

Consequences of a XENONnT/LZ signal for the LHC and thermal dark matter production

in collaboration with S. Baum, R. Catena, J. Conrad, K. Freese
[arXiv:1709.06051, 1712.07969]

Martin B. Krauss

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Hamburg



CHALMERS
UNIVERSITY OF TECHNOLOGY

- After potential DM discovery, what can we learn about DM properties?
 - XENONnT will start 2019
 - LHC Run 3 planned start in 2020, 300 fb^{-1} in 2022
 - Assuming $\mathcal{O}(100)$ XENONnT events in 2021 ($\sim 20 \text{ ton} \times \text{year}$ exposure)
(just below current limits)
 - Non-relativistic EFT and simplified DM models as framework
- What predictions can be made for LHC Run 3 monojet (and dijet) searches?
- Is a discovery compatible with thermal production?
- Using complementarity in DM searches, what can we learn about DM properties?
(mass, couplings, spin,...)

Simplified models & EFT

$$\begin{aligned}
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[Fitzpatrick et al., 2012]

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Simplified models & EFT

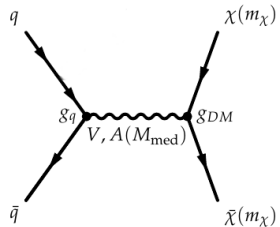
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[Dent et al., 2015]



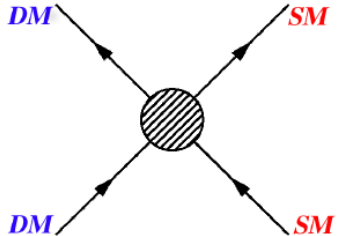
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[Dent et al., 2015]

spin 1/2 DM Coeff. Scalar med.

Vector med.

$$c_1 = \frac{h_1^N \lambda_1}{M_\Phi^2}$$

$$- \frac{h_3^N \lambda_3}{M_G^2}$$

$$c_4$$

$$4 \frac{h_4^N \lambda_4}{M_G^2}$$

$$c_6 = \frac{h_2^N \lambda_2}{M_\Phi^2} \frac{m_N}{m_\chi}$$

$$c_7$$

$$2 \frac{h_4^N \lambda_3}{M_G^2}$$

$$c_8$$

$$-2 \frac{h_3^N \lambda_4}{M_G^2}$$

$$c_9$$

$$-2 \frac{h_3^N \lambda_3}{M_G^2} \frac{m_N}{m_\chi} - 2 \frac{h_3^N \lambda_4}{M_G^2}$$

$$c_{10} = \frac{h_2^N \lambda_1}{M_\Phi^2}$$

$$c_{11} = - \frac{h_1^N \lambda_2}{M_\Phi^2} \frac{m_N}{m_\chi}$$

Benchmark points from direct detection

Direct detection can only constrain

$$M_{\text{eff}} \equiv 0.1 \frac{M_{\text{med}}}{\sqrt{g_q g_{\text{DM}}}} .$$

Assume XENONnT(/LZ) detects $\mathcal{O}(100)$ (S1) signal events with an exposure of

$$\varepsilon = 20 \text{ton} \times \text{year}$$

→ Calculate M_{eff} for various combinations of couplings and mediators.

