Physikalisches Kolloquium February 12, 2020, DESY, Zeuthen

# Study of structure and Transport Properties of Warm Dense Matter using Ultra-Fast X-rayMethods

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# **Talk outline**

- Introduction to Plasma Physics
- Definition of Warm Dense Matter (WDM)
- Applications in Astrophysics
- Applications in Inertial Confinement Fusion (ICF)
- Experimental methods to study WDM
- Introducing large laser facilities
- Ultra-fast laser probes and transport properties of WDM
- Research projects:
  - OMEGA: Warm Dense Matter equation of state and transport
  - DRACO: XANES absorption spectroscopy on heated copper



#### **Plasma Physics** $\rightarrow$ What is "plasma"?



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### Warm Dense Matter

- Solid densities (~1 g/cm<sup>3</sup>), moderate temperatures (0.1 – 100 eV)
- lons strongly coupled and fluid-like, exhibit longrange order
- Electrons fully or partially degenerate
   quantum effects become important
- Planetary cores/dynamos, white dwarfs, crusts of old stars, supernova explosions, etc.
- Also can be created in a lab by high power lasers, inertial confinement fusion (ICF)
- Equation of state (EOS) structure and transport properties largely unknown





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Coupling parameter: ratio of Coulomb & thermal energy

$$\Gamma_{\alpha,\beta} = \frac{Z_{\alpha}Z_{\beta}e^2}{4\pi\varepsilon_0 a_{\alpha,\beta}k_B T_{\alpha,\beta}}$$

 $\rightarrow$  strongly coupled plasma for  $\Gamma_{ii} > 1$ 

Degeneracy parameter: ratio of thermal energy & Fermi energy

 $\theta = \frac{k_B T}{E_F}$ 

 $\rightarrow$  in WDM  $\theta < 1$ 

Most electrons populate the states within the Fermi sea and cannot further reduce the distance between them (Pauli exclusion principle).

## **Motivation for plasma transport & EOS experiments**

- Interiors of planets are often in the WDM regime
- Jupiter's magnetic field generated by a dynamo of metallic hydrogen deep inside

#### **Outstanding questions:**

- Metallization of hydrogen:
  - What pressure? Temperature?
  - Phase transition continuous
  - or discontinuous? (first/second order)

#### H/He phase separation:

- Atmospheric H/He mixing ratio different from protoplanetary nebula
- Condensation of He in the metallic envelope (He harder to ionize than H)
- Helium rain inside Jupiter?
- Diamond rain inside Neptune?
- Rock core?

Guillot T., Hubbard W.B., Stevenson D.J. & Saumon D. 2003, *"The Interior of Jupiter"*, in Jupiter (Eds. F. Bagenal *et al.*)



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### **Motivation for plasma transport experiments**

- Electric conductivity of plasmas inside planetary dynamos
- Convection inside planets and stars
- Radiation transport in stars and supernovae
- Fuel mix and heating processes in ICF
- Fast electron/proton transport in fast ignition







### **Fusion technology with lasers (ICF)**

- Inertial confinement fusion concept to fusion relies heavily on the knowledge of equation of state and transport properties of WDM
- 60 200 laser beams of combined power of 10<sup>12</sup> W (TW)
- Fusion reaction of hydrogen atoms to form Helium at extreme temperatures and densities
- Three main laser approaches:



- Direct drive compression
- Indirect drive with x-rays
- Fast ignition fusion







## **Generating WDM: Dynamic compression**

- Experimental platforms for dynamic compression using lasers
- Typical solid/liquid target compression techniques typically used:
  - Shocks driven directly by laser ablation
  - Shock-and-release 
     → achieve states away from the Hugoniot (few Mbar, 1–100 eV)

#### Main challenges:

- Preheat of material before the shock changing initial conditions
- Shock stability and reproducibility (mainly issue on lasers)
- Gradients in generated plasma conditions



## **Generating WDM: Isochoric heating**

#### Experimental platforms for isochoric heating:

- High energy laser facilities
- Particle accelerators, synchrotrons and free-electron lasers (FEL)

#### Typical heating sources:

- X-rays (laser generated, synchrotron or x-ray FEL)
- Protons (typically laser-acceleration, lower energies than accelerators)
- Electrons (laser generated)
- Direct heating with optical lasers also done, but requires very thin targets
- Allows more homogeneous heating of solid density samples due to better penetration of x-rays or particle sources of dense targets, density is known
- Ultrafast and non-thermal a) processes can be studied with short-pulse heating sources
- Main challenges:
  - Non-uniform plasma blow off
  - Electromagnetic pulses damaging electronics
  - Radiation, nuclear activation



