

Exponentially Light Dark Matter from Coannihilation

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Cargèse

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Coannihilation in a simple model

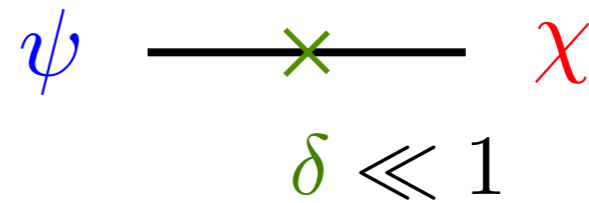
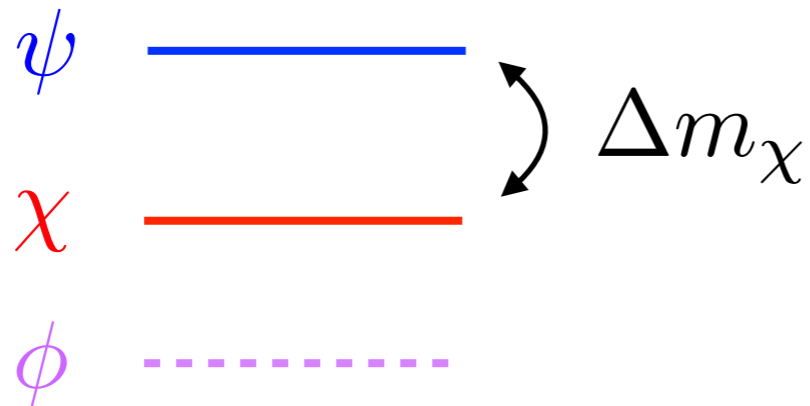
$$\mathcal{L} \supset -\frac{m_\phi^2}{2} \phi^2 - \frac{m_\psi^2}{2} \psi^2 - \frac{y}{2} \phi \psi^2$$

ψ —————

ϕ - - - - -

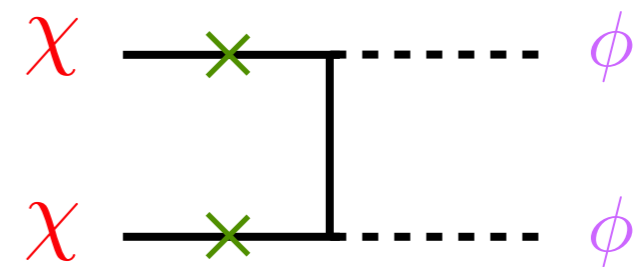
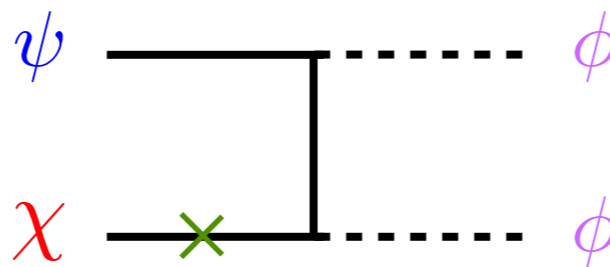
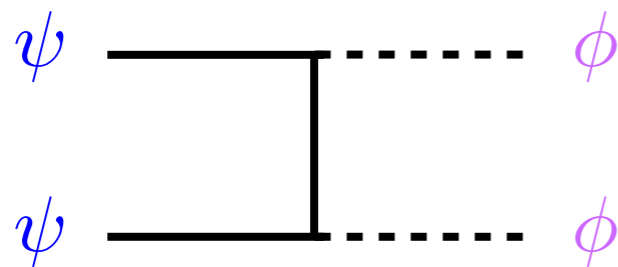
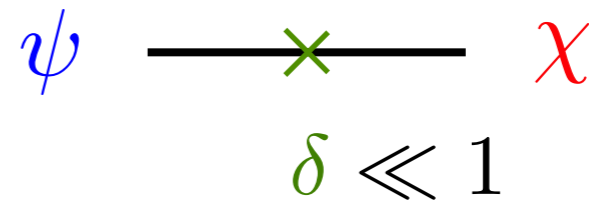
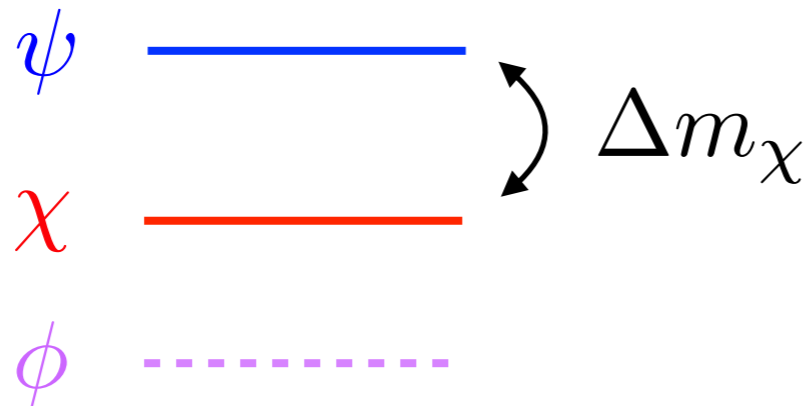
Coannihilation in a simple model

$$\mathcal{L} \supset -\frac{m_\phi^2}{2} \phi^2 - \frac{m_\psi}{2} \psi^2 - \frac{y}{2} \phi \psi^2 - \frac{m_\chi}{2} \chi^2 - \delta m_\chi \chi \psi$$



