

Minestrone: flavour cosmology,
varying Yukawas,
flavour and DM.

I. Baldes 12.6.2018. DESY Workshop Seminar.

References:

1. Flavour and DM.

1501.07268 + refs therein.
(Also see 1505.03862 for an
alternative).

2. Flavour Cosmology / varying Yukawas

hep-ph/0401012

+ DESY papers	1604.04526	(PT)
	1608.03254	(FN)
	1612.02447	(RS)
	1706.08534	(CP)
	1803.08546	(CH)
	1804.07314	(CH)

Also see S.F. King et al. for

GUT flavour models and leptogenesis

(e.g. 1710.03229).

(1)

1. Flavour Portal to DM

SM fermion mass:

FN model: $m_f = a_{ij} \left(\frac{\langle \phi \rangle}{M} \right)^{n_{ij}} \frac{v}{\sqrt{2}}$

from $\mathcal{L} \supset a_{ij} \left(\frac{\langle \phi \rangle}{M} \right)^{n_{ij}} \bar{f}_L f_R \phi$.

Now make DM transform under

G_{flav} . DM mass:

$$m_\chi = b_\chi \left(\frac{\langle \phi \rangle}{M} \right)^{n_\chi} \langle \phi \rangle$$

from $\mathcal{L} \supset b_\chi \left(\frac{\phi}{M} \right)^{n_\chi} \phi \bar{\chi}_L \chi_R$

$$\mathcal{L} \xrightarrow{SSB} \overbrace{n_f \frac{m_f}{\langle \phi \rangle} \bar{f}_L f_R \phi}^{\equiv \lambda_f} + \overbrace{(n_\chi + 1) \frac{m_\chi}{\langle \phi \rangle} \bar{\chi}_L \chi_R \phi}^{\equiv \lambda_\chi}$$

Flavour talks to SM fermions and DM

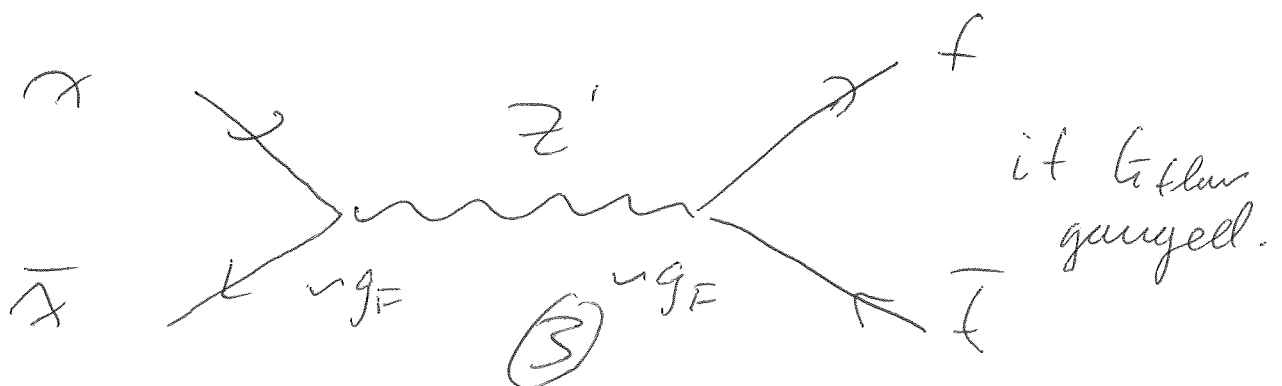
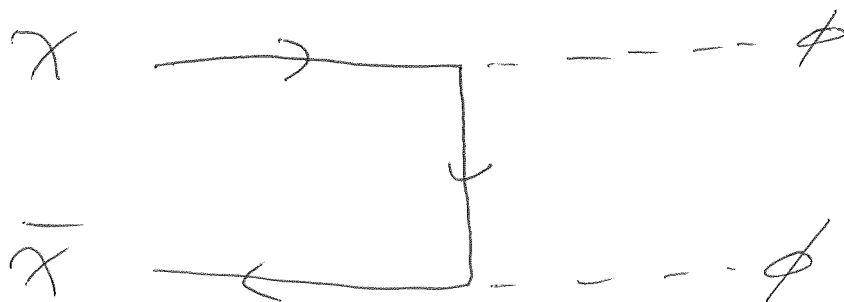
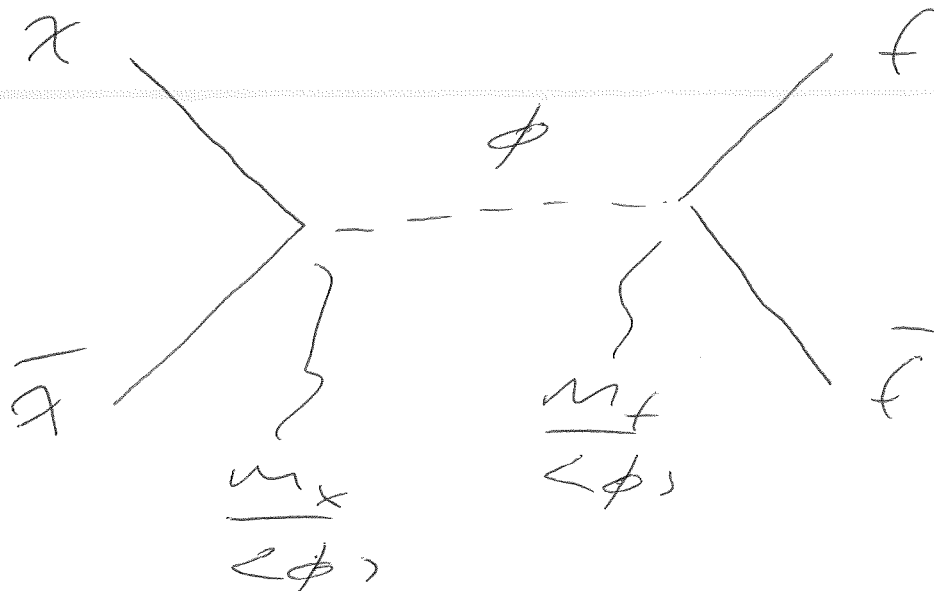
Flavour mass arises from SSB of G_{flav} .

