

Composite Higgs Phenomenology

Astrophysical Signals of Composite Dark Matter

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Mass: From the Higgs to Cosmology

Cargese 2018

Composite Higgs Models

- A compelling solution to the **hierarchy problem**:

$$\mu_{\text{EW}}^2 / \Lambda_{\text{Pl}}^2 \sim 10^{-28} \quad \text{👎}$$

- The Higgs (and possible *other scalars*) are **pseudo-Goldstone bosons**
- They emerge when $G \rightarrow H$ spontaneously
- G must be explicitly broken to allow the generation of a potential for the pGBs

new physics!

$$f \approx \text{TeV}$$

EW physics

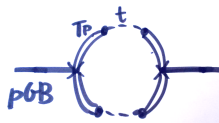
$$v \approx 246 \text{ GeV}$$

Group theory constraints on **EWSB** and the **properties** of the new scalars

A Dark Matter Candidate?

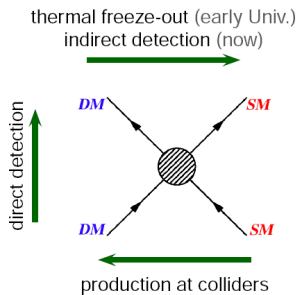
Natural light scalars with a mass at the EW scale:

$$m^2 \sim \frac{y_q^2}{(4\pi)^2} f^2, \quad f \sim \text{TeV}.$$



Justifies “for free” the
WIMP miracle:

$$\begin{aligned} \Omega h^2 &\approx 0.1 \times \left(\frac{3 \times 10^{-26} \text{cm}^3 \text{s}^{-1}}{\langle \sigma v \rangle} \right) \\ &\approx 0.1 \times \left(\frac{\alpha^2 / (200 \text{ GeV})^2}{\langle \sigma v \rangle} \right). \end{aligned}$$



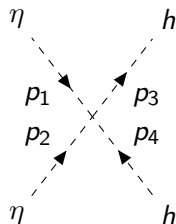
Minimal Model

The potential is dynamically generated:

$$V_{\text{eff}}(h, \eta) = \frac{1}{2}(\mu_h^2 h^2 + \mu_\eta^2 \eta^2) + \frac{\lambda_h}{4} h^4 + \frac{\lambda}{2} h^2 \eta^2 + \frac{\lambda_\eta}{4} \eta^4 + \mathcal{O}((h, \eta)^6)$$

- ✓ **Direct Searches** \Rightarrow Small λ
- ✓ **Observed relic density** \Rightarrow Not so small...

In CHMs:



$$= -2i\lambda - \frac{i}{f^2}(p_1 + p_2) \cdot (p_3 + p_4)$$

New derivative interactions

Case of heavier extra pGBs: *ArXiv 1801.06537*

Complementary Case: work in progress

- Wide region of the parameter space **still to be tested:**

$$\text{DM DM} \rightarrow \kappa \kappa \rightarrow \text{SM SM} ;$$

- The new channel **evades current constraints** ;
- In some models, it can even have a **large annihilation rate** ;
- Emerging of a large number of extra pGBs ($>$ SM scalars) ;
- *Limiting Case:*

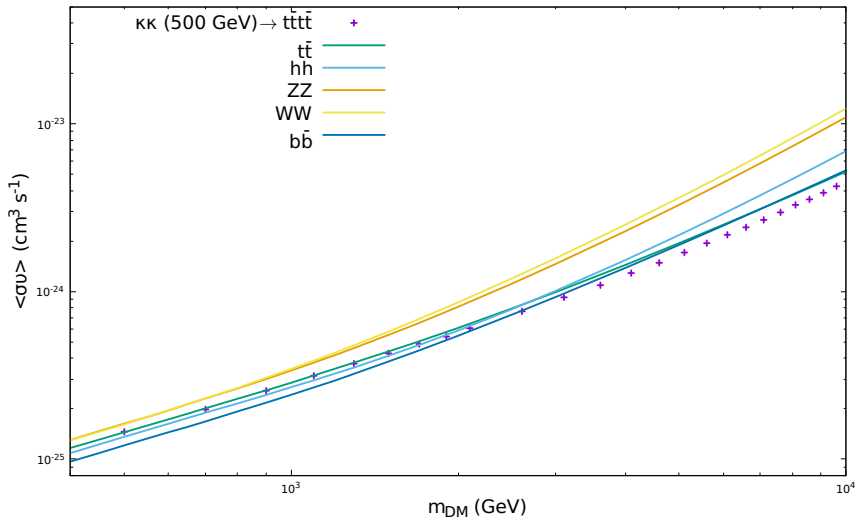
$$\lambda \approx \left(\frac{m_\eta}{f} \right)^2 ,$$

in $\text{DM DM} \rightarrow \text{hh (WW; ZZ)}$, so that the **new channel dominates**.



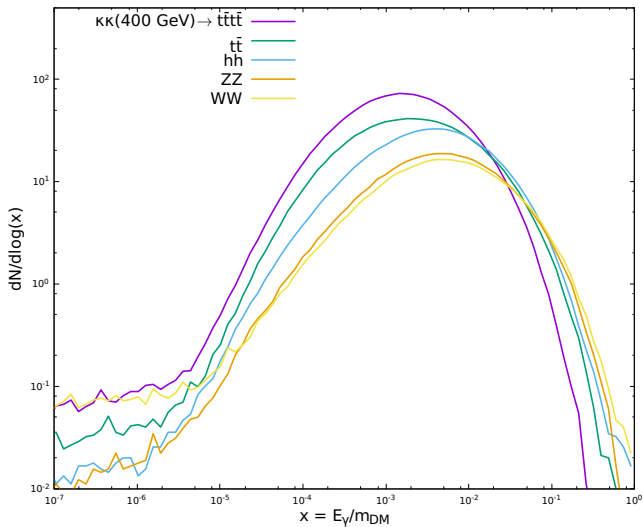
Deserves a proper study!

Fermi Limits



Preliminary Plots — Indirect Detection

γ Spectrum
 $m_{\text{DM}} = 500 \text{ GeV}$



Continuing to investigate new signatures...

Thank you for your attention!

*(special thanks to **LIP** and **COST Action** for the financial support)*