



# Baikal-GVD – the Next Generation Neutrino Telescope in Lake Baikal



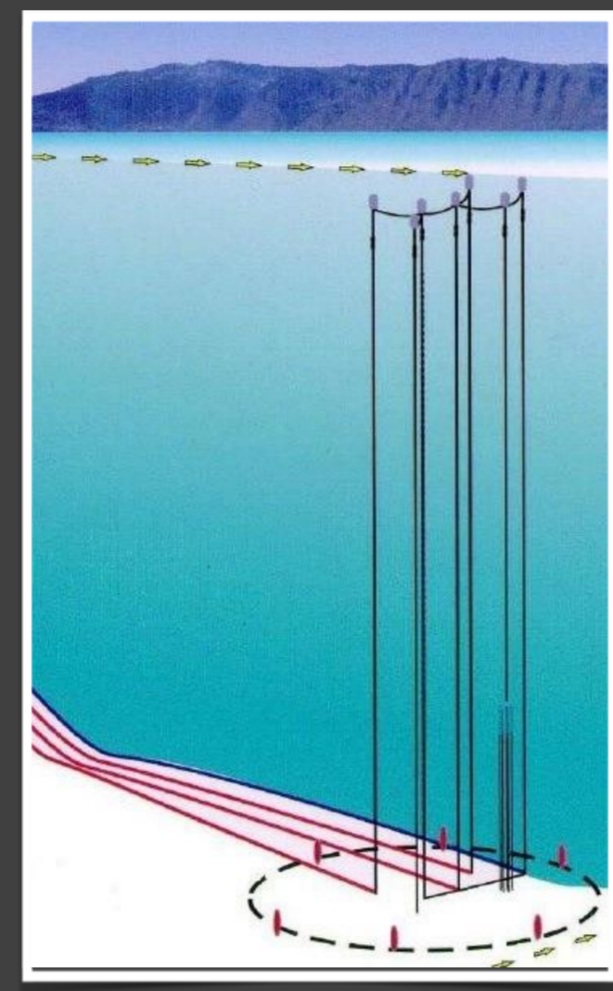
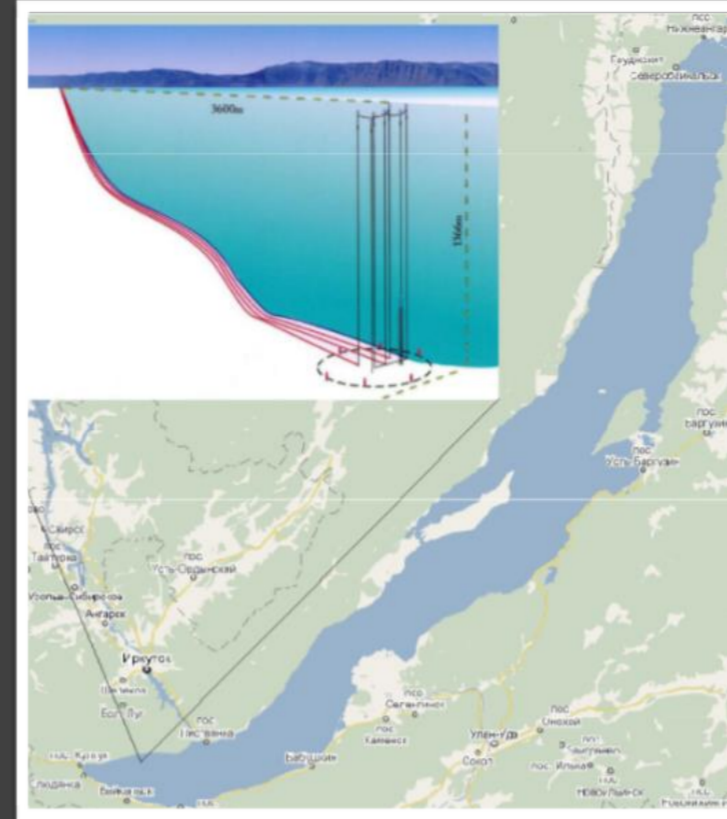
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on behalf of the Baikal Collaboration

## Outline

- Baikal-GVD (Gigaton Volume Detector) Phase I allows to investigate
  - Galactic and extragalactic neutrino point sources in TeV – PeV energy range
  - Diffuse neutrino flux – energy spectrum, local and global anisotropy, flavor content
  - Transient sources like GRB etc.
  - Dark matter – indirect search
  - Exotic particles – monopoles, Q-balls, nuclearities, ...
- The experiment looks like km<sup>3</sup> - scale 3D-array of photo sensors located deep underwater
- Flexible structure allows to rearrange of the main building blocks to change, for example, energy threshold
- High sensitivity and resolution of neutrino energy, direction and flavor content

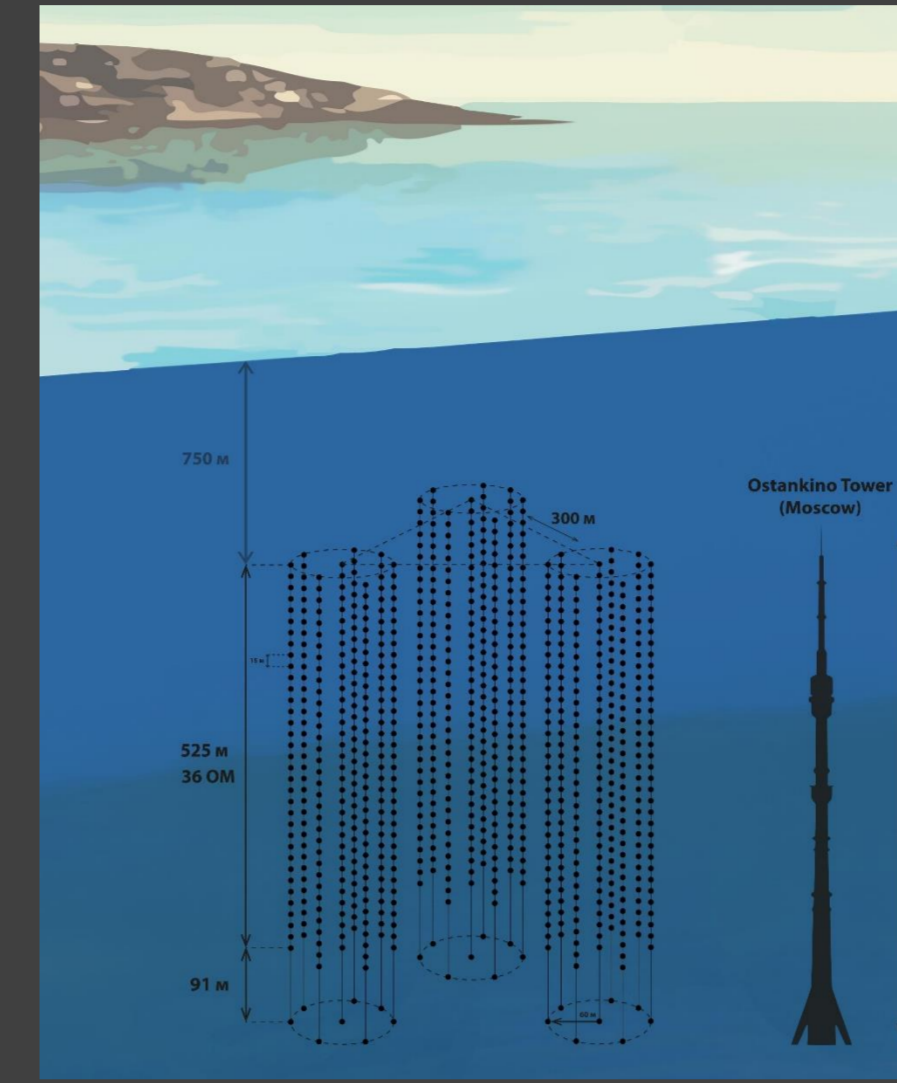
## Site:

