

# Dark matter searches at ATLAS & CMS.

**Priscilla Pani**  
(DESY)

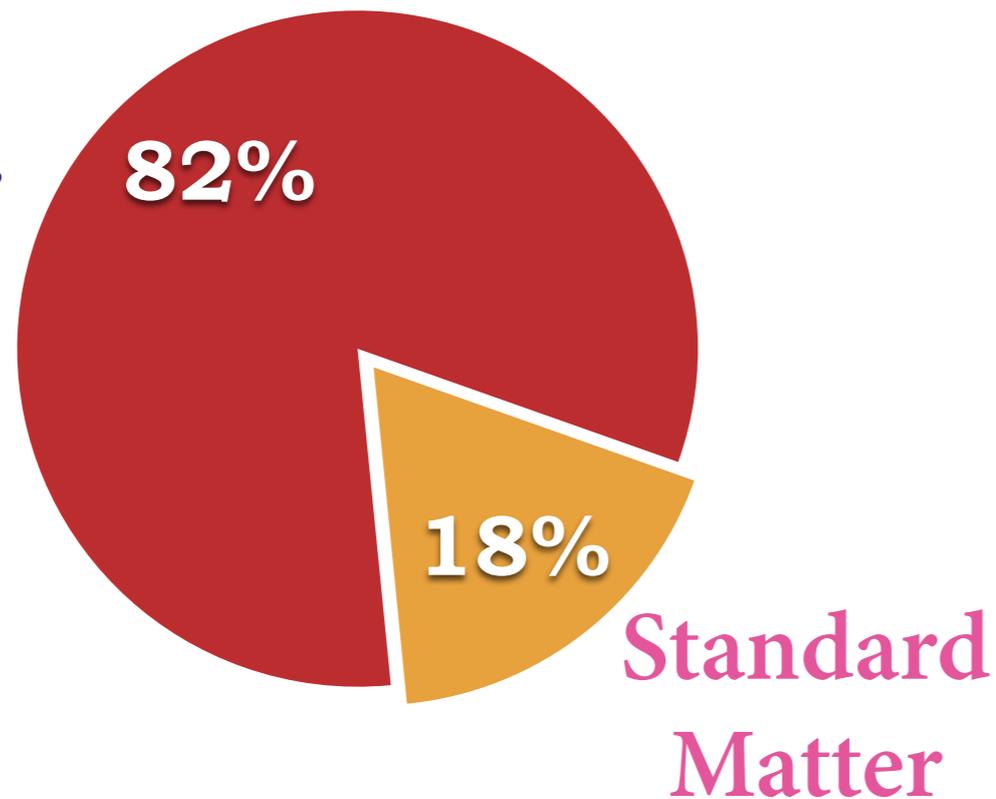


**HELMHOLTZ**  
RESEARCH FOR GRAND CHALLENGES



# The Dark Matter mystery

Dark  
Matter



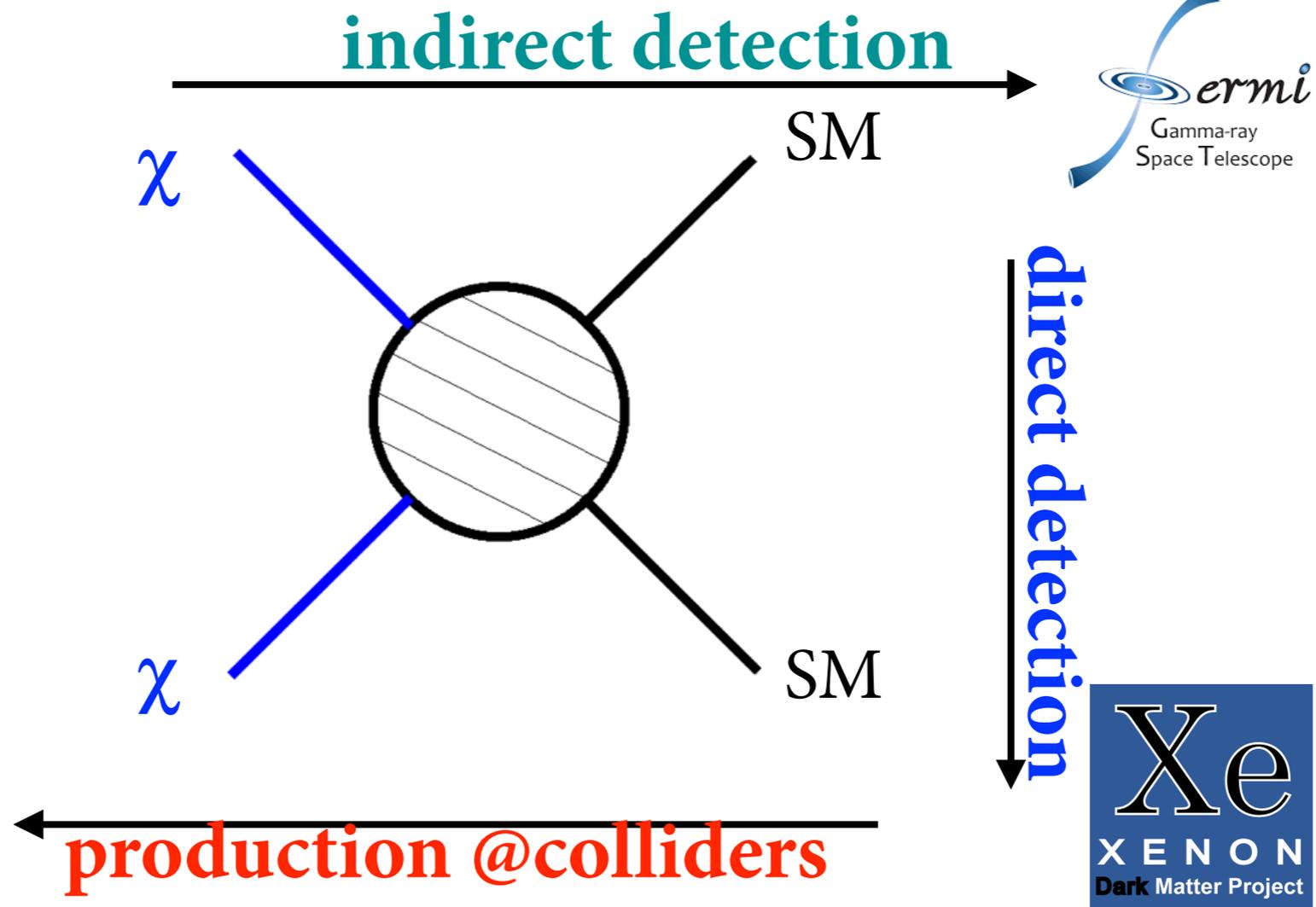
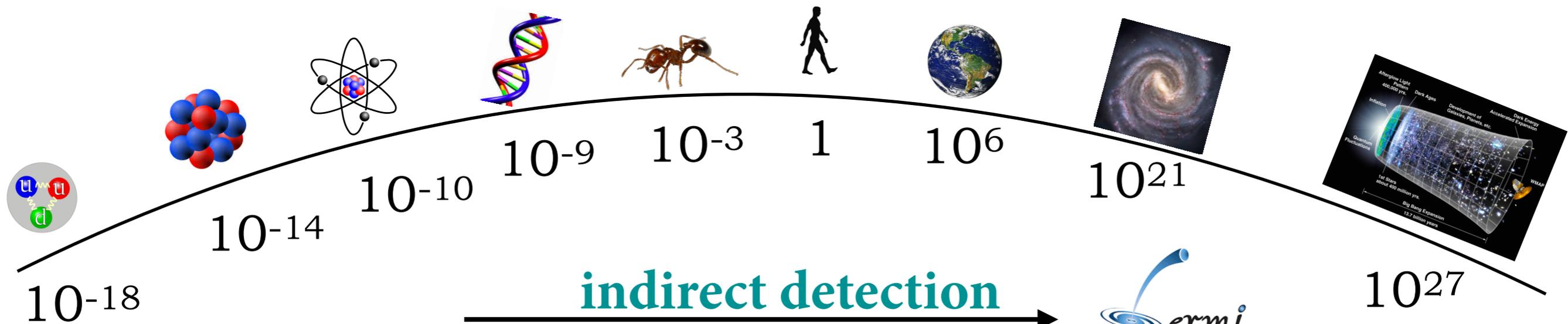
Dark Matter in Galaxy Merger 1E 0657-558



- ❖ Electrically neutral
- ❖ Observed via gravity, massive
- ❖ Weakly interacting
- ❖ Elementary particles created in the early universe

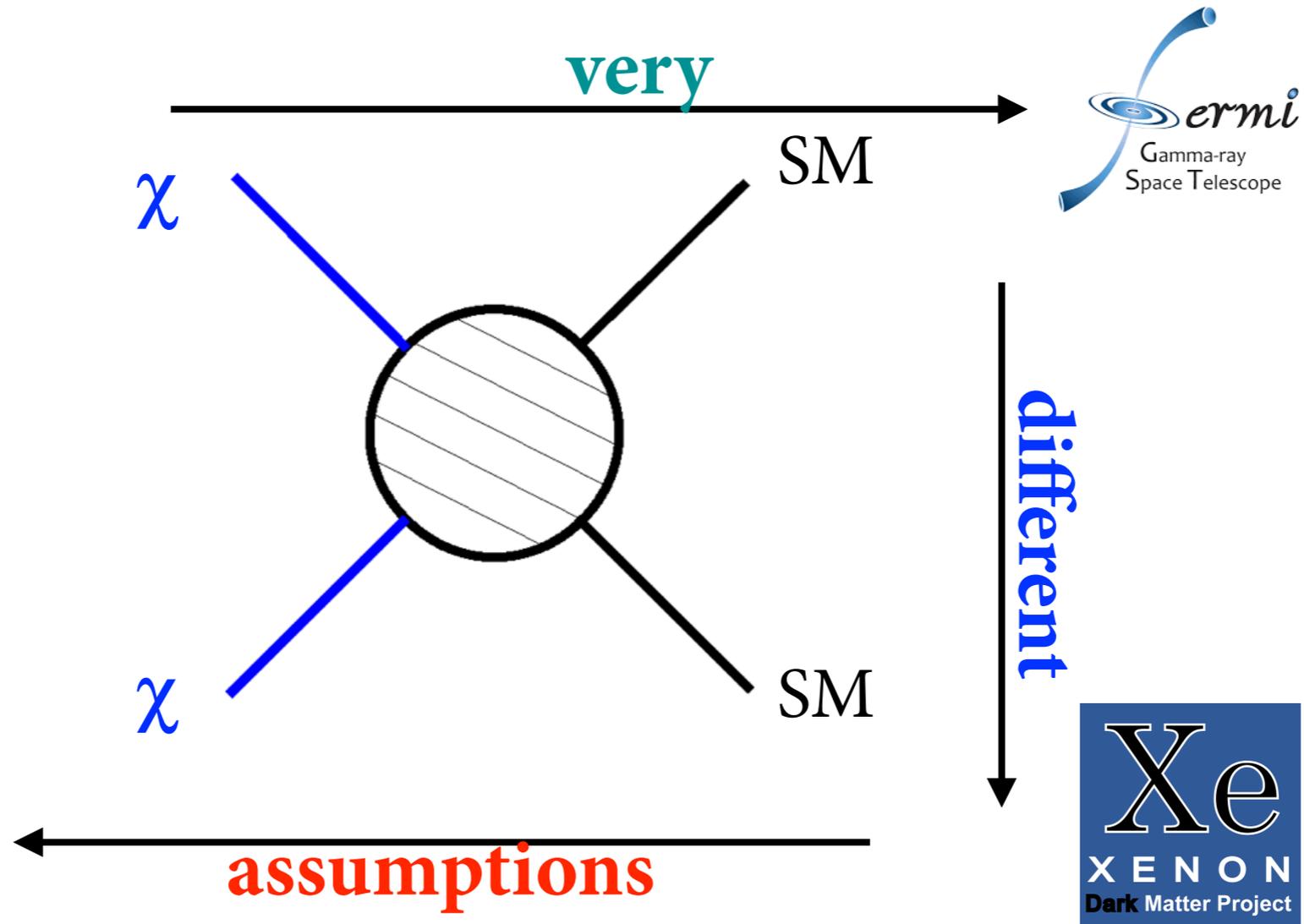
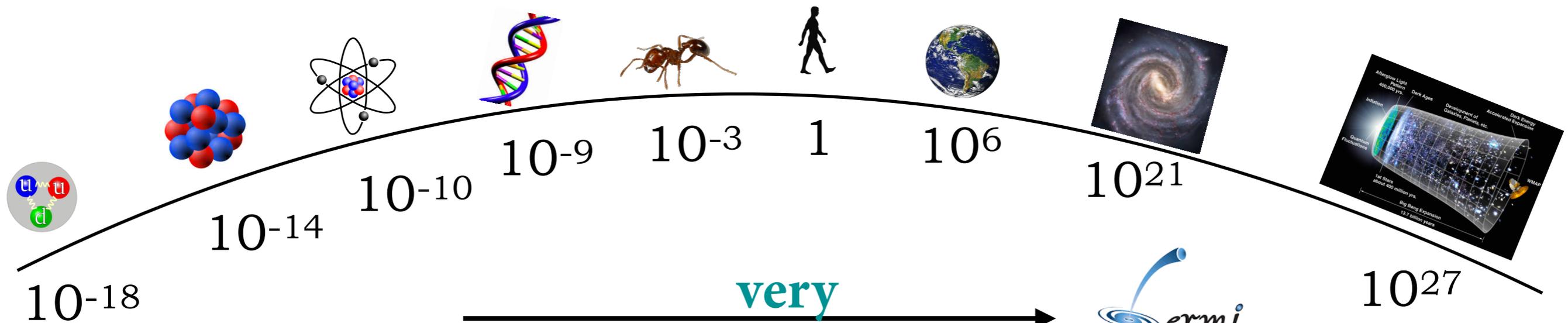
# The Dark Matter quest

universe scales in meters

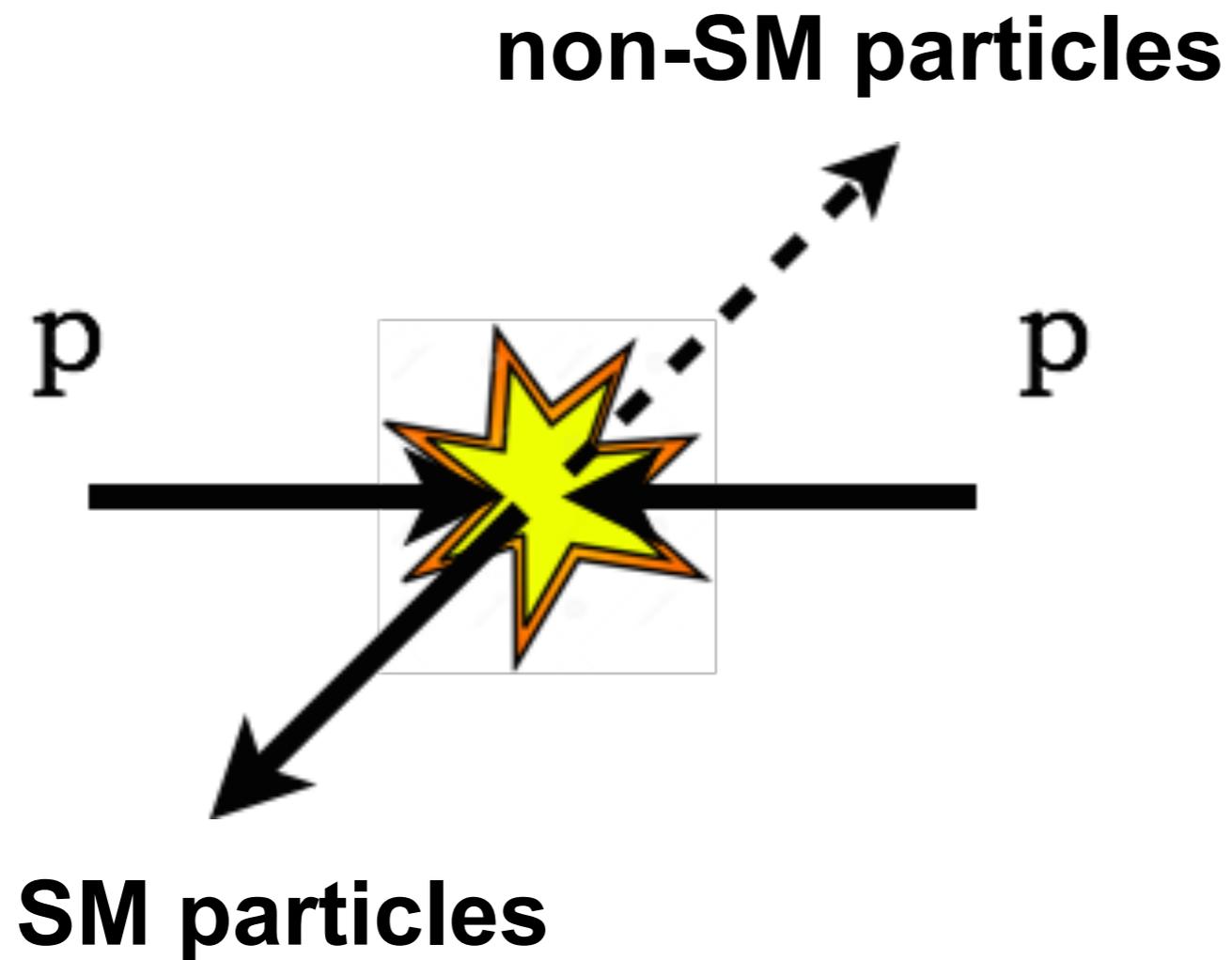


# The Dark Matter quest

universe scales in meters

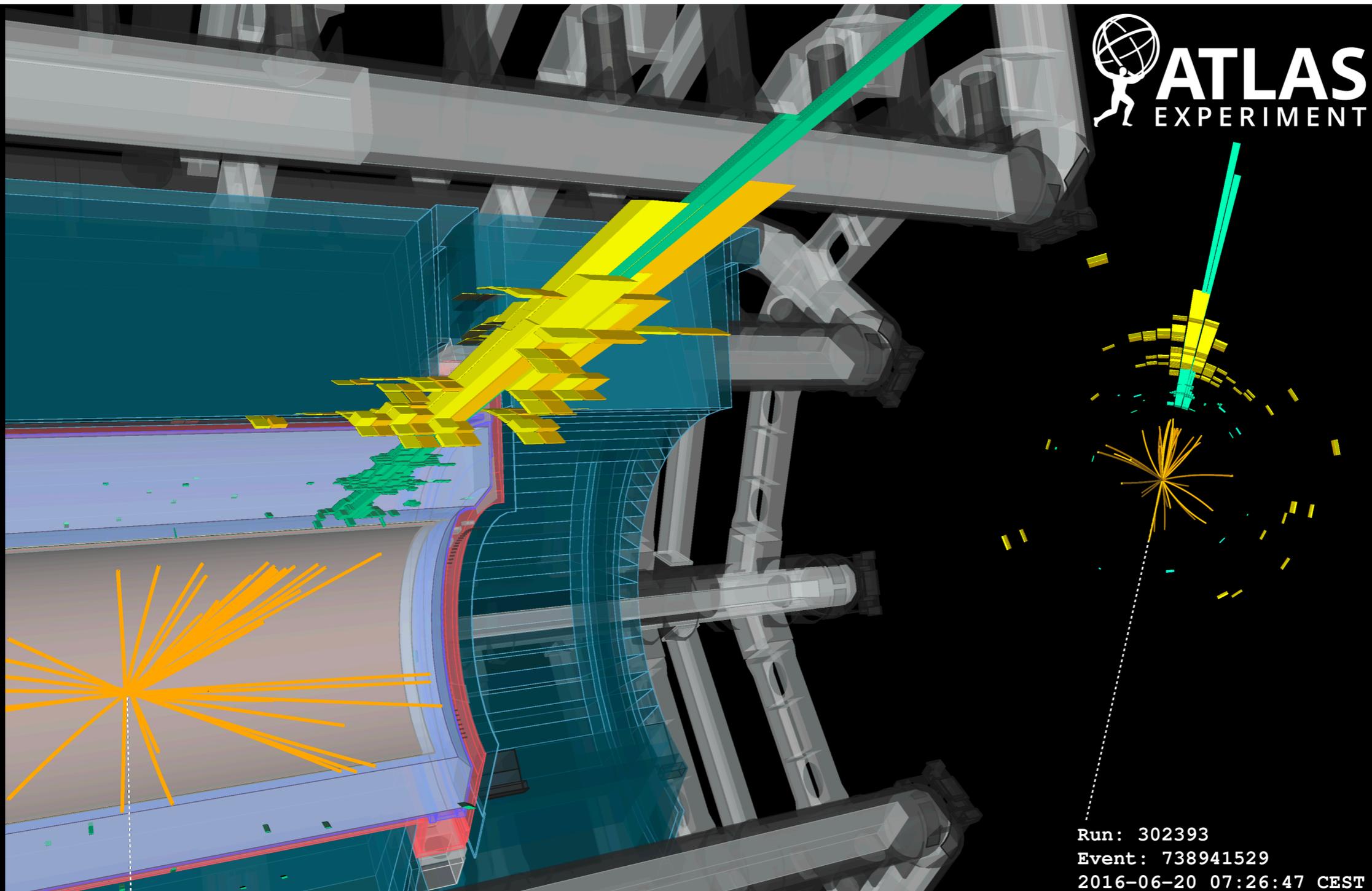


# The collider ansatz

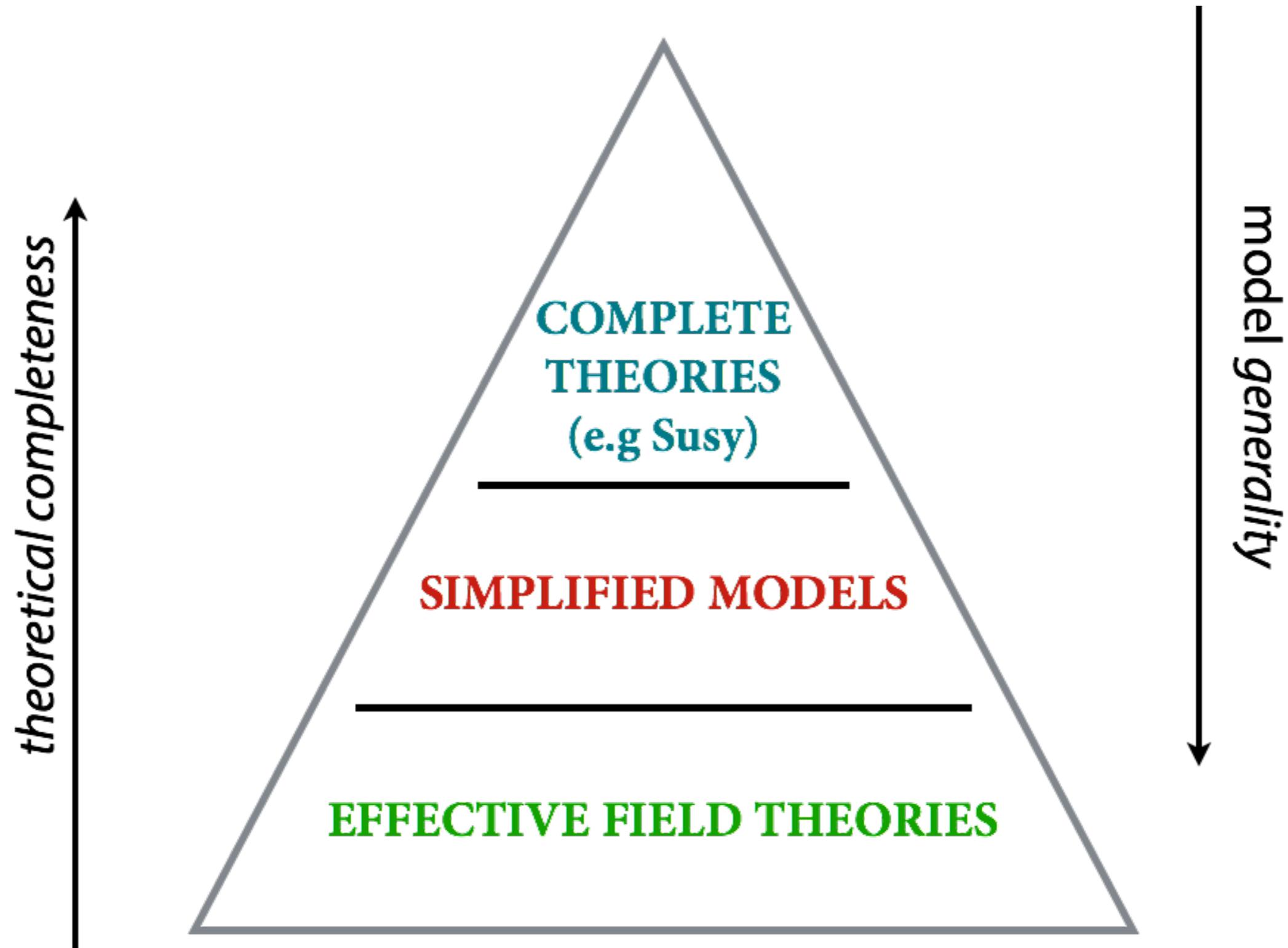


1. Production mechanism / theoretical framework
2. Particles detection and identification

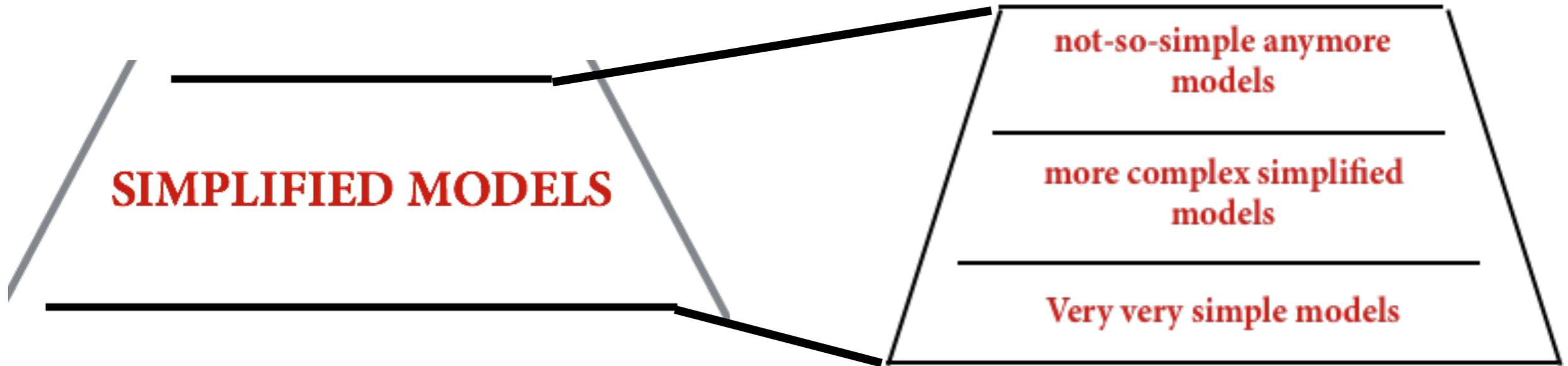
# 1. Production mechanism



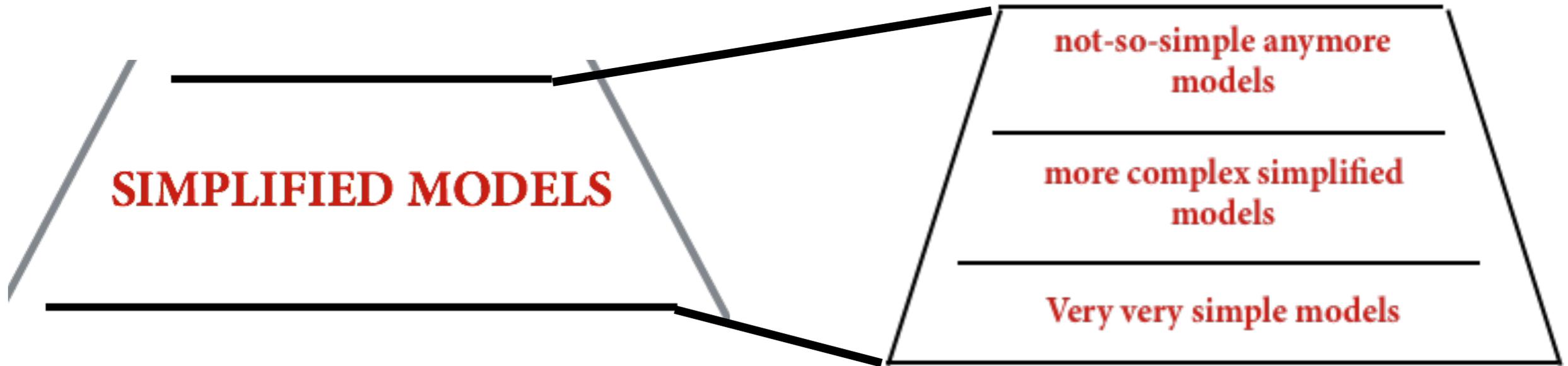
# Theoretical framework



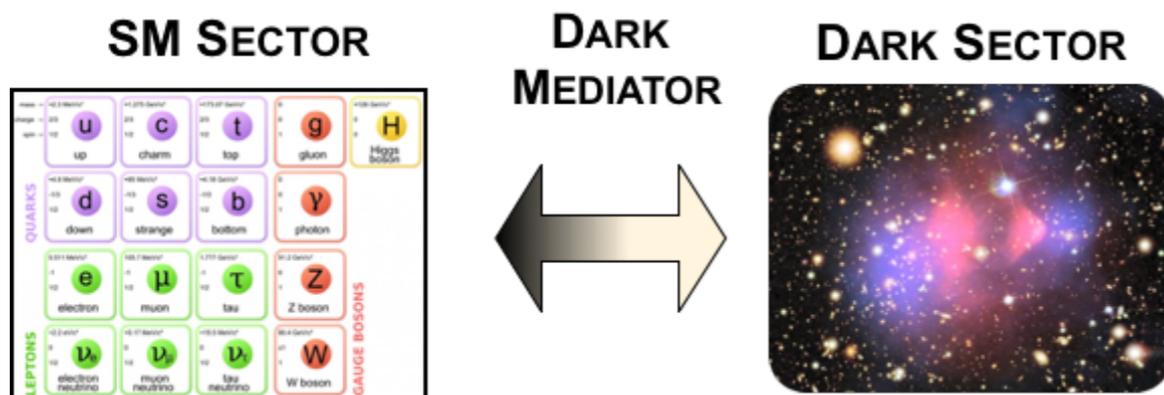
# Theoretical framework



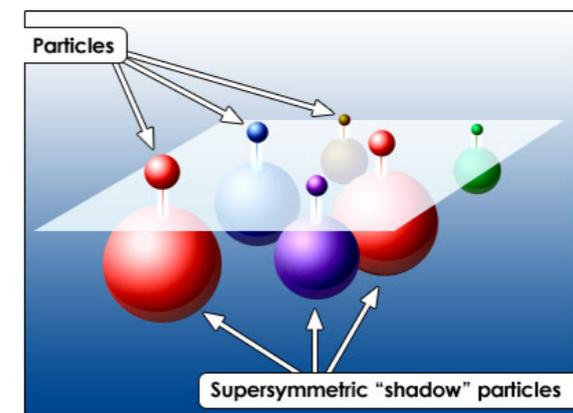
# Theoretical framework



*“Mediator-based DM models”*

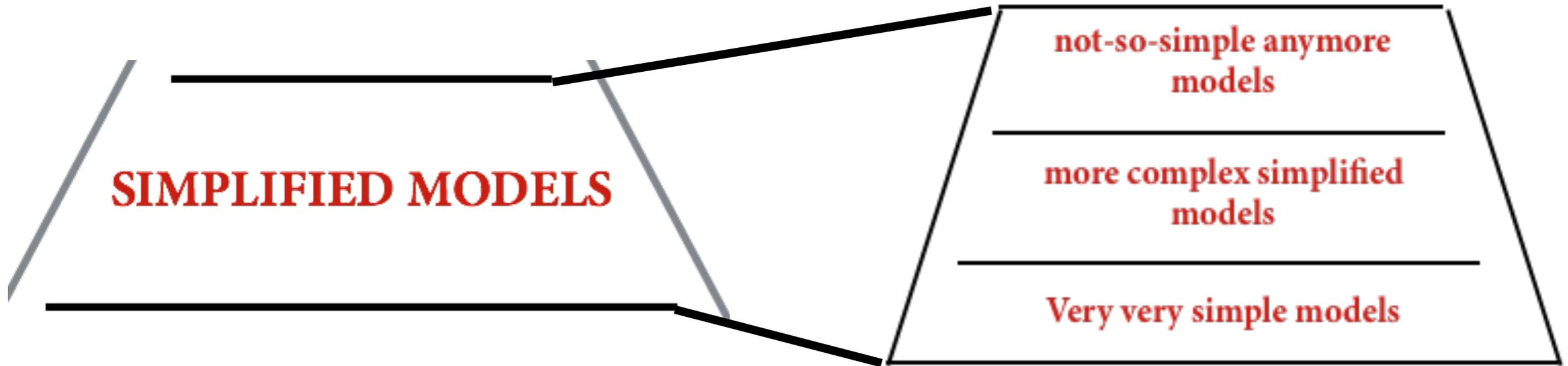


*“SUSY simplified models”*

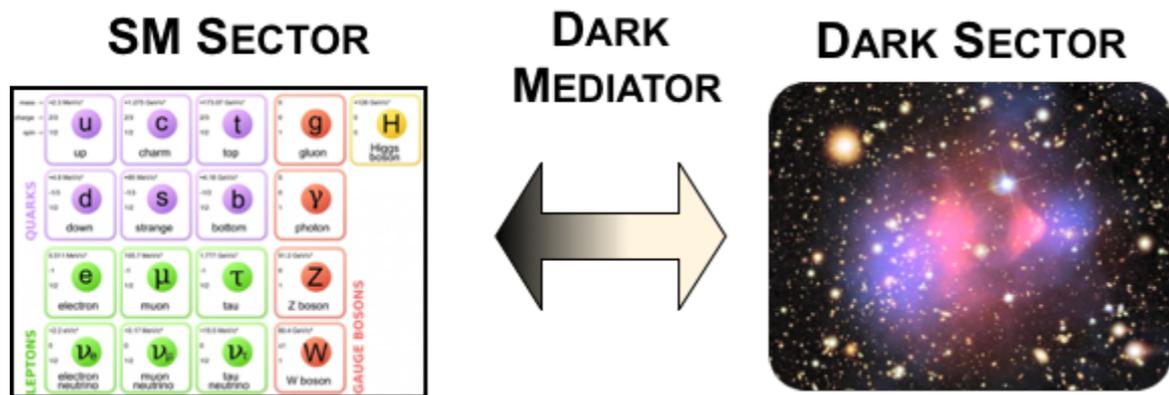


+ *“Higgs Portal DM models”* + axions

# Theoretical framework

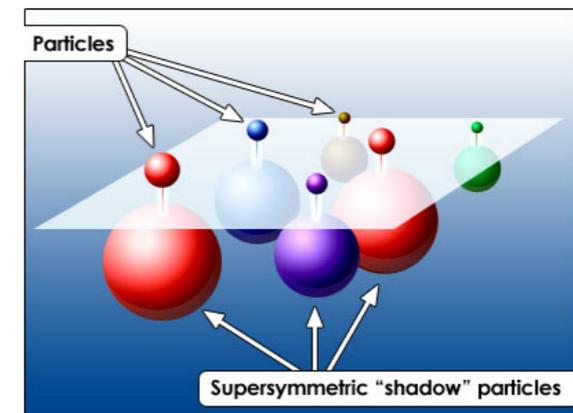


*“Mediator-based DM models”*



**not covered further**

*“SUSY simplified models”*

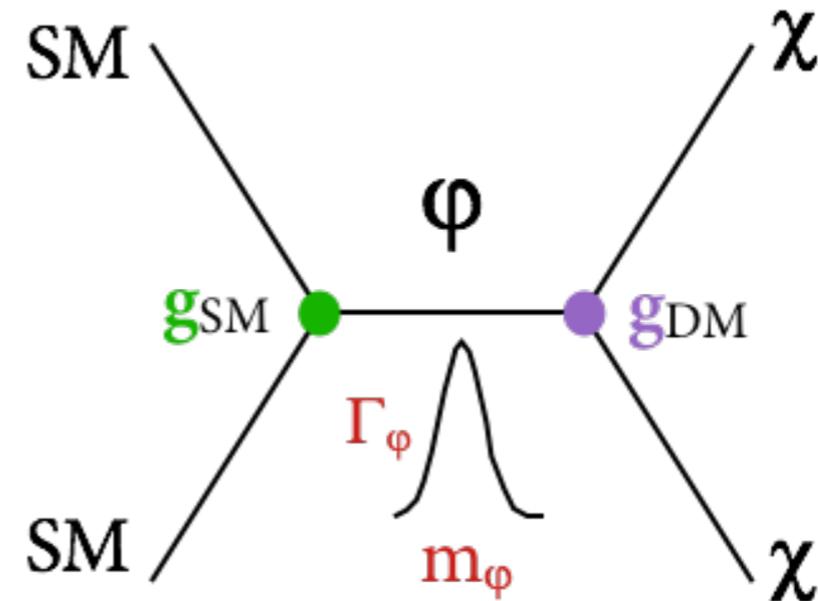


+ *“Higgs Portal DM models”* + axions

# Mediator simplified models

- ★ Reduce a complex model to a simple one with **DM + mediator**
- ★ **Few free parameters:**  $m_\phi$ ,  $m_\chi$ ,  $g_{SM}$ ,  $g_{DM}$ ,  $\Gamma_\phi$
- ★ Nature of mediator and DM can (also) be **systematically classified based on their spin and CP**
- ★ **Very rich phenomenology**

arXiv:1507.00966 (and ref. therein) + [LPCC WG](#)



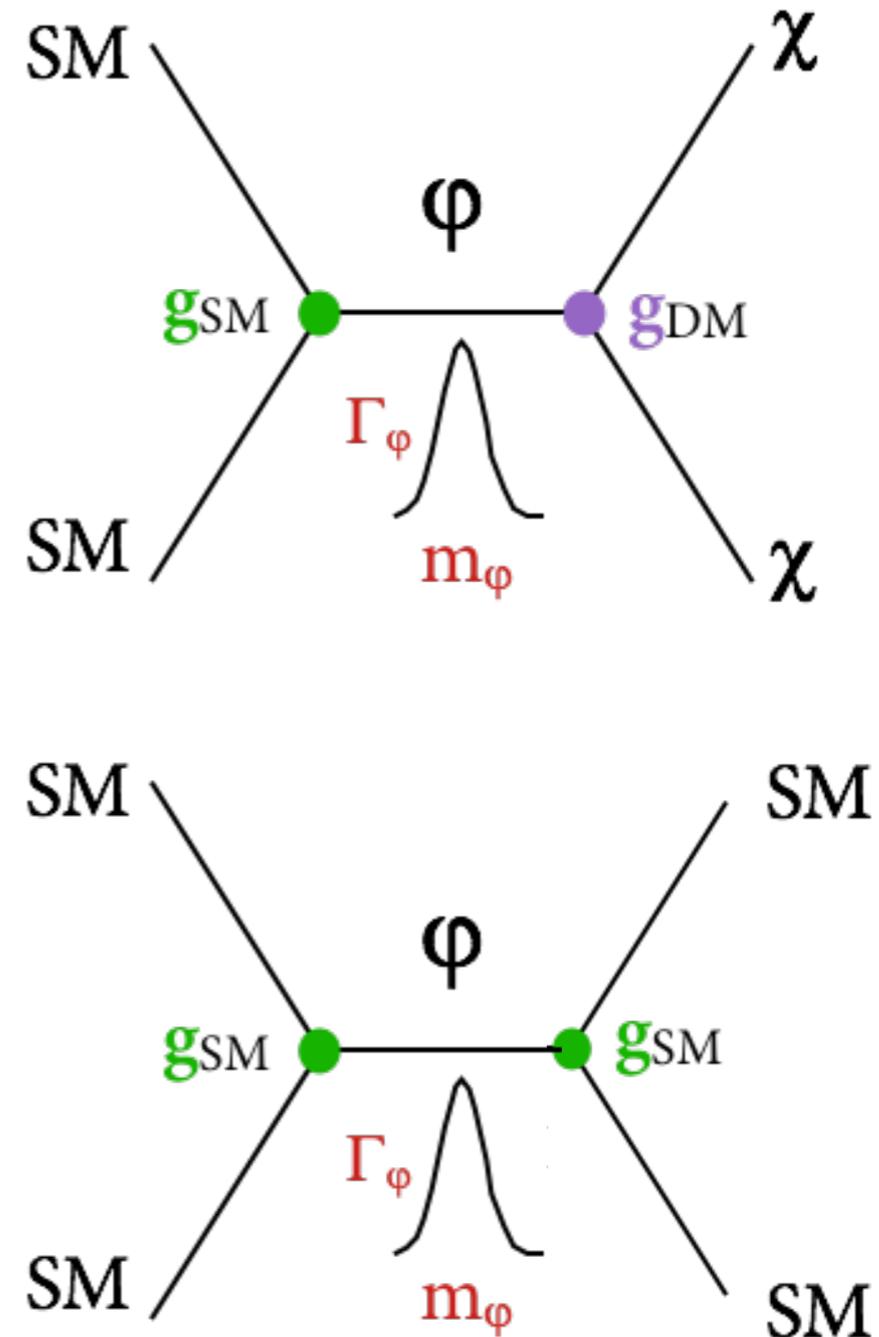
$E_{T\text{-miss}} + X$

**Selected results on spin-0 and spin-1 mediators in the following**

# Mediator simplified models

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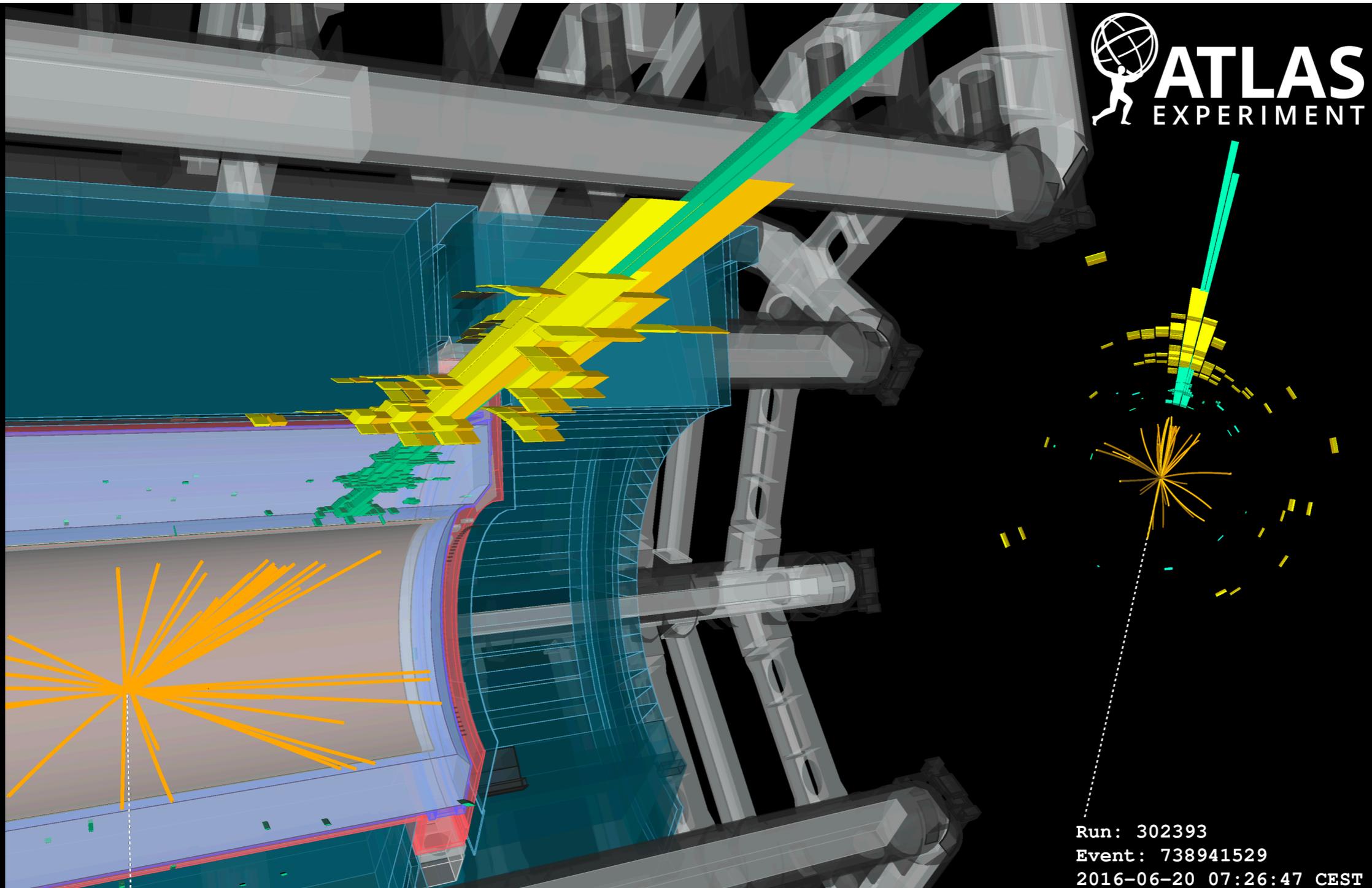


$E_{T\text{-miss}} + X$

Resonance

Selected results on spin-0 and spin-1 mediators in the following

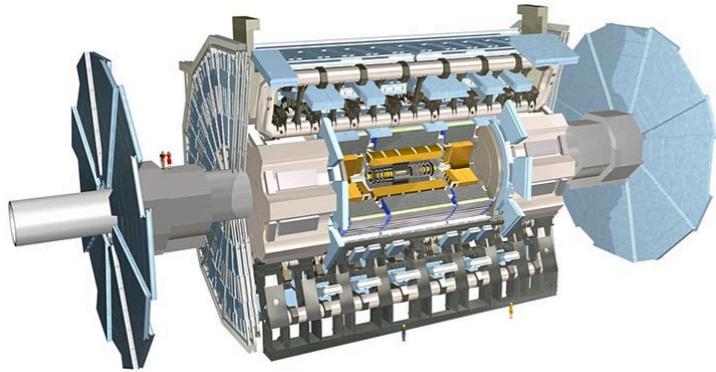
# 2. Detection and identification



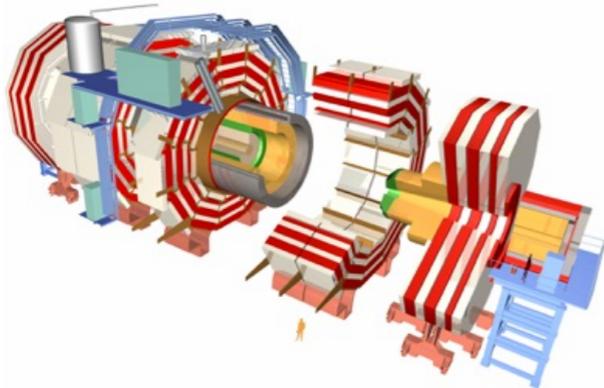
 **ATLAS**  
EXPERIMENT

Run: 302393  
Event: 738941529  
2016-06-20 07:26:47 CEST

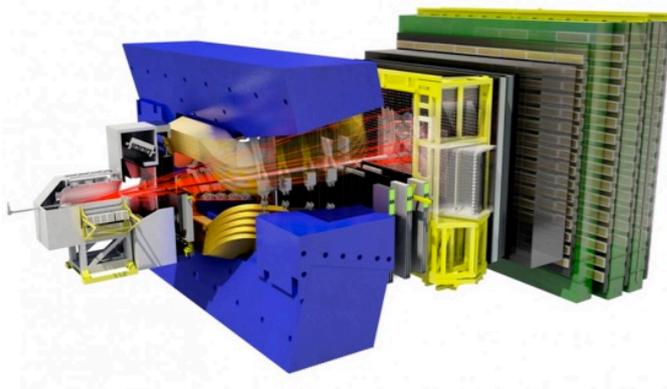
# Summary of DM Collider experiments



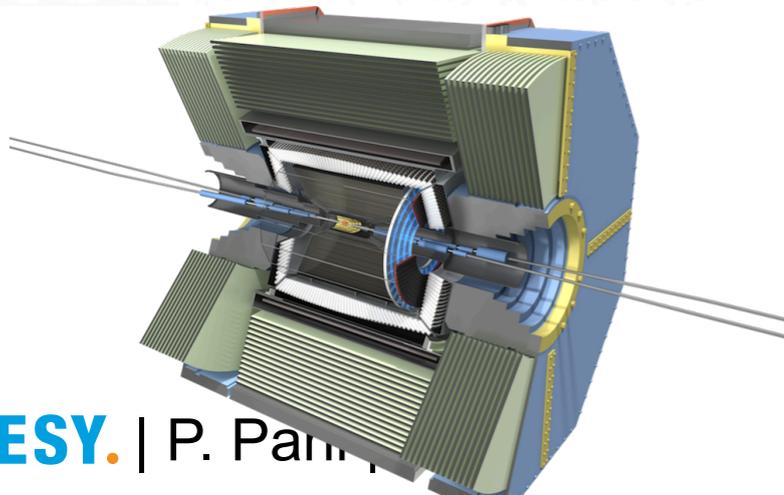
Focus Mediator-models & SUSY  
DM Results [Public Page](#)  
Overview : [DM Summary Paper](#)



