



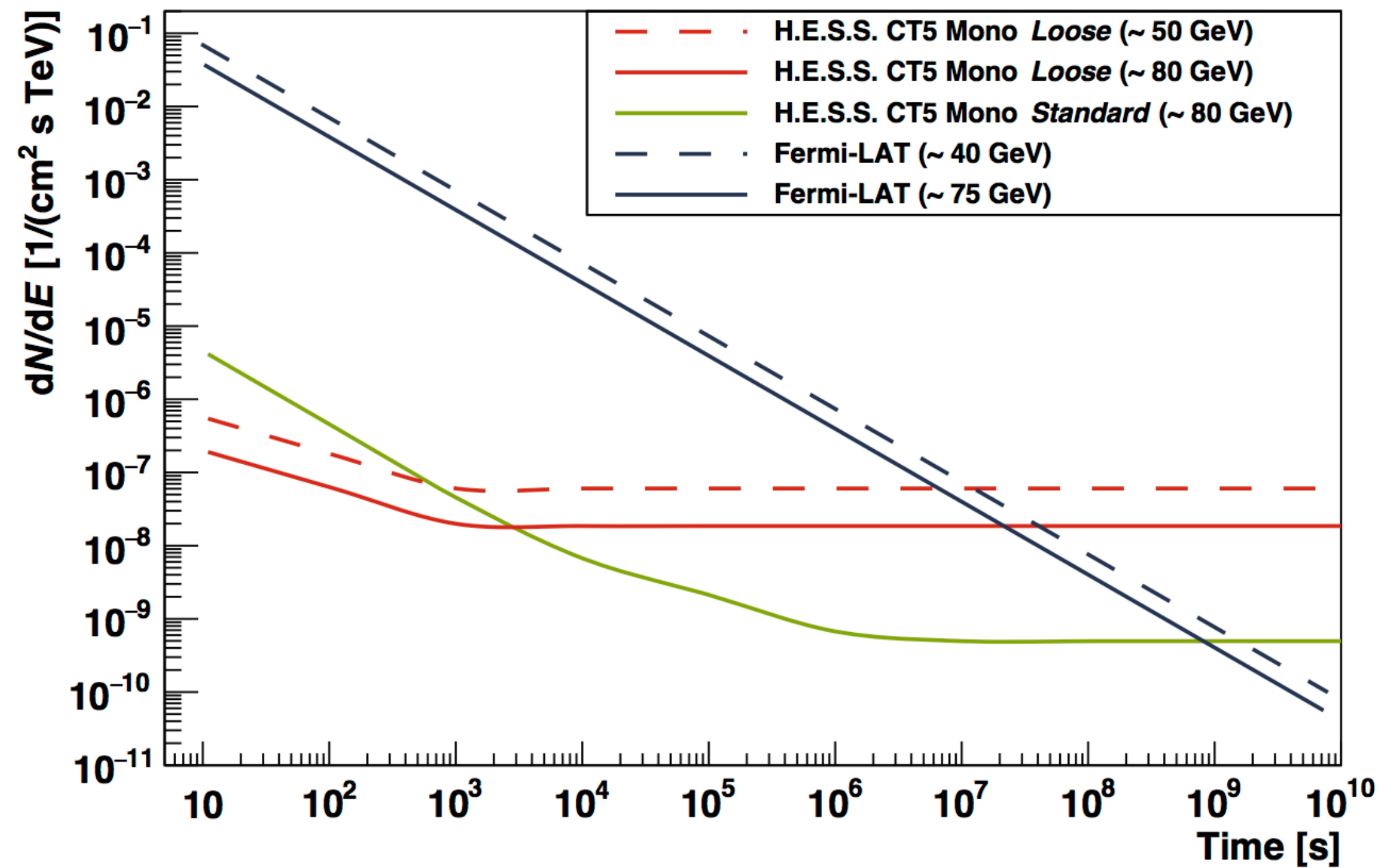
The H.E.S.S. Transients Follow-up System

Clemens Hoischen (Potsdam University),
Kathrin Egberts (Potsdam University) Matthias Füßling (CTAO), Stefan Ohm(DESY)



Transients and IACTs (Follow-up Instruments)

- Not that many known transient phenomena in the IACT energy range:
 - AGN flares & GRBs (since this year!)
 - Main objective: **Discoveries!**



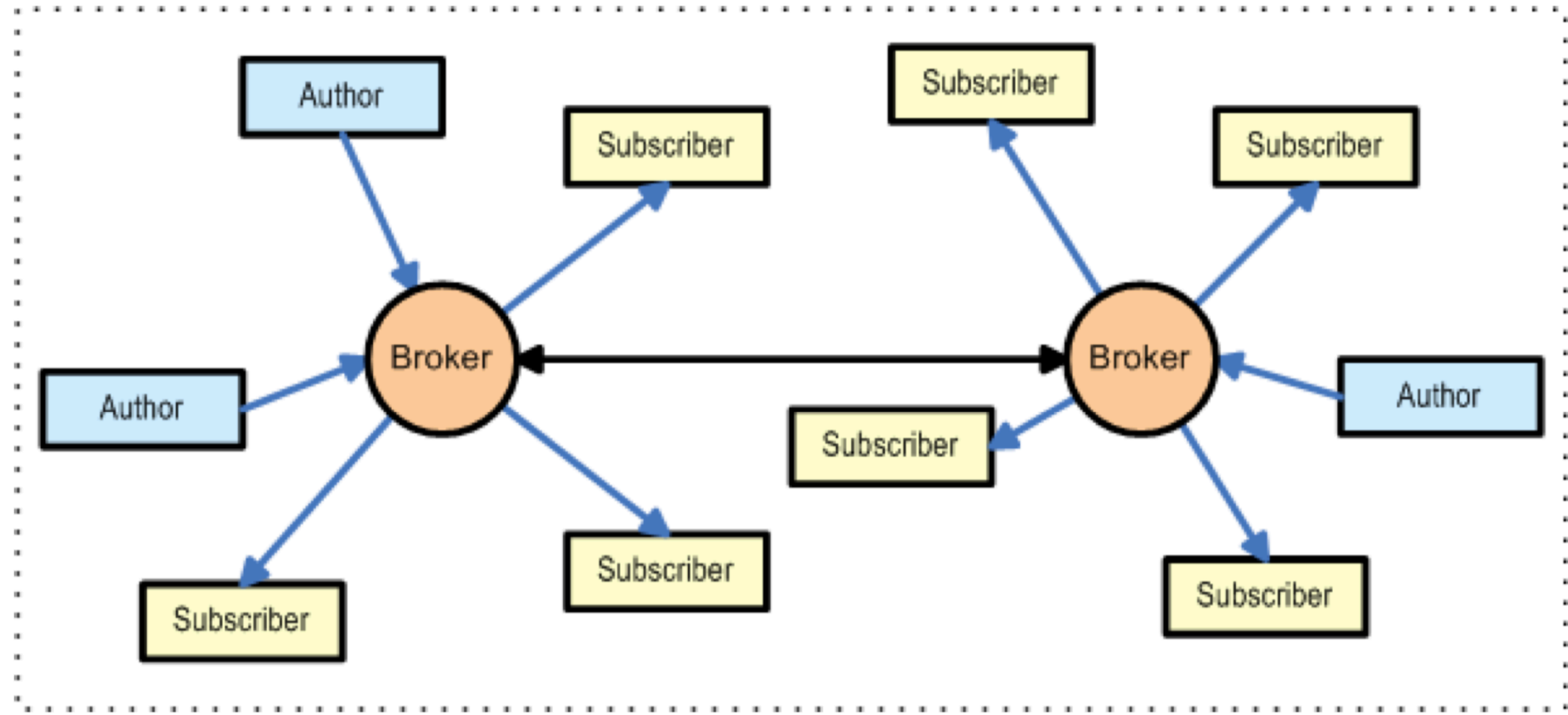
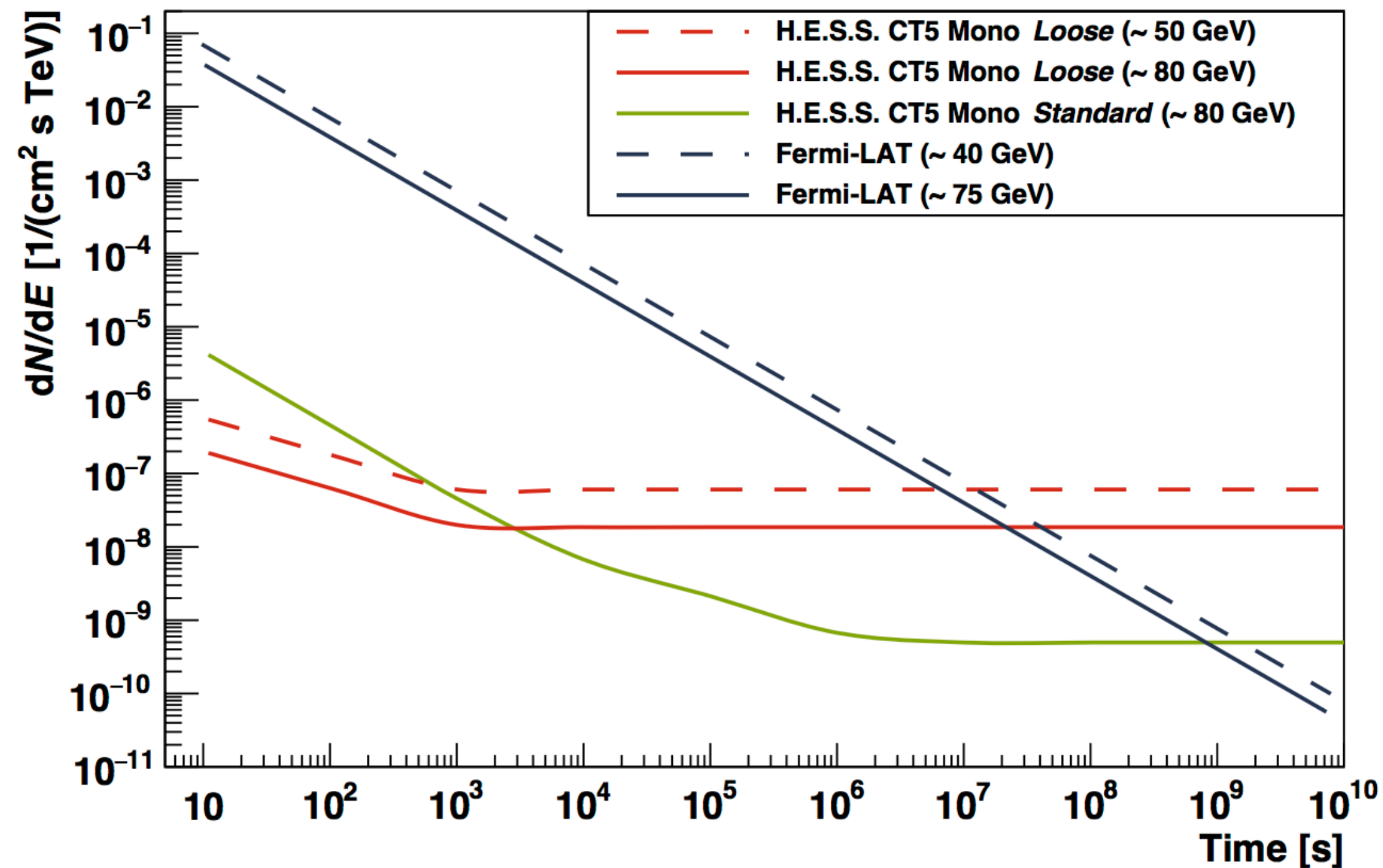
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- Largely dependent on external triggers