- Location N51.765732° E104.415042°
- Depth of 1366 m at only 3.6 km from shore
- High deep water transparency (22 m) and low light scattering (30-50 m)
- Fresh water (simple mechanical solutions, no background from K40)
- No bioluminescence
- The most northern location allows observing the Galactic Center 18 hours per day
- Reliable ice cover (mid February - mid April) as an excellent platform for:
  - bed cable laying
  - telescope deployment and maintenance



## Baikal-GVD Phase I Array:

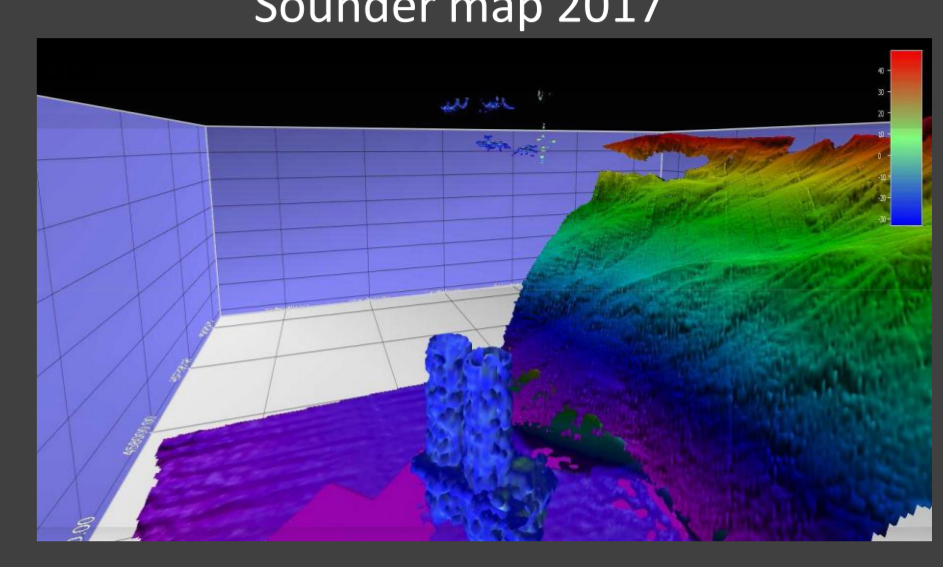
- 2304 Optical Modules (OM) in GVD Phase I
- OMs arranged in 8 Clusters with 8 Strings each
- Depth is 750 – 1275 m (91 m above the bed)
- 15 m between OMs in a String
- 300 m between Clusters



3 Clusters in 2018

## Optical Module:

- 10" PMT R7081HQE, Q<sub>eff</sub> ≈ 0.35
- 17" Glass pressure-resistant sphere
- Underwater 5-pin SubConn connector
- Amplifier, HV DC-DC, controller
- 2 on-board LED flashers for calibration
- Mu-metal cage
- Elastic gel



Sounder map 2017

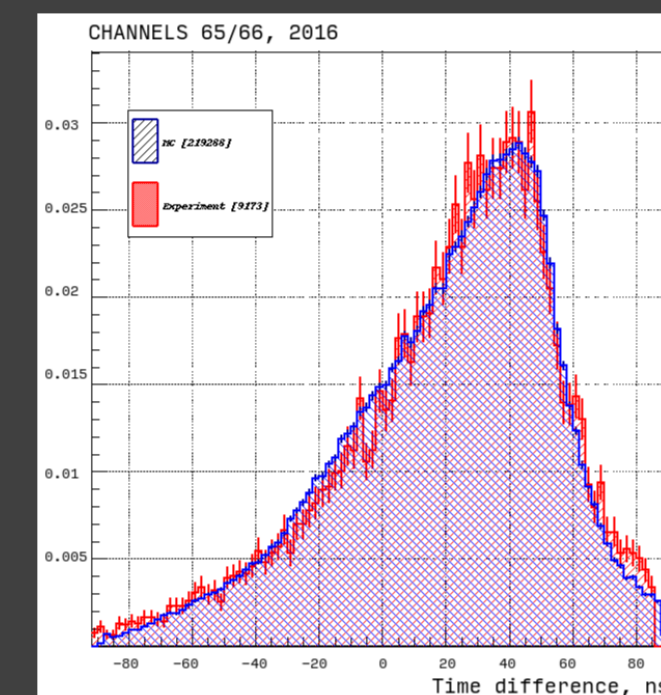
Year	2016	2017	2018	2019	2020	2021
# of Clusters planned	1	2	4	6	8	10
# of Clusters actual	1	2	3	5	7	9
# of Oms	288	576	864	1440	2016	2592

