

Evolution of EUDET-type telescopes and future plans

Common beam telescopes, DAQ and reconstruction tool

Jan Dreyling-Eschweiler (DESY) for the DESY telescope and test beam crew BTTB6, Zurich, Switzerland, January 17th 2018



HELMHOLTZ RESEARCH FOR GRAND CHALLENGES



2 goals of this talk

State of the EUDET-type beam telescope package as an infrastructure

Trigger the community for a future joint effort

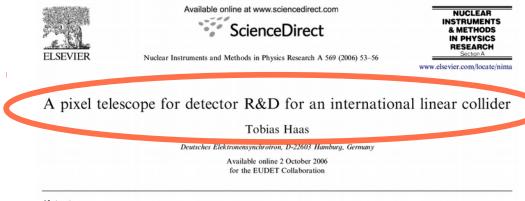
DISCLAIMER: This talk is mainly based on my experience since 2015 and my historical research.



Starting in 2006 with EUDET

Building up an infrastructure for EUropean DETector research & development





Abstract

A beam telescope based on monolithic active pixel technology will be developed within the EUDET collaboration, a coordinated detector R&D programme for a future international linear collider. The telescope will consist of up to six sensor planes. The telescope

Tobias Haas (2006), https://doi.org/10.1016/j.nima.2006.09.011

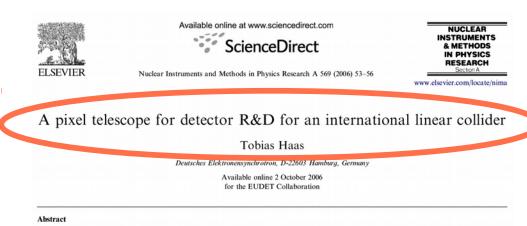
Starting in 2006 with EUDET

Building up an infrastructure for EUropean DETector research & development





- Tool of detector R&D for linear collider (DUTs: CCD, DEPFET, CMOS...)
- Source of 1-6 GeV electrons or other
- Operating in a magnetic field
- Using pixel sensors for 3+3 planes
- Infrastructure for trigger, integration and readout



A beam telescope based on monolithic active pixel technology will be developed within the EUDET collaboration, a coordinated detector R&D programme for a future international linear collider. The telescope will consist of up to six sensor planes. The telescope

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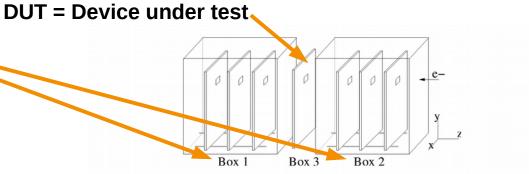
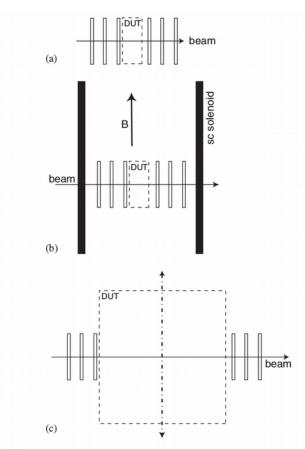


Figure 5: Mechanical concept with three separate areas for the two reference arms and the DUT.

Ingrid Gregor (2007), EUDET-Report-2007-05

Proposed & realized layouts...

... to meet the test needs



Standard setup (e.g. spatial resolution)

> DUT < 0.5m dz = 2-150mm

Magnet setup (e.g. Lorentz angle)

> DUT < 0.3m dz = 2-100mm

Two-arm geometry (e.g. e/γ-separation)

"DUT" > 0.5m dz = 2-150mm







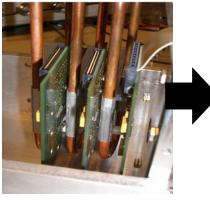


Fig. 2. Three foreseen layouts for the beam telescope; (a) a compact layout for characterization of high precision devices, (b) a compact layout inside a magnetic field and (c) a two-arm geometry for the use with larger DUTs.