Focus Mediator-models & SUSY  
DM Results [EXOTICA](#), [B2G](#)  
Overview (2018): [DM summary plots](#)

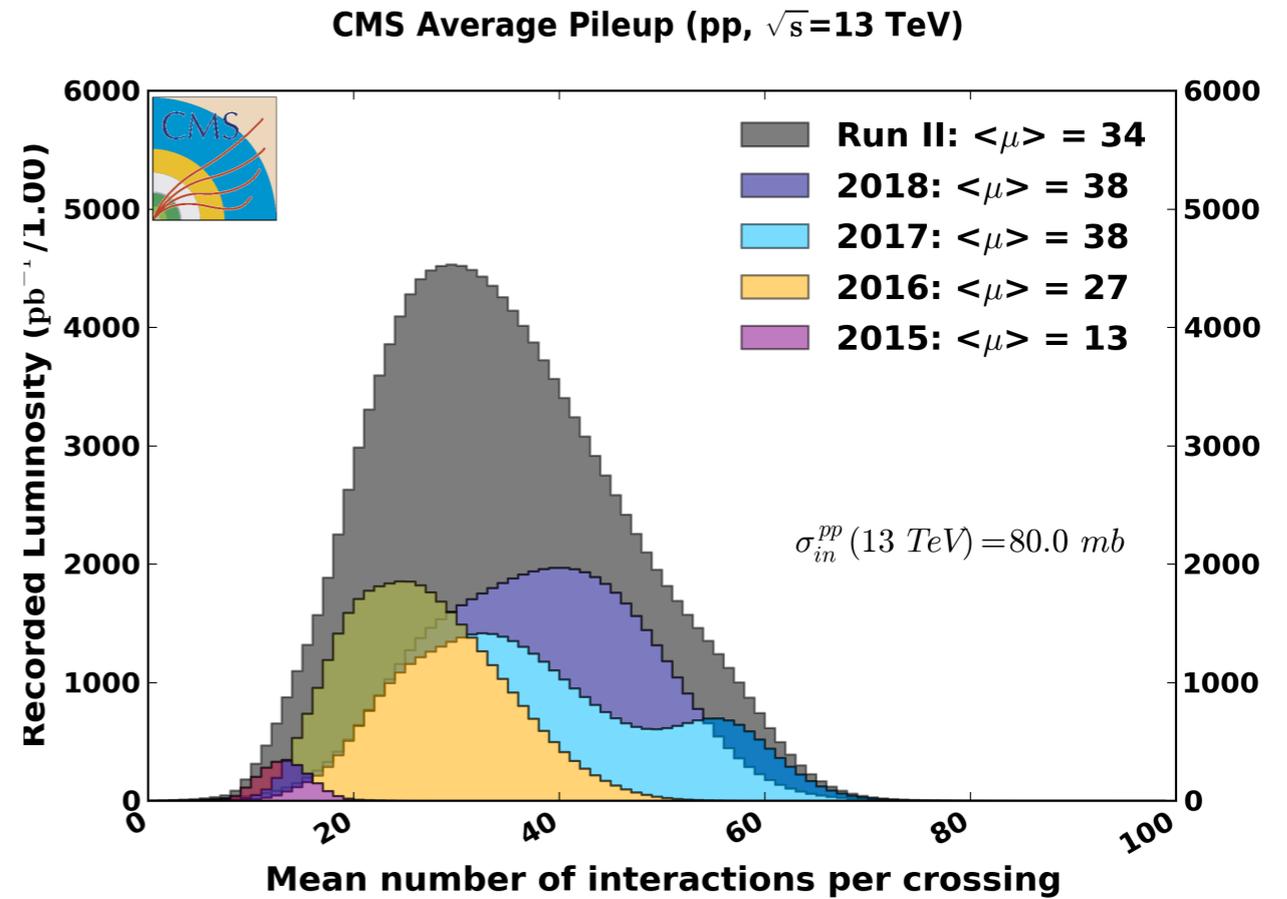
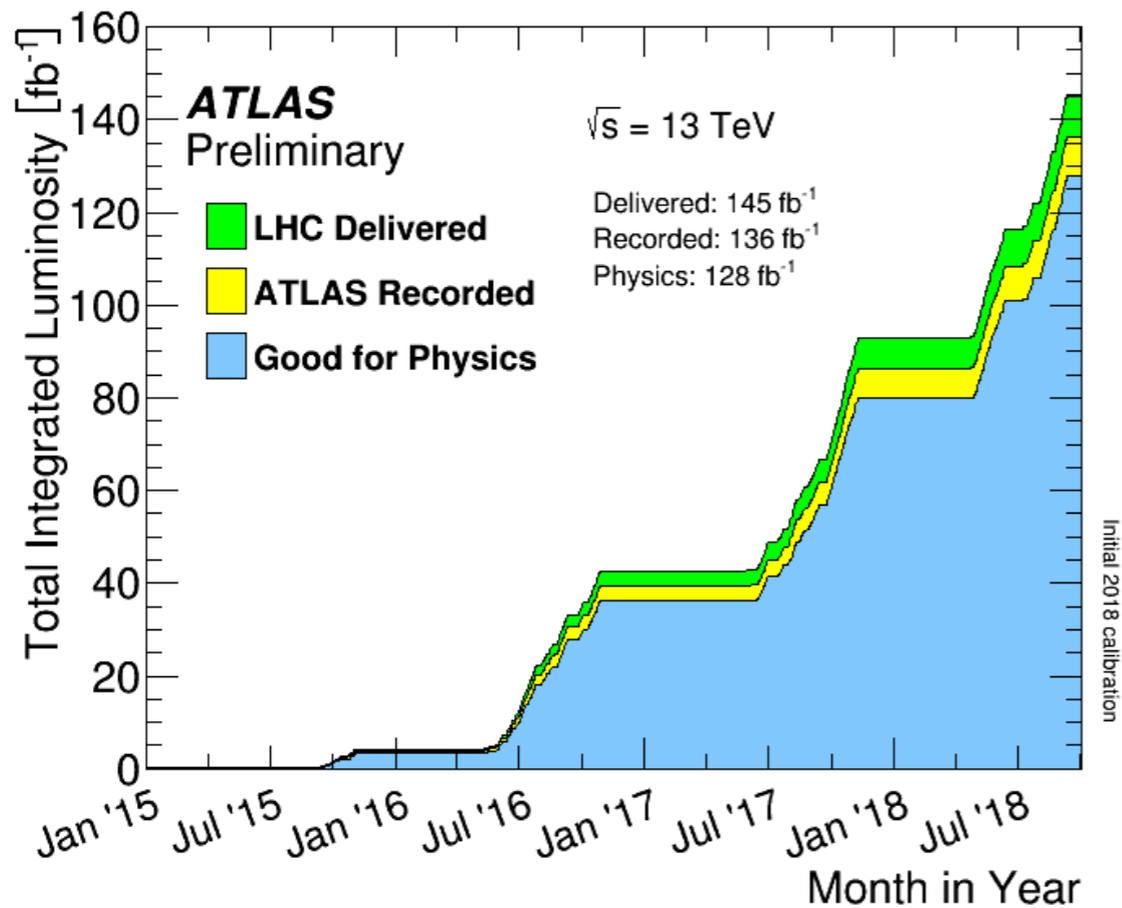


Focus B-mesons, loops, resonance  
DM Results [Public page](#)



Focus B-mesons, dark sector  
DM Results [DMPuzzle2018](#), [BelleII Book](#)

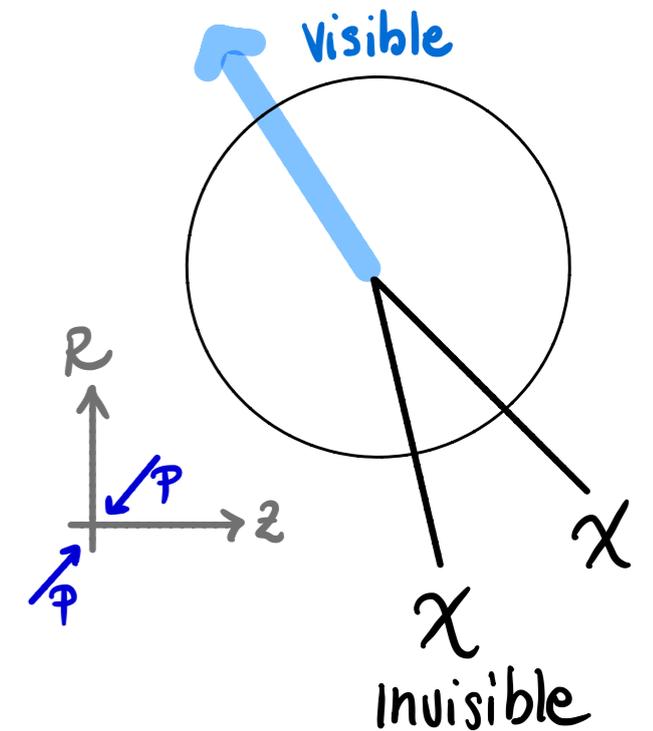
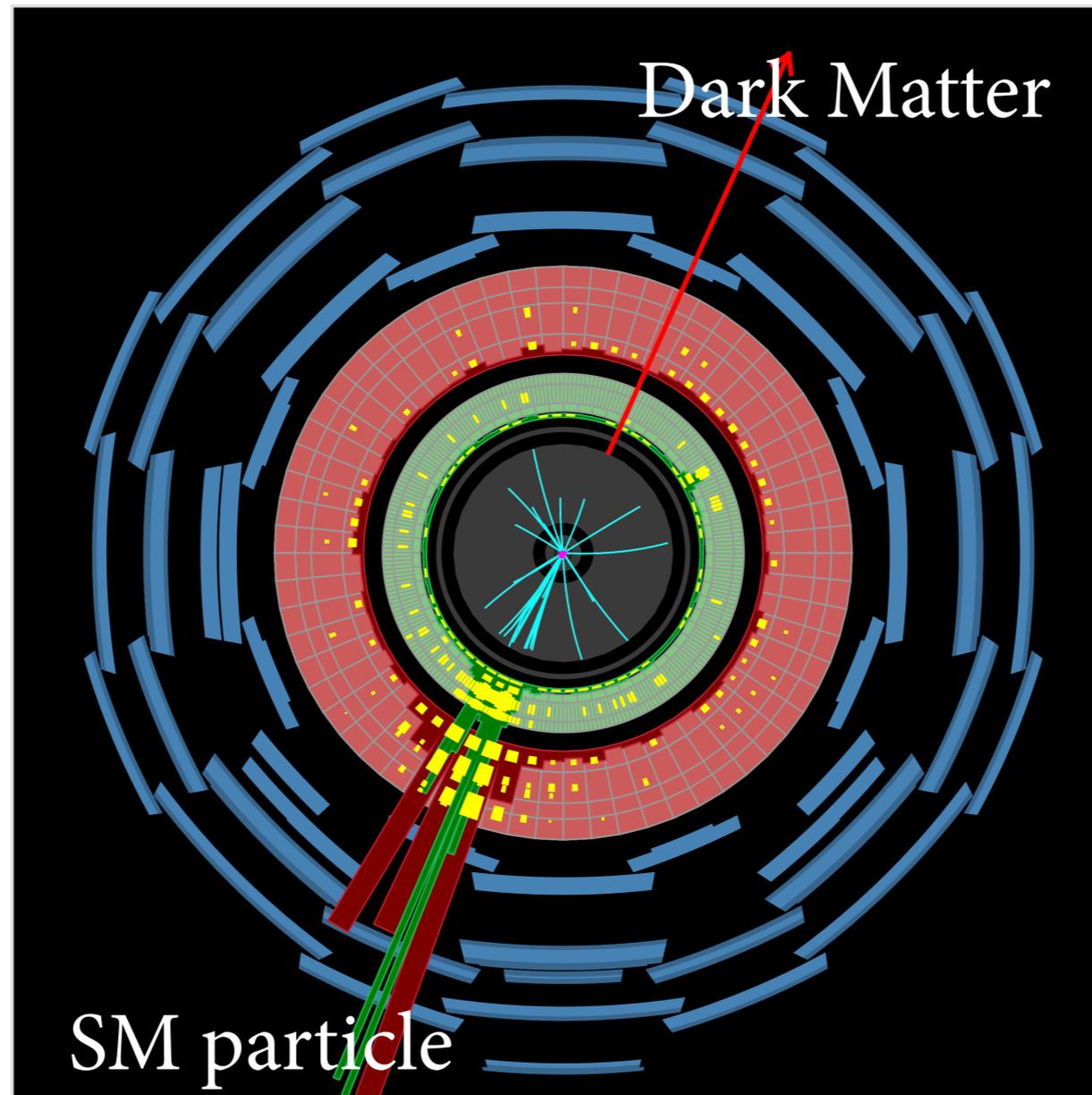
# Timelines and datasets



# Particles detection

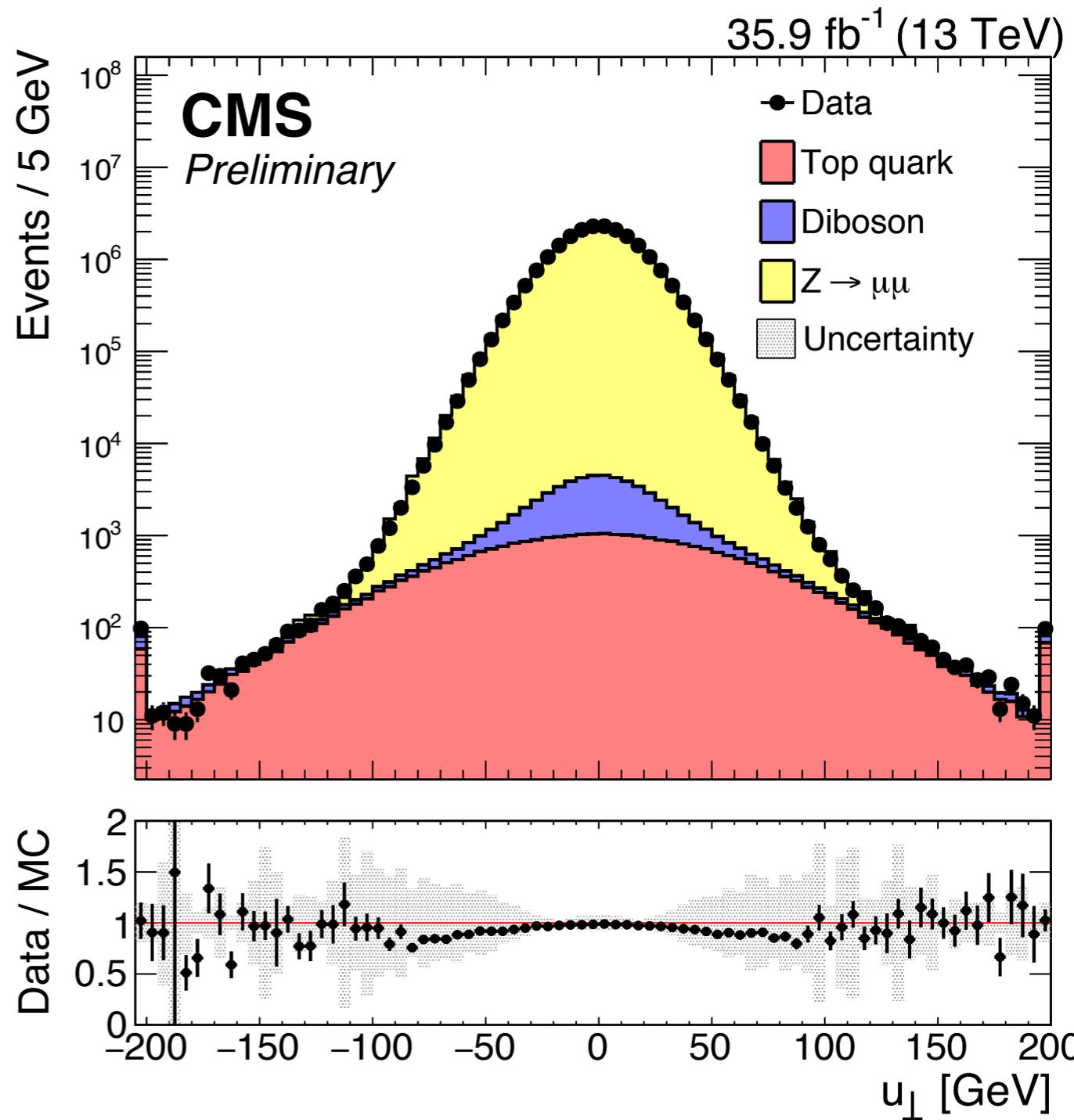
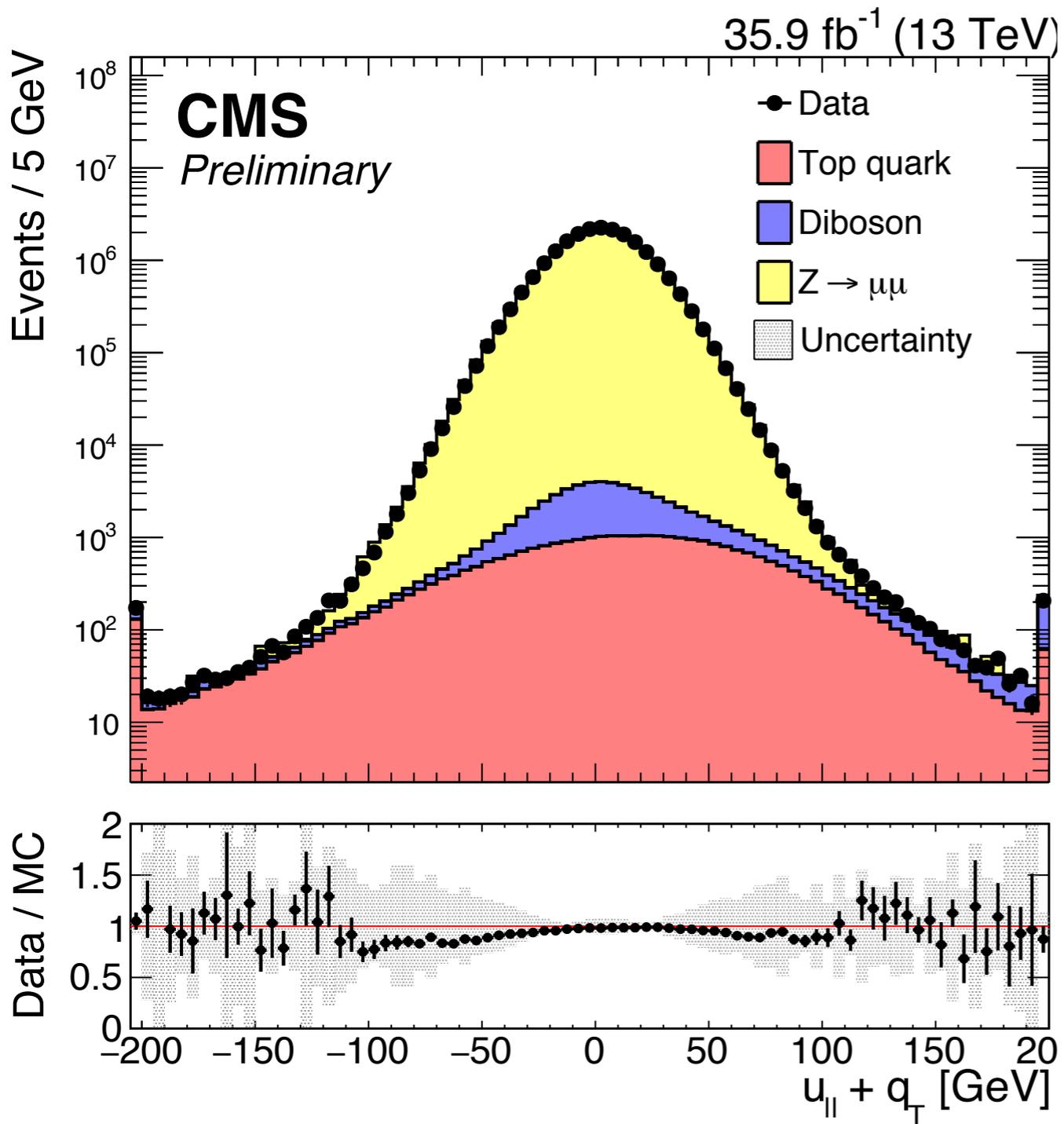
Particles produced in the collision are detected as analogue signals by the sub-detectors, digitised, recorded and reconstructed offline as particle-objects.

- Electrons
- Muons
- Photons
- jets
- b-jets/c-jets
- invisible particles



C. Doglioni TeVPa2018

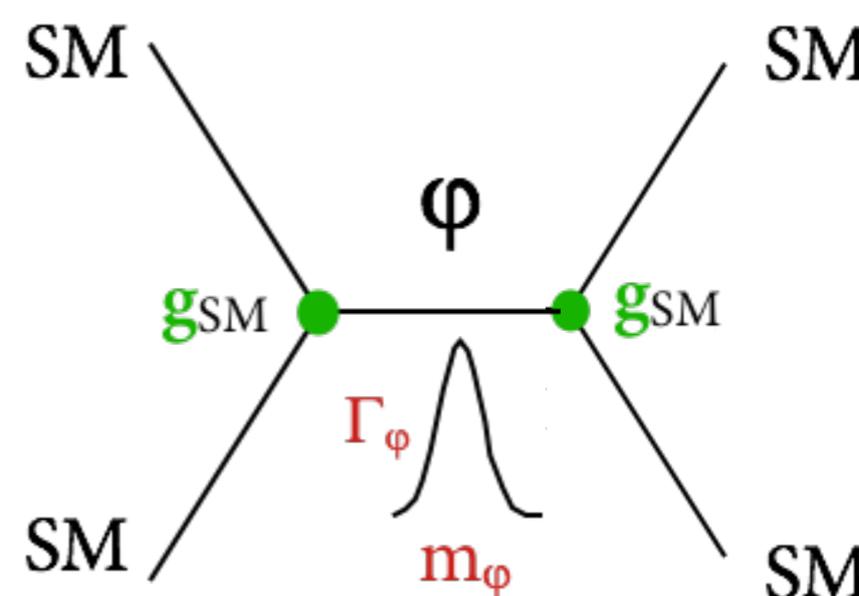
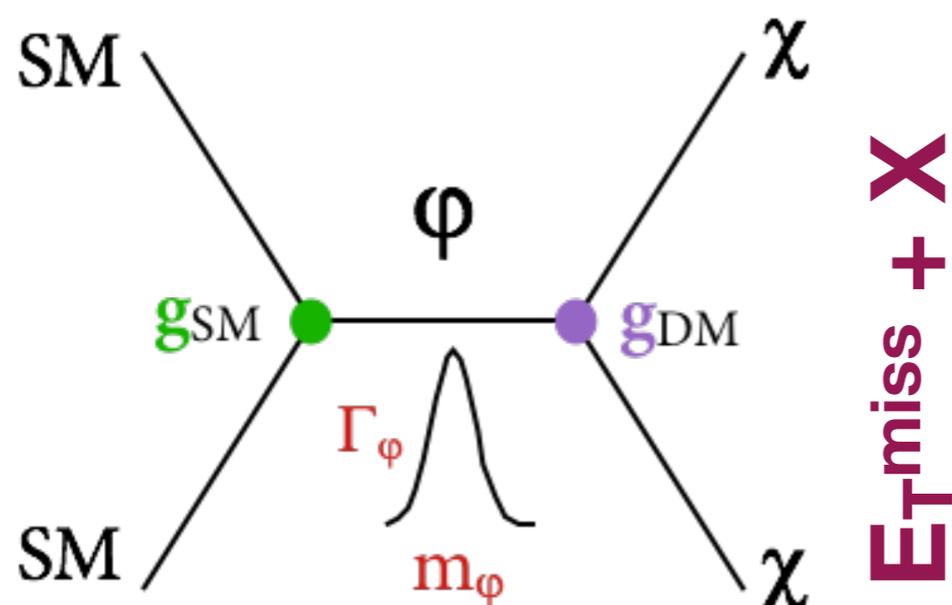
# Missing Energy performance



CMS-PAS-JME-17-001

# Mediator simplified models

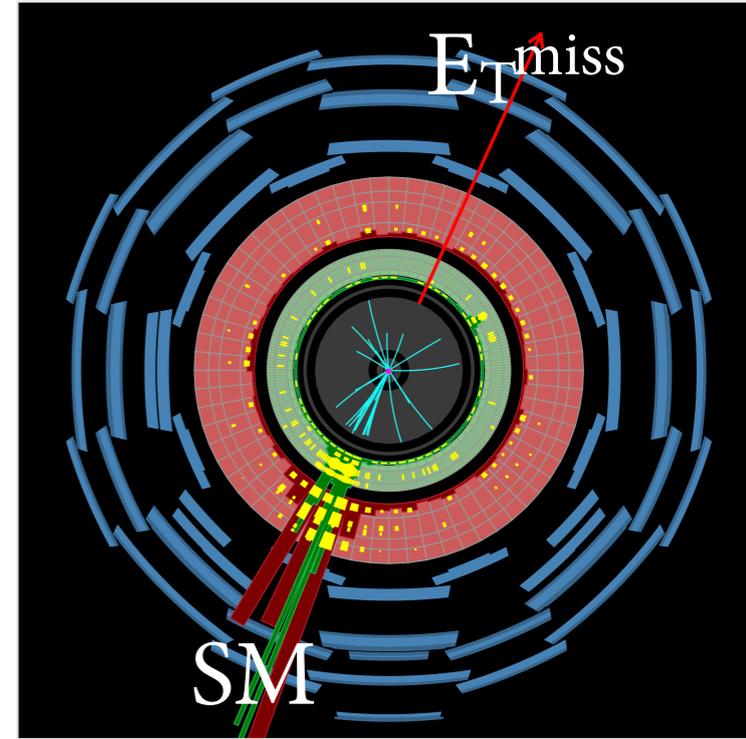
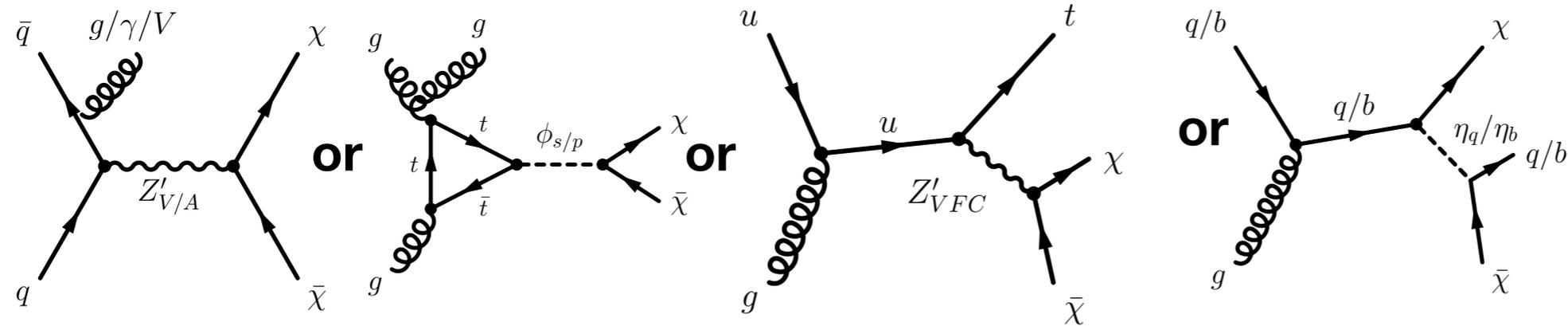
## Reminder



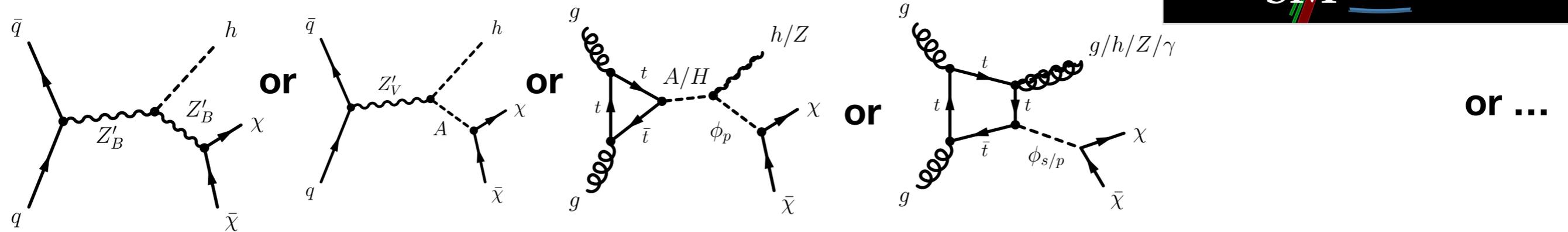
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# Techniques 1 - $E_T^{\text{miss}} + X$

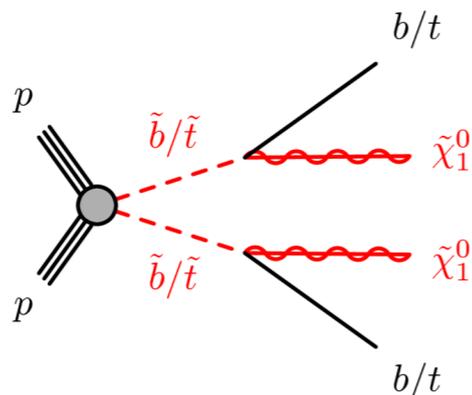
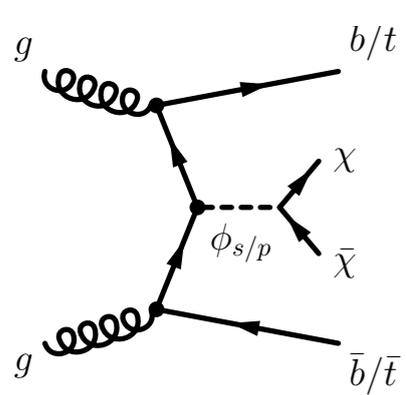
## jet/b-jet/single-t/ $\gamma/V + E_T^{\text{miss}}$



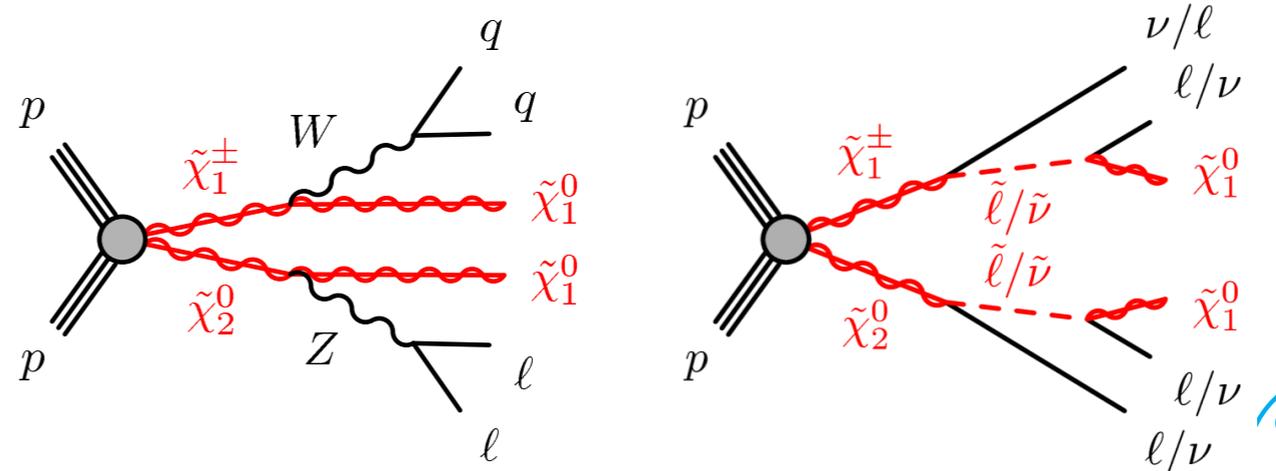
## $h/Z + E_T^{\text{miss}} \rightarrow h(bb/\gamma\gamma)$ Strahlung/decay



## $tt / bb + E_T^{\text{miss}}$

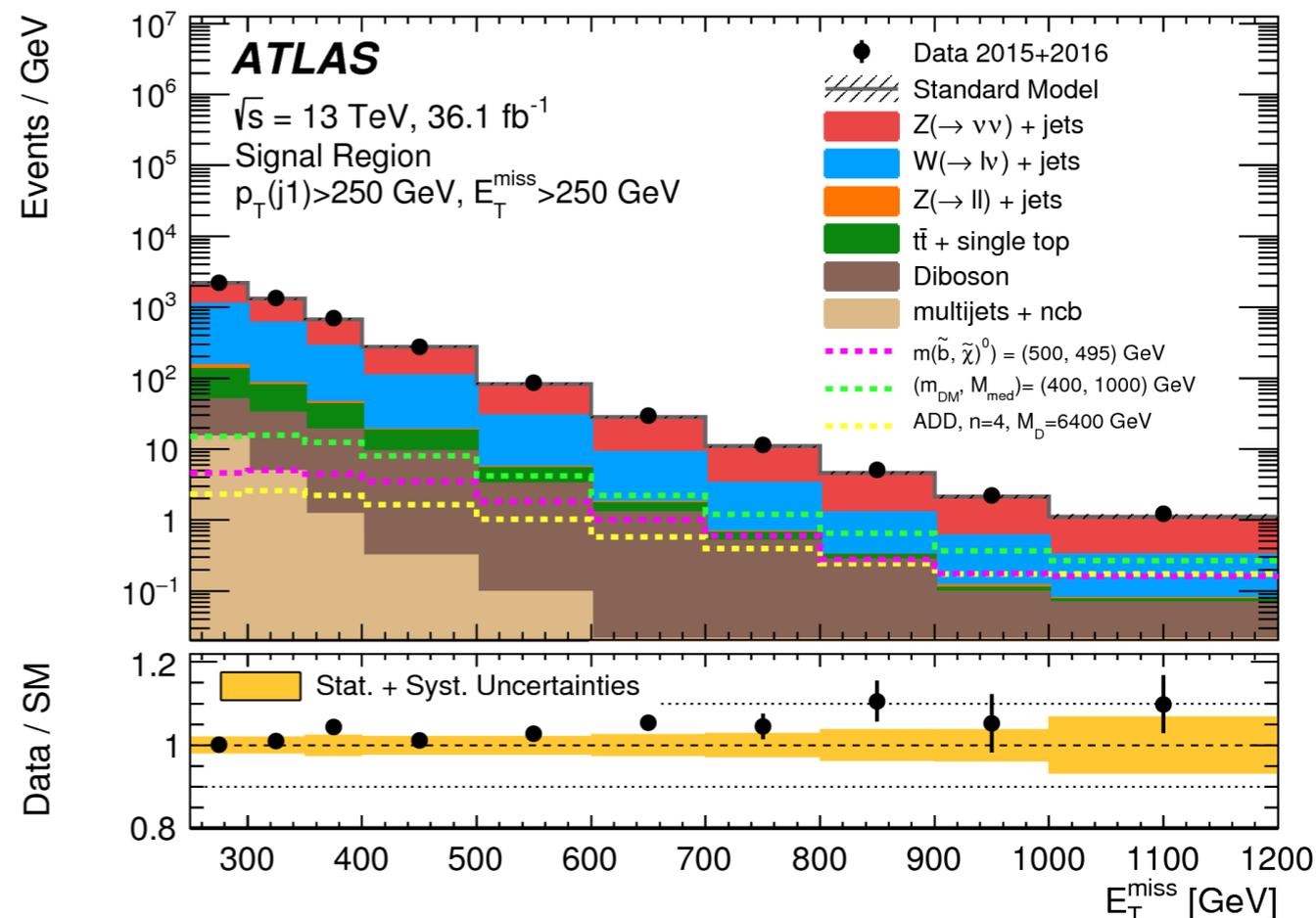


## $2L/2LSS/3L + E_T^{\text{miss}}$

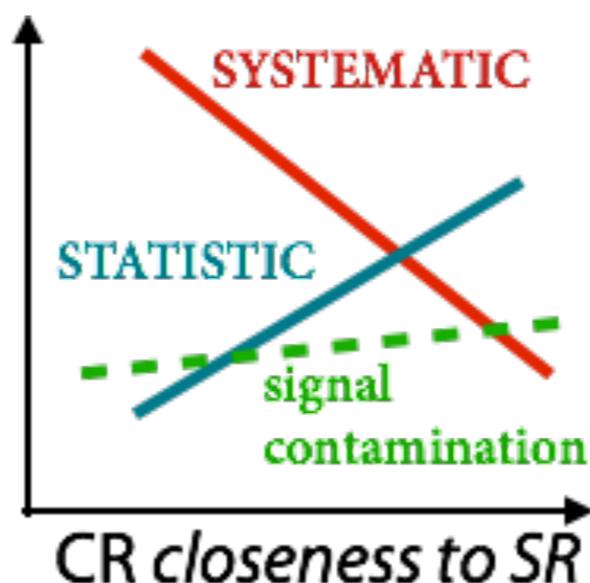
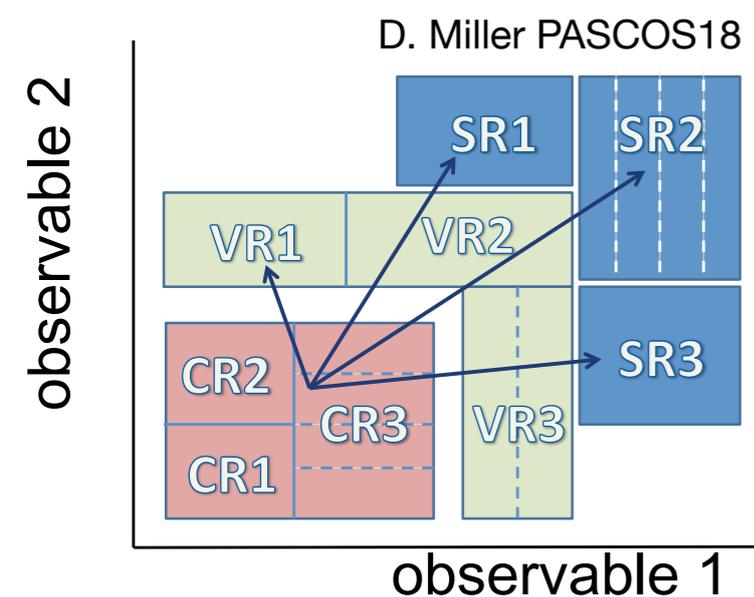


# $E_T^{\text{miss}} + X$ commonalities

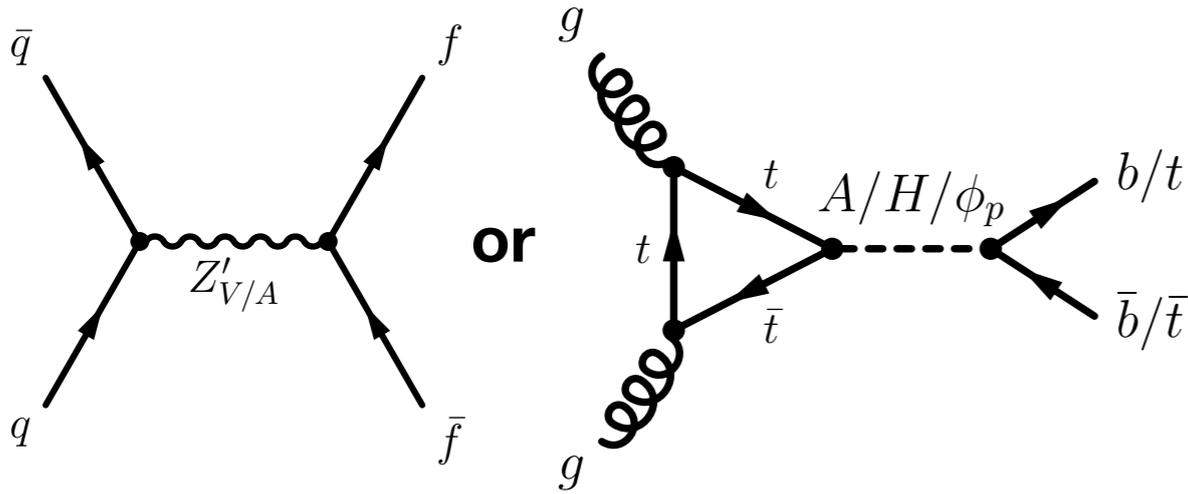
- 1) Definition of a set of Signal enriched Regions (SR)
- 2) Definition of a set of Control Regions (CR) to derive a data-driven normalisation of MC with transfer factors (TF).
- 3) Validation of the TF in the Validation Region (VR)



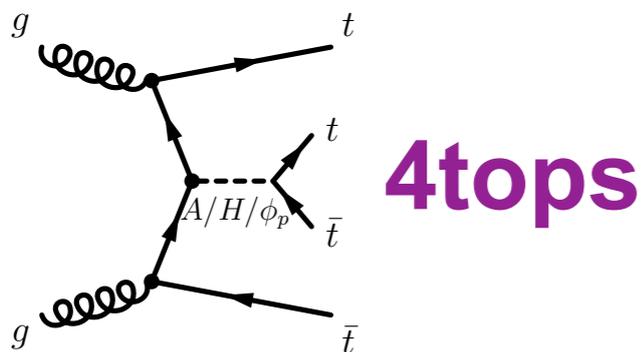
- 4) Unblinding ! check whether an excess is observed (p-value)
- 5) If no excess is found the results are interpreted in terms of limits on selected models.



