

The XENON Dark Matter Project at Gran Sasso National Laboratory



Andrea Molinario

PATRAS 2019

XENON collaboration



Columbia



Rensselaer



Nikhef



Muenster



Stockholm



Mainz



MPIK, Heidelberg



Freiburg



University of Zurich

Zurich



Chicago



UCLA



UCSD



Rice



Purdue



Coimbra



Subatech



LPNHE



LAL



Bologna



INFN



Weizmann



جامعة نورثويورك أبوظبي
NYU ABU DHABI

NYUAD

~160 scientists
27 institutions

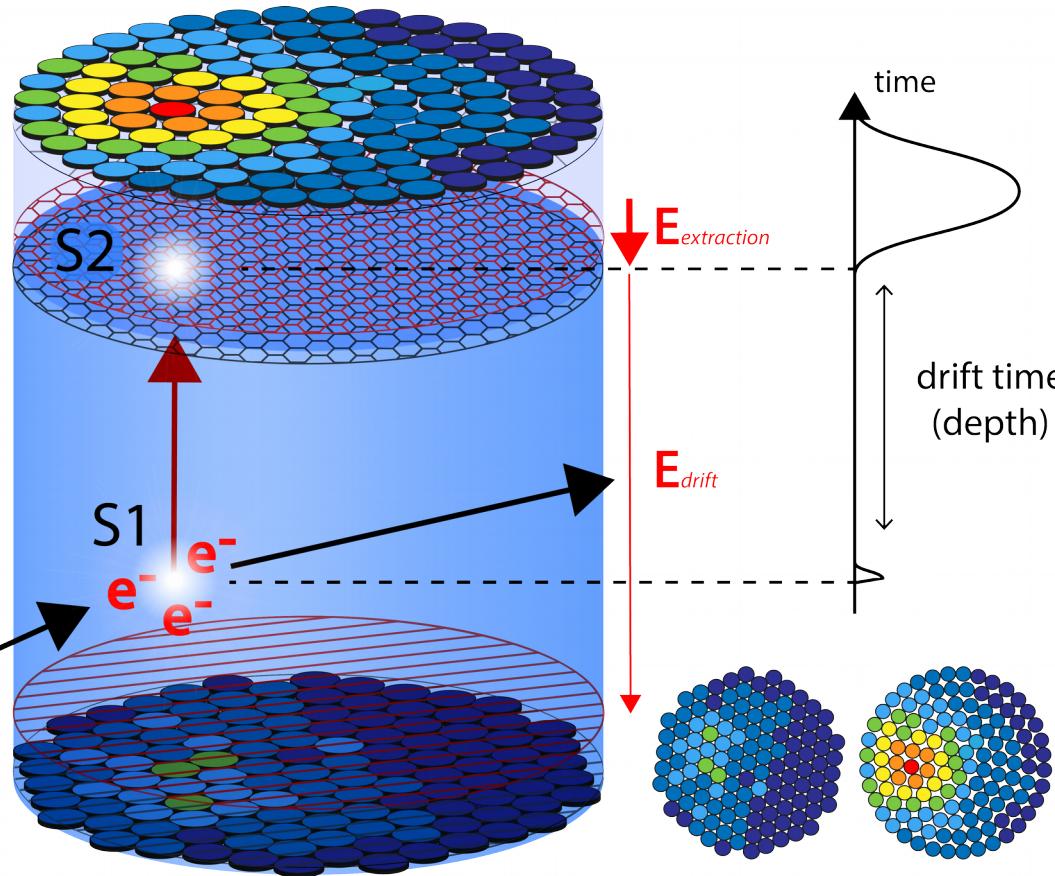


Dual-phase Xenon TPC

G_{Xe}

L_{Xe}

particle



Liquid Xenon

High density, self-shielding

Good scintillator

No long-living radioactive isotope

Time Projection Chamber

3D position

ER/NR discrimination

Multiple scatter rejection

Low energy threshold

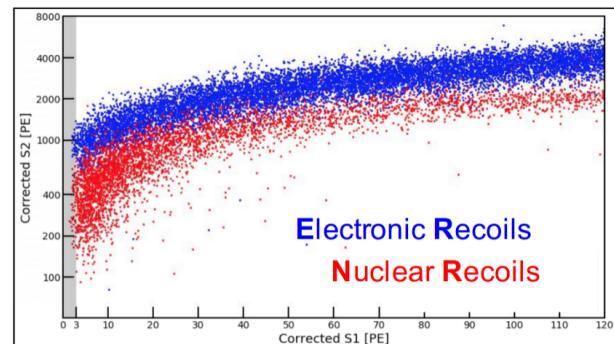
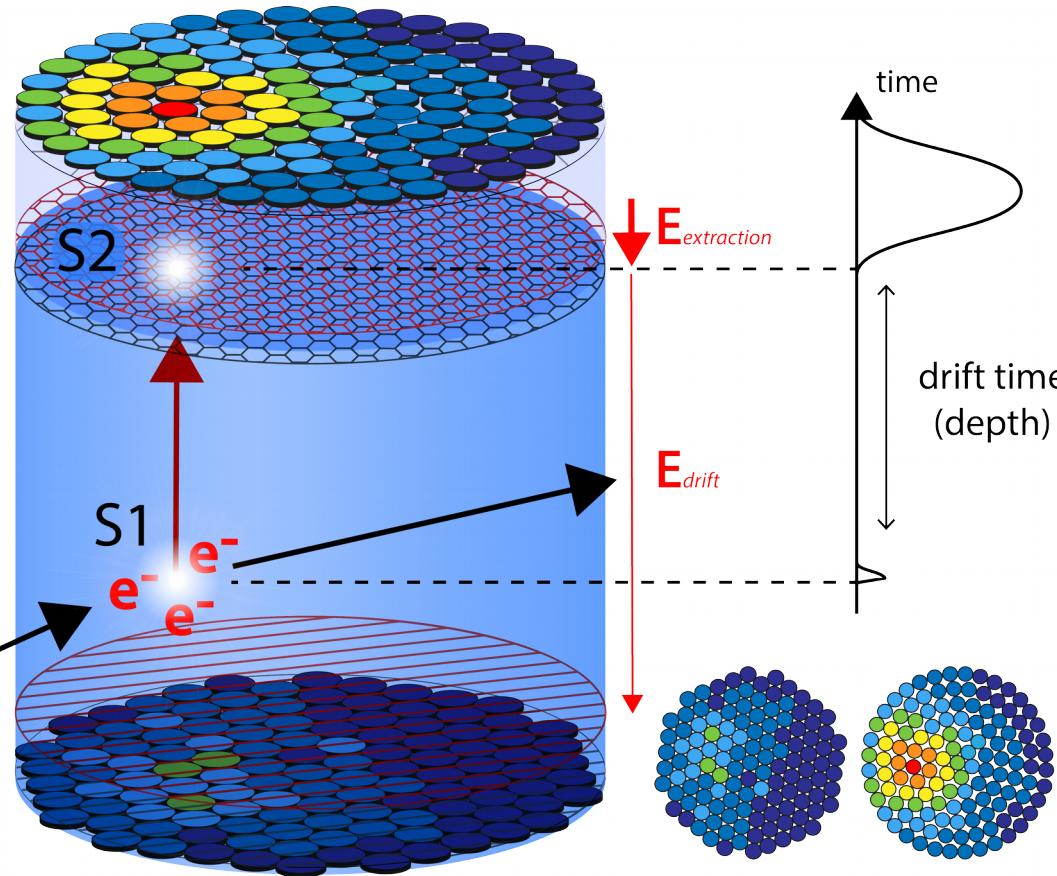
Scalable to multi-ton

Dual-phase Xenon TPC

G_{Xe}

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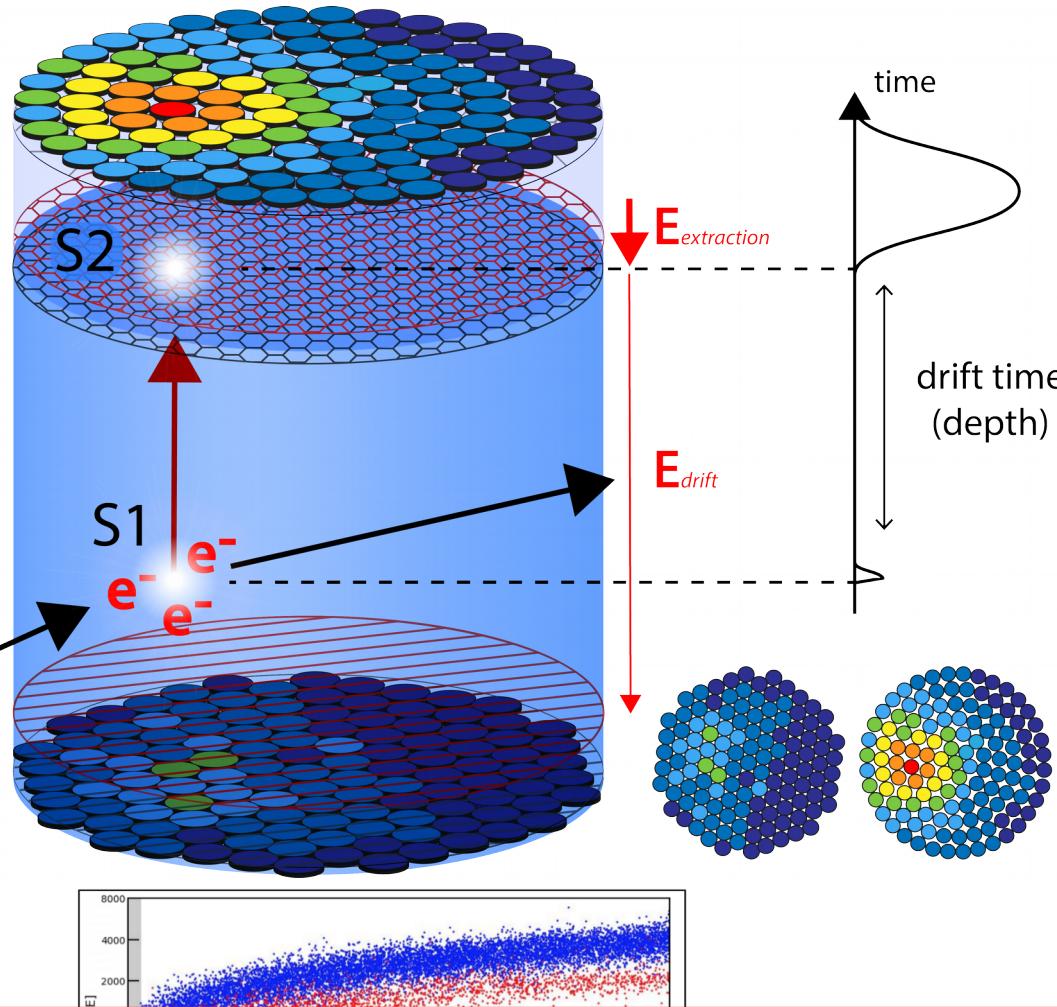
Scalable to multi-ton