T. Haas (2006), https://doi.org/10.1016/j.nima.2006.09.011 I. Gregor (2007), EUDET-Report-2007-05

Device Evolution

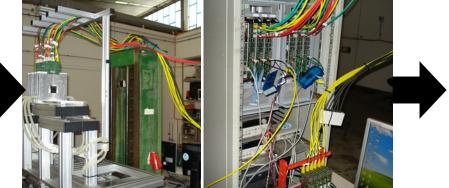
Prototyping – commissioning – development and optimization



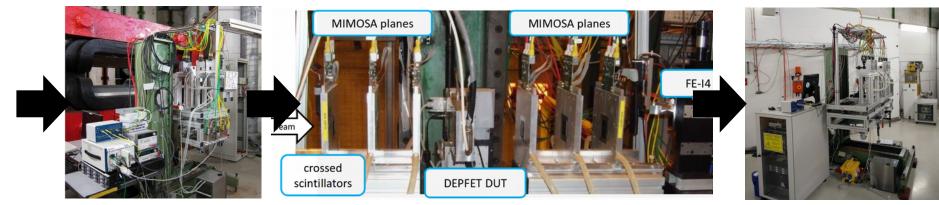


2007 3-plane demonstrator

2009 Mimosa26-prototype



2011 mechanics and EUDRB DAQ



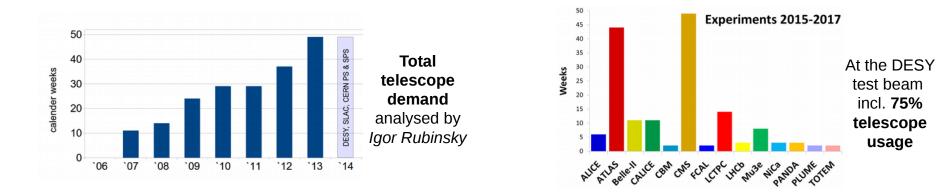
2012 final mechanics and NI DAQ

2012 user-driven extensions

2016 "pre-commercial" optimization

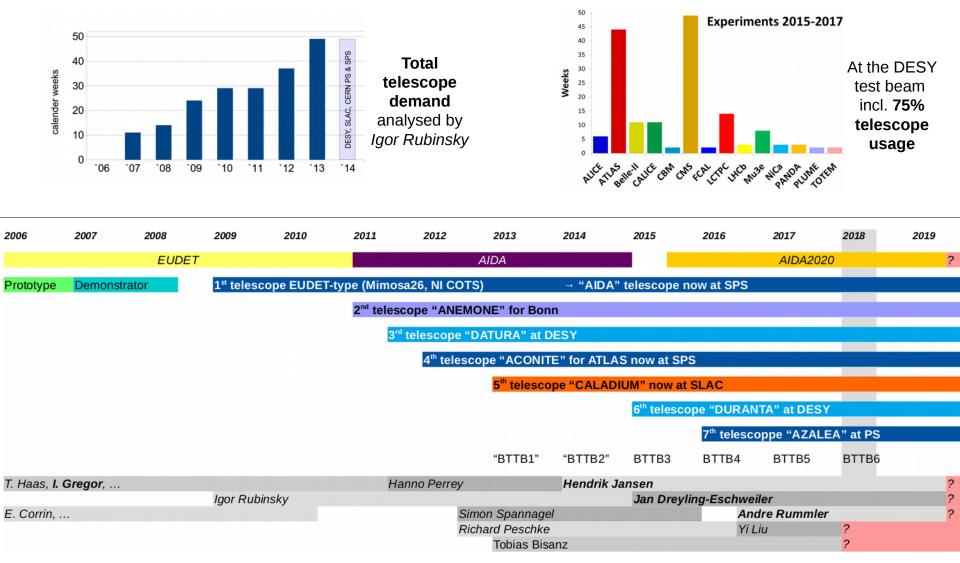
Usage & infrastructure evolution

User demand, devices and people



Usage & infrastructure evolution

User demand, devices and people



EUDET-type telescopes family today

Seven copies around the world at 5 different beam test beam facilities

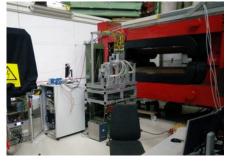




Contact: Carsten Hast



CALADIUM @ SLAC in Stanford, USA



DATURA @ TB21



DESY

DURANTA @ TB22 TB contact: Ralf Diener, Norbert Meyners, Marcel Stanitzki Telescope contact: Hendrik Jansen, Jan Dreyling-Eschweiler

Talk by Mengging

(Thursday)

Talk by Meng-

ging/Uwe

Talk by Ralf (yesterday)







TB contact: Daniel Elsner Telescope contact: David-Leon Pohl



AZALEA @ PS, T10



SPS/PS contact:SHenric WilkensSTelescope contact:André Rummler



AIDA @ SPS, H6B

Talk by Maarten (this morning)



ACONITE @ SPS, H6A Talk by Alexander (this morning)



ANEMONE @ BONN / ELSA

DESY. | EUDET-type telescopes | Jan Dreyling-Eschweiler, 17th January 2018, BTTB6

