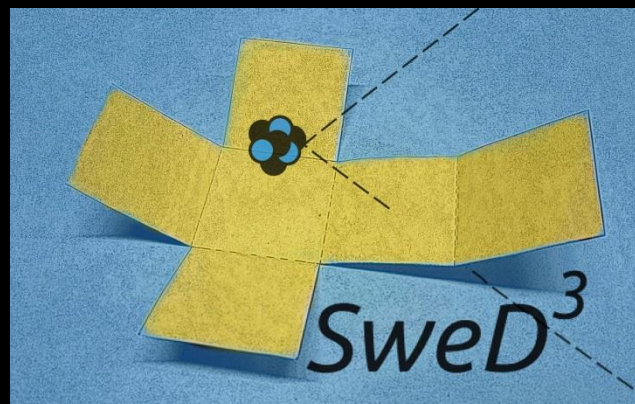
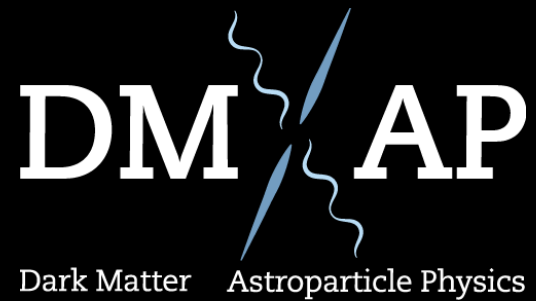


# NEW IDEAS FOR DIRECT DETECTION OF DARK MATTER



Stockholm  
University



Jan Conrad

The puzzle of Dark Matter,  
DESY, October 2018

# LAST TIME IN HAMBURG 2011

## Concluding remarks

- Essentially all bases covered.
- "Stacking" methodology big step forward
  - We have reached the predicted 5yr sensitivity (optimistic) already now
  - currently constraining thermal WIMP cross-section for masses  $< 30$  GeV ( $b\bar{b}$ ,  $\mu\mu$ ) *including J-factor uncertainties*
  - Improvement by around a factor 4-5 until the end of the Fermi-LAT mission (only with dwarfs).
- All-sky diffuse analyses: challenging.
- Line features: robust analysis, clear signature, but detector can fool
- *Close to submission/internal refereeing: Dwarf stacking, DM Satellites, update on line search, further down: clusters, anisotropies, halo.*
- Looking ahead: "vanilla" CTA will be competitive above  $\sim 300$  GeV.

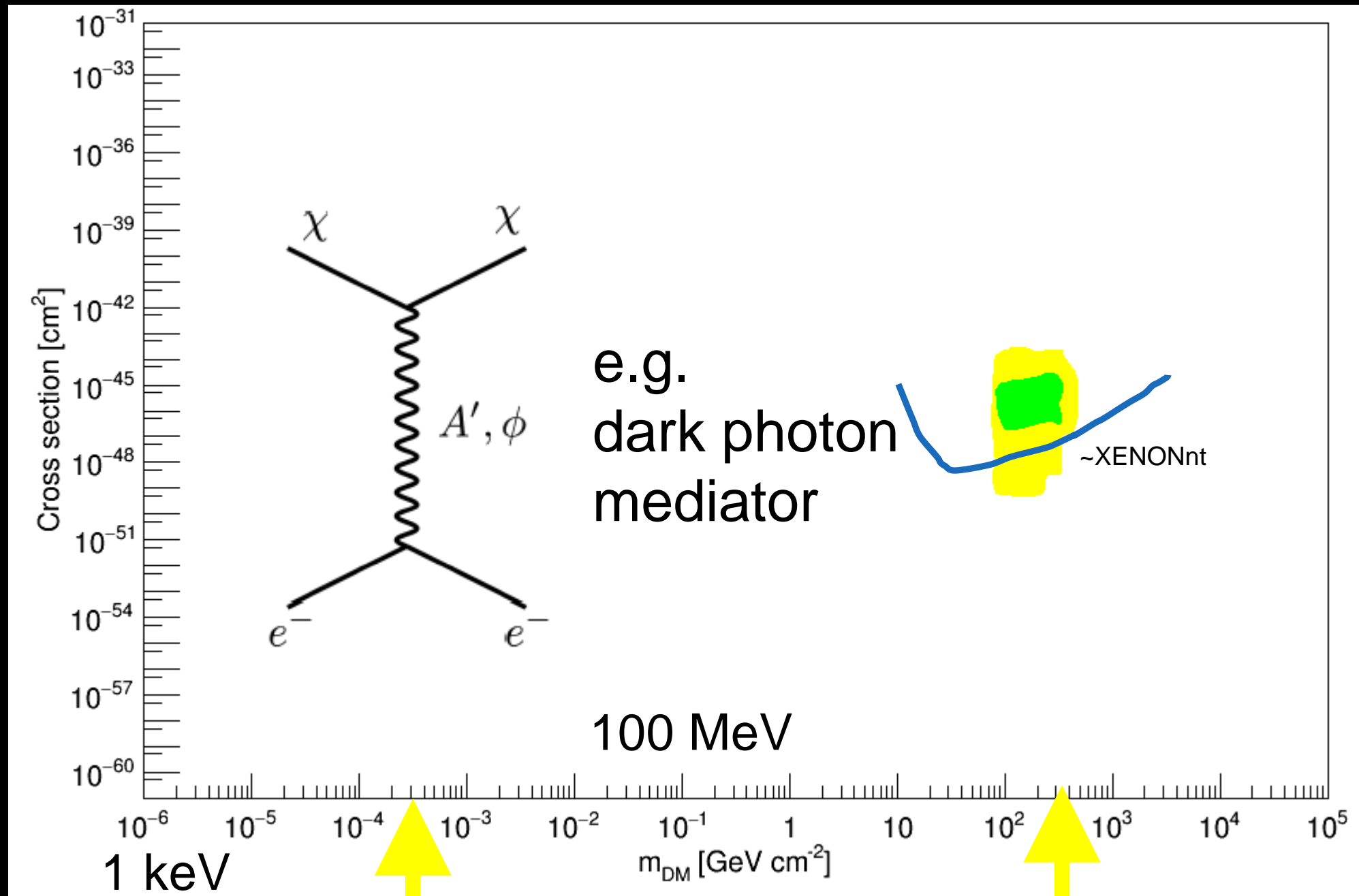
09-08-07

Jan Conrad, Oskar Klein Centre, Stockholm Universitet

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Disclaimer: I have an ongoing bet for XENONnT/LZ to discover WIMPs

# DARK SECTOR DARK MATTER



Super-light dark matter (absorption)

Light dark matter (scattering) == Dark Sector Dark Matter

WIMPs

# KINEMATICS OF DIRECT DETECTION

$$E_D = \frac{q^2}{2m_N} \quad q_{max} = 2m_{DM}v \quad v \sim 10^{-3}$$

- Nucleon scattering

$$E_D \sim 1 \text{ eV} \xrightarrow{\text{yields}} \mathbf{m_{DM} \sim 100 \text{ MeV}}$$

$$E_{kin} = \frac{m_{DM}v^2}{2} \sim 100 \text{ eV}$$

- Electron scattering

$$E_D \sim 1 \text{ eV} \xrightarrow{\text{yields}} \mathbf{m_{DM} \sim 1 \text{ MeV}}$$

$$E_{kin} = \frac{m_{DM}v^2}{2} \sim 1 \text{ eV}$$

**Might want to go  
to ~keV**

# COVERED IN THIS TALK

## ○ Future:

$$n_{DM} = \frac{\rho_{DM}}{m_{DM}} \sim \frac{0.4 \text{ GeV}}{m_{DM}}$$

Main Science Goal	Experiment	Target	Readout	Estimated Timeline
Sub-GeV Dark Matter (Electron Interactions)	SENSEI	Si	charge	ready to start project (2 yr to deploy 100g)
	DAMIC-1K	Si	charge	ongoing R&D 2018 ready to start project (2 yr to deploy 1 kg)
	UA'(1) liquid Xe TPC	Xe	charge	ready to start project (2 yr to deploy 10kg)
	Scintillator w/ TES readout	GaAs(Si,B)	light	2 yr R&D 2020 in sCDMS cryostat
	NICE; NaI/CsI cooled crystals	NaI CsI	light	3 yr R&D 2020 ready to start project
	Ge Detector w/ Avalanche Ionization Amplification	Ge	charge	3 yr R&D 1 yr 10kg detector 1 yr 100kg detector
	PTOLEMY-G3, 2d graphene	graphene	charge directionality	1 yr fab prototype 1 yr data
	supercond. Al cube	Al	heat	10+ yr program

