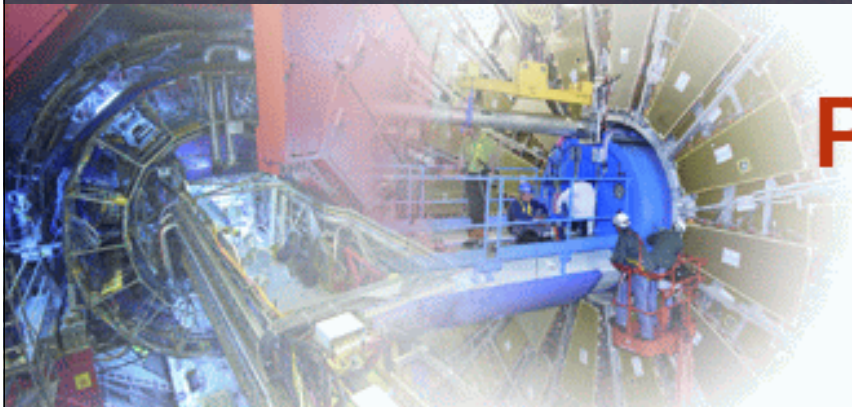


Theory of Beyond the Standard Model Physics

Hitoshi Murayama (IPMU Tokyo & Berkeley)
Physics at the LHC 2010, DESY, Jun 8, 2010

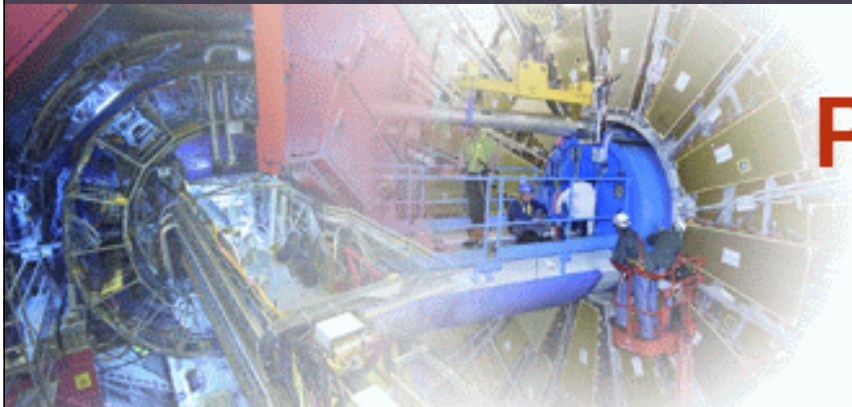


PHYSICS AT LHC 2010

**7 - 12 June 2010
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Theories of Beyond the Standard Model Physics

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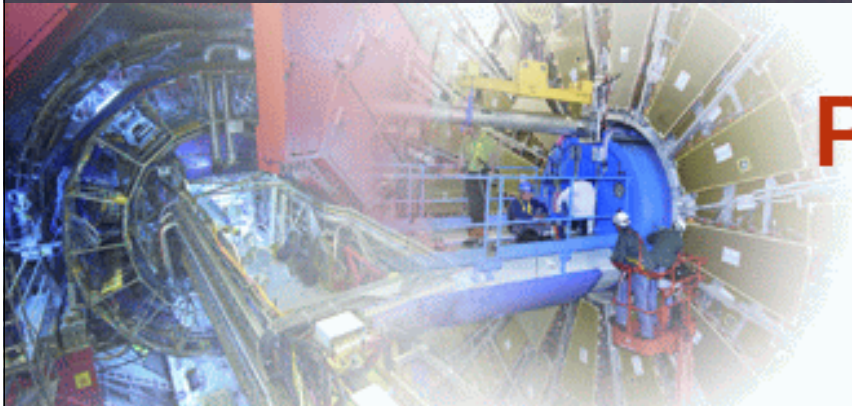


PHYSICS AT LHC 2010

**7 - 12 June 2010
DESY, Hamburg**

How stupid theorists are and Why LHC matters

Hitoshi Murayama (IPMU Tokyo & Berkeley)
Physics at the LHC 2010, DESY, Jun 8, 2010



PHYSICS AT LHC 2010

7 - 12 June 2010

DESY, Hamburg

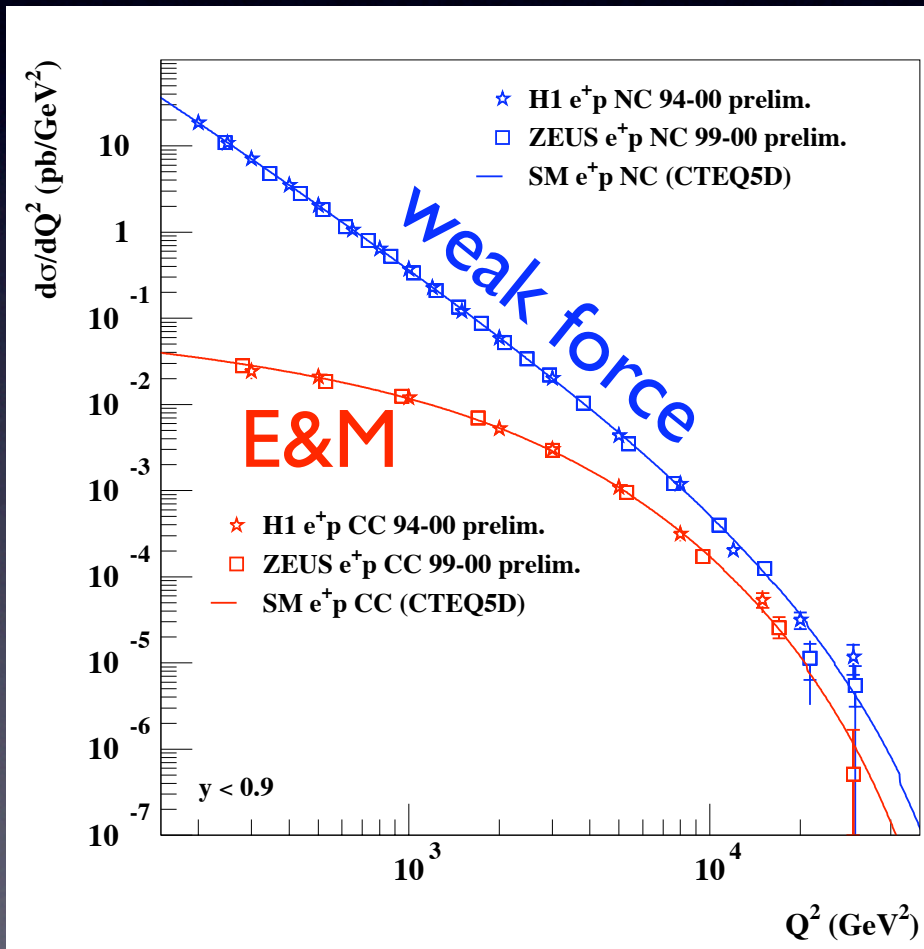
Standard Model

- A **monument** of the 20th century physics
- unifies **quantum mechanics** and **relativity** (but not GR)
- minimal particle content, renormalizable
- explains **1340 pages of Particle Data Group** with **only 19 parameters**
- tested down to 10^{-12} for electron $g_e - 2$
- **the only missing particle is Higgs boson**

$$\begin{aligned}
 \mathcal{L} = & \\
 & -\frac{1}{4g'^4} B_{\mu\nu} B^{\mu\nu} - \frac{1}{4g^2} W_{\mu\nu}^a W^{\mu\nu a} - \frac{1}{4g_s^2} G_{\mu\nu}^a G^{\mu\nu a} \\
 & + \bar{Q}_i i \not{D} Q_i + \bar{u}_i i \not{D} u_i + \bar{d}_i i \not{D} d_i + \bar{L}_i i \not{D} L_i + \bar{e}_i i \not{D} e_i \\
 & + (Y_u^{ij} \bar{Q}_i u_j \tilde{H} + Y_d^{ij} \bar{Q}_i d_j H + Y_l^{ij} \bar{L}_i e_j H + c.c.) \\
 & - \lambda (H^\dagger H)^2 + \lambda v^2 H^\dagger H + \frac{\theta}{64\pi^2} \epsilon^{\mu\nu\rho\sigma} G_{\mu\nu}^a G_{\rho\sigma}^a
 \end{aligned}$$

an imminent unification

HERA ep collider



- Unification of electromagnetic and weak forces
- ⇒ electroweak theory
- Long-term goal since '60s
- We are getting there!
- The main missing link: Higgs boson
- Then aren't we done??

Why BSM?

Once upon a time, there was a hierarchy problem...

- At the end of 19th century: a “crisis” about electron
 - Like charges repel: hard to keep electric charge in a small pack
 - Electron is point-like
 - At least smaller than 10^{-17}cm

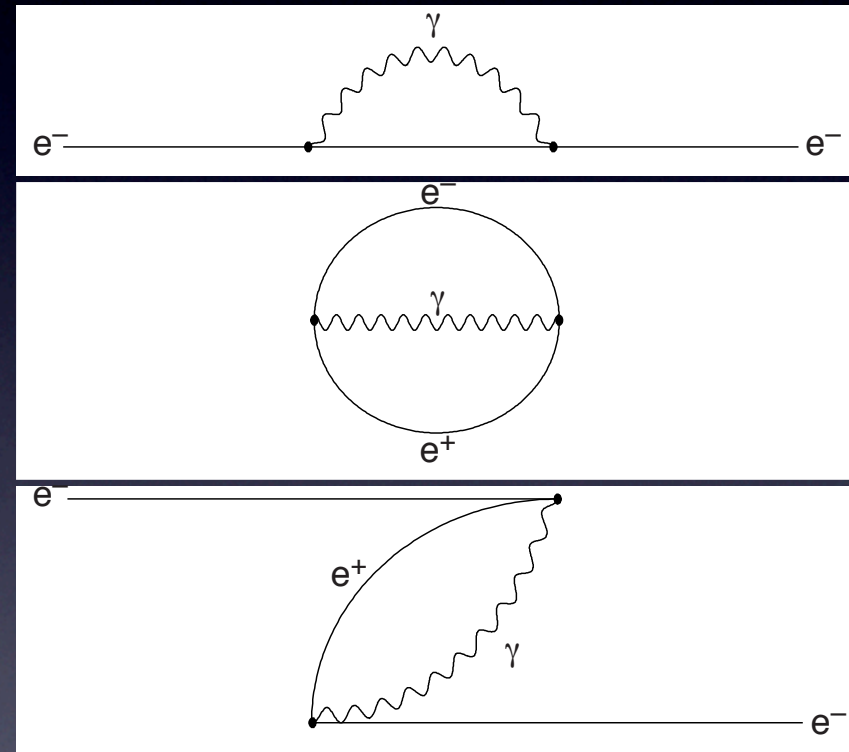
- **Need a lot of energy to keep it small!**

$$\Delta m_e c^2 \sim \frac{e^2}{r_e} \sim \text{GeV} \frac{10^{-17}\text{cm}}{r_e}$$

- Correction $\Delta m_e c^2 > m_e c^2$ for $r_e < 10^{-13}\text{cm}$
- Breakdown of theory of electromagnetism
⇒ **Can't discuss physics below 10^{-13}cm**

Anti-Matter Comes to Rescue by Doubling of #Particles

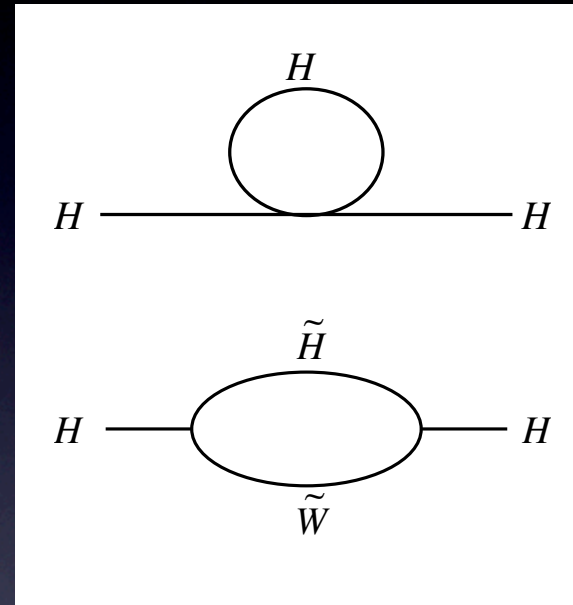
- Electron creates a force to repel itself
- Vacuum bubble of matter anti-matter creation/annihilation
- Electron annihilates the positron in the bubble
⇒ only 10% of mass even
for Planck-size $r_e \sim 10^{-33}$ cm



$$\Delta m_e \sim m_e \frac{\alpha}{4\pi} \log(m_e r_e)$$

History repeats itself?