Operators with larger suppression

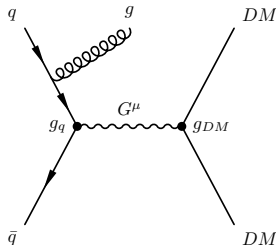
↓
smaller M_{eff}

Benchmark points

Spin 0 DM	Op.	g_q	g_{DM}	M_{eff} [GeV]
	1	h_1	g_1	14564.484
	1	h_3	g_4	10260.217
	7	h_4	g_4	4.509
	10	h_2	g_1	10.706
Spin 1/2 DM	Op.	g_q	g_{DM}	M_{eff} [GeV]
	1	h_1	λ_1	14564.484
	1	h_3	λ_3	7255.068
	4	h_4	λ_4	147.354
	6	h_2	λ_2	0.286
	7	h_4	λ_3	3.188
	8	h_3	λ_4	225.159
	10	h_2	λ_1	10.706
	11	h_1	λ_2	351.589
Spin 1/2 DM	Op.	g_q	g_{DM}	M_{eff} [GeV]
	1	h_1	b_1	14564.484
	1	h_3	b_5	10260.216
	4	h_4	$\Re(b_7)$	188.302
	4	h_4	$\Im(b_7)$	3.215
	5	h_3	$\Im(b_6)$	6.946
	7	h_4	b_5	4.509
	8	h_3	$\Re(b_7)$	287.728
	9	h_4	$\Im(b_6)$	3.674
	10	h_2	b_1	10.706
	11	h_3	$\Im(b_7)$	223.794

Impact on LHC monojet searches

- Translating the $\mathcal{O}(100)$ XENONnT events into regions in the $M_{\text{med}}\text{-}\sigma$ plane
- Mediator necessarily couples to quarks.
→ Can be produced in pp collisions
- Can decay into pair of DM particles (E_{miss}^T)
- Initial state radiation (e.g., gluon)
→ jet in detector

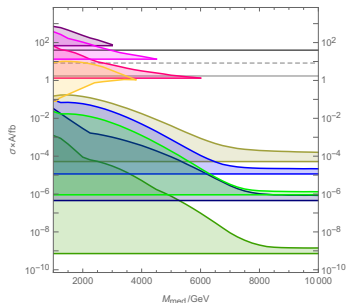


Current Limits and projections

For 12.9 fb^{-1} integrated luminosity \rightarrow monojet limit $\sigma \times \mathcal{A} \approx 40 \text{ fb}$
(Event level with selection cuts).

For projections after Run 3 we consider scaling with L and \sqrt{L} .

Monojet predictions



spin 0 DM

- $\hat{\mathcal{O}}_1(h_1, g_1)$
- $\hat{\mathcal{O}}_1(h_3, g_4)$

spin $\frac{1}{2}$ DM

- $\hat{\mathcal{O}}_1(h_1, \lambda_1)$
- $\hat{\mathcal{O}}_1(h_3, \lambda_3)$
- $\hat{\mathcal{O}}_4(h_4, \lambda_4)$
- $\hat{\mathcal{O}}_8(h_3, \lambda_4)$
- $\hat{\mathcal{O}}_{11}(h_1, \lambda_2)$

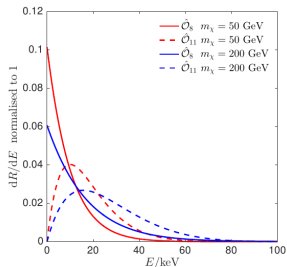
spin 1 DM

- $\hat{\mathcal{O}}_1(h_1, b_1)$
- $\hat{\mathcal{O}}_1(h_3, b_5)$

Limits and projections

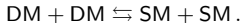
- current limit
- - - projected sensitivity
300 fb⁻¹ (\sqrt{L})
- projected sensitivity
300 fb⁻¹ (L)

Combining spectral information from direct detection with the discovery or lack of discovery of a monojet signal at the LHC can provide important information about the nature of the DM and mediator.



DM thermal production

DM in the early Universe in thermal equilibrium



Boltzmann equation

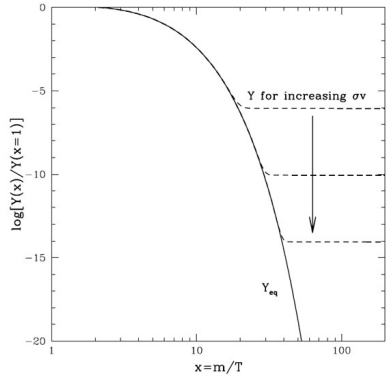
$$\dot{n} + 3Hn = -\langle\sigma v_{M\phi l}\rangle(n^2 - n_{\text{eq}}^2)$$

with the thermally averaged annihilation cross-section

$$\langle\sigma v_{M\phi l}\rangle = \int_0^\infty d\epsilon \mathcal{K}(x, \epsilon) \sigma v_{\text{lab}}$$