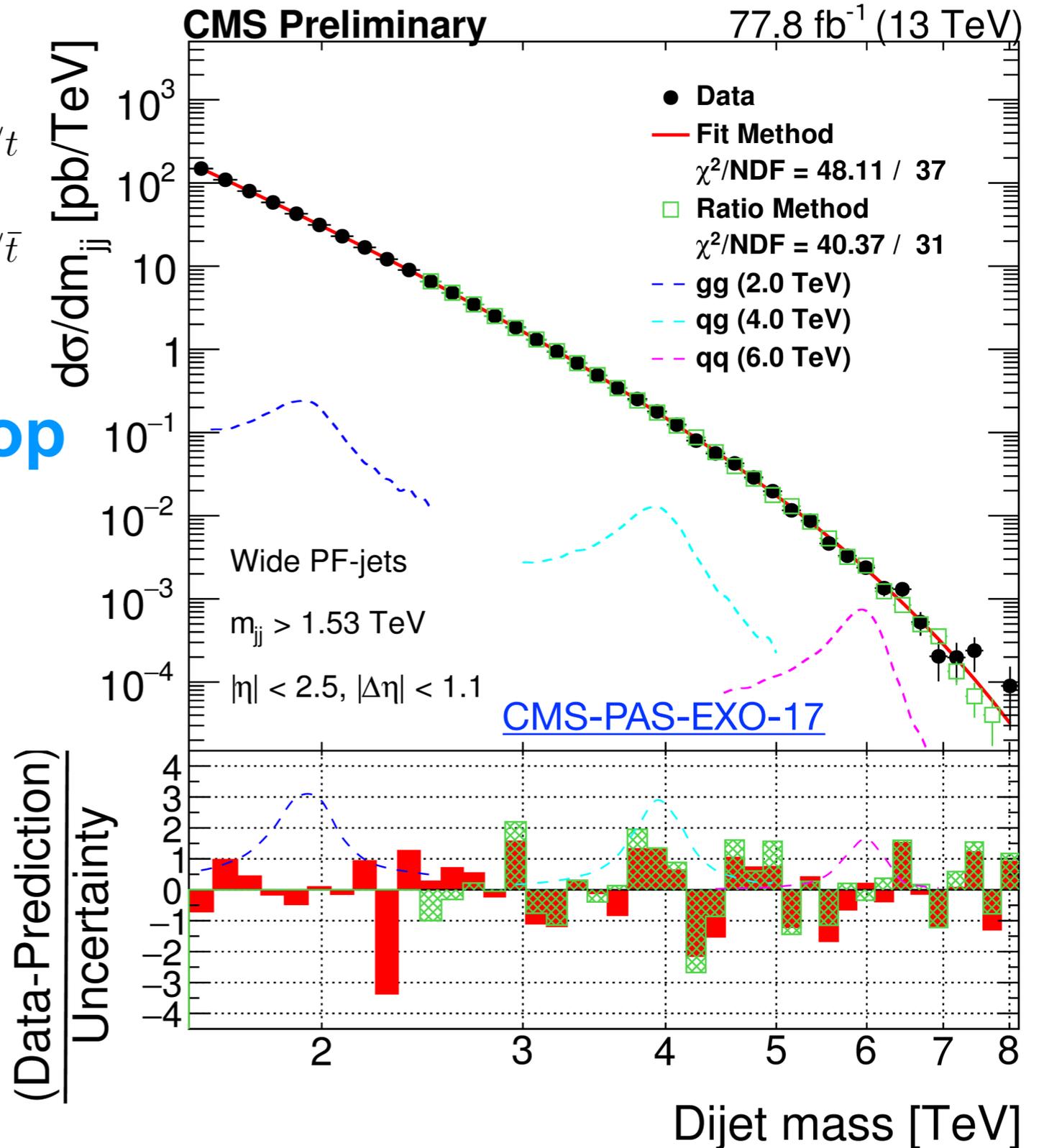
# Techniques 2 - resonances



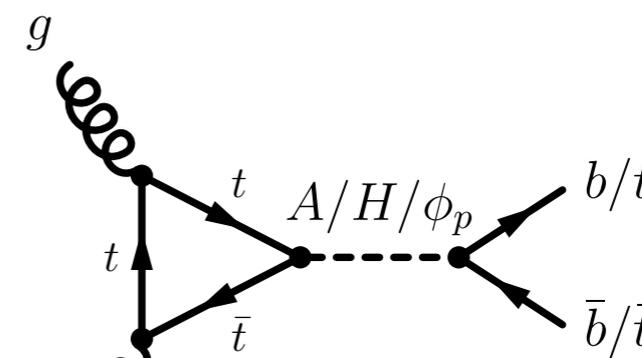
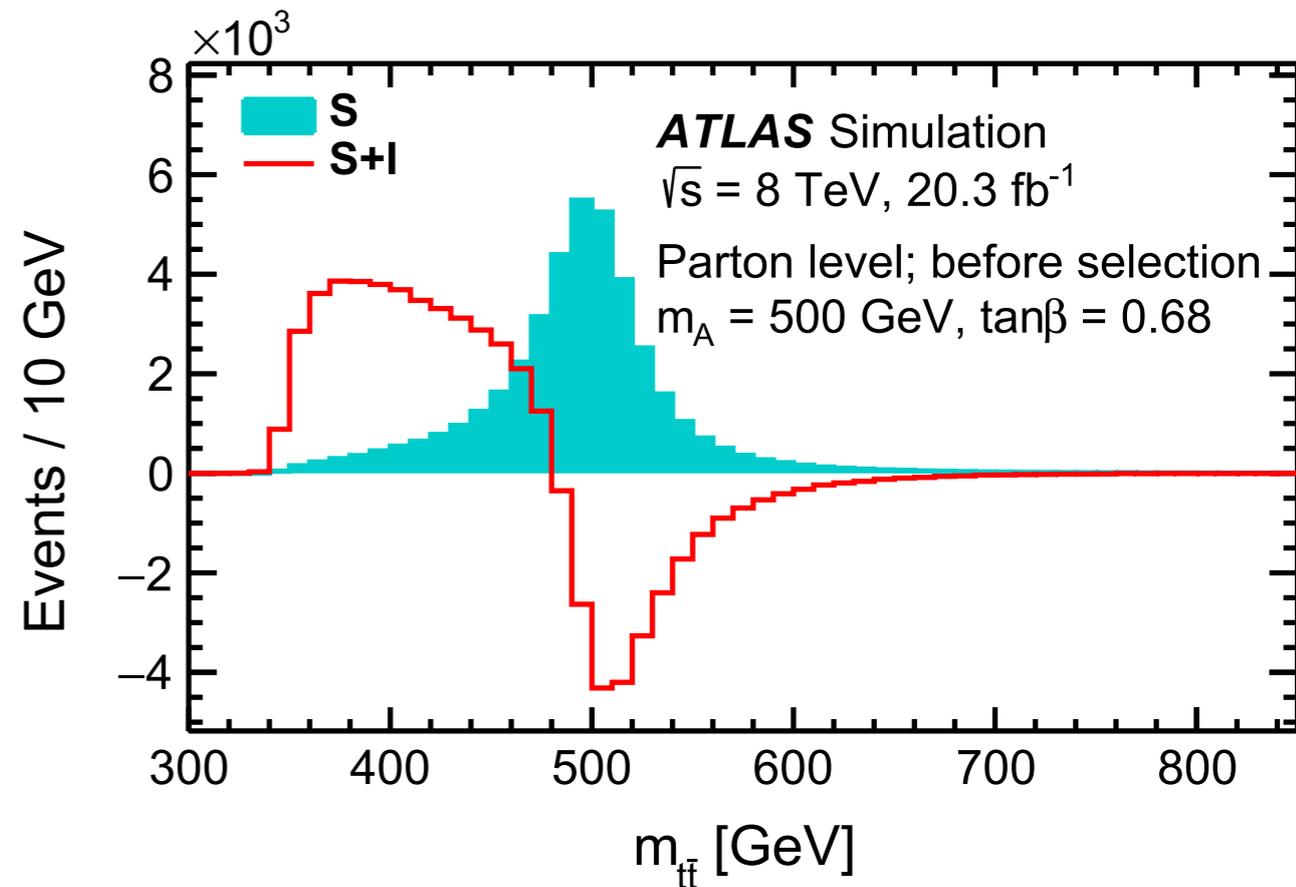
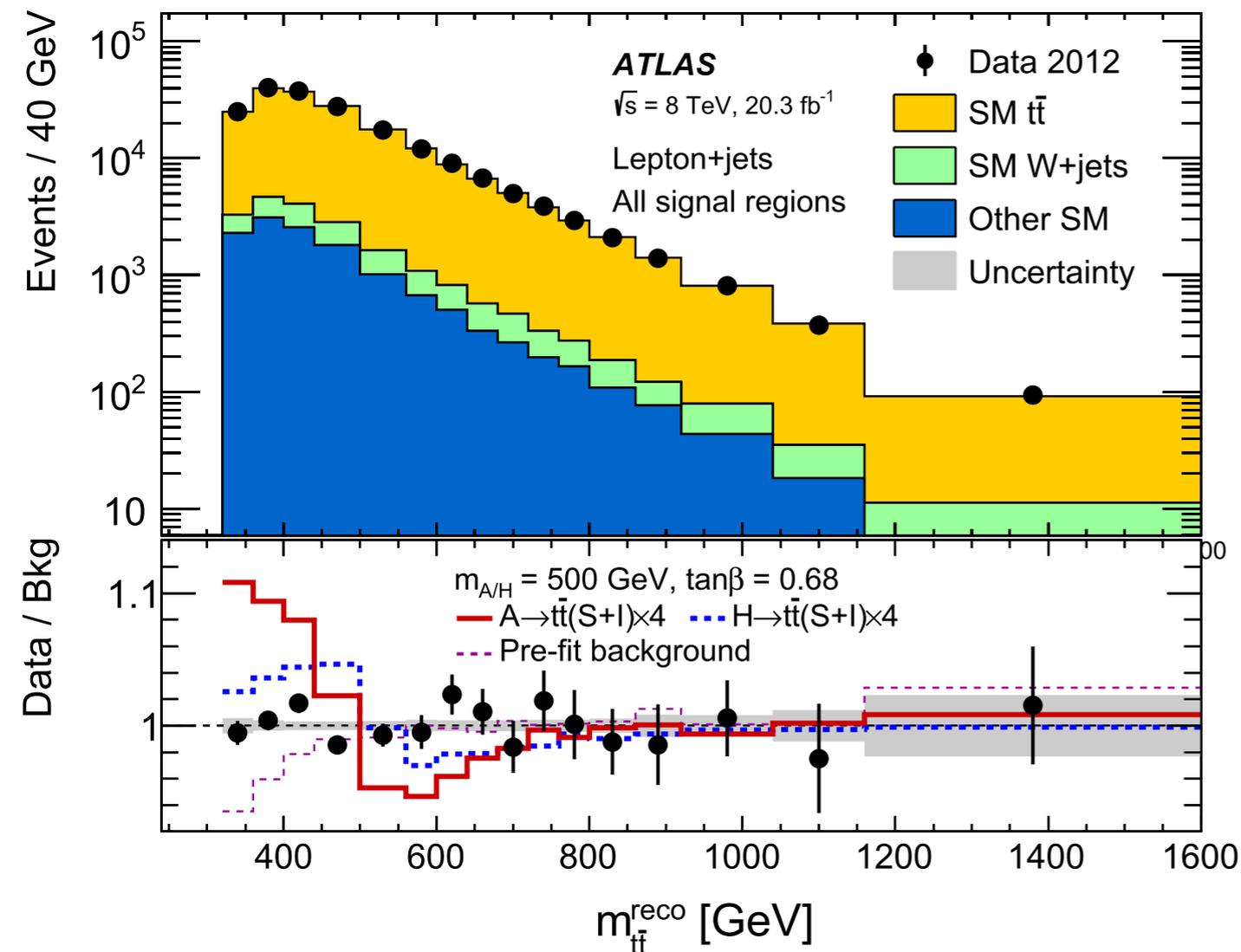
di-jet/di-lepton/di-bjet/di-top



- ★ Low-mass determined by trigger (TLA/di-jet+ISR have the lowest mass sensitivity)
- ★ Sensitive to the mediator width



# A special case ...

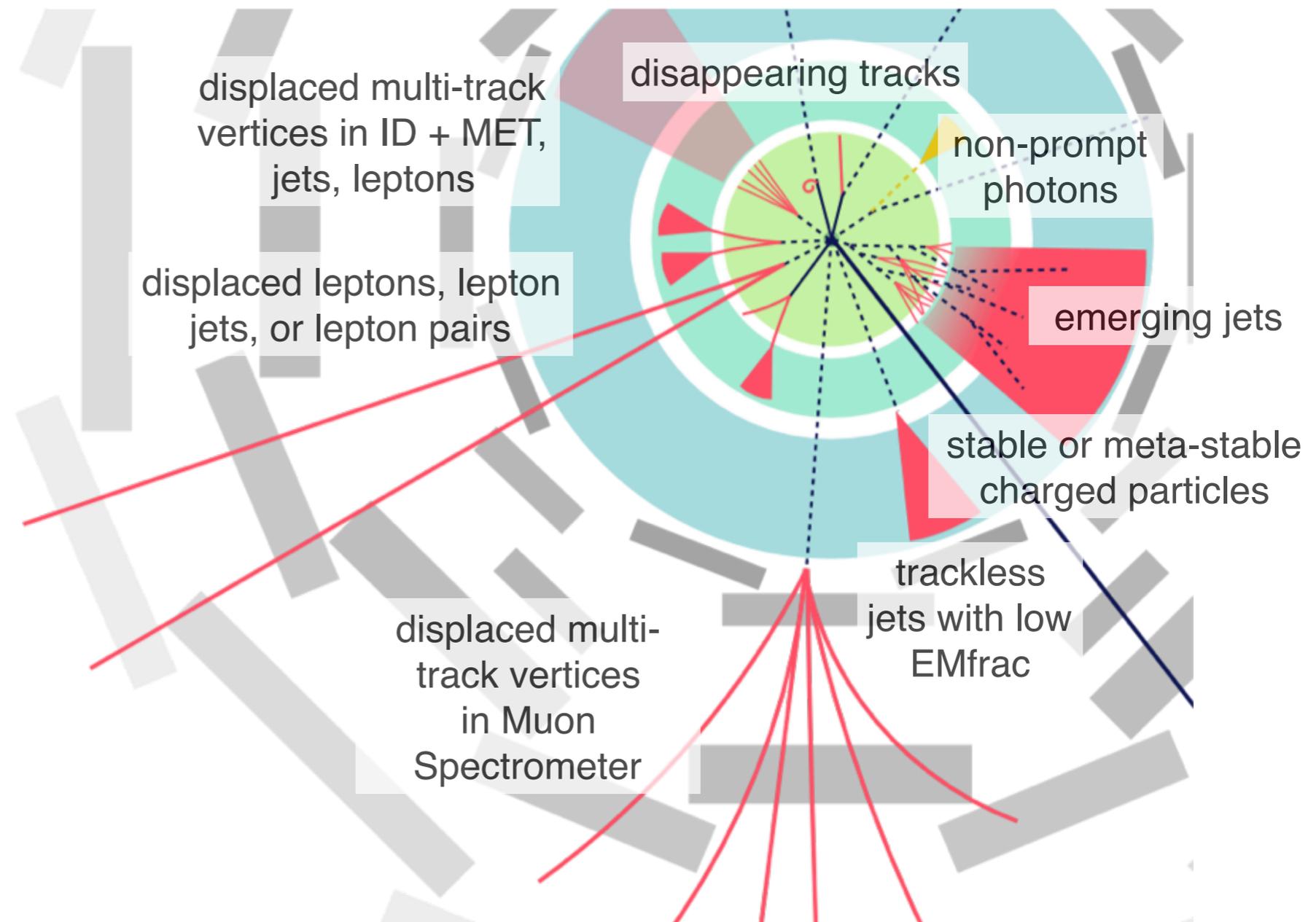


(\*) interference with SM  $t\bar{t}$  to be considered  
 (e.g. [1707.06025](https://arxiv.org/abs/1707.06025))

arXiv:1707.06025

# Techniques 3 - Long Lived Particles

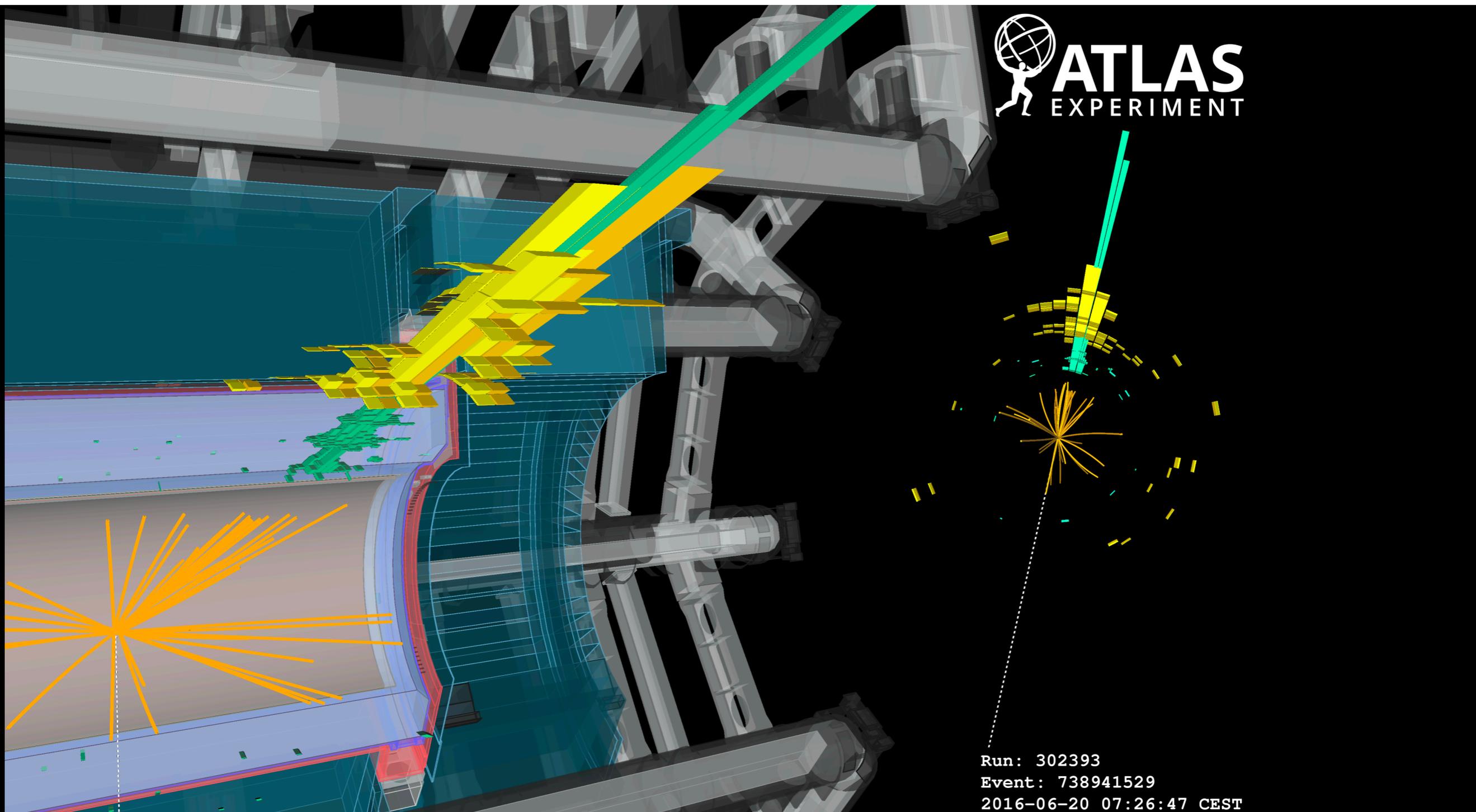
- macroscopic decay length models
- hidden DM
- weak-scale hidden sectors
- SUSY LLPs
- ....



Well established in SUSY, less interpretation in other DM models.

**not covered further**

# 3. Highlights for simplified models

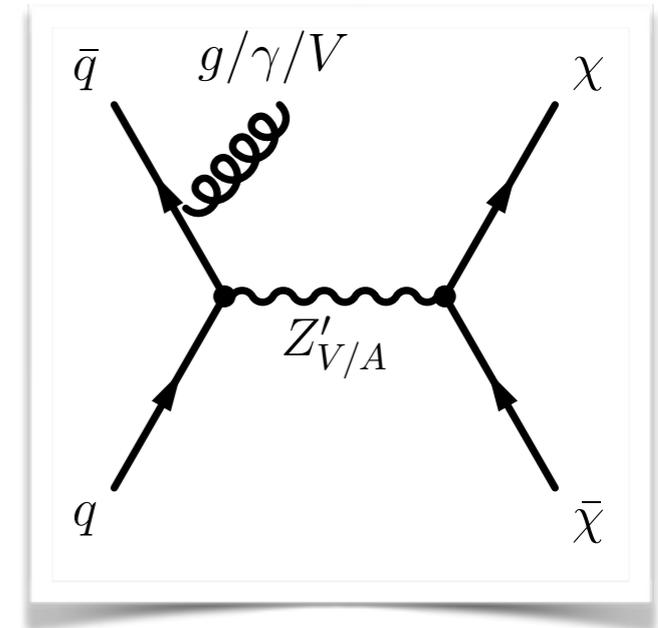
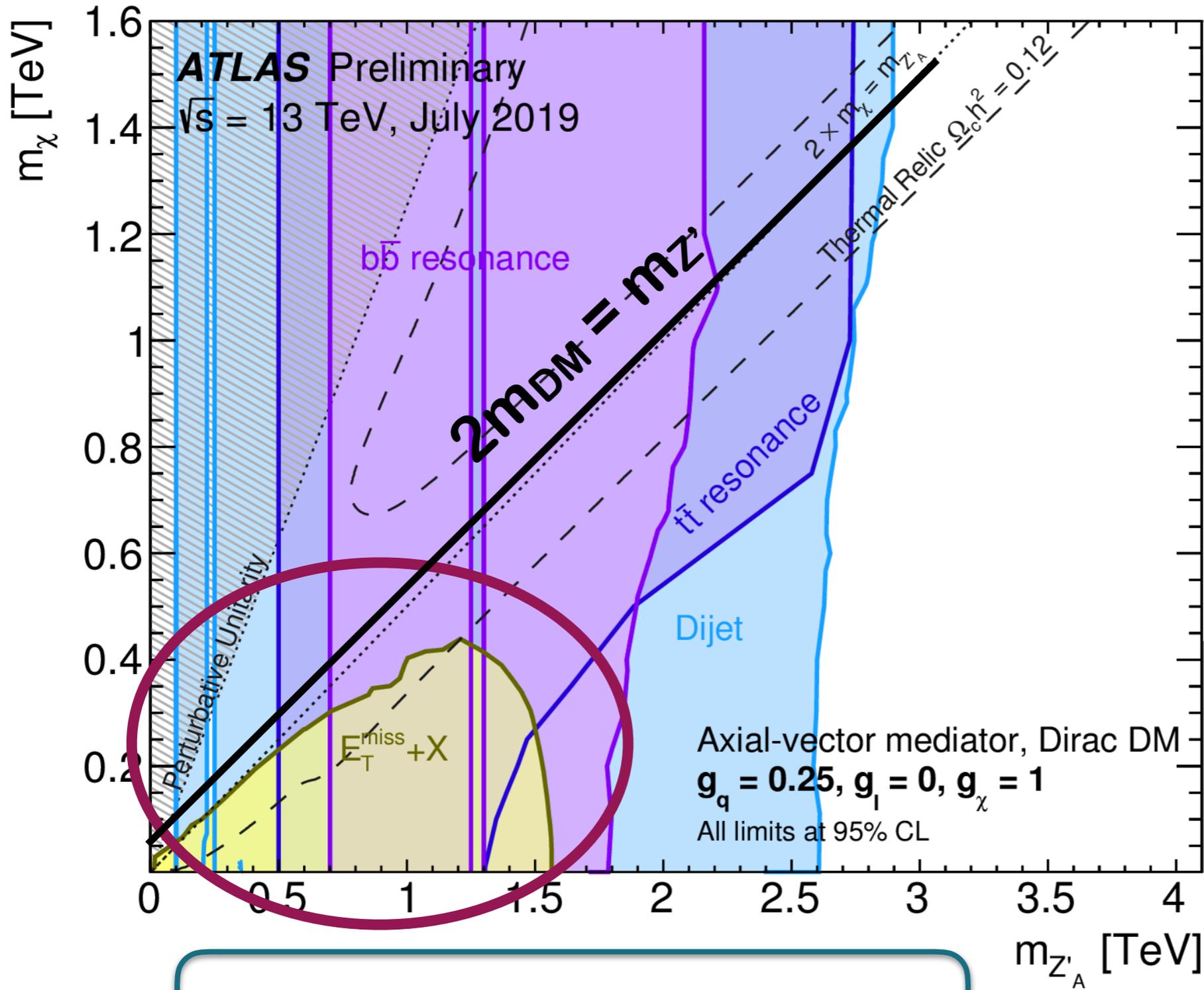


 **ATLAS**  
EXPERIMENT

Run: 302393  
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# Spin-1 mediators - masses

JHEP 05 (2019) 142



**$E_T^{\text{miss}} + X$**

- $E_T^{\text{miss}} + \gamma, 36.1 \text{ fb}^{-1}$   
Eur. Phys. J. C 77 (2017) 393
- $E_T^{\text{miss}} + \text{jet}, 36.1 \text{ fb}^{-1}$   
JHEP 1801 (2018) 126
- $E_T^{\text{miss}} + Z(\text{ll}), 36.1 \text{ fb}^{-1}$   
PLB 776 (2017) 318
- $E_T^{\text{miss}} + V(\text{had}), 36.1 \text{ fb}^{-1}$   
JHEP 10 (2018) 180

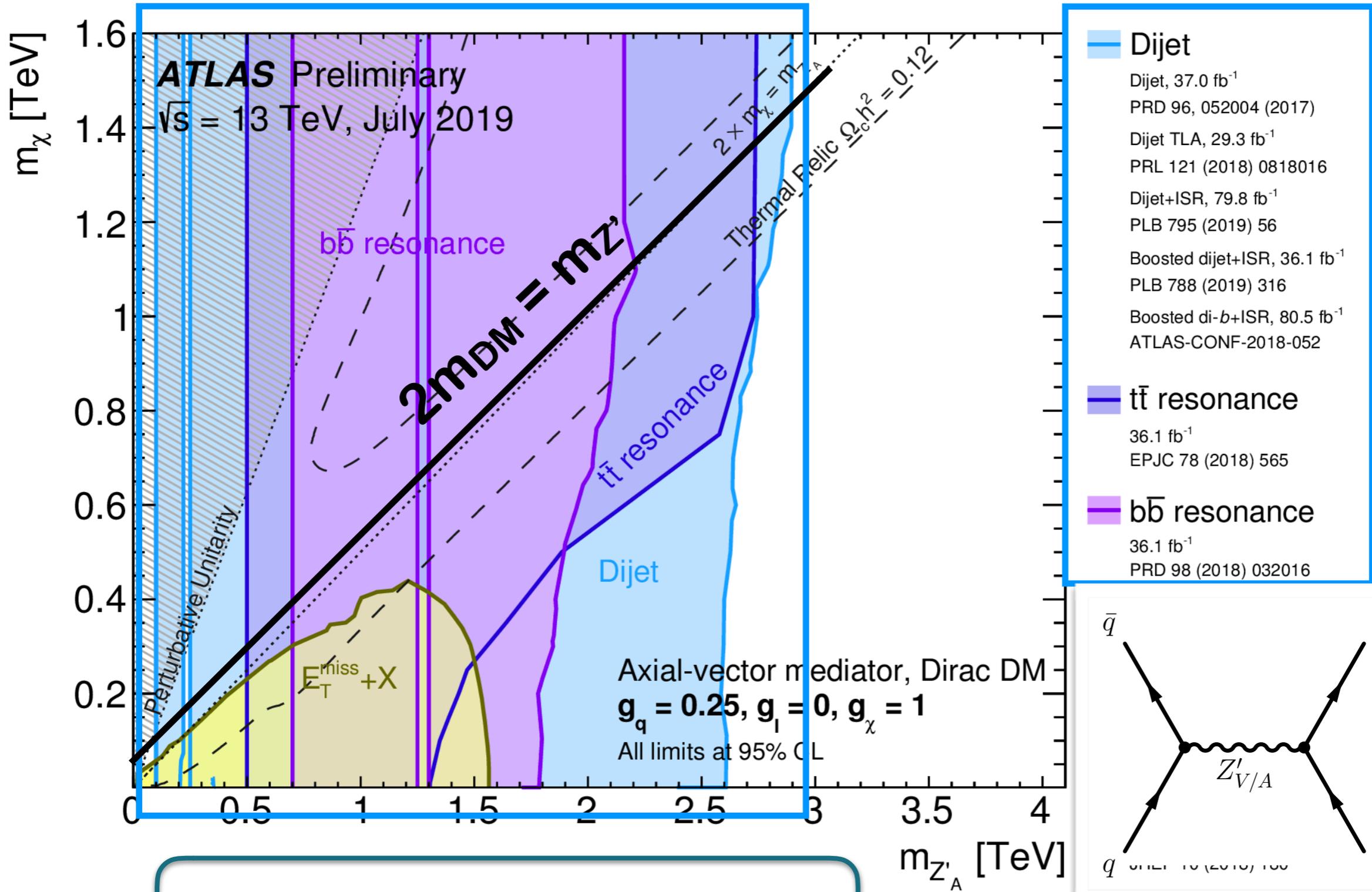
**$E_T^{\text{miss}} + X$**

$g_q = 0.25, g_{\text{lep}} = 0, g_{\text{DM}} = 1$

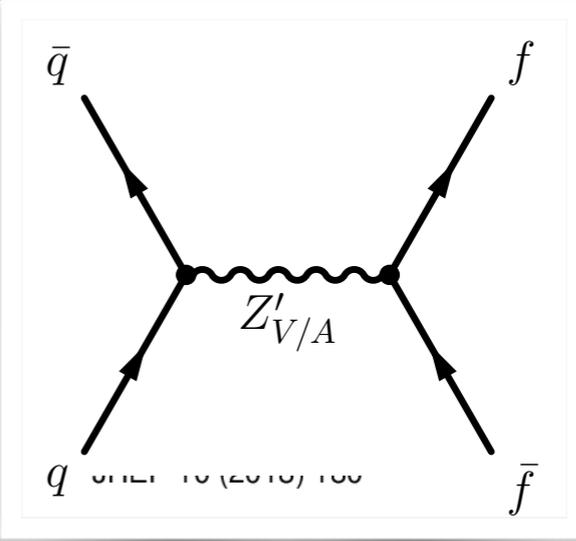
**Axial-vector mediator**

# Spin-1 mediators - masses

Resonance

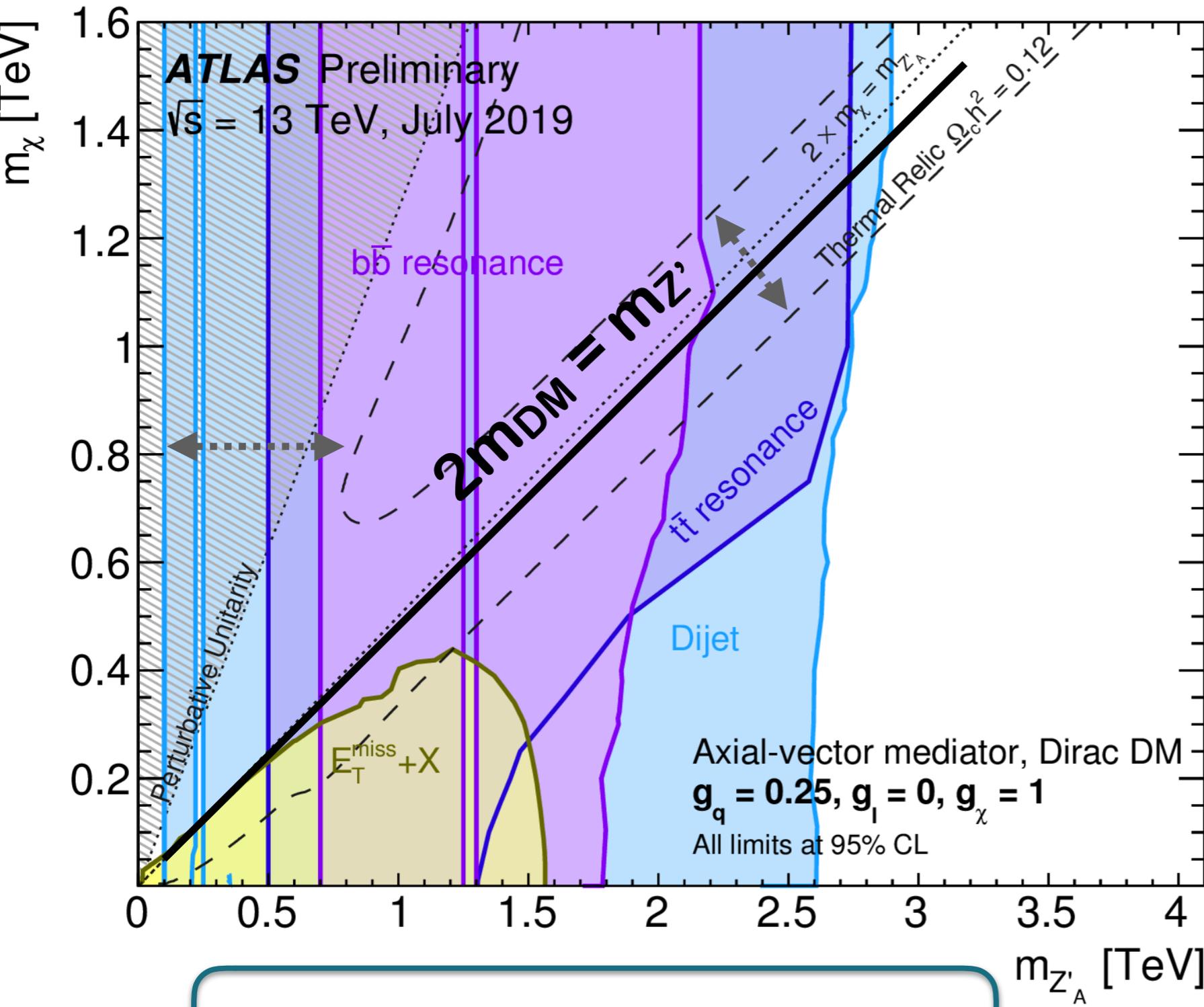


$g_q = 0.25, g_{\text{lep}} = 0, g_{\text{DM}} = 1$



**Axial-vector mediator**

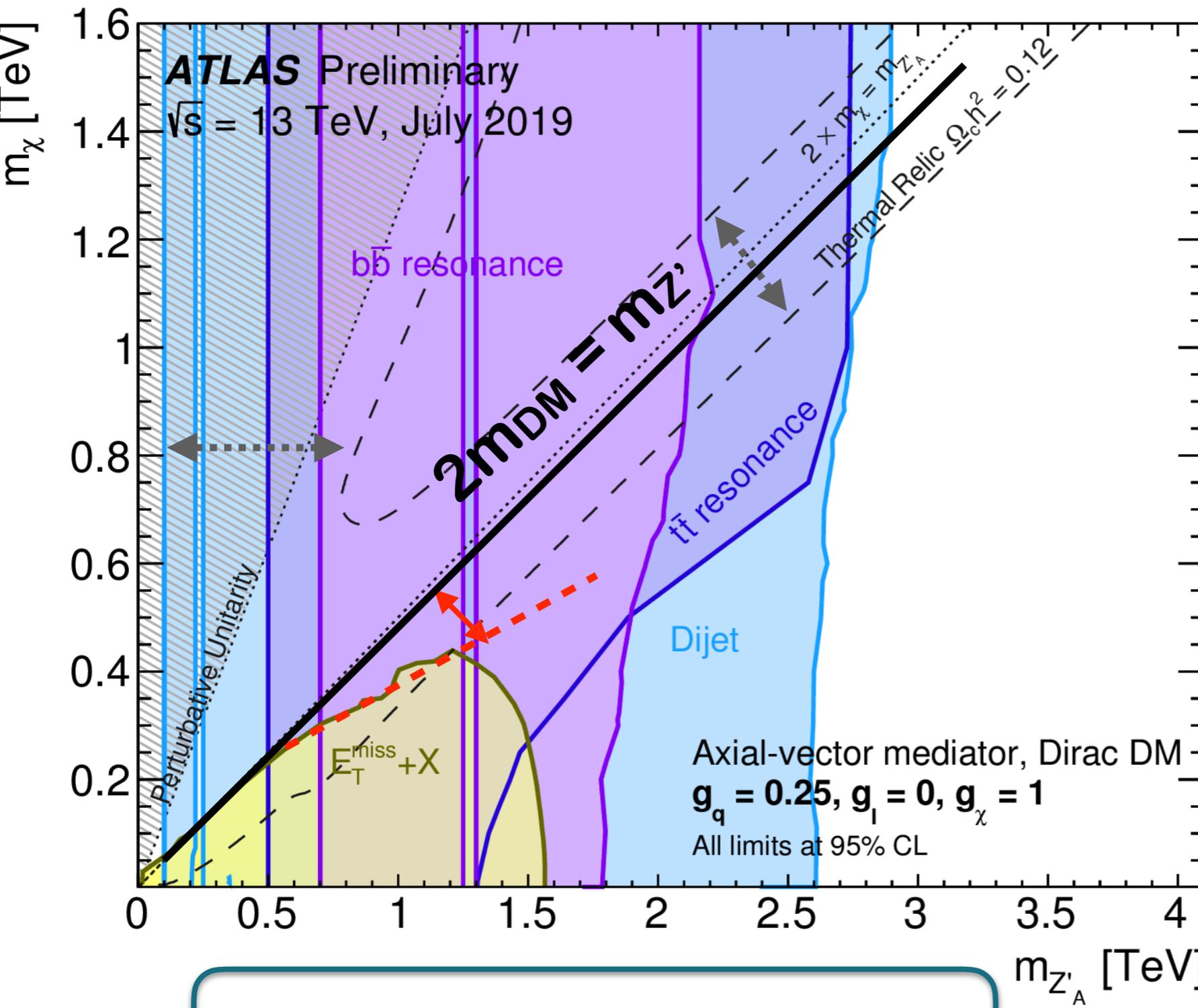
# Spin-1: features explained



$g_q = 0.25, g_{lep} = 0, g_{DM} = 1$

Model's predicted relic density depleted with respect to 0.12 thermal relic

# Spin-1: features explained



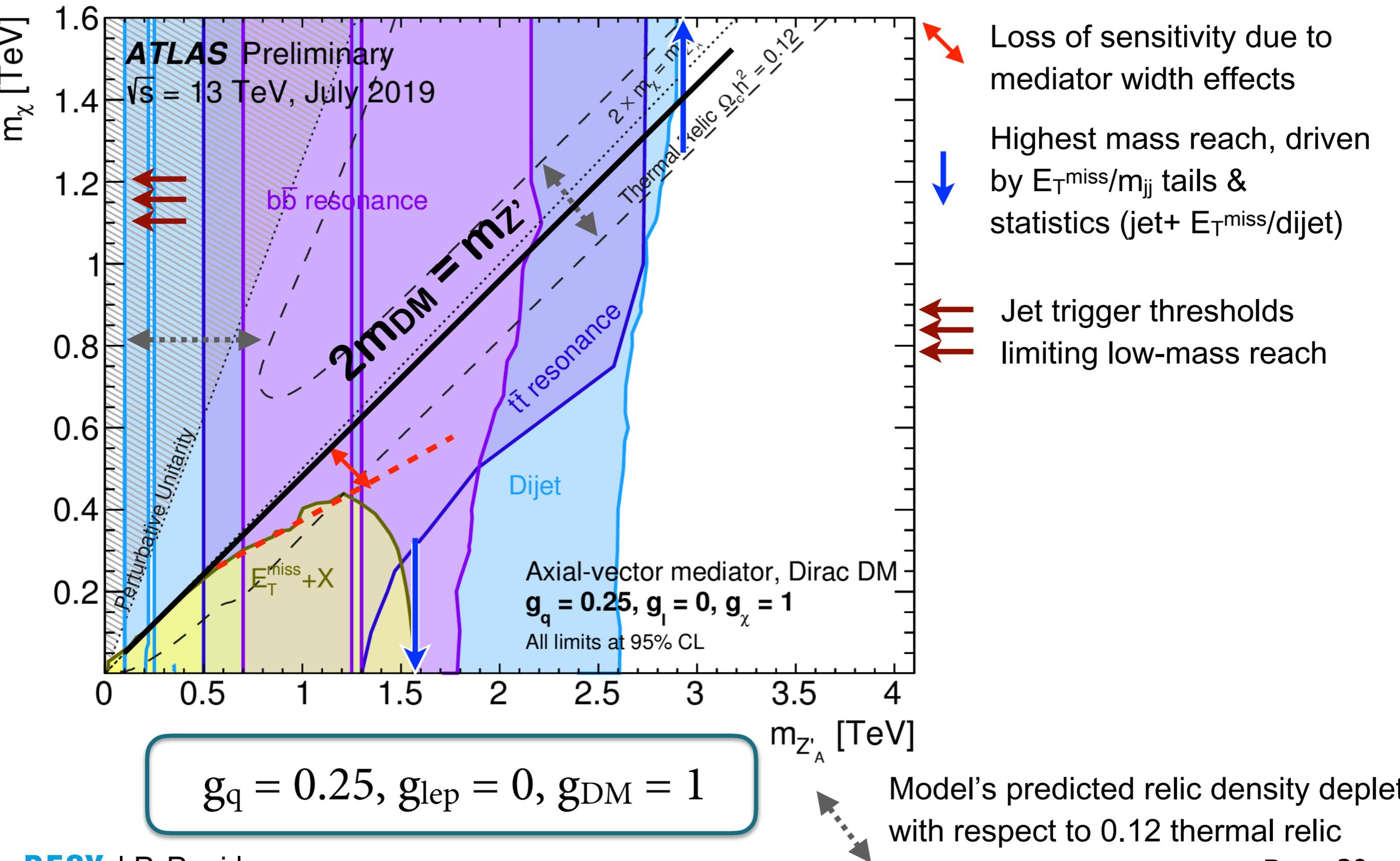
Loss of sensitivity due to mediator width effects

$$g_q = 0.25, g_{\text{lep}} = 0, g_{\text{DM}} = 1$$

Model's predicted relic density depletion with respect to 0.12 thermal relic

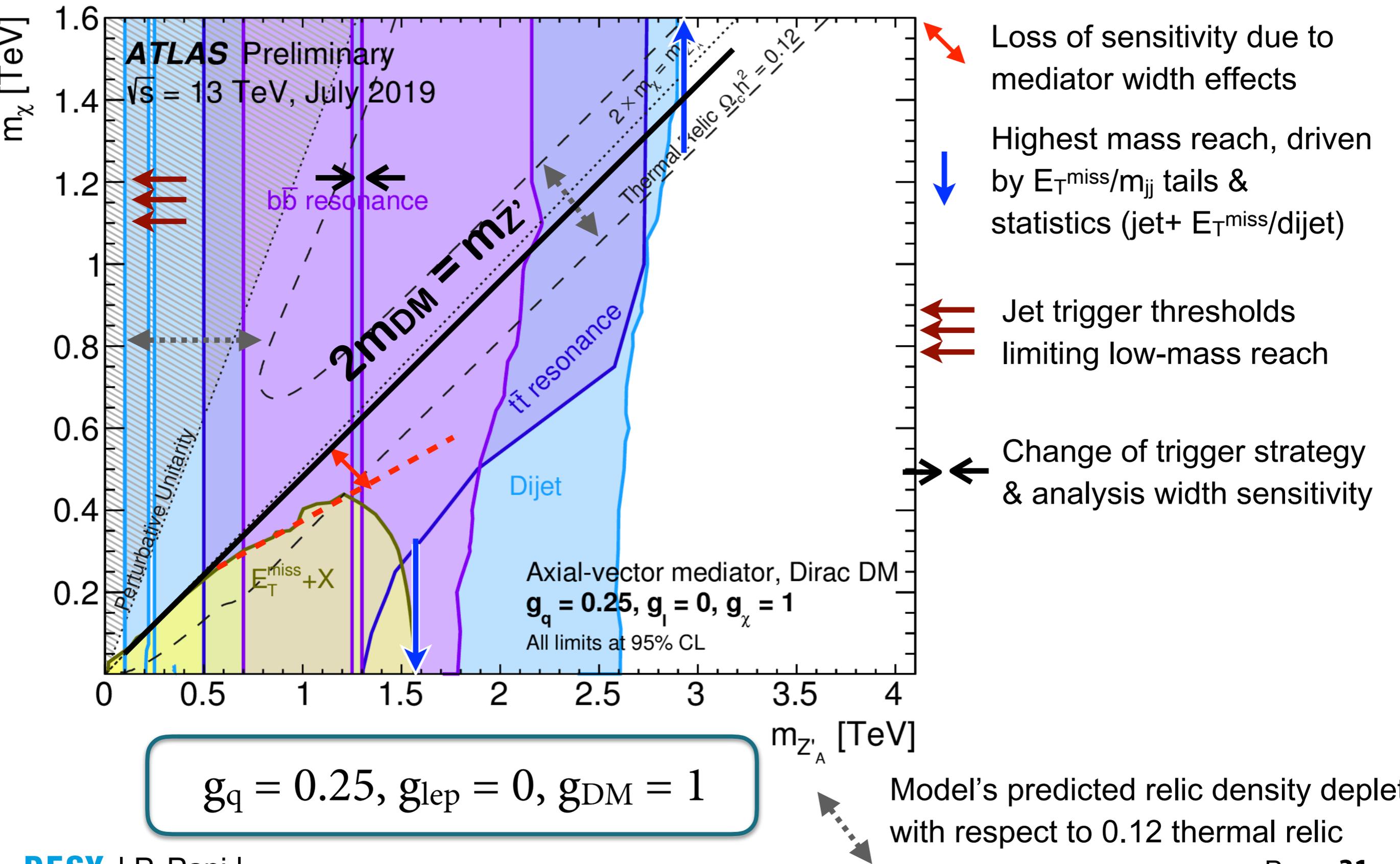


# Spin-1: features explained



$g_q = 0.25, g_{\text{lep}} = 0, g_{\text{DM}} = 1$

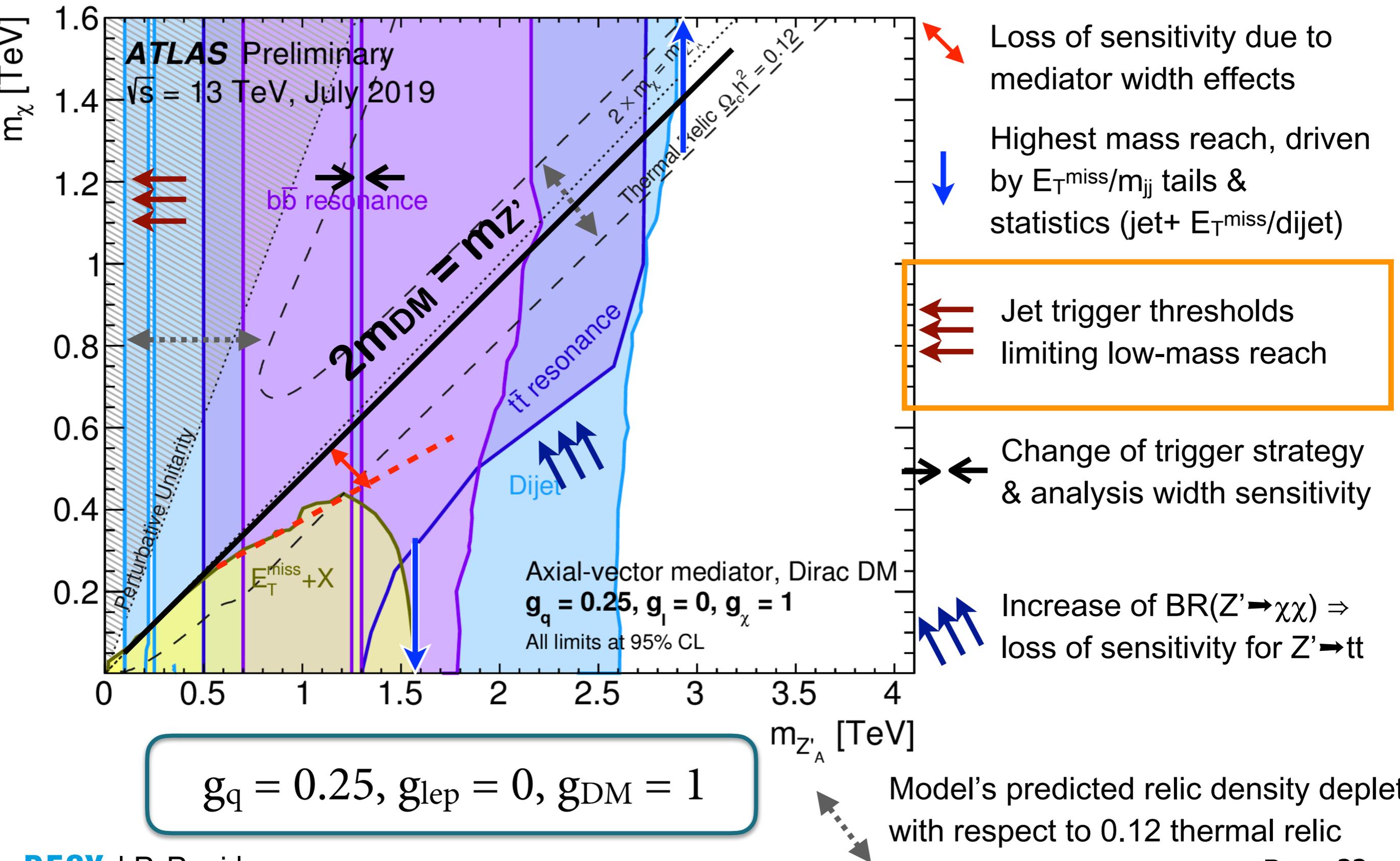
# Spin-1: features explained



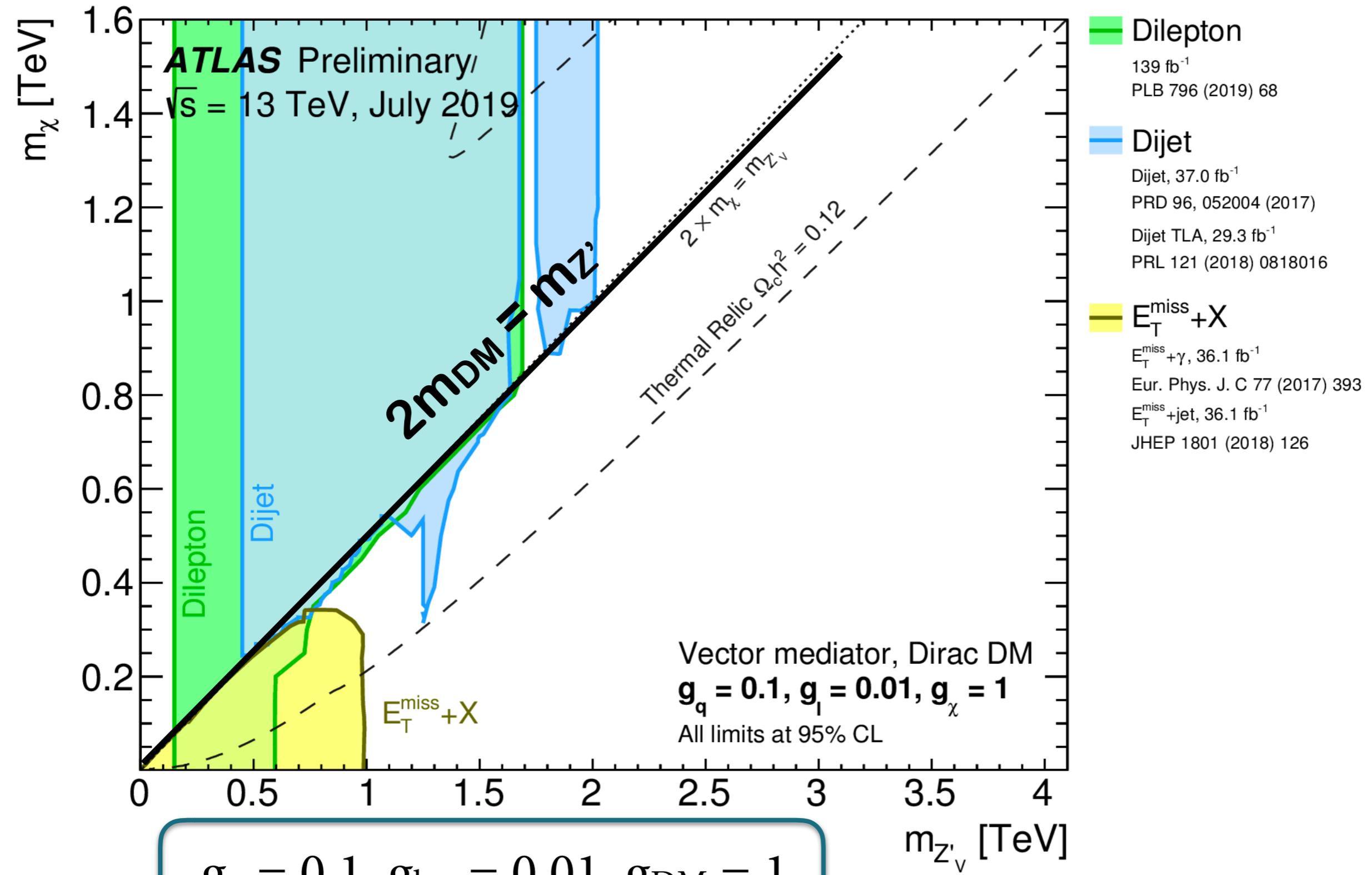
$g_q = 0.25, g_{\text{lep}} = 0, g_{\text{DM}} = 1$



