# **DUNE Near Detector**

**Alfons Weber** 

12<sup>th</sup> Terascale Detector Workshop Dresden, March 2019

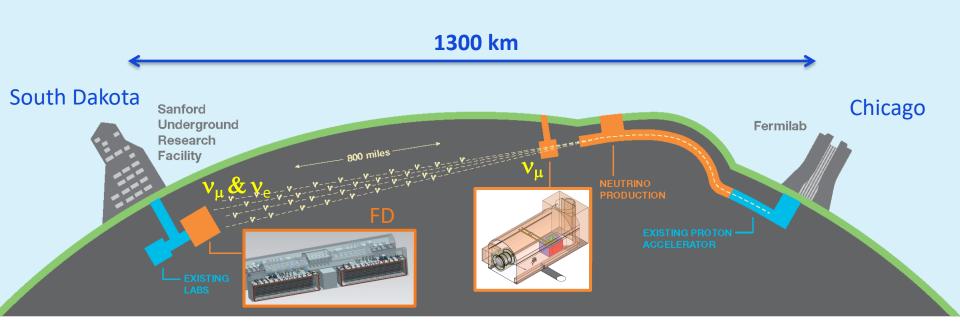






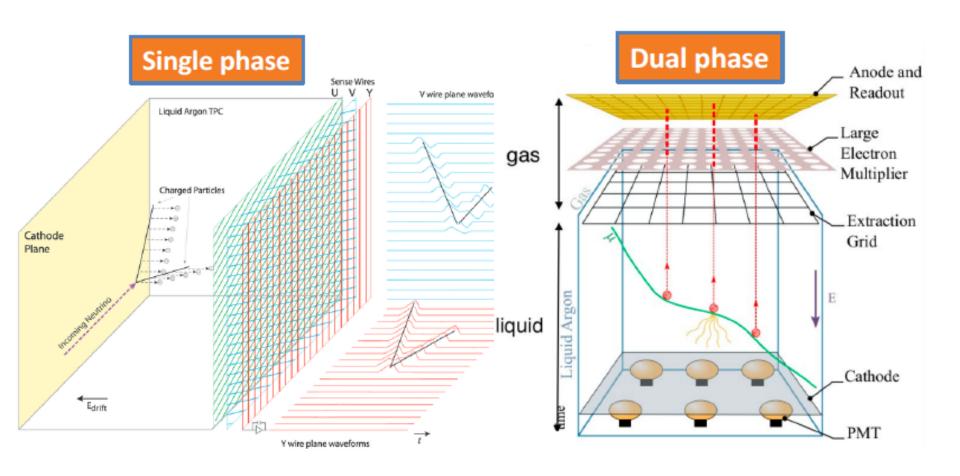
# **DUNE General Setup**

- LBNF/DUNE will consist of
  - An intense 1.2 MW upgradeable ν-beam fired from Fermilab
  - A massive 68 kt (40kt instrumented) deep underground LAr detector in South Dakota and a large Near Detector at Fermilab
  - A large international collaboration





# Far Detector (LArTPC)

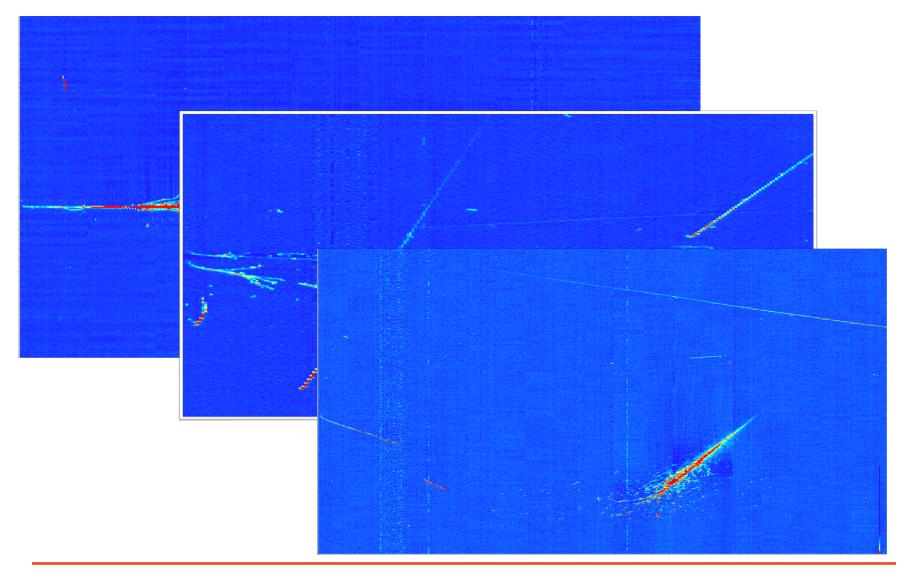








## **Events from ProtoDUNE@CERN**







### **Oscillations for Dummies**

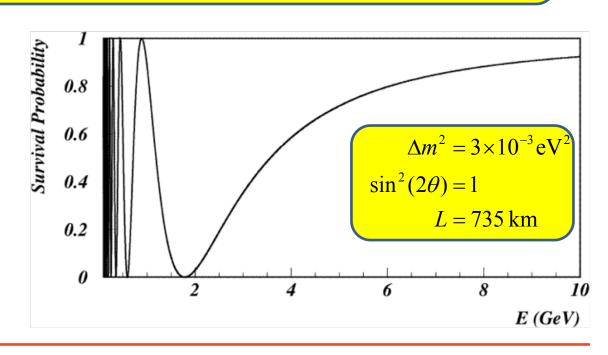
$$V_{\mu} \longrightarrow V_{\mu} \text{ or } V_{\tau}$$

$$V_{\mu} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} v_{1} \\ v_{2} \end{pmatrix} \qquad P(v_{\mu} \to v_{\tau}) = \sin^{2}(2\theta) \sin^{2}\left(\frac{1.27\Delta m^{2}L}{E_{\nu}}\right)$$

- Measure prob.
  - Survival
  - Appearance
- Result

Mar-2019

- Mixing angle
- Mass differences







# The Full Monty

- Life isn't that easy
  - 3 Flavour oscillations
  - Matter effects
- The full formula

$$P(\nu_{\mu} \to \nu_{e}) = \sin^{2}(2\theta) \sin^{2}\left(\frac{\Delta m^{2}L}{4E}\right)$$

$$\begin{split} P(\nu_{\mu} \to \nu_{e}) &= 4C_{13}^{2}S_{13}^{2}S_{23}^{2}\sin^{2}\frac{\Delta m_{31}^{2}L}{4E} \times \left(1 + \frac{2a}{\Delta m_{31}^{2}}\left(1 - 2S_{13}^{2}\right)\right) \\ &+ 8C_{13}^{2}S_{12}S_{13}S_{23}(C_{12}C_{23}\cos\delta - S_{12}S_{13}S_{23})\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\sin\frac{\Delta m_{21}^{2}L}{4E} \\ &- 8C_{13}^{2}C_{12}C_{23}S_{12}S_{13}S_{23}\sin\delta\sin\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\sin\frac{\Delta m_{21}^{2}L}{4E} \\ &+ 4S_{12}^{2}C_{13}^{2}\left\{C_{12}^{2}C_{23}^{2} + S_{12}^{2}S_{23}^{2}S_{13}^{2} - 2C_{12}C_{23}S_{12}S_{23}S_{13}\cos\delta\right\}\sin^{2}\frac{\Delta m_{21}^{2}L}{4E} \\ &- 8C_{13}^{2}S_{13}^{2}S_{23}^{2}\cos\frac{\Delta m_{32}^{2}L}{4E}\sin\frac{\Delta m_{31}^{2}L}{4E}\frac{aL}{4E}\left(1 - 2S_{13}^{2}\right) \end{split}$$







**DUNE Near Detector** 

# **Measuring Oscillations**

Oscillation probabilities

$$P_{\nu_{\mu} \to \nu_{e}}(E_{\nu}) = \frac{\phi_{\nu_{e}}^{far}(E_{\nu})}{\phi_{\nu_{\mu}}^{far,no-osc}(E_{\nu})} = \frac{\phi_{\nu_{e}}^{far}(E_{\nu})}{\phi_{\nu_{\mu}}^{near}(E_{\nu}) * F_{far/near}(E_{\nu})}$$

Well known (1-2%)





### Are there cancellations?

Oscillation signal

$$\frac{dN_{\nu_e}^{far}}{dE_v} / \frac{dN_{\nu_\mu}^{near}}{dE_v} = P_{\nu_\mu \to \nu_e}(E_\nu) * \frac{\sigma_{\nu_e}^{Ar}(E_\nu)}{\sigma_{\nu_\mu}^{Ar}(E_\nu)} * F_{far/near}(E_\nu)$$