Parameterise: $m_\phi = k \langle \phi \rangle$.

Some similarities to Higgs portal:

- Mediator couples preferentially to heavier flavours.
- Correlation b/w DM annihilation and scattering w/ Nuclei.

Annihilation diagrams.



$$\langle \sigma_{\phi}^T V \rangle \sim \frac{\lambda_x^4}{m_x^3} T$$

$$\langle \sigma_{\phi}^{SV} \rangle \sim \frac{\lambda_x^2 \lambda_{\phi}^2 m_x T}{(4m_x^2 - m_{\phi}^2)^2 + \Gamma_{\phi}^2 m_{\phi}^2}$$

$$\langle \sigma_{Z'}^V \rangle \sim \frac{g_F^4 m_x^2}{(m_{Z'}^2 - 4m_x^2)^2 + \Gamma_{Z'}^2 m_{Z'}^2}$$

DM + Nucleon SI scattering:

$$\sigma_{\phi}^{SI} \sim \frac{\lambda_x^2 \lambda_{\phi N}^2 m_{xN}^2}{m_{\phi}^4}, \quad \lambda_{\phi N} \propto \lambda_{\phi}$$

$$\sigma_{Z'}^{SI} \sim \frac{g_F^2 \lambda_{Z'N}^2}{m_{Z'}^4} m_{xN}^2, \quad \lambda_{Z'N} \propto g_F.$$

Explicit Model : $U(1)_F$ FN model

Charges : $(Q_{q_1}, Q_{q_2}, Q_{q_3}) = (3, 2, 0)$

$(Q_{u_1}, Q_{u_2}, Q_{u_3}) = (3, 2, 0)$

$(Q_{d_1}, Q_{d_2}, Q_{d_3}) = (4, 2, 2)$

Large flavor coupling $\phi, \bar{F}_{L,2}, C_{R,1}$

" Z' couplings to lighter generations.

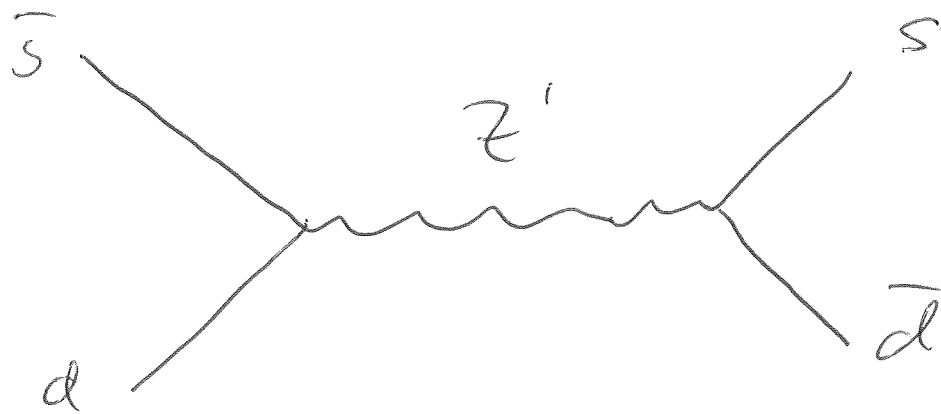
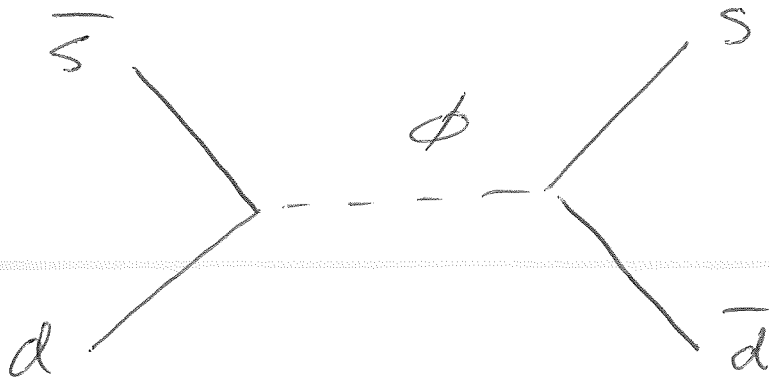
Assume anomalies are cancelled by unspecified field content in hidden and/or messenger sector.

(e.g. flavor messengers are vector-like quarks under SM gauge groups
- not necessarily vector-like under G_F - giving more freedom to cancel the anomalies).