IVOA standards:

- VoEvent (alerts)
- VoEvent Transport Protocol (communication of alerts)

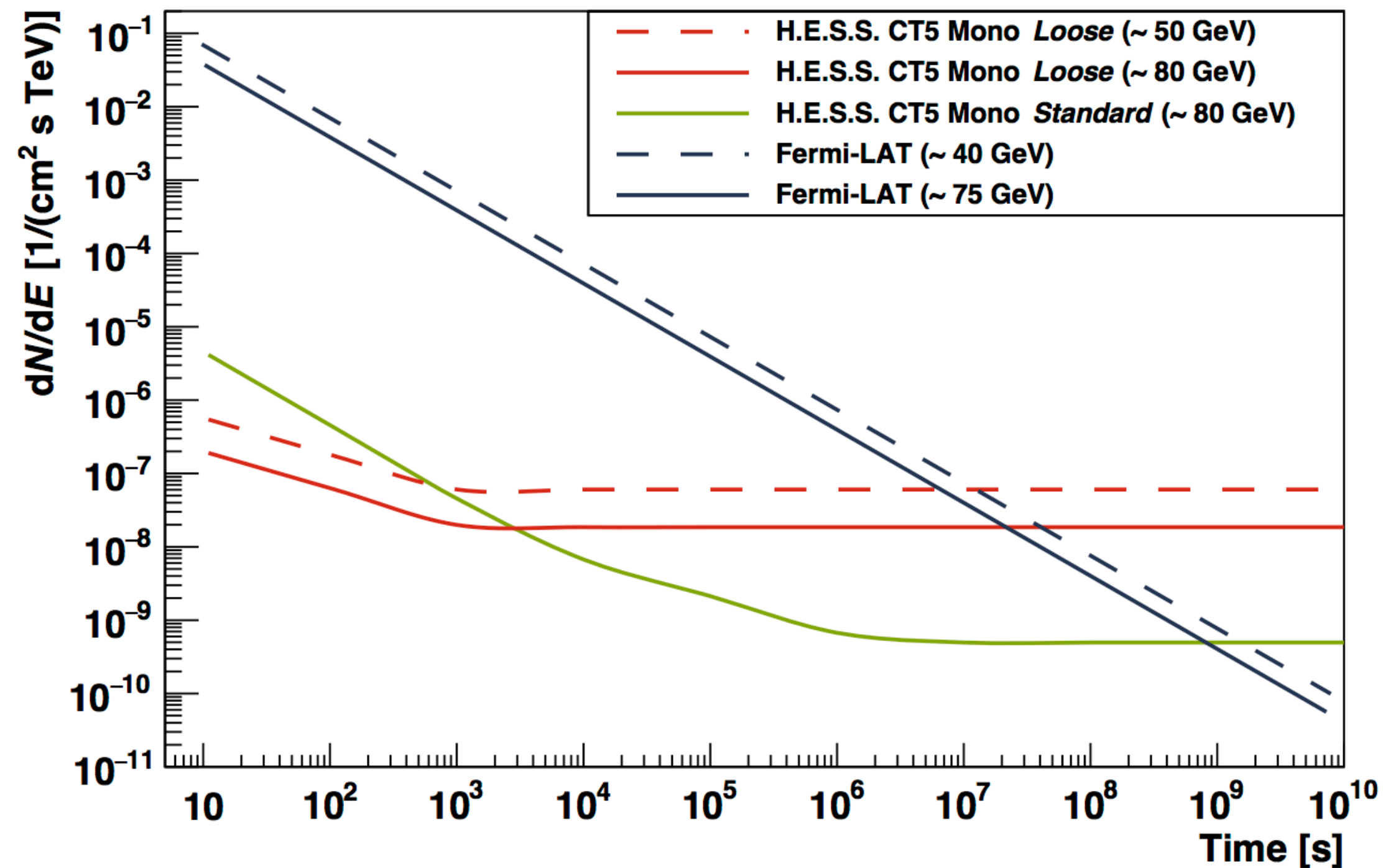


from VoEvent Transport Protocol v1.1 (2009)