Behind this story: Joint effort of experts

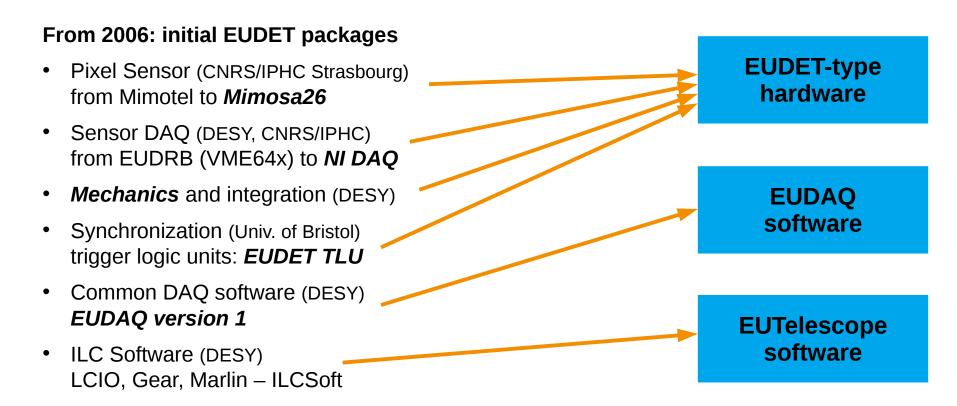
Sensor, trigger and DAQ R&D, integration and development

From 2006: initial EUDET packages

- Pixel Sensor (CNRS/IPHC Strasbourg) from Mimotel to *Mimosa26*
- Sensor DAQ (DESY, CNRS/IPHC) from EUDRB (VME64x) to NI DAQ
- *Mechanics* and integration (DESY)
- Synchronization (Univ. of Bristol) trigger logic units: **EUDET TLU**
- Common DAQ software (DESY)
 EUDAQ version 1
- ILC Software (DESY) LCIO, Gear, Marlin – ILCSoft

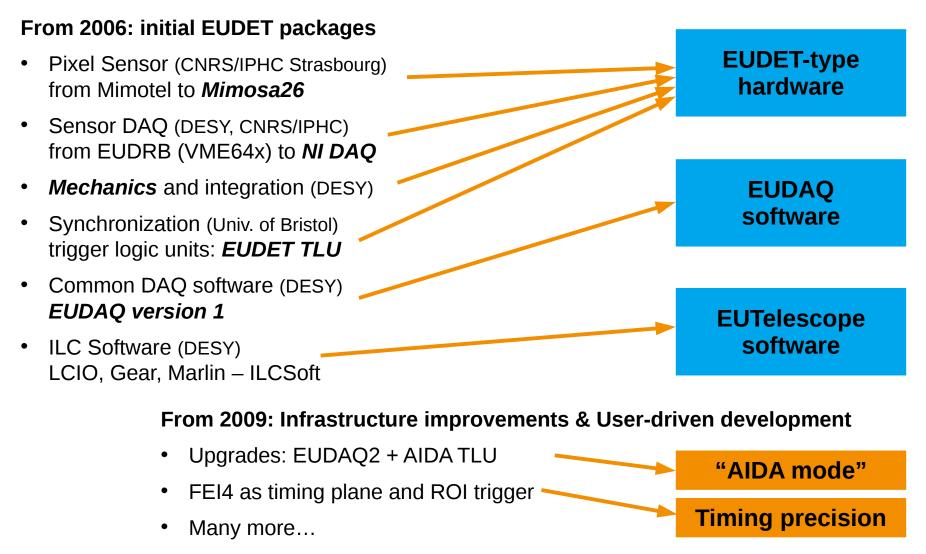
Behind this story: Joint effort of experts

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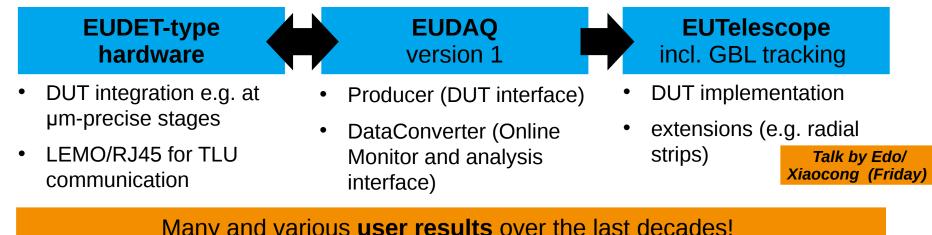
Sensor, trigger and DAQ R&D, integration and development



DESY. | EUDET-type telescopes | Jan Dreyling-Eschweiler, 17th January 2018, BTTB6