Model is largely specified by the parameters

$$M_\phi, m_\chi, R \equiv \frac{M_\phi}{\langle \phi \rangle}$$

Flavour constraints



ΔM_K

$$M_\phi \gtrsim \sqrt{R} \times 580 \text{ GeV}$$

ϵ_K

$$M_\phi \gtrsim \sqrt{R} \sqrt{\text{arg}(V_{12}^d V_{21}^d)} \times 2.3 \text{ TeV}$$

$$M_{Z'} \gtrsim \left(\frac{g_F}{10^{-3}} \right) \times 210 \text{ GeV}$$

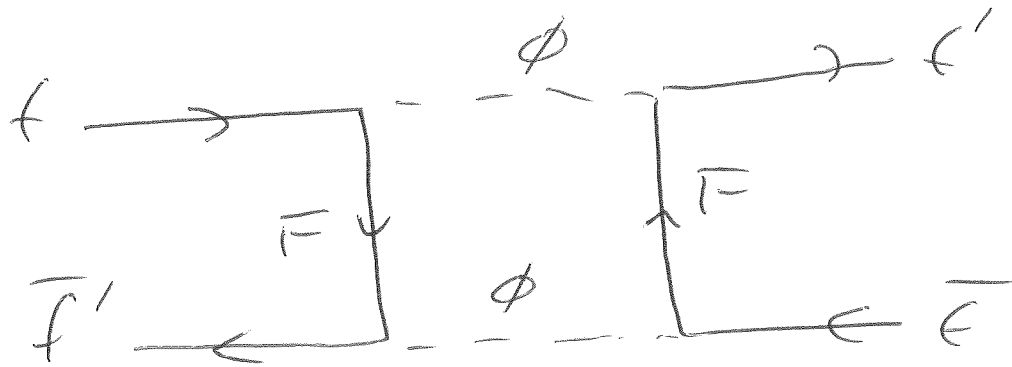
$$M_{Z'} \gtrsim \left(\frac{g_F}{10^{-3}} \right) \sqrt{\text{arg}(V_{12}^d V_{21}^d)}$$

i.e.
 $\langle \phi \rangle \gtrsim 210 \text{ TeV. } \textcircled{6}$

$$\times 3.3 \text{ TeV}$$

UV Messenger fields

→ One-loop $\overline{F} C N C$ (1204.1275)



Limit from $D - \overline{D}$

$$m_\phi \gtrsim C_D \times k \times 2.3 \text{ TeV}$$

from ΣK^0 .

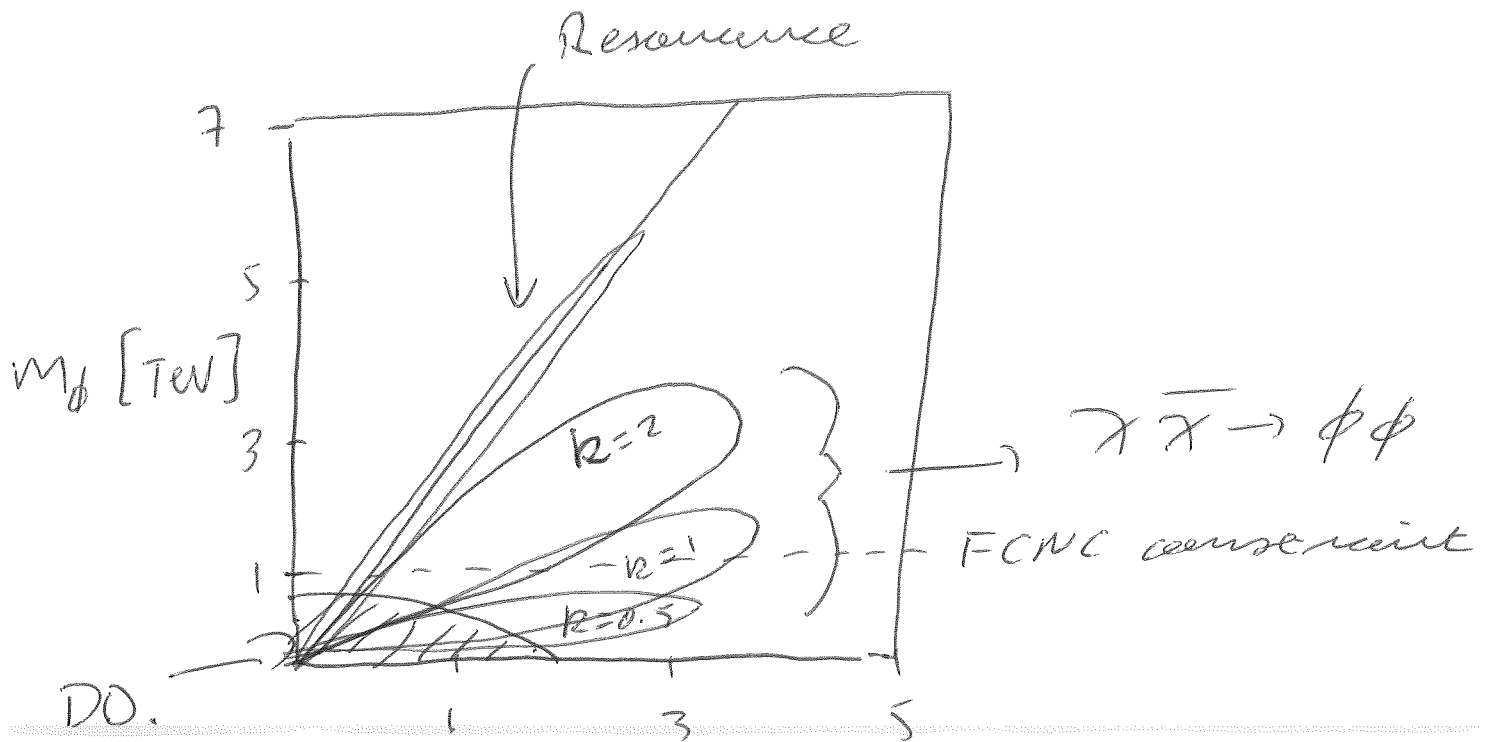
$$m_\phi \gtrsim C_K \times k \times 27 \text{ TeV.}$$

C_D, C_K : product of unknown $O(1)$ coefficients.

