

Leptophilic Dark Matter from Gauged Lepton Number

Phenomenology and Gravitational Wave Signatures

arXiv:1809.09110

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Gauged Lepton Number

[Schwaller, Tait, Vega-Morales (2013)]

SM + RH- ν

- + $U(1)_\ell$ lepton number gauge group $\implies Z'$ gauge boson
- + spontaneous lepton number breaking \implies scalar field ϕ
- + anomaly cancellation \implies 4 exotic leptons
DM candidate

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Phenomenology

- LEP: $m_{Z'} > 200 \text{ GeV}$, $v_\Phi > 1880 \text{ GeV}$
- LHC: Z' searches, Higgs measurements,
exotic lepton searches
- DM: relic density, direct/indirect detection
- 1st order phase transition \implies GW

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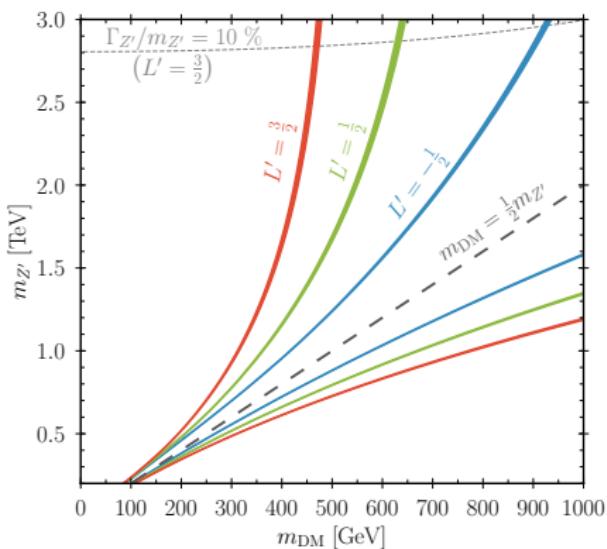
Phenomenology

- LEP: $m_{Z'} > 200 \text{ GeV}$, $v_\Phi > 1880 \text{ GeV}$ $\implies v_\Phi = 2 \text{ TeV}$
- LHC: Z' searches, Higgs measurements, exotic lepton searches \leftarrow see 1809.09110
- DM: relic density, direct/indirect detection \leftarrow next slide
- 1st order phase transition \implies GW \leftarrow main focus

Dark Matter

DM candidate: mostly SM singlet, chiral couplings to Z'

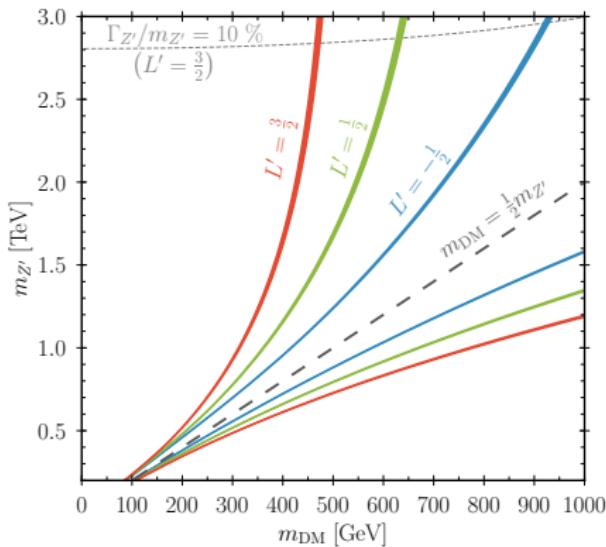
Relic Density



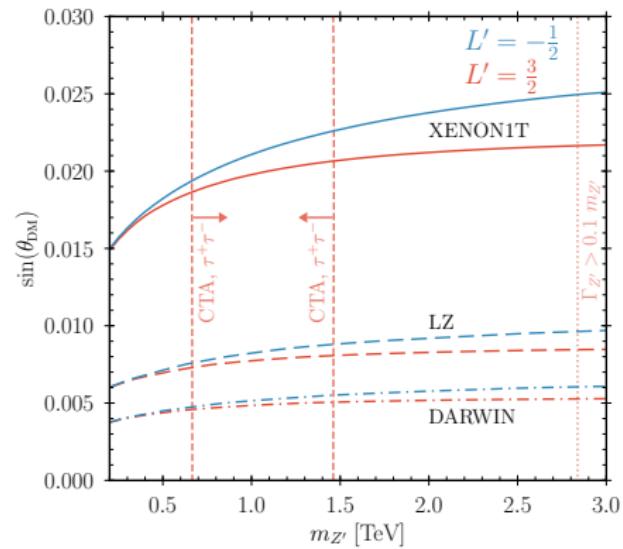
Dark Matter

DM candidate: mostly SM singlet, chiral couplings to Z'
doublet admixture: $\nu_{\text{DM}} = \cos \theta_{\text{DM}} \nu_S + \sin \theta_{\text{DM}} \nu_D$