and

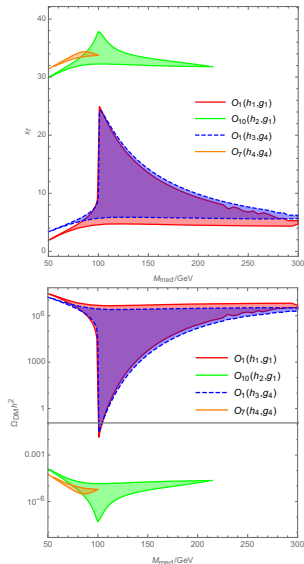
$$x = \frac{m}{T}.$$



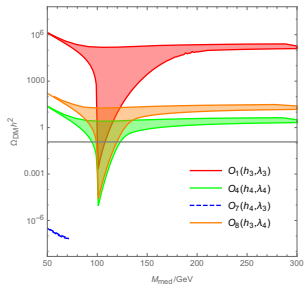
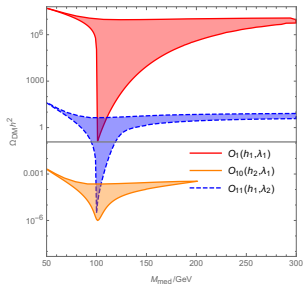
Results for scalar DM

Simplified models corresponding to spin 0 DM.

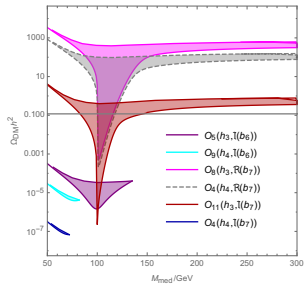
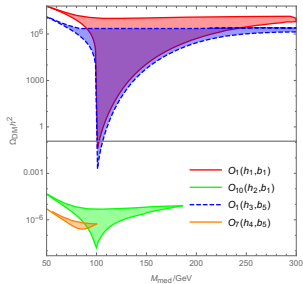
- $\hat{\mathcal{O}}_7(h_4, g_4)$ and $\hat{\mathcal{O}}_{10}(h_2, g_1)$ not compatible with the thermal production mechanism for any value of M_{med} .
- $\Omega_{\text{DM}} h^2$ much smaller than observed.
- $\hat{\mathcal{O}}_1(h_1, g_1)$ and $\hat{\mathcal{O}}_1(h_3, g_4)$ generate values for $\Omega_{\text{DM}} h^2$ which are in general too large
- For $M_{\text{med}} \sim 100$ GeV
 - resonant production of DM
 - compatible with observed relic density AND XENONnT/LZ signal



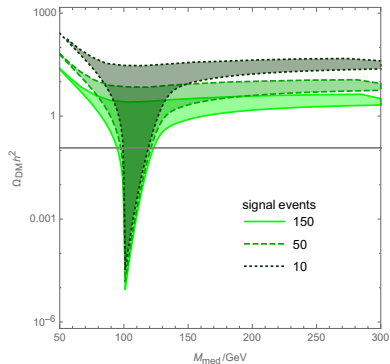
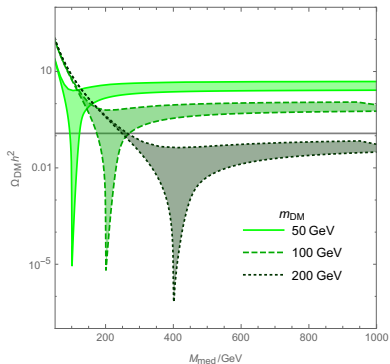
Fermionic DM



Vector DM



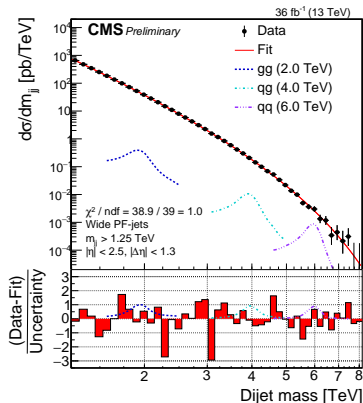
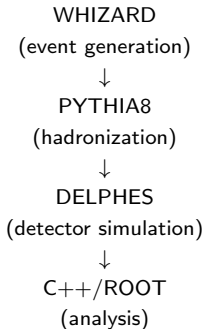
Dependence on m_{DM} and number of signal events