Coannihilation in a simple model

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$$\langle \sigma_{\text{eff}} v \rangle \propto e^{-2\Delta x_f}$$

$$\delta^2 e^{-\Delta x_f}$$

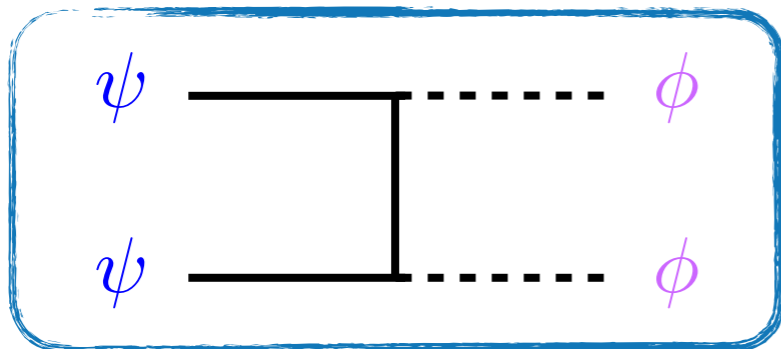
$$\delta^4$$

Coannihilation in a simple model

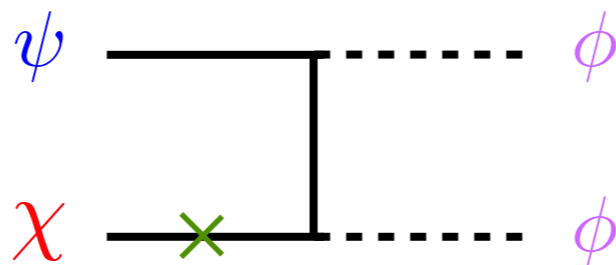
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Sterile coannihilation:

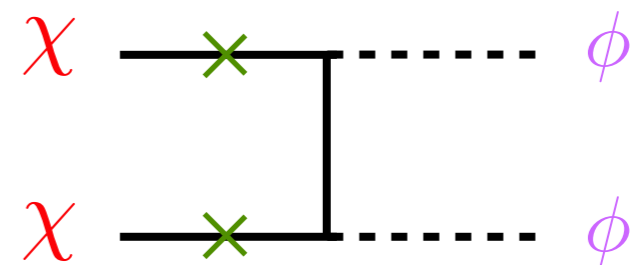
$$\delta^2 \ll e^{-\Delta x_f}$$



$$\langle \sigma_{\text{eff}} v \rangle \propto e^{-2\Delta x_f}$$



$$\delta^2 e^{-\Delta x_f}$$

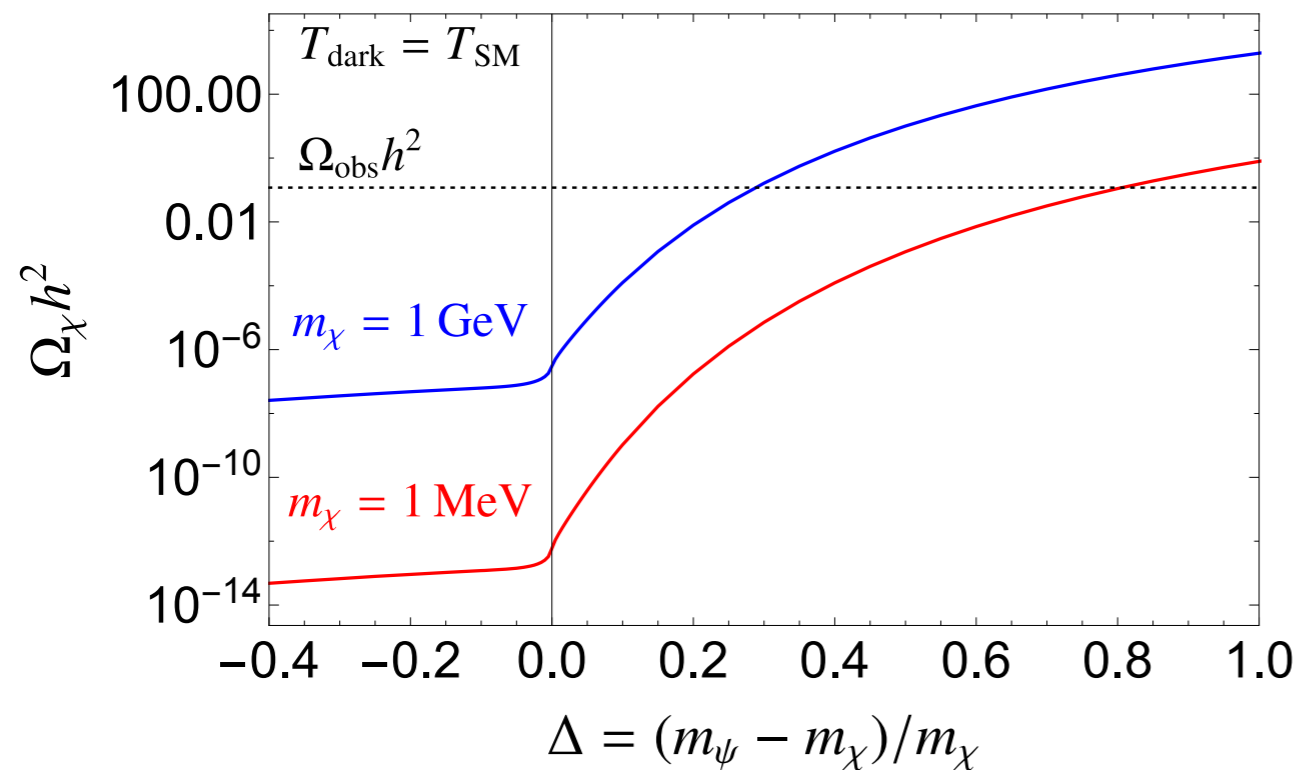


$$\delta^4$$

Sterile coannihilation:

$$\Omega_{\text{DM}} h^2 \sim \frac{1}{T_{\text{EQ}} M_{\text{Pl}} \langle \sigma_{\text{eff}} v \rangle} \sim \frac{1}{T_{\text{EQ}} M_{\text{Pl}}} \frac{e^{2\Delta x_f}}{\sigma_{\psi\psi}}$$

Relic Density vs Mass Splitting



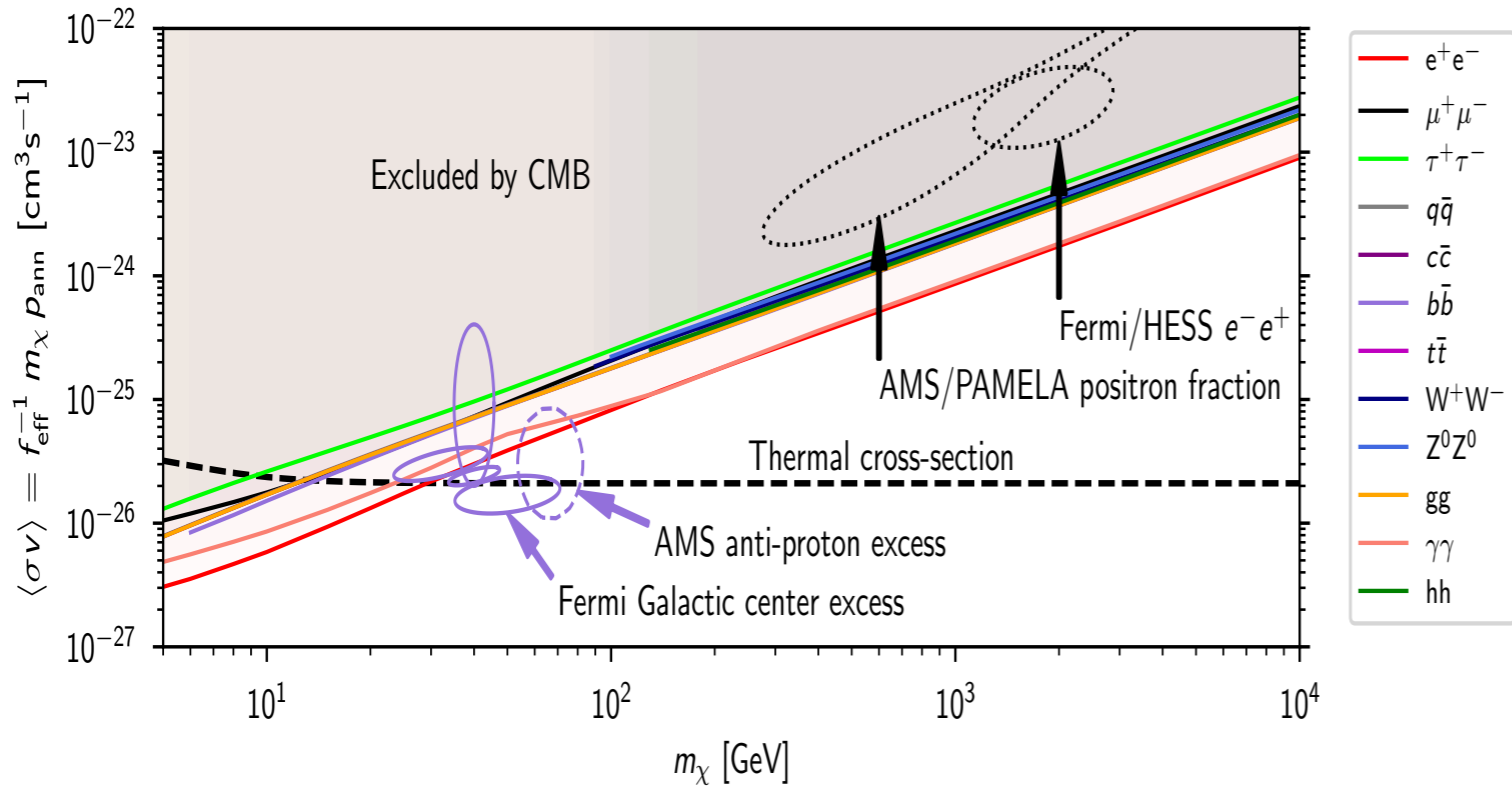
$$m_\chi \sim e^{-\Delta x_f} \sqrt{T_{\text{EQ}} M_{\text{Pl}}}$$