Dual-phase Xenon TPC

G_{Xe}

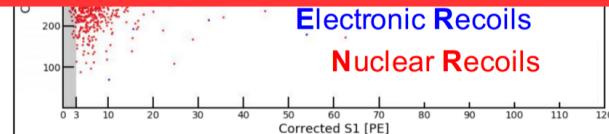
L_{Xe}

particle



Ideal for WIMP and rare processes search

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3D position

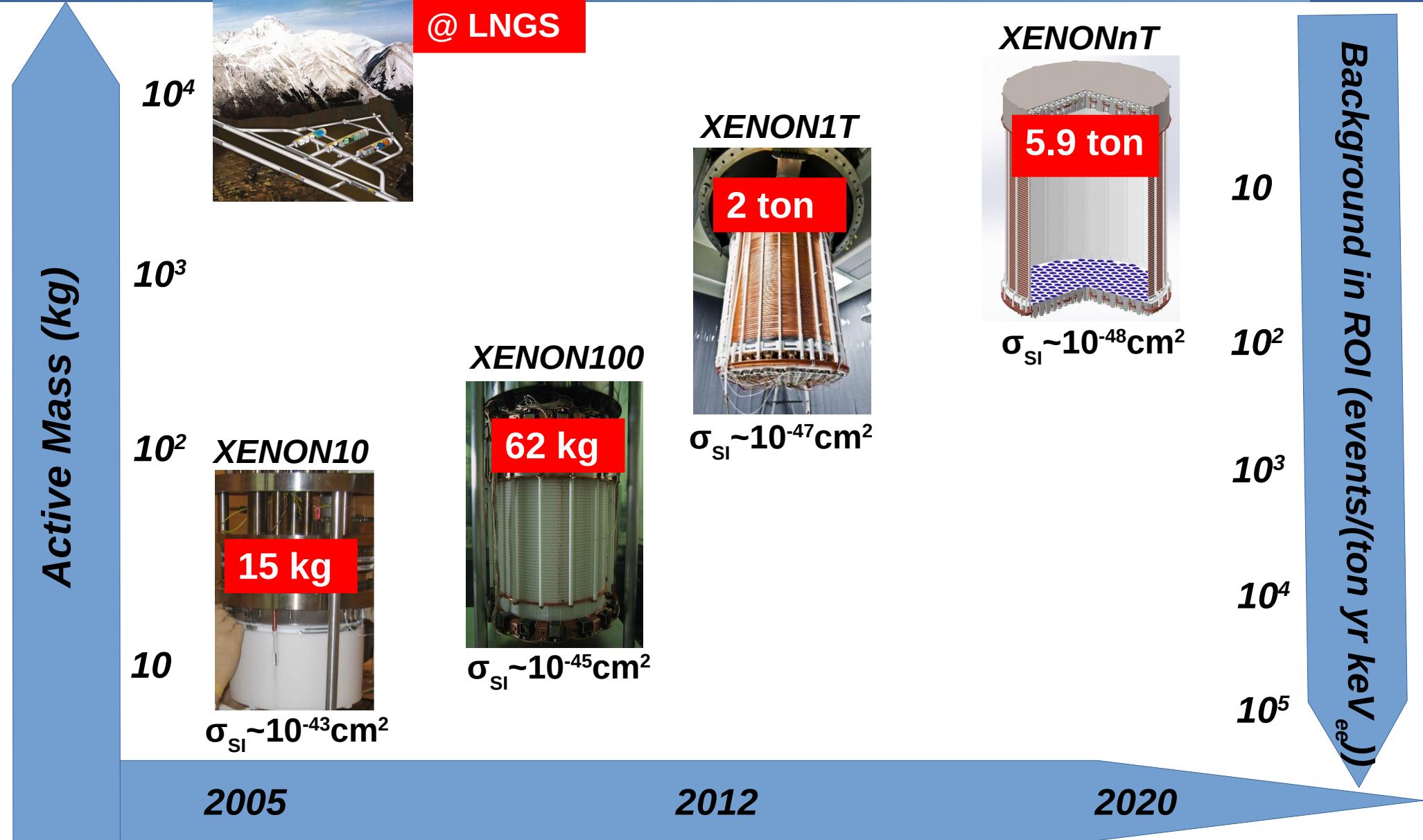
ER/NR discrimination

Multiple scatter rejection

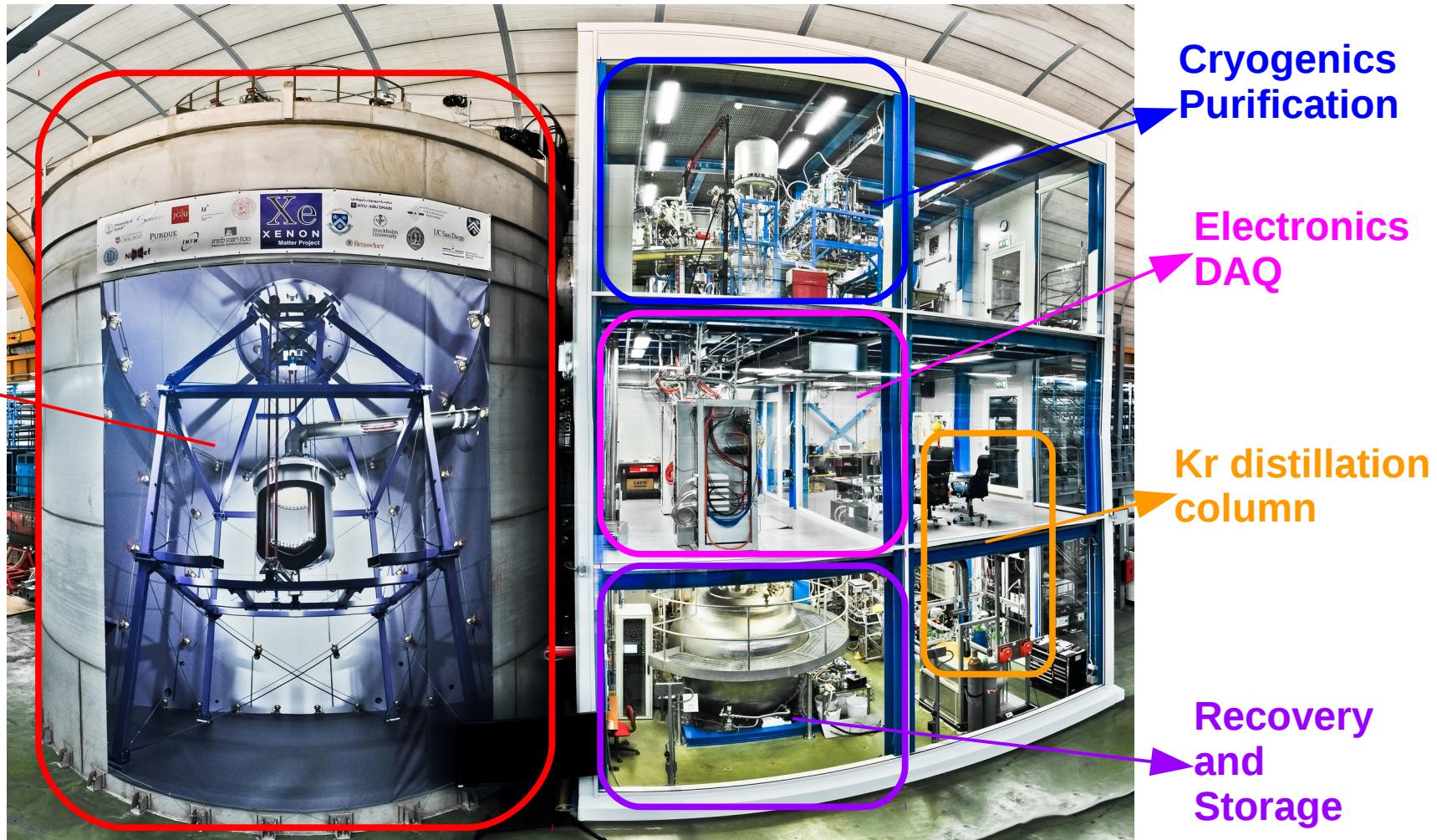
Low energy threshold

Scalable to multi-ton

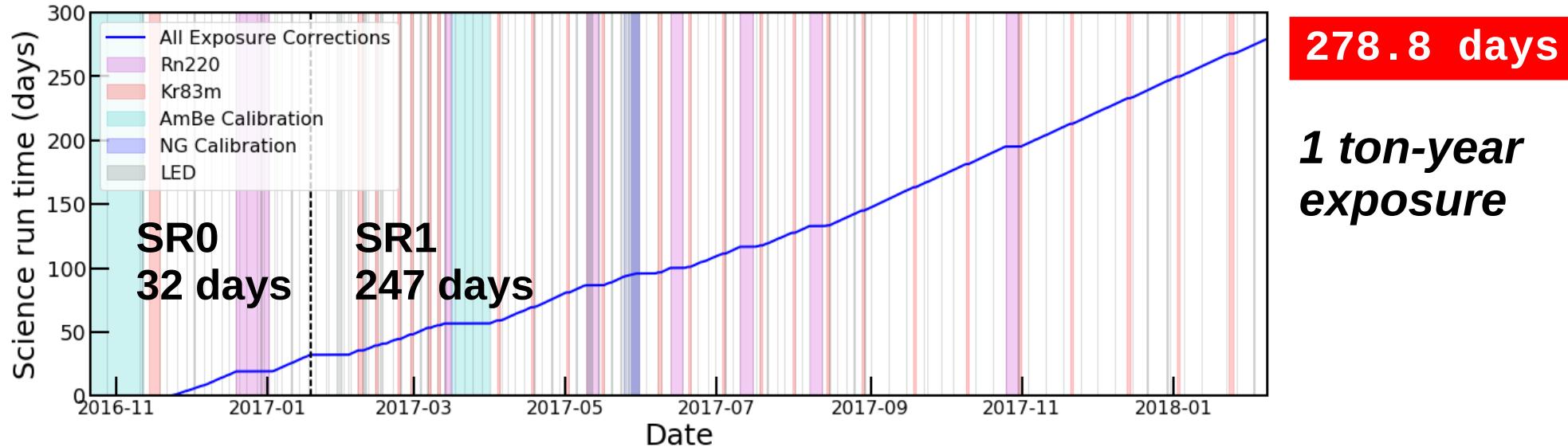
Timeline of the project



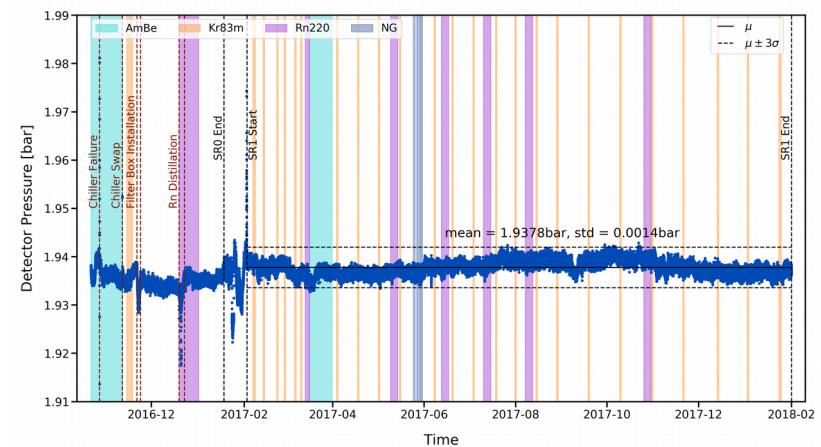
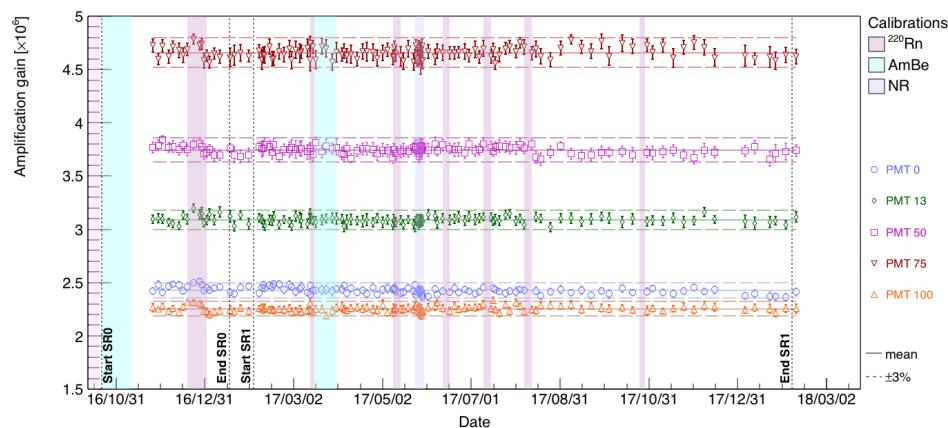
XENON1T



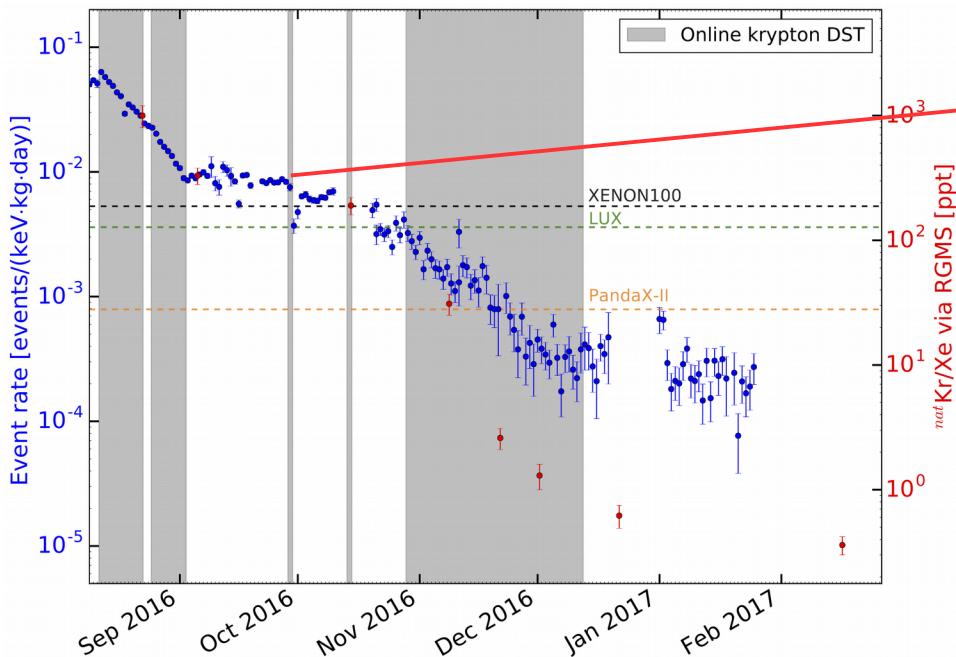
Data taking



Monitoring the stability of the detector and PMTs



Electronic recoil background



Initially Kr-dominated

Kr concentration reduced by distillation from 1 ppb to 0.7 ppt

SR1 ER background dominated by ^{222}Rn (mainly ^{214}Pb β -decay)

