

A global study of the extended scalar singlet model

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Introduction

- A net imbalance between matter (baryons) and antimatter (anti-baryons) number densities as

$$\eta \equiv \frac{n_B - n_{\bar{B}}}{n_\gamma} \approx 10^{-10}. \quad (1)$$

- Theories of baryogenesis try to dynamically generate the asymmetry from symmetric initial conditions in the early universe.
- Electroweak baryogenesis (EWBG): baryogenesis via a strong first-order electroweak phase transition (EWPT) around $T \sim 100$ GeV.
- With $m_h = 125$ GeV, EWPT is *not* first-order in the Standard Model (SM) \rightarrow require new physics beyond the SM (BSM).
- A simple \mathbb{Z}_2 symmetric scalar Higgs portal *cannot* simultaneously account for the observed dark matter (DM) abundance and matter-antimatter asymmetry.
- We study the phenomenology of the extended scalar singlet model with a fermionic DM candidate.

Singlet fermion dark matter model

- Introduce a new real scalar singlet S and a Dirac fermion DM field ψ via

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_S + \mathcal{L}_\psi + \mathcal{L}_{\text{portal}}, \quad (2)$$

where \mathcal{L}_{SM} is the SM Lagrangian,

$$\mathcal{L}_S = \frac{1}{2}(\partial_\mu S)(\partial^\mu S) + \frac{1}{2}\mu_S^2 S^2 + \frac{1}{3}\mu_3 S^3 - \frac{1}{4}\lambda_S S^4, \quad (3)$$

$$\mathcal{L}_\psi = \bar{\psi}(i\cancel{\partial} - \mu_\psi)\psi - g_S \bar{\psi}\psi S, \quad (4)$$

$$\mathcal{L}_{\text{portal}} = -\mu_{\Phi S} \Phi^\dagger \Phi S - \frac{1}{2}\lambda_{\Phi S} \Phi^\dagger \Phi S^2. \quad (5)$$

- A linear term $\mu_1^3 S$ is removed by a constant shift $S \rightarrow S + \sigma$.
- Both $\phi \equiv \langle \Phi^0 \rangle$ and S can acquire non-zero VEVs. We define them as

$$\langle \phi \rangle \equiv v_0 = 246.22 \text{ GeV}, \quad \langle S \rangle = s_0. \quad (6)$$

- After electroweak symmetry breaking (EWSB), $\mathcal{L}_{\text{portal}}$ induces a mixing between $\varphi = \phi - v_0$ and $s = S - s_0 \rightarrow$ physical mass eigenstates are

$$\begin{pmatrix} h \\ H \end{pmatrix} = \begin{pmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} \varphi \\ s \end{pmatrix}, \quad (7)$$

where α is the mixing angle.

- For the tree-level potential to be bounded from below, we require

$$\lambda_\phi > 0, \quad \lambda_S > 0, \quad \lambda_{\Phi S} > -2\sqrt{\lambda_\phi \lambda_S}. \quad (8)$$

- After EWSB, the fermion DM Lagrangian is

$$\mathcal{L}_\psi = \bar{\psi}(i\cancel{\partial} - m_\psi)\psi - g_S \bar{\psi}\psi s, \quad (9)$$

where $m_\psi = \mu_\psi + g_S s_0$.

- With $m_h = 125.13$ GeV, the model contains 7 free parameters

$$m_H, \quad s_0, \quad \mu_3, \quad \lambda_S, \quad \alpha, \quad m_\psi, \quad g_S. \quad (10)$$

Constraints

- Thermal relic density:** require $\Omega_\psi h^2 \leq 0.1188$ from *Planck*. Likelihood is one-sided Gaussian.
- Direct detection:** require $\sigma_{\text{SI}}^{\text{eff}} \leq \sigma_{\text{PandaX-II}}$ where $\sigma_{\text{SI}}^{\text{eff}}$ = effective DM-nucleon cross-section and $\sigma_{\text{PandaX-II}} = 90\%$ C.L. upper limit from PandaX-II. Likelihood is one-sided Gaussian.
- Electroweak baryogenesis:** require $v_c/T_c \geq 0.6$ where $v_c =$ Higgs VEV at the critical temperature T_c . Likelihood is one-sided Gaussian.
- Electroweak precision observables (EWPO):** require

$$\Delta S = 0.04 \pm 0.11, \quad (11)$$

$$\Delta T = 0.09 \pm 0.14, \quad (12)$$

$$\Delta U = -0.02 \pm 0.11. \quad (13)$$

Likelihood is a correlated 3D Gaussian.