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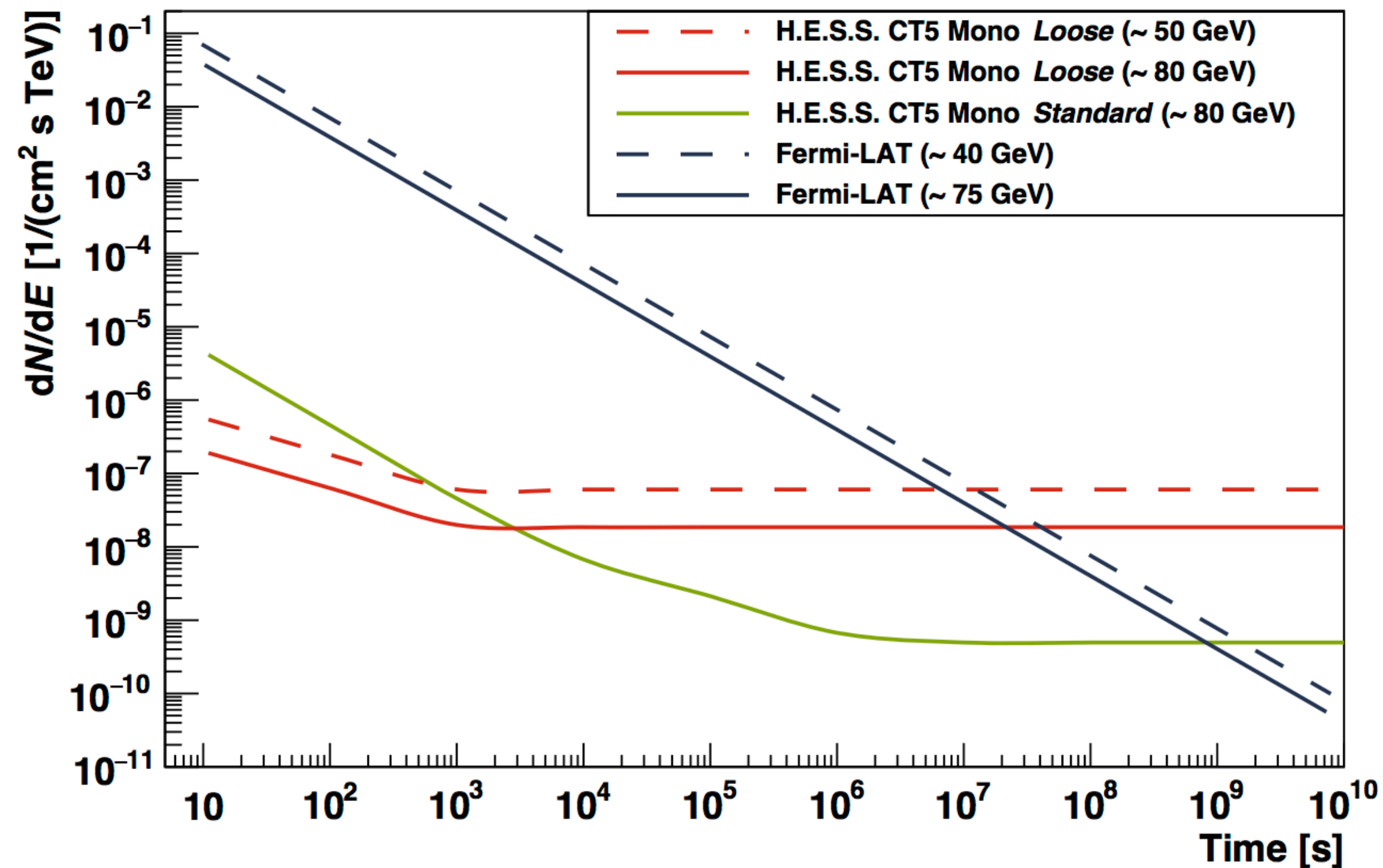
- Largely dependent on external triggers
 - Historically driven by GRB science
 - React as fast as possible
 - Solid decision making needed
 - Clear reaction schemes needed
 - Need to know about results as soon as possible



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First time detection of a GRB at sub-TeV energies; MAGIC detects the GRB 190114C

ATel #12390; *Razmik Mirzoyan on behalf of the MAGIC Collaboration*
on 15 Jan 2019; 01:03 UT
Credential Certification: Razmik Mirzoyan (Razmik.Mirzoyan@mpp.mpg.de)

High Energy Stereoscopic System - an Array of Imaging Air Cherenkov Telescopes



	CT1-4	CT5
# Telescopes	4	1
Mirror Area	107 m ²	614 m ²
PMTs	960	2048
Field of View	5 deg	3.2 deg
Rotation speed	100 deg/min	200 (100) deg/min
Energy threshold	~ 150 GeV	~ 50 GeV
Year of construction	2003	2012

Main Objectives of the System

- Receive and process high-level astrophysical alerts
 - standardised format (VoEvent)
 - matching to H.E.S.S. science cases
 - decision making (can include complex algorithms and scanning patterns)
- Initiation of the follow-up observations
 - controls array of telescopes as fast as possible
 - reaction details dependant on the science case and it's time-scale
 - changing the nominal observation schedule
- Provide Feedback
 - analysis results in real-time
 - decide between prolongation or abortion of the observations
 - alert experts, PIs and the community

Main Objectives of the System

Expert:
Clemens Hoischen
(Potsdam University)

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ToO Alert System

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Matthias Fülling
(CTAO)

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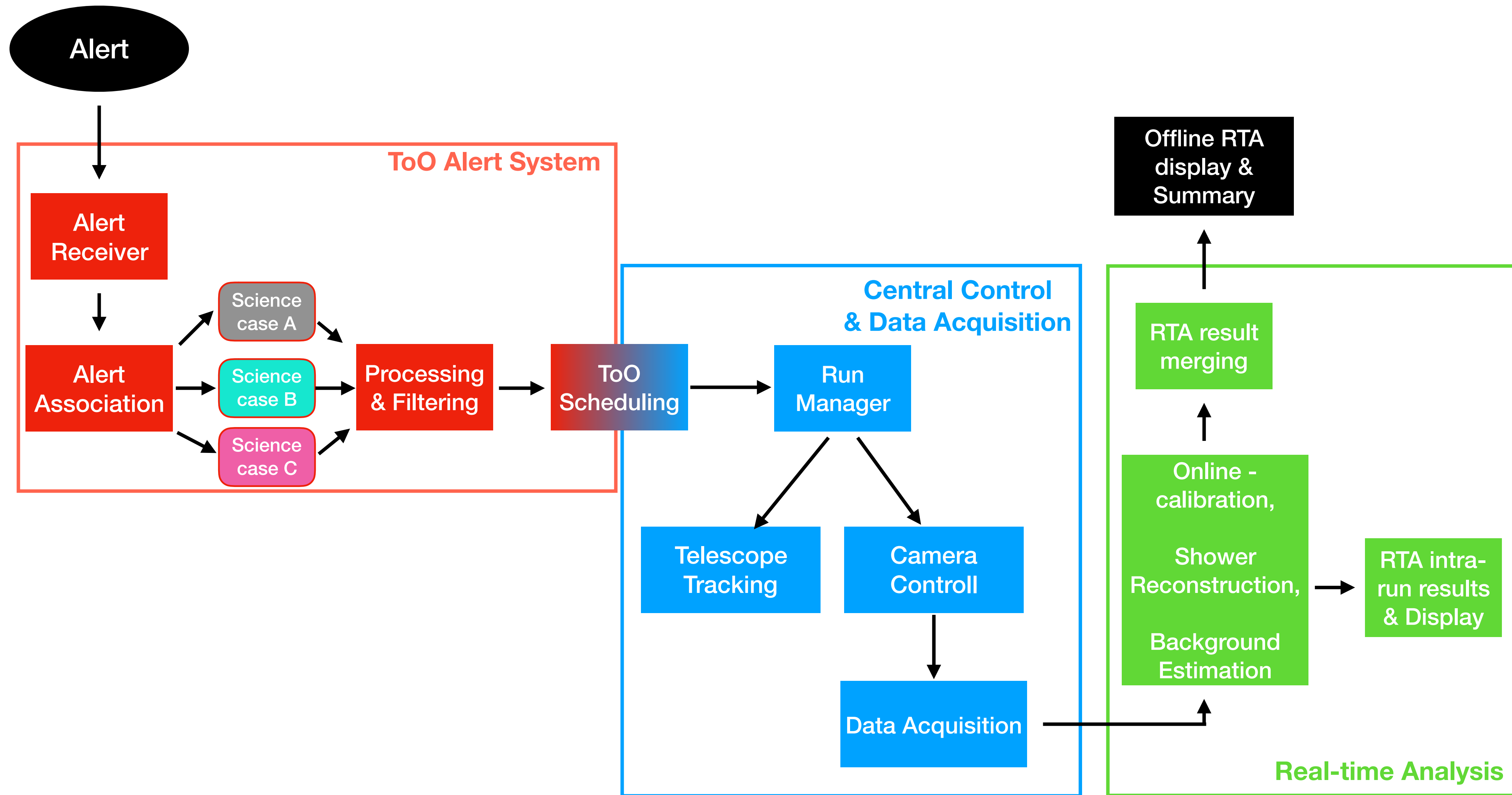
Central Control & Data Acquisition