Timeline

## Muon ΔT distribution

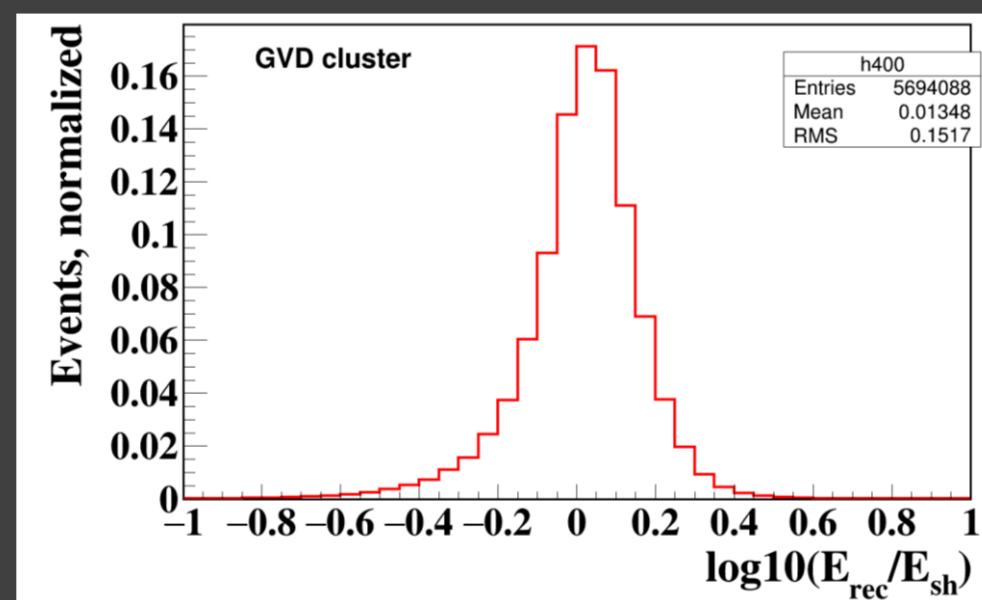
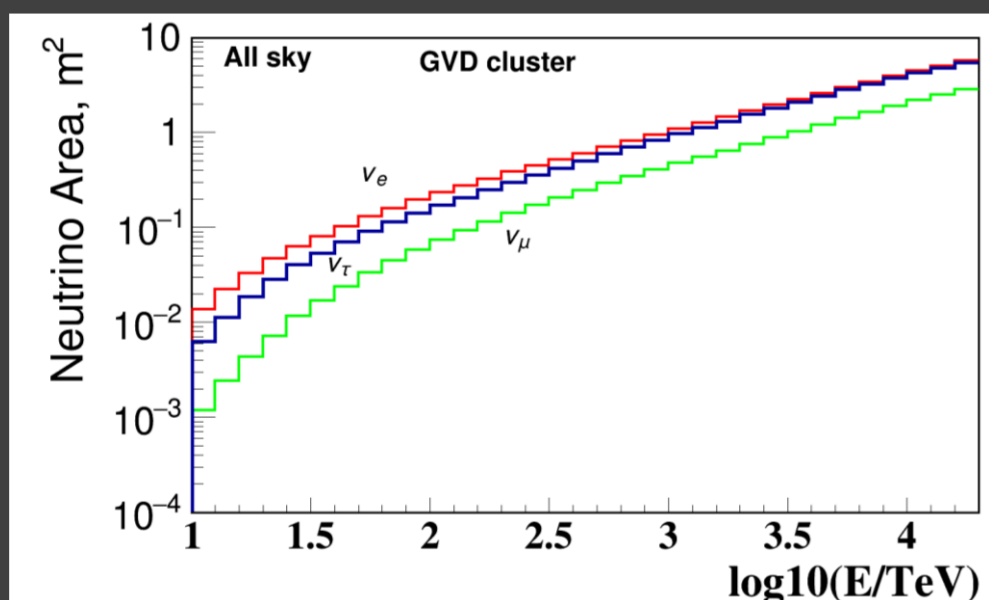
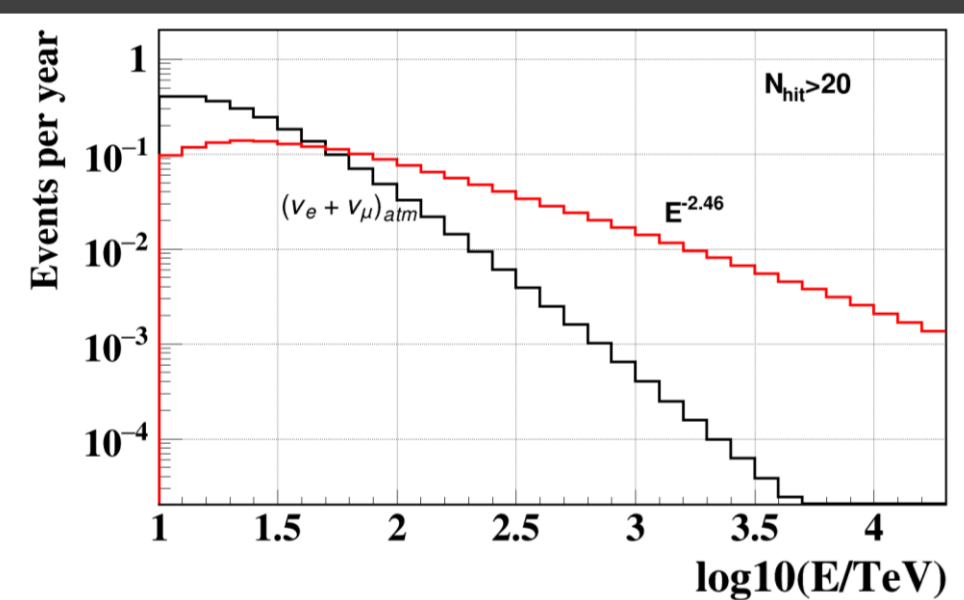
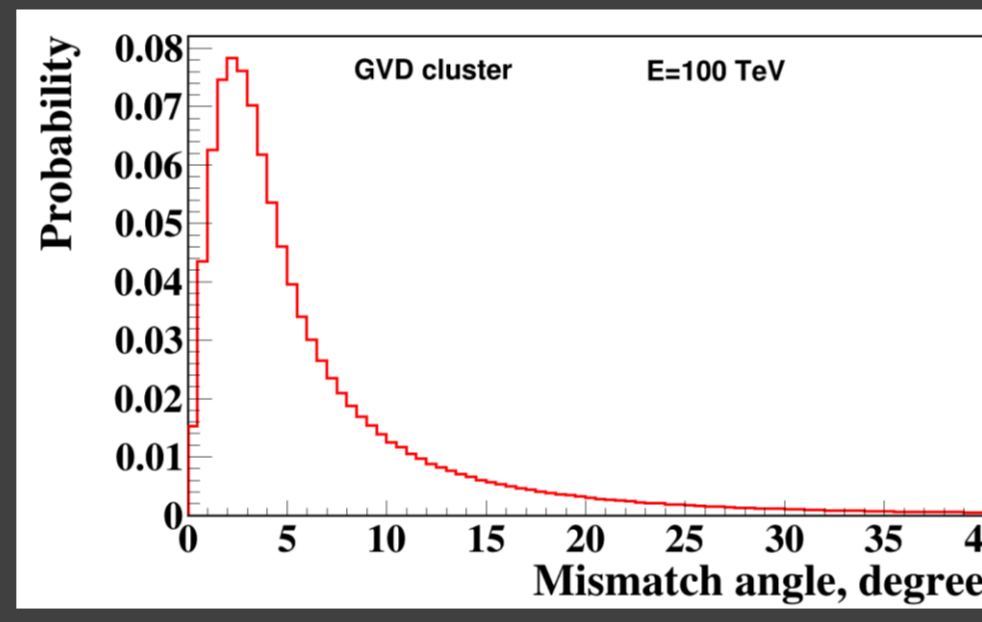
- hit time difference on adjacent OMs

- Muon hit selection:
  - Causality criterion, 20 ns window
  - Three or more OMs in a row must be hit
  - At least one OM of the three must have Q > 3 p.e.
- Serves as a cross check of time calibration with 5 ns precision, OM sensitivities

