

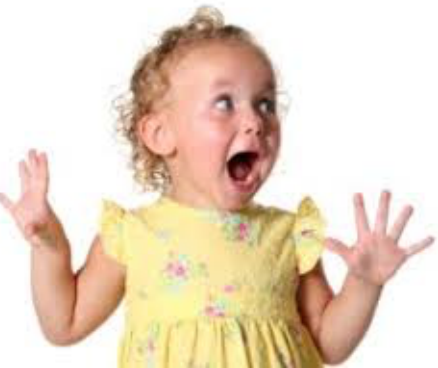
Proposal of New BSM Searches at the LHC and Belle-II

Filippo Sala

DESY Hamburg



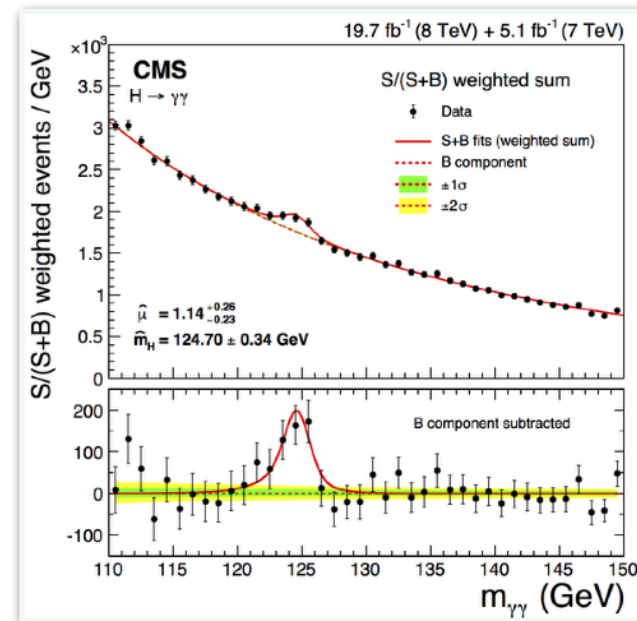
BSM at the LHC?



2010

2012

Bound
Exclusion
...



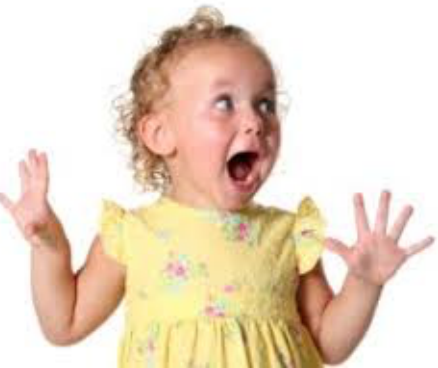
2018

It's Standard Model
Limit

No New Physics
at these energies
...

?

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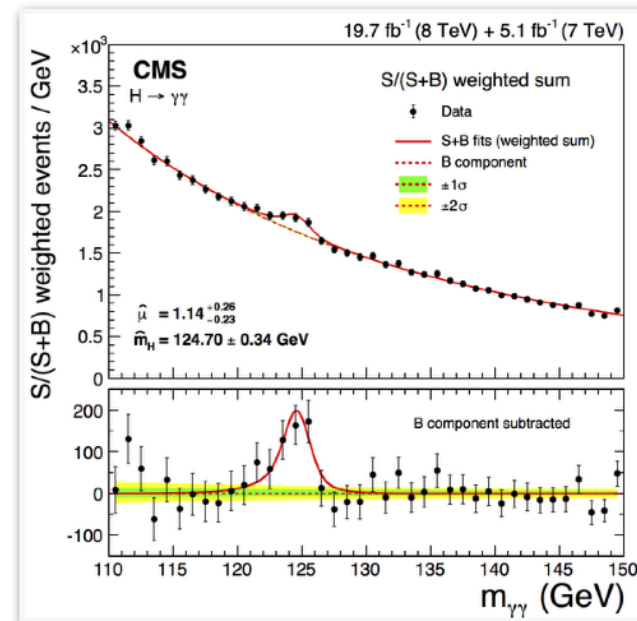


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(Most?) **Solid discovery method** at colliders
Look for peaks in invariant mass distributions

Diphotons

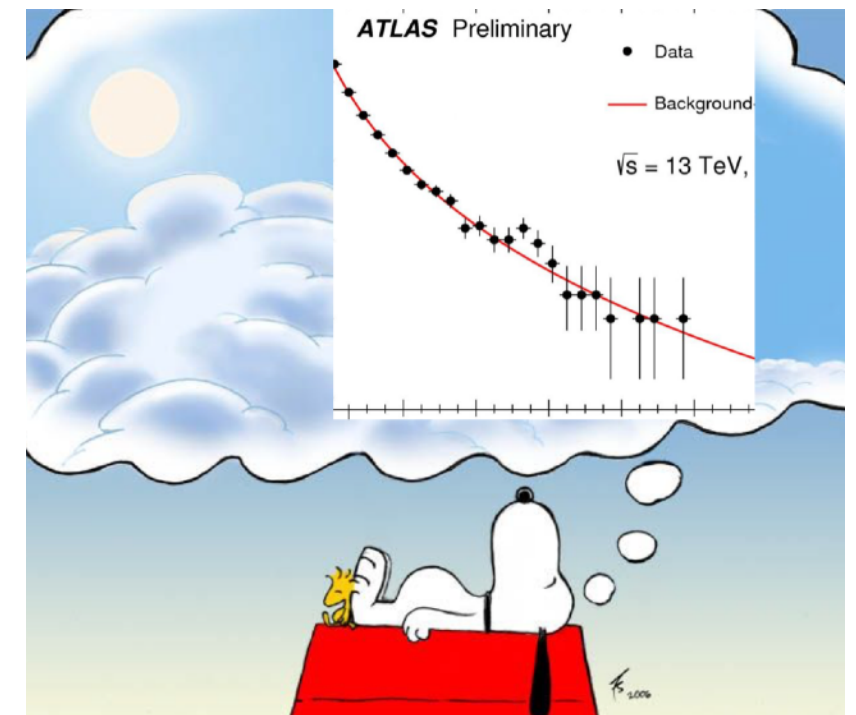
Dijets

Dibosons (W,Z)

Dileptons

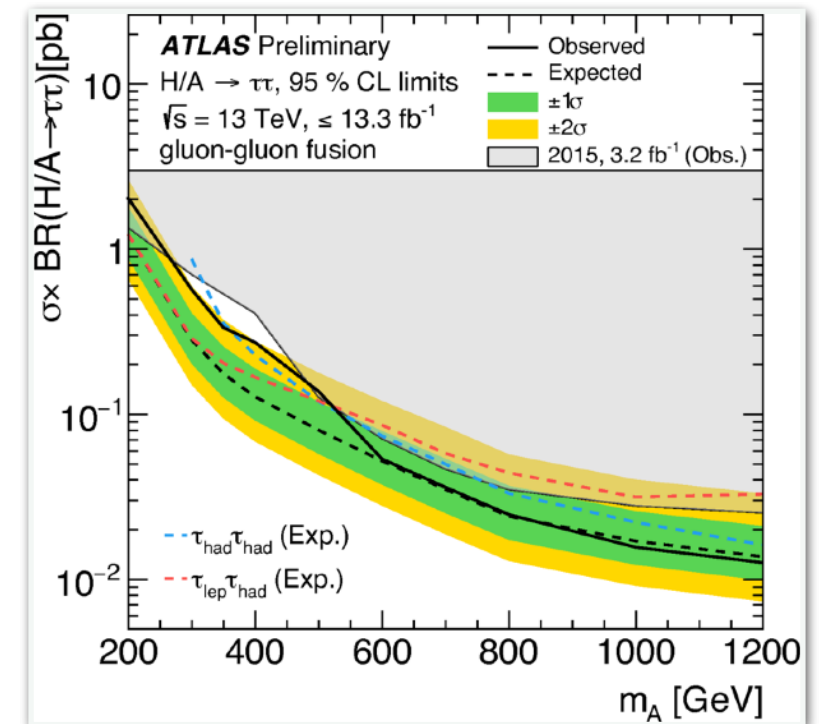
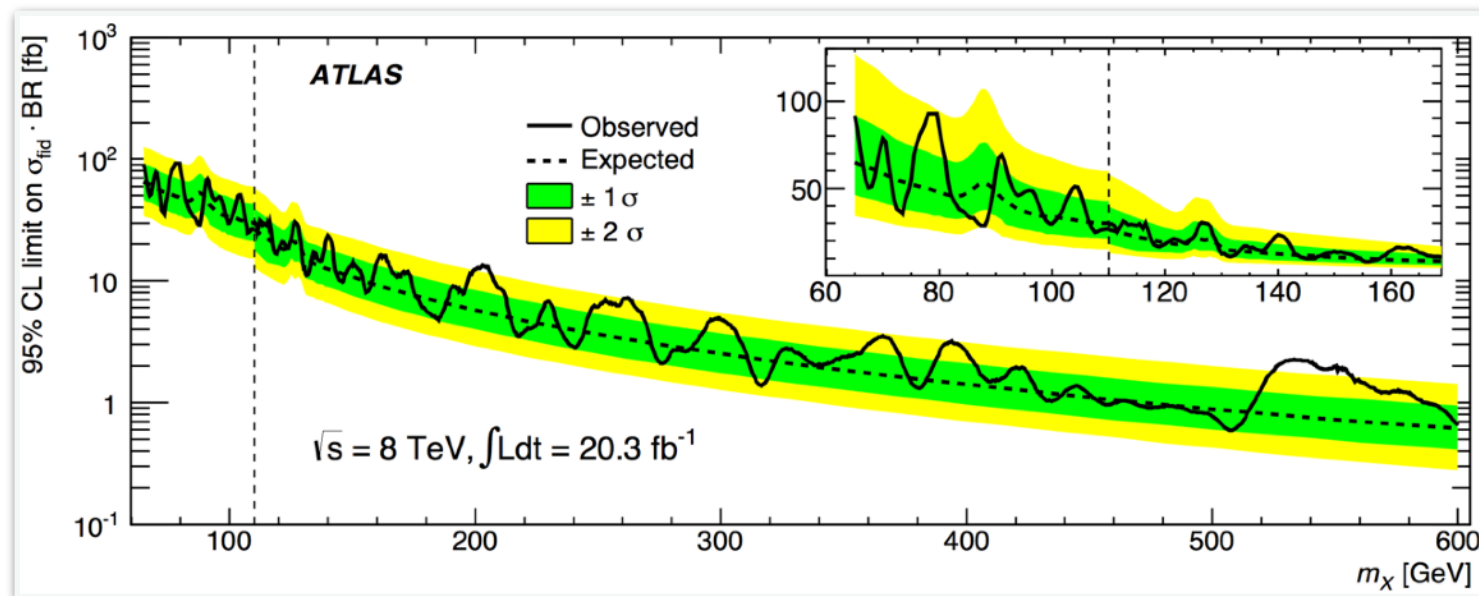
Ditop

DiHiggs



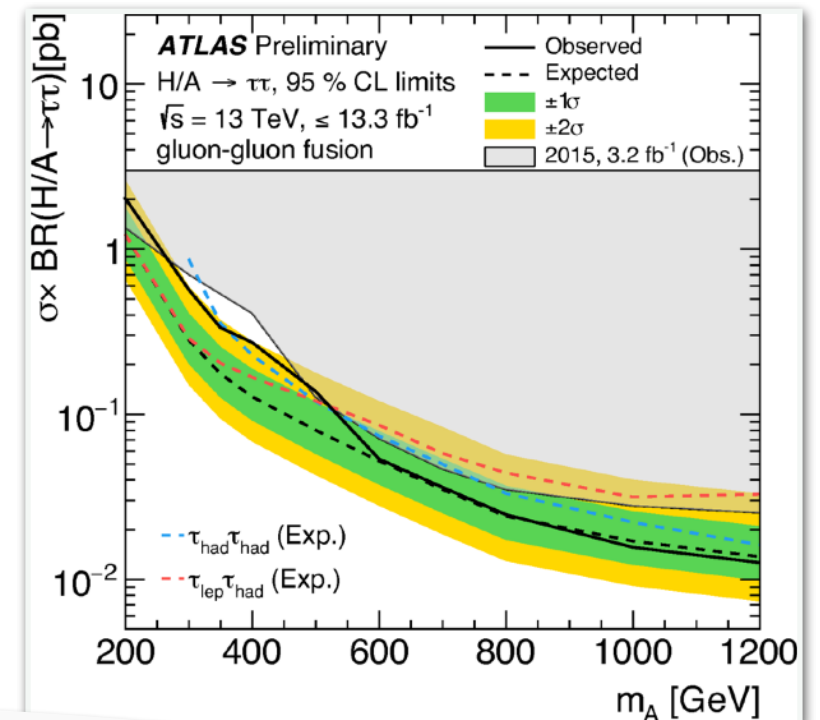
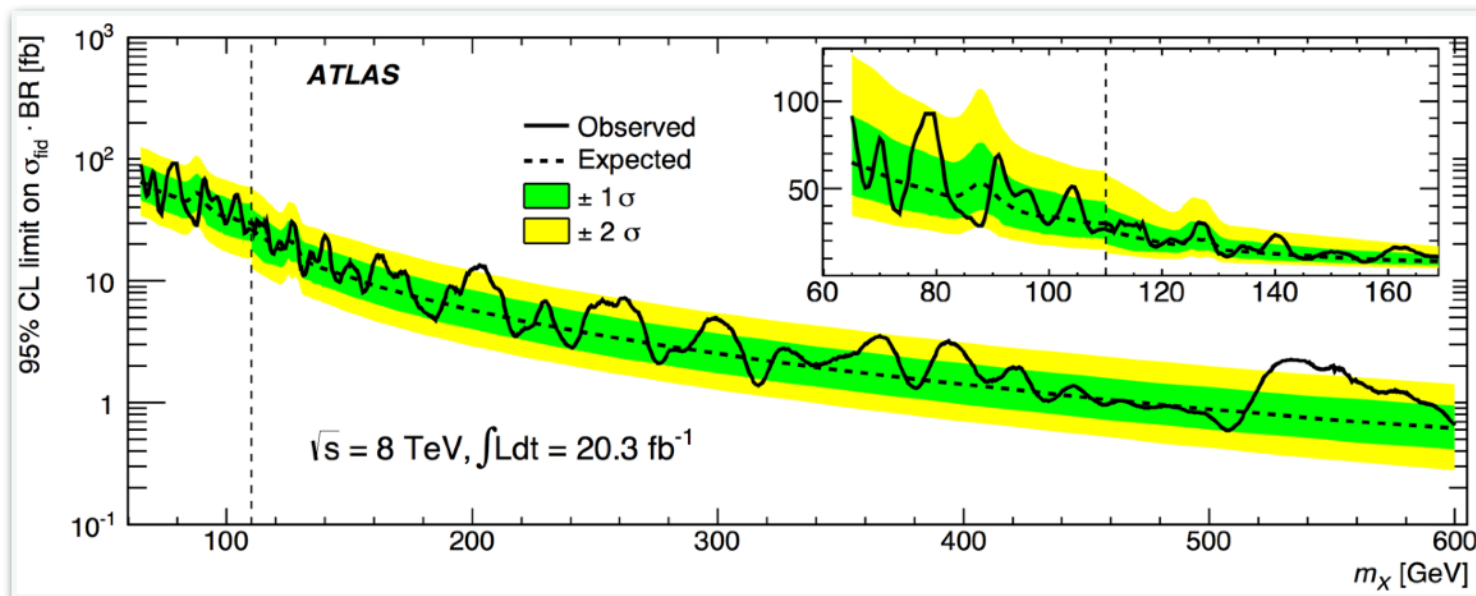
BSM resonances: where to look?

Present searches: $M_{\gamma\gamma, \tau\tau, \dots} > O(100) \text{ GeV}$



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1. Theory bias towards high masses

Why not $M_X < O(100) \text{ GeV}$? 2. “Low-mass already constrained by previous colliders (LEP,...)”

3. “It is very difficult!” Minimal pT cuts, ...

Here: demystify 1. 2. and 3., **new limits** and prospects at LHC

1. Theory bias towards high masses

Why not $M_X < O(100)$ GeV ?

Why low-mass resonances?

LHC is pushing **solutions to SM problems** (hierarchy, flavour, ...) to $M_{\text{BSM}} \gg \text{TeV}$

Richer sectors could exist there

How to test them?

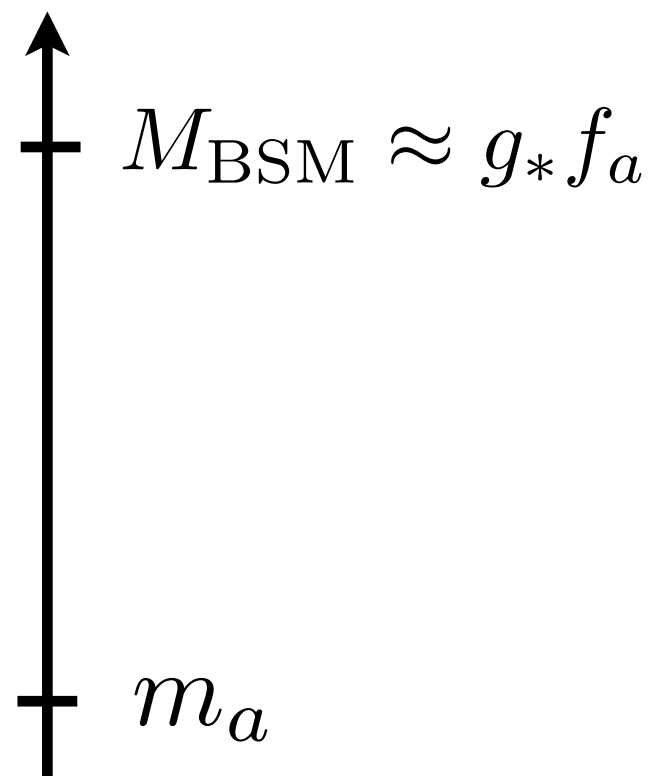
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via ALPs!



aka **P**seudo **G**oldstone Bosons (**PGBs**)



from spontaneous breaking of global symmetry

with a small explicit breaking that controls $m_a \neq 0$

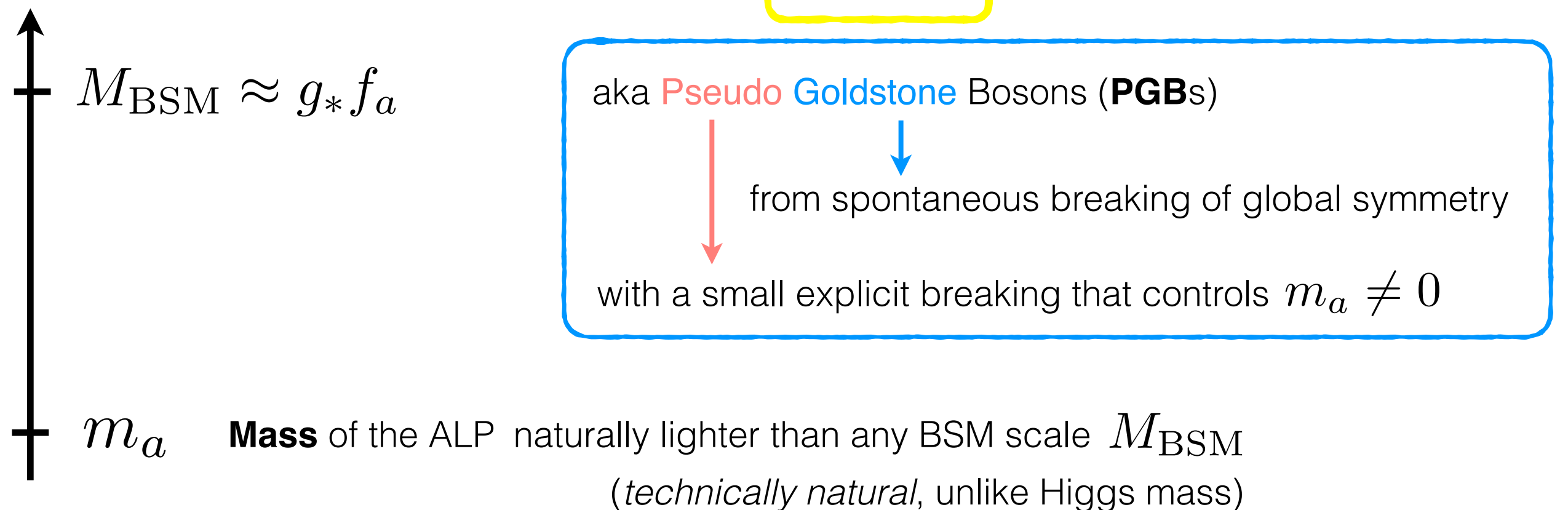
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f_a decay constant controls ALP **Couplings**

$$\text{e.g. } \mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\alpha_s c_3 G\tilde{G} + \alpha_2 c_2 W\tilde{W} + \alpha_1 c_1 B\tilde{B} \right]$$

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$M_{\text{BSM}} \approx g_* f_a \gtrsim \text{TeV}$ ~ from LHC exclusions

from spontaneous breaking of global symmetry

with a small explicit breaking that controls $m_a \neq 0$

$m_a \sim 1 - 100 \text{ GeV}$ ~ an unexplored range

$f_a \sim 0.1 - 100 \text{ TeV}$ of interest for colliders (see rest of talk)

e.g. $\mathcal{L}_{\text{int}} = \frac{1}{4\pi f_a} [\alpha_s c_3 G\tilde{G} + \alpha_2 c_2 W\tilde{W} + \alpha_1 c_1 B\tilde{B}]$

ALPs and BSM: strongly coupled

They already exist: pions from QCD

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“**Just because**” strong sector: vector-like confinement

see e.g. [Kilic Okui Sundrum 0906.0577](#)

[add gauge group that confines at $\gtrsim$ TeV, w/new fermions, vector-like to satisfy EW precision tests]

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Natural strong sector: composite Higgs models

Example: $SO(6)/SO(5)$ has 5 PGB, the Higgs and a singlet η

see e.g. [Gripaios+ 0902.1483](#)
[Redi Tesi 1205.0232](#)

$$\text{No tuning in } \eta \text{ potential} \Rightarrow m_\eta \sim m_h \times \frac{f}{v} \sim 600 \text{ GeV} \times \sqrt{\frac{0.05}{(v/f)^2}}$$

with dependence on top representation

$$\text{e.g. if only bottom contributes: } m_\eta \sim 10 \text{ GeV} \times \sqrt{\frac{0.05}{(v/f)^2}}$$

Larger coset structures have more PGBs

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Less natural composite Higgs models:

DM & GUT

[Bernard+ 1409.7391](#)

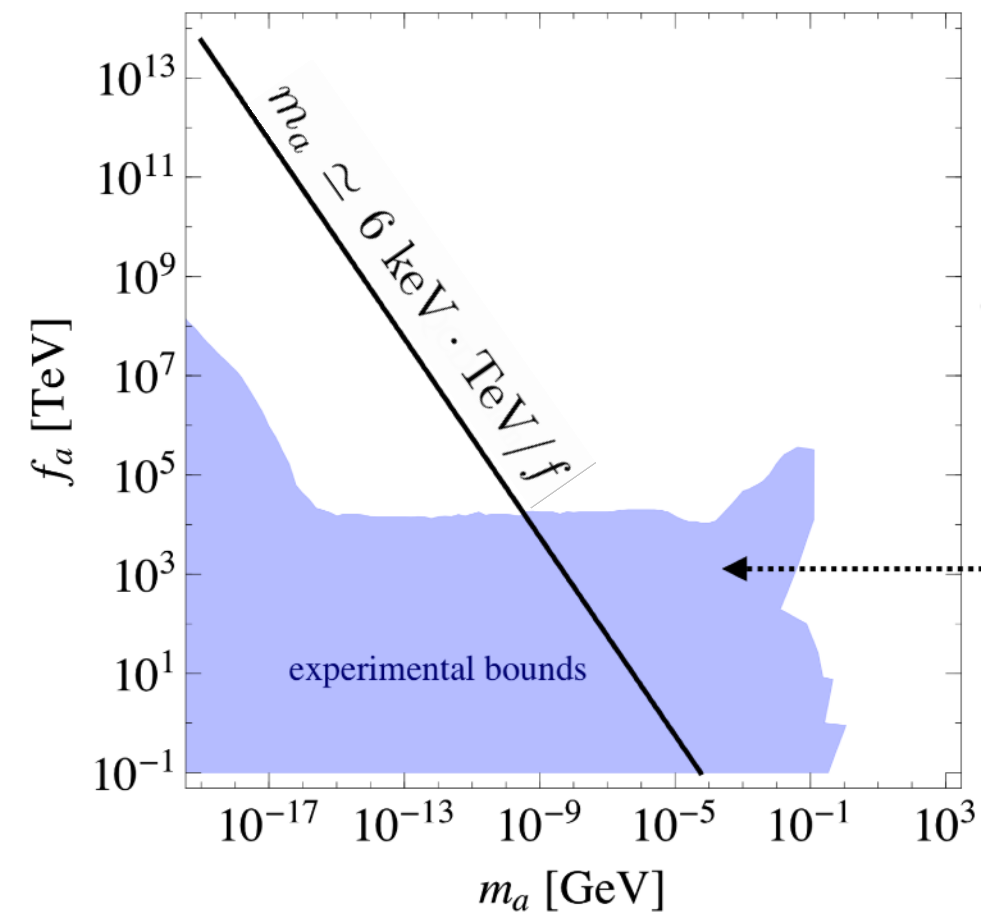
give up on little hierarchy and focus on generate EW & DM scales [Antipin+1410.1817](#)

....

ALPs and BSM: QCD axion

Standard QCD axion points to large f_a and small m_a

$$V_a \simeq -\Lambda_{\text{QCD}}^4 \cos \frac{Na}{f}$$



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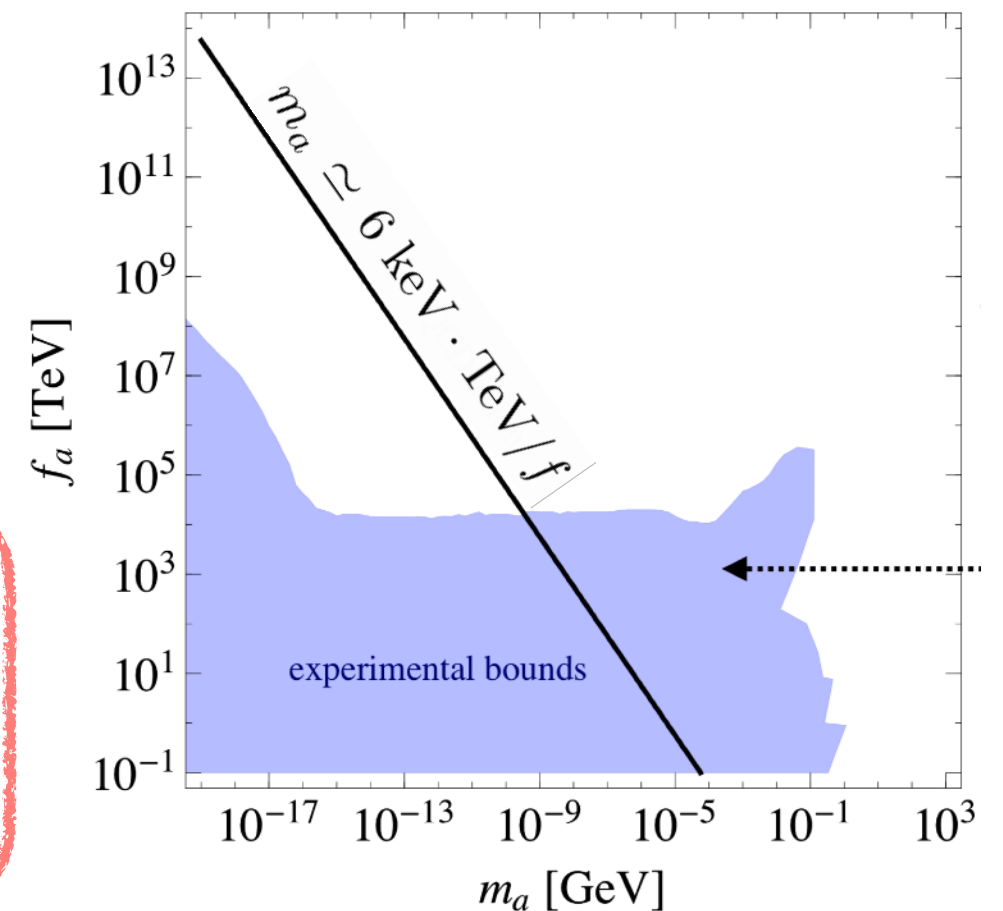
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but: “**Axion Quality**” problem spoils solution to strong CP

$$\Delta V_{PQ} = \lambda_\Delta \frac{\Phi^\Delta}{\Lambda_{\text{UV}}^{\Delta-4}} + \text{h.c.} \quad \Phi = \frac{f_a}{\sqrt{2}} e^{ia/f_a}$$

Kamionkowski March-Russel hep-th/9202003,...