Near muon/electron ratio

1-2% uncertainty

Small theo. uncertainty

$$\frac{dN_{\nu_e}^{near}}{dE_{\nu}} / \frac{dN_{\nu_{\mu}}^{near}}{dE_{\nu}} = \frac{\sigma_{\nu_e}^{Ar}(E_{\nu})}{\sigma_{\nu_{\mu}}^{Ar}(E_{\nu})} * \frac{\phi_{\nu_e}^{near}(E_{\nu})}{\phi_{\nu_{\mu}}^{near}(E_{\nu})}$$

- Need to know
  - Flux & cross section ratios
  - Far/near extrapolation

Not so small uncertainty





# **But in Reality**

$$\frac{\frac{dN_{\nu_e}^{far}}{dE_{rec}}}{\frac{dN_{\nu_{\mu}}^{near}}{dE_{rec}}} = \frac{\int P_{\nu_{\mu} \to \nu_e}(E_{\nu}) * \phi_{\nu_{\mu}}^{near}(E_{\nu}) * F_{far/near}(E_{\nu}) * \sigma_{\nu_e}^{Ar}(E_{\nu}) * T_{\nu_e}^{far}(E_{\nu}, E_{rec}) dE_{\nu}}{\int \phi_{\nu_{\mu}}^{near}(E_{\nu}) * \sigma_{\nu_{\mu}}^{Ar}(E_{\nu}) * T_{\nu_{\mu}}^{near}(E_{\nu}, E_{rec}) dE_{\nu}}$$

- No cancellations
  - Unless you unfold
- Need to understand especially
  - Detector effects in near and far detector
  - Relation of visible to neutrino energy
  - Cross section ratios
  - Near to far flux extrapolation

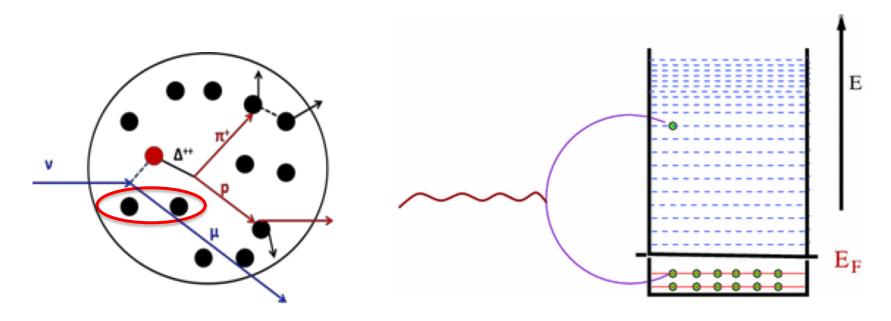




### **Problem of Nuclear Effects**

#### Unknown

- Neutrino energy
- Final state particles (neutrons, nuclear recoil)
- Nuclear effects







#### What does the ND need to measure?

ND Fluxes

$$\phi_{\nu_x}^{near}(E_{\nu})$$

- Prior constrained 5-10%
- Total and differential cross sections on Argon

$$\frac{d^n \sigma_{\nu_{\chi}}^{Ar}}{da \ db \ dc \dots} (E_{\nu})$$
 (Largely unknown)

True to reconstruction "matrix"

$$T_{\nu_x}^{far}(E_v, E_{rec})$$
 and  $T_{\nu_x}^{near}(E_v, E_{rec})$ D

#### Depends on

- Detector effects
- differential cross sections





# **DUNE ND Approach**

- Have a capable Multi-Purpose Detector (MPD) to
  - Constrain flux
    - target nucleus does not matter
  - Measure as many different differential cross sections as possible on argon
    - sensitive to pions, protons, neutrons, electrons, photons
    - Other nuclear targets might be useful (especially H)
- Have a LAr TPC to measure
  - reactions on argon (mostly inclusive)
  - constrain detector effects



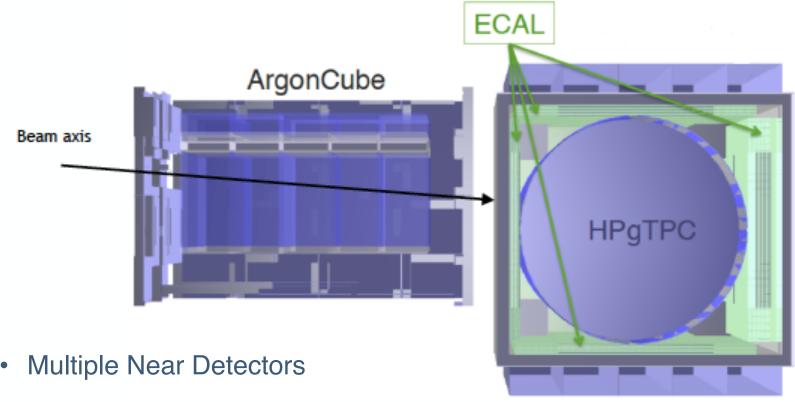
# **ND Detector Systems**

- Highly segmented LAr TPC (LArTPC)
- High Pressure gaseous TPC (HPgTPC)
- 3 Dimensional Scintillator Tracker (3DST)
- Move the detector(s) (DUNE-PRISM)





# **Near Detector Complex**

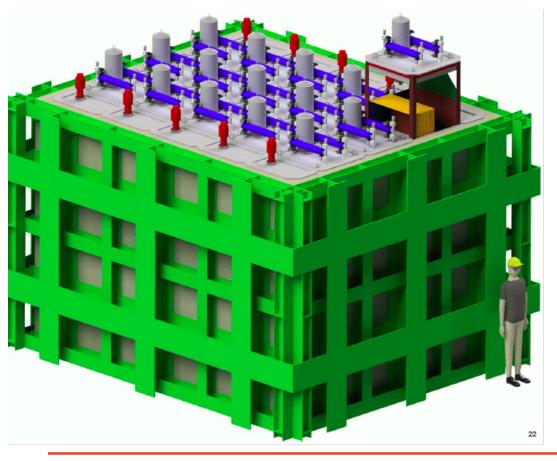


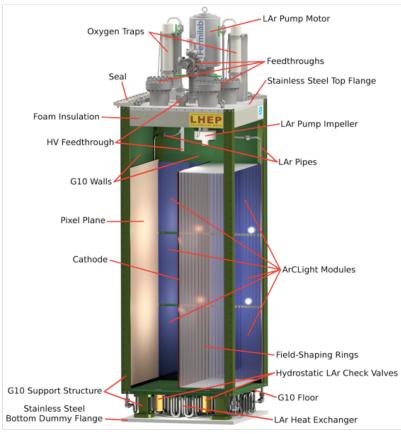
- characterise beam & neutrino interactions & detector response
- LAr TPC (similar to FD)
- High pressure gaseous argon TPC tracker & ECAL & muon systems

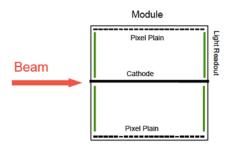




# LATPC (7x5x3 m³)







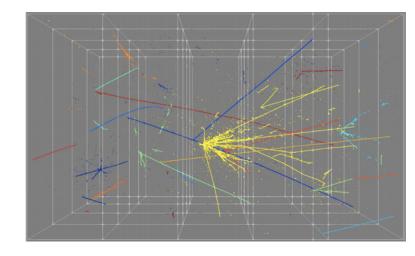


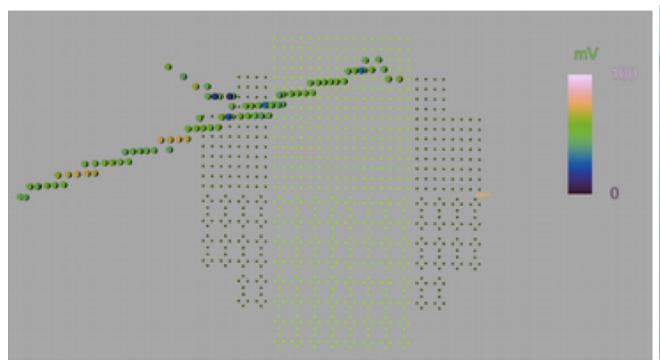


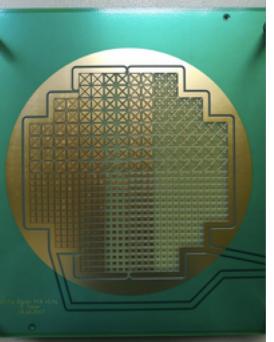


### **Pixel Readout**

- High Intensity requires pixel readout
- Readout pads and dedicated ASIC





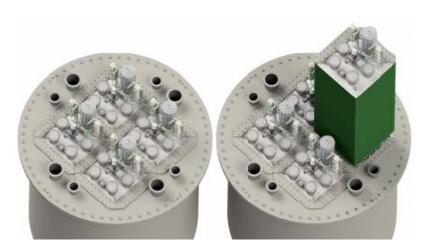


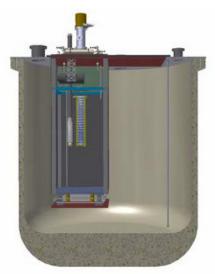






#### **ArgonCube 2X2 prototype (ProtoDUNE-ND)**







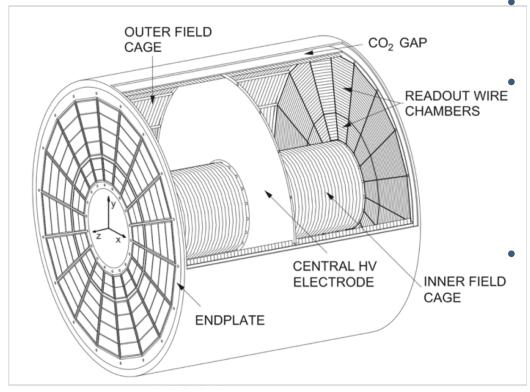
In the laboratory in Bern First cool down done

Will be brought to Fermilab after testing at Bern. To be placed in the NuMI beam MINOS ND Hall





# Multi-purpose detector



HPgTPC pressure vessel surrounded by the 5 coils comprising the Helmholtz coil system.