# Spin-1: features explained



# Spin-1 leptophilic case



E<sub>T</sub><sup>miss</sup> + X Resonance

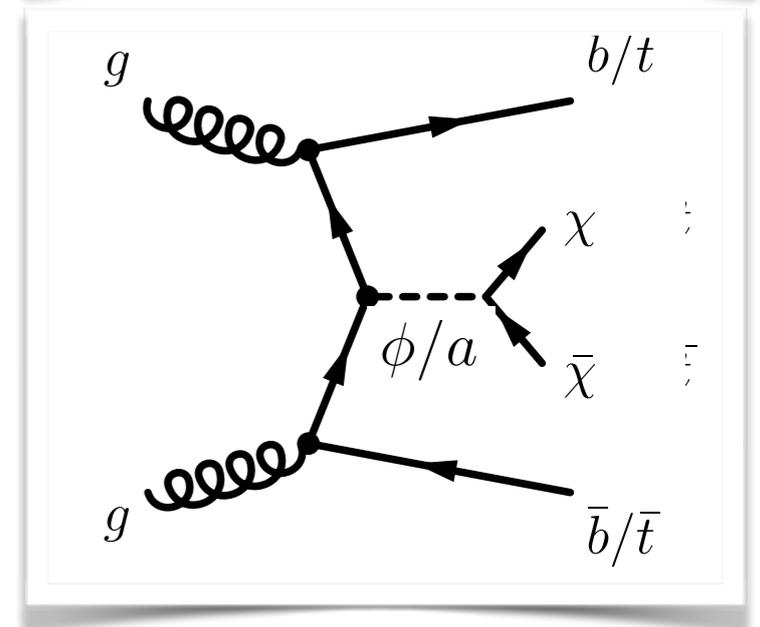
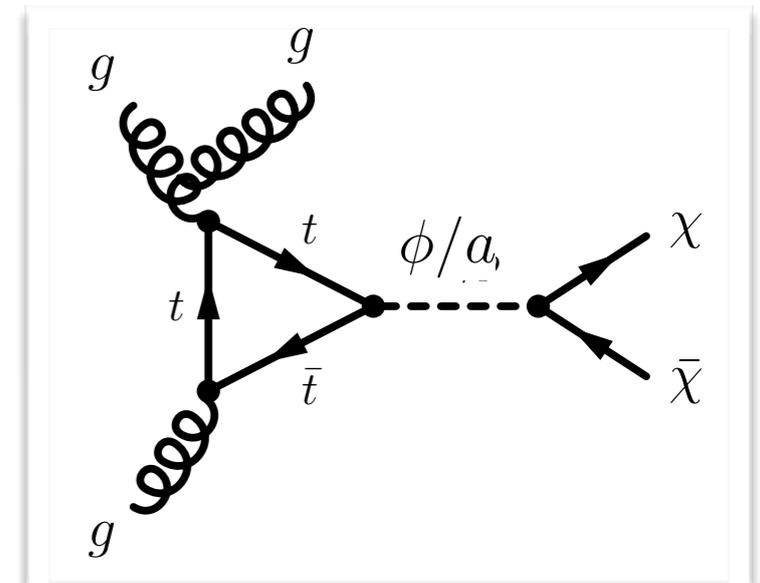
$g_q = 0.1, g_{lep} = 0.01, g_{DM} = 1$

**Vector mediator**

# Spin-0 mediators

$$\mathcal{L} \sim \sum_f i g_v \frac{y_f}{\sqrt{2}} A \bar{f} \gamma^5 f$$

Needed to easily fulfil Flavour Constraints (MFV)



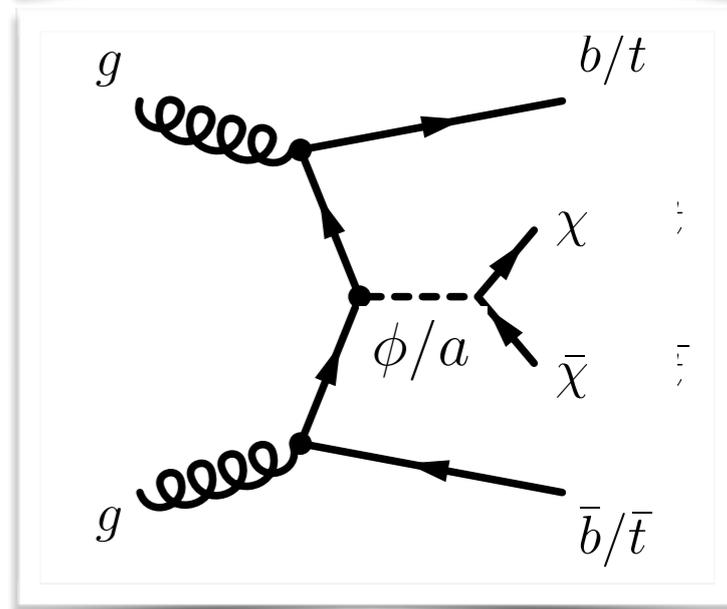
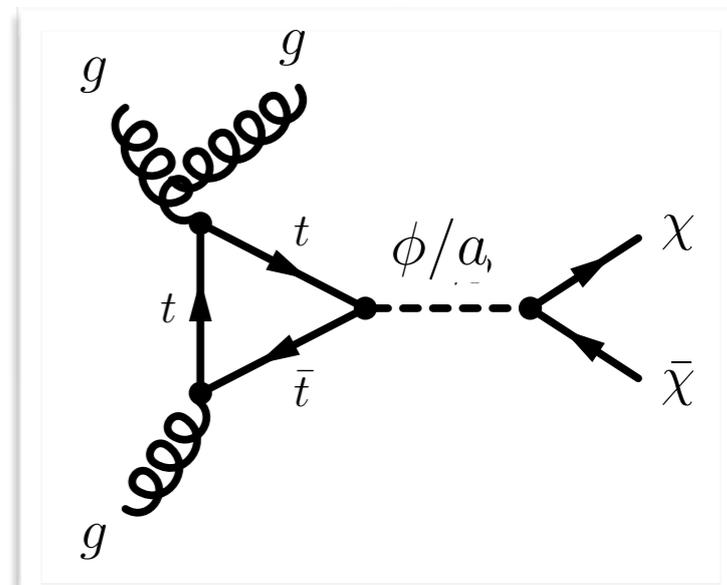
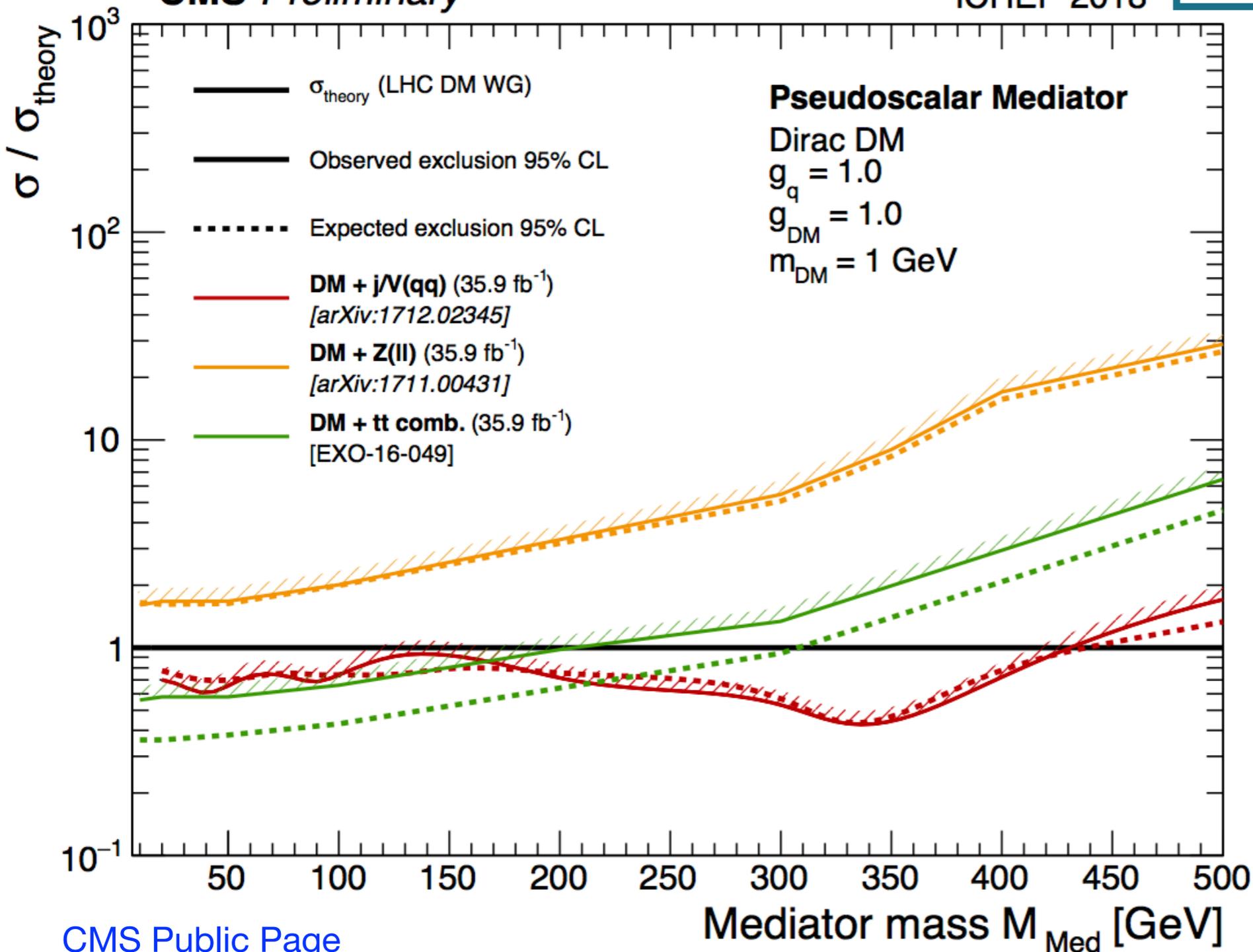
# Spin-0 mediators

$$\mathcal{L} \sim \sum_f i g_v \frac{y_f}{\sqrt{2}} A \bar{f} \gamma^5 f$$

Needed to easily fulfil Flavour Constraints (MFV)

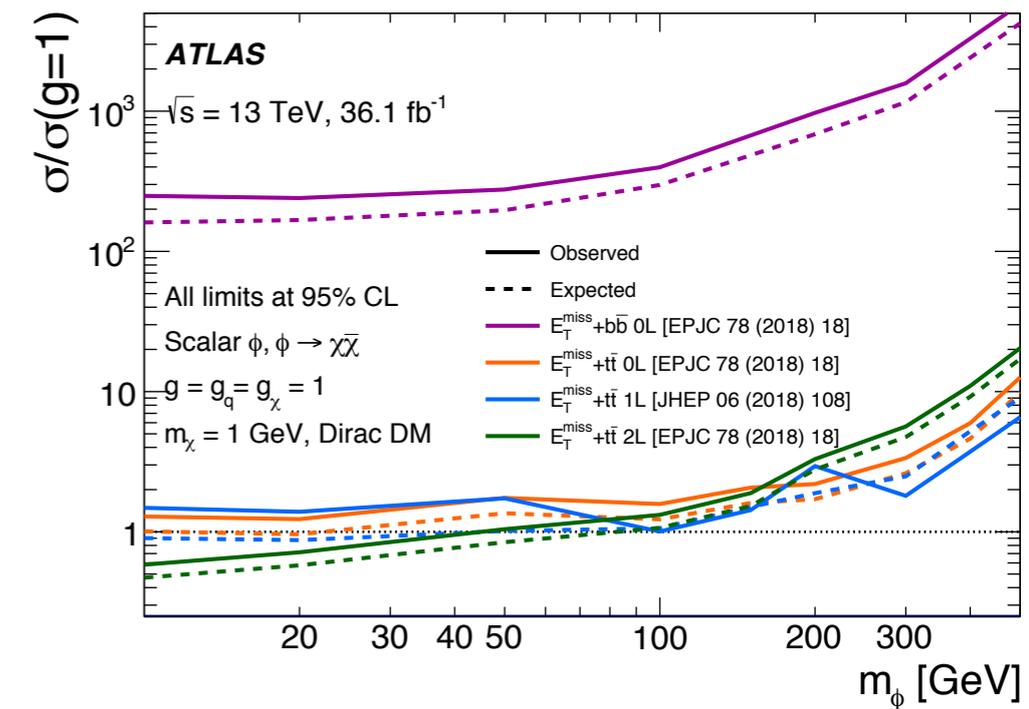
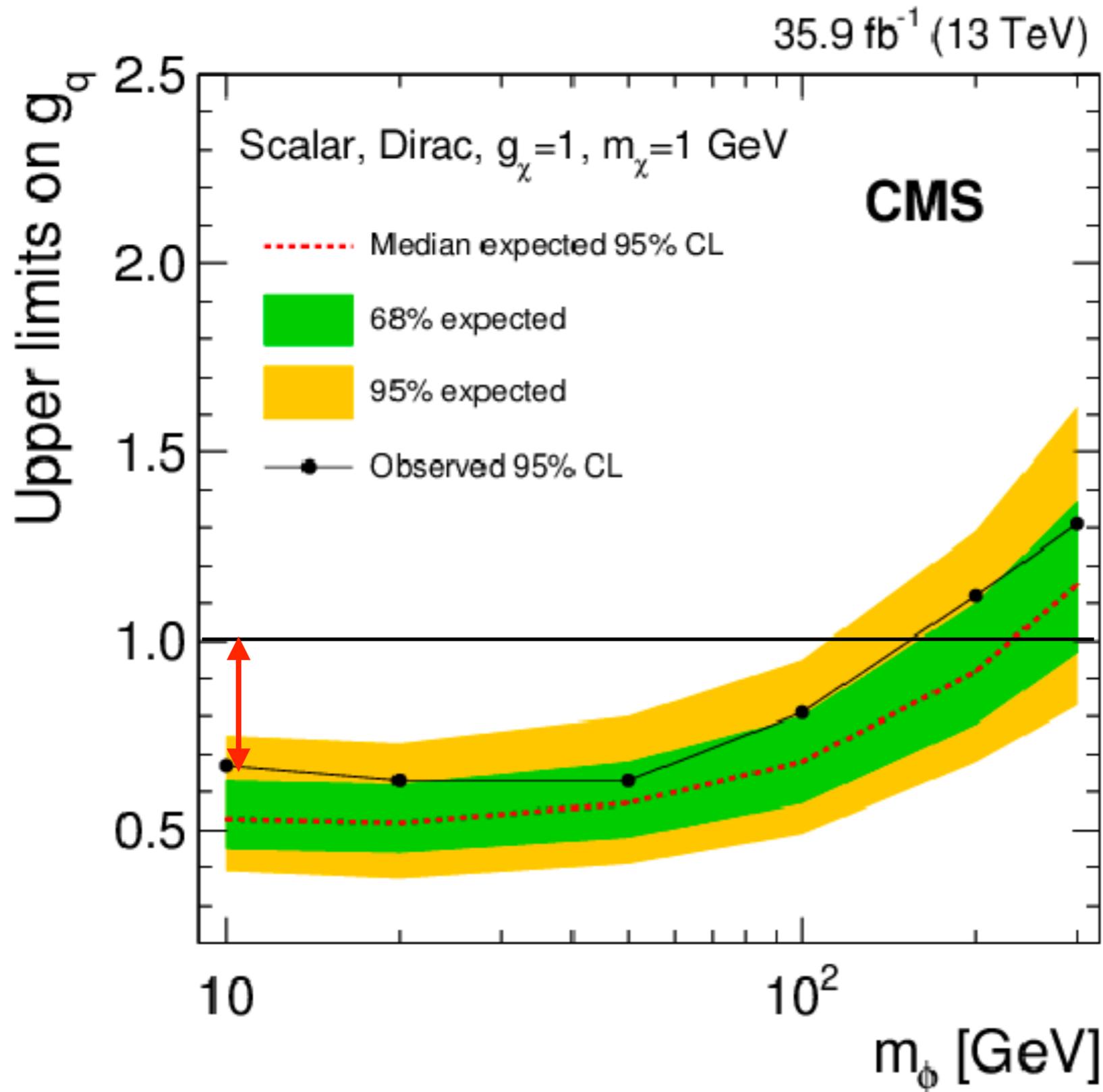
CMS Preliminary

ICHEP 2018

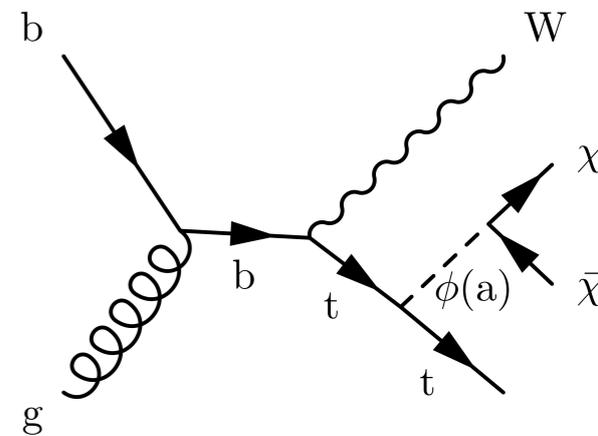
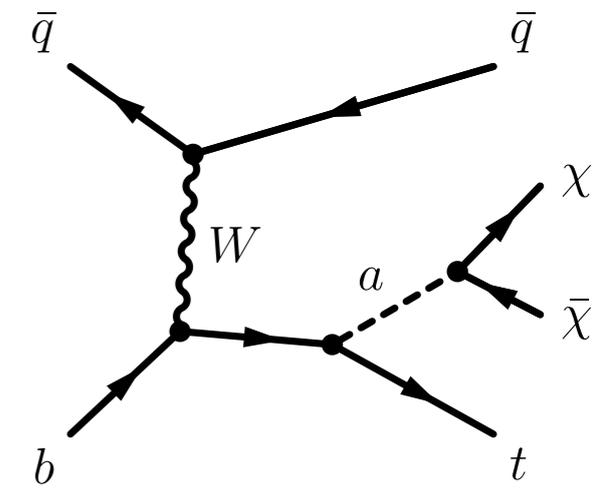
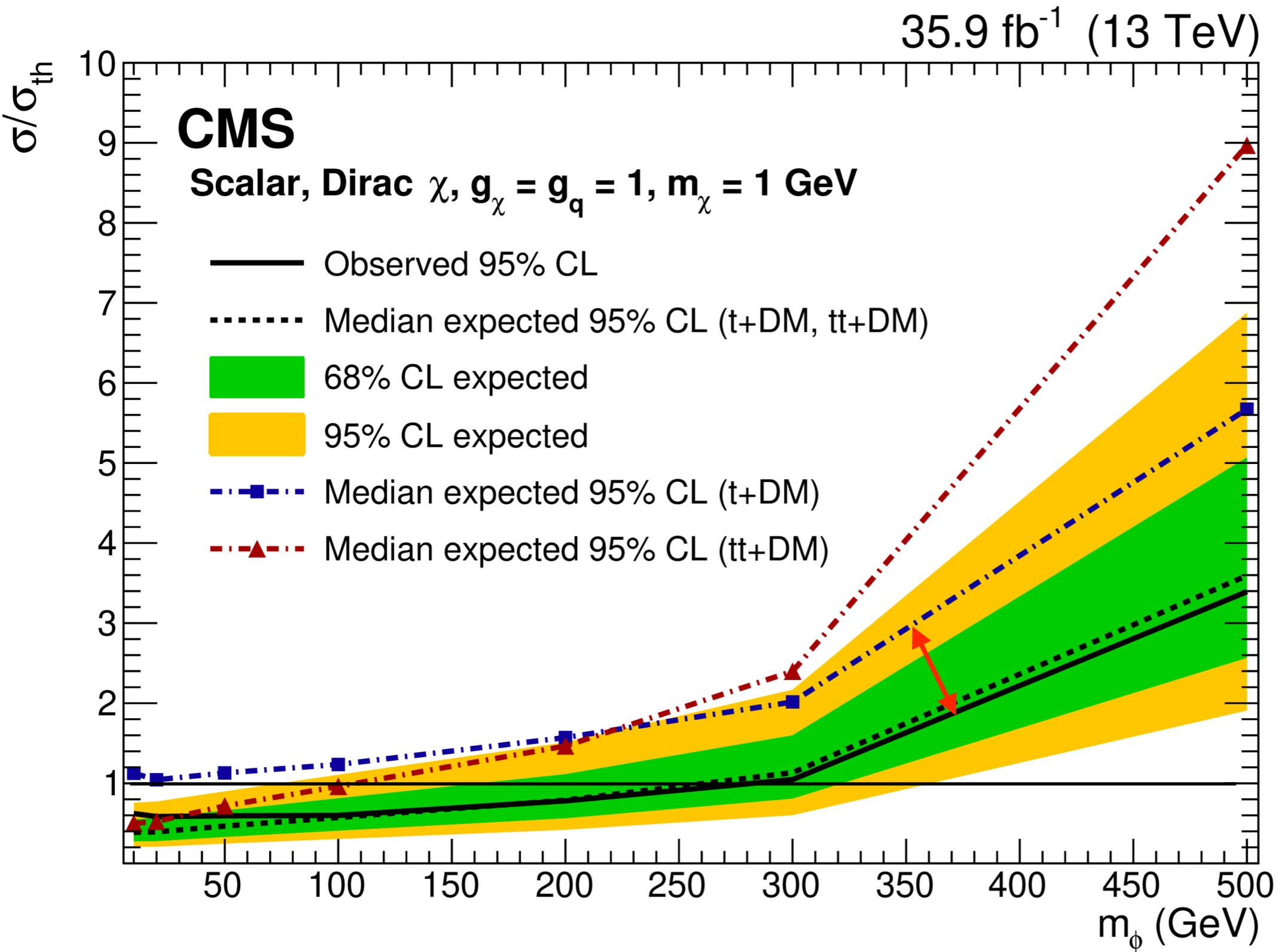


# CMS grand combination

[arXiv:1807.06522](https://arxiv.org/abs/1807.06522)



# Spin-0 with single top



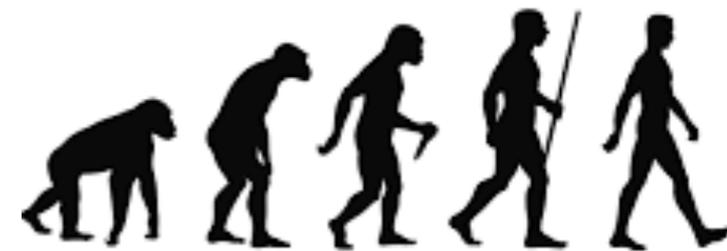
# Considerations on the results

- ★ Simplified models are good phenomenology proxies.
- ★ Simplified models are simplified models.
- ★ Simplified models are not full and complete theories, which might have more complex topologies.
- ★ All exclusions need to be taken with a grain of salt.

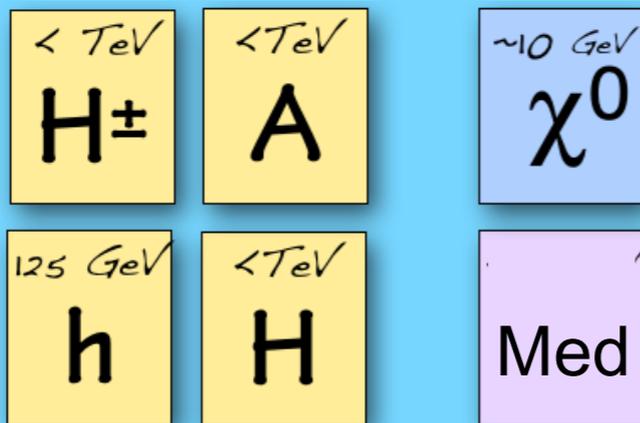




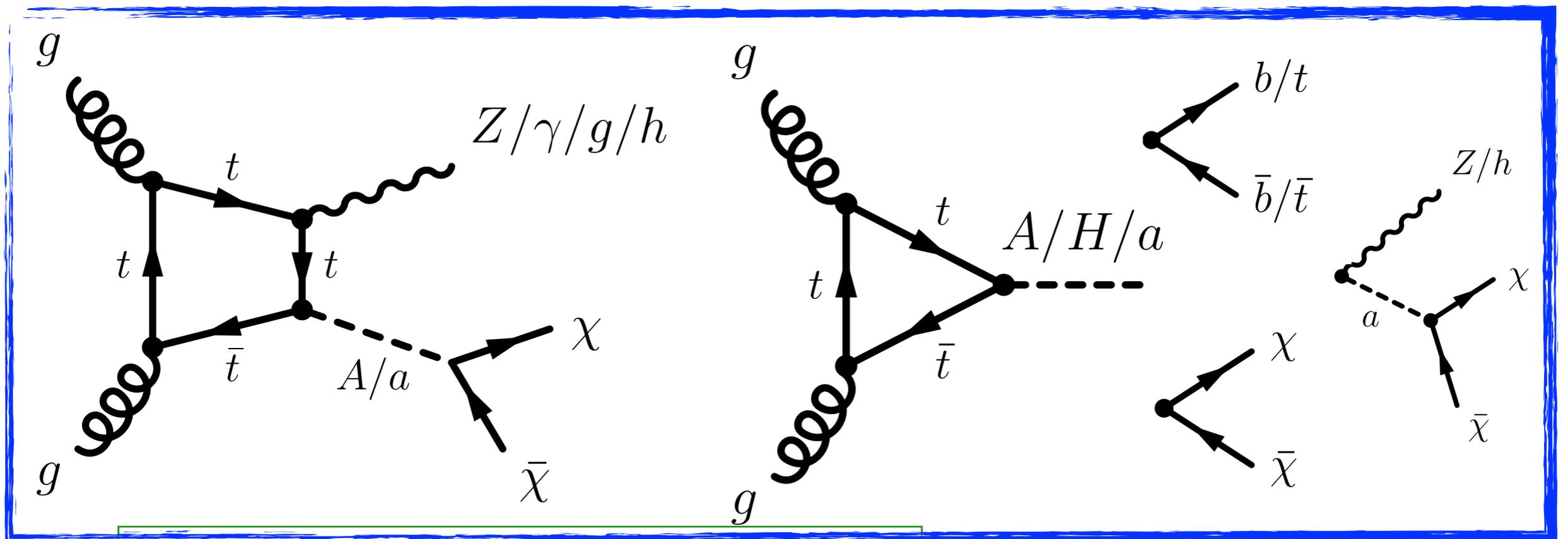
# 2HDM-based models



## 2HDM DM models



★ **Richer phenomenology:**  
Higgs bosons productions and decays, mixing, many final states.

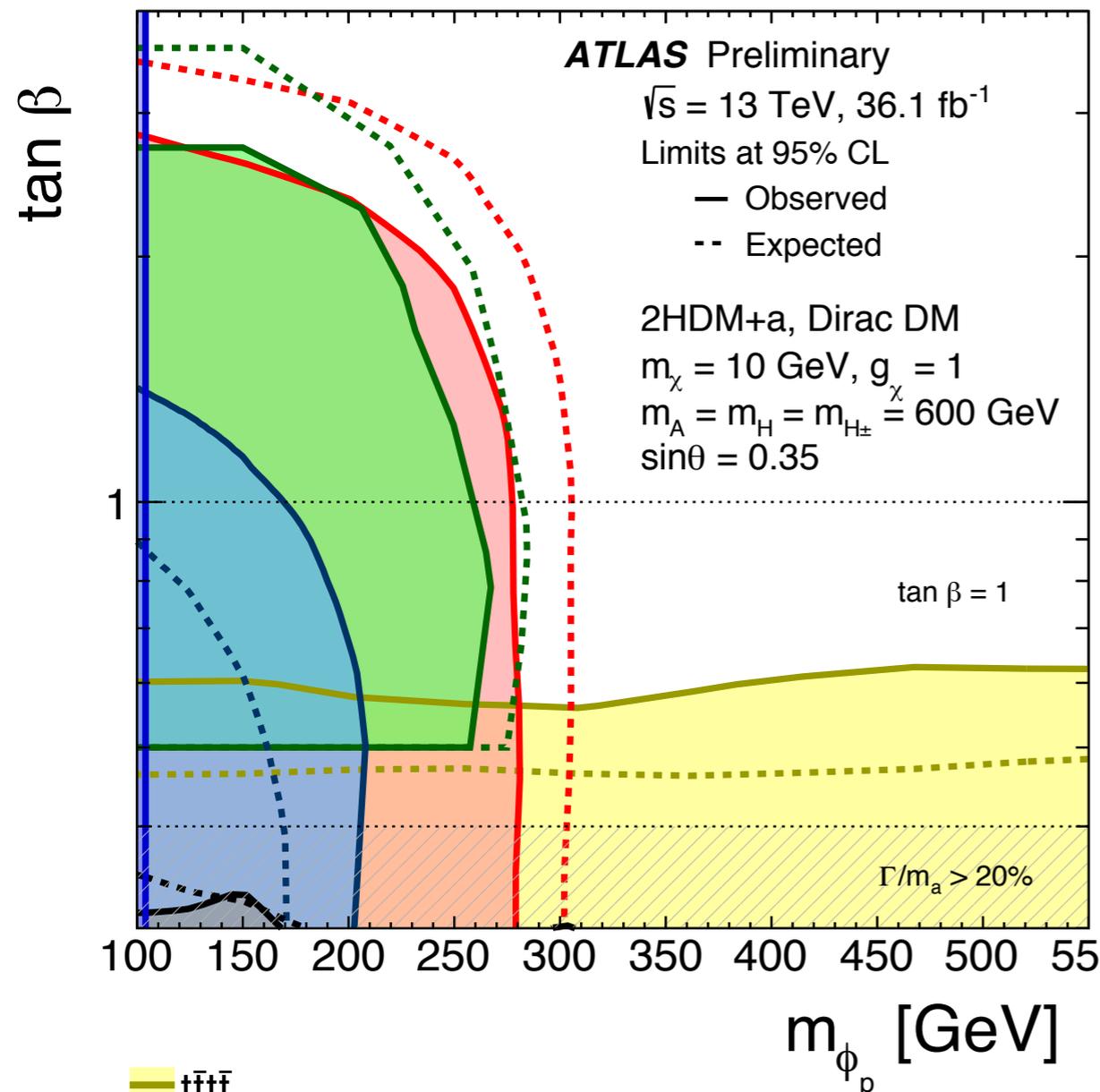


[arxiv:1810.09420](https://arxiv.org/abs/1810.09420) (and ref. therein) + [LPCC WG](#)

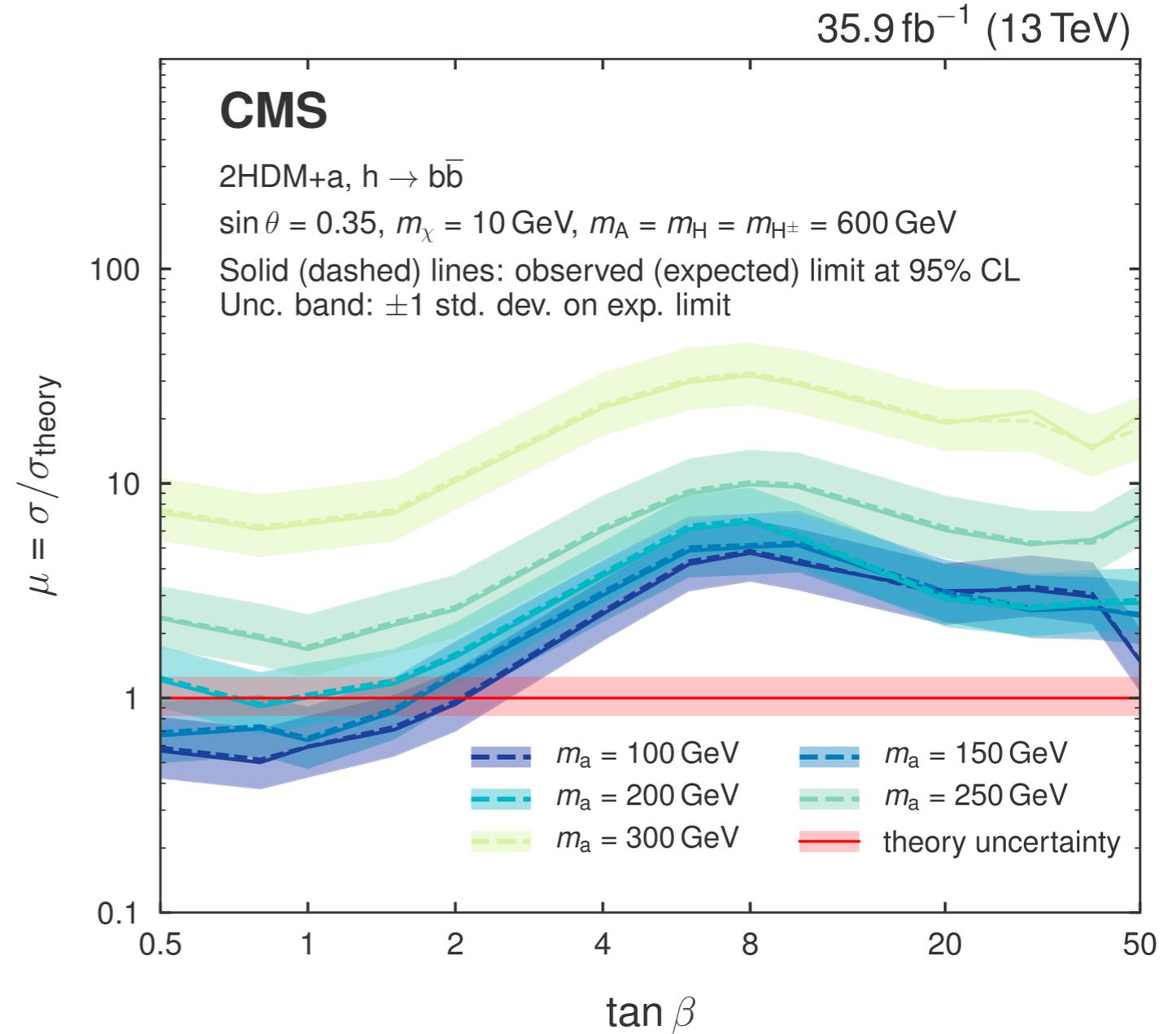
# 2HDM+pseudoscalar models

JHEP 05 (2019) 142

Eur. Phys. J. C 79 (2019) 280



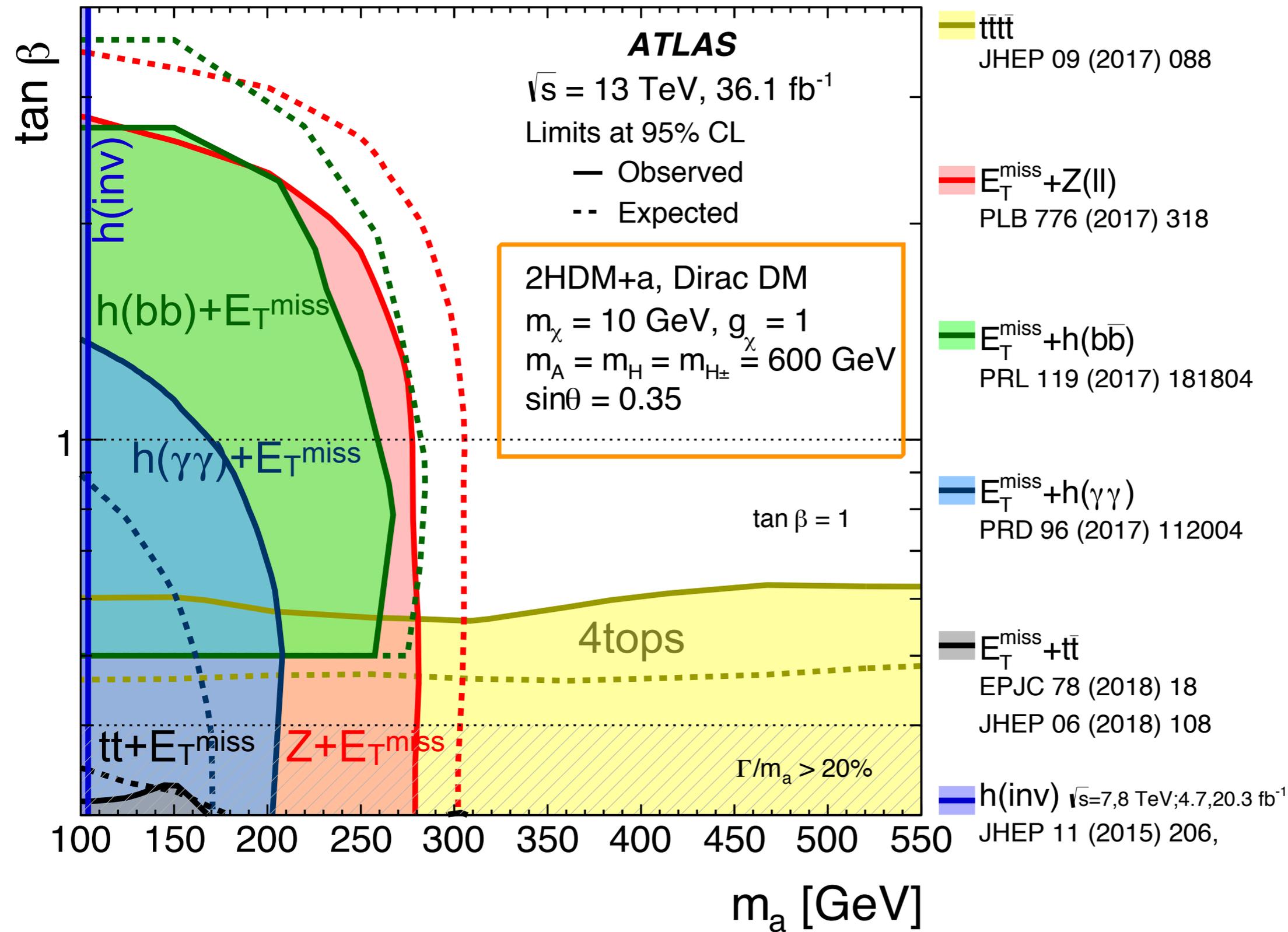
- $t\bar{t}\bar{t}$   
JHEP 09 (2017) 088
- $E_T^{\text{miss}} + h(\gamma\gamma)$   
PRD 96 (2017) 112004
- $E_T^{\text{miss}} + Z(\ell\ell)$   
PLB 776 (2017) 318
- $E_T^{\text{miss}} + t\bar{t}$   
EPJC 78 (2018) 18  
JHEP 06 (2018) 108
- $E_T^{\text{miss}} + h(b\bar{b})$   
PRL 119 (2017) 181804
- $h(\text{inv})$   
JHEP 06 (2018) 108



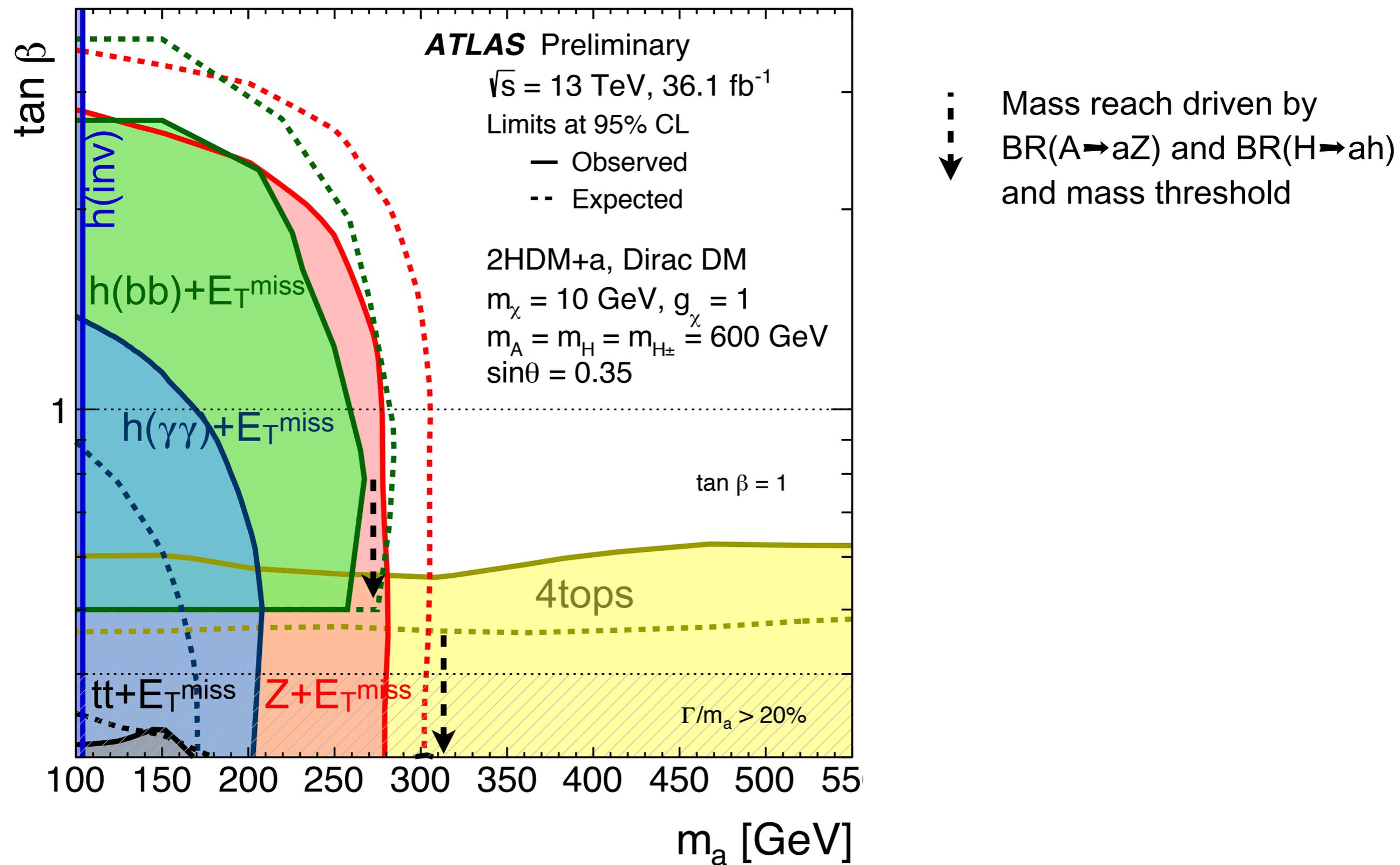
- $m_a = 100 \text{ GeV}$
- $m_a = 150 \text{ GeV}$
- $m_a = 200 \text{ GeV}$
- $m_a = 250 \text{ GeV}$
- $m_a = 300 \text{ GeV}$
- theory uncertainty



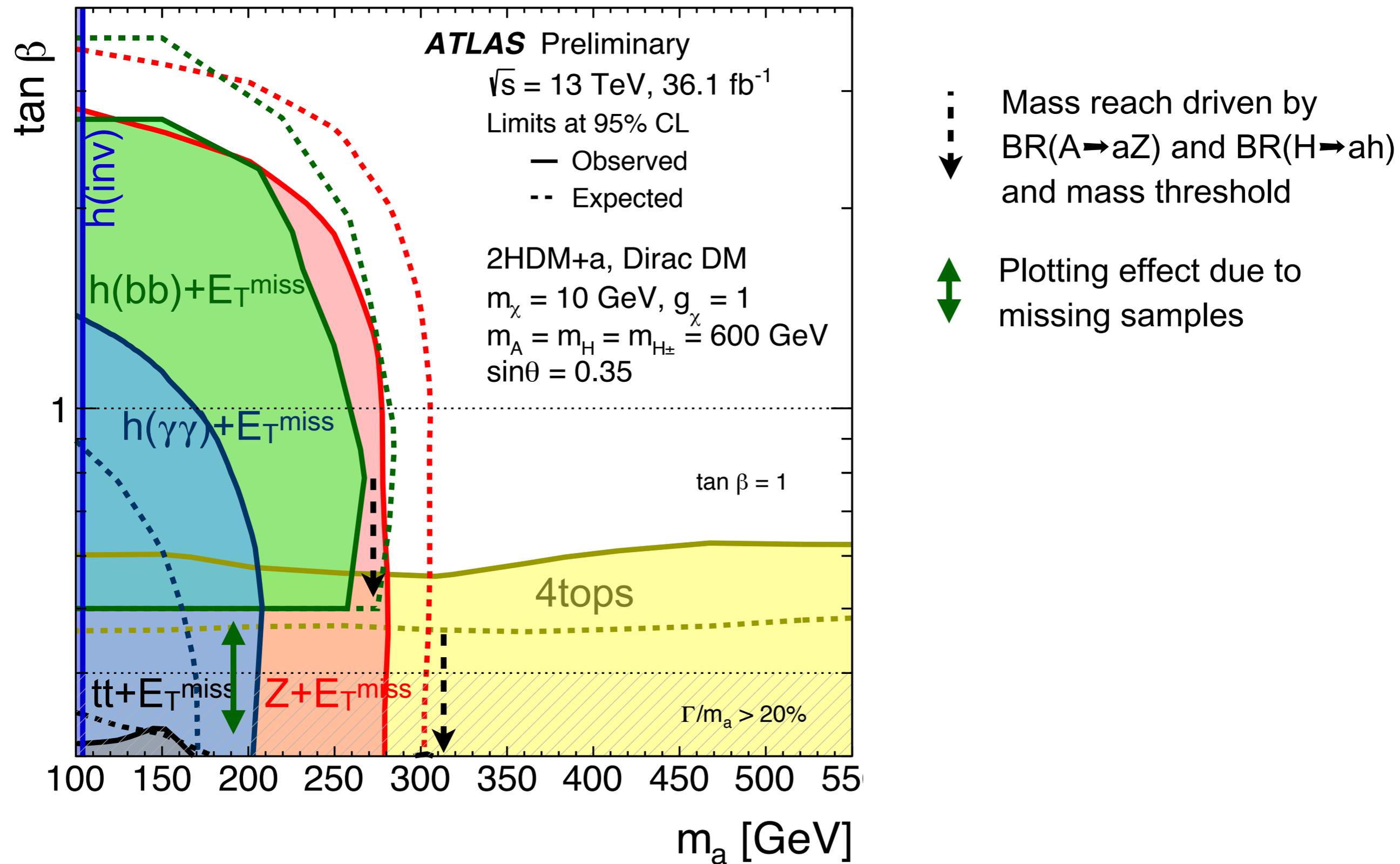
# Results (I)



