

# Invisible Higgs decays in Vector Boson Fusion

Othmane Rifki

*LHC discussion on Dark Matter*

23. Sep 2018

# Outline

- Why search for Higgs to Invisible?
- Vector boson fusion channel (VBF)
- Other Higgs to invisible limits
- Global fits from visible decays
- Higgs portal interpretation

# Motivation

Higgs (H) is a special probe for **Dark Matter (DM)**



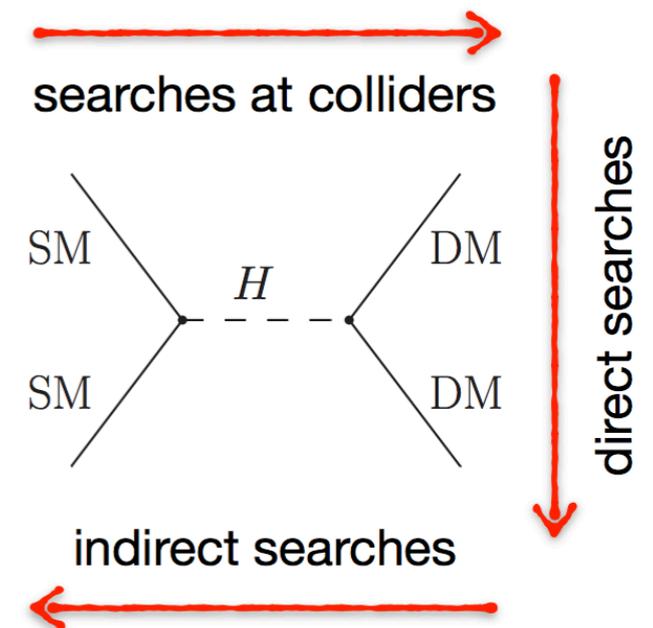
	I	II	III	
Quarks	$u$	$c$	$t$	$\gamma$
	$d$	$s$	$b$	$g$
Leptons	$\nu_e$	$\nu_\mu$	$\nu_\tau$	$Z$
	$e$	$\mu$	$\tau$	$W$

Force Carriers

Three Generations of Matter

- DM may “live” in some Dark Sector and not charged under gauge groups of the SM
- Higgs could have some coupling to the Dark Sector, i.e “Higgs Portal”
- Higgs acting as mediator between SM and DM

➔ **Direct production at the LHC**

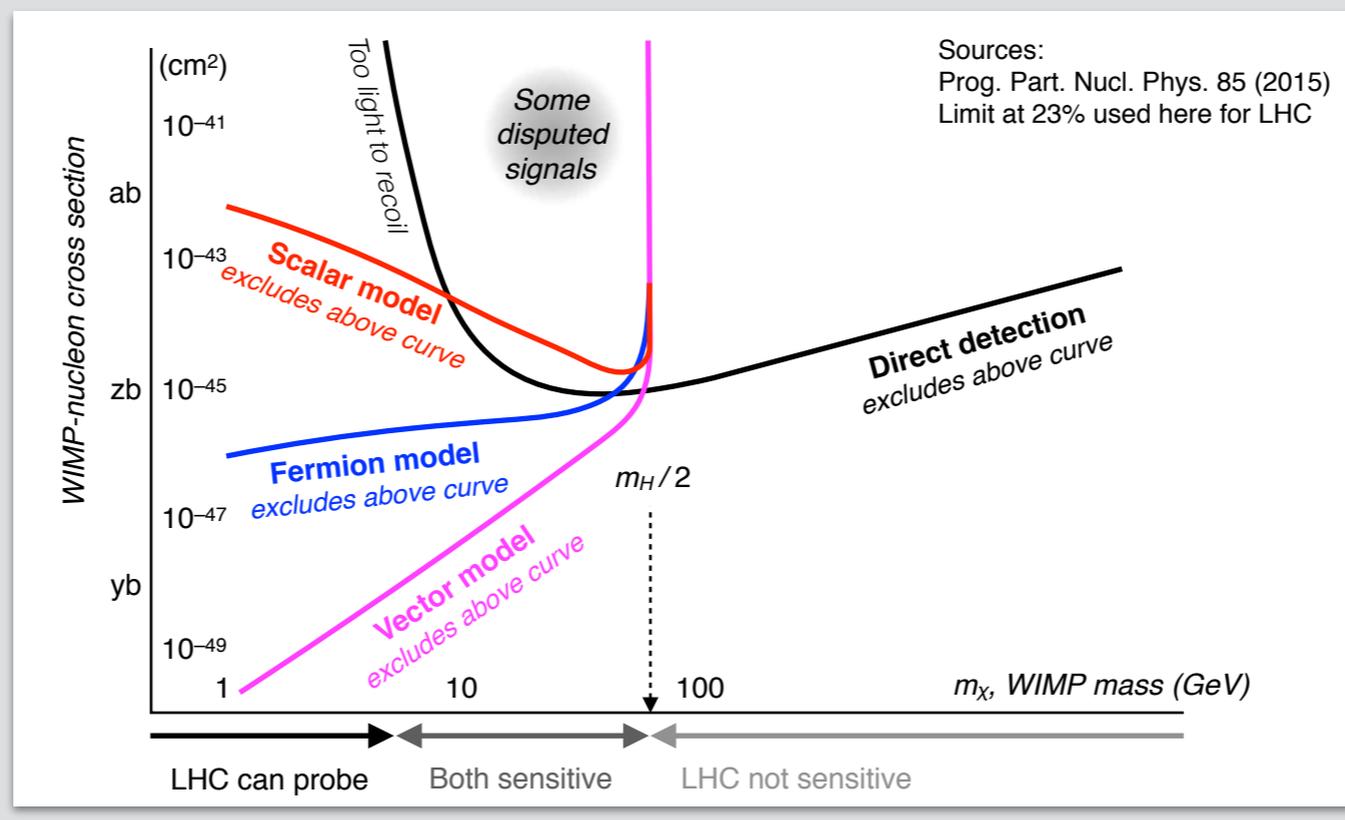


# Motivation

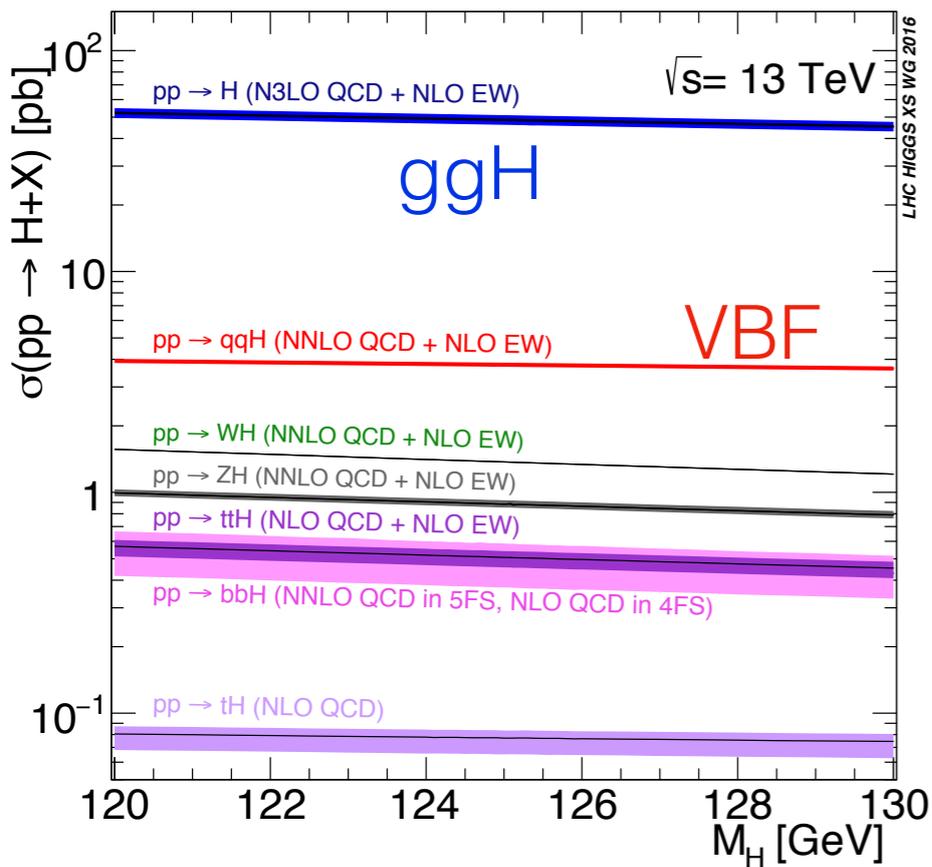
- In the SM,  $H \rightarrow$  Invisible only from  $H \rightarrow ZZ \rightarrow \nu\nu\nu$  with  $B(H \rightarrow \text{inv}) = 0.026 \times 0.20^2 = 0.1 \%$   
**Any deviation would indicate BSM physics!**
- Higgs to Invisible searches are powerful for  $m_{DM} \leq m_H/2$

## Dark matter interpretation of invisible Higgs

Limit on  $B(H \rightarrow \text{invisible})$  with Higgs-DM portal results in colored exclusion curves.

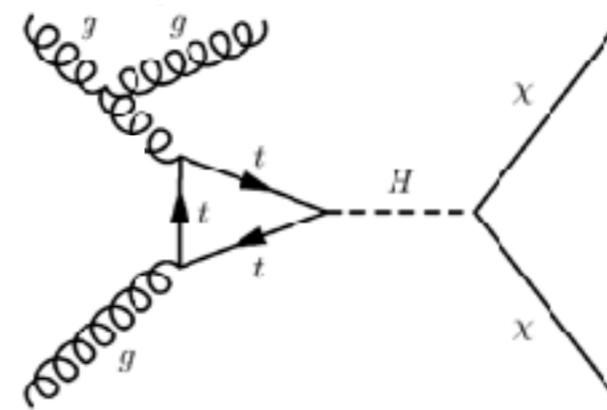


# Why VBF?



ggH production is much larger than VBF

## Why not use ggH?



1. ggH would need to be boosted by an extra gluon in the event  
VBF has some natural boost
2.  $qq > Z > \nu\nu$  background is much larger than ggH production

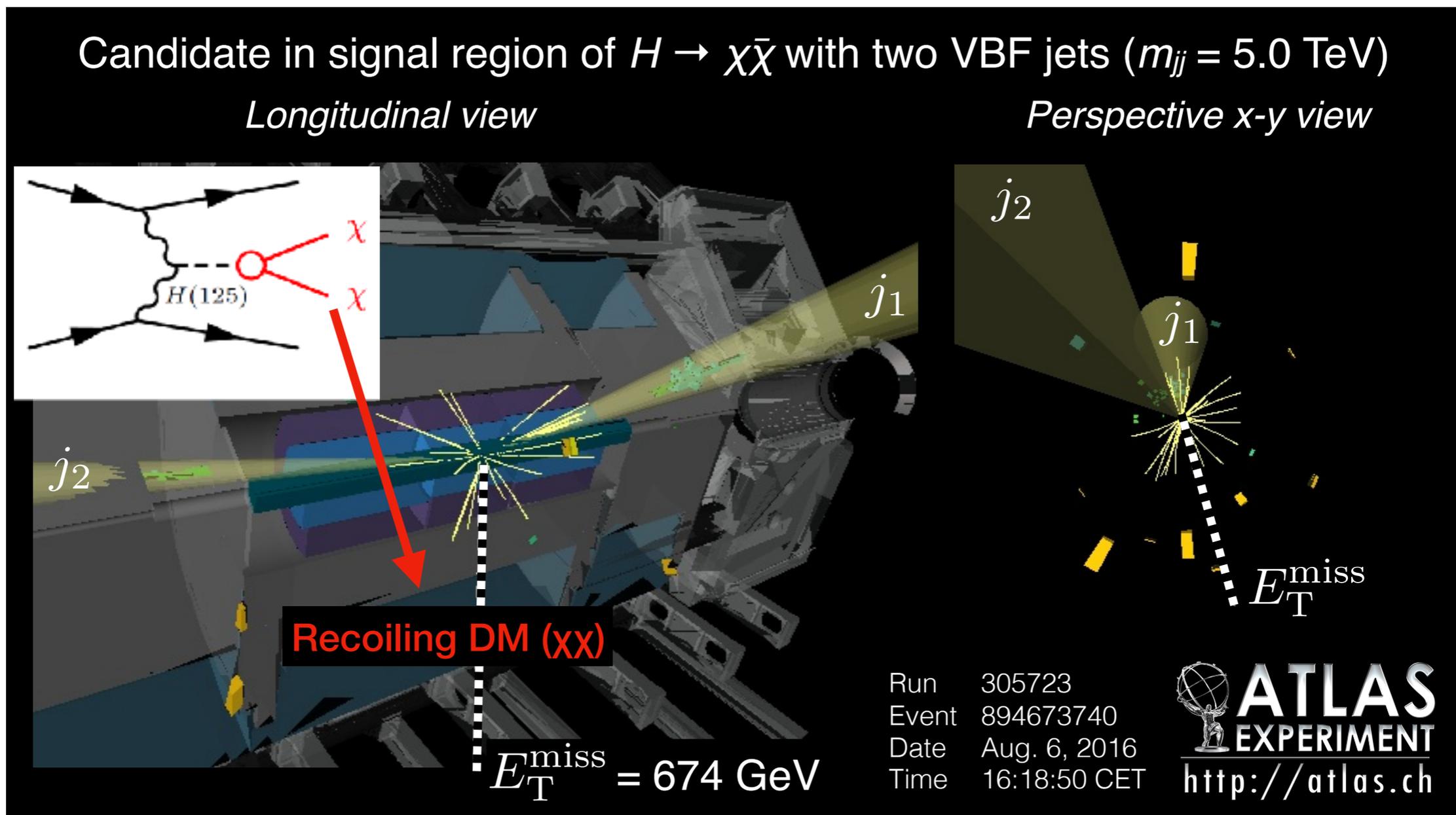
$$\frac{\sigma(\text{ggH})}{\sigma(Z) \times BR(Z \rightarrow \nu\nu)} = \frac{\approx 45 \text{ pb}}{\approx 8422 \text{ pb}}$$

$$\frac{\sigma(\text{VBF H})}{\sigma(\text{VBF Z}) \times BR(Z \rightarrow \nu\nu)} = \frac{\approx 3.8 \text{ pb}}{\approx 2.9 \text{ pb}}$$