Not shown: ECAL and μ tagger.

 10 ATM Ar-CH<sub>4</sub> TPC inside cylindrical pressure vessel

#### **ECAL**

- Scintillator-Pb or Scintillator Cu
- ½ inside pressure vessel, ½ outside

SC Helmhotz coil magnet system

- 3 coils for central field
- 2 bucking coils
- Note: continuing optimization study for NC
- μ tagging system



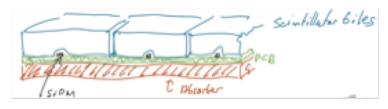


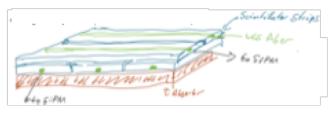


18

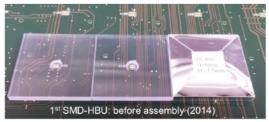
### **ECAL ala CALICE**

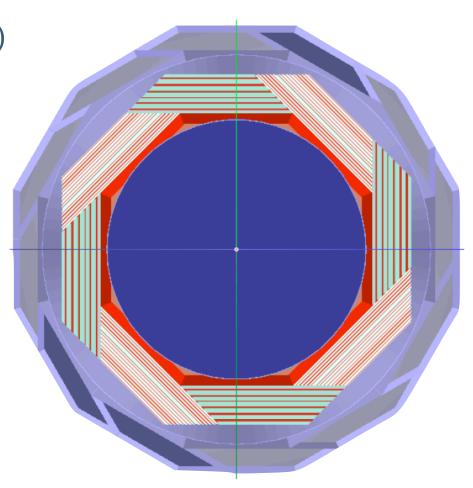
- Surrounds TPC to detect photons and neutrons
- Plastic scintillator tiles (or strips)
- SiPM readout





**CALICE AHCAL concept** 



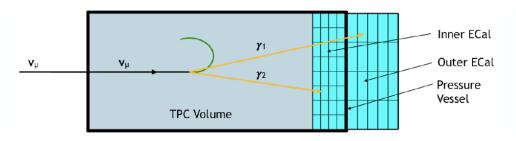




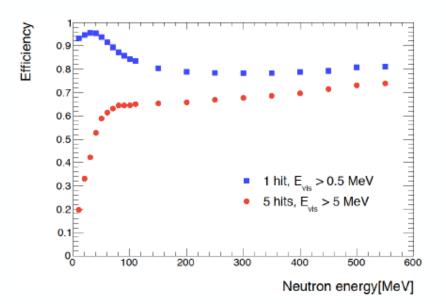


### **ECAL Performance**

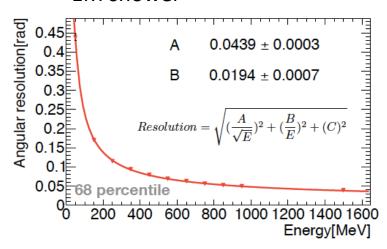
inside and outside pressure vessel (could be all outside, if PV is thin enough)

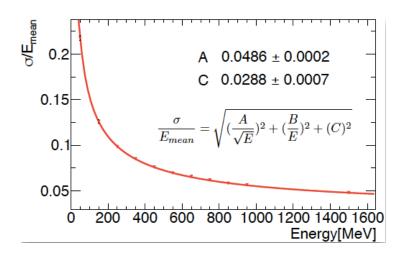


#### neutron



#### EM shower







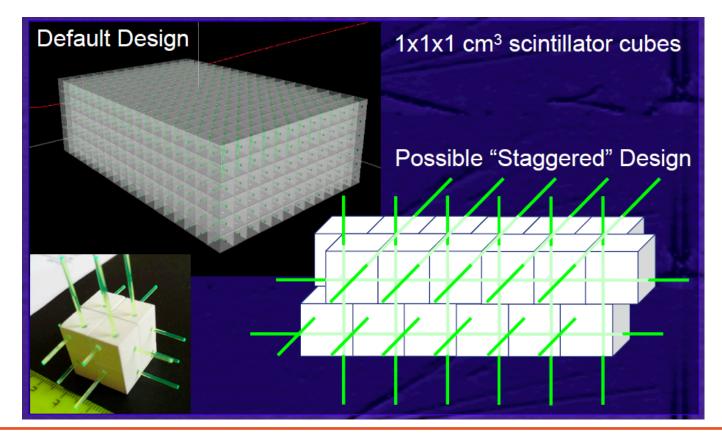




### 3DST

#### Purpose

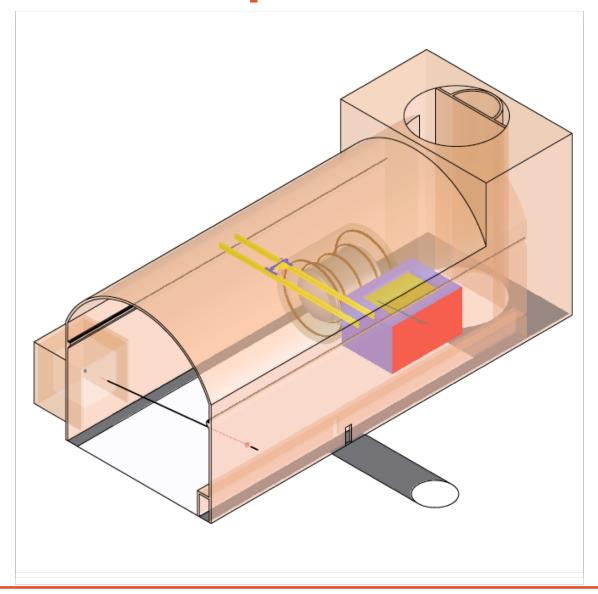
- Beam stability
- Interactions on CH, H-interactions can be identified
- neutrons







# **ND Hall Concept**







#### **DUNE-PRISM**

- Alternative approach
  - Vary flux by changing off axis angle
  - Direct extrapolation
    - Measure product of flux \* cross section \* detector response

$$\frac{dN_{\nu}^{det}}{dE_{rec}} = \int \phi_{\nu}^{det}(E_{\nu}) * \sigma_{\nu}^{target}(E_{\nu}) * D_{\nu_{\mu}}^{det}(E_{\nu}, E_{rec}) dE_{\nu}$$

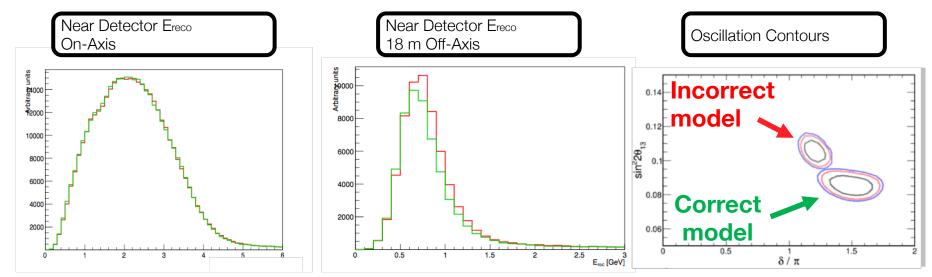
Move LAr detector along different off-axis angles





# **Identify Hidden Systematics**

- Measurements with significantly different flux will help
- Case Study
  - Move 20% of proton energy to neutrons
  - ND (on-axis fit) fixes this by modifying cross section