For interesting pheno need $C_K \sim 0.1$.

Results

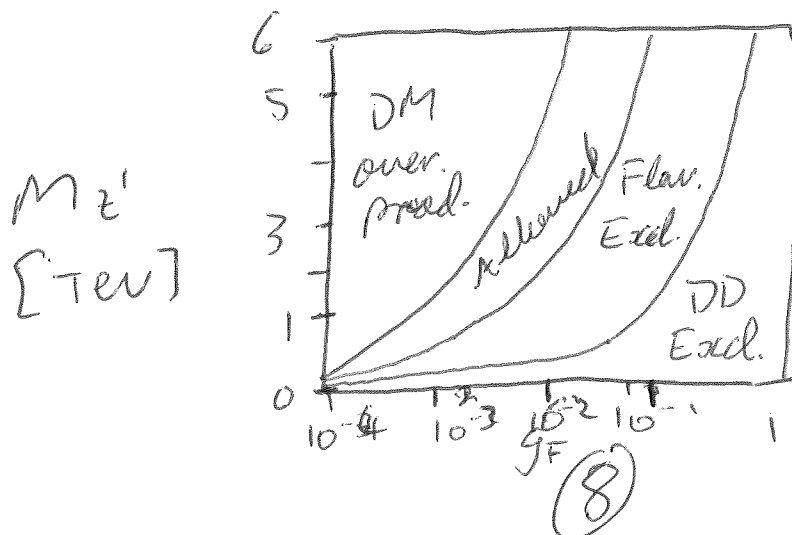
Global $U(1)_F$



DD can be under ν -flavor.
(but many points above).

Gauged case $\langle \phi \rangle \gg v_{EW}$ due to
 Z' FCNC.

→ have to be on resonance
for DM.



② Flavour Cosmology

Motivation from EWBG. SM CP source

$$\epsilon_{CP} \sim \frac{1}{T_n^6 M_W^6} \prod_{\substack{i,j \\ u,c,t}} (m_i^2 - m_j^2) \times \prod_{\substack{i,j \\ d,s,b}} (m_i^2 - m_j^2)$$

$$\times J_{CP}$$

$$\hookrightarrow \sim S_{12} S_{23} S_{13} S_{12} \delta_{CKM}$$

$$\sim 10^{-19}$$

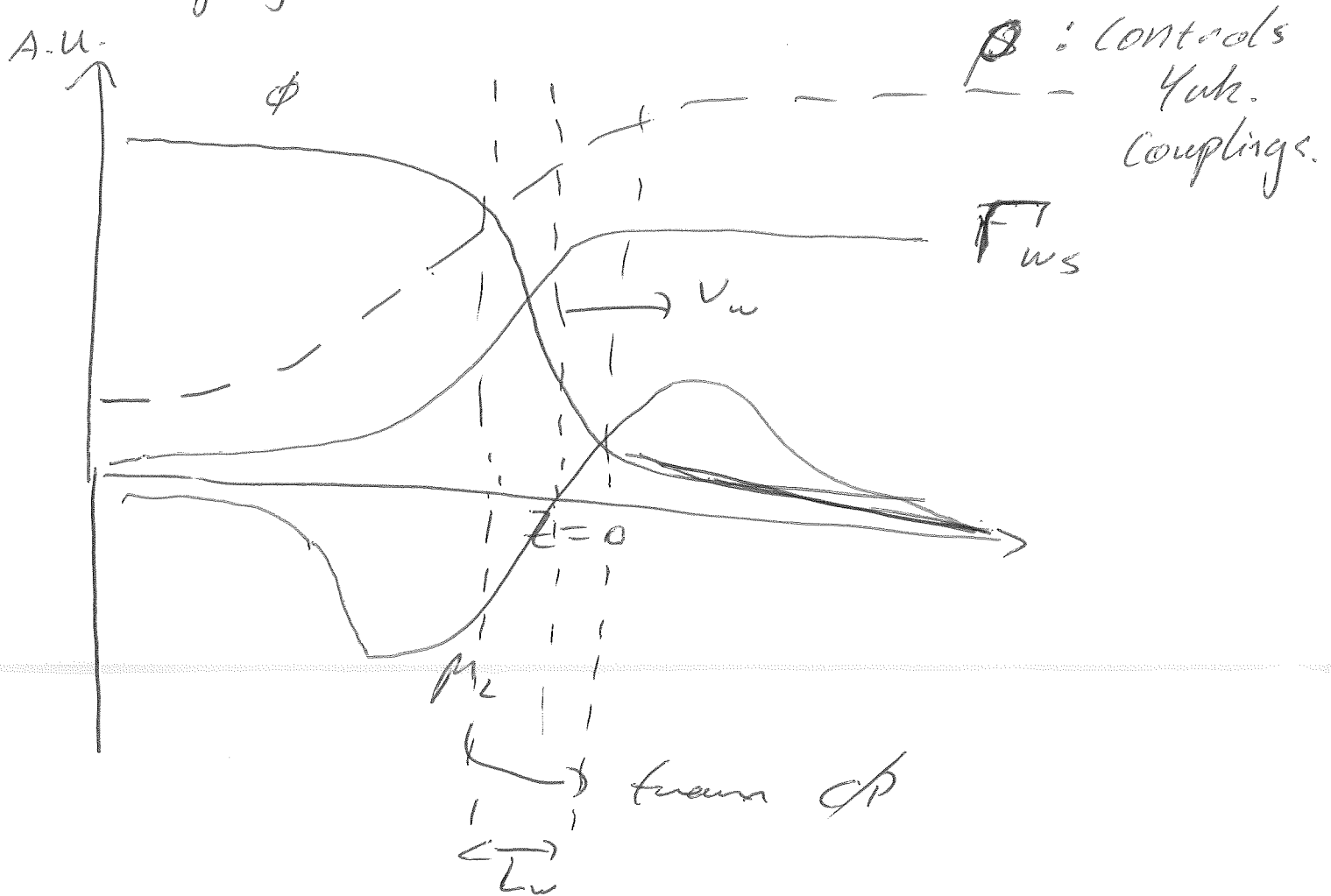
is too small.

