

The local universe, the search for life, and the Physics Nobel Prize 2019

to J. Peebles, and M. Mayor & D. Queloz

*“for contributions to our understanding of the evolution of the universe
and Earth’s place in the cosmos”*

Ansgar Reiners

Georg-August-Universität Göttingen



DESY Colloquium



NASA, Apollo 4, Nov 9, 1967



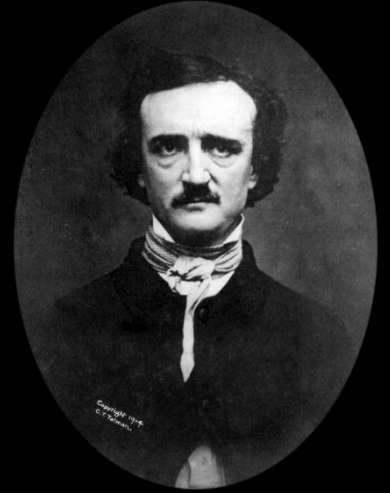
NASA, Apollo 8, Dec 24, 1968



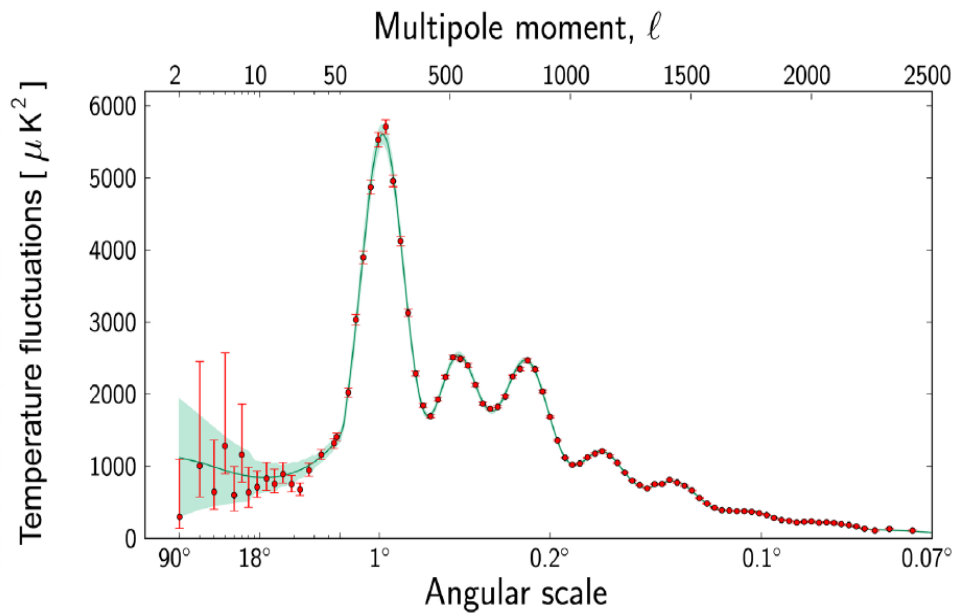
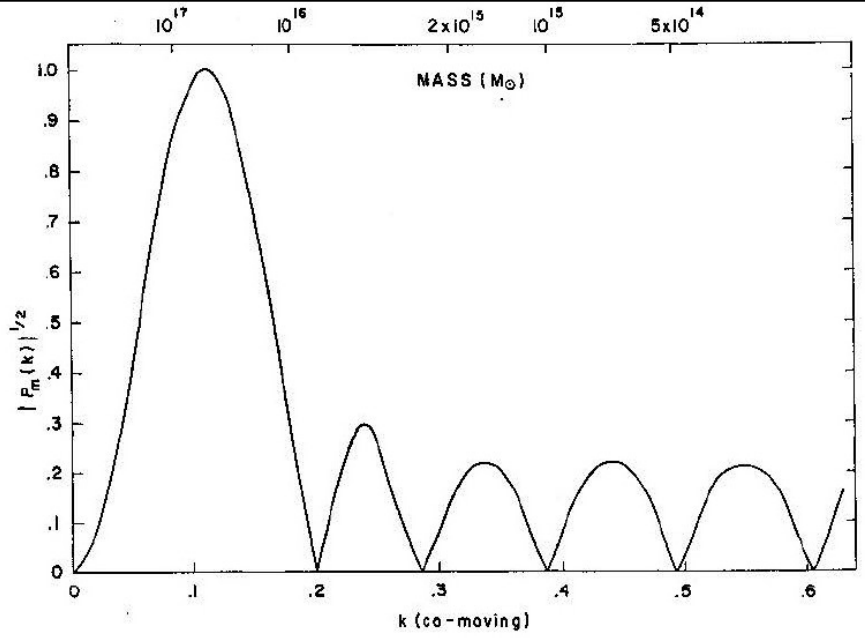
NASA, Apollo 8, Dec 24, 1968

Edgar Allan Poe, *Eureka – A Prose Poem*, 1848:

Let us now endeavor to conceive what Matter must be, when, or if, in its absolute extreme of Simplicity. Here the Reason flies at once to Imparticularity – to a particle – to one particle – a particle of one kind – of one character – of one nature – of one size – of one form – a particle, therefore, "without form and void" – a particle positively a particle at all points – a particle absolutely unique, individual, undivided, and not indivisible [...]

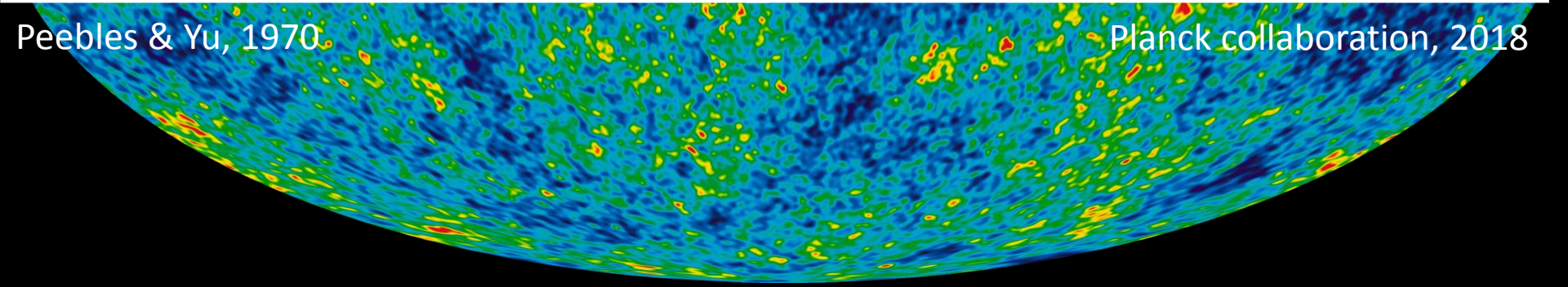


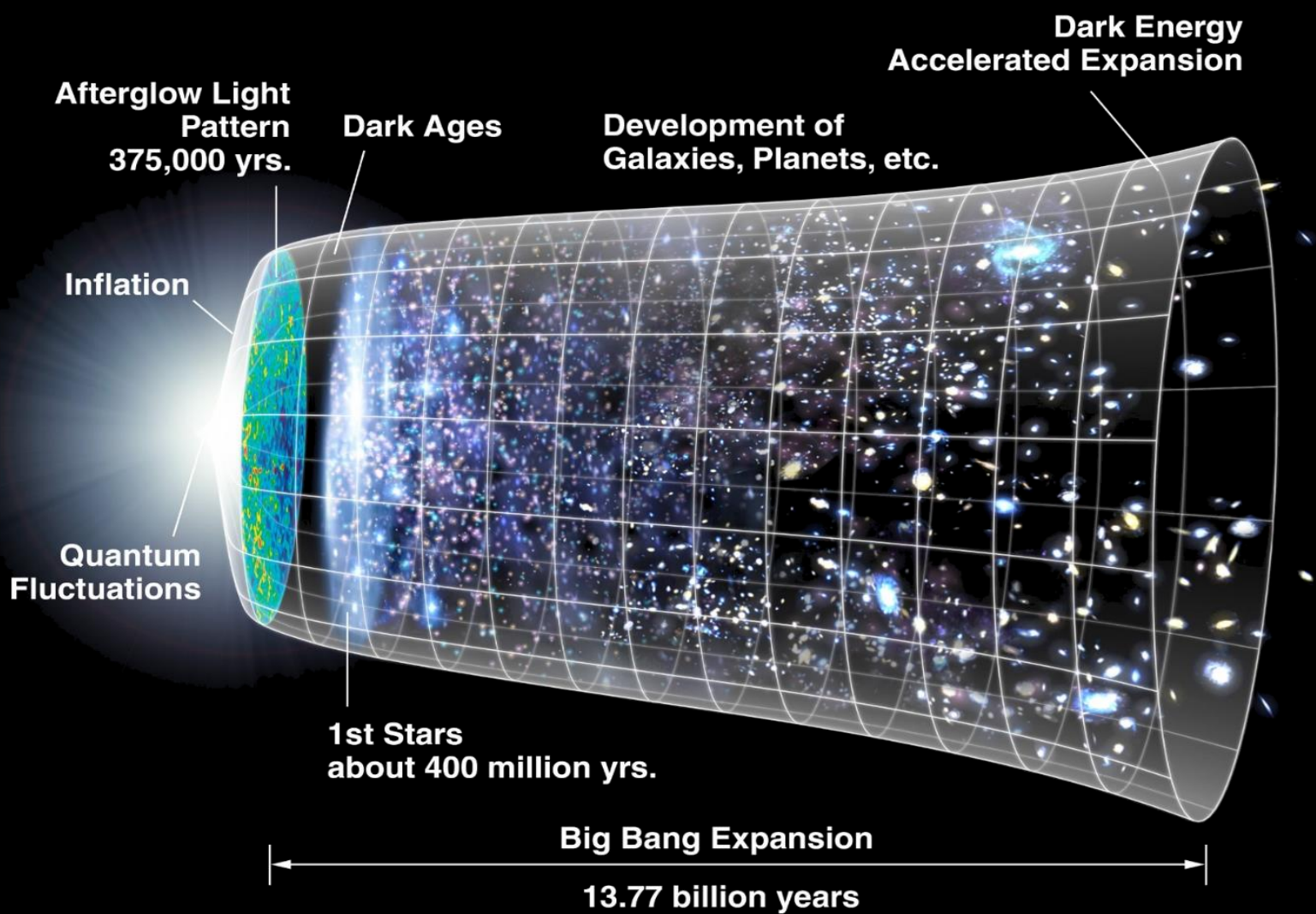
The Primordial Particle



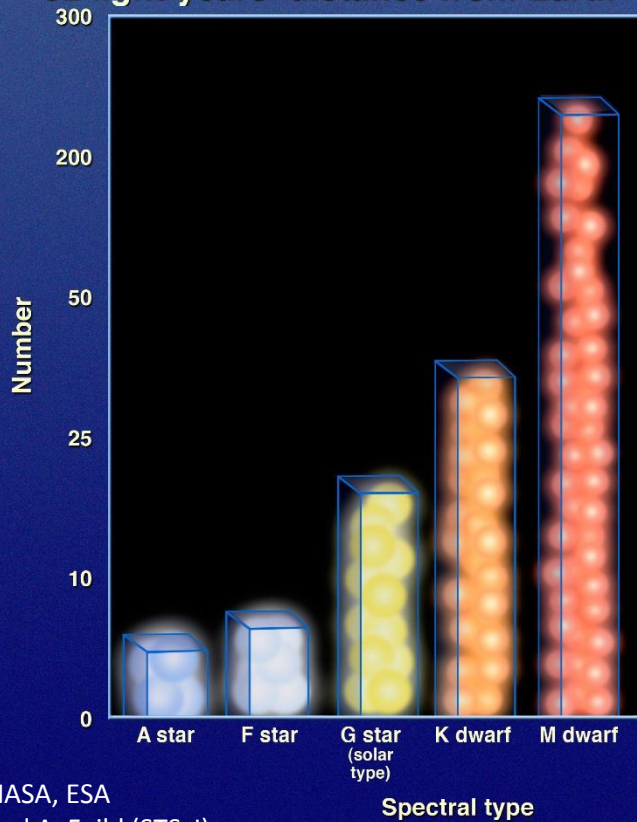
Peebles & Yu, 1970

Planck collaboration, 2018





An inventory of stars within 32 light-years' distance from Earth



NASA, ESA
and A. Feild (STScI)



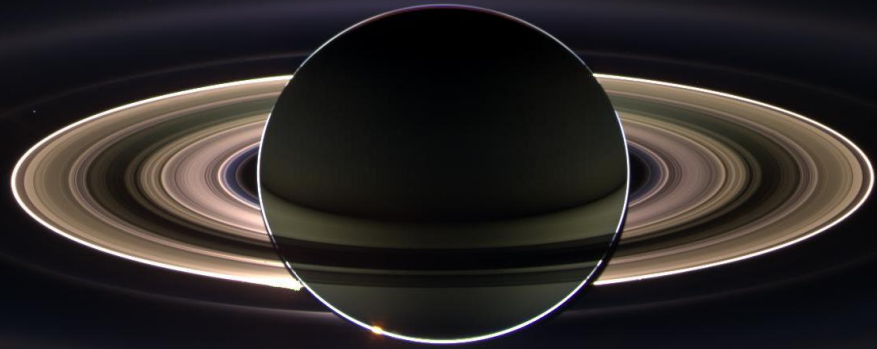
Stars and Brown Dwarfs Within 11 Light Years Base grid is parallel to Galactic plane, at $z = -9.88$ ly. Each grid square measures 2.41×2.41 ly.