*For details and experiments targeting nuclear scattering (e.g CRESST etc)*

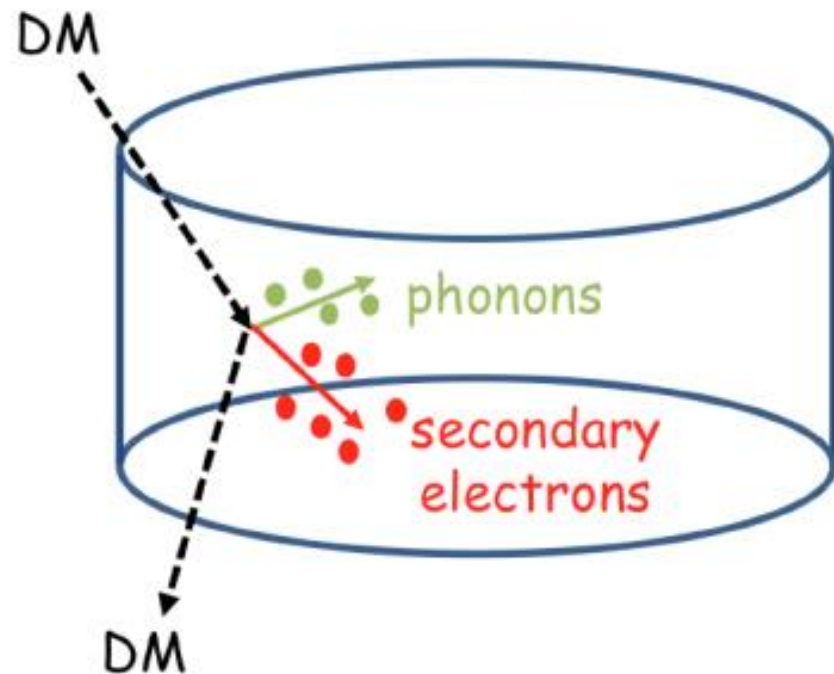
*Battaglieri et al:  
arxiv:1707.04591*

## ○ After the future: Quantum Materials for Dark Matter (QM4DM)

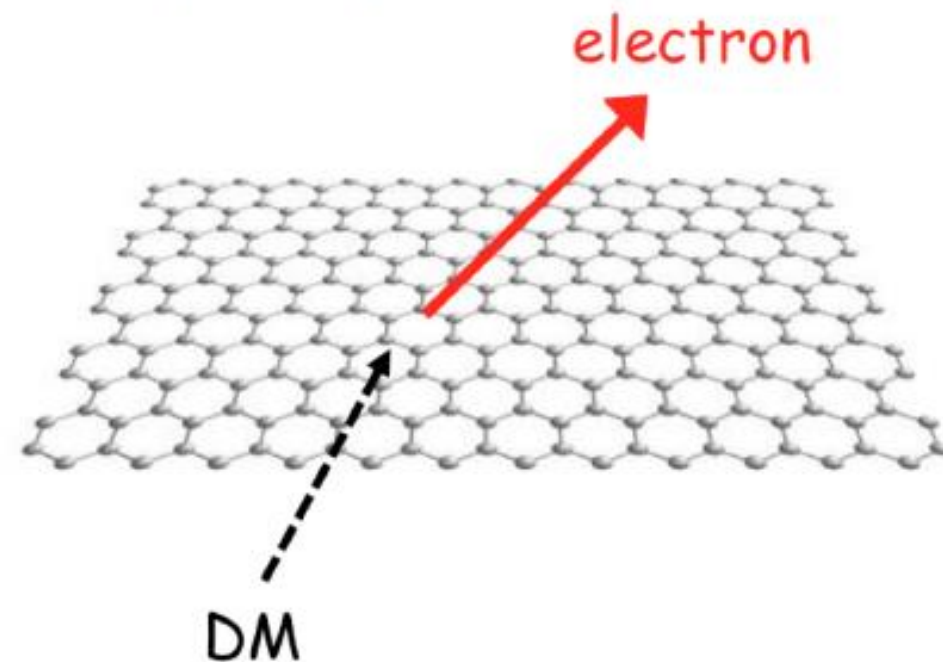
# 2D TARGETS (E.G. GRAPHENE)

*Hochberg et al, Phys. Lett. B 772 (2017) 239*

Detect “secondaries” → loose directional information:



Keep directional information if detect primary





# PTOMELY-G3

- Graphene-based
- Main purpose: detection of the cosmic neutrino background
- Phase I: directional dark matter detection
- Phase II:  $C_{\nu B}$
- Prototype approved for move to LNGS

## PTOLEMY: A Proposal for Thermal Relic Detection of Massive Neutrinos and Directional Detection of MeV Dark Matter

E. Baracchini<sup>3</sup>, M.G. Betti<sup>11</sup>, M. Biasotti<sup>5</sup>, A. Boscá<sup>16</sup>, F. Calle<sup>16</sup>, J. Carabe-Lopez<sup>14</sup>, G. Cavoto<sup>10,11</sup>, C. Chang<sup>22,23</sup>, A.G. Cocco<sup>7</sup>, A.P. Colijn<sup>13</sup>, J. Conrad<sup>18</sup>, N. D'Ambrosio<sup>2</sup>, P.F. de Salas<sup>17</sup>, M. Favrezi<sup>6</sup>, A. Ferella<sup>18</sup>, E. Ferri<sup>6</sup>, P. Garcia-Abia<sup>14</sup>, G. Garcia Gomez-Tejedor<sup>15</sup>, S. Gariazzo<sup>17</sup>, F. Gatti<sup>5</sup>, C. Gentile<sup>25</sup>, A. Giachero<sup>6</sup>, J. Gudmundsson<sup>18</sup>, Y. Hochberg<sup>1</sup>, Y. Kahn<sup>26</sup>, M. Lisanti<sup>26</sup>, C. Mancini-Terracciano<sup>10</sup>, G. Mangano<sup>7</sup>, L.E. Marcucci<sup>9</sup>, C. Mariani<sup>11</sup>, J. Martínez<sup>16</sup>, G. Mazzitelli<sup>4</sup>, M. Messina<sup>20</sup>, A. Molinero-Vela<sup>14</sup>, E. Monticone<sup>12</sup>, A. Nucciotti<sup>6</sup>, F. Pandolfi<sup>10</sup>, S. Pastor<sup>17</sup>, J. Pedrós<sup>16</sup>, C. Pérez de los Heros<sup>19</sup>, O. Pisanti<sup>7,8</sup>, A. Polosa<sup>10,11</sup>, A. Puiu<sup>6</sup>, M. Rajteri<sup>12</sup>, R. Santorelli<sup>14</sup>, K. Schaeffner<sup>3</sup>, C.G. Tully<sup>26</sup>, Y. Raitses<sup>25</sup>, N. Rossi<sup>10</sup>, F. Zhao<sup>26</sup>, K.M. Zurek<sup>21,22</sup>

