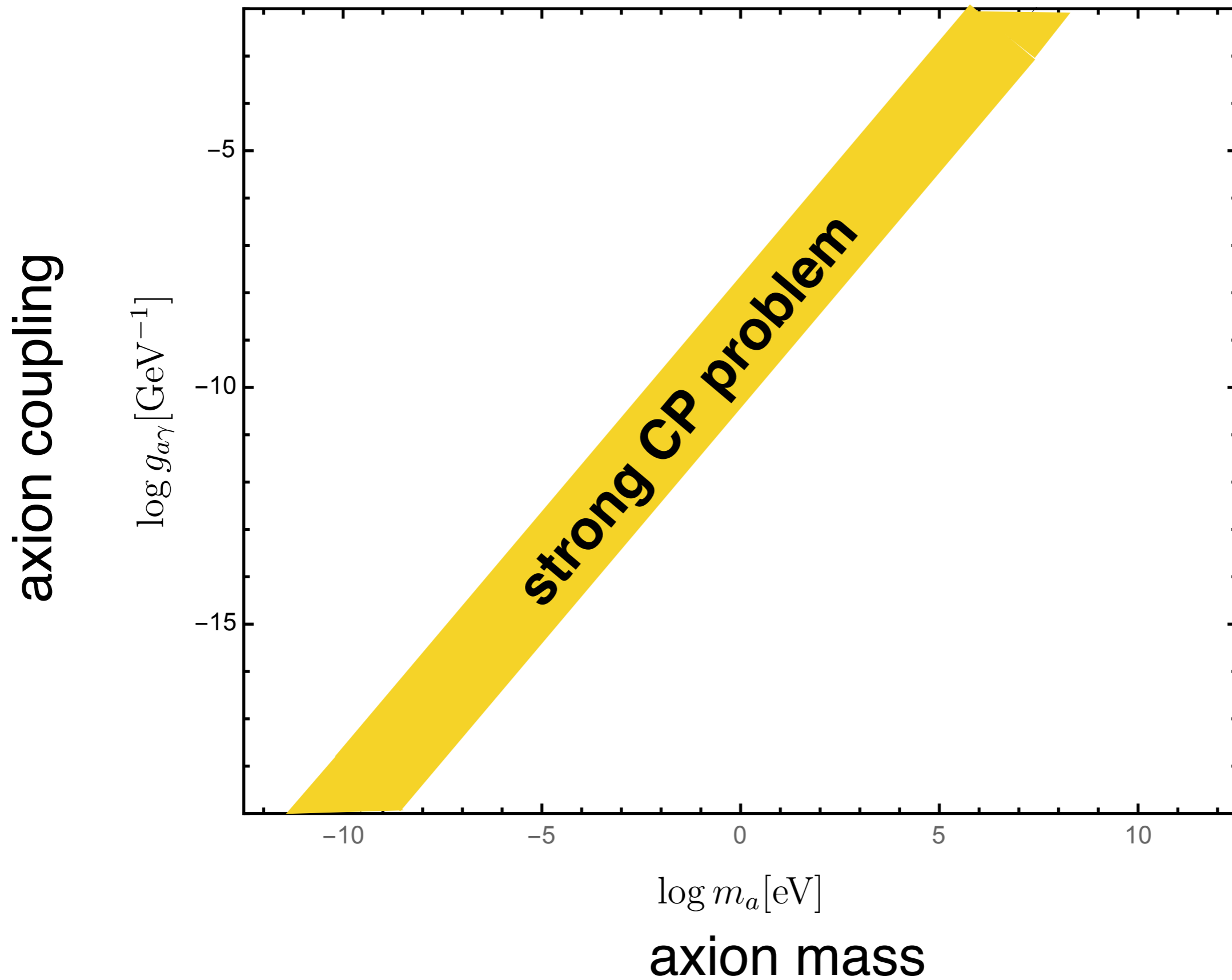


and it's not a new concept! Higgs!