**Focus on processes where signal  $\sim$  background**

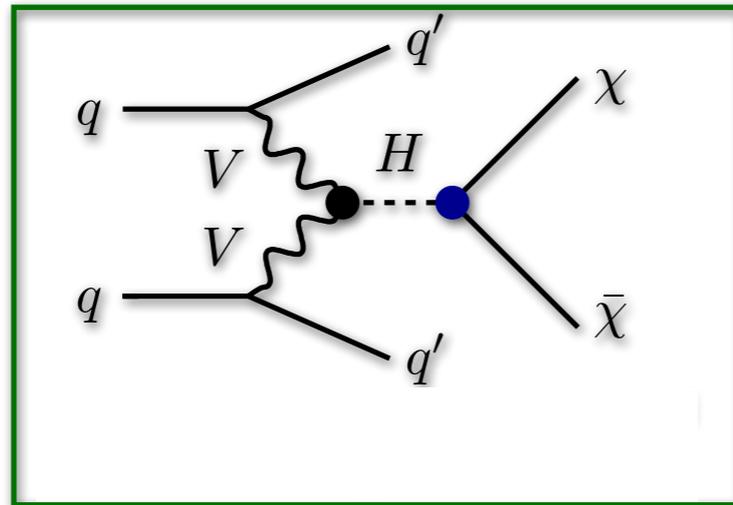
# VBF Topology

- 2 well-separated jets
- Large missing transverse momentum

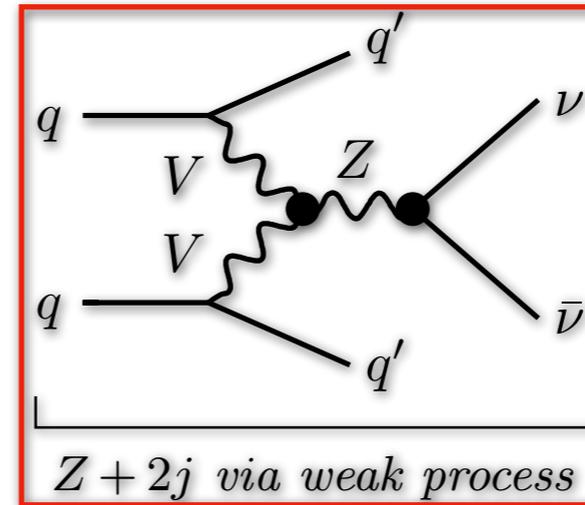


# Signal and Background

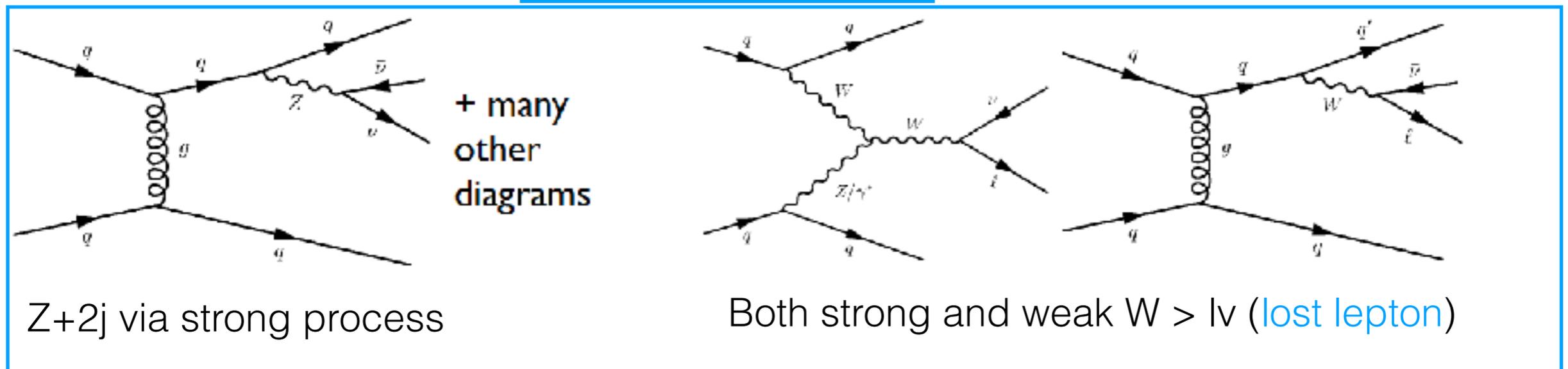
Signal



Irreducible Background



Reducible Background



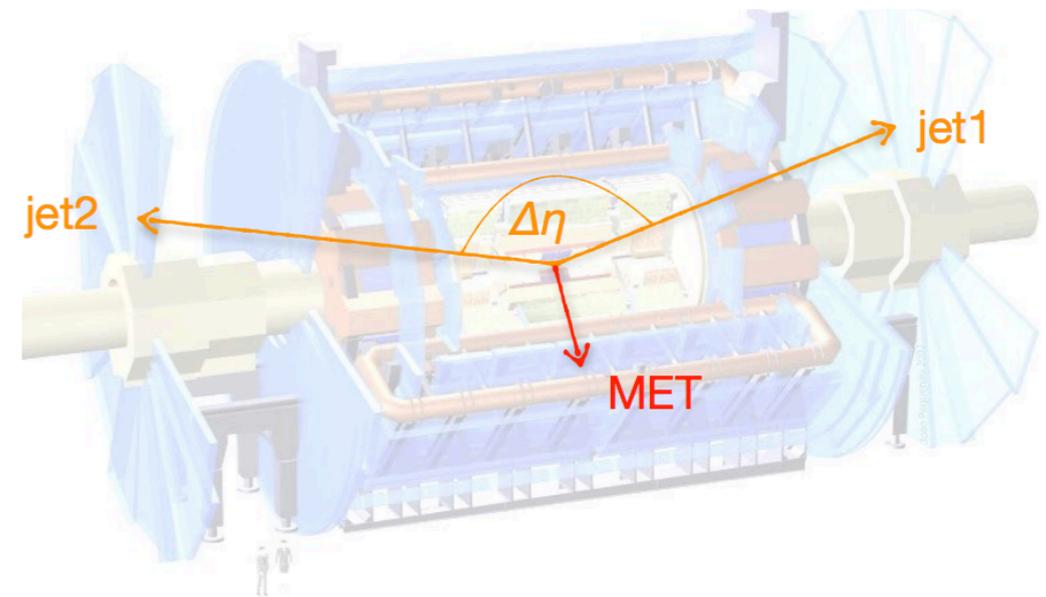
# Selection

## Basic 2-jet+MET Selection

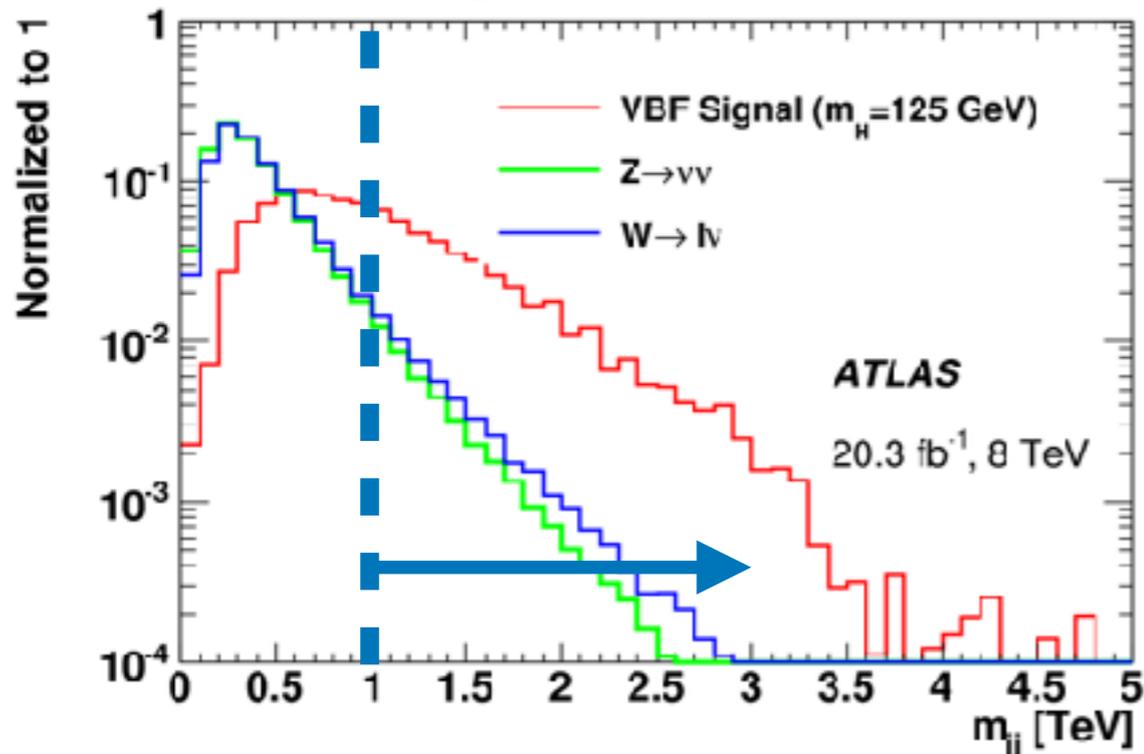
- Jet 1:  $p_T > 80$  GeV
- Jet 2:  $p_T > 50$  GeV
- Veto 3rd jet above 25 GeV
- MET  $> 180$  GeV (determined by trigger)

## MET cleaning

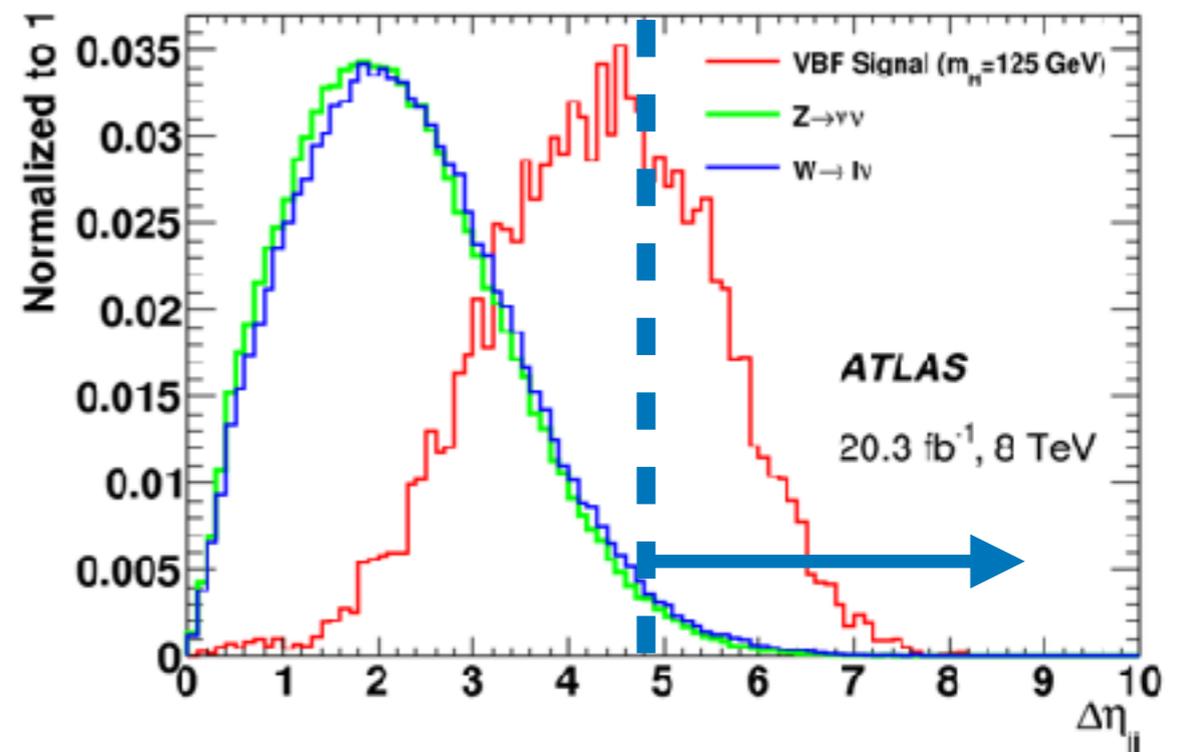
- $|\text{DPhi}(j, \text{MET})| > 1$
- $|\text{DPhi}(j, j)| < 1.8$



