Optical Ring Cavity Search for

Axion Dark Matter

Ippei Obata (ICRR, University of Tokyo)

based on arXiv: 1805.11753

Collaborated with

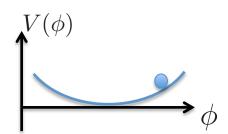
Tomohiro Fujita (Kyoto Univ.)

& Yuta Michimura (KAGRA group, Tokyo Univ.)

Axion (light scalar field) as Dark Matter

Axion evolves according to KG equation

$$\ddot{\phi} + 3H\dot{\phi} + m^2\phi = 0$$
 (background equation)



After m > H (oscillation begins), it starts to oscillate and effectively behaves as a pressureless dust

$$\rho=\frac{1}{2}\dot{\phi}^2+\frac{1}{2}m^2\phi^2\simeq\frac{\rho_0}{a^3}$$

$$P=\frac{1}{2}\dot{\phi}^2-\frac{1}{2}m^2\phi^2\simeq\frac{P_0}{a^3}\sin(2mt)\ \sim 0 \quad \Rightarrow \text{Axion Dark Matter (ADM)}$$

■ Axion dark matter with very low mass can blur the structure on sub-galactic scales.

$$k_J \sim \frac{1}{30~{
m kpc}} \left(\frac{m}{10^{-22} {
m eV}}\right)^{1/2}$$
 $ightarrow$ resolve the "small scale problems"?

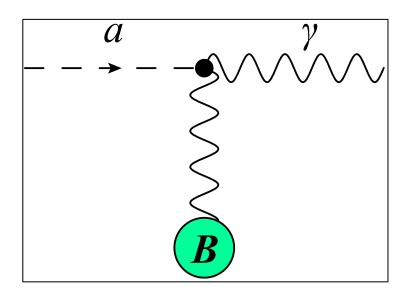
(Jeans length scale)

Search for Axion through photon

(regardless of DM or not)

Axion generally couples to photons via topological term.

$$\mathcal{L}_{a\gamma\gamma} = \frac{g_{a\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu} = g_{a\gamma} a \mathbf{E} \cdot \mathbf{B}$$



Sikivie 1983,...

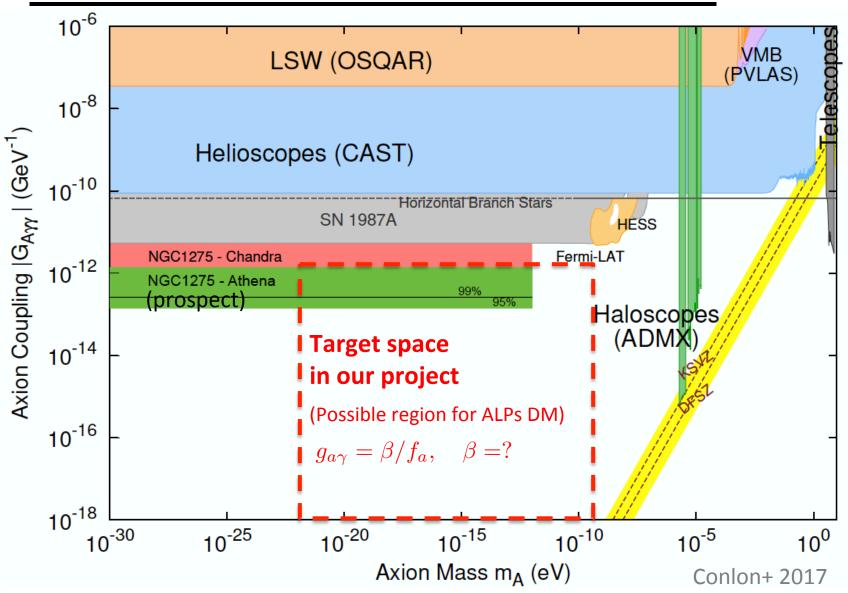
- Axion is converted into photon in the presence of background magnetic field.
- →search for axion-photon convertion with

Laboratory experiments

Astrophysical observations

Axion-photon conversion (Primakoff effect)

Overview of exclusion limits



Axion DM - Photon Coupling (c=1)

Rewriting axion-photon coupling with respect to gauge potential

$$\frac{g_{a\gamma}}{4}aF_{\mu\nu}\tilde{F}^{\mu\nu} = g_{a\gamma}\dot{a}A_i\epsilon_{ijk}\partial_jA_k + (\text{total derivative})$$