Relic Density



Direct Detection



Cosmological Phase Transitions

finite- T corrections restore symmetry at high T

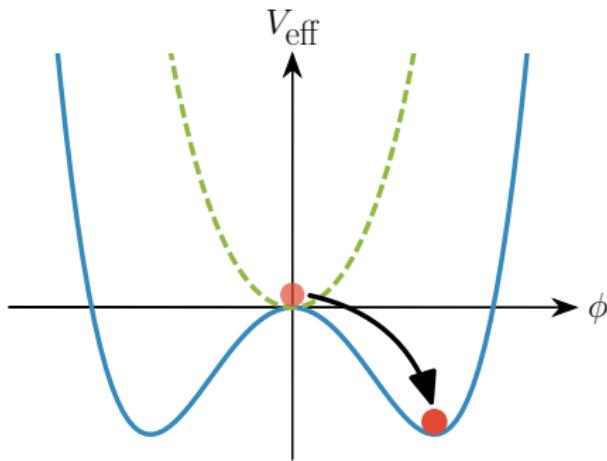
\implies symmetry breaking phase transition in the early Universe

Cosmological Phase Transitions

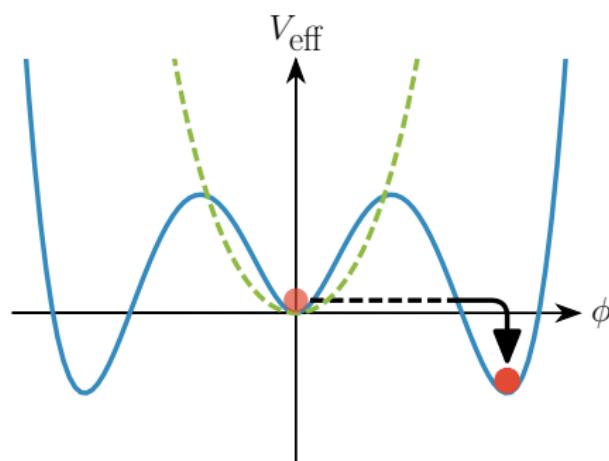
finite- T corrections restore symmetry at high T

⇒ symmetry breaking phase transition in the early Universe

cross-over:



1st-order:

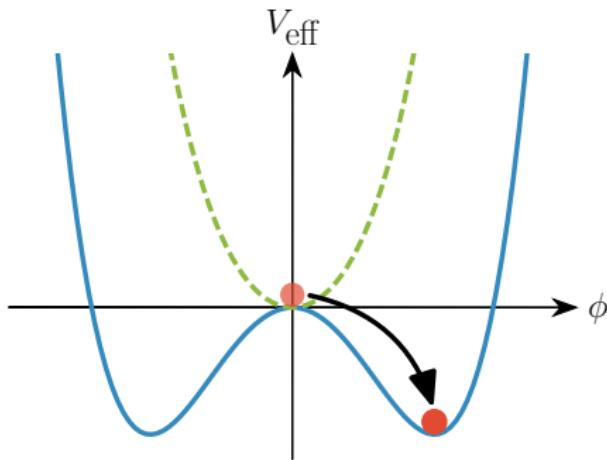


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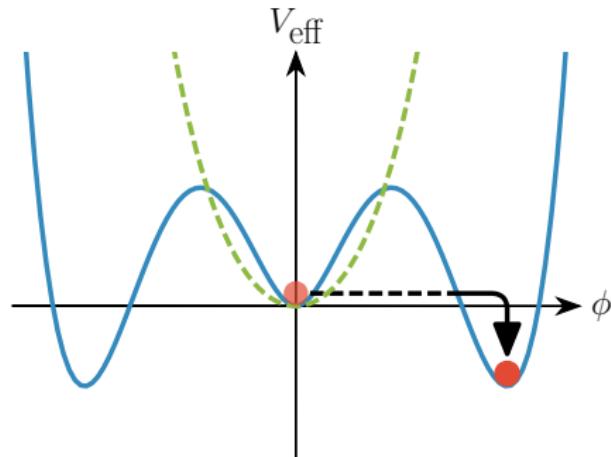
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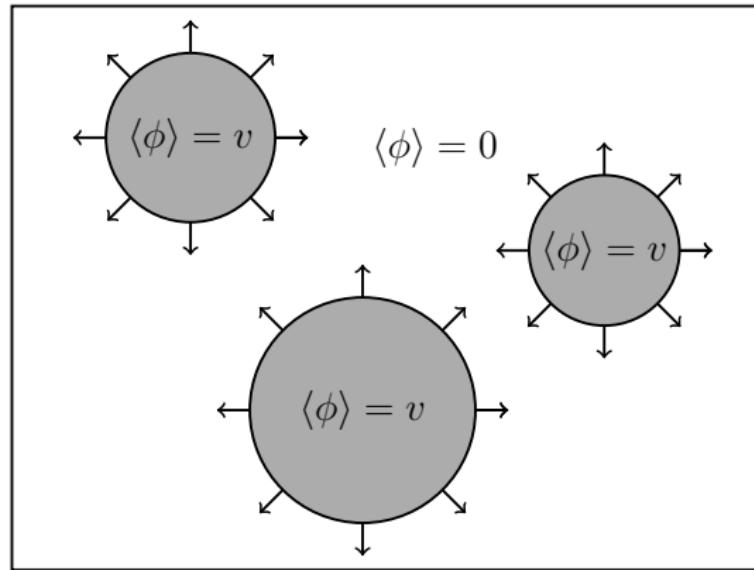
Gravitational Waves only from 1st-order Transition!

1st-Order Phase Transition

high- and low- T minima separated by barrier

⇒ 1st-order PT via tunneling

⇒ bubble nucleation

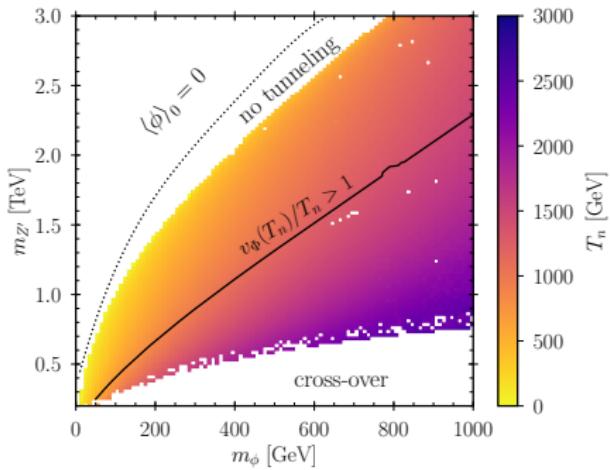


Nucleation Temperature

- nucleation rate \longleftrightarrow Hubble expansion
 $\Gamma(T)$ \longleftrightarrow $H(T)$
- nucleation temperature (T_n): $\Gamma/H^4 \sim 1$

Nucleation Temperature

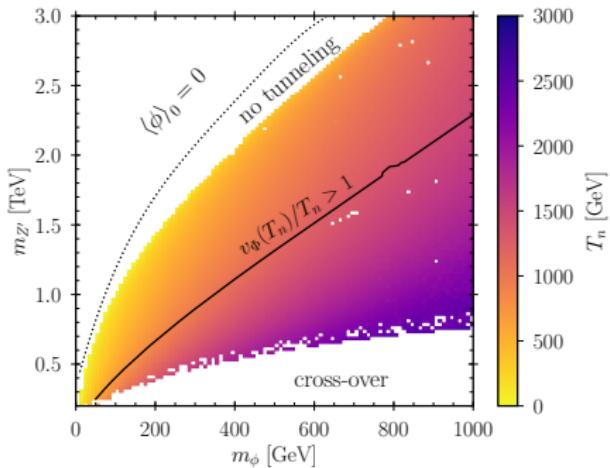
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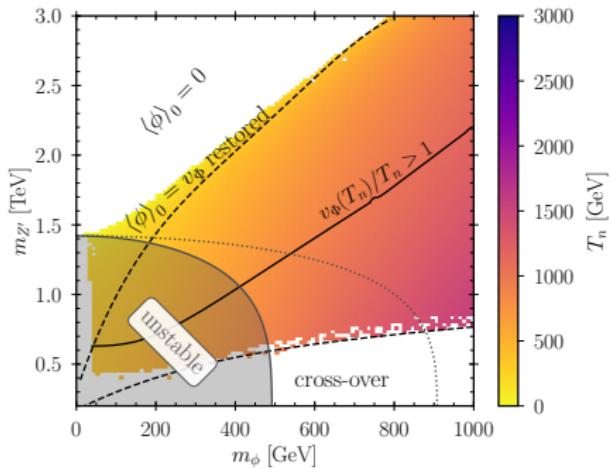
without dark leptons