The common EUDET-type package

Common DUT integration from data taking over track reconstruction to results

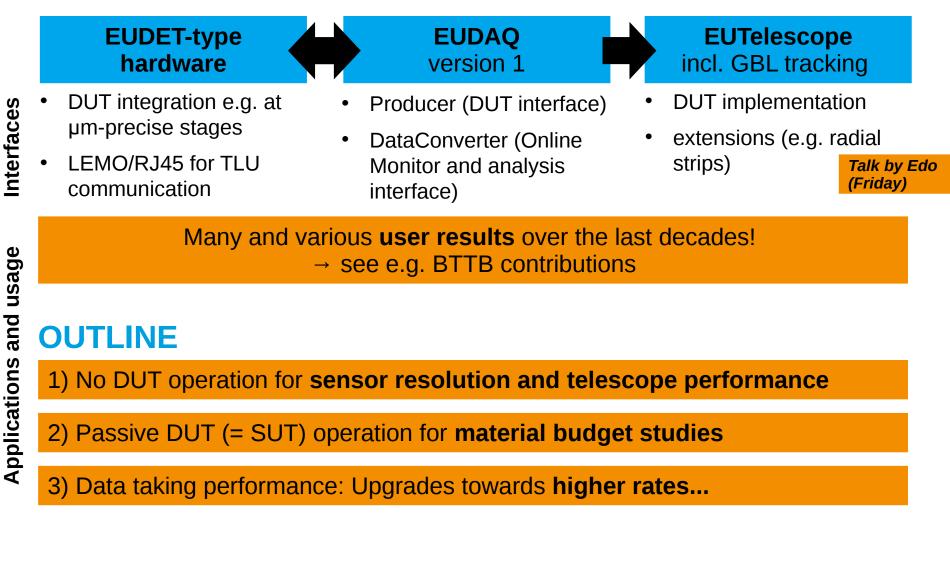


Many and various **user results** over the last decades! → see e.g. BTTB contributions

nterfaces

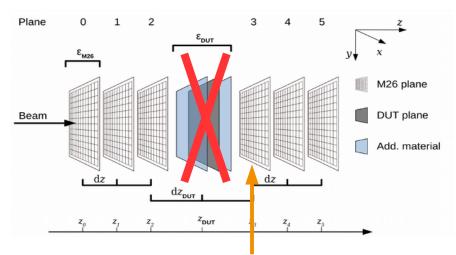
The common EUDET-type package

Common DUT integration from data taking over track reconstruction to results



1) Sensor resolution and response

Mimosa26 characterization

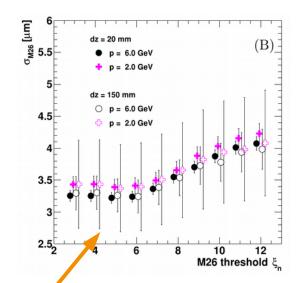


Mimosa26 sensor as DUT

- MAPS (binary AMS 350 nm CMOS)
- 1152x576 pixels, ~20 mm x 10 mm area
- pitch: **18.4 \mum** \rightarrow binary resolution: **5.3 \mum**

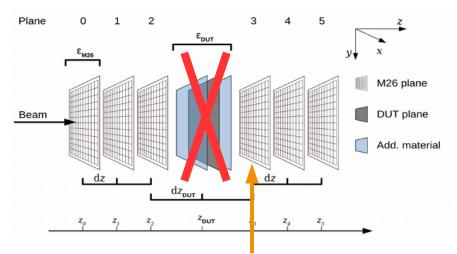
Results using GBL tracking

• avg. intrinsic res. $\sigma_{M26} = 3.24 \ \mu m @$ th. 6



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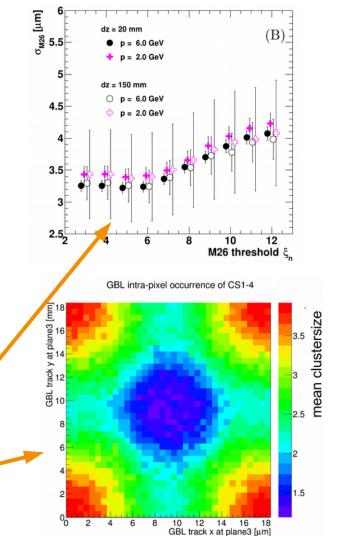


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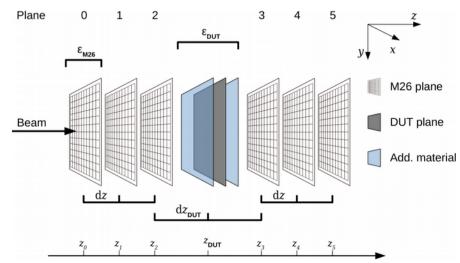
Results using GBL tracking

- avg. intrinsic res. σ_{M26} = **3.24 µm** @ th. 6
- charge sharing \rightarrow cluster \rightarrow intra-pixel distribution
- Allows track resolution predictions...