- Direct Higgs searches:** constraints from the LEP, Tevatron and the LHC. Likelihood is computed using the *HiggsBounds_v4.3.1* package.
- Higgs signal strengths:** constraints from the Higgs signal strength and mass measurements performed at the LHC. Likelihood is computed using the *HiggsSignals_v1.4.0* package.

Preliminary results

Parameter Minimum Maximum Prior type

| Parameter | Minimum | Maximum | Prior type |
|-------------|-----------|---------|------------|
| m_H | 10 GeV | 10 TeV | log |
| s_0 | -1 TeV | 1 TeV | flat |
| μ_3 | -1 TeV | 1 TeV | flat |
| λ_S | 10^{-3} | 10 | log |
| α | 0 | $\pi/2$ | flat |
| m_ψ | 10 GeV | 10 TeV | log |
| g_S | 10^{-3} | 10 | log |

Table 1: Ranges and priors for the free parameters of our model.

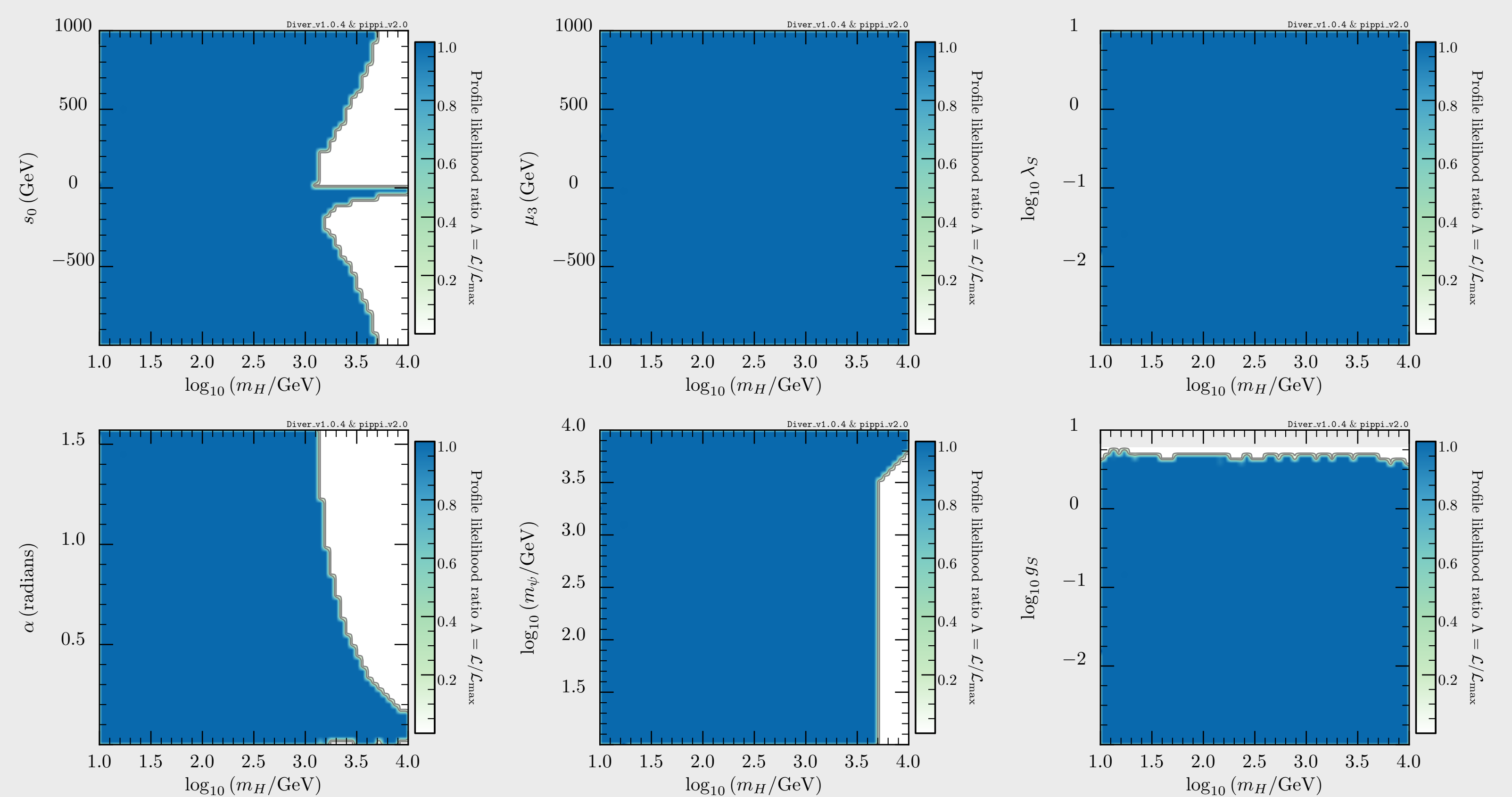


Fig. 1: 7D scan of the model parameter space using *only* the EWBG constraint.