**Exponentially lighter
than the weak scale**

CMB

Planck collaboration, 2018

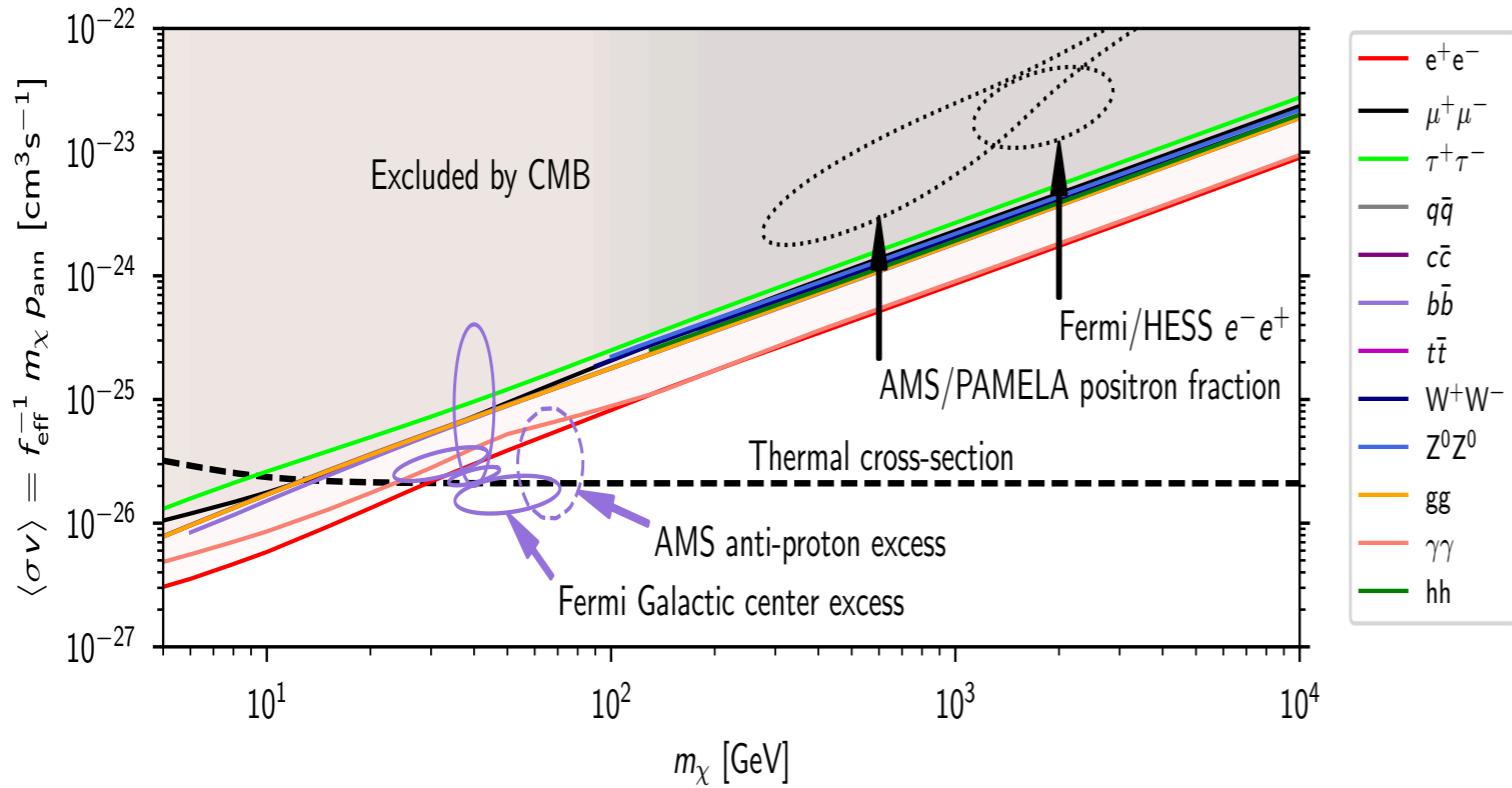
Excluded



CMB

Planck collaboration, 2018

Excluded

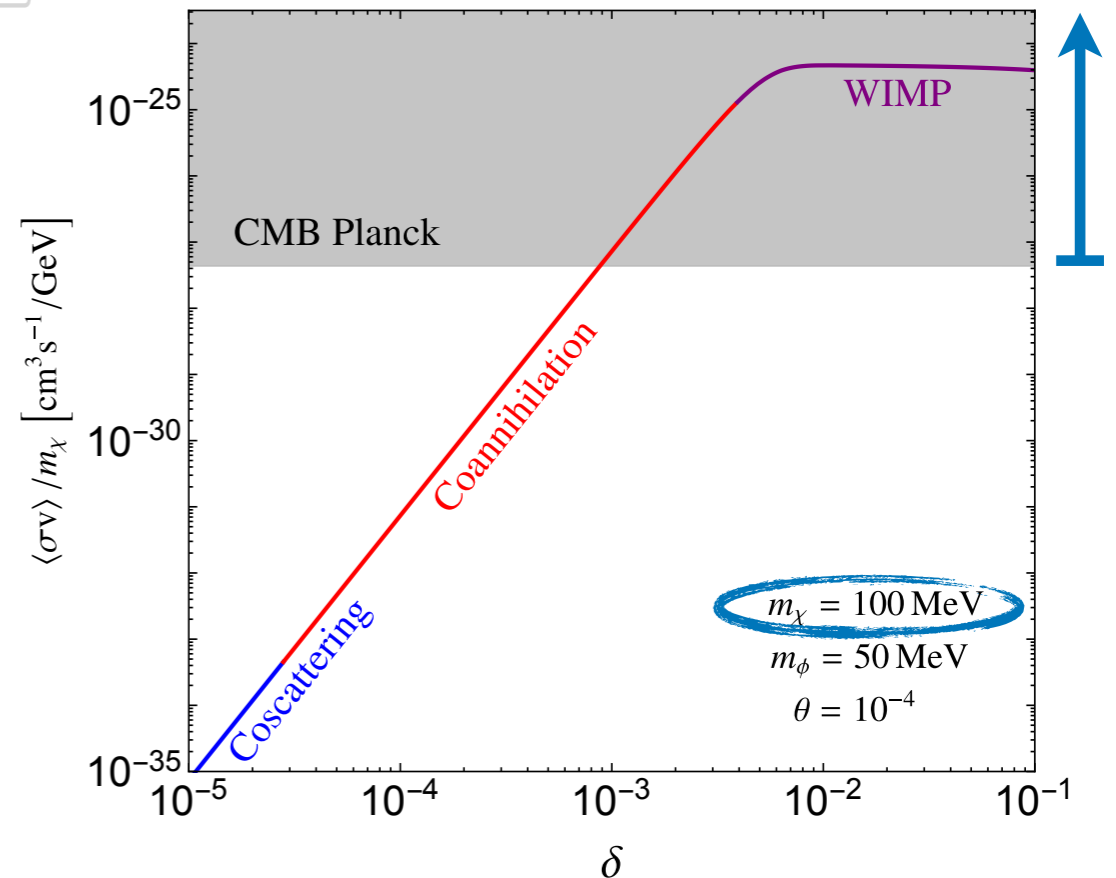


Naturally avoids the CMB bound for sub-GeV DM

Sterile

coannihilation: the relic density is insensitive to the DM coupling

DM annihilation cross-section



Sterile coannihilation realizations

**Annihilation to
SM particles**

**Annihilation to Dark
Sector (DS) particles**

Sterile coannihilation realizations

