My Session

T 106: Suche nach Dunkler Materie 4

Thursday, March 30, 2017, 16:45-18:55, VSH 19

Selection status for this session: not selected

Ð	16:45	T 106.1	Group Report: Why the Fermi GeV gamma-ray excess cannot be explained by dark matter annihilation — •WIM DE
Ð	17:05	T 106.2	Search for Right-Handed Neutrinos from Dark Matter Annihilation with Fermi-LAT and H.E.S.S. Telescopes - $\bullet Migu$
Ð	17:20	T 106.3	Search for Decaying Dark Matter with Astrophysical Muon Neutrinos Measured by IceCube - \bullet Jöran Stettner, The Schoenen, and Christopher Wiebusch for the IceCube collaboration
Ð	17:35	T 106.4	Group Report: Status des ALPS Experiments — • KLAUS ZENKER für die ALPS-II Kollaboration
Ð	17:55	T 106.5	Search for chameleons with an InGrid based X-ray detector at the CAST experiment — KLAUS DESCH, JOCHEN KAMIN SCHMIDT
Ð	18:10	T 106.6	Artificial Neural Networks as event classifiers for an InGrid detector at CAST — KLAUS DESCH, JOCHEN KAMINSKI, CHI
Ð	18:25	T 106.7	Suche nach Dunkler Materie in Ereignissen mit fehlender transversaler Energie und Jets beim ATLAS Experiment — BÜSCHER, KATHARINA JACOBI, MANUEL LORNATUS UND JAN SCHÄFFER
Ð	18:40	T 106.8	WIMP search at the International Linear Collider — • MORITZ HABERMEHL and JENNY LIST



WIMP Search at the International Linear Collider

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DPG-Frühjahrstagung





Münster 30 March 2017





The International Linear Collider

- a future electron positron collider
 - mature technology
 - waiting for political decision in Japan
- centre-of-mass energy: 250 500 GeV (upgrade: 1 TeV)
- $\mathcal{L} = 1.8 \cdot 10^{34} \text{cm}^{-2} \text{s}^{-1}$ (upgrade: $3.6 \cdot 10^{34} \text{cm}^{-2} \text{s}^{-1}$)
- polarised beams: $P(e^-) = \pm 80\%$, $P(e^+) = \pm 30\%$
- 2 detectors: SiD and ILD (International Large Detector)





How well can the ILC explore the WIMP paradigm ?

- Weakly Interacting Massive Particles (WIMPs) are candidates for dark matter
- WIMPs can be searched for
 - directly
 - indirectly
 - at colliders

 \Rightarrow idea: SM particles \rightarrow WIMP pair production

- singlet-like fermion WIMP [arxiv:1604.02230]
- likelihood analysis of
 - Planck, PICO-2L, LUX, XENON100
 - LEP, LHC
 - plus LZ, PICO250 projections
- surviving region
- couplings tested in [-1,1]
 - \Rightarrow "simplified model"

 $10^{2}_{10^{1}}$ $10^{2}_{10^{2}}$ m. (GeV)

diator mass (GeV)

< MAX[3m., 300 GeV]





Theoretical Framework: Effective Operators

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- 1. classify WIMP based on its quantum numbers (spin and weak isospin)
- 2. construct minimal effective Lagrangian \Rightarrow general approach
- example: vector-like fermion WIMP and vector-like operator \Rightarrow only one parameter "energy scale of new physics" $\Lambda = M_{mediator} / \sqrt{g_f g_{\chi}}$





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- example: vector-like fermion WIMP and vector-like operator \Rightarrow only one parameter "energy scale of new physics" $\Lambda = M_{mediator} / \sqrt{g_f g_{\chi}}$
- validity (above grey area)
 - $M_{mediator} \gg \sqrt{s}$ • g_f , $g_\chi \lesssim \sqrt{4\pi}$ (pertubativity)

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 $\Rightarrow \Lambda > 3 m_{\chi} \\ \Rightarrow \Lambda > 300 \text{ GeV}$





WIMP Detection at ILC

• Signal

 WIMP pair production with a photon from initial state radiation e⁺e⁻ → y200

 ${\rm e^+e^-} \rightarrow \chi \chi \gamma$

- quasi model-independent
- single photon in an "empty" detector
 - \rightarrow missing four-momentum
- observables: E_{γ} , θ_{γ}

• Main Background Processes

- Neutrino pairs $e^+e^- \rightarrow \nu \bar{\nu} \gamma$
 - irreducible
 - polarisation: enhance or suppress
- Bhabha scattering $e^+e^- \rightarrow e^+e^-\gamma$
 - huge cross section
 - mimics signal if leptons are undetected





Role of Polarisation

Vector operator, /s = 500 GeV, 500 fb⁻¹, v2016



N _{500fb⁻¹}	unpolarised	+80%, -30%
$ u u \gamma$	2479.19	483.51
$e^+e^-\gamma$	84.74	83.06