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#### Active probes needed for WDM

- High density
- Relatively low temperatures
  - Optical probes cannot penetrate inside WDM
  - ➔ use x-rays, particles

#### X-ray Thomson scattering

- Measurements of temperature, density, ionization state and microscopic structure
- Requires very strong x-ray sources



Scattering vector:  

$$k = (4\pi/\lambda_0)\sin(\theta/2)$$
 $\lambda_s = \sqrt{\frac{k_B T_e^{\text{eff}}}{4\pi e^2 n_e}}$ 

$$S_{ee}^{tot}(k,\omega) = |f_I(k) + q(k)|^2 S_{ii}(k,\omega) + Z_f S_{ee}^0(k,\omega) + Z_c \int \tilde{S}_{ce}(k,\omega-\omega') S_s(k,\omega') d\omega'$$
Bound electrons/following motion of the ions Free/delocalized electrons Bound-free transitions

S. H. Glenzer and R. Redmer, Rev. Mod. Phys. 81, 1625 (2009)

concep



Rayleigh peak: elastic scattering (bound electrons) Compton peak: inelastic scattering (free/metallic electrons)





**Rayleigh peak:** elastic scattering (bound electrons) **Compton peak:** inelastic scattering (free/metallic electrons)

#### Determine the elements of EOS from features of scattering spectra:

• Electron temperature (T<sub>e</sub>) from width of the inelastic peak (or detailed balance)



$$\frac{S(\mathbf{k},\omega)}{S(-\mathbf{k},-\omega)} = e^{-\frac{\hbar\omega}{k_B T_e}}$$





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- Electron density (n<sub>e</sub>) from downshift of a plasmon peak (collective scattering)



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- Ion temperature (T<sub>i</sub>) from elastic scattering strength









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- Ion temperature (T<sub>i</sub>) from elastic scattering strength
- Average ionization state (Z) from intensity ratio of the Rayleigh and Compton peaks

 $\Delta \omega = -\frac{\hbar^2 k^2}{2m_e} \pm \mathbf{k}.\mathbf{v}$ 

$$\frac{S(\mathbf{k},\omega)}{S(-\mathbf{k},-\omega)} = e^{-\frac{\hbar\omega}{k_B T_e}}$$





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- Ion temperature (T<sub>i</sub>) from elastic scattering strength
- Average ionization state (Z) from intensity ratio of the Rayleigh and Compton peaks
- Atomic structure from bound-free tail contribution



$$\frac{S(\mathbf{k},\omega)}{S(-\mathbf{k},-\omega)} = e^{-\frac{\hbar\omega}{k_B T_e}}$$



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## **Optical diagnostics for WDM**

#### VISAR (Velocity Interferometry of Any Reflector)

- Measures velocity of moving shockwave
- Used to infer pressure and density in shocked material (WDM state)

#### SOP (Streaked Optical Pyrometry)

- Measures optical emission from shock
- If calibrated absolute temperature measurement can be extracted from the intensity measurement by comparison with blackbody radiation





P. M. Celliers *et al.*, Rev. Sci. Instrum. **75**, 4916 (2004) J. E. Miller *et al.*, Rev. Sci. Instrum. **78**, 034903 (2007)

## **Going to big lasers: OMEGA**





# **Novel experimental platform at OMEGA**

- Warm Dense Matter with conditions relevant to ICF and Jovian planets interiors was generated by 14–15 OMEGA beams (University of Rochester) with laser drives of varied pulse duration and intensity: up to 3 × 10<sup>15</sup> W/cm<sup>2</sup>
- A versatile platform for shock-and-release experiments, classical EOS measurements as well as novel approach to study transport in WDM
- Targets are planar samples of various materials including diamond, graphite, CH foams and aerogels
- WDM/plasma state diagnosed using:
  - Imaging x-ray Thomson scattering (XRTS)
  - X-ray radiography (V backlighter)
  - Velocity interferometry (VISAR)
  - Streaked optical pyrometry (SOP)

- K. Falk et al., Phys. Plasmas 21, 056309 (2014)
- K. Falk *et al.*, Phys. Rev. E **90**, 033107 (2014)
- K. Falk et al., PPCF 59, 014050 (2017)
- K. Falk et al., Phys. Rev. Lett. **120**, 025002 (2018)



K. Falk *et al.*, Phys. Rev. Lett. **112**, 155003 (2014)

# **Experimental configuration**

#### **VISAR Target viewed from TIM5**



TIM5-ASBO telescope



#### **IXTS** Target viewed from top





# Imaging x-ray Thomson Spectrometer (IXTS)

- A novel diagnostic particularly suitable for shock-release experiments developed by LANL and the University of Michigan to operate in a standard TIM setup
- Toroidal Ge crystal used to provide spatial and spectral resolution in the same image



# X-ray scattering setup

- Probe x-rays Ni He-alpha @ 7.8 keV
- Narrow-band Ni He-alpha x-rays created by 10 laser beams incident onto a Ni foil @  $2-3 \times 10^{15}$  W/cm<sup>2</sup>
- Collimated by a 200 µm diameter Ta slit, placed 500  $\mu$ m from the shocked target

Target surface normal



Ni K- $\alpha$  line





5 µm Ni backlighter

Target + Be filter.