$m_{jj} > 1$  TeV

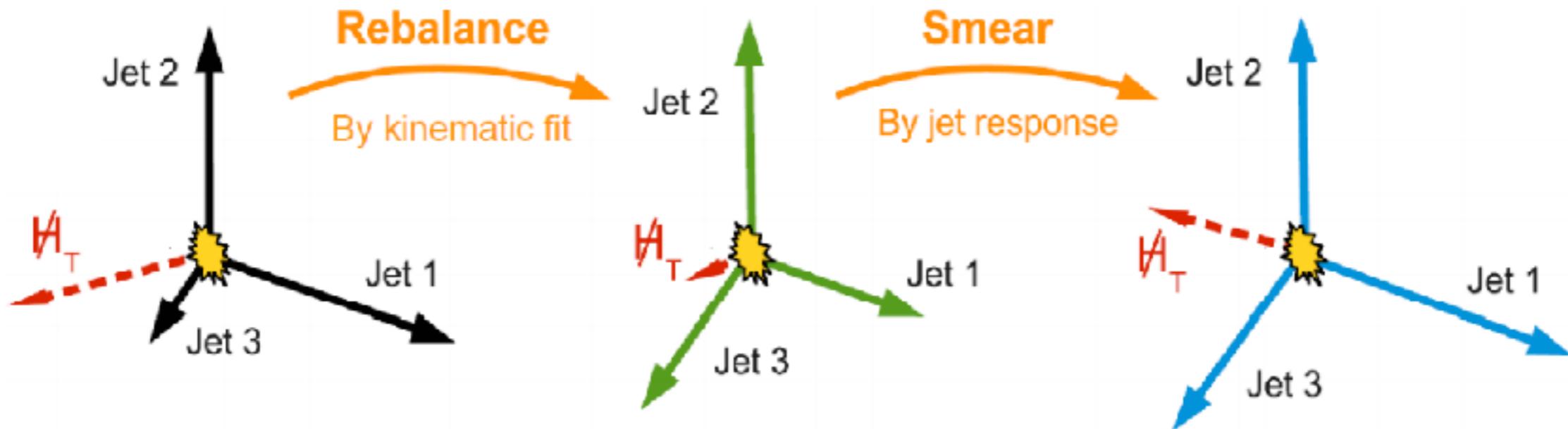
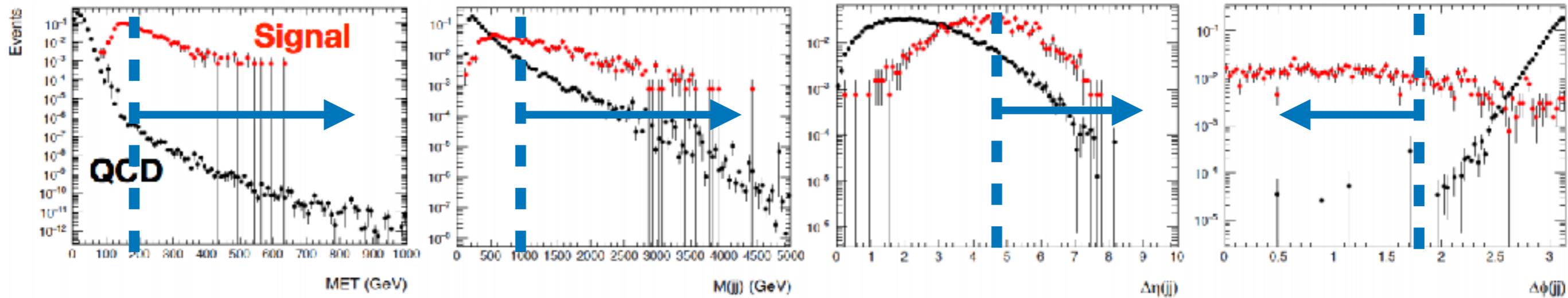


$\Delta\eta_{jj} > 4.8$



# Multijet Background

QCD multijet background is small for analysis selection  
 Difficult to estimate since it is due to instrumental effect



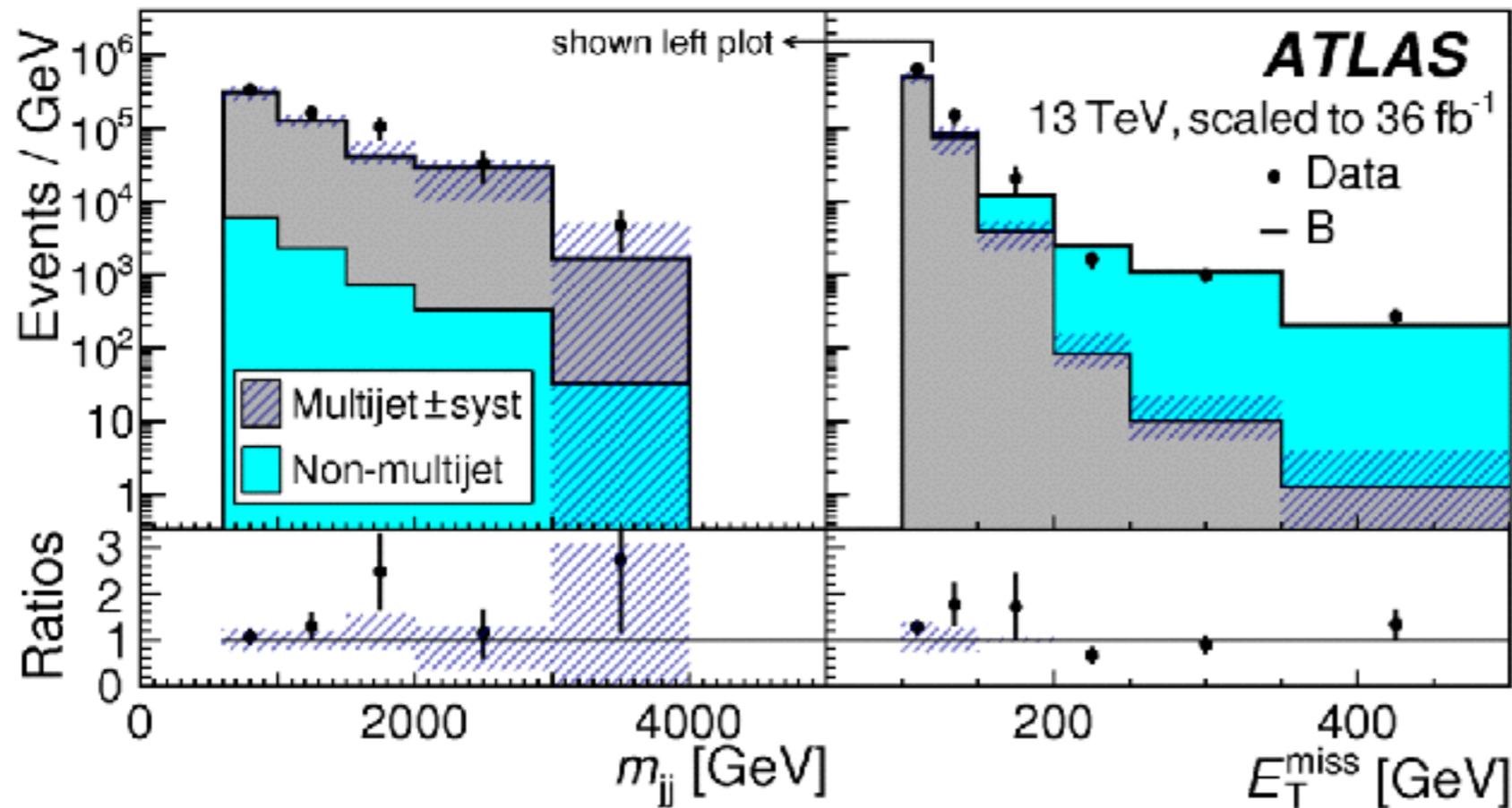
Seed: Single jet triggers

Vary jet  $p_T$ , minimize MET  
 Based on jet resolution  
 Sample of events with no MET

Randomly vary jet  $p_T$   
 according to known  
 resolution

# Multijet Background

## Multijet background under control



Multijet events also contribute to regions with electrons if one jet is mis-identified as electron

- Additional MET significance requirement (high and low MET sig. used in fit)

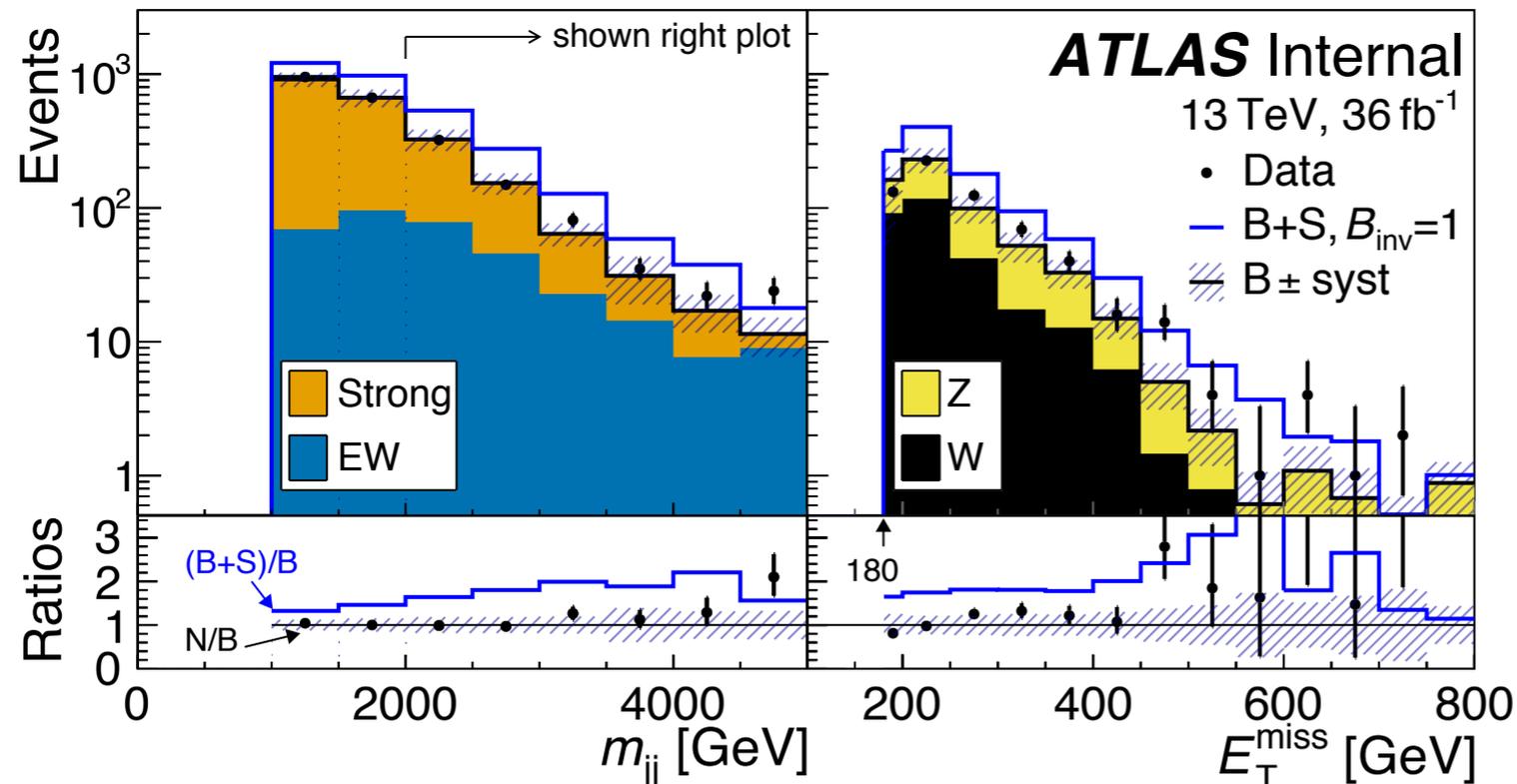
$$E_T^{\text{miss}} \text{ sig.} = \frac{E_T^{\text{miss}}}{\sqrt{p_T(j1) + p_T(j2) + p_T(el)}}$$

# Backgrounds

W and Z backgrounds are ~50/50

Small top and diboson background from MC  
 Multijet reduced by SR cuts  
 Background estimation next...

Description	SR
	Yield [EW]
$N$ , observed	2252
$B$ , expected	2243
$Z \rightarrow \nu\nu$	1111 [18%]
$Z \rightarrow ee, \mu\mu$	12 [ 9%]
$Z \rightarrow \tau\tau$	10 [16%]
$W \rightarrow e\nu, \mu\nu$	540 [16%]
$W \rightarrow \tau\nu$	533 [20%]
Other	36
$S$ , signal	1070
VBF	930
Gluon fusion	140



Signal for 100% branching fraction

Larger EW fraction at higher  $M_{jj}$

# Background Modelling

**Statistical uncertainty:**  $1/\sqrt{2252} \sim 2\%$

**Systematic uncertainty:**  $\sim 20\%$

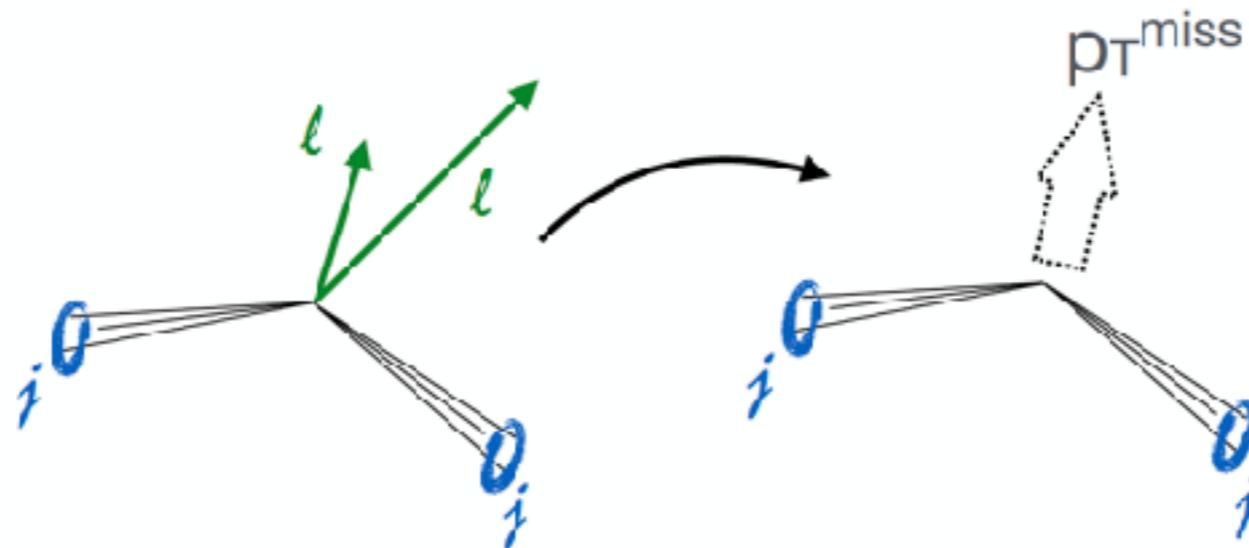
- Modelling of underlying physics of W or Z + 2 jets
- Jet uncertainties (forward jets, jet veto)