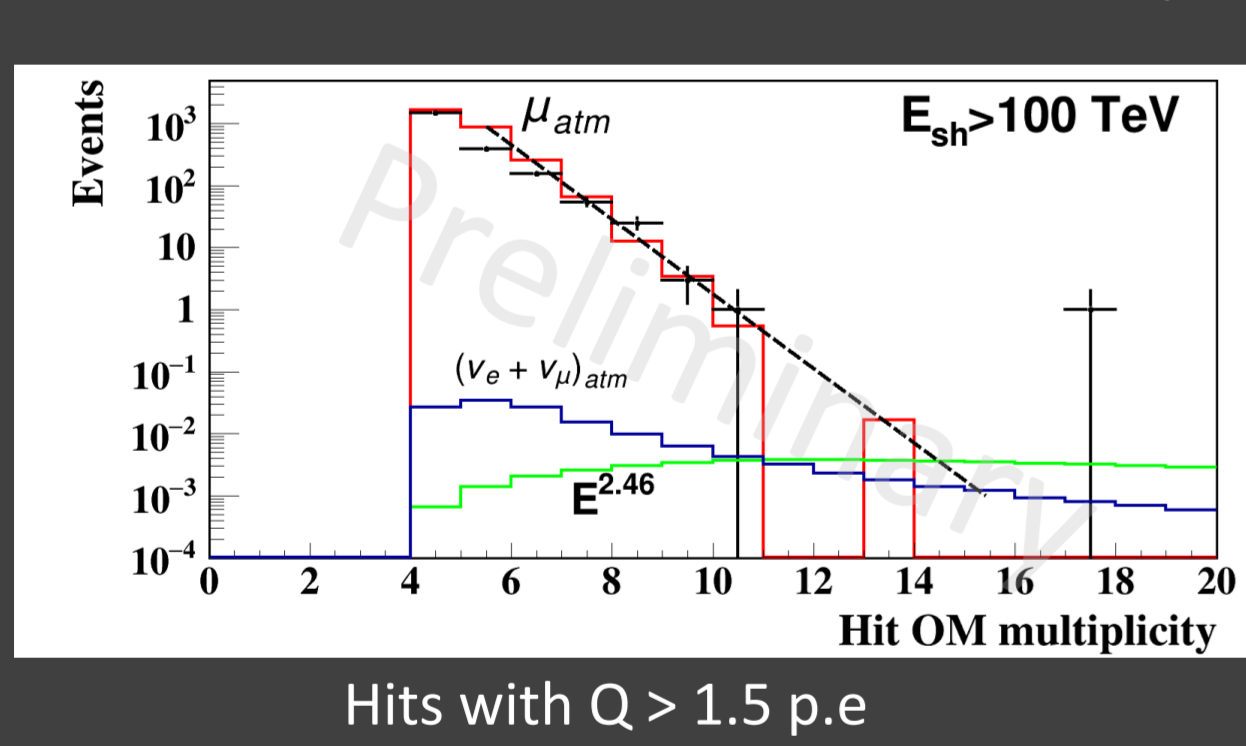


## Cascade detection in one cluster

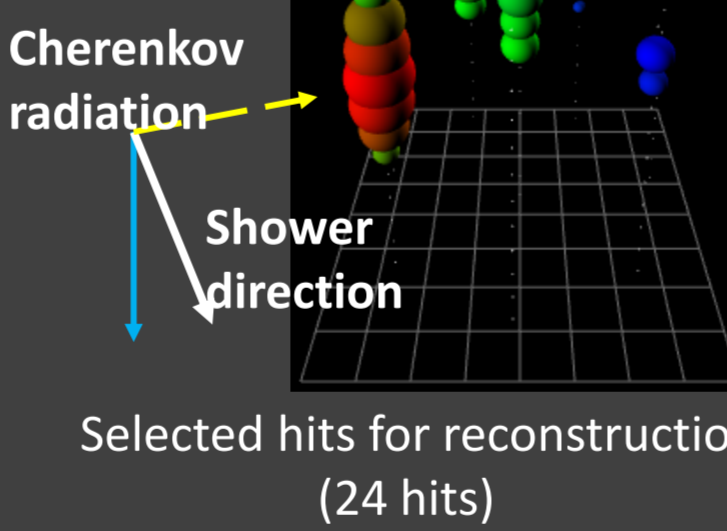
- Directional resolution ≈ 3° - 4° (median)
- Energy resolution δE/E ≈ 30% (averaged by E<sup>-2</sup> ν<sub>e</sub> spectrum)
- ≈ 0.6 events/yr (E<sub>sh</sub> > 100 TeV, N<sub>hit</sub> > 20)
- S<sub>cl</sub> ≈ (0.05 - 0.1) S<sub>IC</sub> - HESE eff. area