Source	Fraction [%]	Mitigation strategy	ER background in the ROI
^{222}Rn	85	S2/S1, material selection	$82^{+5}_{-3}(\text{syst}) \pm 3(\text{stat})$ events/(ton yr keV _{ee})
Solar ν	5	S2/S1	
^{85}Kr	4	S2/S1, distillation	
Materials	4	S2/S1, material selection, fiduc.	Lowest ER background for a dark matter detector
^{136}Xe	1	S2/S1	

Nuclear recoil background

Source

Radiogenic neutrons (from materials)

CEvNS (mainly ^8B solar ν)

Cosmogenic neutrons

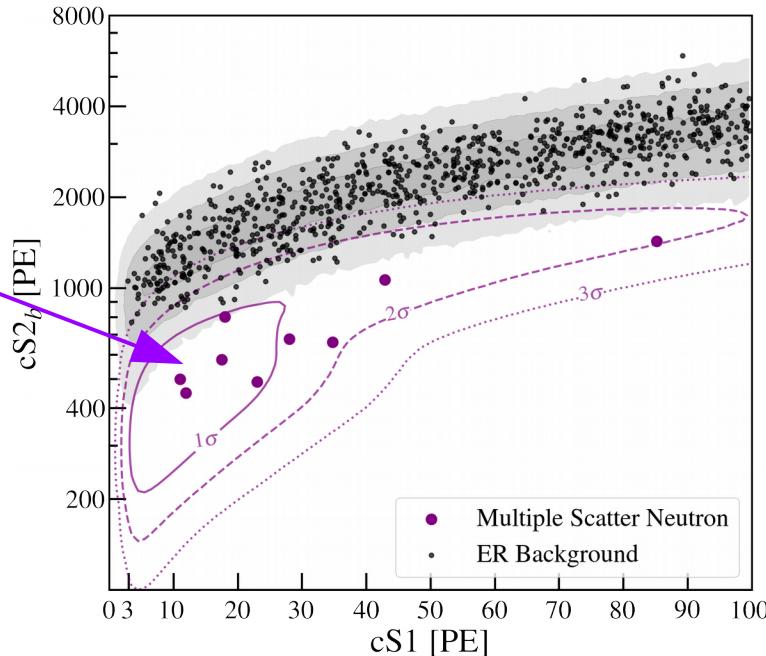
Dedicated search for multiple scatter events found 9 candidates with (6.4 ± 3.2) expected

Constrain the expected single-scatter neutron event rate

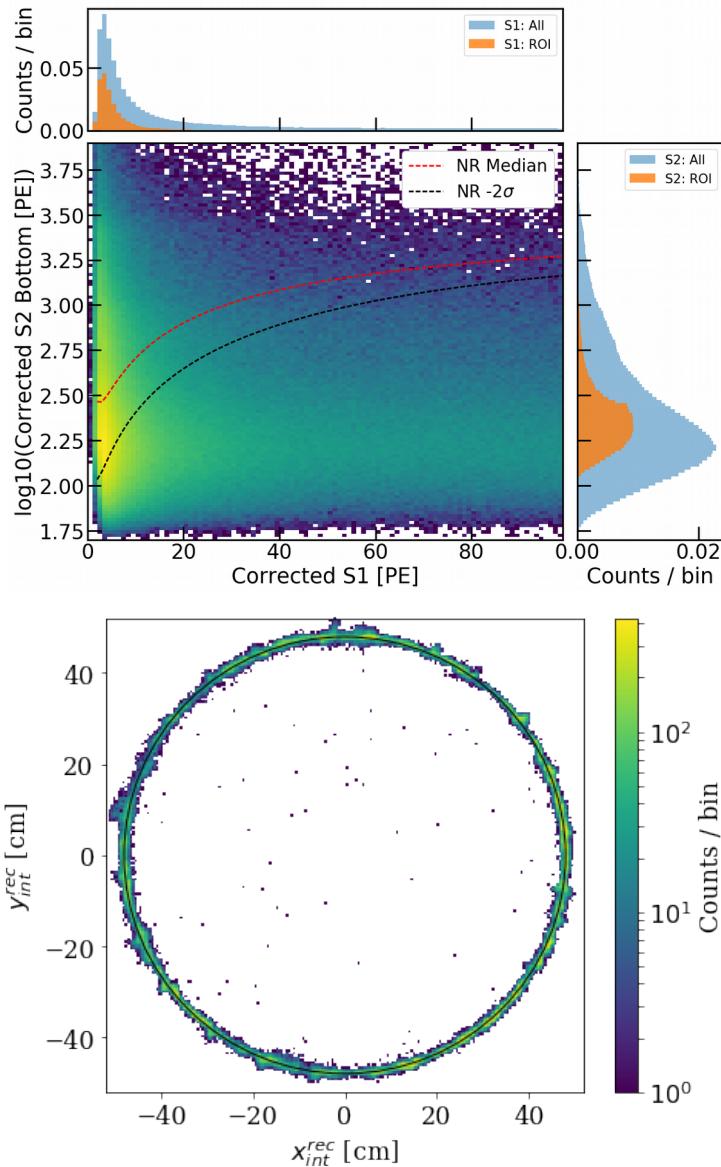
Mitigation strategy

Material selection, reject multiple scatter, fiducialization

Muon Veto, reject multiple scatter, fiducialization



Other backgrounds



Accidental coincidences

Random pairing of lone S1 and S2

Background model derived from data
and used in likelihood estimation

Surface events

^{222}Rn progeny
plate-out on the
inner surface of
PTFE panels

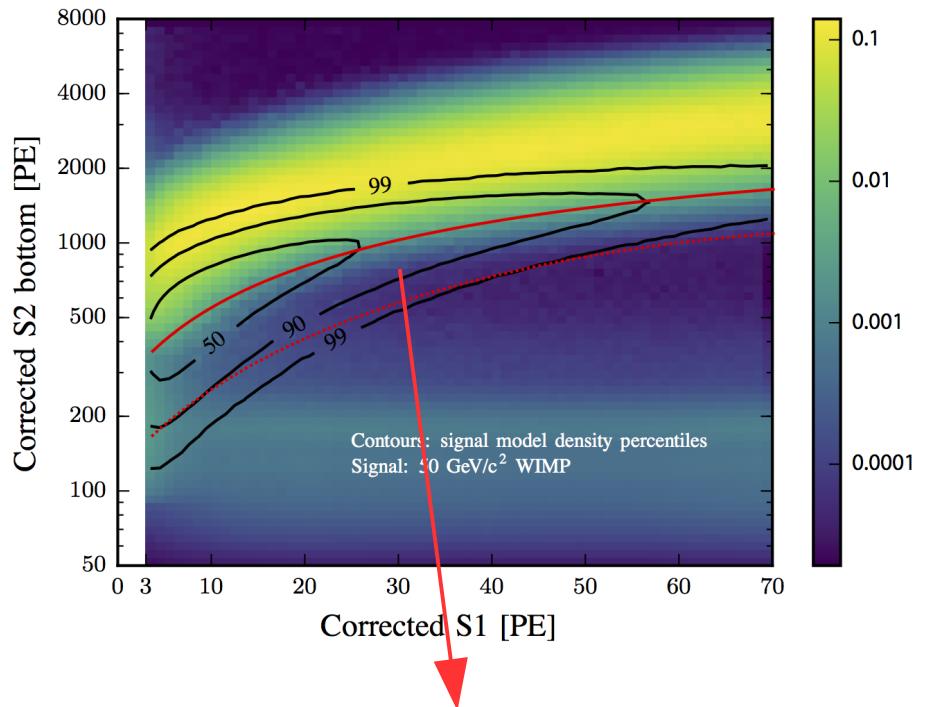
Charge loss
which reduces
S2 size

Events
shifted in
NR band

Data-driven background model

Background predictions

ROI corresponds in average to
 $[4.9, 40.9]$ keV_{nr} ($[1.4, 10.6]$ keV_{ee})



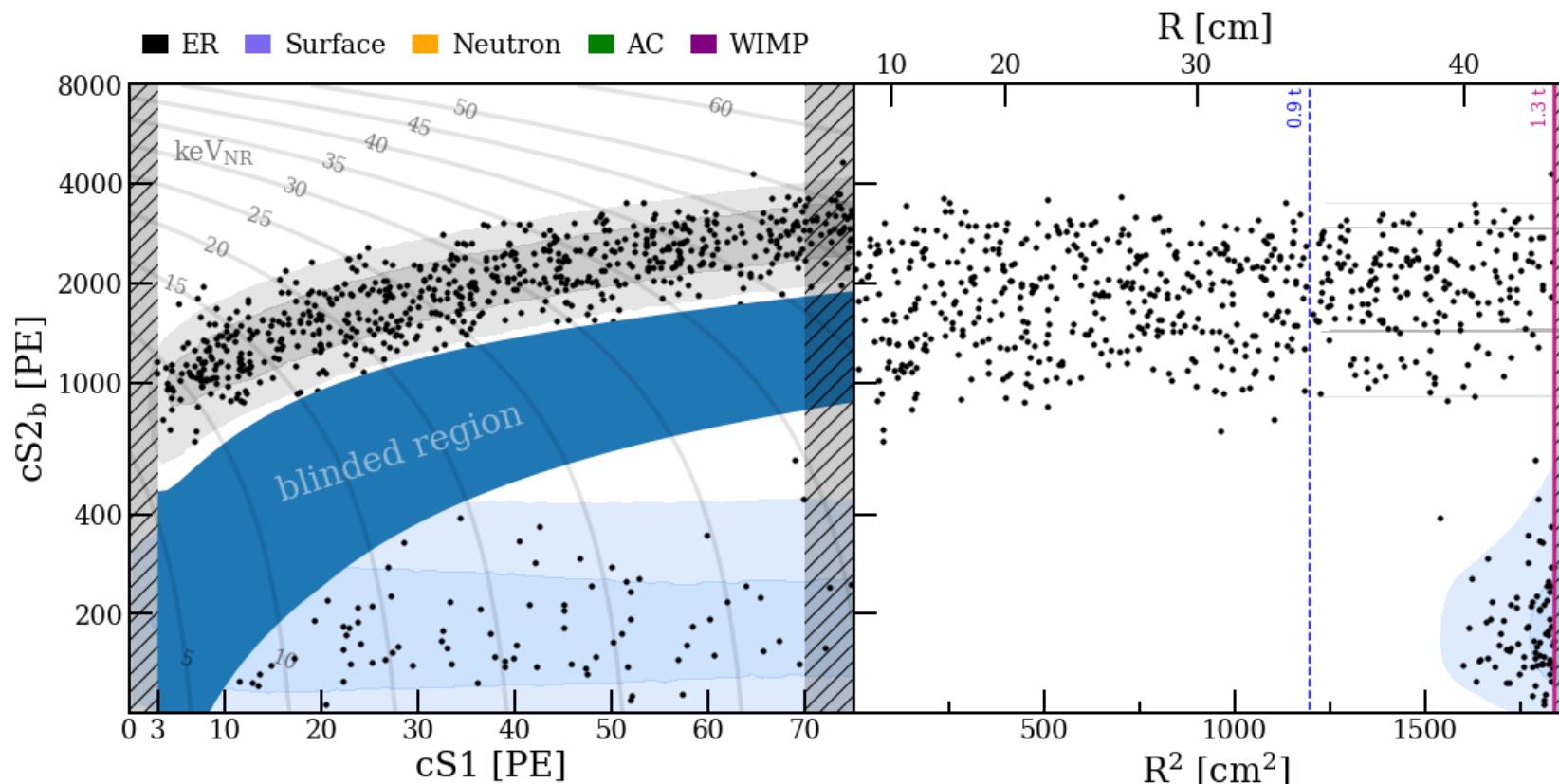
NR reference region
50% NR acceptance with 99.75% ER rejection

