# The evolution of APFEL: APEL APEL The evolution of APFEL:

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#### Motivation

- Since its born, APFEL has undergone a large number of developments:
  - FONLL structure functions,
  - NLO QED evolution,
  - lepton PDFs,
  - Scale variations,
  - intrinsic charm,
  - displaced thresholds,
  - MS masses,
  - $\bullet$  small-x resummation,
  - interface to the NNPDF code,
  - **...**
- Way beyond the purposes for which it was conceived:
  - very large memory footprint,
  - non-optimal "convenience" solutions for the new modules,
  - hard to maintain.
- APFEL is written in FORTRAN77 that is not suitable for large projects:
  - lack of modularity,
  - non-optimal (built-in) memory management.
- Compelling reasons to rewrite APFEL keeping in mind its applications.

- Concerning the language, C++ was a somewhat natural choice:
  - modularity ensured by the object-oriented nature,
  - **dynamical** allocation of the **memory**,
  - used for the new-generation tools (e.g. LHAPDF) and thus easier **interface**,
  - powerful features coming with the C++11/17 standard.
- The code **design** was driven by a profound rethinking of the strategy:
  - the main application field is collinear/TMD factorisation.
  - In this context, many relevant quantities are computed as **convolutions**:

$$M(x) = \int_{x}^{1} \frac{dy}{y} O(y) d\left(\frac{x}{y}\right) = \int_{x}^{1} \frac{dz}{z} O\left(\frac{x}{z}\right) d(z) \equiv O(x) \otimes d(x)$$

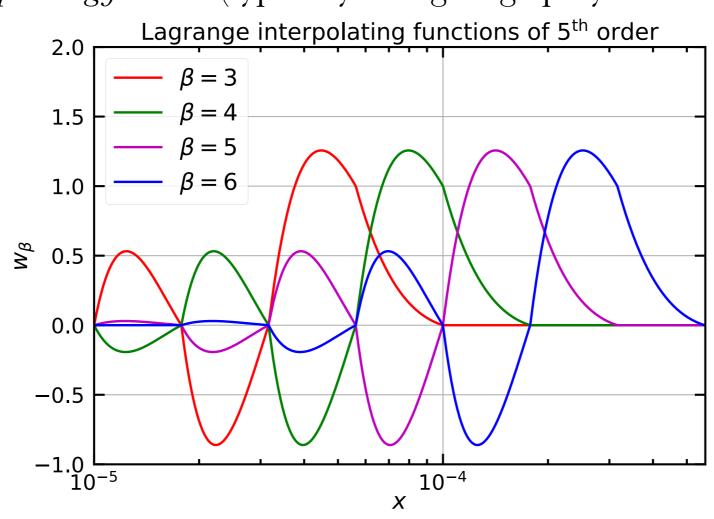
**operator** *O*: typically a complicated object slow to compute: *e.g.* a **perturbative** hard cross section. **d**: typically a fast-to-access function: *e.g.* a **non-perturbative** PDF or a FF.

- $\bullet$  Adopt the **x-space** (as opposed to *N*-space) formalism:
  - $\bullet$  most of the results are available in x-space,
  - no restriction on the parameterisations.
- The purpose is to make convolutions **fast**.

- Define an **interpolation grid** in x with N+1 nodes  $g \equiv \{z_0, \ldots, z_N\}$
- $\bullet$  Use the interpolation formula for the **distribution** d:

$$\frac{d}{d}(z) = \sum_{\beta=0}^{N} w_{\beta}(z) \frac{d}{\beta} \quad \text{with} \quad \frac{d}{\beta} = \frac{d}{d}(x_{\beta})$$

•  $w_{\beta}(z)$  interpolating function (typically a Lagrange polynomial of some degree n).



• piecewise function different from zero over n+1 intervals around  $\beta$ . Zero elsewhere. Hard to integrate.