Simple solution: increase Yukawas in the early universe!

The Yukawas today are set during the EWPT.



Varying Yukawas:



$$\partial_z n_B = \frac{3}{2} v_w^{-1} P_{ws} (N_c \mu_L T^2 - A n_B)$$

$$P_{ws} = 10^{-6} T e^{-a \phi(z)/T}$$

$$a \approx 37 = \frac{E_{sph.}}{v_{EW}}$$

$$\text{In SM } A = \frac{15}{2}$$

$$\mu_L = \frac{\mu_{tL} + \mu_{bL} + \mu_{cL} + \mu_{sL} + \mu_{uL} + \mu_{dL}}{2}$$

To determine $\mu_c(z)$ solve
Transparent Eq's.

∇p_{source}
↓

$$A(z) \cdot r'(z) + B(z) \cdot r(z) = \bar{S}(z)$$

|

Set by dynamics and interactions of
the particles

$$\Gamma = (\mu_1, \dots, \mu_n, u_1, \dots, u_n)$$

|
chem. pot. local vel.
in plasma.

$$\eta_B = \frac{\eta_B(-\infty)}{\leq}$$

$$= \frac{135 M_c}{4\pi^2 v_w g_* T} \int_{-\infty}^{\infty} dz \rho_{ws} \mu_c \exp \left[-\frac{3}{2} \frac{L}{v_w} \int_{-\infty}^z dz \rho_{ws} \right]$$

$$\sim \frac{\rho_{ws} \mu_c L_w}{g_* T} \sim 10^{-8} \frac{M_c}{T} \quad \text{for } L_w \sim \frac{1}{T}$$

It turns out dominant term w/ varying Yuk's
is: $S \sim \text{Im} [V^\dagger m^\dagger m V]_{ii}$.
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FN - Froggatt - Nielsen.

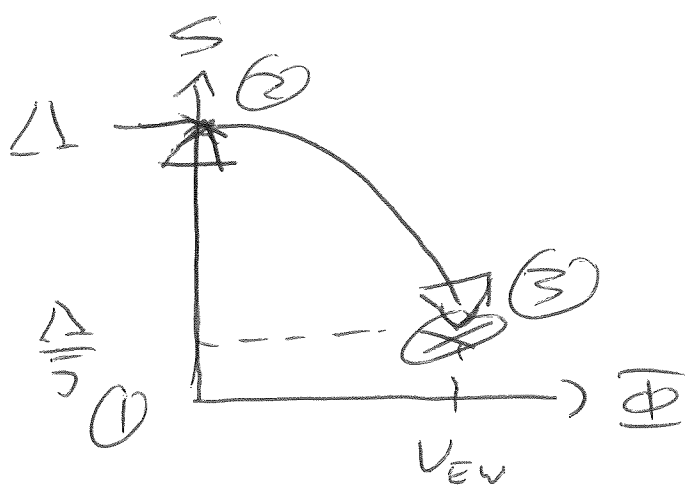
$$\mathcal{L} \supset y_{ij} \left(\frac{S}{\Lambda} \right)^{n_{ij}} \overline{Q_{iL}} \Phi Q_{jR}$$

$$\Rightarrow y_{\text{eff}} \sim y_{ij} \left(\frac{\langle S \rangle}{\Lambda} \right)^{n_{ij}}$$

Need $\langle S \rangle \ll \Lambda$ before EWP

$$\langle S \rangle = \frac{\Lambda}{5} \text{ today}$$

(to fit quark masses/mixing).



① For $T \gg \Lambda$
in symmetric
phase.

② At $T \sim \Lambda$
evolve to
 $\langle S \rangle \sim \Lambda$.

③ At EWP

$$\langle S \rangle : \Lambda \rightarrow \frac{\Lambda}{5}$$

$$\langle \Phi \rangle : 0 \rightarrow V_{EW} = 246 \text{ GeV.}$$

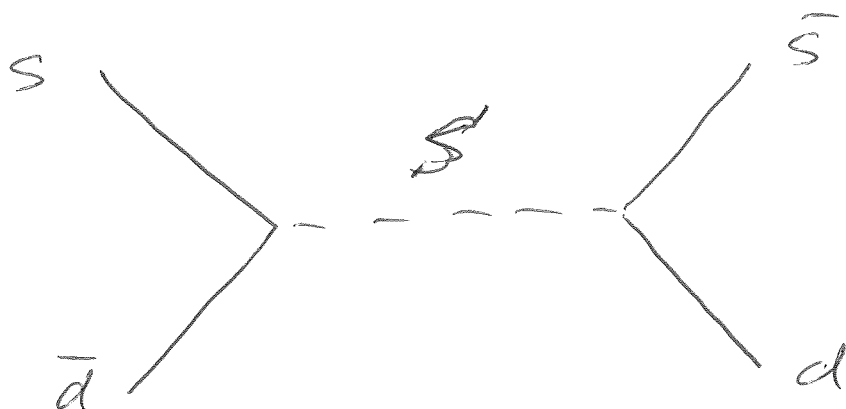
$$|V(\Phi=0, S=\Lambda) - V(\Phi=V_{EW}, S=\Lambda_{s/5})|$$

$$\sim M_s^2 \Lambda_s^2 - (M_\phi^2 V_{EW}^2 + M_s^2 \Lambda_s^2 / 25)$$

$$\Rightarrow \text{Need } M_s^2 \Lambda_s^2 \sim M_\phi^2 V_{EW}^2 \quad (*)$$

(for careful derivation see 1608.03254)

But $K - \bar{K}$



requires $m_s \Lambda_s \gtrsim (\text{few TeV})^2$

$\rightarrow$ incompatible with $(*)$.