Background model in 4 dimensions:
S1, S2, R, Z

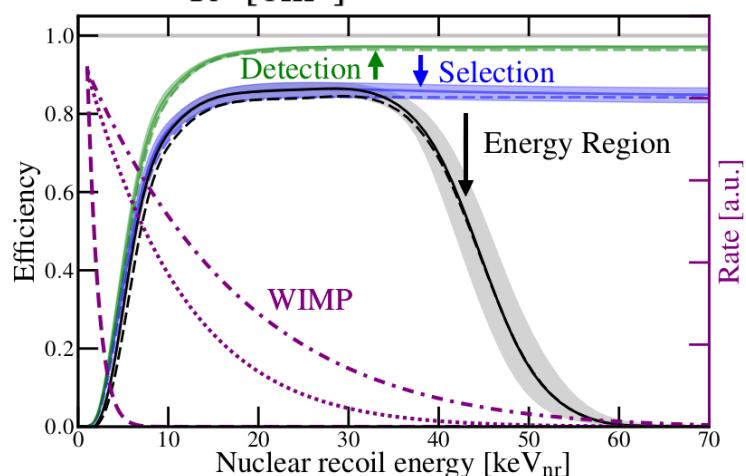
Expected events / bin	Mass (ton)	1.3	1.3	0.9
	(cS1, cS2 _b)	Full	Reference	Reference
ER		627 ± 18	1.62 ± 0.30	1.12 ± 0.21
Neutron		1.43 ± 0.66	0.77 ± 0.35	0.41 ± 0.19
CE ν NS		0.05 ± 0.01	0.03 ± 0.01	0.02
AC		$0.47^{+0.27}_{-0.00}$	$0.10^{+0.06}_{-0.00}$	$0.06^{+0.03}_{-0.00}$
Surface		106 ± 8	4.84 ± 0.40	0.02
Total BG		735 ± 20	7.36 ± 0.61	1.62 ± 0.28

Statistical inference in 1.3 t fiducial volume and full (S1, S2) space

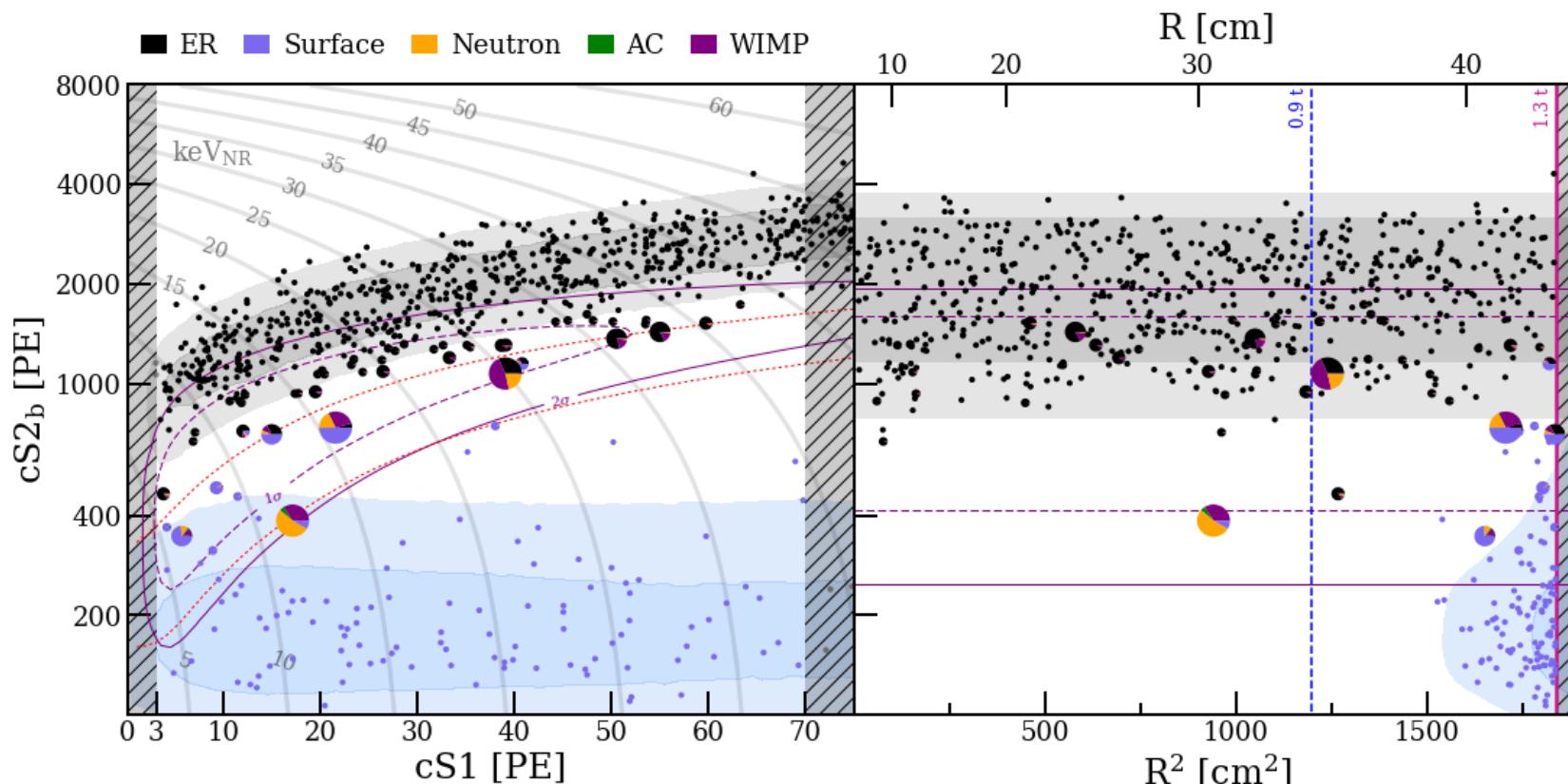
SI-WIMP result



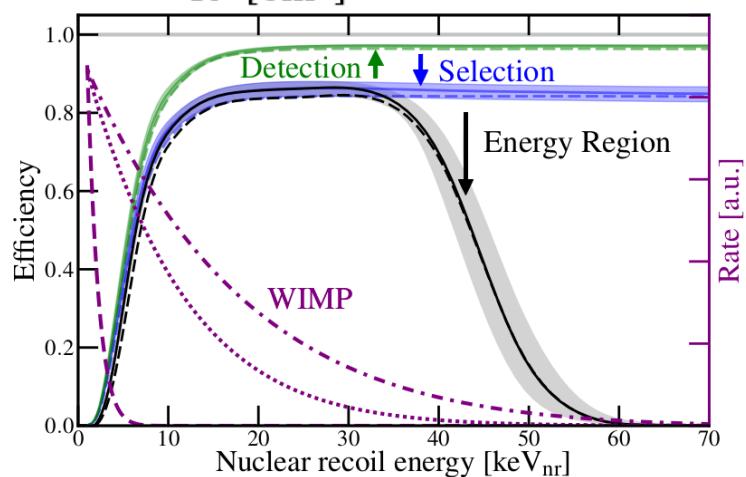
All selection criteria were defined before unblinding



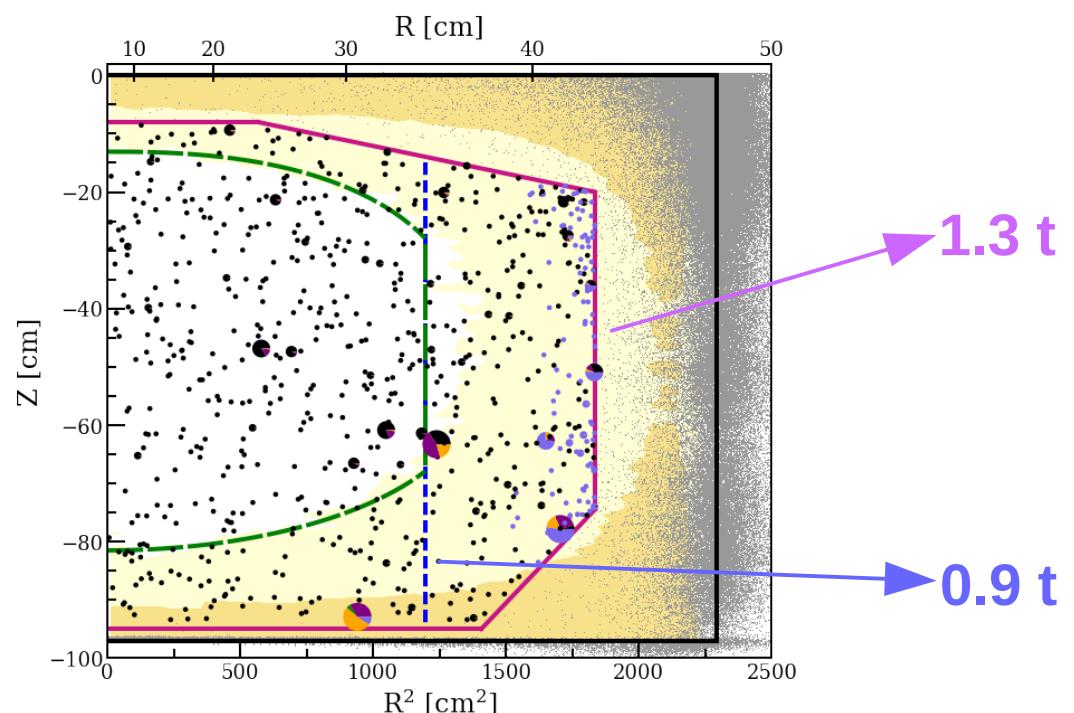
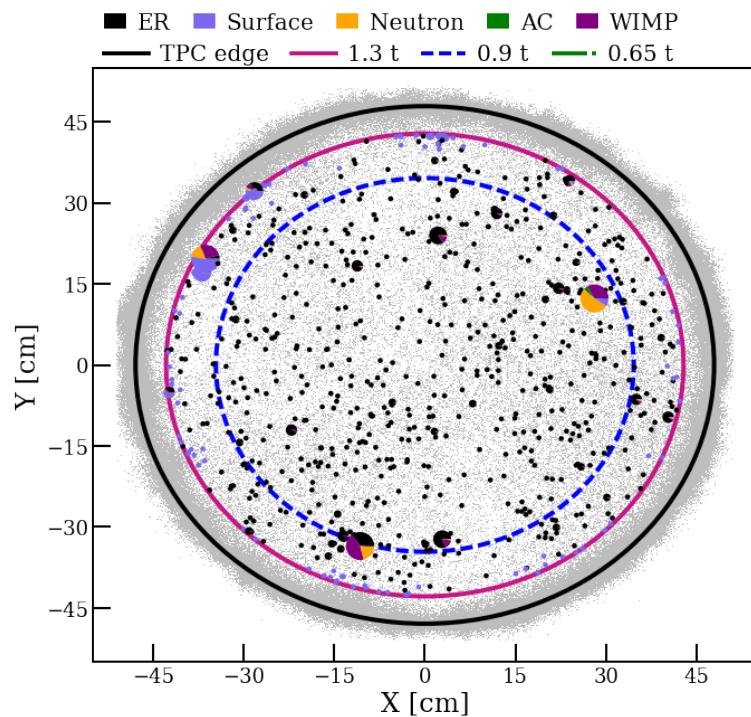
SI-WIMP result



Events that pass all cuts are shown
They are shown as pie charts
representing the best-fit probabilities of
the background and signal (200 GeV
WIMP) components at each event



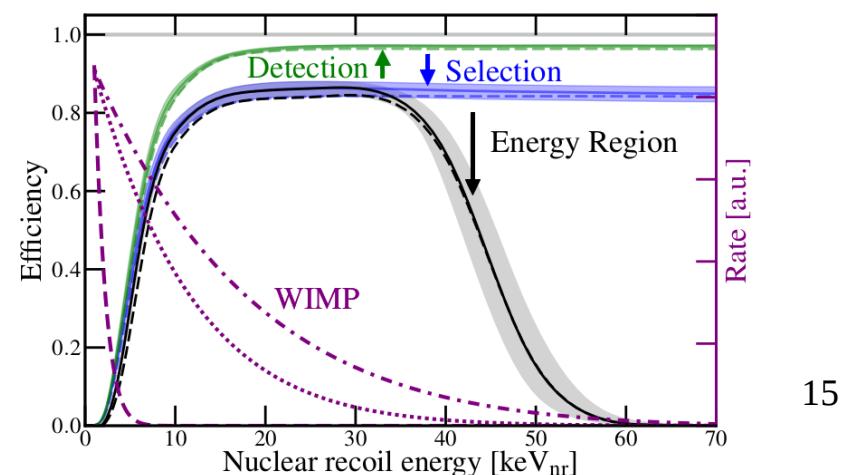
SI-WIMP result



Performed unbinned profile likelihood, model uncertainties included as nuisance parameters