On-axis all seems fine, but off-axis disagreement

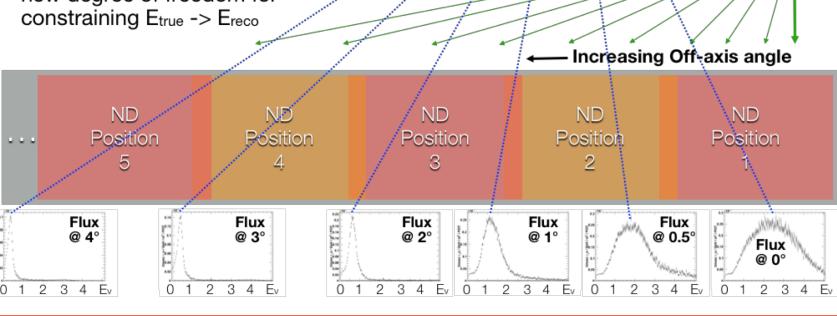




24

# **PRISM Approach**

- A major challenge for DUNE is determining the E<sub>true</sub> -> E<sub>reco</sub> matrix (i.e. not just the ratio)
  - Energy loss due to neutrons, threshold effects, particle ID (e.g. pion mass), etc.
- Making measurements at a variety of off-axis angles provides an entirely new degree of freedom for constraining E<sub>true</sub> -> E<sub>reco</sub>





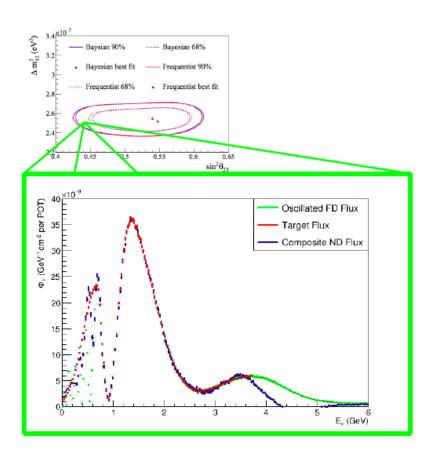


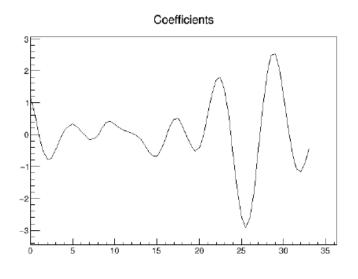


Neutrino Beam

Etrue (GeV)

### Use linear combination





- $\sin^2 \theta_{23} = 0.43$
- Also able to disentangle flux and cross section effects
  - Flux integrated measurements with totally different flux

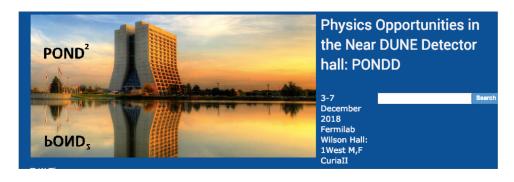




# ND Physics

- Main function of ND complex
  - Support the oscillation physics program
    - Cross section
    - Flux
    - LAr detector systematics
- Physics program of itself
  - Sterile neutrinos
  - Dark photons, dark matter
  - Fundamental cross sections
  - Nuclear physics
  - (limited by your imagination only)

**DUNE Near Detector** 



https://indico.fnal.gov/event/18430/

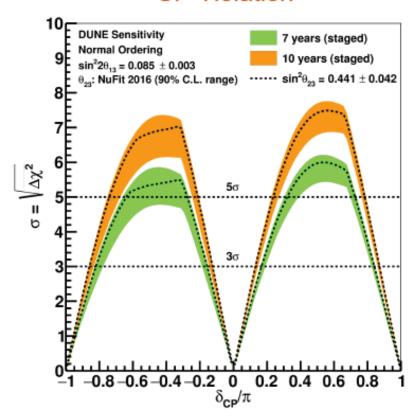




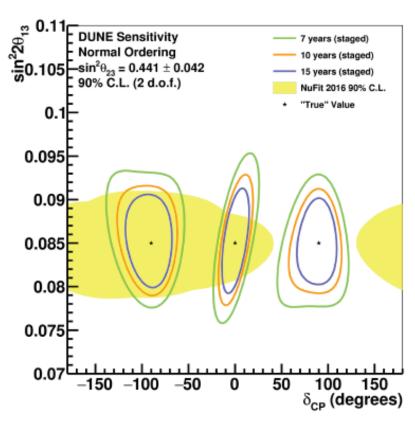
27

# **Oscillation Highlights**

#### **CP Violation**



Width of band indicates variation in possible central values of  $\theta_{23}$ 



Simultaneous measurement of neutrino mixing angles and  $\delta_{CP}$ 





# **Summary and Conclusion**

- DUNE is an neutrino facility with an exciting physics program
- Capable Near Detector suite
  - essential to achieve the oscillations
  - Physics program of itself
- Different detectors to do different jobs
  - LAr TPC to understand detector effects
  - GAr TPC to understand cross sections
  - ECAL to detect neutrals (n, photon)
  - 3DST to understand neutrons and beam stability
  - DUNE-PRISM to measure different fluxes
- Many open questions
  - Get involved and have fun





# Backup

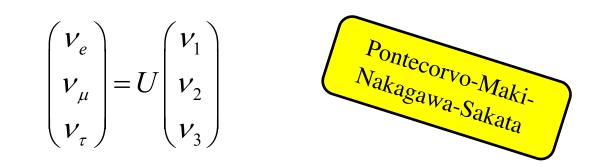




### **Neutrino Mixing The PMNS Matrix**

- Assume that neutrinos do have mass:
  - mass eigenstates ≠ weak interaction eigenstates
  - Analogue to CKM-Matrix in quark sector!

$$\begin{pmatrix} v_e \\ v_\mu \\ v_\tau \end{pmatrix} = U \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$



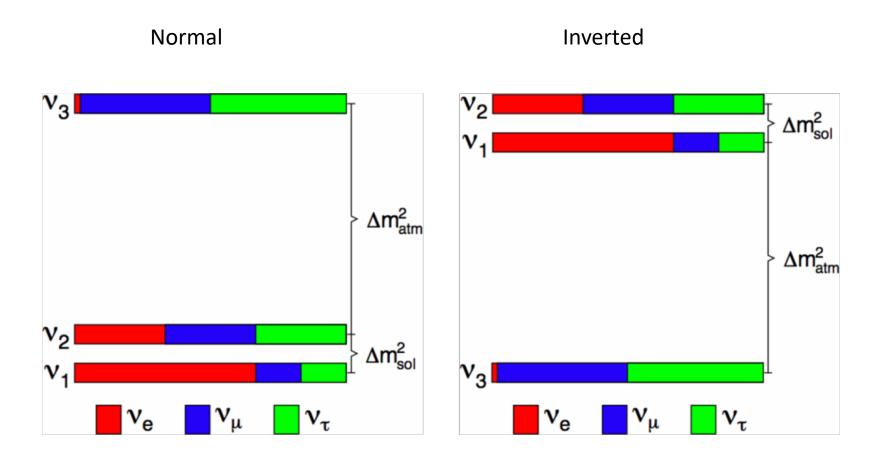
$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\delta_{2}} & 0 \\ 0 & 0 & e^{i\delta_{3}} \end{pmatrix}$$

with  $c_{ij} = \cos(\theta_{ij})$ ,  $s_{ij} = \sin(\theta_{ij})$ ,  $\theta_{ij} = \text{mixing angle and } \Delta m_{ij}^2 = \text{mass}^2$  difference





# **Mass Ordering**







### **Matter Effects**

Simplified treatment: two neutrinos only

In vacuum

$$P(\nu_{\mu} \to \nu_{e}) = \sin^{2}(2\theta) \sin^{2}\left(\frac{\Delta m^{2}L}{4E}\right)$$

in matter

$$P(\nu_{\mu} \to \nu_{e}) = \sin^{2}(2\theta_{m})\sin^{2}\left(\frac{\Delta m_{m}^{2}L}{4E}\right)$$
with  $\sin(2\theta_{m}) = \frac{\sin(2\theta)}{\sqrt{(\cos 2\theta - A)^{2} - \sin^{2}(2\theta)}}$ 

$$\Delta m_{m}^{2} = \Delta m^{2}\sqrt{(\cos 2\theta - A)^{2} - \sin^{2}(2\theta)}$$

$$A = \pm \frac{2\sqrt{2}G_{F}N_{e}E}{\Delta m^{2}}$$

- Matter modifies oscillation probability
  - Sign of mass difference matters (opposite for anti-v)
  - Larger effect at higher energies







#### Schedule/Timeline

#### ★ Costs and technical schedule are understood

- Multiple independent reviews
- FD excavation started

#### **★** Schedule based on a realistic funding profile

- DOE planning line (including large contingency)
- Planned CERN contributions
- Anticipated international contributions

#### **★ International Key Milestones:**

- 2017: start of construction at SURF
- **2018:** operation of two large-scale prototypes at CERN
- 2019: International approval of DUNE funding matrix
- **2021:** start of installation of first 17-kt far detector module
- **2024:** start of operation of 17-kt far detector module
- 2026: start of beam operation (1.2 MW) with two 17-kt FD modules





34

### The DUNE Collaboration

- 1144 collaborators from 178 institutions in 32 countries
- 622 faculty/scientists, 191 postdocs, 106 engineers, 5 computing professionals, 220 PhD students
- Growing at a rate of about 100 collaborators/year