- Spectral Type O
- Spectral Type B
- Spectral Type A
- Spectral Type F
- Spectral Type G
- Spectral Type K
- Spectral Type M
- Brown Dwarf
- White Dwarf





NASA, Voyager 1, Feb 4, 1990



SEARCHING FOR INTERSTELLAR COMMUNICATIONS

By GIUSEPPE COCCONI* and PHILIP MORRISON†

Cornell University, Ithaca, New York

NO theories yet exist which enable a reliable estimate of the probabilities of (1) planet formation ; (2) origin of life ; (3) evolution of societies possessing advanced scientific capabilities. In the absence of such theories, our environment suggests that stars of the main sequence with a lifetime of many billions of years can possess planets, that of a small set of such planets two (Earth and very probably Mars) support life, that life on one such planet includes a society recently capable of considerable scientific investigation. The lifetime of such societies is not known ; but it seems unwarranted to deny that among such societies some might maintain themselves for times very long compared to the time of human history, perhaps for times comparable with geological time. It follows, then, that near some star rather like the Sun there are civilizations with scientific interests and with technical possibilities much greater than those now available to us.

* Now on leave at CERN, Geneva.

† Now on leave at the Imperial College of Science and Technology, London, S.W.7.

To the beings of such a society, our Sun must appear as a likely site for the evolution of a new society. It is highly probable that for a long time they will have been expecting the development of science near the Sun. We shall assume that long ago they established a channel of communication that would one day become known to us, and that they look forward patiently to the answering signals from the Sun which would make known to them that a new society has entered the community of intelligence. What sort of a channel would it be ?

The Optimum Channel

Interstellar communication across the galactic plasma without dispersion in direction and flight-time is practical, so far as we know, only with electromagnetic waves.

Since the object of those who operate the source is to find a newly evolved society, we may presume that the channel used will be one that places a minimum burden of frequency and angular discrimi-



A search for life on Earth from the Galileo spacecraft

**Carl Sagan^{*}, W. Reid Thompson^{*}, Robert Carlson[†], Donald Gurnett[‡]
& Charles Hord[§]**

^{*} Laboratory for Planetary Studies, Cornell University, Ithaca, New York 14853, USA

[†] Atmospheric and Cometary Sciences Section, Jet Propulsion Laboratory, Pasadena, California 91109, USA

[‡] Department of Physics and Astronomy, University of Iowa, Iowa City, Iowa 52242-1479, USA

[§] Laboratory for Atmospheric and Space Physics, University of Colorado, Boulder, Colorado 80309, USA

In its December 1990 fly-by of Earth, the Galileo spacecraft found evidence of abundant gaseous oxygen, a widely distributed surface pigment with a sharp absorption edge in the red part of the visible spectrum, and atmospheric methane in extreme thermodynamic disequilibrium; together, these are strongly suggestive of life on Earth. Moreover, the presence of narrow-band, pulsed, amplitude-modulated radio transmission seems uniquely attributable to intelligence. These observations constitute a control experiment for the search for extraterrestrial life by modern interplanetary spacecraft.



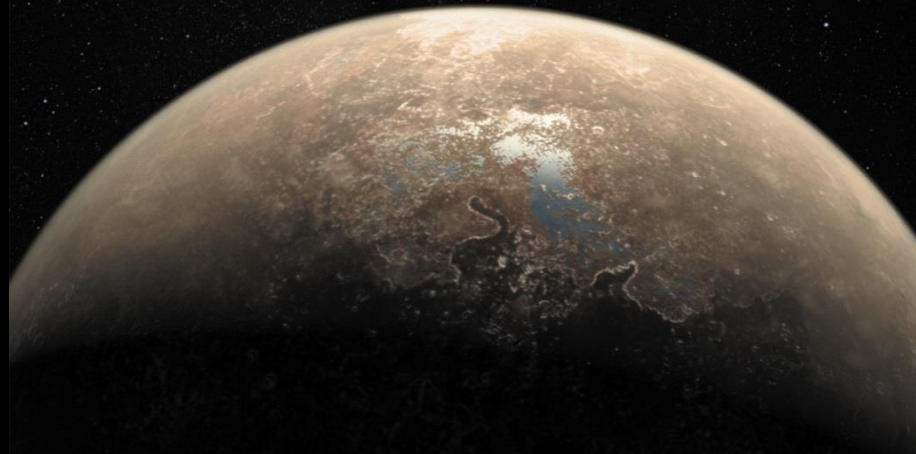
2001 – Space Odyssey

nature

THE INTERNATIONAL WEEKLY JOURNAL OF SCIENCE

NEAR HORIZON

*A warm terrestrial planet in orbit around
Proxima Centauri, closest star to the Sun* **PAGES 408 & 437**



[NATURE.COM/NATURE](https://www.nature.com/nature)

25 August 2016 £10

Vol. 536, No. 7617

PALE RED DOT



— Mercury's Orbit

Sun —

Habitable Zone

Proxima b Orbit

Proxima Centauri

Period: 11.186 days
Minimum mass: 1.27 Earth masses

Mass: 0.12 solar masses
Luminosity: 0.00155 solar luminosities
Rotation period: 83 days
Temperature: 2800 Celsius
Distance to Earth: 4.23 light-years

New Scientist

WEEKLY August 27 - September 2, 2016

SHADOW WORLD
The untouchable layer
of reality all around us

GENETIC RECODE
Blueprint of life gets
a total rewrite

TIME TWISTER
Quantum trick reverses
cause and effect

BAD MEDICINE Why so much health advice turns out to be wrong

SPECIAL REPORT: PROXIMA B

WE'VE FOUND AN EARTH-LIKE PLANET AROUND OUR NEAREST STAR

Should we go there?