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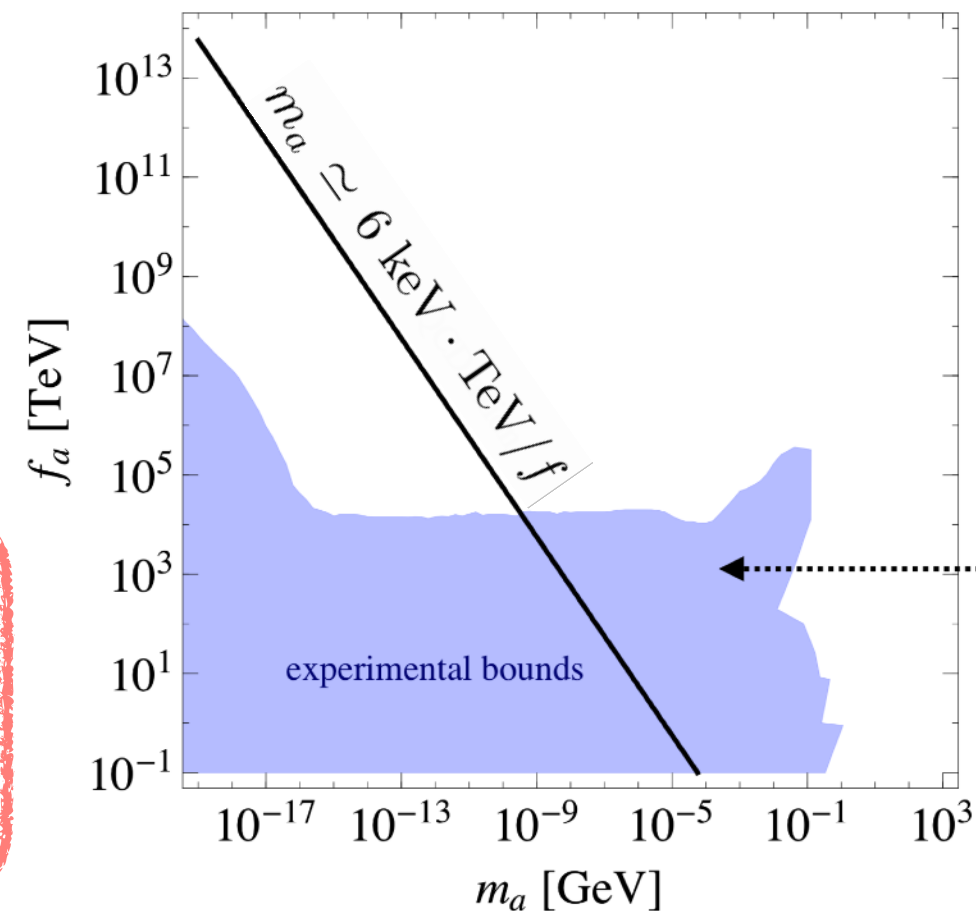
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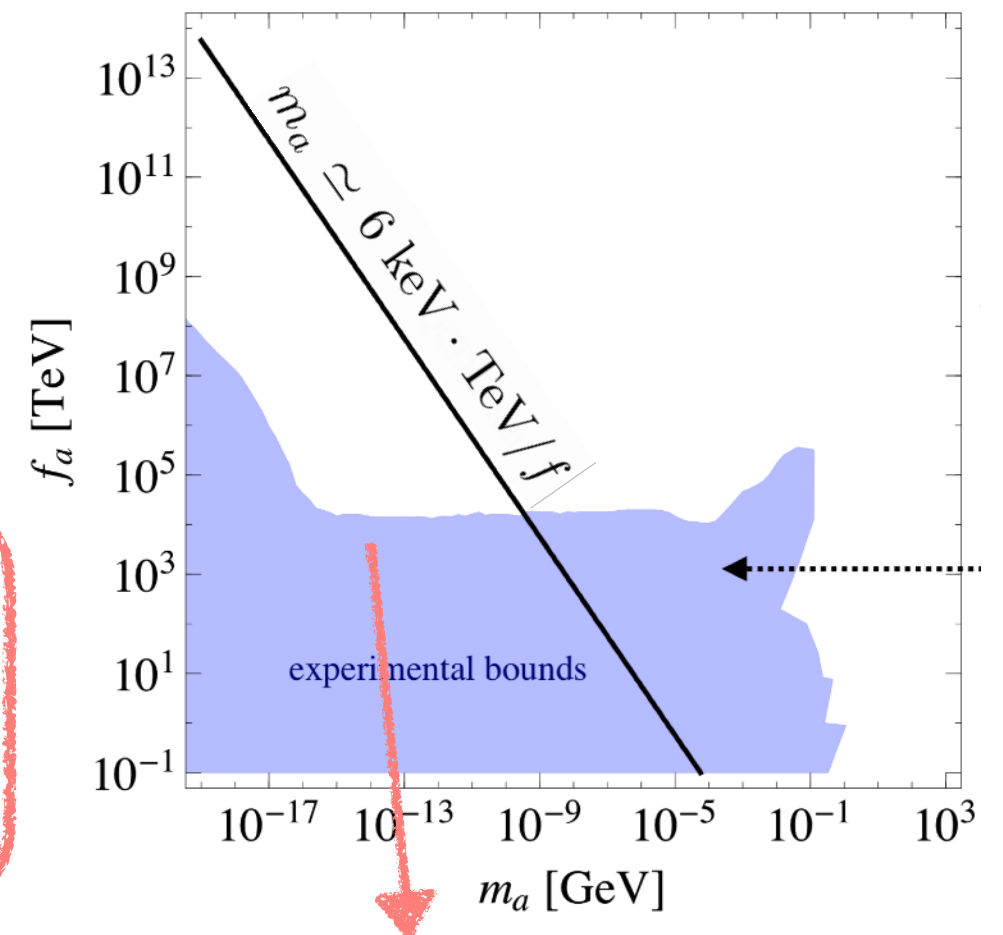
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1. Push Δ to much larger values

... Redi Sato 1602.05427

Duerr+ 1712.01841

2. Model building for smaller f_a
and larger m_a

Rubakov hep-ph/9703409

M.K Gaillard+ 1805.06465

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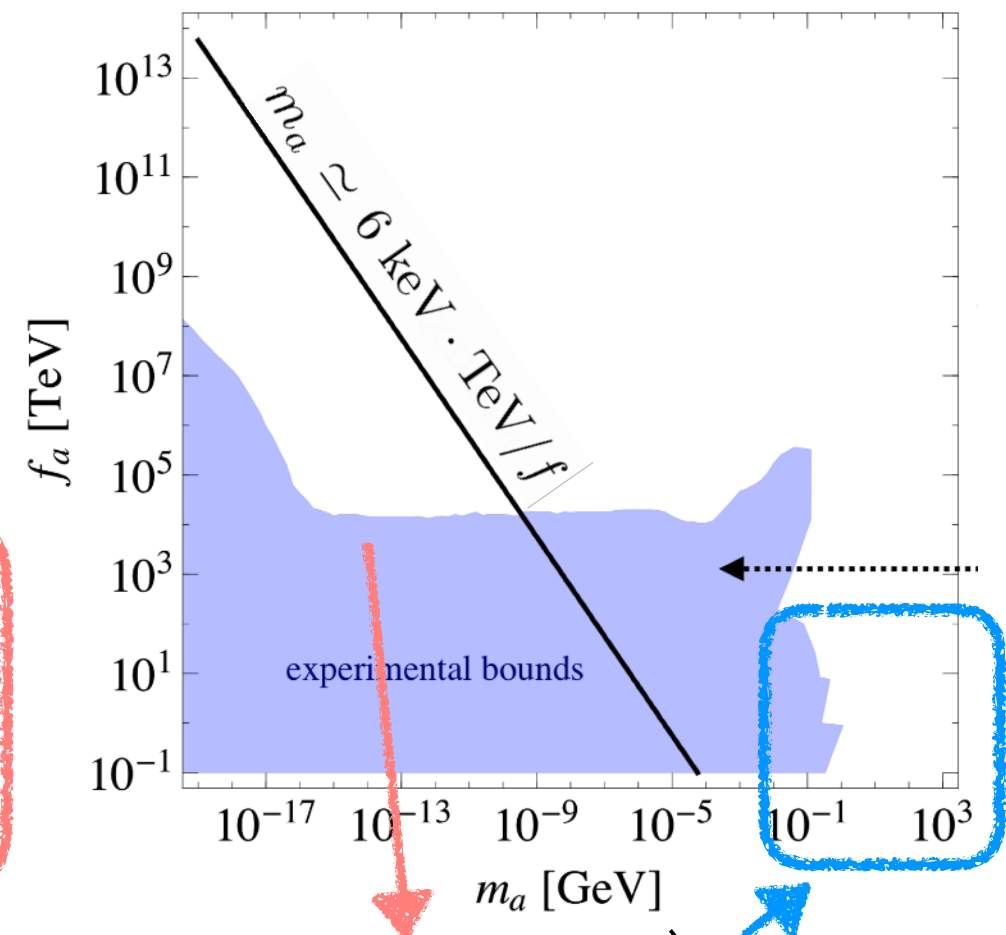
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strong CP solved here!

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Rubakov hep-ph/9703409
M.K Gaillard+ 1805.06465

ALPs and BSM: Dark Matter

Pseudoscalar-mediated Dark Matter

$$\mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\alpha_s c_3 G\tilde{G} + g_* \Phi \psi \tilde{\psi} \right]$$

$$\Phi = \frac{f_a}{\sqrt{2}} e^{ia/f_a}$$

Observed relic abundance for

$$m_\psi \simeq 4.6 \text{ TeV} \frac{c_3}{10} \left(\frac{g_*}{3} \right)^2 \Rightarrow f \simeq 1.9 \text{ TeV} \frac{3}{g_*}$$

$$m_\psi = g_* f / \sqrt{2}$$

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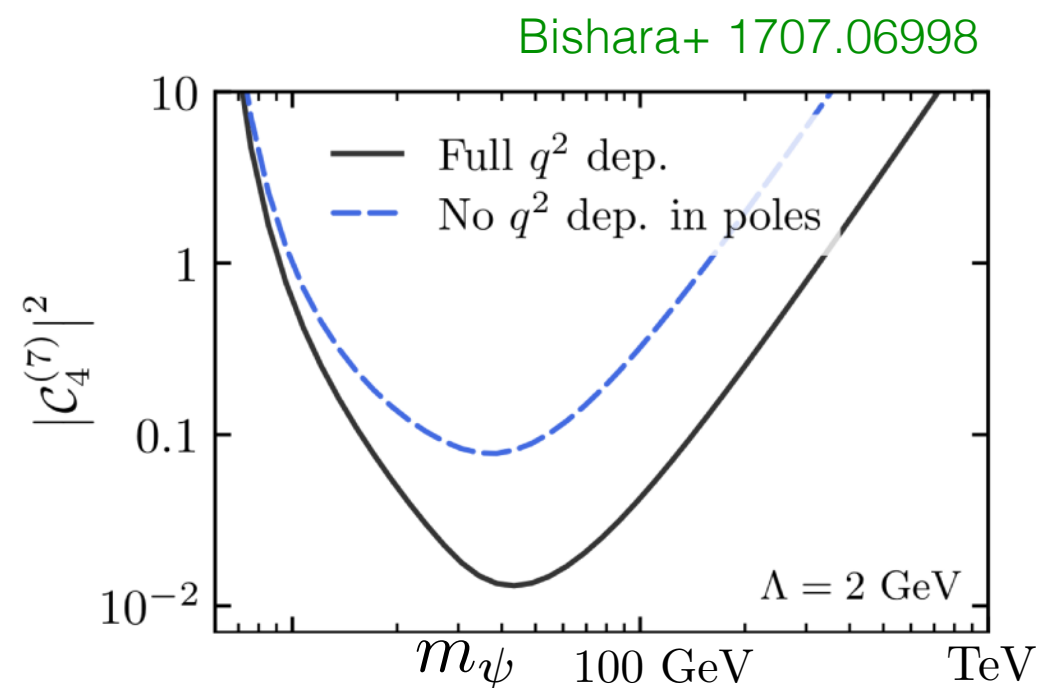
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Direct Detection completely irrelevant

Indirect Detection not yet competitive
(caveat: should compute Sommerfeld)

Colliders are needed to test this scenario!



$$Q_4^{(7)} = \frac{\alpha_s}{8\pi} (\bar{\chi} i \gamma_5 \chi) G^{a\mu\nu} \tilde{G}_{\mu\nu}^a$$

ALPs and BSM: SUSY R-axion I

$N = 1$ SUSY always accompanied by a continuous $U(1)_R = \text{"R-symmetry"}$

$$R : \theta_\alpha \rightarrow e^{i\epsilon} \theta_\alpha \quad [R, Q] = -Q$$

R-charge assignments:

$$\Phi = \phi + \sqrt{2}\theta \psi + \theta^2 F$$

$$r_\phi = r_\Phi$$

$$r_\psi = r_\Phi - 1$$

$$r_F = r_\Phi - 2$$

Vector superfields are real $\Rightarrow$ **gauginos** have $r_\lambda = 1$

Lagrangian $\mathcal{L}$ R-symmetric $\Rightarrow R(W) = 2$

($\Leftarrow$ if Kahler canonical)

$$\mathcal{L} \supset \int d^2\theta W + \text{c.c.}$$

W superpotential

ALPs and BSM: SUSY R-axion II

Nelson-Seiberg NPB416 (1994)

- i) SUSY broken in global minimum
- ii) superpotential W “generic”
(i.e. contains all terms not forbidden by symmetries)

$\Rightarrow$ Lagrangian respects a $U(1)_R$

$U(1)_R$ needs to be broken because of

gaugino masses $\mathcal{L} \supset m_\lambda \lambda\lambda \quad [r_\lambda = 1]$

EW symmetry breaking & Higgsino masses $\mathcal{L} \supset B_\mu H_u H_d + \text{c.c.} \quad \& \quad W \supset \mu H_u H_d$

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Massless Goldstone in the spectrum R-axion a

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$\rightarrow$ Tune CC to zero: explicit breaking of $U(1)_R$

$$m_a^2 \sim (10 \text{ MeV})^2 \times \frac{M_{\text{SUSY}}}{10 \text{ TeV}} \times \frac{m_{3/2}}{\text{eV}} \quad \text{Bagger+} \\ \text{hep-ph/9405345}$$

$\rightarrow$ Metastable vacuum Intriligator Seiberg Shih 2007

....

$$m_a \ll M_{\text{SUSY}}$$

light SUSY particle by symmetry

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light SUSY particle **by symmetry**

Could be first sign of SUSY at colliders!

Bellazzini Mariotti Redigolo FS Serra 1702.02152

R-axion- as Dark Matter mediator

Spectrum à la gauge mediation

Gravitino DM

$$m_{3/2} = F/(\sqrt{3}M_{\text{Pl}}) \simeq 11 \text{ meV} \cdot (g_*/3) \cdot (f/4 \text{ TeV})^2$$


Cannot make observed DM for these values of parameter

(motivated by naturalness of Fermi scale)

Respects all bounds from cosmology and collider

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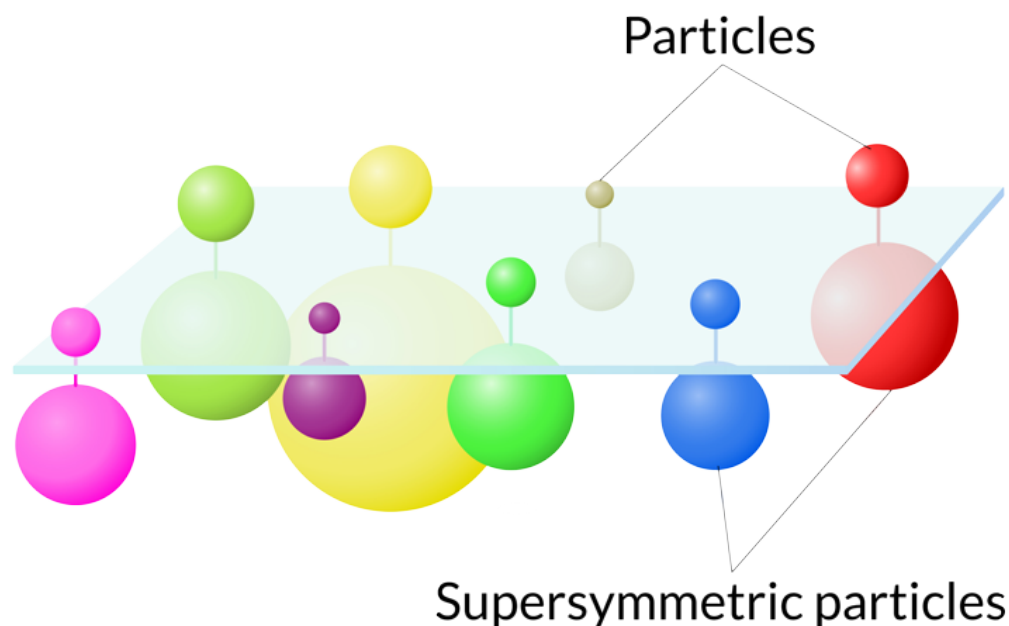
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DM from messenger or SUSY breaking sectors

Dimopoulos Giudice Pomarol hep-ph/9607225

R-axion is natural portal! Mardon Nomura Thaler 0905.3749
Fan Thaler Wang 1004.0008

Motivated realisation of ALP-portal Dark Matter

Theory summary

Pseudo Goldstone bosons (ALPs) with

$$m_a \sim 1 - 100 \text{ GeV}$$

$$f_a \sim 0.1 - 100 \text{ TeV}$$

arise in several motivated models (**QCD axion**, **DM**, **SUSY**, **CHM**,...)

Pheno observation

Coupling to gluons $\sim a G \tilde{G}$ often neglected in ALPs pheno

but mandatory for **QCD axion**

natural in **CHM** (e.g. loops of tops, ...)

as well as in **SUSY**! (e.g. loops of tops, gluinos, ...)

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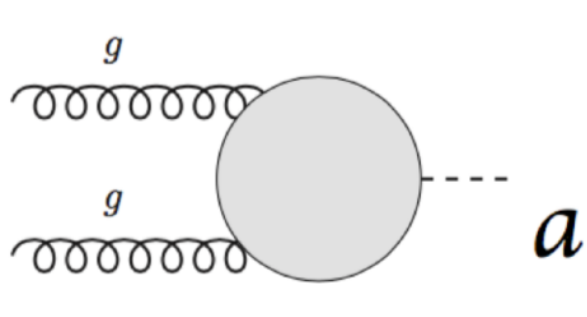
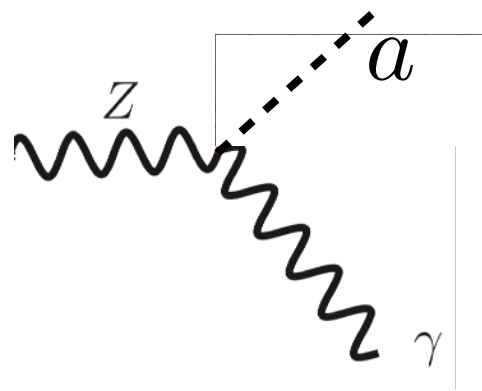
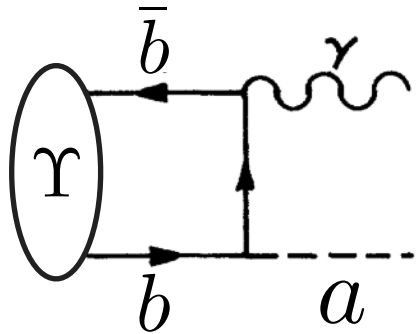
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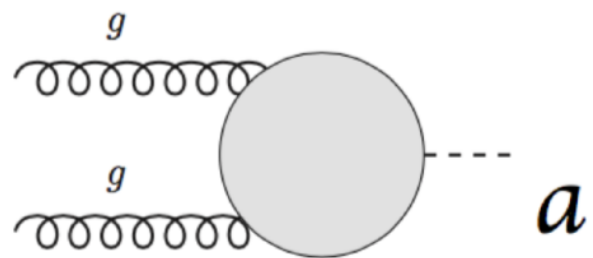
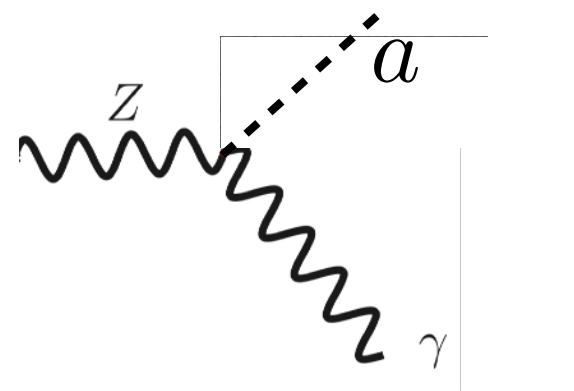
Present ALP mass coverage

$$\mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\alpha_s c_3 G\tilde{G} + \alpha_2 c_2 W\tilde{W} + \alpha_1 c_1 B\tilde{B} \right] + iC_f m_f \frac{a}{f_a} \bar{f} \gamma_5 f + C_h v \left(\frac{\partial_\mu a}{f_a} \right)^2 h + \dots$$



.....

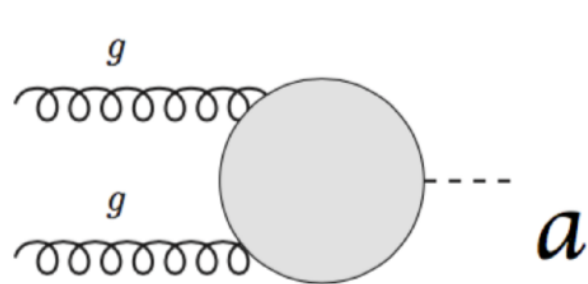
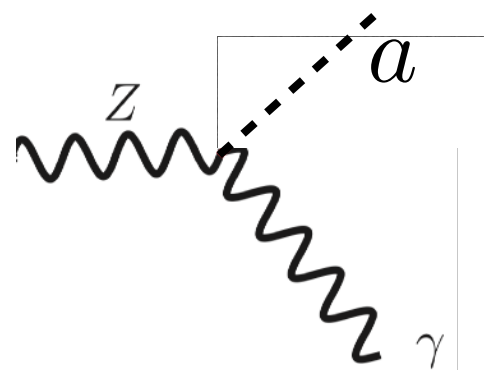
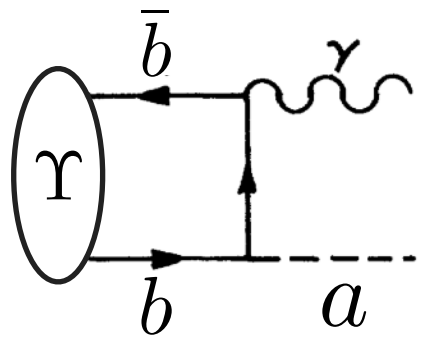
A Feynman diagram showing the production of a photon pair ($\gamma\gamma$) from a quark-antiquark pair ($q\bar{q}$). On the left, a vertical oval labeled Υ represents the initial state. Two horizontal lines emerge from it: a top line with an arrow pointing left labeled $\bar{b}$, and a bottom line with an arrow pointing right labeled b . These lines extend to the right and then turn upwards to meet at a vertex. From this vertex, a wavy line labeled γ extends to the right. The bottom line continues as a dashed line to the right, labeled a .



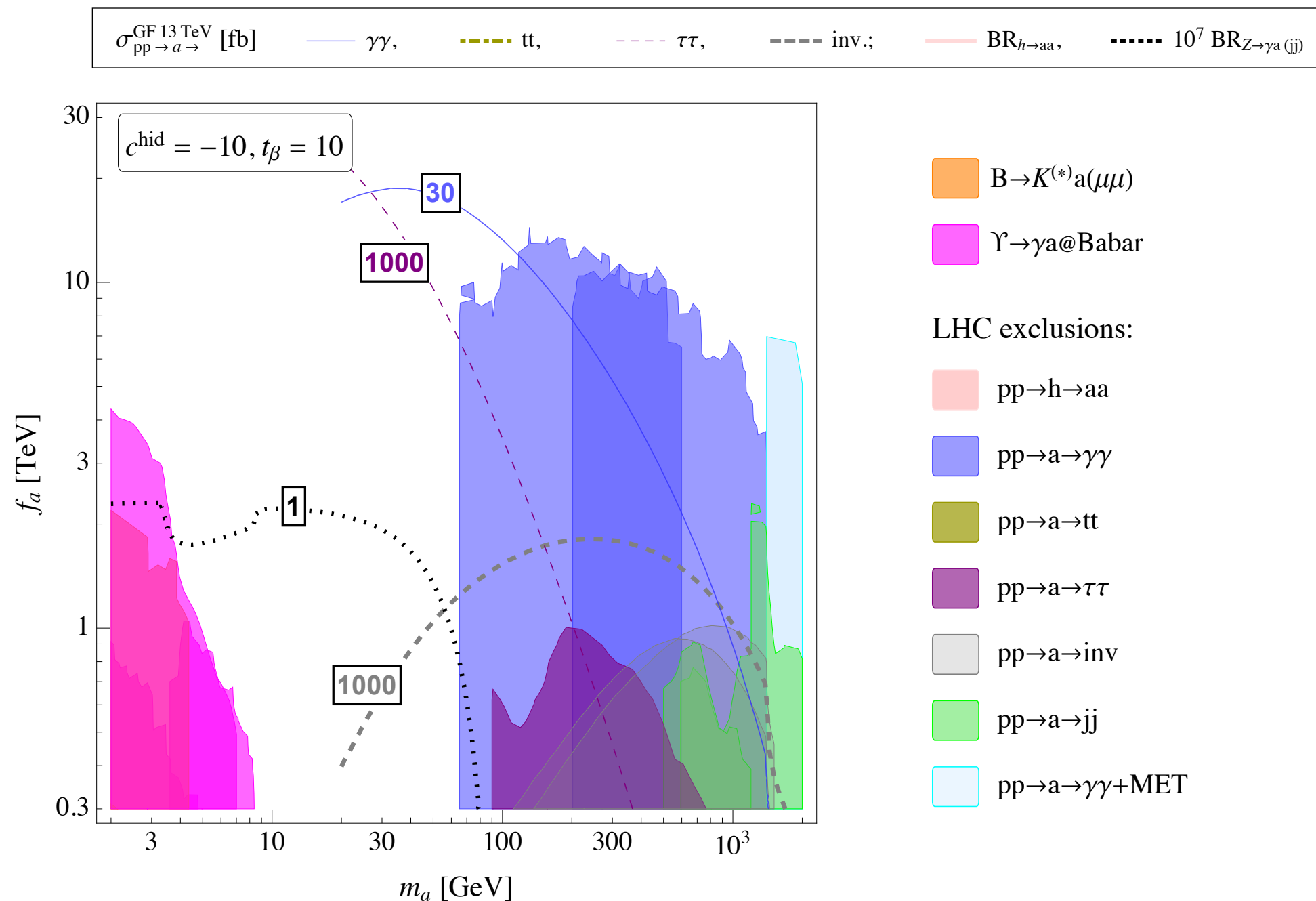
pp→a→γγ+MET

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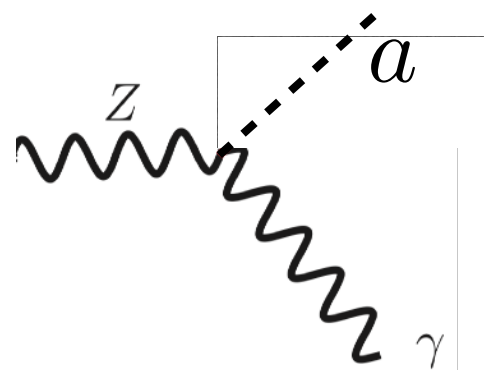
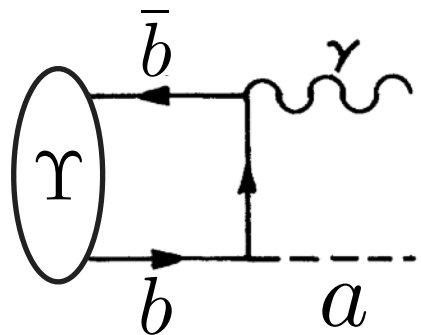


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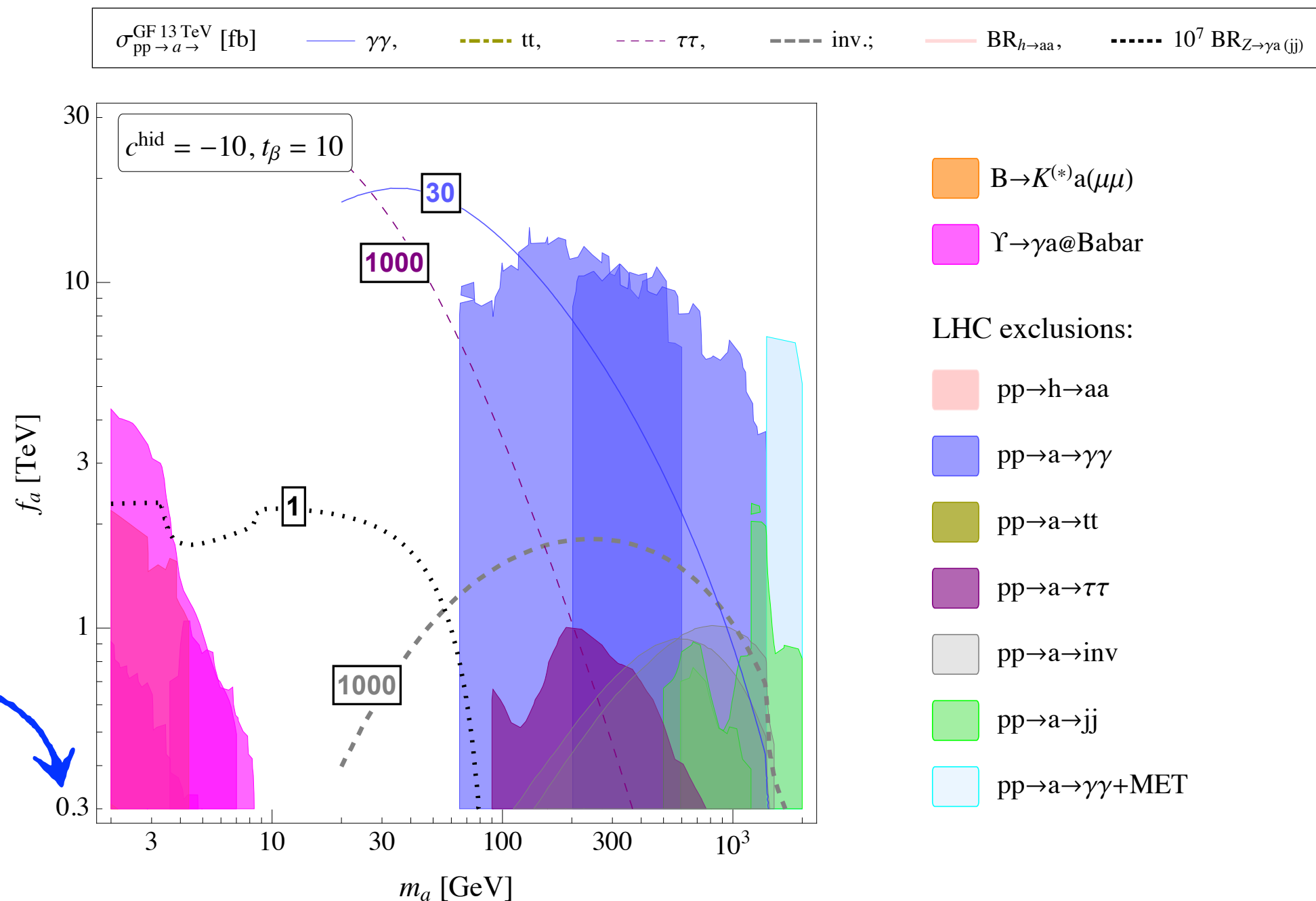


Present ALP mass coverage

$$\mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\alpha_s c_3 G\tilde{G} + \alpha_2 c_2 W\tilde{W} + \alpha_1 c_1 B\tilde{B} \right] + iC_f m_f \frac{a}{f_a} \bar{f} \gamma_5 f + C_h v \left(\frac{\partial_\mu a}{f_a} \right)^2 h + \dots$$

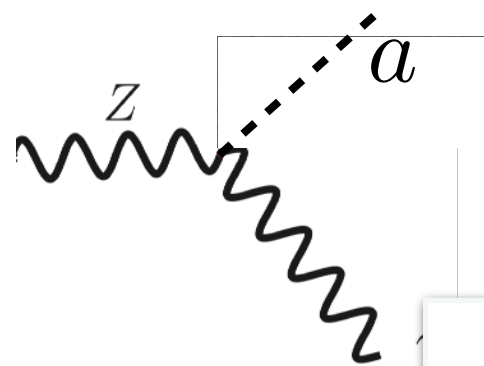
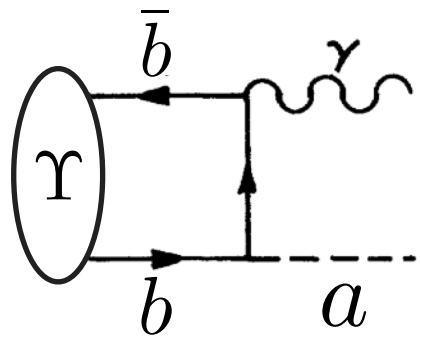