Possible sol'n: multiple flavours.

~~One~~ universal Yuk's today.

Other varies Yuk's during EWPT

but has negligible VEV early $\rightarrow$ FCNCs
new loop process.

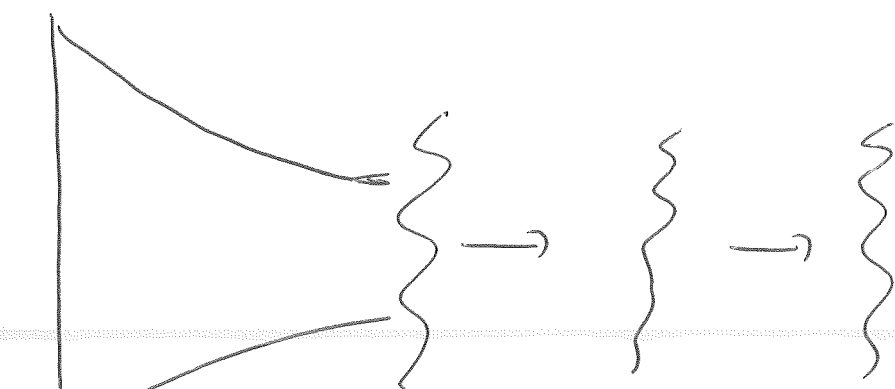
(13)

RS - Randall Sundrum

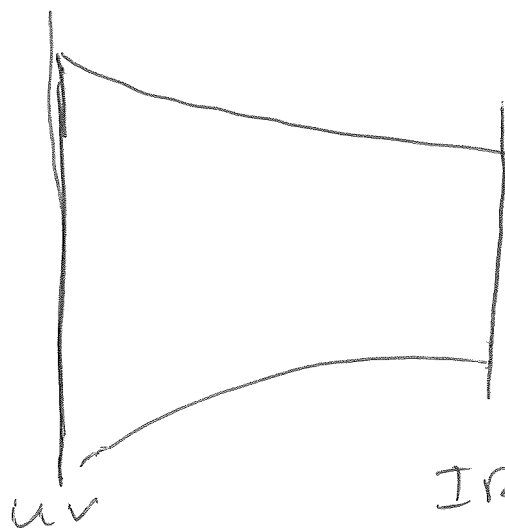
$$ds^2 = e^{-2ky} \eta_{\mu\nu} dx^\mu dx^\nu - dy^2$$

$$y \in [0, y_{IR}]$$

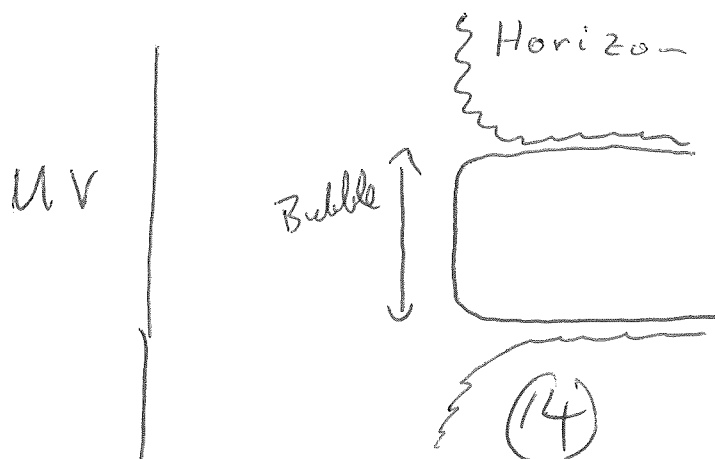
But at high T : AdS - Schwarzschild!



Horizon moves to ∞
as T drops (or rapidly
during the PT).