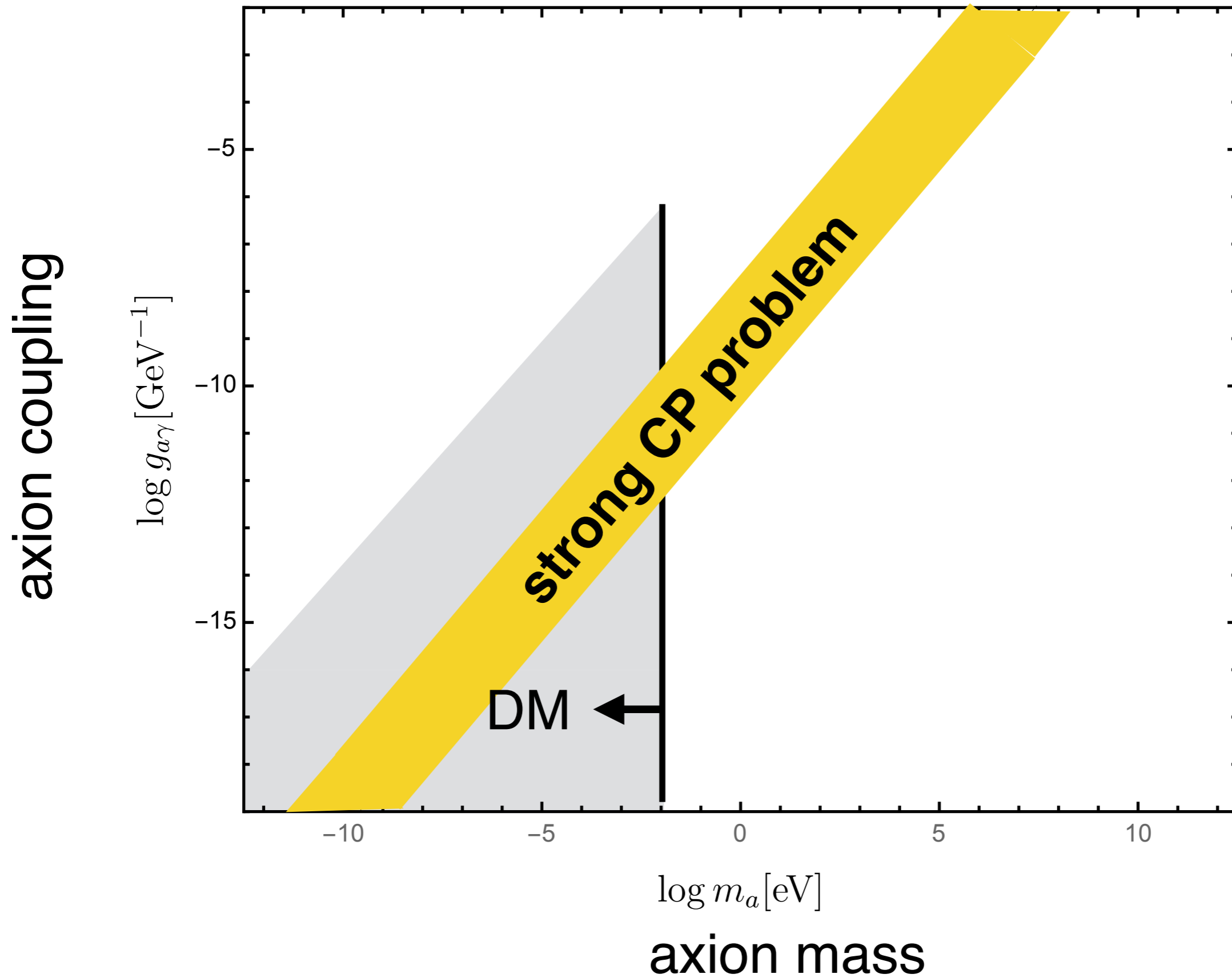
where to search for axions



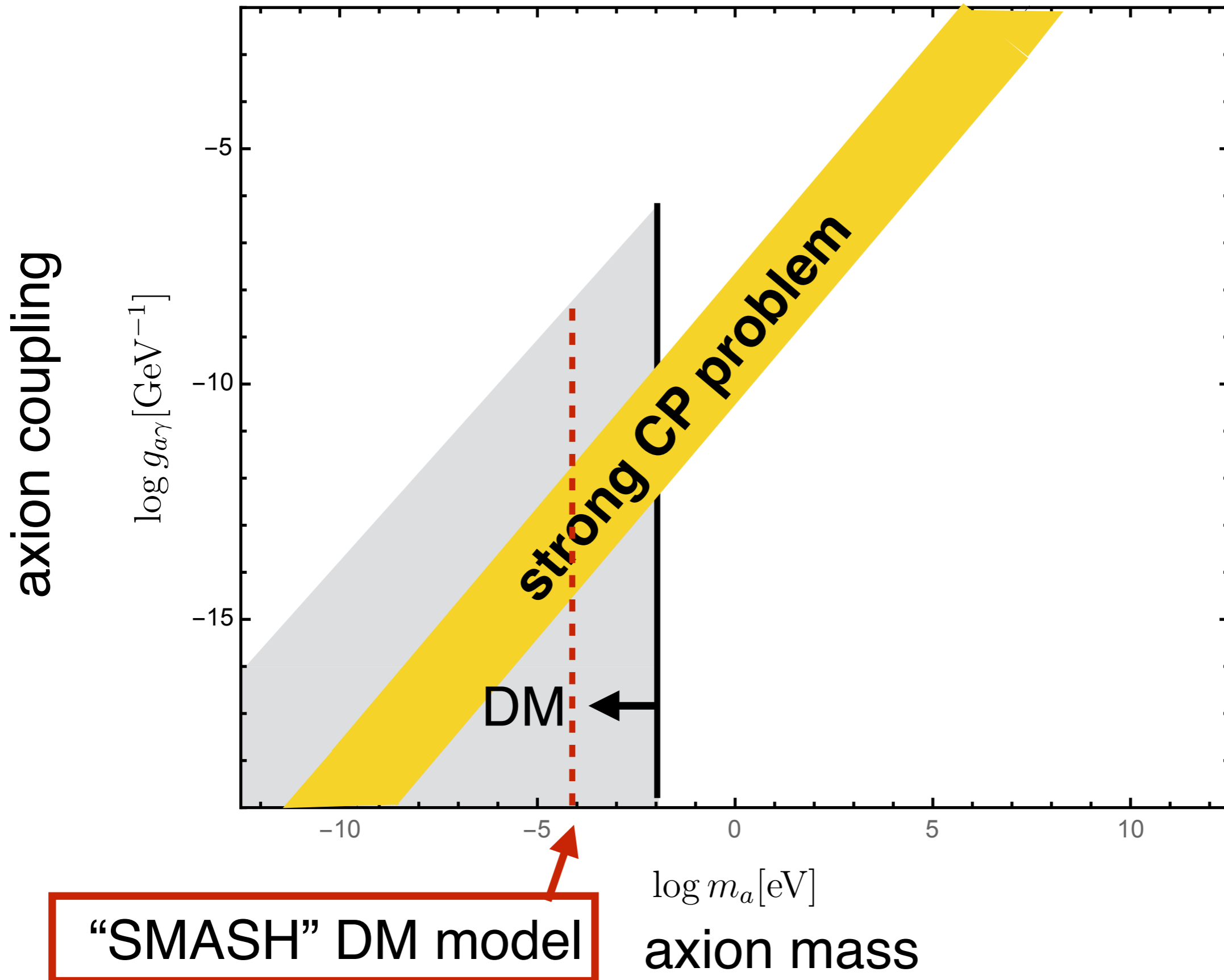
where to search for axions



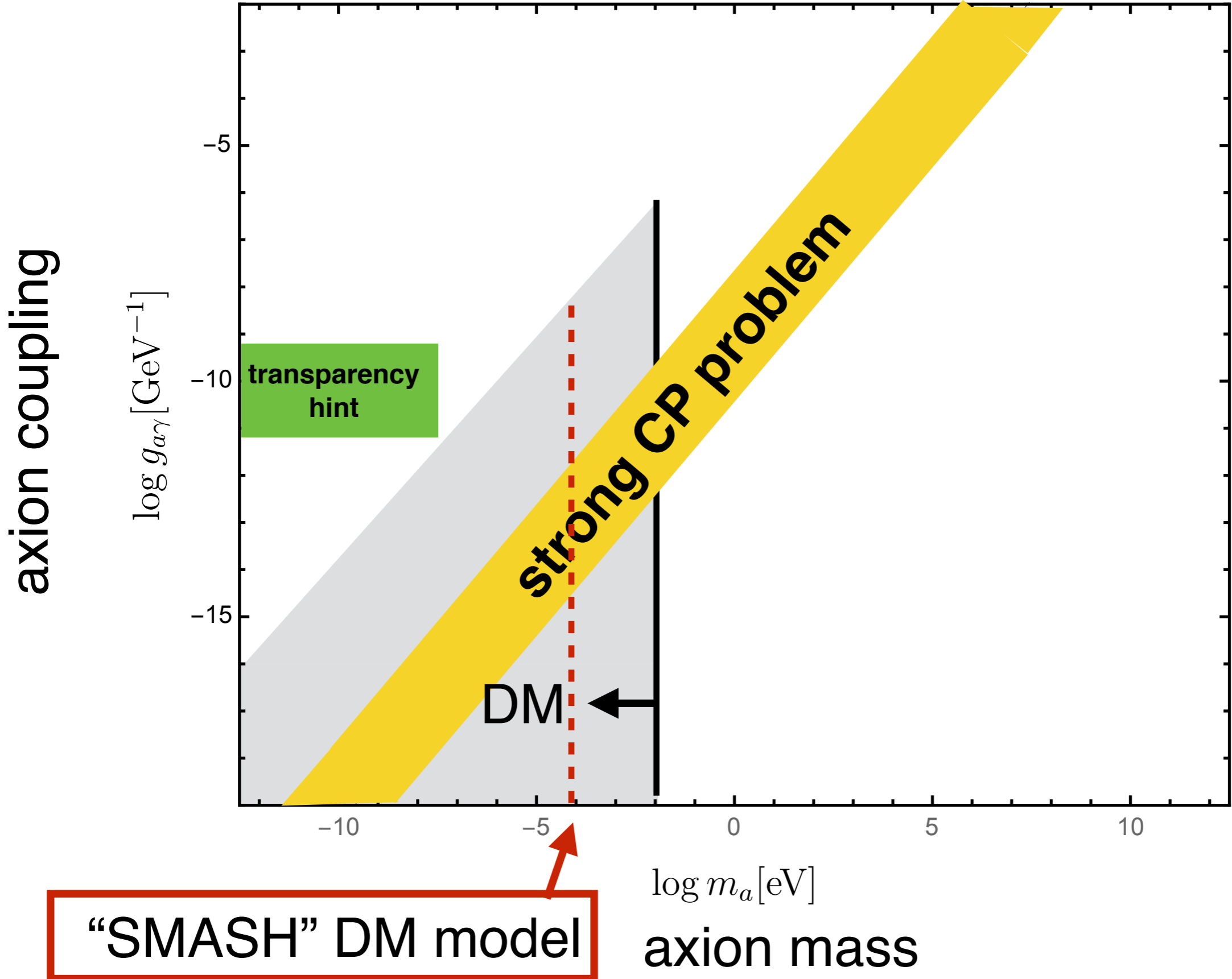
where to search for axions



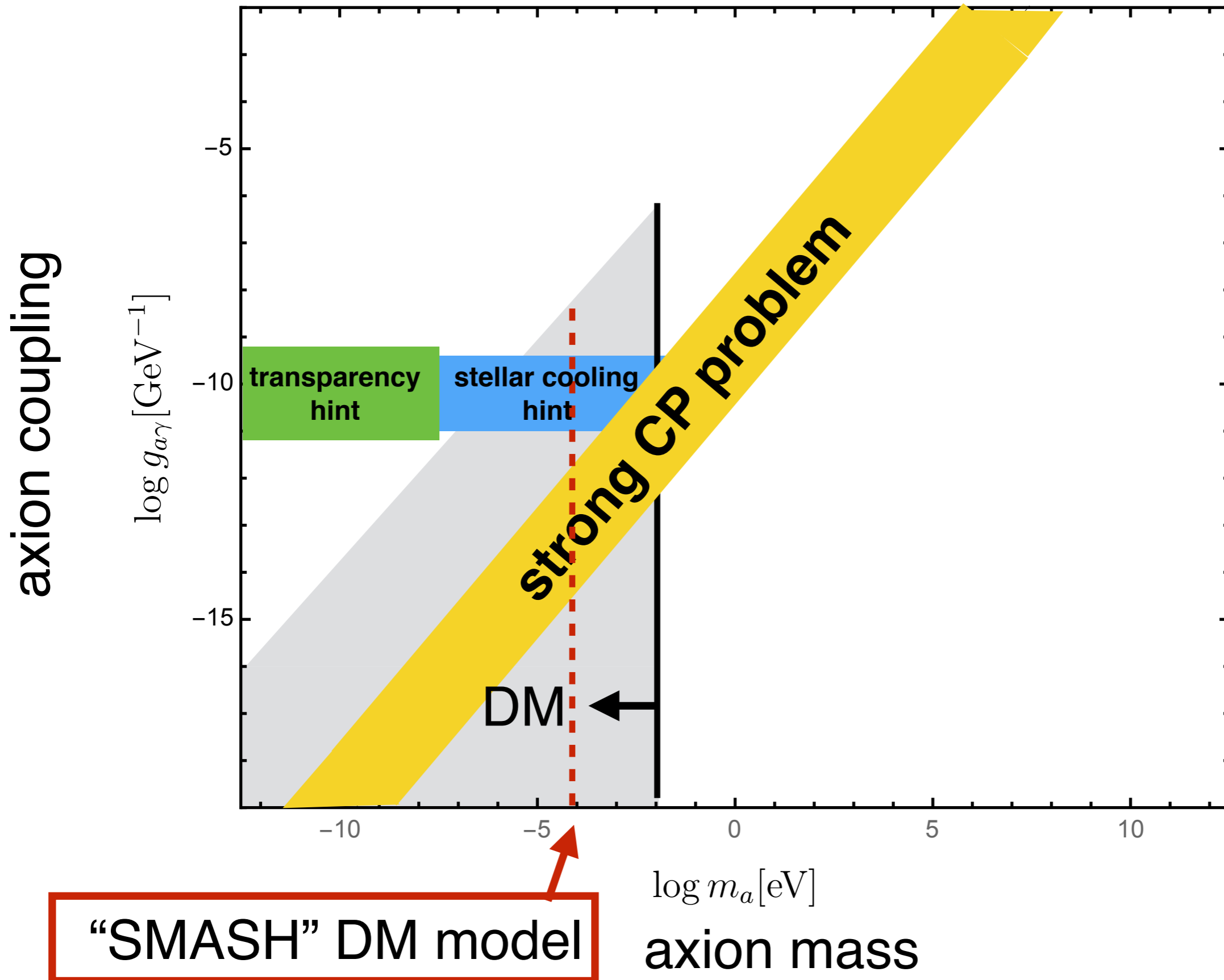
where to search for axions



where to search for axions

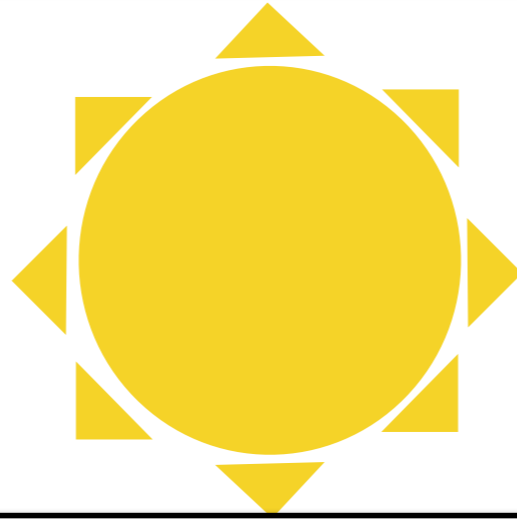


where to search for axions



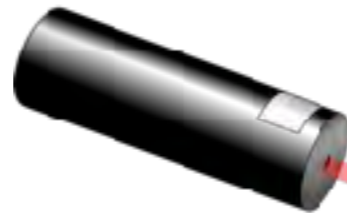
how to search for axions

- **helioscope**



- search for axions emitted by the sun

- **laboratory**



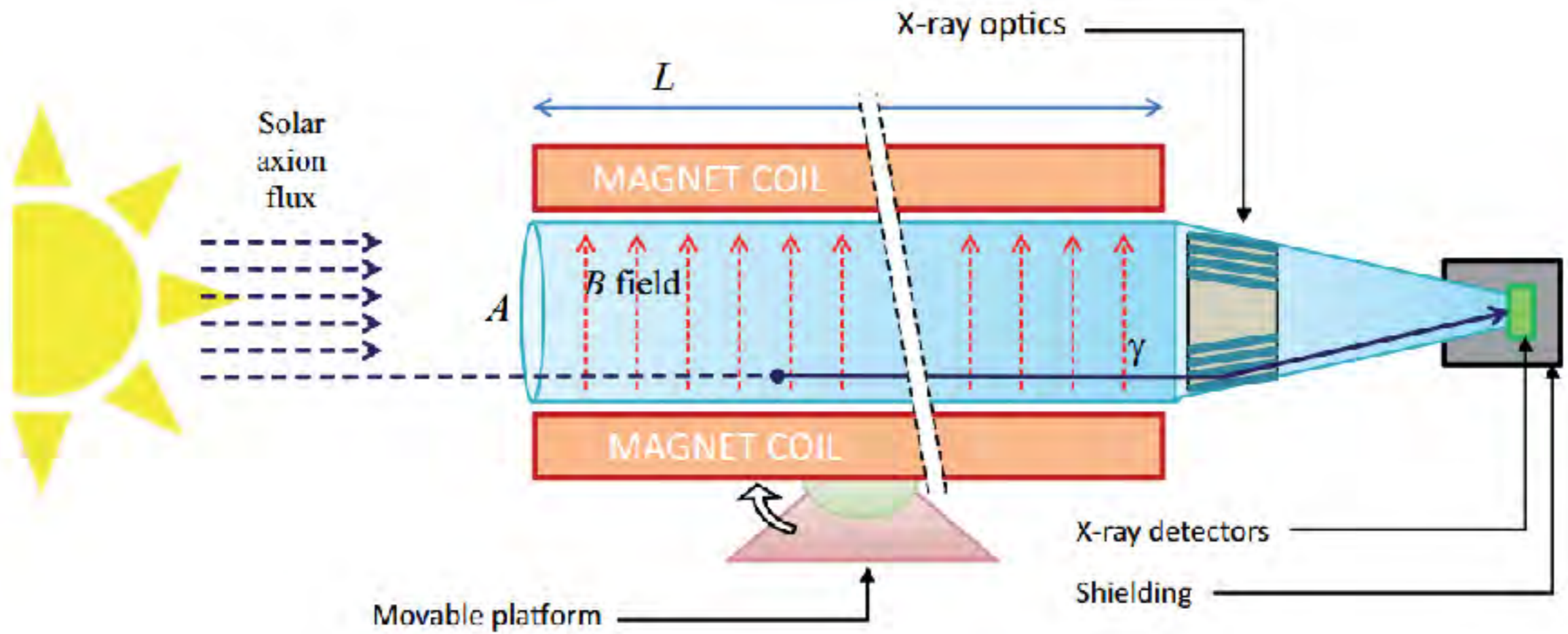
- produce axions with a laser and a B-field

- **haloscope**



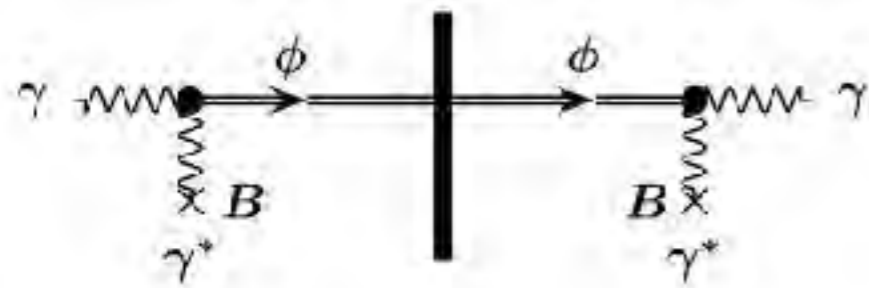
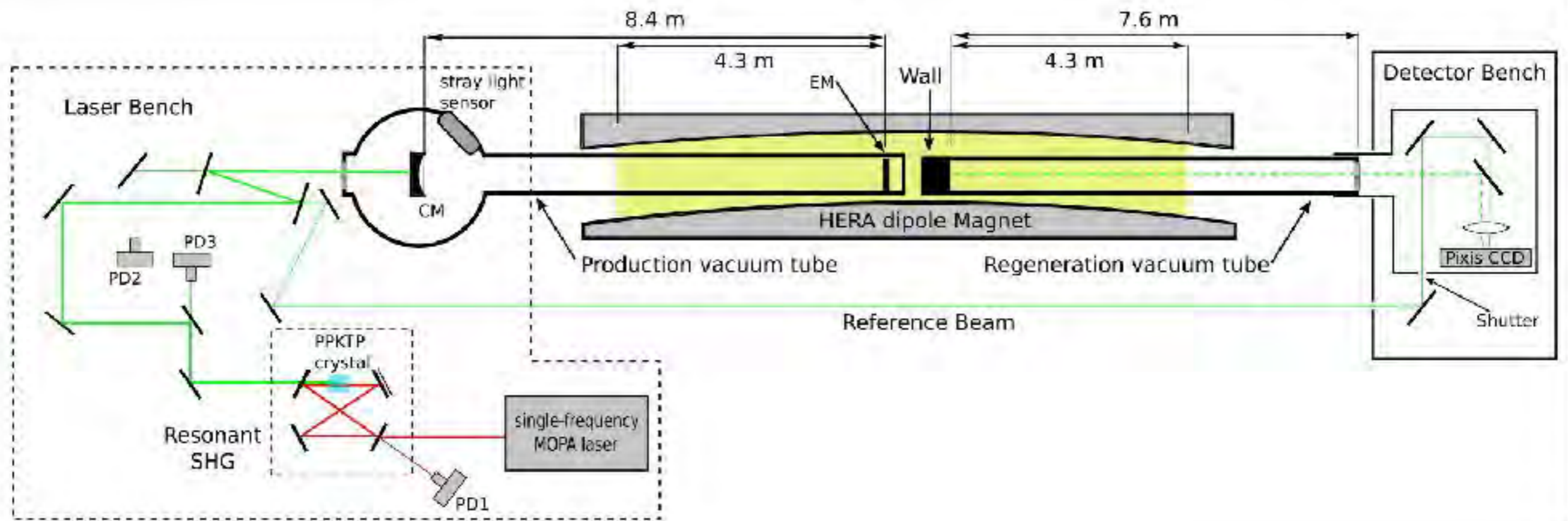
- find axions in the galactic DM halo

helioscope: "CAST"



