

# Neutrino - LHC working group

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Neutrinos: tiny masses

+ large mixings

$$\Delta m_{atm}^2 \simeq 3 \cdot 10^{-3} \text{ eV}^2$$

$$|\tan \theta_{atm}|^2 \simeq 1$$

$$\Delta m_{sol}^2 \simeq 7 \cdot 10^{-5} \text{ eV}^2$$

$$|\tan \theta_{sol}|^2 \simeq 0.4$$

$${}^3\text{H decay: } m_\nu \lesssim 2 \text{ eV}$$

$$|U_{e3}|^2 \lesssim 0.05$$

strong bounds for charged leptons

$$BR(\mu \rightarrow e\gamma) \lesssim 1.2 \cdot 10^{-11}$$

$$BR(\mu^- \rightarrow e^- e^+ e^-) \lesssim 10^{-12}$$

$$BR(\tau \rightarrow e\gamma) \lesssim 1.1 \cdot 10^{-7}$$

$$BR(\tau \rightarrow \mu\gamma) \lesssim 6.8 \cdot 10^{-8}$$

$$BR(\tau \rightarrow lll') \lesssim O(10^{-8}) \quad (l, l' = e, \mu)$$

$$|d_e| \lesssim 10^{-27} \text{ e cm}, \quad |d_\mu| \lesssim 1.5 \cdot 10^{-18} \text{ e cm}, \quad |d_\tau| \lesssim 1.5 \cdot 10^{-16} \text{ e cm}$$

SUSY contributions to anomalous magnetic moments

$$|\Delta a_e| \leq 10^{-12}, \quad 0 \leq \Delta a_\mu \leq 43 \cdot 10^{-10}, \quad |\Delta a_\tau| \leq 0.058$$

Neutrino masses due to

$$\frac{f}{\Lambda} (H_u L)(H_u L)$$

Seesaw I: postulates very heavy  $\nu_R$ :

$$m_\nu \simeq -(Y_\nu^T v_u) M_R^{-1} (Y_\nu v_u)$$

$$\hat{m}_\nu = U^T \cdot m_\nu \cdot U$$

expect:  $10^7 \text{ GeV} \lesssim M_{R_i} \lesssim 10^{14} \text{ GeV}$ .

much more parameters than observables in  $\nu$ -sector

\* P. Minkowski, *Phys. Lett. B* **67** (1977) 421; T. Yanagida, KEK-report 79-18 (1979);  
M. Gell-Mann, P. Ramond, R. Slansky, in *Supergravity*, North Holland (1979), p. 315;  
R.N. Mohapatra and G. Senjanovic, *Phys. Rev. Lett.* **44** 912 (1980).

## Superpotential

$$W = Y_e^{ji} \widehat{L}_i \widehat{H}_d \widehat{E}_j^c + Y_\nu^{ji} \widehat{L}_i \widehat{H}_u \widehat{N}_j^c + M_i \widehat{N}_i^c \widehat{N}_i^c$$

RGE running with mSUGRA boundaries:

$$(\Delta M_{\tilde{L}}^2)_{ij} = -\frac{1}{8\pi^2} (3m_0^2 + A_0^2) (Y_\nu^\dagger L Y_\nu)_{ij}$$

$$(\Delta A_l)_{ij} = -\frac{3}{8\pi^2} A_0 Y_{l_i} (Y_\nu^\dagger L Y_\nu)_{ij}$$