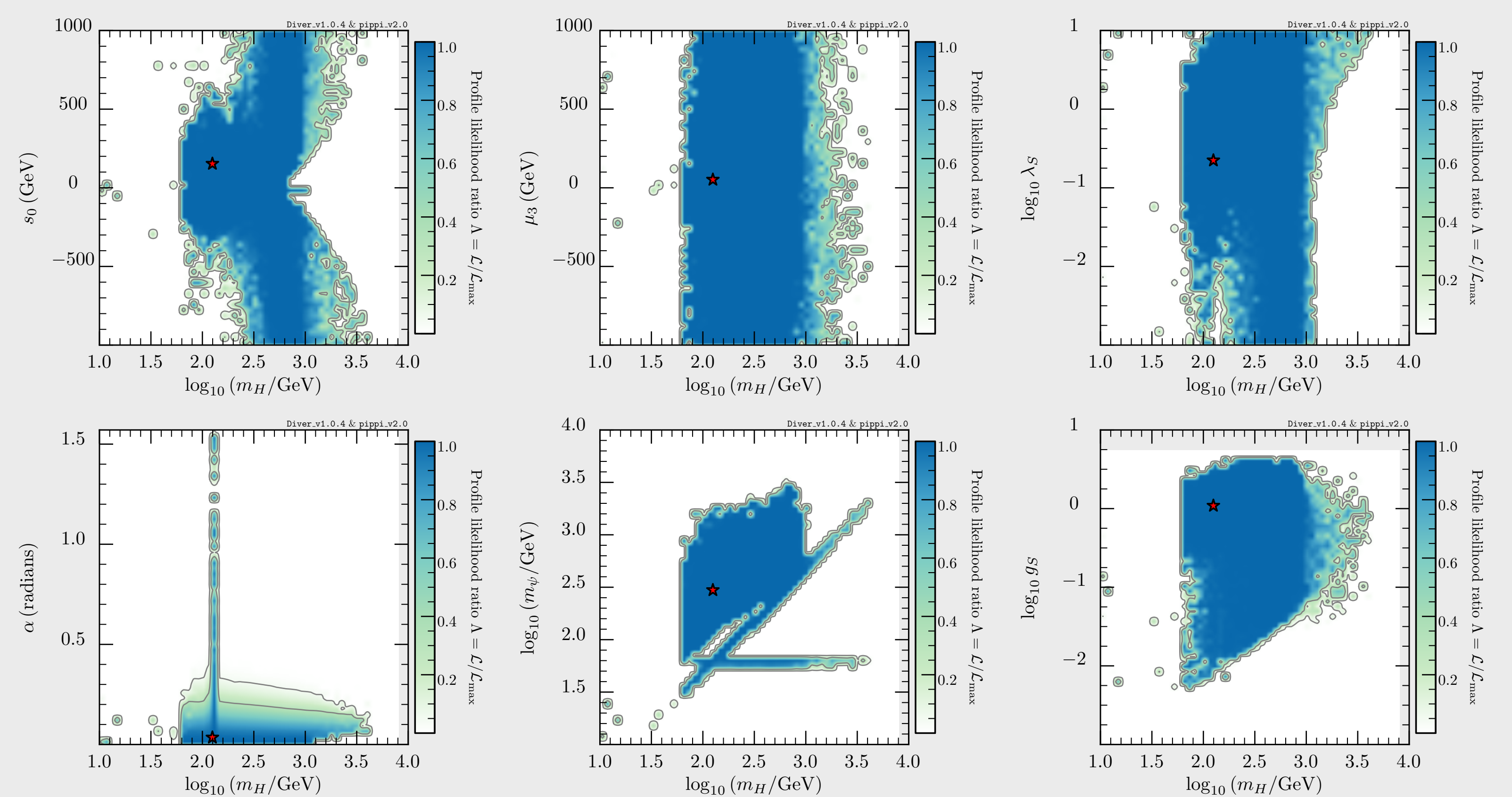


Fig. 2: 7D scan of the model parameter space using all available constraints.