Expert:
Stefan Ohm
(DESY)

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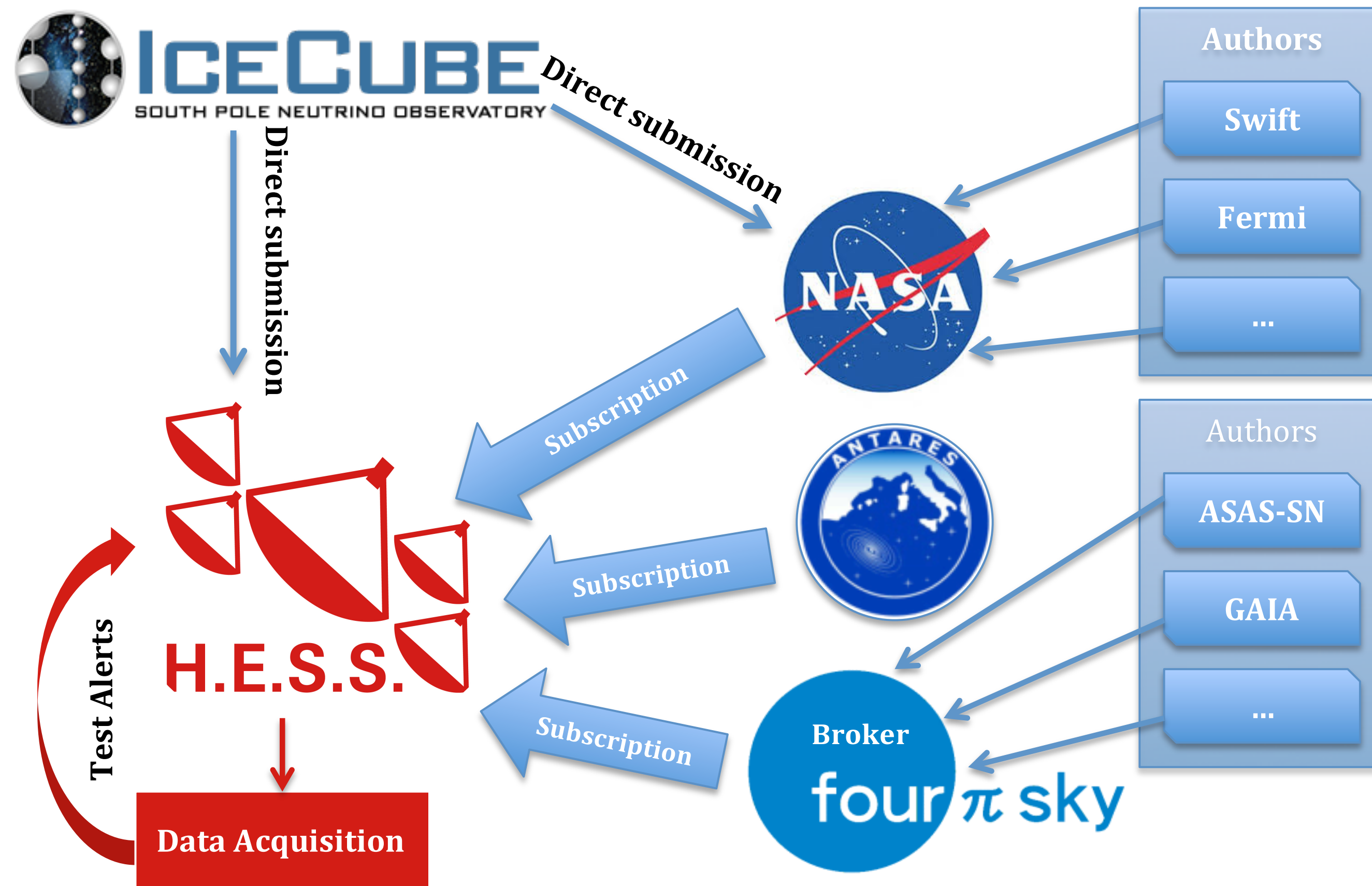
Real-time Analysis

The H.E.S.S. Transients Follow-up System - Overview



Receiving Alerts

- Based on *comet*¹ broker - implementation of VoEvent Transport Protocol with many helpful features
- Subscribed to GCN, 4piSky, Antares and whitelisted IceCube IPs for direct submission of alerts
- Broker is monitored with *monit*² to guarantee uptime of the receiver.



Alert statistics per month

- Receives ~ 50k alerts
- Stores ~18k
- 1.5 k non-test alerts
- ~ 50 alerts of interest for H.E.S.S.

¹ see comet.readthedocs.io/

² monit is a watchdog which automatically restarts a process if it crashes.

Matching Alerts to Science Cases

- Initial parsing of alerts
 - identifying the experiment, alert type, ... done with voevent-parse¹.
 - identification of the alert type is driven by the IVORN (unique alert identifier)

■ A single incoming alert will be matched with all science case configurations which are related to this alert.

- Processed in prioritised order (proposal grade & urgency of timely follow-up)
- Each science case has its own configuration file
- Easy to register additional science cases with a new configuration

Alert Type	Matched Science Cases
LAT_Updated_Pos	GRB_prompt, GRB_afterglow
LAT_Pos_Gnd	GRB_prompt
LAT_Offline_Pos	GRB_afterglow
BAT_GRB_Pos	GRB_prompt, SGRAXP_prompt, GRB_afterglow, GRB...
gwnet_LVC_#S_Preliminary	GravitationalWave
gwnet_LVC_#S_Initial	GravitationalWave
gwnet_LVC_#S_Update	GravitationalWave, GravitationalWave
GBM_Gnd_Pos	GRB_prompt, GRB_afterglow_long, GRB_afterglow
GBM_Fin_Pos	GRB_prompt, GRB_afterglow
IceCube_HESE#	Neutrino_prompt, Neutrino_afterglow
ICECUBE_EHE_AMON	Neutrino_prompt, Neutrino_afterglow
IceCube_GFU	Neutrino_GFU_Prompt, Neutrino_GFU_Afterglow,
gwnet_LVC_#MS_Preliminary	GravitationalWave_Test, GravitationalWave_Test
gwnet_LVC_#MS_Initial	GravitationalWave_Test, GravitationalWave_Test
MAXI_Unknown	Flaring_Star_All, Flaring_Star_Prompt, Flarin...
MAXI_Known	Flaring_Star_Afterglow, Flaring_Star_Prompt, ...
HESS_FireDrill	FireDrill