## Search for cascades in 2015 (41.6 live days)



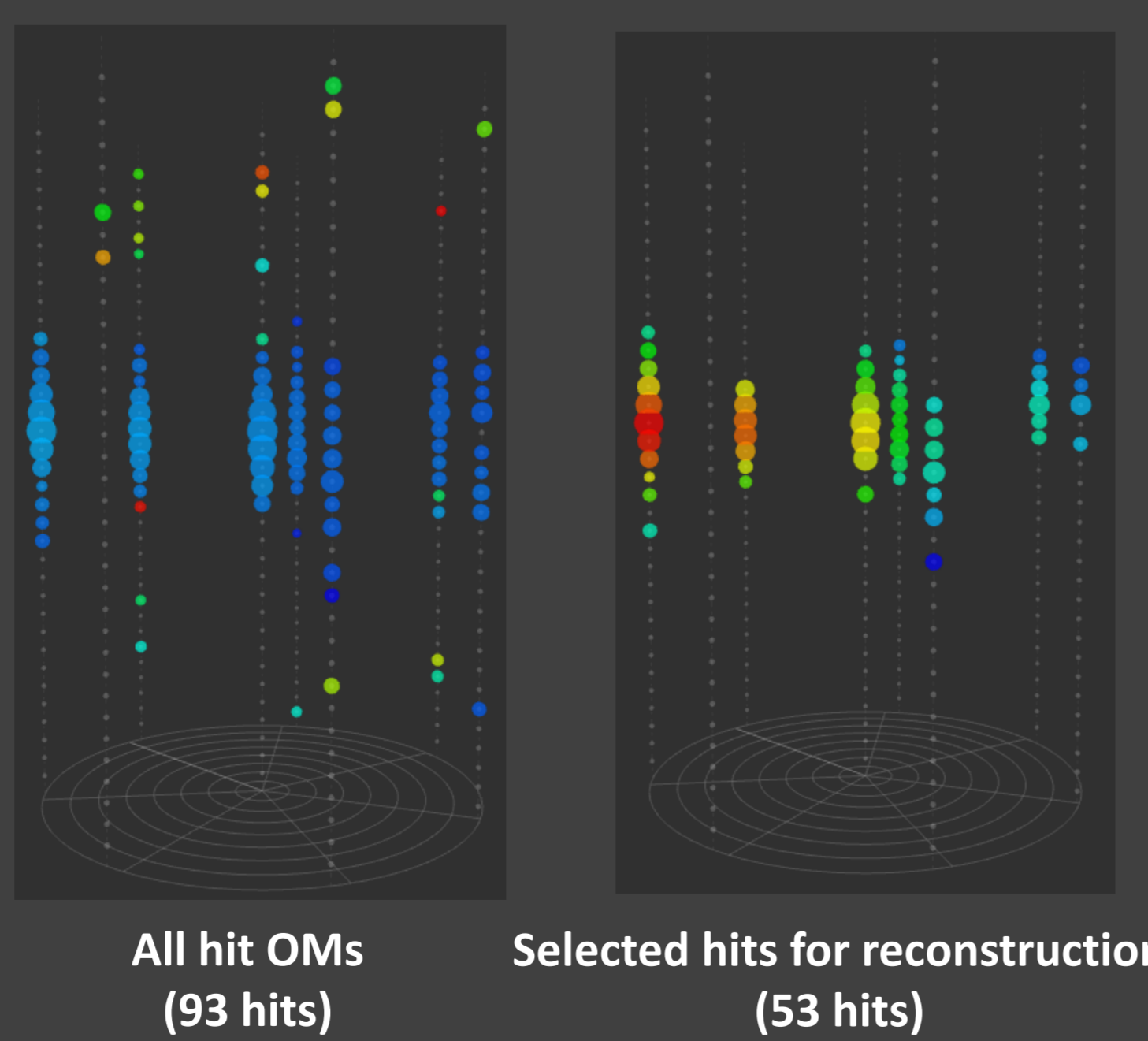
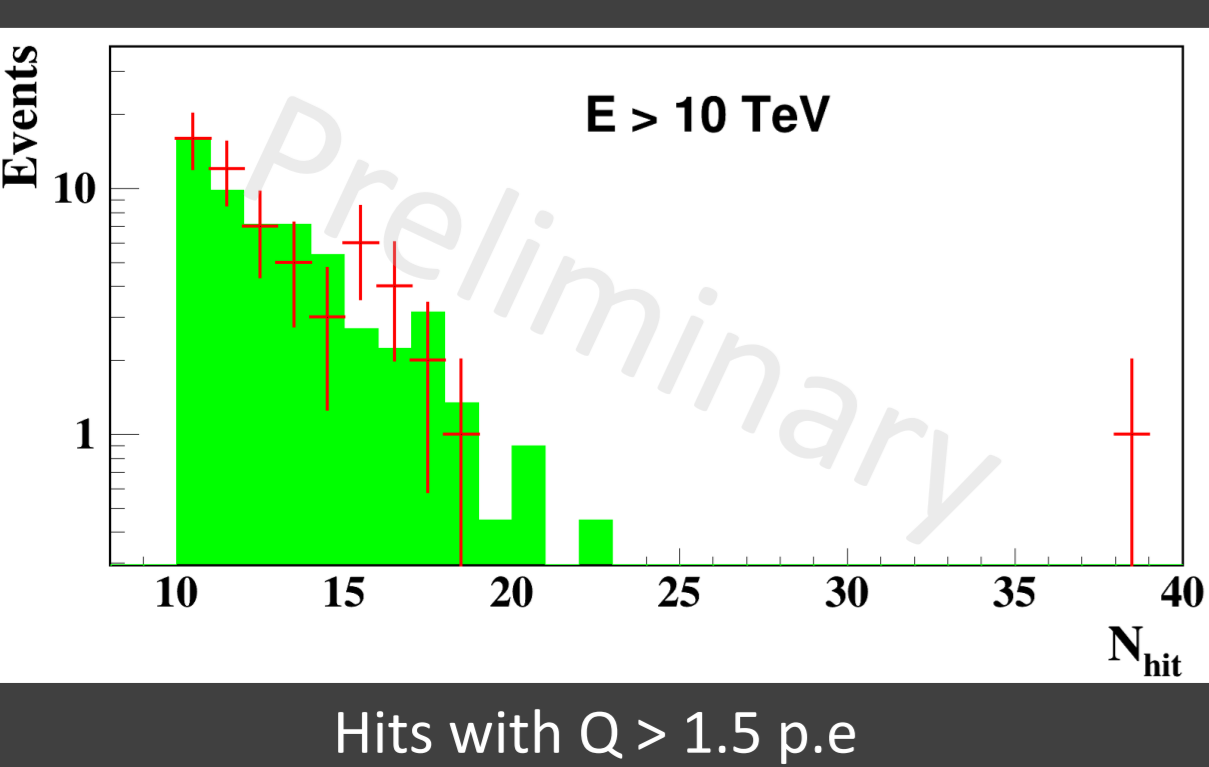
Cuts	Events	Rejection
Reconstruction of coordinates (Q > 1.5 p.e)	1 171 077	1
χ <sup>2</sup> < 2	316229	1/3.7
L <sub>a</sub> < 10 & η > 0	12931	1/90
E > 30 TeV	1291	1/900
E > 60 TeV	859	1/1360
E > 100 TeV	539	1/2000



	E <sup>-2.46</sup>	atm. ν <sub>e</sub>	atm. ν <sub>μ</sub>	atm. ν (total)
Probability of N <sub>hit</sub> > 16	0.047	0.0015	0.0026	0.0041

## Search for cascades in 2016 (182.0 live days)

Cuts	Events	Rejection
Coordinates reconstruction & N <sub>hit</sub> > 9	577495	1
χ <sup>2</sup> < 4	2405	1/240
Energy reconstruction		
L <sub>a</sub> < 20	374	1/6.4
η > 0	159	1/2.4
E > 10 TeV	57	1/2.8
E > 100 TeV	5	1/11.4
Total rejection factor:		1/115499



E=157 TeV, θ = 57°, φ = 249°,  
x=-25m, y=-37m, z=11m, ρ=44m

## Baikal Site Infrastructure:

- Railroad, Power line on the shore
- The new data taking center at the array site has been installed in 2016.
- The new shore lab was created on the site during summer 2017.
- The building in Baikalsk town is ready for the local lab and temporary OMs storage for detector maintenance and upgrade.



## The optical modules production facility in Dubna:

- The facility allows to produce and test up to 12 OM per day
- We need to produce and send to the site 600 OM per season



## Search for muon neutrino in 2016 (33 live days)

### Muon Neutrino Selection

Polar angle distribution of muons selected with the requirement of at least 6 hits OM's at 3 strings. Comparison of reconstructed events in obtained Data and simulated atmospheric muon flux generated with CORSIKA QGSJET