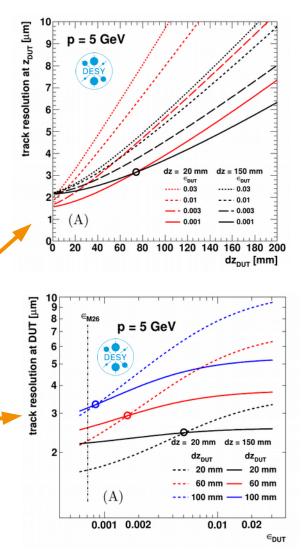
1) Resolution predictions for users' setup

Telescope characterization



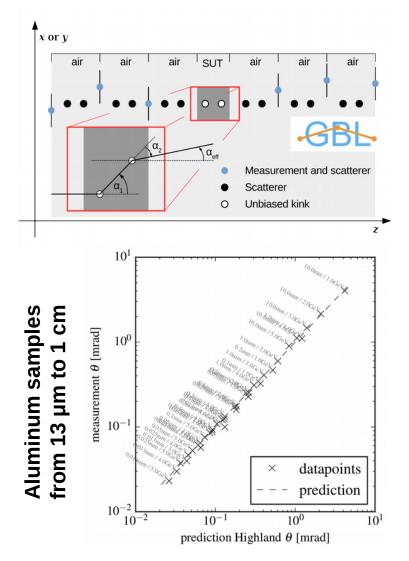
High track resolution through

- high energy and less material
- small telescope-DUT distance dz_{DUT}
- "thin" DUT: narrow setup (dz ~ 20mm)
 "thick" DUT: wide setup (dz ~ 150mm)
- → Calculate your setup: goo.gl/kR9VGS
- depends on multiple scattering ...



2) Material Budget studies

DUT → Sample under test (SUT)

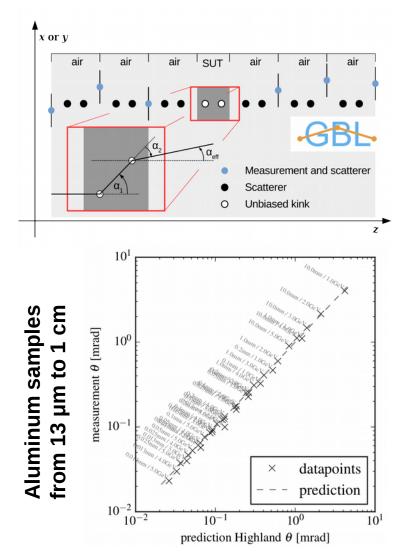


Getting scattering angle Θ

- Track model for GBL, e.g. two scatterers
- Kink angle (α) distribution
- Θ width of distribution

2) Material Budget studies

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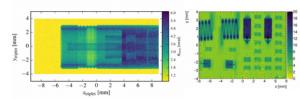


Getting scattering angle Θ

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Applications

- X0 measurements
- Imaging (2D or 3D)



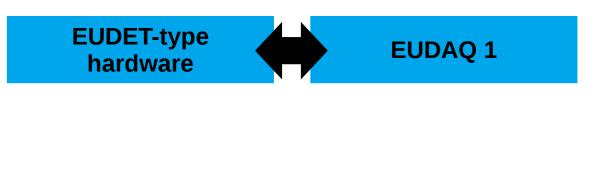
Talk by Jan-Hendrik Tutorial by Michaela Talk by Hendrik/Paul

Collaborative effort

- X0 scattering day on 30th Nov. 2017
- Exchange with (Benjamin Schwenker, Ulf Stolzenberg) Uni Göttingen