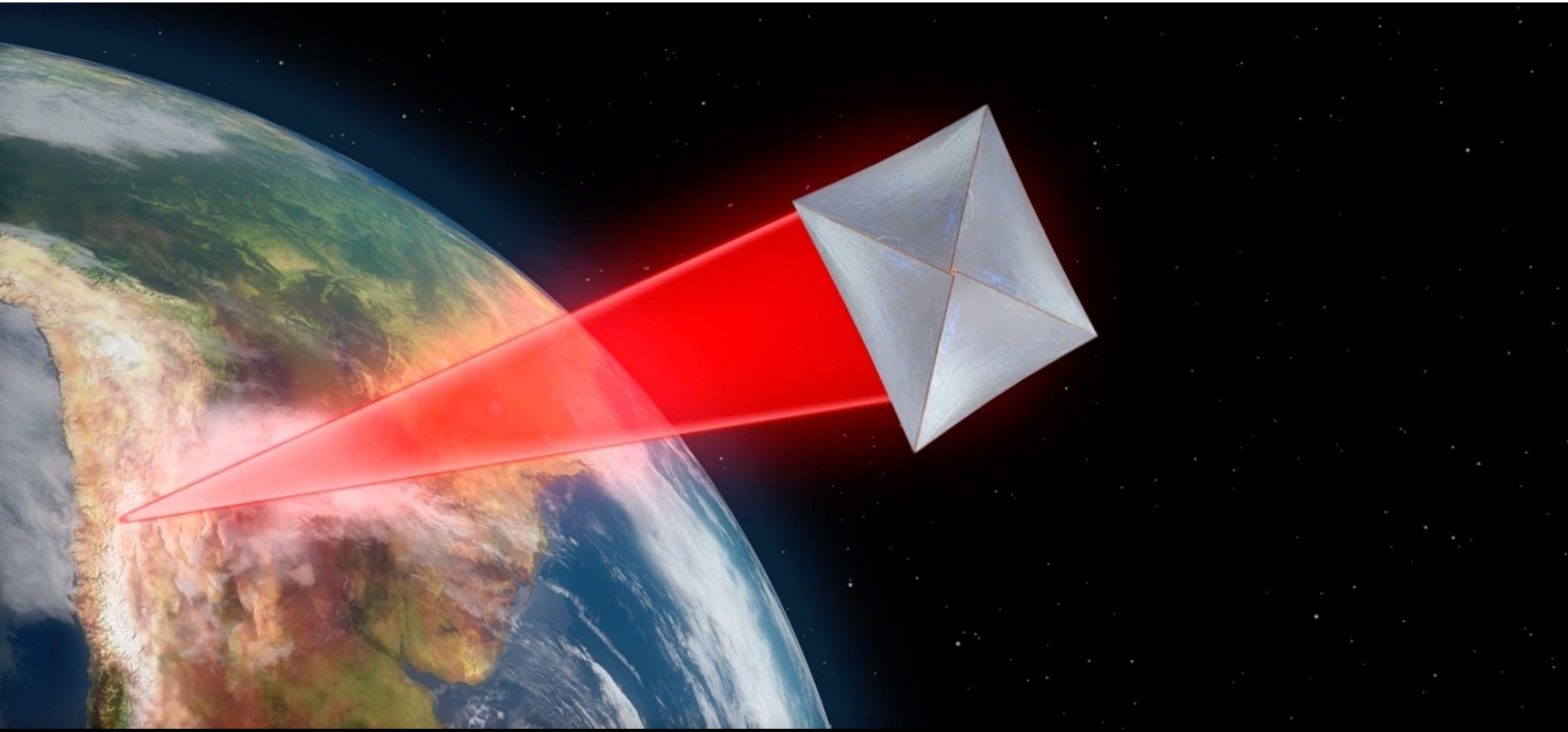
No3088 US\$5.95 CAN\$5.95

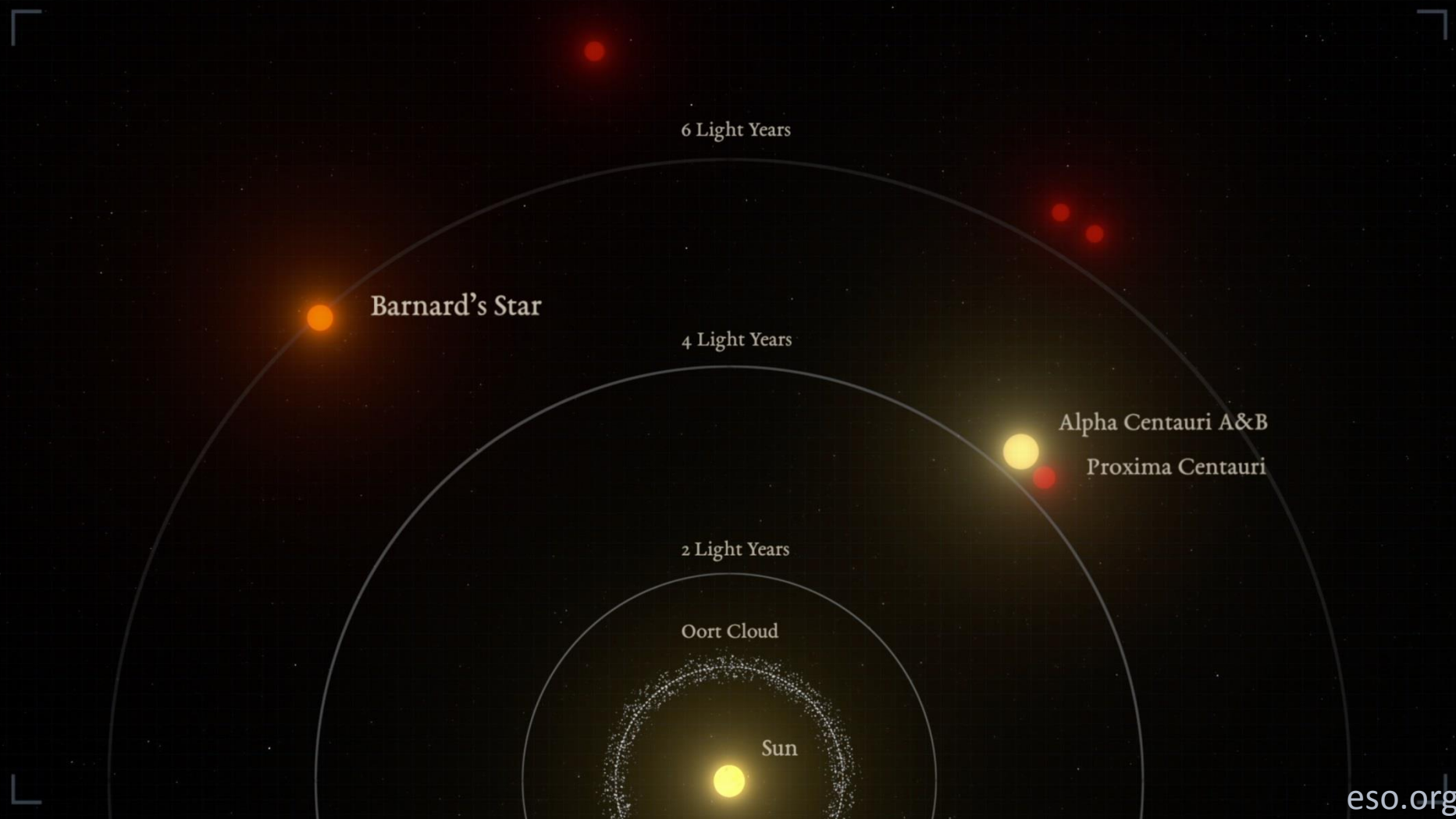


Science and technology news
www.newscientist.com
US jobs in science



BREAKTHROUGH STARSHOT





6 Light Years

Barnard's Star

4 Light Years

Alpha Centauri A&B

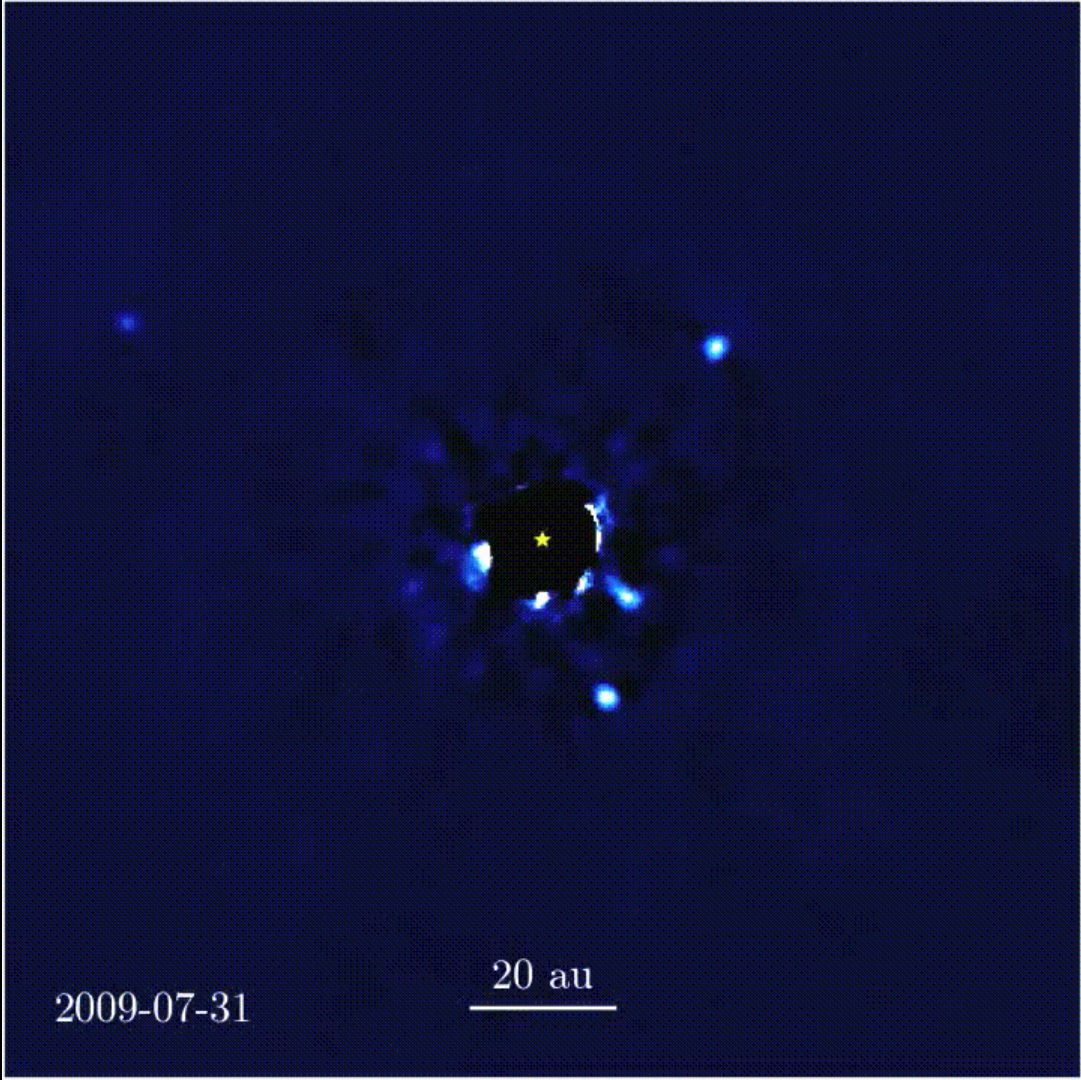
Proxima Centauri

2 Light Years

Oort Cloud

Sun

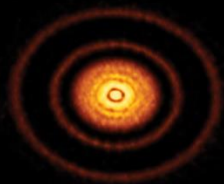




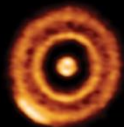
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J. Wang, C. Marois



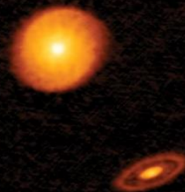
AS 209



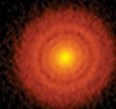
HD 143006



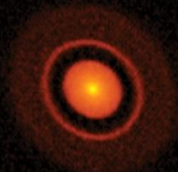
IM Lup



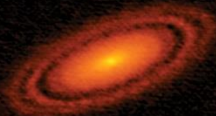
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RU Lup



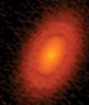
Elias 24



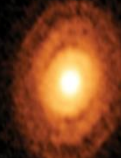
DoAr 25



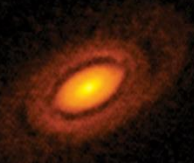
DoAr 33



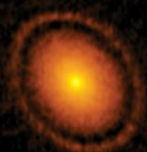
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Wa Oph 6



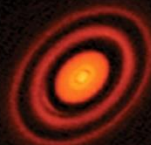
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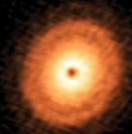
GW Lup