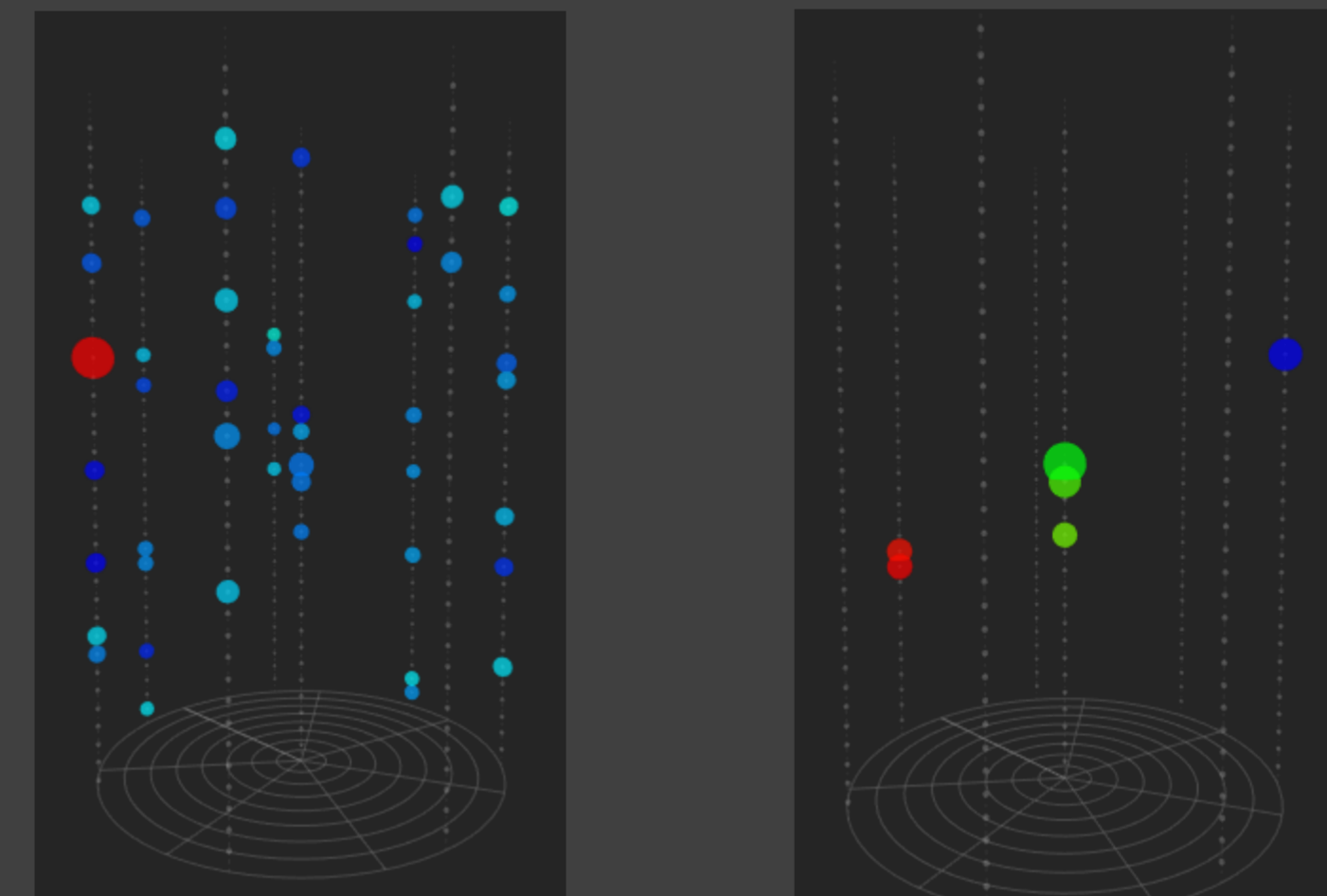
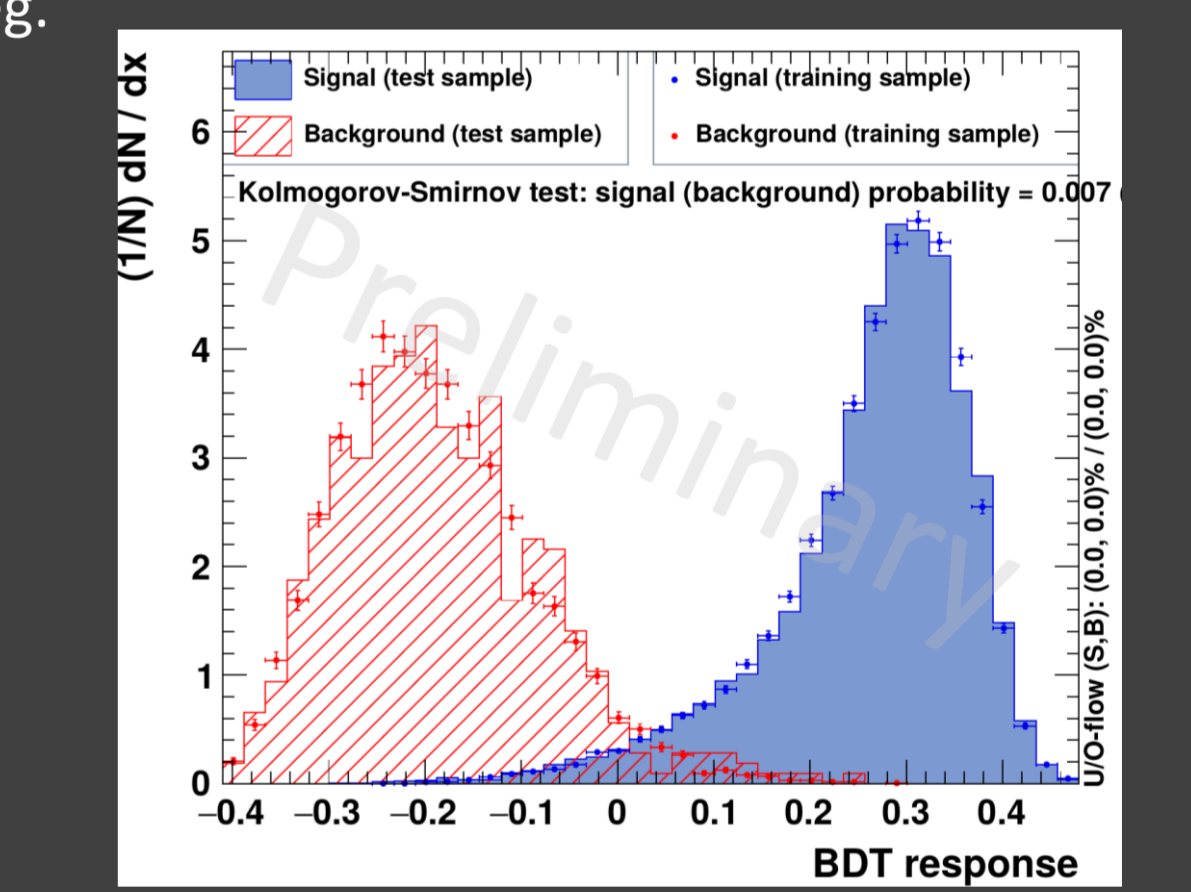
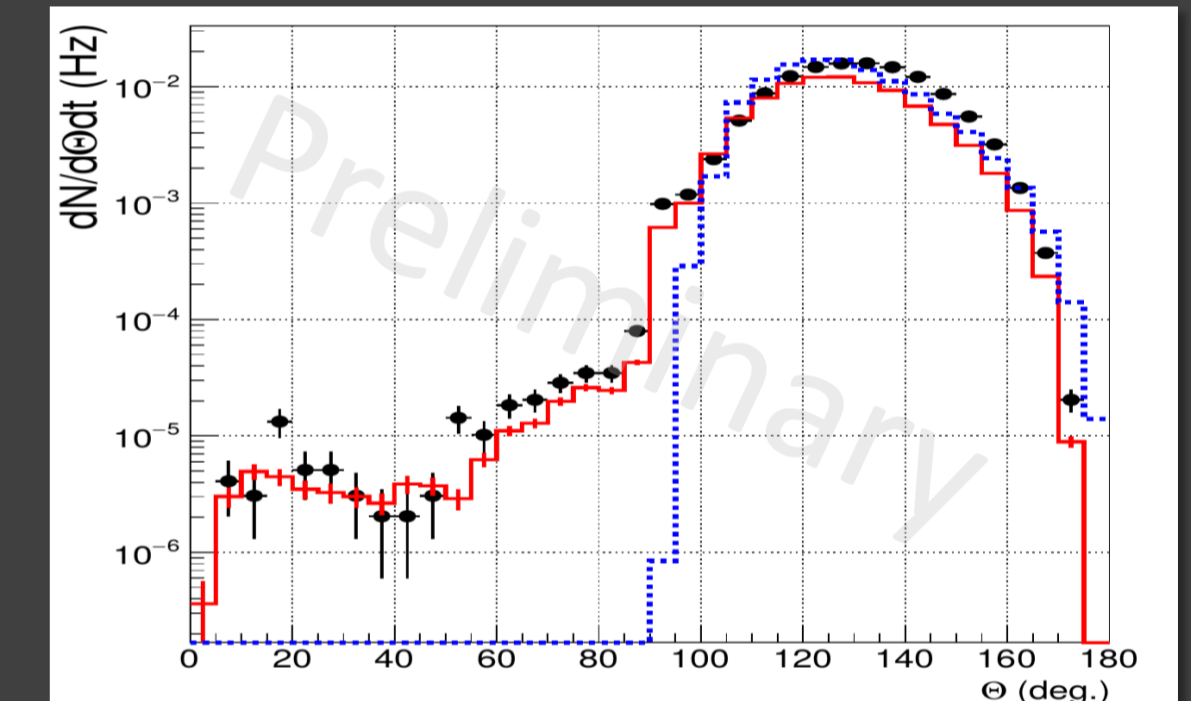
### Atmospheric background suppression

After track reconstruction and cuts on quality variables have been done, Boosted decision tree (BDT) was used. BDT is trained on events reconstructed as upgoing with 0 < θ < 80 deg. 30k signal events, 9k background evts.