• Compute the integral of the **operator** O with the interp. functions:

$$O_{\alpha\beta} \equiv \int_{x_{\alpha}}^{1} \frac{dy}{y} O(y) w_{\beta} \left(\frac{x_{\alpha}}{y}\right)$$

such that:

$$M_{\alpha} \equiv \sum_{\beta=0}^{N} O_{\alpha\beta} d_{\beta}$$
 with  $M(x) = \sum_{\alpha=0}^{N} w_{\alpha}(x) M_{\alpha}$ 

- This reduces convolutions to multiplications between a matrices and vectors: **linear algebra**.
- Therefore, the **three main ingredients** of of APFEL++ are:
  - 1. the **interpolation grid** g along with the interpolating functions,
  - 2. the **distribution**  $d_{\beta}$ ,
  - 3. the **operator**  $O_{\alpha\beta}$ .
- They can be **encapsulated** in **C++ objects** to compute convolutions.

- An additional complication is given by the **flavour structure**:
  - distributions are vectors in flavour space,
  - operators are either matrices or vectors in the flavour space.

$$\frac{df_{i\alpha}}{d\ln\mu^2} = \sum_{j,\beta} P_{\alpha\beta}^{ij} f_{j\beta}$$

$$F_{lpha} = \sum_{j,eta} C^j_{lphaeta} f_{jeta}$$

#### **DGLAP** equations

DIS structure functions

- Matrices in flavour space are often sparse.
- Define **sets of objects** with a **flavour map**:
  - associate one operator to one distribution, e.g.:

$$P_{qq}, P_{gq} \rightarrow \Sigma$$
 $P_{qg}, P_{gg} \rightarrow g$ 
 $P^v \rightarrow V$ 
 $P^+ \rightarrow T_{3,8,15,24,35}$ 
 $P^- \rightarrow V_{3,8,15,24,35}$ 

- the same operator can be assigned to more than a distribution and viceversa.
- avoid multiplications by zero.

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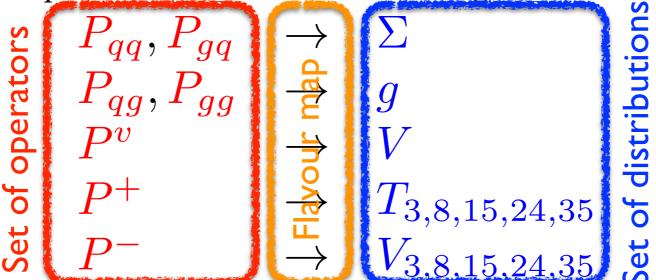
$$\frac{df_{i\alpha}}{d\ln\mu^2} = \sum_{j,\beta} P_{\alpha\beta}^{ij} f_{j\beta}$$

$$m{F}_{lpha} = \sum_{j,eta} m{C}_{lphaeta}^{j} m{f}_{jeta}$$

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- the same operator can be assigned to more than a distribution and viceversa.
- avoid multiplications by zero.

- Define **multiplication** between sets of distributions and operators:
  - overload multiplication operator in C++.
  - Making convolutions with a flavour structure becomes very **easy**.

```
//_____
Set<Distribution> Dglap::Derivative(int const& nf, double const& t, Set<Distribution> const& f) const
{
   return _SplittingFunctions(nf, exp(t/2)) * f;
}
```

- Define multiplication between sets of distributions and operators:
  - overload multiplication operator in C++.
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Overloaded multiplication: takes care of convolutions and flavour structure

- The flavour structure is completely defined by the flavour map:
  - same procedure for convolutions in any flavour basis,
  - sets of operators and distributions multiplied only if they **share** the same map,
  - $\bullet$  easy to account for  $n_f$  dependence.