**Annihilation to
SM particles**

**Annihilation to Dark
Sector (DS) particles**

Need coupling to radiation

$$T_{\text{DM}} \sim a^{-1}$$

1. Introduce dark light degrees of freedom

$$T_{\text{dark}} \neq T_{\text{SM}}$$

2. Couple the Dark Sector to the SM

$$T_{\text{dark}} = T_{\text{SM}}$$

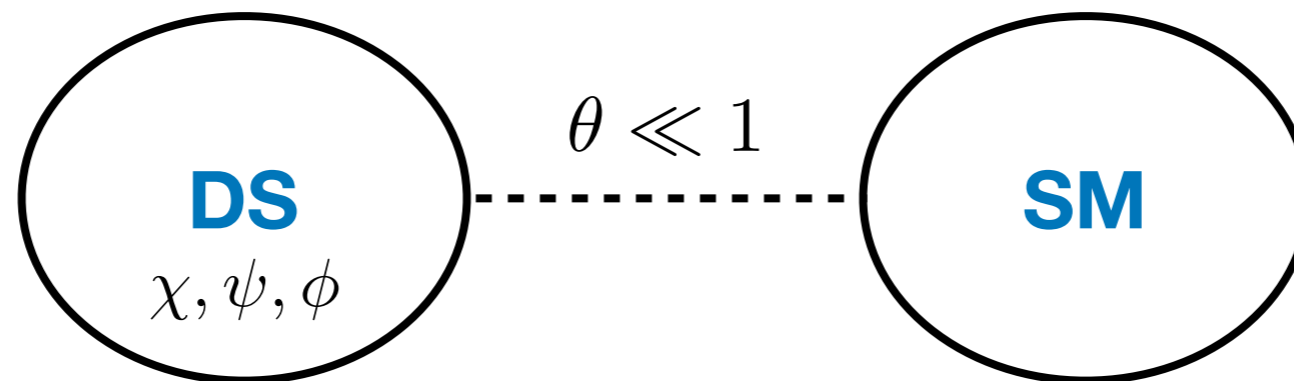
Backup

Coupling to the SM: Higgs portal

$$V \supset -\frac{m_H^2}{2} + \frac{\lambda}{4}|H|^4 + \frac{\mu_\phi^2}{2}\phi^2 + \frac{\lambda_\phi}{4!}\phi^4 + a_\phi\phi|H|^2$$