Collaborating Institutions

**Sep 2018** 



Armenia (3), Brazil (29), Bulgaria (1), Canada (1), CERN (32), Chile (3), China (5), Colombia (13), Czech Republic (11), Spain (34), Finland (4), France (23), Greece (4), India (45), Iran (2), Italy (63), Japan (7), Madagascar (8), Mexico (8), The Netherlands (4), Paraguay (4), Peru (8), Poland (6), Portugal (7), Romania (7), Russia (10), South Korea (4), Sweden (1), Switzerland (35), Turkey (2), UK (136), Ukraine (4), USA (621)







## **The DUNE Collaboration**

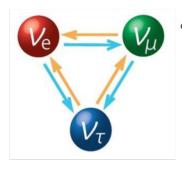




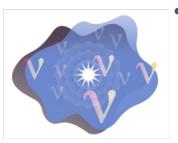




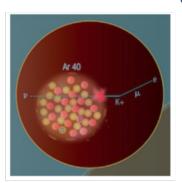
## **Physics Program**



- Neutrino Oscillations
  - Search for leptonic CP violation
  - Determine neutrino mass ordering
  - Precision PMNS measurements



- Supernova Physics
  - Observation of time and flavour profile provides insight into collapse and evolution of supernova
  - Unique sensitivity to electron neutrinos



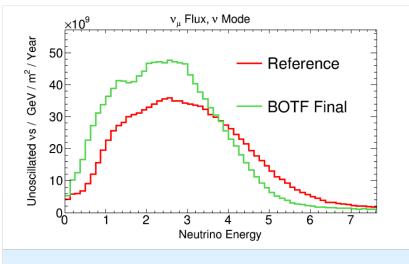
- Baryon number violation
  - Predicted by many BSM theories
  - LAr TPC technology well-suited to certain proton decay channels (e.g.,  $p \rightarrow K + \overline{\nu}$ )
  - $\Delta$ (B-L) ≠ 0 channels accessible (*e.g.*, n $\rightarrow$  $\bar{n}$ )



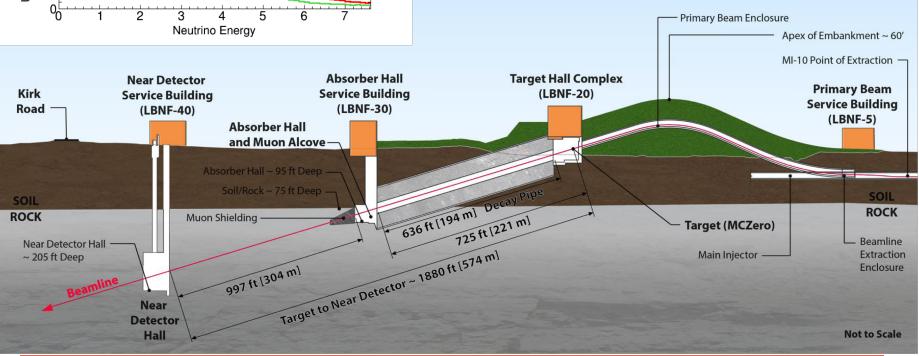




#### Beam



- Proton beam energy 60-120 GeV
- Power 1.2 MW → 2.4 MW
- Neutrinos and anti-neutrinos



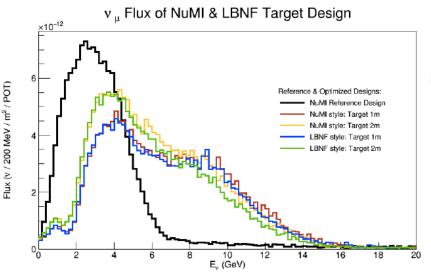


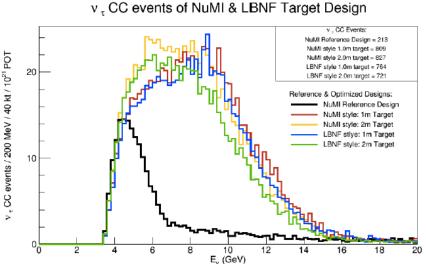




# **High Energy Tune**

- Can change the flux by changing
  - Target positions
  - Horns (shape, position, current)





- Physics
  - Tau appearance
  - NSI

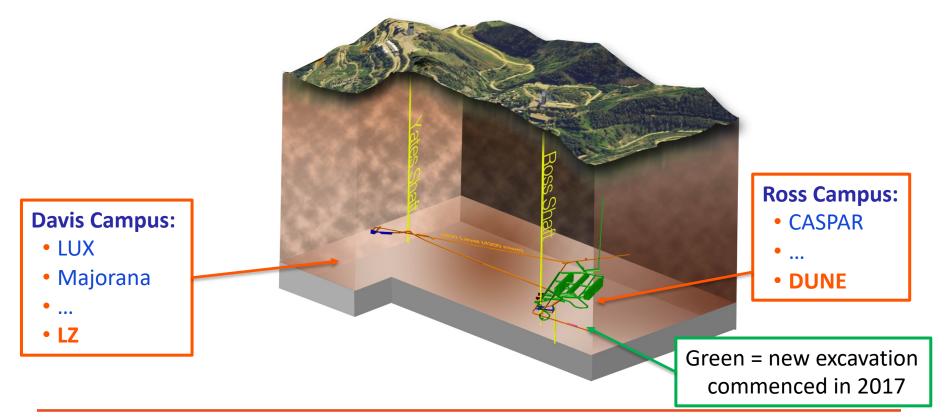




### **Underground Laboratory SURF**

#### **DUNE Far Detector site**

- Sanford Underground Research Facility (SURF), South Dakota
- Four caverns on 4850 level (~ 1 mile underground)



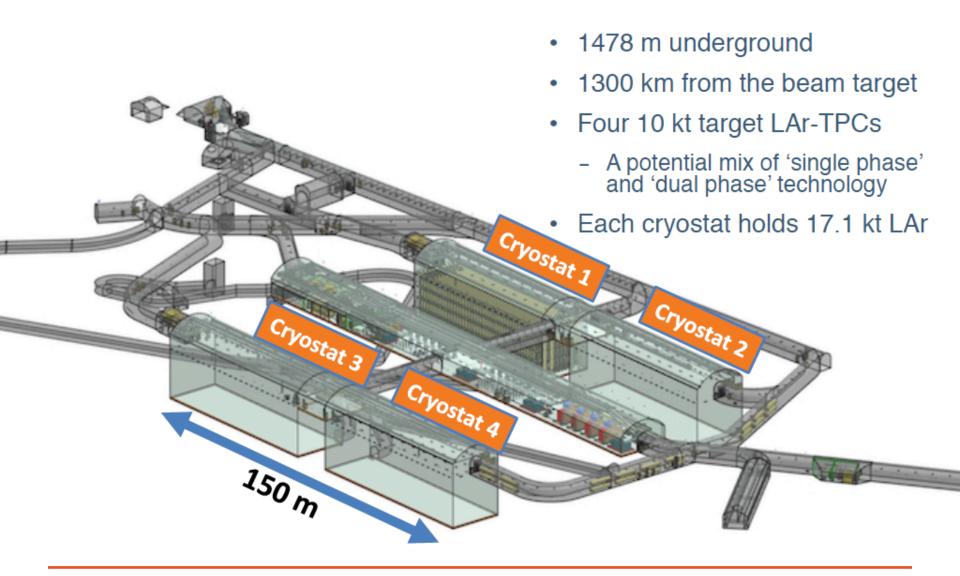






40

#### **DUNE Far Detector**



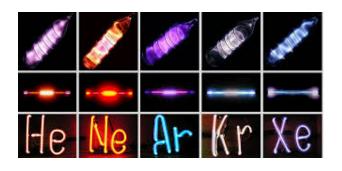






## Why Liquid Argon?

- **Dense**: 40% denser than water
- Cheap: abundant (1% of atmos.)
- lonizes easily:
   55,000 electrons/cm
- Excellent scintillation:
   20,000 photons/MeV
   (@ 500 V/cm)









#### Flux Measurements

- Can be done on any target
- Needs known cross section

 $\phi_{\nu_x}^{near}(E_{\nu})$ 

- Proposed measurements
  - Neutrino-electron scattering
    - (electro-)weak cross section calculable with high precision
    - Cross section is very small
  - Low-v method
    - Cross section is independent of energy in limit where neutrino energy is FS lepton energy. No hadronic energy ( $\nu$ )
    - Cross section depends on ability to measure all hadronic energy including neutrons
    - Breaks down at low energy (ν<Ε<sub>cut</sub>)





### Flux Measurements (II)

**Coherent Scattering** 

$$\nu_l + N \to l^- + \pi^+ + N$$

$$\overline{\nu}_l + N \to l^+ + \pi^- + N$$

$$\nu_l + N \to \nu_l + \pi^0 + N$$

$$\overline{\nu}_l + N \to \overline{\nu}_l + \pi^0 + N$$

- Does not depend on internal nuclear structure
  - Long/short range correlation
  - Final state interactions
- Cross section
  - Not as well known
  - Small (momentum transfer to nucleus is small)
- Allows flavour identification