- Higgs also repels itself
- Double #particles again
⇒ superpartners
- “Vacuum bubbles” of superpartners cancel the energy required to contain Higgs boson in itself
- Standard Model made consistent with whatever physics at shorter distances



$$\Delta m_H^2 \sim \frac{\alpha}{4\pi} m_{SUSY}^2 \log(m_H r_H)$$

Three Directions

History repeats itself

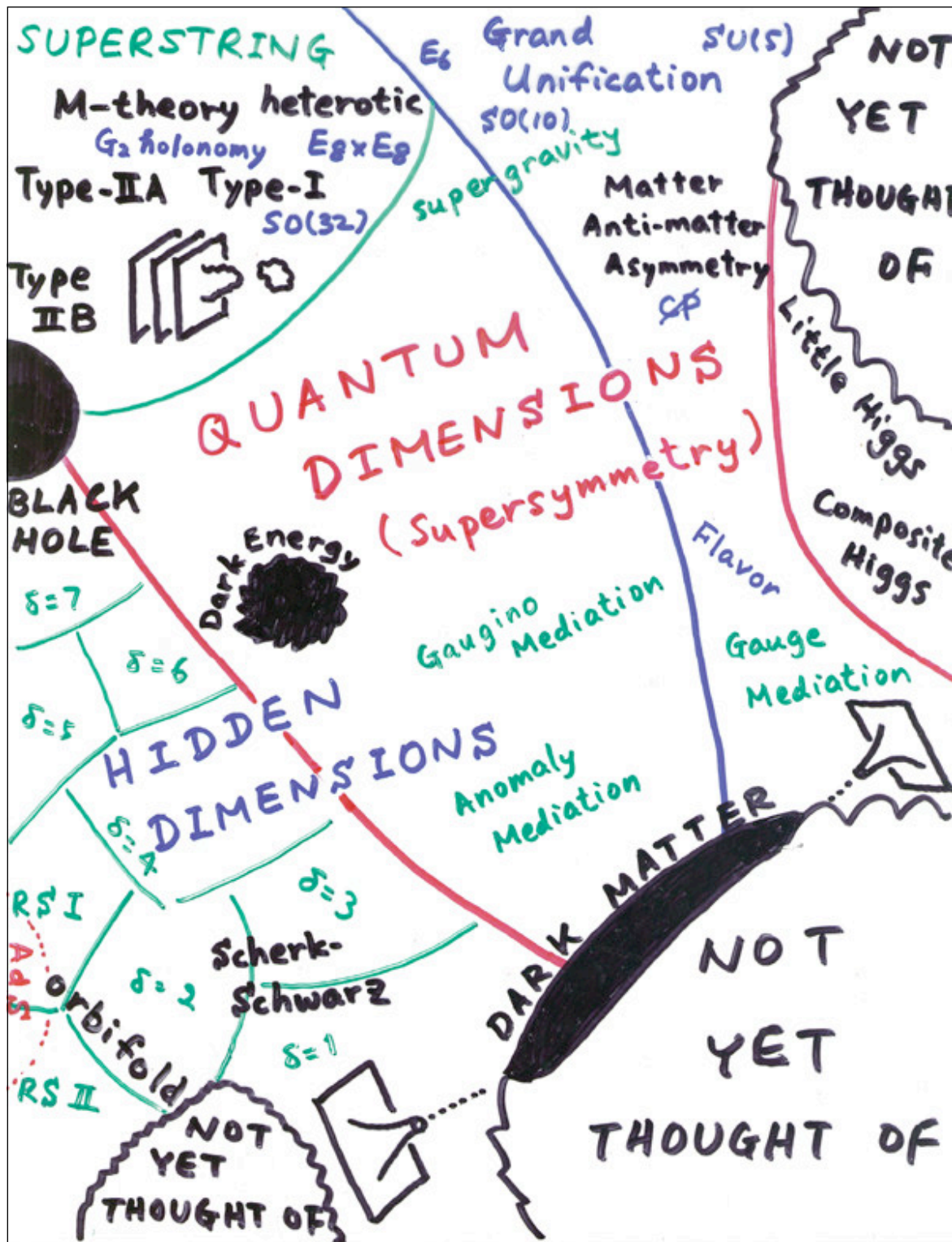
- Crisis with electron solved by anti-matter
- Double #particles again \Rightarrow supersymmetry

Learn from Cooper pairs

- Cooper pairs composite made of two electrons
- Higgs boson may be fermion-pair composite
 \Rightarrow technicolor

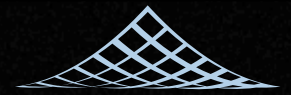
Physics as we know it ends at TeV


- Ultimate scale of physics: quantum gravity
- May have quantum gravity at TeV
 \Rightarrow hidden dimensions (0.1 mm to 10^{-17} cm)



- We really don't know what is going on at TeV
- Can we zoom in onto a point on this map?

Growing Concern among theorists



- No established deviations in
 - precision electroweak
 - flavor physics
 - LEP/Tevatron searches
- Is nature fine-tuned?
- after all, cosmological constant tuned 10^{-120}
- maybe there isn't anything beyond the Standard Model?  There definitely is!

Five empirical evidences

- Since 1998, it became clear that there are **at least five missing pieces in the SM**

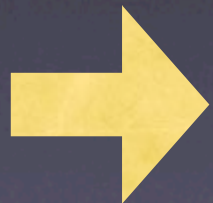


- **non-baryonic dark matter**

- **neutrino mass**

- **accelerated expansion of the Universe**

- **apparently acausal density fluctuations**



- **baryon asymmetry**

The New Minimal SM

Davoudiasl, Kitano, Li, HM

$$L_\Lambda = (2.3 \times 10^{-3} \text{ eV})^4 \quad \text{dark energy, dof}=0$$

$$L_S = \frac{1}{2} \partial_\mu S \partial^\mu S - \frac{1}{2} m_S^2 S^2 - \frac{k}{2} |H|^2 S^2 - \frac{h}{4!} S^4. \quad \text{dark matter, dof}=1$$

$$L_\varphi = \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - \frac{1}{2} m^2 \varphi^2 - \frac{\mu}{3!} \varphi^3 - \frac{\kappa}{4!} \varphi^4. \quad \text{inflation, dof}=1$$

$$L_N = \bar{N}_\alpha i \not{\partial} N_\alpha - \left(\frac{M_\alpha}{2} N_\alpha N_\alpha + h_v^{\alpha i} N_\alpha L_i \tilde{H} + c.c. \right). \quad \text{neutrino mass, dof}=4$$

$$L_{RH} = -\mu_1 \varphi |H|^2 - \mu_2 \varphi S^2 - \kappa_H \varphi^2 |H|^2 - \kappa_S \varphi^2 S^2 - (y_N^{\alpha\beta} \varphi N_\alpha N_\beta + c.c.).$$

$$L_{NMSM} = L_{MSM} + L_S + L_\Lambda + L_N + L_\varphi + L_{RH}$$

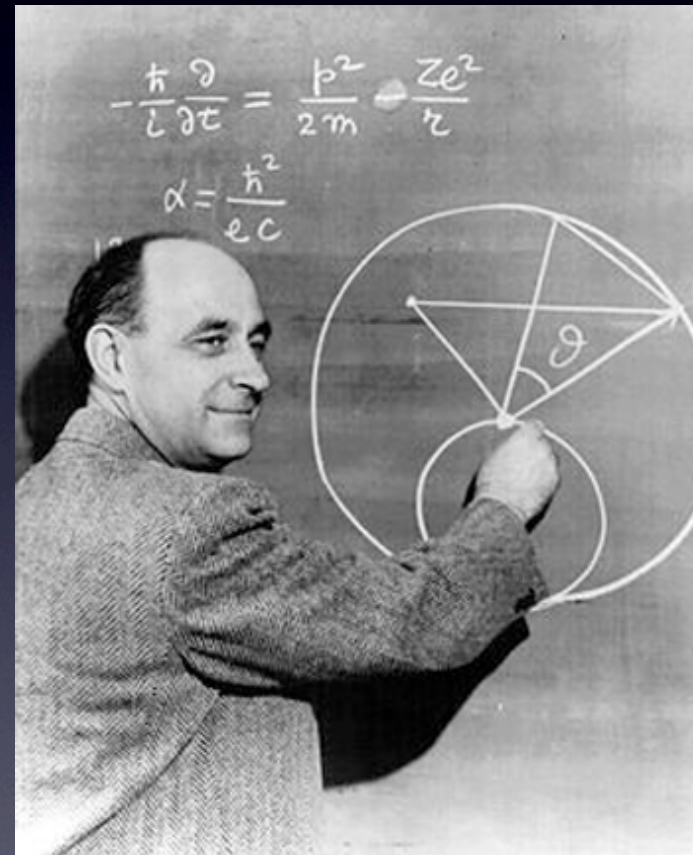
can incorporate all known established data
even small modification affects LHC data
invisible Higgs, direct detection of DM

New Era

- ~1900 reached atomic scale $10^{-8}\text{cm} \approx \alpha/m_e$
- ~1970 reached strong scale $10^{-13}\text{cm} \approx M e^{-2\pi/\alpha_s} b_0$
- ~2010 will reach weak scale $10^{-17}\text{cm} = \text{TeV}^{-1}$
- known since Fermi (1933), finally there!
- presumably it is also a derived scale
 - from SUSY breaking? extra dimensions? string theory?
- If so, **we expect rich spectrum** of new particles!
- We'll start with Higgs boson(s)

Fermi's dream era

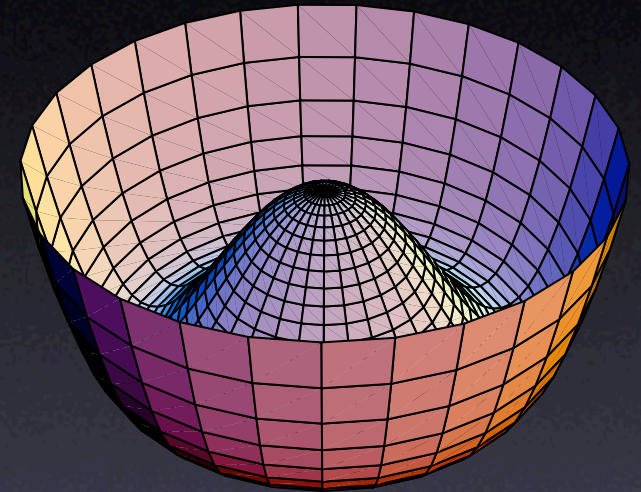
- Fermi formulated the first theory of the weak force (1932)
- *The required energy scale to study the problem known since then: $\sim \text{TeV}$*
- We are finally getting there!