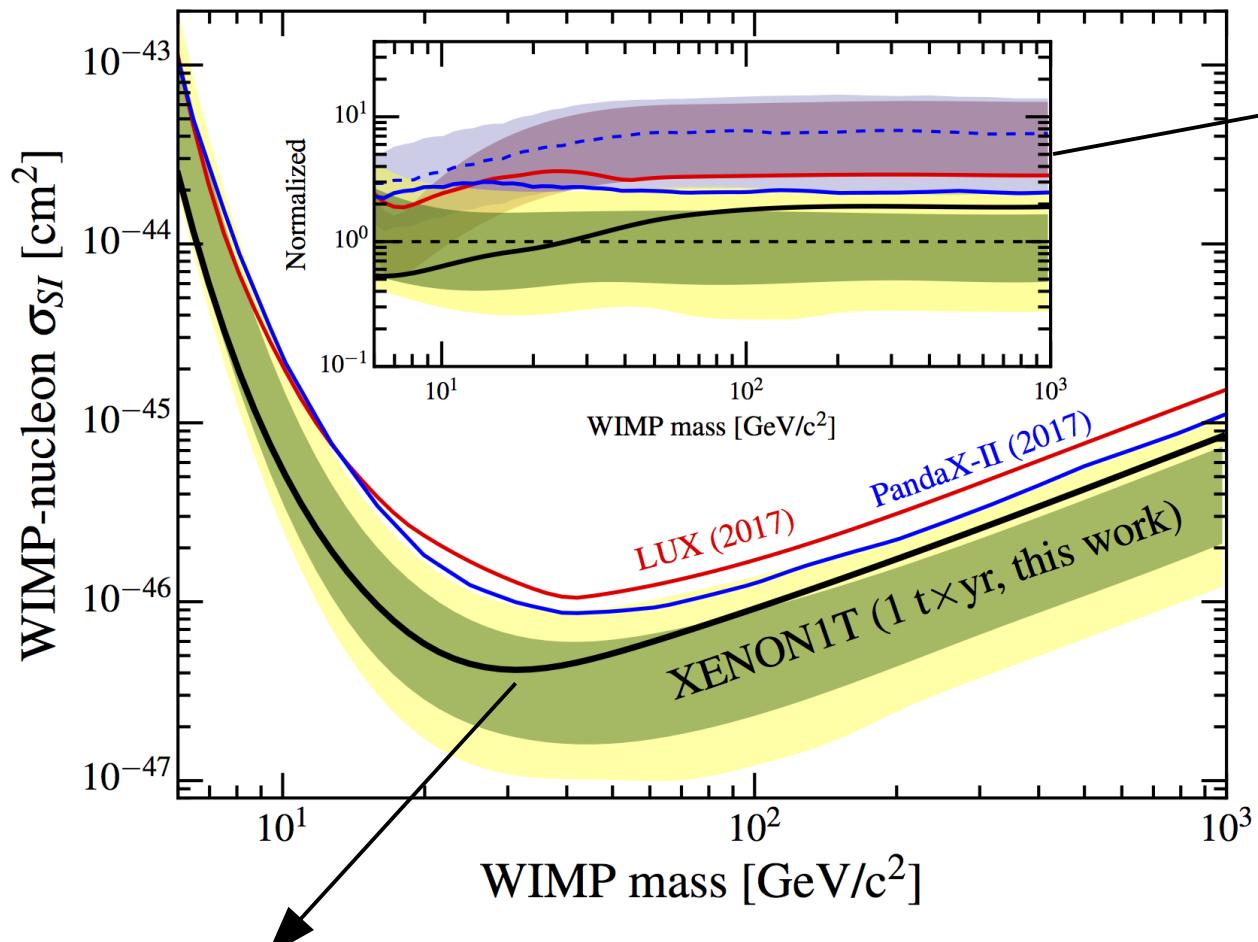
Maximum radius of 1.3 t fiducial volume set by surface event contribution.

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SI-WIMP result

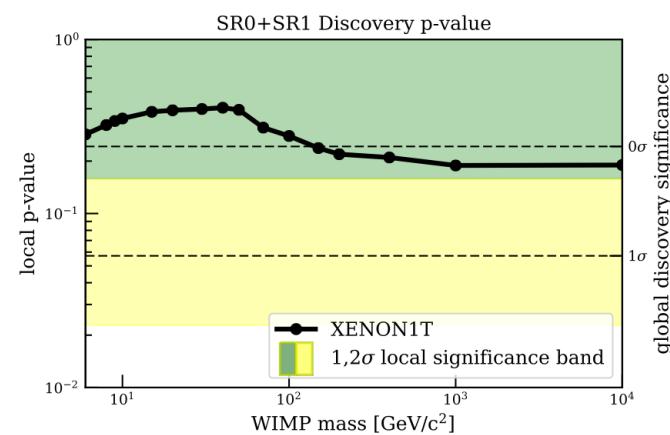
E. Aprile et al., Phys. Rev. Lett. 121, 111302 (2018)



$\sigma_{SI} < 4.1 \times 10^{-47}$ cm 2 (90% C.L.) @ 30 GeV/c 2

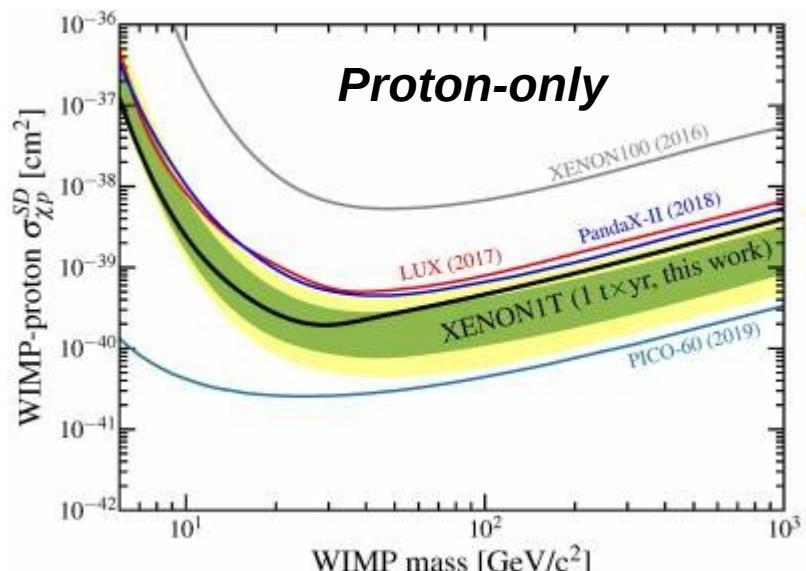
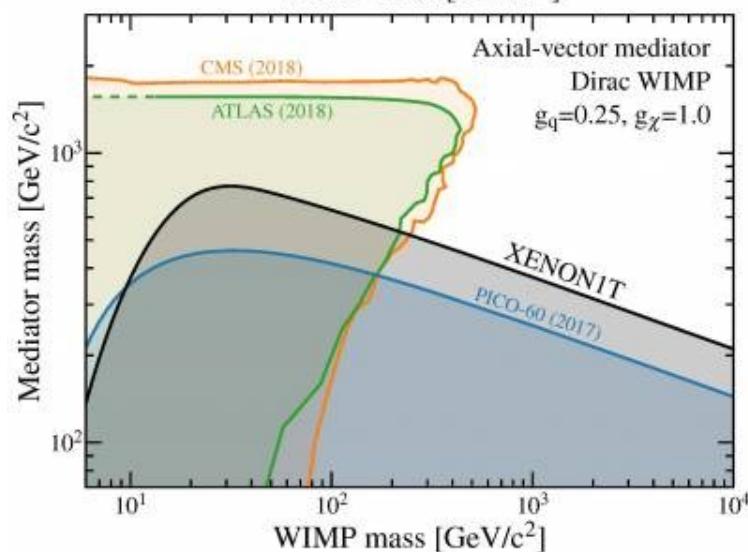
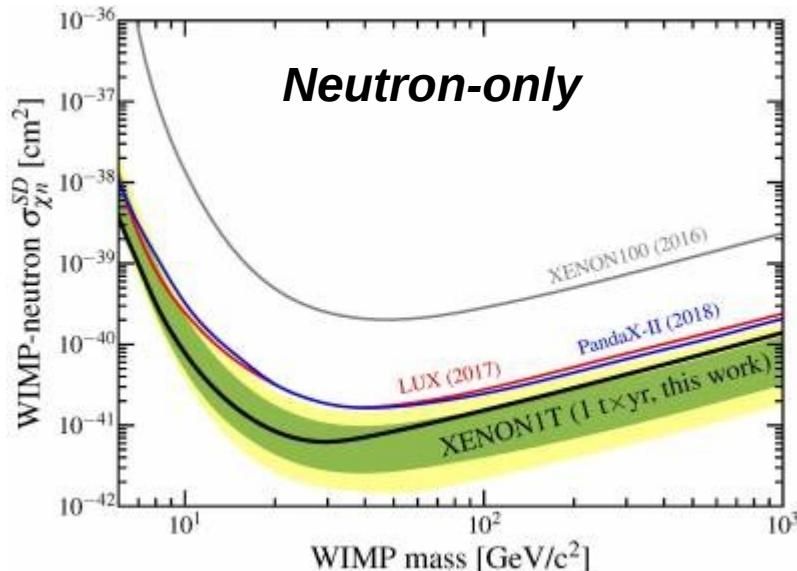
Median sensitivity 7 times better than previous experiments

No significant excess ($>3\sigma$) in the 1.3 tons fiducial volume at any WIMP mass



SD-WIMP result

E. Aprile et al., Phys. Rev. Lett. 122, 141301 (2019)



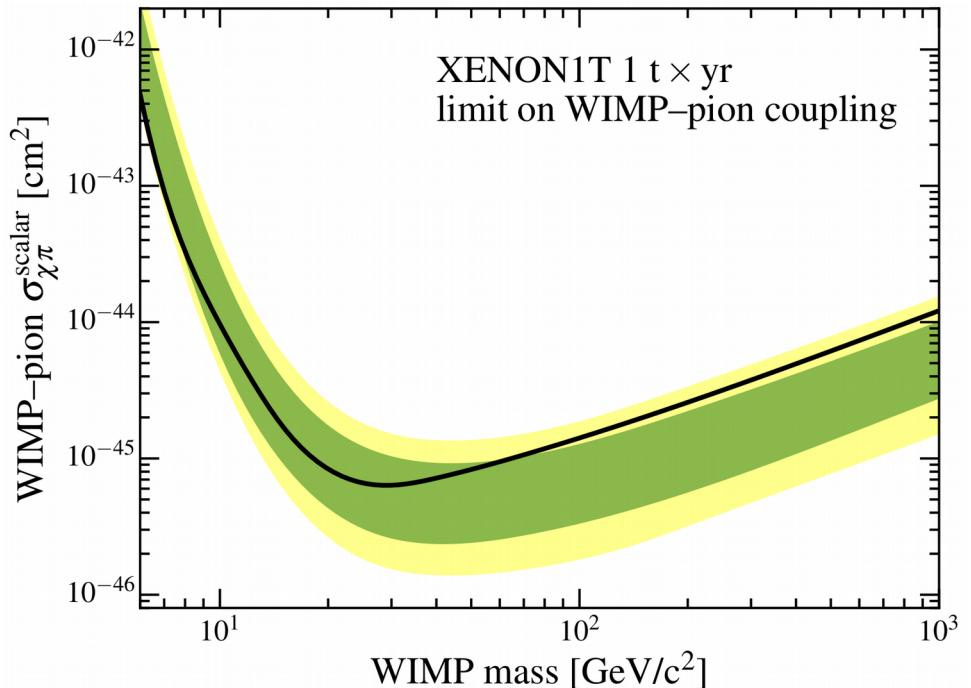
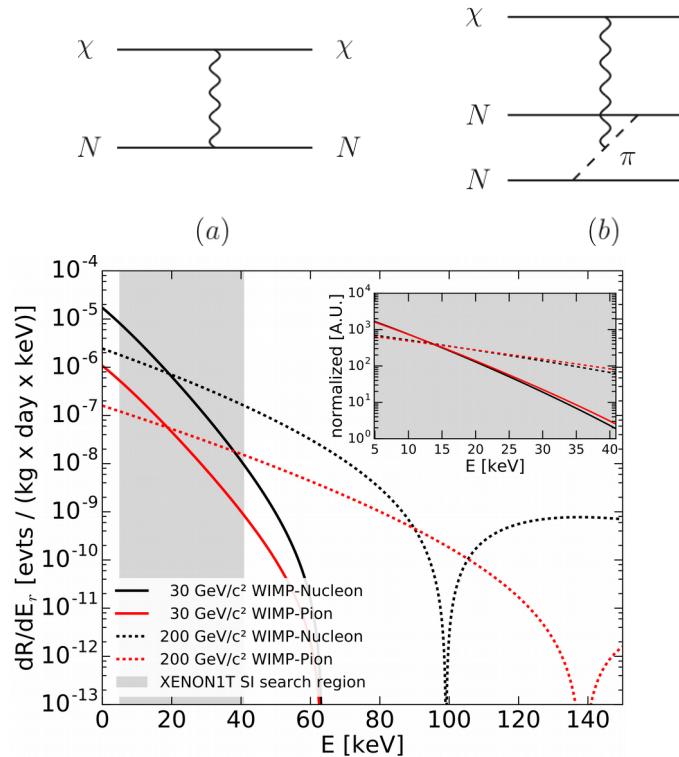
Same event selection criteria
for a SD search