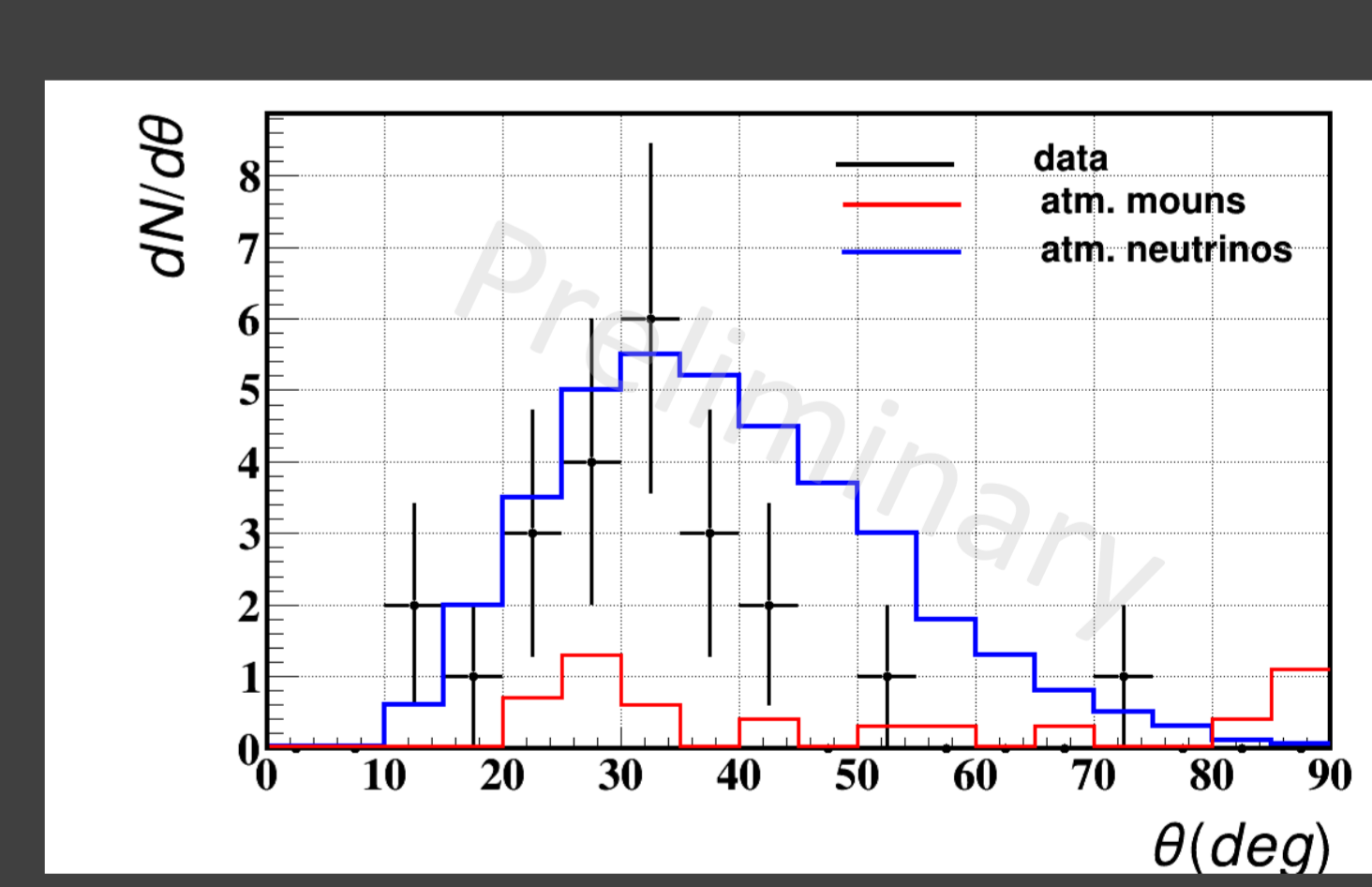
- BDT reponse cut:
  - > 0.20, 80% signal efficiency
  - > 0.25, 65%
  - > 0.30, 40%

### Result

Angular distribution for BDT > 0.2 cut  
23 events were selected in the signal region  
≈ 3 events – expected bkg. from atm. muons  
≈ 36 events – expected signal from atm. neutrinos