3) Data-taking performance

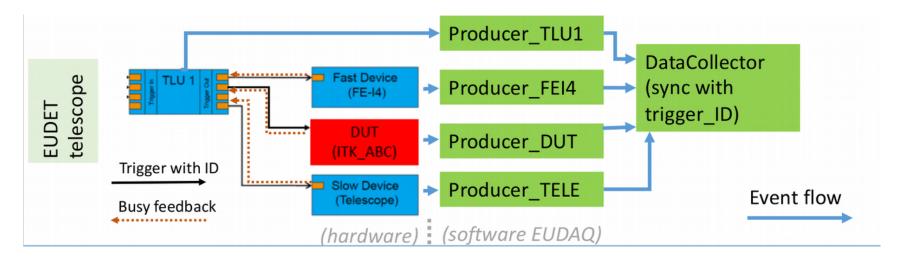




Strategy

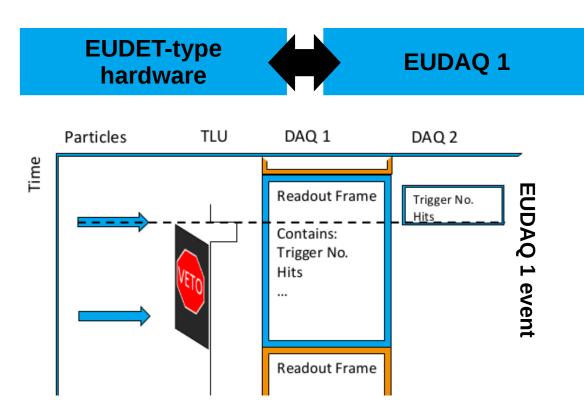
- \rightarrow trigger from TLU
- \rightarrow busy for read-out from DUTs
- \rightarrow one data collector

The data taking is limited by the slowest device.



Exemplary setup

3) Data-taking performance



MimosaDAQ:

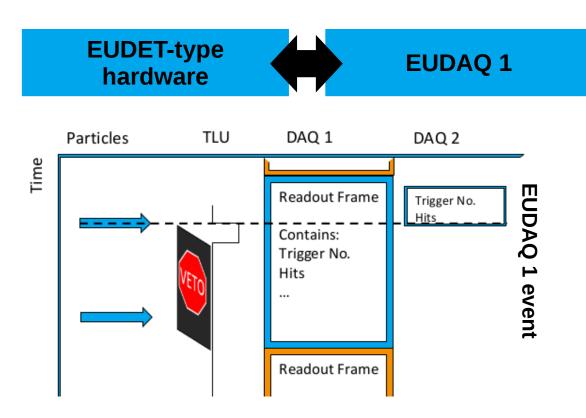
- Mimosa26 is continuously read-out line by line
- Two subsequent Mimosa frames are necessary
- \rightarrow Readout time: 2x115.2 µs (= 4.3 kHz)

Strategy

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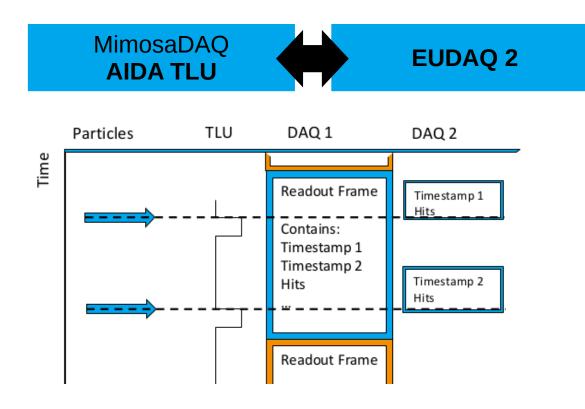
The data taking is limited by the slowest device.

EUDET-type telescope performance:

- MimosaDAQ
- EUDET TLU
- EUDAQ 1
- ~ 2 kHz trigger/event rate at beam conditions

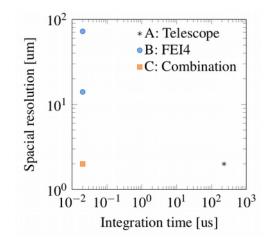
3) AIDA mode: Getting faster...

Towards a higher time resolution



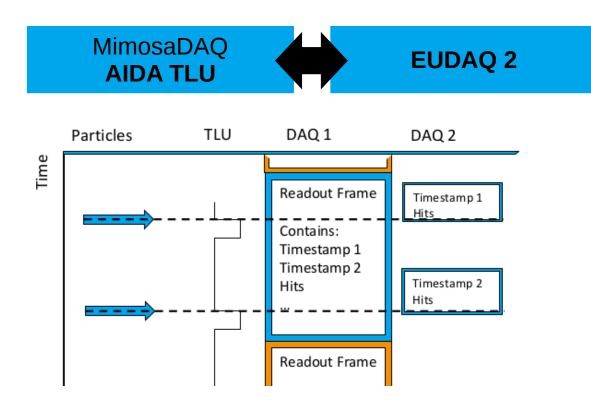
Strategy

- → *triggers* from TLU
- \rightarrow sync. by a common clock
- → opt. *multiple* data collectors