Equation of motion for photon:
$$\ddot{A}_i - \nabla^2 A_i + g_{a\gamma} \dot{a} \epsilon_{ijk} \partial_j A_k = 0,$$
 Background axion field:
$$a(t) = a_0 \cos(mt + \delta_\tau)$$

Decomposing photons into two helicity modes, one can obtain

$$\ddot{A}_k^{\pm} + \omega_{\pm}^2 A_k^{\pm} = 0, \\ \omega_{\pm}^2 \equiv k^2 \left(1 \pm \frac{g_{a\gamma} a_0 m}{k} \sin(mt + \delta_{\tau}) \right)$$
 Braking parity symmetry in dispersion relation
$$A_i(t, x) = \sum_{\lambda = \pm} \int \frac{dk}{(2\pi)^3} A_k^{\lambda}(t) e_i^{\lambda}(\hat{k}) e^{ik \cdot x}$$

$$A_i(t, \boldsymbol{x}) = \sum_{\lambda = \pm} \int \frac{d\boldsymbol{k}}{(2\pi)^3} A_{\boldsymbol{k}}^{\lambda}(t) e_i^{\lambda}(\hat{\boldsymbol{k}}) e^{i\boldsymbol{k}\cdot\hat{\boldsymbol{x}}}$$

Estimation of Phase Velocity

Phase velocity of each polarized photon reads

$$c_{\pm} \equiv \frac{\omega_{\pm}}{k} = \left(1 \pm \frac{g_{a\gamma}a_0m}{k}\sin(mt + \delta_{\tau})\right)^{1/2}$$

■ So we get the difference of phase velocity and its amplitude

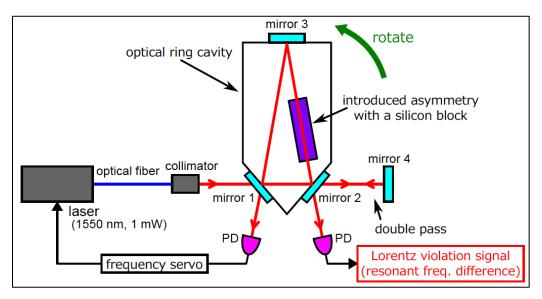
$$\delta c \equiv |c_+ - c_-| = \delta c_0 \sin(mt + \delta_\tau)$$

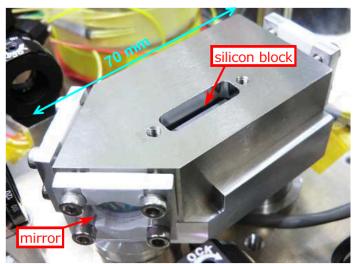
$$\frac{\delta c_0}{c} \simeq 2.7 \times 10^{-24} \left(\frac{g_{a\gamma}}{10^{-12} \text{ GeV}^{-1}} \right) \left(\frac{\lambda}{1550 \text{ nm}} \right) \left(\frac{\rho_a}{0.3 \text{ GeV/cm}^3} \right)^{1/2}$$

Is it possible to search for this small difference by using optical methods?

Optical Ring Cavity Experiments

Bayes+ 2011, Michimura+ 2013,...



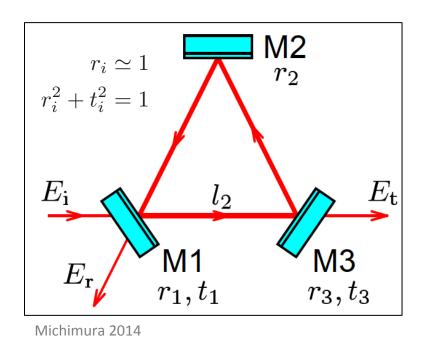


Michimura+ 2016

- It has been proposed to probe the (anisotropic) Lorentz violation signal by measuring the difference of resonant frequency between two optical paths.
- The sensitivity is mainly determined by quantum noise by common noise rejection.

$$\delta c/c \lesssim 10^{-15} \ (L=0.1 \ \mathrm{m}, \ F=10^2, \ P=1 \ \mathrm{mW})$$
 (Michimura+ 2013)