*Barrachini et al*  
*1808.01892*

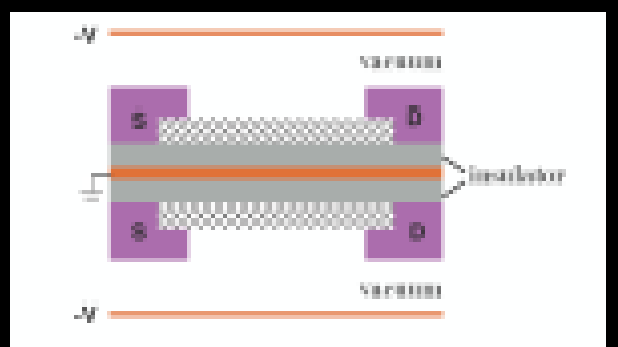
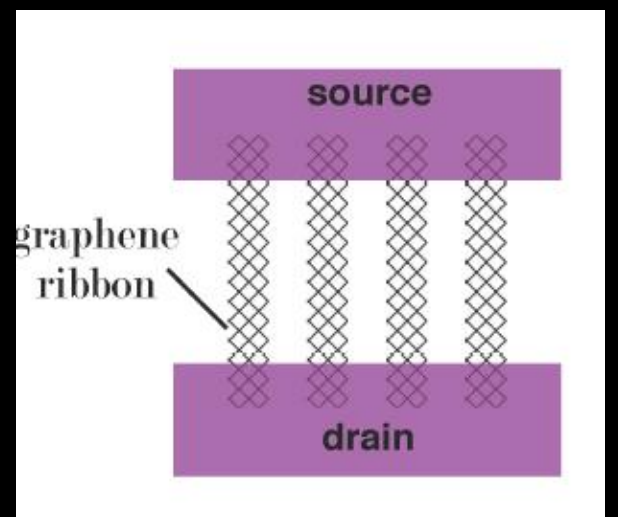
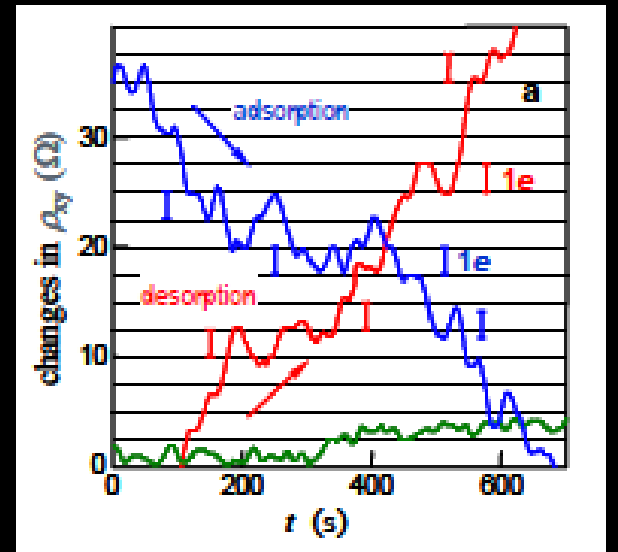


# GRAPHENE FIELD EFFECT TRANSISTOR (G-FET)

- The graphene nobel prize was for noting noticable conductivity changes of graphene (in FET configuration) -for added or removed single electrons.

*Geim et al, 2004*

- Small band gap introduced by cutting into ribbons ( $\sim .8\text{eV}/\text{width}(\text{nm})$ ), i.e. meV gap
- removing/adding electron makes FET conductive  $\rightarrow$  source/drain current.

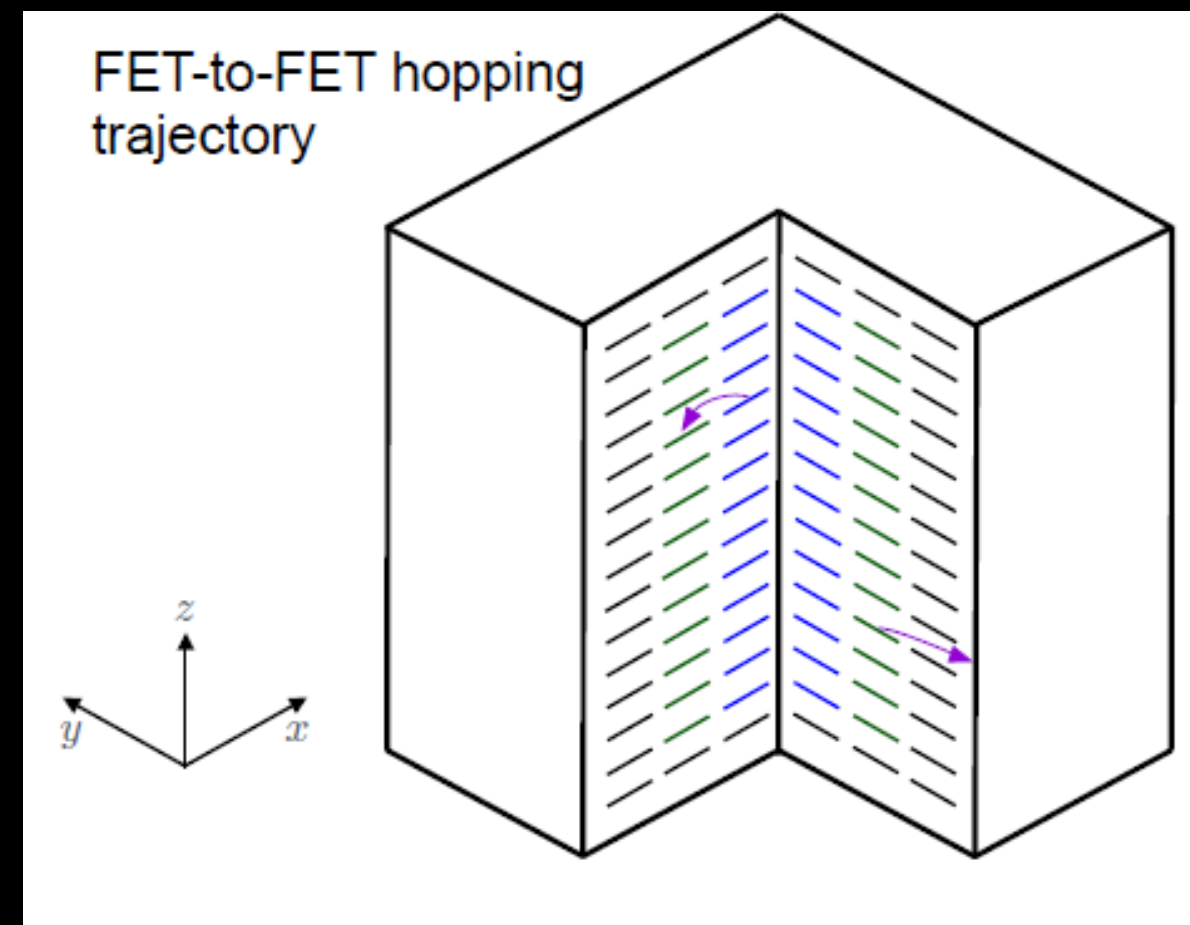
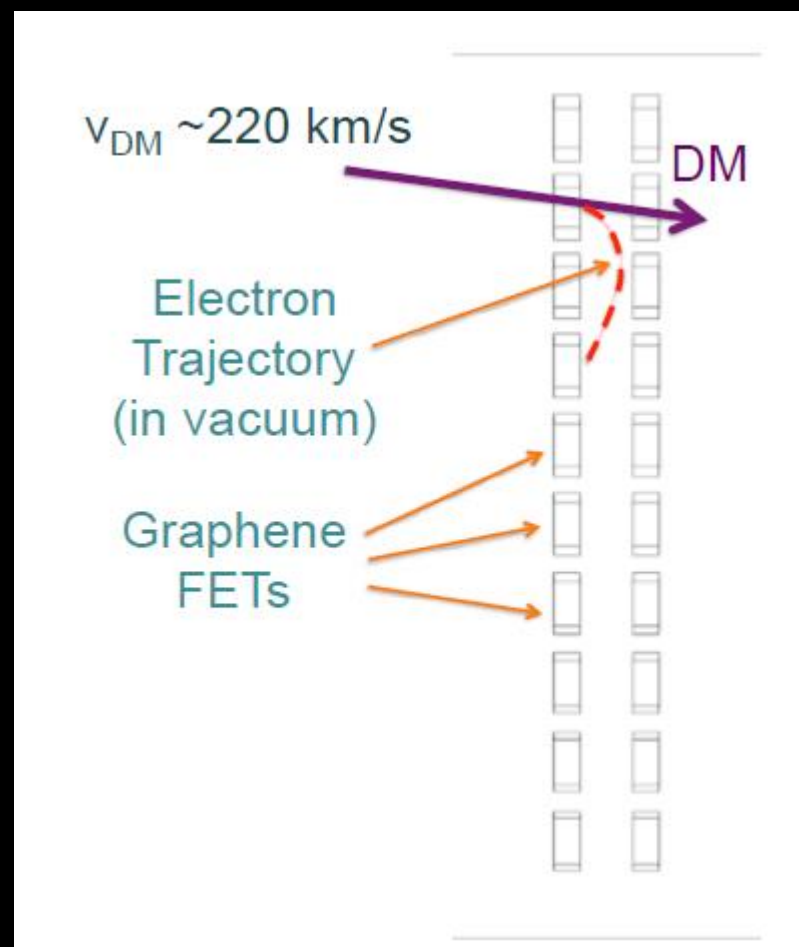