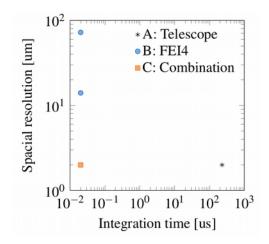
3) AIDA TLU and EUDAQ2

Towards a higher time resolution



Strategy

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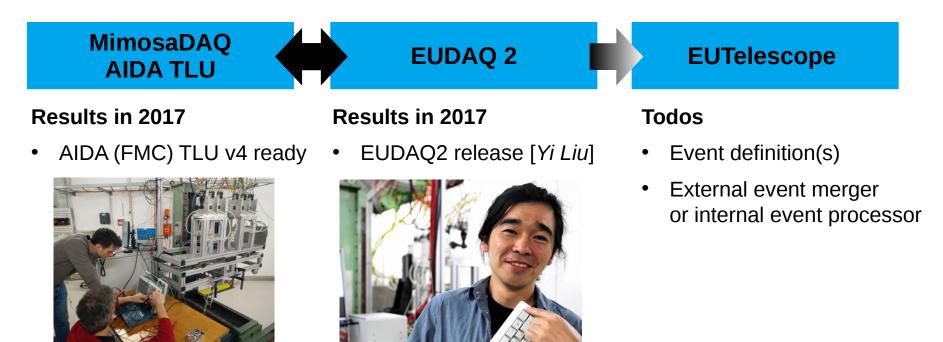
Recent problems & solutions during realisation:

- DUT-DAQ communications has to be updated → EUDET mode option
- Limited by slowest device → *individual busy* option
- → backward compatible for "EUDET-mode-DUTs" with new EUDAQ2 flexibility

Talk by David on the AIDA TLU

3) Results in 2017 and next steps...

... for a higher trigger rate and time-resolved track rate



Next steps

- Further tests for recommended user setup this February
- New Mimosa DAQ incl. FEI4 DAQ? → MMC3 board

Portal & References

telescopes.desy.de

Reference

• eudaq.github.io

Reference

Talk by Yannick

eutelescope.web.cern.ch

Lessons learnt

A personal collection

What was good?

- Common infrastructures has enabled detector R&D
- Mimosa26
 - Spatial resolution
 - Material budget \rightarrow thickness 50 µm
- EUDAQ as common DAQ
- Central coordination, user support and user-driven development
- Community development → BTTB
- Sustainability in Frameworks
 - WP15.2: infrastructure for tracking
 - WP5: common DAQ O AIDA
 - DESY HEP strategy



Lessons learnt

A personal collection

What was good?

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 - WP15.2: infrastructure for tracking
 - WP5: common DAQ
 - DESY HEP strategy



What can be improved?

- Sensor/Timing
 - New technologies (faster, ...)
 - Radiation hardness
- Flexible and robust mechanics
- Reconstruction SW
 - EUTelescope accessibility
 - Scope of reconstruction software
- Data and software interfaces
- Manpower on software maintenance and documentation

Summary and Outlook

Summary

- A common beam telescope infrastructure has been developed in the last decade
- The infrastructure provides
 - common setup and user interfaces
 - high-precision tracking
 - Modular and user-driven development
- The infrastructure is maintained and supported by dedicated experts

Summary and Outlook

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Future Plans

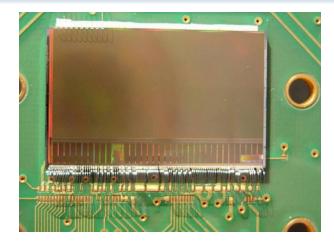
- User release for AIDA TLU EUDAQ2 EUTelescope package
- New MimosaDAQ?
 → 8.7 kHz frame rate
- Support in 19/20
- User needs/wishes, JRA Telescope 202x

Forum this evening



Mimosa 26 pixel sensors

- AMS 350 nm CMOS
- 18.4 um x 18.4 um
- 1152 x 576 pixels
- Roughly 10 mm x 20 mm

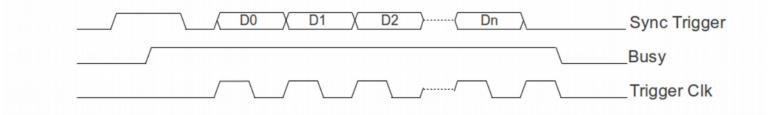


- Thickness: specs 50 um, measurement (55 ± 3) um
- HR epitaxial layer of 20 um thickness
- Binary resolution 5.3 um, improved by charge sharing
- Protected by 25 um Kapton on each side
- Material budget of sensor plus Kapton: $\varepsilon = 7.5e-4$