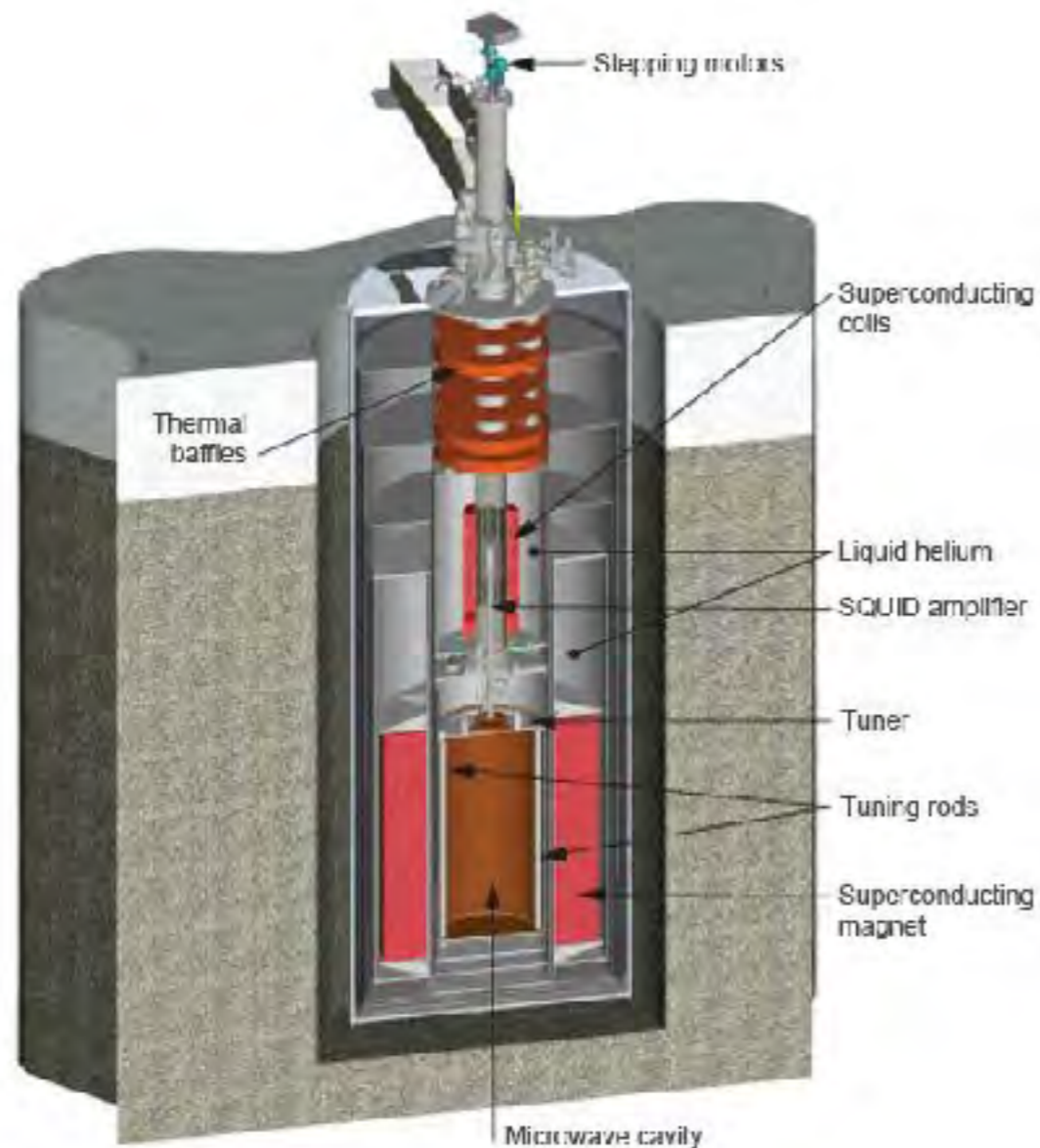
- in operation since 2003

laboratory: "ALPS"

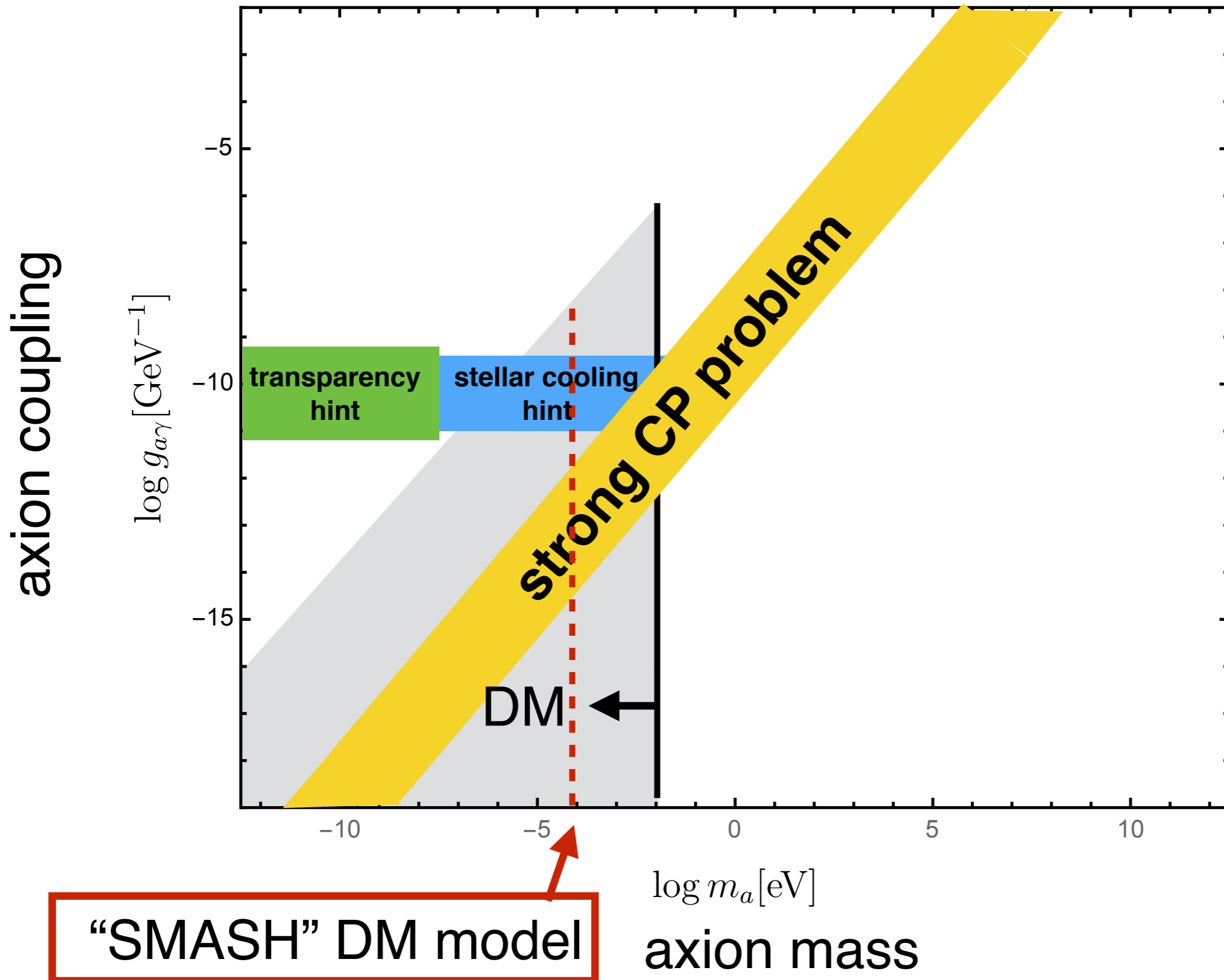


haloscope: “ADMX”

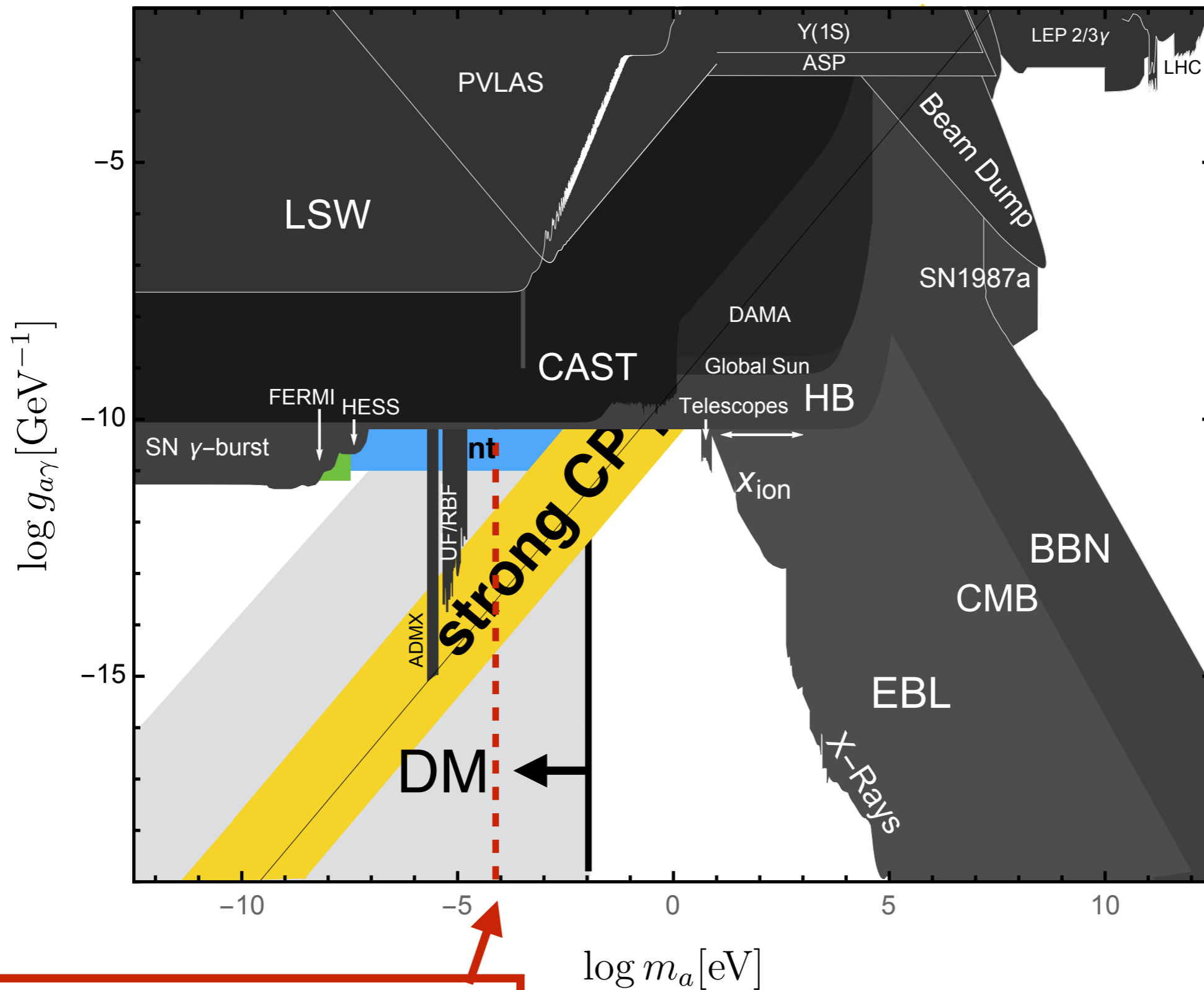
- the galactic DM halo is ...**here!**
- axions in a B field emit photons
 - **too weak to detect directly**
- enhanced with resonator