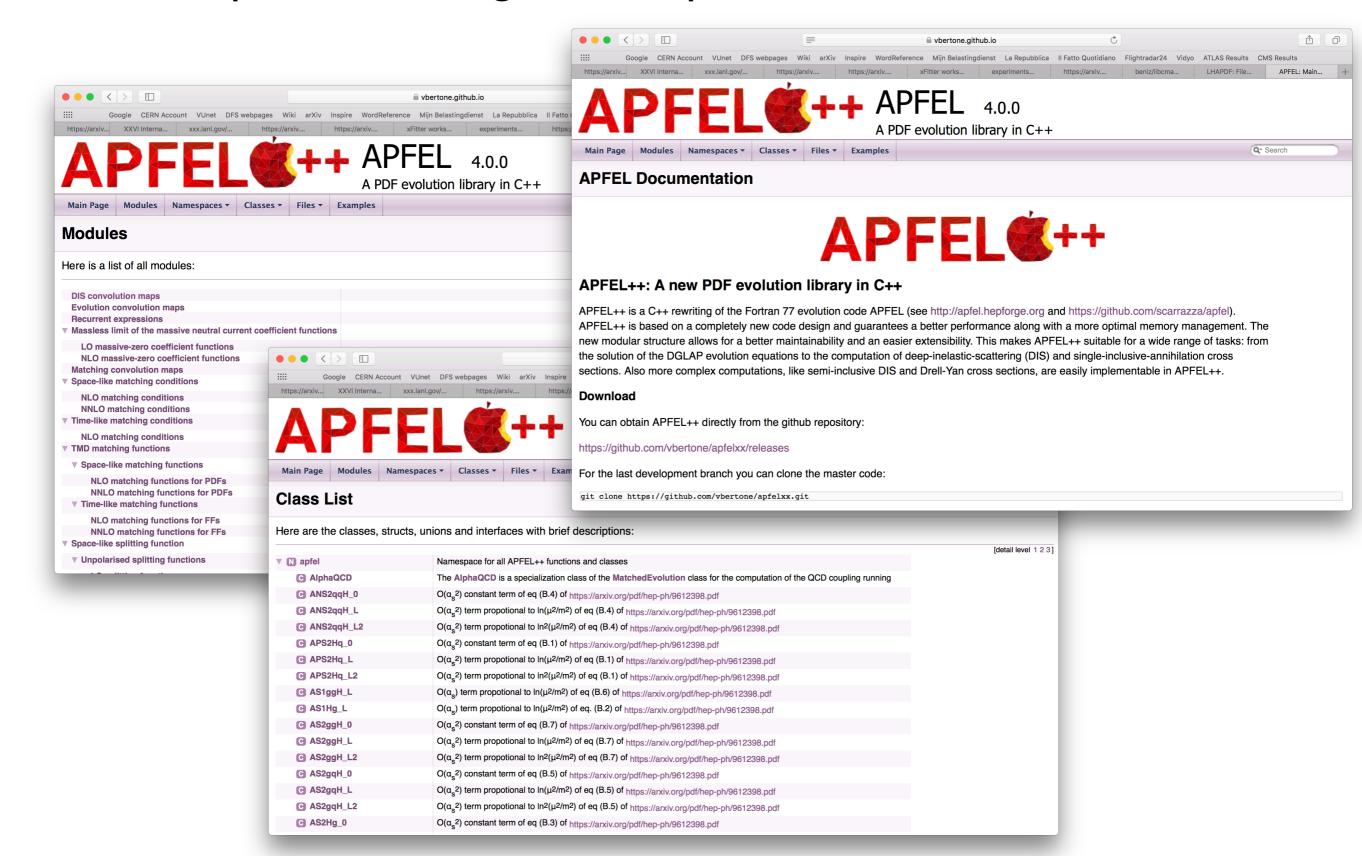
• Use (e.g.) 4th order Runge-Kutta to solve systems of **ordinary** differential equations:

```
\begin{cases} \frac{d\mathbf{y}}{dt} = \mathbf{F}(t, \mathbf{y}) \\ \mathbf{y}(t_0) = \mathbf{y}_0 \end{cases}
                                                   Template function...
      template<class U>
                                                    ... that returns a std::function...
      function<U(double const&, U const&, double const&)>
      rk4(function<U(double const& t, U const& Obj)> const&(f) ... of a std::function
) -> U{ return
                                U const& dy4
```

Very same function used to solve **both** the DGLAP and the  $\alpha_s$  RGE.