**MC systematics are too large to exploit the full statistics!**

**Solution:** Use Z $\rightarrow$ ll and W $\rightarrow$ lv data with found leptons to model Z $\rightarrow$ vv and W $\rightarrow$ lv (with lost lepton)

Define control regions:

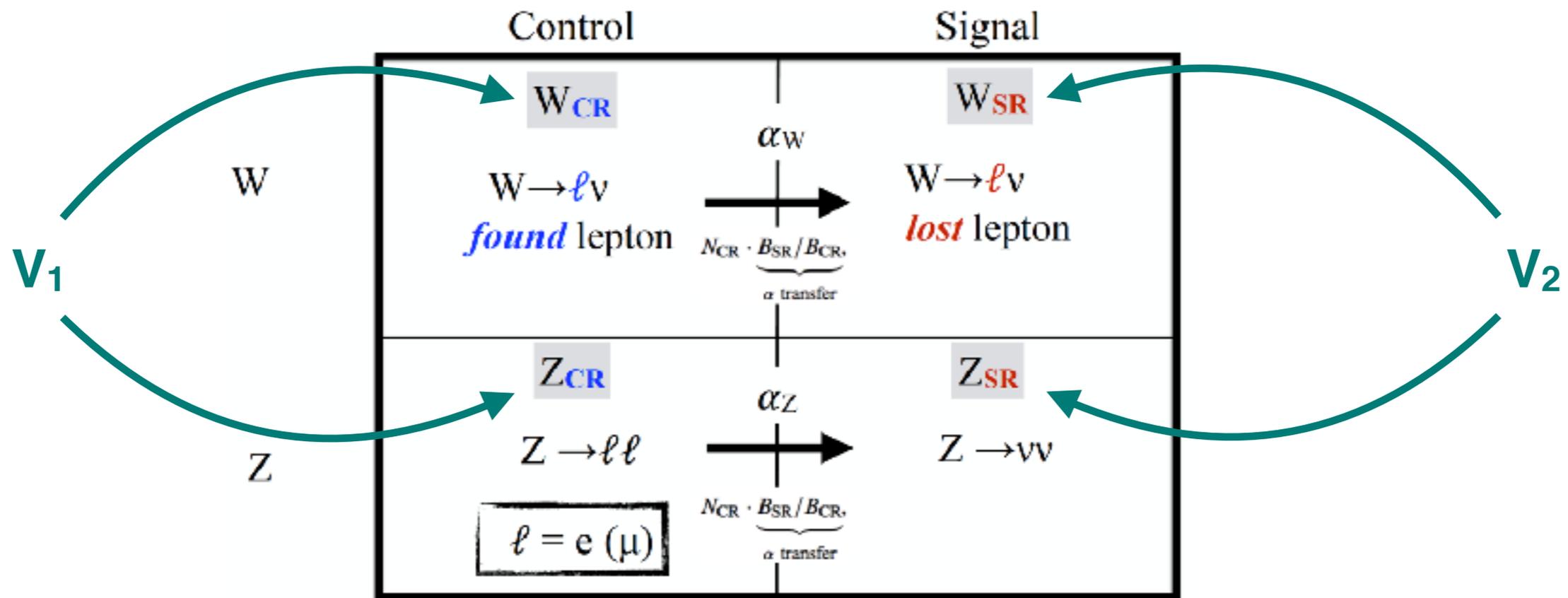
- Z CR: 2-leptons, use **ll-system** instead of MET
- W CR: 1-lepton, add **1-lepton** p<sub>T</sub> to the MET



# Background Modelling

Use precise measurement of  $V_1$  to get an improved prediction for  $V_2$ :

$$\frac{d\sigma(V_2)}{dx} = \left[ \frac{d\sigma(V_2)/dx}{d\sigma(V_1)/dx} \right]_{\text{theory}} \times \left[ \frac{d\sigma(V_1)}{dx} \right]_{\text{measured}}$$



Simultaneous fit to **signal region** and **Z and W control regions** leads to cancellation of important theory uncertainties

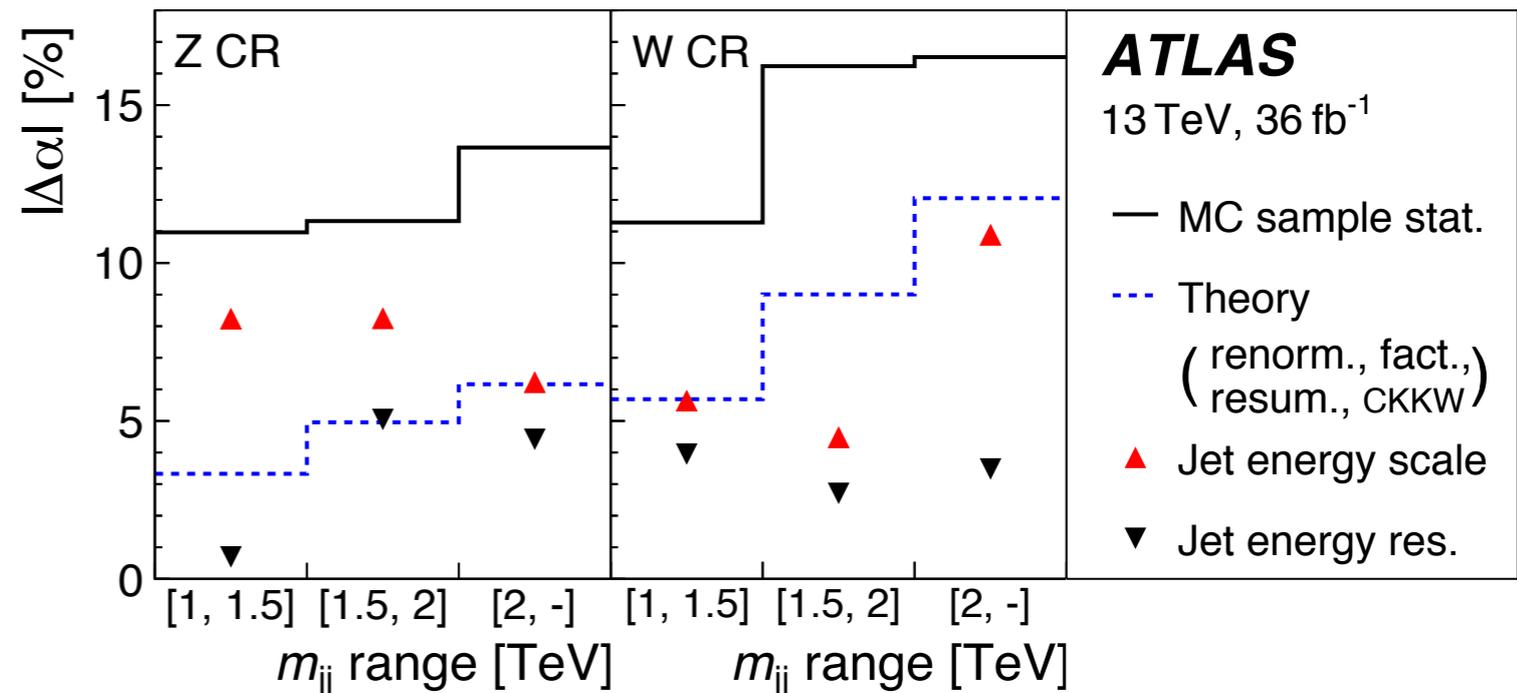
# Uncertainties

Systematics reduction with ratio method:

Source	$\mathcal{B}_{\text{inv}}$ improve. [%]		Yields, $\alpha$ changes (%)				
	$\Delta$	visual	$S$	$B_{\text{SR}}^Z$	$B_{\text{CR}}^Z$	$\alpha_Z$	$\alpha_W$
<b>Experimental (<math>\dagger</math>)</b>							
Jet energy scale	10	—	12	7	8	8	6
Jet energy resol.	2	+	2	0	1	1	4
$E_T^{\text{miss}}$ soft term	1	+	2	2	2	2	2
Lepton id., veto	2	+	-	-	-	-	4
Pileup distrib.	1	+	3	1	2	3	1
Luminosity	0	+	2	2	2	-	-
<b>Theoretical (<math>\ddagger</math>)</b>							
Resum. scale	1	+	-	2	3	0	2
Renorm., fact.	2	+	-	20	19	1	2
CKKW matching	4	+	-	2	3	1	5
PDF	0	+	1	1	2	1	1
3 <sup>rd</sup> jet veto	2	+	7	-	-	-	-
<b>Statistical</b>							
MC sample ( $\star$ )	12	—	4	5	9	10	9
Data sample	21	—	6	5	12	12	6
<b>Combined</b>							
All $\dagger$ sources	17	—					
All $\ddagger$ sources	10	—					
Combine $\dagger, \ddagger$	28	—					
Combine $\dagger, \ddagger, \star$	42	—					

QCD scale variations are treated as correlated

Many uncertainties due to MC stat!



$$B_{\text{SR}}^{\text{est}} = B_{\text{SR}} \cdot \underbrace{N_{\text{CR}}/B_{\text{CR}}}_{k \text{ normalization}} = N_{\text{CR}} \cdot \underbrace{B_{\text{SR}}/B_{\text{CR}}}_{\alpha \text{ factor}}$$

# Fit Model

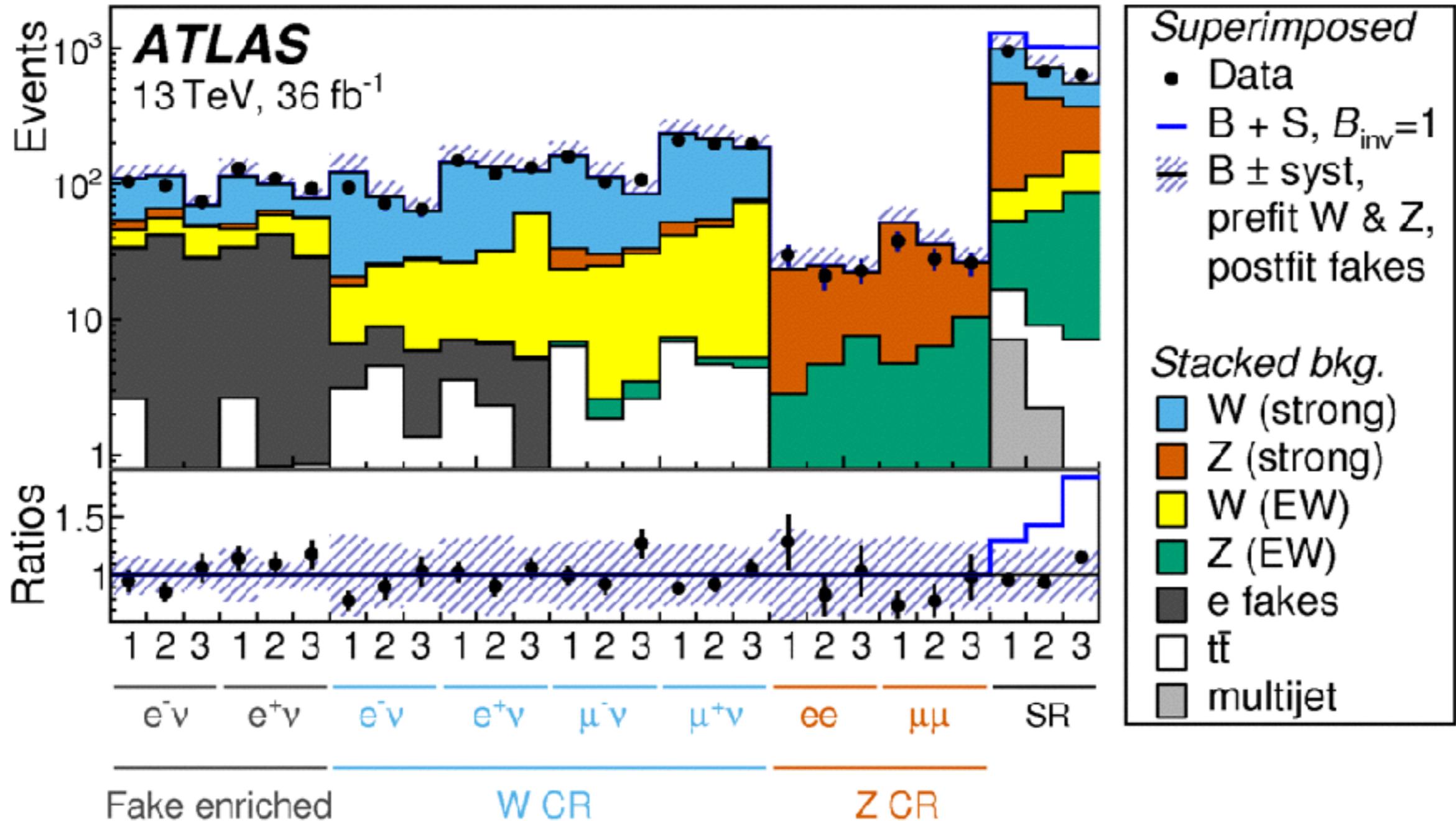
Global likelihood fit using **8 control region bins** for each **signal region bin**

- Signal region bins defined in  $M_{jj}$ : **1-1.5TeV, 1.5 - 2TeV, > 2TeV**
- Each  **$M_{jj}$  bin is treated independently**
- **Separate normalisation factors for W and Z processes**: 3x2 factors
- 1 normalisation factor for mis-identified electrons
- 1 signal yield correlated across all bins
- Systematics implemented as correlated Gaussian constrained nuisance parameters