# DETECTOR CONFIGURATION

Stacked planar arrays of G-FETs

Directionality by FET-to-FET propagation



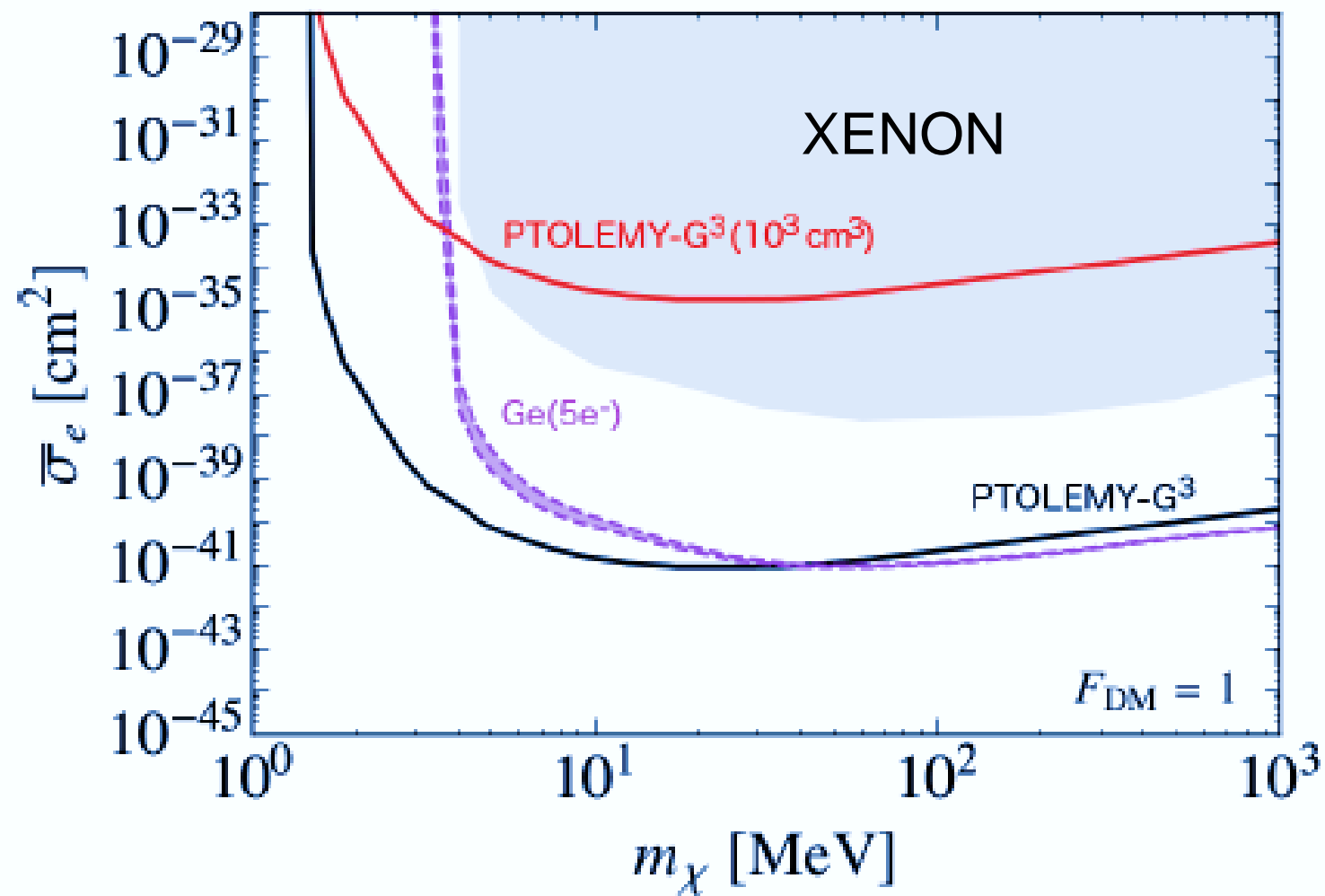
$$1\text{kg} \sim 10^{10}\text{cm}^2 \sim 10^9\text{cm}^3$$

# BACKGROUNDS

- Large surface → cosmic rays → underground (LNGS)
- Can you imagine what is the main problem?
  - source identified with low enough contamination ( $10^{-21}$ )
- High radio-purity wafer-level fabrication proven by e.g. CDMS  
*Jastram et al. NIM-A 772: 14-25*

Controllable at face value.

# REACH



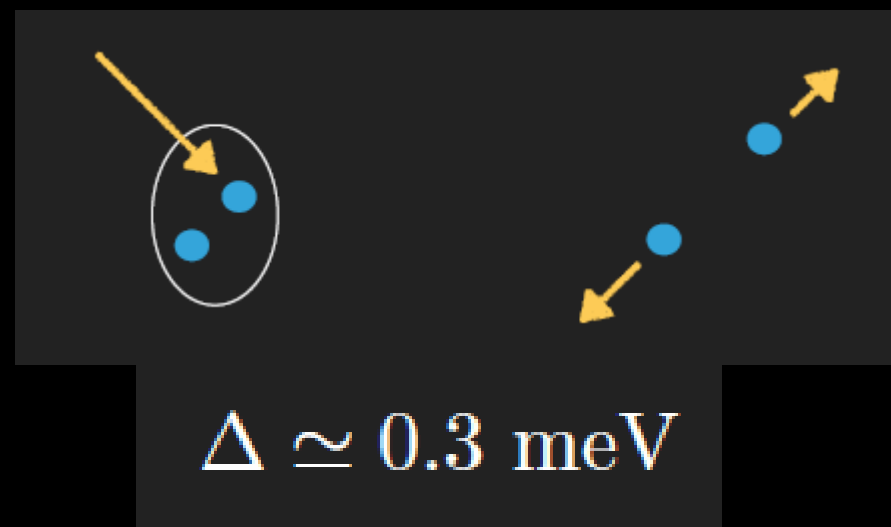
Background-free 95% C.L. sensitivity for  
graphene 1-kg-year (black) and  $10^3$  cm<sup>3</sup>  
( $10^4$  cm<sup>2</sup>) (orange)

