

New Limits on Axion-Photon Coupling Constant for Solar Axions

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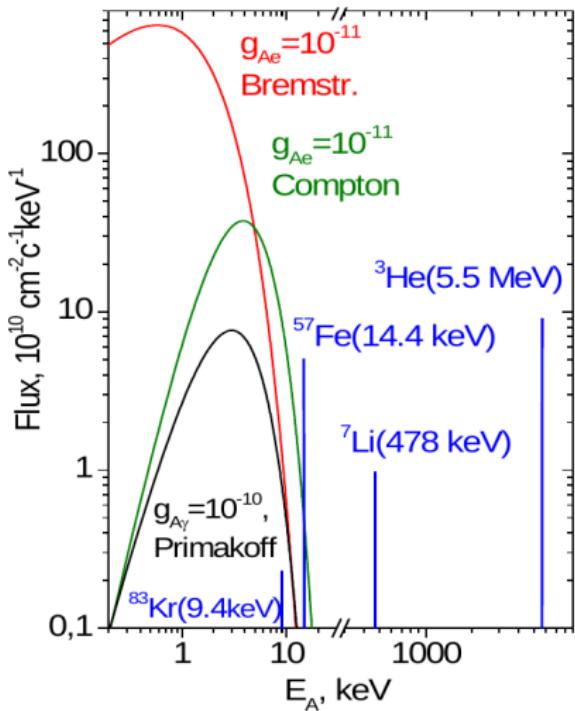
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The 14th Patras Workshop
on Axions, WIMPs and WISPs
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Axion Searches at PNPI

- Series of experiments devoted to Solar Axion Searches
- **Resonant absorption** by atomic nuclei:
 - ^{57}Fe : *A.Derbin et.al, Eur.Phys.J.C62:755-760, 2009*
 - ^{169}Tm : *A.Derbin et.al, Phys.Rev.D83:023505, 2011*
 - ^{83}Kr *Yu. Gavril'yuk et.al, JETP.Lett.107:10, p. 617, 2018*
- **Axio-electric effect:**
 - Si: *A. Derbin et.al, JETP Lett., 95, 379, 2012*
 - BiGeO: *Derbin et.al., Eur.Phys.J.C74:3035, 2014*

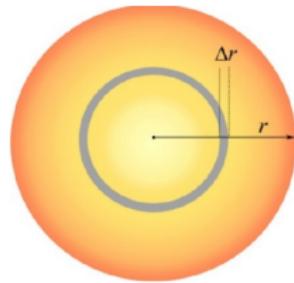
Solar Axion Production



- **Nuclear reactions:**
 - ${}^7\text{Be} + e^- = Li^* + \gamma$;
 $Li^* = Li + A$
(478 keV)
 - $p + d = {}^3\text{He} + A$
(5.5 MeV)
- **Thermal excitation** of M1 nuclear levels
- **Primakoff effect**
- **Axio-recombination**
Bremmstrahlung
Compton

Flux of ^{83}Kr Solar Axions

- Axion flux Φ_A depends on level energy $E = 9.4 \text{ keV}$, temperature distribution $T(r)$, level lifetime $\tau_\gamma = 3.6 \mu\text{s}$, ^{83}Kr solar abundance $N(r)$, and photon/axion emission ratio ω_A/ω_γ



$$\Phi_A = \int N(r) \frac{2 \exp(-E_\gamma/kT(r))}{1 + 2 \exp(-E_\gamma/kT(r))} \frac{\omega_A}{\tau_\gamma \omega_\gamma} dr$$

$$\Phi_A(E_{M1}) = 5.97 \times 10^{23} \left(\frac{\omega_A}{\omega_\gamma} \right) \text{ cm}^{-2} \cdot \text{s}^{-1} \cdot \text{keV}^{-1}$$

Experimental Setup



- Cylindrical proportional counter with **Cu** casing (13 mm).
- Inner diameter: 137 mm
Length: 595 mm
Volume: 8.77 l
- Pressure: 5.6 bar
 ^{83}Kr mass: 101 g
- Setup was equipped with passive shielding: Cu (20 cm), Pb (20 cm), polyethylene (8 cm)

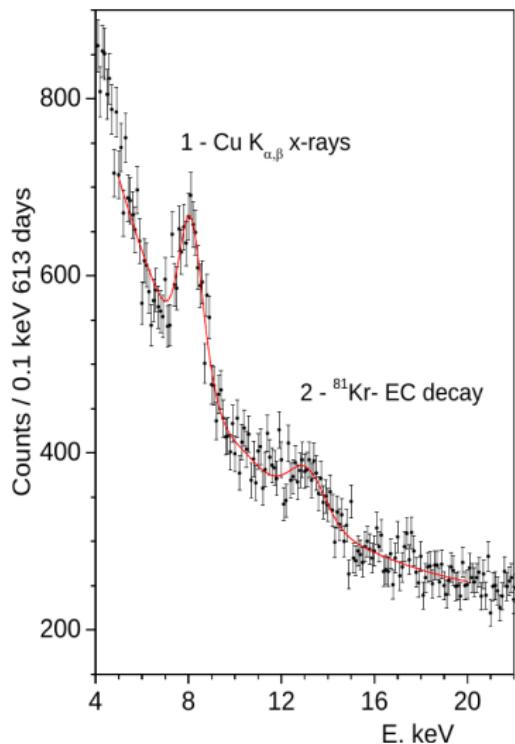
BNO Facility (INR RAS)