**27 bins, 9 free parameters, 1 signal yield**

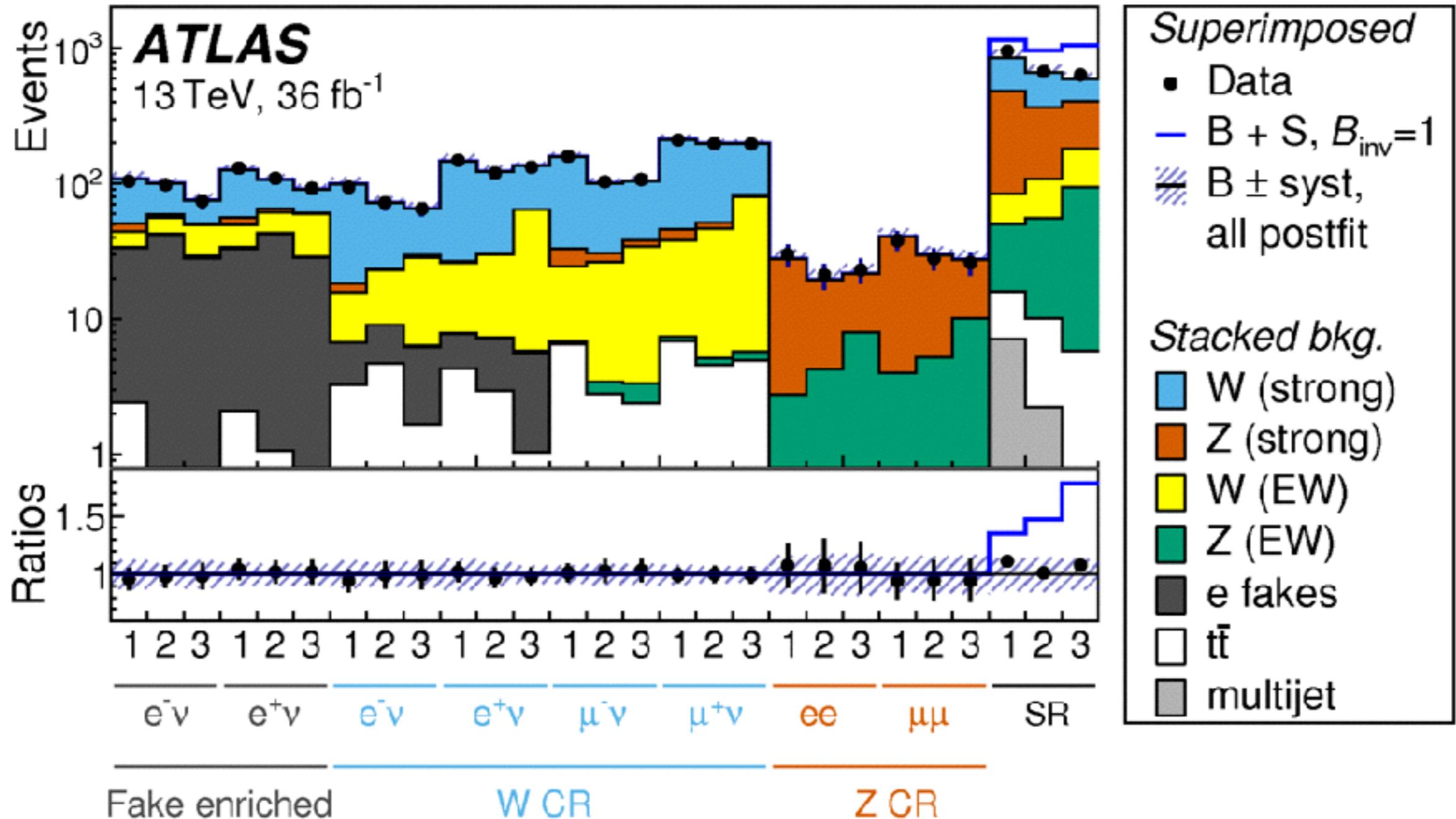
# Pre-fit

Global likelihood fit with 27 bins, 9 free parameters, 1 signal yield



# Post-fit

Global likelihood fit with 27 bins, 9 free parameters, 1 signal yield

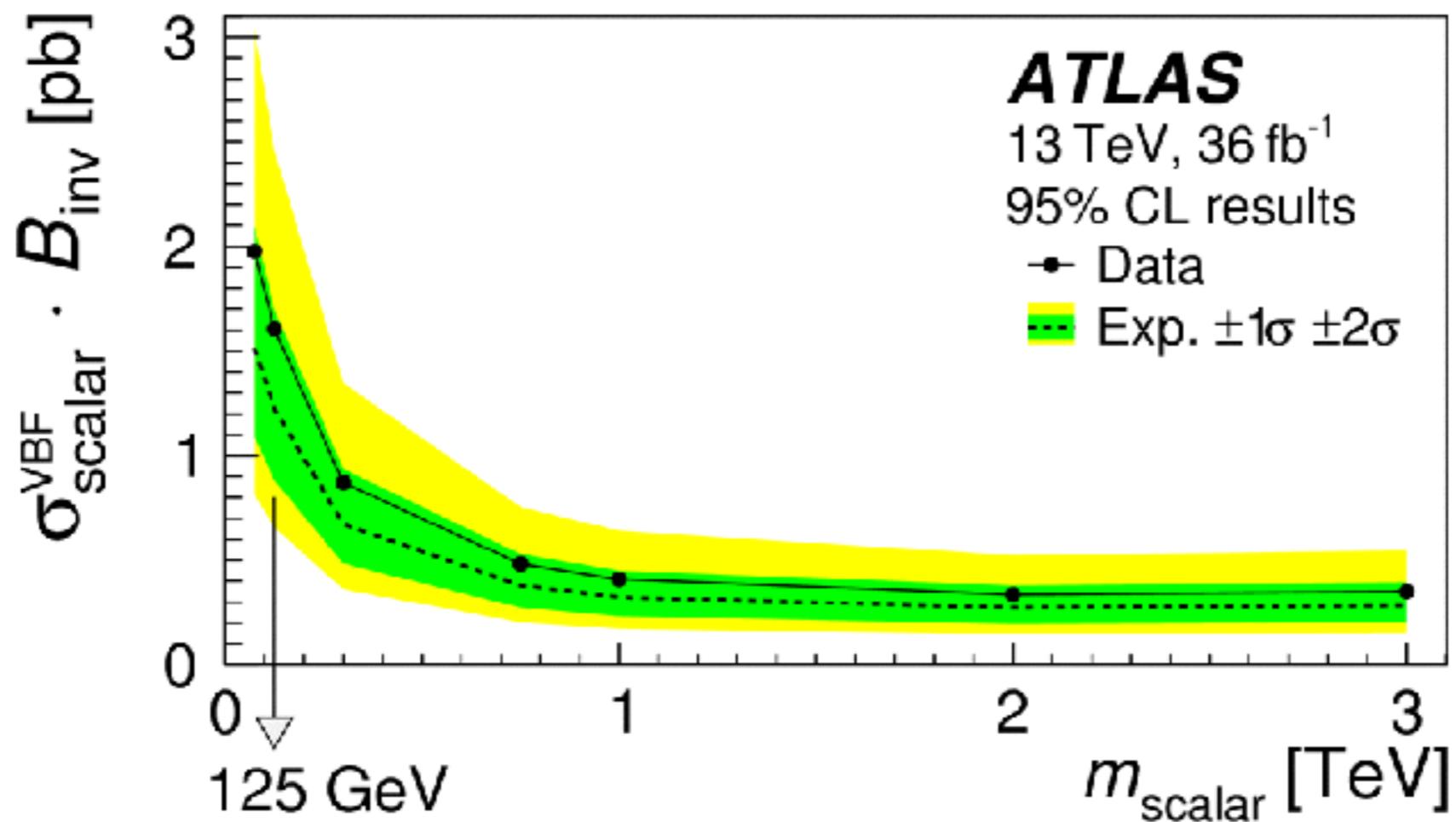


# Result

Upper bound on branching fraction of Higgs to Invisible at 95% CL:

**Observed = 0.37**

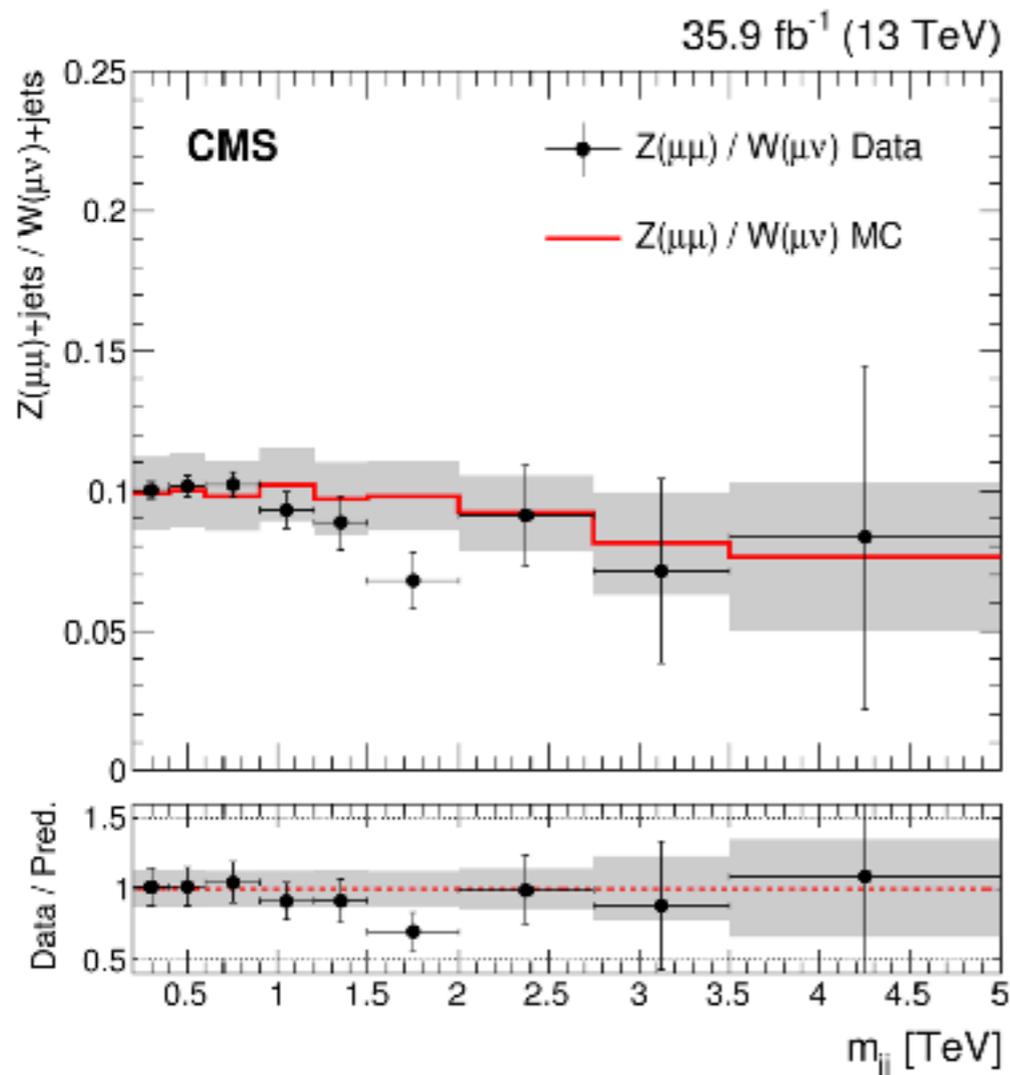
**Expected = 0.28**



# CMS Result

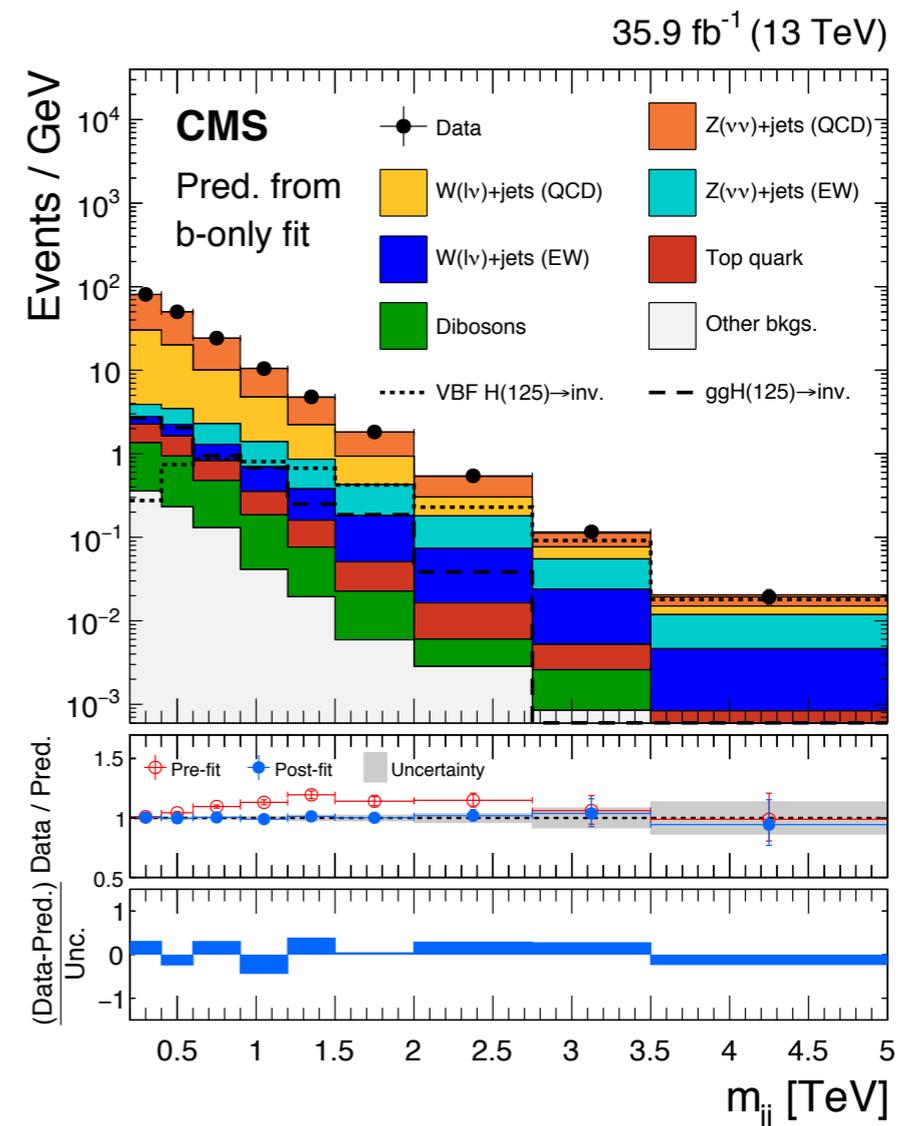
Cut-and-Count

**Observed = 0.58**  
**Expected = 0.30**



Shape-fit MC

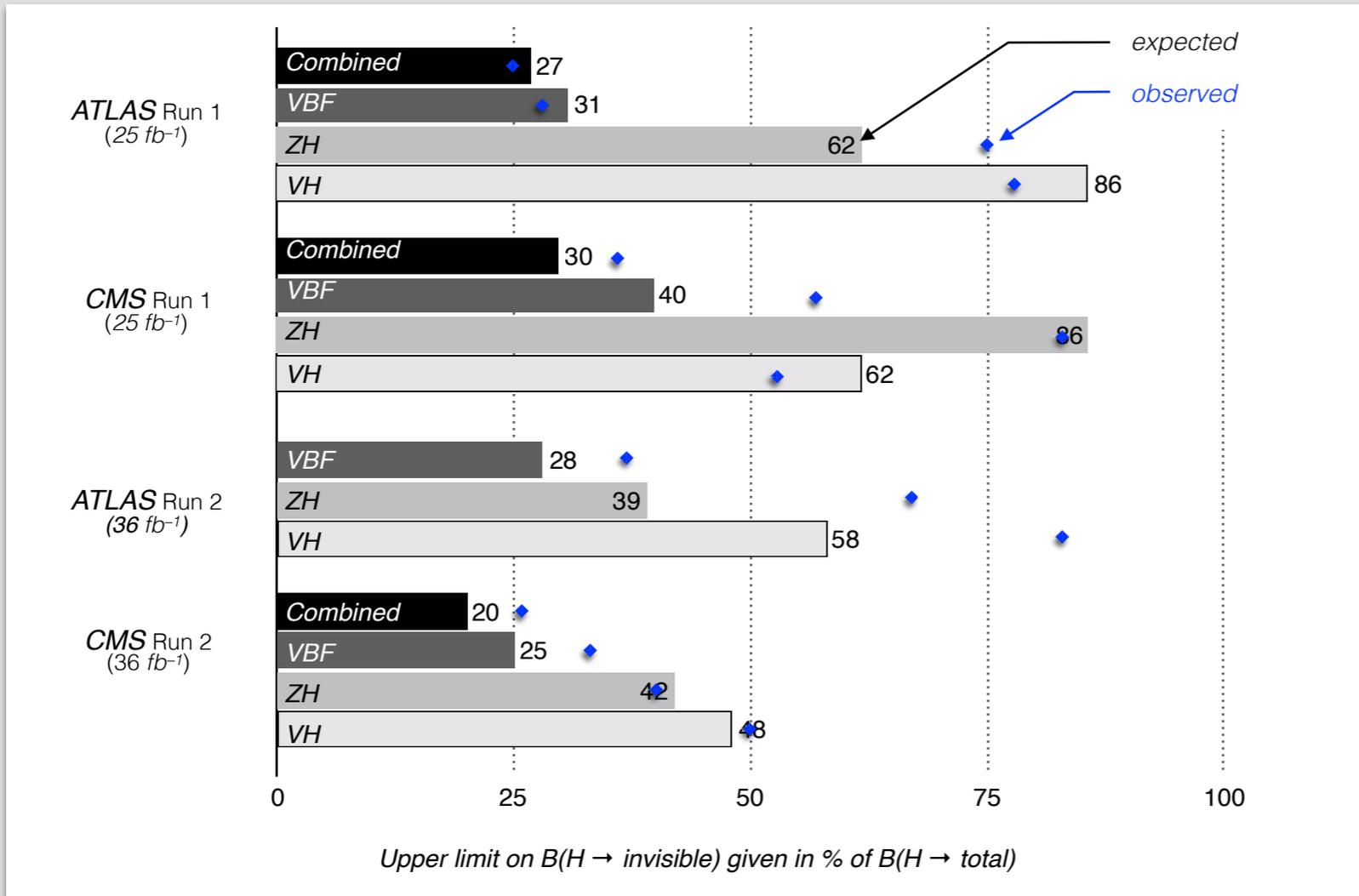
**Observed = 0.33**  
**Expected = 0.25**



# Other Higgs to Invisible

## Invisible Higgs decays comparisons

For upper limits, smaller is better. 95% conf. level. Selected competitive results are shown.



Sources:

J. High Energy Phys.

11 (2015) 206

01 (2016) 172

02 (2017) 135

Phys. Letters B

776(2018)318–337

arXiv

1809.06682

1809.05937

**If a signal is observed, then it will have to be in multiple places!**

# Global Higgs fits

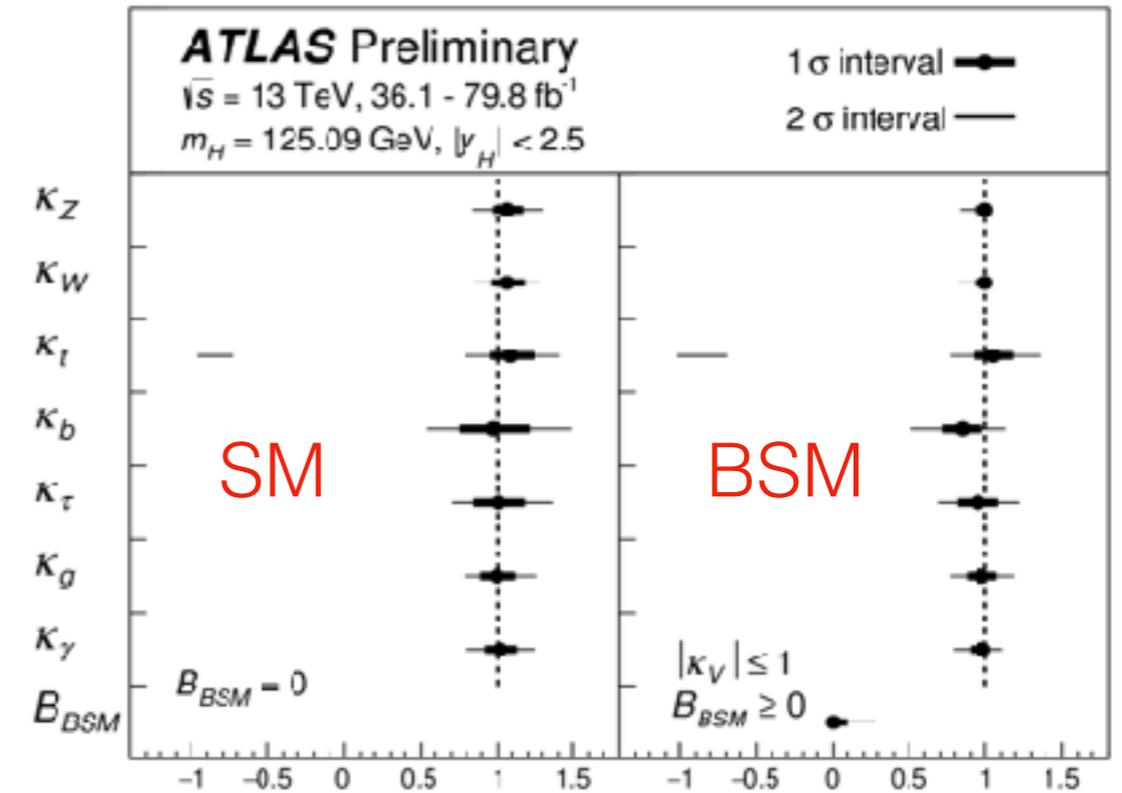
There is another way to constrain Higgs to invisible...

**Use the higgs visible decays!**

Higgs decays would be suppressed if the Higgs have an additional decay mode

$$\sigma(\text{ggH}) \times BR(H \rightarrow WW) = \sigma(\text{ggH}) \frac{\Gamma_{WW}}{\Gamma_{\text{total}}}$$

$$\Gamma_{\text{total}} = \Gamma_{bb} + \Gamma_{WW} + \Gamma_{ZZ} + \dots + \Gamma_{\text{BSM}}$$



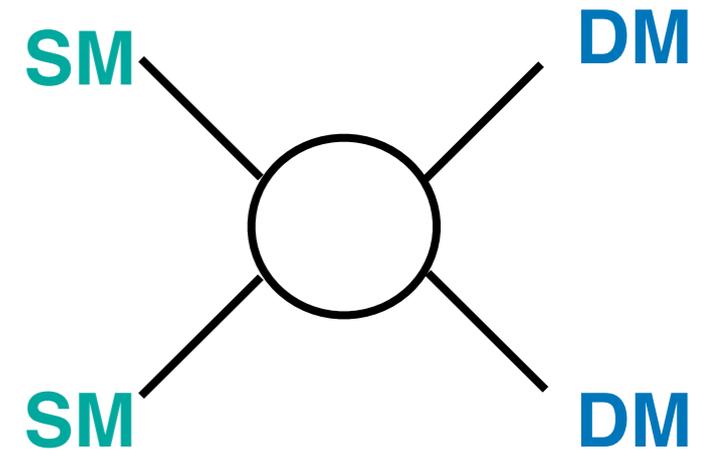
Do a global fit using all the measurements to get limits on the  $k_x$  parameters

Decay channels	$\kappa_i$ assumption	Upper limit on $BR_{\text{inv}}$	
		Obs.	Exp.
Invisible decays	$\kappa_{W,Z,g} = 1$	0.25	0.27
Visible decays	$\kappa_{W,Z} \leq 1$	0.49	0.48
<b>Inv. &amp; vis. decays</b>	<b>None</b>	<b>0.23</b>	<b>0.24</b>
Inv. & vis. decays	$\kappa_{W,Z} \leq 1$	0.23	0.23

Combination of direct searches  
 Indirect limit from visible processes  
 Combination of everything under two assumptions

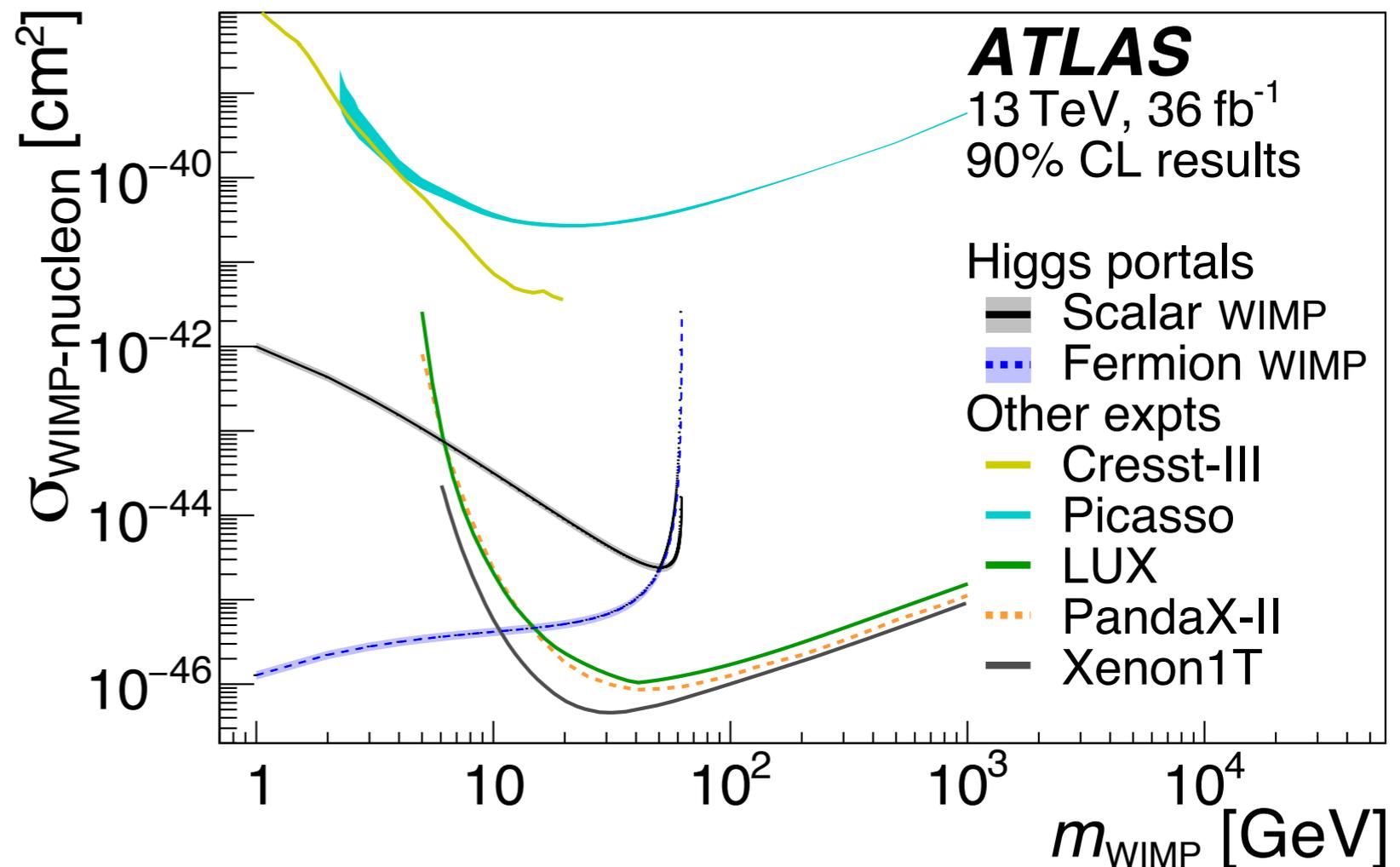
# Higgs Portal Interpretation

**Higgs to invisible provides strong constraint on DM for  $m_{\text{DM}} < m_H/2$**



Strong complementarity with direct dark matter searches:

- H to invisible not sensitive above  $m_H/2$
- Direct DM not sensitive below  $m_{\text{DM}} \sim 10$  GeV
- Overlap in 10 - 60 GeV



Assuming DM couples to SM via Higgs only...