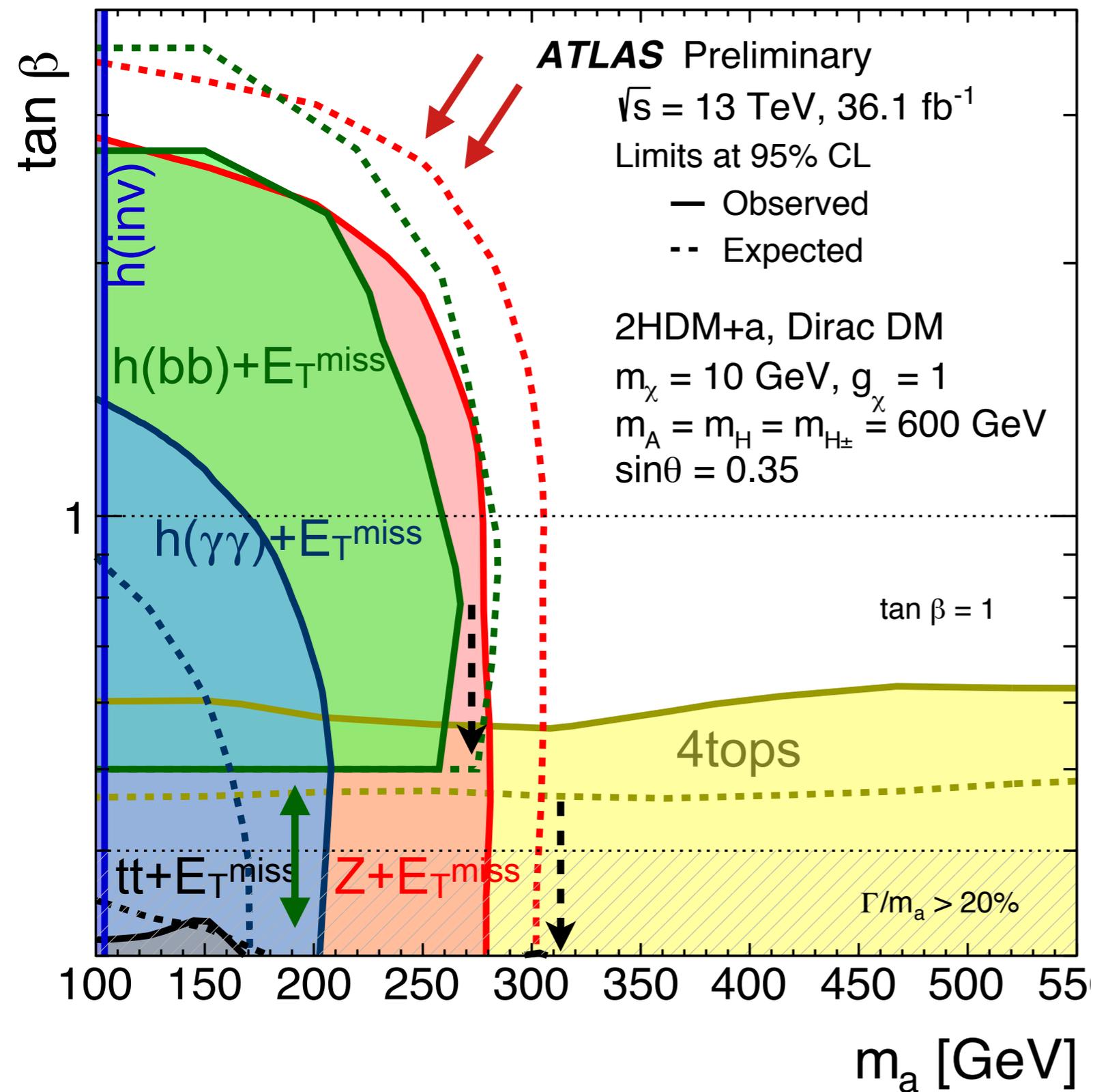
# Results (I)



# Results (I)

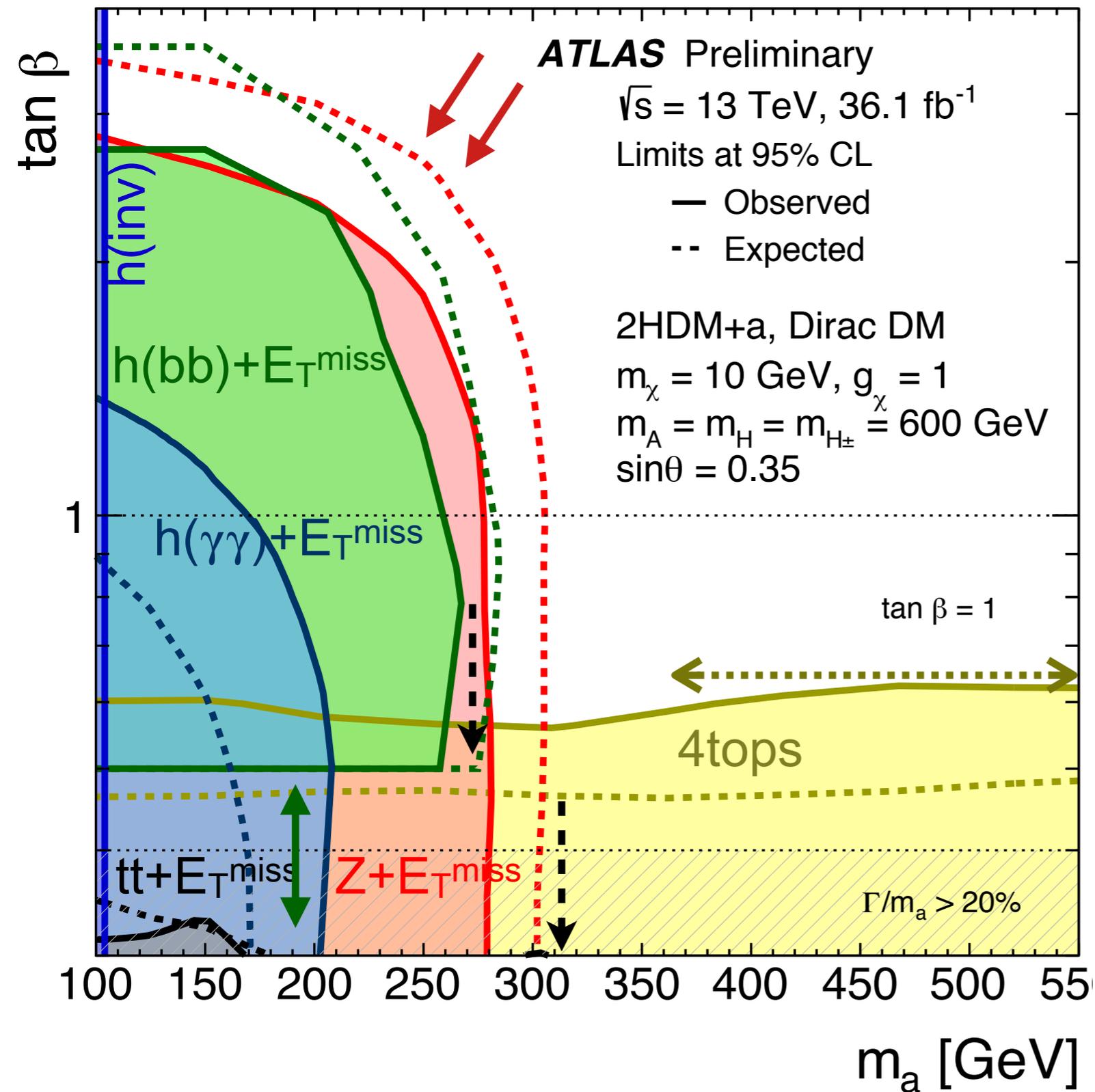


# Results (I)

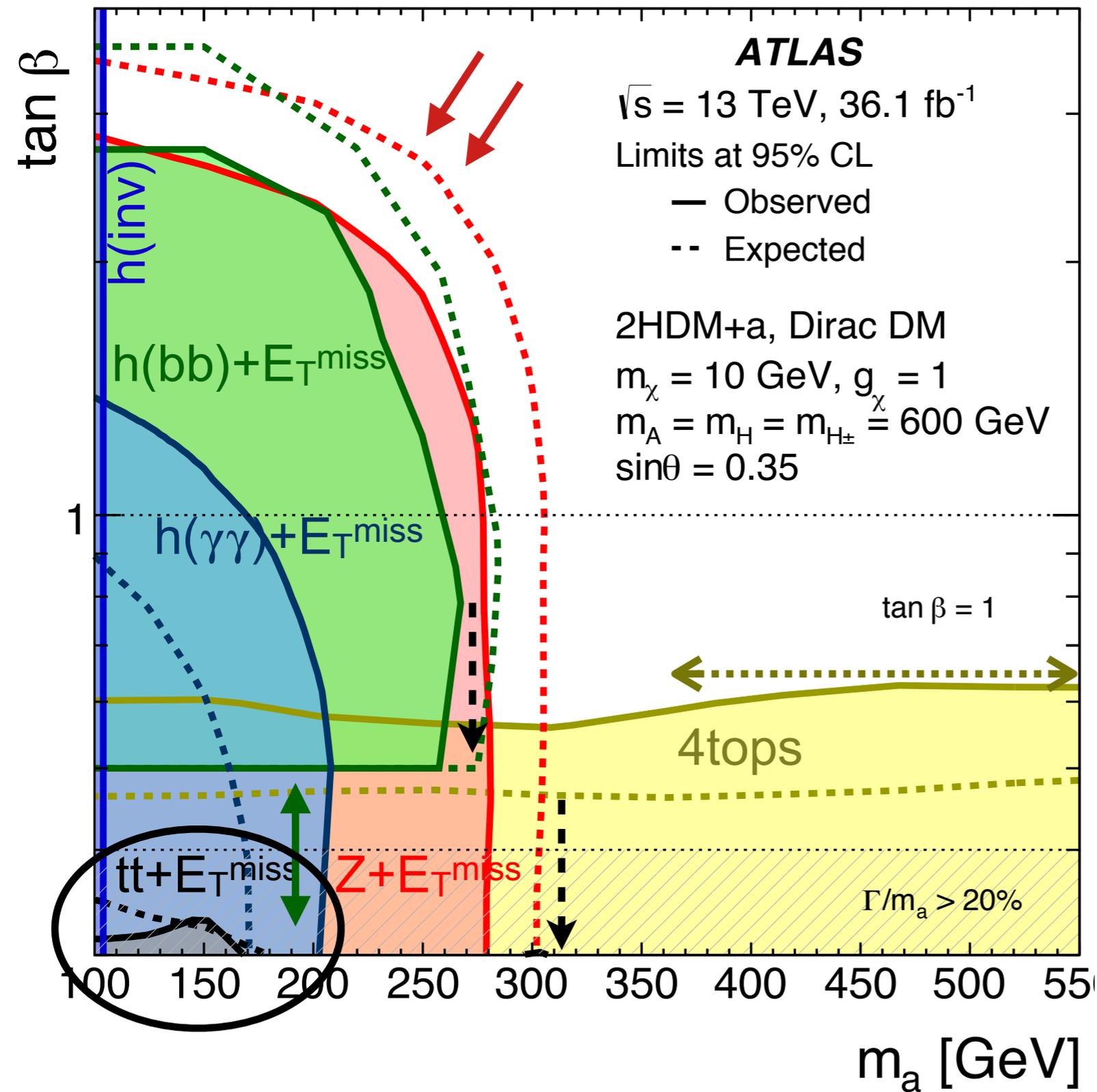


- Mass reach driven by  $\text{BR}(A \rightarrow aZ)$  and  $\text{BR}(H \rightarrow ah)$  and mass threshold
- Plotting effect due to missing samples
- Loss of sensitivity due to missing  $bb$ -initiated production

# Results (I)

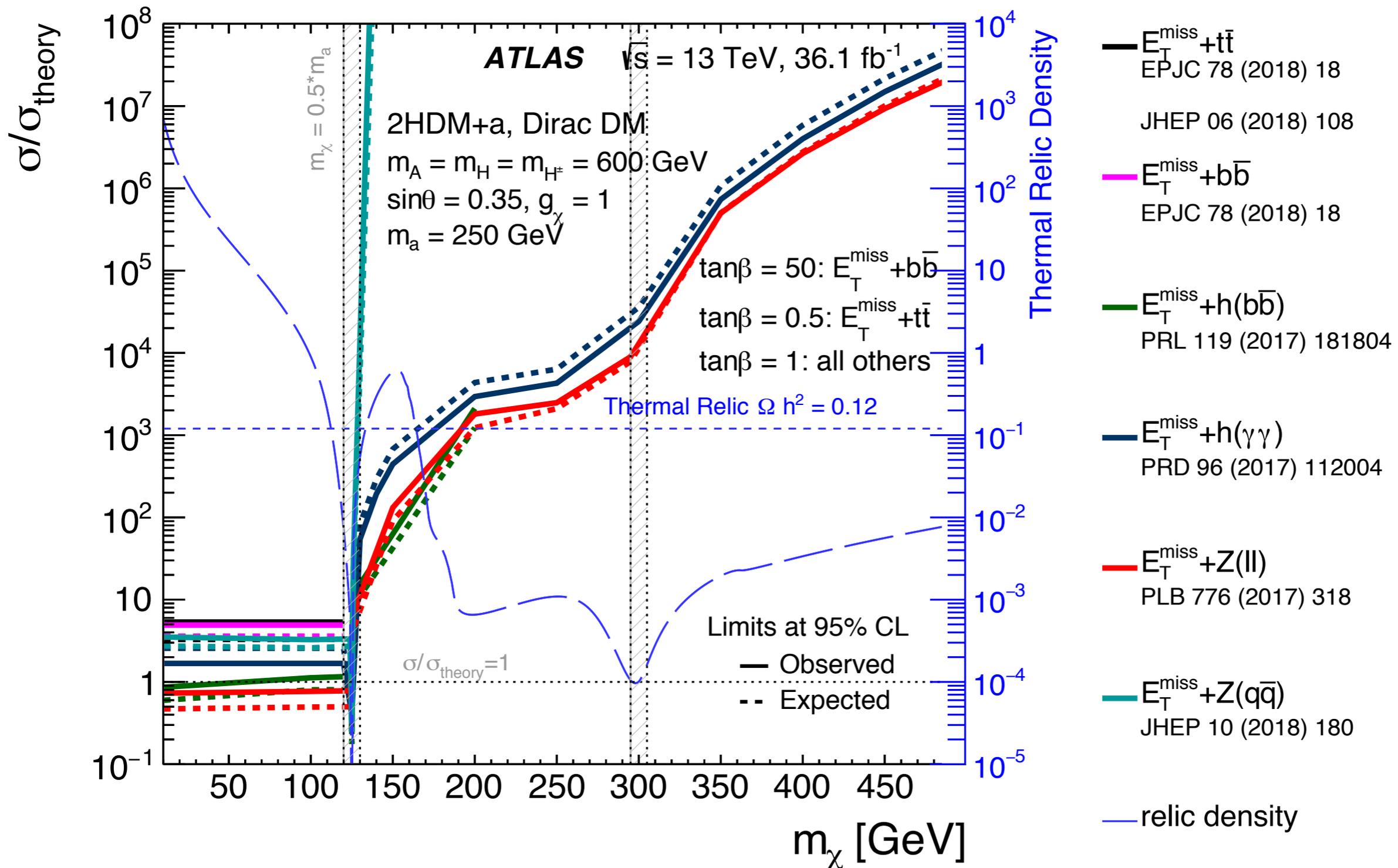


# Results (I)

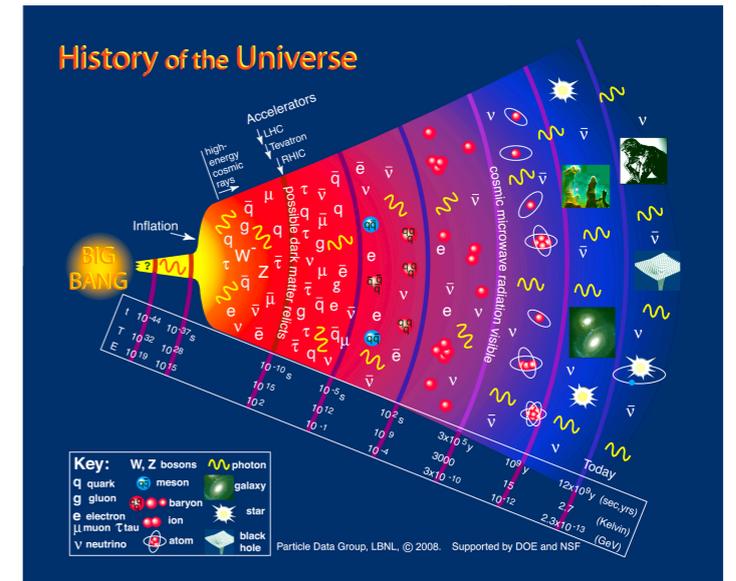
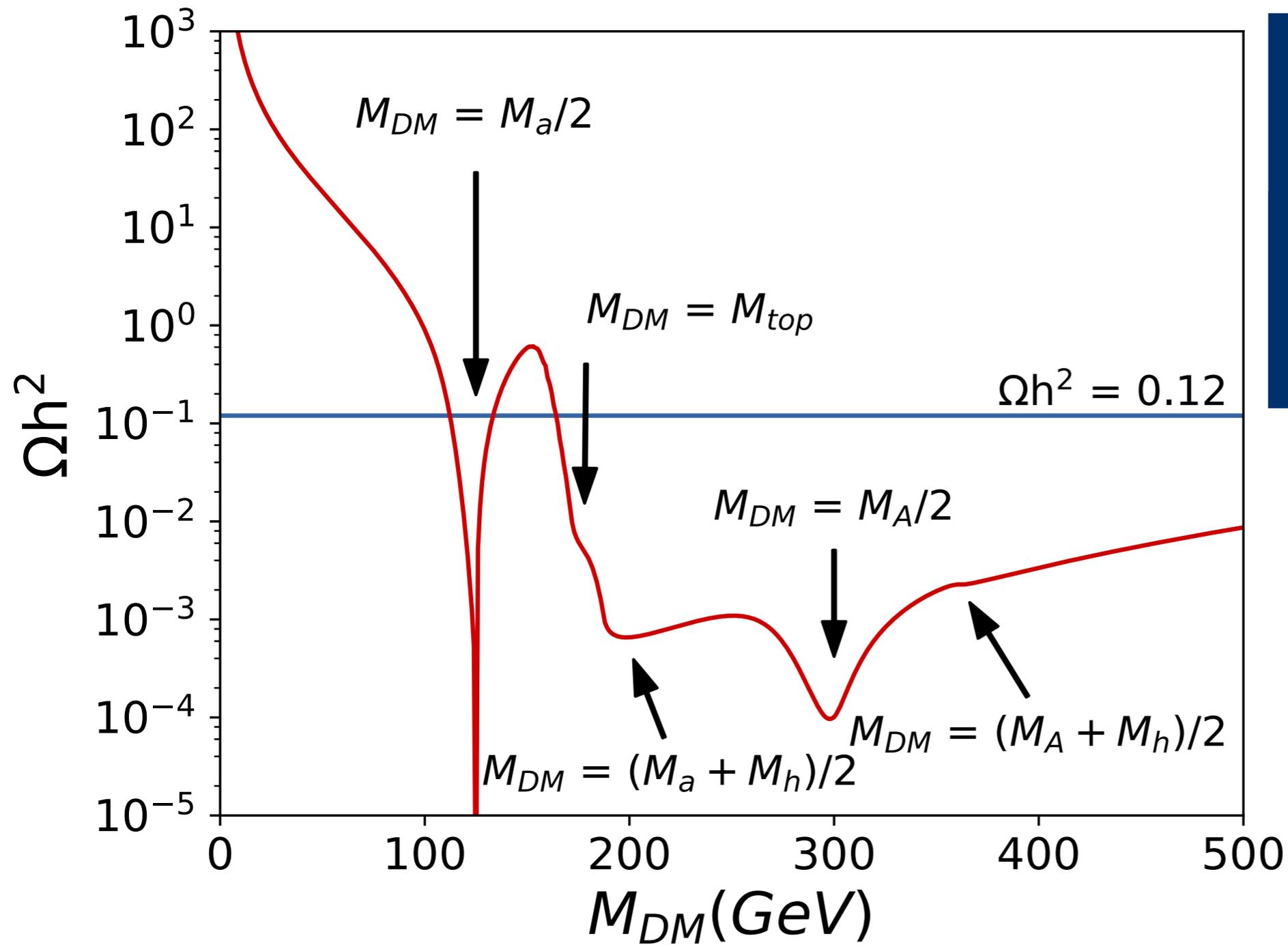


- Mass reach driven by  $\text{BR}(A \rightarrow aZ)$  and  $\text{BR}(H \rightarrow ah)$  and mass threshold
- Plotting effect due to missing samples
- Loss of sensitivity due to missing  $bb$ -initiated production
- Sensitivity dominated by  $ttA/H(tt) \Rightarrow$  independent on  $m(a)$
- Limited by the choice of  $\sin\theta$

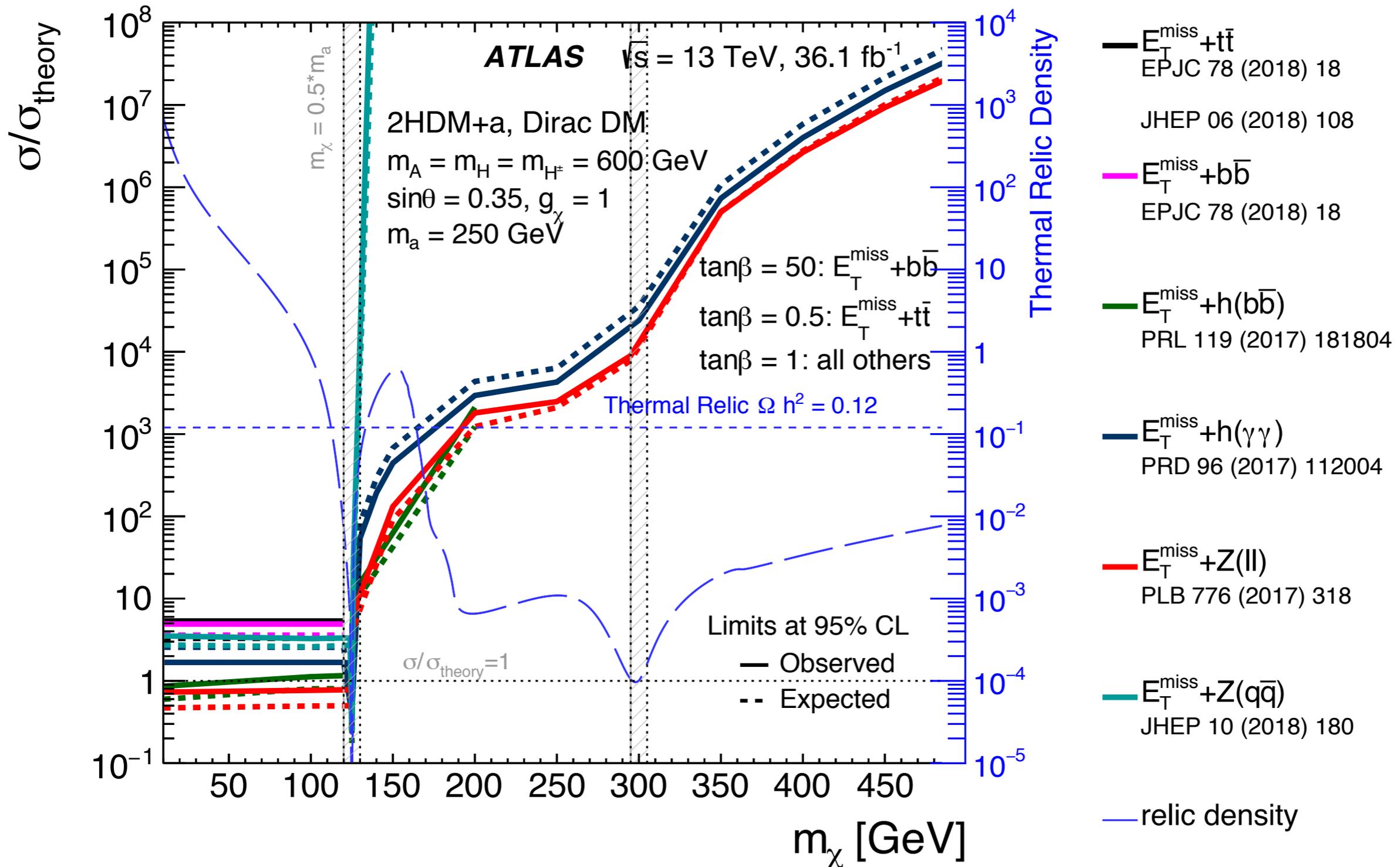
# Relic density perspective



# Understanding the relic prediction



# Relic density perspective



# Further considerations

## where to from here?

- ★ Many results with the full Run-2 datasets still in preparation but we can already plan ahead: leave no stone unturned!
- ★ [HL-LHC Yellow Report](#) shows many projection on searches evolution in the next data-taking periods, reaching higher higher DM & mediator masses
- ★ [LPCC DMWG](#) working on establishing additional “less simplified” frameworks



My personal take:

1) NEW TRIGGERS 2) NEW SIGNATURES

# Recording more/better data

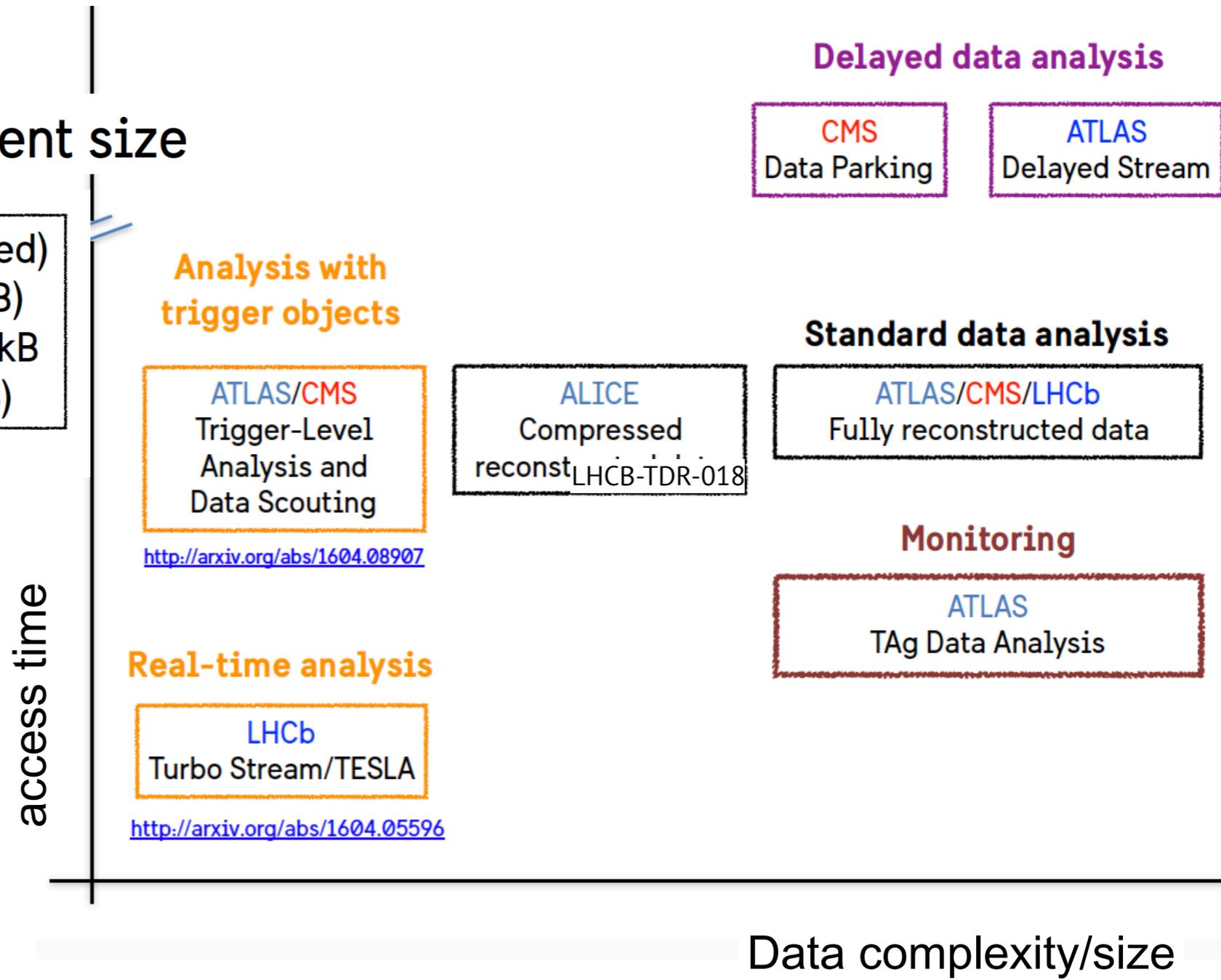
**Limited by:**

- fast read-out of  $\sim 100M$  detector channels
- computing resources (reconstruction)
- disk storage (saving for further processing)
- everyone else's favourite physics channel

**Bandwidth = Event rate x Event size**

LHC: 40 MHz  
 ATLAS: 1 kHz  
 LHCb: 12.5 kHz  
 CMS: 1 kHz

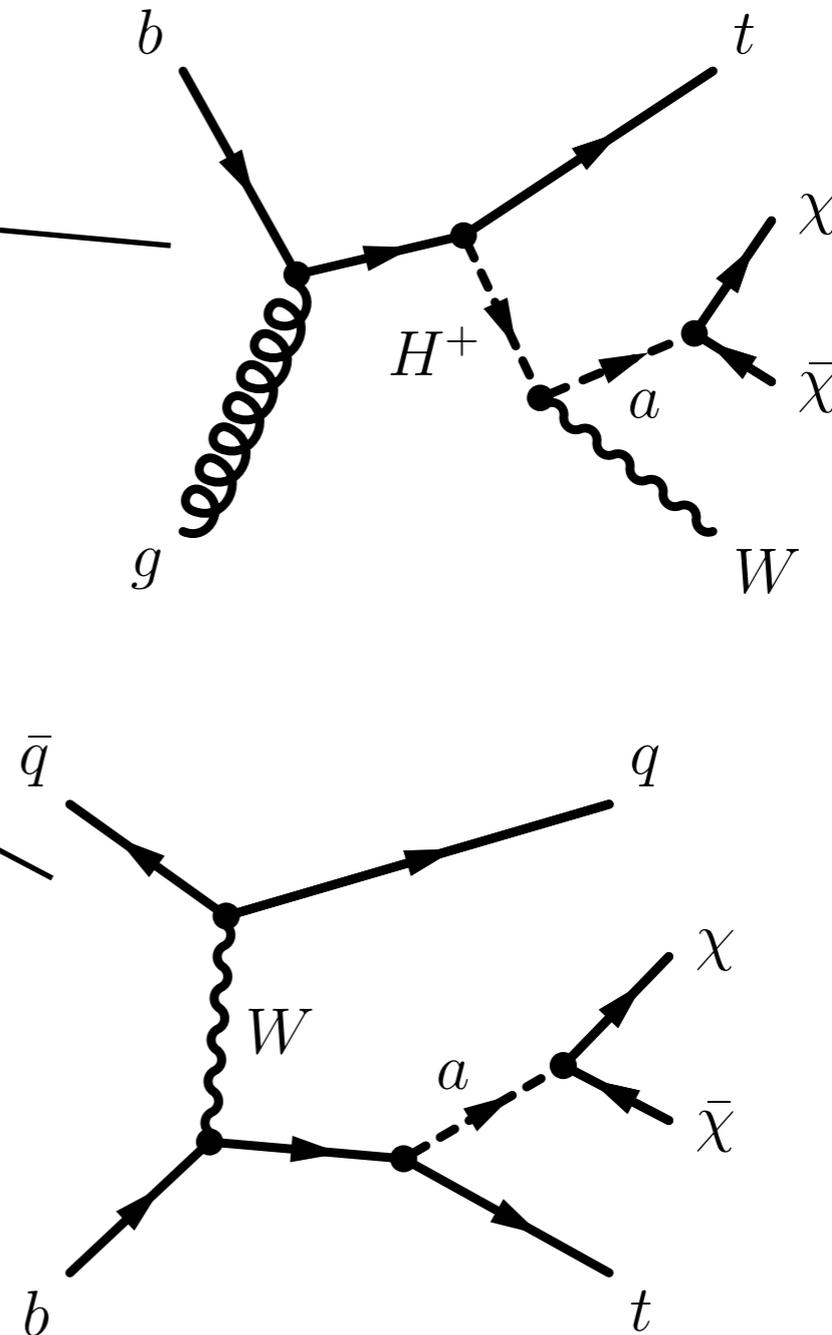
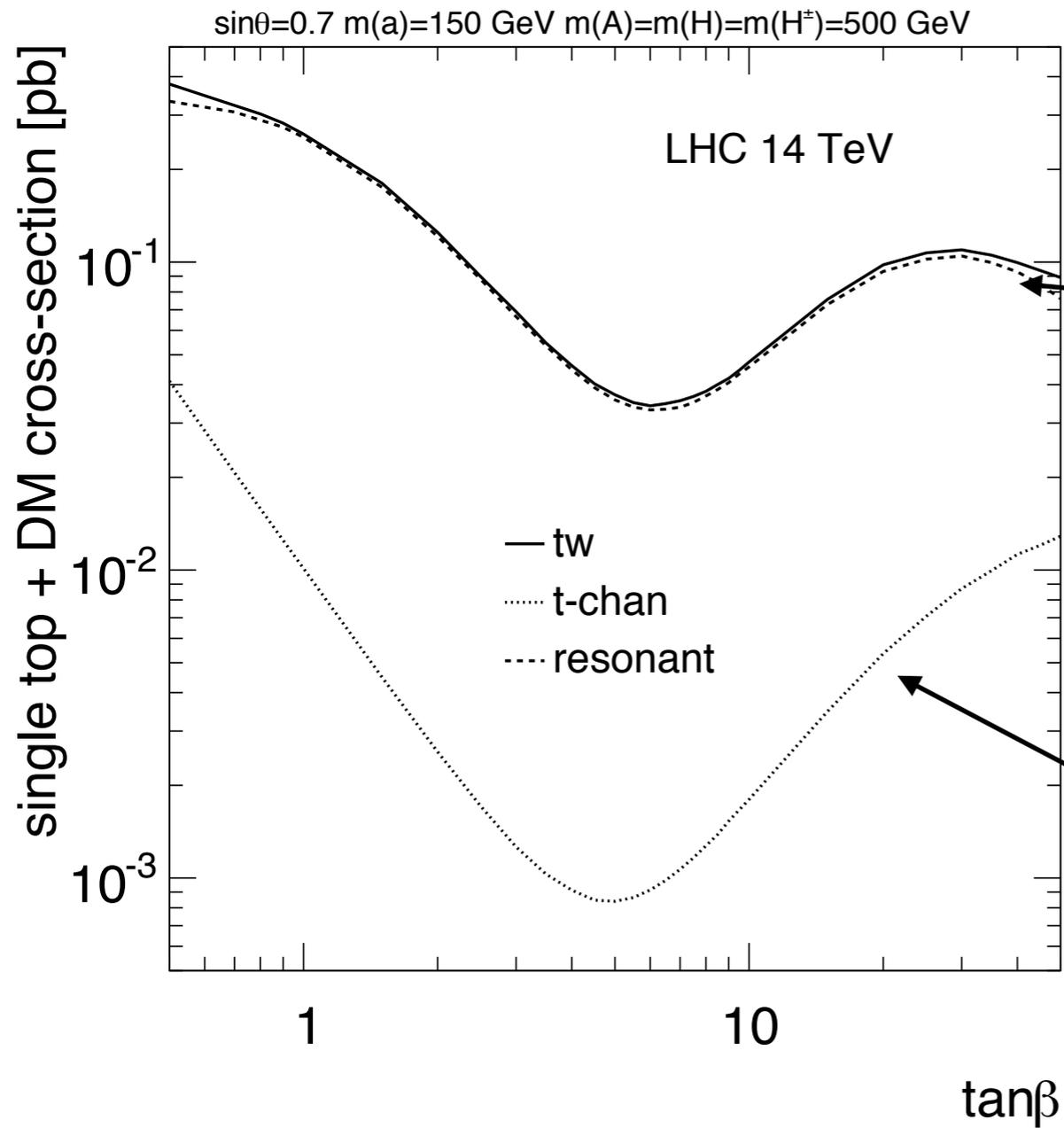
(Reconstructed)  
 ATLAS:  $\sim 100$  MB  
 LHCb:  $\sim 100$  kB  
 CMS:  $\sim 100$  MB



We need more of this !

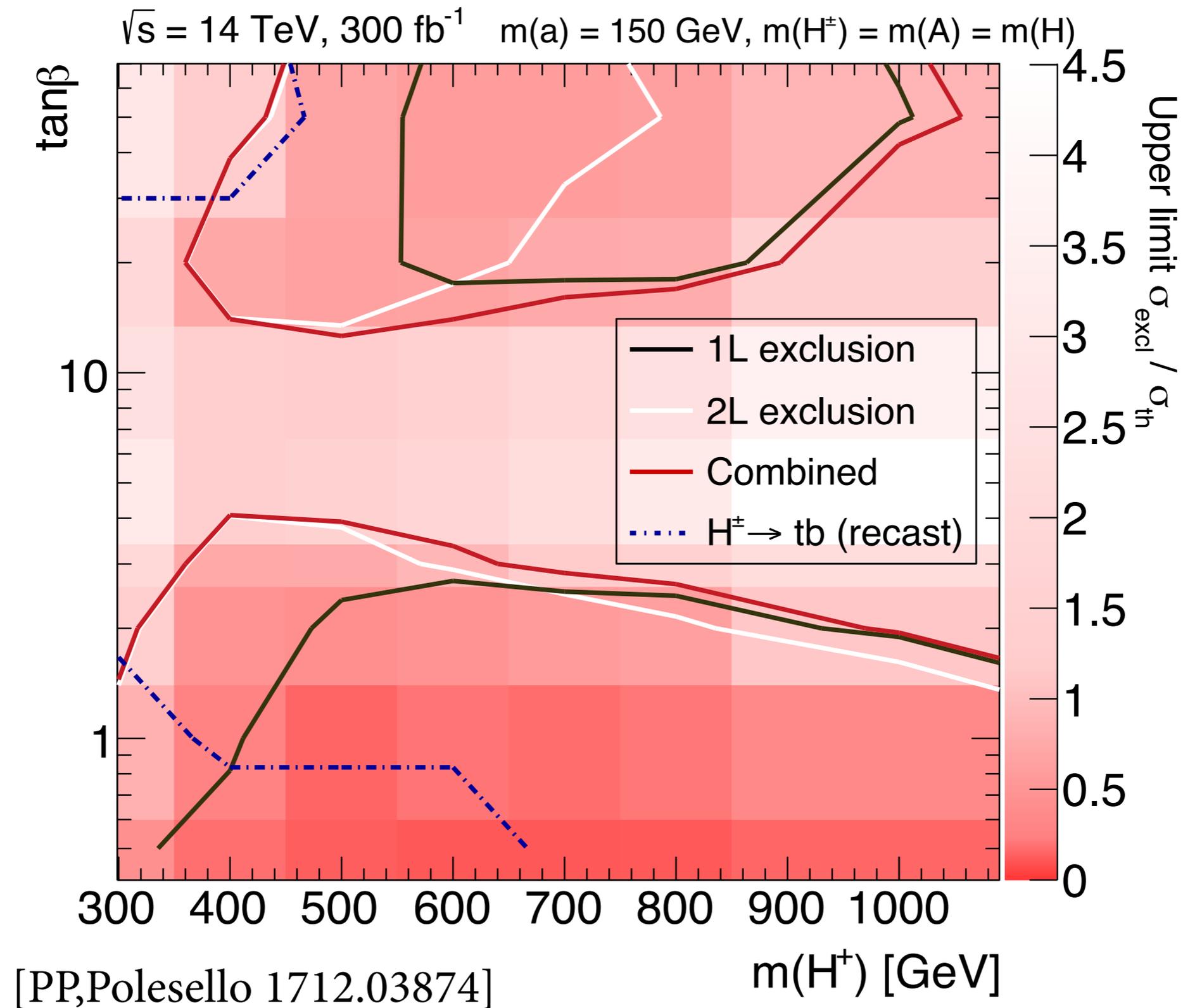
# New signatures

$< TeV$ <b><math>H^\pm</math></b>	$< TeV$ <b><math>A</math></b>	$\sim 10 GeV$ <b><math>\chi^0</math></b>
$125 GeV$ <b><math>h</math></b>	$< TeV$ <b><math>H</math></b>	$\sim 10^2 GeV$ <b><math>a</math></b>

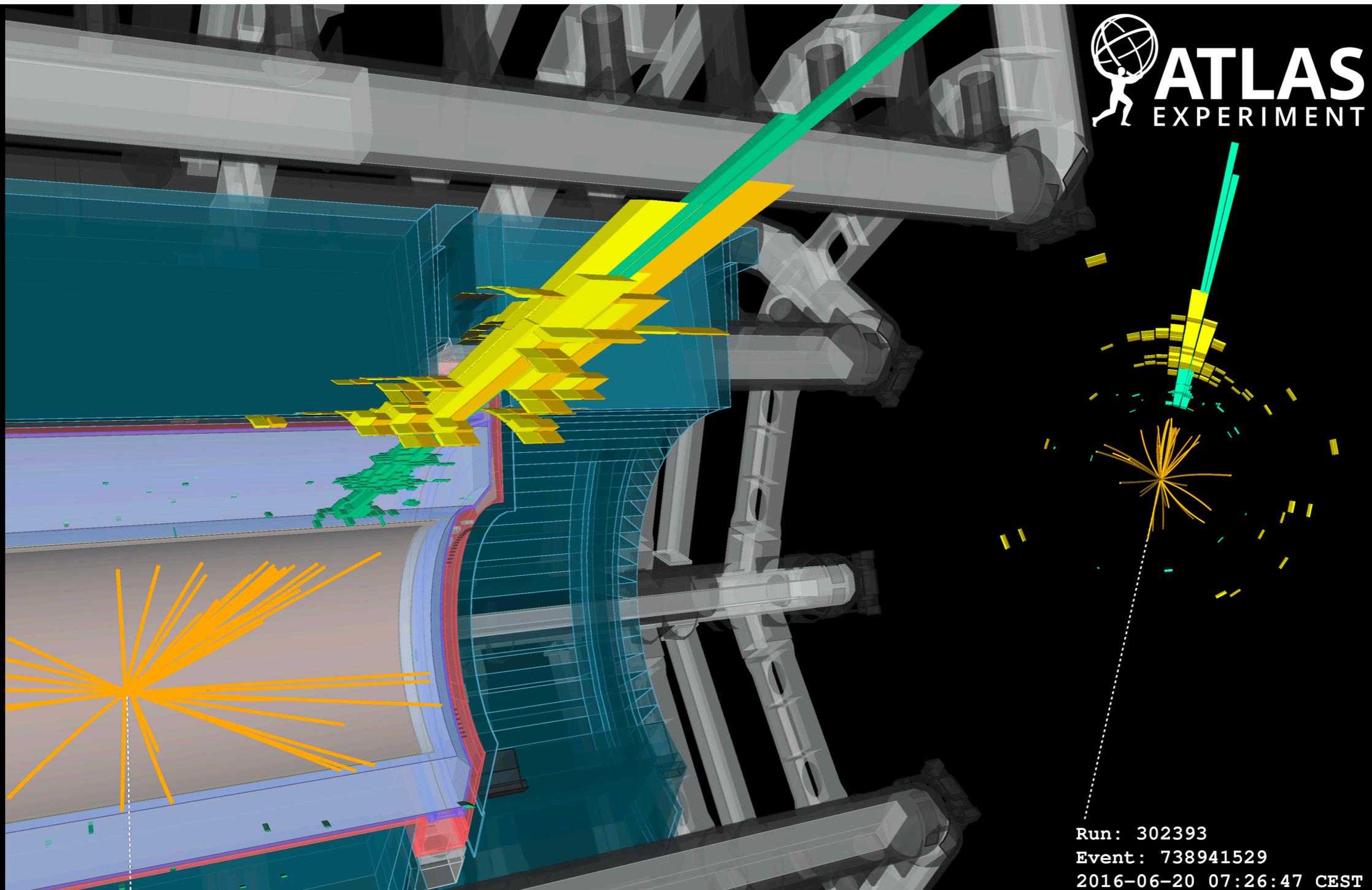


[PP,Polesello 1712.03874]

# Sensitivity forecast



# Bonus: Dark Energy

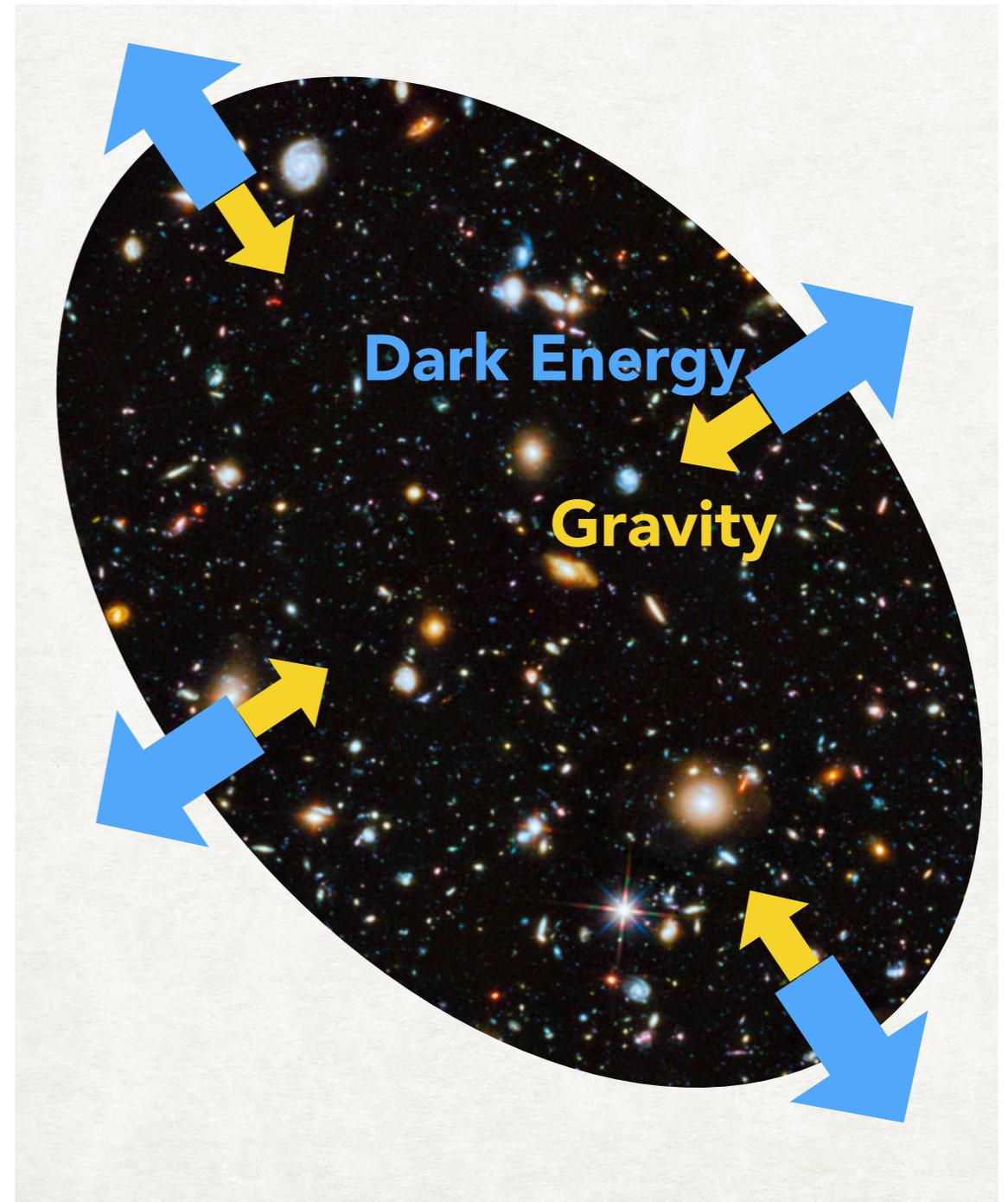


 **ATLAS**  
EXPERIMENT

Run: 302393  
Event: 738941529  
2016-06-20 07:26:47 CEST

# Dark Energy

- ★ Dark Energy = universe accelerated expansion
- ★ Big **unanswered question** in cosmology and particle physics
  - new particle or modified gravity?
  - constant or dynamic?
  - interacting or not?
  - microscopic nature?
- ★ **no leading candidate theory**

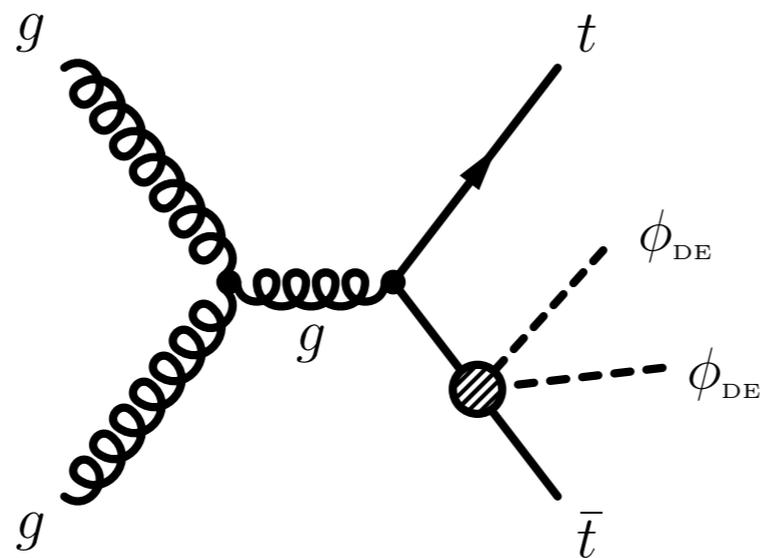


# Horndeski EFT model

$$\mathcal{L}_{\text{SM}} + \sum_{i=1}^9 \frac{c_i}{M_i^{(d-4)}} \mathcal{O}_i^{(d)},$$