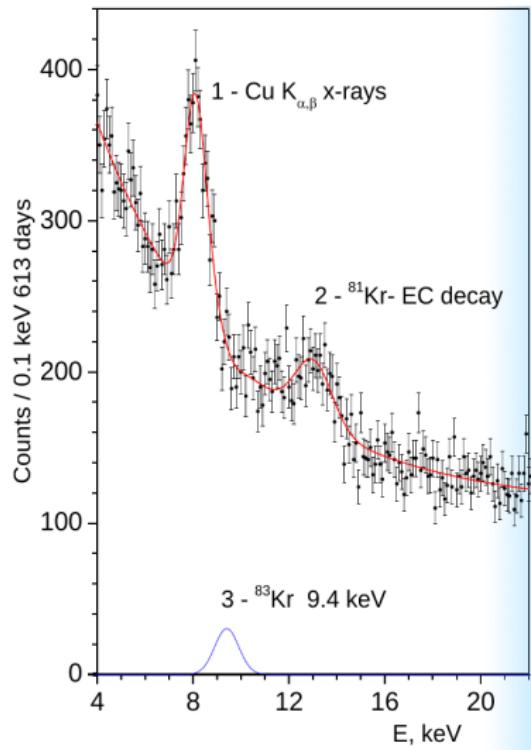
- Underground low-background facility (4200 m.w.e.)
- Muon flux $2.6 \text{ m}^{-2} \cdot \text{d}^{-1}$ — reduced by $\sim 10^7$ times compared to surface rate.

Results from Phase II

Spectra obtained during **613 days** of measurements

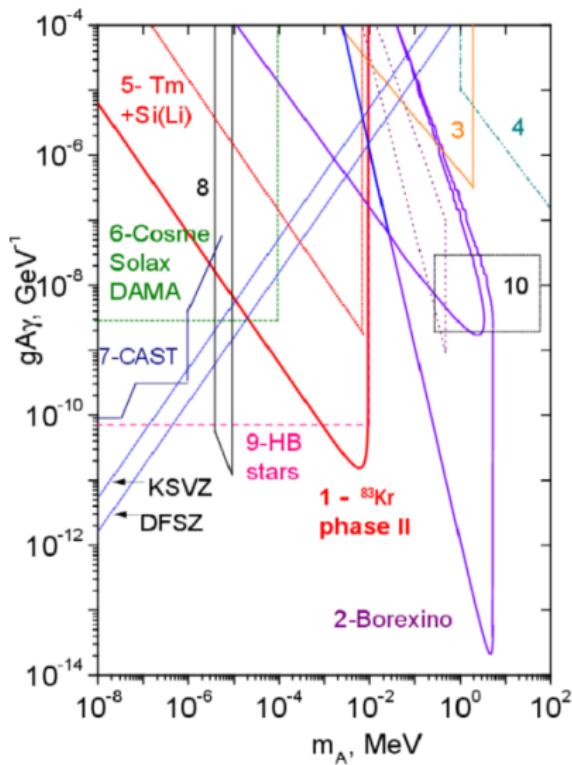


Full spectrum

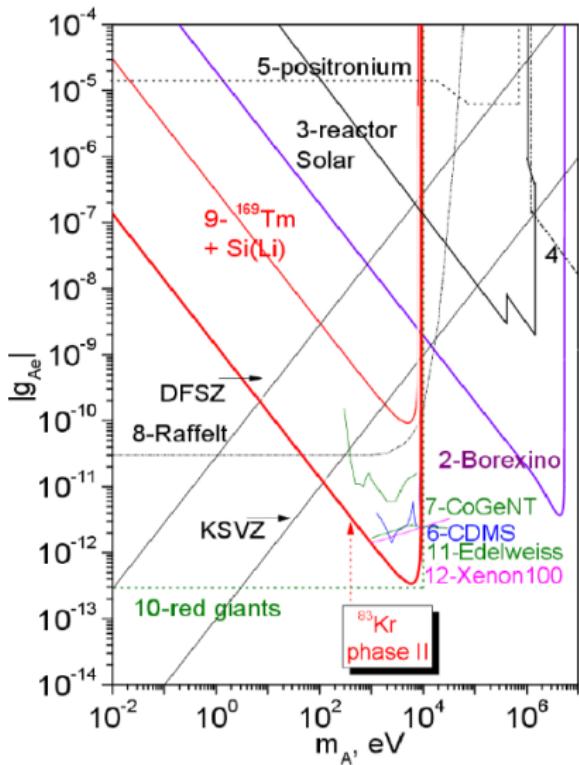


Cuts applied

Obtained Limits (95% c.l.)



$$|g_{A\gamma} \times m_A| \leq 6.26 \times 10^{-8}$$



$$|g_{Ae} \times m_A| \leq 1.38 \times 10^{-9}$$