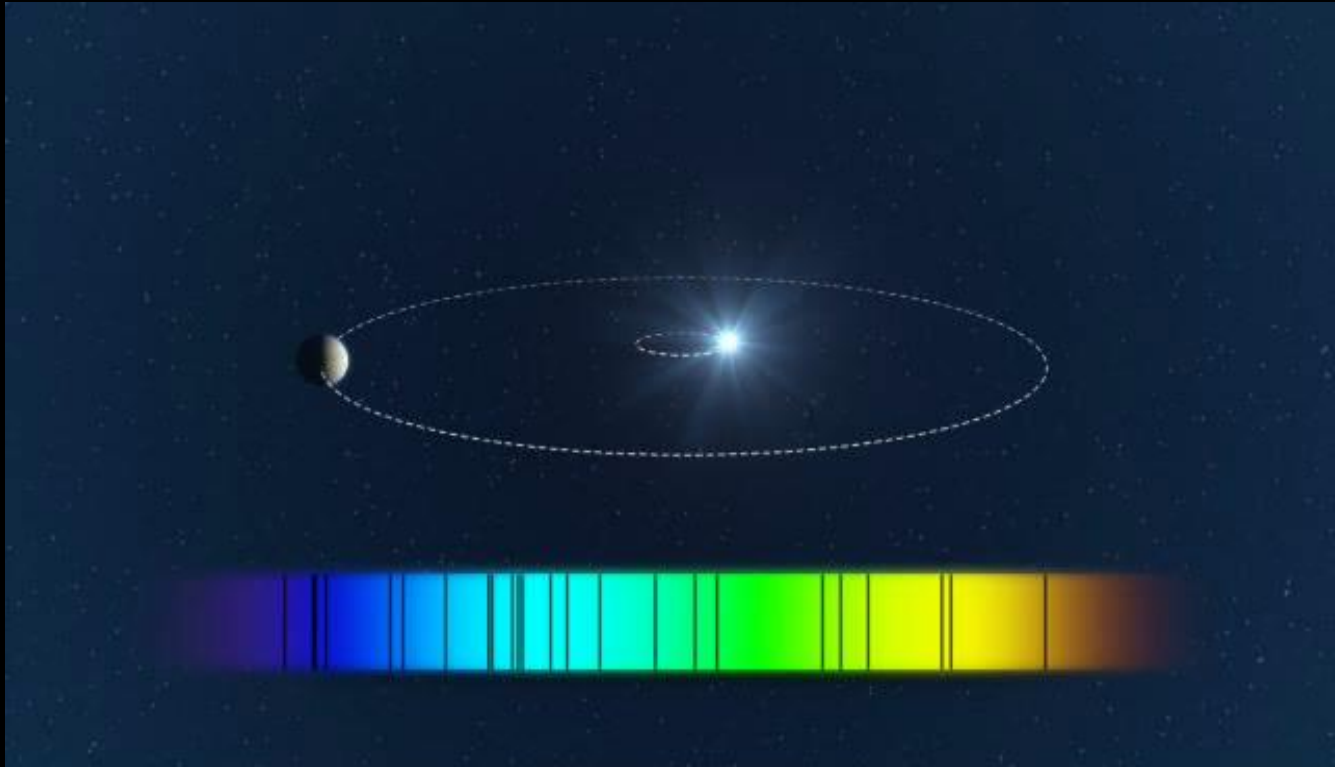
HD 142666



HD 163296

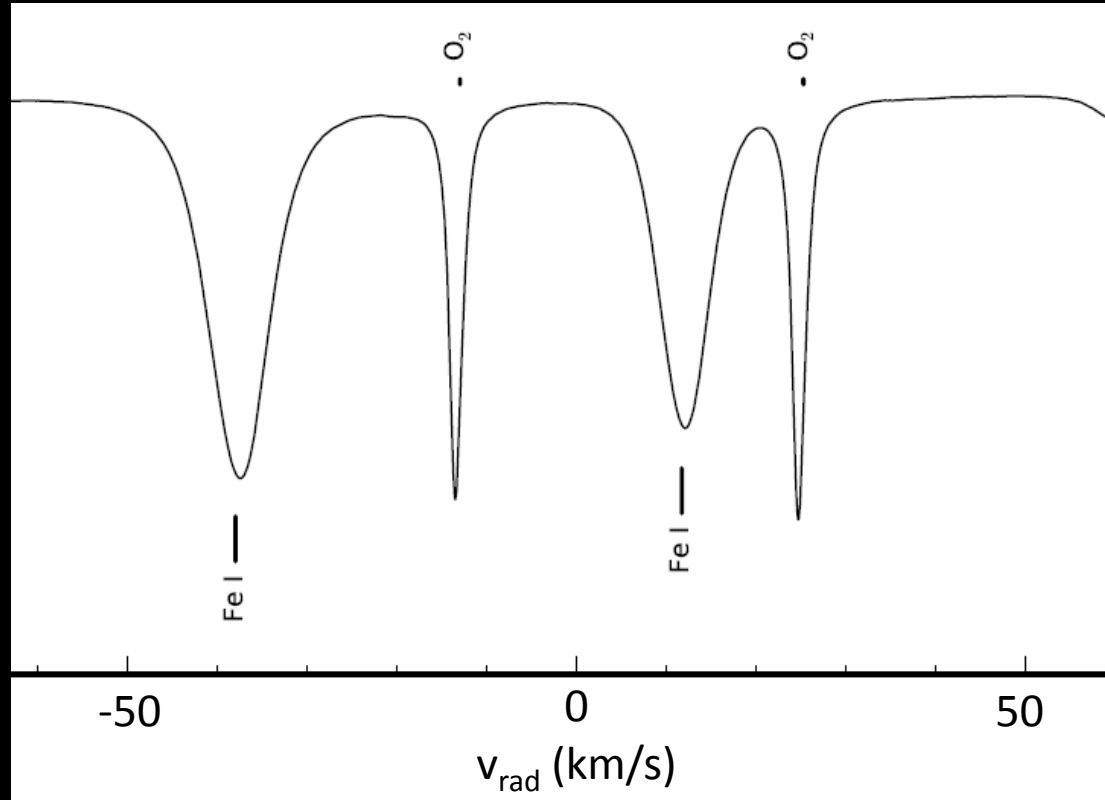


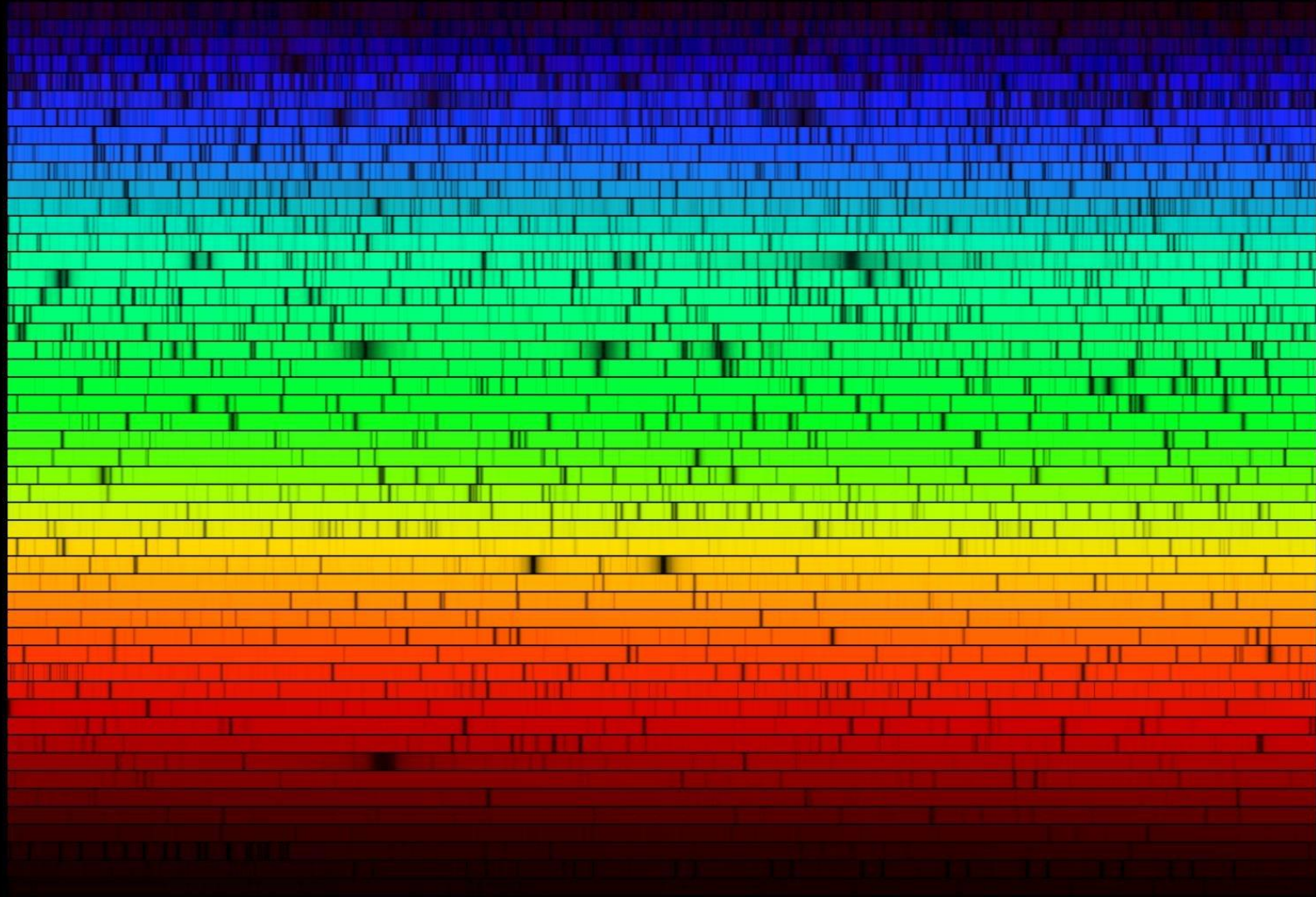
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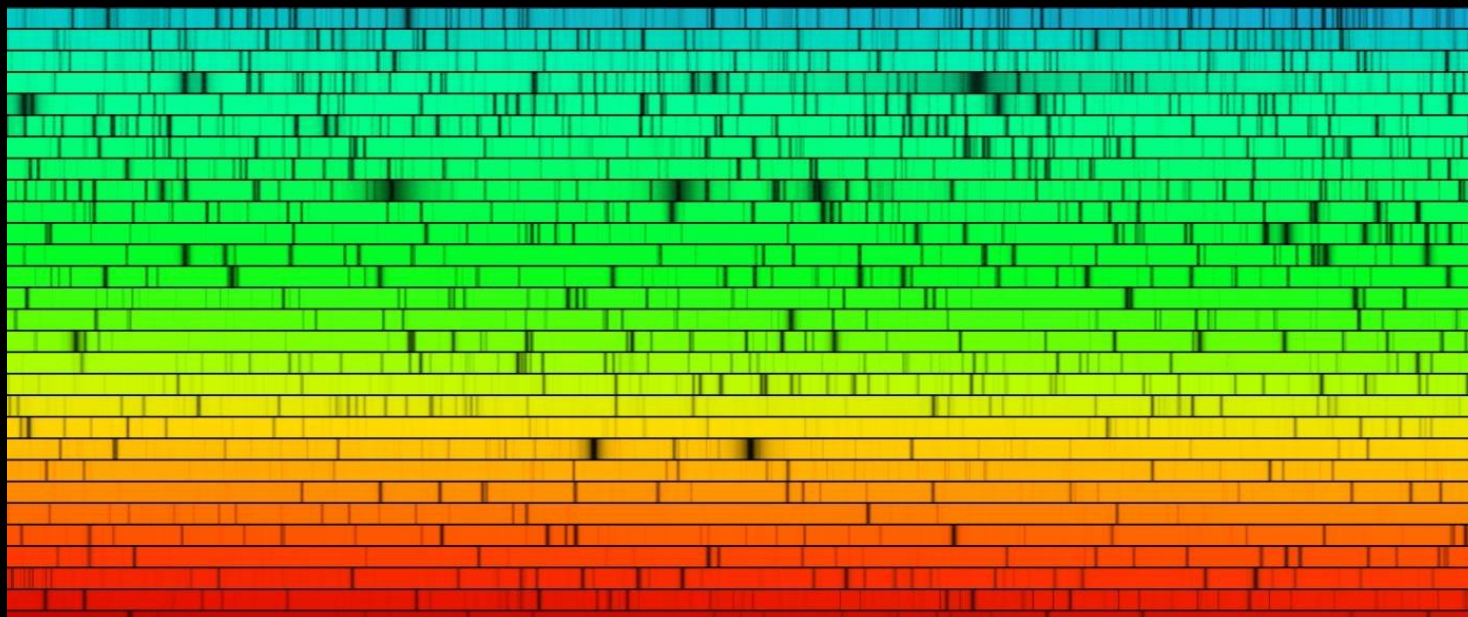


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A stellar spectral line is a few km/s wide







produced by γ -ray photon-photon scattering in a hot, dense accretion disk from which some positrons escape to annihilate

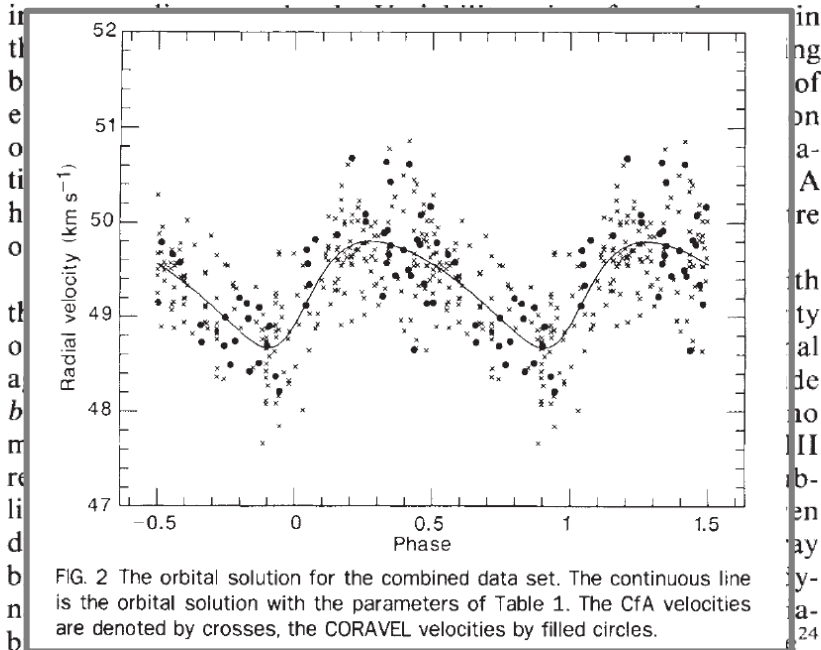


FIG. 2 The orbital solution for the combined data set. The continuous line is the orbital solution with the parameters of Table 1. The CfA velocities are denoted by crosses, the CORAVEL velocities by filled circles.

(the orbital period of GX1+4 is unknown) or from an episodic ejection of mass from the M giant, smothering the X-ray source²⁶. The GRIS data are being searched for the characteristic 2-min pulsation period of such a source.

The unseen companion of HD114762: a probable brown dwarf

David W. Latham*, Tsevi Mazeh†, Robert P. Stefanik*, Michel Mayor‡ & Gilbert Burki‡

* Harvard-Smithsonian Center for Astrophysics, 60 Garden Street, Cambridge, Massachusetts 02138, USA

† School of Physics and Astronomy, Raymond and Beverly Sackler Faculty of Exact Science, Tel Aviv University, Tel Aviv 69978, Israel

‡ Observatoire de Genève, Chemin des Maillettes 51, CH-1290 Sauverny, Switzerland

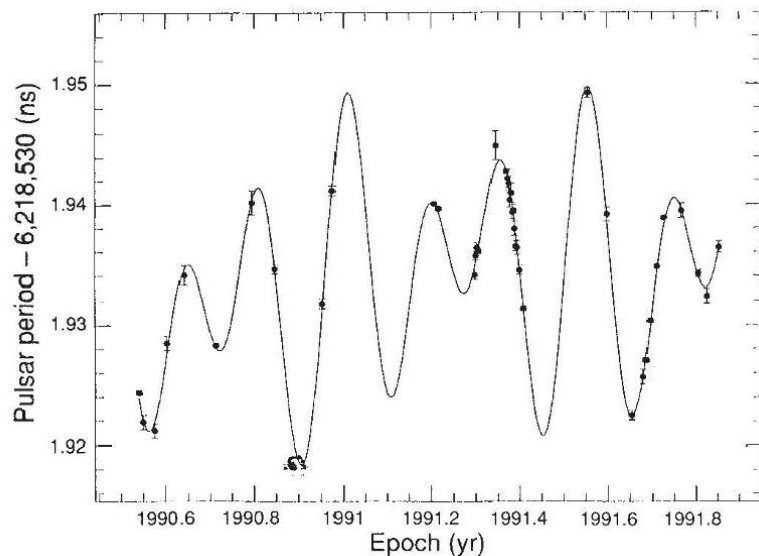
BROWN dwarfs are substellar objects with too little mass to ignite hydrogen in their cores. Despite considerable effort to detect brown dwarfs astrometrically¹⁻⁴, photometrically⁴⁻⁹, and spectroscopically¹⁰⁻¹², only a few good candidates have been discovered. Here we present spectroscopic evidence for a probable brown-dwarf companion to the solar-type star HD114762. This star undergoes periodic variations in radial velocity which we attribute to orbital motion resulting from the presence of an unseen companion. The rather short period of 84 days places the companion in an orbit similar to that of Mercury around the Sun, whereas the rather

A planetary system around the millisecond pulsar PSR1257+12

A. Wolszczan* & D. A. Frail†

* National Astronomy and Ionosphere Center, Arecibo Observatory, Arecibo, Puerto Rico 00613, USA

† National Radio Astronomy Observatory, Socorro, New Mexico 87801, USA



scope at a frequency of 430 MHz (ref. 2). The characteristics of this survey and the details of data analysis are described else-

one as a post-fit residual, implying that the pulse arrival times of PSR1257+12 are indeed affected by two independent periodicities. Further detailed analysis has shown that the periodicities are independent of radio frequency and that other millisecond pulsars routinely observed at Arecibo with the same data acquisition equipment show no such effect in their timing residuals.