$$(\Delta M_{\tilde{E}}^2)_{ij} = 0$$

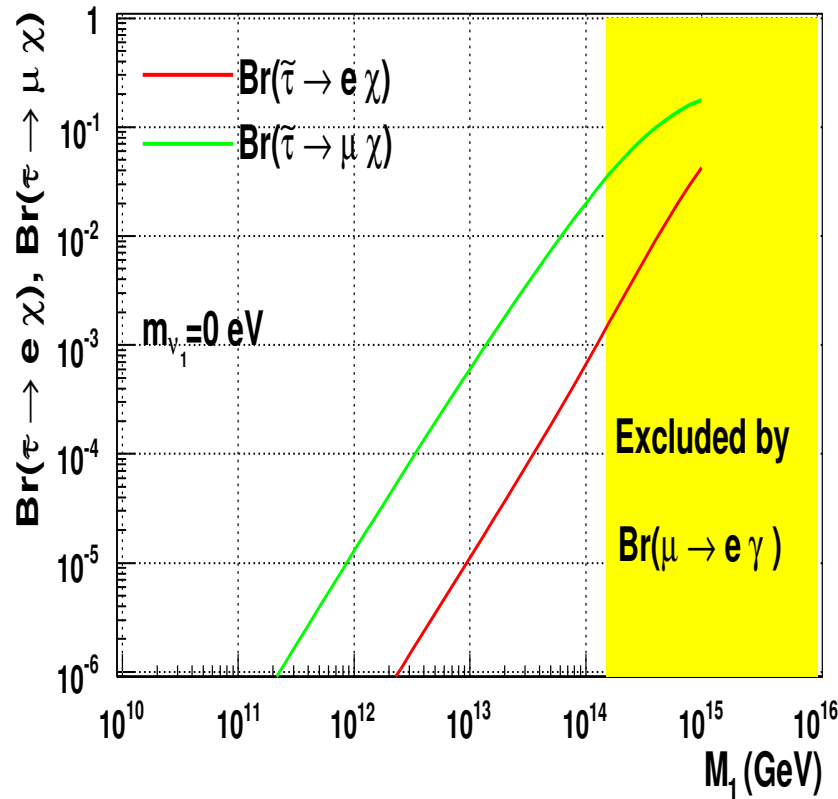
$$L_{kl} = \log\left(\frac{M_X}{M_{R_k}}\right) \delta_{kl}$$

## Consequence

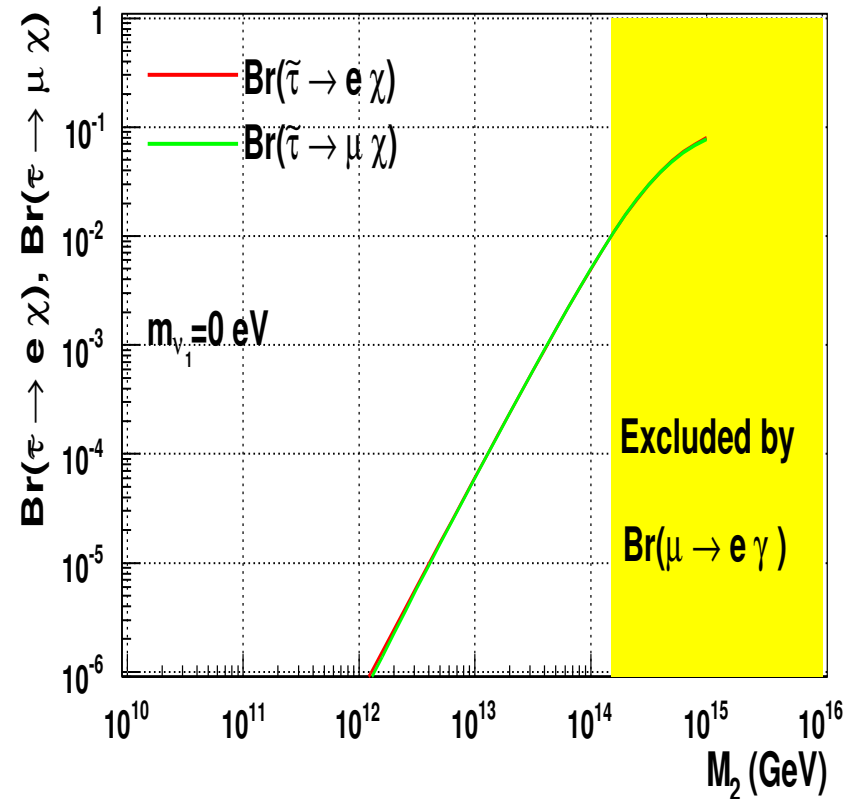
$$l_j \rightarrow l_i \gamma, l_i l_k^+ l_r^-$$