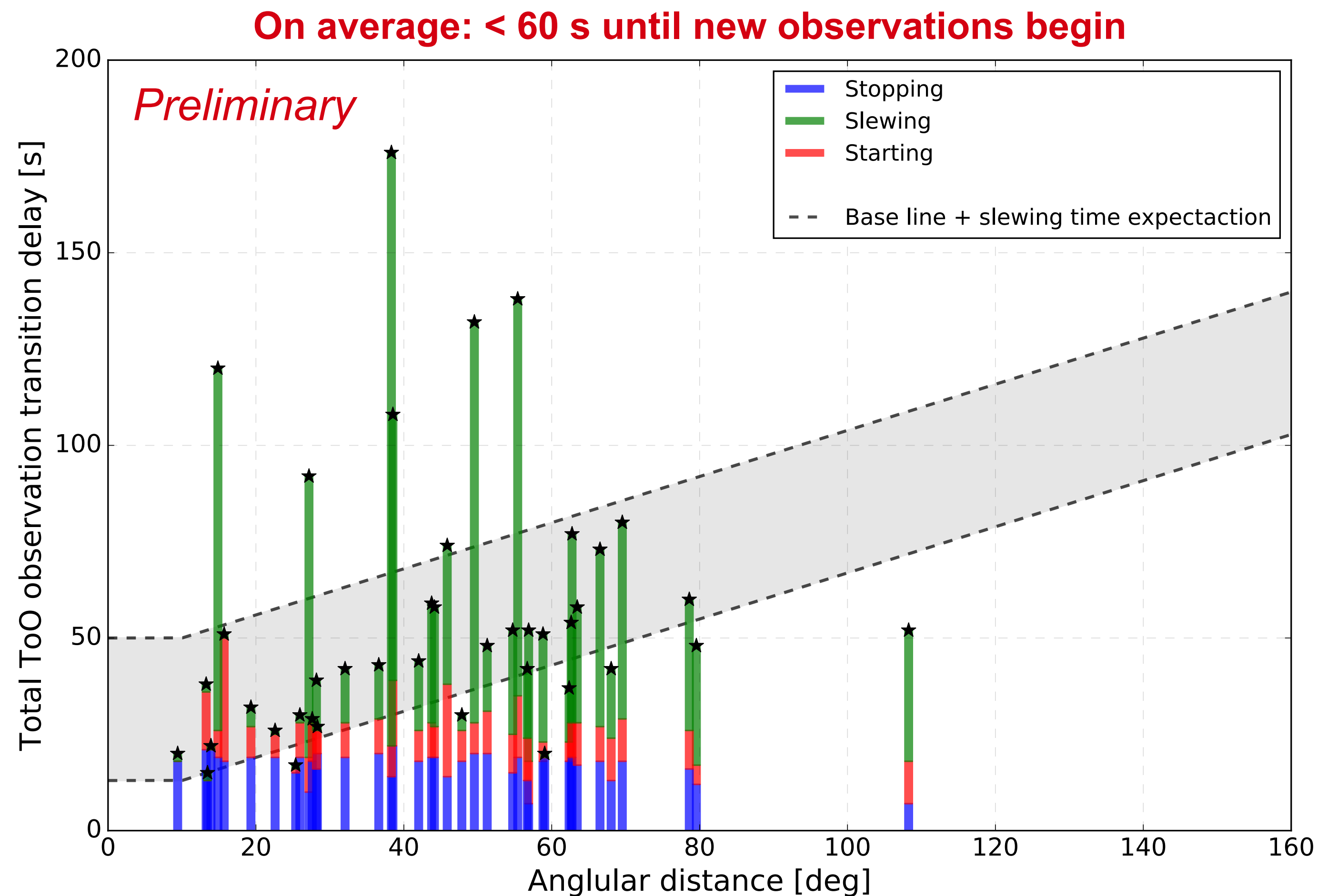
¹ see <https://voevent-parse.readthedocs.io/en/stable/#>

Scheduling ToO Observations

- Two main modes:
 - **Prompt:** if the position of the ToO is visible within the next 10 minutes.
 - Fully automatic reaction
 - Always takes precedence over nominal schedule
 - transition between observations with a **special ToO mode**
 - Shifters are alerted by sound, pop-ups and email (including instructions)
 - Experts and PIs are alerted by email.

Also handle corner cases:

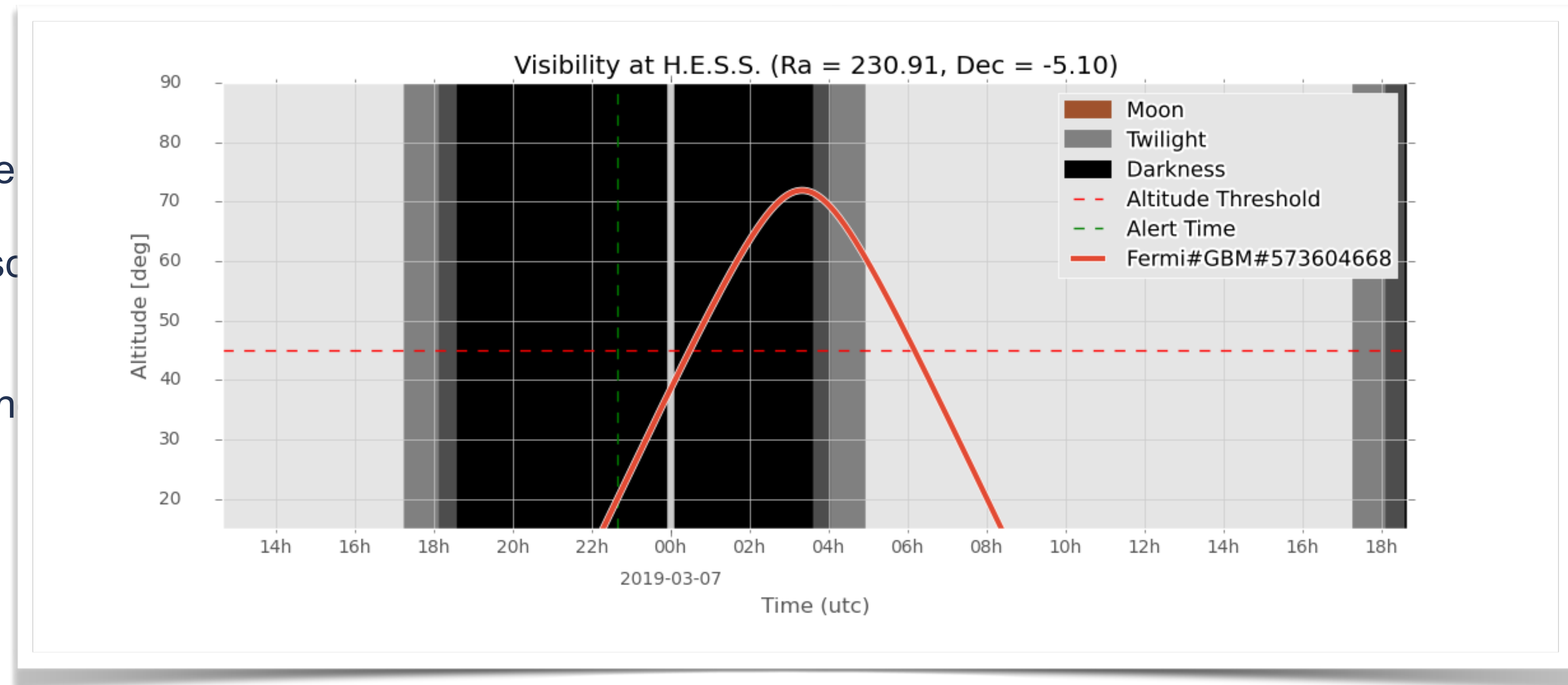
- Coordinate updates
- new alerts with higher priority
- alert position is within the current FoV