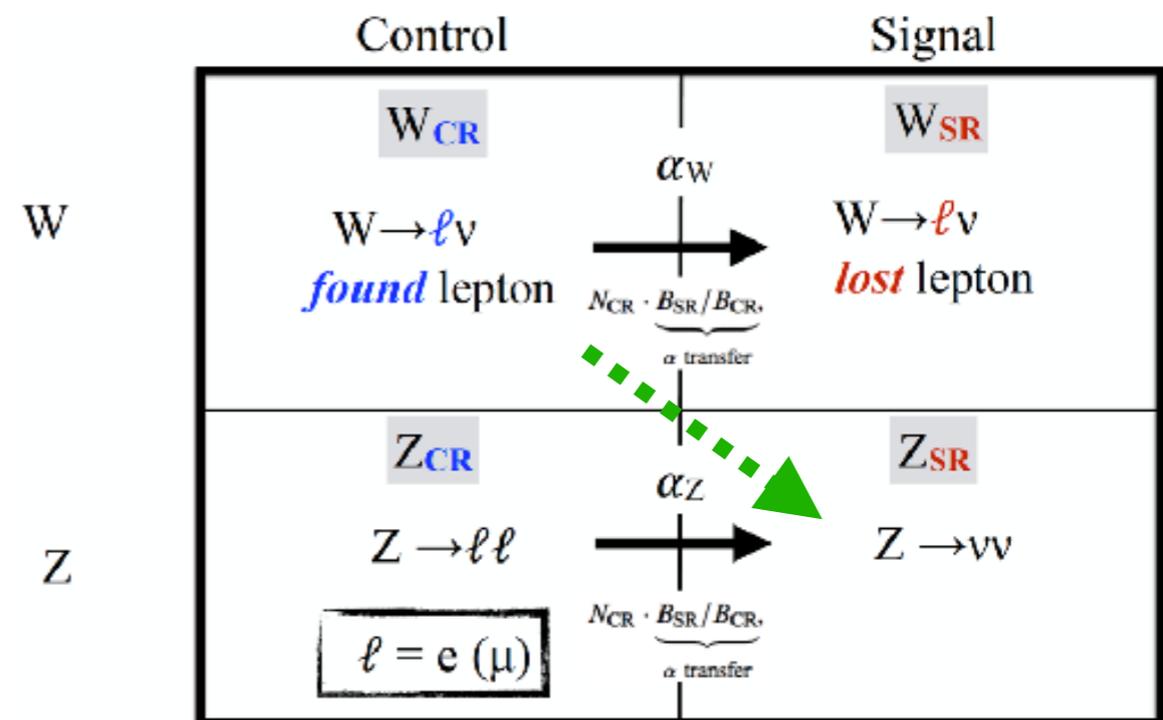
# Extrapolating between W and Z

- $Z > \ell\ell$  statistics is really poor due to small branching ratio ( $\ell\ell:7\%$  vs.  $\nu\nu:20\%$ )
- Using 166  $Z > \ell\ell$  events to model 1111  $Z > \nu\nu$  events!
- W events give x10 more stats.

Why don't we use W's?

We do not know how to correlate theory uncertainties when extrapolating between EWK+QCD W and Z in the presence of non-trivial VBF selection

Description	SR	W CR	Z CR
	Yield [EW]	Yield [EW]	Yield [EW]
$N$ , observed	2252	1602	166
$B$ , expected	2243	1648	183
$Z \rightarrow \nu\nu$	1111 [18%]	-	-



# Theorist of the month

[arXiv:1705.04664](https://arxiv.org/abs/1705.04664)

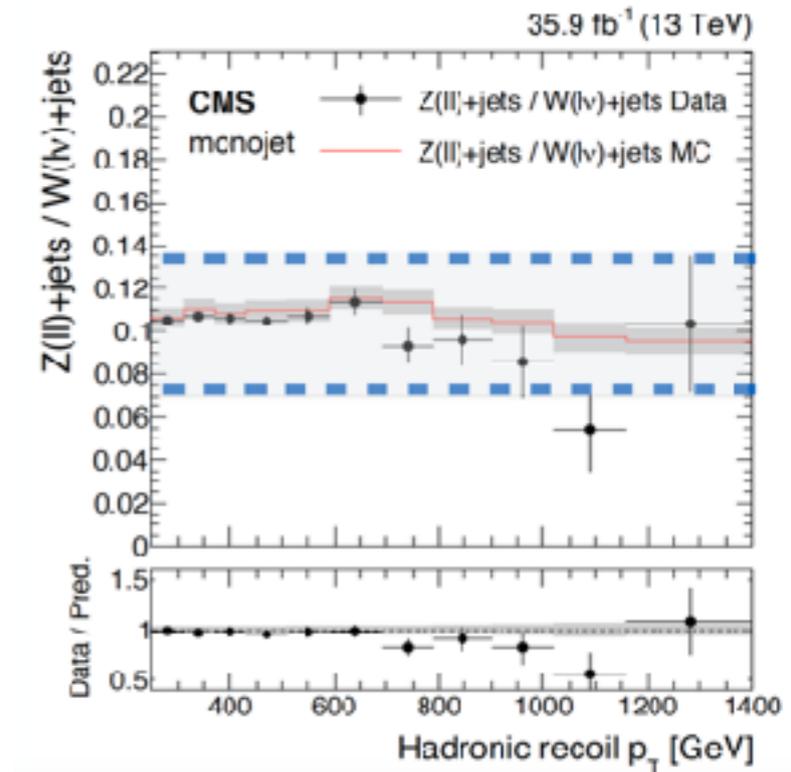
## Precise predictions for $V$ +jets dark matter backgrounds

J. M. Lindert<sup>1</sup>, S. Pozzorini<sup>2</sup>, R. Boughezal<sup>3</sup>, J. M. Campbell<sup>4</sup>, A. Denner<sup>5</sup>, . . .

Combination of state-of-the-art predictions (N)NLO QCD+(N)NLO EW in order to match (future) experimental sensitivities (1-10% accuracy in the few hundred GeV-TeV range)

**Jonas Lindert**

Theorist of the month of October  
**At DESY the week of October 22nd**



# The puzzle of Dark Matter – assembling the pieces

Symposium on joint interpretation of collider, direct and indirect dark matter searches

29-31 October 2018  
DESY, Hamburg



#### Invited speakers

Laura Covi (Göttingen)  
Caterina Doglioni (Lund)  
Hans Kraus (Oxford)  
Greg Landsberg (Brown)  
Javier Rico (UA Barcelona)  
Matthias Schott (Mainz)

Abstract submission welcome  
for Young Scientist Forum!  
Student support available.

Local organizing committee:  
Kelly Beernaert, Katharina Behr,  
Alexander Grohsjean, Sarah Heim,  
Elisa Püschel

[www.desy.de/dm2018](http://www.desy.de/dm2018)

AKADEMIE DER  
WISSENSCHAFTEN  
IN HAMBURG



We would be happy to  
welcome you to our  
**Symposium at DESY**

- Combination and comparisons of Dark Matter searches at **colliders, direct and indirect experiments**
- Including input from **theory**
- Exciting line-up of talks and topics
- Young Scientist Forum
- Dark Matter show at the planetarium (outreach)
- **Registration deadline October 10th**

# Backup

# Fit model

$$B_{\text{SR}}^{\text{best}} = B_{\text{SR}} \cdot \underbrace{N_{\text{CR}}/B_{\text{CR}}}_{k \text{ normalization}} = N_{\text{CR}} \cdot \underbrace{B_{\text{SR}}/B_{\text{CR}}}_{\alpha \text{ factor}}$$

Summary of fit model, each bin of  $m(jj)$  is treated separately according to the model shown here.

	SR	ee	$\mu\mu$	e MET sig. > 4	e MET sig. < 4	$\mu$
<b>Signal</b>	$\mu \times S$					
<b>Z</b>	$k_Z \times B_Z$					
<b>W+jets</b>	$k_W \times B_W$					
<b>mis-ID</b>				$\beta$	$R \times \beta$	

Actually binned  
in charge

$B_{W(Z)}$ : prediction from MC

$R$ : ratio from loose not tight

$\beta$ : normalization of mis-ID component

$m_{jj}$	$k_W$	$k_Z$	$\beta$
1.0 – 1.5 TeV	$0.91 \pm 0.18$	$1.10 \pm 0.27$	$3.63 \pm 1.76$
1.5 – 2.0 TeV	$0.94 \pm 0.18$	$0.98 \pm 0.23$	$4.0 \pm 1.47$
> 2.0 TeV	$1.07 \pm 0.19$	$1.13 \pm 0.27$	$6.8 \pm 2.13$

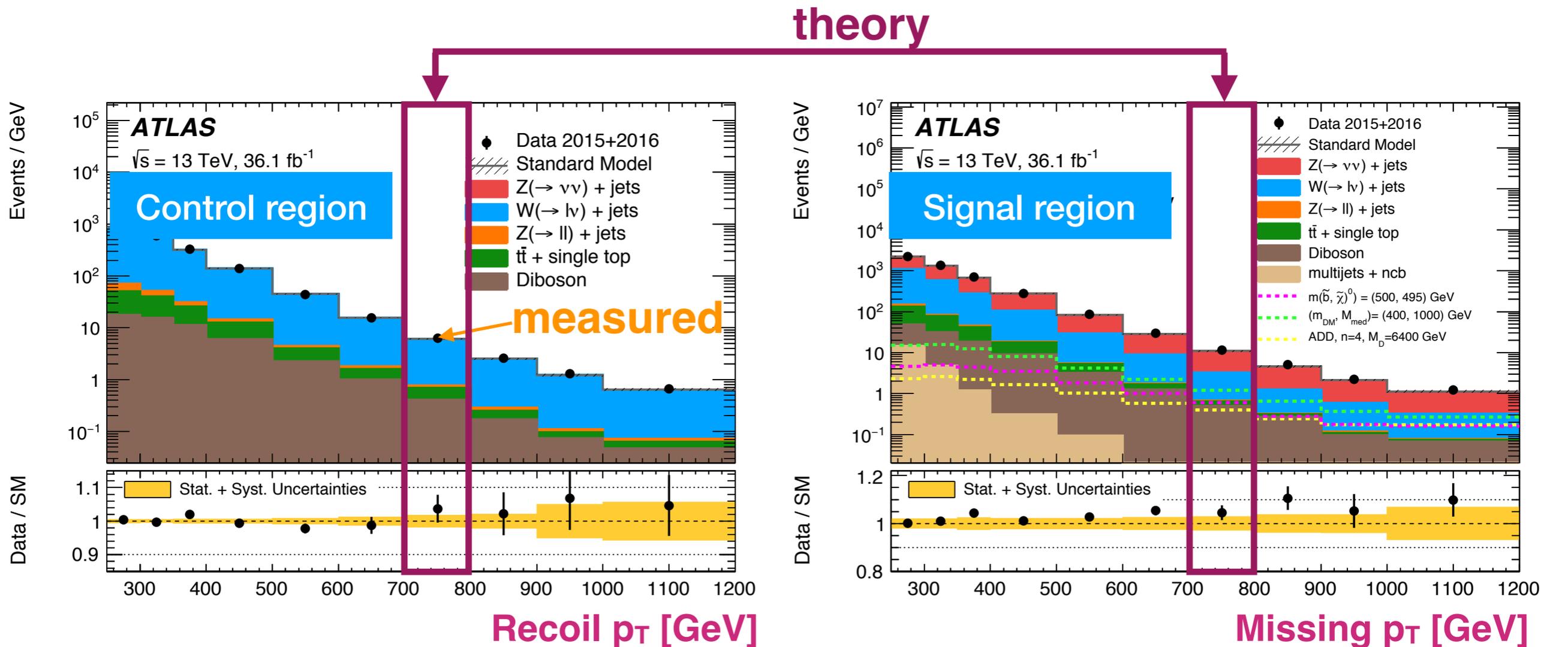
# Yields

Description	SR	W CR	Z CR
	Yield [EW]	Yield [EW]	Yield [EW]
$N$ , observed	2252	1602	166
$B$ , expected	2243	1648	183
$Z \rightarrow \nu\nu$	1111 [18%]	-	-
$Z \rightarrow ee, \mu\mu$	12 [ 9%]	38 [ 9%]	181 [23%]
$Z \rightarrow \tau\tau$	10 [16%]	11 [16%]	-
$W \rightarrow e\nu, \mu\nu$	540 [16%]	1400 [30%]	-
$W \rightarrow \tau\nu$	533 [20%]	130 [34%]	-
Other	36	67	2
$S$ , signal	1070	-	-
VBF	930	-	-
Gluon fusion	140	-	-

	1 - 1.5TeV	1.5 - 2TeV	> 2TeV
<b>Observed</b>	952	667	633
<b>Background</b>	$850 \pm 113$	$660 \pm 90$	$590 \pm 81$
<b>Signal</b>	300	310	460
<b>S/B</b>	0.4	0.5	0.8