LEP only constrains smaller values of f_a

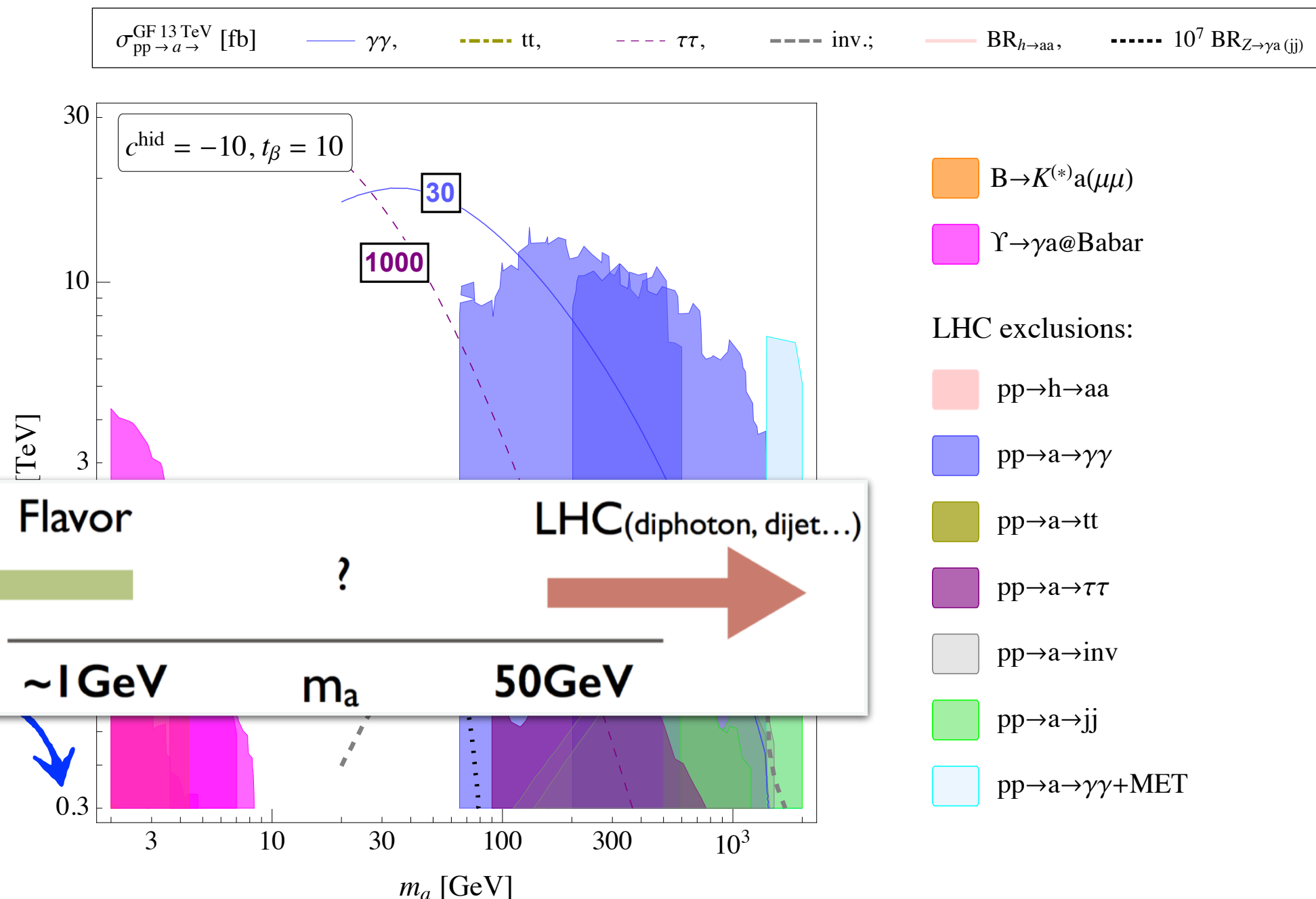


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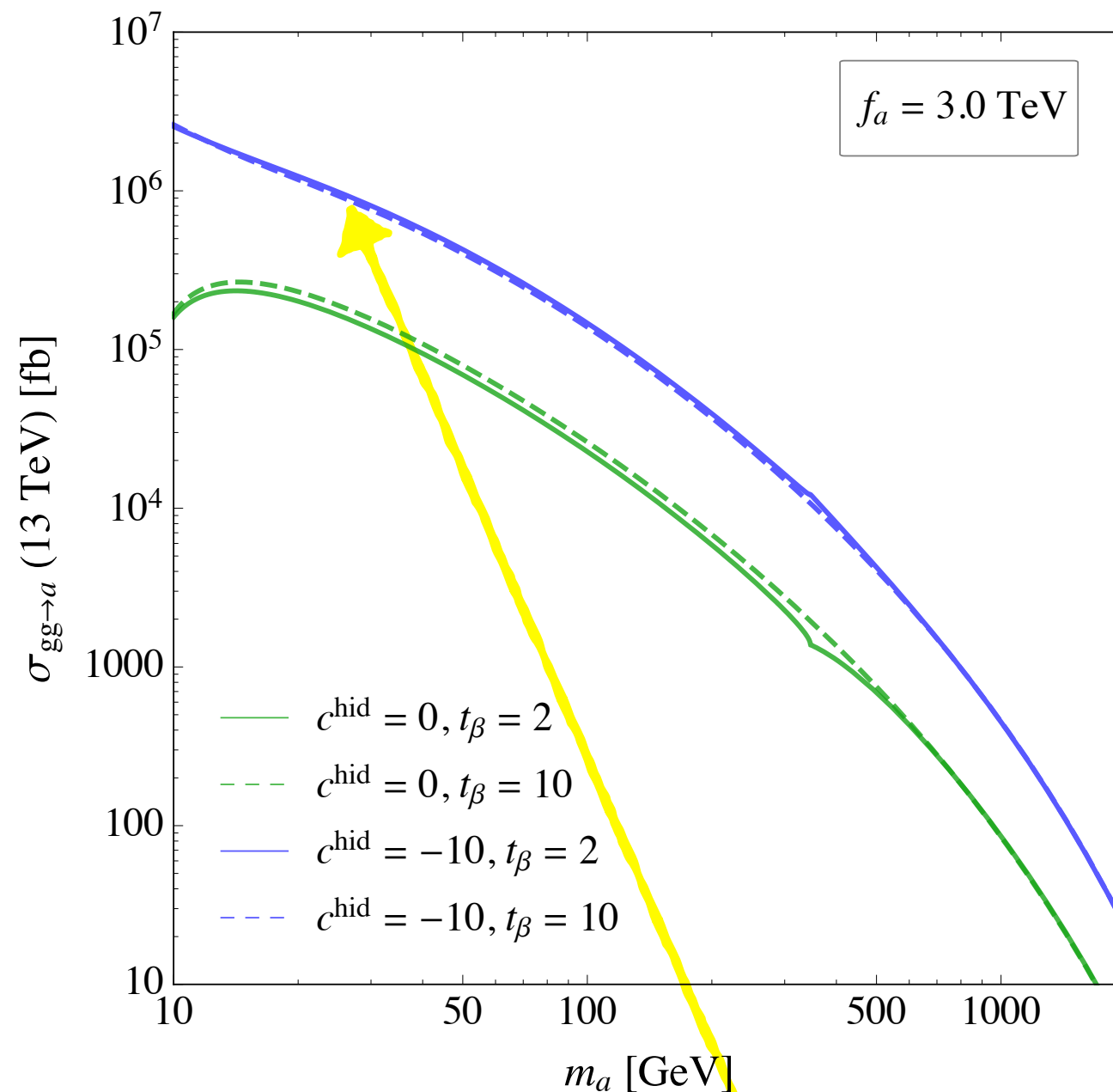
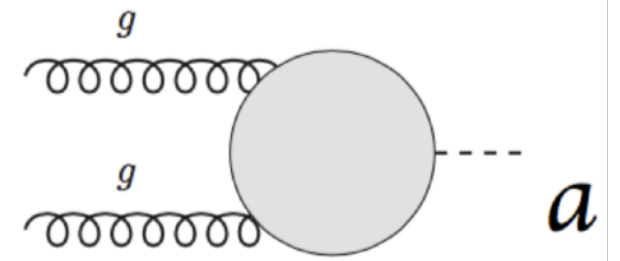
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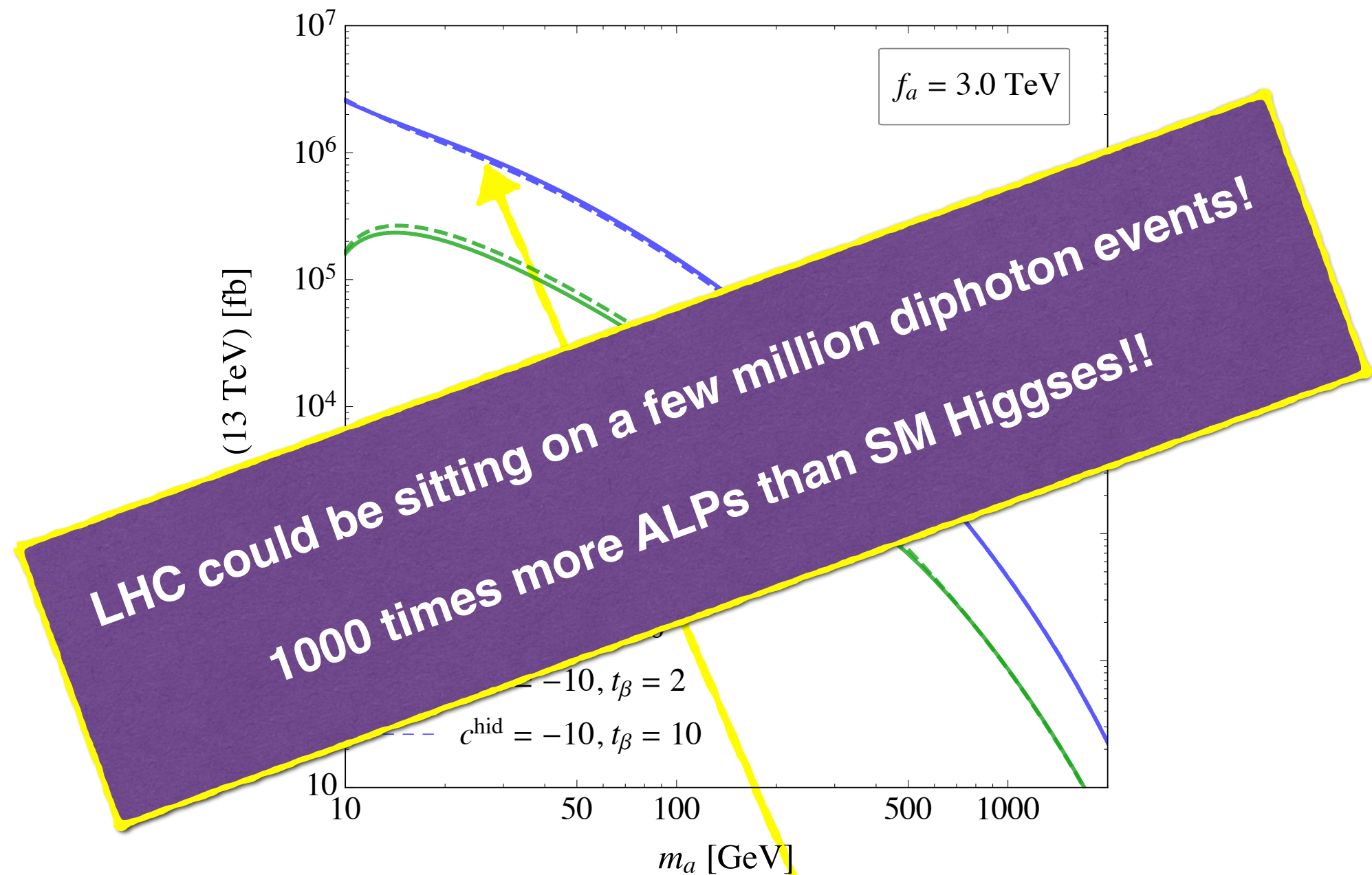
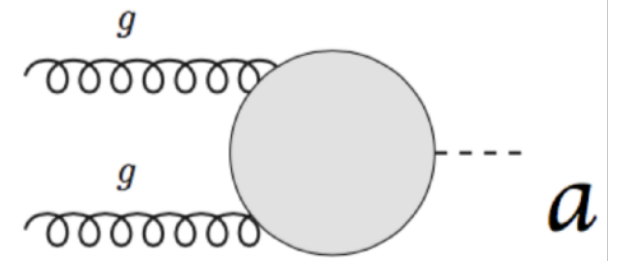


ALP production at the LHC



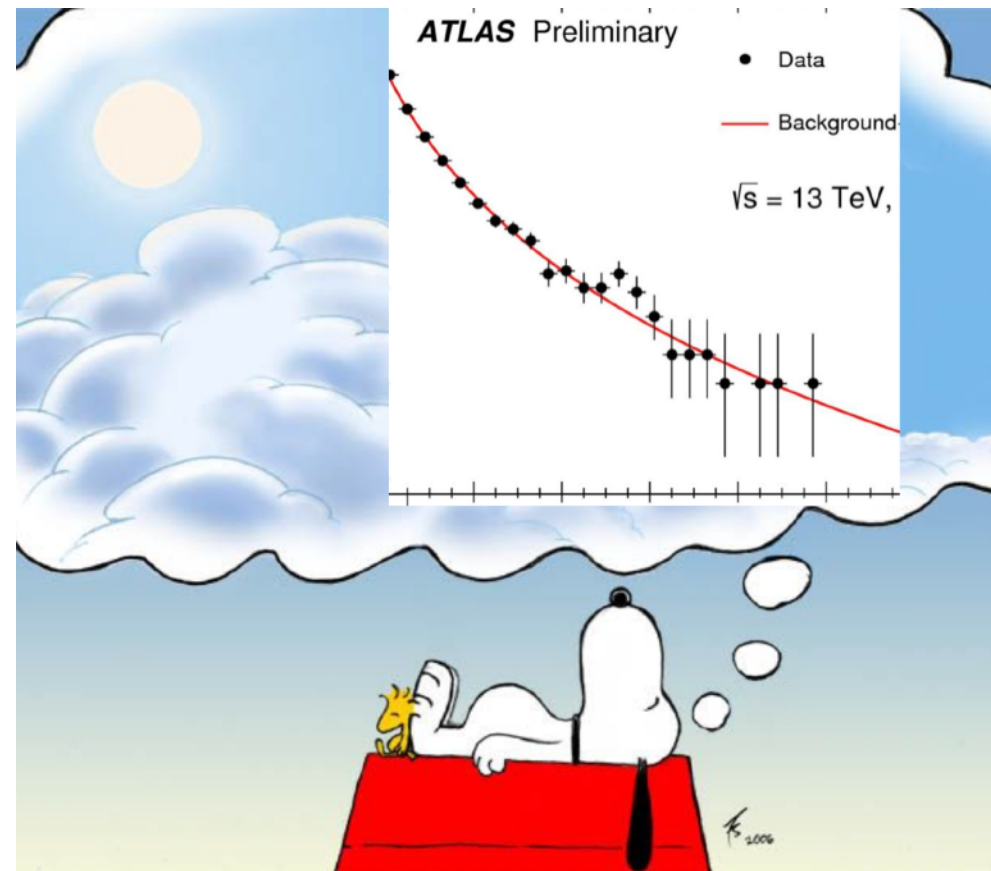
Production cross sections of $\sim 10^5 \text{ pb}$ are still allowed! [$f_a \approx 300 \text{ GeV}$]

ALP production at the LHC



Production cross sections of $\sim 10^5 \text{ pb}$ are still allowed! [$f_a \approx 300 \text{ GeV}$]

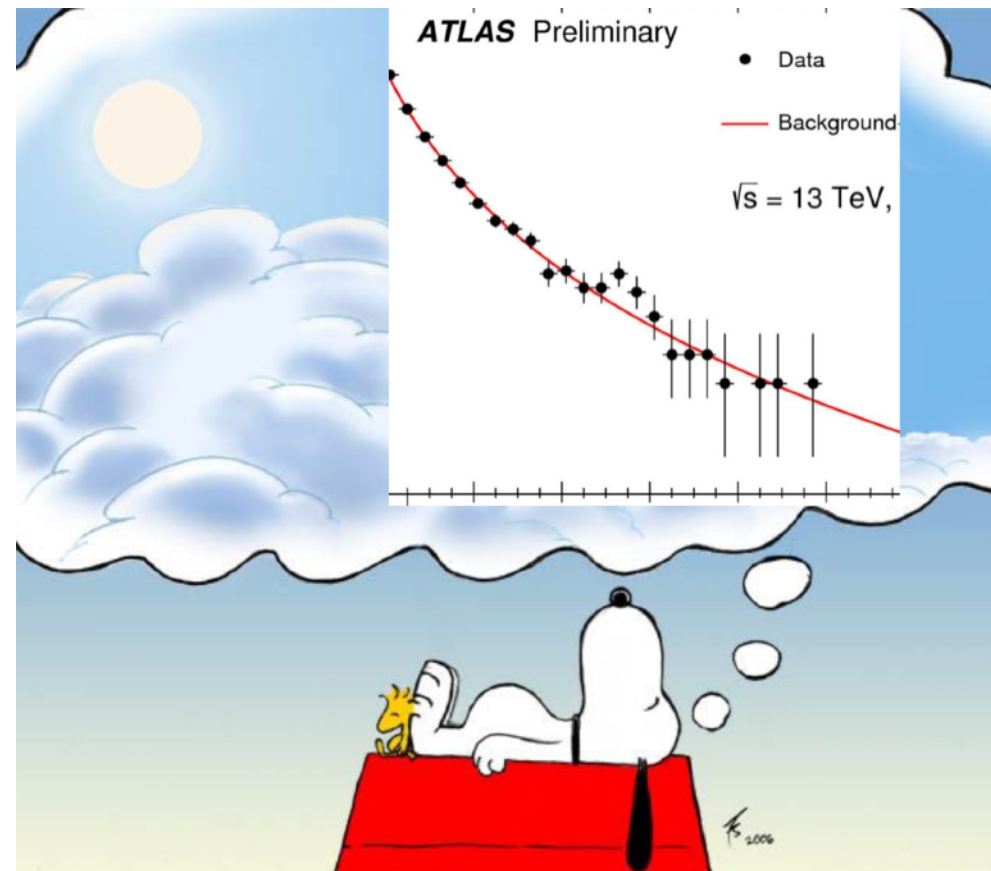
May our dreams come true?



1. Theory bias towards high masses

Why not $M_X < O(100) \text{ GeV}$? 2. “Low-mass already constrained by previous colliders (LEP,...)”

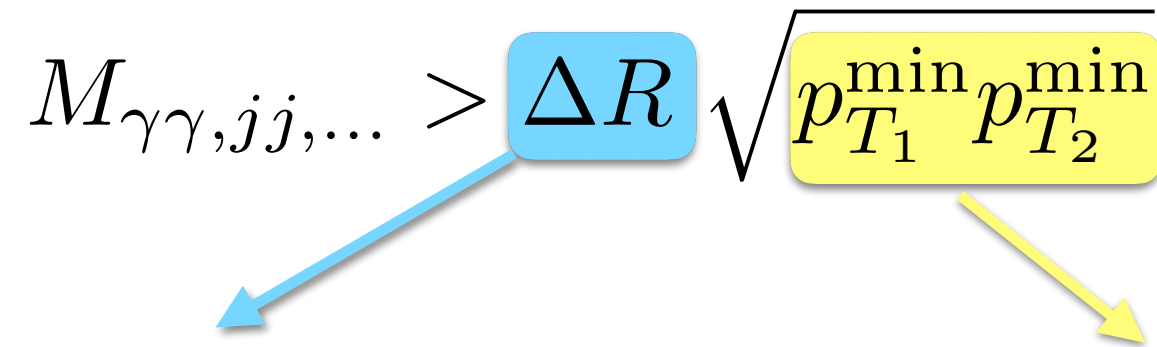
May our dreams come true?



Why not $M_X < O(100) \text{ GeV}$?

3. “It is very difficult!” Minimal pT cuts, ...

Why difficult to go below ~ 100 GeV?

$$M_{\gamma\gamma,jj,\dots} > \Delta R \sqrt{p_{T_1}^{\min} p_{T_2}^{\min}}$$


Isolation of photon/jet/... $\Delta R \equiv \sqrt{\Delta\eta^2 + \Delta\phi^2}$

Minimal cuts on transverse momenta

Two ways to lower $M_{\gamma\gamma}$

 Lower ΔR

 Lower $p_T^{\min}$

Lower $p_T^{\min}$?

D0 ($\sigma_{\gamma\gamma}$)	$p\bar{p} \rightarrow a \rightarrow \gamma\gamma$	4.2 fb^{-1}	1.96 TeV	$p_{T_1, T_2} > 21, 20 \text{ GeV}$	[1]	$\Delta R \gtrsim 0.4$	
CDF ($\sigma_{\gamma\gamma}$)	$p\bar{p} \rightarrow a \rightarrow \gamma\gamma$	5.36 fb^{-1}	1.96 TeV	$p_{T_1, T_2} > 17, 15 \text{ GeV}$	[2]		
ATLAS	$pp \rightarrow a \rightarrow \gamma\gamma$	4.9 fb^{-1}	7 TeV	$p_{T_1, T_2} > 25, 22 \text{ GeV}$	[8]		9.4 GeV
ATLAS	$pp \rightarrow a \rightarrow \gamma\gamma$	20.2 fb^{-1}	8 TeV	$p_{T_1, T_2} > 40, 30 \text{ GeV}$	[9]		13.9 GeV
CMS	$pp \rightarrow a \rightarrow \gamma\gamma$	5.0 fb^{-1}	7 TeV	$p_{T_1, T_2} > 40, 25 \text{ GeV}$	[10]		14.2 GeV

LHC pT cuts in diphoton cross section measurements
 but LHC diphoton searches do not reach such low masses

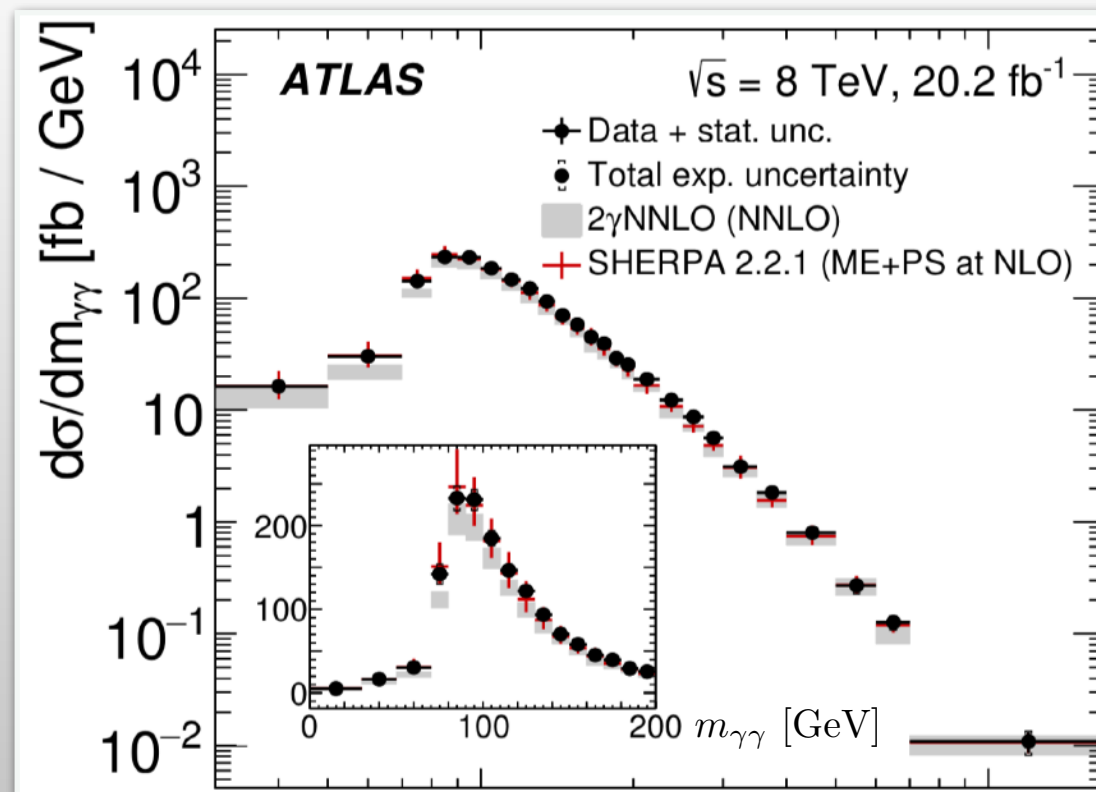
$m_{\gamma\gamma}^{\text{MIN}}$

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Why? Background Shape

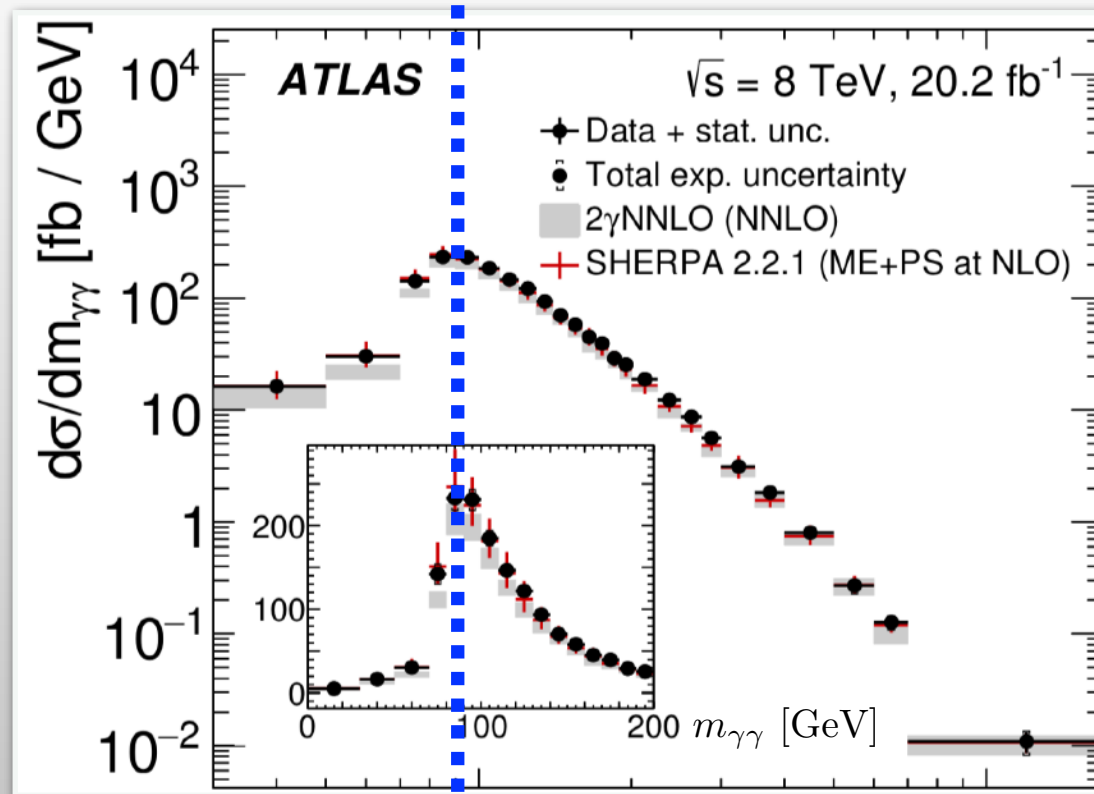
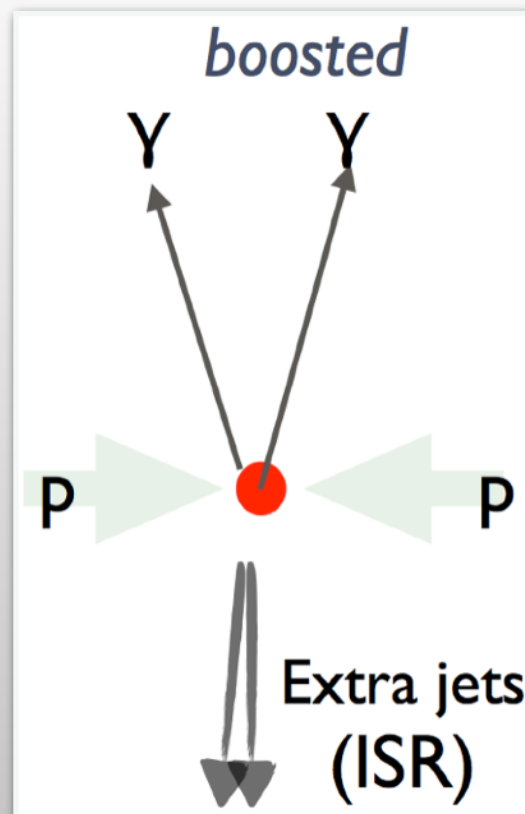


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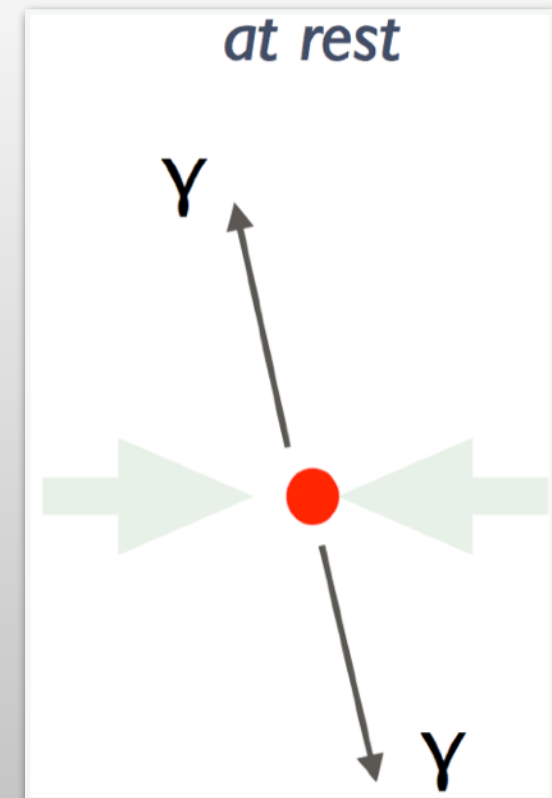
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						9.4 GeV 13.9 GeV 14.2 GeV

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Why? Background Shape



$$m_{\gamma\gamma} \simeq 2 p_T^{\min}$$

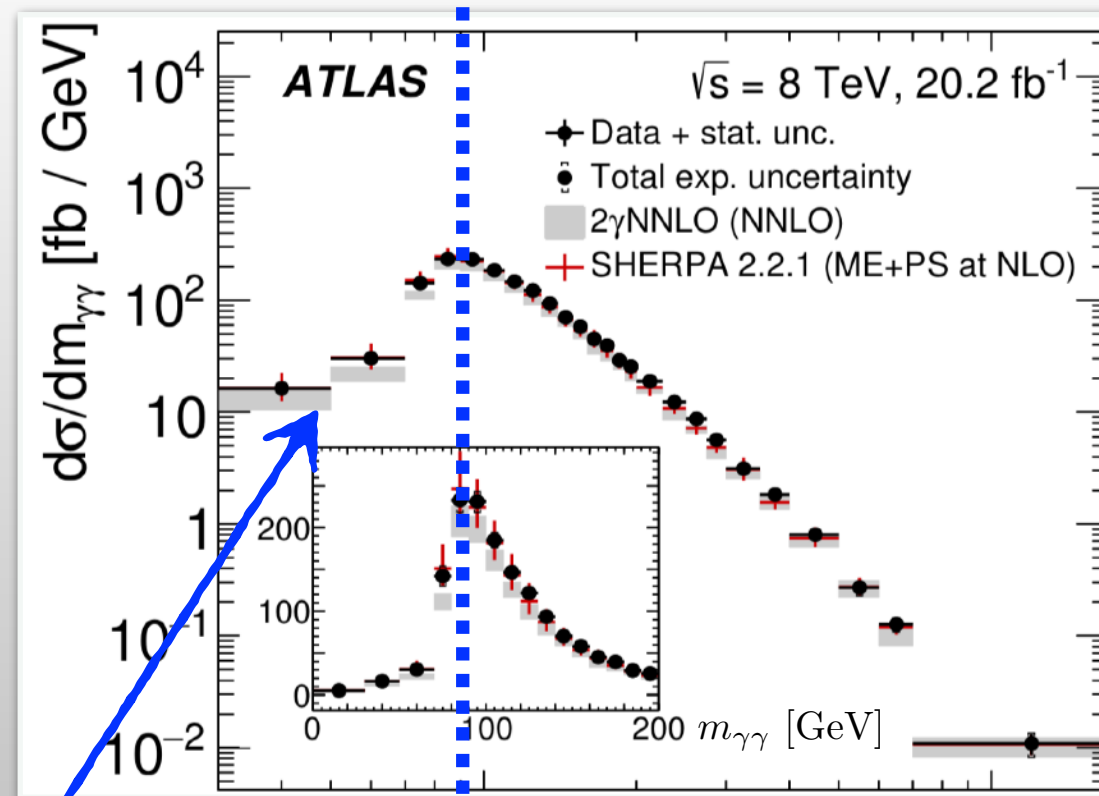


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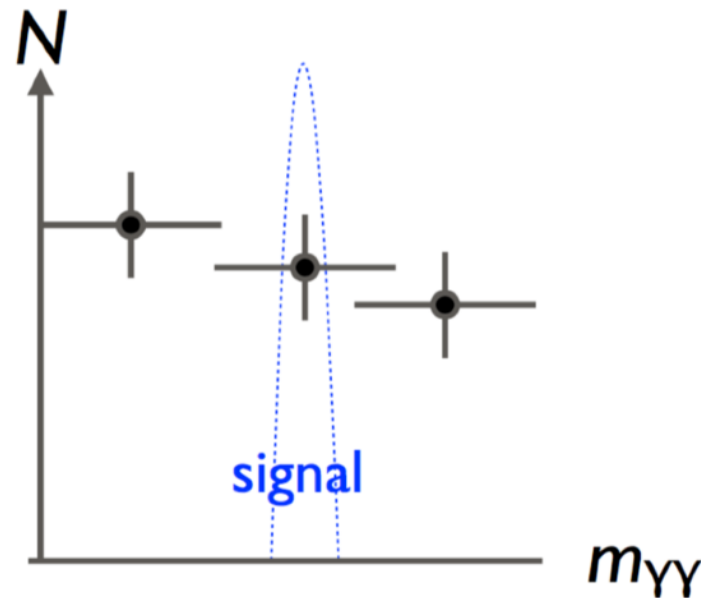


$$m_{\gamma\gamma} \simeq 2 p_{T1}^{\min}$$

Below p_T cuts: Background has a structure, so **data-driven estimates are difficult**

New $\gamma\gamma$ Bound & Sensitivities

Starting point: inclusive **diphoton cross section measurements** @ ATLAS7,8 and CMS7



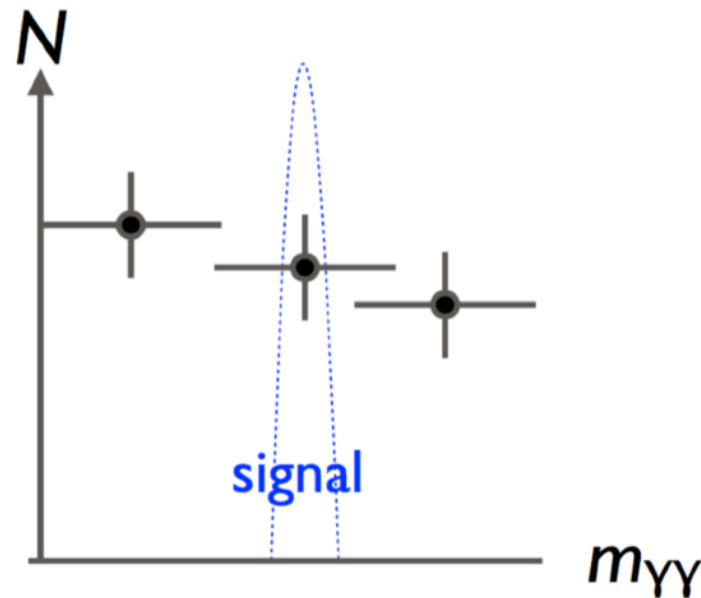
New Bound we assume zero knowledge of bkg

$$N_{\text{bin}}^{\text{signal}} < N_{\text{bin}}^{\text{meas.}} (1 + 2 \Delta_{\text{bin}})$$

experimental rel. uncertainty

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$$N_{\text{bin}}^{\text{signal}} < N_{\text{bin}}^{\text{meas.}} (1 + 2 \Delta_{\text{bin}})$$

experimental rel. uncertainty

Current LHC reach?