Most stringent limit on WIMP-
neutron scattering cross section

Exclude new parameter space in isoscalar
theory with axial-vector mediator

WIMP-Pion coupling

E. Aprile et al., Phys. Rev. Lett. 122, 071301 (2019)

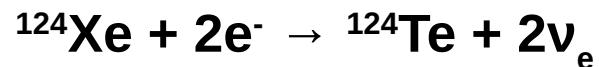
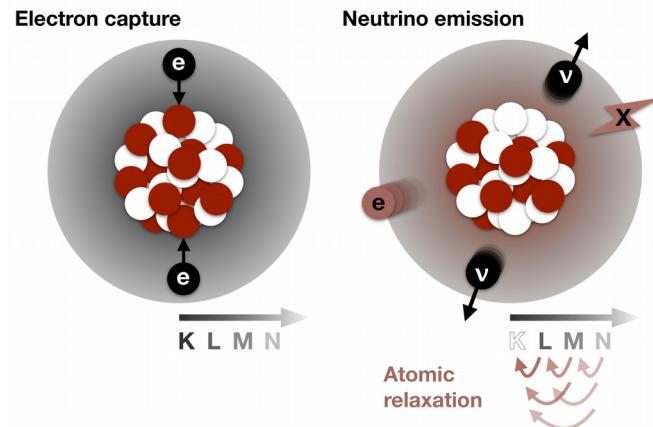


Coupling of WIMP with virtual pion-current between two nucleons

Same falling exponential differential recoil spectrum as WIMP-nucleon interaction

Limit setting as in SI analysis

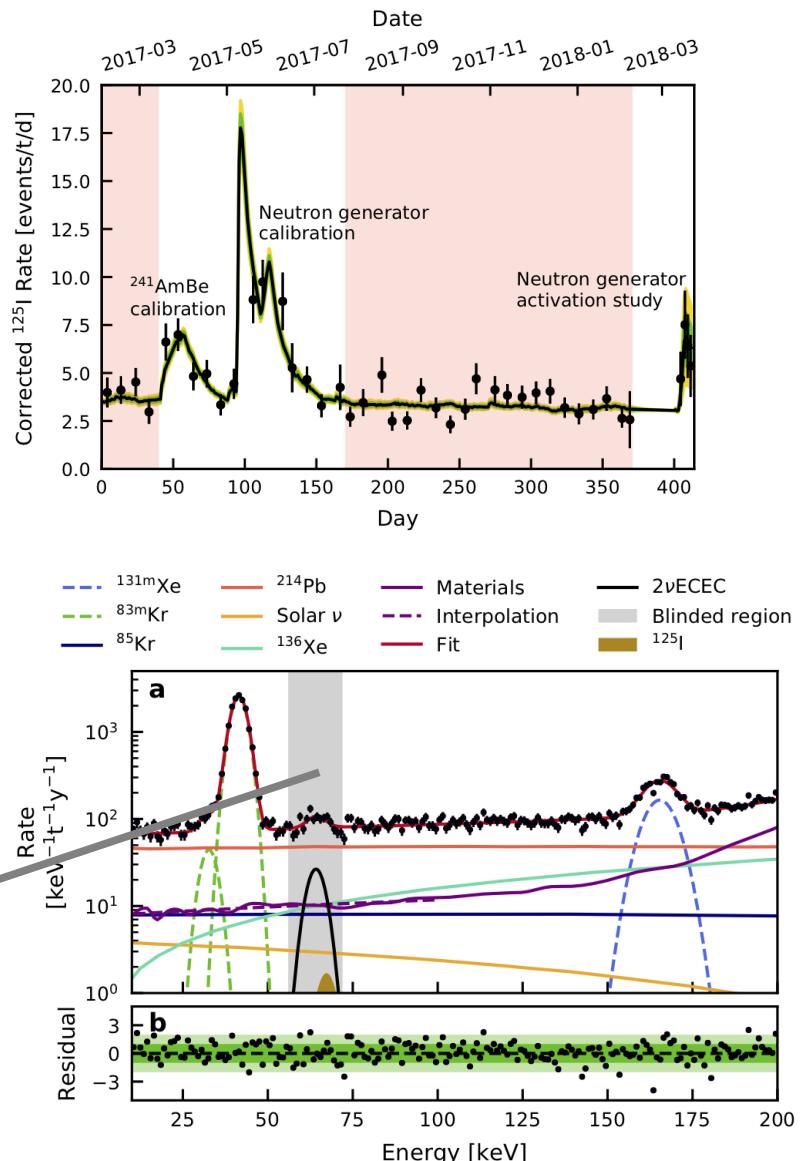
^{124}Xe Double Electron Capture



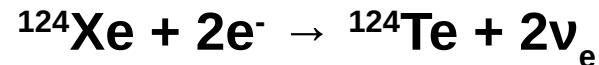
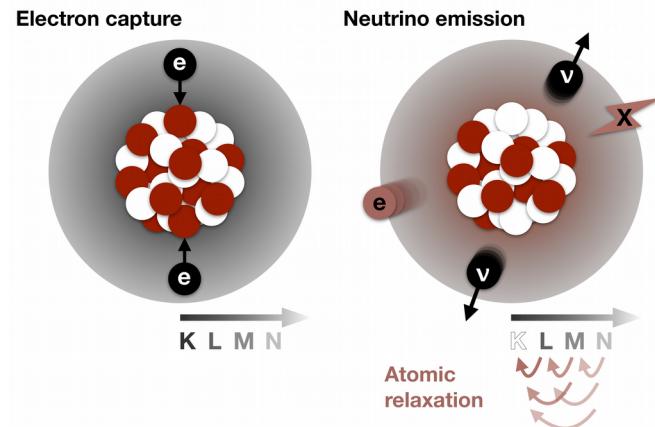
Detection of X-rays and
Auger electrons
Total energy (64.3 ± 0.6) keV

Background from
 ^{125}I produced by
 ^{124}Xe activation

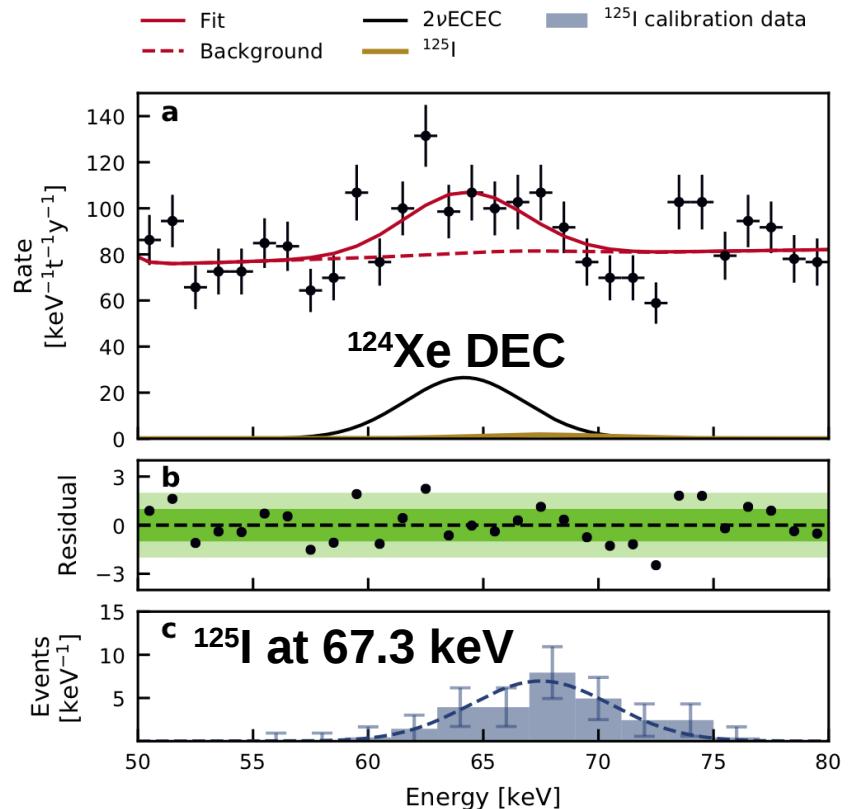
Blinded [56-72] keV
region



^{124}Xe Double Electron Capture



Detection of X-rays and
Auger electrons
Total energy (64.3 ± 0.6) keV

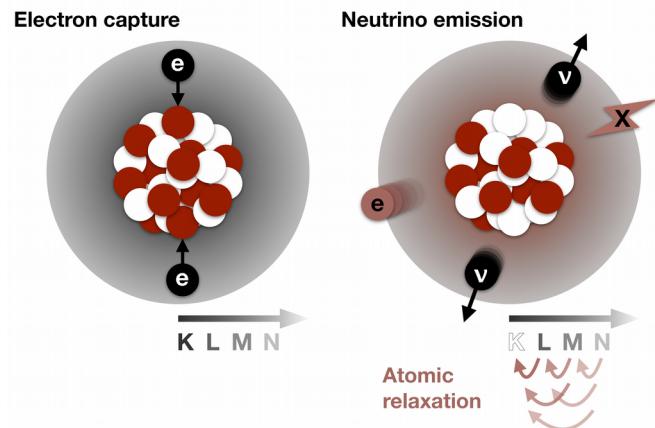


Detected peak at (64.2 ± 0.5) keV with 4.4σ significance

Measured half-life of the process $T_{1/2} = (1.8 \pm 0.5_{\text{stat}} \pm 0.1_{\text{sys}}) \times 10^{22} \text{ y}$

E. Aprile et al., Nature 568 (2019), no. 7753, 532-535

^{124}Xe Double Electron Capture

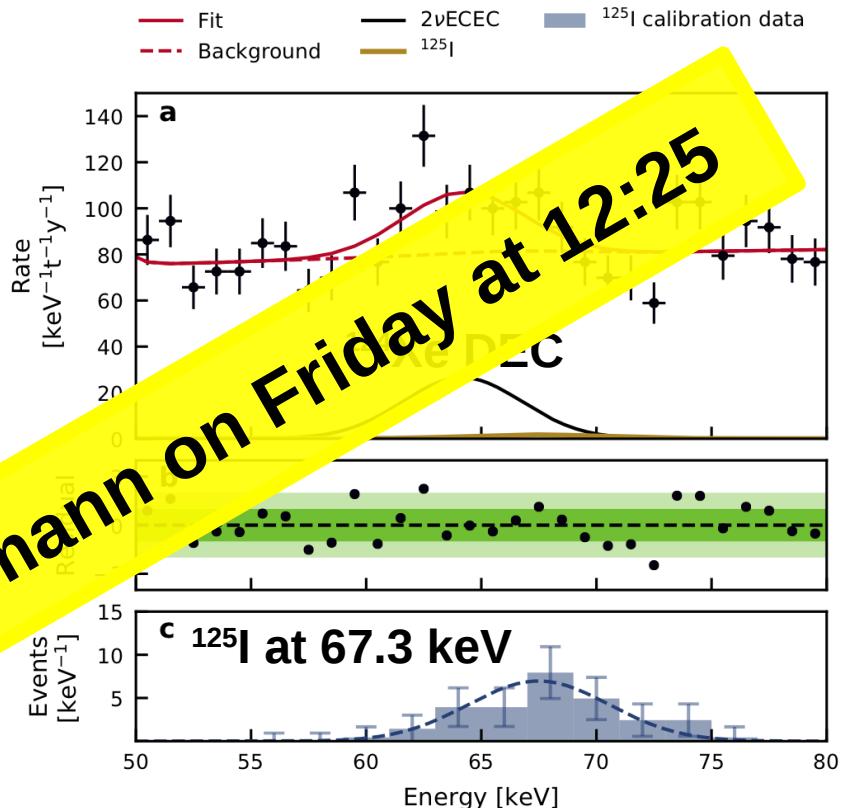


Detection of X-rays and Auger electrons
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R. Aprile et al., Nature 568 (2019), no. 7753, 532-535