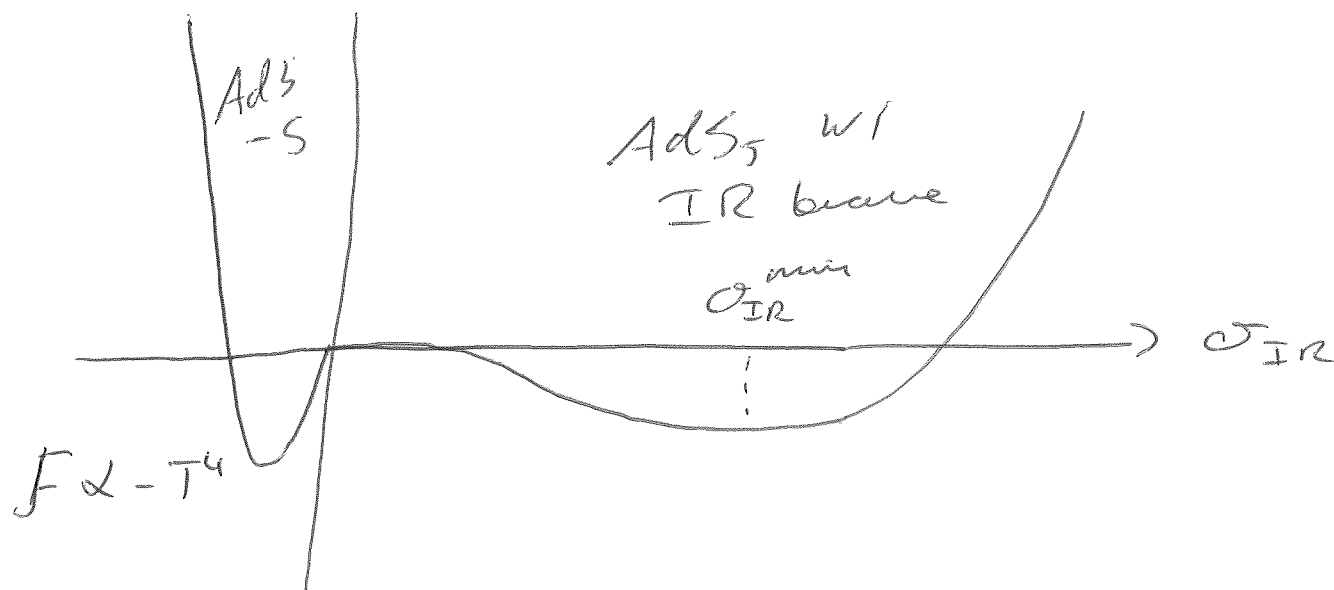


IR brane moves in from
 ∞ .



Picture of
the PT.

Use Goldberger - Wise mechanism
to stabilise extra dimension.



$$\sigma_{IR} \equiv e^{-k y_{IR}} \quad \swarrow \text{position of}$$

$$\sigma_{IR}^{min} \sim M_{Pl} \sim \text{TeV}.$$

It turns out the potential is very shallow and slowly varying.

$$V(\sigma_{IR}) \sim \sigma_{IR}^4 \times f(\sigma_{IR}^\epsilon)$$

$$\text{where } \epsilon \equiv \sqrt{4 + \frac{m_{GW}^2}{R^2}} - 2 \sim \frac{1}{20}$$

$$\text{Higgs: } V(H) = \lambda (|H|^2 - v_P^2 \sigma_{IR}^2)^2$$

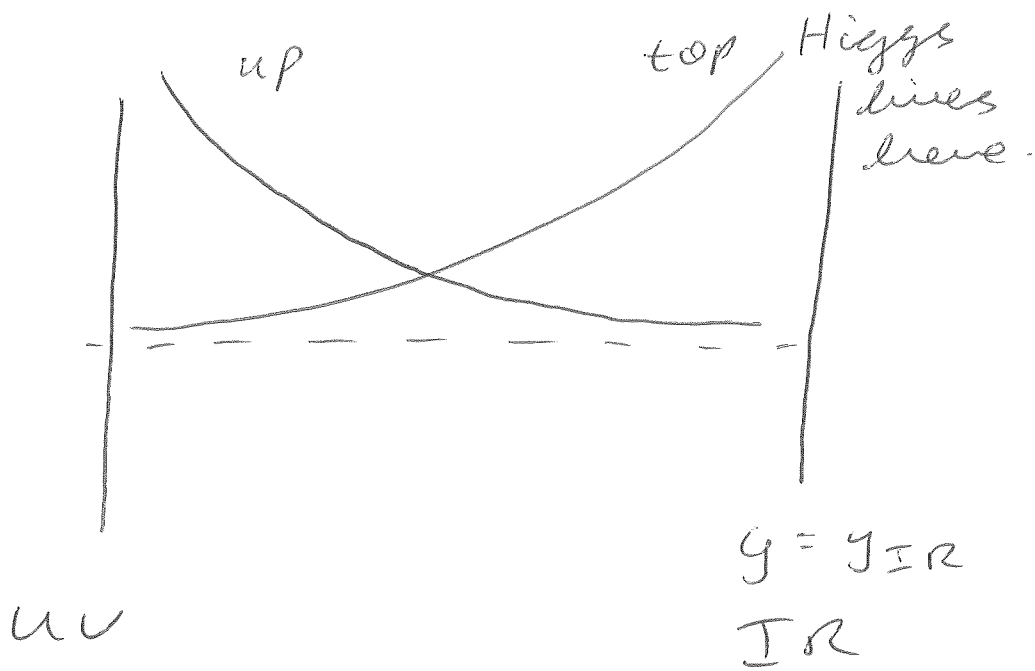
$$\hookrightarrow \sim M_{Pl}$$

$$\text{So } \langle H \rangle = v_{EW} \frac{\sigma_{IR}}{\sigma_{IR}^{min}}$$

depends on the
vacuum.

(15)

Feynman



$$S \supset - \int d^5 x \sqrt{g} \bar{Q} k Q$$

Profile $f^{(0)}(y) = N e^{(2-c_0)y}$

$c_0 > 1/2 \rightarrow$ UV localised

$c_0 < 1/2 \rightarrow$ IR localised.

4D effective Yuk.

$$S \supset \int d^5x \sqrt{g} \delta(y - y_{IR}) \lambda \tilde{H} \bar{Q}^{(0)} u_R$$

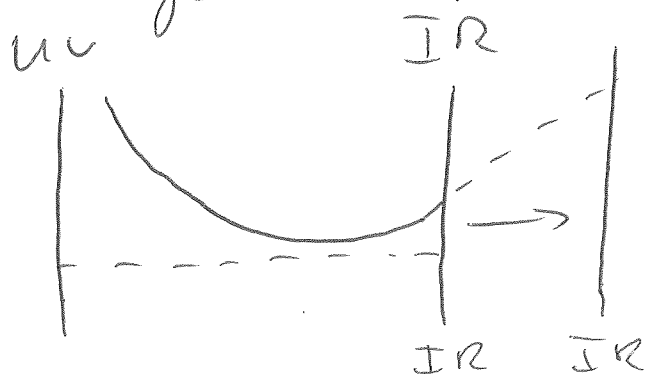
$$\tilde{H} = \sigma_{IR} H$$

But: overlap decreases as IR brane moves away.