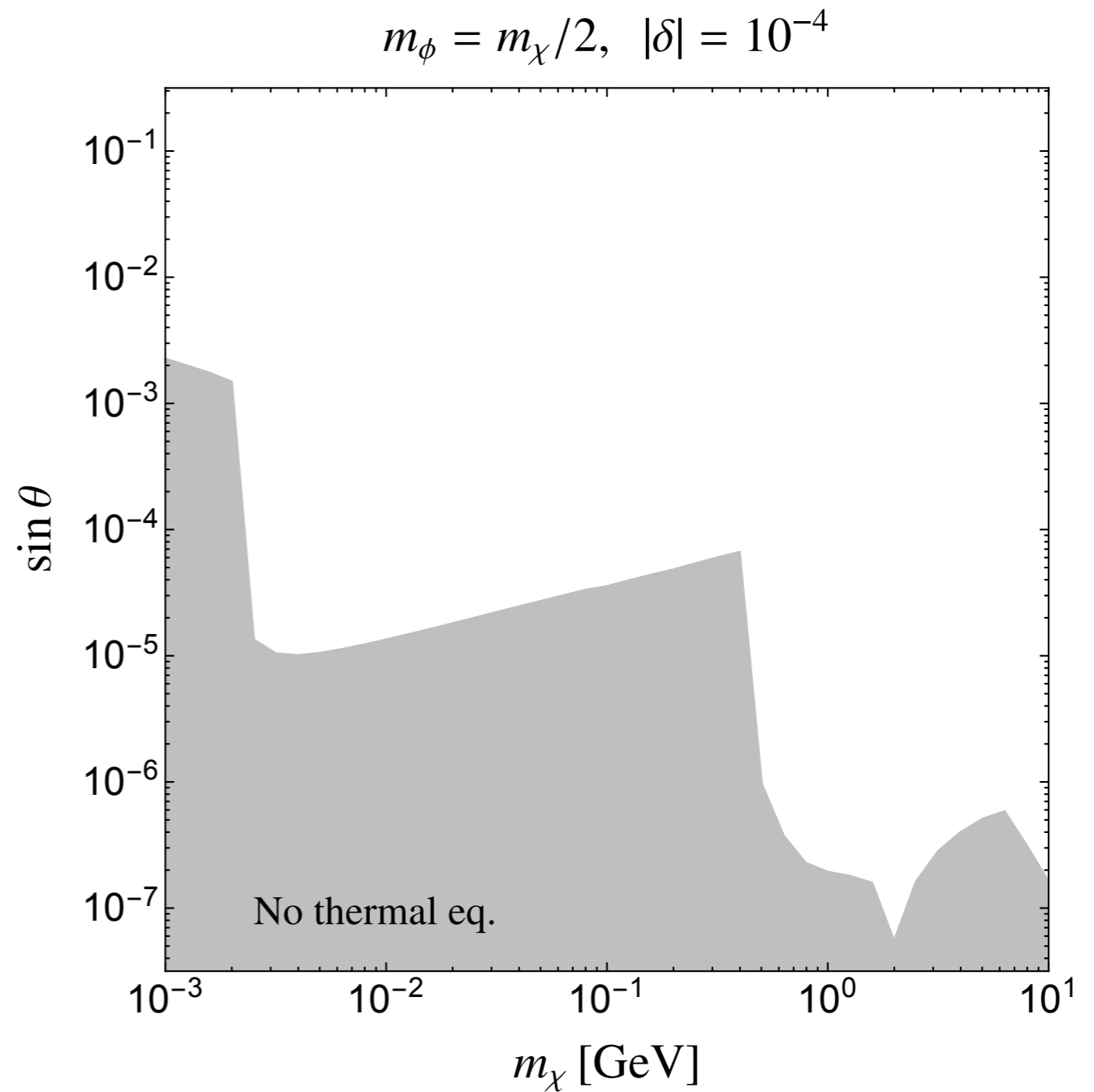
The DM scalar mixes with the SM Higgs

$$\tan 2\theta = \frac{4a_\phi v}{\lambda v^2 - \lambda_\phi v_\phi^2 - 2\mu_\phi^2}$$