Dedicated talk by S. Lindemann on Friday at 12:25

Ongoing analysis

S2-only analysis

WIMP search with Migdal effect

ALPs, Super WIMPs,
Dark photons, Solar Axions

Annual modulation

0v $\beta\beta$ of ^{136}Xe

^{37}Ar calibration



Ongoing analysis

S2-only analysis

WIMP search with Migdal effect

Low energy Electronic Recoils

ALPs, Super WIMPs,
Dark photons, Solar Axions

Annual modulation

$0\nu\beta\beta$ of ^{136}Xe

^{37}Ar calibration



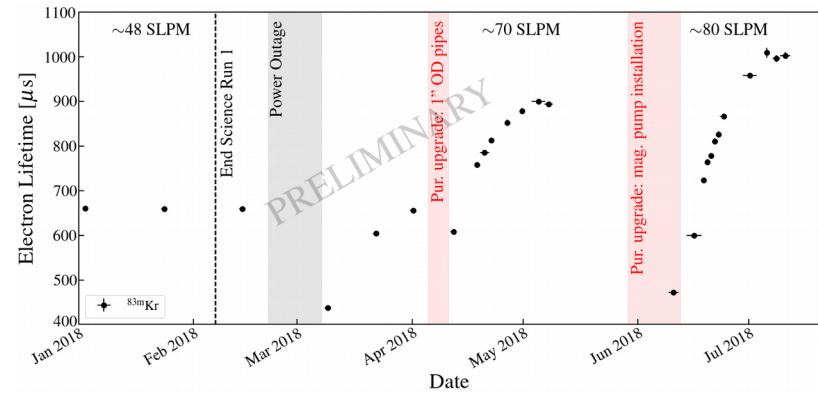
Tests after SR1

Upgrade of purification system

New magnetic pump

Increased purification of gas flow

1 ms electron lifetime reached

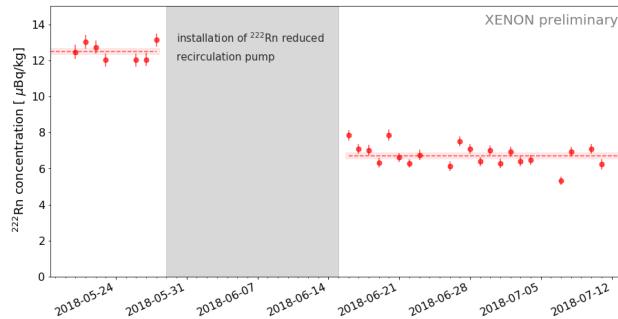


Rn-removal

With new magnetic pump Radon reduced by 45%

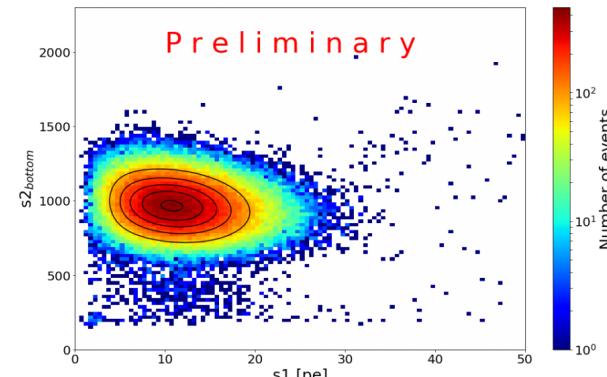
Rn distillation tested, another 30% reduction

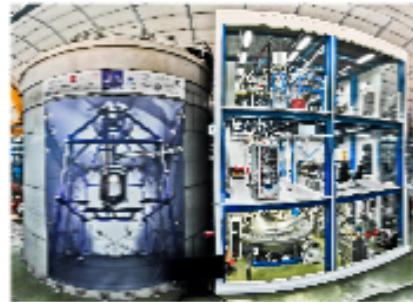
Factor 4 above XENONnT goal ($1\mu\text{Bq/kg}$)



^{37}Ar calibration

Test of new calibration source for low energy ER (2.8 keV, 0.27 keV)





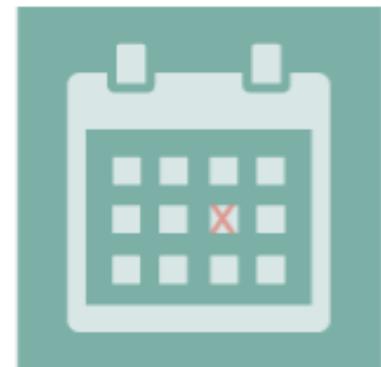
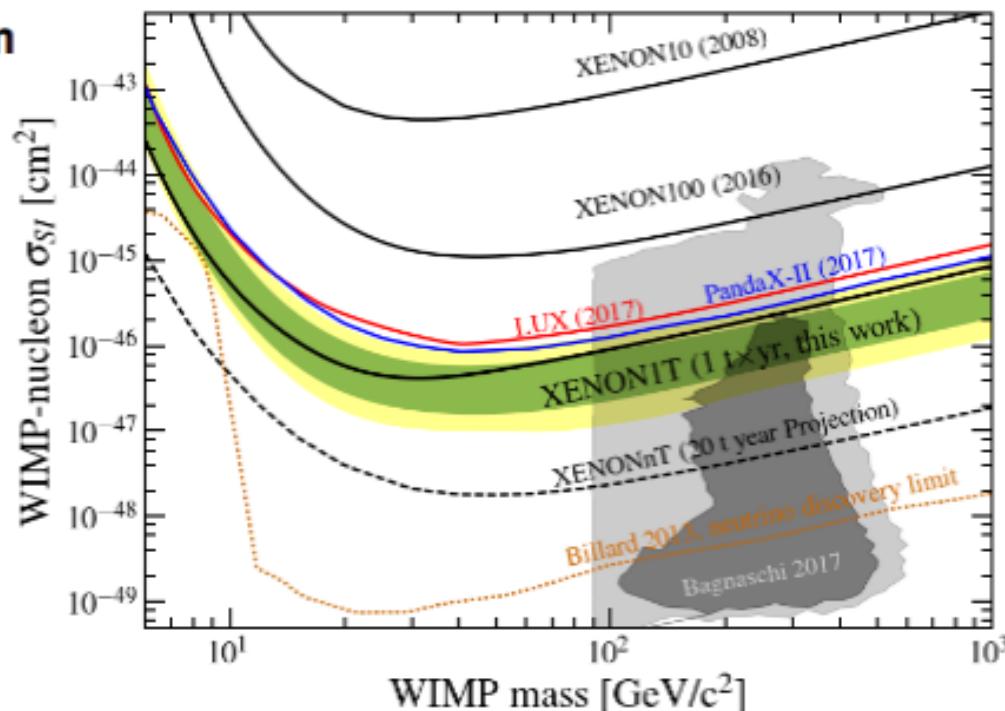
**Use most of
XENON1T subsystem**



**Increase fiducial
mass by factor 4**

To increase sensitivity by one order of magnitude:

$$\sigma_{SI} \sim 10^{-48} \text{ cm}^2$$

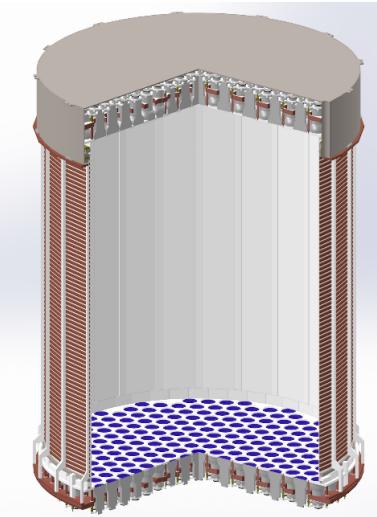


**Construction
in 2019**



**Reduce background
by factor 10**

New features XENONnT



NEW TPC

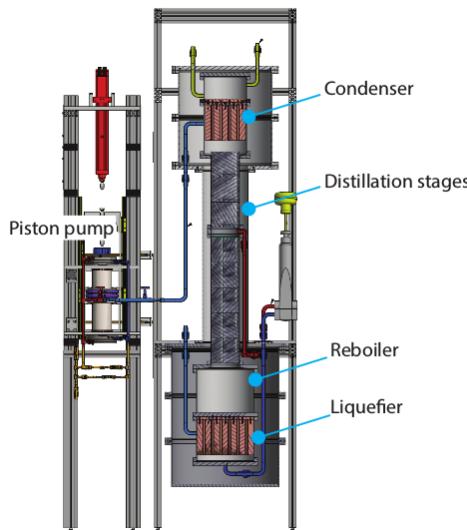
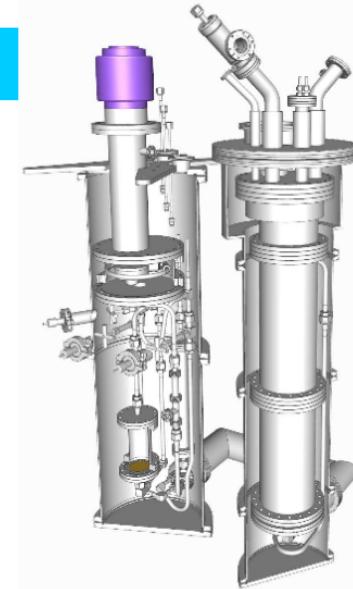
494 PMTs

1.5 m height
1.3 m diameter

LXE PURIFICATION

Much faster
purification
speed

Possible to purify
the 8 t of Xe in a
reasonable time



RADON DISTILLATION COLUMN

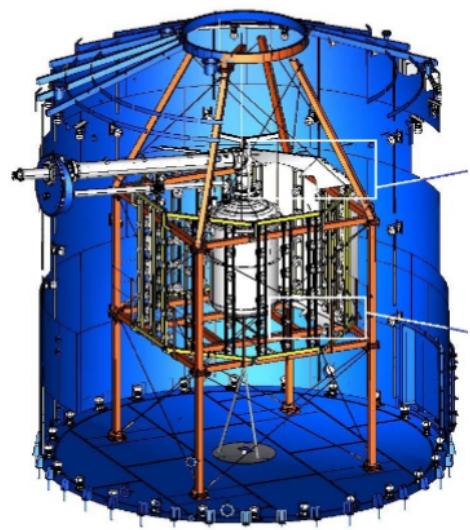
Goal 1 $\mu\text{Bq}/\text{kg}$
Rn contamination

Rn distillation
already tested
in XENON1T

NEUTRON VETO

0.2% Gd-doped water

120 additional
PMTs around
cryostat



Conclusions

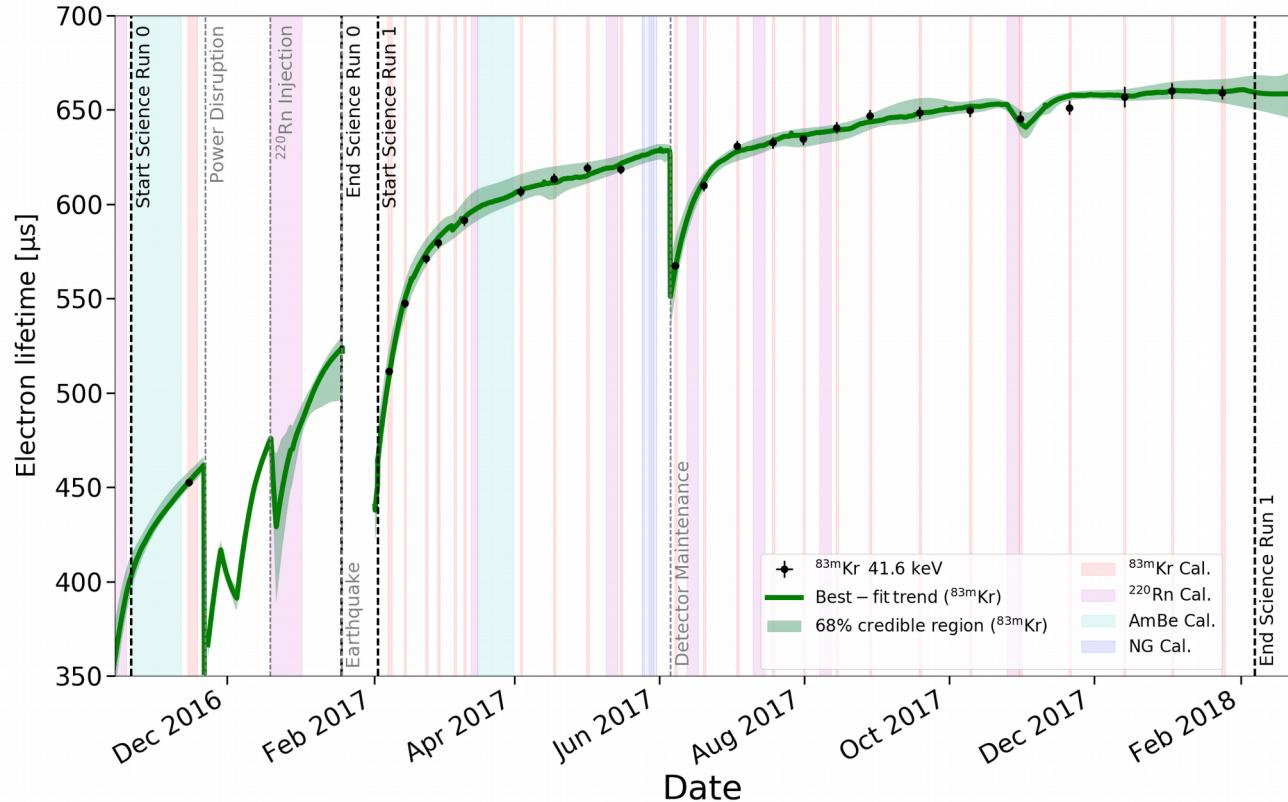
XENON1T reached 1 ton-year exposure with the lowest ER background for a dark matter detector