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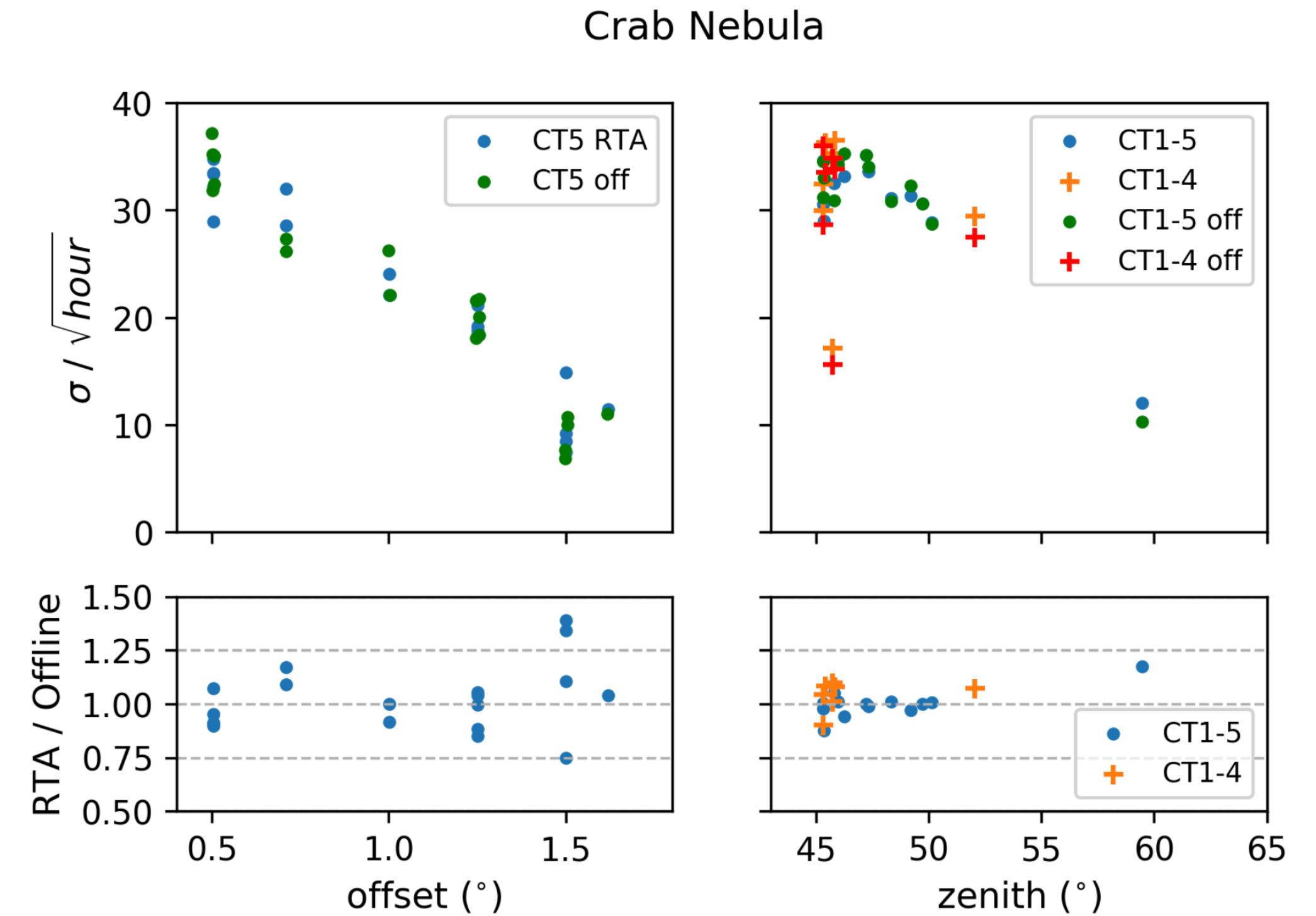
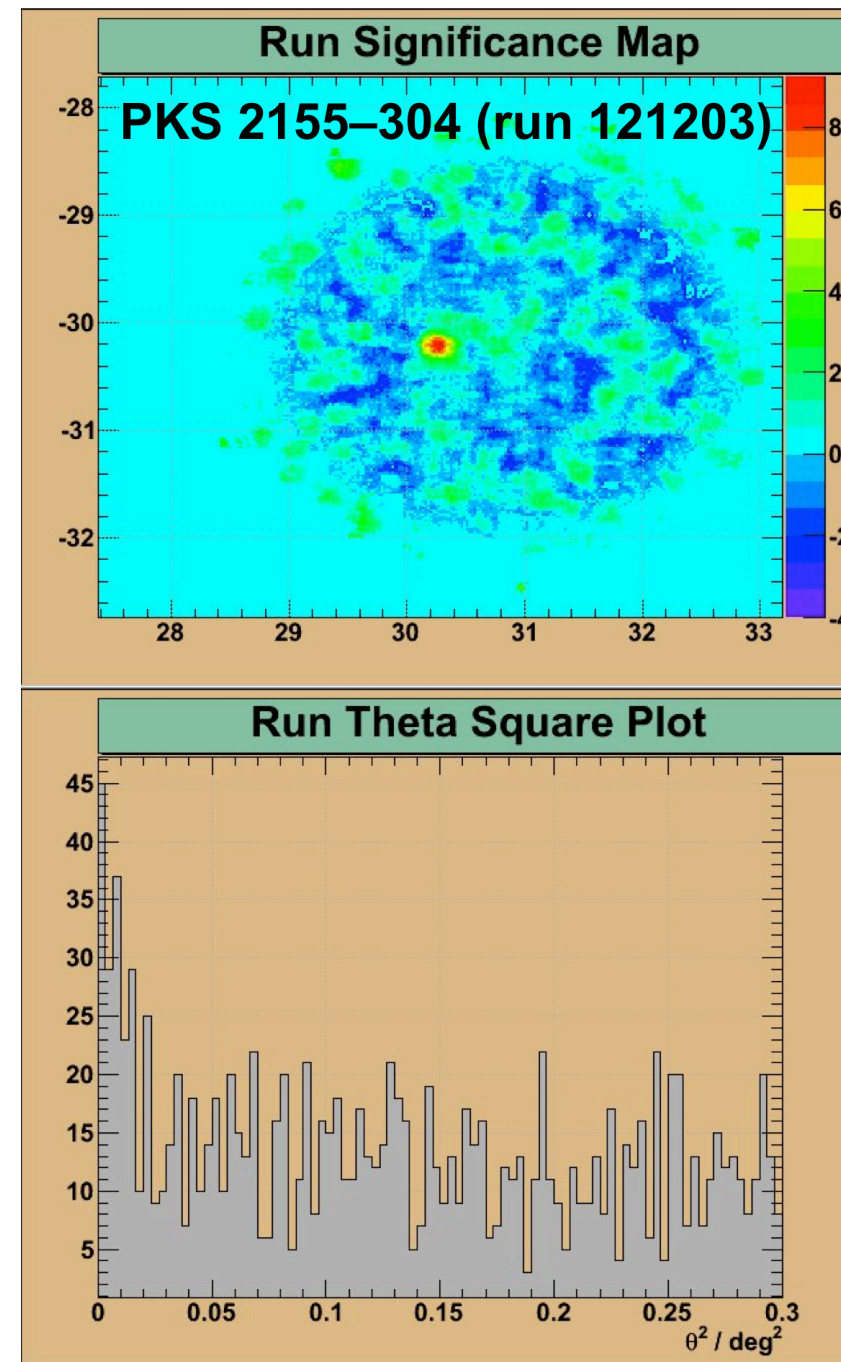


■ **Afterglow:** observation starts later than 10 minutes after the event

- Shifters are alerted by pop-ups, sound and emails (including instructions)
- Experts and PIs are alerted by email.
- Observations are initiated with the help of scripts (given in the instructions) by the shifters.
- Allows to Consult PIs and Experts on-call by phone.
- Also applies to alerts during the day.

Real Time Analysis

- Simplified live-calibration scheme
- Image reconstruction in different operation modes (mono and stereo)
- Background subtraction
- Shows live:
 - Θ^2 histograms
 - Sky maps
- Merging of consecutive observations
- Entered in a database
 - allows to review results from the last night(s)
- Used to decide if observations should be prolonged
 - during the same night (e.g. AGN monitoring)
 - for the next night(s)
- Used to inform the community quickly (currently via ATels)



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H.E.S.S. detection of a strong VHE activity from the blazar 3C 279

ATel #11239; **Mathieu de Naurois for the H. E.S. S. collaboration**
on 28 Jan 2018; 15:00 UT

Credential Certification: Jean-Philippe Lenain (jlenain@in2p3.fr)

Subjects: Gamma Ray, VHE, Request for Observations, AGN, Blazar, Quasar

Referred to by ATel #: 11246, 11464, 11680

H.E.S.S. and ATOM detection of renewed activity of the FSRQ 3C 279

ATel #11680; **Mathieu de Naurois for the H. E.S. S. Collaboration**
on 4 Jun 2018; 14:05 UT

Credential Certification: Michael Zacharias (m.zacharias@lsw.uni-heidelberg.de)

Subjects: Optical, Gamma Ray, >GeV, VHE, Request for Observations, AGN, Black Hole, Blazar, Quasar

Referred to by ATel #: 11687

H.E.S.S. follow-up of IceCube-170922A

ATel #10787; **Mathieu de Naurois for the H. E.S. S. collaboration**
on 27 Sep 2017; 14:33 UT

Credential Certification: Fabian Schüssler (fabian.schussler@cea.fr)