#### Doxygen documentation

https://vbertone.github.io/apfelxx/html/index.html



# Convoluting operators

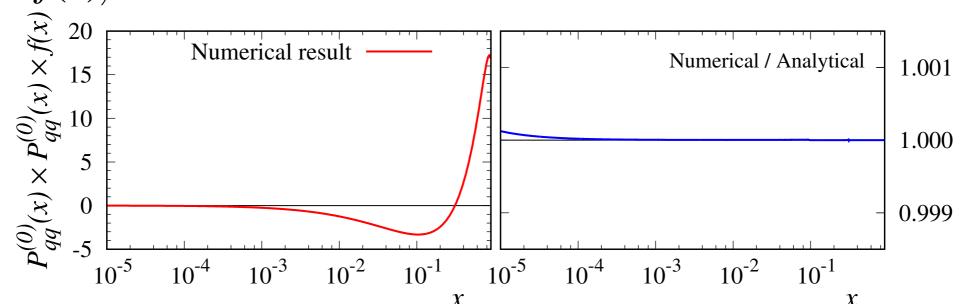
- An operation that is often needed is the convolution between operators:
  - involved in the computation of factorisation scale variations,
  - computation of the PDF evolution operator.

$$M(x) = \underbrace{O^{(1)}(x) \otimes O^{(2)}(x)}_{O(x)} \otimes d(x) \rightarrow M_{\alpha} = \sum_{\beta} \sum_{\gamma} O^{(1)}_{\alpha\gamma} O^{(2)}_{\gamma\beta} d_{\beta}$$

Consider for example:

$$P_{qq}^{(0)}(x) = \left(\frac{1+x^2}{1-x}\right)_{+} \quad \text{such that} \quad P_{qq}^{(0)}(x) \otimes P_{qq}^{(0)}(x) = \left(\frac{4(x^2+1)\ln(1-x)+x^2+5}{1-x}\right)_{+} \\ - \frac{(3x^2+1)\ln x}{1-x} - 4 + \left(\frac{9}{4} - \frac{2\pi^2}{3}\right)\delta(1-x)$$

• Compare the numerical convolution with the analytic result (using a test function f(x)):



#### **Evolution operator**

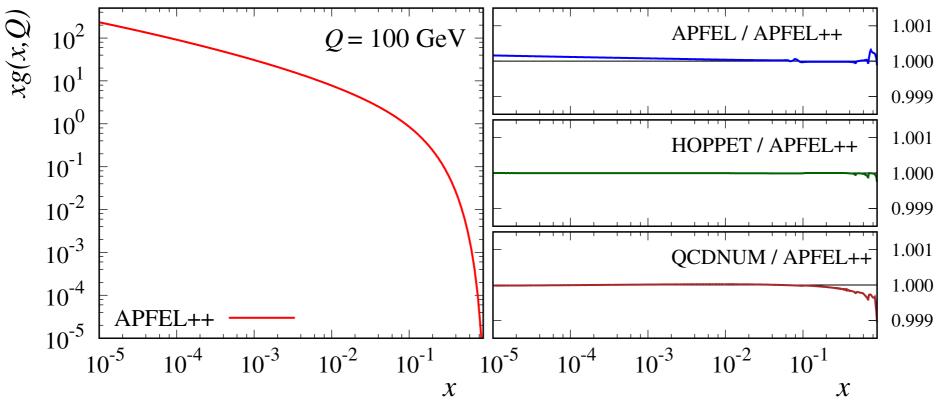
• The DGLAP can be written in terms the **evolution operator**:

$$\begin{cases} \frac{d}{d \ln \mu^2} \Gamma_{\alpha\beta}^{ij}(\mu_0, \mu) = \sum_{k,\gamma} P_{\alpha\gamma}^{ik}(\mu) \Gamma_{\gamma\beta}^{kj}(\mu_0, \mu) \\ \Gamma_{\alpha\beta}^{ij}(\mu_0, \mu_0) = \delta_{ij} \delta_{\alpha\beta} \end{cases}$$