Millisecond pulsars are extremely stable rotators. Systematic timing observations of objects like the 1.5-ms pulsar 1937+21 (ref. 6) have not revealed any timing noise, quasiperiodic TOA variations or 'glitches' at the level often found in the population of younger pulsars and believed to be related to neutron star seismology⁷. The frequency independence of the amplitude of

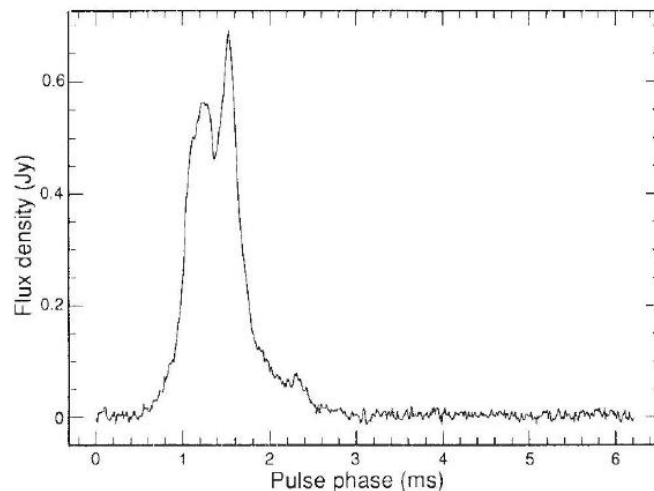
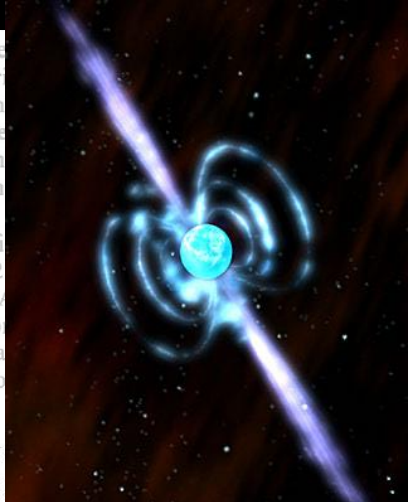
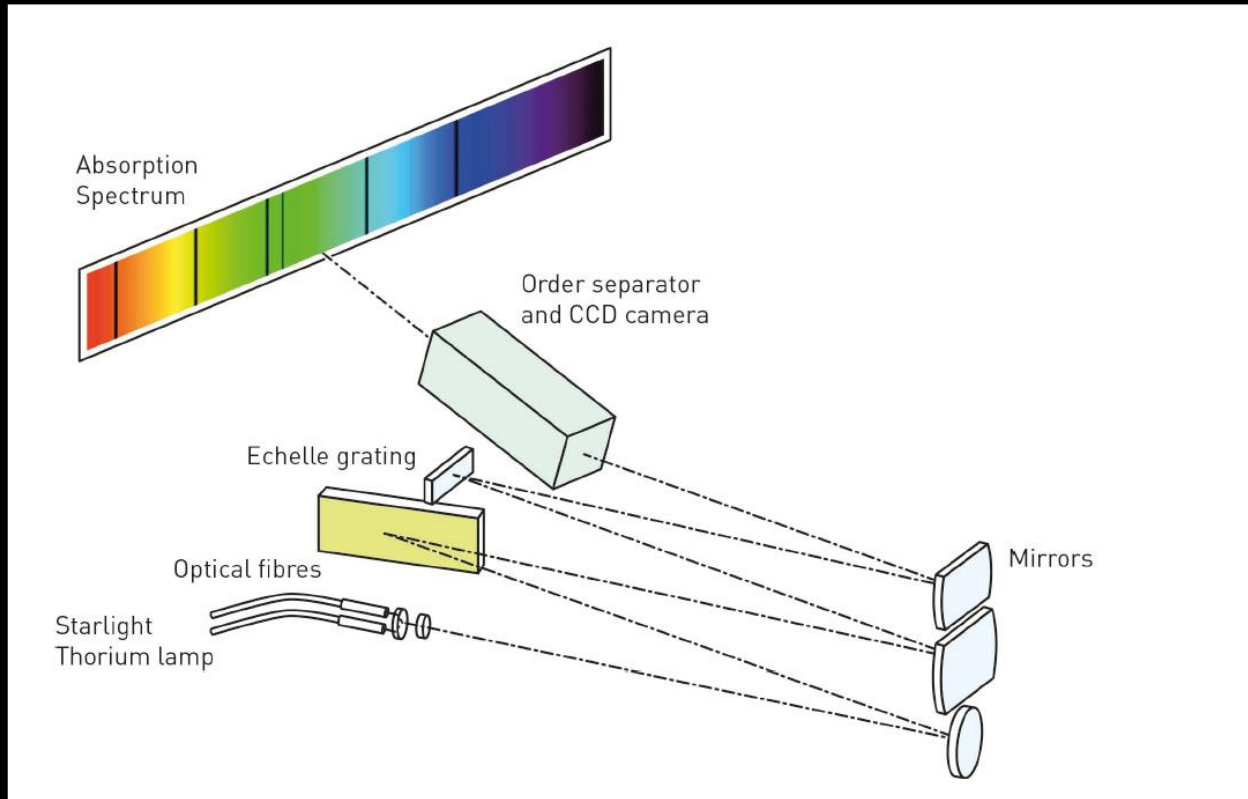


FIG. 1. The average pulse profile of PSR1257+12 at 430 MHz. The effective time resolution is $\sim 12 \mu\text{s}$.





adapted from Baranne et al., 1996

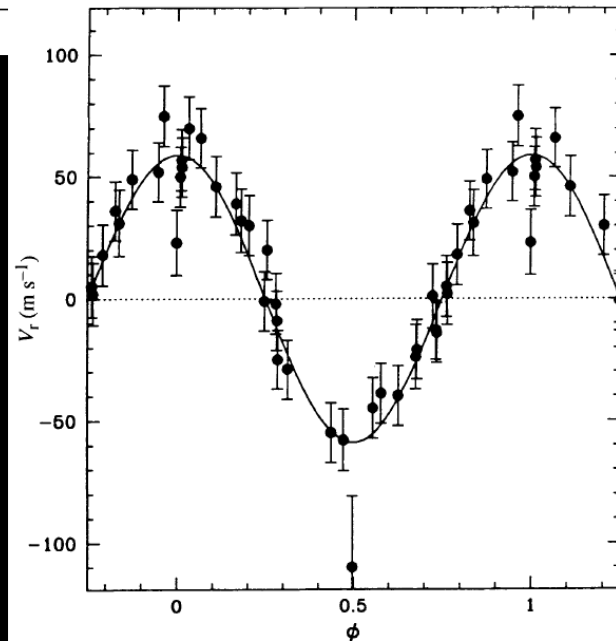
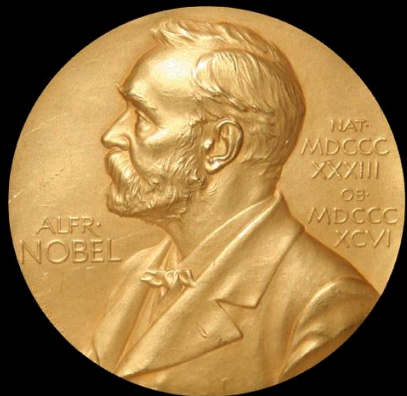
A Jupiter-mass companion to a solar-type star

Michel Mayor & Didier Queloz

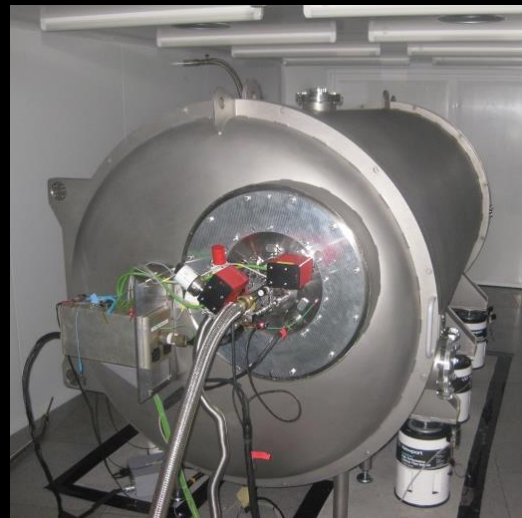
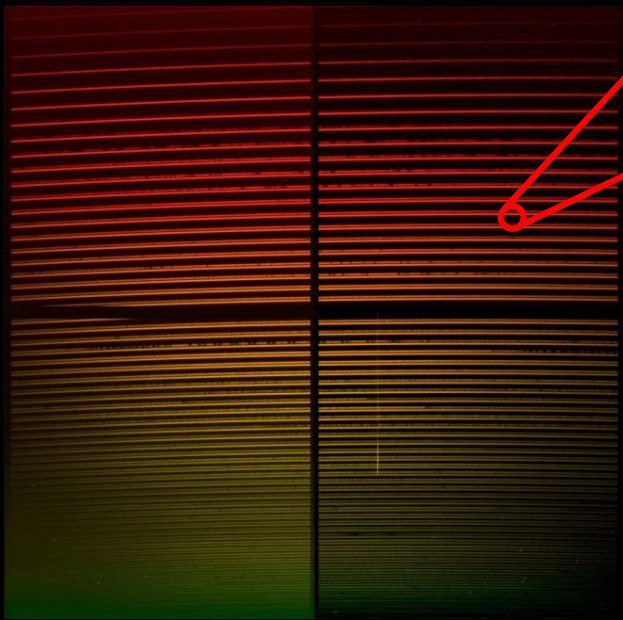
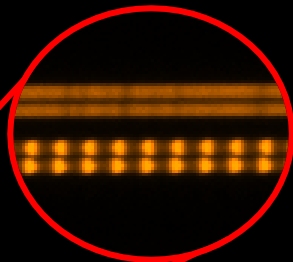
Geneva Observatory, 51 Chemin des Maillettes, CH-1290 Sauverny, Switzerland

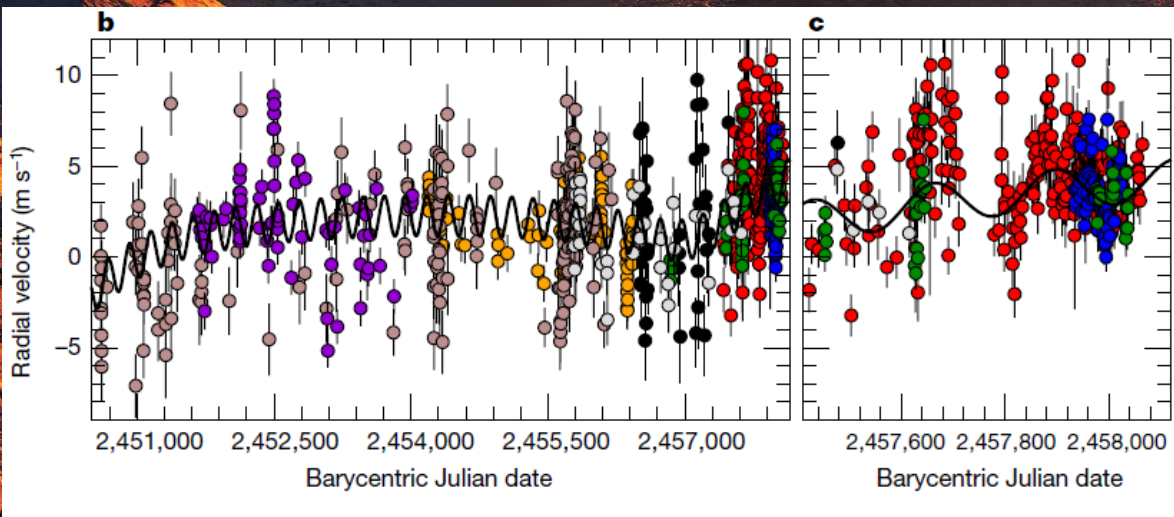
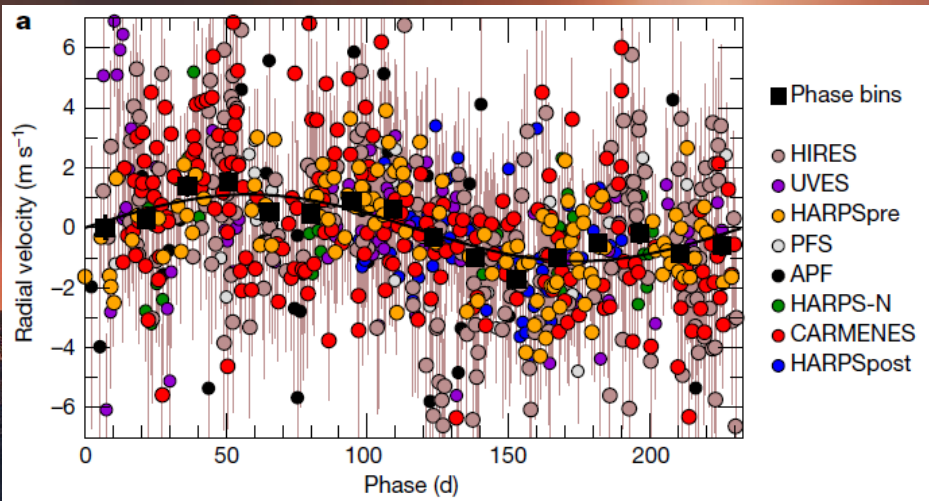
The presence of a Jupiter-mass companion to the star 51 Pegasi is inferred from observations of periodic variations in the star's radial velocity. The companion lies only about eight million kilometres from the star, which would be well inside the orbit of Mercury in our Solar System. This object might be a gas-giant planet that has migrated to this location through orbital evolution, or from the radiative stripping of a brown dwarf.