Most stringent limit for WIMP-nucleon SI cross section was set for WIMP masses greater than $6 \text{ GeV}/c^2$

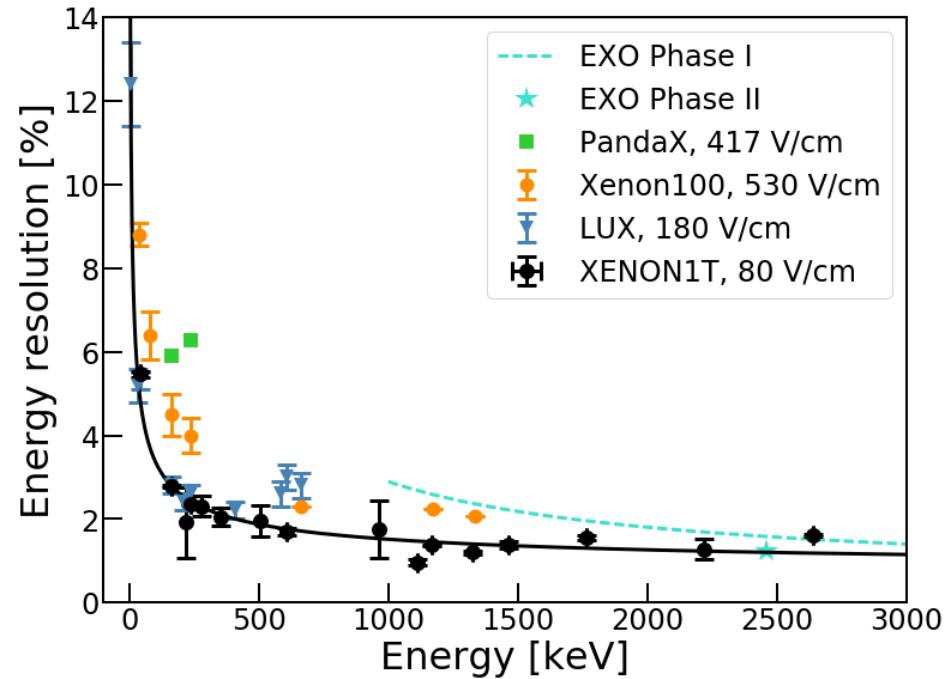
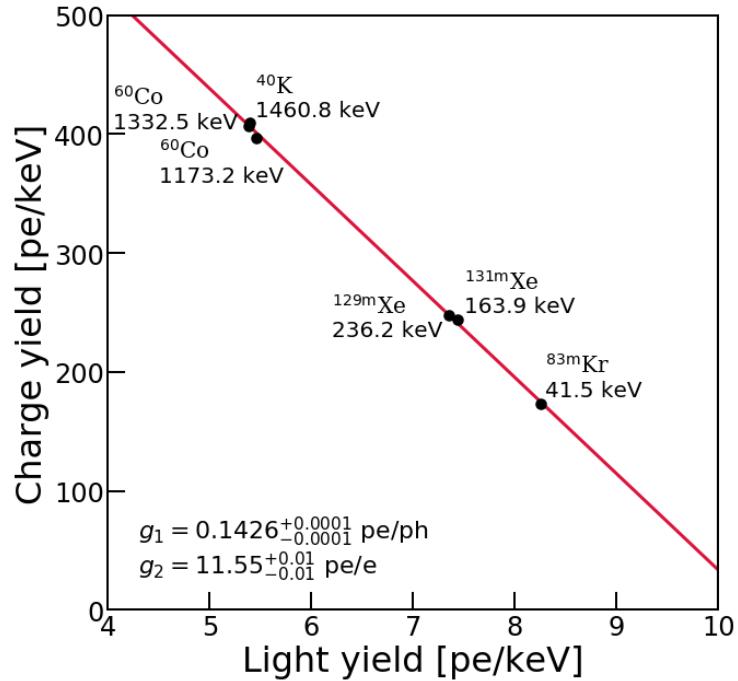
First detection of double electron capture of ^{124}Xe , longest half-life ever measured

Upgrade to XENONnT is ongoing, expected to start data taking by the end of 2019

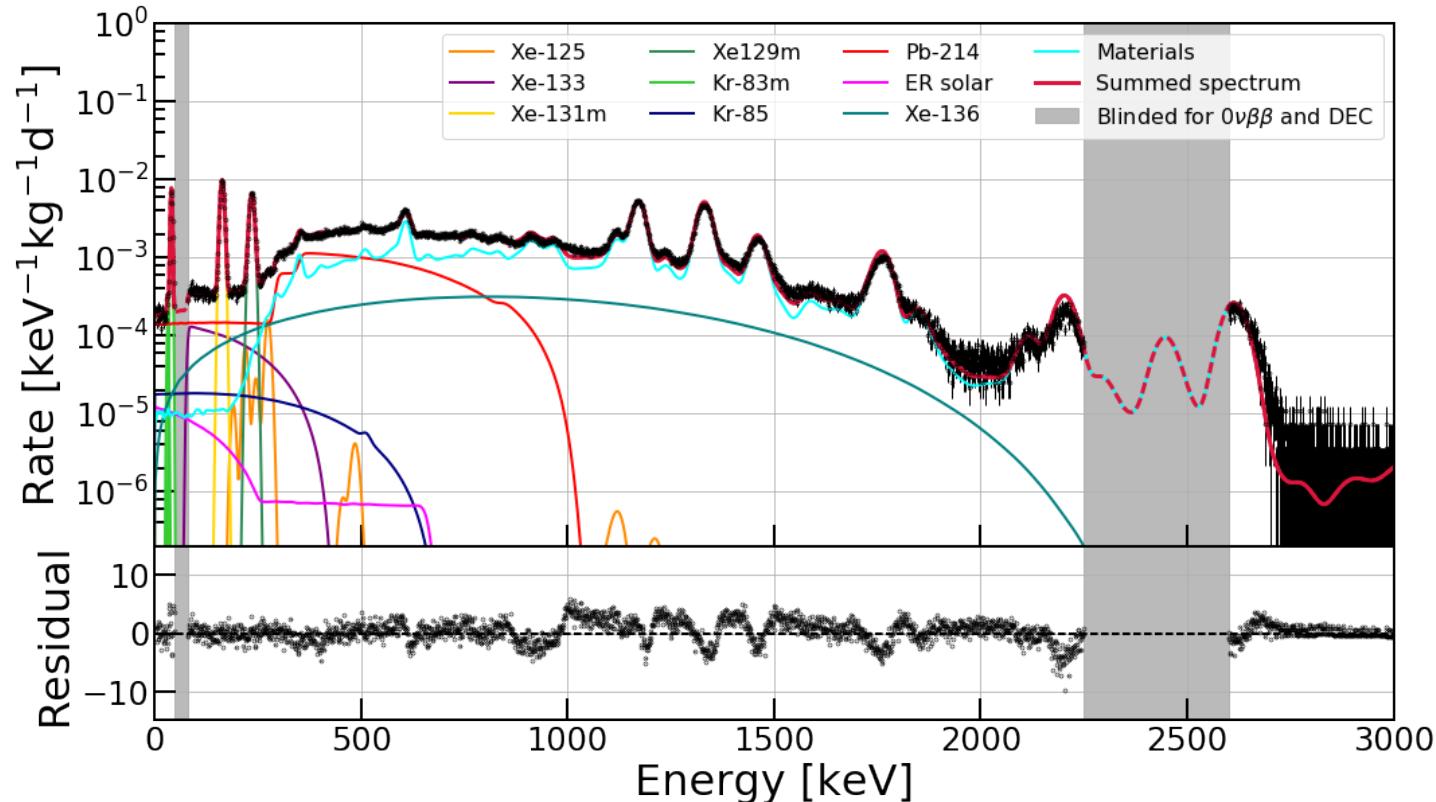
Calibrations (1)



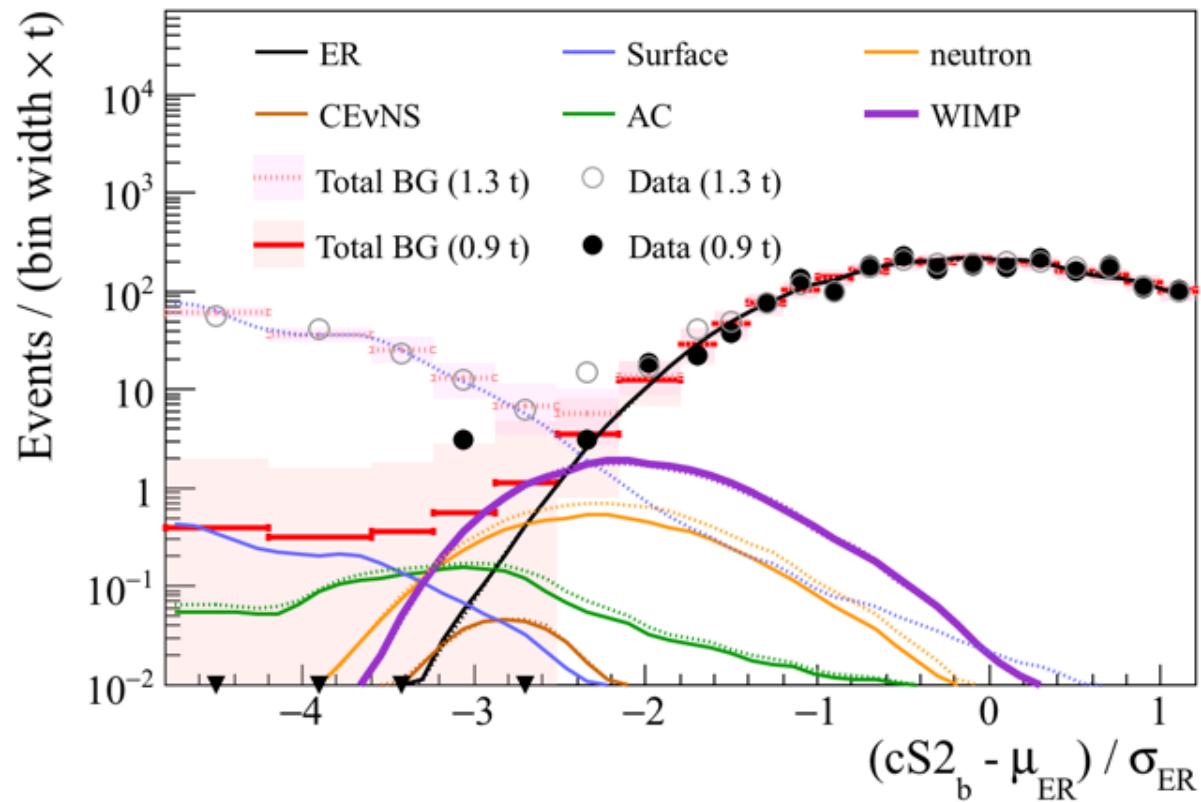
Calibrations (2)



Data – MC matching



SI-WIMP result



$0\nu\beta\beta$ decay

