PBHs + Higgs stability



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ROOM: 208A, BUILDING 2A

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About myself

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▶ Born on 27.4.1987 in Hiroshima

▶ Ph.D. Physics @ Univ. of Tokyo; 3.2015, supervised by **Koichi Hamaguchi**

- ▶ Ist Postdoc @ Kavli IPMU; 4.2015-9.2017
- 2nd Postdoc @ DESY; 10.2017

My Interests

Interplay between Particle Physics & Cosmology

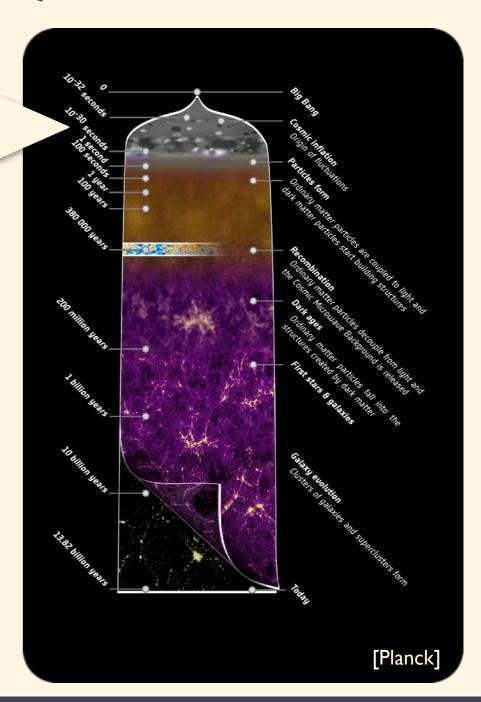
Dynamics of Early Universe from the viewpoint of Quantum Fields under extreme conditions.

Mostly before BBN... t < Is

- Inflation, (P) Reheating, Thermalization
- Dark Matter production
- Baryon asymmetry of Universe
- Phase transition, classical lumps

- Primordial Black Holes (PBHs)
- Higgs stability

Revisit the old ideas... Suppose that BSM is far away from us.



Primordial Black Holes

What is "Primordial" Black Hole?

- ▶ Black Hole formed **BEFORE** any astrophysical objects exist (even in **RD** era!).
- → Need special mechanisms for Gravity > Pressure. (i.e., need large density perturb ~ 0.5)

Why is it interesting?

- ▶ Non-particle candidate of DM
- Candidate of gravitational waves observed by **LIGO**
- ▶ Can constrain **other** DM candidates even if PBHs are subdominant.
- **)**

My interests → Implication on Particle Physics

- PBHs from Inflation: concrete models; Can we probe them?

 [Inomata, Kawasaki, KM, Tada, Yanagida; 1611.06130; Inomata, Kawasaki, KM, Yanagida; 1711.06129; cf) 1711.08956]
- Coexistence of other DM candidates for subdominant PBHs.
- Are there other novel mechanisms (more) closely tied with BSM?

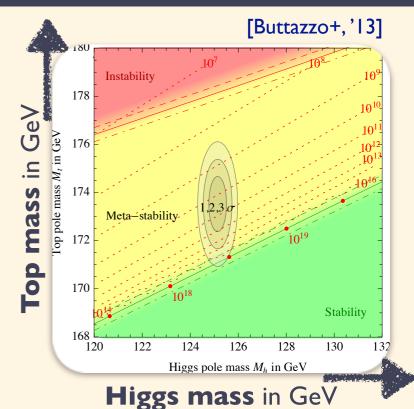
Higgs stability

Suppose new physics is far away...

- Run the SM up to, say, Planck scale.
- → Our Electroweak Vacuum tends to be **metastable**!
 - ♣ Life time @ vacuum » Current age of Universe

v.s. Cosmological environment

Can we live in the metastable vacuum? If not, we need BSM below the instability scale.



- v.s. High temperature plasma → reduce the allowed region, but still viable @ best fit.
 [e.g. Espinosa+, '08; Rose+, '15]
- ▶ v.s. Inflation (w/ inflaton mass & scale » instability scale)
- → Need a (tiny) Higgs-inflaton coupling to stabilize the Vacuum during inflation. [Westphal+; Espinosa+]
- → However, this very coupling may **destabilize it after inflation!** (via preheating)

[Ema, **KM**, Nakayama; 1602.00483, 1706.08920]

- Are there other interesting cosmological environment?
 - Evaporating Black Holes? [KM, Yamada; 1706.04523]
 - What happens in a more generic non-equil. dist., say from the end of inflation to thermalization?