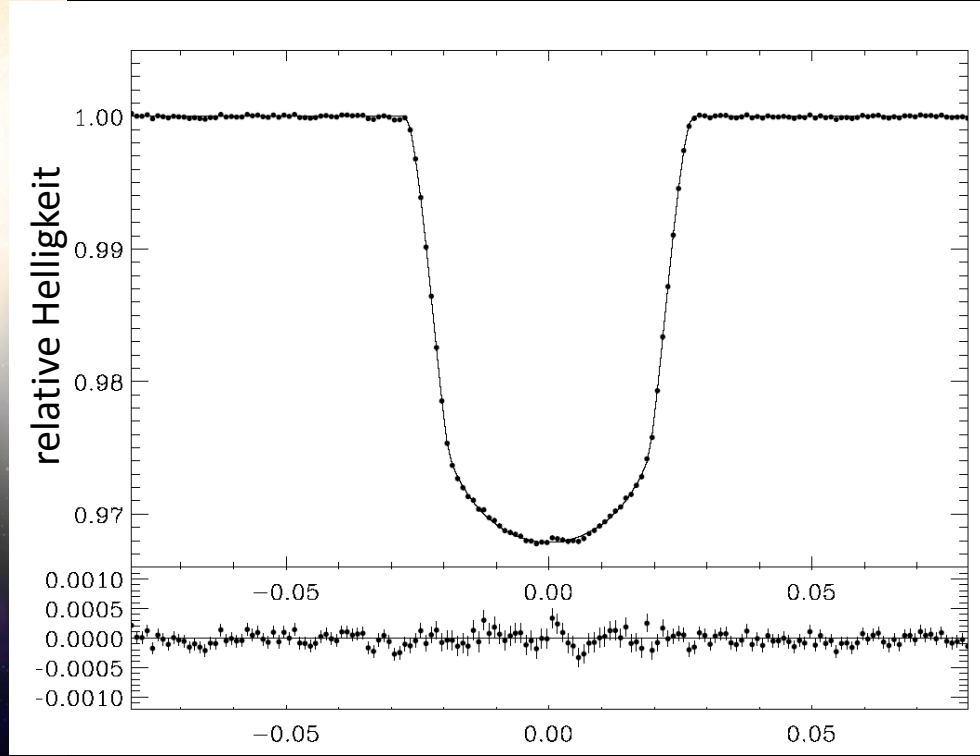
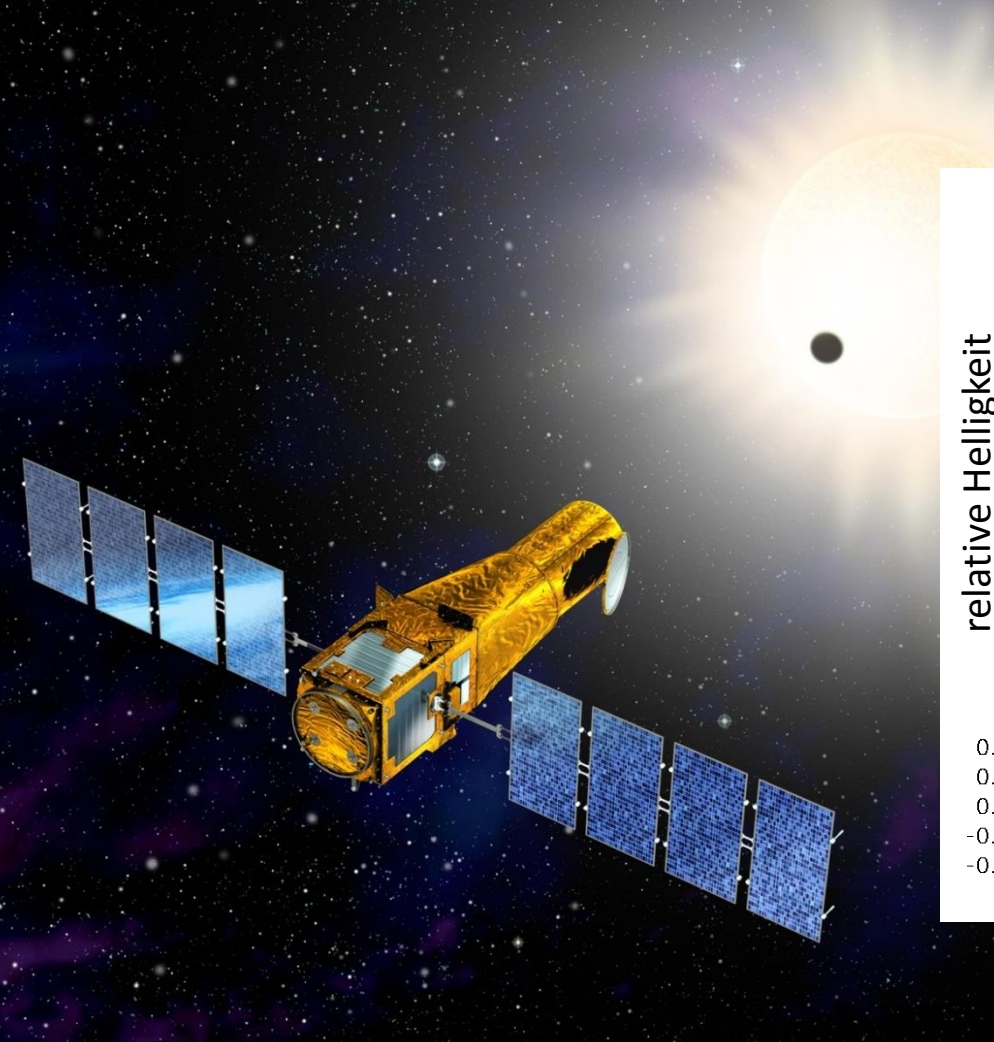
NATURE · VOL 378 · 23 NOVEMBER 1995





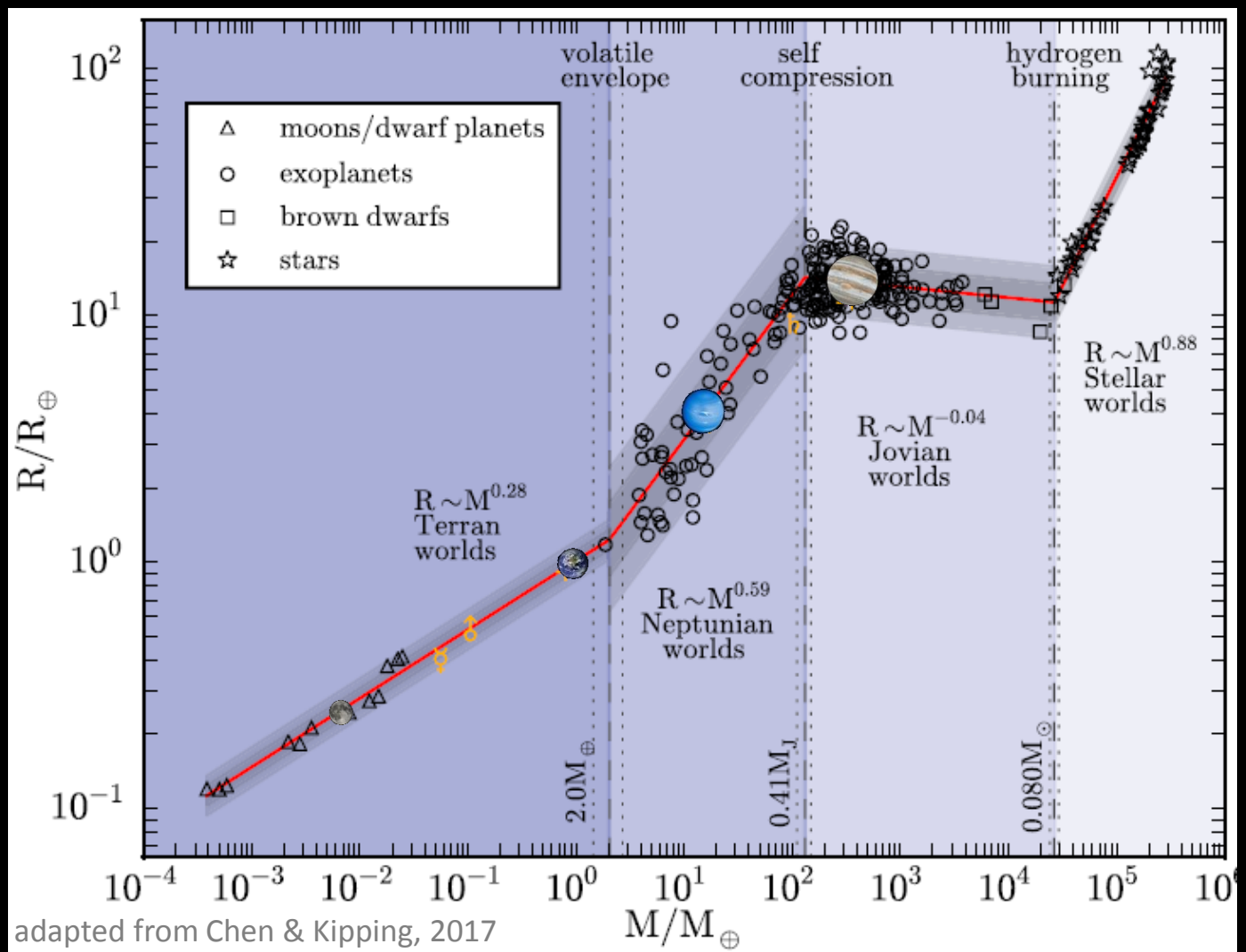




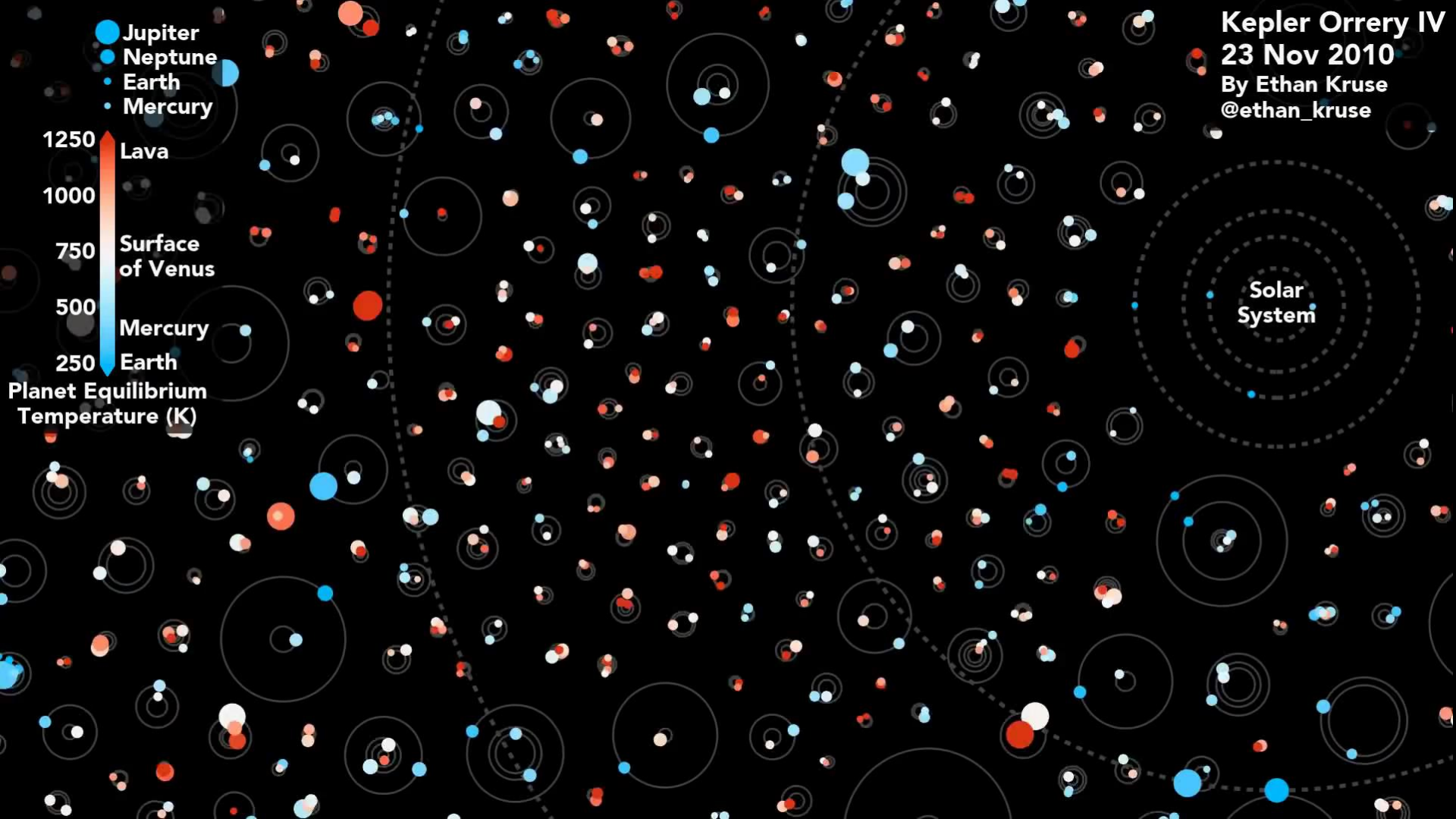


Alonso et al., 2008

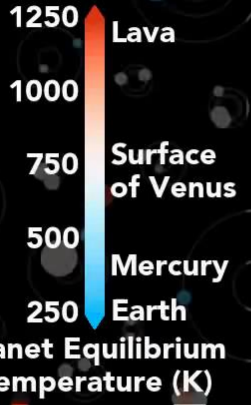
With mass and radius known, we begin to understand the physics