current constraints



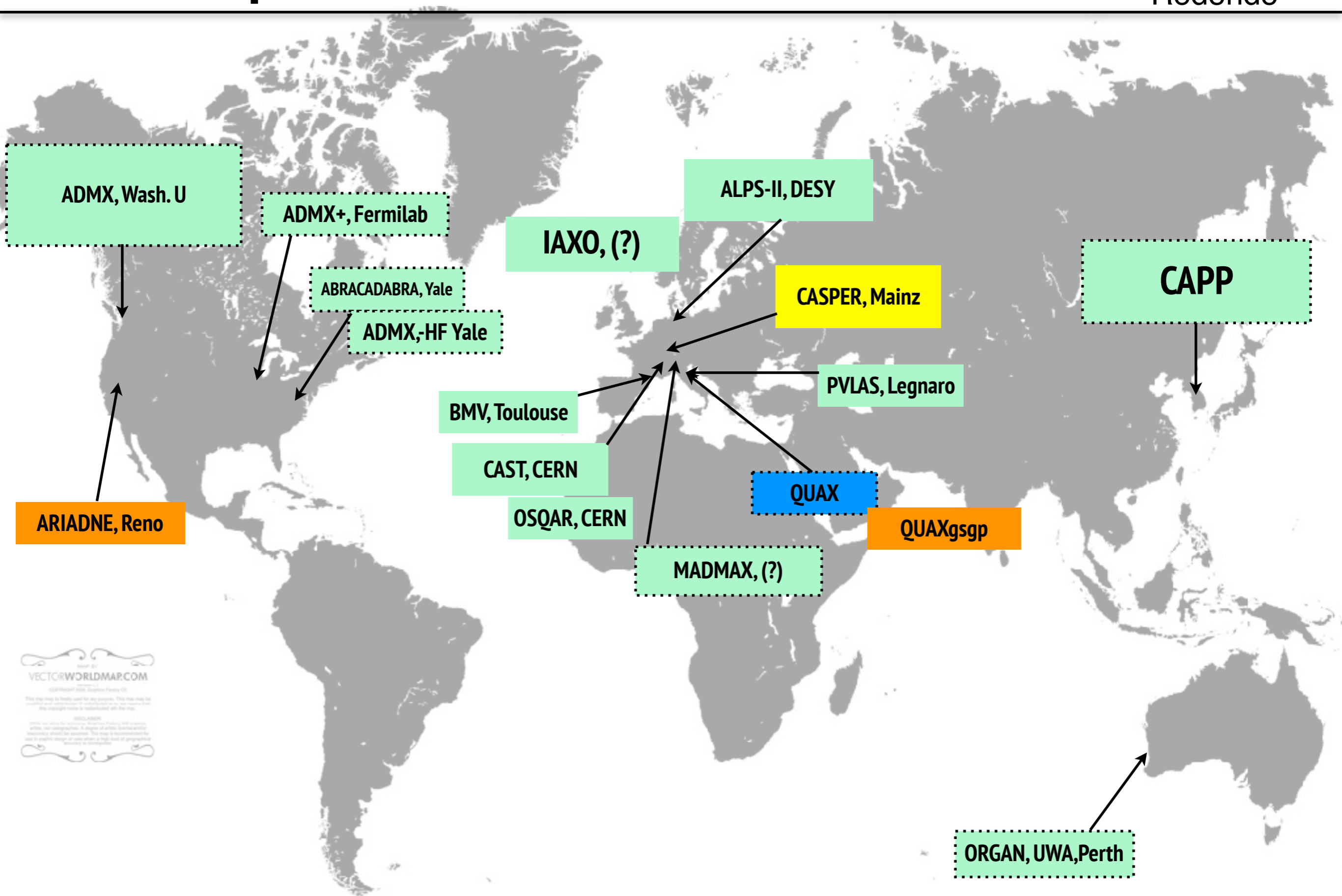
axion coupling



“SMASH” DM model

future experiments

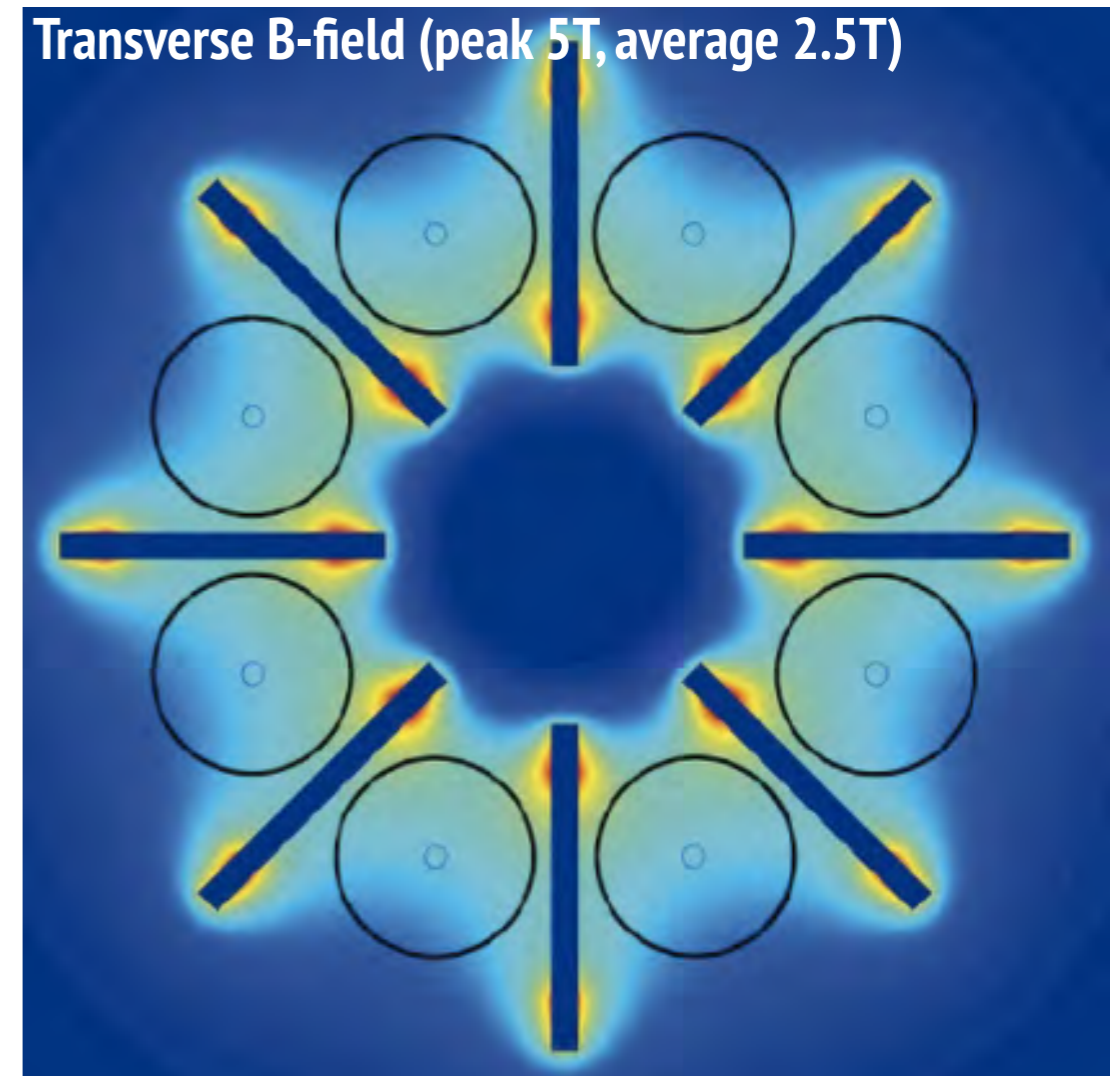
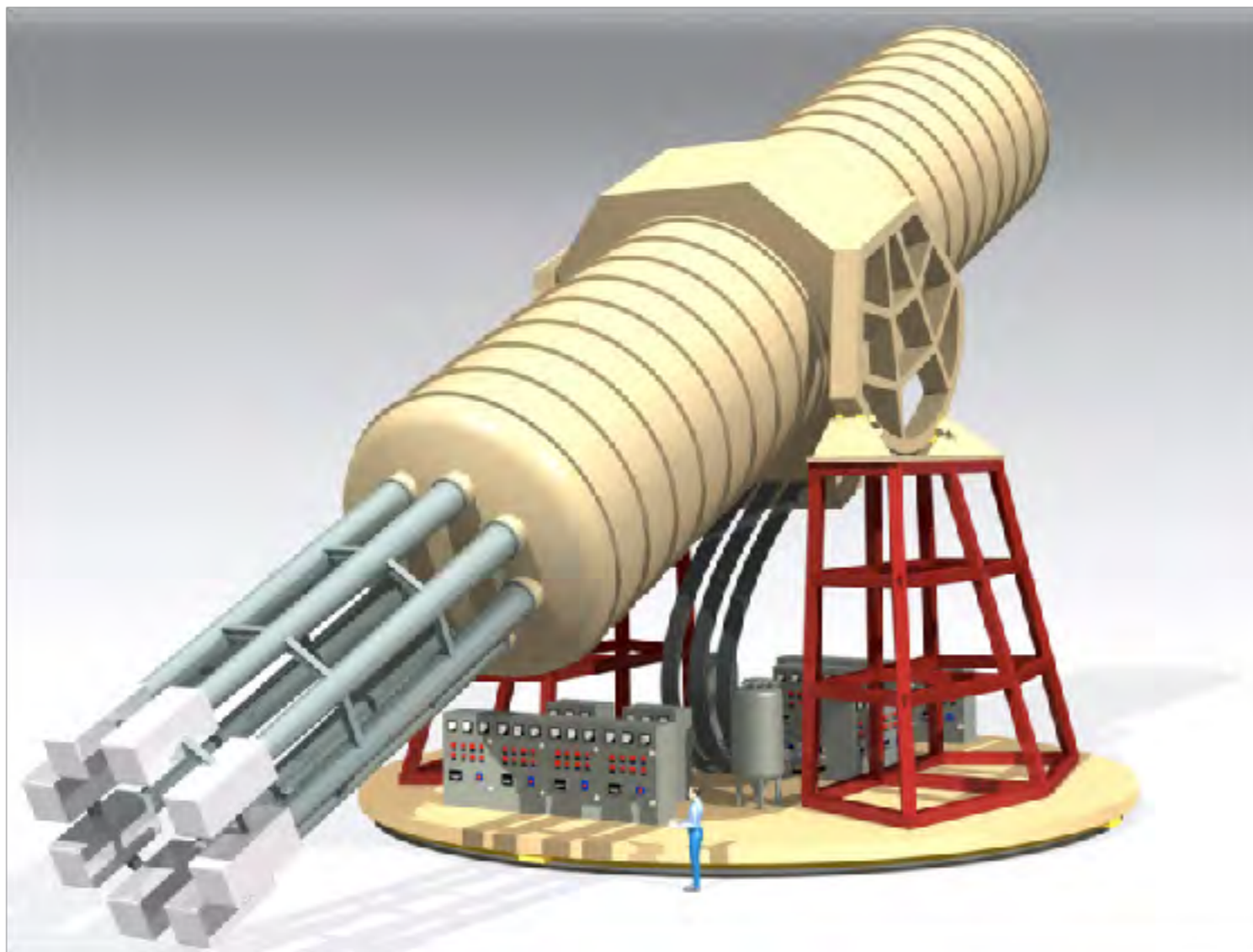
map from J. Redondo