Mechanism of Resonant Cavity



■ The transmittance of cavity is given by

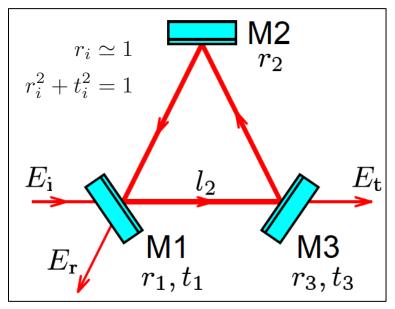
$$\begin{aligned}
& \Phi_2 = l_2 \omega / c_3 \\
E_t &= E_i t_1 t_3 e^{-i\phi_2} + E_i t_1 r_3 r_2 r_1 t_3 e^{-i(\phi + \phi_2)} + \dots \\
&= E_i \frac{t_1 t_3 e^{-i\phi_2}}{1 - r_1 r_2 r_3 e^{-i\phi}} \\
P_t &= |E_t|^2 \\
&= \frac{(t_1 t_3)^2}{(1 - r_1 r_2 r_3)^2 + 4 r_1 r_2 r_3 \sin^2(\phi / 2)} |E_i|^2
\end{aligned}$$

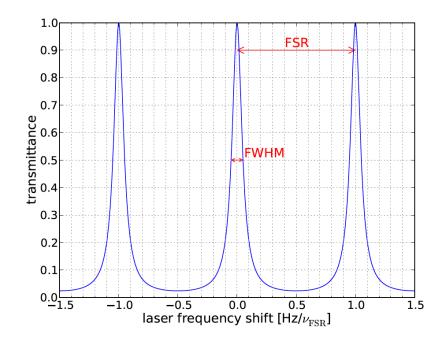
The resonance occurs when the phase of laser beam satisfies

$$\phi = 2\pi m \quad (m=1,2,3,...) \quad \longleftrightarrow \quad \nu = \frac{mc}{L}$$
 (resonant frequency)

At this time, the phase of reflected photon rapidly changes.

Mechanism of Resonant Cavity





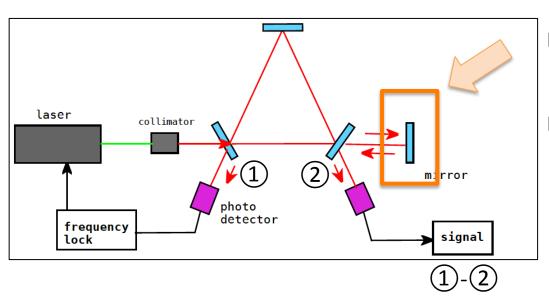
Michimura 2014

(The sharpness of the peak is characterized by the finesse)

$$\mathcal{F} \equiv \frac{\nu_{\text{FSR}}}{\nu_{\text{FWHM}}} = \frac{\pi \sqrt{r_1 r_2 r_3}}{1 - r_1 r_2 r_3} \gg 1 \quad (r \to 1)$$

However, only in the above setup we cannot distinguish the variation of light speed from environmental noises (fluctuations of laser intensity, frequency, cavity length, etc.)

Double Pass Configuration



- We prepare the mirror at far right.
- The transmitted beam is reflected off and goes back to the cavity.

■ We get the signal from two counterpropagating optical paths and measure its difference.

Most environmental noises change two resonant frequencies coherently

→The difference of two signals will reject these common noise fluctuations (Quantum shot noise becomes the primary source of noise)

Reconsidering AxionDM-Photon...

Phase velocity of each polarized photon reads

$$c_{\pm} \equiv \frac{\omega_{\pm}}{k} = \left(1 \pm \frac{g_{a\gamma}a_0m}{k}\sin(mt + \delta_{\tau})\right)^{1/2}$$

So we get the difference of phase velocity and its amplitude

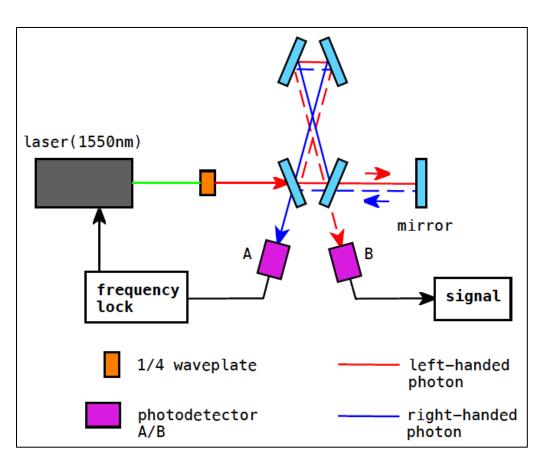
$$\delta c \equiv |c_+ - c_-| = \delta c_0 \sin(mt + \delta_\tau)$$

$$\frac{\delta c_0}{c} \simeq 2.7 \times 10^{-24} \left(\frac{g_{a\gamma}}{10^{-12} \text{ GeV}^{-1}} \right) \left(\frac{\lambda}{1550 \text{ nm}} \right) \left(\frac{\rho_a}{0.3 \text{ GeV/cm}^3} \right)^{1/2}$$

✓ We apply the optical ring cavity measurement to the detection of axion DM-photon coupling by improving previous experimental setup!