Shocked target

Au cap

# **Shock-and-release experiments**

#### Planar targets with diamond and graphite used

 0.2 g/cm<sup>3</sup> silica aerogel foam was used as a low density pressure standard, i.e. release material with well-known Hugoniot:

 $u_s = 1.2027u_p + 0.507$ 

- 40 μm step on foam used for shock breakout timing
- When shock transits into the foam a release wave travels back through the C sample creating WDM
- Pressure in diamond/graphite determined from Hugoniot relations:

$$P_{release} = \rho_0 U_s U_p$$



Knudson *et al.*, J. Appl. Phys. **97**, 073514 (2005) Page 25



# **Experimental results**

- Radiography measured the release wave density
- Imaging XRTS used for temperature measurement
- VISAR and SOP used to obtain shock velocity
  - Shock velocity was used to calculate the pressure from experimentally determined relationship



0.8 - 0.8 - 100 - 123 - 100 - 123 - 100 - 123 -

Cemperature (e

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# **First model-independent EOS measurement**

- This was the first model independent full equation of state (EOS) measurement (T,  $\rho$  and P) in WDM to date
- Experimental data were compared with SESAME tables:
- Comparison with ab-initio QMD models provided excellent agreement



### **Transport properties with ultrafast probes**

- Transport coefficients in plasma are linked to non-equilibrium/ultra-fast processes
  - Conductivity, diffusion, mixing of material, particle stopping powers, etc.
  - High temporal resolution (femto- or pico-seconds) is required
- Ultrafast x-ray/particle sources for heating and probing:
  - Nano/micro-structured targets for K-alpha production with short-pulse lasers
  - Laser wakefield acceleration
  - Proton acceleration with lasers
- Multi-beam fs laser facilities
- Development of new instruments





# **DRACO laser facility and HIBEF**

- HZDR has an on-site multiple beam short-pulse laser facility:
  - 200 TW (6 J, 30 fs) laser system with reliable betatron setup
  - Additional 1 PW (30 J, 30 fs) laser available since 2017
  - PENELOPE laser system (150 J, 150 fs) under development (expected in 2020)
- HZDR has direct links to the HIBEF beamline at the European XFEL:
  - Multi-beam laser facilities, magnetic field generators available, XFEL operational



# **Ultrafast melting of Warm Dense Cu**

- Study ultrafast structural transformation in WDM
- Non-thermal melting, lattice disordering, bond hardening
- Utilize unique laser-driven x-ray beams
- Develop suitable x-ray spectrometer



F. C. Kabeer, Phys. Rev. B 89, 100301 (2014)



#### Laser wakefield acceleration for electron acceleration:

- A femtosecond laser pulse drives a plasma wave in a gas target that creates "bubbles" free of electrons with a strong potential in which injected electrons are accelerated
- Laser-driven betatron source:
  - Driven by oscillation of electrons in strong electrostatic fields inside the "bubbles"
  - Oscillatory motion emits bremstrahlung with synchrotron-like "betatron" spectrum
- Low divergence (tens of mrad) laser **Betatron x-ray beam** Short duration (tens of fs) Broadband spectrum with  $E_{crit} \sim 1 - 100 \text{ keV}$ Gas jet electrons laser pulse x-ŕays Plasma wake

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# X-ray absorption spectroscopy (XANES)

- X-ray Absorption Near Edge Spectroscopy (XANES)
- Absorption peaks above edges correspond to spatial distribution of ions
- Photoionized electron ejected into continuum state with low residual energy
  - ⇒ Scattering by the surrounding ions
     ⇒ Absorption cross section reflects the ionic structure of the matter.
- Slope of K-edge corresponds to density of free states ~ temperature.
- Has to be modelled by simulations (e.g. Abinit, DFT-MD)



Page 37

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F. Dorchies et al., Phys. Rev. Lett. 107, 245006 (2011)