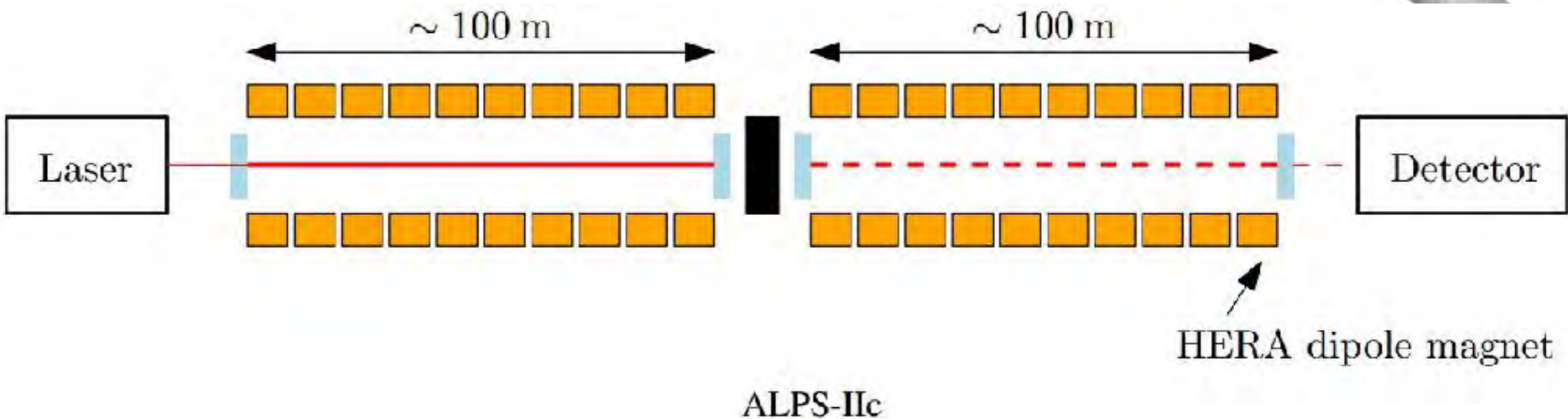
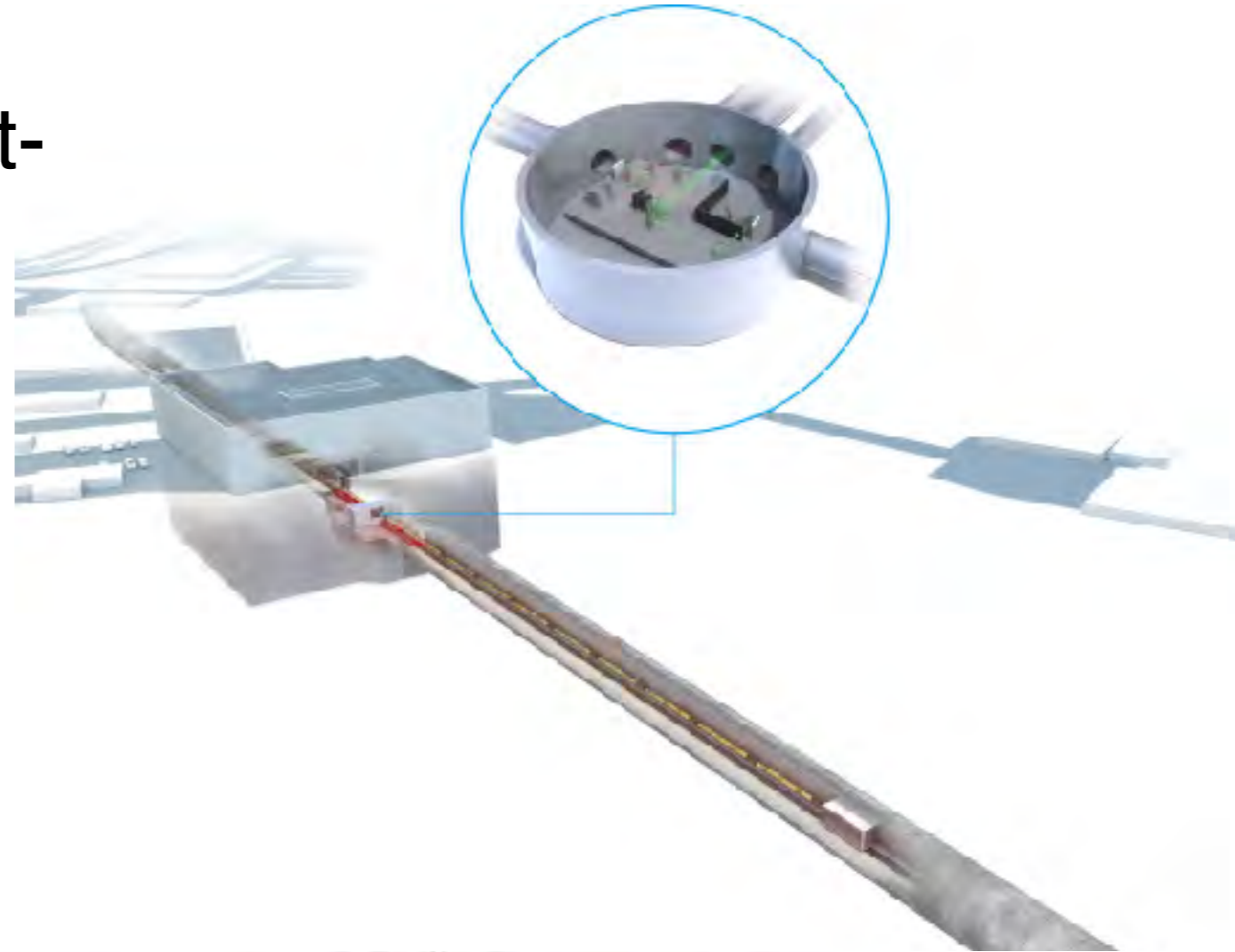
Map by
VECTORWORLD.MAR.COM
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accuracy is not required.

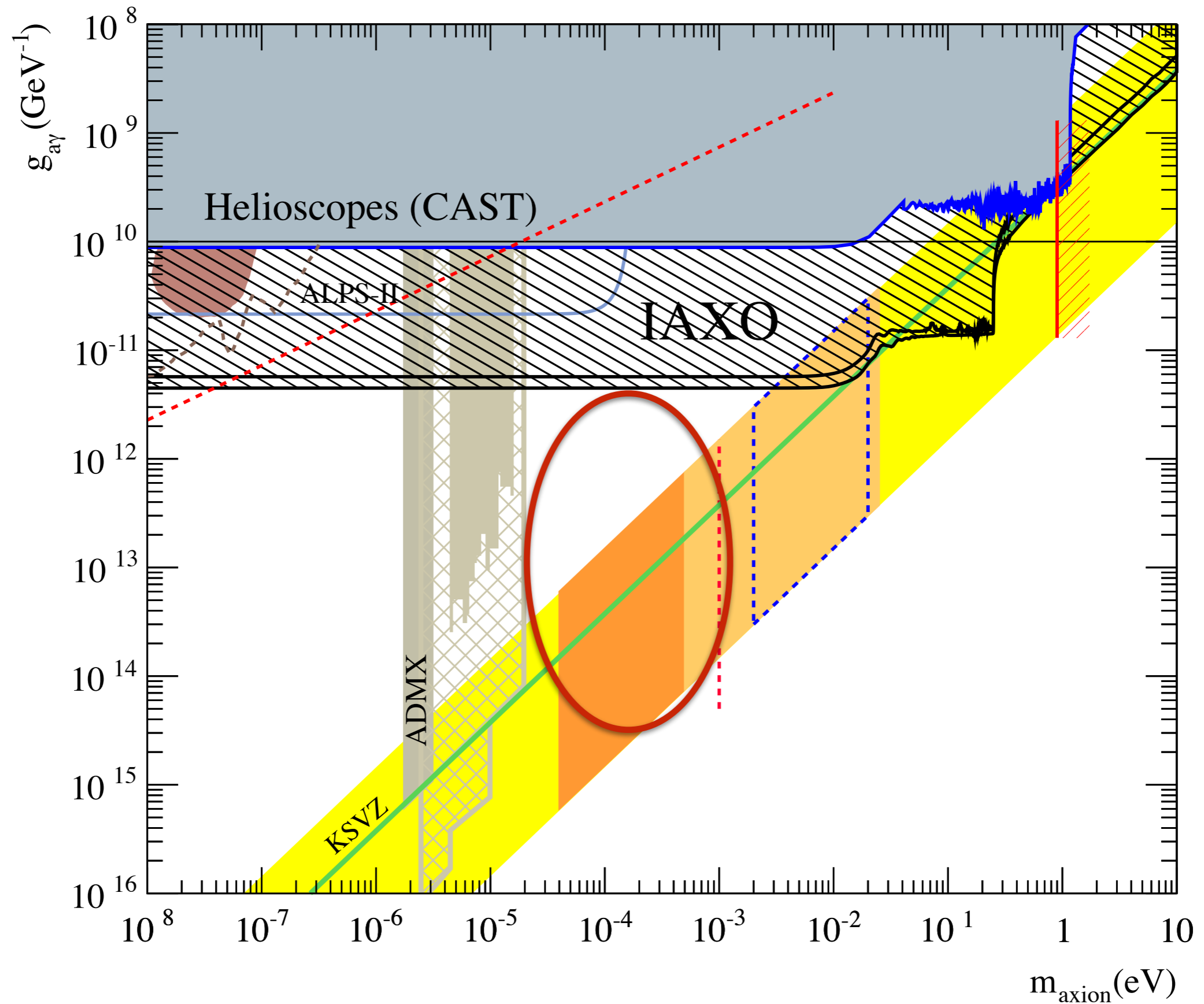
- IAXO: next generation helioscope
 - larger, stronger magnet
- when and how not yet clear

- more than one order of magnitude improvement over CAST



- ALPS-II: next generation light-shining-through-the-wall
 - more magnets, better laser, better detector
 - resonator in regeneration volume
- clear schedule at DESY
 - several stages: ALPS II a-c
 - factor 3000 improvement





axion DM modifies
maxwell equations:

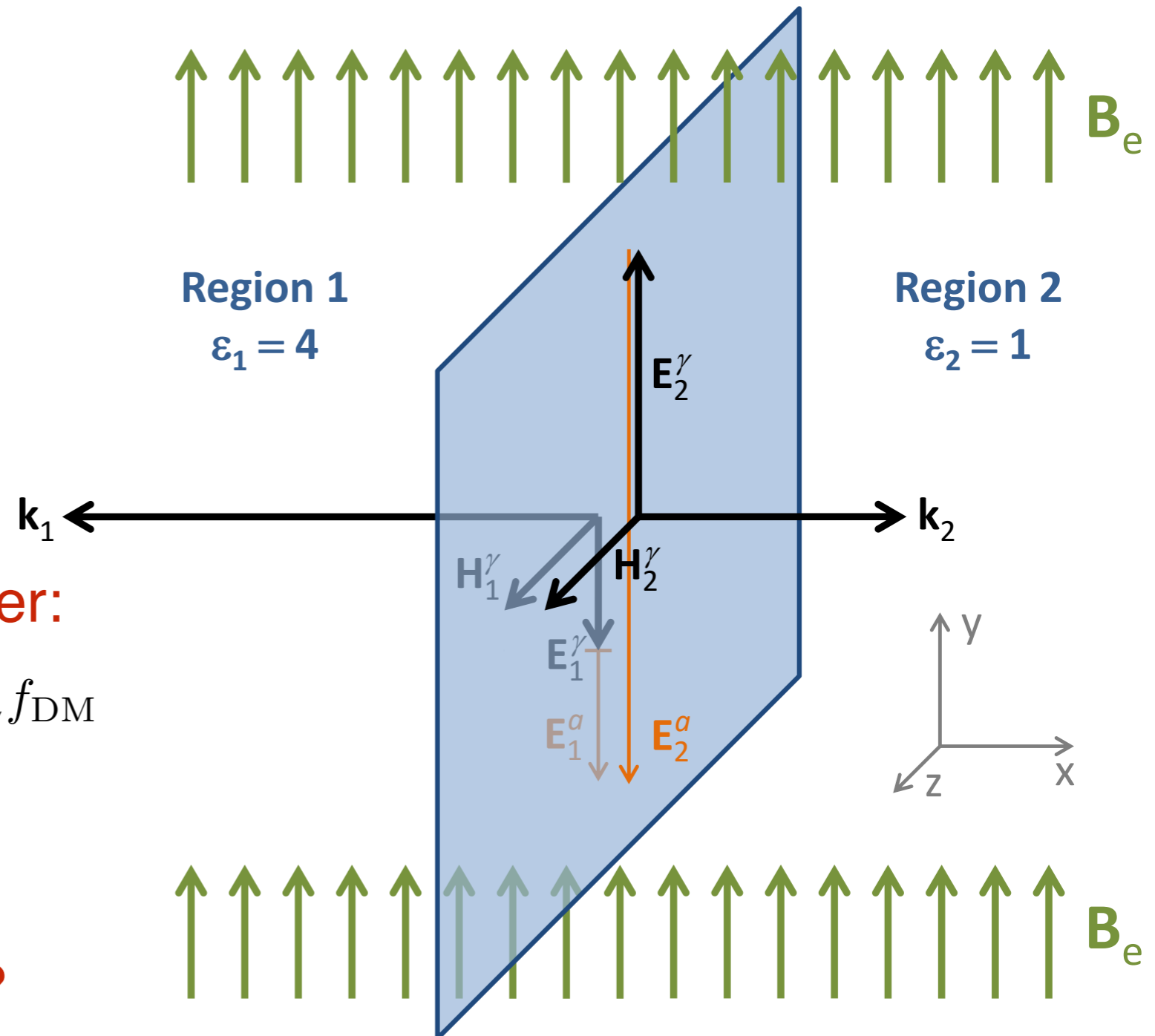
- EM radiation emitted at dielectric transition region

- power emission by one layer:

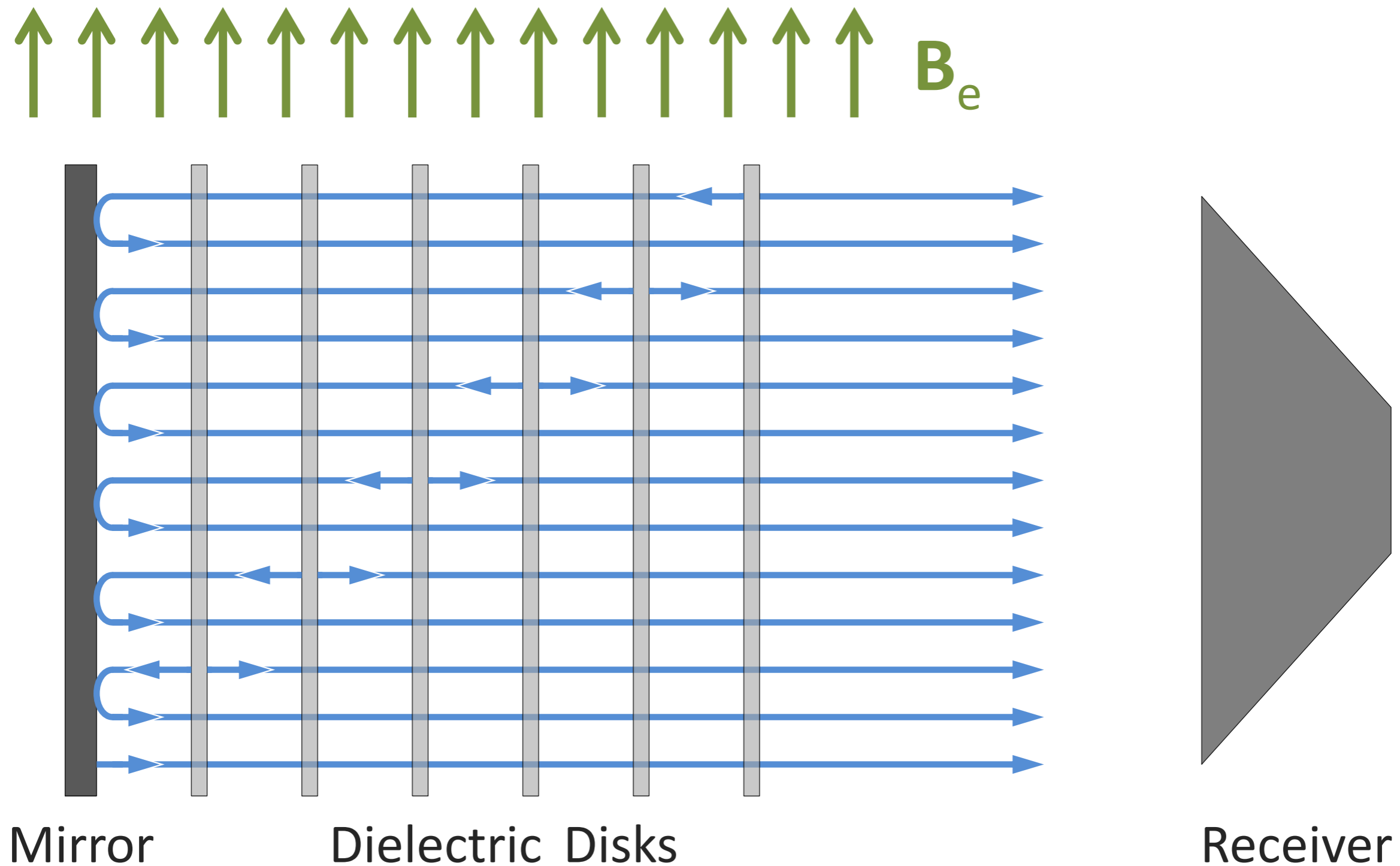
$$\frac{P}{A} = 2.2 \times 10^{-27} \frac{\text{W}}{\text{m}^2} \left(\frac{B_e}{10\text{T}} \right)^2 C_{a\gamma}^2 f_{\text{DM}}$$

- too small to detect?
- stronger field? larger area?

➡ resonator

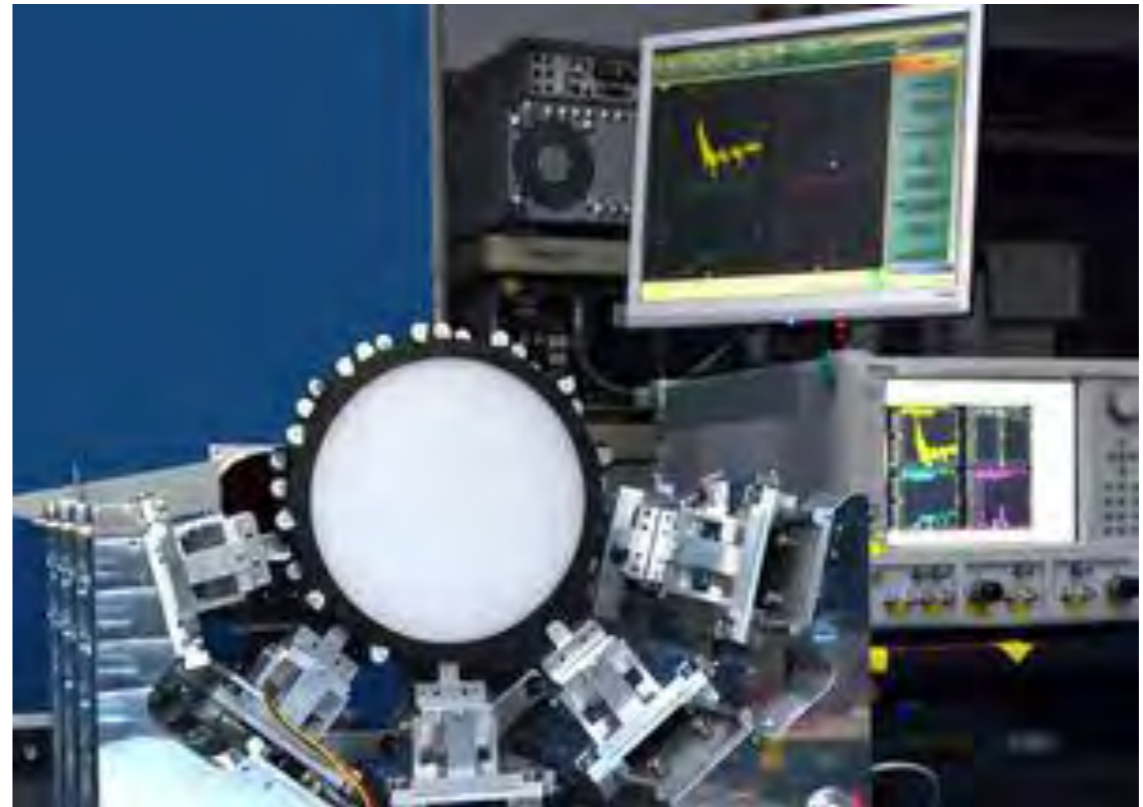
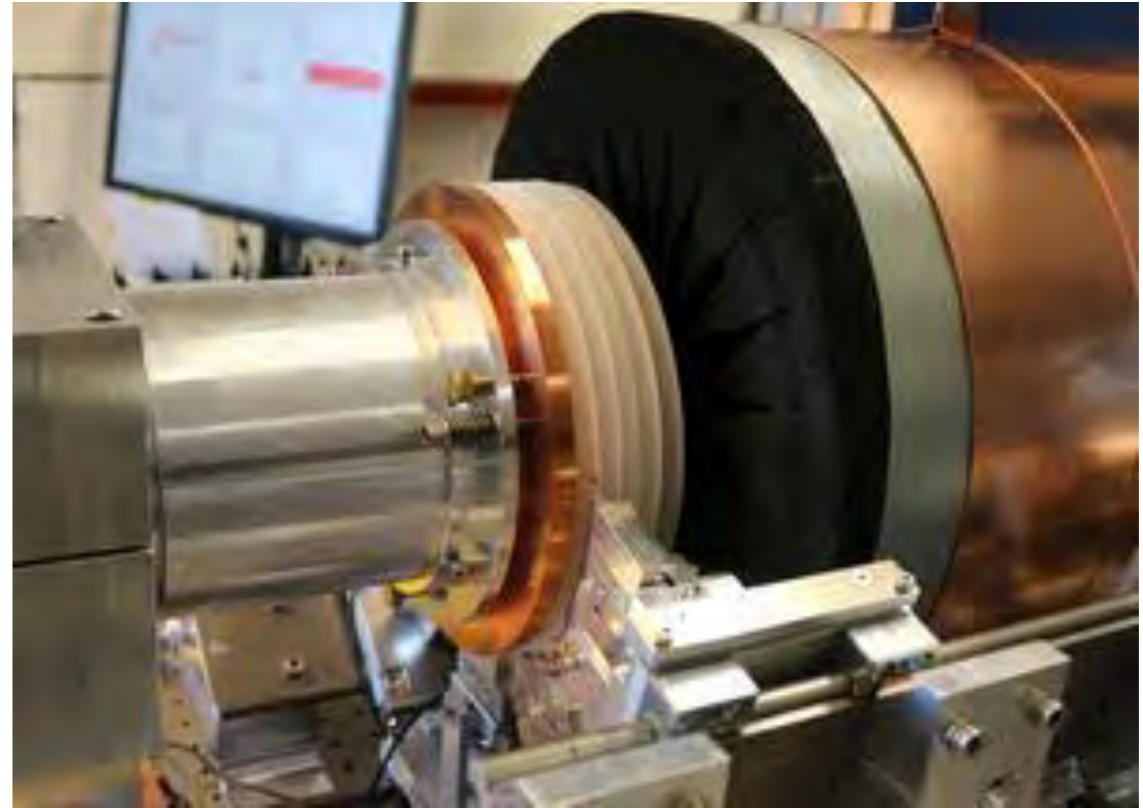


- resonator: multiple layers



MADMAX experiment

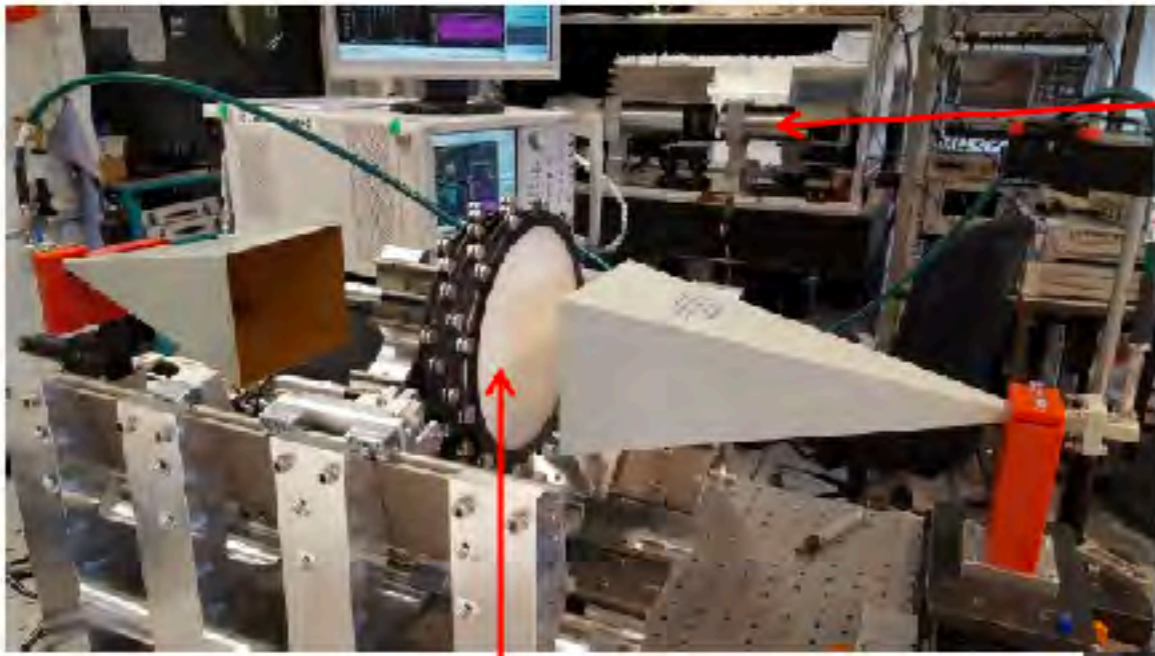
- MADMAX: magnetized dish and mirror axion experiment:
 - dark matter axion search in region 50 - 400 μeV
 - multiple dielectric layers in strong B field
 - overcome limitations from microwave cavities



test setup in Munich

O. Reimann / MPP

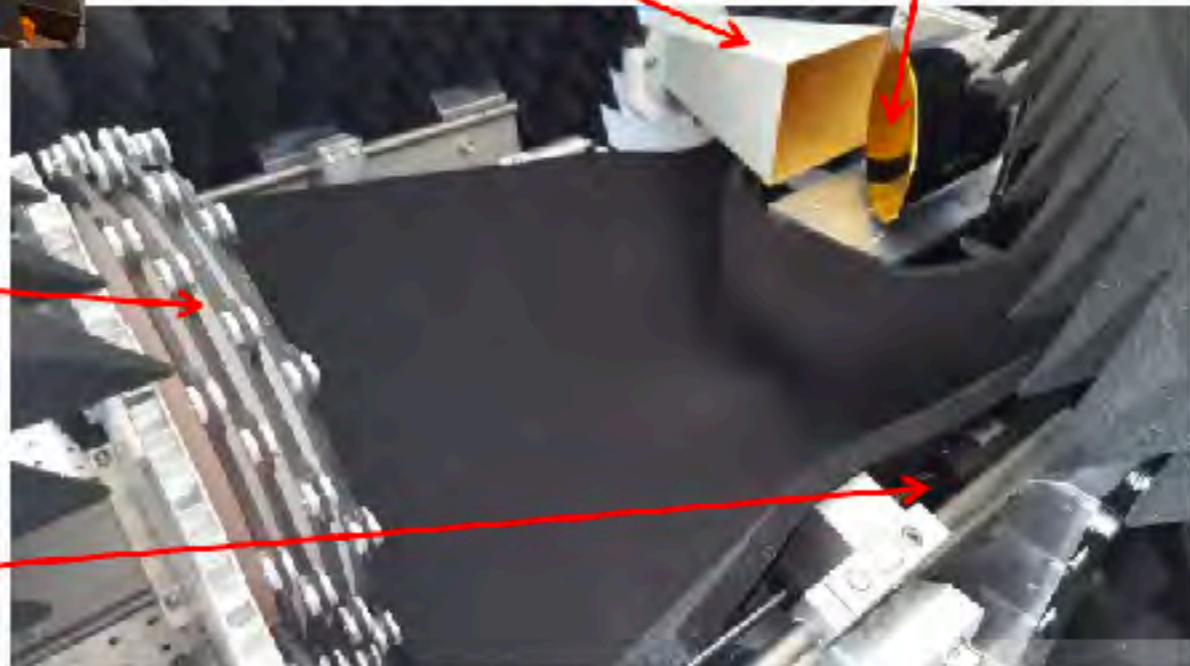
- The real device (200mm sapphire disks):