e+

 e^+

W

~~~^.

 $Z^0/\gamma$ 

 $\overline{\nu}_{e}$ 

 $e^+$ 

- background
  - neutrinos can be suppressed for right-handed electrons and left-handed positrons
- WIMPs
  - type of interaction can be tested



10

15

20 vears

2500 v 2000

1500

- extrapolation of sensitivity from full simulation
  - reachable  $\Lambda$  at different  $\sqrt{s}$  and integrated luminosities
  - for small M $_\chi$  (< 100 GeV)
  - allows to give estimates for sensitivity
    - for different time scales
    - for different running scenarios





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• for 
$$\sqrt{s} = 500 \,\text{GeV}$$
:

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- goal: realistic estimate of systematics and background



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Luminosity Spectrum: "our PDF"



lepton collider: requirement of high instantaneous luminosity  $\Rightarrow$  beams are highly collimated



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lepton collider: requirement of high instantaneous luminosity

- $\Rightarrow$  beams are highly collimated
- $\Rightarrow$  strong electro-magnetic fields
- $\Rightarrow$  both bunches are focussed in the field of the other beam



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- $\Rightarrow$  e<sup>+</sup>/e<sup>-</sup> emit synchrotron radiation ("beamstrahlung")
- $\Rightarrow$  ... and on average lose a few percent of their energy Moritz Habermehl  $\mid$  WIMP Search at the ILC  $\mid$  DPG  $\mid$  30 March 2017



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 updated analysis and exclusion limits are underway
 Electron Phase Space



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- previous Bhabha sample does **not** cover all phase space at low polar angles→ fixed
- updated analysis and exclusion limits are underway
   Electron Phase Space



# Summary: WIMP searches at the International Linear Collider









- WIMPs are still among favorite candidates for dark matter
  - ILC covers unique parameter space
  - complementary sensitivity to LHC searches
- goal: realistic estimate of sensitivity
  - taking systematics into account:
    - e.g. luminosity spectrum
  - detailed modelling of the particle phase space
  - $\Rightarrow$  input to
    - detector design (hermeticity in forward region)
    - machine design (luminosity spectrum)
- update of simulation study underway











#### References

- International Linear Collider
  - Technical Design Report: arXiv:1306.6327
  - Operating Scenarios: arXiv:1506.07830
- Simulation Study
  - EFT interpretation, vector-like mediator: PhD thesis of Andrii Chaus, Université Paris-sud 11, 2014PA112300
  - Cosmological interpretation: arXiv:1206.6639
- Likelihood analysis
  - arXiv:1604.02230
  - arXiv:1603.07387
  - arXiv:1407.1859



## Modelling of Signal and Background

- generated using WHIZARD 2.2.8
  - polarised beams
  - beam spectrum
  - photon modelling:
    - in matrix element " $\nu \bar{\nu} \gamma$ "  $\Rightarrow$  correct E,  $\theta$
    - as dedicated ISR parametrisation
       ISR implementation: all orders of soft-collinear photons, first three orders of hard-collinear photons ⇒ best cross-section

 $\Rightarrow$  no double counting

- signal:  $\chi \chi \gamma$ 
  - reweight  $\nu\bar{\nu}\gamma$  according to WIMP mass, spin, ...
- background:
  - $\bullet \ \nu \bar{\nu} + {\rm n} \gamma$
  - $e^+e^- + n\gamma$  (Bhabha scattering)
- full Geant4 based ILD simulation







## Signal Definition and Background Rejection

- signal definition (single photon plus missing energy)
  - $E_{\gamma} > 10 \text{ GeV}$
  - $E_{\gamma} < 220~{
    m GeV}$  (Z return at 242 GeV: avoid large background)
  - $|\cos heta_{\gamma}| < 0.98$  (tracking needed to distinguish  $\gamma$  from e^{-/+})
  - $\Rightarrow$  Bhabhas: hard photon boosts leptons in detector
- empty detector
  - veto events with track with  $p_T > 3 \text{ GeV}$
  - additional visible energy < 20 GeV
  - no  $e^+/e^-$  in forward region  $\Rightarrow$  no cluster in BeamCal
- $\Rightarrow$  retains 90% of signal
- $\Rightarrow$  Bhabha background rejection improved by factor 15



| $e^+e^-\gamma$                                               | C Bartels | new analysis |  |  |  |
|--------------------------------------------------------------|-----------|--------------|--|--|--|
|                                                              | 01.10/    |              |  |  |  |
| рт                                                           | 21.1%     | 26.1%        |  |  |  |
| E <sub>vis</sub>                                             | 16.0%     | 1.9%         |  |  |  |
| BeamCal                                                      | 0.29%     | 0.02%        |  |  |  |
| Moritz Habermehl   WIMP Search at the ILC   DPG   30 March 2 |           |              |  |  |  |



## Higher Sensitivity with Improved Bhabha Rejection



lower Bhabha background than in previous ILD analysis by A.Chaus

sensitivity is improved by 300 GeV  $\hat{=}$  by 15% for right-handed electrons and left-handed positron

for "vector" operator,  $\sqrt{s} = 500$  GeV, 500 fb<sup>-1</sup>

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## ILC Candidate Site: Kitakami





#### Results for an "Axial-Vector" Mediator





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## Measuring the WIMP Mass





- $E_{\gamma}$ : shape information is used
- range depends on  $M_\chi$  and  $\sqrt{s}$
- lepton collider  $\rightarrow$  initial state is known
- $\sqrt{s}$  known  $ightarrow M_{\chi}$
- uncertainty on  $\sqrt{s}$  dominates accuracy of  $M_{\chi}$





lepton collider: requirement of high instantaneous luminosity

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- $\Rightarrow$   $e^+/e^-$  emit synchrotron radiation ("beamstrahlung")
- $\Rightarrow$  ... and on average lose a few percent of their energy
- $\Rightarrow$  this energy distribution leads to systematic uncertainy on  $M_\chi$
- $\Rightarrow$  beamstrahlung photons generate  $e^+/e^-$  pairs



### Bhabha Scattering Background and the BeamCal

- if leptons are not detected
   → mimics mono-photon signal
- forward region of detector important: BeamCal (6 mrad  $< \theta <$  40 mrad  $\Leftrightarrow$  3.91  $< \eta <$  5.85)
- beamstrahlung photons generate  $e^+/e^-$  pairs  $\rightarrow$  energy deposition in detector  $\rightarrow \epsilon \ll 1$  at very low angles
- reconstruction of Bhabha leptons in BeamCal reduces background by factor 60 - 100

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