$$\tilde{l}_j \rightarrow l_i \tilde{\chi}_s^0$$

$$\tilde{\chi}_s^0 \rightarrow l_i \tilde{l}_k$$



degenerate  $\nu_R$



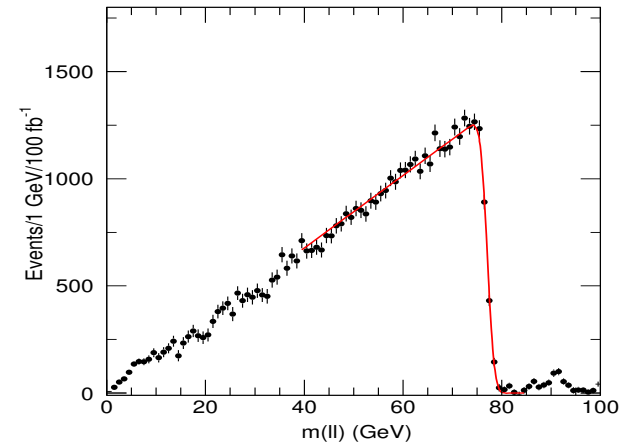
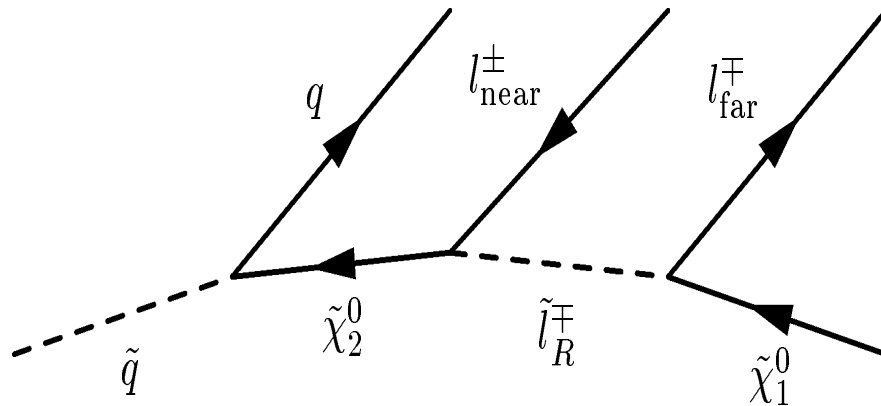
hierarchical  $\nu_R$

$(M_1 = M_3 = 10^{10} \text{ GeV})$

SPS3 ( $M_0 = 90 \text{ GeV}$ ,  $M_{1/2} = 400 \text{ GeV}$ ,  $A_0 = 0 \text{ GeV}$ ,  $\tan \beta = 10$ ,  $\mu > 0$ )