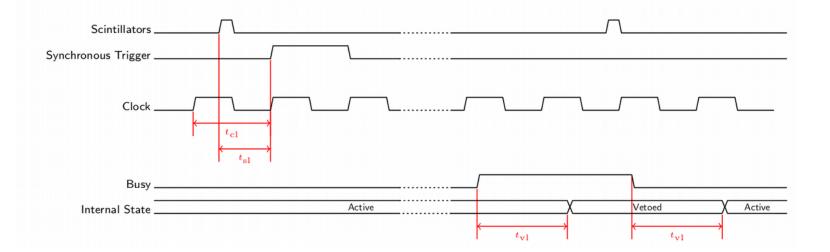


Trigger Logic Unit EUDET and AIDA TLU

EUDET Handshake (Trigger/busy with trigger number)



Synchronous mode (New for AIDA TLU)



Timing of the telescope

Rolling Shutter and combination with FEI4

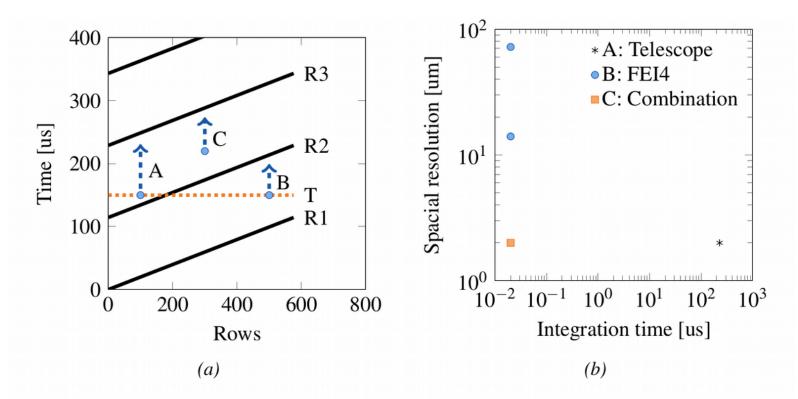


Figure 6.2 (a) Rolling shutter readout of the MIMOSA26 sensor. [R1,R2,R3] indicate the readout cycles. Hits are indicated with [A,B,C]. [T] trigger time.

(b) Pixel size versus the integration time of the specific readout system. [A] Performance of the telescope with six MIMOSA26 planes. Integration time $\approx 2 \times 114.5 \,\mu$ s. Spacial resolution $\approx 2 \,\mu$ m. [B] ATLAS FEI4 APIX sensor. Integration time 25 ns. Spacial resolution $\approx 14 \times 72 \,\mu$ m². [C] combination of both system into one setup.

Rates, rates, rates!

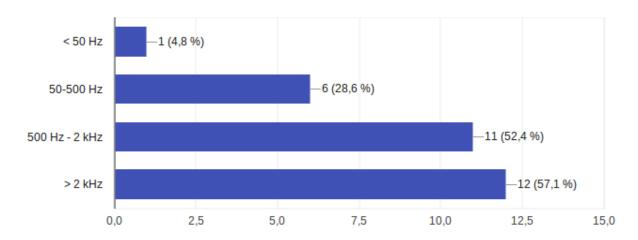
What rate are we taking about?

- Particle rate (source spectrum)
- Trigger rate (HW)
- Track rate (SW, after reco)



What rates are you interested in?

21 Antworten





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Rates and timing at a telescope

Two goals for a telescope device (incl. reconstruction software)

- **Rates** of reconstructed tracks:
 - For scattering measurements, X0 studies, ...
 - Depends on:
 - Source (particle flux/structure) and sensor (RO and geometry) define the track rate [#/frame] = particle flux [#/s/m^2] * time resolution [s/frame] * active area [m^2] * efficiency
 - **Reco software** (Track finding algorithm) defines the reco limit [#/frame] = maximal track rate [#/frame]
- Timing of reconstructed tracks:
 - For efficiency measurements, ...
 - Depends on:
 - Sensor: time resolution / information



Status and limit of the EUDET-tscope

one telescope frame

#	EUDA Q	TLU	Telescope	#Trigger	Max track rate	Timing referenc e	Max timed- track rate
1	1	EUDET (trigger ID, global busy)	Mimosa26 (115.2us, 2cm^2)	1	Reco limit	FEI4 (25ns, 4cm^2)	1
2	2	AIDA mixed- mode	Mimosa26 (115.2us, 2cm^2)	n	Reco limit	FEI4 (25ns, 4cm^2)	min(n, Reco limit)
3	2	AIDA (common clk, Asynch.)	Mimosa26 (115.2us, 2cm^2)	n	Reco limit	FEI4 (25ns, 4cm^2)	min(n, Reco limit)

Estimated limit (if hits are uniformly distributed)

Reco limit (e.g. ~30 tracks per frame) * frame rate (~ 8.7 kHz) = limit of track rate ~ 250 kHz



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