Our Cavity Design

1805.11753



✓ This basement can measure
the difference of resonant
frequencies between left and
right-handed polarized photon.

$$\frac{\delta \nu}{\nu} = \frac{\delta c}{c} = 3 \times 10^{-24} \left(\frac{g_{a\gamma}}{10^{-12} \text{GeV}^{-1}} \right) \sin(mt + \delta_{\tau})$$

- ✓ A null-experiment sensitive to the axion-photon coupling.
- ✓ Due to a double pass configuration, common fluctuation noises become irrelevant.

Sensitivity curve is primarily given by shot noise : $\sqrt{S_{
m shot}} = \sqrt{\frac{\lambda}{4\pi P}\left(\frac{1}{t_{
m m}^2} + \omega^2\right)}, \quad t_r = \frac{L\mathcal{F}}{\pi}$

Sensitivity Estimation

■ Signal-to-Noise ratio (SNR) is given by

$$\delta c \equiv |c_{+} - c_{-}| = \delta c_0 \sin(mt + \delta_{\tau})$$

(SNR) =
$$\frac{\sqrt{T}}{2\sqrt{S_{\text{shot}}}} \frac{\delta c_0}{c} \quad \sqrt{S_{\text{shot}}} = \sqrt{\frac{\lambda}{4\pi P} \left(\frac{1}{t_r^2} + \omega^2\right)}, \quad t_r = \frac{L\mathcal{F}}{\pi}$$

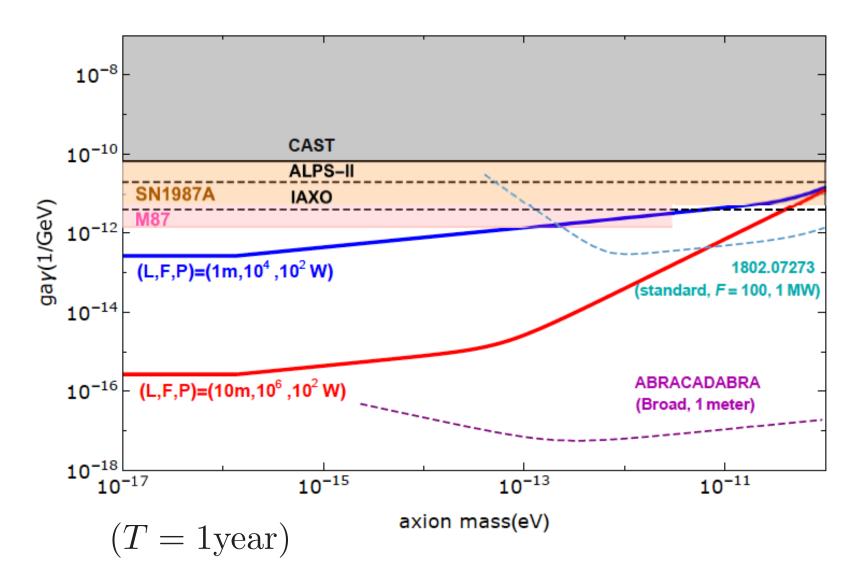
until
$$T < au = rac{2\pi}{mv^2} \sim 1 \mathrm{yr} \left(rac{10^{-16} \mathrm{eV}}{m}
ight)$$
 (coherent time of ADM)

- lacksquare If T is longer than au, SNR is modified as ${
 m (SNR)}=rac{(T au)^{1/4}}{2\sqrt{S_{
 m shot}}}rac{\delta c_0}{c}$
- ✓ Finally the sensitivity of axion-photon coupling leads to

$$g_{a\gamma} \lesssim \frac{2}{2.7 \times 10^{-12}} \sqrt{\frac{S_{\text{shot}(\omega)}}{T}} \text{ (GeV)}^{-1} \quad (T < \tau)$$
$$\lesssim \frac{2}{2.7 \times 10^{-12}} \sqrt{\frac{S_{\text{shot}(\omega)}}{(T\tau)^{1/2}}} \text{ (GeV)}^{-1} \quad (T \ge \tau)$$

Projected Sensitivity Reach

1805.11753



Summary & Outlook

- We proposed a novel experiment to probe a coupling of ADM to photon by using optical ring cavity.
- We prepared two sets of experimental parameters and demonstrated that both of them can reach sensitivities beyond the current constraints.
- We have to investigate the technical noises at low frequency since the common mode rejection ratio is finite in actually.
- ☐ We're starting to operate the cavity with feasible parameters.