M. Hirsch et al. Phys. Rev. D 78 (2008) 013006

G. Polesello

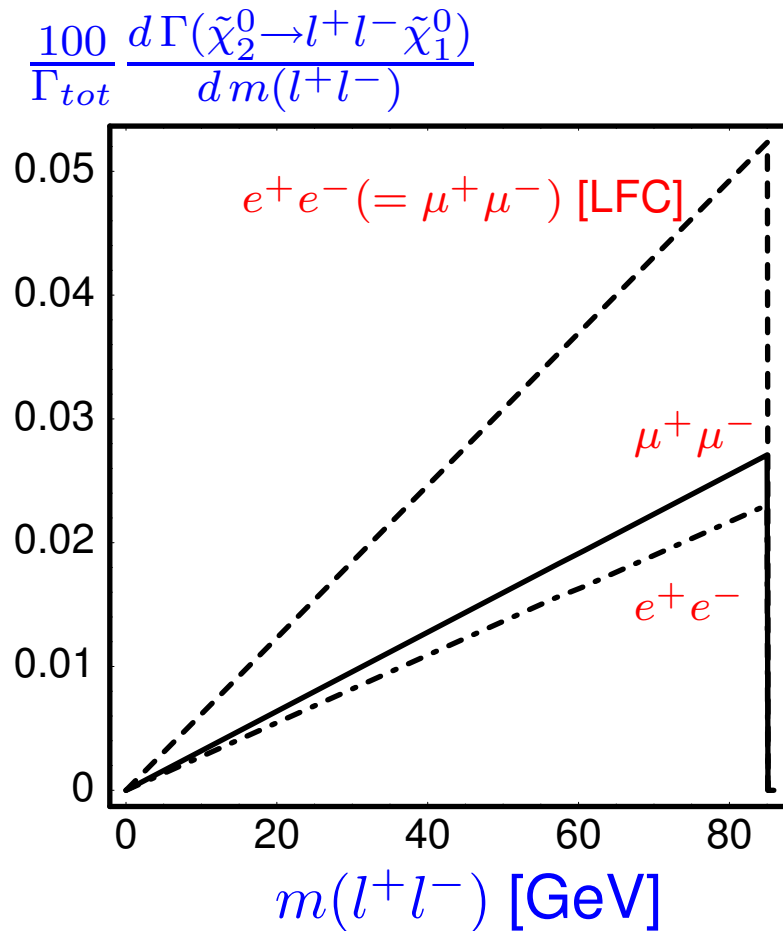
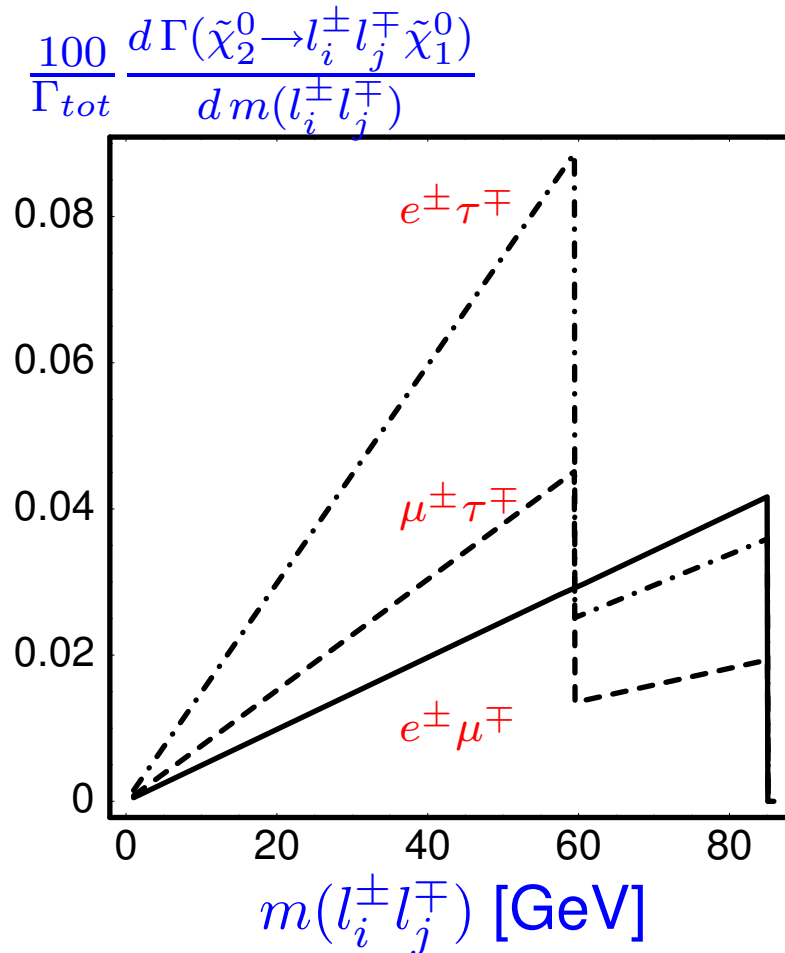


5 kinematical observables depending on 4 SUSY masses

e.g.:  $m(ll) = 77.02 \pm 0.05 \pm 0.08$   
 $\Rightarrow$  mass determination within 2-5%

For background suppression

$$N(e^+e^-) + N(\mu^+\mu^-) - N(e^+\mu^-) - N(\mu^+e^-)$$



A. Bartl et al., Eur. Phys. J. C 46 (2006) 783

R-parity:  $(-1)^{3(B-L)+2S}$

bilinear R-parity violation:

$$W = W_{MSSM} + \epsilon_i \hat{L}_i \hat{H}_u$$

⇒ mixings between SM and SUSY particles

Gravitino as dark matter\*

generic prediction of GMSB : light Gravitino LSP

NLSP:  $\tilde{\chi}_1^0$  oder  $\tilde{l}_R$  ( $l = e, \mu, \tau$ )

\*S. Borgani, A. Masiero, M. Yamaguchi, PLB**386** (1996) 189

F. Takayama and M. Yamaguchi, PLB **485** (2000) 388

M. Hirsch, W. P., D. Restrepo, JHEP **0503**, 062 (2005)



## Neutrino physics\*:

- neutrino masses via  $\nu$ - $\tilde{\chi}_i^0$  mixing
- neutrino mixing angles in terms of R-parity violating couplings