Dijet searches

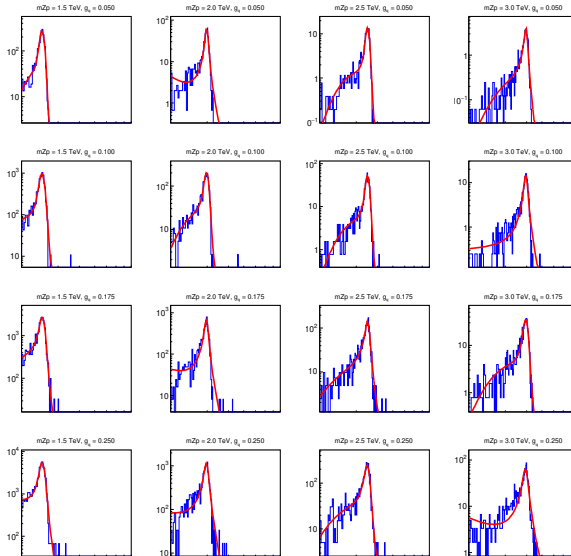
- Instead of pair of DM, mediator can decay in pair of quarks
→ Pair of jets in the detector
- Reconstruct mediator mass from jet invariant mass m_{jj}

Dijet Simulation:

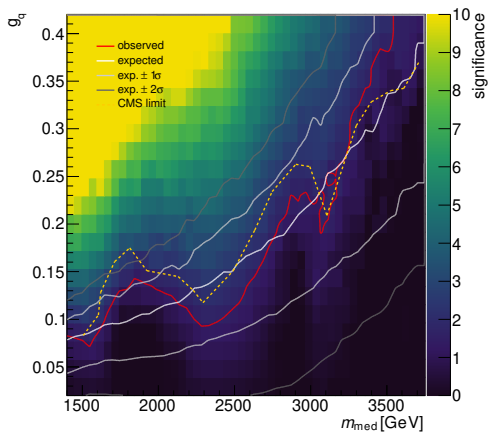


[CMS PAS EXO-16-056]

Simulated Signal



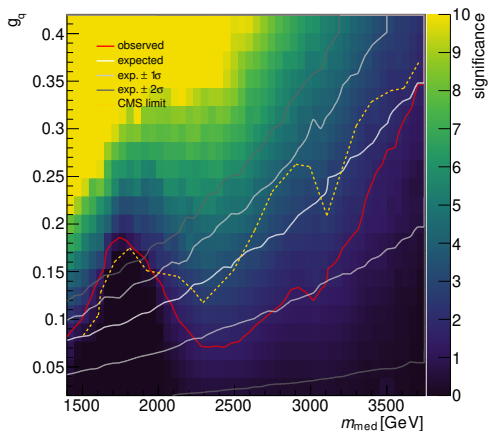
preliminary



95% C.L. exclusion limits for vector mediator

$$36 \text{ fb}^{-1} (\sqrt{s} = 13 \text{ TeV})$$

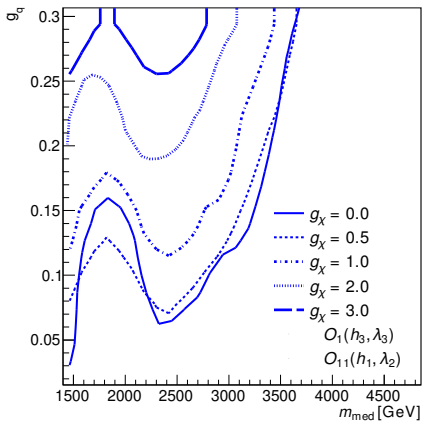
preliminary



95% C.L. exclusion limits for vector mediator
(sideband fit)
 $36 \text{ fb}^{-1} (\sqrt{s} = 13 \text{ TeV})$