Program to launch end of 2019

# SUPERCONDUCTING DETECTORS

*Hochberg+ JHEP 1608 (2016) 057*

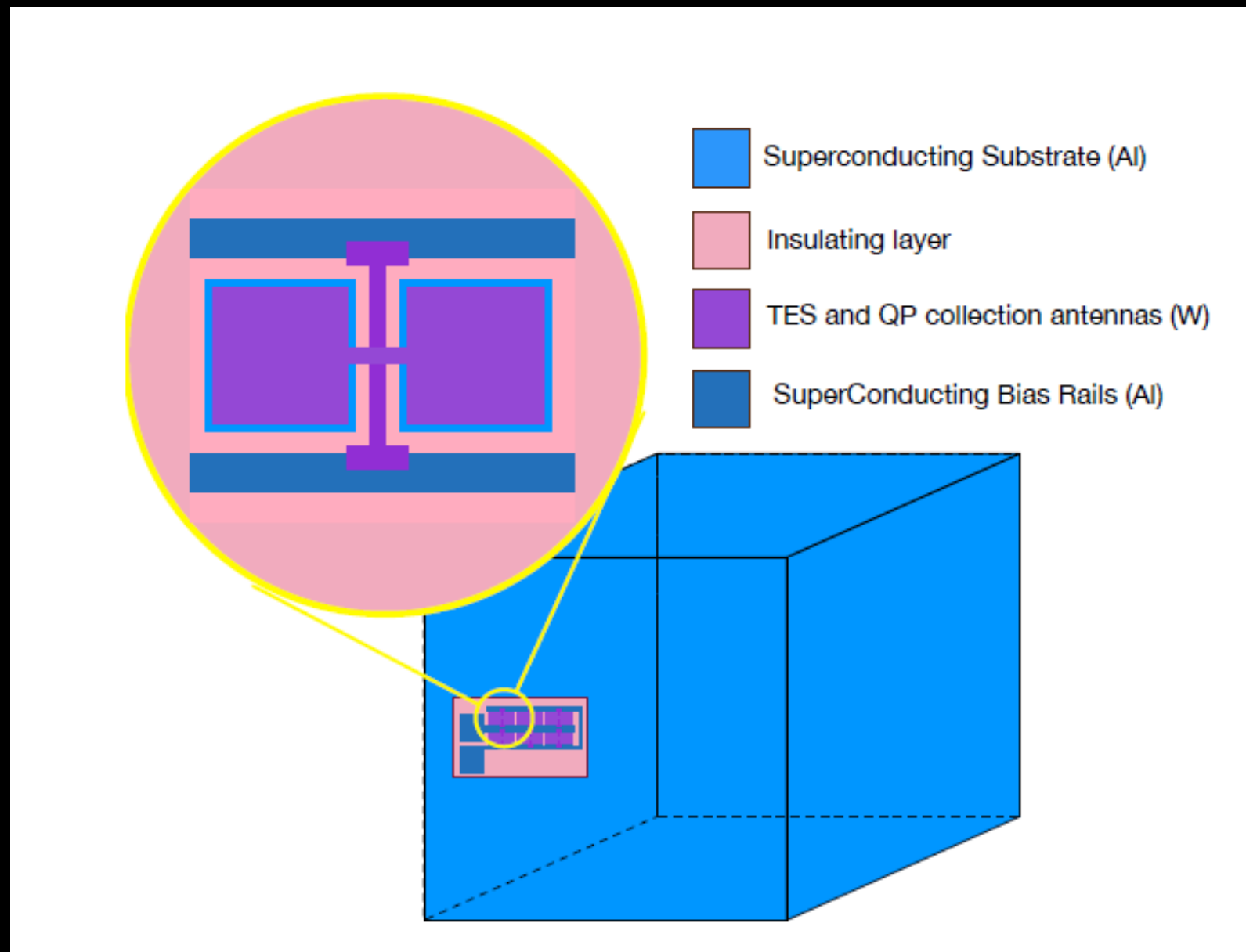
- Simple idea:
- "band-gap" becomes binding energy of Cooper pairs



- interaction produces long-lived ( $\sim \mu\text{s}$ ) quasi-particles and phonons

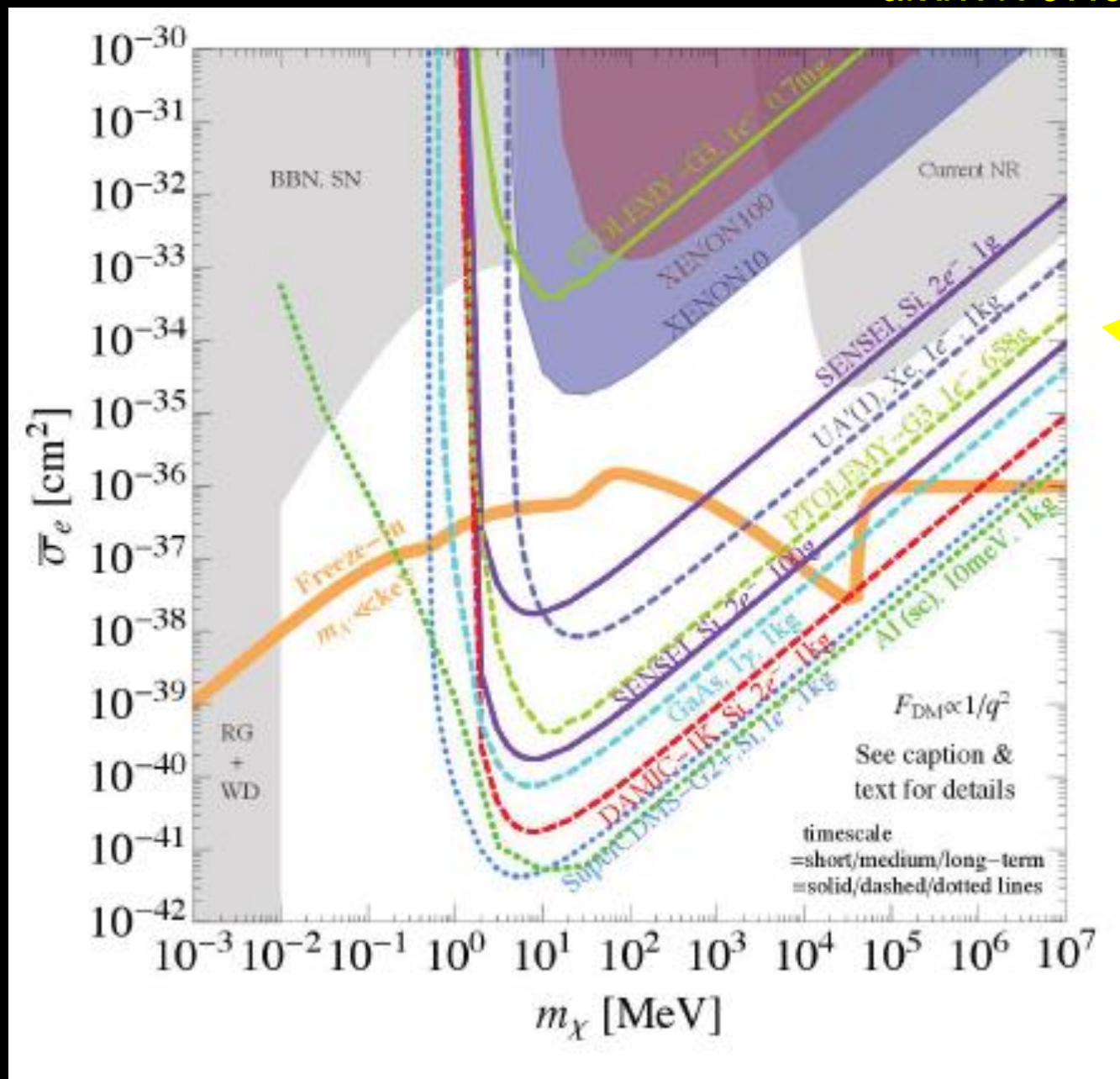
# SUPERCONDUCTING DETECTORS

*Hochberg+ JHEP 1608 (2016) 057*



# REACH

Battaglieri et al:  
arxiv:1707.04591



Why is  
this not  
better?