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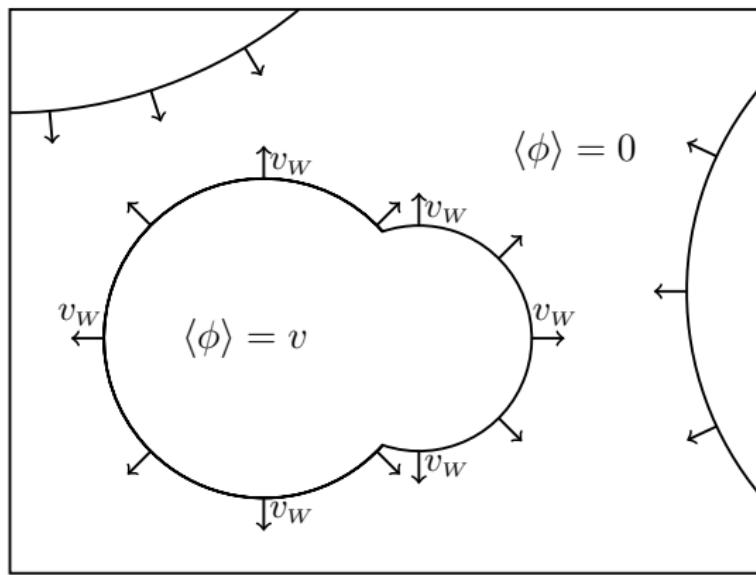


$m_{DM} = 500$ GeV, $m_{HL} = 1$ TeV

Gravitational Waves

GW spectrum: $h^2\Omega_{\text{GW}}(f) \simeq h^2\Omega_\phi(f) + h^2\Omega_{\text{sw}}(f) + h^2\Omega_{\text{turb}}(f)$

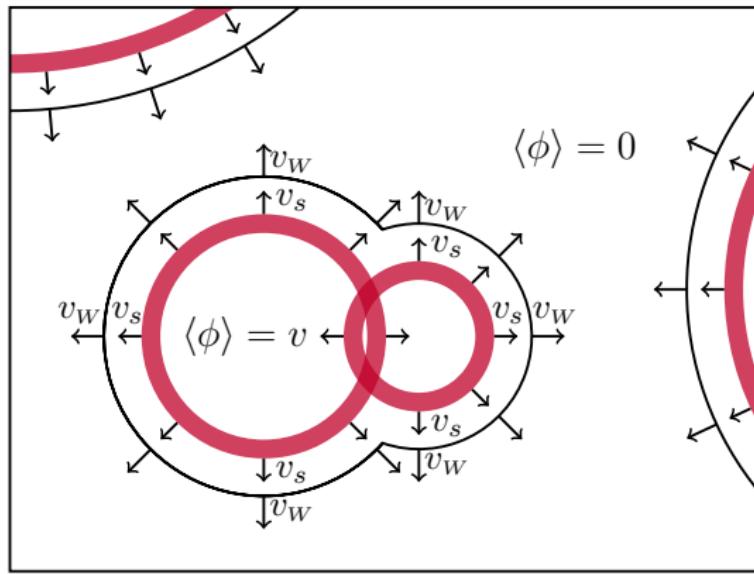
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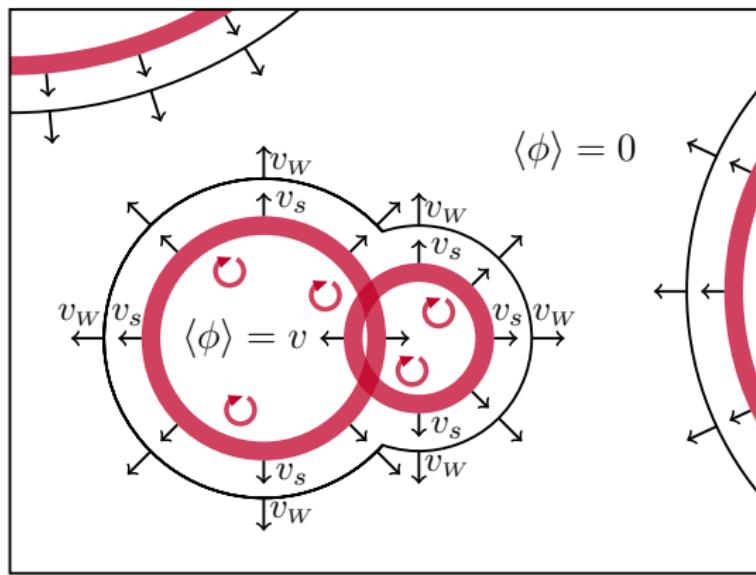
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- $h^2\Omega_{\text{sw}}(f)$: sound waves in the plasma