Subjects: VHE, Neutrinos

Referred to by ATel #: 10799, 10817, 10830, 10833, 10844, 11419

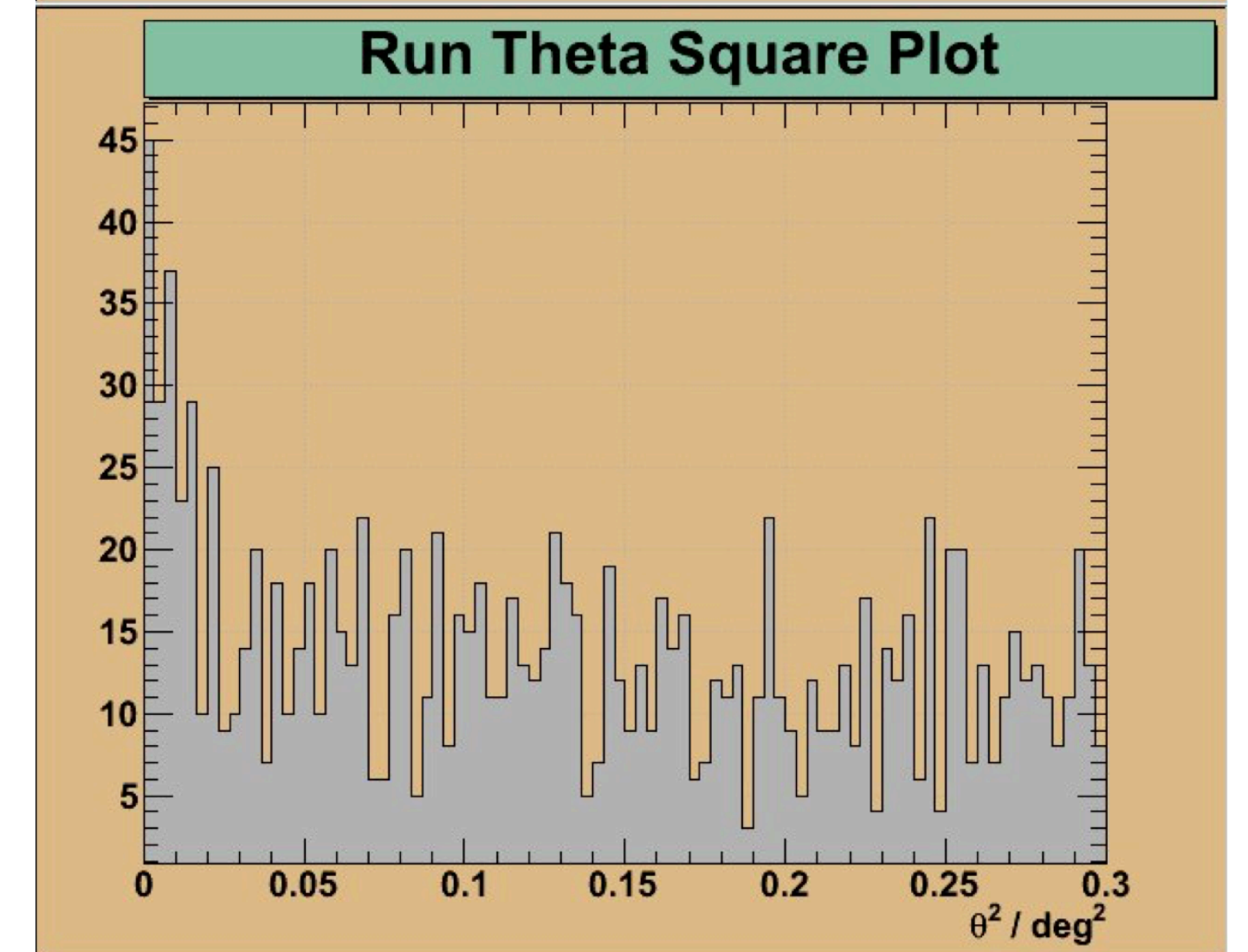
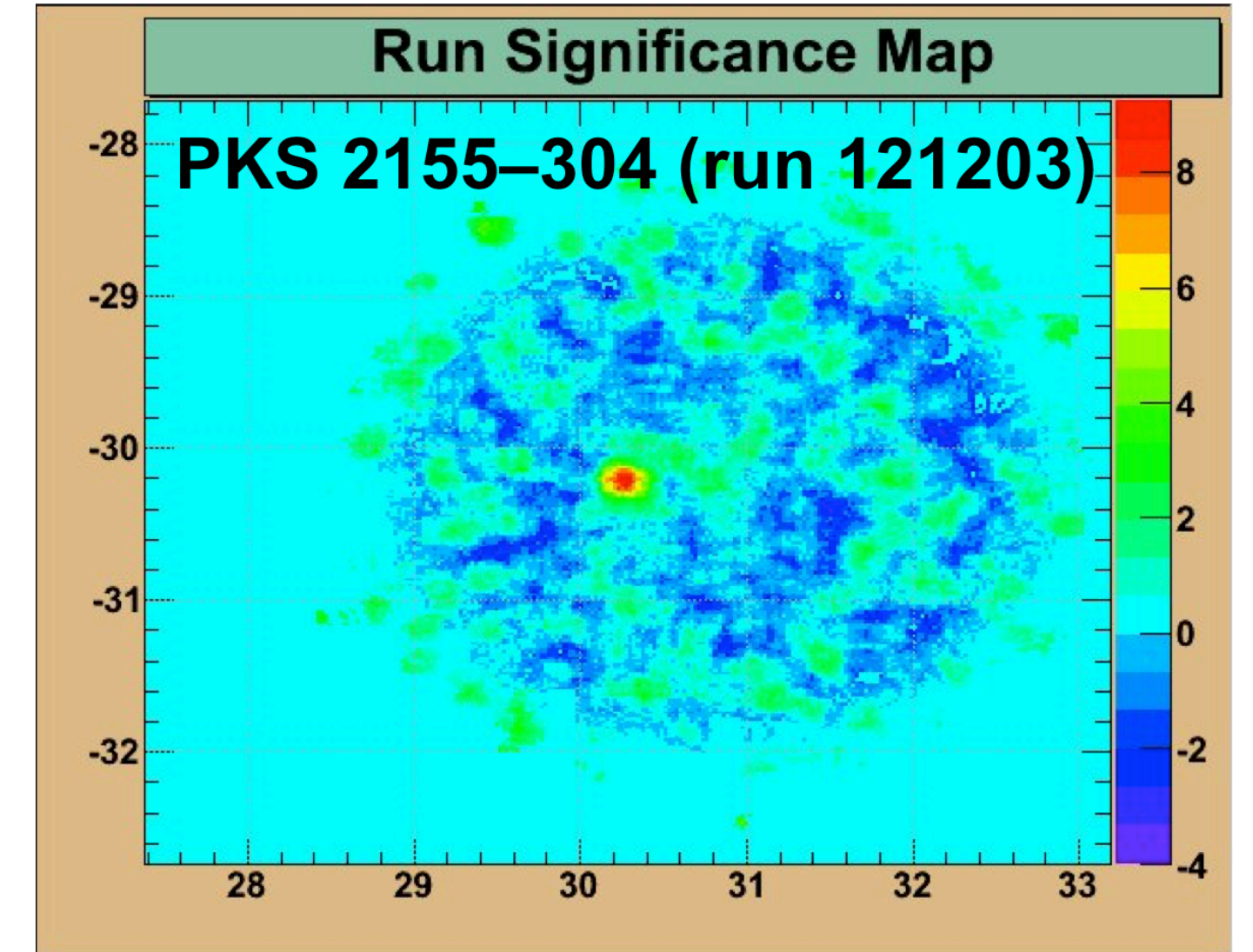
Link between Transients Alert System and Real Time Analysis

■ Currently:

- manually via the shifters monitoring the significance plots
- Thresholds defined by PIs based on Flux estimates