If bump-hunt on steep slope or if Monte Carlos improve:

$$N_{\text{bin}}^{\text{signal}} < N_{\text{bin}}^{\text{meas.}} \times 2 \Delta_{\text{bin}}$$

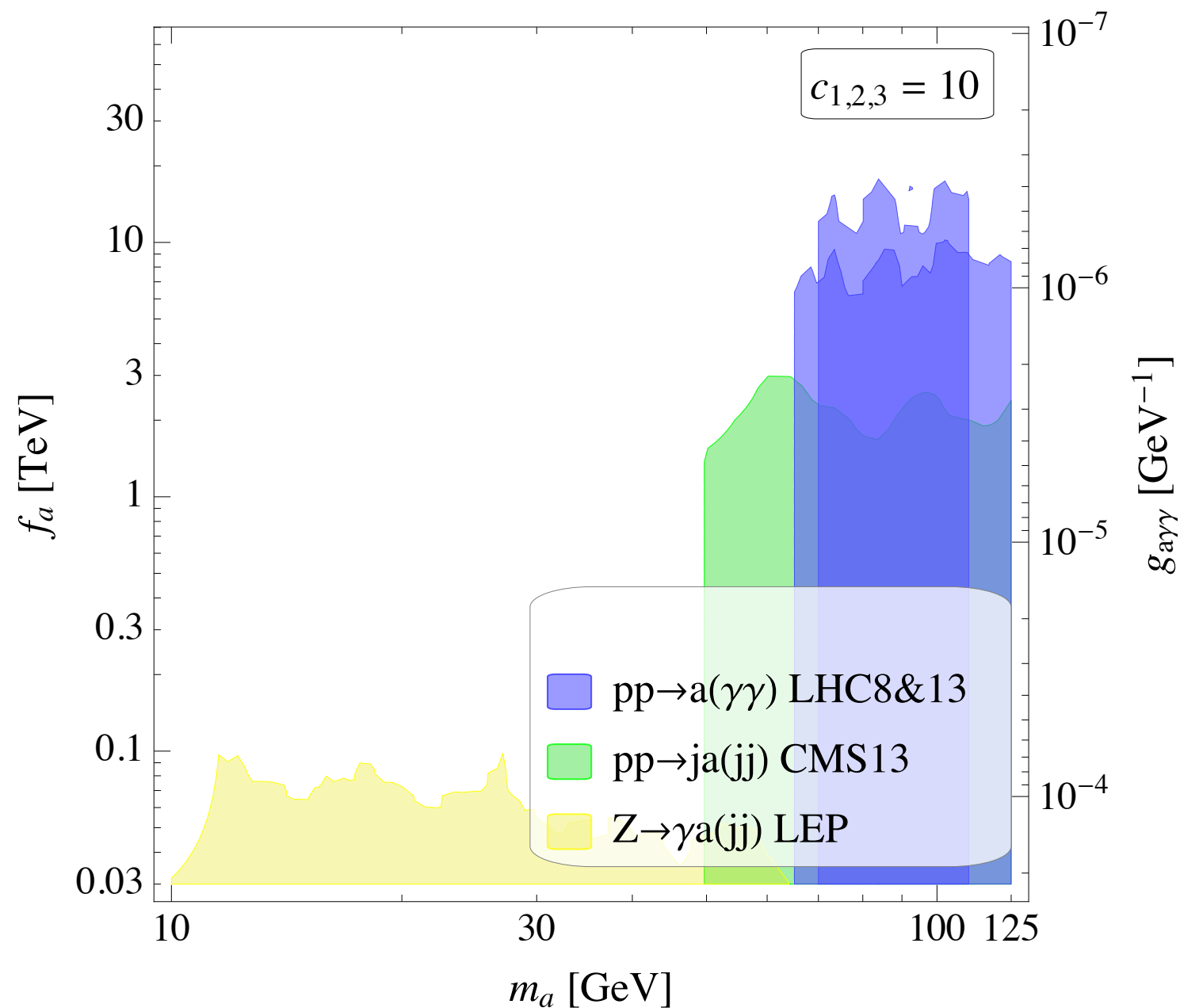
(we take data = SM prediction, and optimise binning)

Future LHC Reach we rescale by simulation of bkg with same cuts at different energies
[Madgraph+Pythia+Delphes]

Impact on ALP parameter space

$$\mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\alpha_s c_3 G\tilde{G} + \alpha_2 c_2 W\tilde{W} + \alpha_1 c_1 B\tilde{B} \right]$$

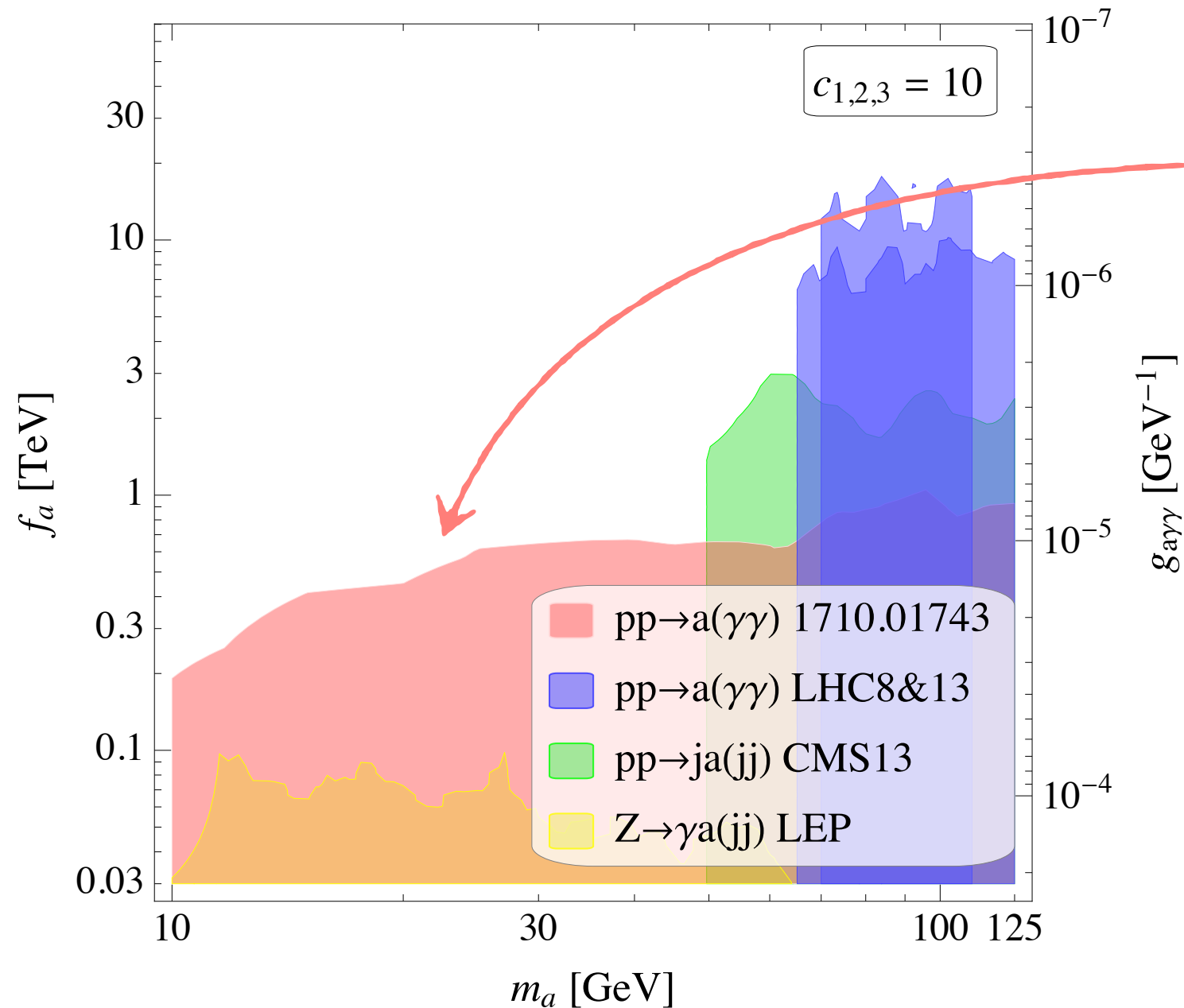
$$\alpha_1 = \frac{5}{3}\alpha_y$$



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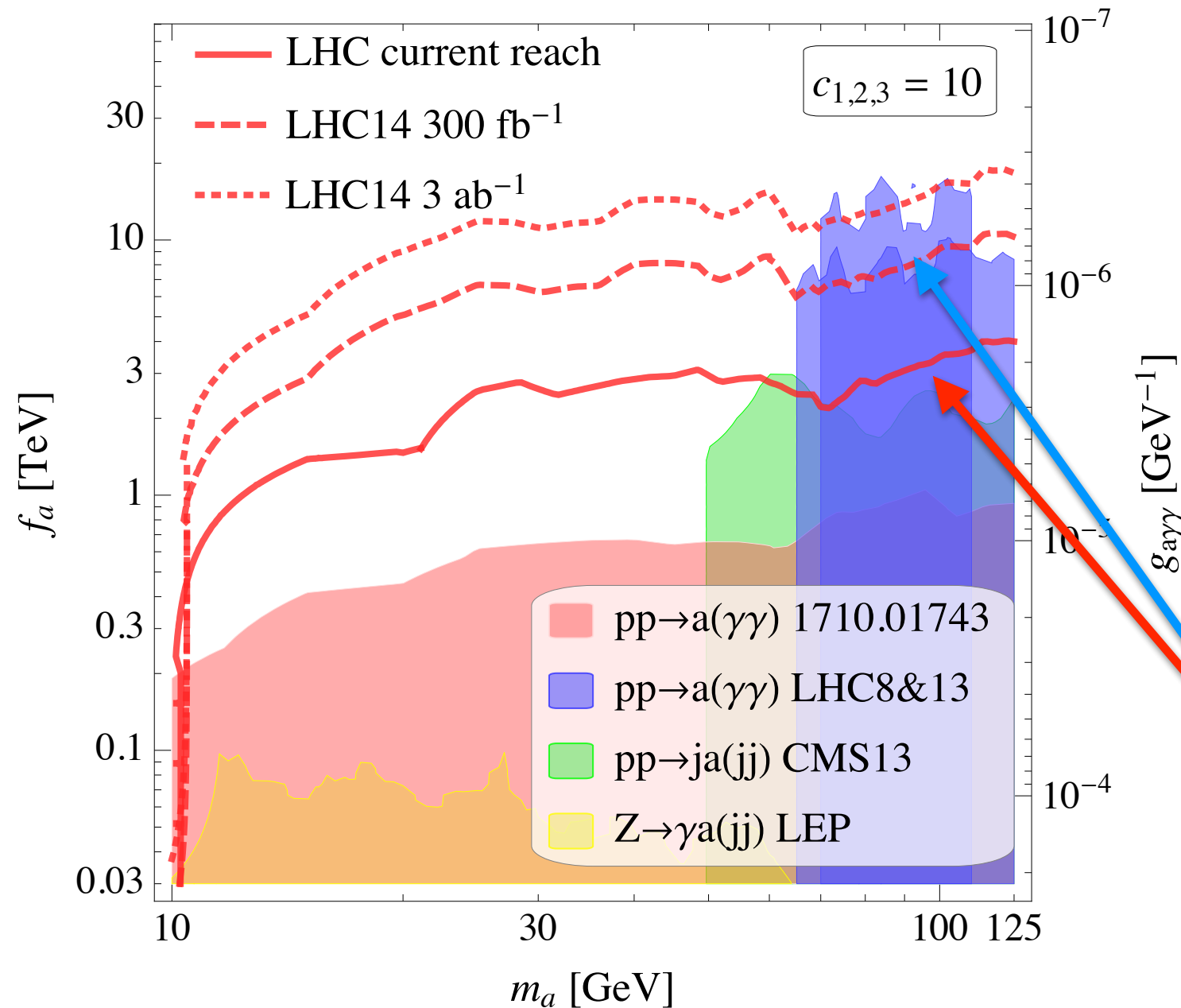


Our simple **bound**
is (by far) the strongest one!

Impact on ALP parameter space

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Memo:

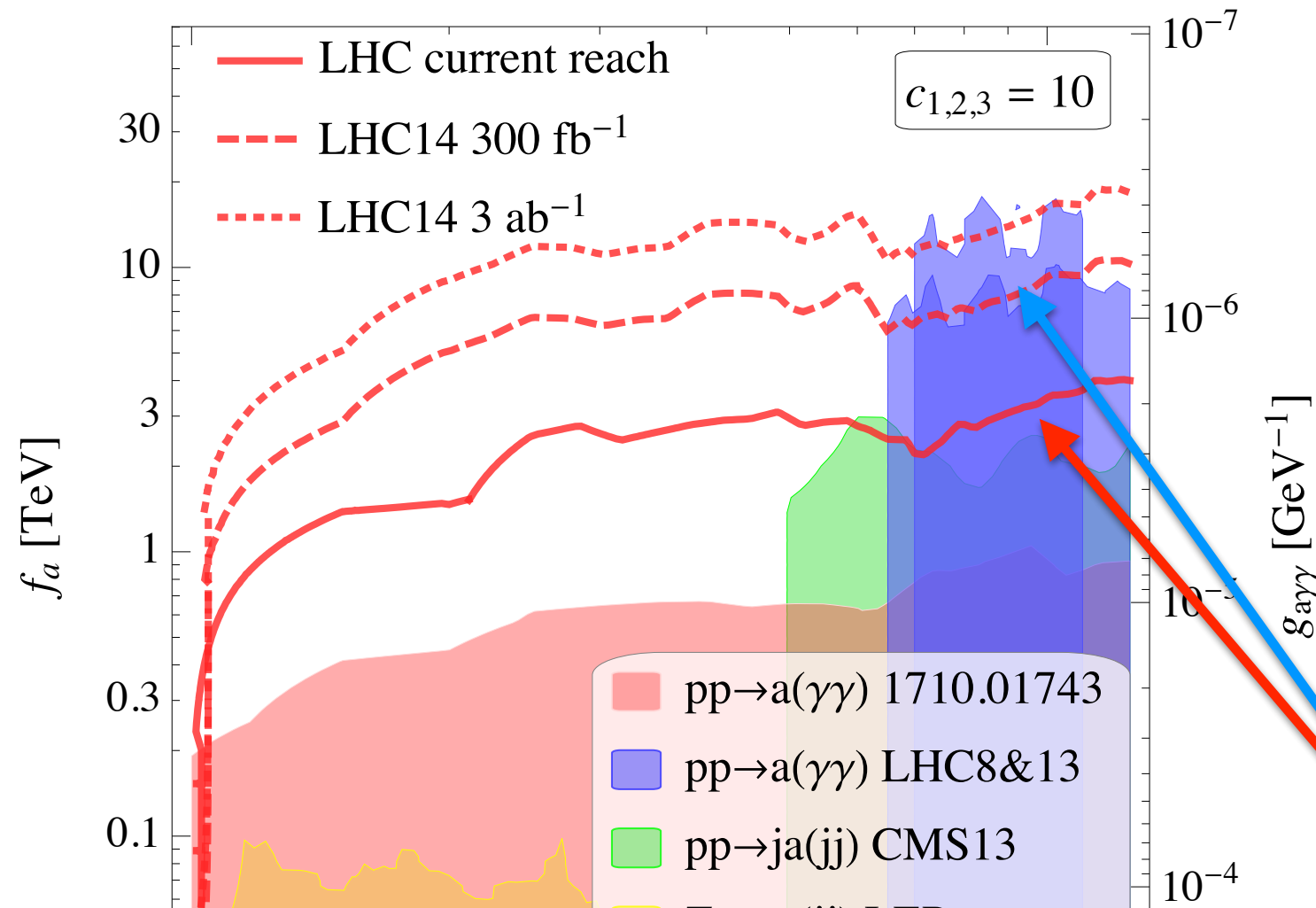
reach assumes bump hunt on steep slope or improvement in SM prediction

But at least we have been conservative for masses that LHC already explored

Impact on ALP parameter space

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Memo:

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But at least we have been conservative
already explored

Could be useful to **non-ALP scenarios!**

Light Z' (e.g. as DM mediators)

Extra Higgses [Delgado+ 1603.00962](#)

....

So, what should ATLAS & CMS do?

Study so far shows clearly an unexplored potential, but

Challenge: need improvement in Monte Carlo to predict the backgrounds or being able to hunt bumps on steep data

$$M_{\gamma\gamma,jj,\dots} > \Delta R \sqrt{p_{T_1}^{\min} p_{T_2}^{\min}}$$

■ Lower ΔR

and look for resonances recoiling against a jet (or any ISR)

Background is smooth and one does not need Monte Carlos!

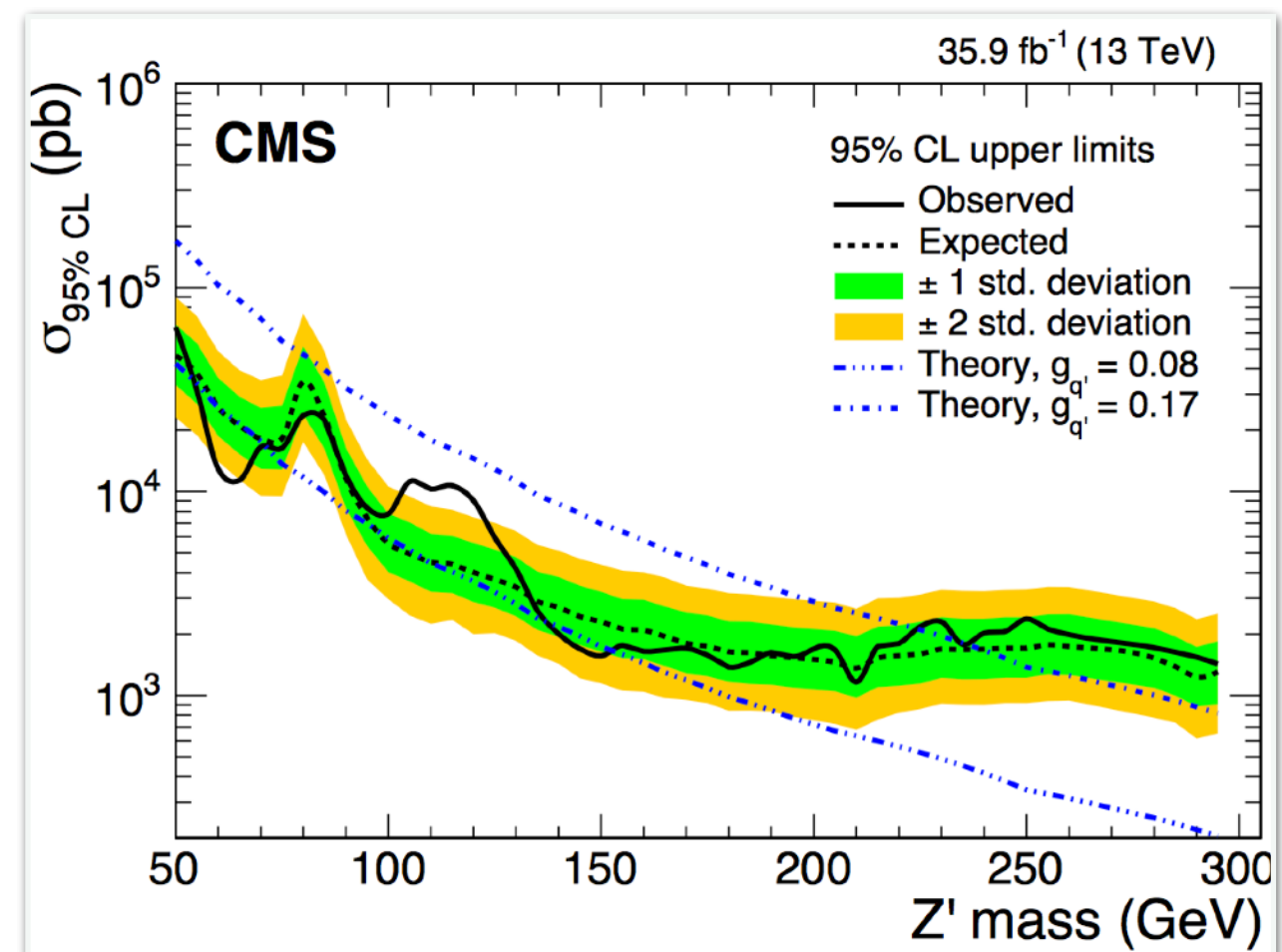
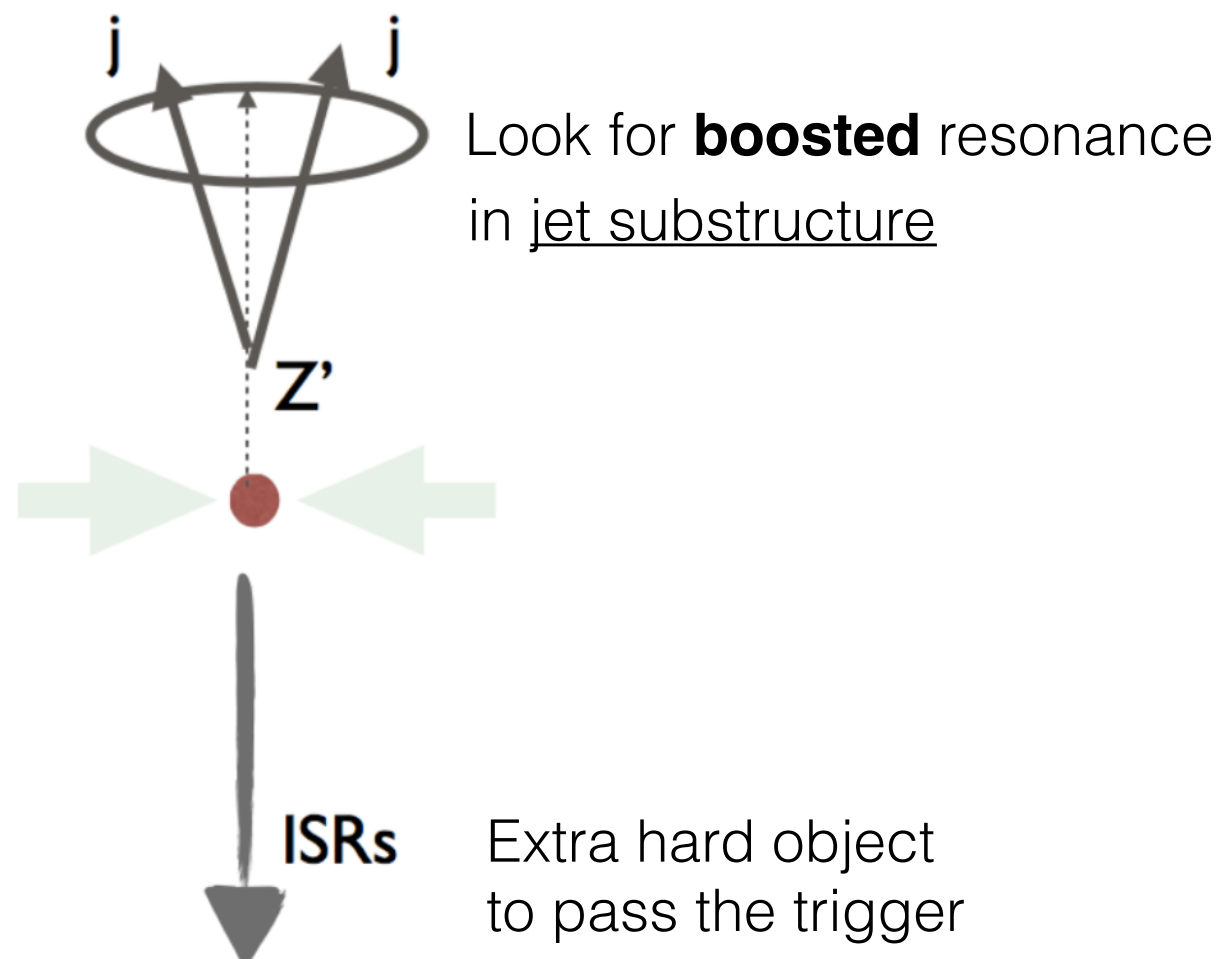
Lower ΔR : The Present

				m_{jj}^{MIN}	
CMS	$pp \rightarrow a \rightarrow jj$	18.8 fb^{-1}	8 TeV	500 GeV	[38]
ATLAS	$pp \rightarrow a \rightarrow jj$	20.3 fb^{-1}	8 TeV	350 GeV	[39]
CMS	$pp \rightarrow a \rightarrow jj$	12.9 fb^{-1}	13 TeV	600 GeV	[40]
ATLAS	$pp \rightarrow a \rightarrow jj$	3.4 fb^{-1}	13 TeV	450 GeV	[41]
CMS	$pp \rightarrow ja \rightarrow jjj$	35.9 fb^{-1}	13 TeV	50 GeV	[42]

Done recently in dijet, tremendous improvement in mass reach!

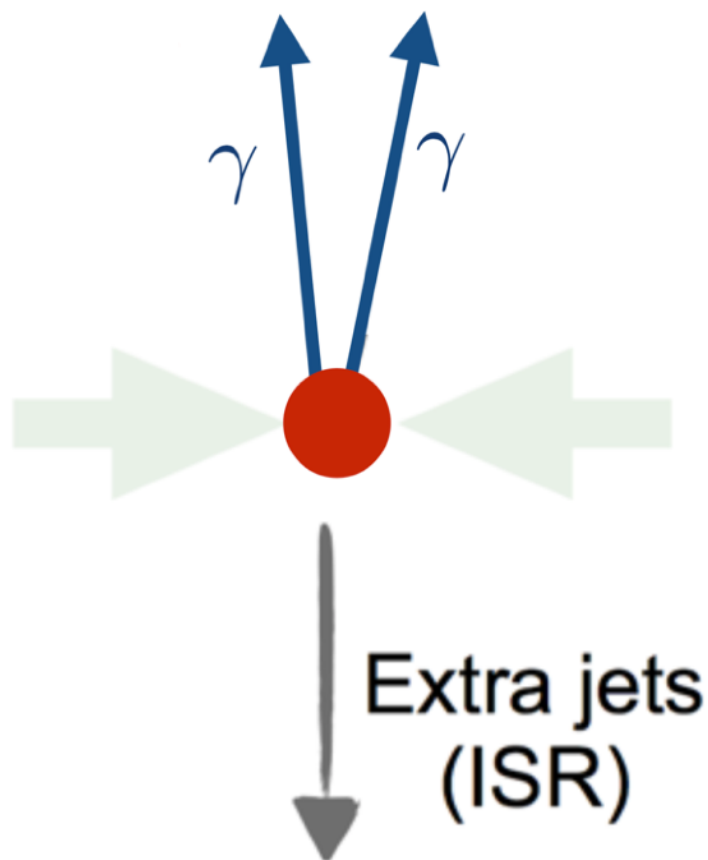
CMS 1710.00159

ATLAS 1801.08769



Lower ΔR : The Future

Low Mariotti Redigolo FS Tobioka in progress



They managed with jets, why not with photons?

Diphotons boosted vs a hard jet $p_T^a > 500 \text{ GeV}$

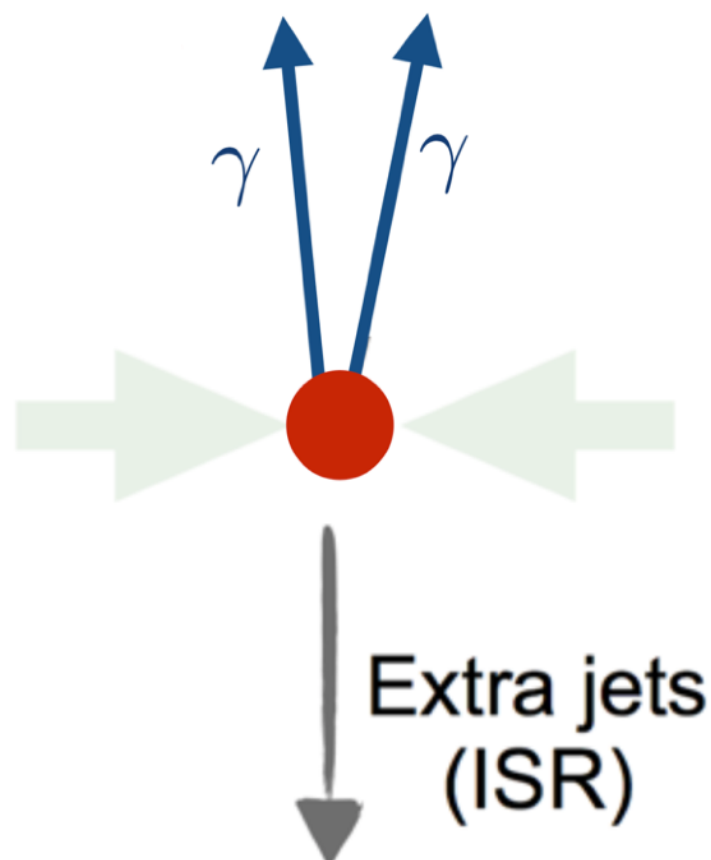
Photons get collimated $\Delta R \simeq \frac{2m_a}{p_T^a}$

Standard isolation rejects signal

$$\sum_{i \neq \gamma_{\text{test}}}^{\Delta R < R_{\text{iso}}} p_{T,i} < \# \quad \sum_{i \neq \gamma_{\text{test}}}^{\Delta R < R_{\text{iso}}} p_{T,i} / E_{T,\gamma_{\text{test}}} < \#$$

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Modify isolation! (NB: simpler than photon-jet substructures...)