## Neutralino decays:

- dominant R-parity violating decays:  $\tilde{\chi}_1^0 \rightarrow W^\pm l_i^\mp$ ,  $\tilde{\chi}_1^0 \rightarrow Z\nu_i$ ,  $\tilde{\chi}_1^0 \rightarrow \nu\tau^+ l_i^-$
- R-parity conserving decay:  $\Gamma(\tilde{\chi}_1^0 \rightarrow \tilde{G}\gamma) \propto \left(\frac{m_{\tilde{\chi}_1^0}}{100 \text{ GeV}}\right)^5 \left(\frac{100 \text{ eV}}{m_{3/2}}\right)^2$
- $\Gamma(\tilde{\chi}_1^0) \simeq (10^{-4}-10^{-2}) \text{ eV} \Rightarrow$  decay length of  $O(10 \mu\text{m}) - O(1 \text{ mm})$

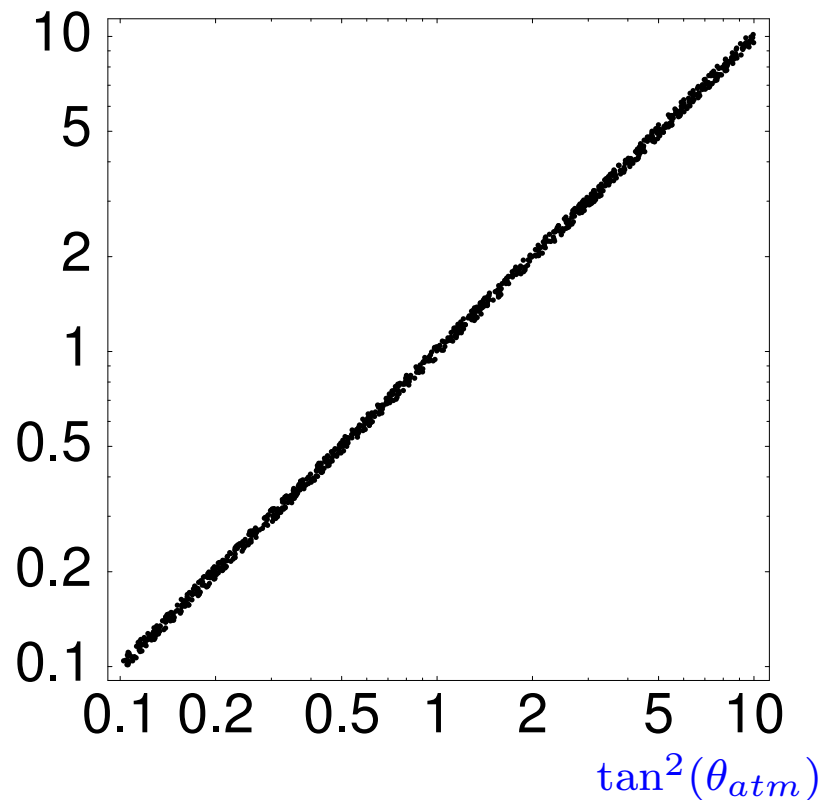
## Gravitino:

- $\tilde{G}$  eventually decays into  $\nu\gamma$   
but:  $\tau(\tilde{G}) \sim O(10^{30})$  · life time of the universe