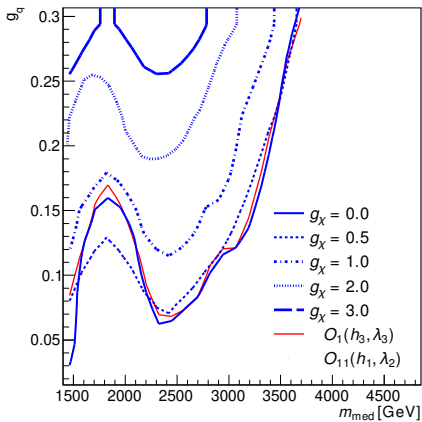
Limits for $g_\chi \neq 0$

preliminary

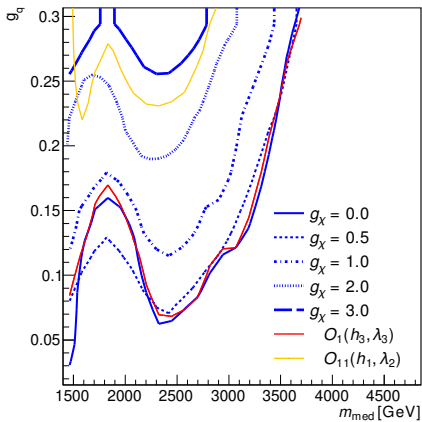


Limits for $g_\chi \neq 0$

preliminary

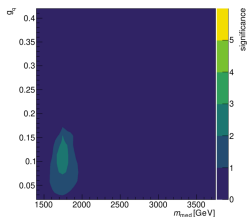


preliminary

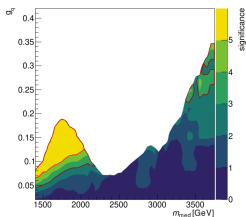


Dijet discovery potential

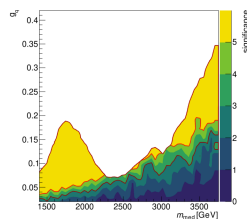
preliminary



36 fb^{-1} (current)



300 fb^{-1} (LHC Run 3)



3000 fb^{-1} (HL-LHC)

Conclusions

- If DM is a WIMP \rightarrow good chance of discovery with next generation of detectors
- Signal at XENONnT/LZ \rightarrow valuable information beyond DM mass and interaction strength
- Predictions for DM searches at the LHC
- Test compatibility with thermal production mechanism
- For most models only resonant production possible ($M_{\text{med}} \simeq 2m_{\text{DM}}$.)
- Analysis will be extended to dijets (work in progress)

Using complementarity in DM searches, we can learn more about DM properties (couplings, spin,...).

Backup-Slides

Model selection with XENONnT

Two types of spectra:

Type A: maximum at $E=0$ ($q=0$)

Type B: maximum at $E \neq 0$ ($q \neq 0$)

Canonical SI and SD interactions are of type A.