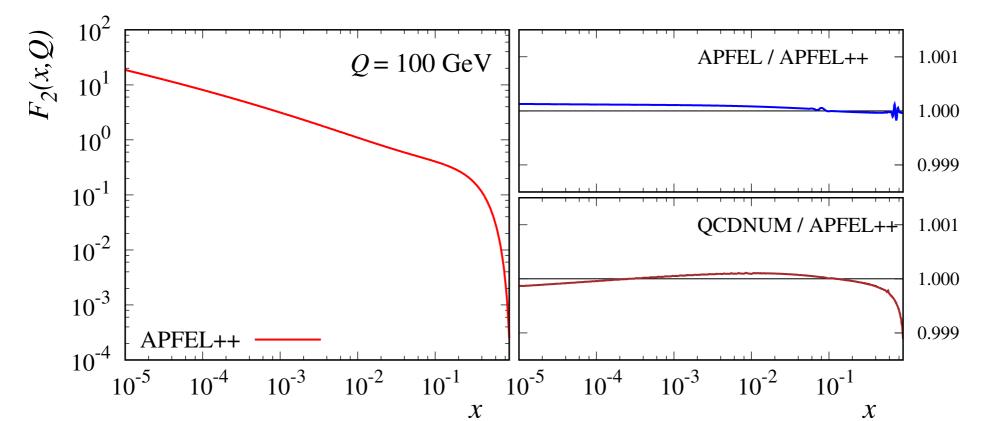
- The evolution operator can be used to evolve any initial scale PDF:
  - harder to compute than evolving PDFs directly,
  - it has to be computed **only once** (for each  $\mu_0$  and  $\mu$ ).
- This object is used for the construction of the **APFELgrid** tables:
  - extremely hard to compute in APFEL (Fortran),
  - very easy with APFEL++.

### Applications

• DGLAP evolution at NNLO:



• DIS structure functions at NNLO:



# PDF evolution performance

Comparison between different codes:

NNLO QCD evolution ~200 points in x ~50 points in Q	Initialisation [s]	Interpolate PDFs 10 <sup>6</sup> times [s]
APFEL++	<b>0.4</b> (0.27 in. + 0.13 tab.)	0.6
APFEL	2.4	1.9
HOPPET	0.4	1.3
QCDNUM	8.7	1.3

• Compare APFEL++ to LHAPDF when interpolating a **std::map**:

PDF set: NNPDF31_nlo_as_0118	Interpolate a PDF map 10 <sup>5</sup> times [s]
APFEL++	0.7
LHAPDF	0.5