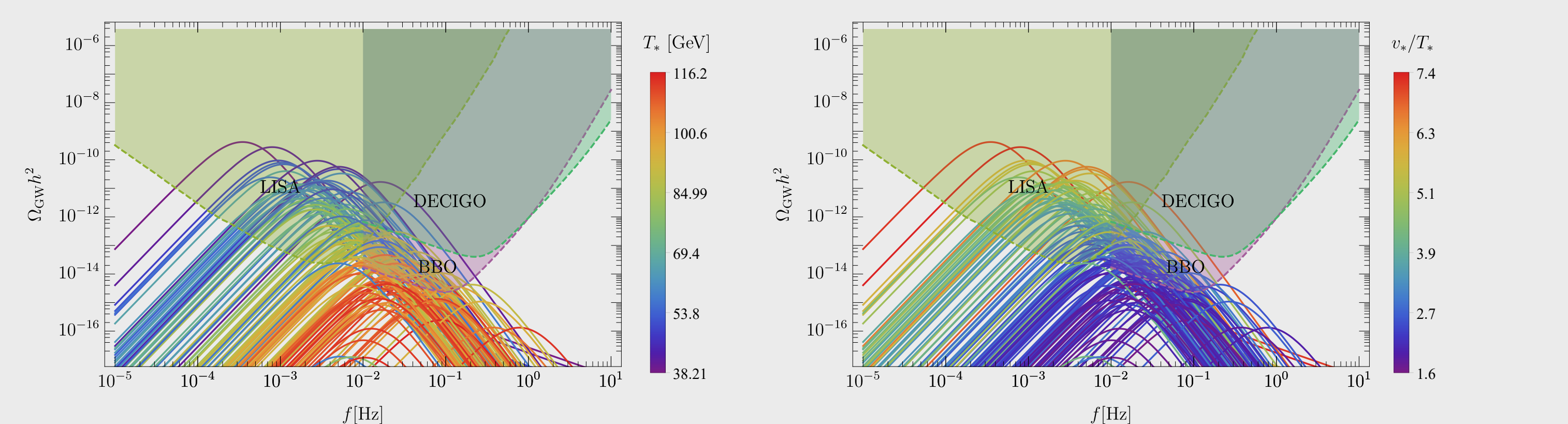
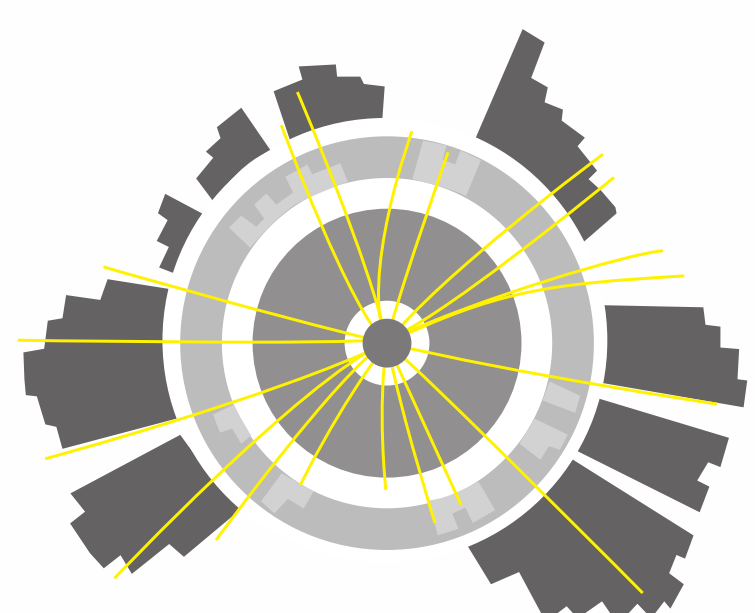


Fig. 3: Gravitational wave (GW) spectra of viable points along with the projected sensitivities of future GW experiments such as LISA, DECIGO and BBO.

Conclusions

- A successful EWBG is viable in all parts of the model parameter space provided $g_S \lesssim 5.62$.
- Combined constraints from the DM relic density, PandaX-II limit, EWBG, EWPO and Higgs searches at colliders places an upper limit on m_H , m_ψ and g_S .
- GW spectra of viable points are often within reach of future GW experiments.



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