Light dark photon mediator, not good for  
accelerators.



# AFTER THE FUTURE:QM4DM

- Advances in solid state physics allow to quantum-engineer materials

- Detect small energy deposition

*Hochberg et al*

- Directional sensitivity

- Optical properties

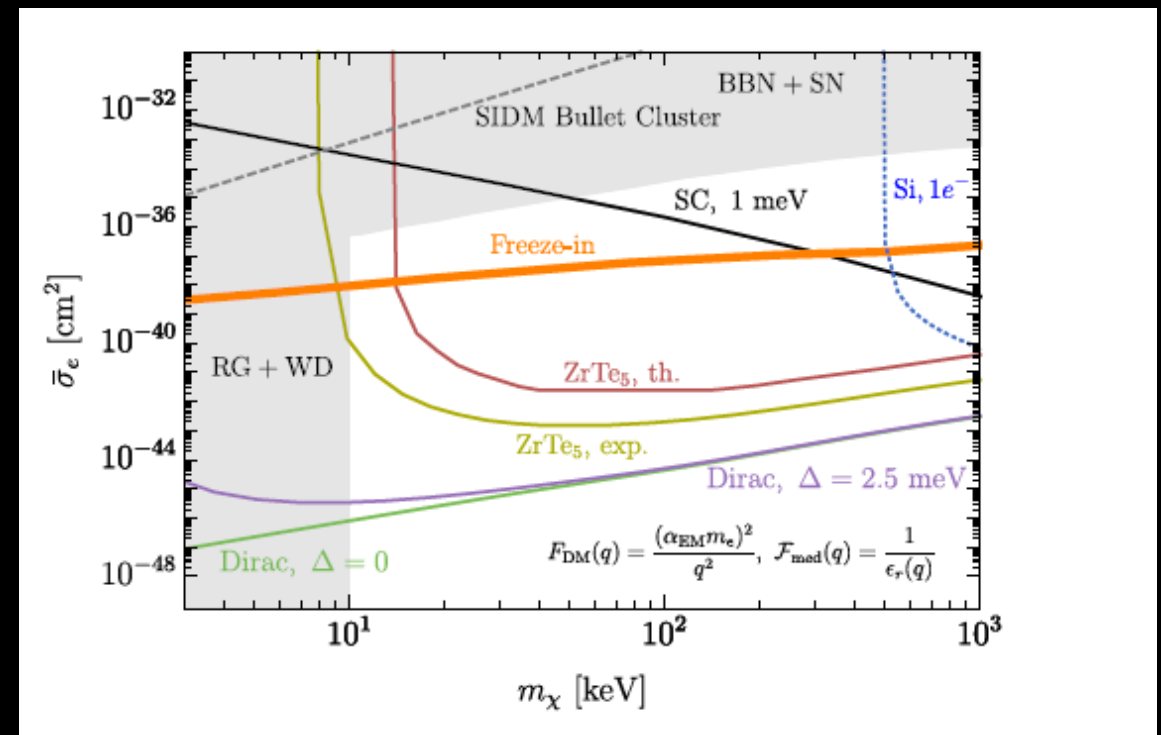
- Magnetic properties ( AXIONS)

- Boundary conditions:

- Affordability(graphene is at 1 USD/cm<sup>2</sup>)

- Chemical proerties, safety etc,etc,etc

- Example: Dirac material (c.f. graphene) with tunable meV band gap and optical properties



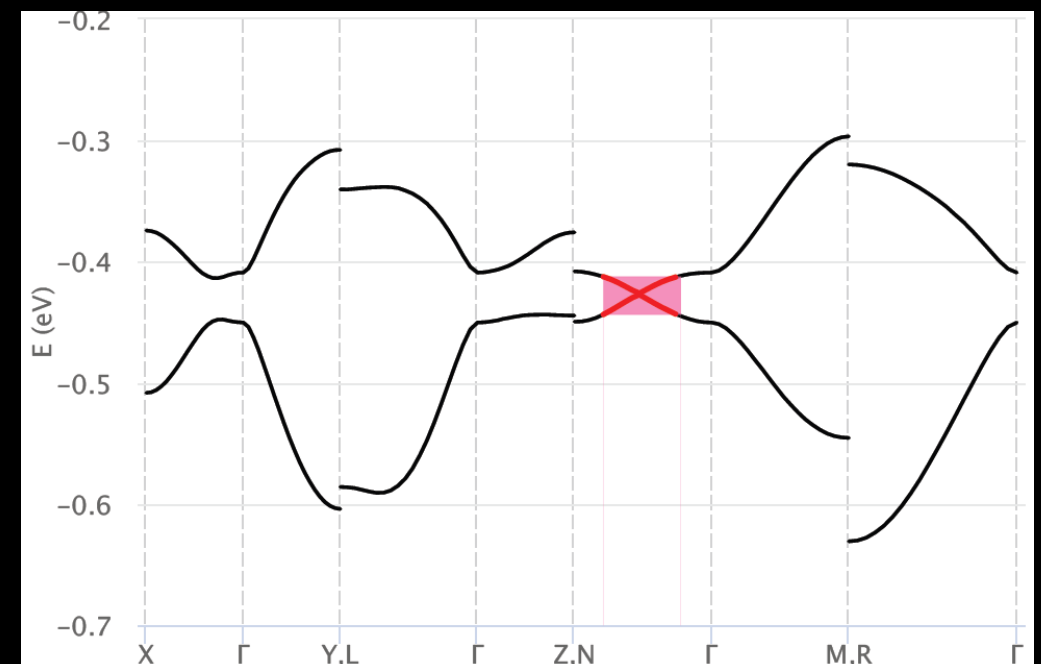
# A STRAW MAN EXAMPLE

- Consider target wish: small band gap
- Can we systematically look for a material that has this property?