It works!

$$\sum_{i \neq \gamma_{\text{test}}, \gamma_1}^{\Delta R < R_{\text{iso}}} p_{T,i}$$

$$\sum_{i \neq \gamma_{\text{test}}, \gamma_1}^{\Delta R < R_{\text{iso}}} p_{T,i} / E_{T,\gamma_{\text{test}}}$$



final checks and optimisations in progress

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Preliminary

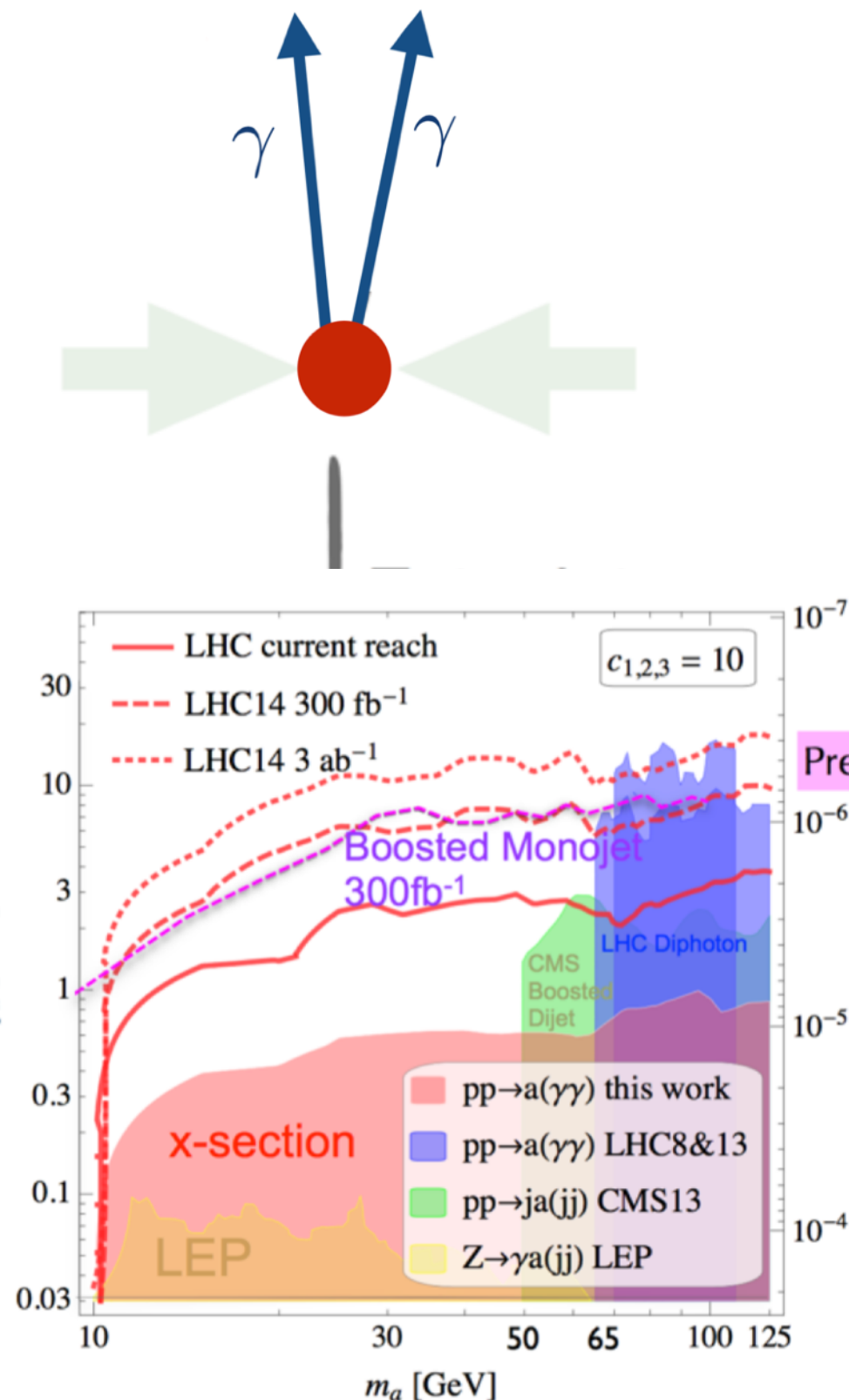
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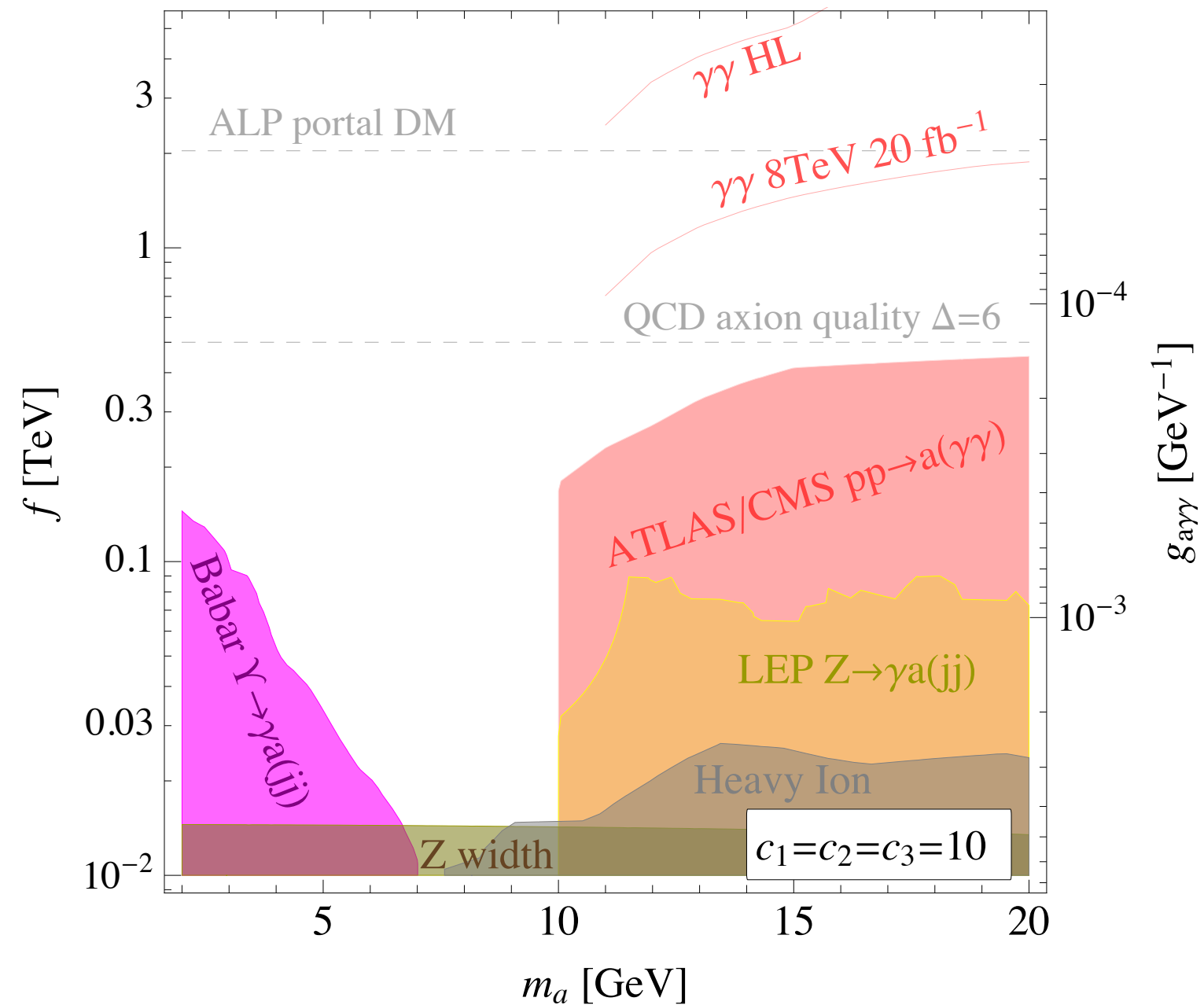
$$\sum_{i \neq \gamma_{\text{test}}, \gamma_1}^{\Delta R < R_{\text{iso}}} p_{T,i} \quad \sum_{i \neq \gamma_{\text{test}}, \gamma_1}^{\Delta R < R_{\text{iso}}} p_{T,i} / E_{T,\gamma_{\text{test}}}$$



final checks and optimisations in progress



Even smaller masses?

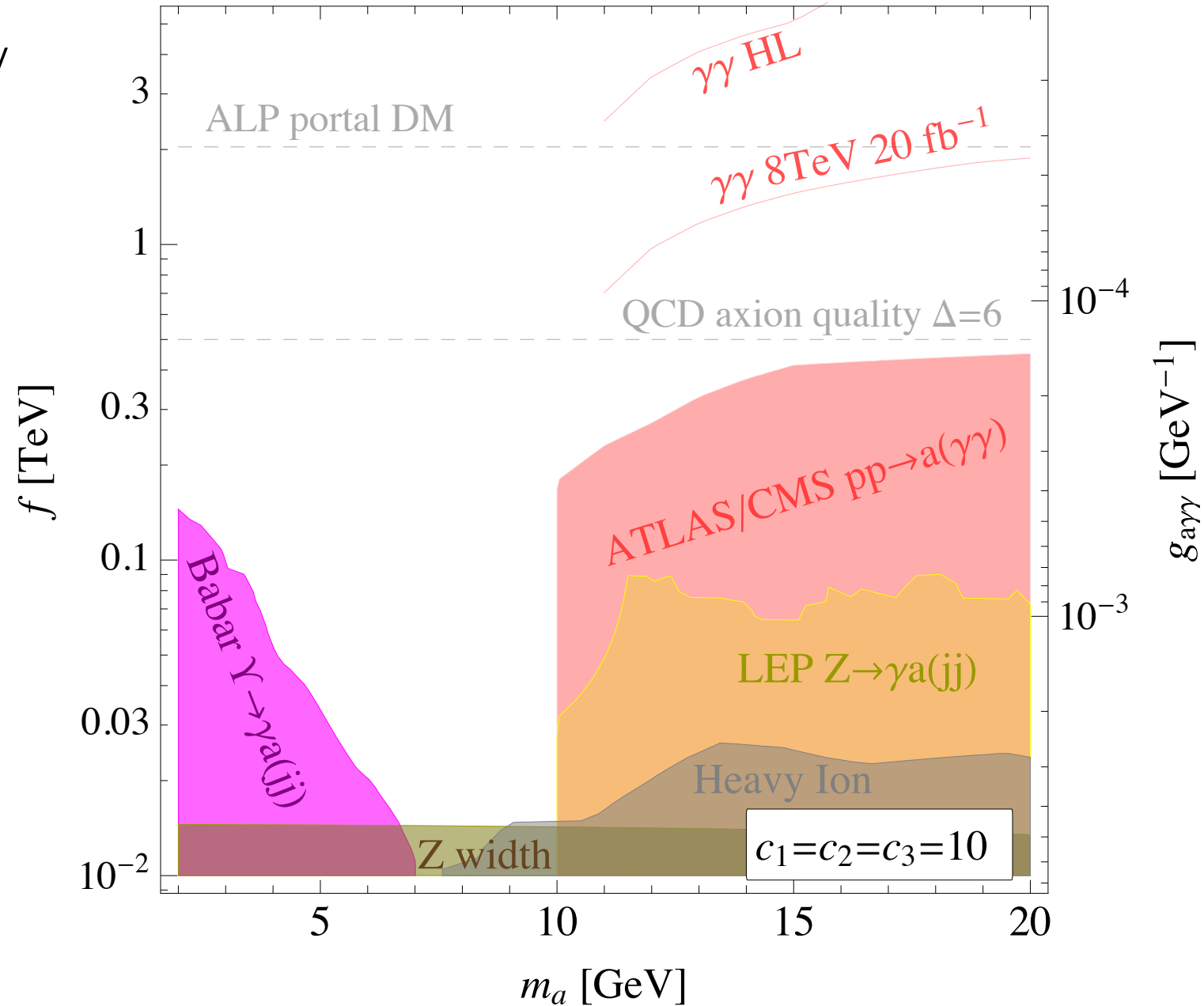
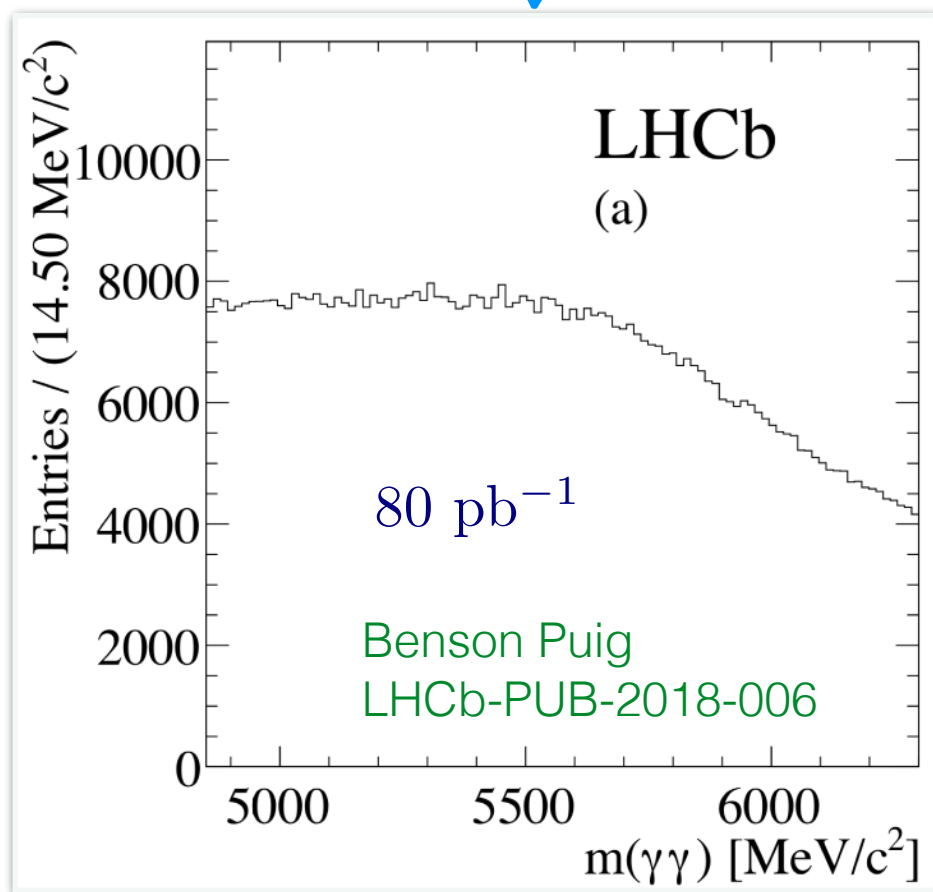


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Even smaller masses? LHCb

LHCb aims at competitive search for $B_s \rightarrow \gamma\gamma$

They published note with search strategy

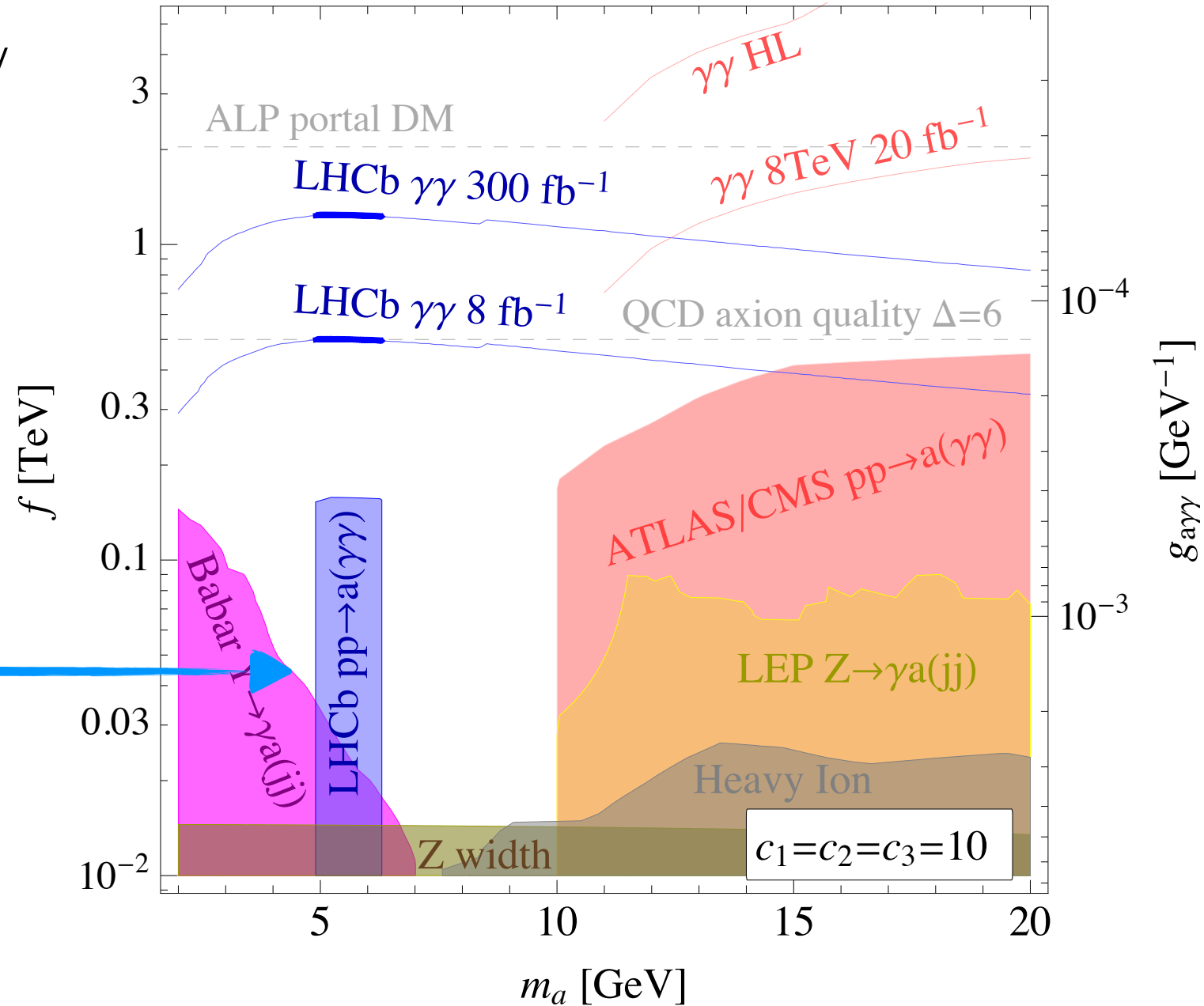
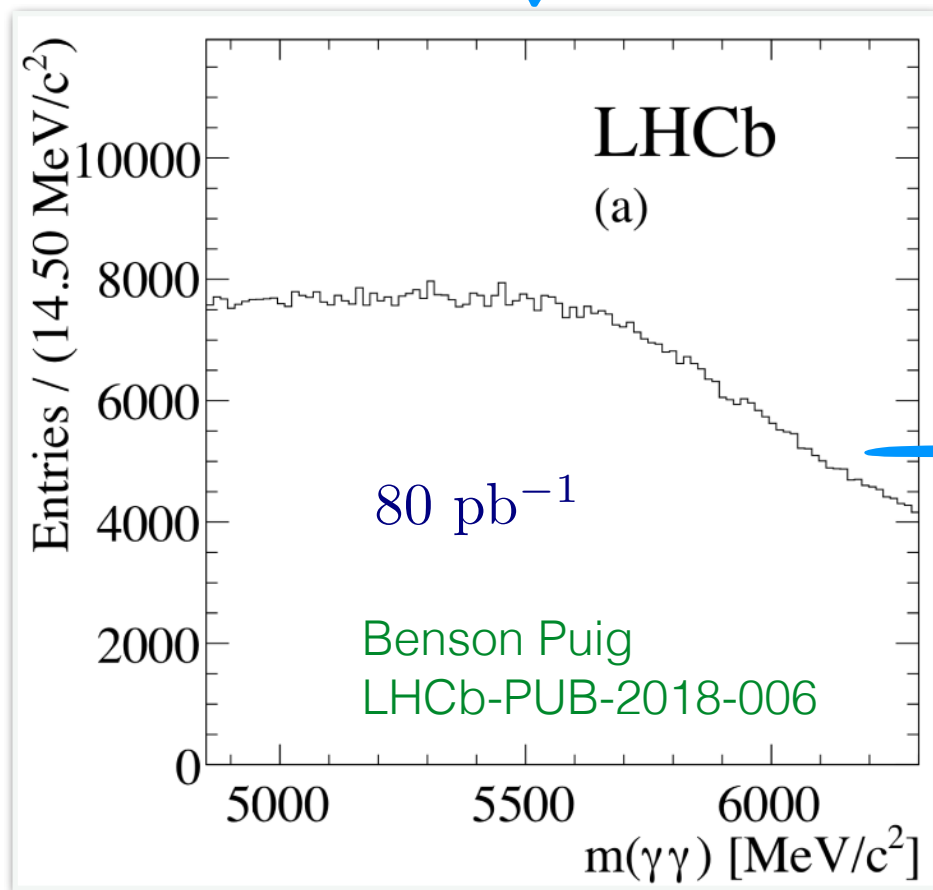


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Even smaller masses? Belle-II

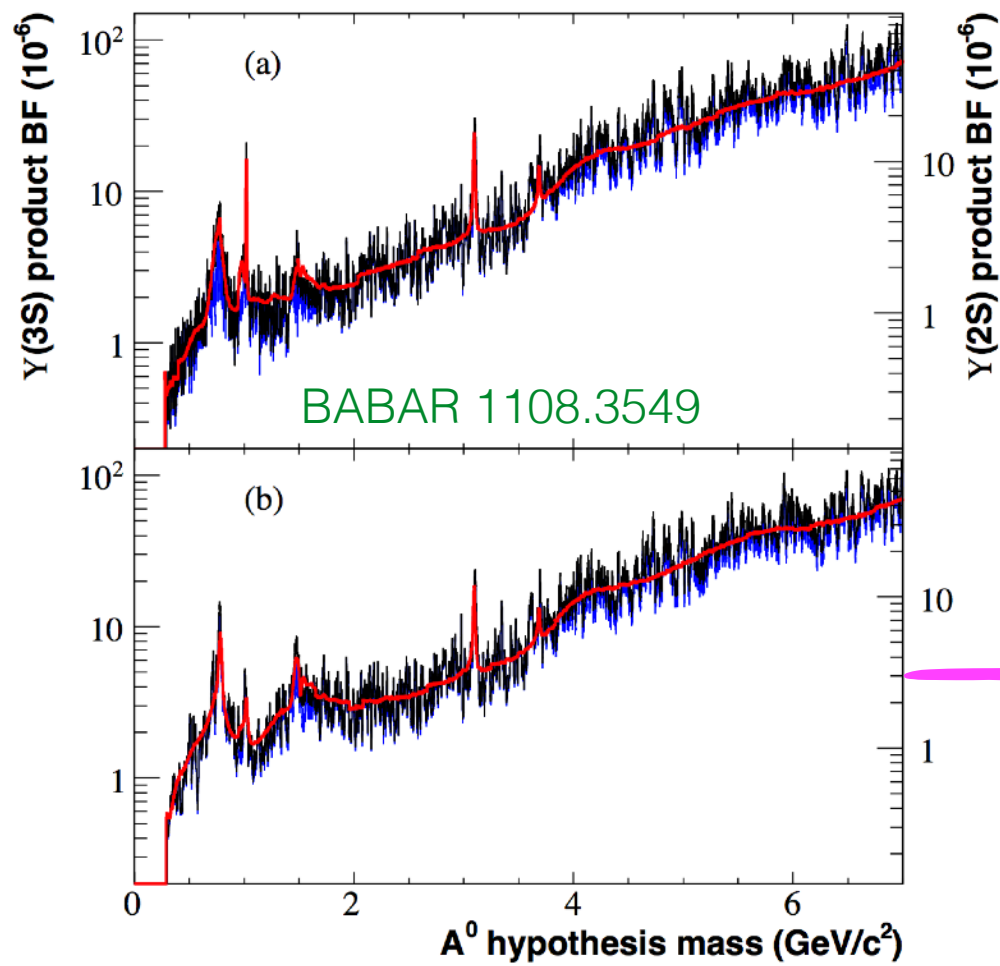
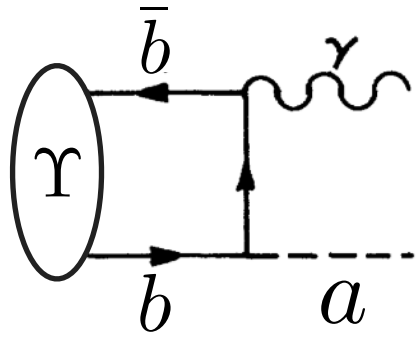
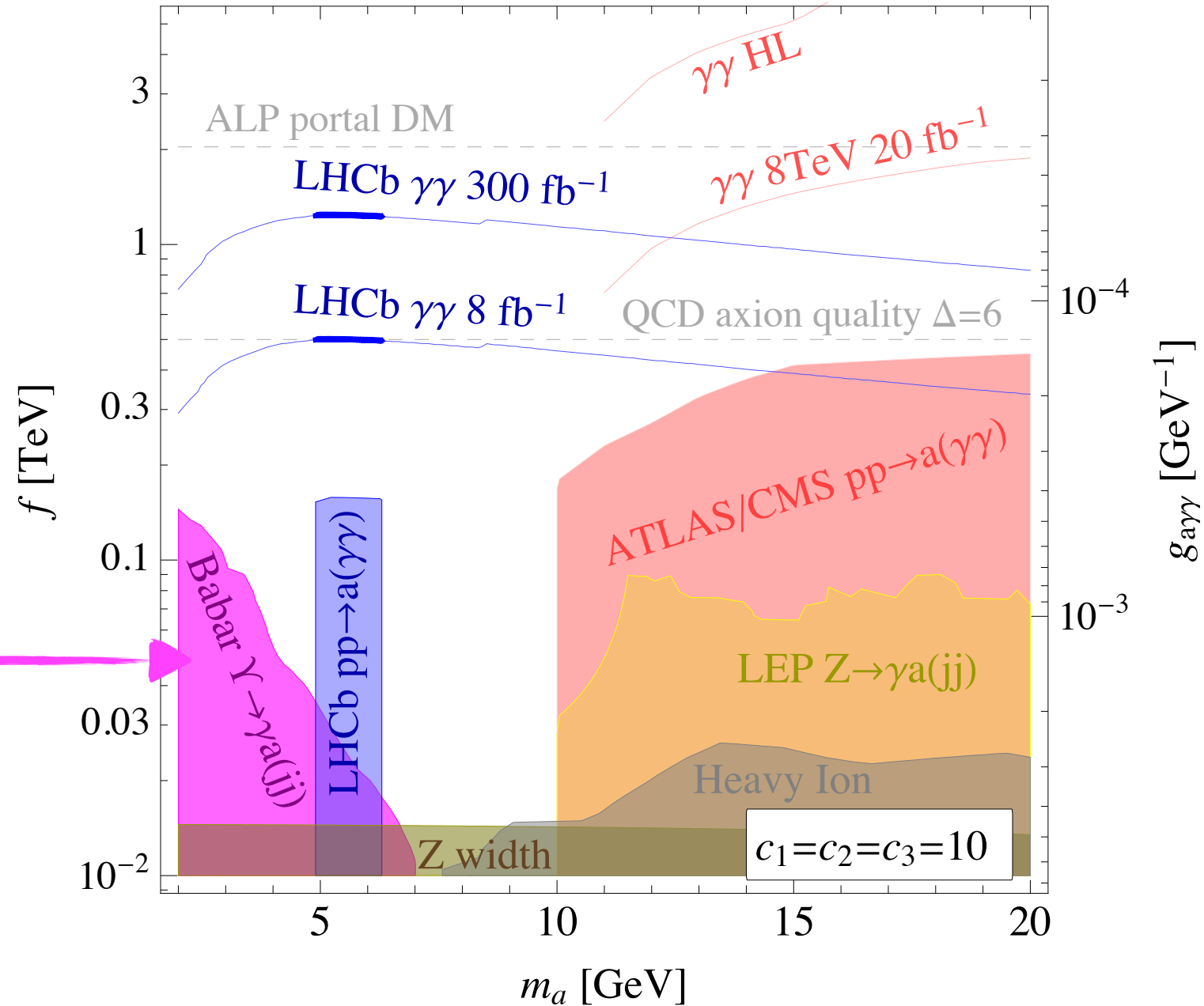


FIG. 5: 90% CL upper limits on product branching fractions (BF) (left axis) $\mathcal{B}(\Upsilon(3S) \rightarrow \gamma A^0) \cdot \mathcal{B}(A^0 \rightarrow \text{hadrons})$ and (right axis) $\mathcal{B}(\Upsilon(2S) \rightarrow \gamma A^0) \cdot \mathcal{B}(A^0 \rightarrow \text{hadrons})$, for (a) CP-all analysis, and (b) CP-odd analysis. The overlaid curves in red are the limits expected from simulated experiments, while the blue curves are the limits from statistical errors only. The



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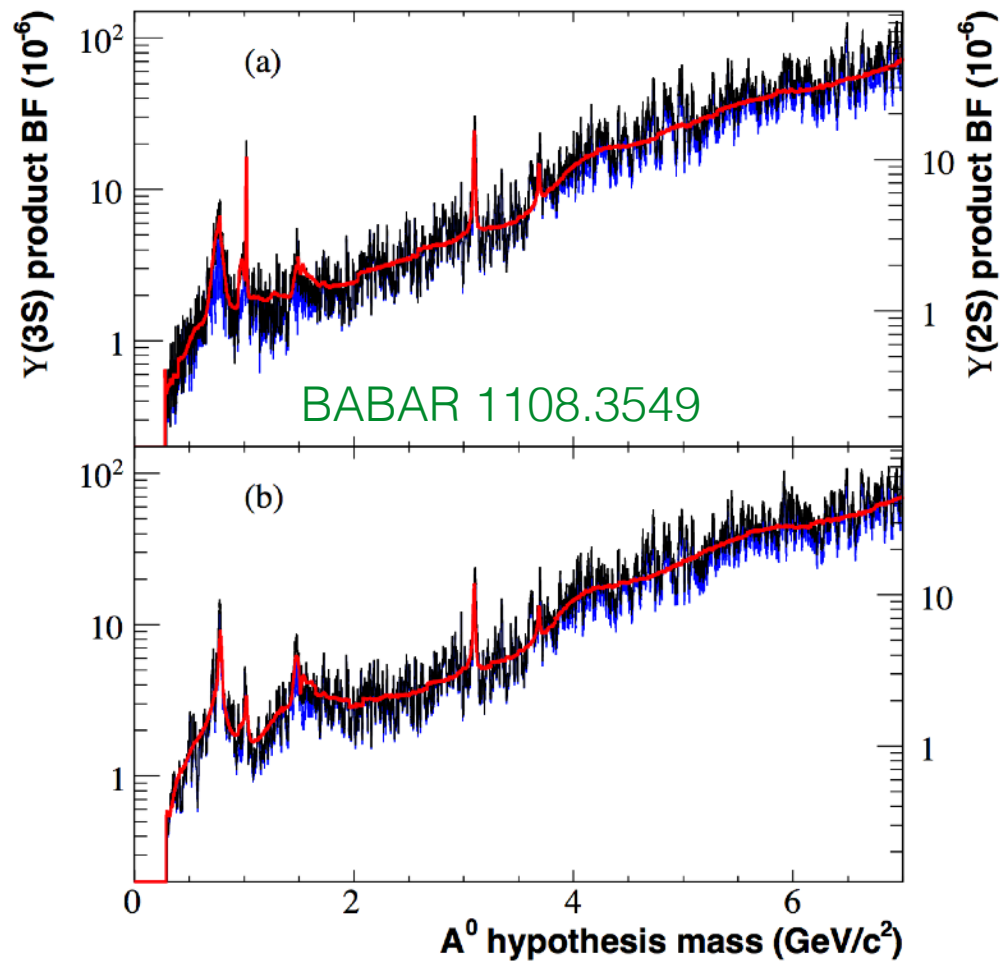
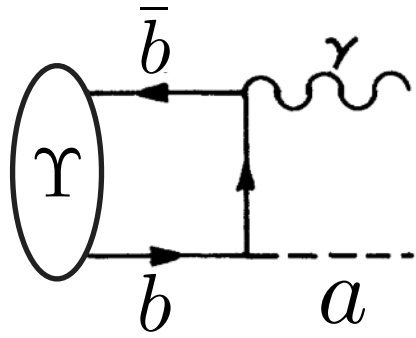
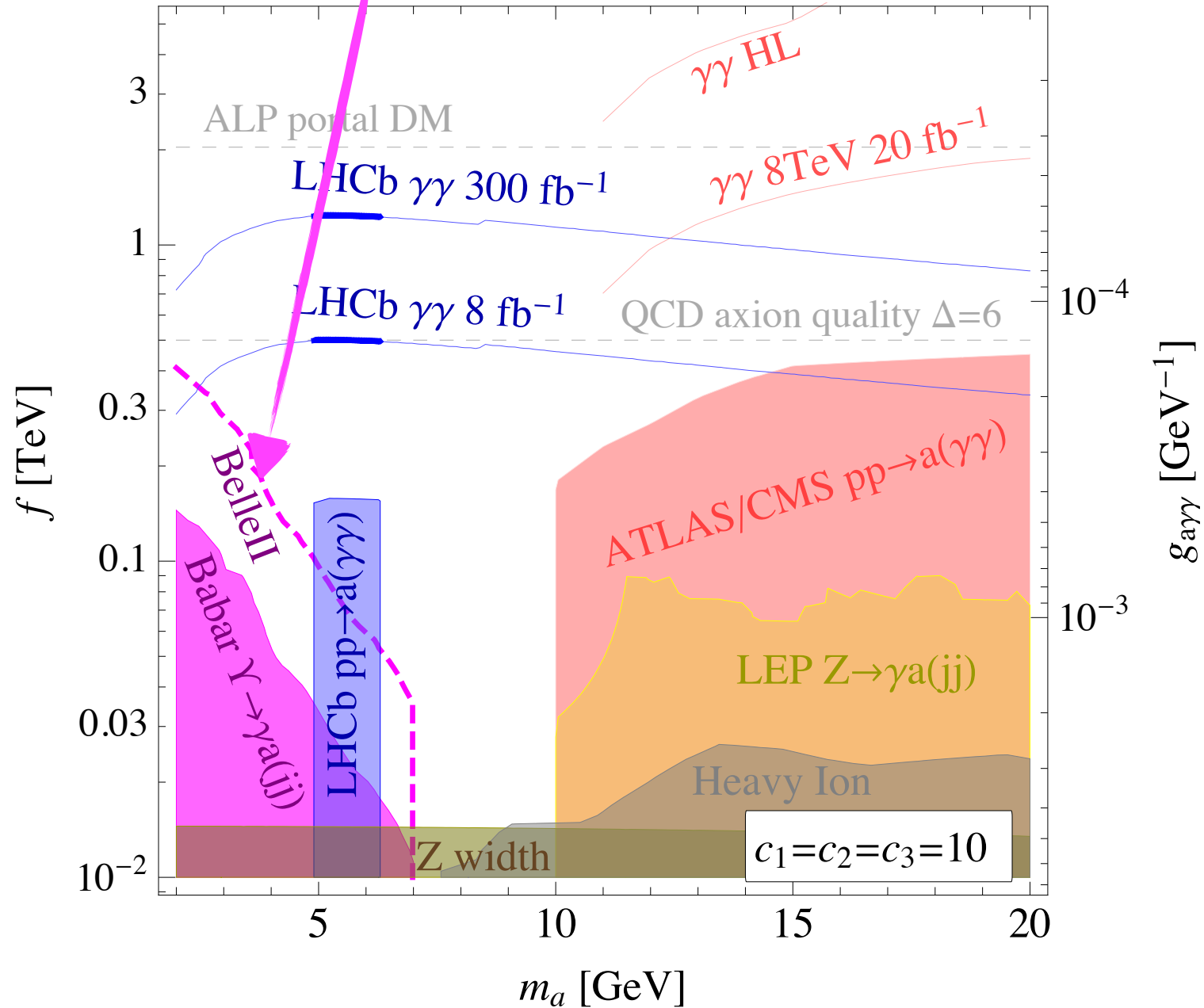