## Flux Measurements (III)

- Measurement on hydrogen
  - Interactions on hydrogen can be identified
    - no initial state transverse momentum
  - Produce ∆-resonances

• 
$$v_l + p \to l^- + \Delta^{++} \to l^- + p + \pi^+$$

• 
$$\bar{v}_l + p \to l^+ + \Delta^0 \to l^+ + p + \pi^-$$

- Small-ish cross section at low energies
- No final state interaction.
- But even cross section on free nucleus has uncertainties
  - Higher resonances, non-resonant contributions, interference





#### **Cross Section and Smearing Matrix**

We don't know how to disentangle them

$$\sigma_{\nu_e}^{Ar}(E_{\nu}) * T_{\nu_e}^{far}(E_{\nu}, E_{rec})$$

- Unknown initial state:
  - Neutrino energy and "nucleon" momentum
- Missing final state particles/energy:
  - neutrons, low-energy protons, (nuclear recoil)
- Detector smearing
- Approach
  - Measure as many exclusive differential cross sections as possible
  - Tune/build a model
  - Take differential cross section and reco2true from tuned MC and detector simulation





### **Cross Section and Smearing Matrix**

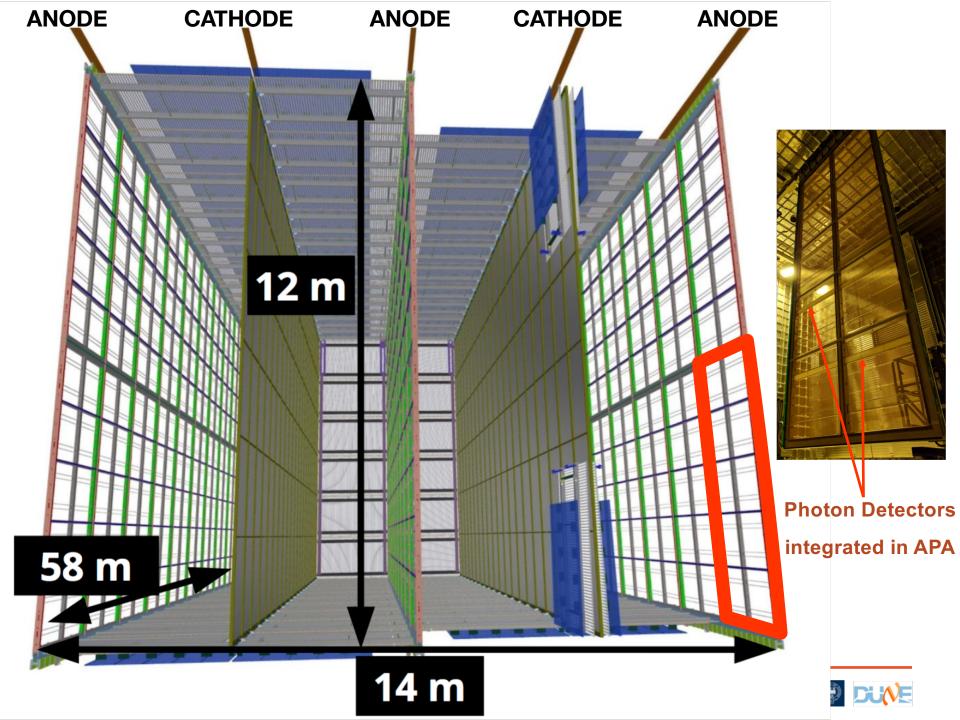
- Can't be measured as function of true neutrino energy
  - Neutrino and nuclear energies are unknown
  - Invisible final state particles
    - neutron, nuclear recoil, FSI
  - Is intrinsically model dependent
- Only possible flux integrated event rates

$$\frac{dN}{dX_{rec}} = \int \phi_{\nu_{\mu}}^{near}(E_{\nu}) \frac{d\sigma_{\nu_{\mu}}^{Ar}}{dX}(E_{\nu}) T_{\nu_{\mu}}^{near}(E_{\nu}, X, X_{rec}) dE_{\nu} dX$$