*Geilhufe, Oolsthorn, Kahlhoefer, Ferella, Conrad, Balatsky*  
*Physica Status Solidi RRL* 2018, 1800293

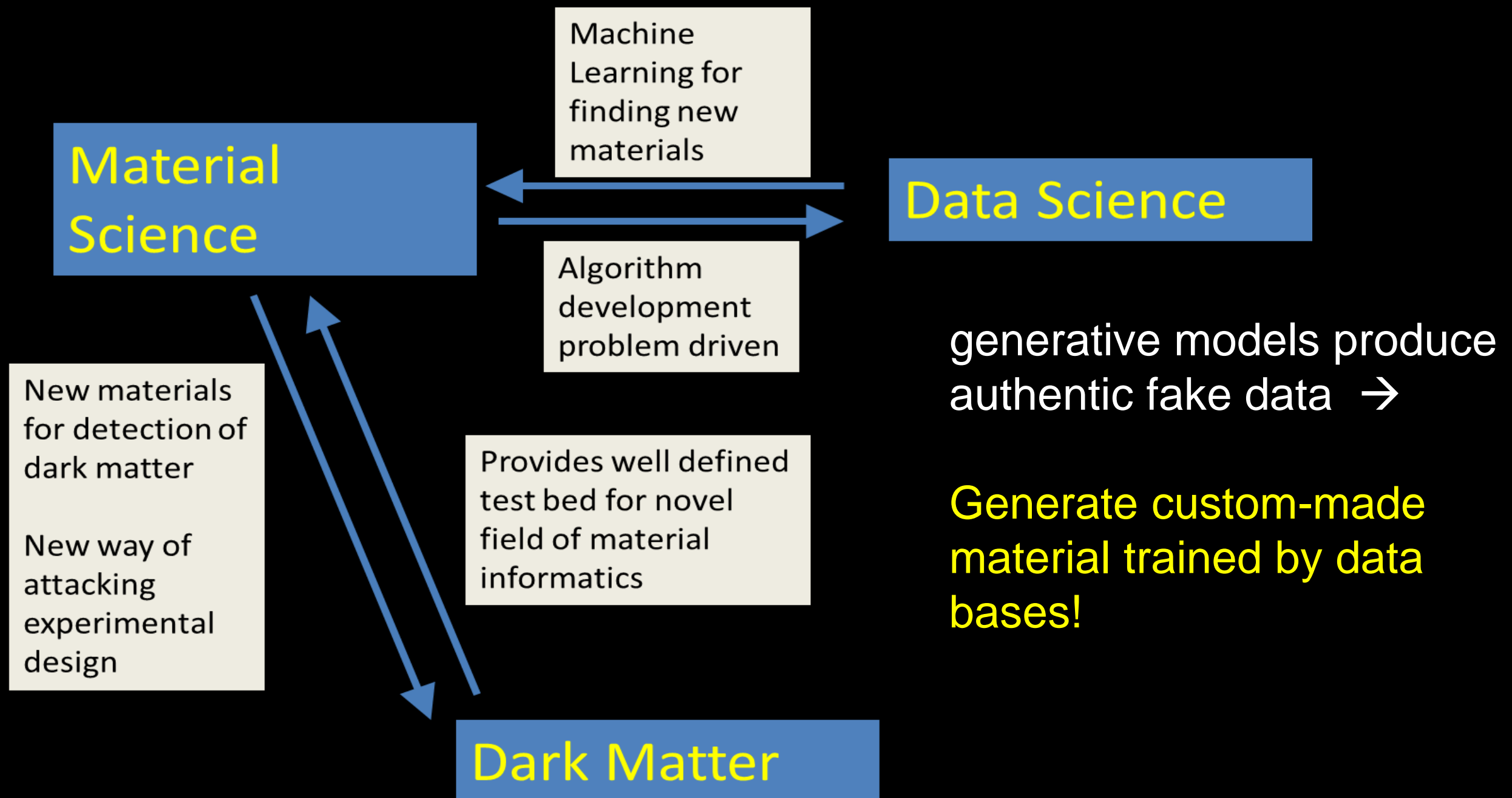
# ORGANIC MATERIALS DATA BASE

- Free to use:  
[omdb.diracmaterials.org](http://omdb.diracmaterials.org)
- band structures, density of states  
calculated by density functional  
theory
- $\approx 27.000$  Materials stored
- 3D organic crystals
- Advanced data mining tools:  
found  $\sim 50$  meV band gap



# QM4DM

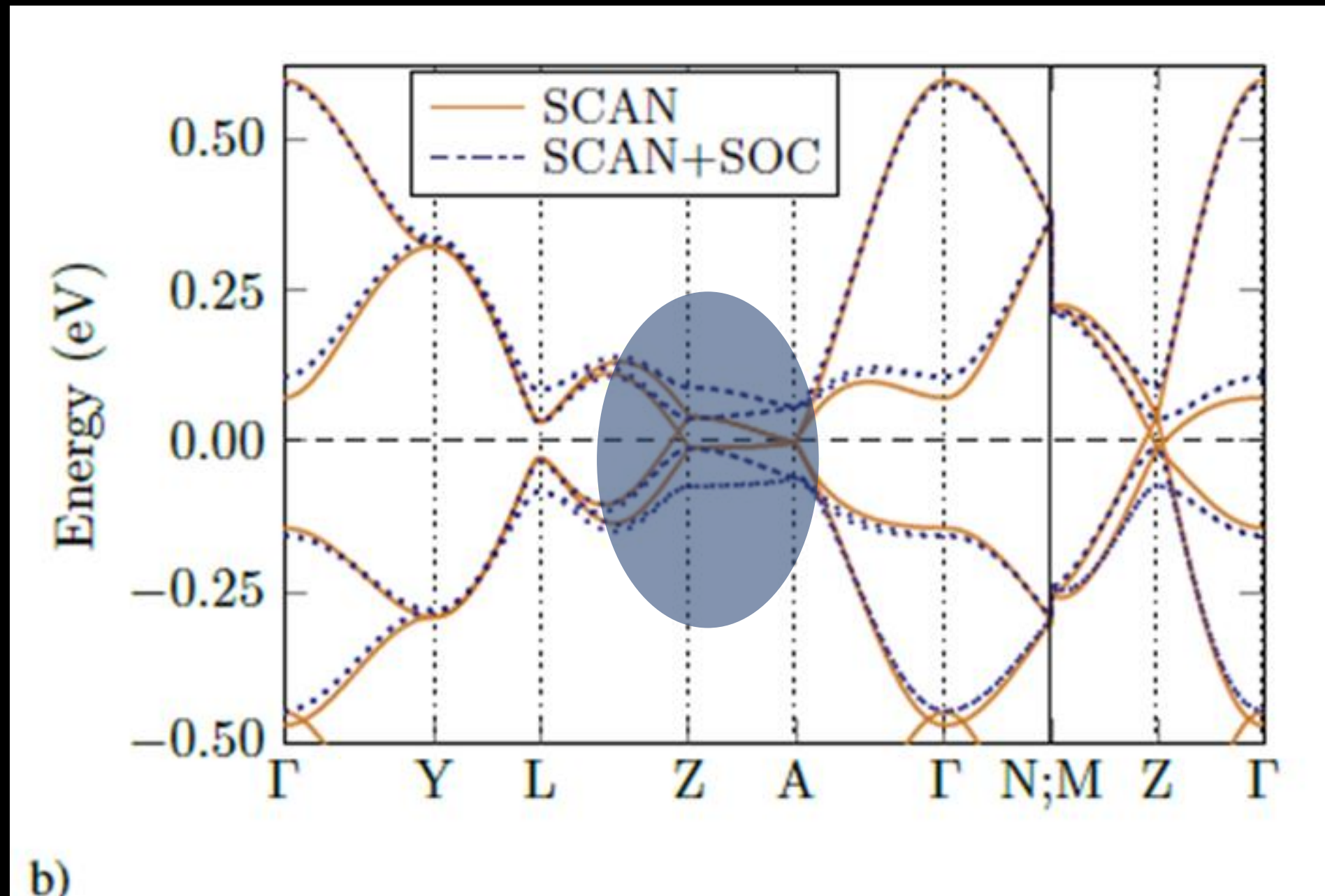
Collaboration with  
Alexander Balatsky (NORDITA)



# CONCLUSIONS

- New parameter space in direct detection is started to be explored → dark sector dark matter in the sub-GeV range → technology is the drive
- A dozen or so experiments realized within the next five to ten years → we don't know what is going to work!
- Exciting times leveraging recent advances in solid state physics → tailor-made target materials.