Waveguide system
(for background reduction)

Receiver horn

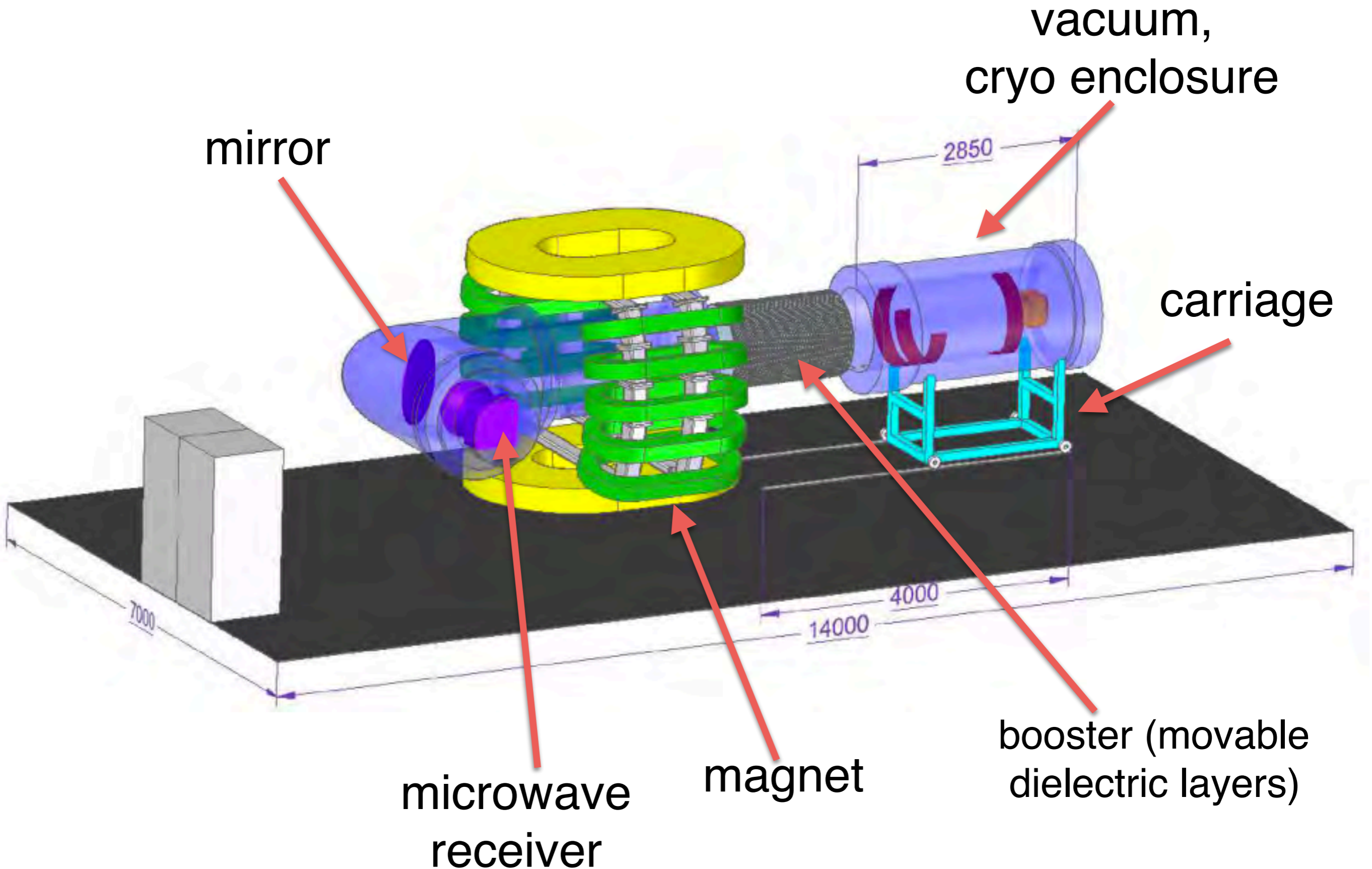
Parabolic mirror



Resonator (adjustable)
5(4) disks, sapphire

Drive motor
(100nm accuracy)

MADMAX prototype plan

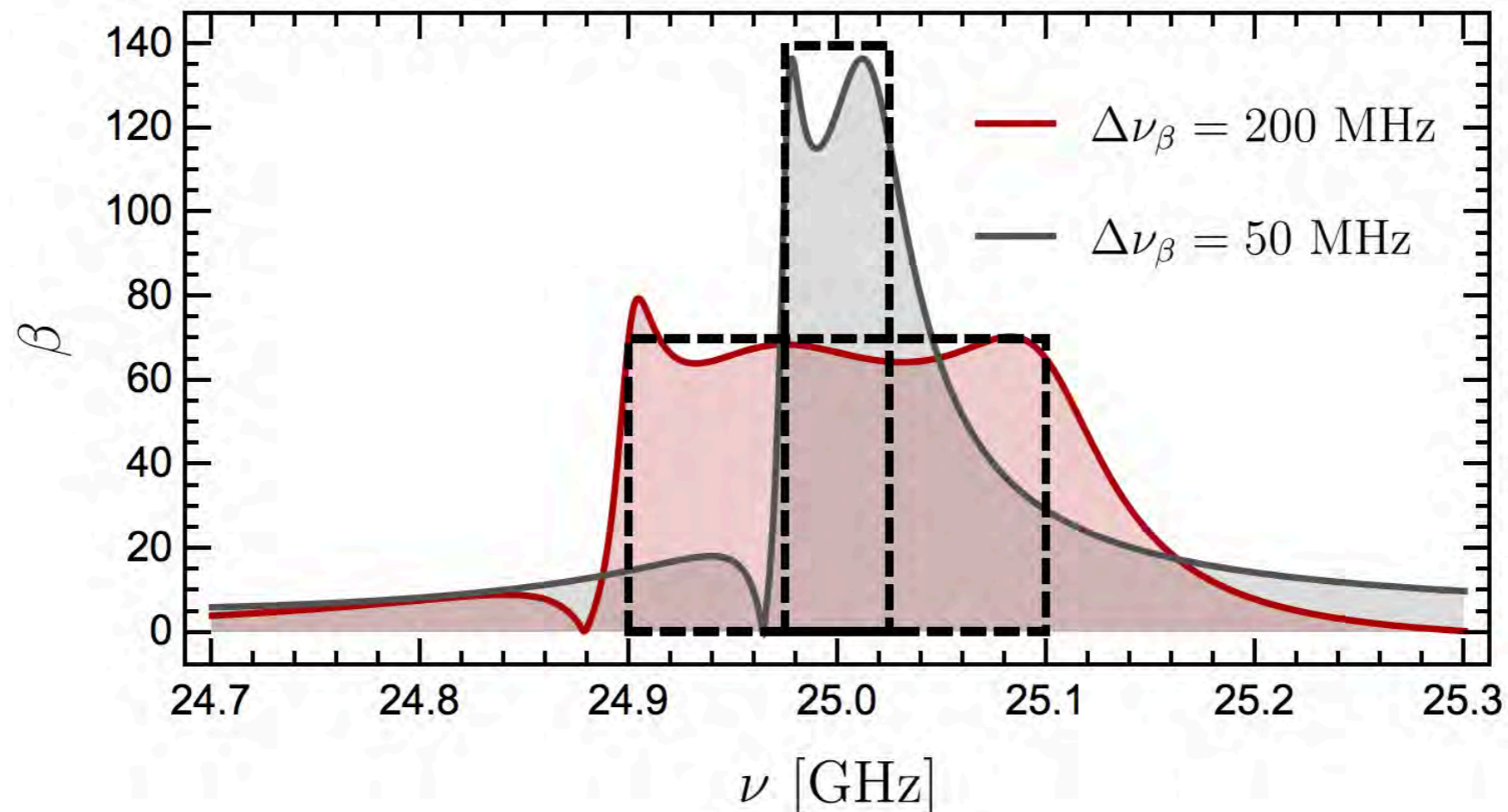


operating principle

- equidistant layers:
 - large boost, good S/N
 - narrow frequency range
 - frequent disk -repositioning required
- slight misalignment of layers:
 - smaller boost, worse S/N
 - broad frequency range
 - less repositioning

→ trade-off for optimal sensitivity

- all disks need individual high-precision adjustment

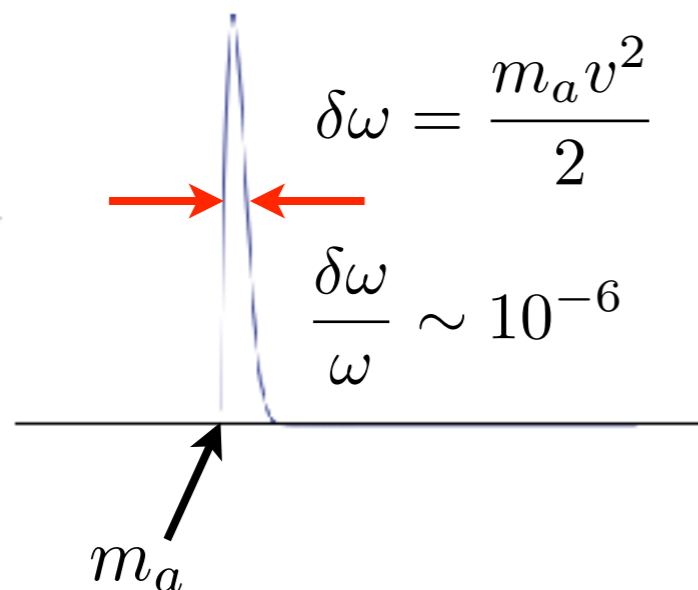


MADMAX experiment

- challenges:

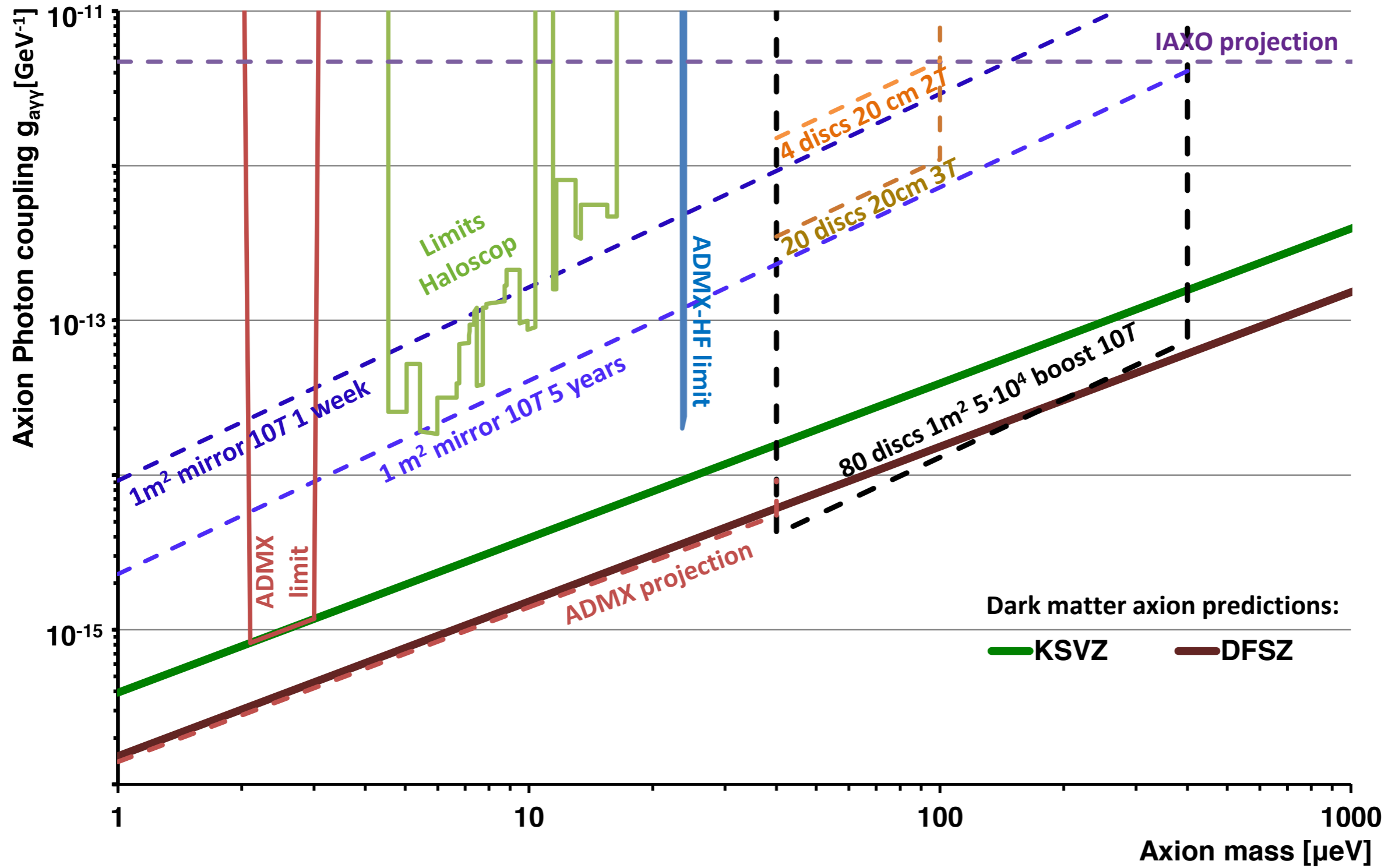
- huge and strong magnet (never built before)
- large, thin dielectric media to be moved around with high precision (in vacuum, strong field)
- tiny signal, unknown frequency
- is DM here or out there?
- coherence:

$$\omega \simeq m_a(1 + v^2/2 + \dots)$$



coherence length

$$\delta L \sim \frac{1}{\delta p} \sim 20\text{m} \left(\frac{10^{-5}\text{eV}}{m_a} \right)$$