Gravitational Waves

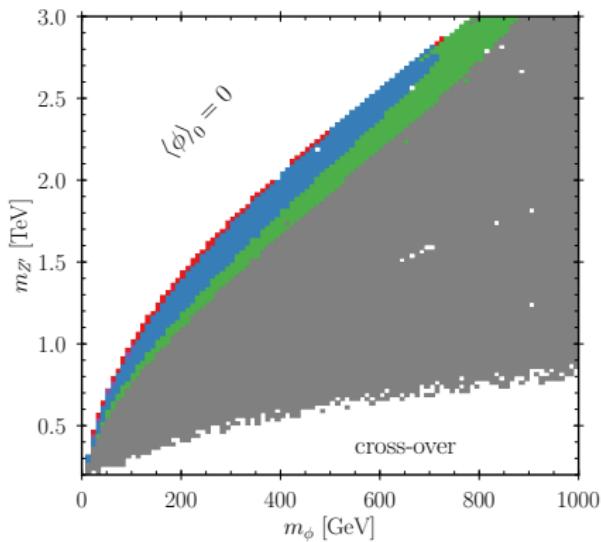
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- $h^2\Omega_\phi(f)$: collision of bubble walls
- $h^2\Omega_{\text{sw}}(f)$: sound waves in the plasma
- $h^2\Omega_{\text{turb}}(f)$: turbulence, vortical fluid motion



Detectability

neglecting dark leptons



LISA

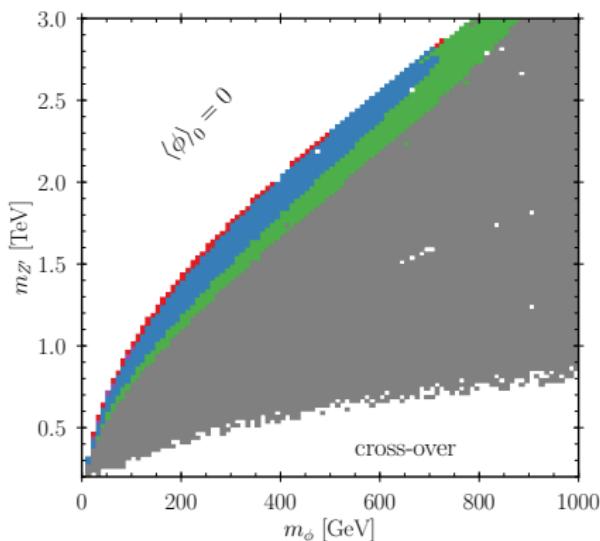
B-DECIGO

DECIGO

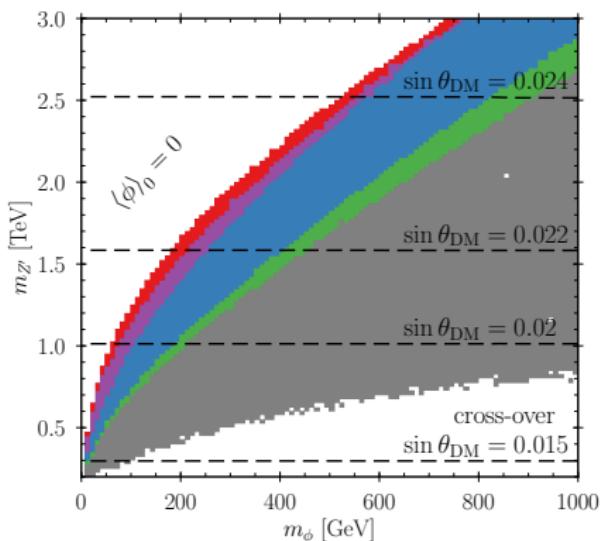
BBO

Detectability

neglecting dark leptons



including dark leptons



$$h^2 \Omega_{\text{DM}} = 0.1198, m_{\text{HL}} = 1.5 \times m_{\text{DM}}$$

LISA

B-DECIGO

DECIGO

BBO

Summary

- SM + $U(1)_\ell$ + vector-like leptons (provide DM candidate)
- Constraints:
 - LEP-2: $v_\Phi > 1880$ GeV
 - LHC: Higgs measurements, Z' searches
 - Direct Detection: mixing angles
- ℓ breaking PT can be 1st order
- generated stochastic GW background can be probed by future experiments (LISA, B-DECIGO, DECIGO, BBO)
- exotic leptons significantly enhance detectability

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Thank you for your attention!