# ORGANIC DIRAC MATERIAL

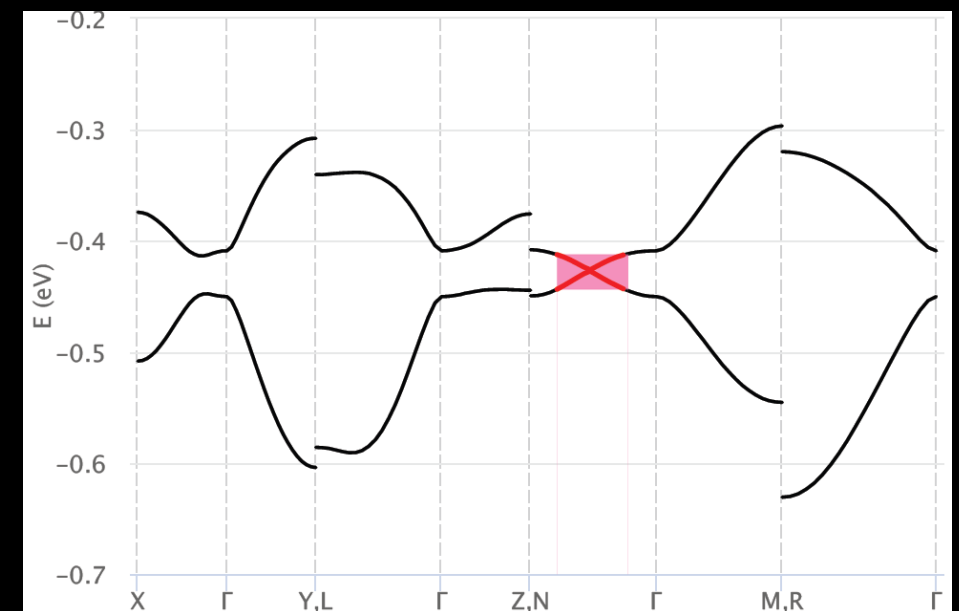
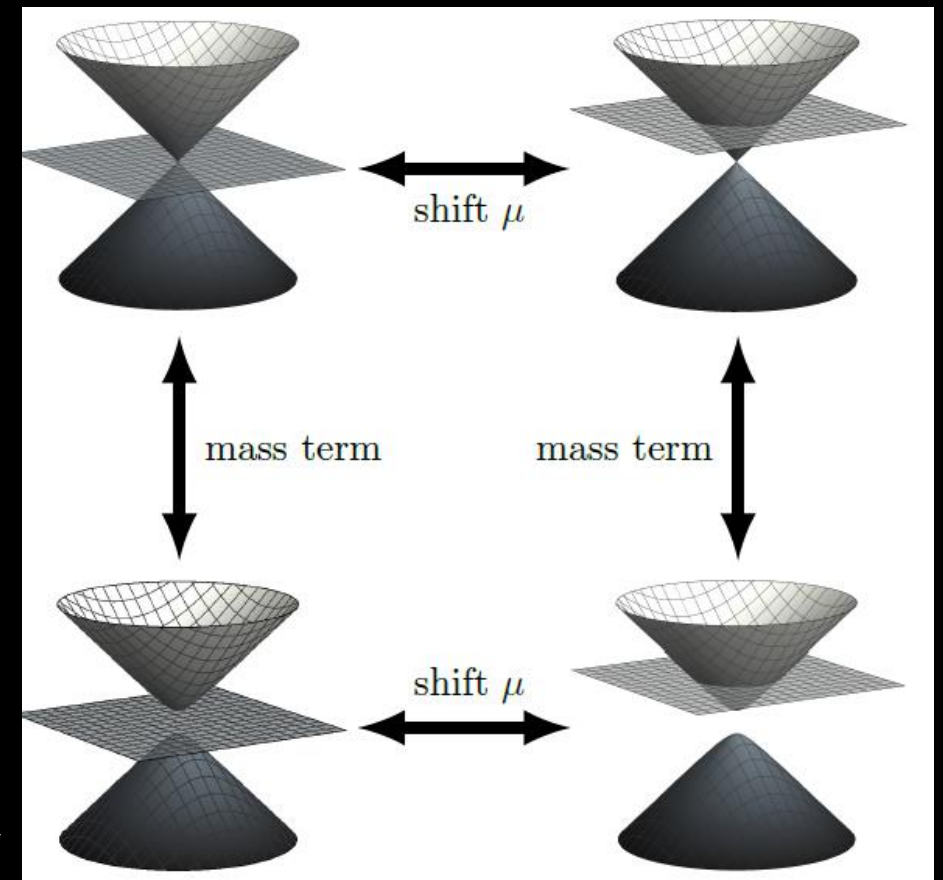


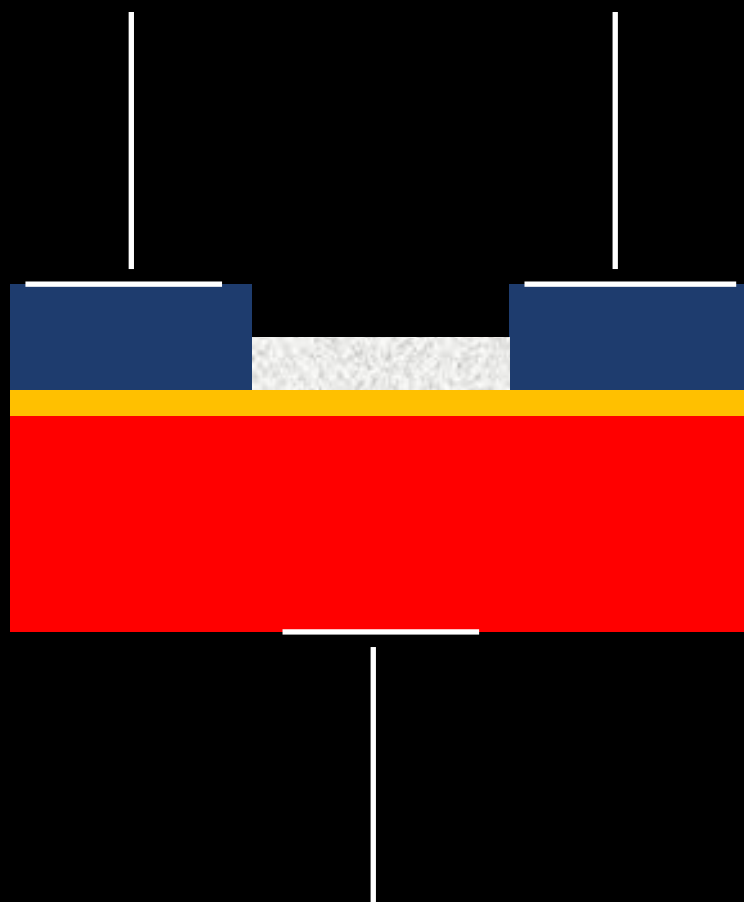
*Linear crossing gapped out by spin-orbit coupling  
Band gap  $\sim 50$  meV, much smaller gaps should  
be possible*



# EXAMPLE 1: DIRAC MATERIALS

- Dirac materials suggested for MeV DM detection (e.g. Hochberg et al Phys.Rev. D97 (2018) no.1, 015004)
- well defined energy momentum relation (Dirac Fermions)
- nodes in spectrum protected by symmetry  
→ symmetry breaking induces tiny gaps
- Examples: graphene, topological insulators, Weyl semimetals, . . .
- OMDB: pattern matching for linear crossing or reference density of states





# THE THEORETICAL FRAMEWORK

- Scalar DM, dark photon (massive, so better  $Z'$ ) mediator (production: freeze-out)
- Dirac Fermion, dark photon mediator (production: asymmetry)
- ELDER (elastic scattering via dark photon exchange)
- SIMP (Strongly Interacting Massive Particles)  $3 \rightarrow 2$  annihilations
- Majorana Fermion with velocity suppressed interactions.
- Freeze in, dark photon exchange.
- Dark photon dark matter