Kepler Orrery IV
23 Nov 2010
By Ethan Kruse
@ethan_kruse

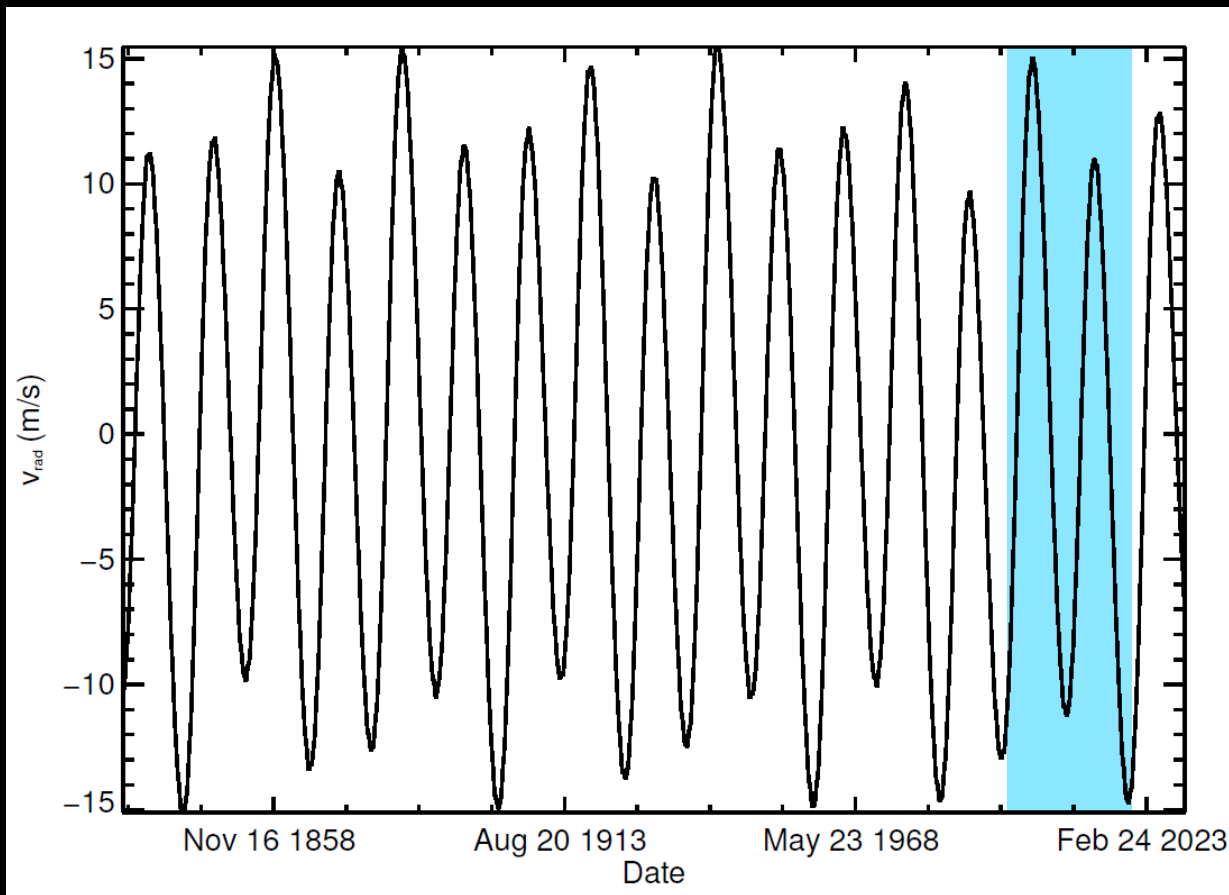


- Jupiter
- Neptune
- Earth
- Mercury



Solar System

The solar system is a challenging one

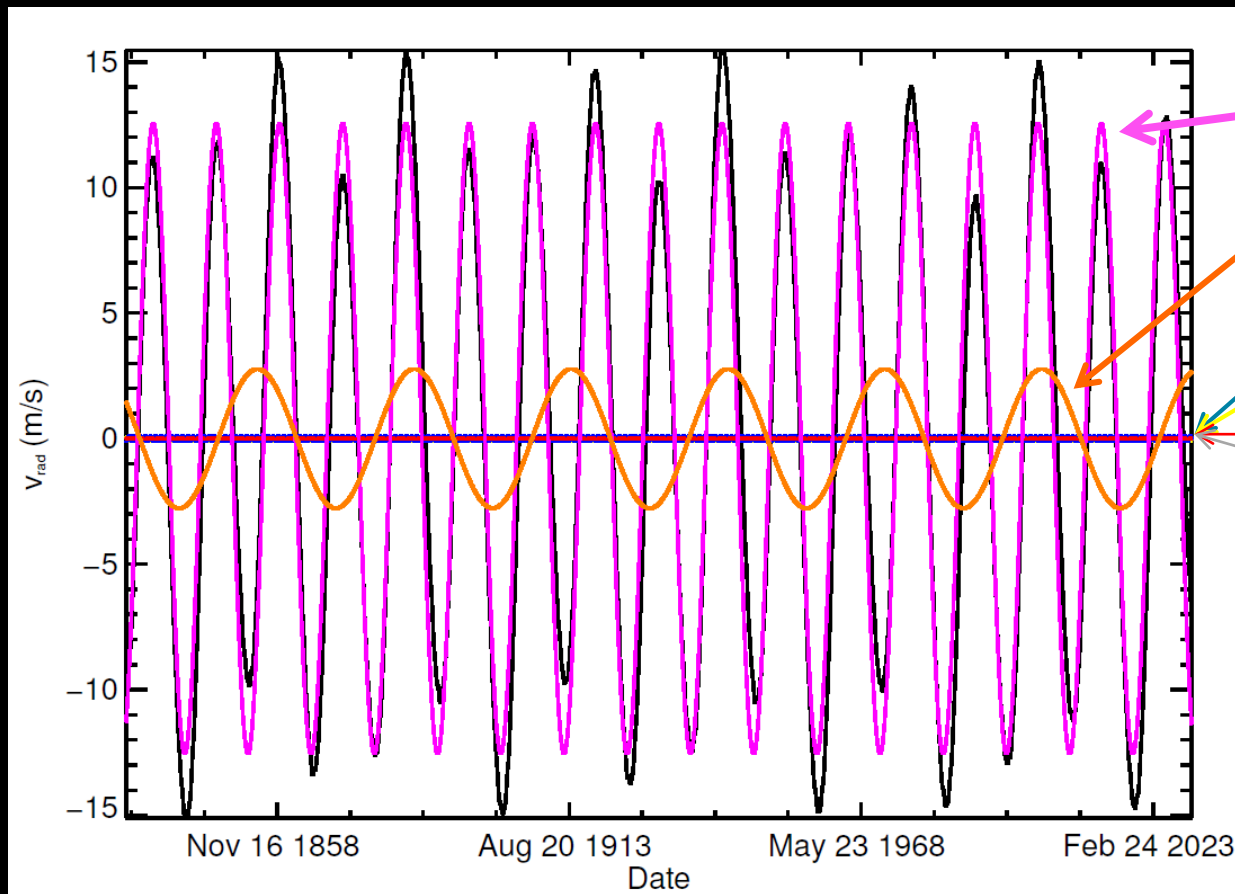


Velocities of the Sun
relative to
Solar System barycenter

data from
JPL/Horizons

„planet hunting epoch“

The solar system is a challenging one



Jupiter 12 m/s

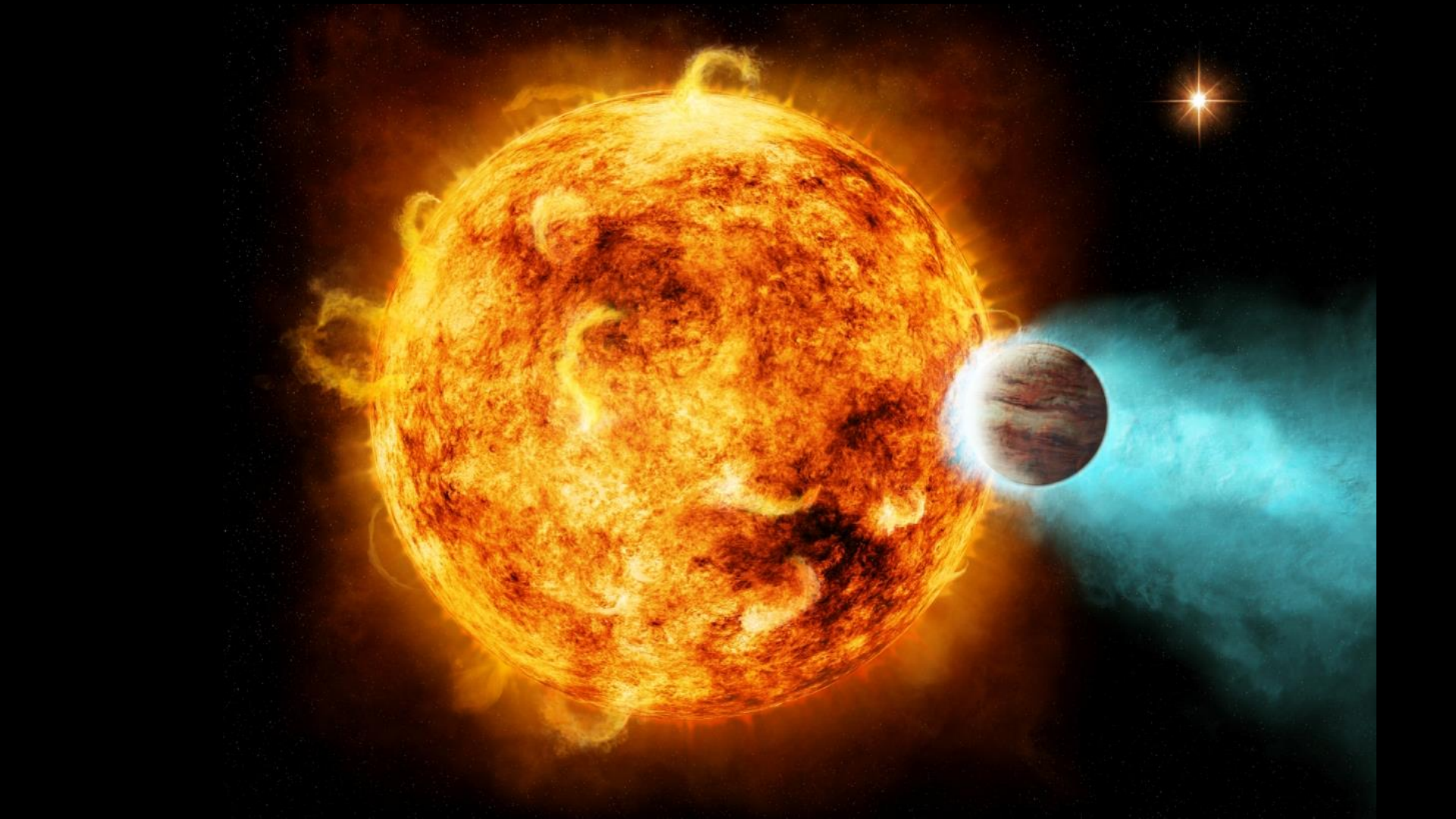
Saturn 2.5 m/s

Earth 9 cm/s

Venus 8.5 cm/s

Mars 8 mm/s

Mercury 8 mm/s



Molecules in planetary atmospheres



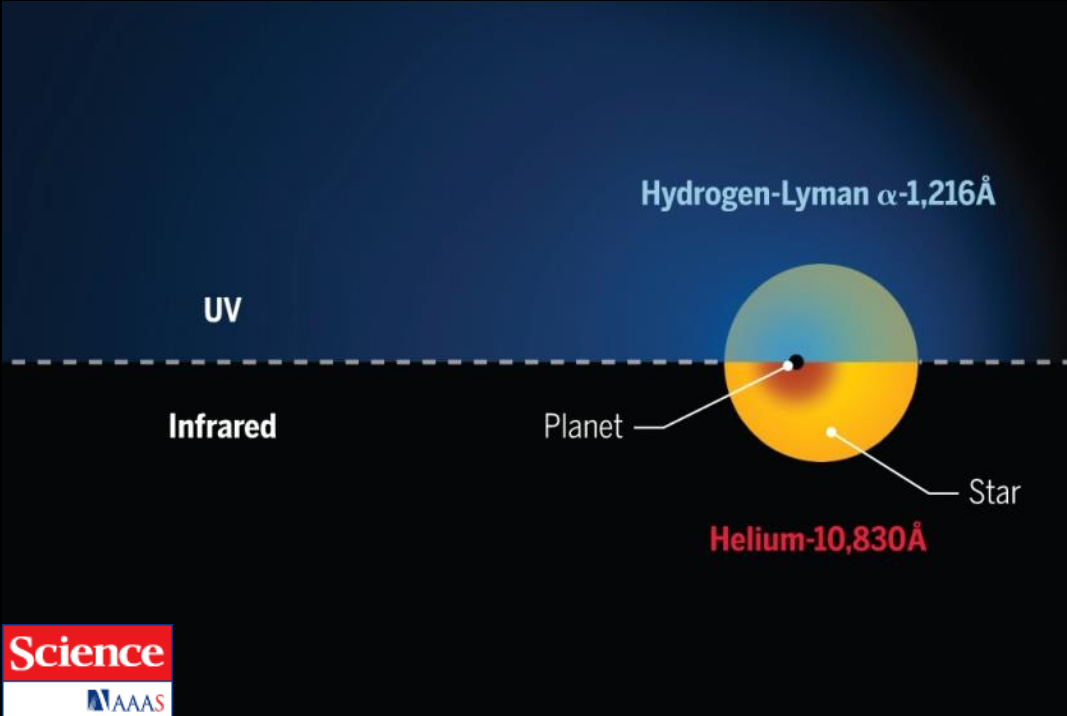
ASTRONOMY