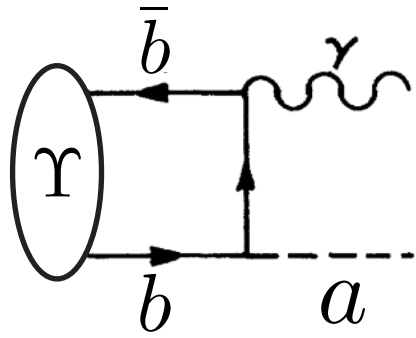
FIG. 5: 90% CL upper limits on product branching fractions (BF) (left axis) $\mathcal{B}(\Upsilon(3S) \rightarrow \gamma A^0) \cdot \mathcal{B}(A^0 \rightarrow \text{hadrons})$ and (right axis) $\mathcal{B}(\Upsilon(2S) \rightarrow \gamma A^0) \cdot \mathcal{B}(A^0 \rightarrow \text{hadrons})$, for (a) CP-all analysis, and (b) CP-odd analysis. The overlaid curves in red are the limits expected from simulated experiments, while the blue curves are the limits from statistical errors only. The

Assumes Belle-II produces **100x** more $\Upsilon(3S)$ than BABAR ($\Upsilon(1S, 2S)$ good as well)

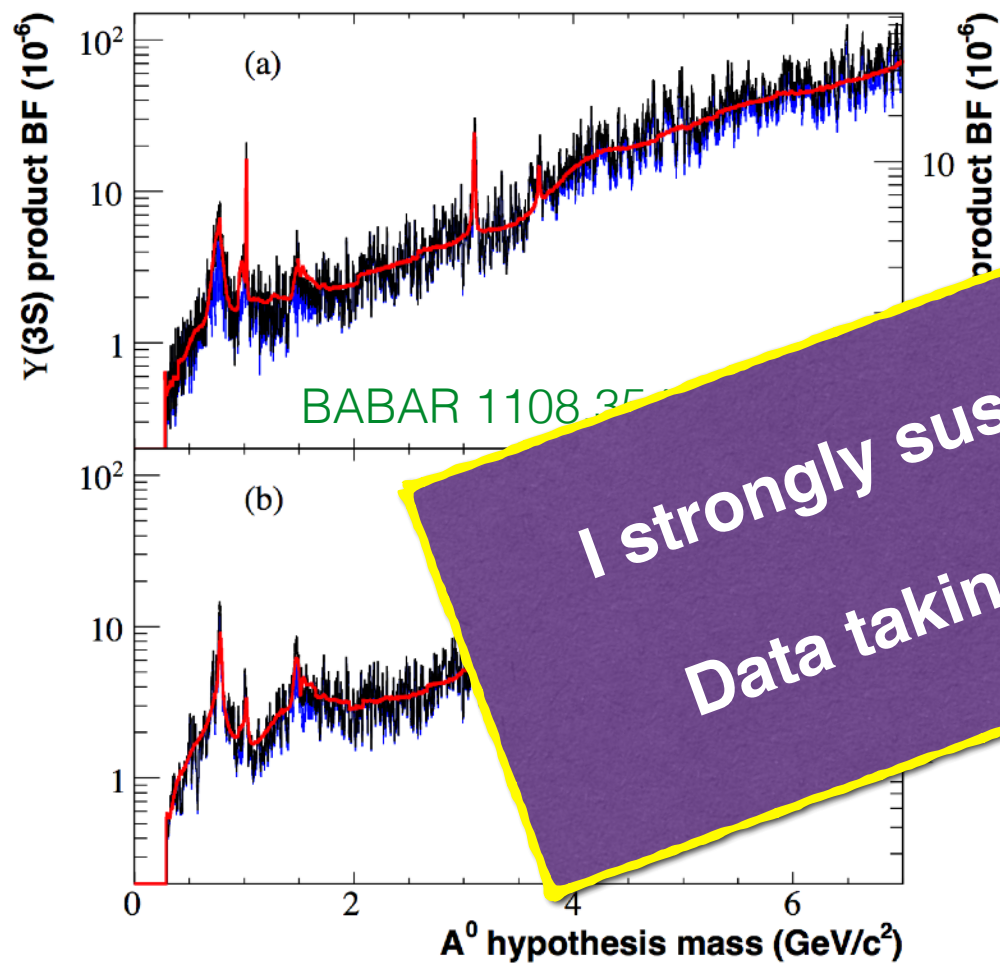


$$\mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\alpha_s c_3 G \tilde{G} + \alpha_2 c_2 W \tilde{W} + \alpha_1 c_1 B \tilde{B} \right]$$

Even smaller masses? Belle-II



Assumes Belle-II produces **100x** more $\Upsilon(3S)$ than BABAR ($\Upsilon(1S, 2S)$ good as well)



I strongly suspect that Belle-II could do better
Data taking started, should tell them soon!

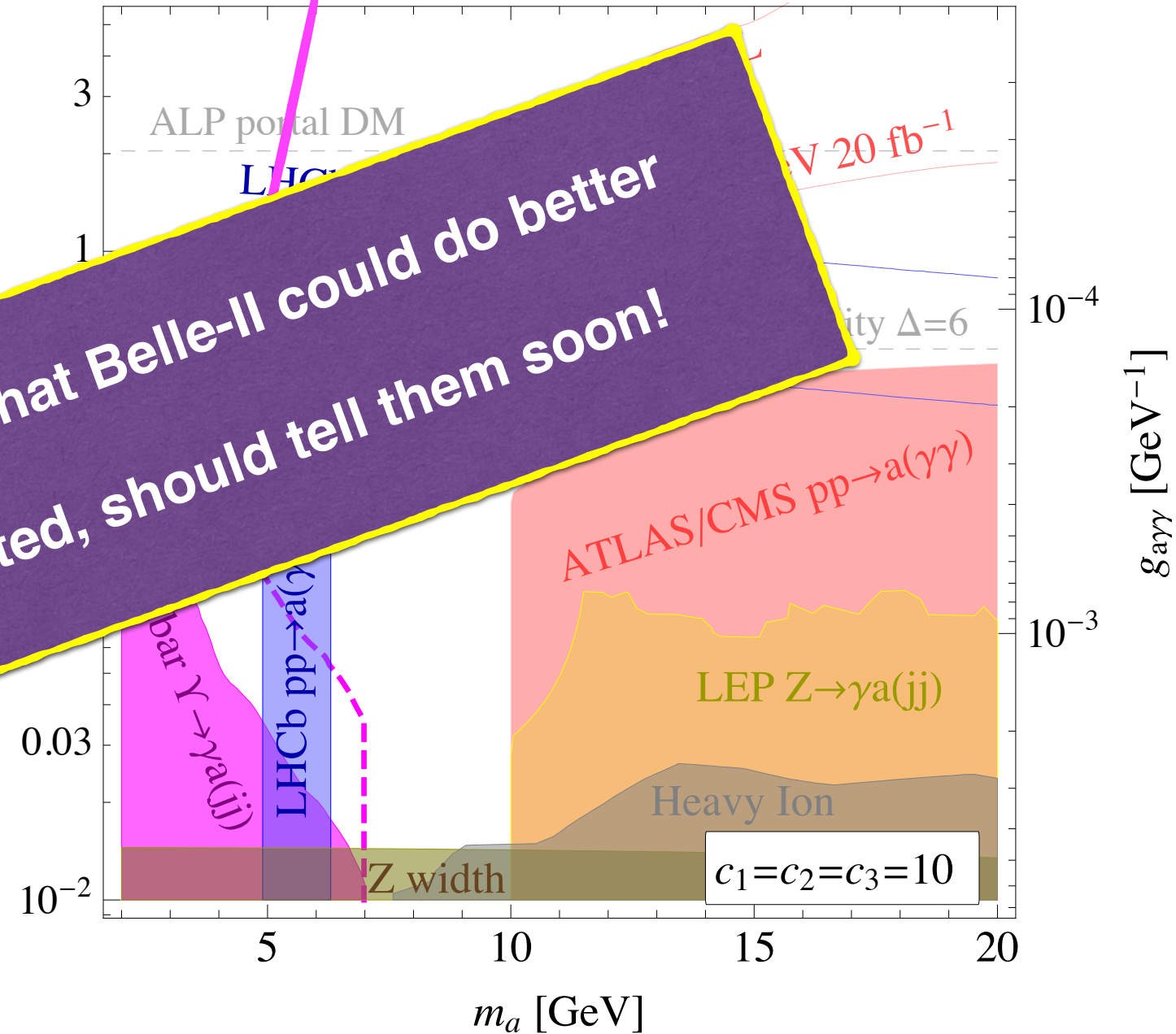


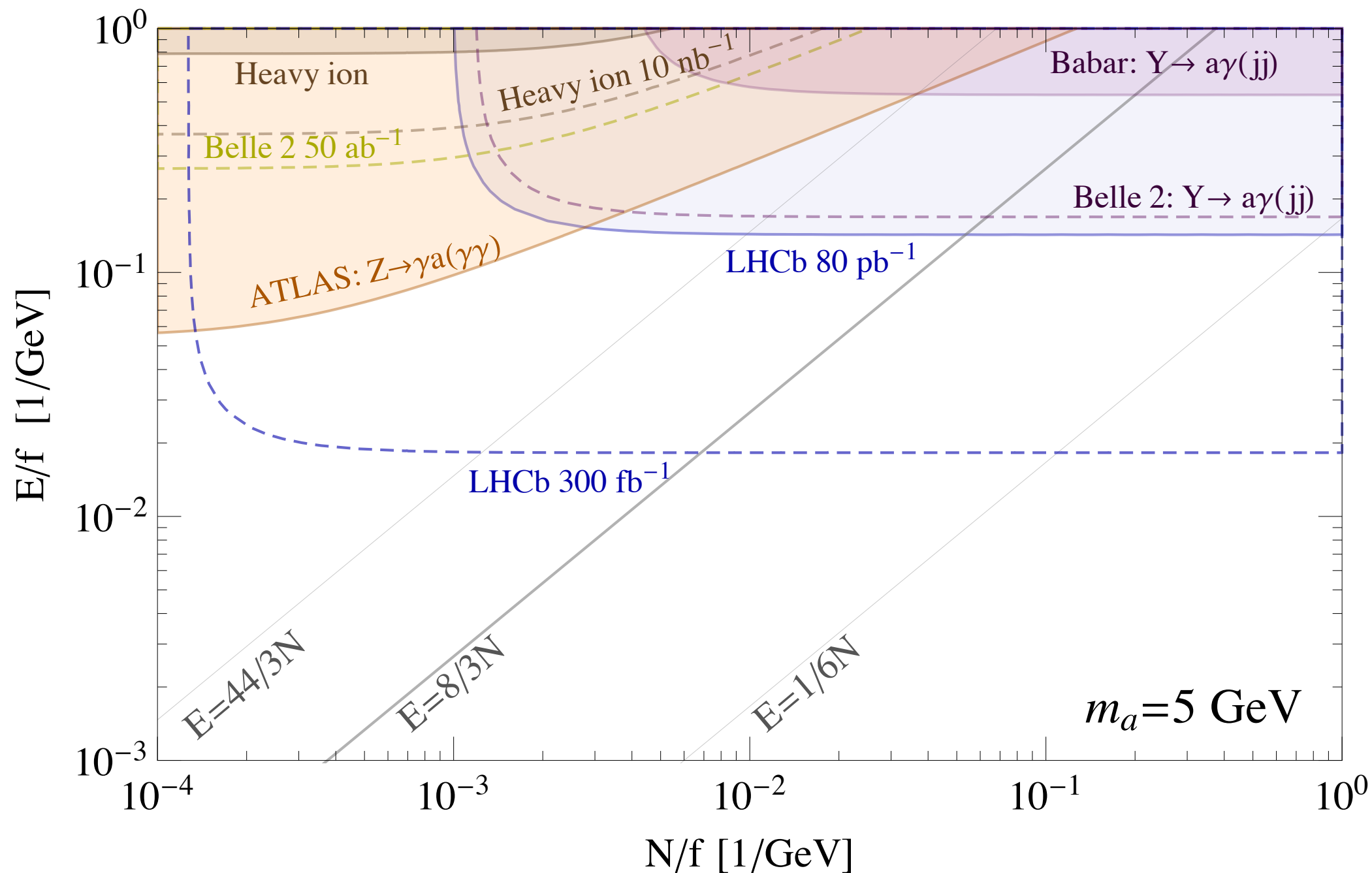
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What if different couplings?

Cid-Vidal Mariotti Redigolo FS Tobioka 1810.xxxxx

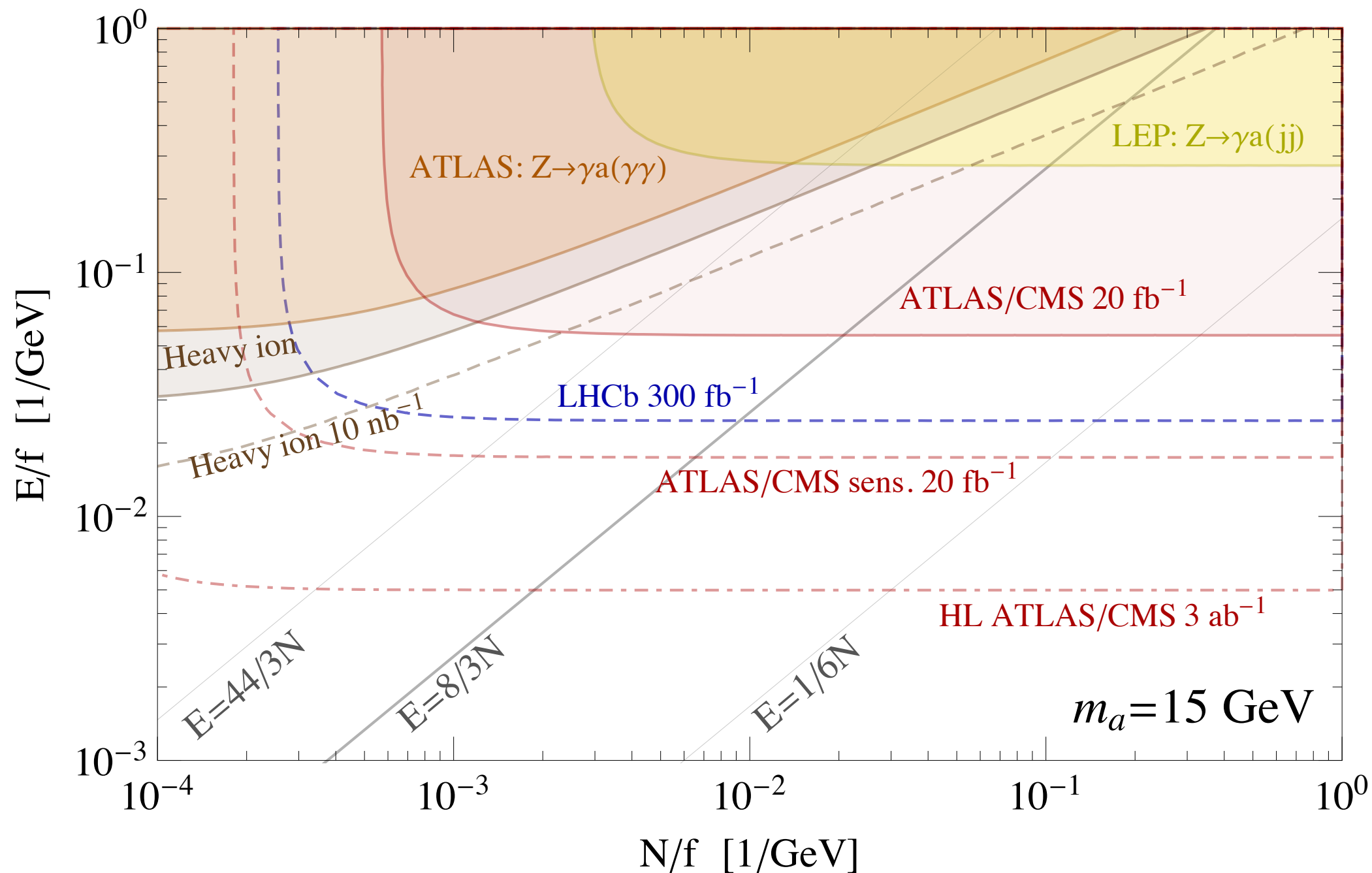
$$\mathcal{L}_{\text{eff}} \supset \frac{N\alpha_3}{4\pi} \frac{a}{f} G_{\mu\nu} \tilde{G}^{\mu\nu} + \frac{E\alpha_{\text{em}}}{4\pi} \frac{a}{f} F_{\mu\nu} \tilde{F}^{\mu\nu}$$



What if different couplings?

Cid-Vidal Mariotti Redigolo FS Tobioka 1810.xxxxx

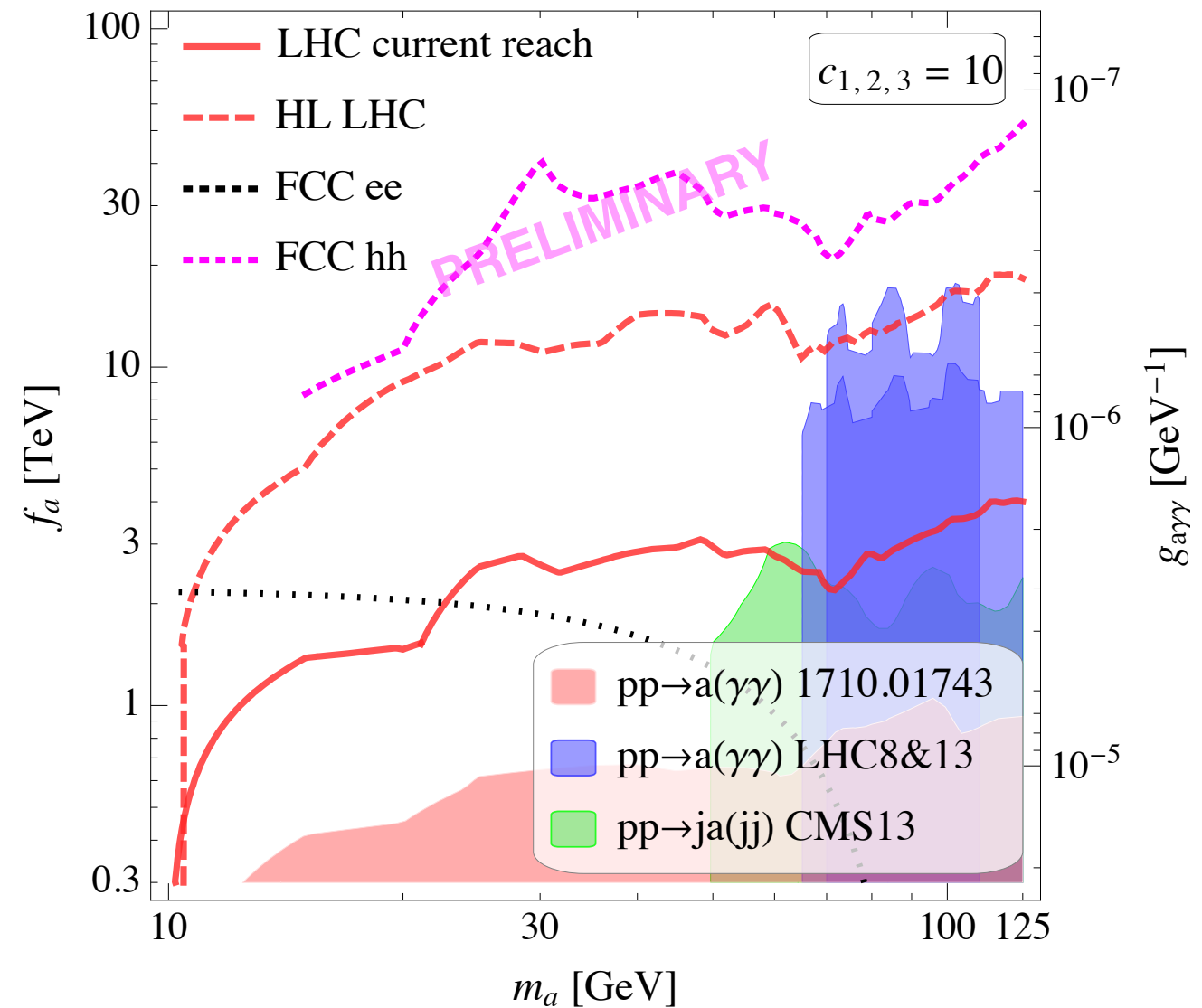
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Even more future?

FCC ee reach computed rescaling
LEP limits on $\text{BR}[Z \rightarrow \gamma a(jj)]$
and assuming 10^{12} Z bosons

If $aG\tilde{G}$ switched on
HL-LHC wins over FCC ee



Even more future?

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If $aG\tilde{G}$ switched on
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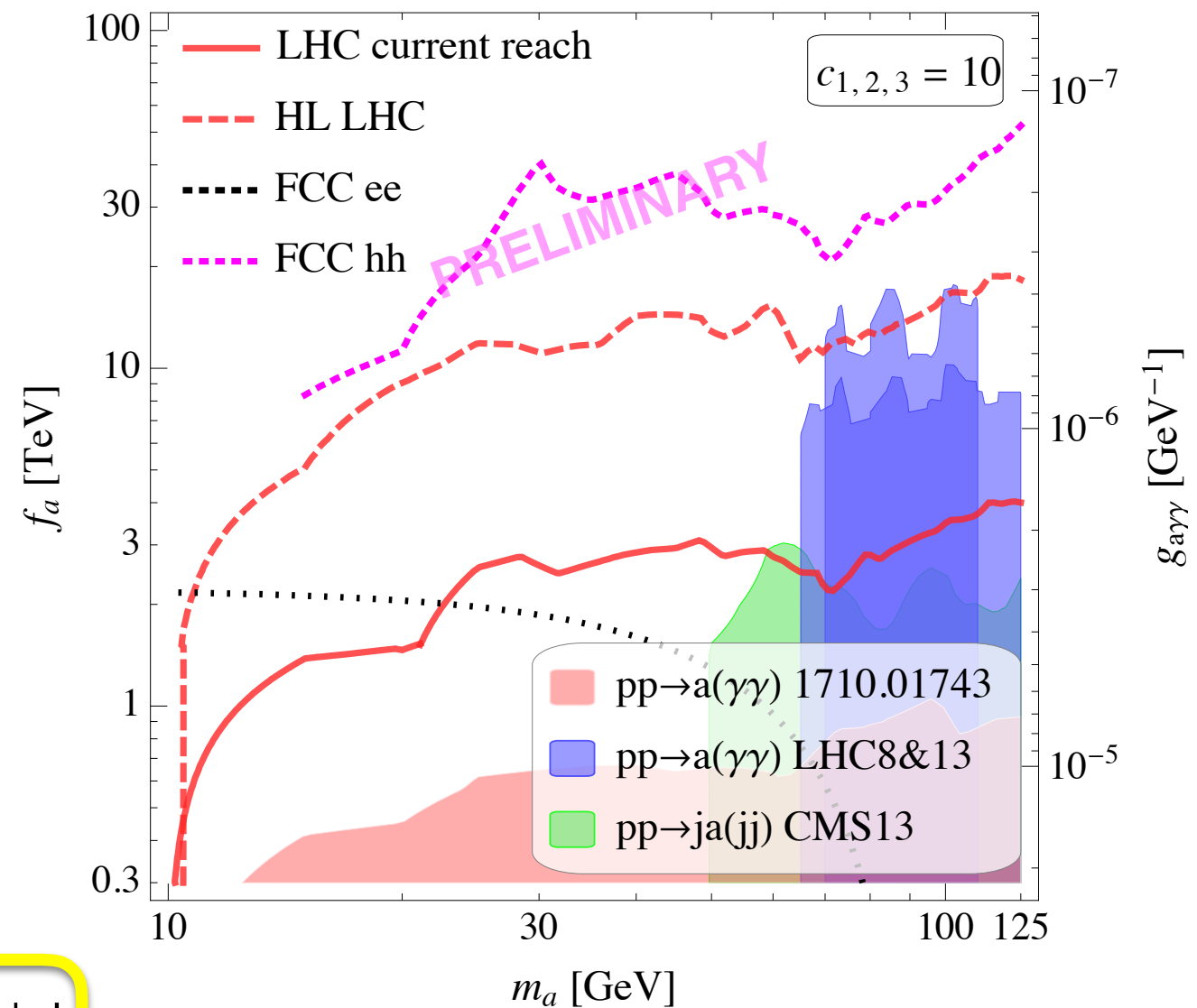
FCC hh reach computed like LHC14 one
[4., from simulations, Lumi=3 ab^{-1}]
NB. pT cuts as in ATLAS8 (30, 40 GeV)

Still, speculative even for FCC standards!

this search has not even been performed at 8 TeV

at 100 TeV game could be very different (larger boosts,...)

Thoughts in progress...



New BSM searches at LHC & Belle-II

■ **ALPs** at the **LHC** are theoretically well motivated (e.g. heavy **QCD axion**, SUSY **R-axion**,...)

■ Existing searches do not go below 50-70 GeV mass

■ But they could!

Lower masses in boosted dijet searches

We set **strongest bound on ALPs** from $\gamma\gamma$

We propose concrete new searches at **ATLAS&CMS**, **LHCb** and **Belle-II**

Mariotti Redigolo FS Tobioka 1710.01743

+ Cid Vidal 1810.xxxxx

+ Low in progress

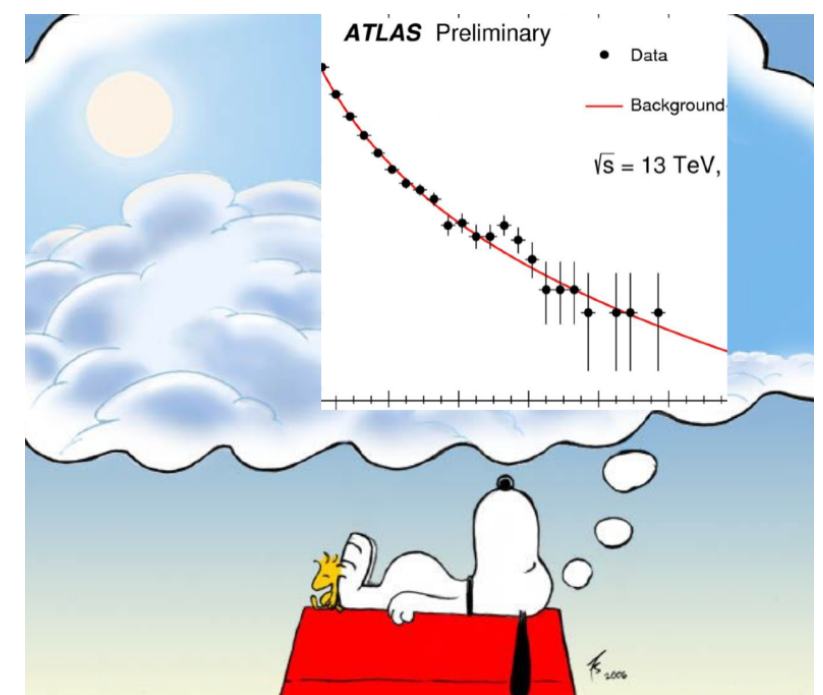
■ **NEXT:** LHC & Belle-II perform the actual searches!

Other LHC (e.g. coupling with fermions)?

Cacciapaglia+1710.11142

Lot of other Belle-II

....