# Mosaic HOPG crystal spectrometers



- Composed of small crystallites with plane slightly tilted from the surface ( $\sim$  $0.1 - 1^{\circ})$
- High reflectivity
- Mosaic focusing
- Angular resolution







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# **Preliminary experiments in Lund**

- Lund Laser Center, Sweden
- Ti:Sapphire,  $\sim$  750 mJ, 37 fs, 14  $\mu$ m spot, 1.5  $\mu$ m Cu foil
- Production of 200 MeV electrons
- Detected signal ~ 30000 photons per shot





# **HOPG spectrometer performance**

normalized transmission

- Crystal spectrometer showed excellent spectral 5.5 eV (E/dE = 1600)
- Single hit analysis
- Angular resolution
- ⇒ distinguish two parts of target



M. Šmíd et al., Rev. Sci. Instrum. 88, 063102 (2017)



Page 40

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# X-ray absorption spectroscopy (XANES)

- 500 µm large spot was directly heated by small part of the laser
- Second part of x-ray beam penetrated unperturbed Cu target
- Signal accumulated over ~ 200 shots
- Statistically significant decrease of XANES features on heated part.
- Experimental proof of ultra-fast Cu melting







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# **Permanent setup at HZDR**

- DRACO (4-5 J, 30 fs, 150 TW)
- 400 MeV electrons
- protons gas target (He+N) HOPG crystal WDM target (Cu) proton target (Au) heater beam (15J, 30fs)
- Achieve higher energy x-ray spectra and stronger signal
- New permanent setup for WDM studies
- Novel thin HOPG-flex crystal spectrometer for better resolution
- Spring 2020: 1 PW laser for direct laser and isochoric x-ray/proton heating

![](_page_42_Picture_8.jpeg)

# Summary

- Warm dense matter (WDM) is a difficult state do describe theoretically
  - Applications in in ICF, planetary science, materials and theoretical physics
  - Transport properties largely unknown with great current interest
  - Requires active x-ray probing
- Novel platform for WDM EOS and transport measurement at OMEGA
  - Imaging X-ray Thomson scattering, VISAR/SOP, X-ray radiography
  - First full model-independent EOS measurement in WDM
- Experiments with short pulse lasers (DRACO facility)
  - LWFA betatron radiation source for ultrafast absorption spectroscopy
  - Heating of Cu samples with lasers or x-rays/protons
  - Study non-thermal melting and bond-hardening
  - New instrument design (x-ray spectrometers, detectors)

![](_page_43_Picture_13.jpeg)

# **OMEGA experiment collaboration**

J. F. Benage, L. A. Collins, C. L. Fryer, G. Kagan, J. D. Kress, D. S. Montgomery, B. Srinivasan, R. G. Watt, J. R. Williams and D. W. Schmidt

Los Alamos National Laboratory

E. J. Gamboa, P. A. Keiter, S. Klein, R. P. Drake University of Michigan

C. McCoy and T. R. Boehly University of Rochester

**P. Tzeferacos and D. Lamb** University of Chicago

Milan Holec Lawrence Livermore National Laboratory

Katerina Falk, Michal Šmíd Helmholtz-Zentrum Dresden-Rossendorf

Technical and scientific staff at the Laboratory for Laser Energetics, University of Rochester

![](_page_44_Picture_9.jpeg)

![](_page_44_Picture_10.jpeg)

![](_page_44_Picture_11.jpeg)

![](_page_44_Picture_12.jpeg)

![](_page_44_Picture_13.jpeg)

![](_page_44_Picture_14.jpeg)

![](_page_44_Picture_15.jpeg)

# Lund/DRACO experiment collaboration

### M. Šmíd, J. Couperus, A. Köhler, T. Kurz, J. Vorberger, P. Perez-Martin, X. Pan, A. Irman, K. Falk Helmholtz-Zentrum Dresden-Rossendorf, Germany

### I. Gallardo Gonzáles, M. Hansson, O. Lundh

Department of Physics, Lund Laser Center, Lund, Sweden

### J. C. Wood, S. P. D. Mangles

John Adams Institute for Accelerator Science, Imperial College London, United Kingdom

Lund experiment funded by LASERLAB.

![](_page_45_Picture_7.jpeg)

![](_page_45_Picture_8.jpeg)

![](_page_45_Picture_9.jpeg)

![](_page_45_Picture_10.jpeg)

# Thank you for your attention! Any questions?

# Katerina Falk

Helmholtz Young Investigator Group Leader

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![](_page_46_Picture_4.jpeg)

HELMHOLTZ RESEARCH FOR GRAND CHALLENGES

![](_page_46_Picture_6.jpeg)

![](_page_46_Picture_7.jpeg)