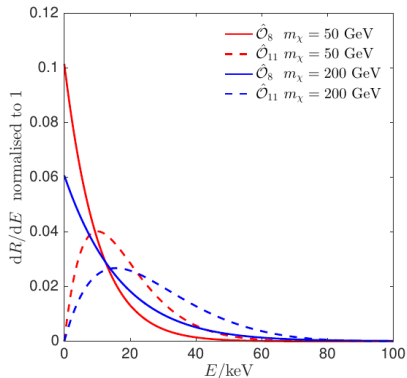
Use test statistic for model selection

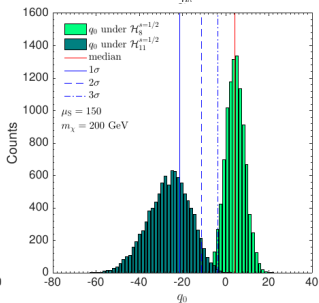
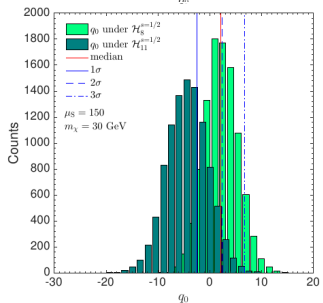
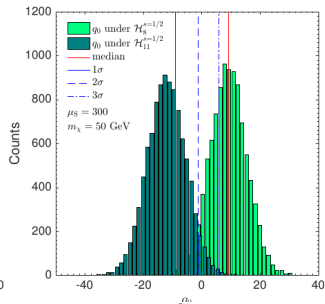
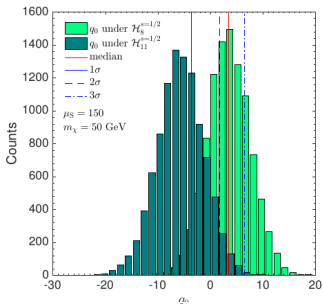
$$q_0 = -2 \ln \left[\frac{\mathcal{L}(d | \hat{\Theta}_0, \mathcal{H}_0)}{\mathcal{L}(d | \hat{\Theta}_a, \mathcal{H}_a)} \right]$$

Assumptions:

neglect operator evolution and chiral EFT corrections,

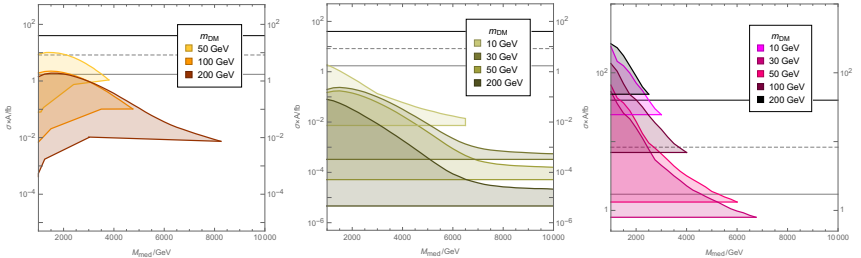
no charged mediators and universal quark-mediator couplings



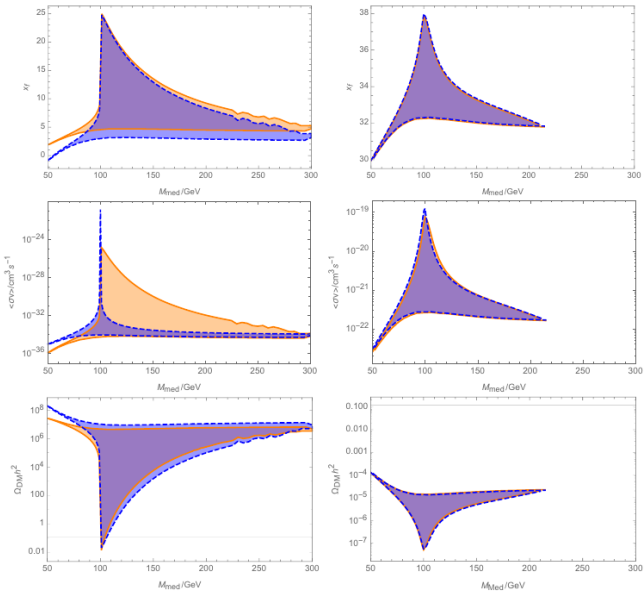


10000 pseudo-experiments each

Dependence on m_{DM}



Regions in the $M_{\text{med}} - (\sigma \times \mathcal{A})$ plane that are compatible with the detection of $\mathcal{O}(100)$ signal events at XENONnT for three representative simplified models, namely $\hat{\mathcal{O}}_1(h_3, b_5)$, $\hat{\mathcal{O}}_1(h_1, b_1)$ and $\hat{\mathcal{O}}_{11}(h_1, \lambda_2)$, and for the DM particle masses $m_{\text{DM}} = 10, 30, 50, 100$ and 200 GeV. Where the cases $m_{\text{DM}} = 30$ GeV and $m_{\text{DM}} = 100$ GeV are omitted, they only marginally differ from the $m_{\text{DM}} = 50$ GeV case.



Comparison of the models $\hat{\mathcal{O}}_1(h_1, g_1)$ (left) and $\hat{\mathcal{O}}_{10}(h_2, g_1)$ (right)