Back up

More on R-axion

PGB from SUSY: R-symmetry

N = 1 SUSY always accompanied by a continuous $U(1)_R$ = “R-symmetry”

$$R : \theta_\alpha \rightarrow e^{i\epsilon} \theta_\alpha \quad [R, Q] = -Q$$

R-charge assignments:

$$\Phi = \phi + \sqrt{2}\theta \psi + \theta^2 F$$

$$r_\phi = r_\Phi$$

$$r_\psi = r_\Phi - 1$$

$$r_F = r_\Phi - 2$$

Vector superfields are real $\Rightarrow$ **gauginos** have **$r_\lambda = 1$**

Lagrangian $\mathcal{L}$ R-symmetric $\Rightarrow R(W) = 2$

($\Leftarrow$ if Kahler canonical)

$$\mathcal{L} \supset \int d^2\theta W + \text{c.c.}$$

W superpotential

A strongly coupled "UV" completion

Very low energy SUSY breaking F

motivated by:

Naturalness + Higgs mass [Gherghetta Pomarol 1107.4697](#)

+ LHC exclusions [Buckley et al. 1610.08059](#)

Gravitino cosmology [Ibe Yanagida 1608.01610](#)

needs a strongly coupled sector

so that $m_{\lambda_i} \sim \frac{g_i^2}{g_*^2} m_*$ OK with LHC bounds

m_* mass gap of the hidden sector
(e.g. mass of messengers in gauge mediation)

$g_* > 1$ coupling between hidden sector states

$$m_*$$

$$m_{\text{soft}} \approx \frac{g^2}{g_*^2} m_*$$

$$m_a \approx \sqrt{\epsilon_R} m_*$$

$$m_G = \frac{F}{\sqrt{3} M_{Pl}}$$

SUSY Naive Dimensional Analysis

$$M_{\text{SUSY}} \sim m_* \sim g_* f \quad f_a \sim f$$

$$F \sim g_* f^2 \quad w_R \sim g_* f^3$$

inspired by
[Cohen et al. 1997](#)
[Luty 1998](#)
[Giudice+ 2007](#)

$a \rightarrow GG$ saturates the upper bound

The R-axion pheno Lagrangian-I

Komargodski Seiberg 0907.2441

Tool: constrained superfield formalism

$$X = \frac{G^2}{2F_X} + \sqrt{2}\theta G + \theta^2 F_X$$

$$\mathcal{R} = e^{i\mathcal{A}/f_a} = e^{ia/f_a + O(aG, \dots)}$$

satisfy the constraints $\begin{cases} X^2 = 0 \\ X(R^\dagger R - 1) = 0 \end{cases}$

~ analogous to ordinary Goldstones $U^\dagger U = 1 \quad U = e^{i\pi}$

Most general effective Lagrangian:

$$r_X = 2 \quad r_{\mathcal{R}} = 1$$

$$\mathcal{L}_{G+a} = \int d^4\theta (X^\dagger X + f_a^2 \mathcal{R}^\dagger \mathcal{R}) + \int d^2\theta (FX + w_R \mathcal{R}^2) + \text{c.c.}$$

Absent for any other axion

$$-\frac{w_R}{f_a F^2} \Box a \bar{G} i \gamma_5 G$$

First pheno prediction (valid for any UV completion!):

R-axion decays to missing energy

$$w_R < \frac{1}{2} f_a F$$

$$\Gamma_{a \rightarrow GG} < \frac{1}{32\pi} \frac{m_a^5}{F^2}$$

Dine Festuccia Komargodski 0910.2527

see also Bellazzini 1605.06111

R-axion pheno overview

Tool: constrained superfield formalism

Komargodski Seiberg 0907.2441

$$X = \frac{G^2}{2F_X} + \sqrt{2}\theta G + \theta^2 F_X$$

satisfy the constraints

$$\begin{cases} X^2 = 0 \\ X(R^\dagger R - 1) = 0 \end{cases}$$

$$\mathcal{R} = e^{i\mathcal{A}/f_a} = e^{ia/f_a} + O(a G, \dots)$$

~ analogous to
ordinary Goldstones

$$U^\dagger U = 1 \quad U = e^{i\pi}$$

$$r_{\mathcal{R}} = 1 \quad r_X = 2 \quad r_{\mathcal{W}} = 1$$

$$\mathcal{L}_{\text{gauge}} = \int d^2\theta \left(\frac{1}{4} - i g_i^2 \frac{c_i^{\text{hid}}}{16\pi^2} \mathcal{A} \right) \mathcal{W}_i^2 - \int d^2\theta \frac{m_{\lambda_i}}{2F} X \mathcal{R}^{-2} \mathcal{W}_i^2 + \text{c.c.}$$

$$\frac{g_i^2 c_i^{\text{hid}}}{16\pi^2} \frac{a}{f_a} F^i \tilde{F}^i$$

$$g_i^2 \frac{c_i^{\text{eff}}}{16\pi^2} \frac{\partial_\mu a}{f_a} \bar{\lambda}_i \gamma_\mu \gamma_5 \lambda_i - i \frac{m_{\lambda_i}}{f_a} a \bar{\lambda}_i \gamma_5 \lambda_i$$

$$R_H \equiv r_{H_u} + r_{H_d}$$

$$\mathcal{L}_{\text{Higgs}} \supset \int d^4\theta \left(\frac{\mu}{F} X^\dagger H_u H_d \mathcal{R}^{2-R_H} - \frac{B_\mu}{F^2} X^\dagger X H_u H_d \mathcal{R}^{-R_H} + \text{c.c.} \right)$$

$$-i a R_H \left(c_\beta^2 \frac{m_u}{f_a} \bar{u} \gamma_5 u + s_\beta^2 \frac{m_d}{f_a} \bar{d} \gamma_5 d + s_\beta^2 \frac{m_\ell}{f_a} \bar{\ell} \gamma_5 \ell \right)$$

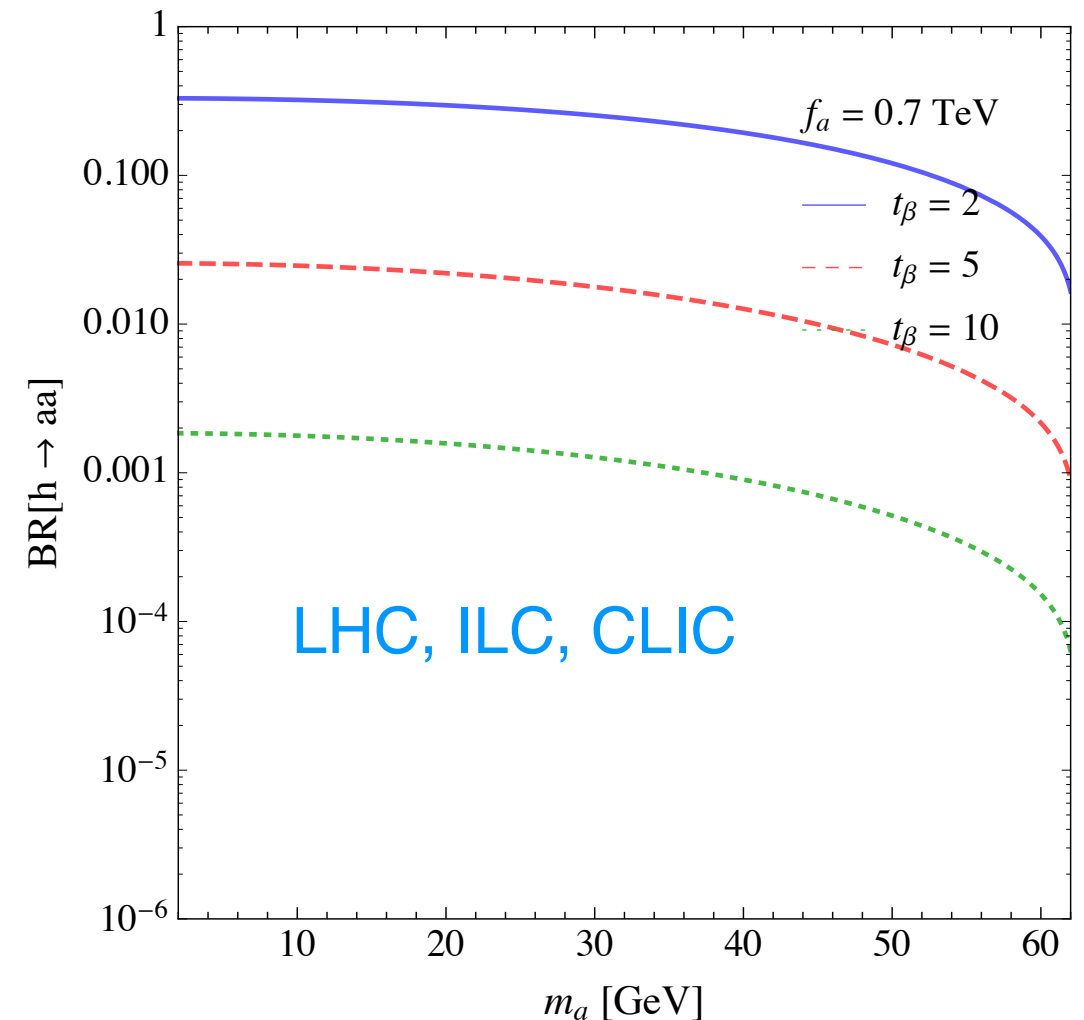
$$\frac{\delta^2}{v} (\partial_\mu a)^2 h$$

$$\delta = R_H \frac{v}{f_a} \frac{s_{2\beta}}{2}$$

a from decays of h, Υ and B

$$\mathcal{L}_{ha^2} = \frac{\delta^2}{v} (\partial_\mu a)^2 h$$

$$\delta = R_H \frac{v}{f_a} \frac{s_{2\beta}}{2}$$



$$\text{BR}_{\Upsilon \rightarrow \gamma a} \simeq 3 - 5 \times 10^{-5} \left(\frac{\text{TeV}}{f_a} \right)^2$$

since Wilczek PRL39 (1977)

experiments: **BABAR**
Belle-II

$$\text{BR}_{B \rightarrow K a, K^* a} \simeq 3 - 5 \times 10^{-4} \left(\frac{\text{TeV}}{f_a} \right)^2$$

see Hall Wise 1981, Freytsis Ligeti Thaler 0911.5355

LHCb
Belle, Belle-II

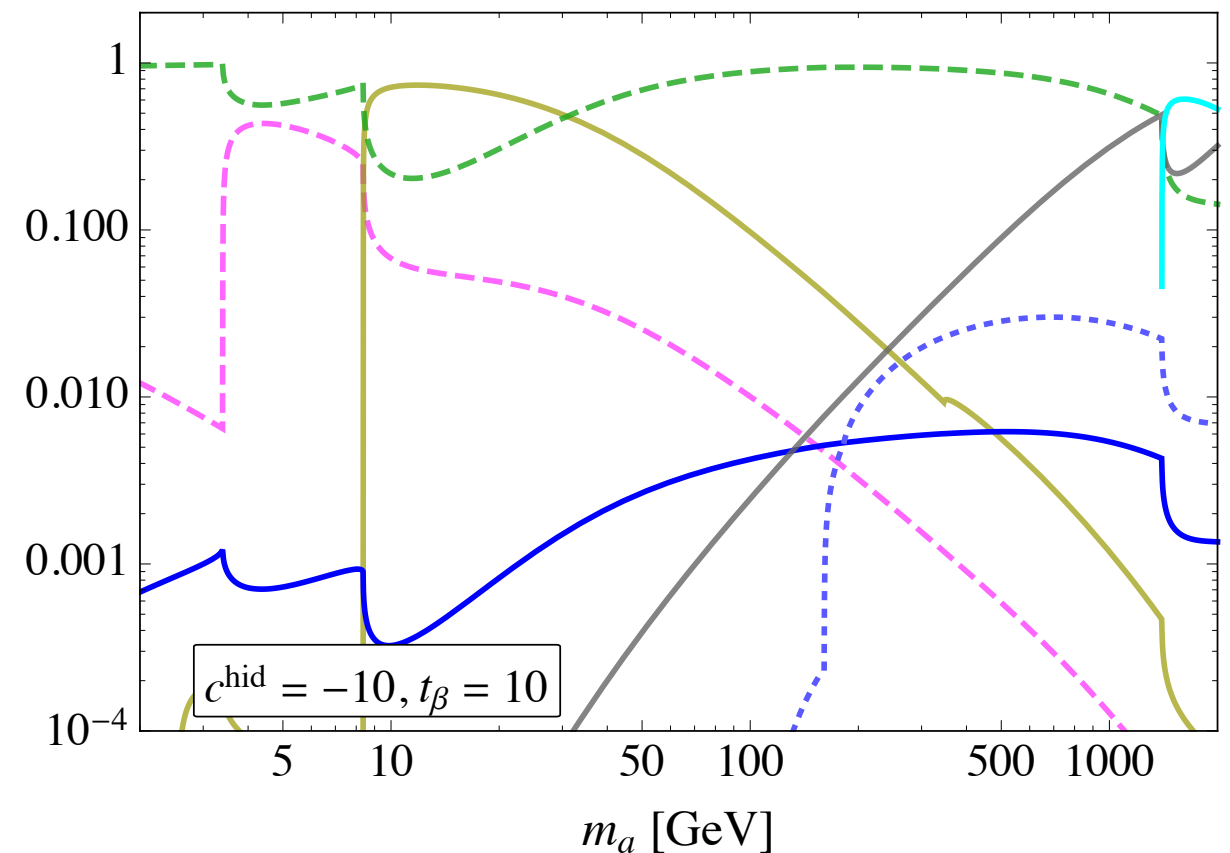
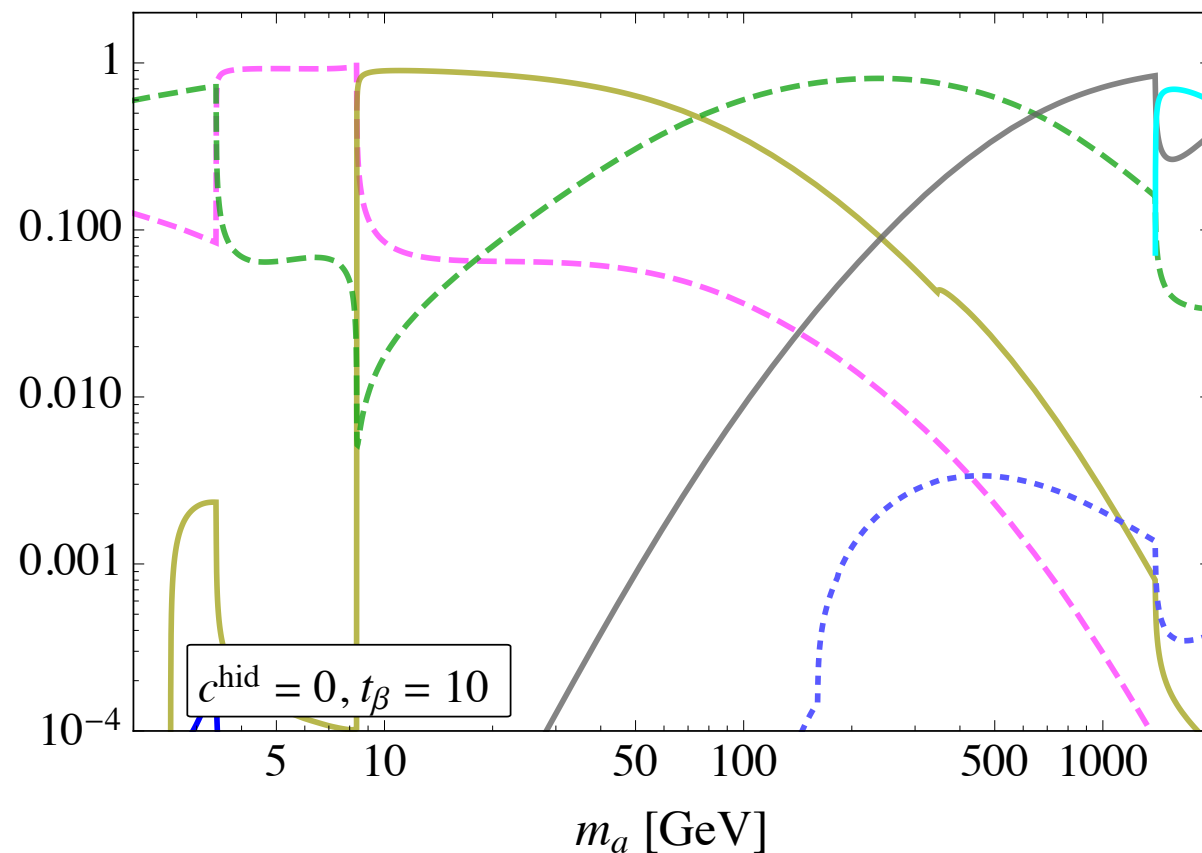
R axion branching ratios

Both plots: $t_\beta = 10$

No anomaly

Large anomalies

BRs: $\mu\mu + \tau\tau$, $cc + bb + tt$, $\gamma\gamma$, jj , $WW + ZZ + Z\gamma$, $inv.$, $\gamma\gamma+MET$



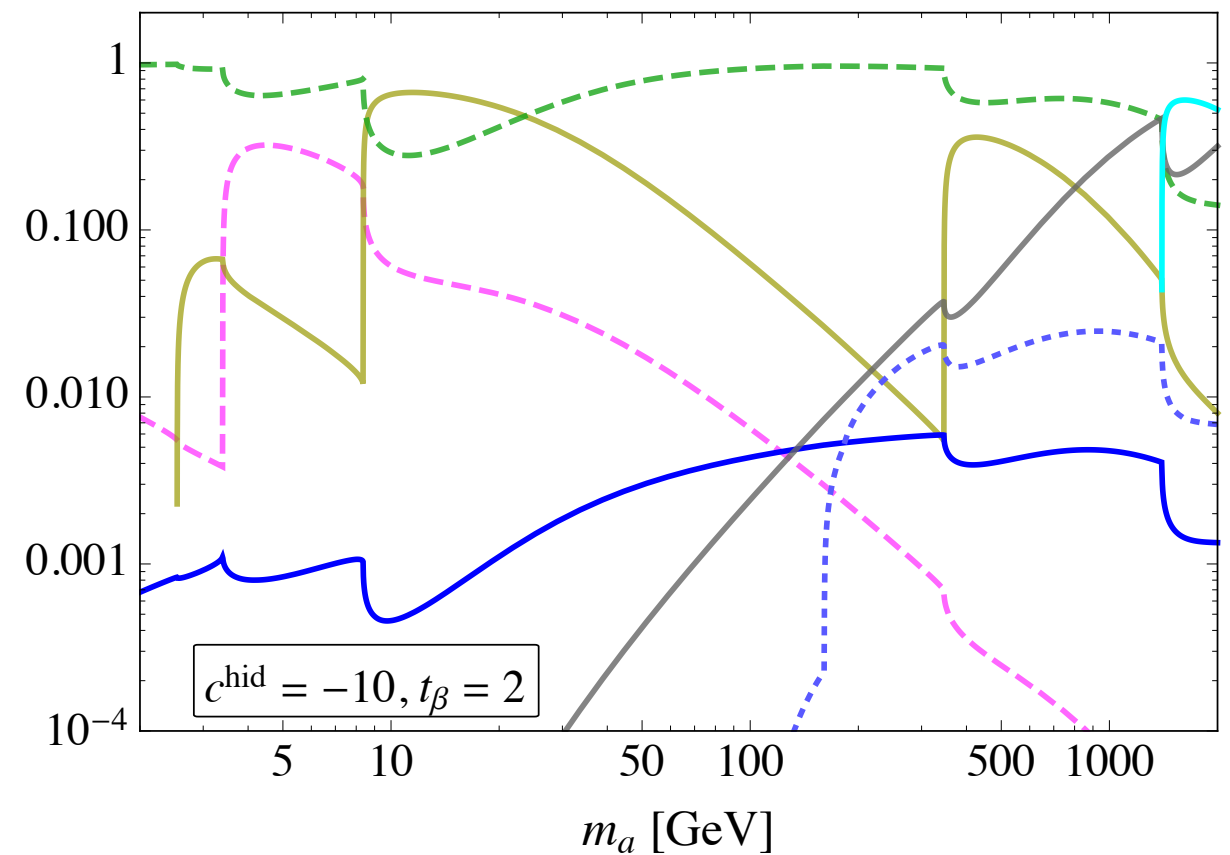
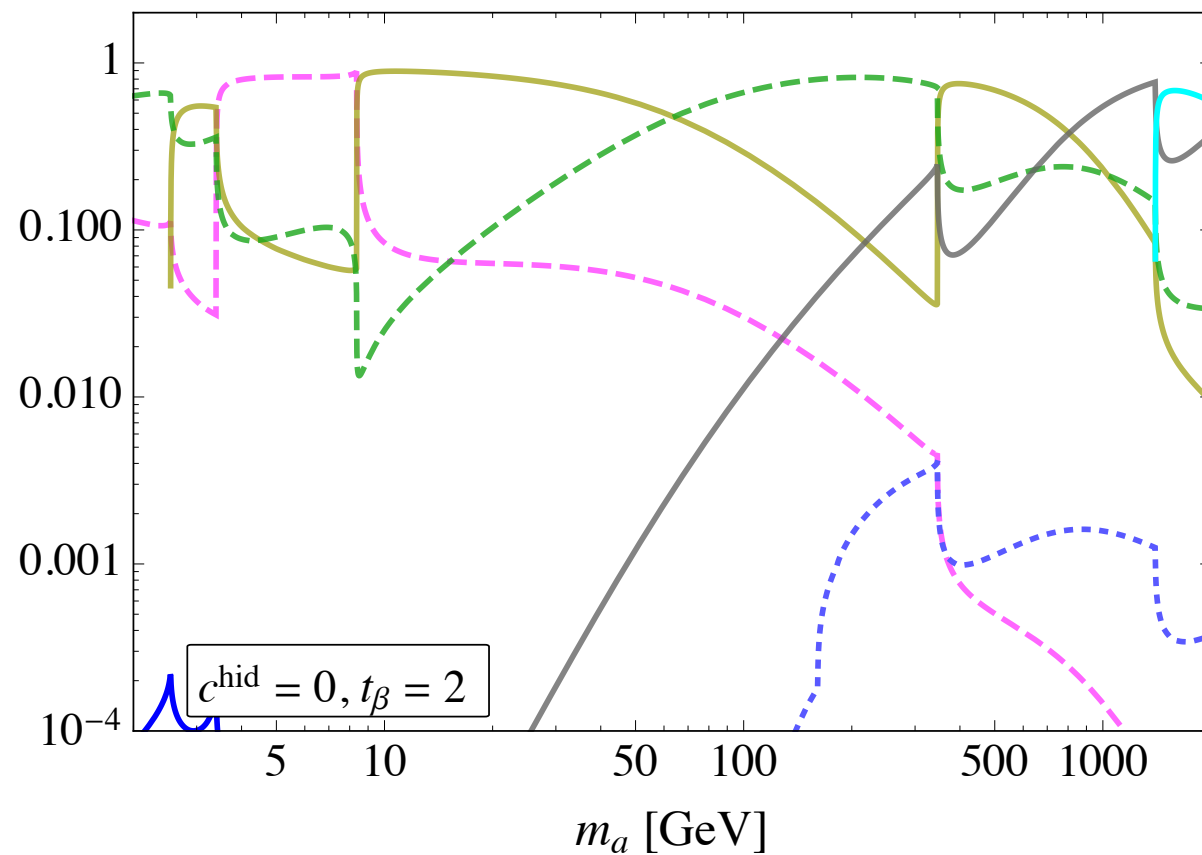
R axion branching ratios

Both plots: $t_\beta = 2$

No anomaly

Large anomalies

BRs: $\mu\mu + \tau\tau$, $cc + bb + tt$, $\gamma\gamma$, jj , $WW + ZZ + Z\gamma$, $inv.$, $\gamma\gamma+MET$



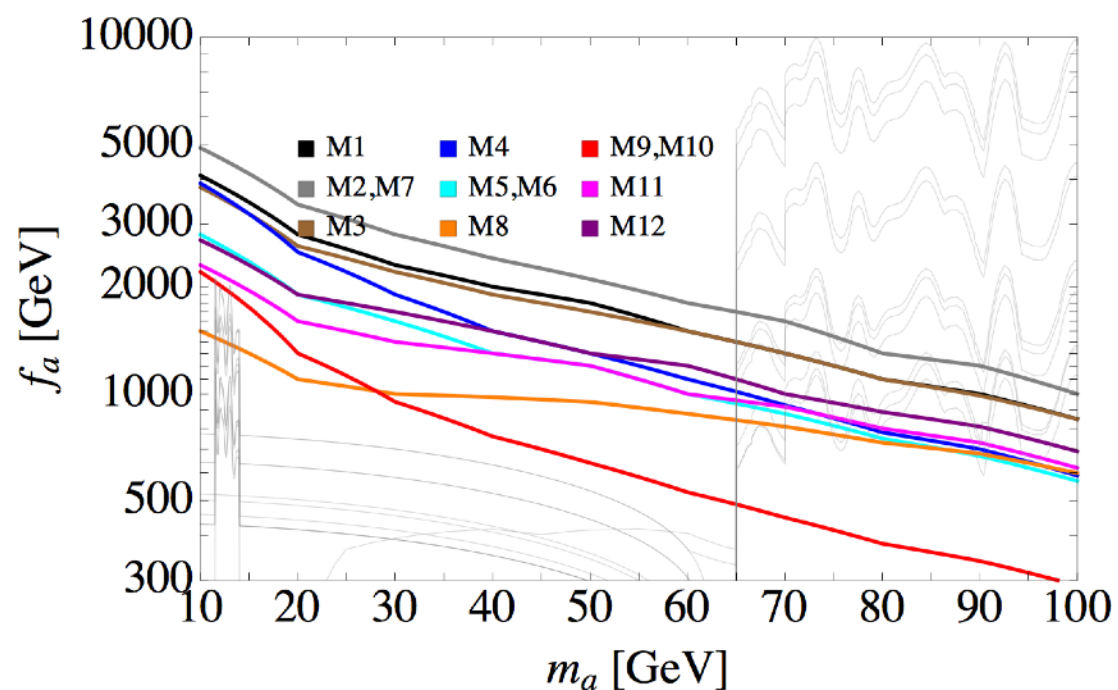
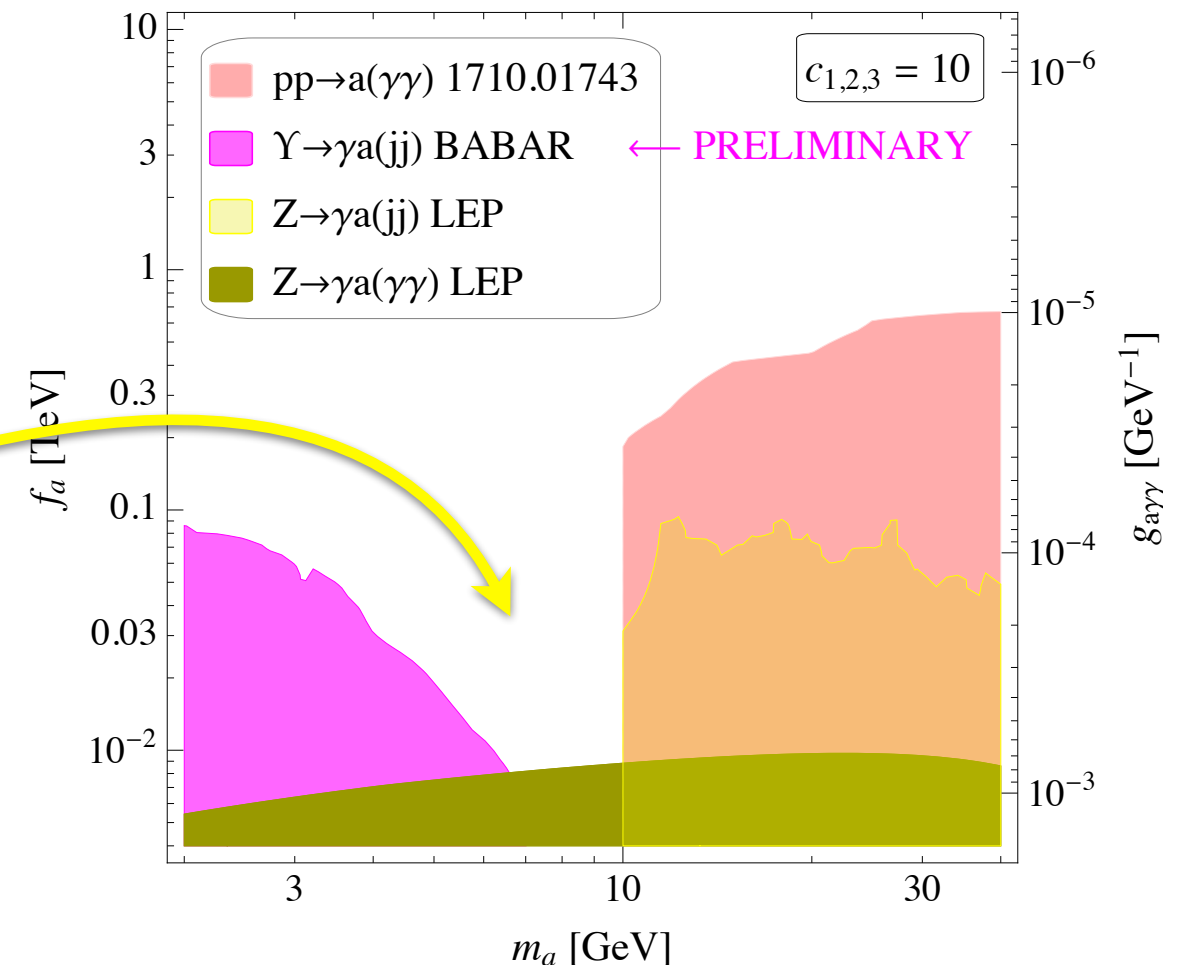
More on low-mass $\gamma\gamma$

Other ways to low-mass resonances?

$m_{\gamma\gamma} < 10 \text{ GeV}$ at the LHCb?

work in progress...

Big hole for $4 \text{ GeV} \lesssim m_a \lesssim 10 \text{ GeV}$



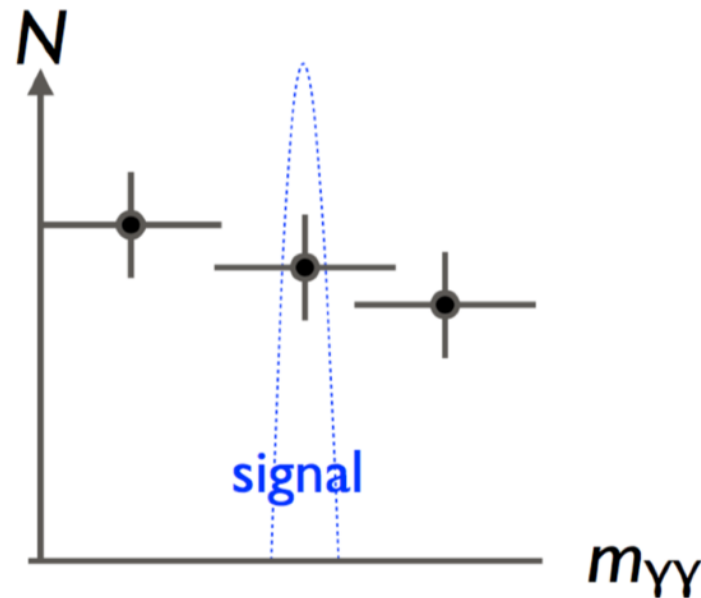
Difermions, e.g. ditaus?