Higgs
=Cosmic Superconductor

Post-Higgs Problem

- robust discovery by ATLAS/CMS
- We see “what” is condensed
- **But we still don’t know “why”**
- Two problems:
 - Why anything is condensed at all
 - Why is the scale of condensation $\sim \text{TeV} \ll M_{\text{pl}} = 10^{15} \text{TeV}$
- **Explanation most likely to be at $\sim \text{TeV}$ scale because this is the relevant energy scale**



Three Directions

Supersymmetry

- Higgs just one of *many* scalar bosons
- SUSY loops make m_h^2 negative

Higgsless/composite

- Higgs bound state of elementary fermions
- condenses because of strong attractive force

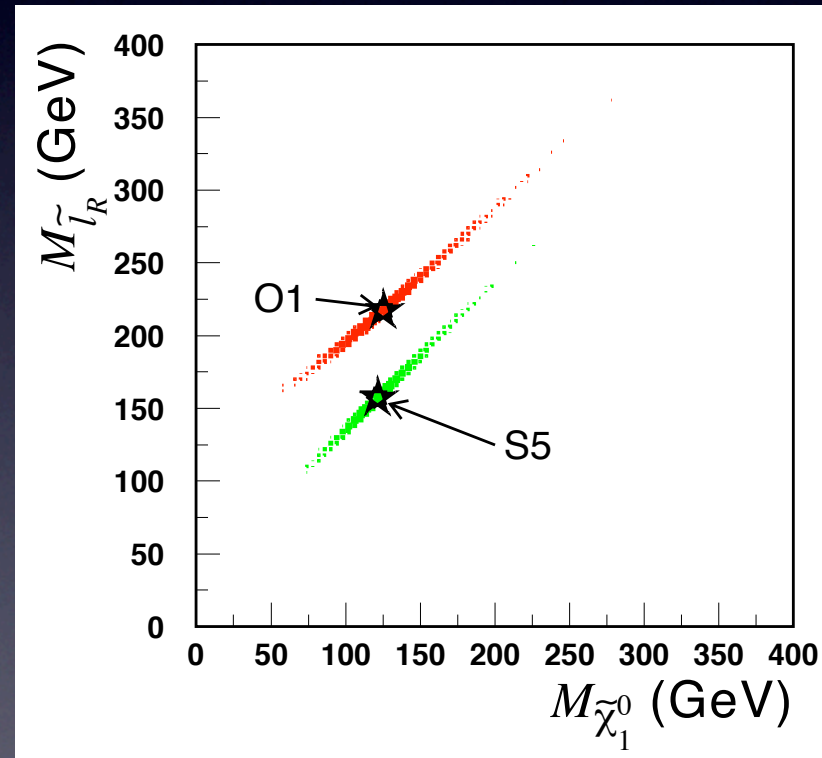
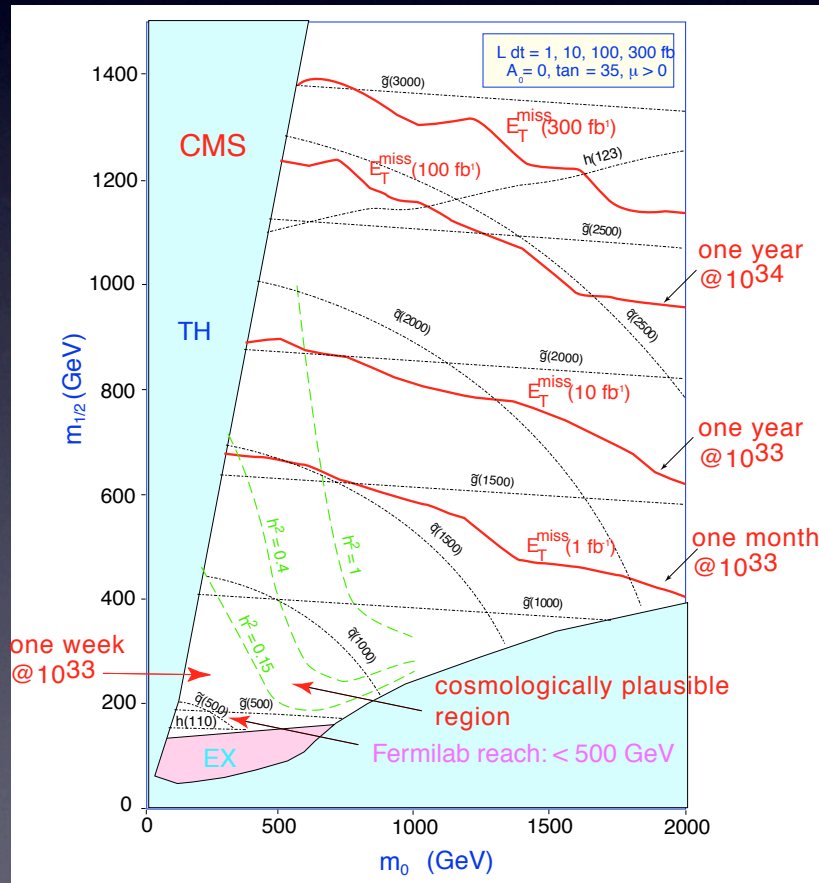
Extra dimension

- Higgs spinning in extra dimensions
- new forces from particles running in extra D

Supersymmetry

discover supersymmetry

Can do many measurements at LHC



The New York Times

July 23, 2015

The Other Half of the World Discovered

Geneva, Switzerland

As an example, supersymmetry

“New-York Times level” confidence

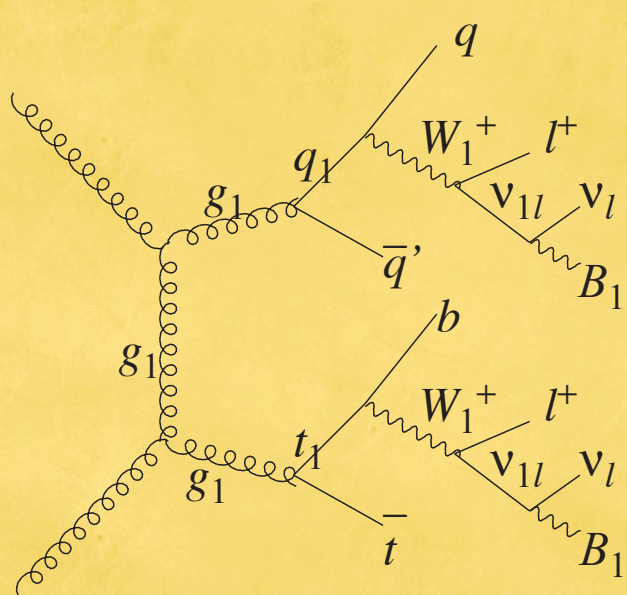
still a long way to

“Halliday-Resnick” level confidence

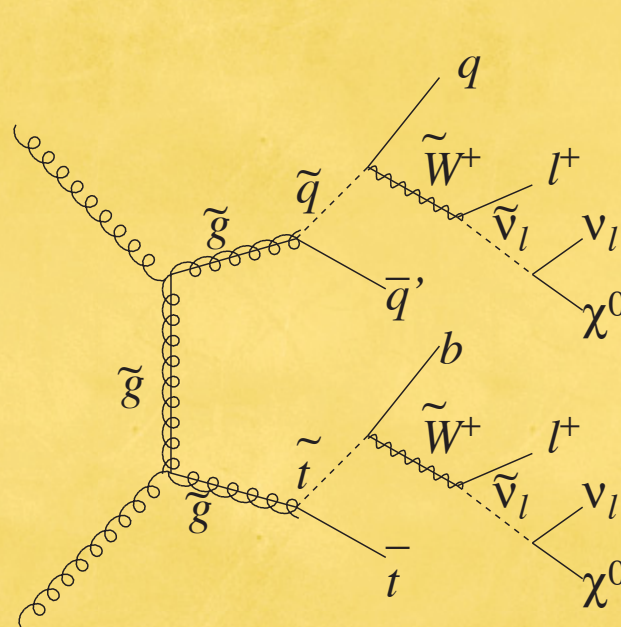
“We have learned that all particles we observe have unique partners of different spin and statistics, called superpartners, that make our theory of elementary particles valid to small distances.”

New physics looks alike

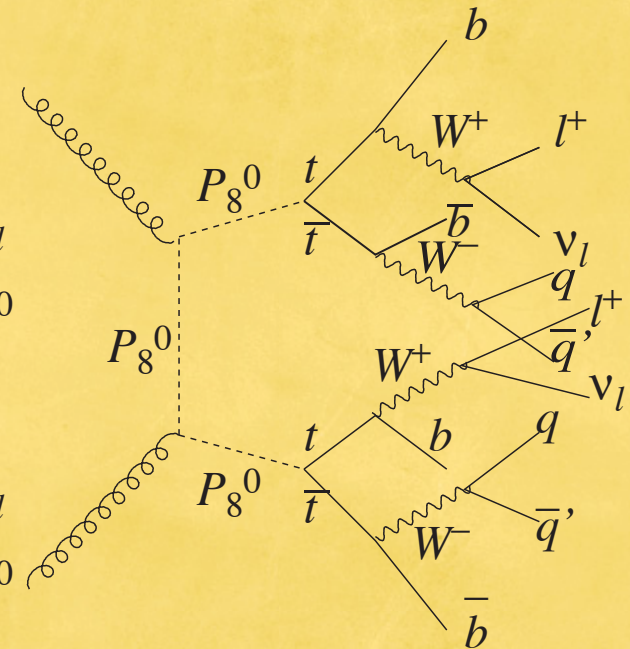
missing E_T , multiple jets, b-jets, (like-sign) leptons



UED
spin 1



SUSY
spin 1/2



technicolor
spin 0

+4th generation, little Higgs with T-parity

- spectroscopy
- precision mass, BR measurements
- need to do this *without* assuming the underlying model!

Squarks

$J=0?$

PDG 2016

The following data are averaged over all light flavors, presumably u, d, s, c with both chiralities. For flavor-tagged data, see listings for Stop and Sbottom. Most results assume minimal supergravity, an untested hypothesis with only five parameters. Alternative interpretation as extra dimensional particles is possible. See KK particle listing.

SQUARK MASS

<u>VALUE (GeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
538±10	OUR FIT		mSUGRA assumptions
532±11	¹ ABBIENDI 11D	CMS	Missing ET with mSUGRA assumptions
541±14	² ADLER 110	ATLAS	Missing ET with mSUGRA assumptions
••• We do not use the following data for averages, fits, limits, etc •••			
652±105	³ ABBIENDI 11K	CMS	extended mSUGRA with 5 more parameters

¹ABBIENDI 11D assumes minimal supergravity in the fits to the data of jets and missing energies and set $A_0=0$ and $\tan\beta = 3$. See Fig. 5 of the paper for other choices of A_0 and $\tan\beta$. The result is correlated with the gluino mass M_3 . See listing for gluino.

²ADLER 110 uses the same set of assumptions as ABBIENDI 11D, but with $\tan\beta = 5$.

³ABBIENDI 11K extends minimal supergravity by allowing for different scalar masses-squared for $H_u, H_d, 5^*$ and 10 scalars at the GUT scale.

SQUARK DECAY MODES

<u>MODE</u>	<u>BR(%)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
j+miss	32±5	ABE 10U	ATLAS	
j l+miss	73±10	ABE 10U	ATLAS	lepton universality
j e+miss	22±8	ABE 10U	ATLAS	
j μ +miss	25±7	ABE 10U	ATLAS	
q χ^+	seen	ABE 10U	ATLAS	

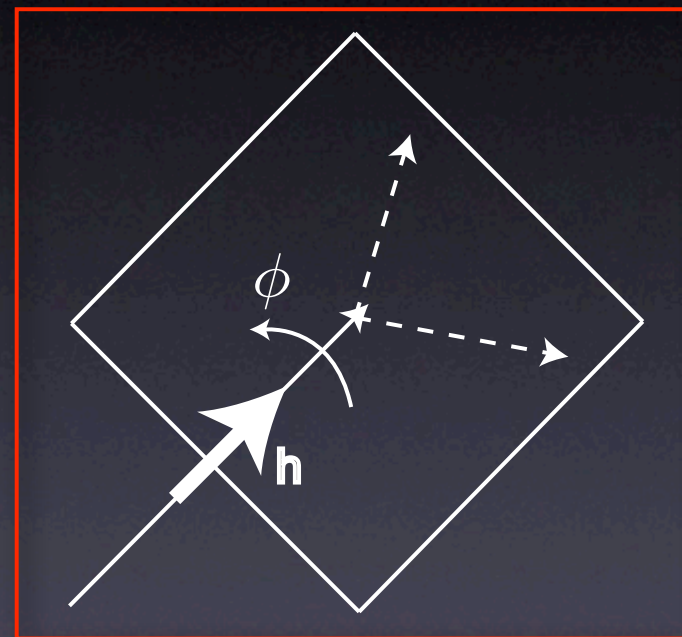
Helicity and phase

- Decay of particle with spin h along the momentum axis
- Rotations about z-axis of decay plane given by

$$\begin{aligned} \mathcal{M} &\propto e^{iJ_z \phi} \\ J_z &= \frac{(\vec{s} + \vec{x} \times \vec{p}) \cdot \vec{p}}{|\vec{p}|} \\ &= \frac{\vec{s} \cdot \vec{p}}{|\vec{p}|} = h \end{aligned}$$

- rotational invariance: a single helicity state has flat distribution in ϕ :

$$|e^{ih\phi}|^2 = 1$$



Quantum Interference among helicities

(with M. Buckley, W. Klemm, and V. Rantal)

- If particles produced in multiple helicities:

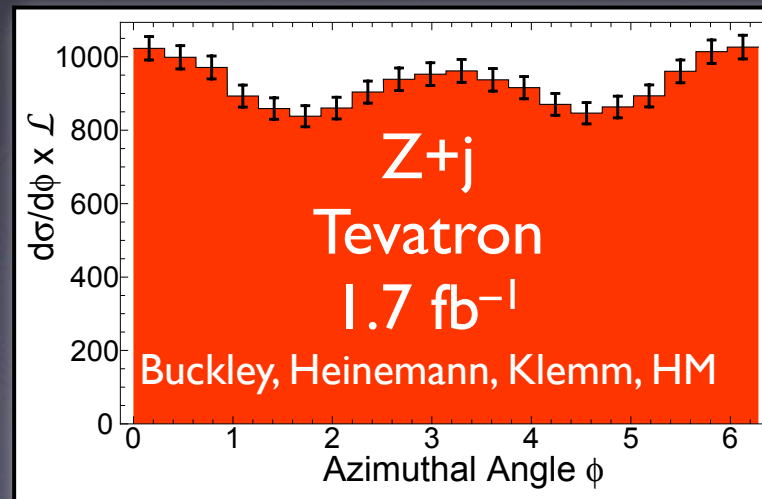
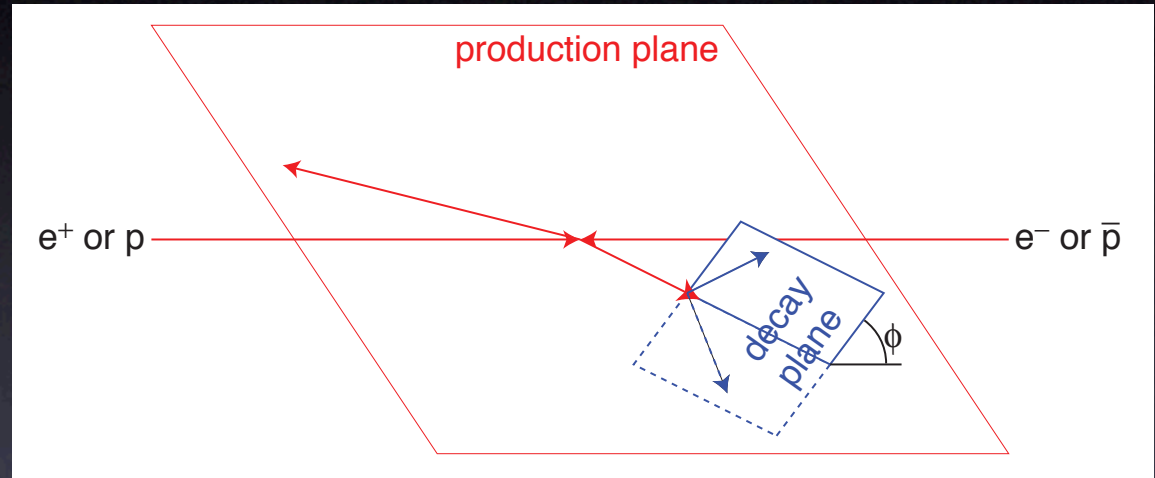
$$\sigma \propto \left| \sum \mathcal{M}_{prod.} \mathcal{M}_{decay} \right|^2$$

$$\mathcal{M}_{decay} = e^{ih\phi} \mathcal{M}_{decay}(h, \phi = 0)$$

- Different helicities interfere once they decay!
- ϕ dependence of cross section tells us what helicities contributed to the interference.
- Can measure only helicity differences (akin to neutrino oscillation)

Definition of the azimuthal angle

- Beam and produced particles span the production plane
- Parent particle and its decay products span the decay plane
- azimuth is the relative angle between two planes



should be
easy at
LHC2011!

KK graviton

(with Vikram Rentala)

- LHC: $pp \rightarrow G+j$ show $\cos(4\phi)$! Doable for TeV KK graviton with $> 100 \text{ fb}^{-1}$

$\sigma_{\text{signal}} = 230 \text{ fb}$ vs $\sigma_{\text{BG}} = 0.15 \text{ fb}$

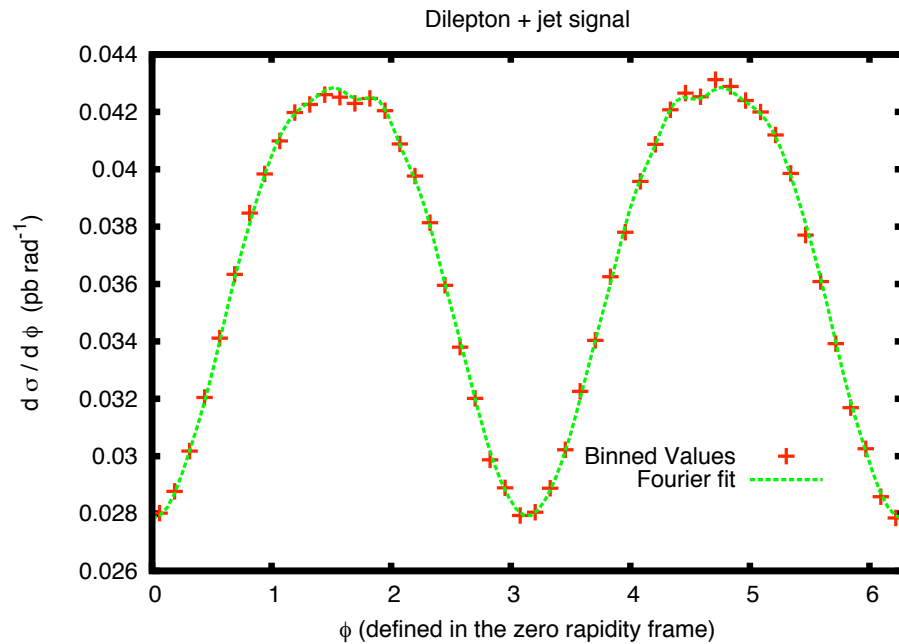


FIG. 5: Differential distribution ($\frac{d\sigma}{d\phi}$) for $m_1 = 1 \text{ TeV}$ and $c = 0.1$. A strong $\cos(2\phi)$ mode can be seen but there is also a $\cos(4\phi)$ component. The Fourier fit is shown in green.

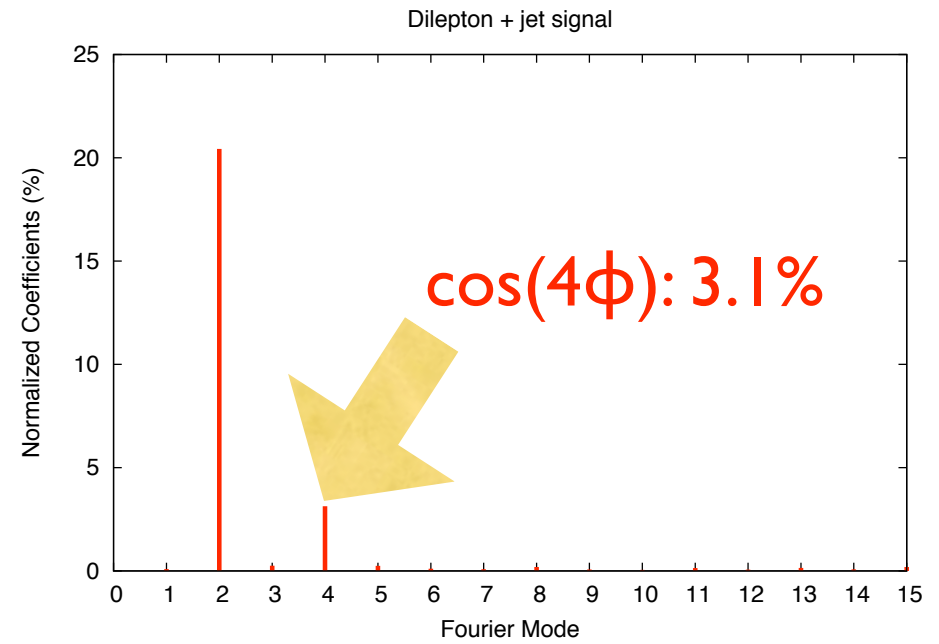
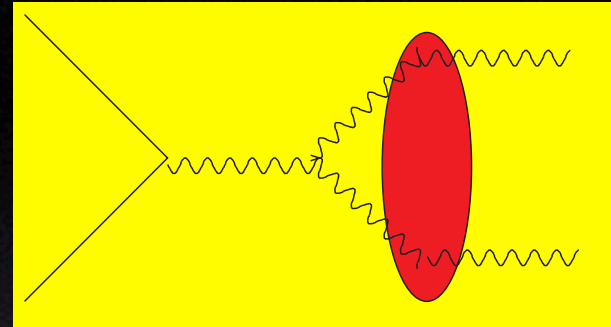
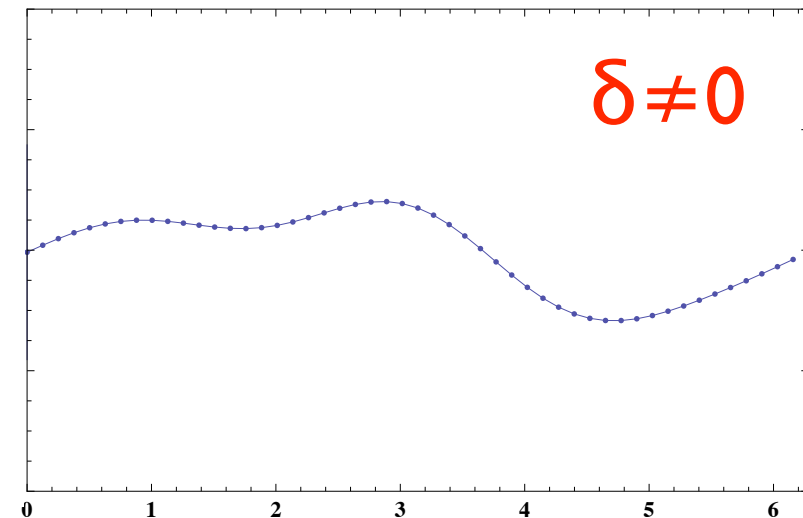
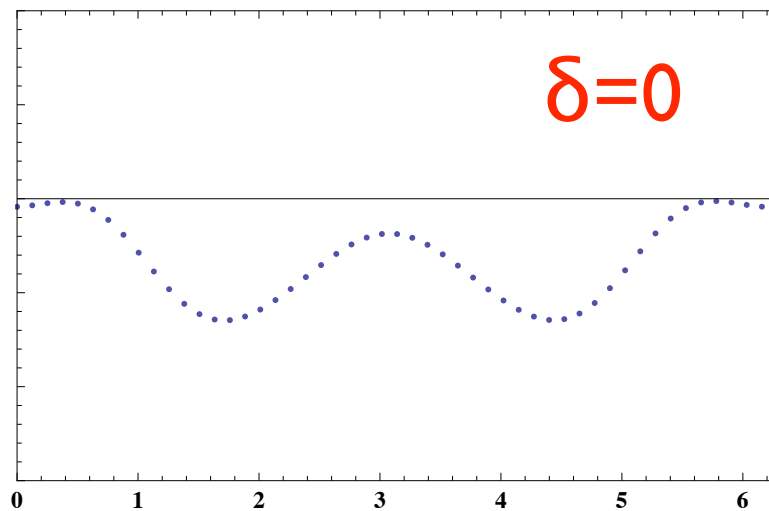


FIG. 6: Discrete Fourier Transform of the binned differential cross-section shown in Figure 5 corresponding to $N = 50$ bins. The errors are negligible and disappear in the limit of large statistics.