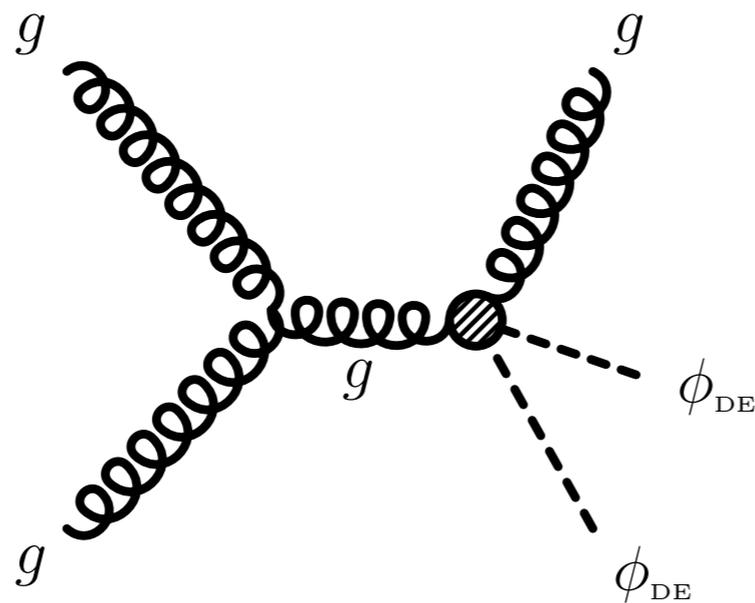
1 scalar field  $\phi_{\text{DE}}$  coupled to gravity

$$\mathcal{L}_1 = \frac{\partial_\mu \phi \partial^\mu \phi}{M_1^4} T_\nu^\nu,$$



→ tops +  $E_{\text{T}}^{\text{miss}}$

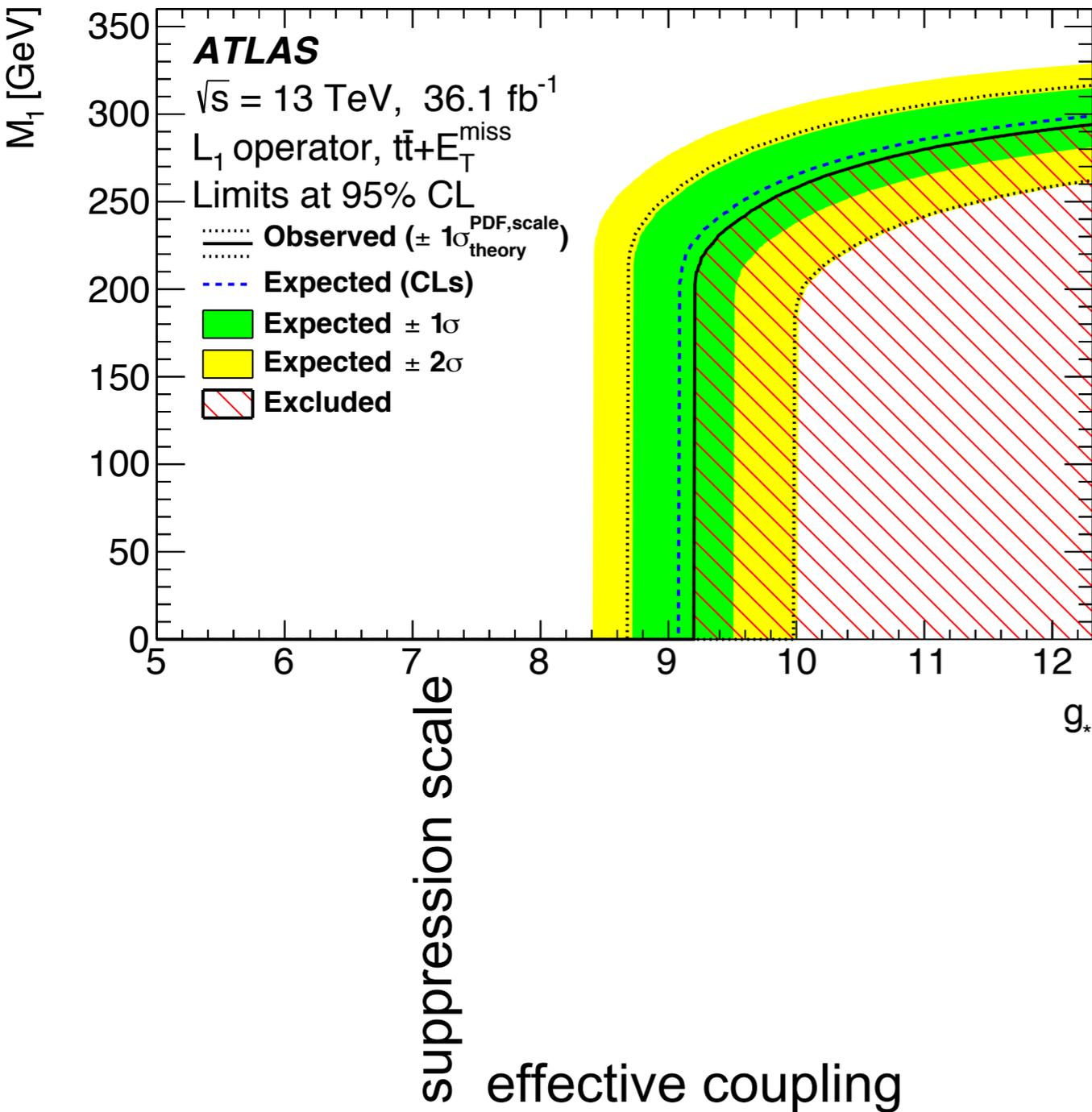
$$\mathcal{L}_2 = \frac{\partial_\mu \phi \partial_\nu \phi}{M_2^4} T^{\mu\nu},$$



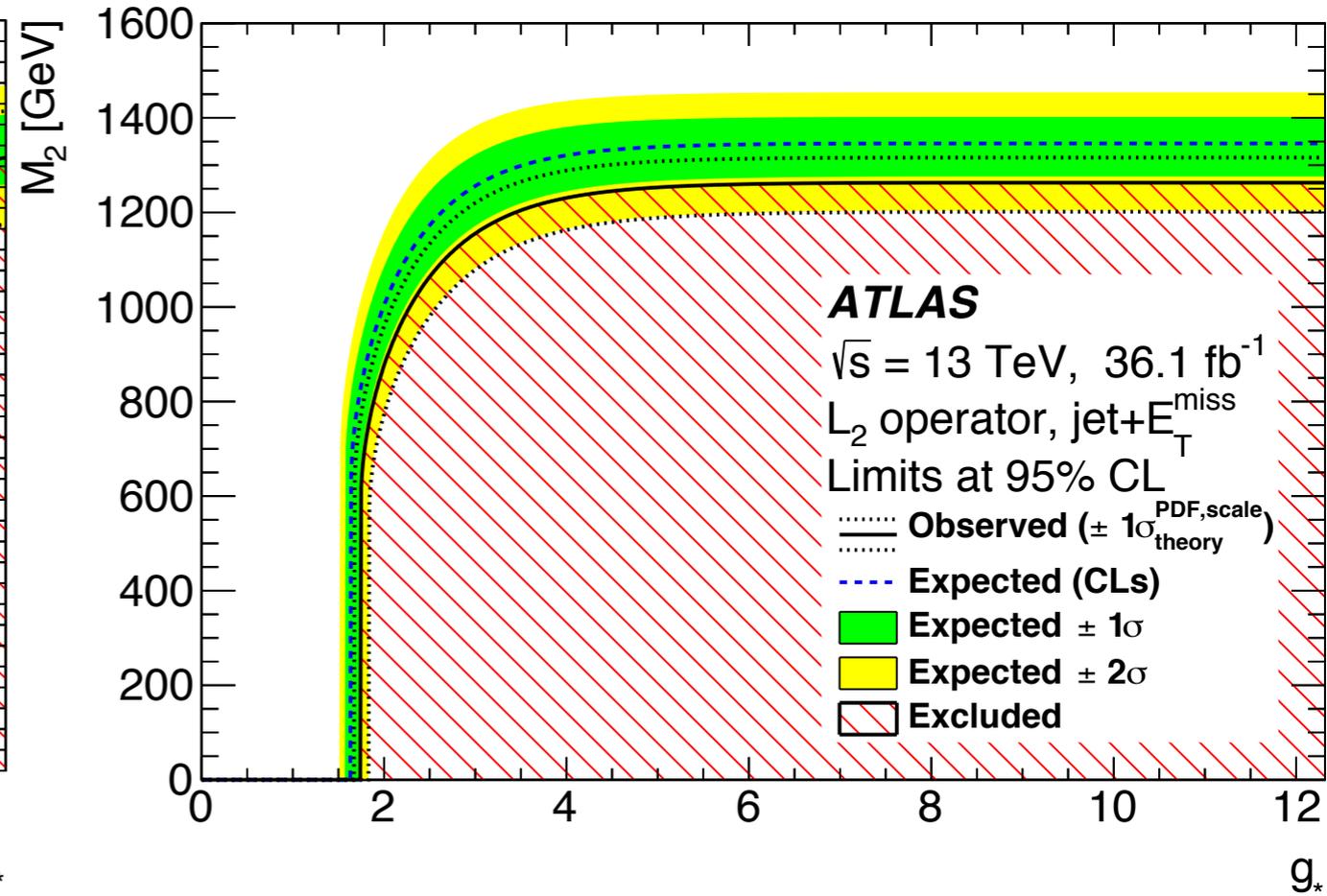
→ jet +  $E_{\text{T}}^{\text{miss}}$

# Results

tops +  $E_T^{\text{miss}}$



jet +  $E_T^{\text{miss}}$



# Conclusion - Cheat sheet

## DM-mediator searches

Signature	Dataset	Reference
Di-lepton resonance	139 fb <sup>-1</sup>	<a href="#">1903.06248</a>
Di-jet, Di-jet + ISR,	139 fb <sup>-1</sup>	<a href="#">1901.10917</a> , <a href="#">ATLAS-CONF-2019-007</a> , <a href="#">1808.03124</a>
Di-bjet	80 fb <sup>-1</sup>	<a href="#">ATLAS-CONF-2018-052</a>
Di-jet + leptons	80 fb <sup>-1</sup>	<a href="#">ATLAS-CONF-2018-015</a>
Dijet + photons	36 fb <sup>-1</sup>	<a href="#">1905.10331</a>
Etmiss + Higgs	36 fb <sup>-1</sup>	<a href="#">1908.01713</a>
Etmiss + t/ttbar	36 fb <sup>-1</sup>	<a href="#">1901.01553</a>
Etmiss + jet	36 fb <sup>-1</sup>	<a href="#">1712.02345</a>
H invisible	36 fb <sup>-1</sup>	<a href="#">Phys. Rev. Lett. 122 (2019) 231801</a>
ATLAS DM summary	36 fb <sup>-1</sup>	<a href="#">JHEP 05 (2019) 142</a>

# Thanks for your attention!

## Contact

**DESY.** Deutsches  
Elektronen-Synchrotron

[www.desy.de](http://www.desy.de)

Dr. Priscilla Pani

ATLAS Group Campus Zeuthen

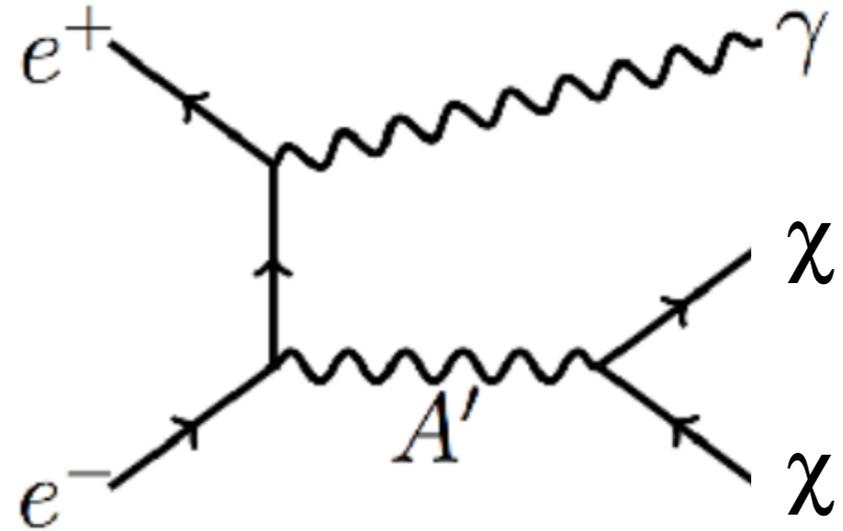
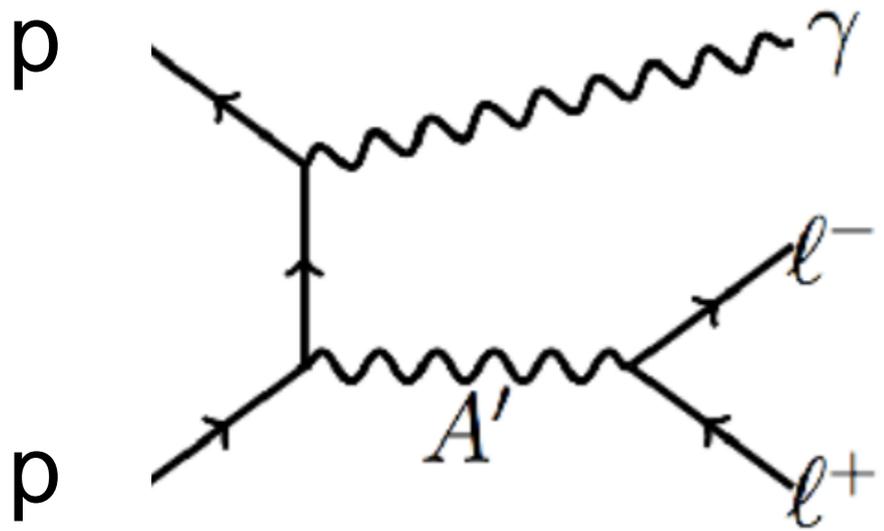
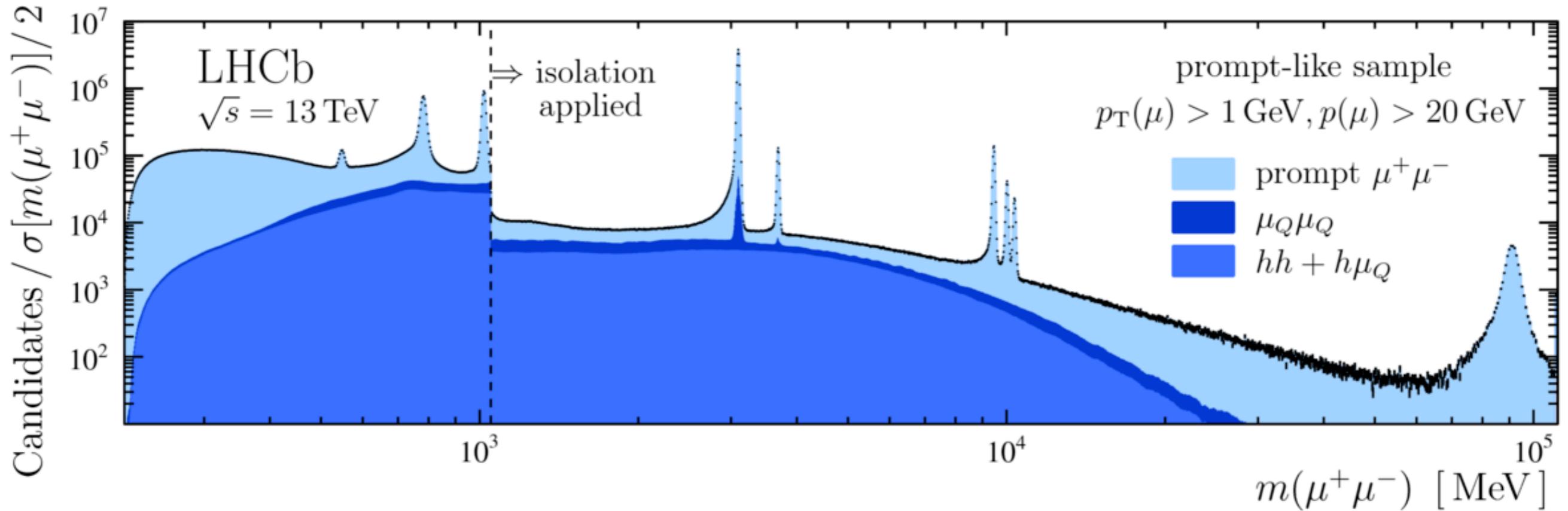
[priscilla.pani@desy.de](mailto:priscilla.pani@desy.de)

[https://atlas.desy.de/external\\_grants/priscilla\\_pani\\_yig/](https://atlas.desy.de/external_grants/priscilla_pani_yig/)

# Backup



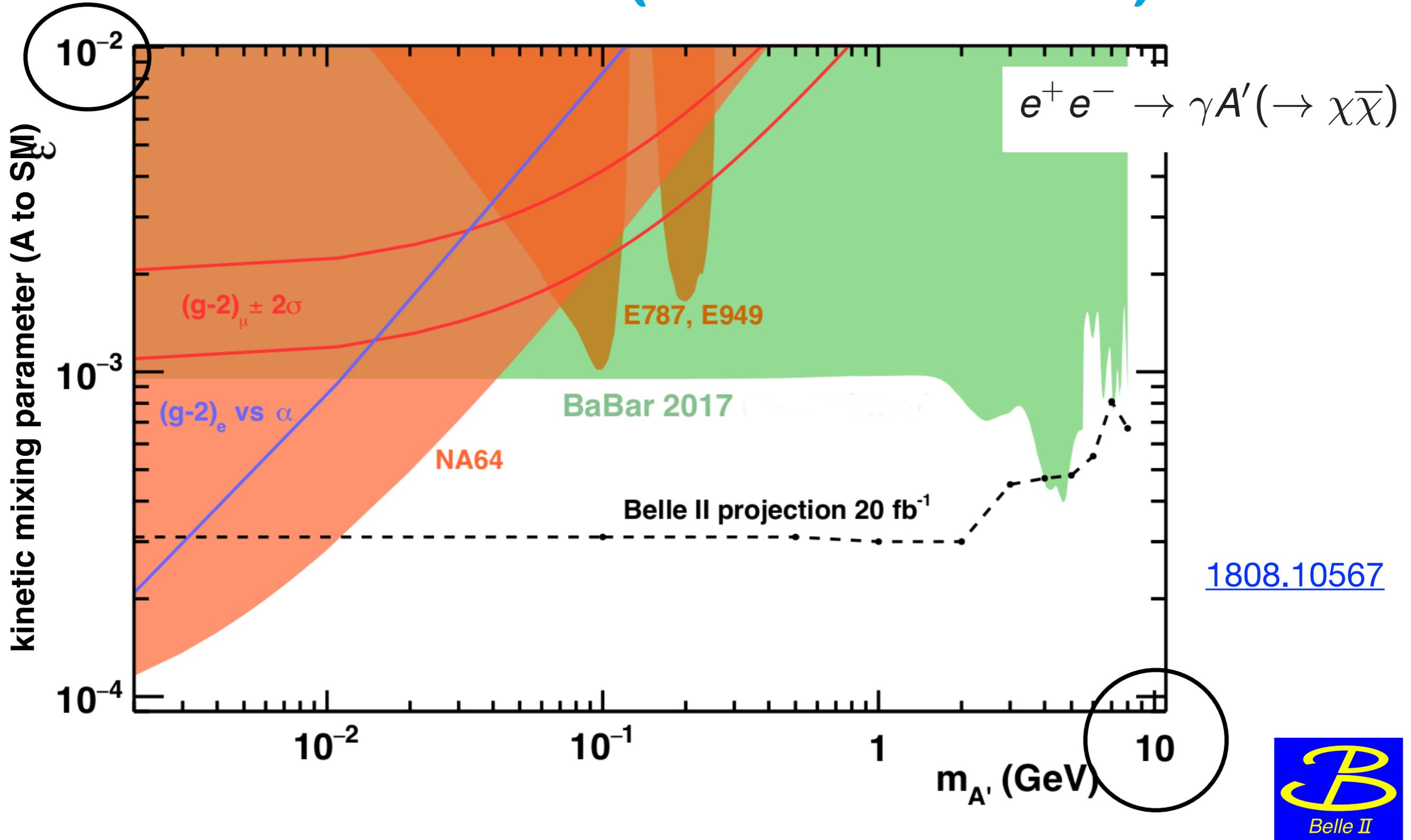
# A word on Dark Photons



See more details on Thursday's parallel session (DM15)

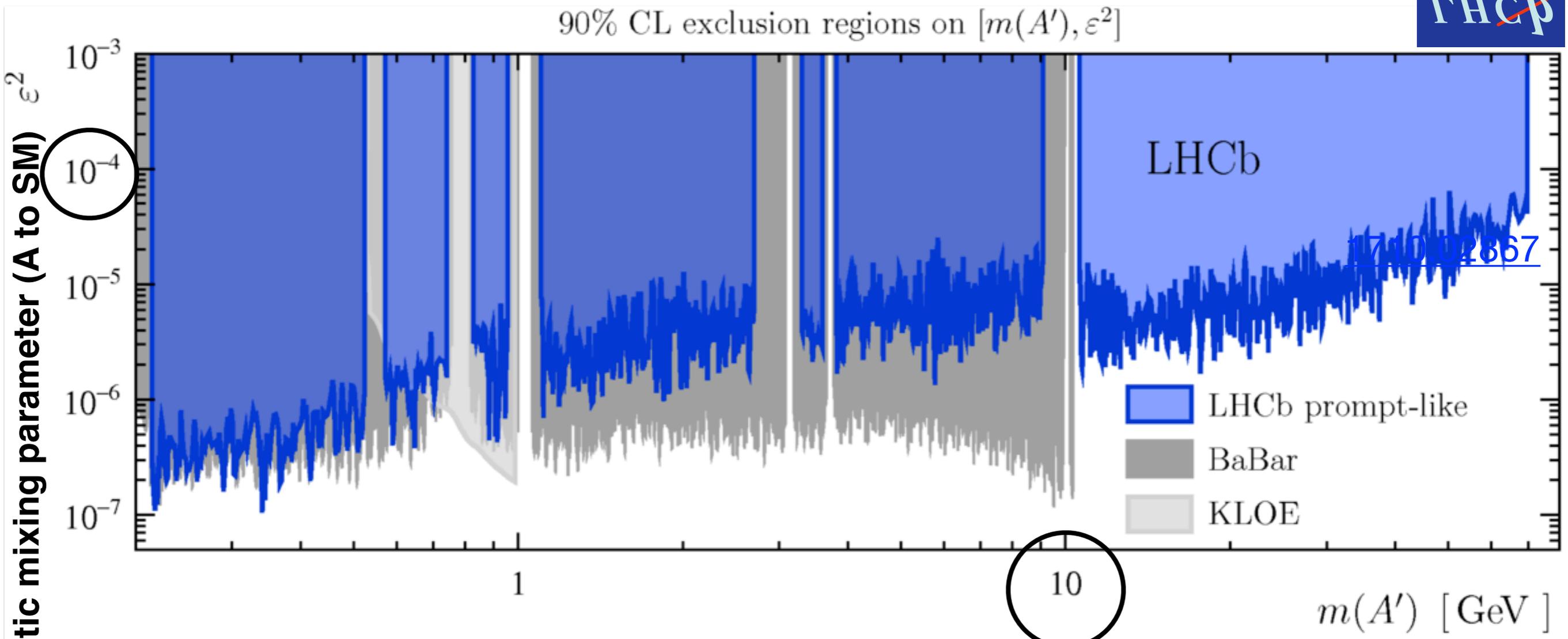


# Dark Photons (low masses)



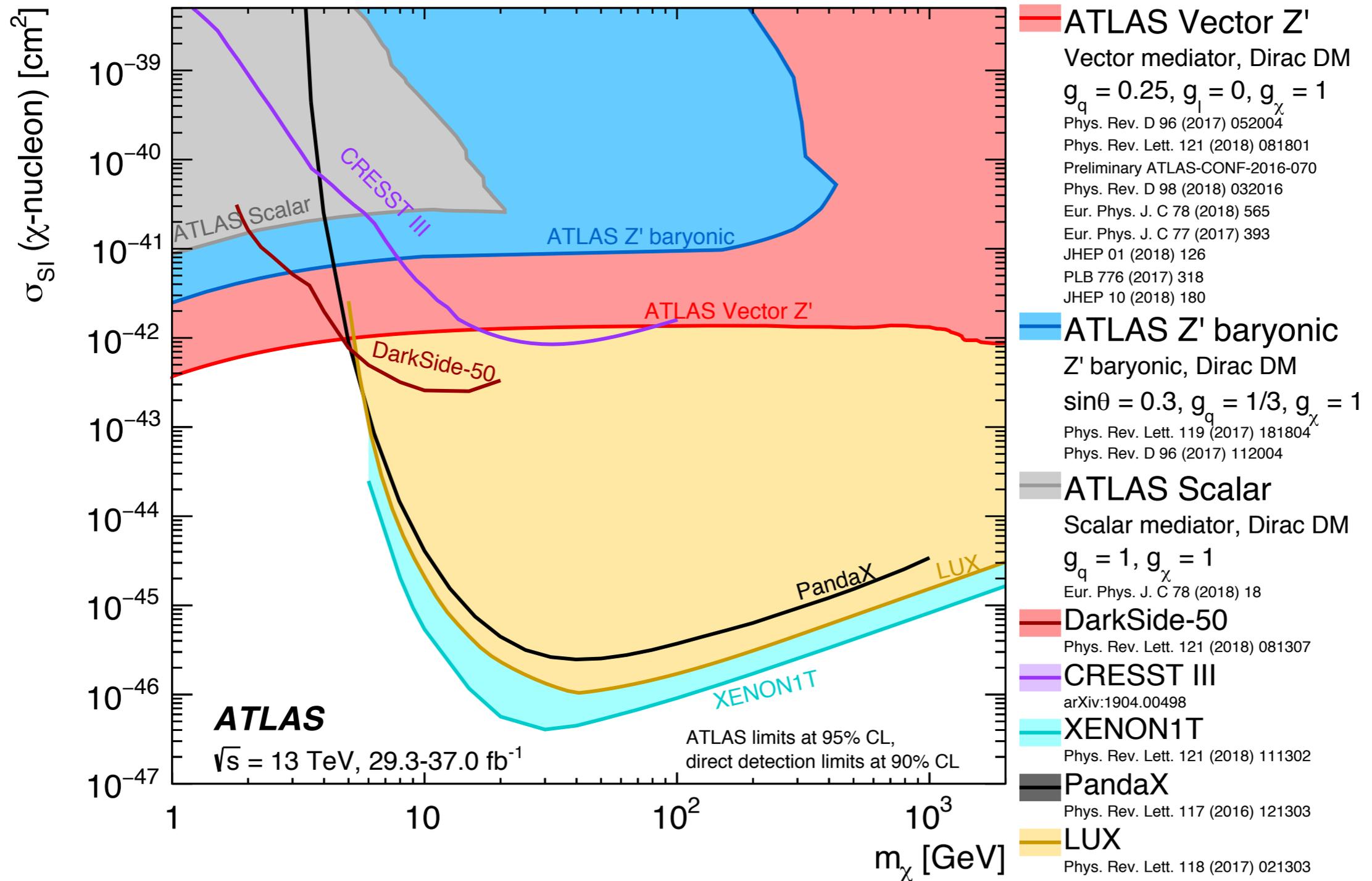
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# Dark Photons (low masses)

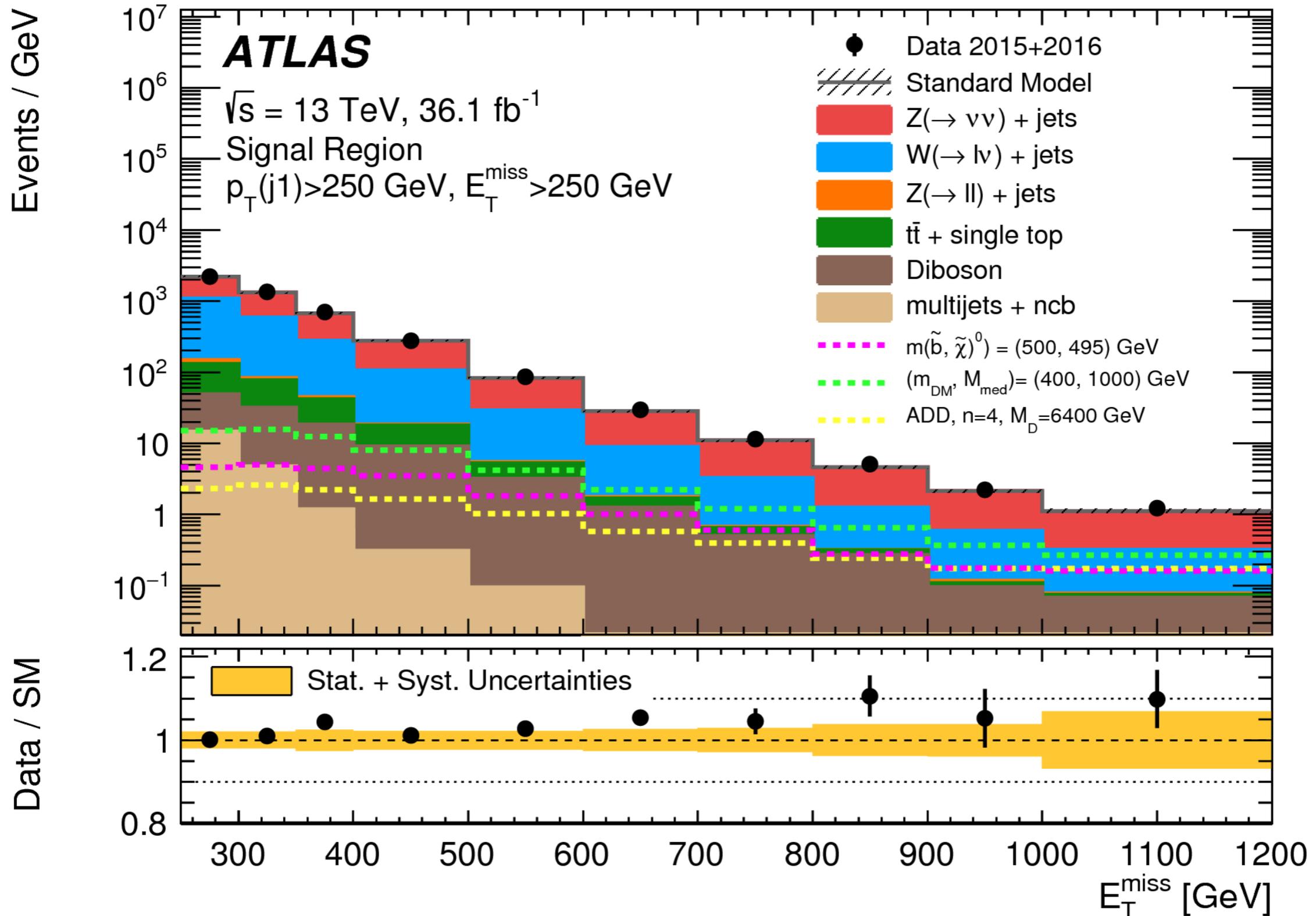


# Comparing to direct detection

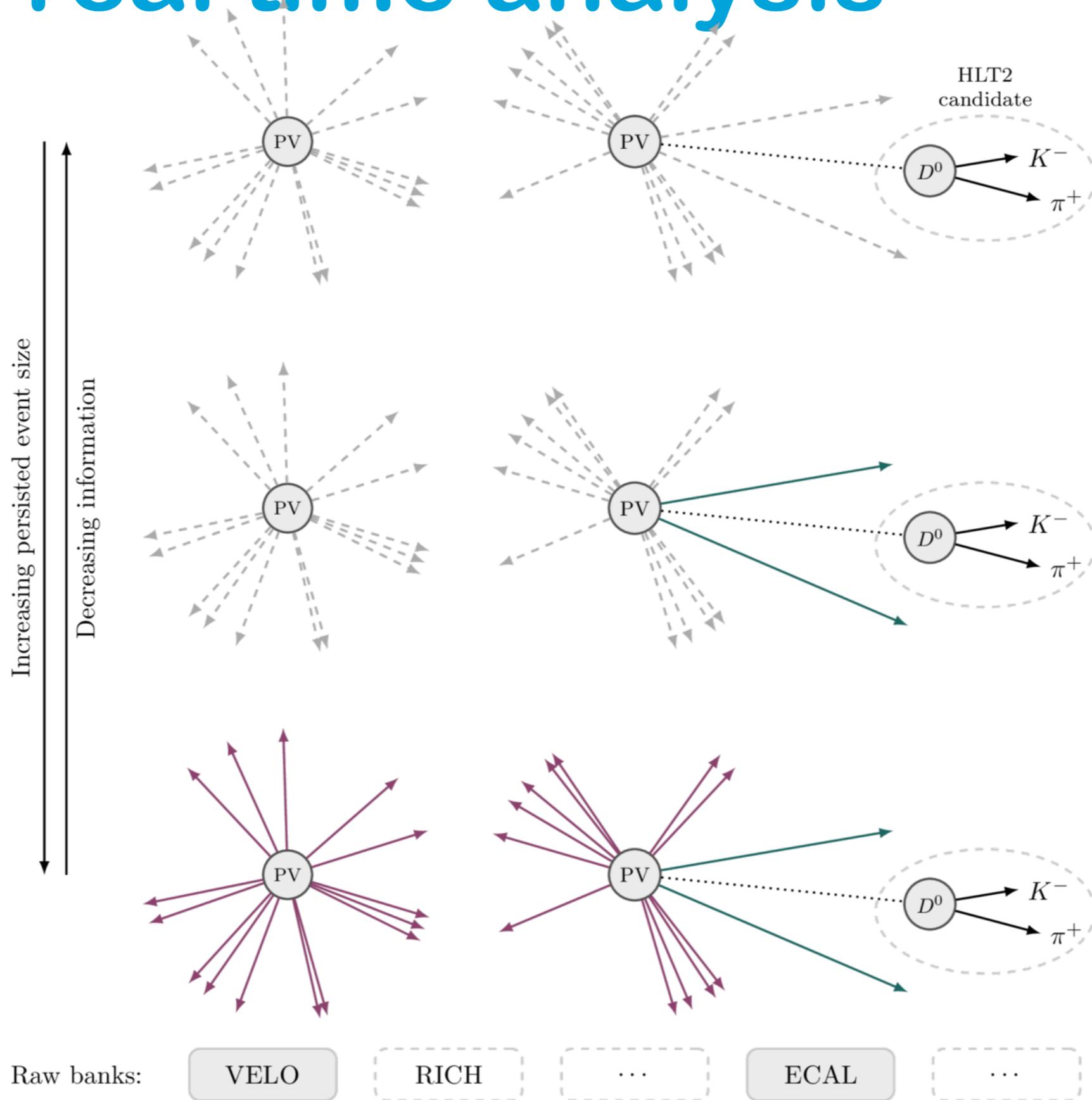
“The plot”



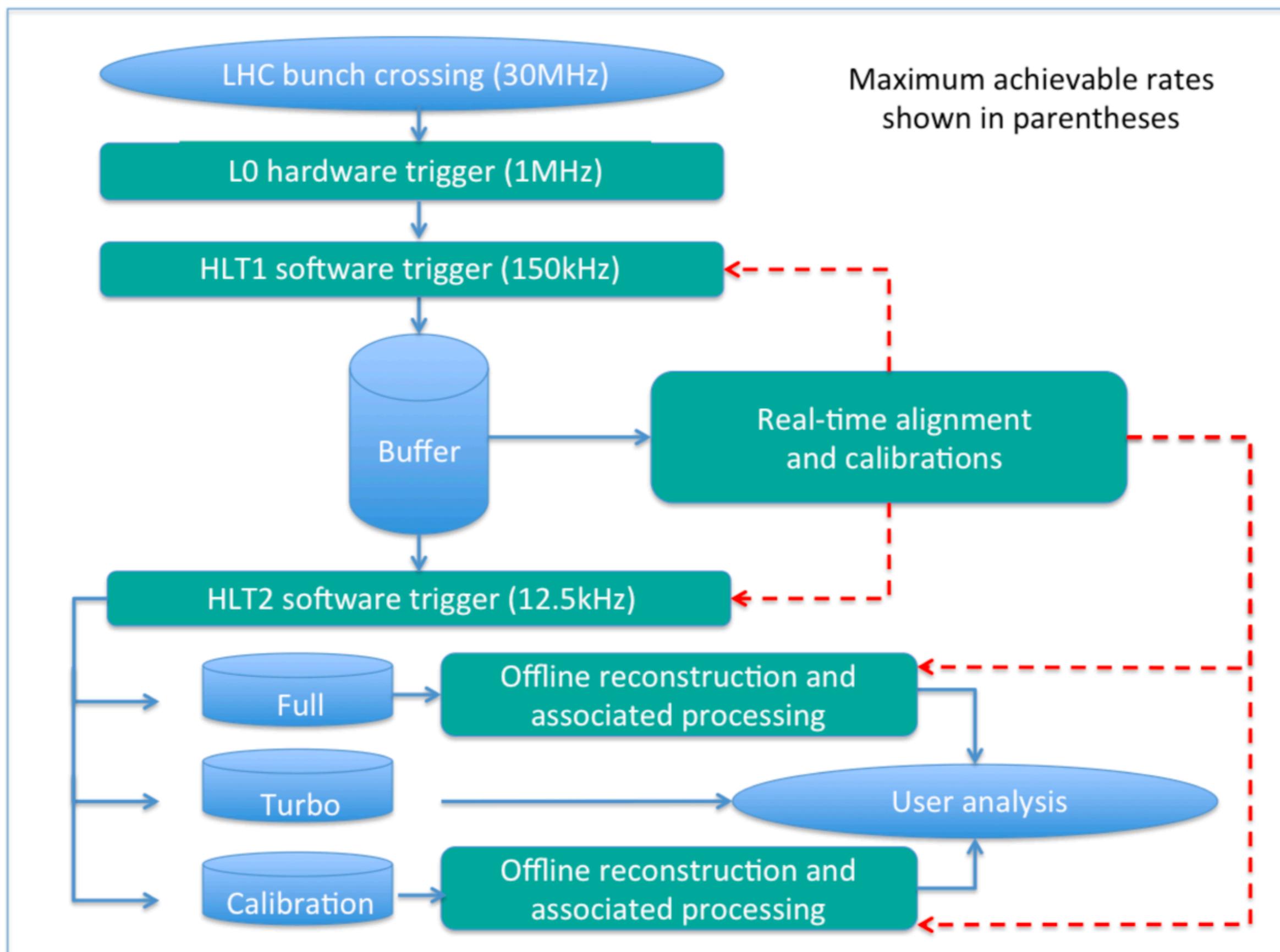
# mono-jet SR



# LHC real time analysis

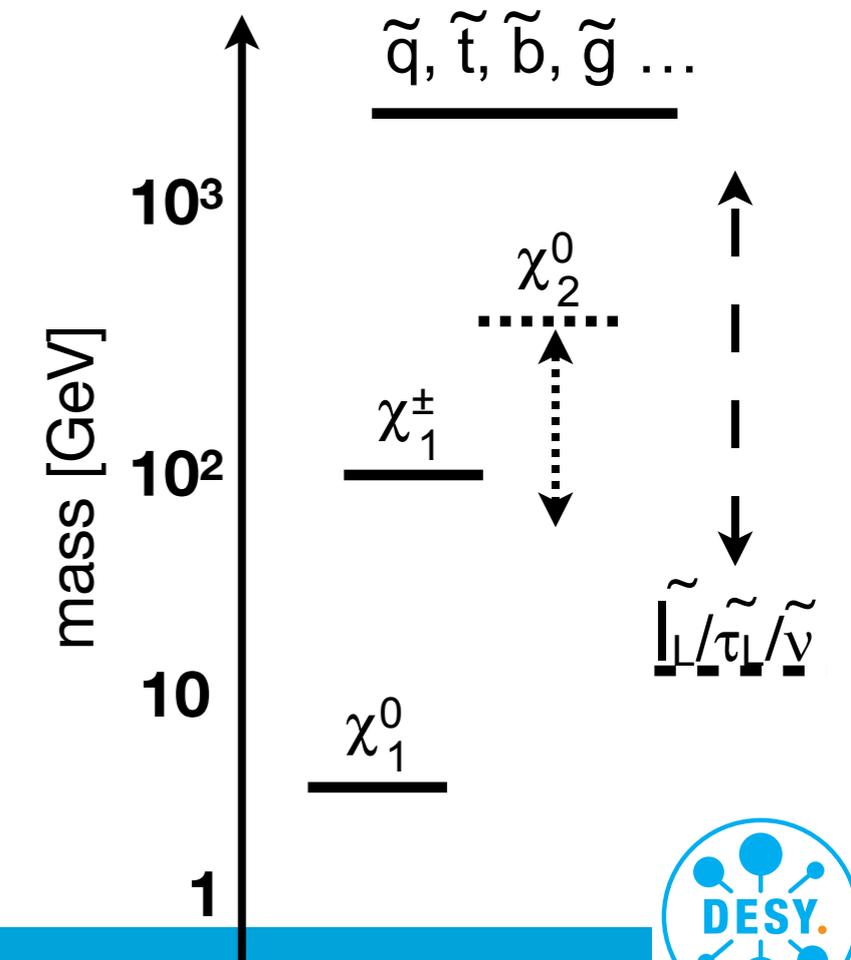
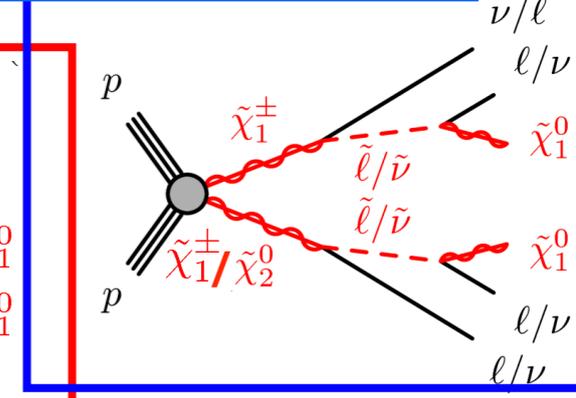
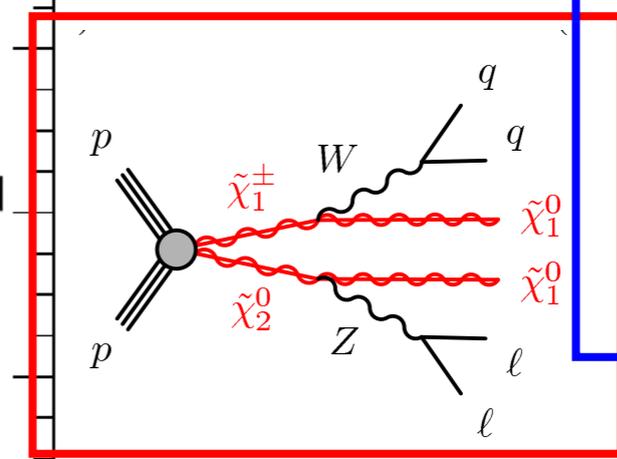
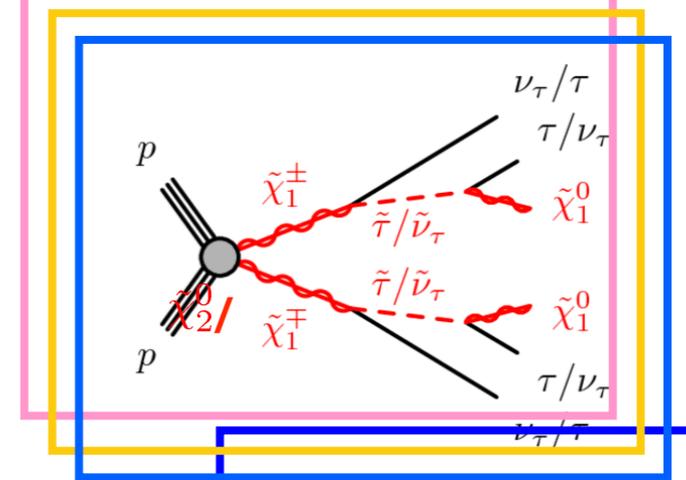
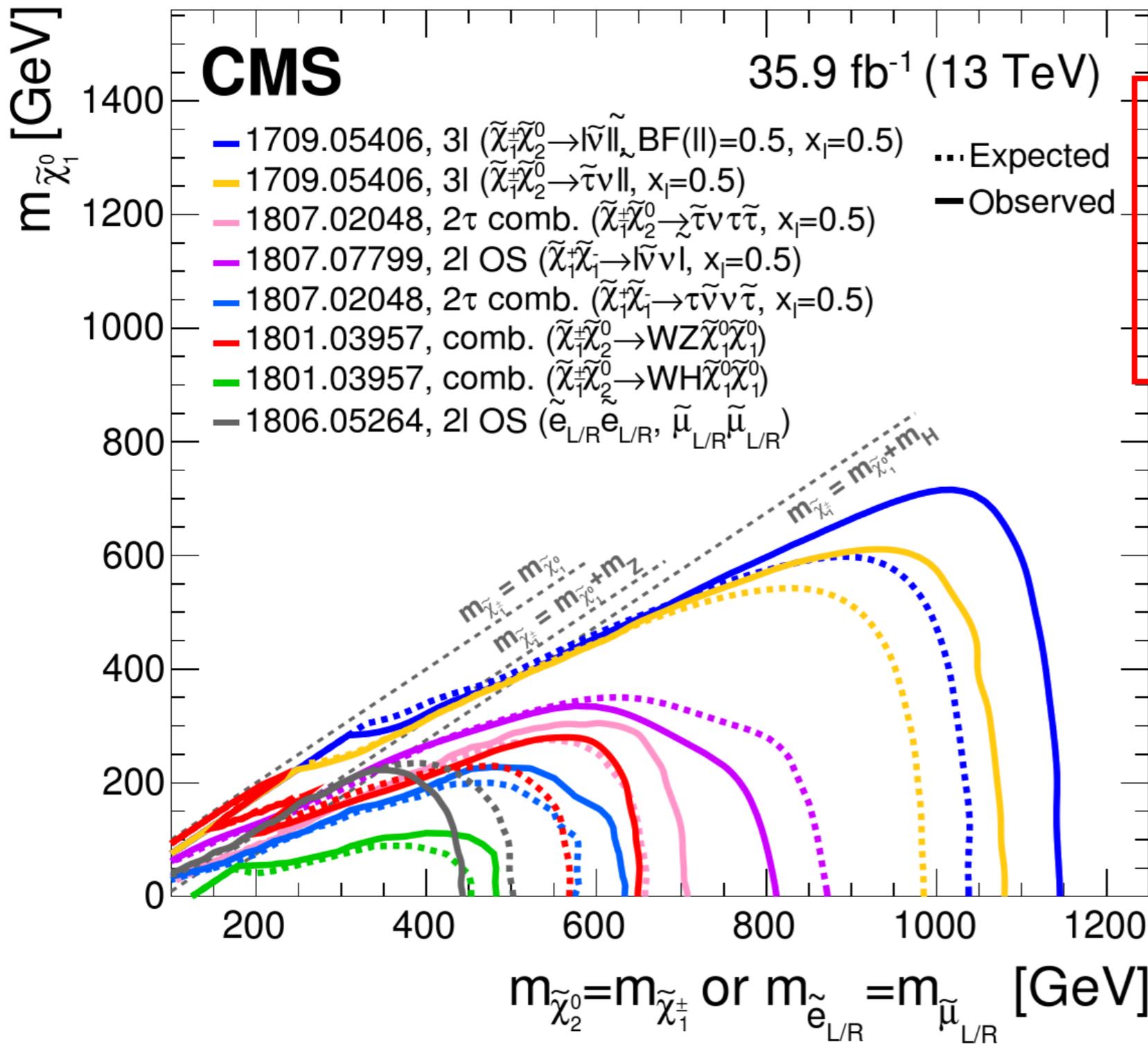