Cacciapaglia Ferretti Flacke Serodio 1710.11142

NB. sensitivities not based on data
definitely worth investigating

New $\gamma\gamma$ Bound & Sensitivities

Starting point: inclusive **diphoton cross section measurements** @ ATLAS7,8 and CMS7



1. New Bound we assume zero knowledge of bkg

$$N_{\text{bin}}^{\text{signal}} < N_{\text{bin}}^{\text{meas.}} (1 + 2 \Delta_{\text{bin}})$$

experimental rel. uncertainty

2. Reach we assume data = SM prediction

$$N_{\text{bin}}^{\text{signal}} < N_{\text{bin}}^{\text{meas.}} \times 2 \Delta_{\text{bin}}$$

3. Reach with smarter bins (= **2.**, where we reduce ~ 10 GeV bins to mass resolution of ~ 3 GeV)

4. Reach we simulate bkg with same cuts at different energies [Madgraph+Pythia+Delphes]

$$N_{\text{bin}}^{\text{signal}}(E_{\text{high}}) < N_{\text{bin}}^{\text{signal}}(E_{\text{low}}) \cdot \sqrt{\frac{L_{\text{high}}}{L_{\text{low}}} \cdot \frac{\sigma_{\text{high}}^{\text{MC}}}{\sigma_{\text{low}}^{\text{MC}}}}$$

Reach of **3.**

Simulated bkg cross sections

Low-mass analyses we found

Experiment	Process	Lumi	$\sqrt{s}$	low mass reach	ref.
LEPI	$e^+e^- \rightarrow Z \rightarrow \gamma a \rightarrow \gamma jj$	12 pb^{-1}	Z-pole	10 GeV	[29]
LEPI	$e^+e^- \rightarrow Z \rightarrow \gamma a \rightarrow \gamma\gamma\gamma$	78 pb^{-1}	Z-pole	3 GeV	[30]
LEPII	$e^+e^- \rightarrow Z^*, \gamma^* \rightarrow \gamma a \rightarrow \gamma jj$	$9.7, 10.1, 47.7 \text{ pb}^{-1}$	161, 172, 183 GeV	60 GeV	[31]
LEPII	$e^+e^- \rightarrow Z^*, \gamma^* \rightarrow \gamma a \rightarrow \gamma\gamma\gamma$	$9.7, 10.1, 47.7 \text{ pb}^{-1}$	161, 172, 183 GeV	60 GeV	[31, 32]
LEPII	$e^+e^- \rightarrow Z^*, \gamma^* \rightarrow Z a \rightarrow jj\gamma\gamma$	$9.7, 10.1, 47.7 \text{ pb}^{-1}$	161, 172, 183 GeV	60 GeV	[31]
D0/CDF	$p\bar{p} \rightarrow a \rightarrow \gamma\gamma$	$7/8.2 \text{ fb}^{-1}$	1.96 TeV	100 GeV	[33]
ATLAS	$pp \rightarrow a \rightarrow \gamma\gamma$	20.3 fb^{-1}	8 TeV	65 GeV	[34]
CMS	$pp \rightarrow a \rightarrow \gamma\gamma$	19.7 fb^{-1}	8 TeV	80 GeV	[35]
CMS	$pp \rightarrow a \rightarrow \gamma\gamma$	19.7 fb^{-1}	8 TeV	150 GeV	[36]
CMS	$pp \rightarrow a \rightarrow \gamma\gamma$	35.9 fb^{-1}	13 TeV	70 GeV	[37]
CMS	$pp \rightarrow a \rightarrow jj$	18.8 fb^{-1}	8 TeV	500 GeV	[38]
ATLAS	$pp \rightarrow a \rightarrow jj$	20.3 fb^{-1}	8 TeV	350 GeV	[39]
CMS	$pp \rightarrow a \rightarrow jj$	12.9 fb^{-1}	13 TeV	600 GeV	[40]
ATLAS	$pp \rightarrow a \rightarrow jj$	3.4 fb^{-1}	13 TeV	450 GeV	[41]
CMS	$pp \rightarrow ja \rightarrow jjj$	35.9 fb^{-1}	13 TeV	50 GeV	[42]
UA2	$p\bar{p} \rightarrow a \rightarrow \gamma\gamma$	13.2 pb^{-1}	0.63 TeV	17.9 GeV	[43]
D0	$p\bar{p} \rightarrow a \rightarrow \gamma\gamma$	4.2 fb^{-1}	1.96 TeV	8.2 GeV	[44]
CDF	$p\bar{p} \rightarrow a \rightarrow \gamma\gamma$	5.36 fb^{-1}	1.96 TeV	6.4 GeV	[45, 46]
ATLAS	$pp \rightarrow a \rightarrow \gamma\gamma$	4.9 fb^{-1}	7 TeV	9.4 GeV	[8]
CMS	$pp \rightarrow a \rightarrow \gamma\gamma$	5.0 fb^{-1}	7 TeV	14.2 GeV	[10]
ATLAS	$pp \rightarrow a \rightarrow \gamma\gamma$	20.2 fb^{-1}	8 TeV	13.9 GeV	[9]

Signal efficiencies and cross section

$$\epsilon_S(m_a) = \frac{\sigma_{\gamma\gamma}^{\text{MCcuts}}(m_a, s)}{C_s \sigma_{\gamma\gamma}^{\text{LO}}(m_a, s)}$$

$\sigma_{\gamma\gamma}^{\text{MCcuts}}$ Simulated w/Madgraph+Pythia+Delphes
matched up to 2 extra jets

$\sigma_{\gamma\gamma}^{\text{LO}}$ reproduces up to a constant factor C_s the shape of $\sigma_{\gamma\gamma}^{\text{MCtot}}$ for $m_{\gamma\gamma} \gtrsim 60$ GeV (i.e. sufficiently far from the sum of the minimal detector p_T cuts on the photons). A constant factor $C_s \equiv \sigma_{\gamma\gamma}^{\text{MCtot}}(s)/\sigma_{\gamma\gamma}^{\text{LO}}(s)$ is hence included in Eq. (5) and we obtain $C_{7\text{TeV}} \simeq C_{8\text{TeV}} \simeq 0.85$ while $C_{2\text{TeV}} \simeq 1$ at the Tevatron center of mass energy. The

$$\sigma_{\gamma\gamma}^{\text{th}}(m_a, s) = \frac{K_\sigma}{K_g} \cdot \sigma_{\gamma\gamma}^{\text{LO}}(m_a, s), \quad (\text{A1})$$

where we work in the approximation $\Gamma_{\text{tot}} \simeq \Gamma_{gg}$ (which is excellent in the parameter space that we have studied), and where

$$\sigma_{\gamma\gamma}^{\text{LO}}(m_a, s) = \frac{1}{m_a s} C_{gg}(m_a^2/s) \cdot \Gamma_{\gamma\gamma}, \quad (\text{A2})$$

$$C_{gg} = \frac{\pi^2}{8} \int_{m_a^2/s}^1 \frac{dx}{x} f_g(x) f_g\left(\frac{m_a^2}{sx}\right), \quad (\text{A3})$$

where $f_g(x)$ is the gluon PDF from the MSTW2008nn1o68 set [58], where we fix the pdf scale $q = m_a$. We work with

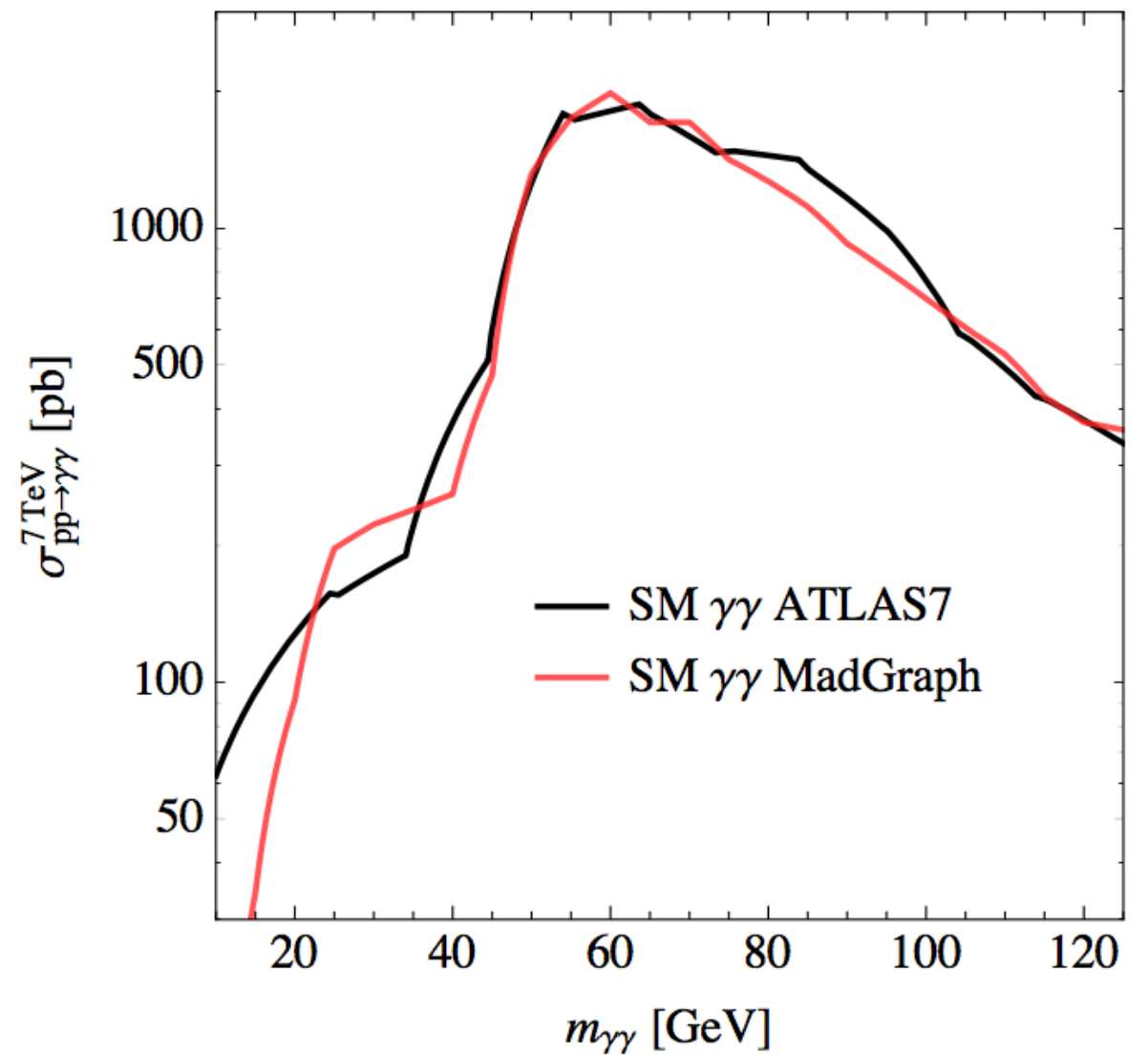
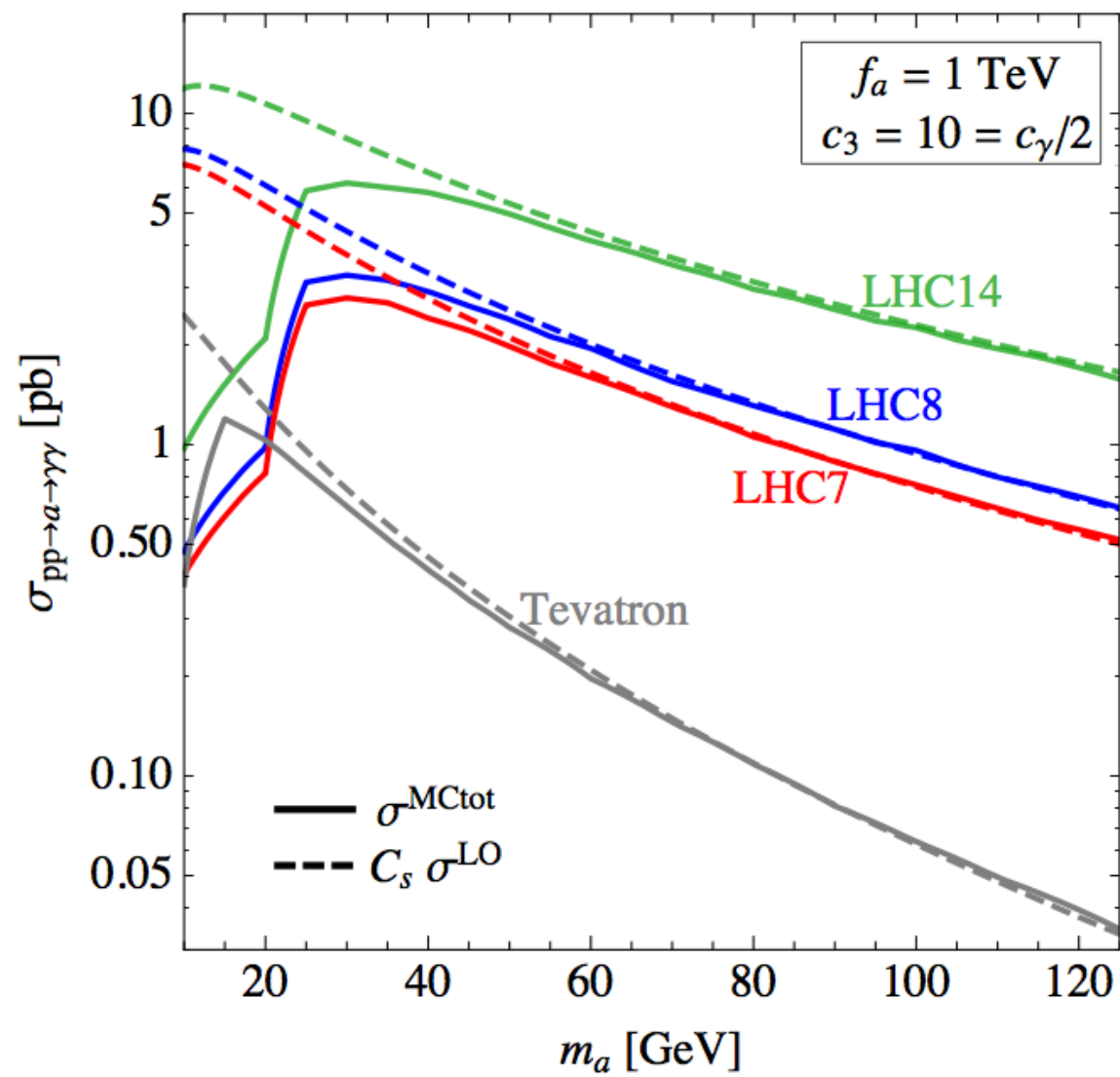
$$K_\sigma = 3.7 \quad \text{from ggHiggs v3.5}$$

Bonvini et al. 2013-2016

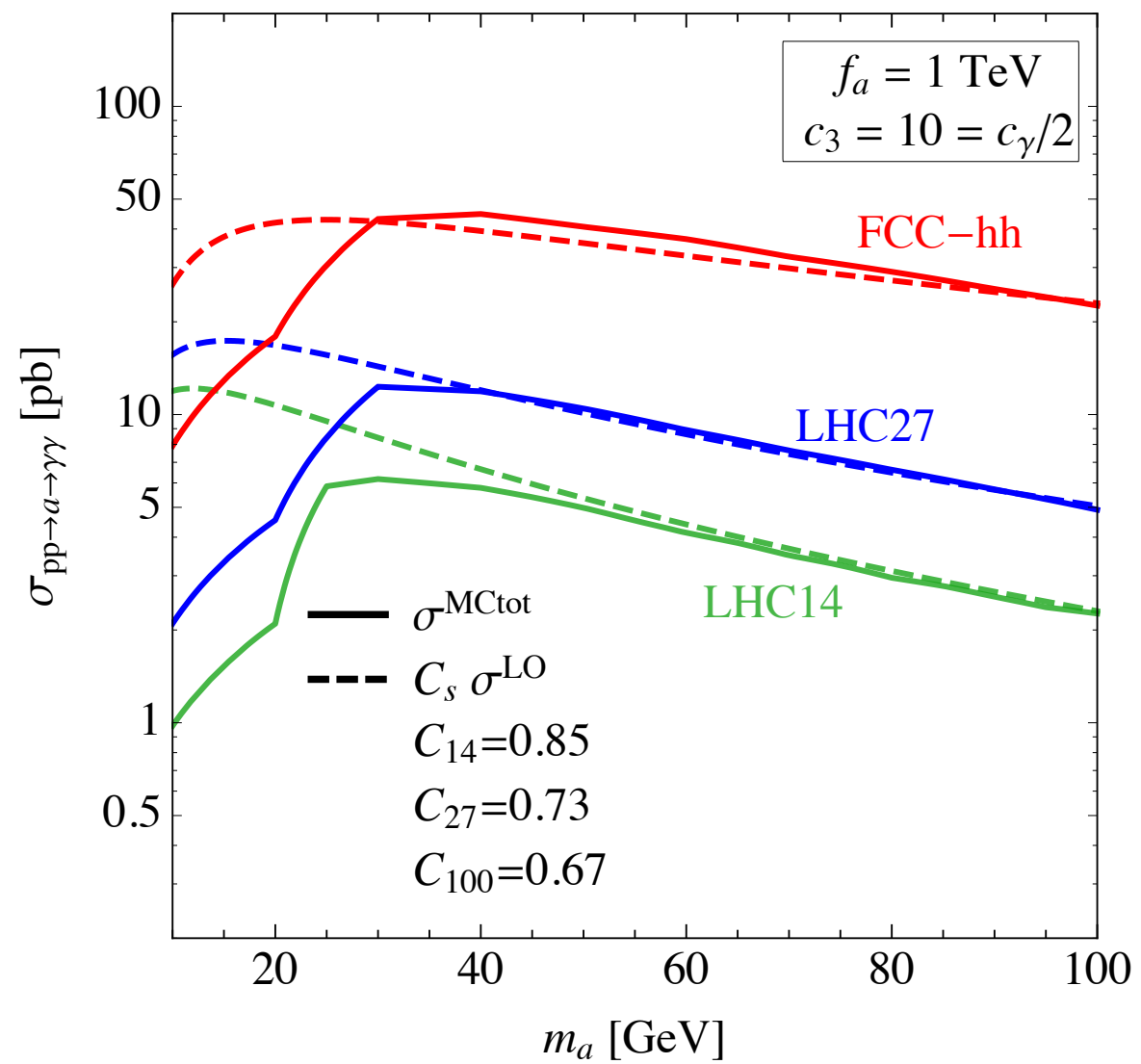
$$K_g = 2.1$$

m_a in GeV	10	20	30	40	50	60	70	80	90	100	110	120
ϵ_S for $\sigma_{7\text{TeV}}$ ATLAS [8]	0	0.008	0.022	0.040	0.137	0.293	0.409	0.465	0.486	0.533	0.619	0.637
ϵ_S for $\sigma_{7\text{TeV}}$ CMS [10]	0	0.002	0.010	0.020	0.030	0.058	0.156	0.319	0.424	0.499	0.532	0.570
ϵ_S for $\sigma_{8\text{TeV}}$ ATLAS [9]	0	0.0007	0.008	0.014	0.024	0.037	0.071	0.233	0.347	0.419	0.452	0.484
ϵ_S for $\sigma_{2\text{TeV}}$ CDF [45, 46]	0.001	0.007	0.026	0.143	0.212	0.241	0.276	0.275	0.283	0.3	0.319	0.327
ϵ_S for $\sigma_{2\text{TeV}}$ D0 [44]	0	0.002	0.008	0.018	0.114	0.169	0.208	0.21	0.217	0.234	0.244	0.252

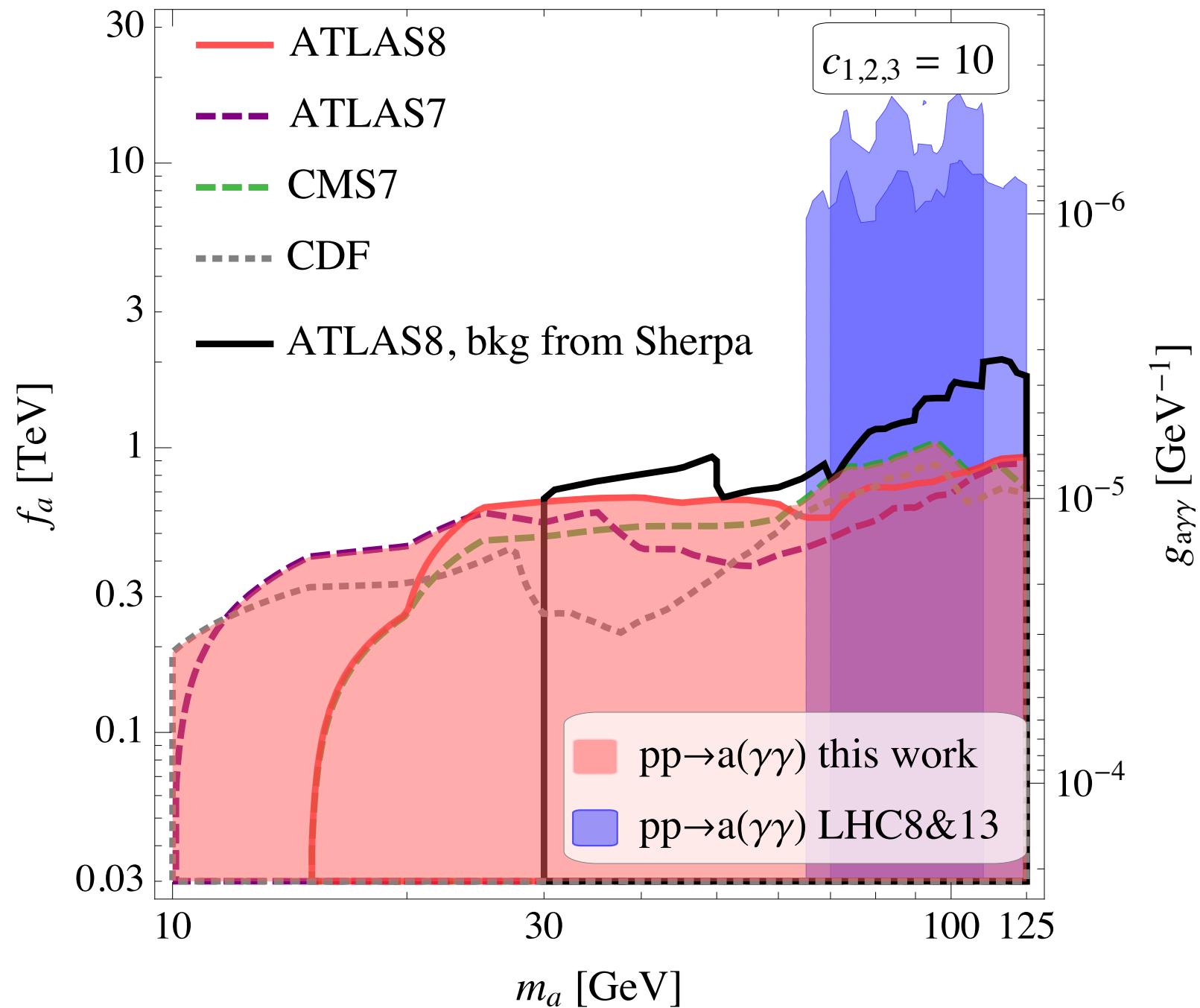
Validation



Validation



Interplay of LHC and Tevatron



FCC ee reach computed rescaling
LEP limits on $\text{BR}[Z \rightarrow \gamma a(jj)]$
and assuming 10^{12} Z bosons

If $aG\tilde{G}$ switched on
HL-LHC wins over FCC ee

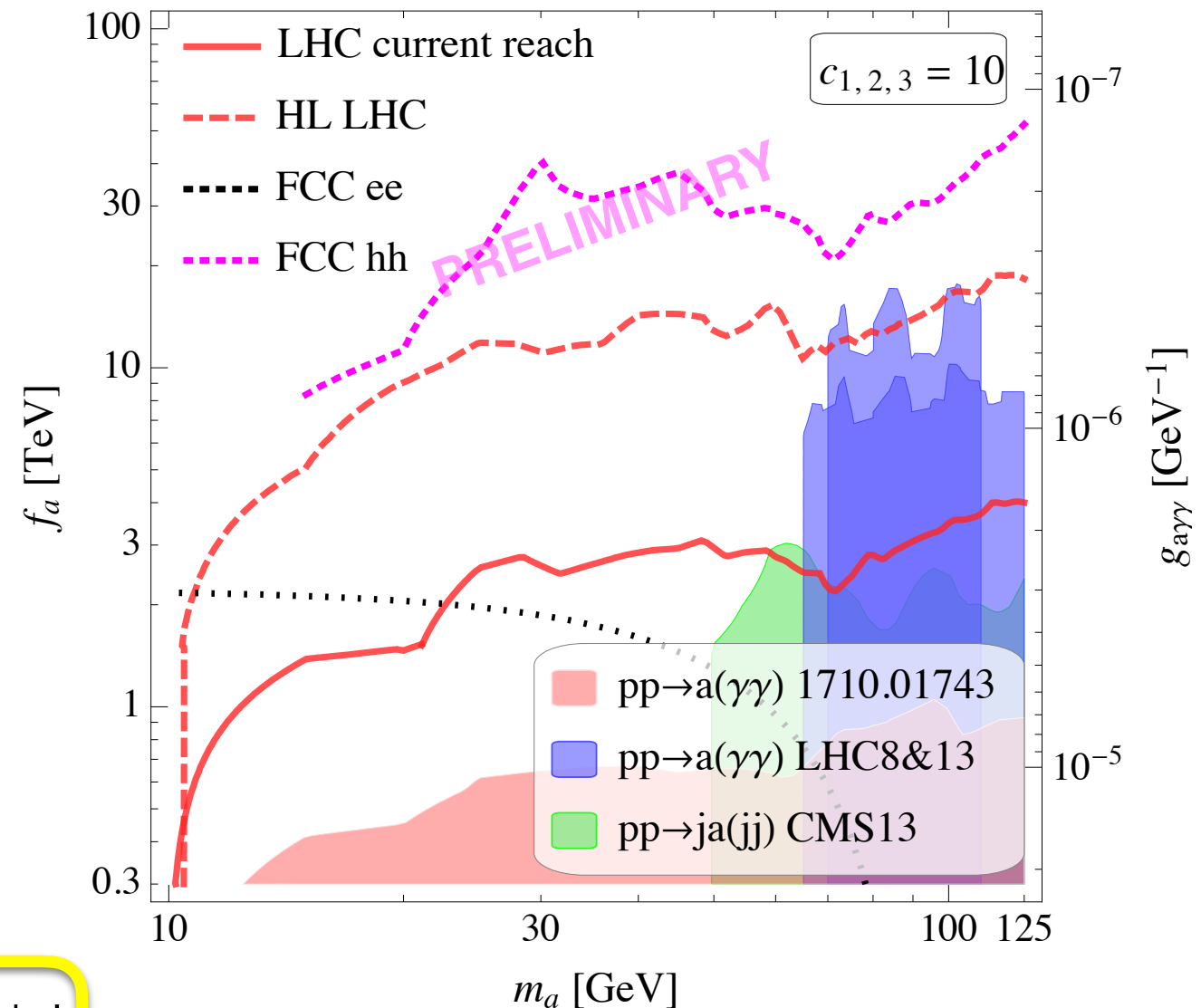
FCC hh reach computed like LHC14 one
[4., from simulations, Lumi=3 ab^{-1}]
NB. pT cuts as in ATLAS8 (30, 40 GeV)

Still, speculative even for FCC standards!

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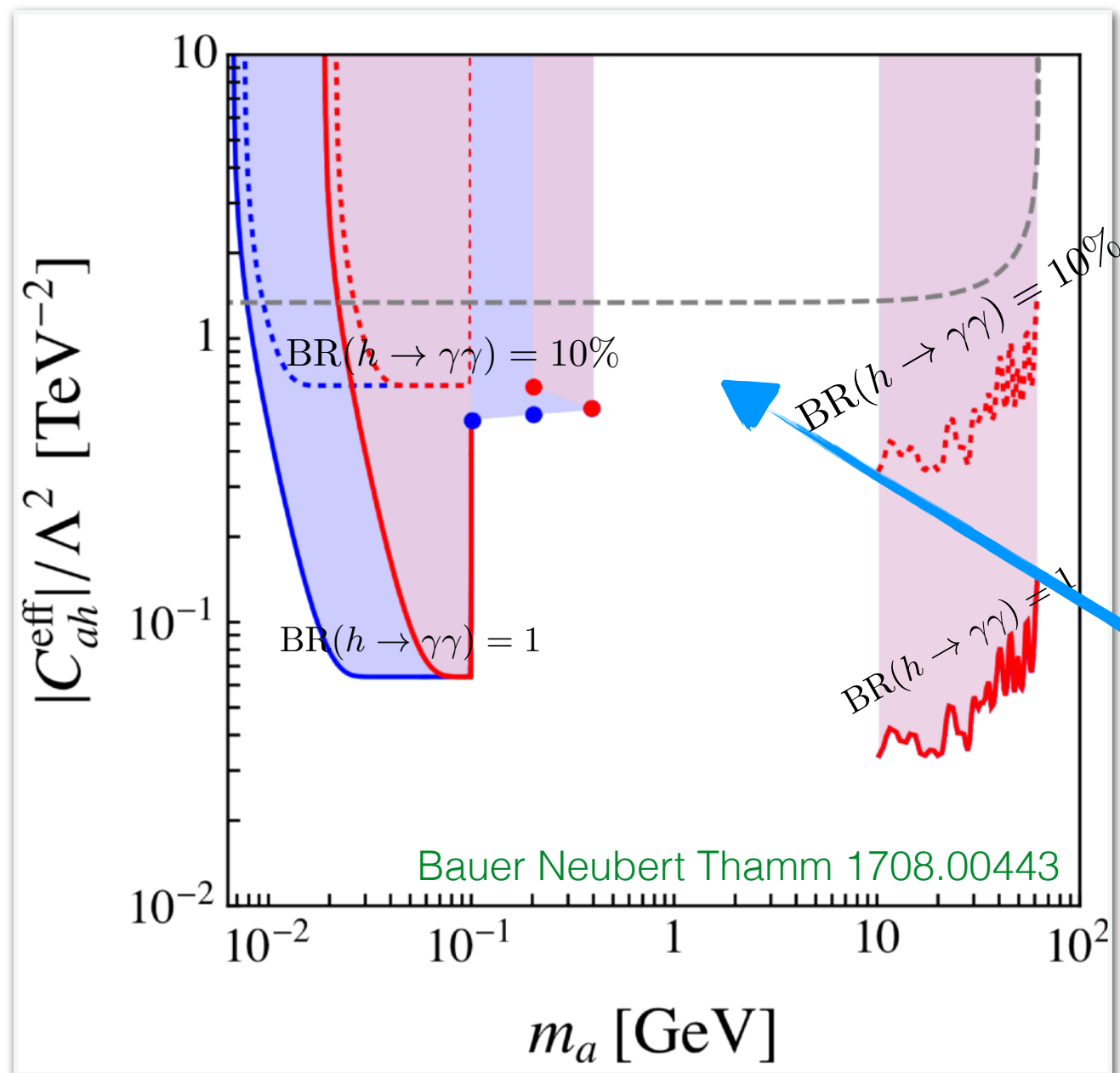
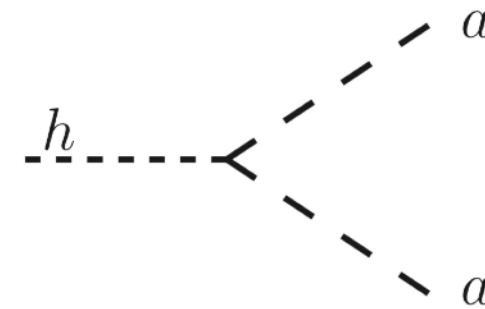
Thoughts in progress...



ALPs without gluons

LHC & other ALP couplings (no gluons)

A selected example: **Higgs decays** to 4 photons



$$\mathcal{L}_{\text{eff}}^{D \geq 6} = \frac{C_{ah}}{\Lambda^2} (\partial_\mu a)(\partial^\mu a) \phi^\dagger \phi$$

$$\mathcal{L}_{\text{eff}}^{D \leq 5} \ni e^2 C_{\gamma\gamma} \frac{a}{\Lambda} F_{\mu\nu} \tilde{F}^{\mu\nu}$$

Modified Diphoton Isolation
could help fill this gap?

For more on ALP at colliders see e.g.

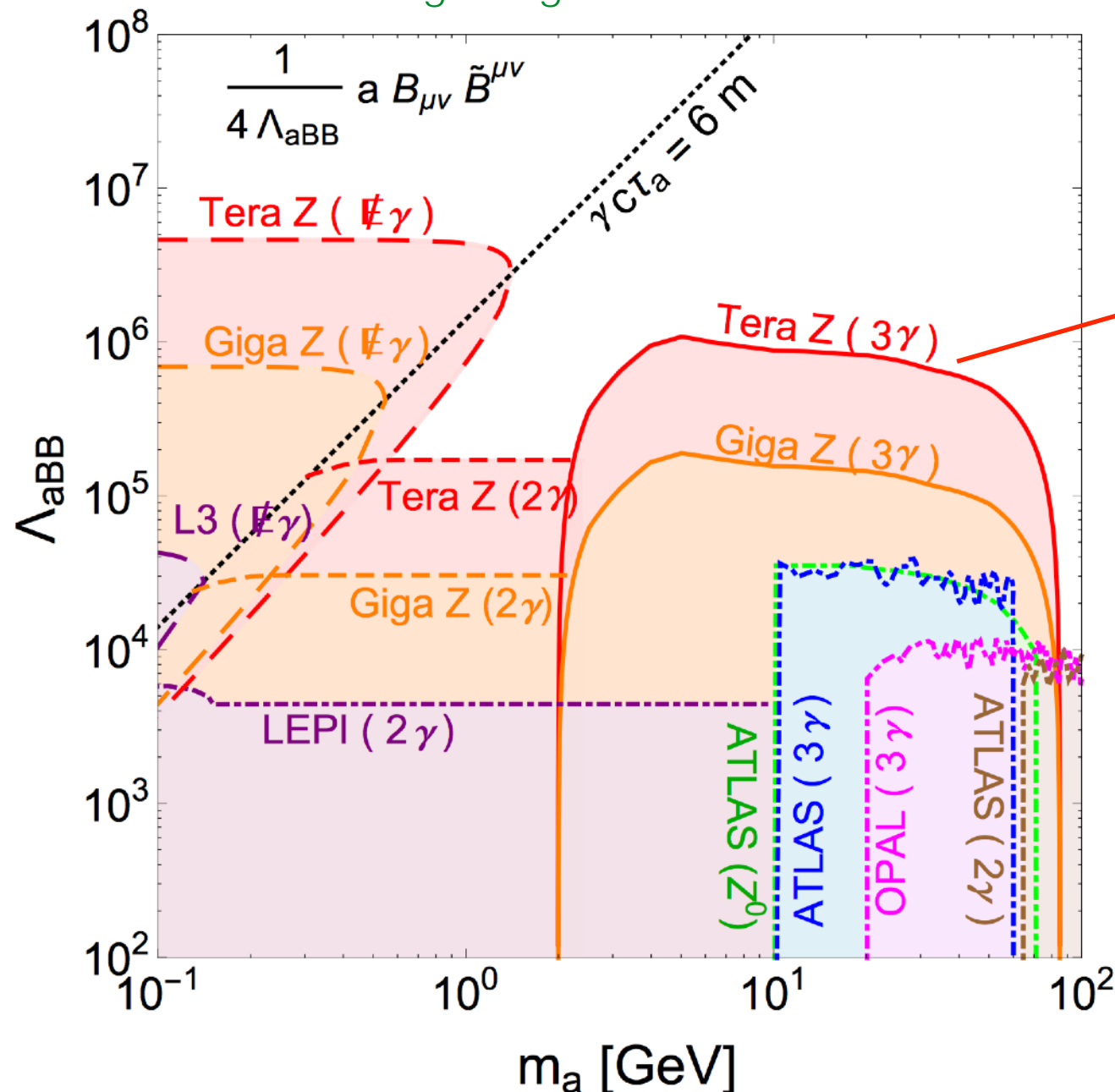
Mimasu Sanz 1409.4792, ,,,, Bauer+1808.10323

FCC-ee with no gluon coupling

$$\mathcal{L}_{\text{int}} = \frac{a}{4\pi f_a} \left[\cancel{\alpha_s c_2 G\tilde{G}} + \alpha_2 c_2 W\tilde{W} + \alpha_1 c_1 B\tilde{B} \right]$$

$$\alpha_1 = \frac{5}{3}\alpha_y$$

Liu Wang Wang Xue 1712.07237



To compare with previous slides:

$$\Lambda_{\text{aBB}} = \frac{\pi}{c_1 \alpha_1} f_a \simeq 20 f_a \frac{10}{c_1}$$

For Tera Z $\text{BR}[Z \rightarrow \gamma a(\gamma\gamma)] \lesssim 3 \times 10^{-9}$
[current LEP limit $\lesssim 5 \times 10^{-6}$]

FCC ee could reach $f_a \lesssim 100 \text{ TeV}$

FCC hh VBF ?

Associated production ?

Photon fusion ?

???