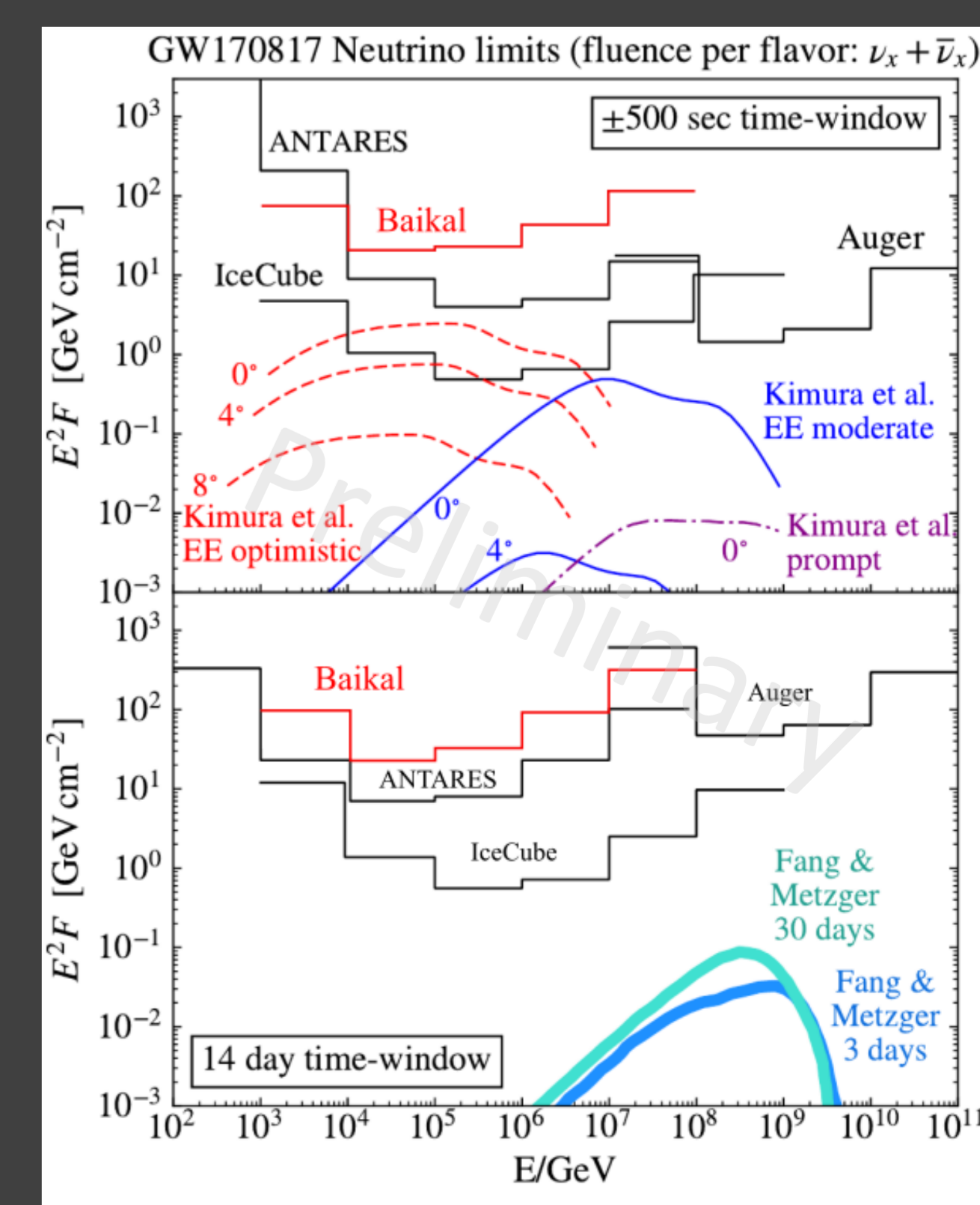
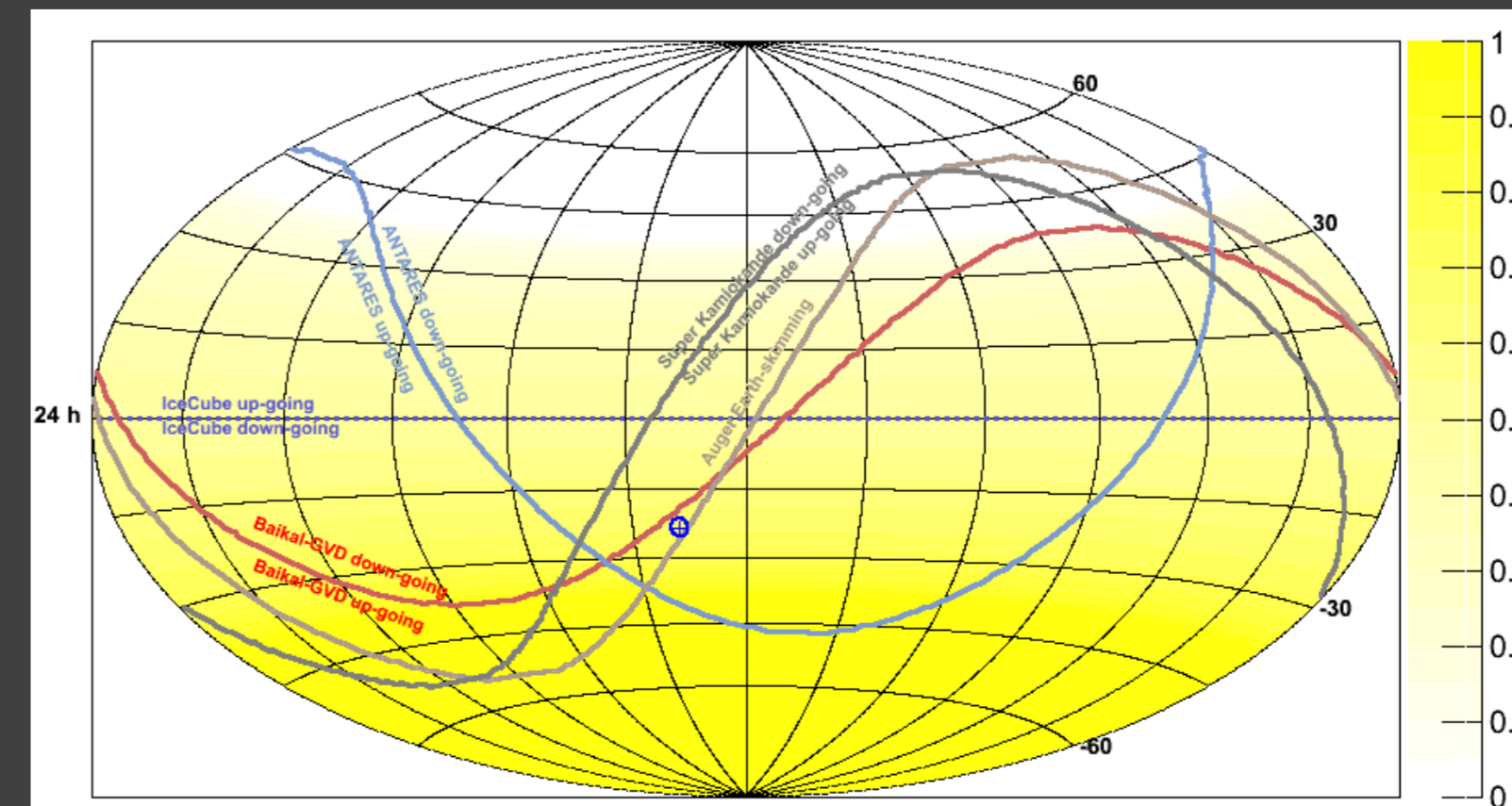


Before hit selection After hit selection (θ = 142.0°)



## Search for high-energy neutrinos associated with the GW170817

- GW: 17.08.2017, 12:41:04 UTC = 1502973664 sec UNIX
- NGS4993 at ~ 40 Mpc, equatorial coordinates α (J2000.0) = 13h 09m 48s.085, β (J2000.0) = -23°22'53".343
- Zenith angle of the source at registration time: 93.3°
- No neutrino events associated with GW170817 using cascade mode within both ± 500 sec and 14 days are observed
- Assuming E-2 spectral behavior and equal fluence in all flavors the upper limits at 90% C.L. are obtained on the neutrino



1. A.D. Avrorin et al., "Gigaton Volume Detector (GVD) in Lake Baikal: status of the project", XVII International Workshop on Neutrino Telescope, 13-17 March 2017, Venezia, Italy, <https://pos.sissa.it/301/962/pdf>

3. A.D. Avrorin et al., "The optical module of Baikal-GVD", 2016. 10 pp., Published in Phys.Part.Nucl.Lett. 13 (2016) no.6, 737-746, DOI: 10.1134/S1547477116060029.  
4. A.D. Avrorin et al., "Data acquisition system for the Baikal-GVD neutrino telescope", Published in Phys.Part.Nucl. 47 (2016) no.6, 933-937, DOI: 10.1134/S1063779616060058.  
5. A.D. Avrorin et al., "Cascades in GVD", 35th ICRC-2017 12-20 июля 2017, Bexco, Busan, S.Korea, <https://pos.sissa.it/301/962/pdf>