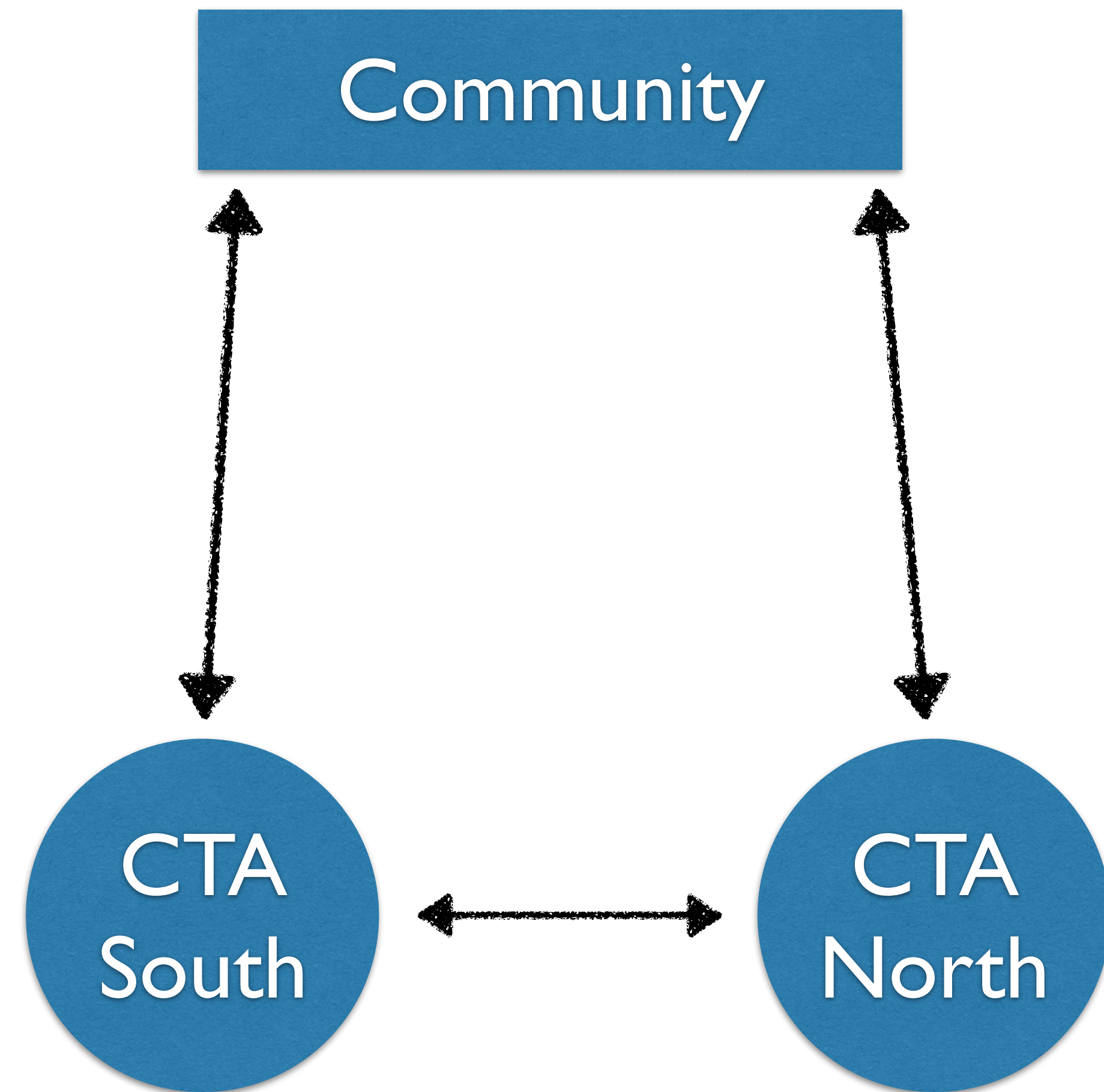
■ Anticipated

- Generate VOEvent for each observation run
- Inject it into the Transients Alert System
- Decide on follow-up observations depending on criteria for known and unknown sources



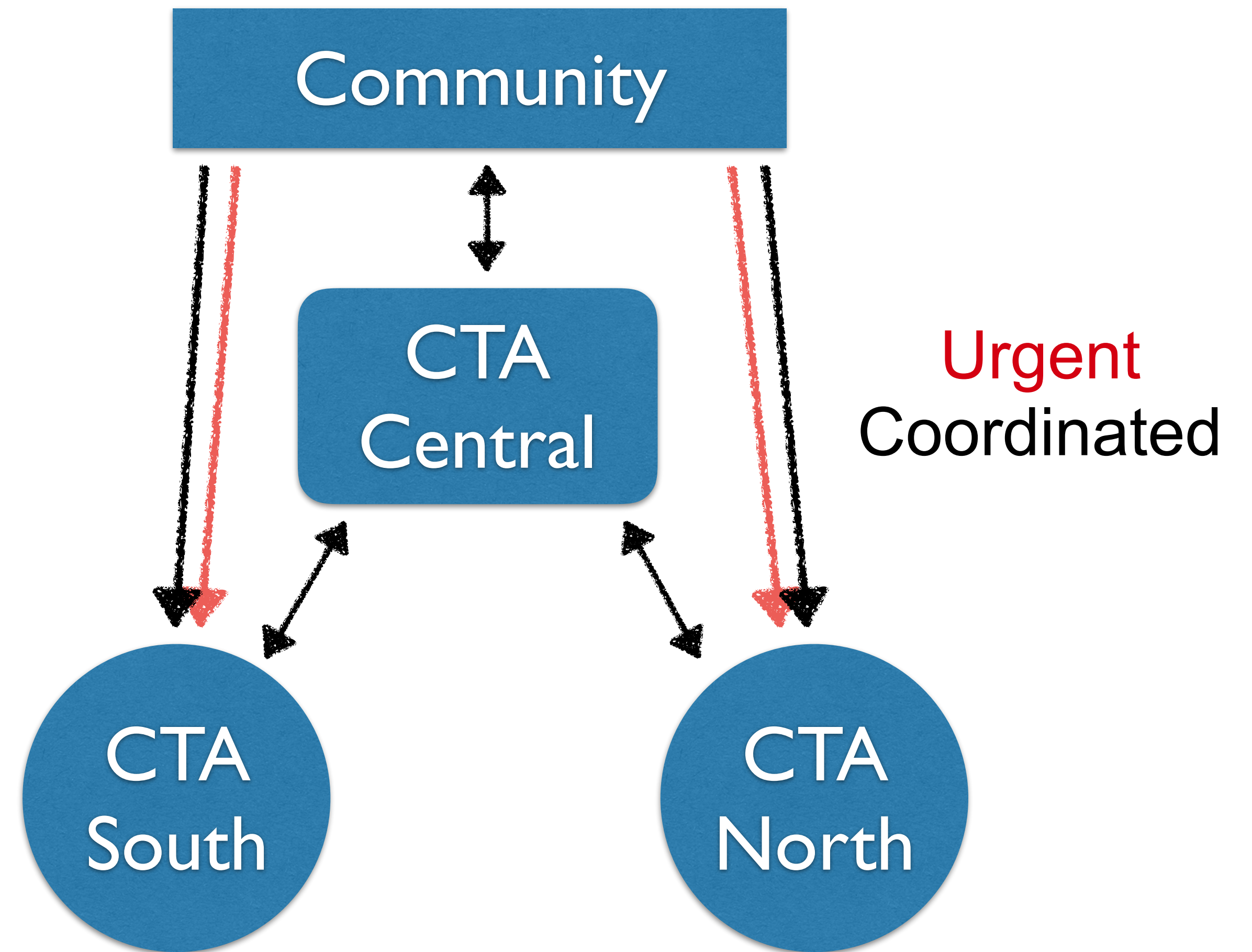
Differences in light of CTA (my view)

- Observation-wise
 - Up to 8 sub-arrays (at the same time)
 - Two sites which (in some cases) need coordination



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- Observation-wise
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 - Two sites which (in some cases) need coordination
- Observatory-wise
 - Open Proposals
 - Data Access/Policy
 - Schedule Policies (open or not)
 - Configurability for a TAC
- Realtime Analysis
 - Currently going for discrete alerts
 - Also affected by Data Policies
- Correlation of data streams
 - Difficult to imagine inside of CTA
 - Clearly interesting as external service
 - Currently not anticipated by the RTA team (afaik)
 - Depends on Data policies