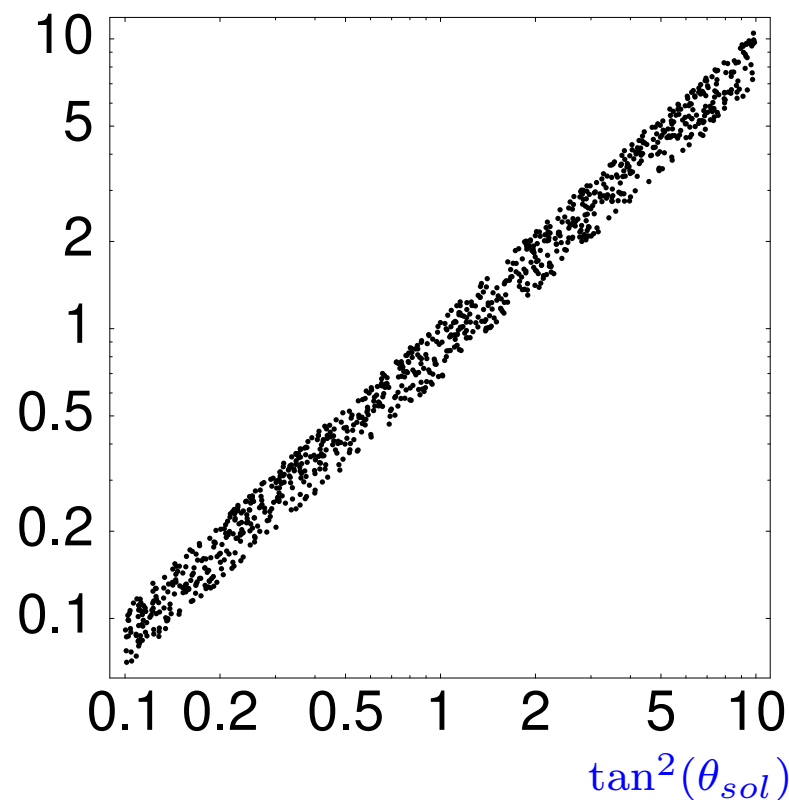
\* M. Hirsch, et al. Phys. Rev. D **62**, 113008 (2000)

## Correlations

$$\text{BR}(\tilde{\chi}_1^0 \rightarrow W\mu) / \text{BR}(\tilde{\chi}_1^0 \rightarrow W\tau)$$

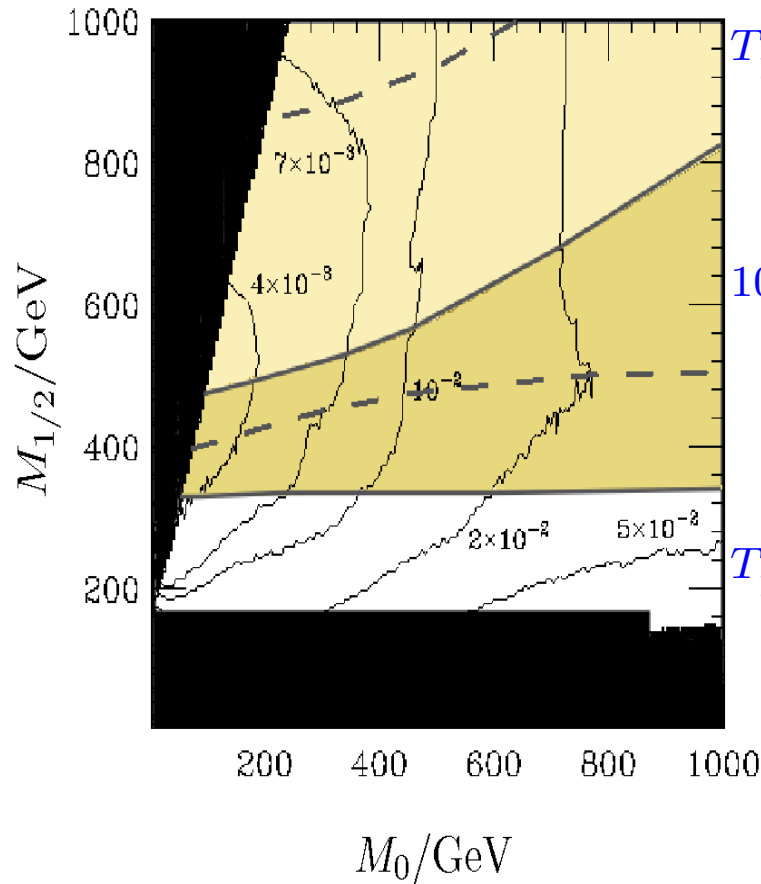
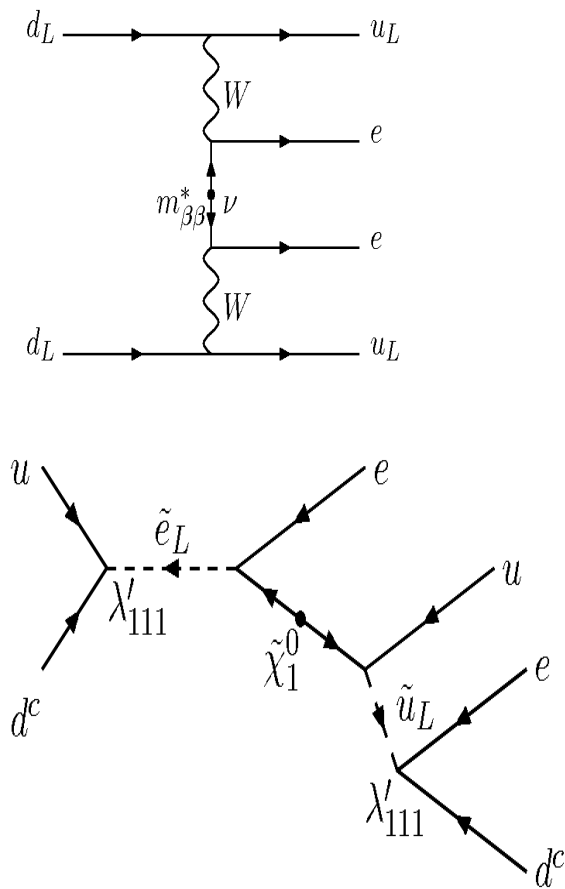