Escaping atmospheres of extrasolar planets

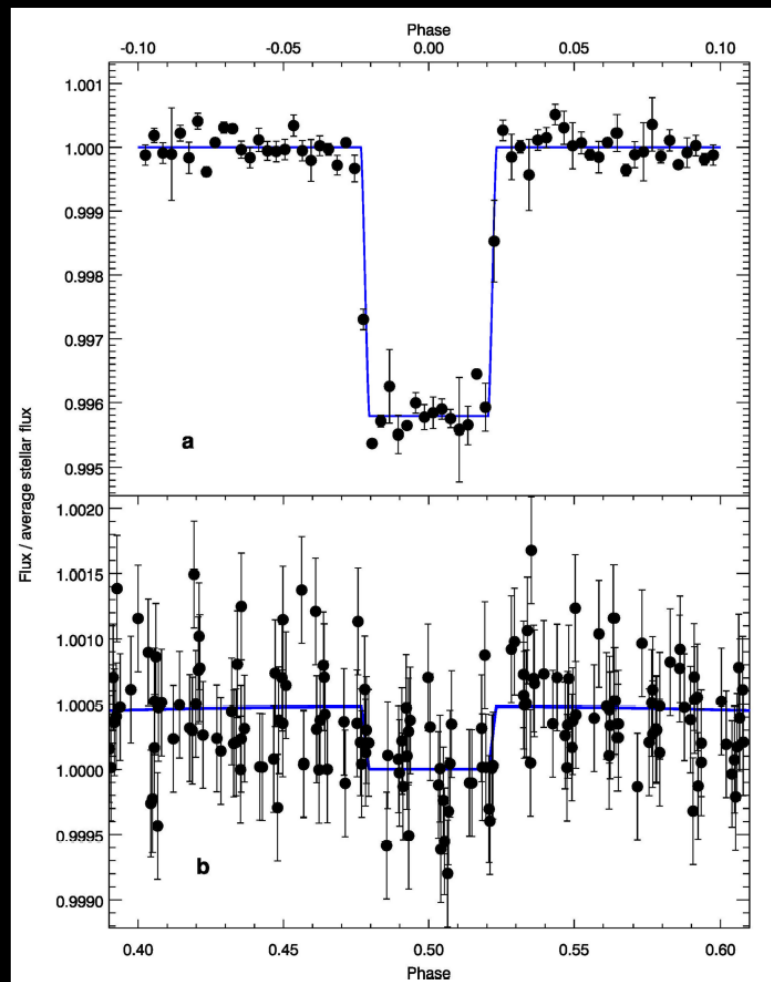
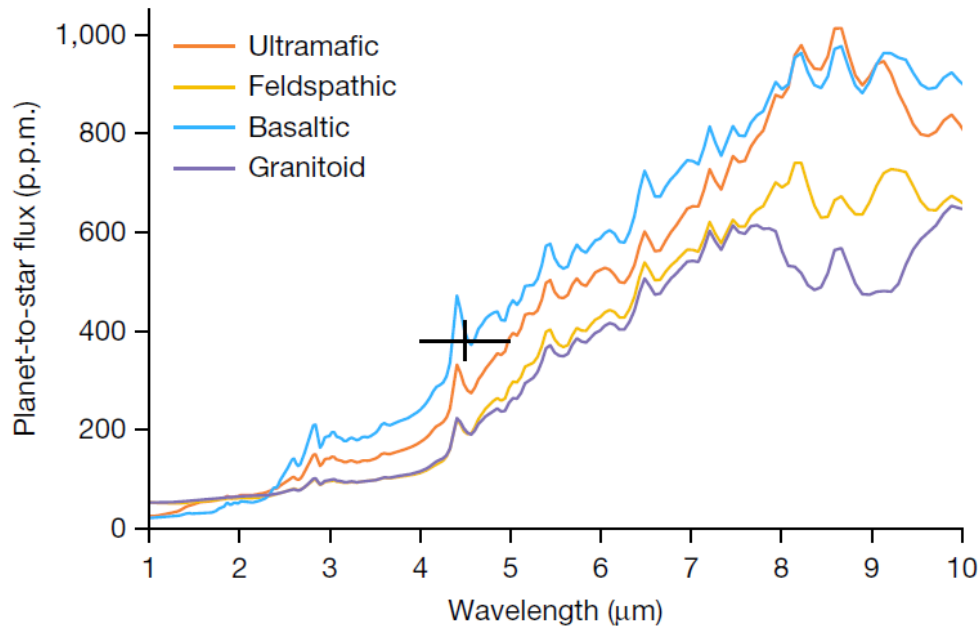
The study of helium absorption opens a new window on escaping exo-atmospheres

By Matteo Brogi

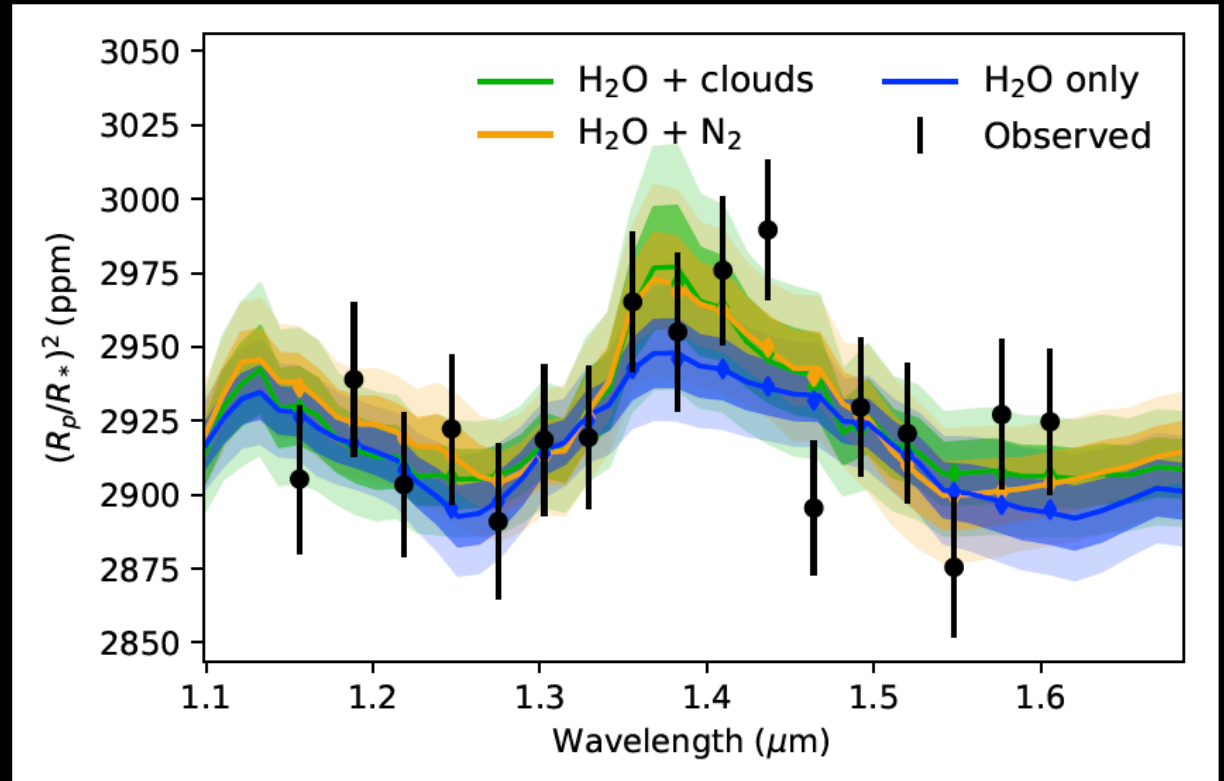
21 DECEMBER 2018 • VOL 362 ISSUE 6421



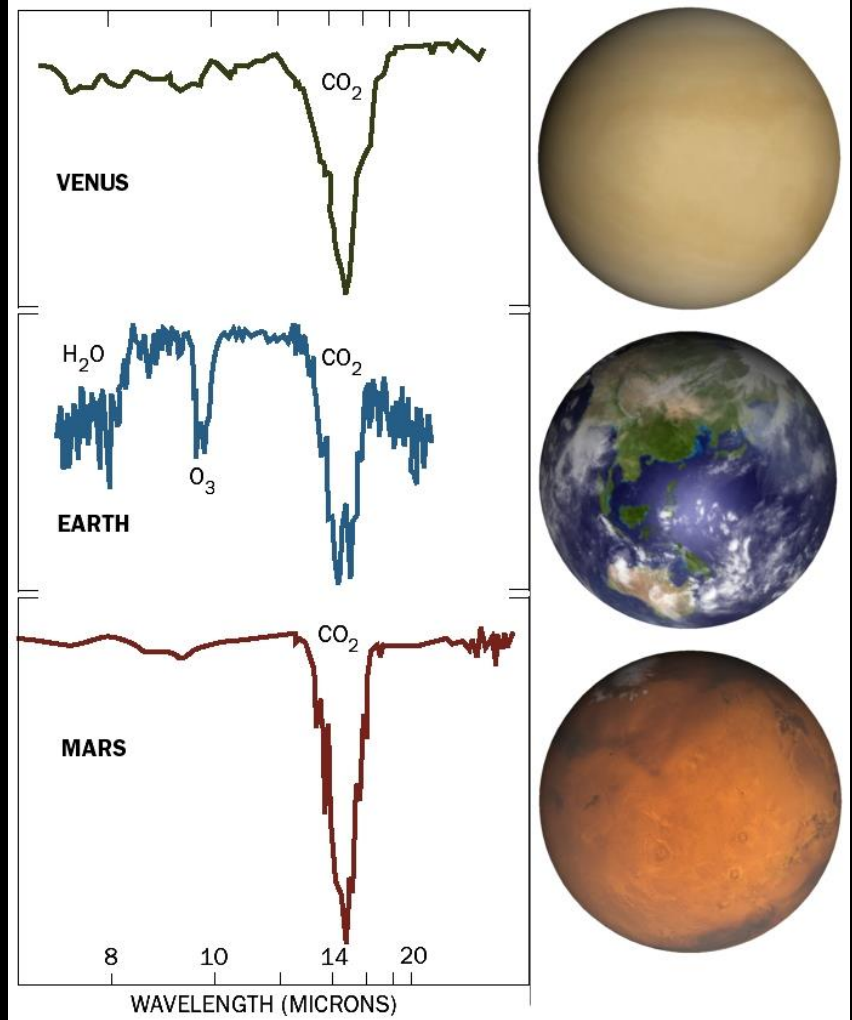
Super-Earth transit spectroscopy (no thick clouds)



Water in Super-Earth atmosphere



Fingerprint of life?



Our place in the universe

1. **Begin to see local universe planet population**
2. **Many stars and planets very different from solar system**
3. **So far, no Earth 2.0 – but many surprises**