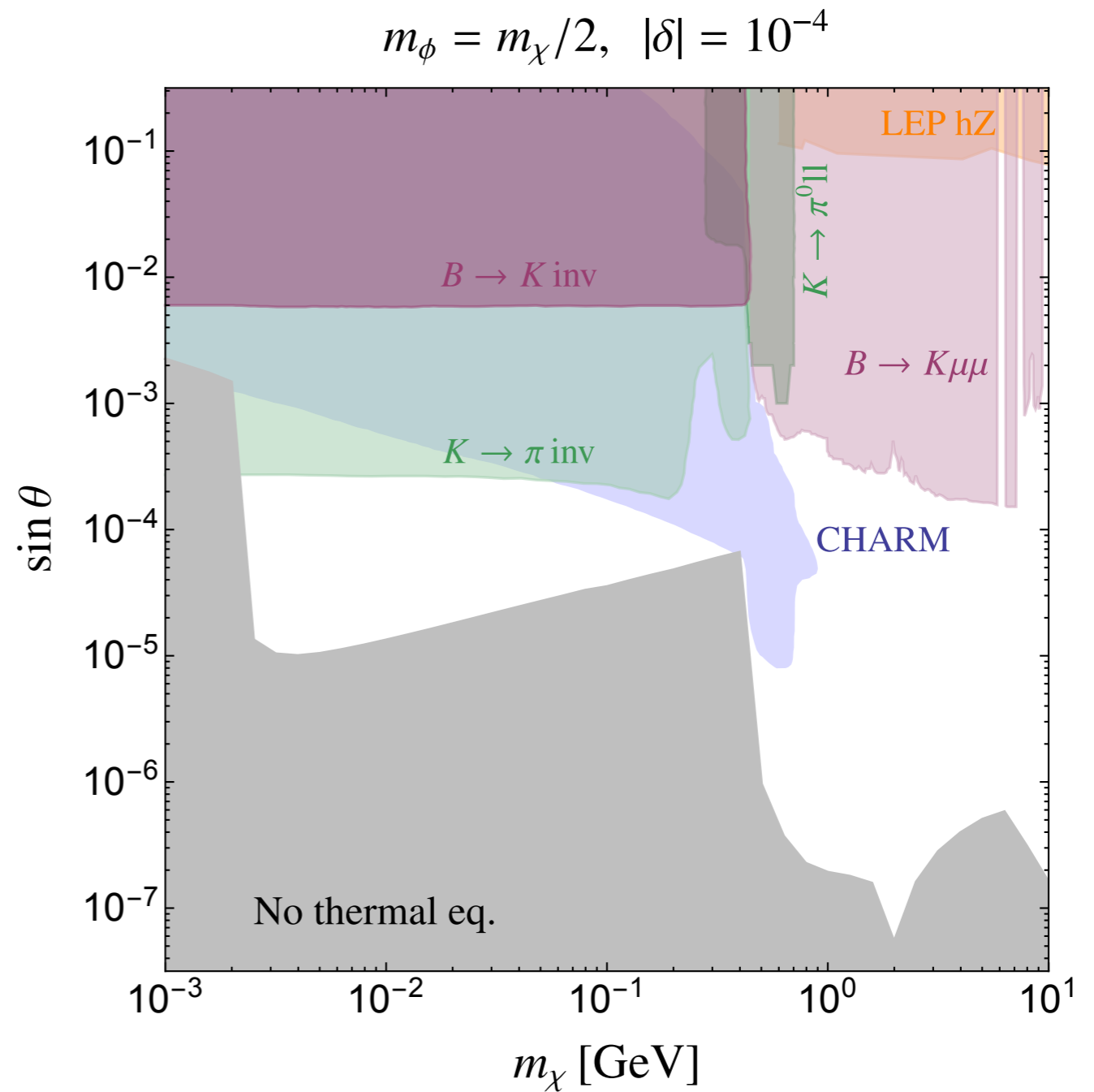
Higgs portal phenomenology

- Require DS and SM equilibrium at DM freeze-out



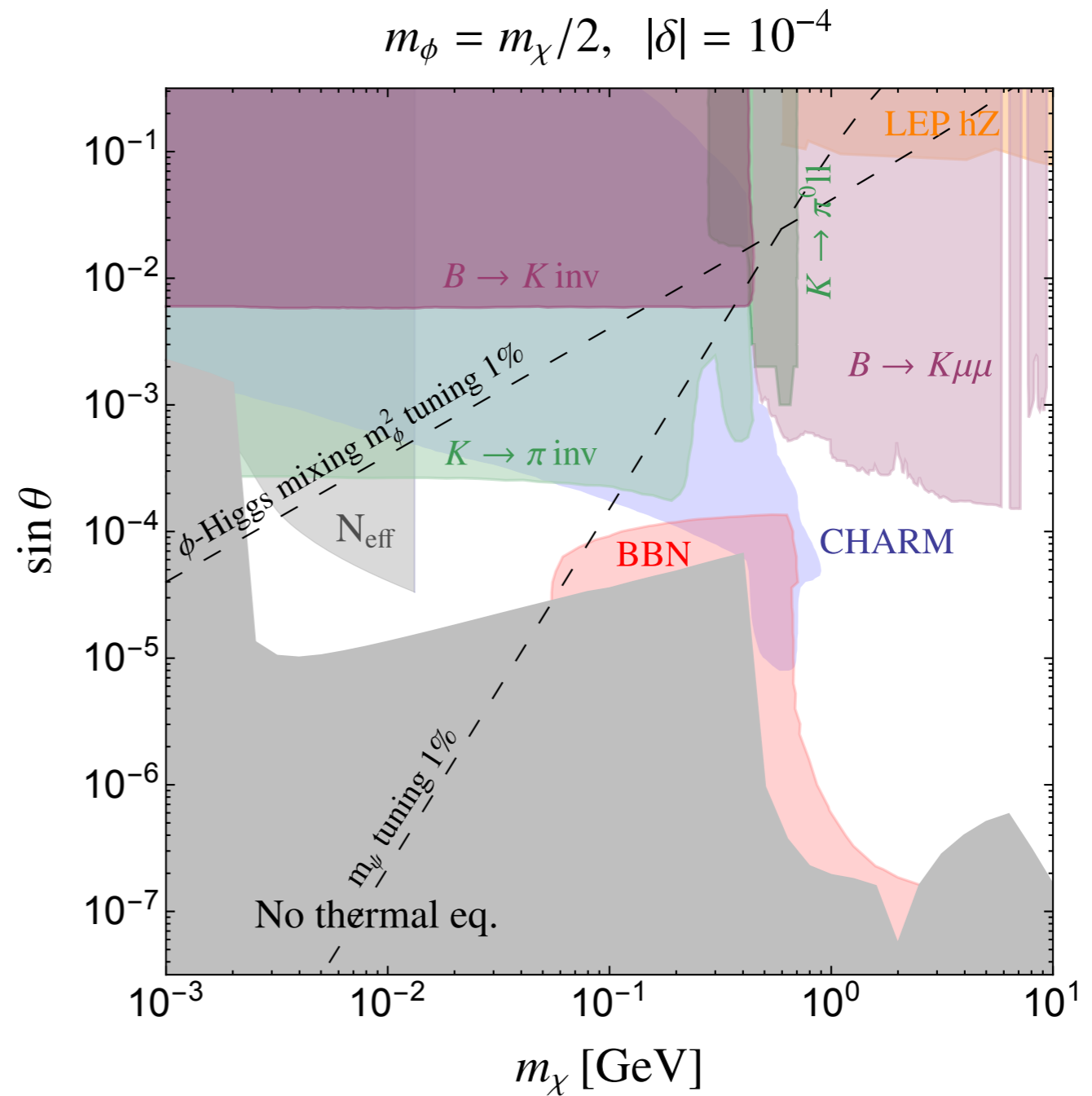
Higgs portal phenomenology

- Require DS and SM equilibrium at DM freeze-out
- Constraints from meson decays on the ϕ -Higgs mixing



Higgs portal phenomenology

- Require DS and SM equilibrium at DM freeze-out
- Constraints from meson decays on the ϕ -Higgs mixing
- CMB constraints on light degrees of freedom (N_{eff})
- BBN constraints on late ψ decay



Higgs portal phenomenology

- Future experiments looking for long-lived ϕ can probe new parameter space