$$\mathcal{A}(u\bar{d} \rightarrow W_L^+ Z_L) = |\mathcal{A}_{LL}| e^{i\delta}$$



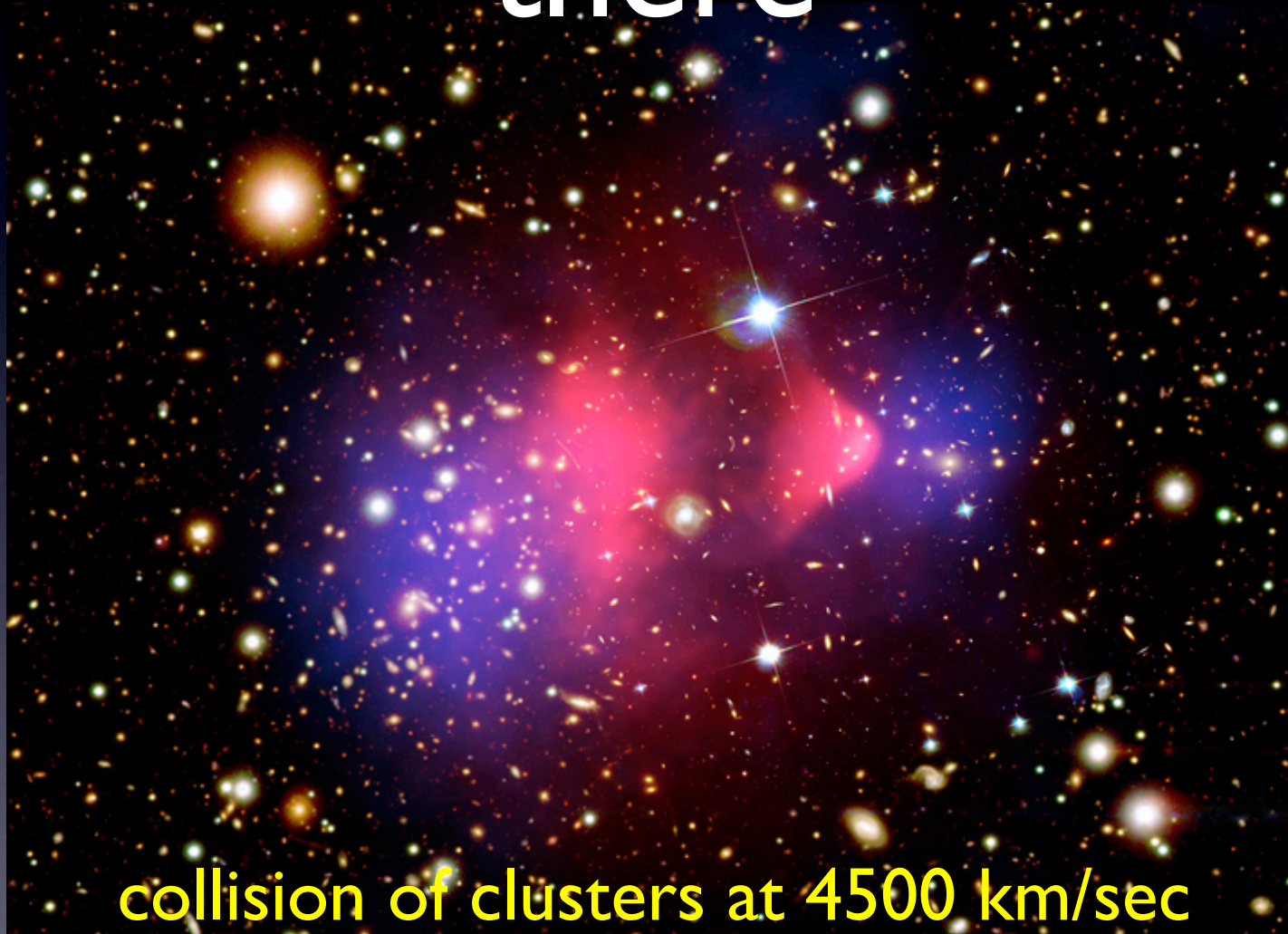
- “rescattering phase” (cf. $K \rightarrow \pi\pi$)
- interference with $W_T Z_T$ after decay



with Vikram Rentala, Jing Shu

Dark Matter

You don't want to be
there



collision of clusters at 4500 km/sec

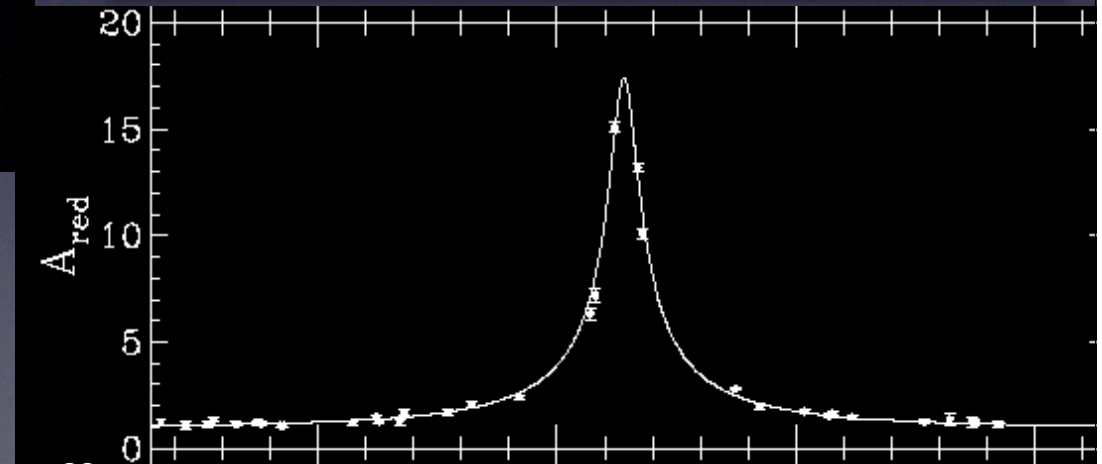
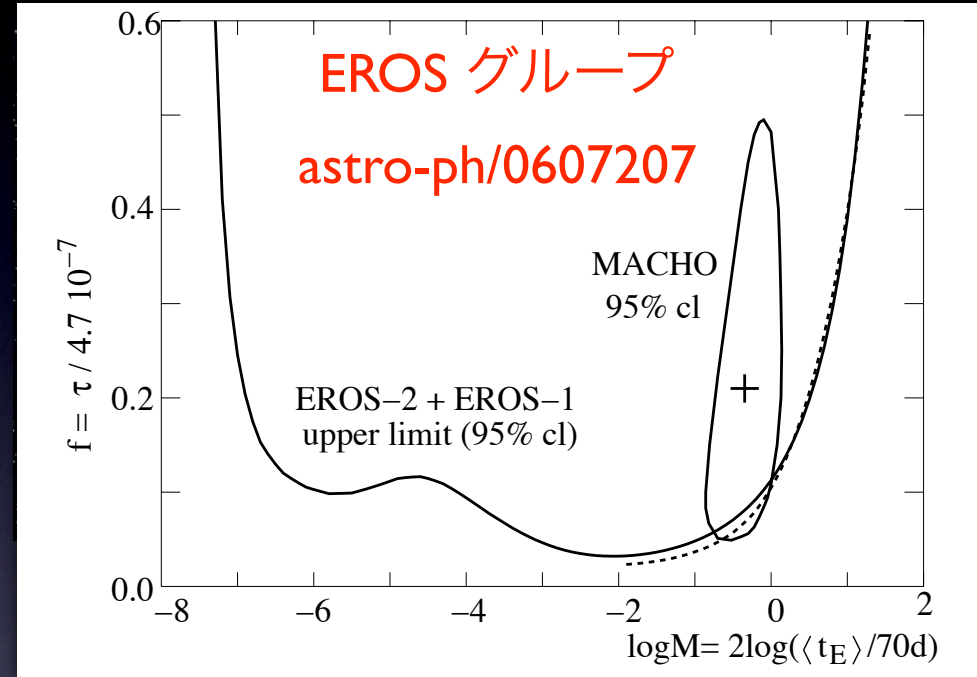


Dim Stars?

Search for **MACHOs**
(Massive Compact Halo Objects)

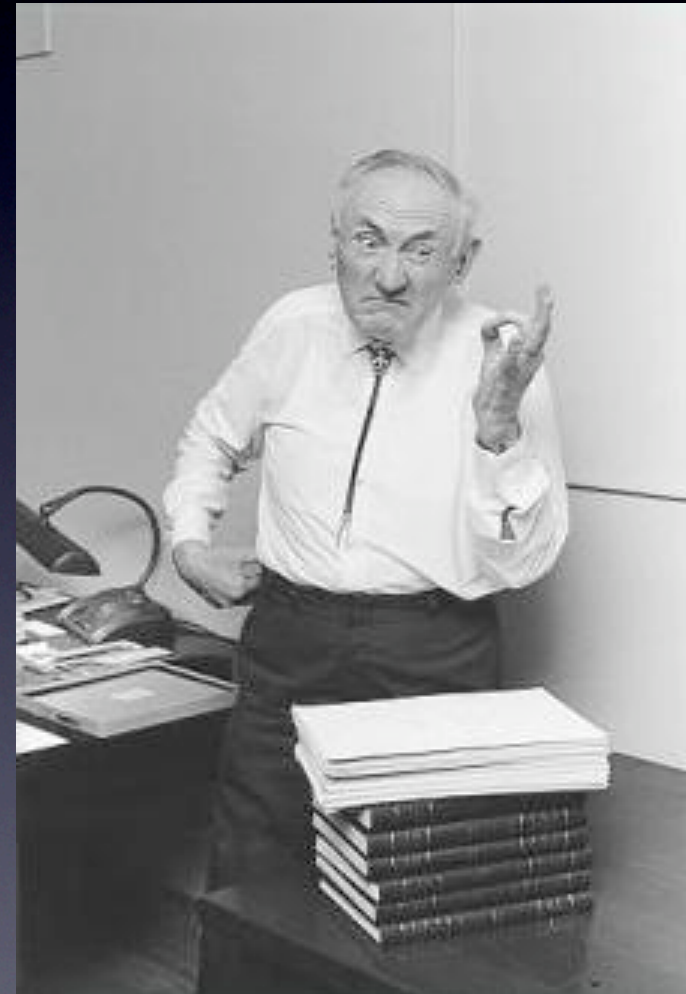


Not enough of them!

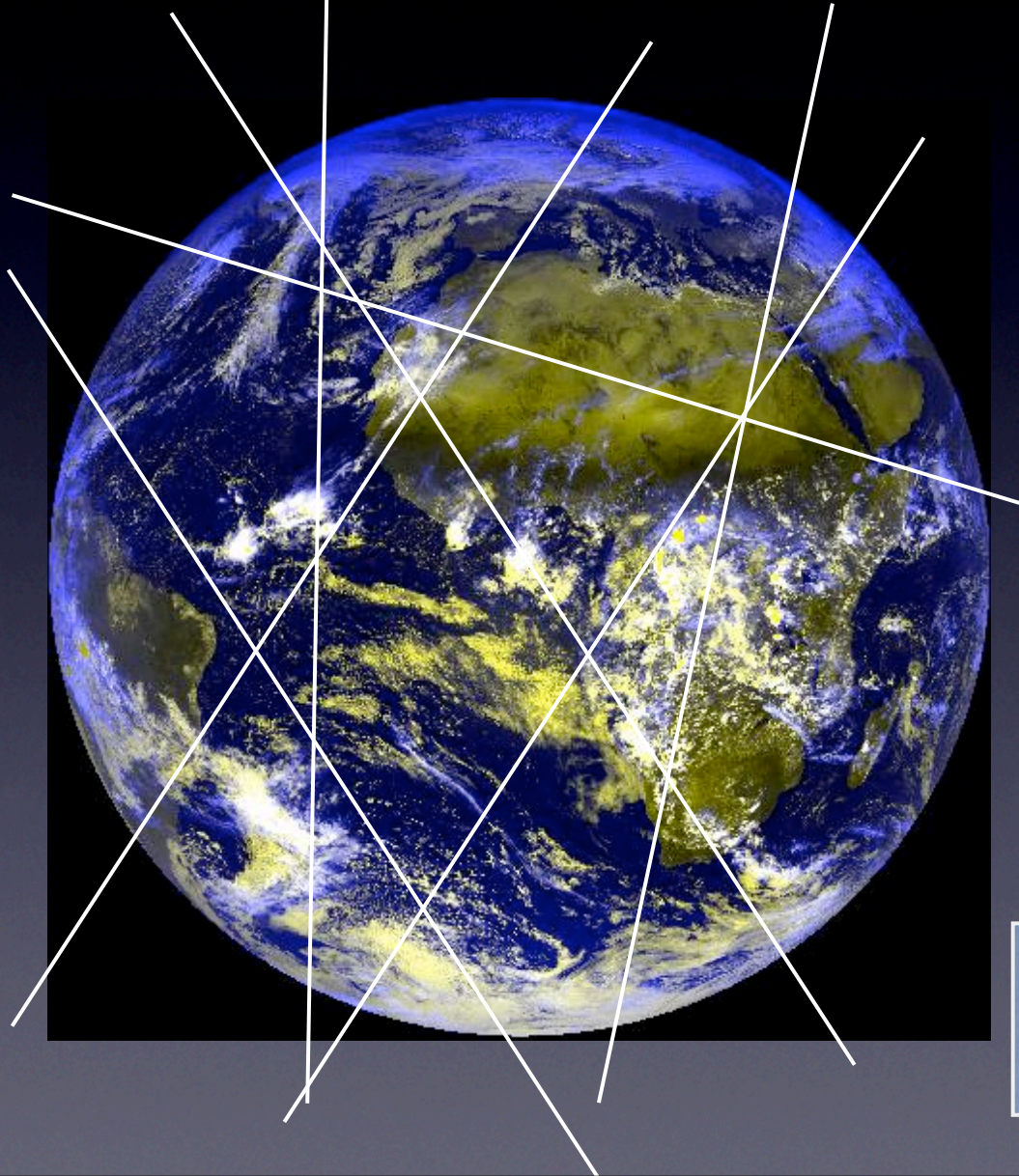


Mass Limits

- moves in $V = G_N \frac{Mm}{r} \frac{\hbar^2}{\hbar^2}$
- “Bohr radius” $r_B = \frac{\hbar^2}{G_N M m^2}$
- if too light, won't fit inside the galaxy!
- 10^{-31} GeV to 10^{50} GeV
- narrowed it down to within 81 orders of magnitude
- a big progress in 70 years since Zwicky



MACHO \Rightarrow WIMP



- Probably **WIMP** (Weakly Interacting Massive Particle)
- Stable heavy particle produced in early Universe, **left-over from near-complete annihilation**

$$\Omega_M = \frac{0.756(n+1)x_f^{n+1}}{g^{1/2}\sigma_{ann}M_{Pl}^3} \frac{3s_0}{8\pi H_0^2} \approx \frac{\alpha^2/(TeV)^2}{\sigma_{ann}}$$

invisible dark matter?