$$\text{BR}(\tilde{\chi}_1^0 \rightarrow \nu e \tau) / \text{BR}(\tilde{\chi}_1^0 \rightarrow \nu \mu \tau)$$



W. P. et al., Phys. Rev. D **63**, 115004 (2001)

regions for  $5\sigma$  discovery



$$T_{1/2}^{0\nu\beta\beta} > 1 \times 10^{27} \text{ yrs}$$

$$100 > T_{1/2}^{0\nu\beta\beta} / 10^{25} \text{ yrs} > 1.9$$

$$T_{1/2}^{0\nu\beta\beta} < 1.9 \cdot 10^{25} \text{ yrs}$$

$L = 10 \text{ fb}^{-1}$  at 14 TeV,  $A_0 = 0$ ,  $\tan \beta = 10$ ,  $\mu > 0$

B. C. Allanach, C. H. Kom, H. Päs, arXiv:0903.0347 (hep-ph)

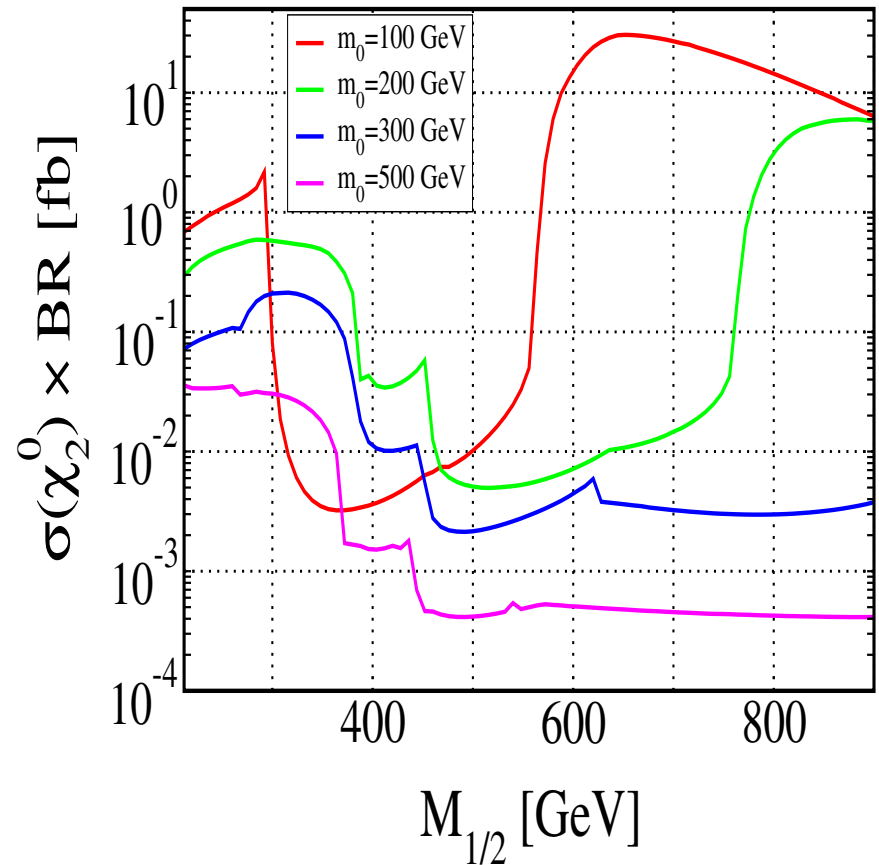
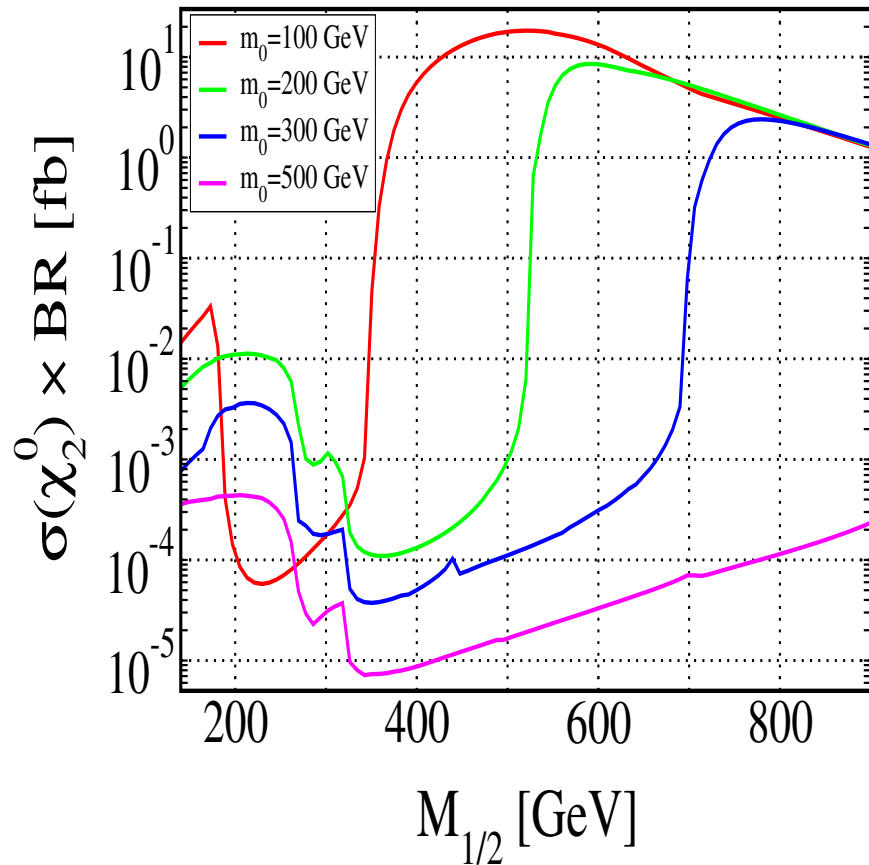
## Aachen, Dortmund, Würzburg

- rare lepton decays at LHC, such as  $\tau \rightarrow \mu^- \mu^- \mu^+$
- LFV signals at LHC, e.g. cascades containing  $\tilde{\chi}_2^0 \rightarrow \mu \tau \tilde{\chi}_1^0$
- predictions of  $\nu$ -mass models for rare lepton decays and properties of new particles (not only SUSY)
- to which extent does a combination of LHC data with data from other experiments allow one to pin down the underlying model?

## Kick-off Workshop, 24./25. Nov. in Würzburg

- <https://indico.desy.de/conferenceDisplay.py?confId=2403>
- [www.terascale.de](http://www.terascale.de) → Schools and workshops

Meeting on Friday at 14:00, 'Großer Gästespeiseraum'



$$\sigma(pp \rightarrow \tilde{\chi}_2^0) \times BR(\chi_2^0 \rightarrow \sum_{i,j} \tilde{l}_i l_j \rightarrow \mu^\pm \tau^\mp \tilde{\chi}_1^0)$$

$$A_0 = 0, \tan \beta = 10, \mu > 0 \text{ (Seesaw II: } \lambda_1 = 0.02, \lambda_2 = 0.5)$$

J.N. Esteves et al., arXiv:0903.1408