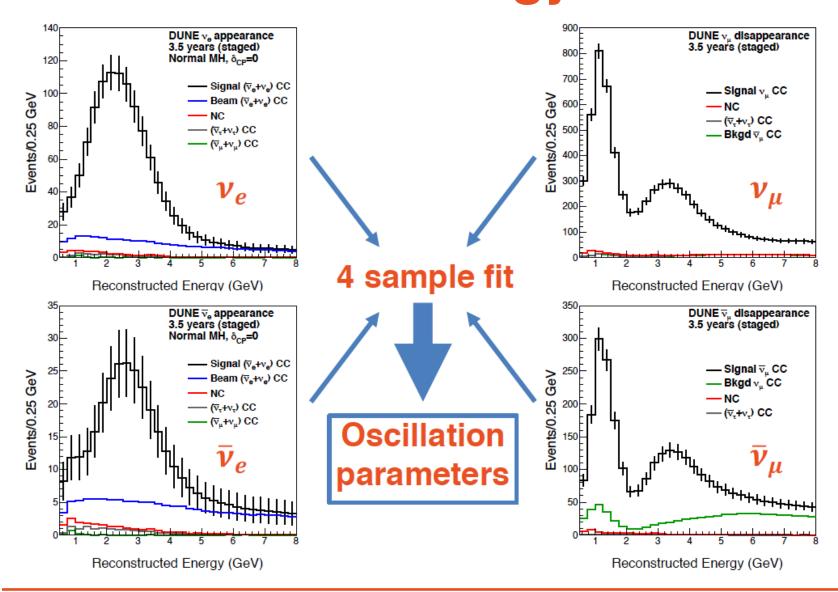
New standard in neutrino physics







### **Measurement Strategy**



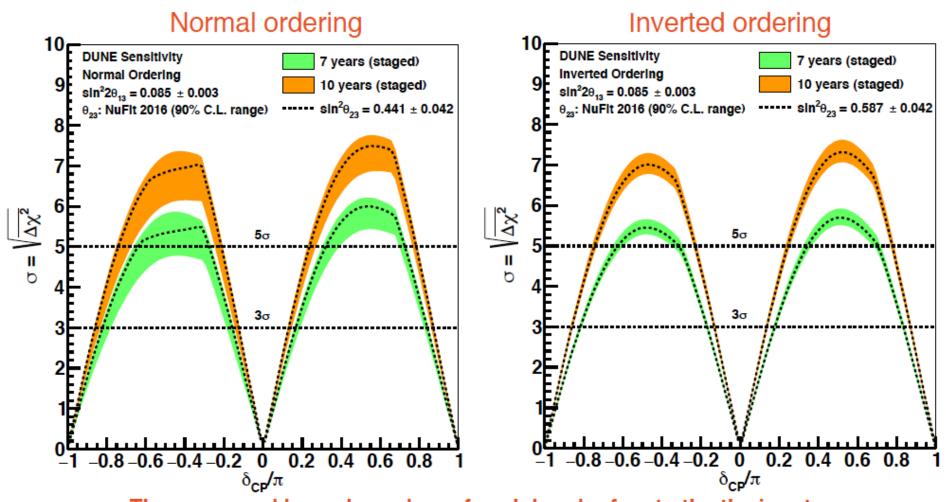
**DUNE Near Detector** 







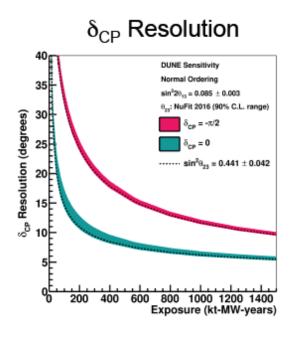
# **CP Sensitivity**

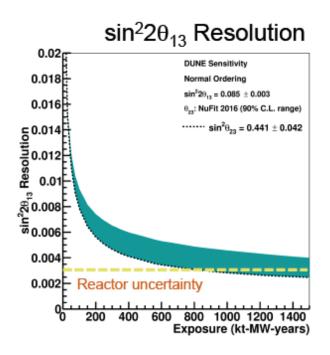


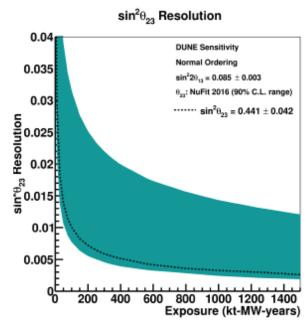




## **Oscillation Highlights**



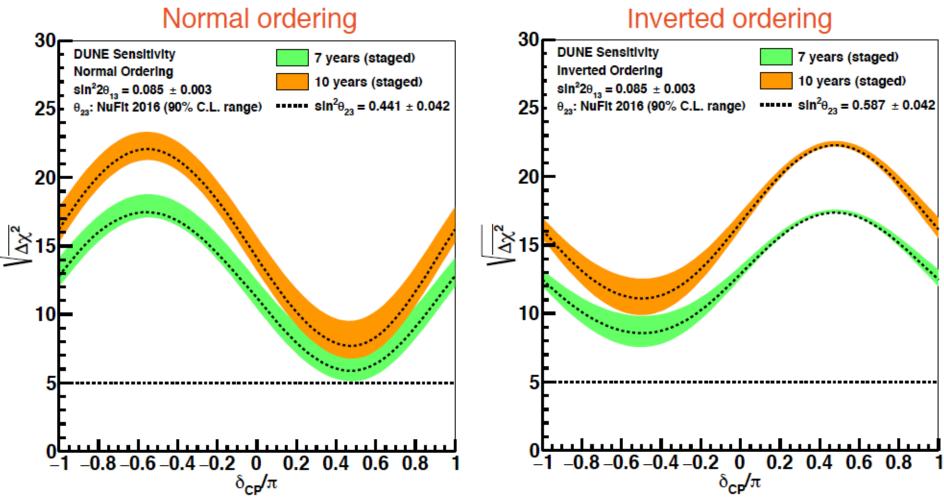








## **Mass Ordering**



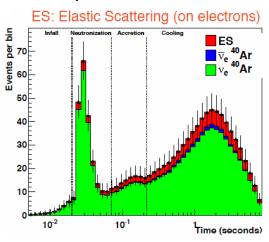
The upper and lower boundary of each band refers to the the input  $\theta_{23}$  maximum and minimum respectively