One possibility: modify bulk mass term to depend on GW scalar

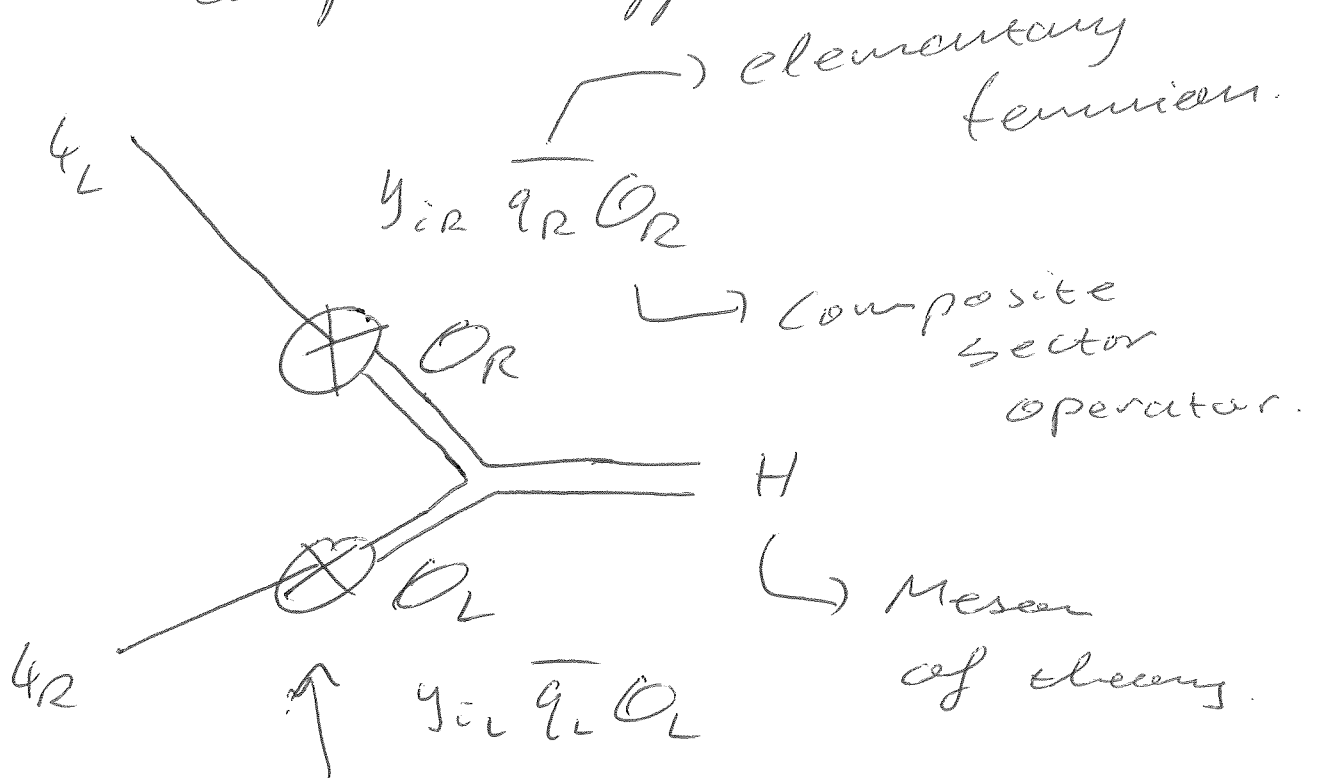
$$c_0 k \bar{Q} Q \rightarrow \rho_0 \phi_{GW} \bar{Q} Q$$

Profile $f_L^{(0)}$ turns around at some point.



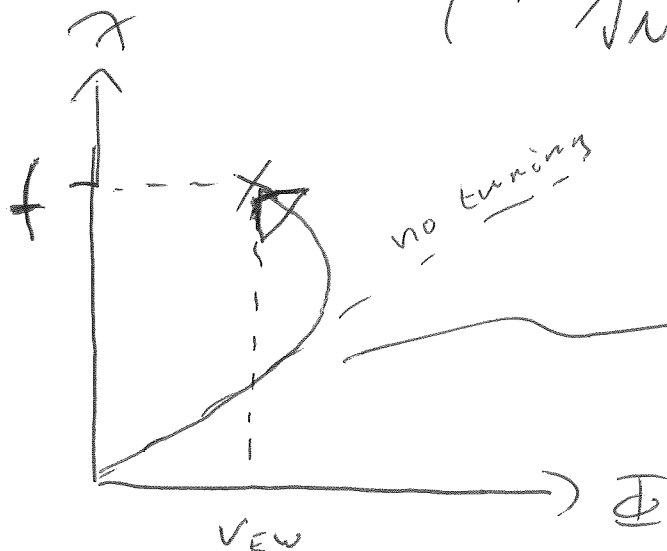
$\rightarrow$ Larger Yuk's in sym. EW phase.

CH - Composite Higgs.



Mixing depends on dilaton χ $\rightarrow$ meson or glueball.
which sets confinement scale.

$$\lambda_g(\chi) \sim \frac{g_L(\chi) g_R(\chi)}{(4\pi/\sqrt{N})}$$



Simultaneous confinement and EWSB.

non-trivial $\phi(\chi)$

~~helps~~

$\rightarrow$ required for BAH.

Some tuning required for $VEW < f$.

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