# LHCb real time model



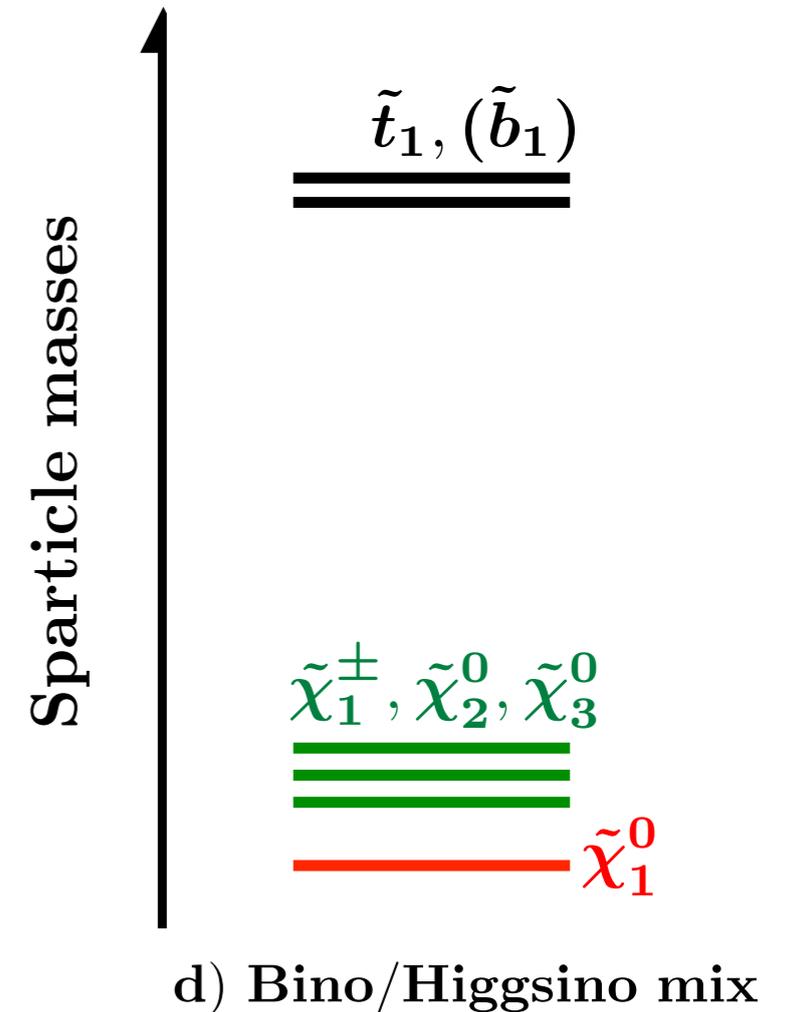
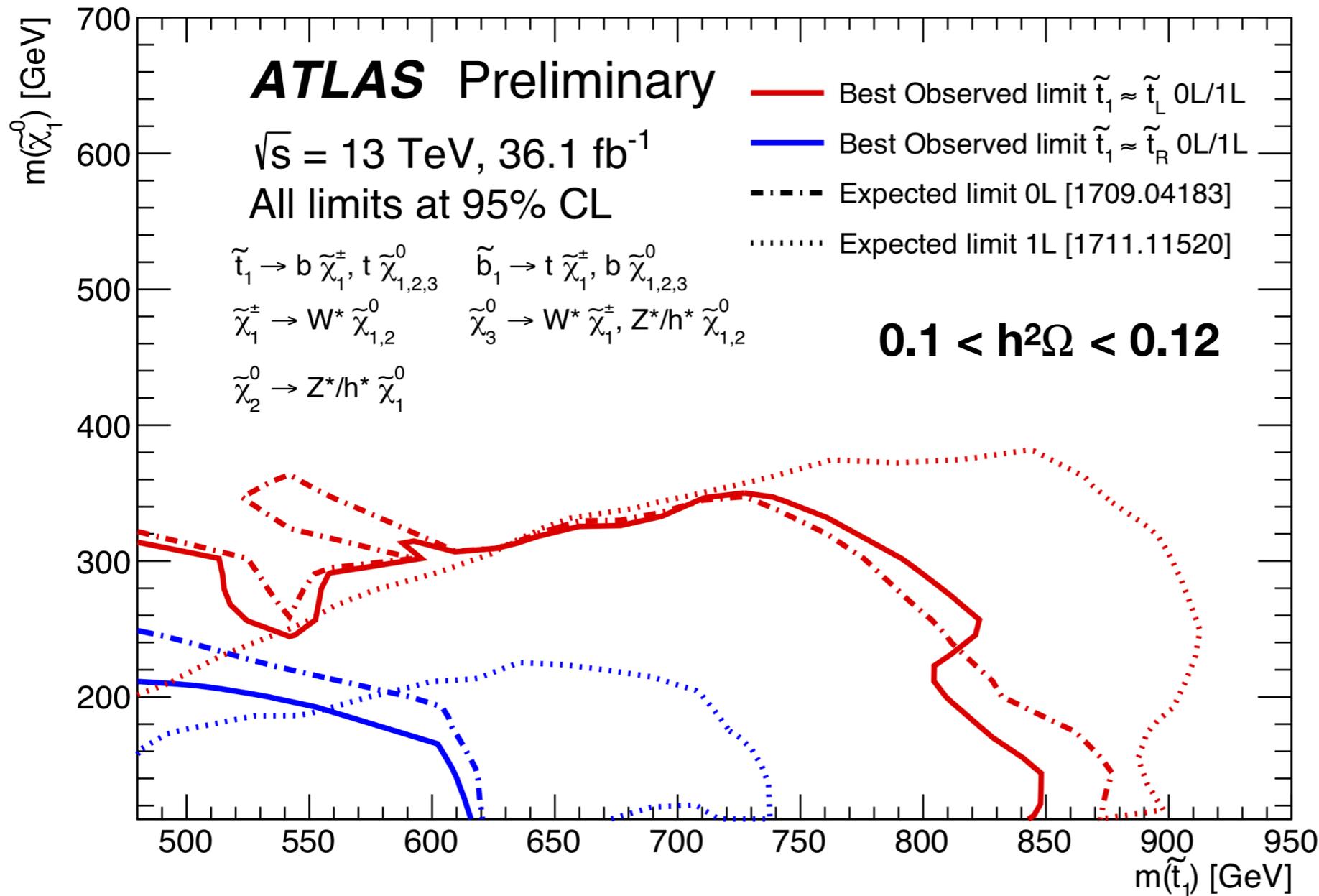
# SUSY: Electroweak

$$pp \rightarrow \tilde{\chi}_2^0 \tilde{\chi}_1^\pm, pp \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-, pp \rightarrow \tilde{l} \tilde{l} \quad \text{July 2018}$$

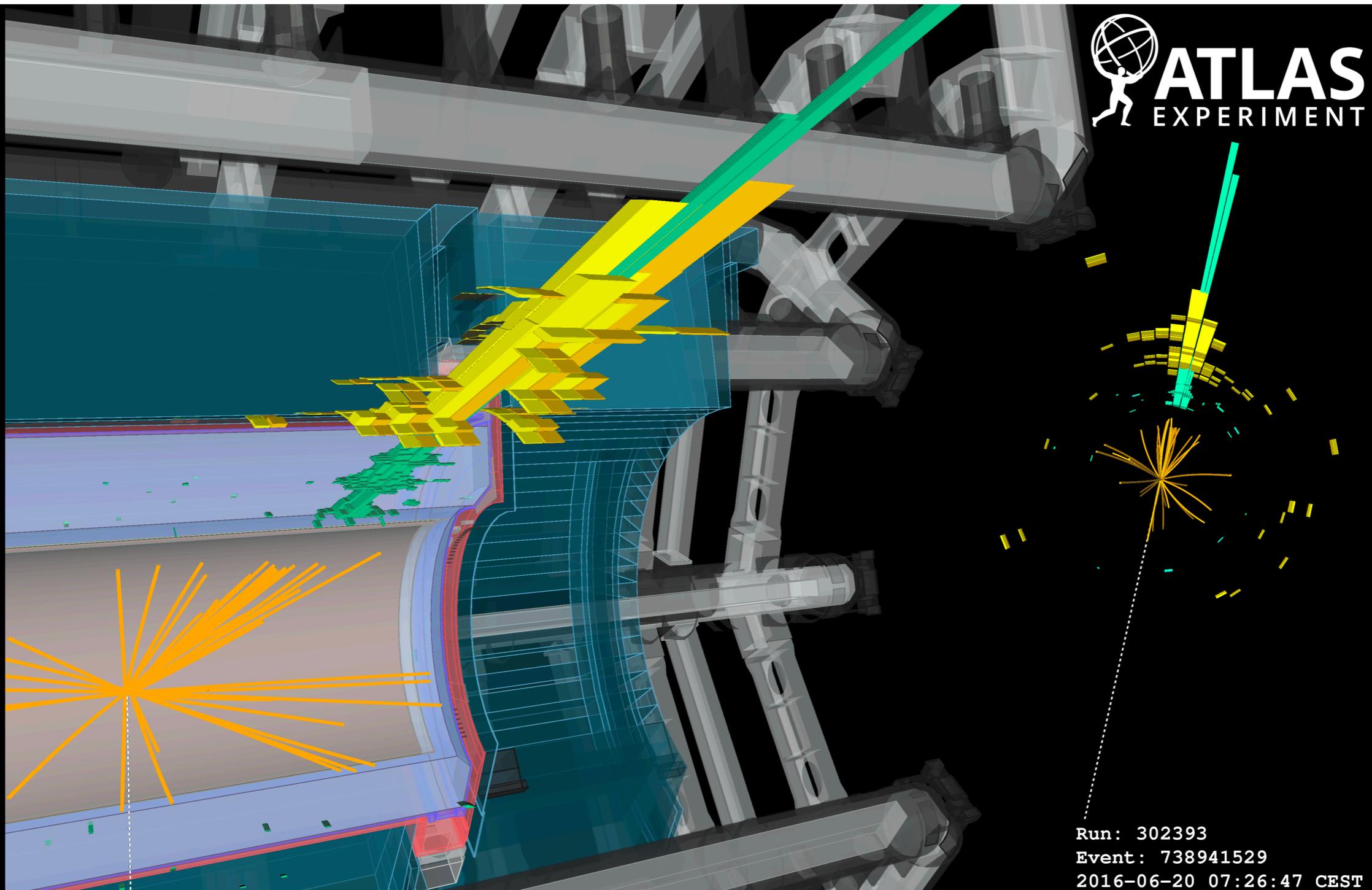


# SUSY: 3rd generation & DM

Bino/Higgsino Mix Model:  $\tilde{t}_1, \tilde{t}_2, \tilde{b}_1, \tilde{b}_2$  production,  $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 20-50$  GeV, March 2018

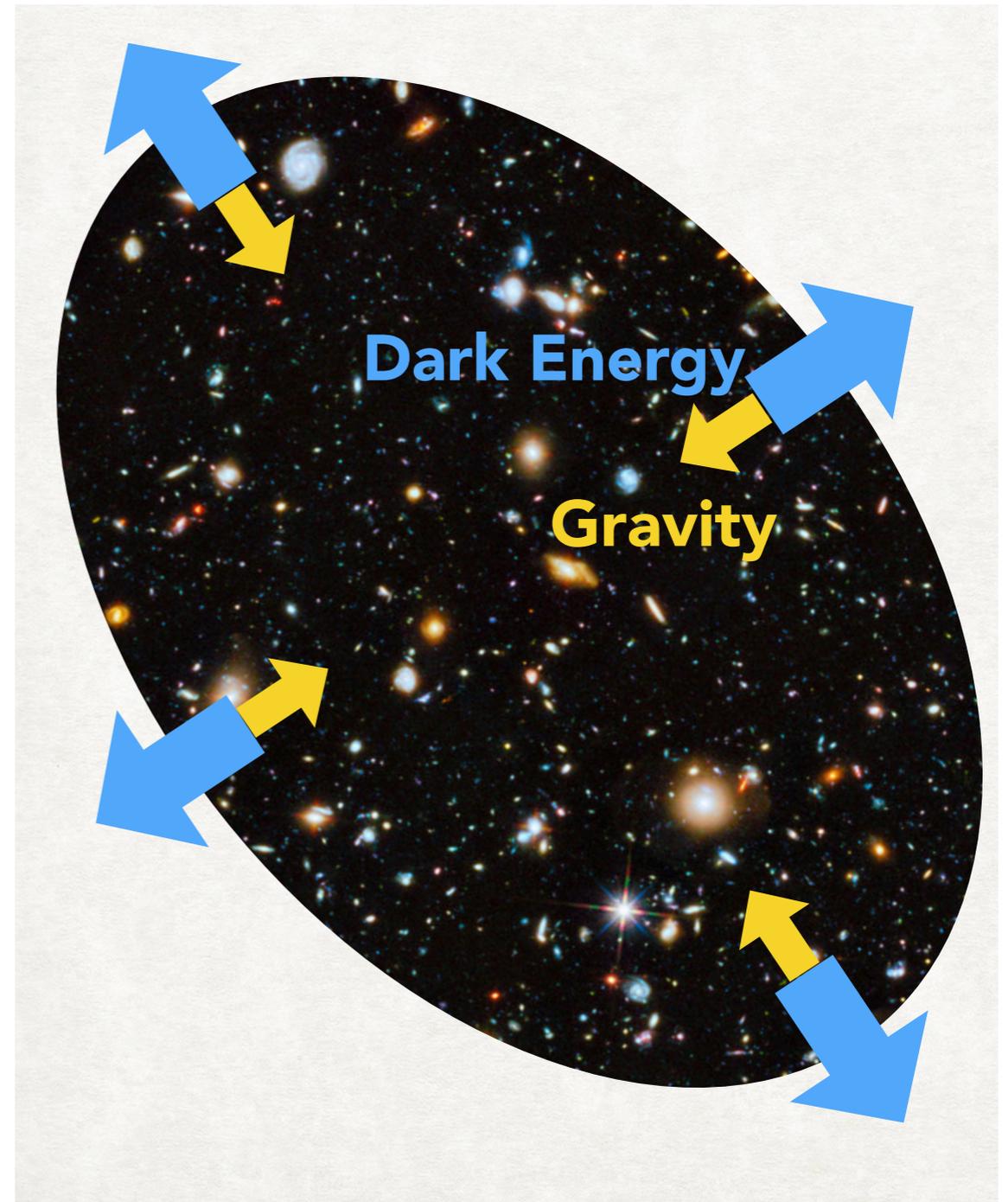


# Bonus: Dark Energy



# Dark Energy

- ★ Dark Energy = universe accelerated expansion
- ★ Big **unanswered question** in cosmology and particle physics
  - new particle or modified gravity?
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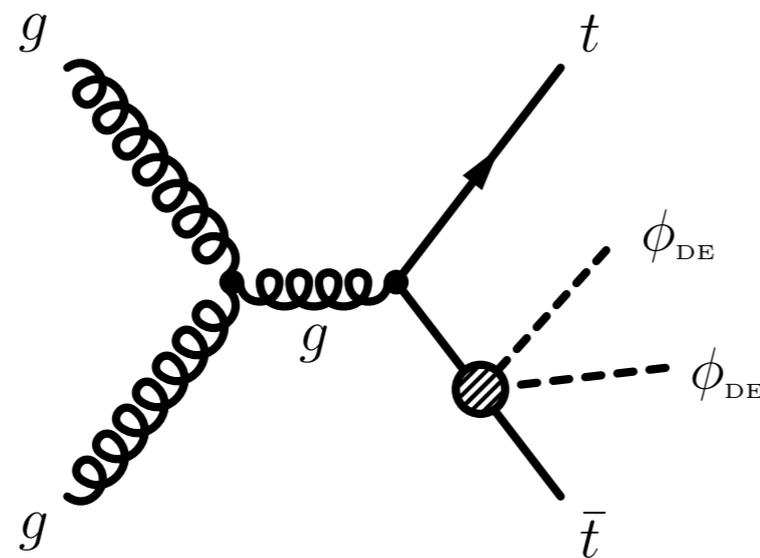


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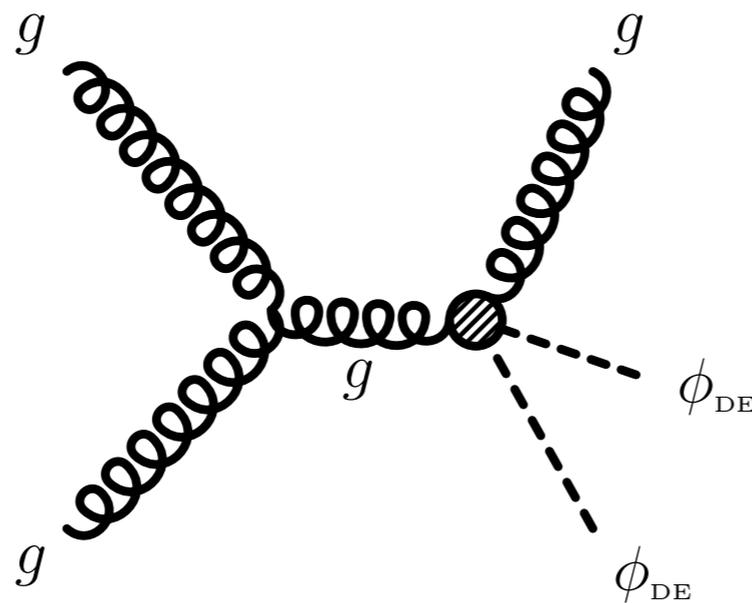
1 scalar field  $\phi_{\text{DE}}$  coupled to gravity

$$\mathcal{L}_1 = \frac{\partial_\mu \phi \partial^\mu \phi}{M_1^4} T_\nu^\nu,$$



→ tops +  $E_{\text{T}}^{\text{miss}}$

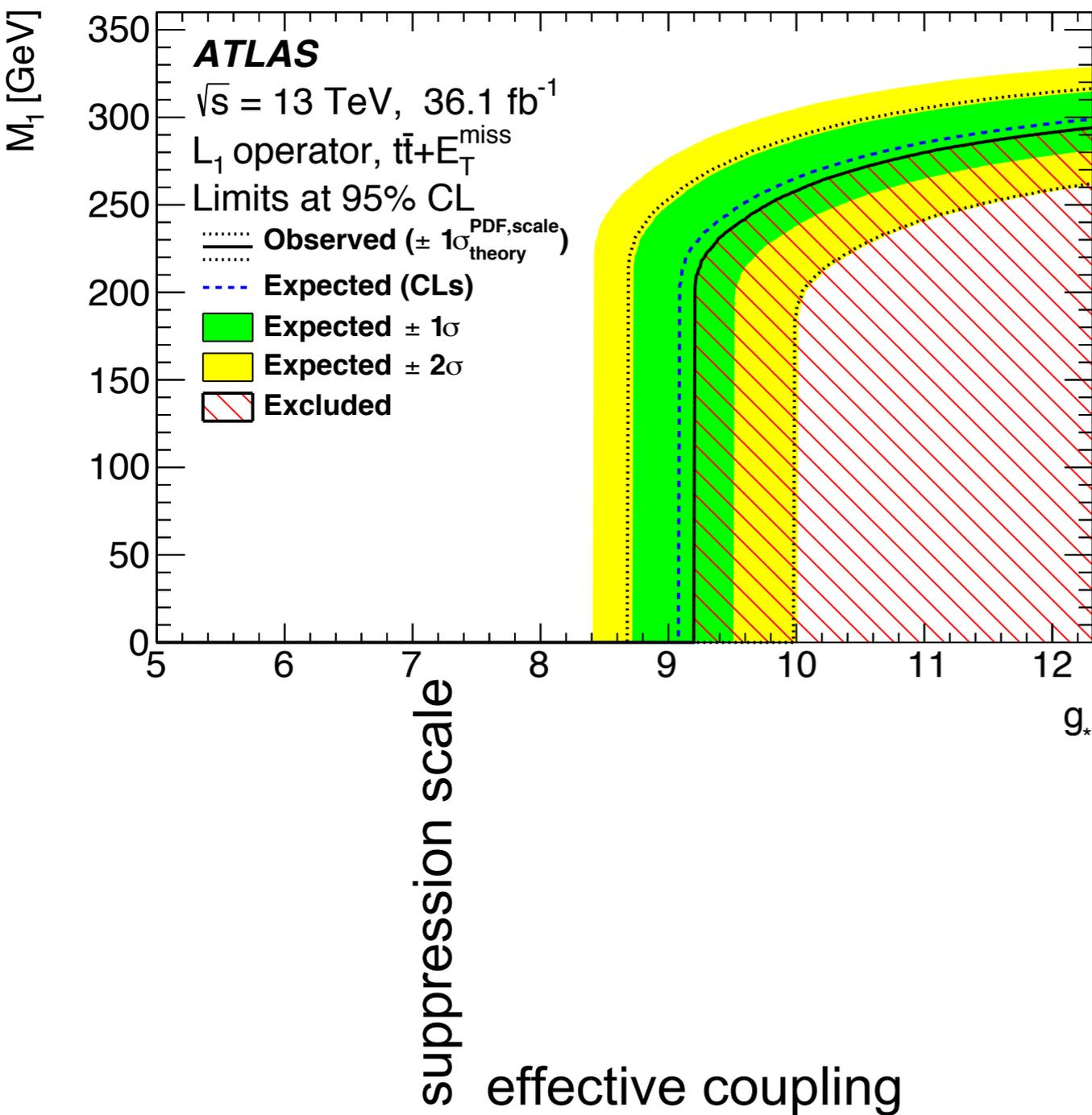
$$\mathcal{L}_2 = \frac{\partial_\mu \phi \partial_\nu \phi}{M_2^4} T^{\mu\nu},$$



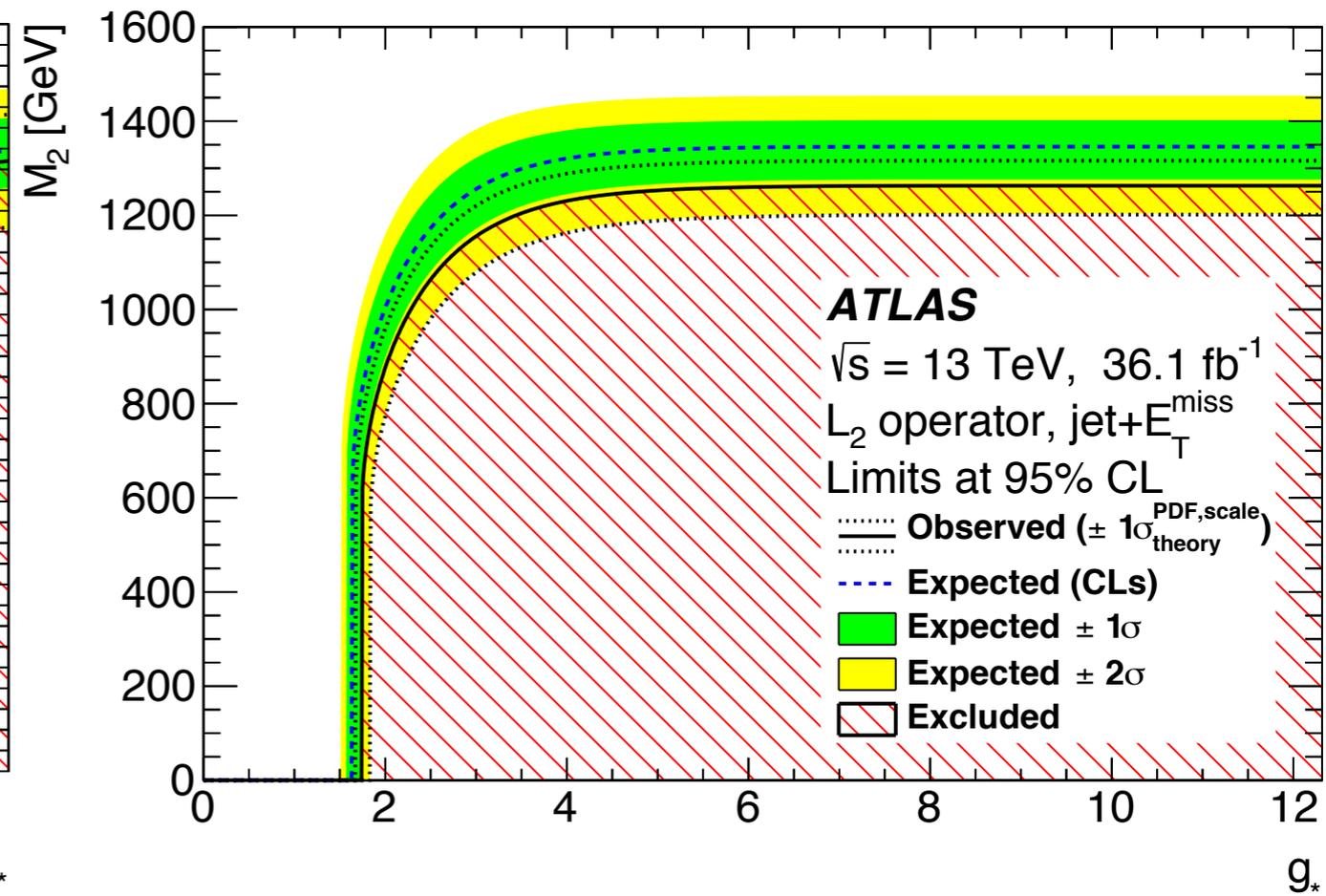
→ jet +  $E_{\text{T}}^{\text{miss}}$

# Results

tops +  $E_T^{\text{miss}}$

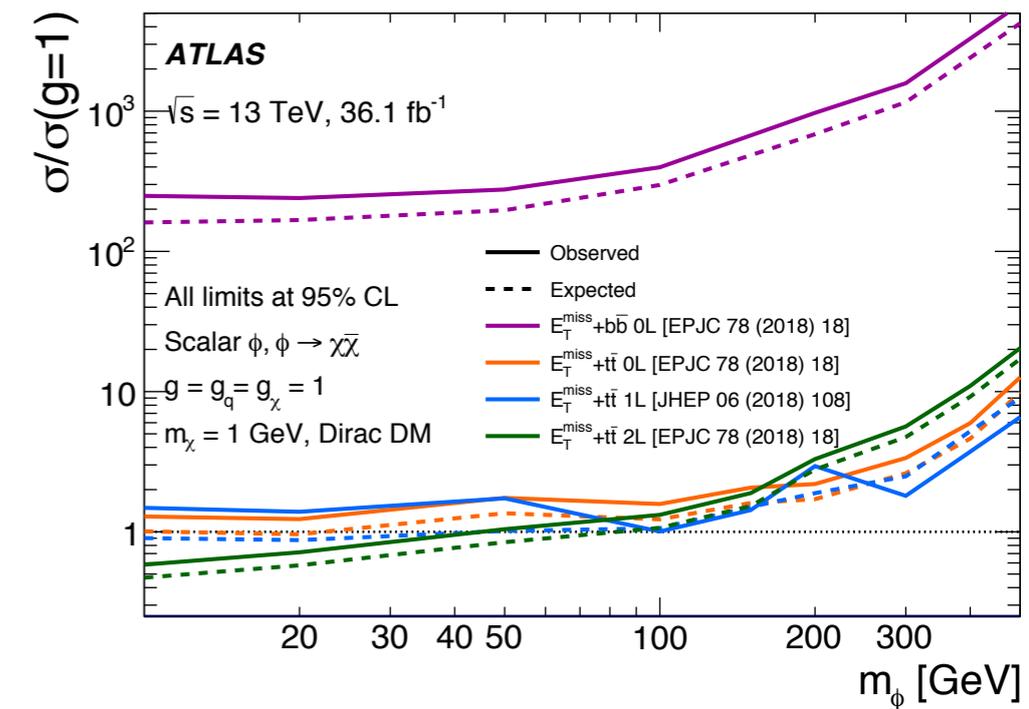
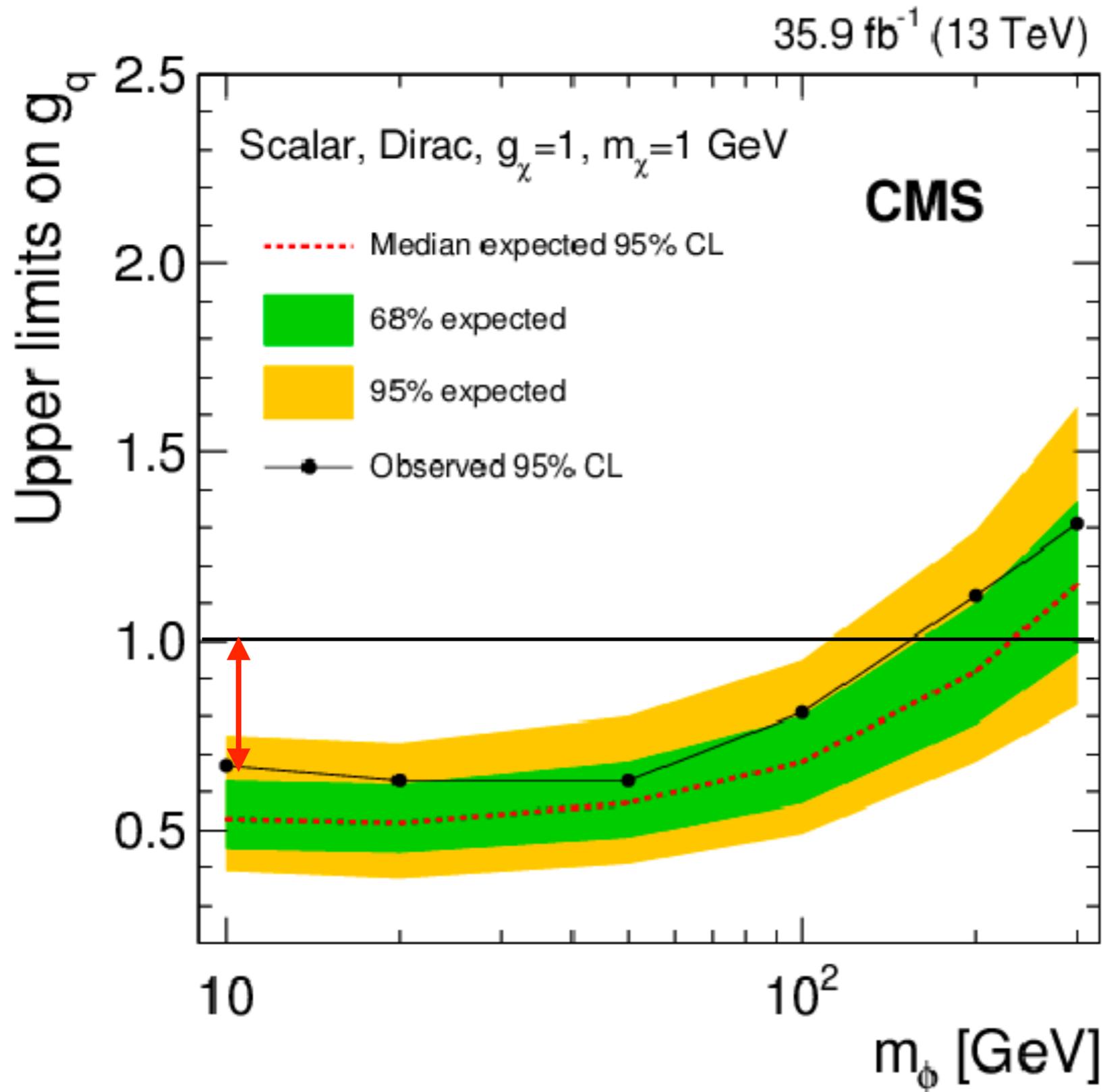


jet +  $E_T^{\text{miss}}$



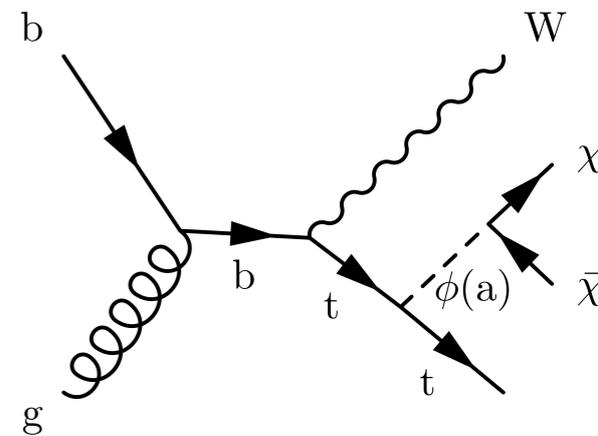
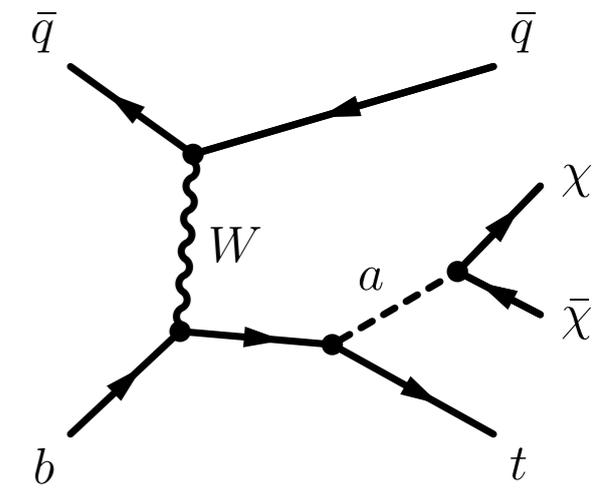
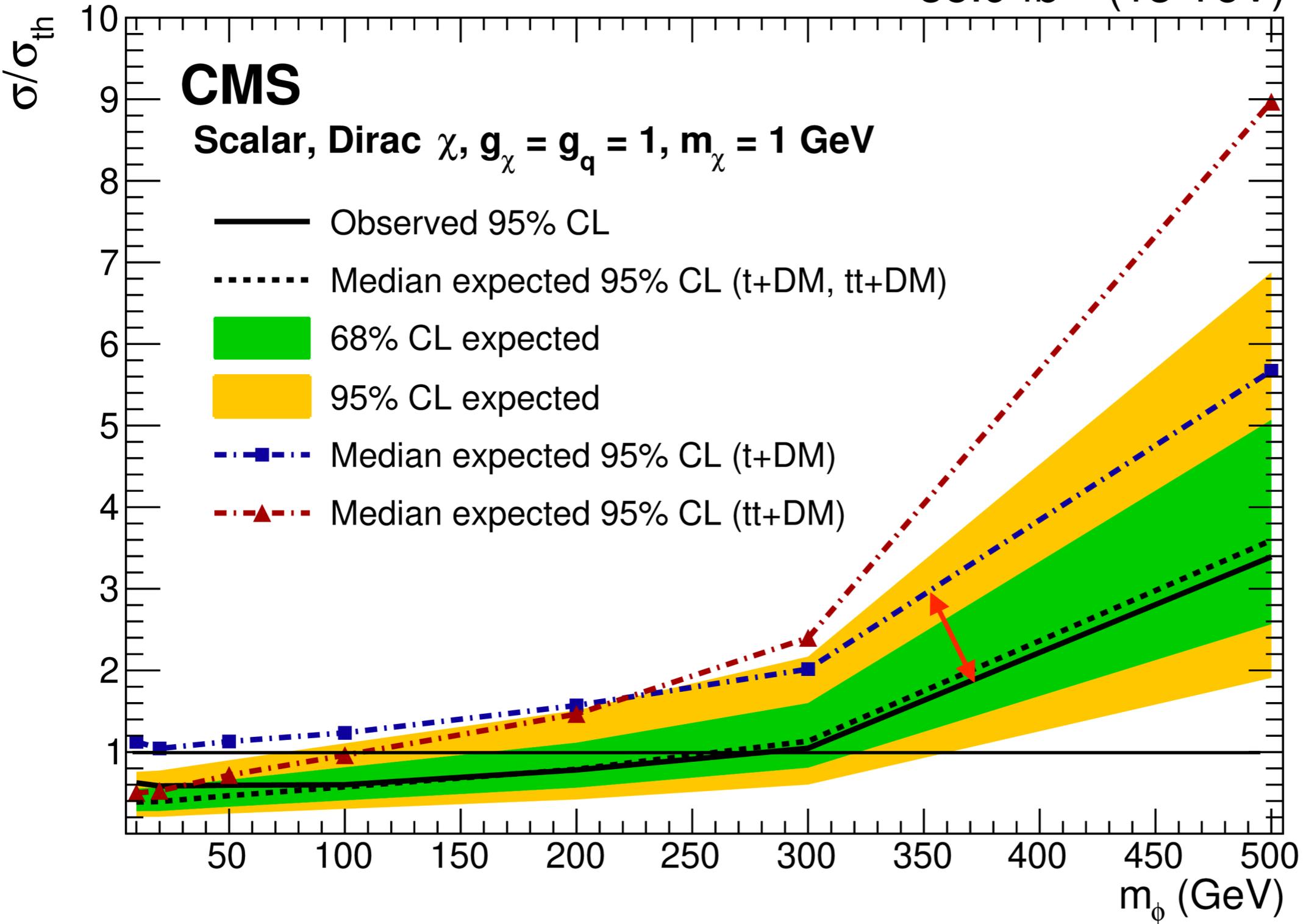
# CMS grand combination

[arXiv:1807.06522](https://arxiv.org/abs/1807.06522)



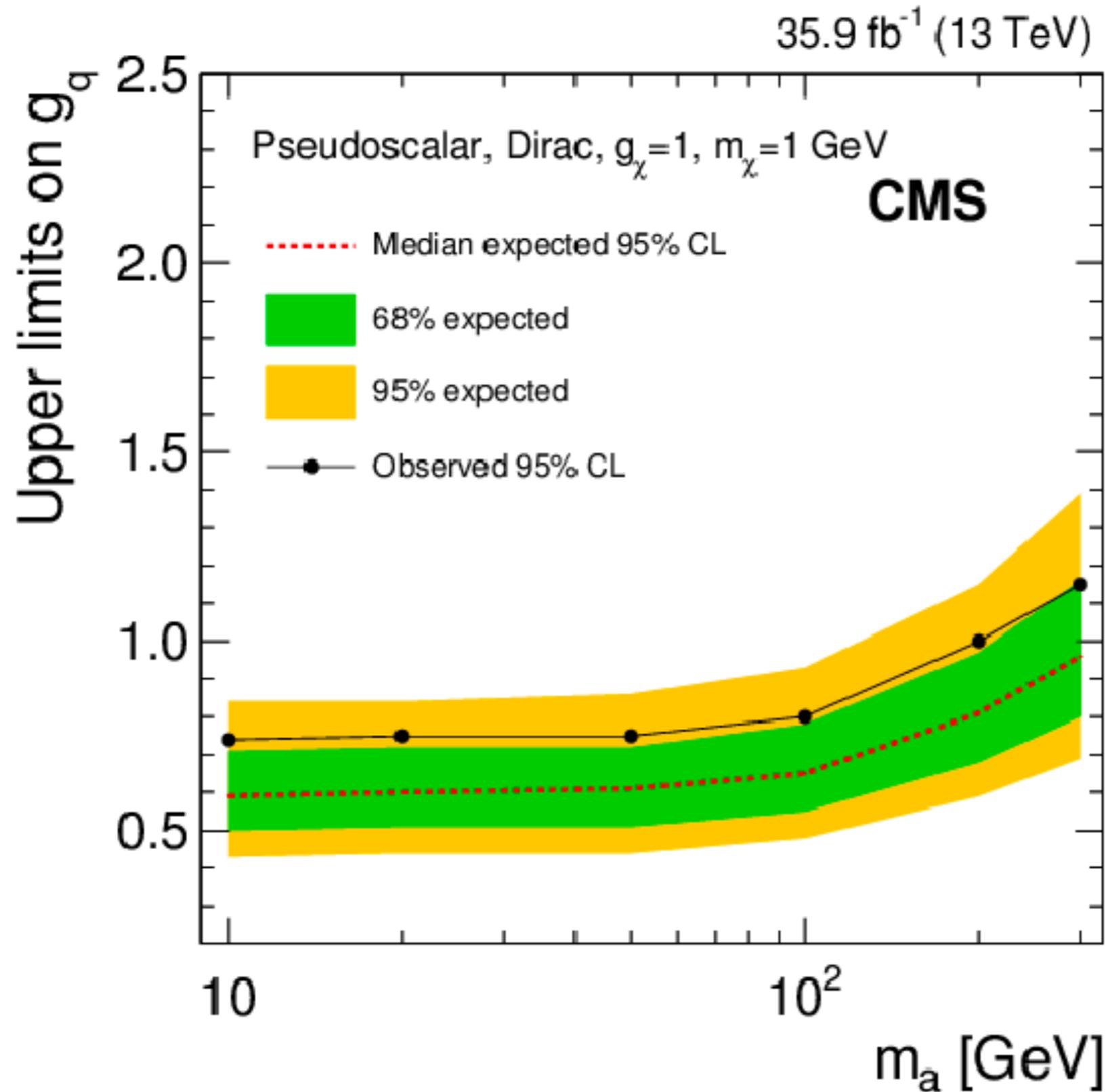
# Spin-0 with single top

35.9 fb<sup>-1</sup> (13 TeV)



# CMS combination Pseudo

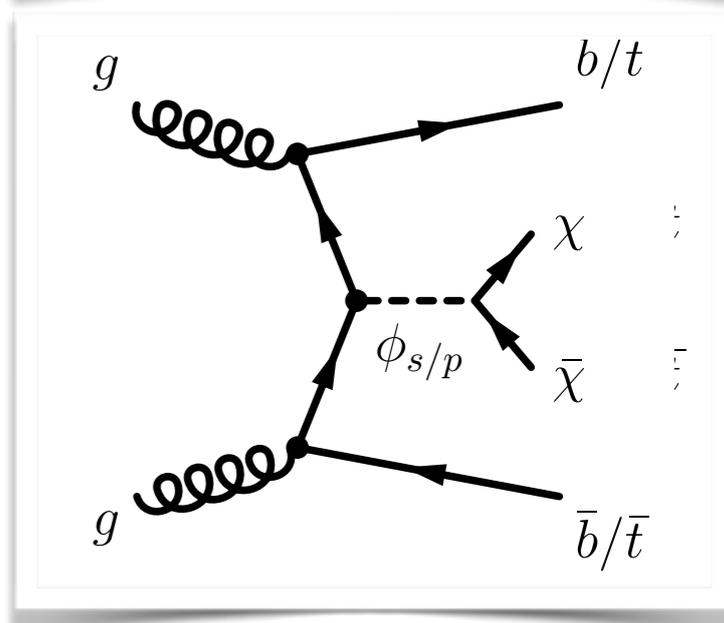
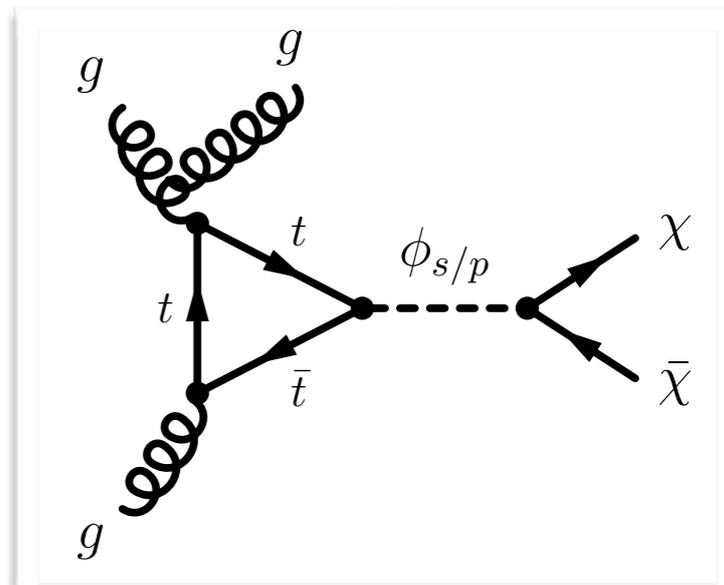
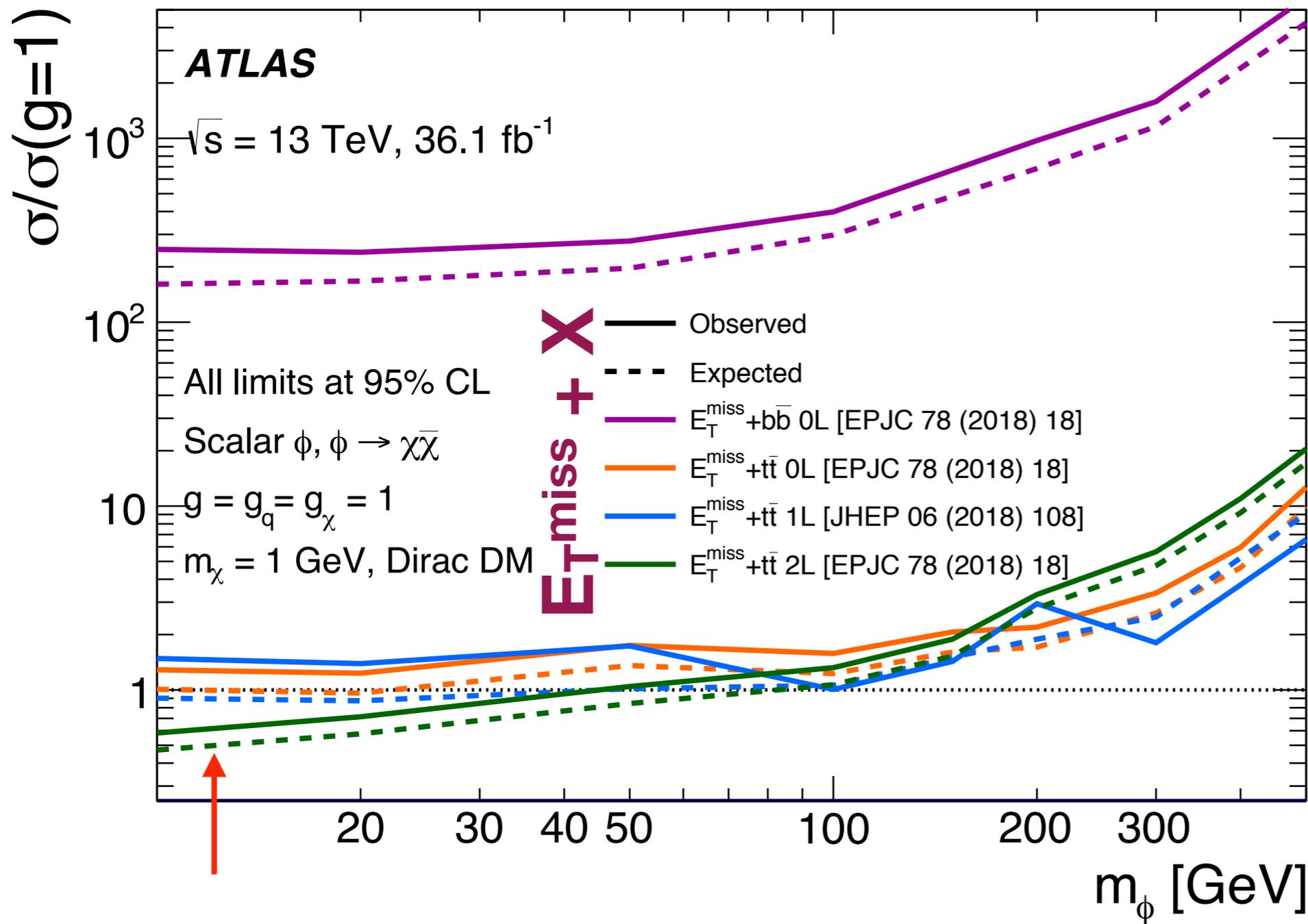
[arXiv:1807.06522](https://arxiv.org/abs/1807.06522)