# APFEL through LHAPDF

- Idea: delegate LHAPDF to interpolate over the  $(x, Q^2)$  grid:
  - more performing,
  - standardise the access to PDFs.

```
// APFEL++ default EvolutionSetup object
apfel::EvolutionSetup es{};

// Feed it to the initialisation class of APFEL++
apfel::InitialiseEvolution ev{es, true};

// Construct pointer to LHAPDF::PDF object
LHAPDF::PDF* distAP = mkPDF(ev);

// Call PDFs
const std::map<int, double> PDFmap = distAP->xfxQ(x, mu);
```

```
APFEL++ evolution:
AlphaQCD(0) = 1.1638e-01
          u-ubar
                      d-dbar
                               2(ubr+dbr)
                                             c+cbar
                                                         aluon
1.0e-05 5.0597e-03 2.9000e-03 2.9868e+01
                                          1.3409e+01
                                                      1.8825e+02
1.0e-04 2.0510e-02 1.1486e-02 1.4191e+01 6.0136e+00
                                                     8.0275e+01
1.0e-03 7.7062e-02 4.3014e-02 6.0966e+00
                                          2.3549e+00
                                                      2.9215e+01
1.0e-02 2.4305e-01 1.3669e-01 2.2041e+00
                                          7.1400e-01
                                                     7.9998e+00
1.0e-01 5.1440e-01 2.4974e-01 3.9254e-01 7.5071e-02 8.9562e-01
3.0e-01 3.3154e-01 1.1785e-01 3.5867e-02 6.7898e-03 9.6761e-02
5.0e-01 1.0288e-01 2.6959e-02 2.4053e-03
                                          5.4369e-04 9.5518e-03
       1.4110e-02 2.6031e-03 5.1163e-05
                                          1.3984e-05
9.0e-01 1.8836e-04 2.4841e-05 2.1042e-08 8.2414e-09 7.3548e-07
```

```
LHAPDF (tabulated) evolution:
AlphaQCD(0) = 1.1638e-01
                               2(ubr+dbr)
                      d-dbar
                                             c+cbar
                                                         gluon
          u-ubar
1.0e-05 5.0549e-03 2.9037e-03 2.9871e+01 1.3411e+01
1.0e-04 2.0510e-02 1.1491e-02 1.4193e+01 6.0140e+00
                                          2.3550e+00
1.0e-03 7.7066e-02 4.3017e-02 6.0971e+00
1.0e-02 2.4306e-01 1.3670e-01 2.2043e+00
                                          7.1472e-01
                                                      8.0004e+00
1.0e-01 5.1438e-01 2.4974e-01 3.9226e-01 7.5312e-02
3.0e-01 3.3153e-01 1.1785e-01 3.5865e-02 6.8228e-03
                                                     9.6755e-02
5.0e-01 1.0288e-01 2.6957e-02 2.4027e-03
                                          5.4586e-04
7.0e-01 1.4109e-02 2.6030e-03 5.1163e-05 1.4071e-05
9.0e-01 1.8834e-04 2.4838e-05 2.1027e-08 8.3325e-09
```

#### Old functionalities

- The FORTRAN version of APFEL implements a **very large** number of functionalities.
- in Currently working to implement all of them also in APFEL++.
- Missing functionalities in APFEL++ to be implemented:
  - QED corrections,
  - intrinsic charm,
  - MS masses,
  - small-x resummation (need to interface APFEL++ to HELL),
  - scale variations,
  - "minor" functionalities:
    - target mass corrections,
    - different solutions for the DGLAP and coupling evolutions (?).

#### New functionalities

- I have already started using APFEL++ for tasks difficult to implement in (or even out of reach) for the FORTRAN version.
- Examples are:
  - Semi-Inclusive DIS (SIDIS) in collinear factorisation:
    - double convolution with time- and space-like evolution at the same time.
  - TMD phenomenology:
    - evolution and matching,
    - $\bullet$  Drell-Yan and SIDIS  $q_T$  distributions.
  - DGLAP evolution with splitting explicitly depending the factorisation scale:
    - e.g. "Physical"-scheme evolution (by Martin and Ryskin).
  - Transversity distributions (PDFs and FFs).

#### SIDIS in collinear factorisation

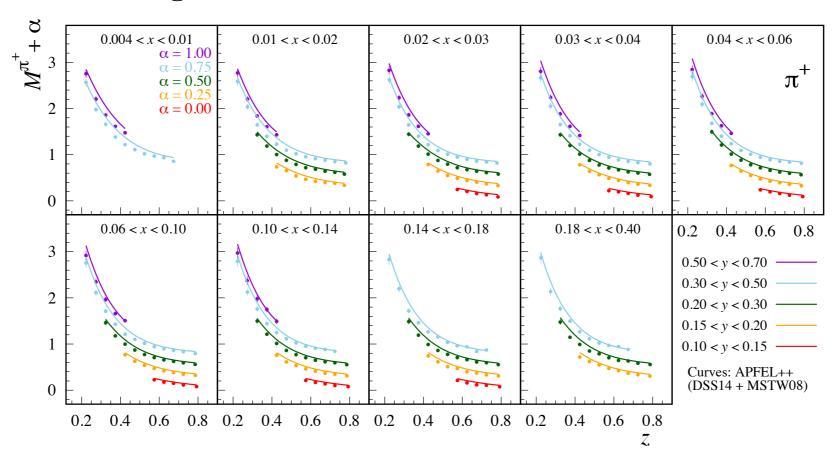
 $\bullet$  SIDIS cross sections (integrated over  $q_T$ ) have this structure:

$$D(x,z) = \int_{x}^{1} \frac{d\xi}{\xi} \int_{z}^{1} \frac{d\zeta}{\zeta} O\left(\frac{x}{\xi}, \frac{z}{\zeta}\right) d^{(1)}(\xi) d^{(2)}(\zeta)$$

• But the hard cross sections (at least up to NLO) factorise as:

$$O(x,z) = \sum_{i} K_{i}C_{i}^{(1)}(x)C_{i}^{(2)}(z)$$

Combination of single convolutions:



Next, I will also try with Drell-Yan cross sections.

# TMD Evolution (PDFs)

$$F_{f/P}(x, \mathbf{b}_T; \mu, \zeta) = \sum_j C_{f/j}(x, b_*; \mu_b, \zeta_F) \otimes f_{j/P}(x, \mu_b) : A$$

$$\times \exp\left\{K(b_*; \mu_b) \ln \frac{\sqrt{\zeta_F}}{\mu_b} + \int_{\mu_b}^{\mu} \frac{d\mu'}{\mu'} \left[\gamma_F - \gamma_K \ln \frac{\sqrt{\zeta_F}}{\mu'}\right]\right\} : B$$

$$\times \exp\left\{g_{j/P}(x,b_T) + g_K(b_T) \ln \frac{\sqrt{\zeta_F}}{\sqrt{\zeta_{F,0}}}\right\} : C$$

### TMD Evolution (PDFs)

$$F_{f/P}(x, \mathbf{b}_T; \mu, \zeta) = \sum_{j} C_{f/j}(x, b_*; \mu_b, \zeta_F) \otimes f_{j/P}(x, \mu_b)$$

$$\times \left[ \exp \left\{ K(b_*; \mu_b) \ln \frac{\sqrt{\zeta_F}}{\mu_b} + \int_{\mu_b}^{\mu} \frac{d\mu'}{\mu'} \left[ \gamma_F - \gamma_K \ln \frac{\sqrt{\zeta_F}}{\mu'} \right] \right\} \right] : B$$

$$\times \left[ \exp \left\{ g_{j/P}(x, b_T) + g_K(b_T) \ln \frac{\sqrt{\zeta_F}}{\sqrt{\zeta_F}} \right\} \right] : C$$

- $b_{\rm T} \ll 1/\Lambda_{\rm QCD}$
- matching to the collinear region
- factorises as hard and non-perturbative
- numerically cumbersome
- precompute using APFEL

- CS evolution
- perturbative

- matching between the small and large  $b_{\rm T}$
- non perturbative
- parametrised and fitted to data

#### SIDIS in TMD factorisation

In SIDIS, what enters the computation of the cross sections is:

$$\mathcal{L}_{\text{SIDIS}} = \int \frac{d^2 \mathbf{b}_T}{(2\pi)^2} e^{-i\mathbf{q}_T \cdot \mathbf{b}_T} F_{f/P}(x, \mathbf{b}_T; \mu, \zeta_F) D_{H/f}(x, \mathbf{b}_T; \mu, \zeta_D)$$