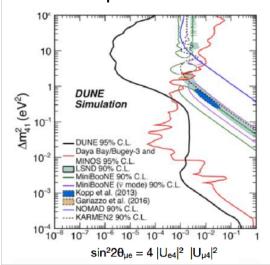


## **Other Physics**

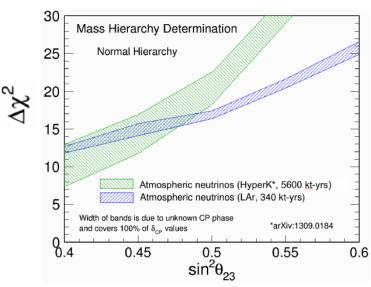
#### supernova



#### atmospherics



#### atmospherics



- Dark matter
- Large extra dimensions
- Dark photons
- NS interactions



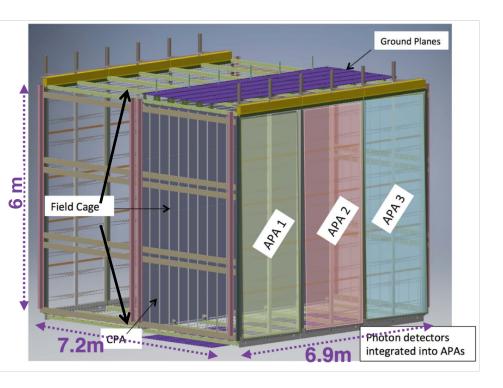




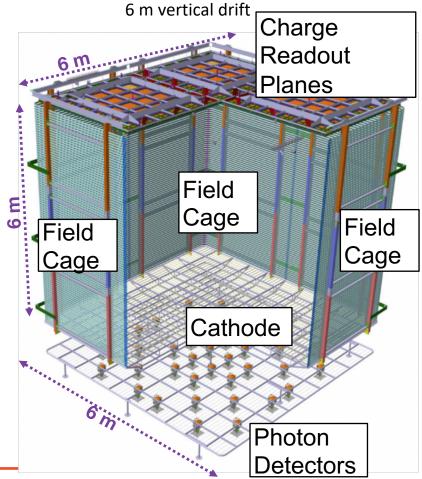
# **Two Technologies**

#### Single-Phase

3.6 m horizontal drift



#### **Dual-Phase**









# **July 2018**









# **Empty Cryostat**

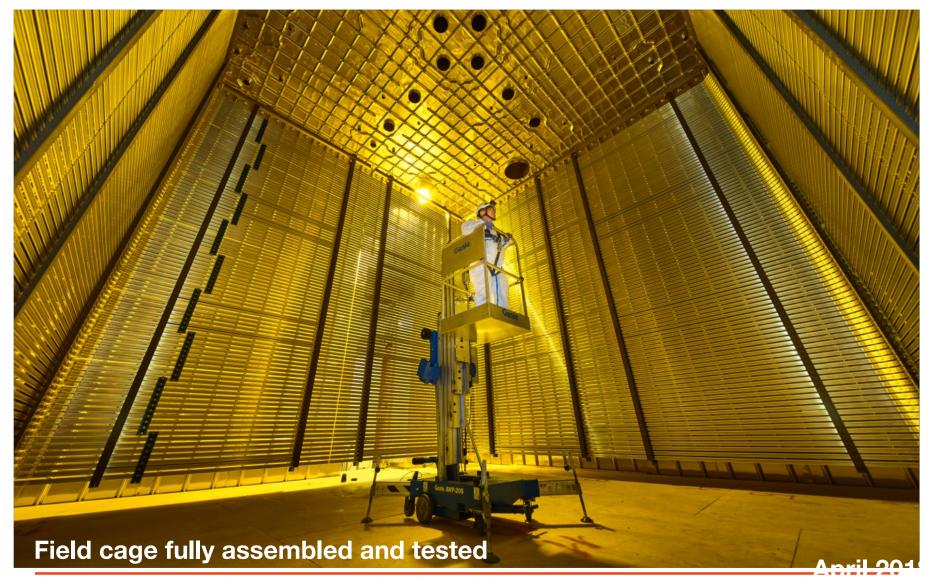








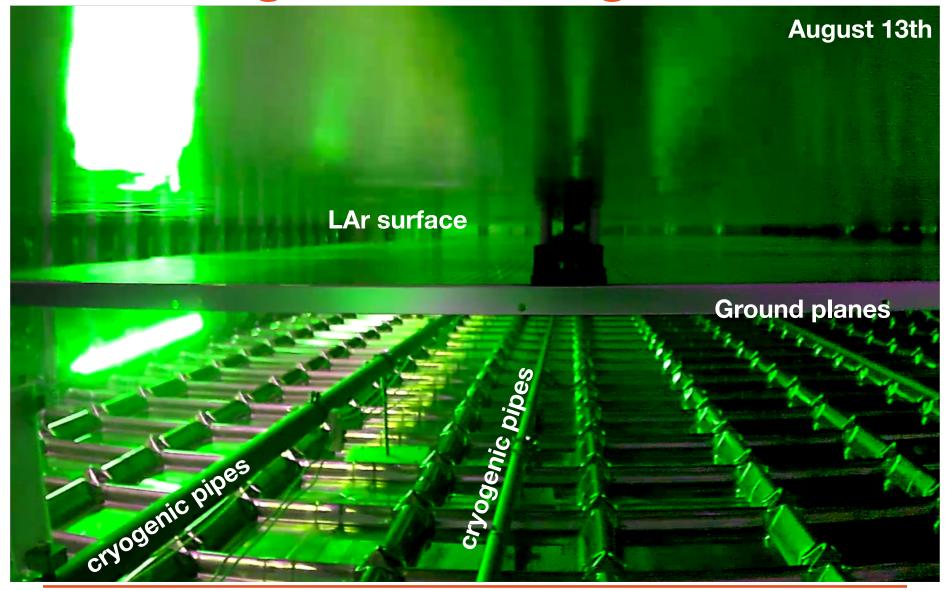
### **ProtoDUNE-DP**







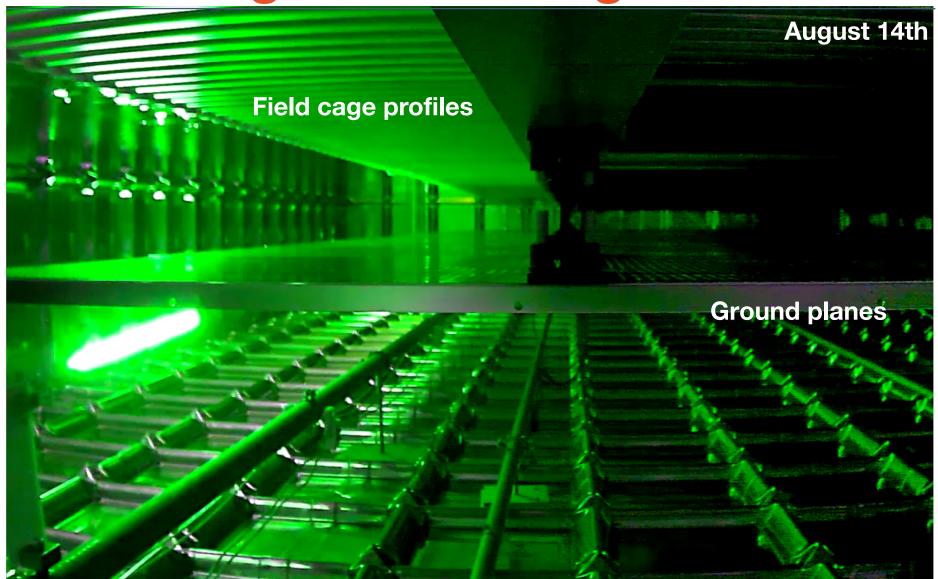
## Yellow light becomes green







# Yellow light becomes green

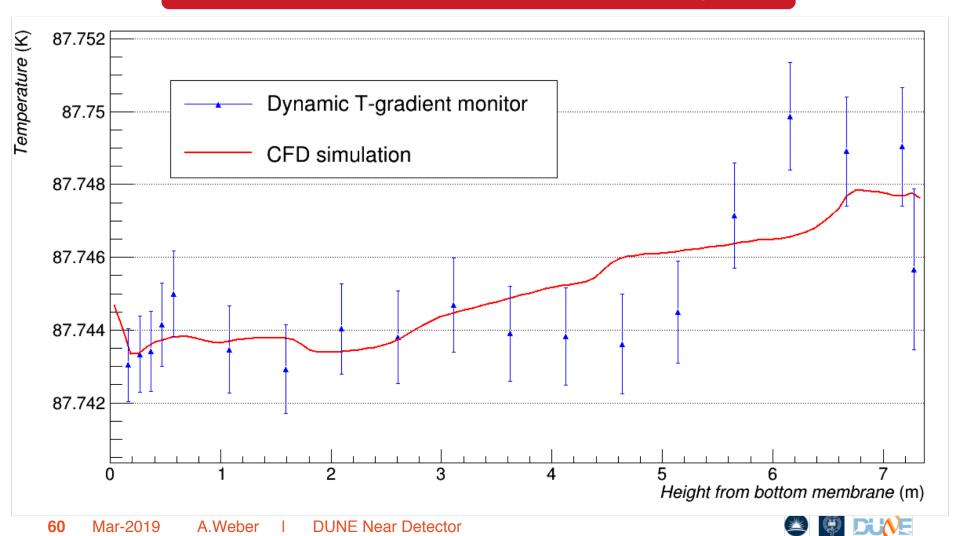




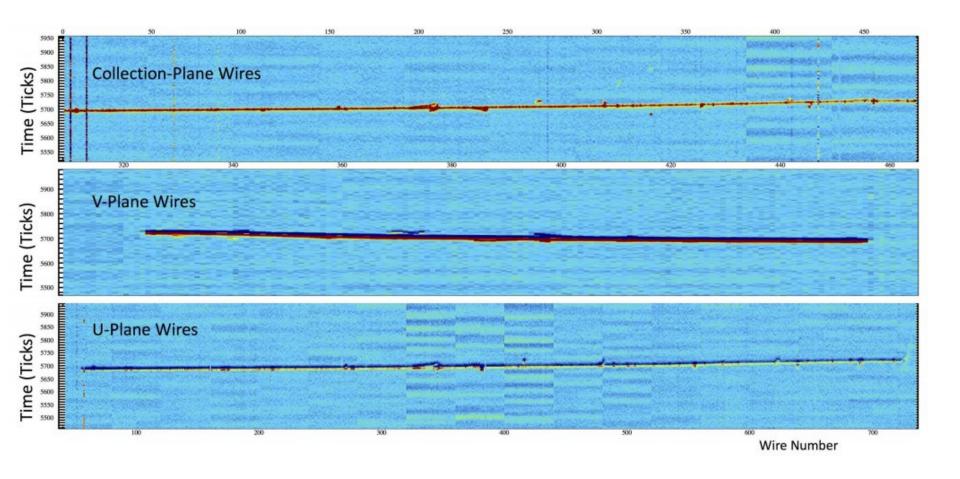


## Liquid Argon temperature

#### **Temperature varies < 0.01 K across the cryostat**



#### **The First Event**

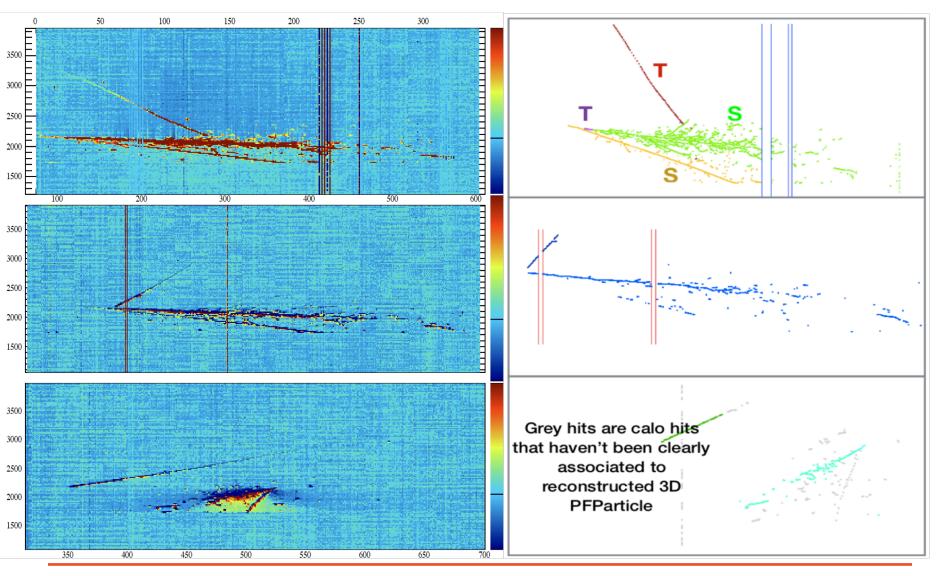








### **Automatic Reconstruction**



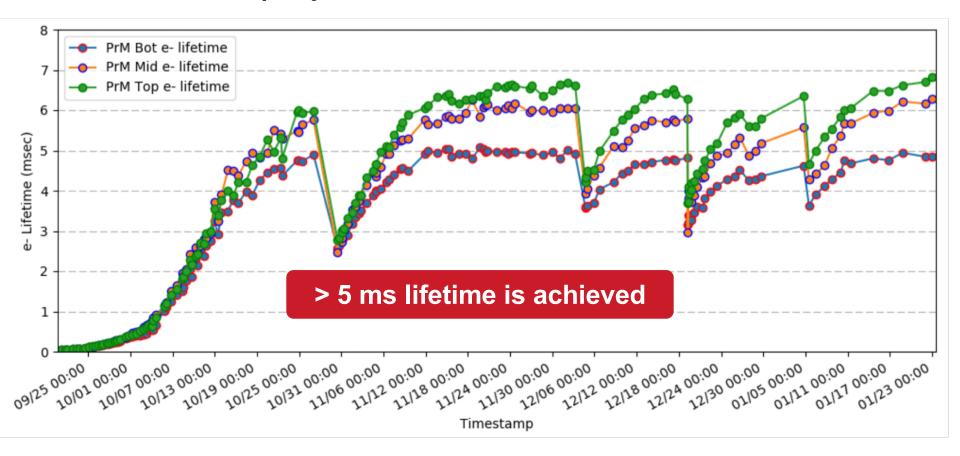




62

# **Liquid Argon Purity**

#### The purity is measured as the electron lifetime



**Electrons need 3 ms to cross the drift volume** 



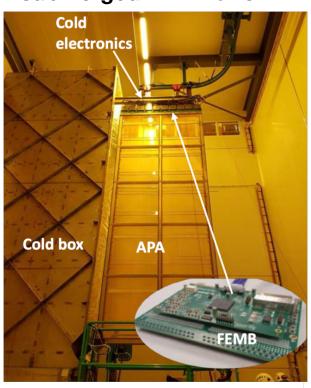


#### **APAs and Cold Electronics**

Exceptionally low noise operation and scalable cryostat design

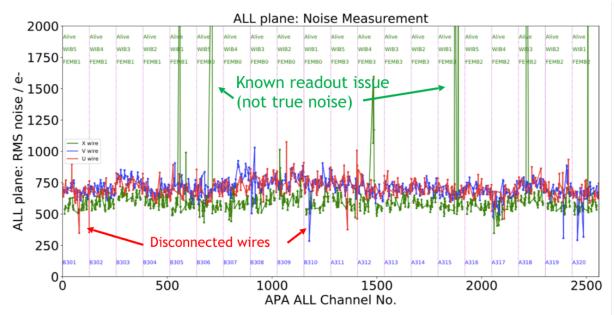
~ 15000 wires, only 4 channels dead (0.03%)

#### **Electronics on top of APAs** submerged in LAr at 87 K



 $ENC < 750 e^- \longrightarrow S/N \sim 20$ 

meets DUNE requirements (S/N>10)











#### **Very preliminary !!!**

#### dE/dx for 1 GeV/c beam protons