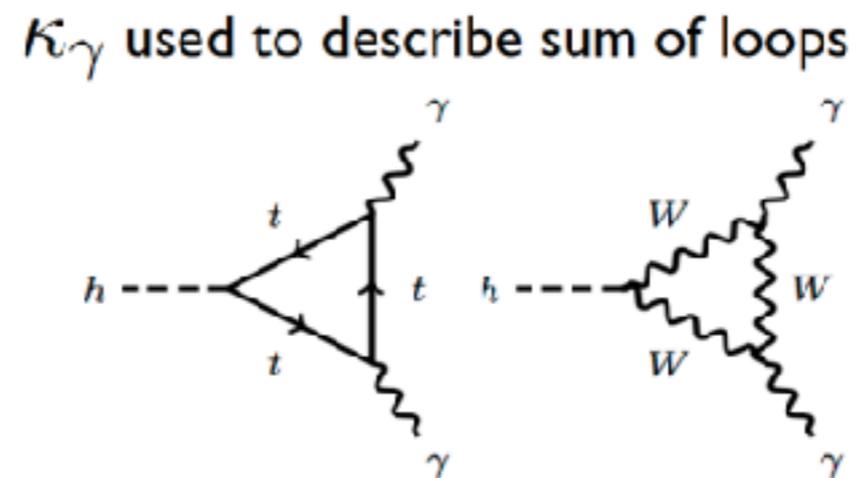
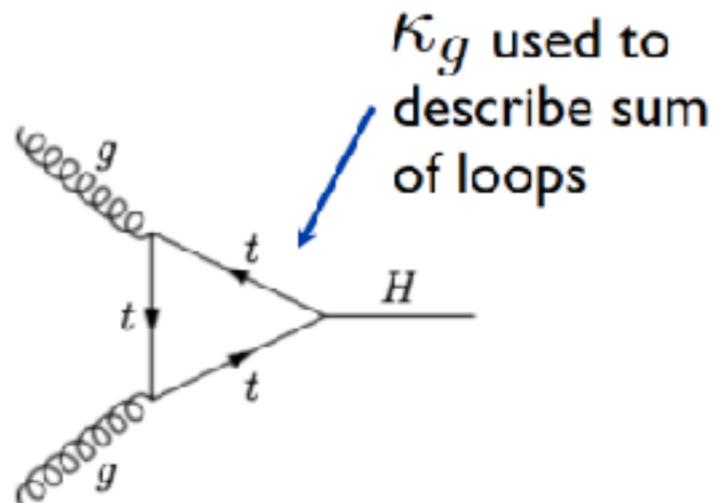
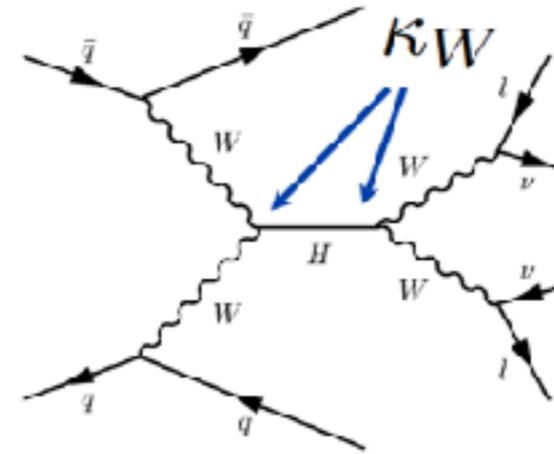
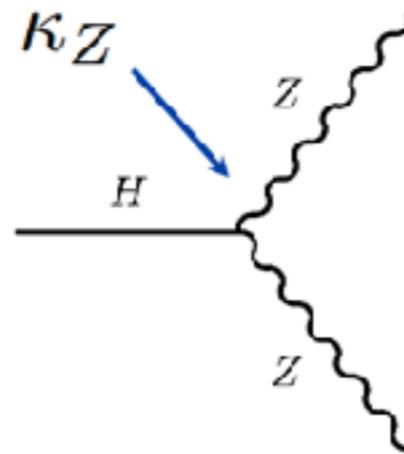
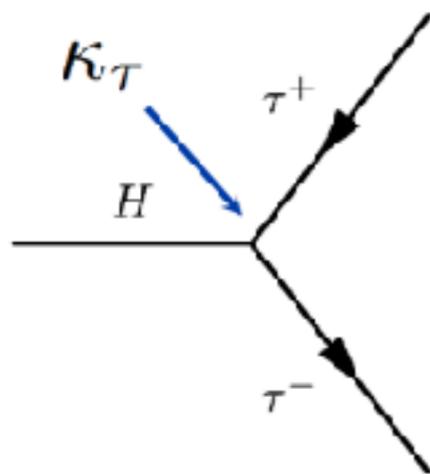
# Background Modelling

$$\frac{d\sigma(V_2)}{dx} = \left[ \frac{d\sigma(V_2)/dx}{d\sigma(V_1)/dx} \right]_{\text{theory}} \times \left[ \frac{d\sigma(V_1)}{dx} \right]_{\text{measured}}$$



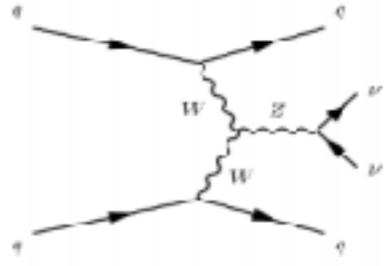
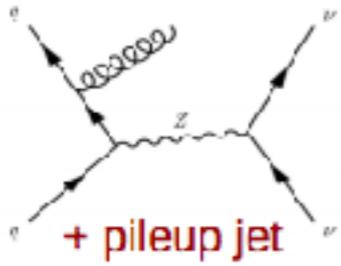
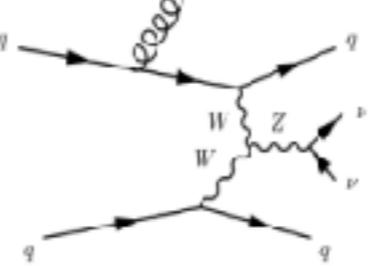
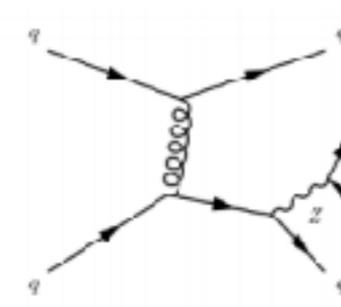
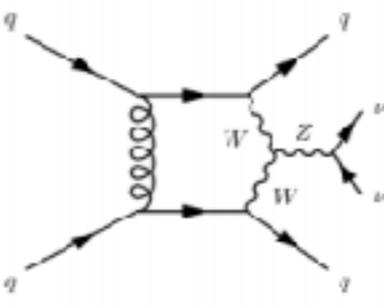
# Global Higgs fits

For each Higgs coupling add a parameter  $\kappa$  which describes its deviation from the SM (SM when  $\kappa_x=1$ )



$$k_h^2 = \Gamma_h / \Gamma_{h,SM} = \sum_j k_j^2 BR_j / (1 - BR_{inv})$$

# W/Z + Jets

Order	QCD	EWK		
	$\alpha_{EW}^2$	$\alpha_{EW}^3$	$\alpha_{EW}^4$	$\alpha_{EW}^5$
$\alpha_s^0$	Negligible in our SR/CR	Doesn't Exist		Interference Only
$\alpha_s^1$	 + pileup jet	Interference Only		Interference Only
$\alpha_s^2$		Interference Only		Higher Order
$\alpha_s^3$		Interference Only	Higher Order	Higher Order

# W/Z + Jets

Order	Sherpa 2.2.1	Madgraph5	Sherpa 2.2.1	$\alpha_{EW}^5$
	$\alpha_{EW}^2$	$\alpha_{EW}^3$	$\alpha_{EW}^4$	
$\alpha_s^0$				
$\alpha_s^1$		Interference Only		
$\alpha_s^2$				
$\alpha_s^3$				

Higher Order →



Higher Order →



# Simulation

CMS

**QCD( $\alpha_{EW}^2$ ):**

MG5\_aMC@NLO at LO in  $\alpha_s$  for up to 4 partons

**EWK( $\alpha_{EW}^4$ ):**MG5\_aMC@NLO at LO in  $\alpha_s$  for up to 2 partons

**Corrections:** apply pT-mjj dependent NLO corrections to the V+jets NLO and EWK

ATLAS

**QCD( $\alpha_{EW}^2$ ):** Sherpa NLO in  $\alpha_s$  for the first 2 partons, LO for up to 4 partons

**EWK( $\alpha_{EW}^4$ ):** Sherpa LO in  $\alpha_s$  for up to 3 partons

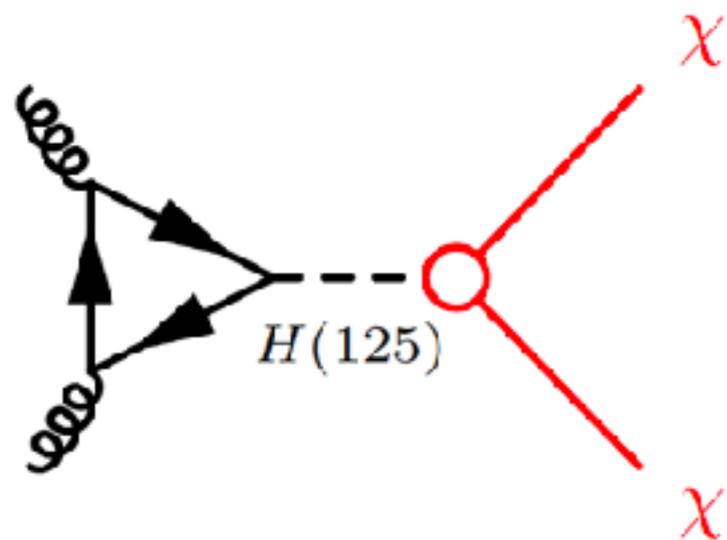
**Interference( $\alpha_{EW}^3$ ):** modelled with MG5\_aMC@NLO

# Comparisons with CMS

Requirement	CMS cut-and-count	CMS Shape	ATLAS	Background
$e(\mu)(\tau)$	< 10 (10) (18) GeV		< 7 (7) (-) GeV	Z(l $\nu$ ), W(l $\nu$ )
Jet $p_T$	> 80 (40) GeV		> 80 (50) GeV	All
Jet Pileup Removal	PFlow		JVT > 0.59	QCD multijet
Third Jet Veto	—		< 25 GeV	ggH
<b>MET</b>	<b>&gt; 250 GeV</b>		<b>&gt; 180 GeV</b>	<b>QCD, tt, W+jets</b>
MHT	—		> 150 GeV	QCD multijet
$\Delta\phi(jj)$	< 1.5		< 1.8	Z( $\nu\nu$ ), W(l $\nu$ )
$\Delta\phi(j, MET)$	0.5		1.0	QCD multijet
<b>m(jj)</b>	<b>&gt; 1300 GeV</b>	<b>&gt; 200 GeV</b>	<b>1-1.15, 1.5-2, &gt;2 TeV</b>	<b>Z(<math>\nu\nu</math>), W(l<math>\nu</math>)</b>
$\Delta\eta(jj)$	> 4	> 1	> 4.8	Z( $\nu\nu$ ), W(l $\nu$ )
$\gamma$ , b-jet	Veto		—	top, $\gamma$ , $V\gamma$

# Search Channels

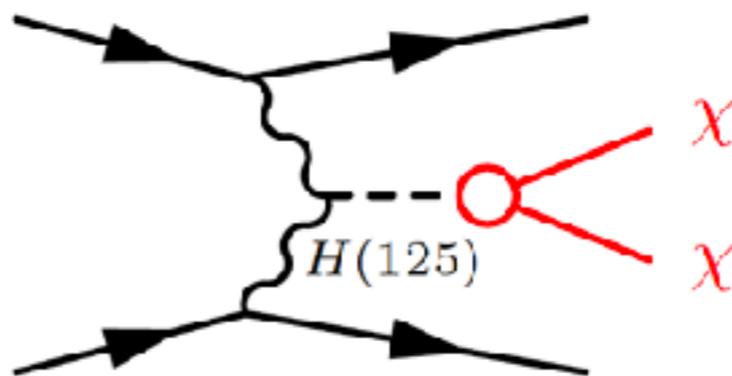
Gluon fusion: 49 pb



- Use an ISR jet for triggering
- Looked for by Mono-jet

**Least sensitive**

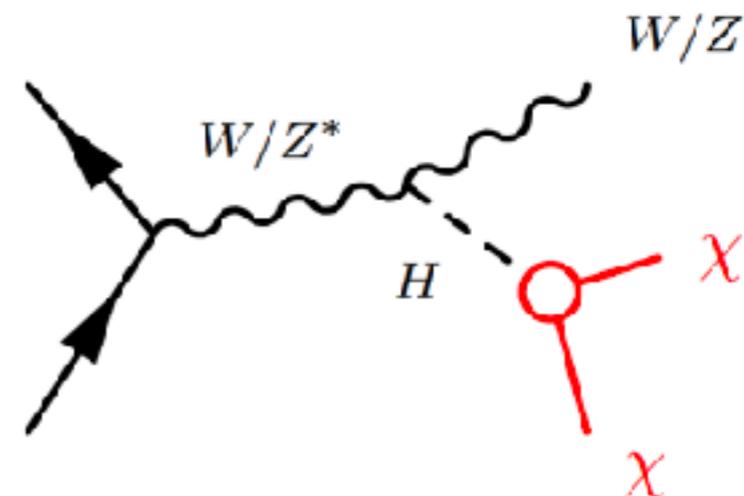
VBF: 3.8 pb



- Tag using forward jets with large  $\Delta\eta(jj)$
- Dedicated analysis

**Most sensitive**

$VH$ : 2.3 pb



- Tag using  $Z(\ell\ell)$  or jet from  $V(\text{had})$
- Looked for by  $V + \text{MET}$

**Intermediate**