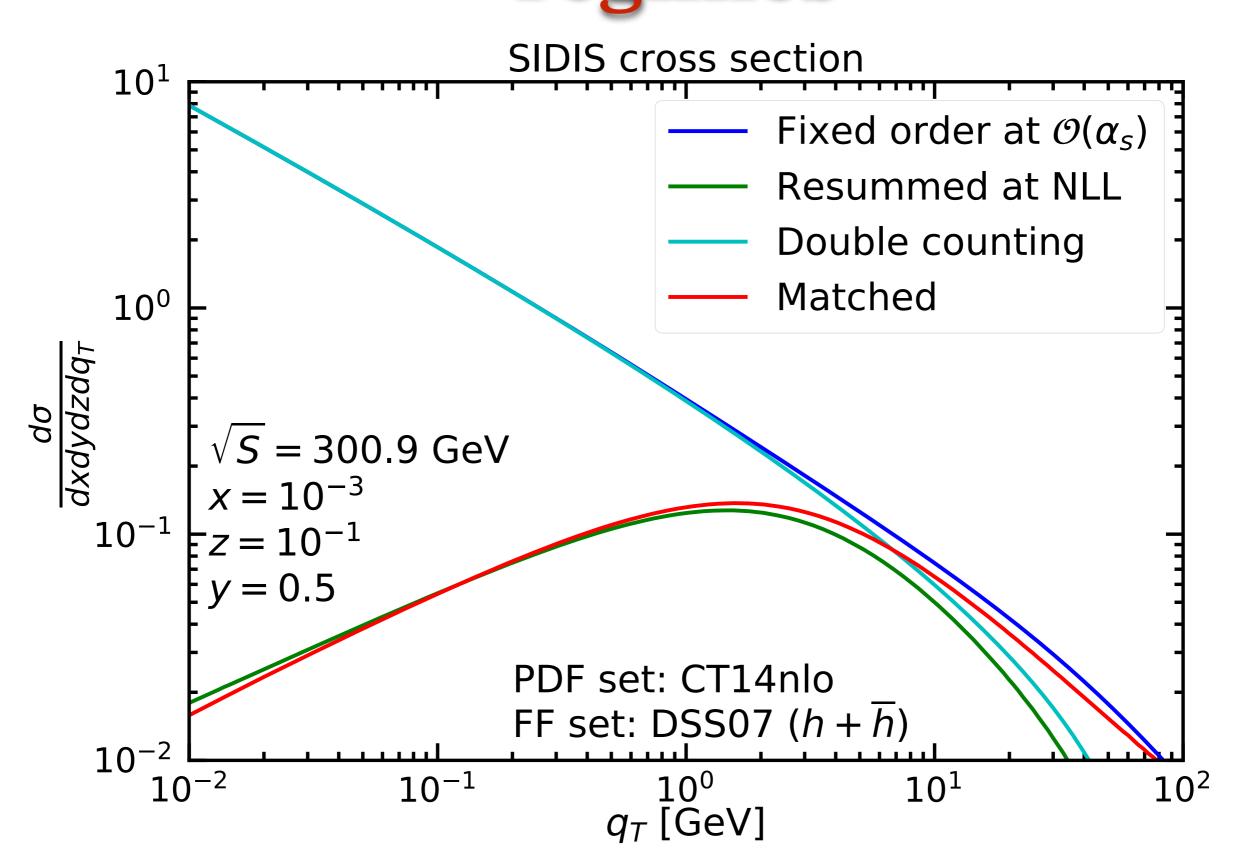
#### Fourier transform

**PDFs** 

**FFs** 

- The ingredients are:
  - a set of evolved TMD-PDFs,
  - a set of evolved TMD-FFs,
  - of the Fourier transform of its product.
- Complex set of tasks that have to be performed optimally
- APFEL provides the ideal environment for this computation:
  - fast and accurate interpolation techniques,
  - precomputation of the time consuming bits.

# Matching collinear and TMD regimes



#### Plans for the future

- Interface to yaml for parsing of evolution parameters.
- Interface to APFELgrid.
- Interface to **APPLgrid/FastNLO**:
  - Drell-Yan and SIDIS cross sections (?).
- PDF evol. and structure functions in a "OO" fashion useful for **xFitter**:
  - many possible evolution and structure functions available at the same time,
  - assign different evolutions to different datasets (e.g. H-VFNS),
  - fit PDFs and FFs at the same time (space- and time-like evolution).