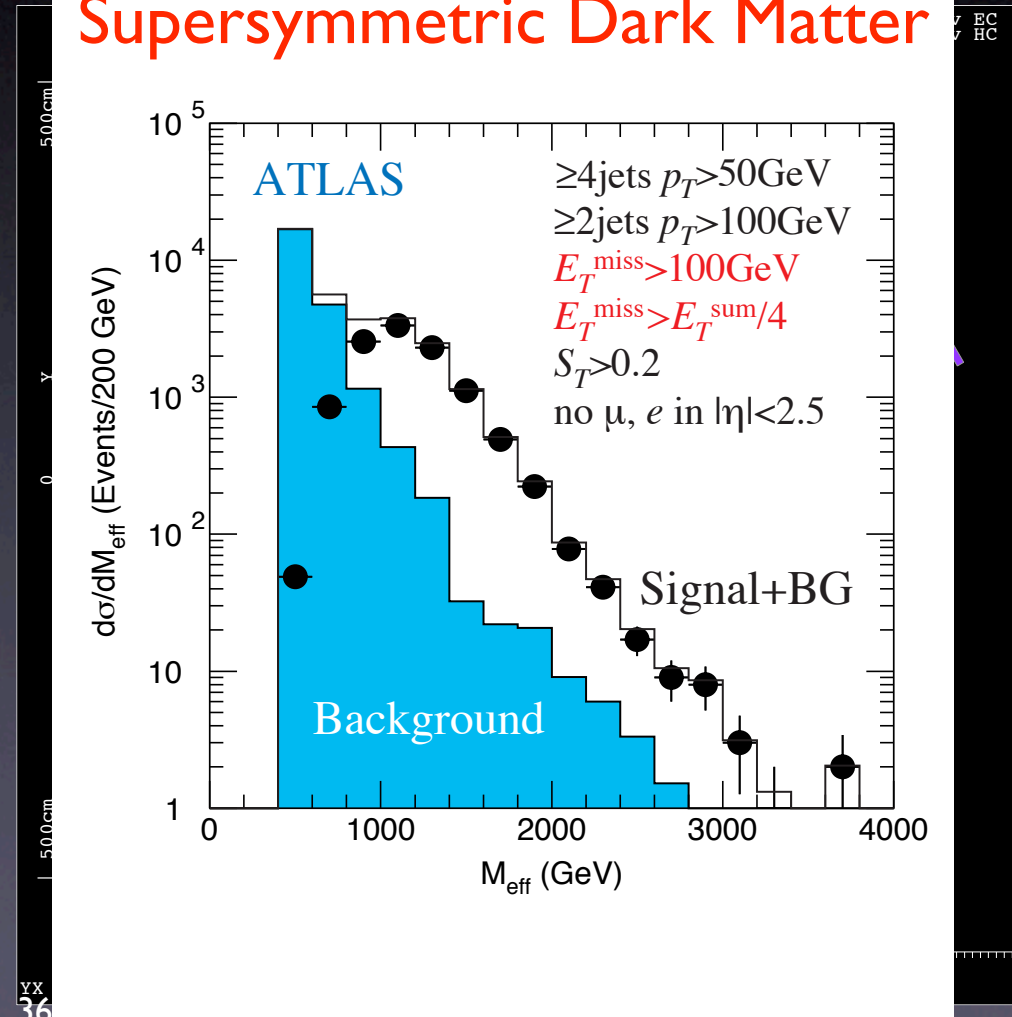
- Mimic Big Bang in the lab
- Hope to create invisible Dark Matter particles
- Look for events where energy and momenta are unbalanced

“missing energy” E_{miss}

- Something is escaping the detector

⇒ Dark Matter!?

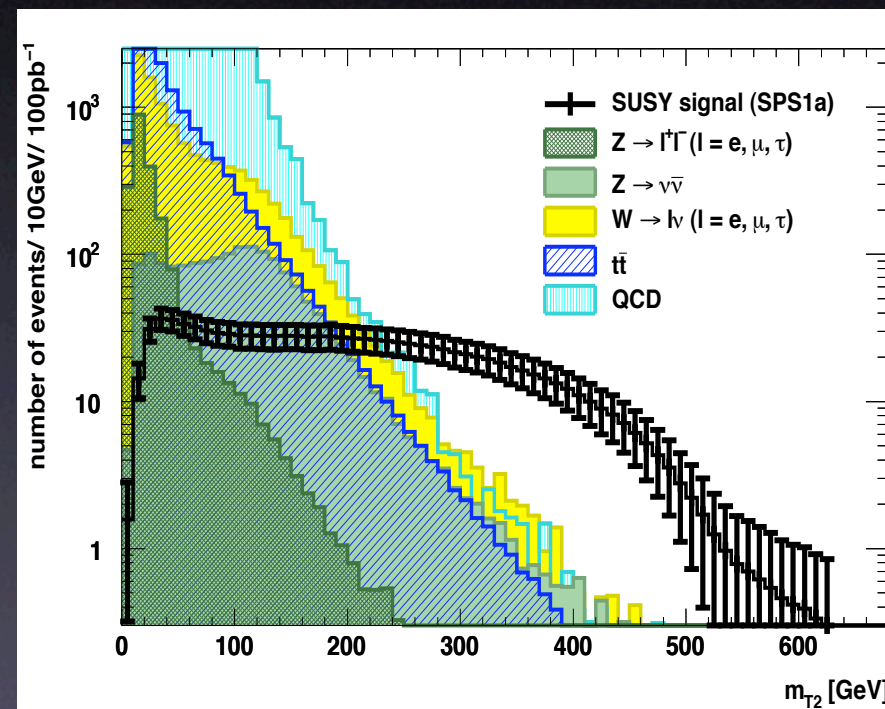
Supersymmetric Dark Matter



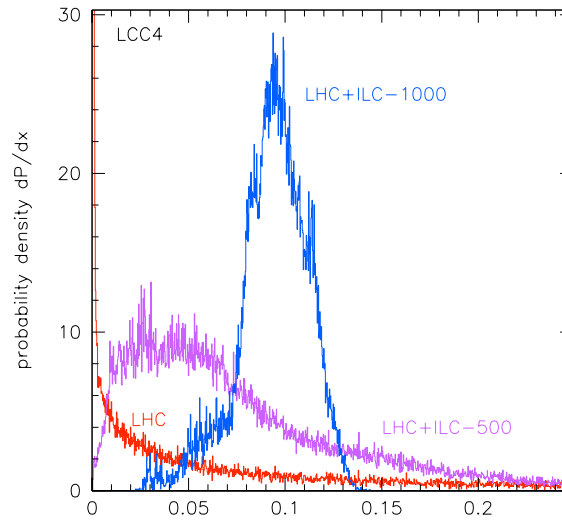
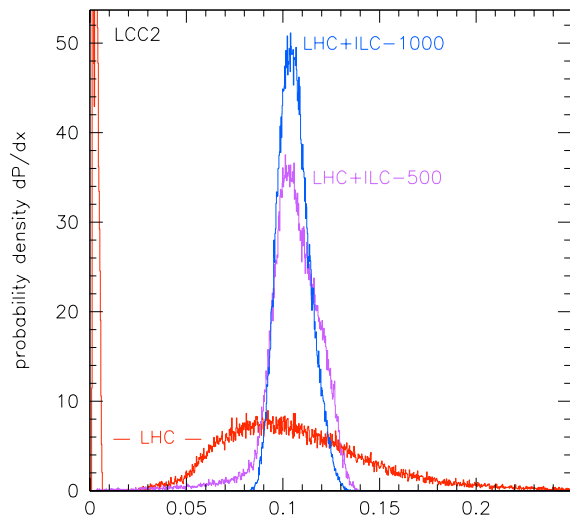
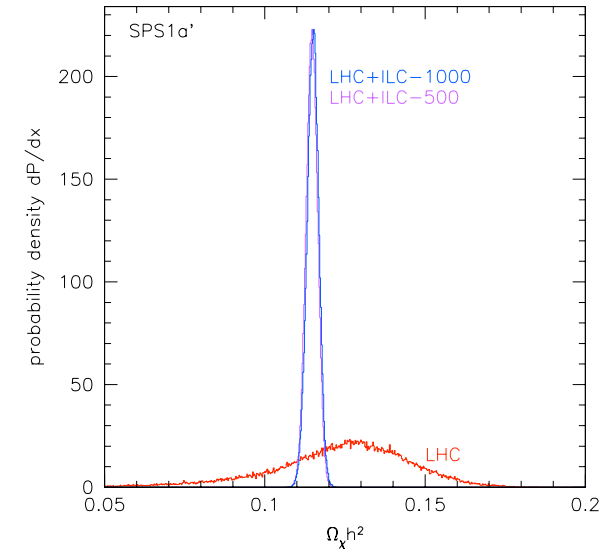
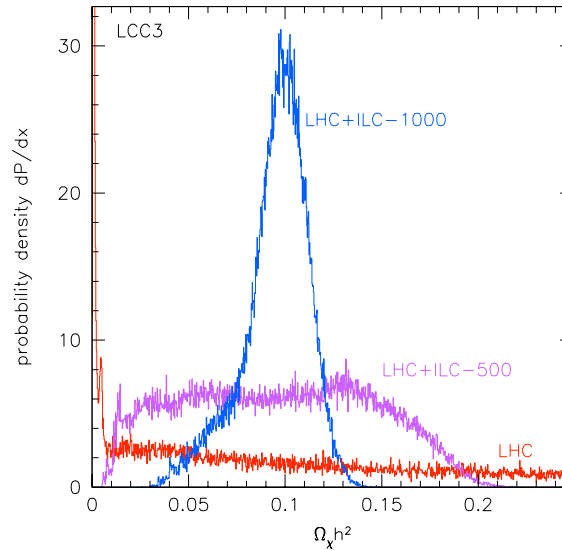
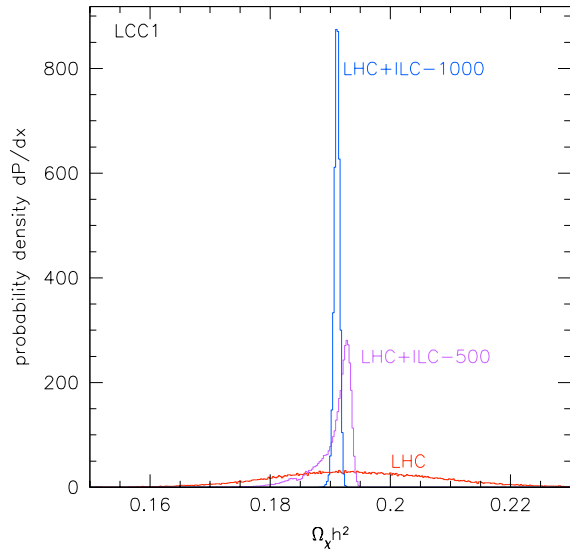
m_{T2} for discovery

- still room for improving the search strategy
- a quicker discovery with m_{T2} than ΣE_T

(A. Barr, C. Gwenlan)



Omega from colliders



SUSY case study
 Baltz, Battaglia, Peskin,
 Wizansky hep-ph/0602187

what Dark Matter is?