MADMAX white paper

- UHH & DESY have signed the MADMAX white paper:

**A new experimental approach to probe
QCD Axion Dark Matter in the mass
range above $40 \mu\text{eV}$**

The MADMAX interest group:

P. Brun^a A. Caldwell^b L. Chevalier^a G. Dvali^{b,c} E. Garutti^d
C. Gooch^b A. Hambarzumjan^b S. Knirck^b M. Kramer^c H. Krüger^f
T. Lasserre^a A. Lindner^f B. Majorovits^{b,1} C. Martens^f A. Millar^b
G. Raffelt^b J. Redondo^{g,2} O. Reimann^b A. Schmidt^d F. Simon^b
F. Steffen^b G. Wieching^c

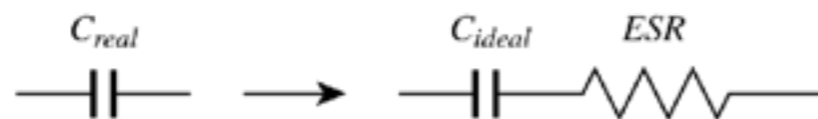
- madmax website:

<https://www.mpp.mpg.de/forschung/astroteilchenphysik-und-kosmologie/madmax-suche-nach-axionen-als-dunkler-materie/>

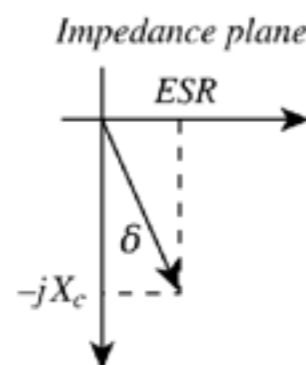
dielectric material

- **Problem:** find the ideal dielectric material to obtain
 - high boost factors
 - over a large surface
- ideal dielectric has:
 - High dielectric constant ($\epsilon > 10$) for large axion/photon conversion factor
 - Low loss ($\tan \delta < 10^{-5}$) in order to reduce photon loss

real dielectric = ideal capacitor + equivalent series resistance (ESR)

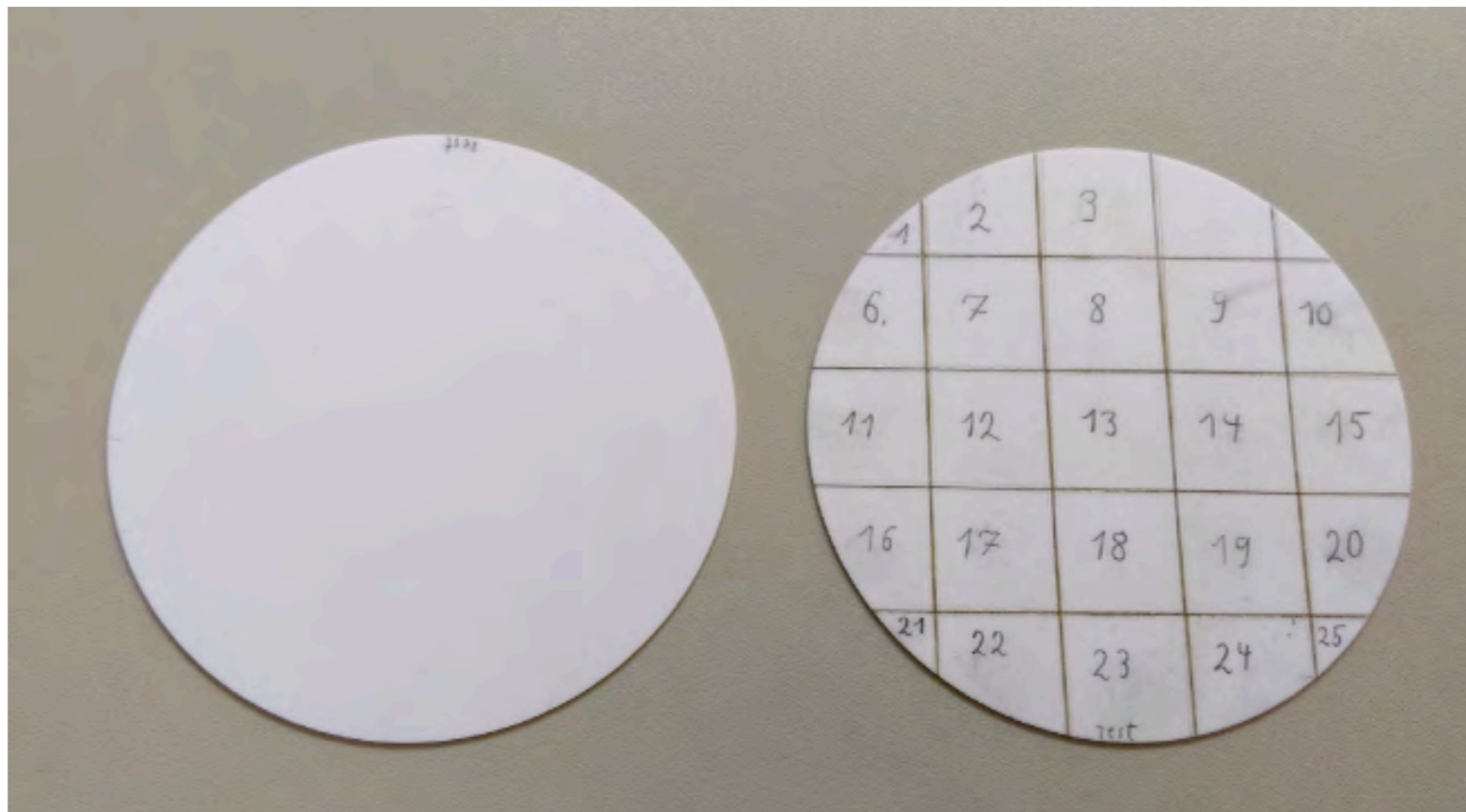


ESR should be minimum, i.e. $\tan \delta$ should be small



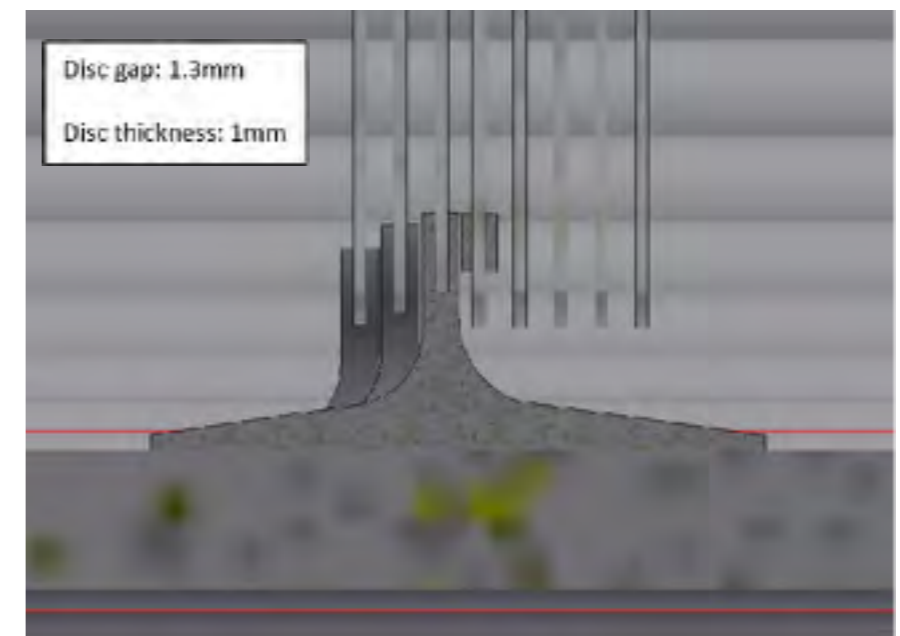
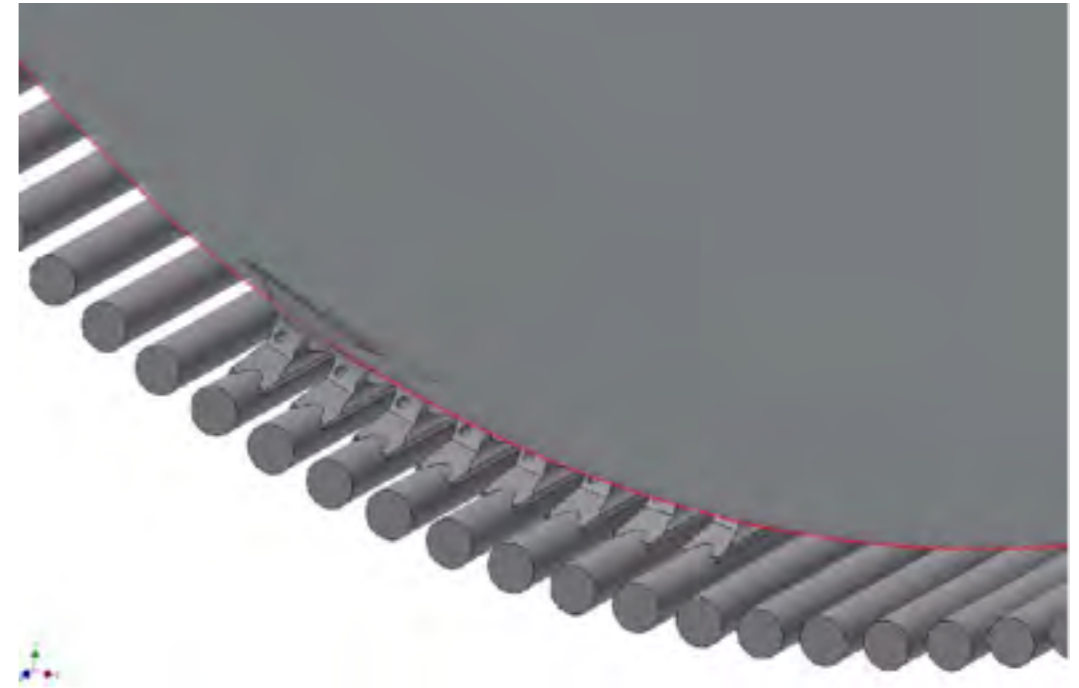
test of dielectric disk tiling

- 1 m² dielectric crystals cannot be grown (today)
- Solution: tiling
 - how to cut dielectric crystals ? (bridle)
 - how to glue ?
 - how to test dielectric properties after glueing ?



ideas for disk positioning system

- sliding rods to support & adjust position
- 3-points fixation / disc, 80 discs
- variable discs distance 1-20 mm
- rod movement ~ 1.6 m



goals of prototype in HH

- Test the scaling of the test system at MPP Munich to a full 80 discs booster system
- Test the mechanical alignment system
- Investigate behaviour of different dielectric material in a cryogenic environment
 - (and with high magnetic field)
- Check the agreement of simulations & measurements, including boost factor, transmissivity and reflectivity
- Study the required precision and stability of the mechanical alignment system and flatness of disc surface
- First test with a 4 T magnet
- First physics run with reduced sensitivity to obtain exclusion limits on Axion models

“There are viable theories and there are natural and elegant theories.

However, all viable, natural and elegant theories contain dark-matter axions” — Ann Nelson

- three types of experiments searching for axions:
 - helioscopes
 - haloscopes
 - light-shining-through-wall
- all experiments will experience dramatic improvements in the coming years
 - better magnets, better detectors, better ideas
- the region where axions solve both QCD and DM problems will **soon be covered by MADMAX**

SFB Lectures

14 Jul 2017

DESY Hamburg (Room 2, Building 2a), 14:30

Axel Lindner: Axion/ALPs in Experiments
[\[indico\]](#)

7 Jul 2017

DESY Hamburg (Room 2, Building 2a), 14:30

Andreas Ringwald: Axion/ALPs in Astrophysics and Cosmology
[\[indico\]](#)

30 Jun 2017

DESY Hamburg (Room 2, Building 2a), 14:30

Andreas Ringwald: Axion/ALPs in Particle Physics
[\[indico\]](#)

backup