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Tau neutrino appearance with KM3NeT / ORCA

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Measuring tau neutrinos in ORCA

Strong experimental tests of the unitarity of the 3-flavour PMNS mixing matrix are very challenging, as direct observations of tau neutrino oscillations are difficult. A statistically highly-significant detection of tau neutrinos from $v_{\mu} \rightarrow v_{\tau}$ oscillations of atmospheric neutrinos could make an important contribution to further constrain the tau-related matrix elements. The discovery of a non-unitary 3v-mixing would be a clear sign for new physics, e.g. new types of neutrino interactions or the presence of sterile neutrinos. ORCA is a megaton-sized water Cherenkov detector for GeV-energy atmospheric neutrinos under construction in the Mediterranean Sea. About 3000 tau neutrinos per year, generated in flavour oscillations, will be detected in ORCA on a statistical basis. Tau neutrino interactions in ORCA result dominantly in events characterized by photon emission from particle cascades located close to the interaction vertex (shower-like events), while track-like muon events contribute <20%. Strong event-type identification is thus essential to quantify the tauinduced shower-like events.



ORCA: a water Cherenkov detector in the deep sea

KM3Ne1

- Significance map for tau neutrino appearance in the atmospheric neutrino flux after traversing Earth
- Most of the tau neutrino appearance signal is expected at 10 – 30 GeV reconstructed neutrino energy and with up-going direction



Random Decision Forest

- decision trees trained on random subsets of events and high-level reconstruction variables



Event selection

- Well-reconstructed events with reconstructed vertex close to or inside the instrumented volume
- Random Decision Forest classifier allows suppression of pure noise and atmospheric muon contamination to 3%-level with a neutrino efficiency of ~95%

Fitting procedure

- Using parametrised detector response obtained from simulated data
- Simultaneous fit of shower- and tracksample
- Asimov dataset approach with $\Delta \chi^2$ fit in the reconstructed energy and zenith-angle plane

Systematics included in fit

- Oscillation parameters: θ_{13} , θ_{23} , ΔM^2 , δ_{CP}
- Flux / cross section: v/anti-v ratio, spectral index
- Normalisation terms: v_u-CC, v_e-CC, NC
- Assumed true oscillation parameters: θ_{13} =8.42°, θ_{23} =45°, others taken from PDG 2014



Expected sensitivity

Normalisation = 1 indicates PMNS assuming unitary mixing



Non-appearance (=0) exclusion at 5σ -level possible within two months of operation with full ORCA detector

Fit result robust against θ_{23} and assumed mass ordering \rightarrow expect first physics result during construction phase

In case of significant deviation from unitary mixing: possibility to exclude unitarity assumption in the longer term

Very competitive sensitivity of ORCA: 20% constraint at 3σ on tau normalisation after one year

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