- cosmological measurement of dark matter

- abundance $\propto \sigma_{\text{ann}}^{-1}$

- detection experiments

- scattering cross section

- production at colliders

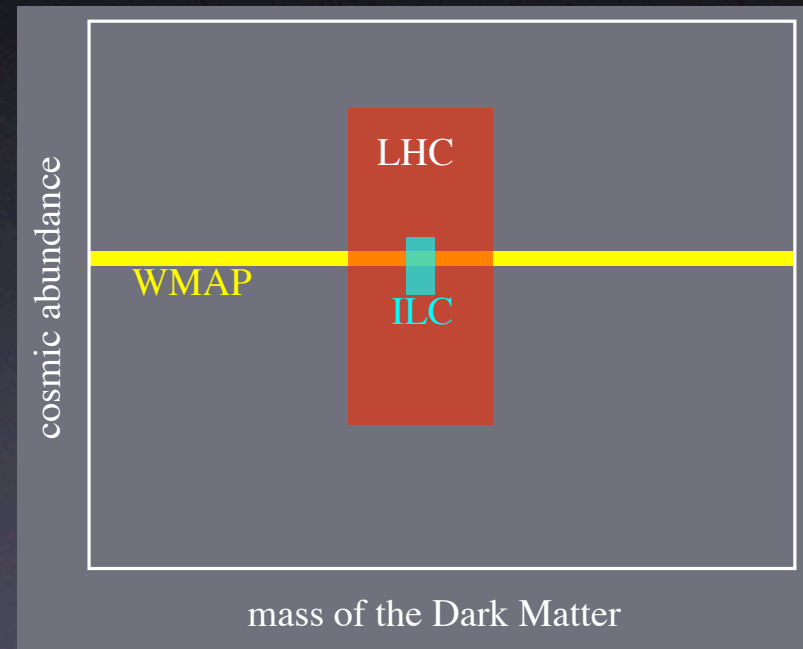
- mass, couplings

- can calculate cross sections

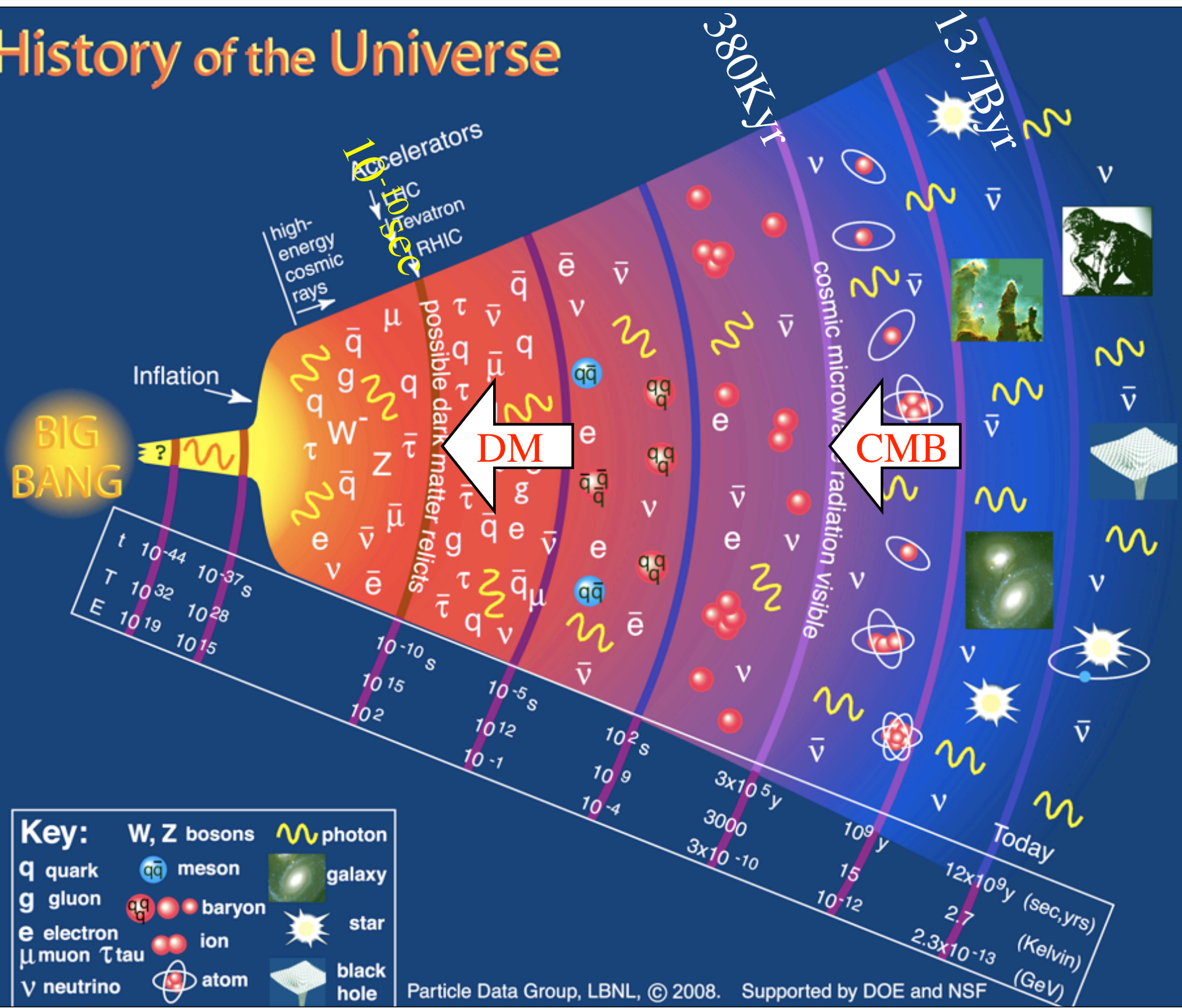
- If they agree with each other:

⇒ Will know *what Dark Matter is*

⇒ Will understand universe back to $t \sim 10^{-10}$ sec



History of the Universe



Anti-Matter

#8

STAR TREK

DEEP SPACE NINE

ANTIMATTER



With a dangerous cargo at stake, Commander Sisko must battle a band of hijackers!



John Vornholt

BESTSELLING AUTHOR OF DIGITAL FORTRESS

DAN BROWN



A NOVEL

ANGELS & DEMONS

"A breathless, real-time adventure... Exciting, fast-paced, with an unusually high IQ." —*San Francisco Chronicle*

Beginning of Universe

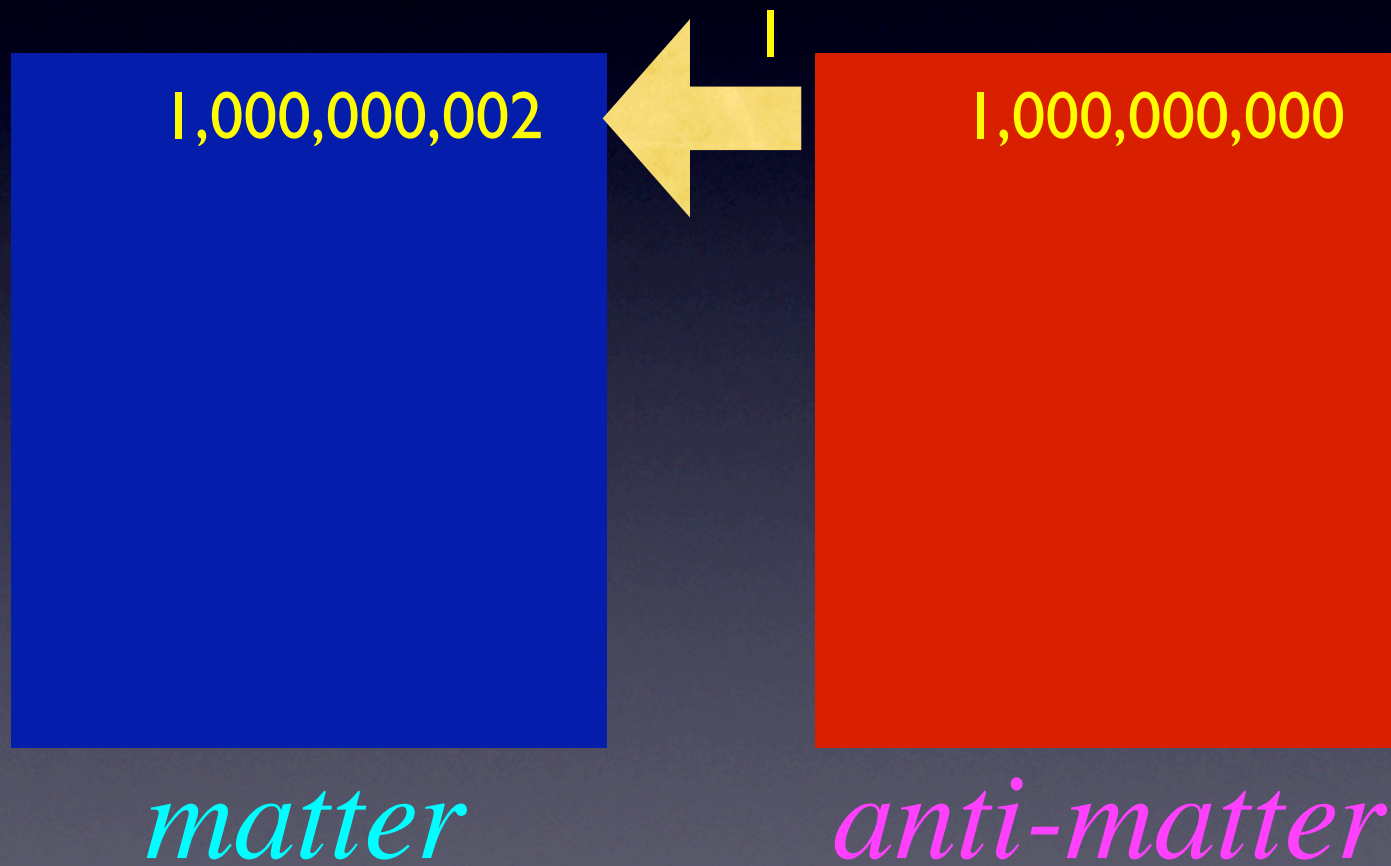
1,000,000,001

matter

1,000,000,001

anti-matter

fraction of second later



turned a billionth of anti-matter to matter

Universe Now

2

•

us

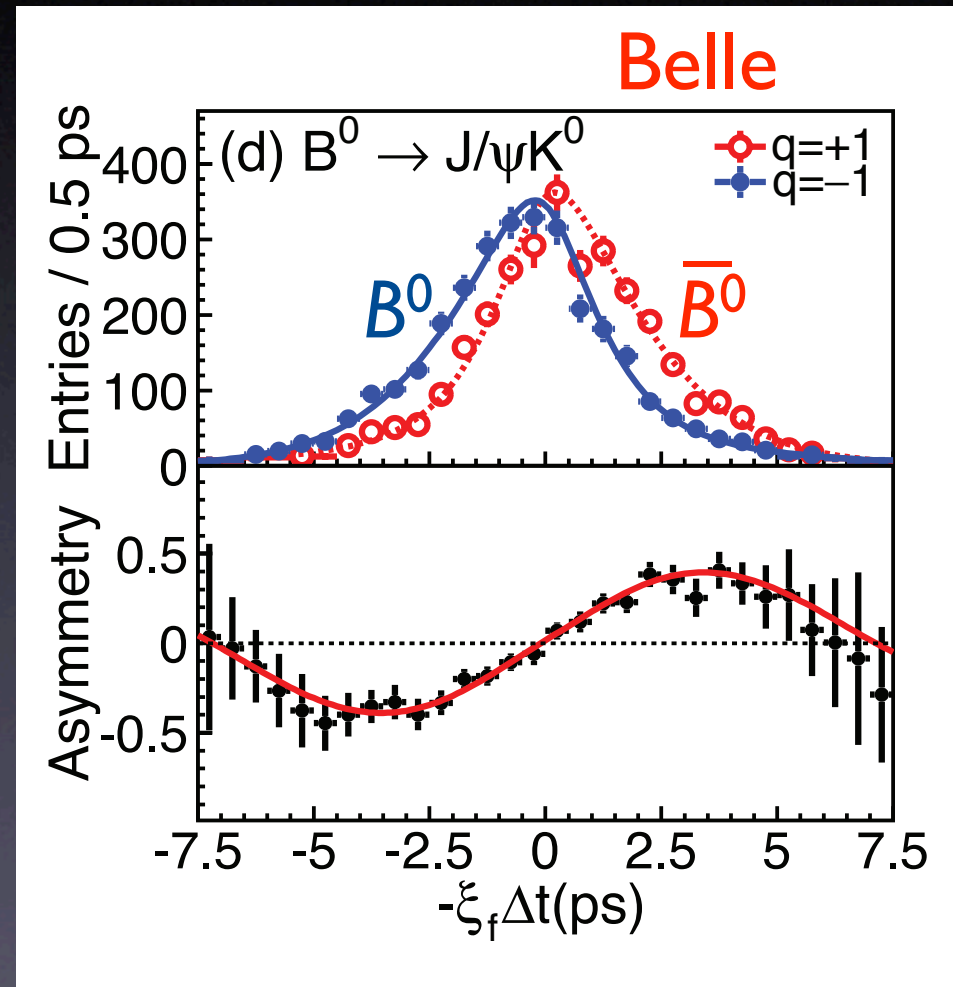
matter

anti-matter

This must be how we survived the Big Bang!

CP Violation

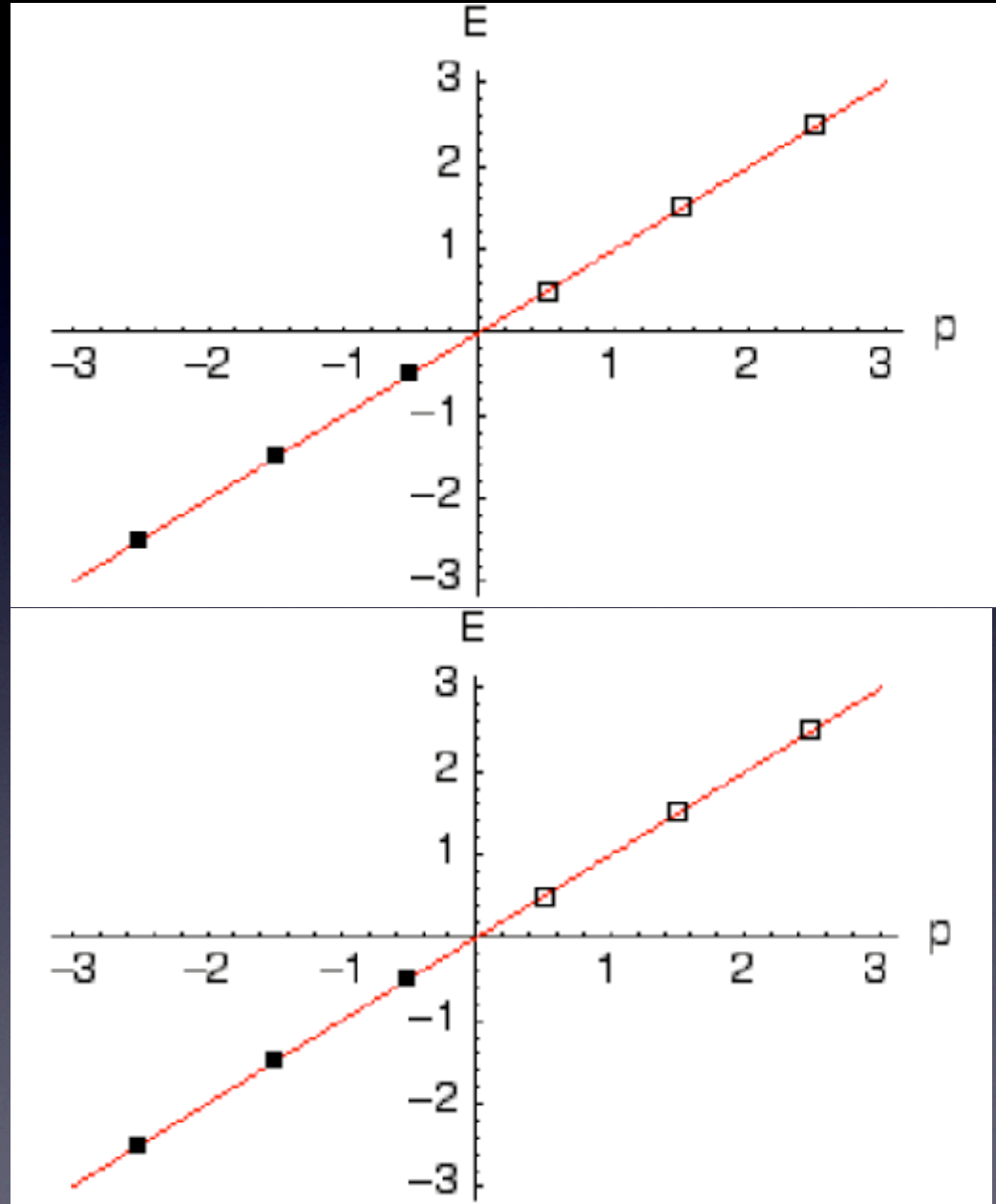
- Is anti-matter the exact mirror of matter?
- 1964 discovery of CP violation
- But only one system, hard to tell what is going on.
- 2001, 2002 Two new CP-violating phenomena
- But CP violation observed so far is too small by a factor of 10^{-16} to explain the absence of anti-matter
- Need new particles, new CP phases
- we don't know the energy scale



SM violates B

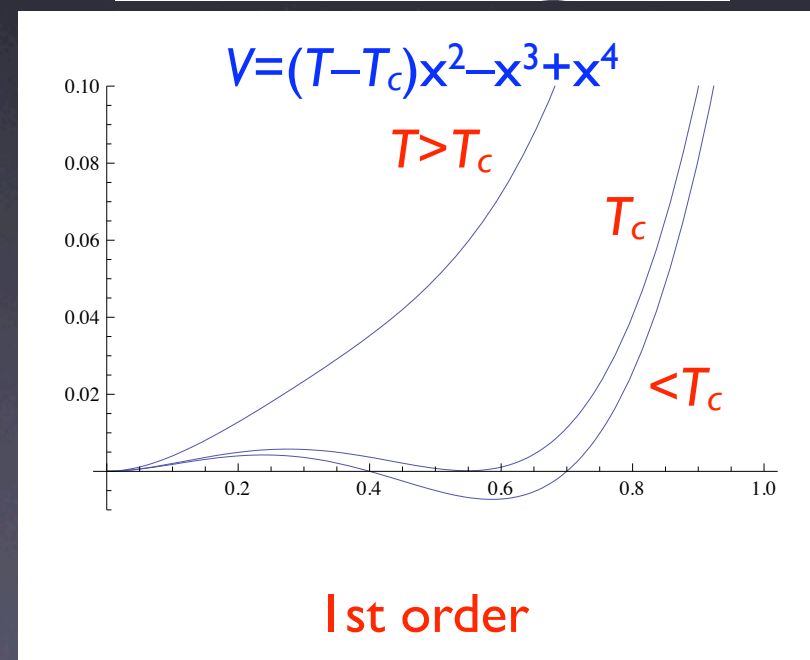
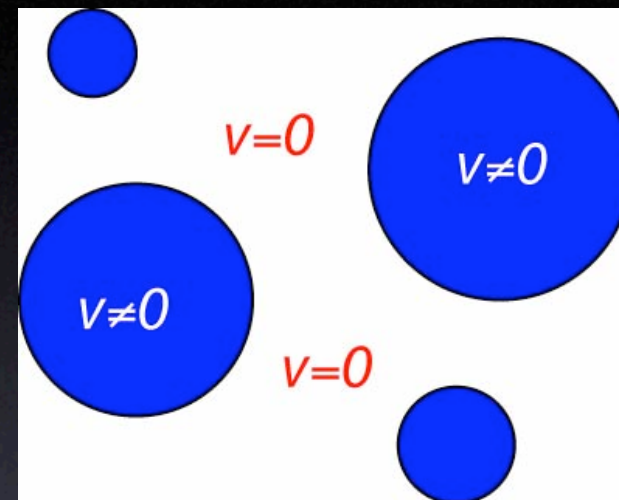
- W and Z bosons massless at high temperature
- W field fluctuates just like in thermal plasma
- solve Dirac equation in the presence of the fluctuating W field

change $\#q, \#l$



Sakharov's conditions

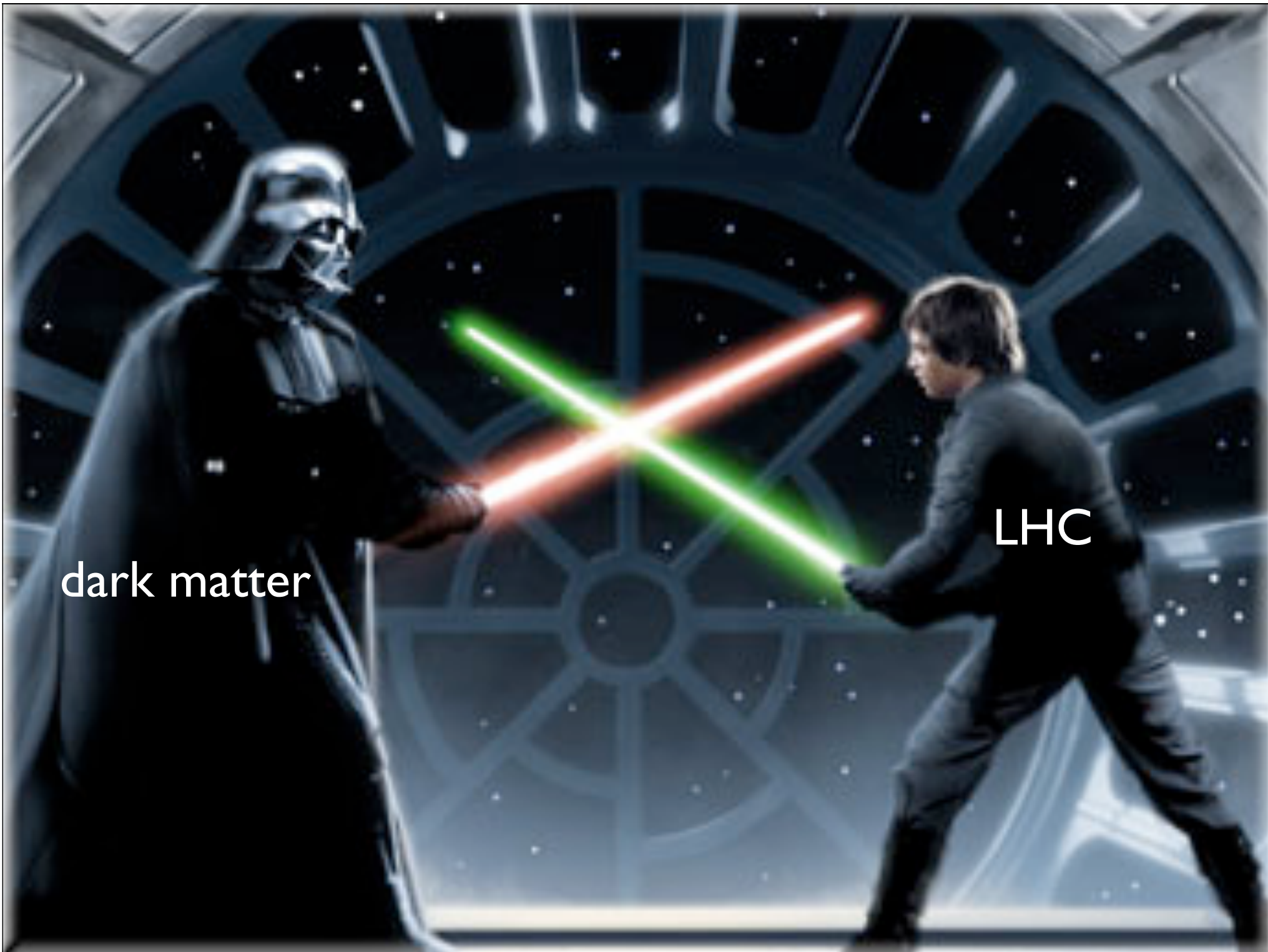
- ✓ Need B violation
- Need **CP violation**
 - new particles
- Need **out-of-equilibrium**
 - **bubbles** in first-order phase transition
 - need extra “**light**” **bosons** in loops
 - e.g. stop ~ 160 GeV, extra Higgs bosons
 - should be in the LHC range
 - or H^6 potential?



Conclusions

- We are entering the **new era**
 - twice in century opportunity!
- Standard Model is definitely not the whole story: **five evidences**
- **LHC well prepared** to address some of the five, physics behind Higgs BEC
- no guarantee, and that's why we need data

Can't wait to see the data from LHC!



dark matter

LHC