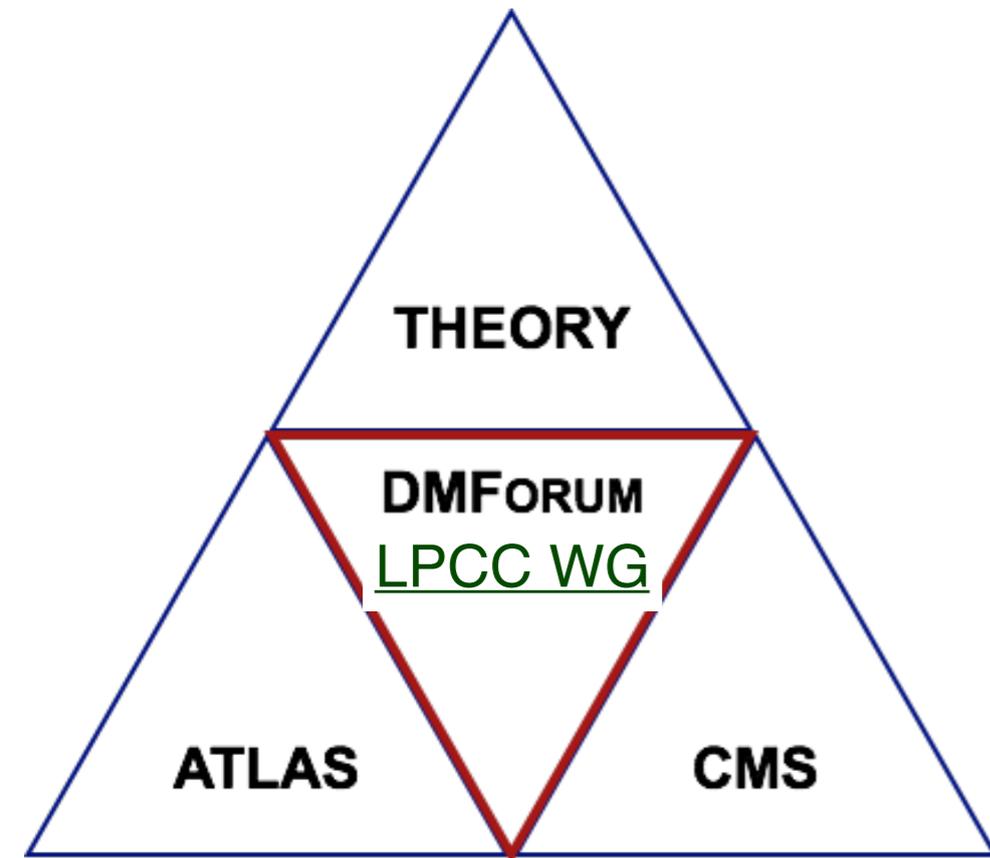
# Spin-0 mediators

$$\mathcal{L} \sim \sum_f ig_v \frac{y_f}{\sqrt{2}} A \bar{f} \gamma^5 f$$

Needed to easily fulfil Flavour Constraints (MFV)



# An inter-community achievement



## Simplified Models for Dark Matter Searches at the LHC

Jalal Abdallah, Henrique Araujo, Alexandre Arbey, Adi Ashkenazi, Alexander Belyaev, Joshua Berger, Celine Boehm,

[Phys. Dark Univ. 9-10 \(2015\) 8-23](#)

## Dark Matter Benchmark Models for Early LHC Run-2 Searches: Report of the ATLAS/CMS Dark Matter Forum

Daniel Abercrombie, Nural Akchurin, Ece Akilli, Juan Alcaraz Maestre, Brandon Allen, Barbara Alvarez Gonzalez, Jeremy

[arXiv:1507.00966](#)

## Recommendations on presenting LHC searches for missing transverse energy signals using simplified $s$ -channel models of dark matter

Antonio Boveia, Oliver Buchmueller, Giorgio Busoni, Francesco D'Eramo, Albert De Roeck, Andrea De Simone, Caterina

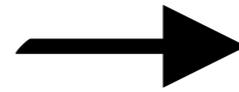
[arXiv:1603.04156](#)

★ Simplified Models are the Run II paradigm:

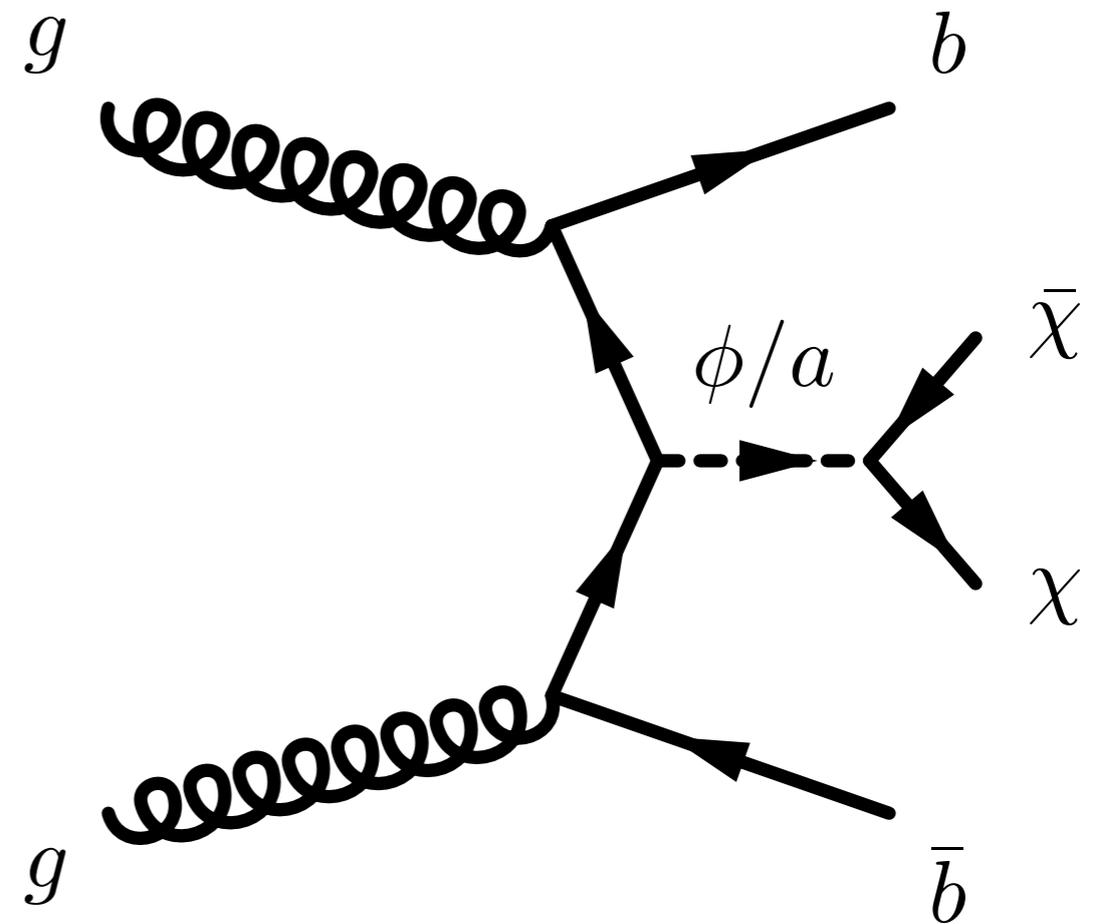
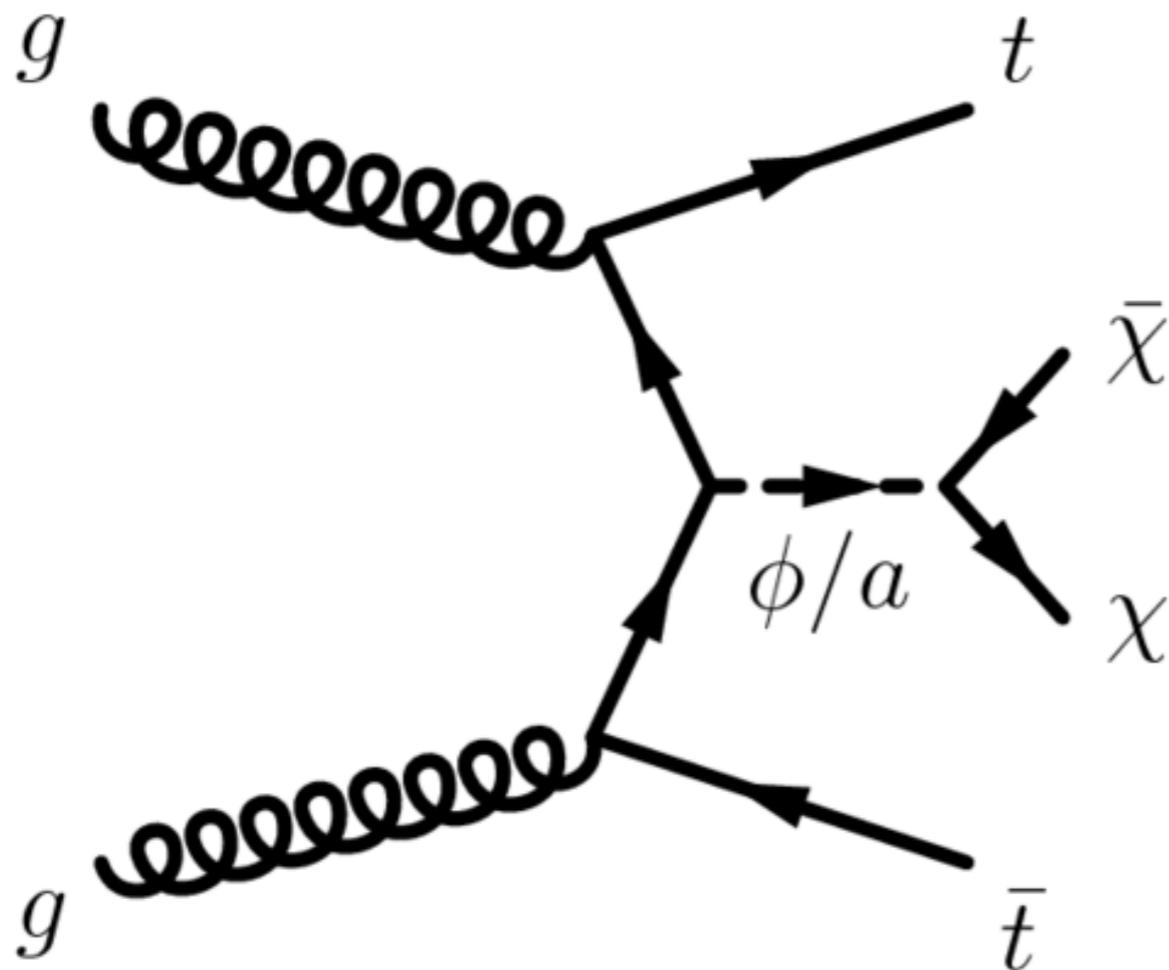
- theoretically self consistent
- minimal and motivated assumptions
- good phenomenology proxies

# Exploring the dark sector with heavy quarks

$$\mathcal{L} \sim \sum_f i g_v \frac{y_f}{\sqrt{2}} A \bar{f} \gamma^5 f$$

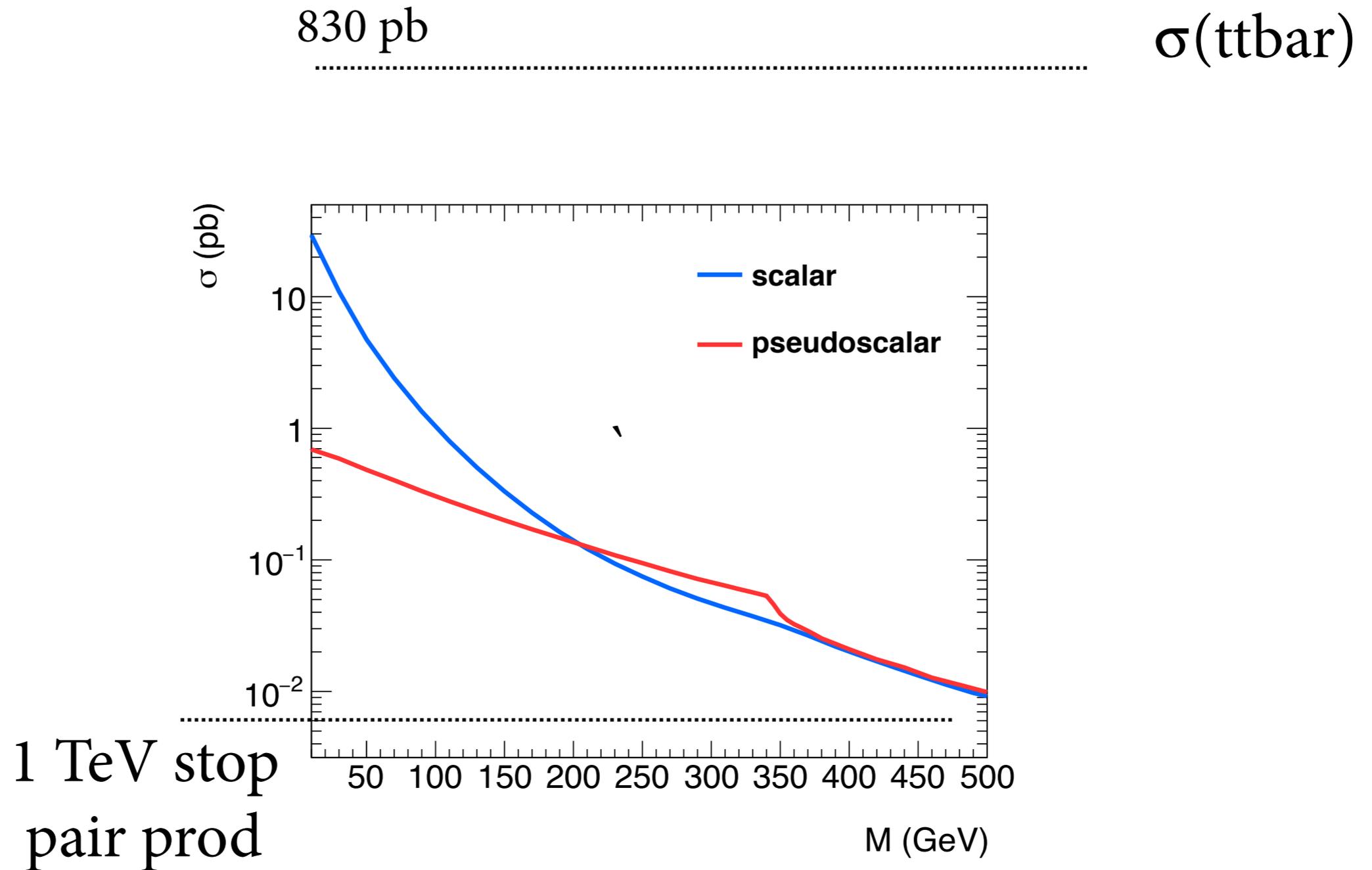


Enhanced cross-section for  
tops and bottoms



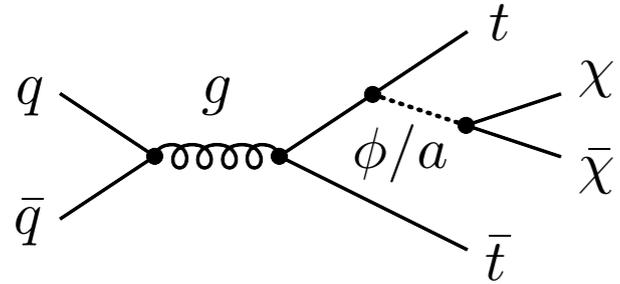
arXiv:1710.11412 and ATLAS-CONF-2017-037

# Understanding the signal



[Haisch,PP,Polesello 2017]

# Understanding the signal

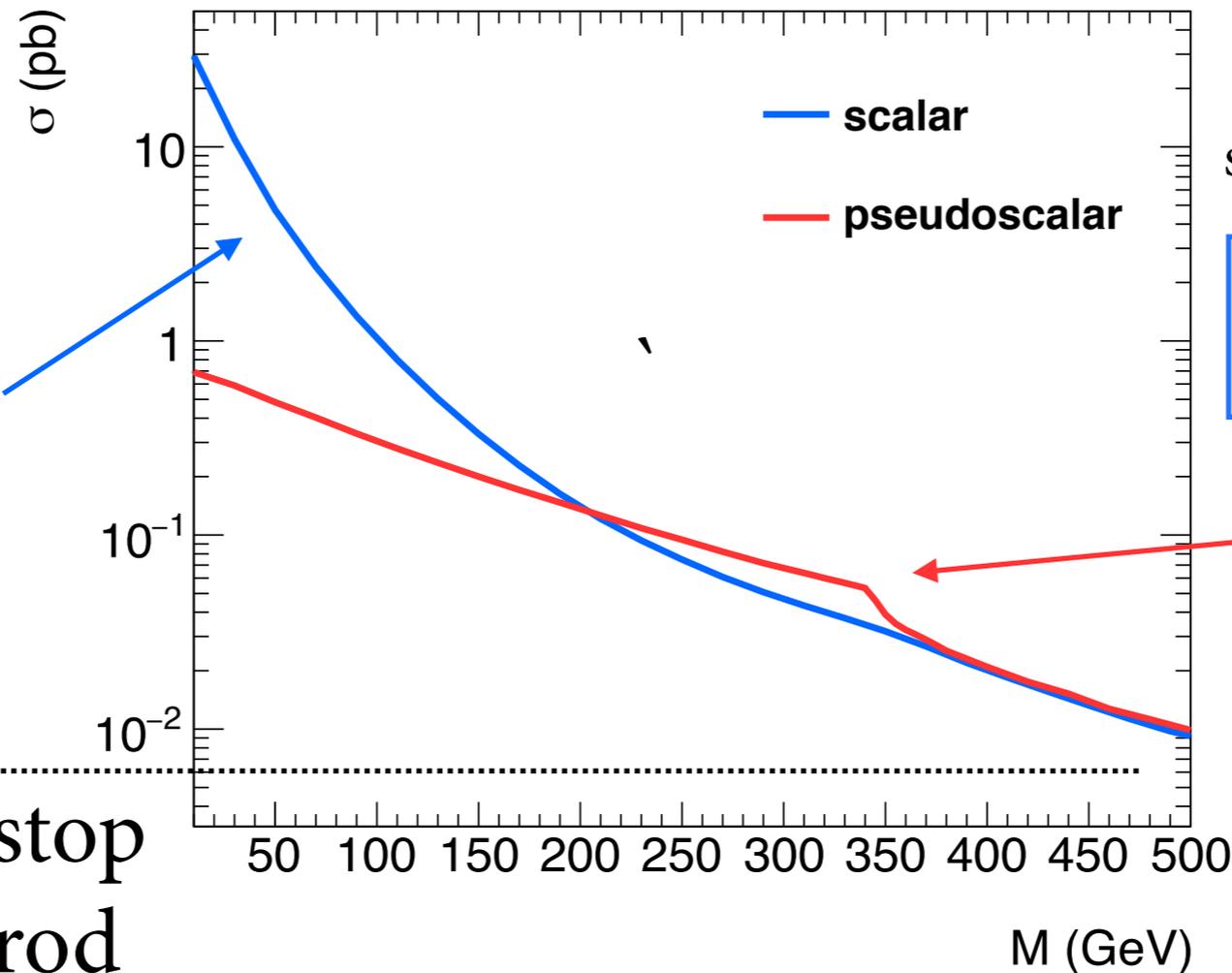


830 pb

$\sigma(tt\bar{b}ar)$

$$\underline{f_{t \rightarrow \phi}(x) \sim 1/x}$$

soft-enhancement  
term



spin and color-averaged ME

$$\overline{\sum} |\mathcal{M}(t\bar{t} \rightarrow \phi)|^2 = \frac{g_t^2 s}{12} \beta^2$$

$$\beta = \sqrt{1 - 4m_t^2/s}$$

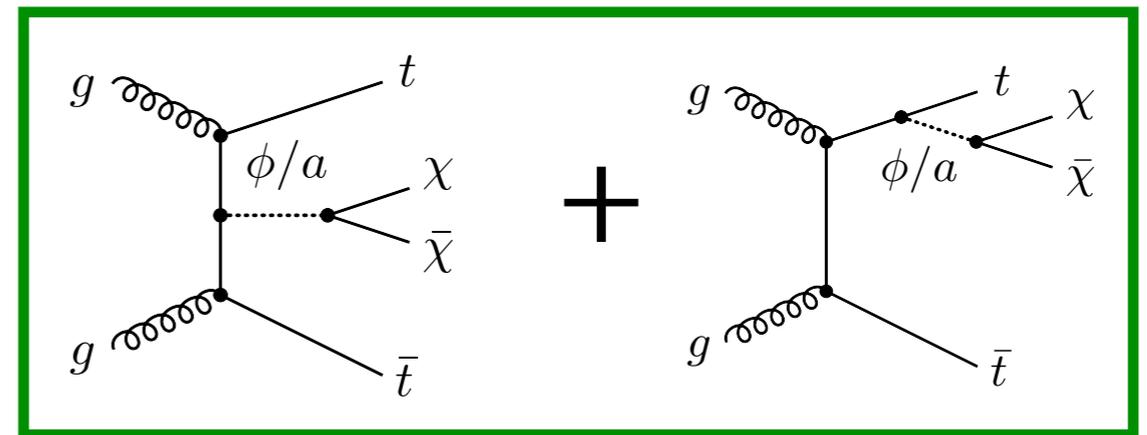
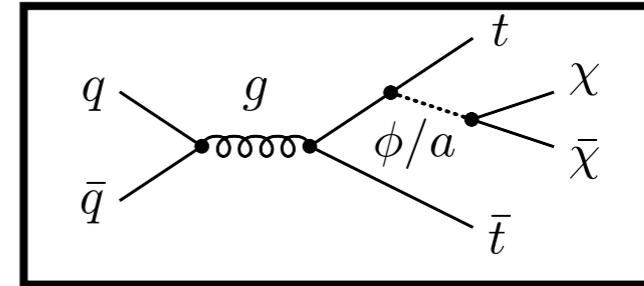
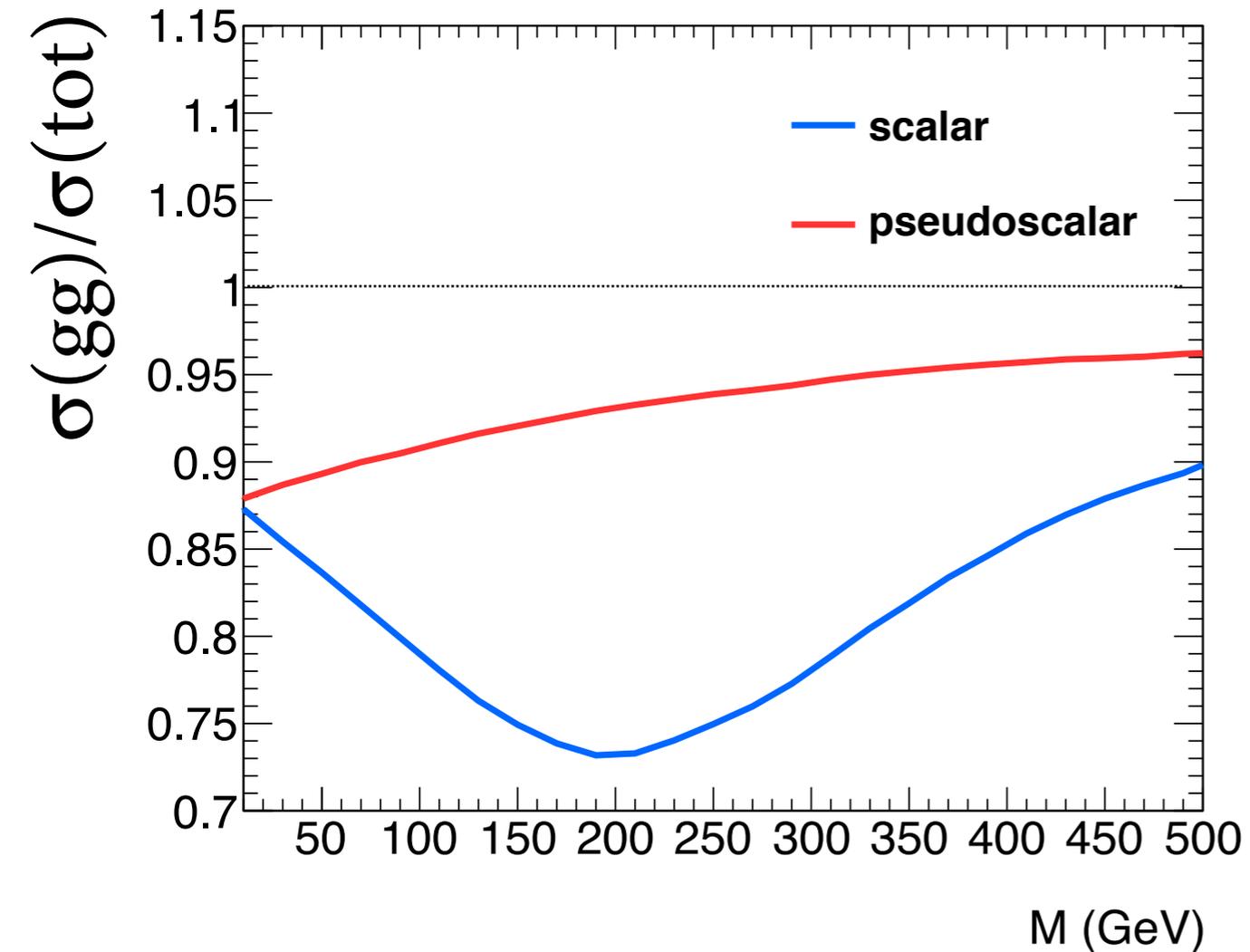
top-threshold  
suppression

$$\phi/a \rightarrow t\bar{t}$$

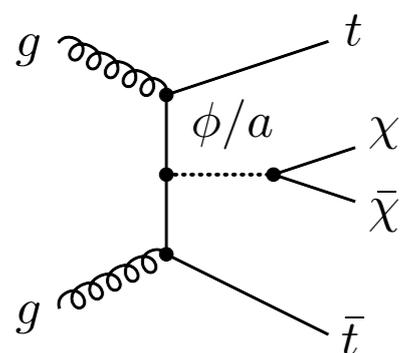
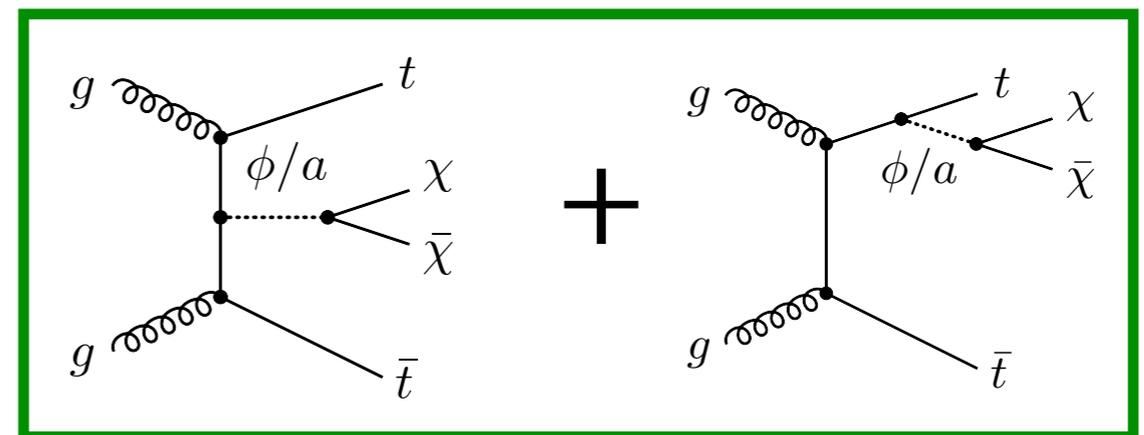
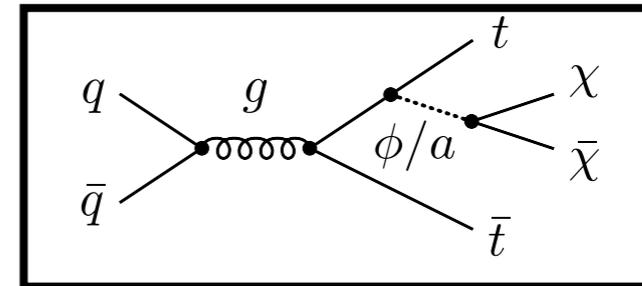
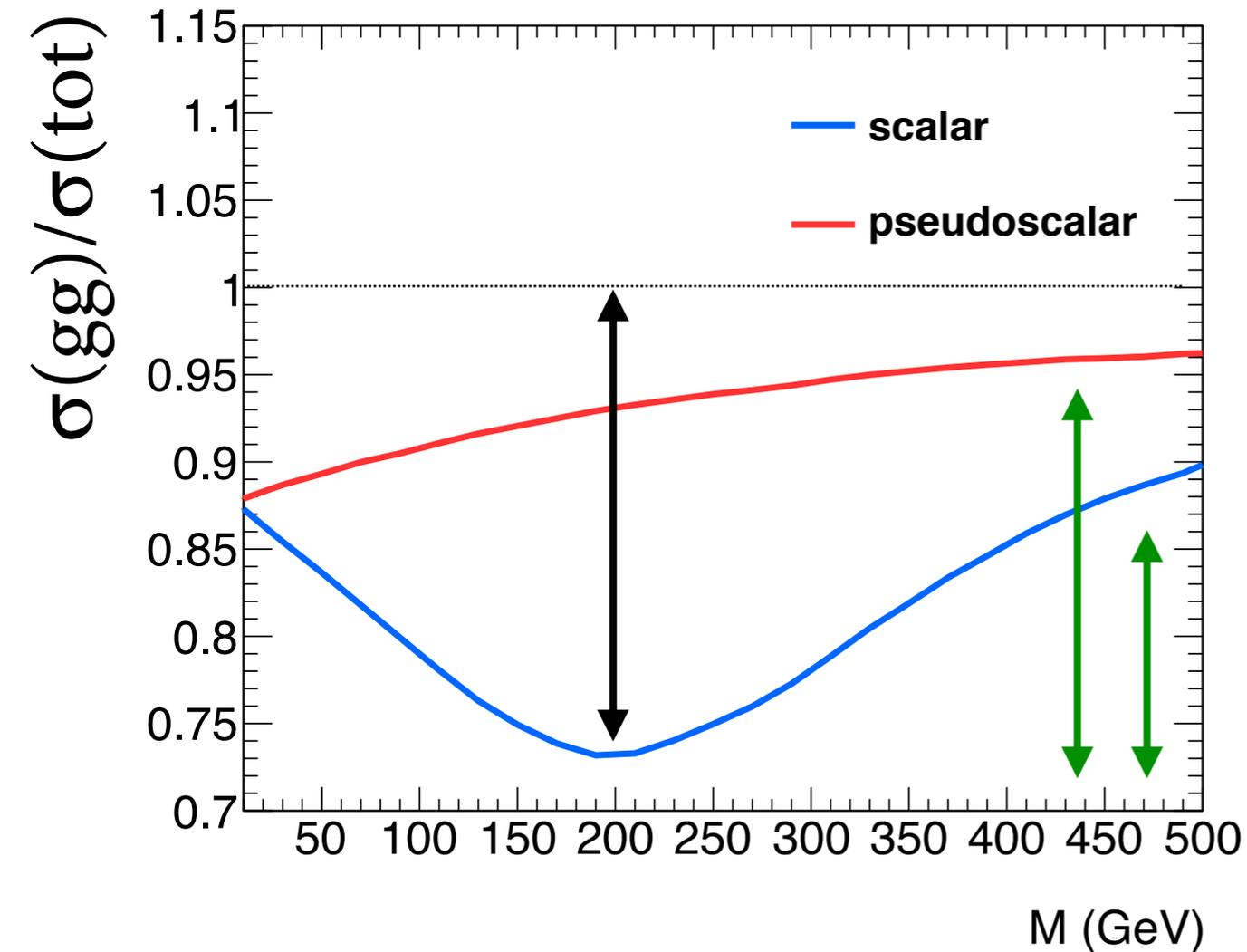
1 TeV stop  
pair prod

[Haisch,PP,Polesello 2017]

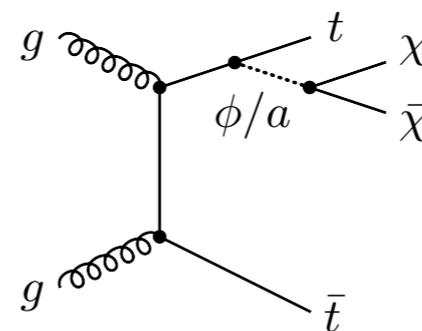
# Understanding the signal



# Understanding the signal



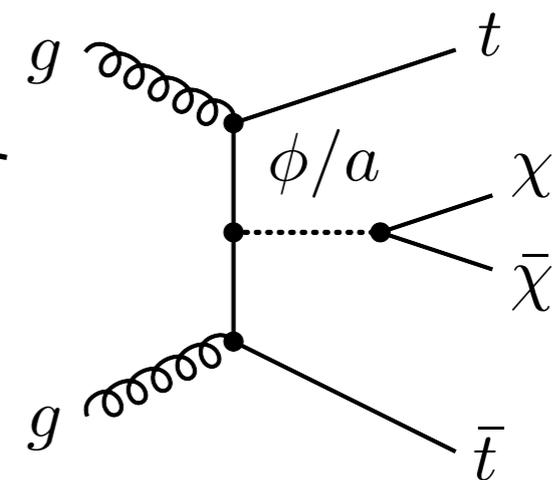
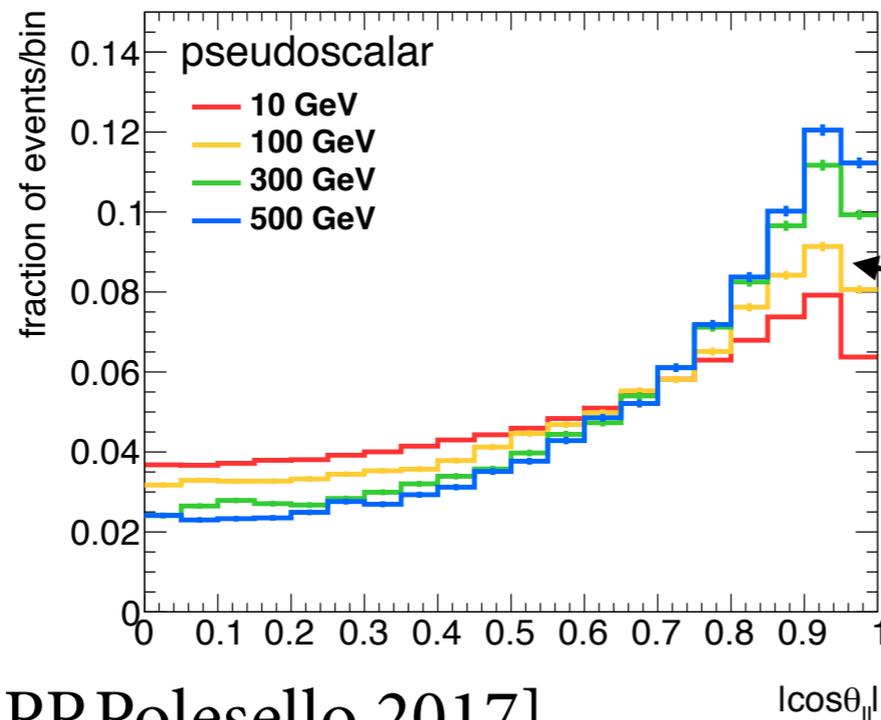
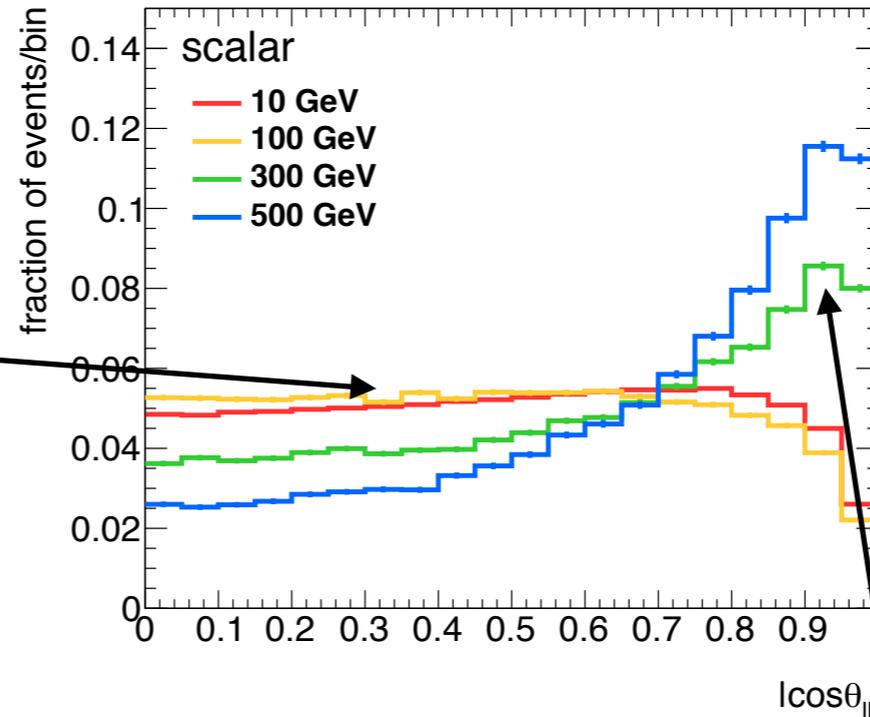
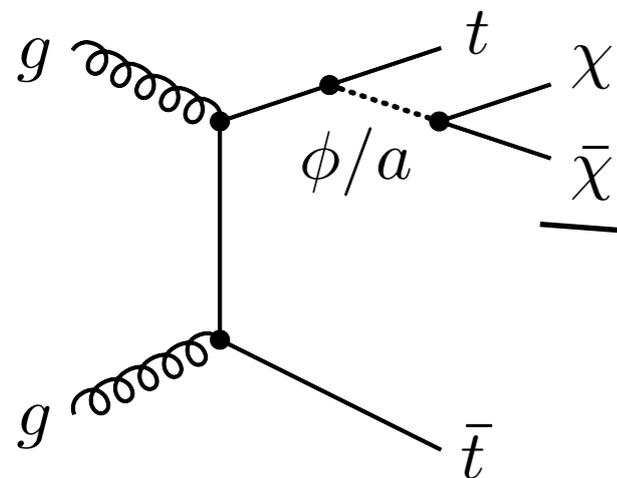
: dominant at  
high mass and  
low-mass  $a$



: dominant at  
low-mass  $\phi$

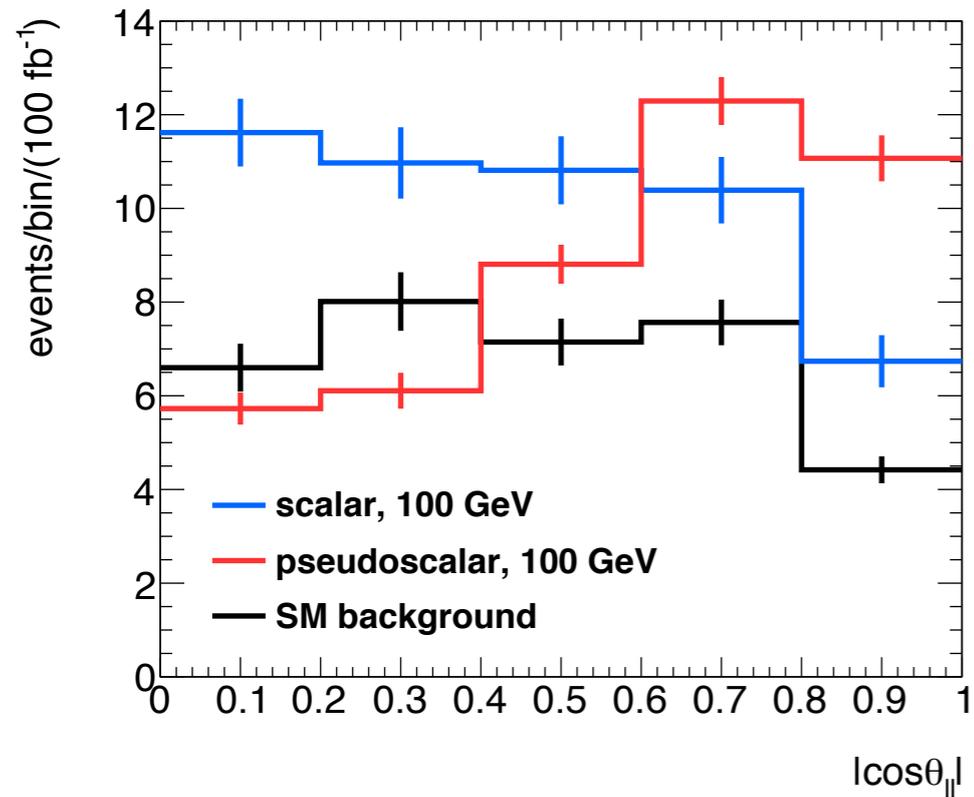
# Future perspectives on the results

$$\cos \theta_{t\bar{t}} \equiv \tanh(\Delta\eta_{t\bar{t}}/2)$$

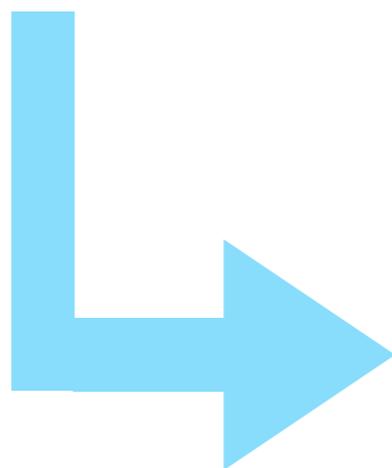


[Haisch,PP,Polesello 2017]

# Run 3 and HL-LHC outlook

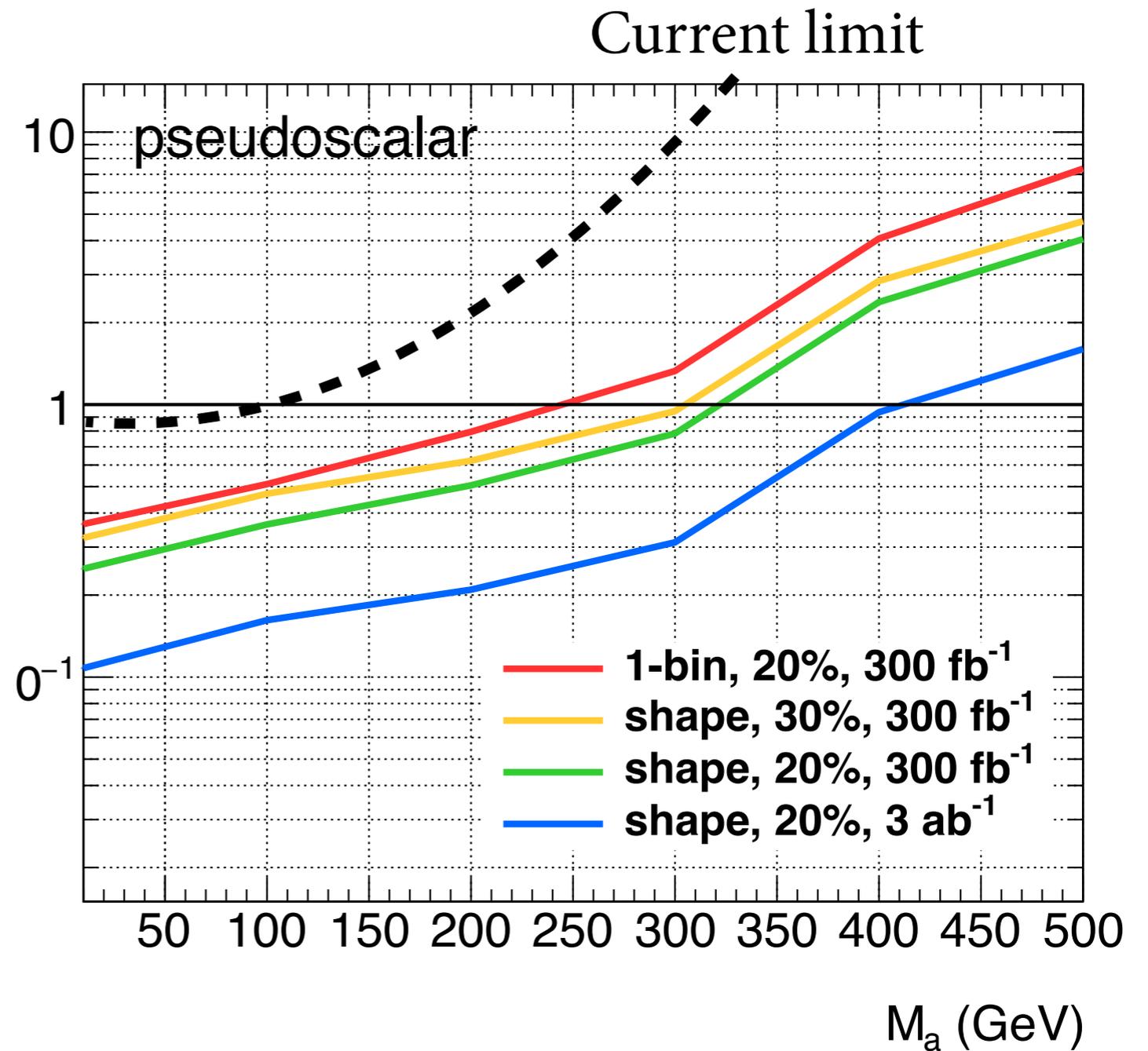
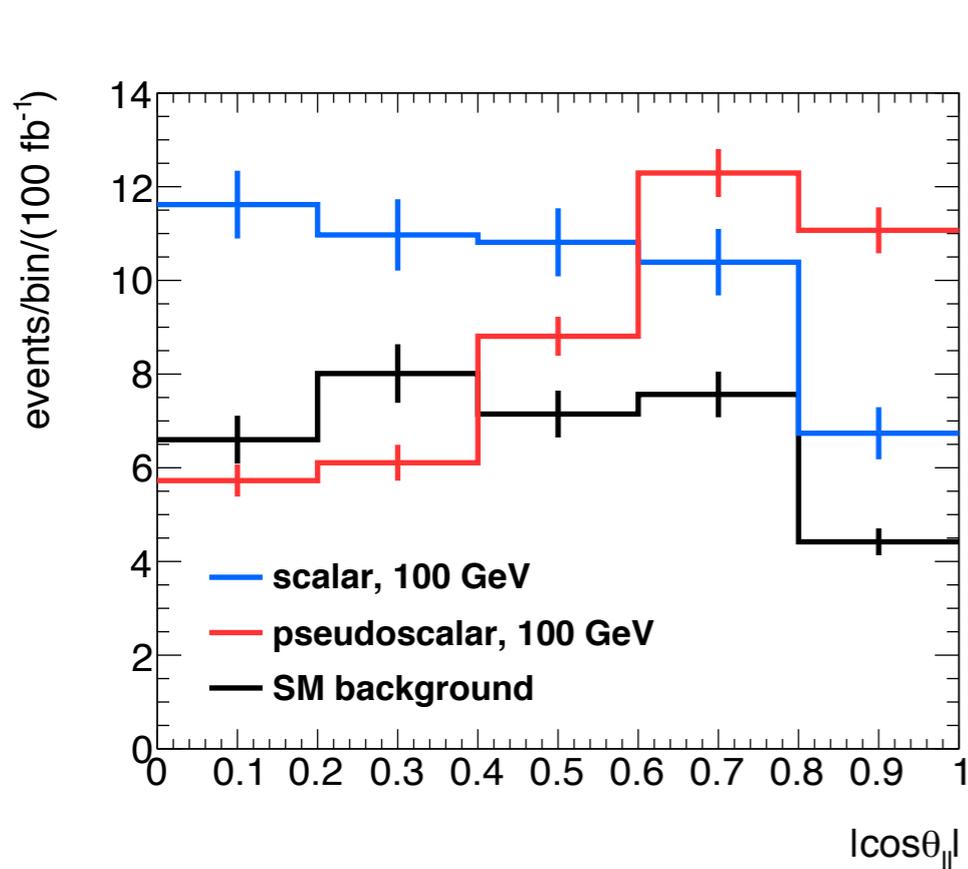


Multi-bin  
fit

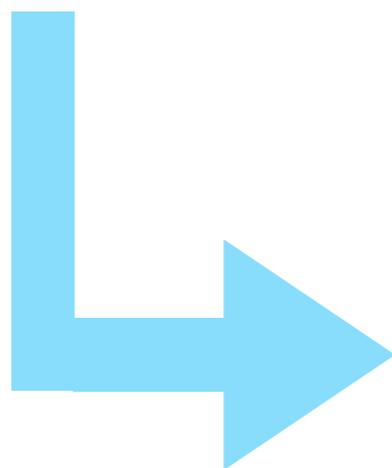


[Haisch,PP,Polesello 2017]

# Run 3 and HL-LHC outlook

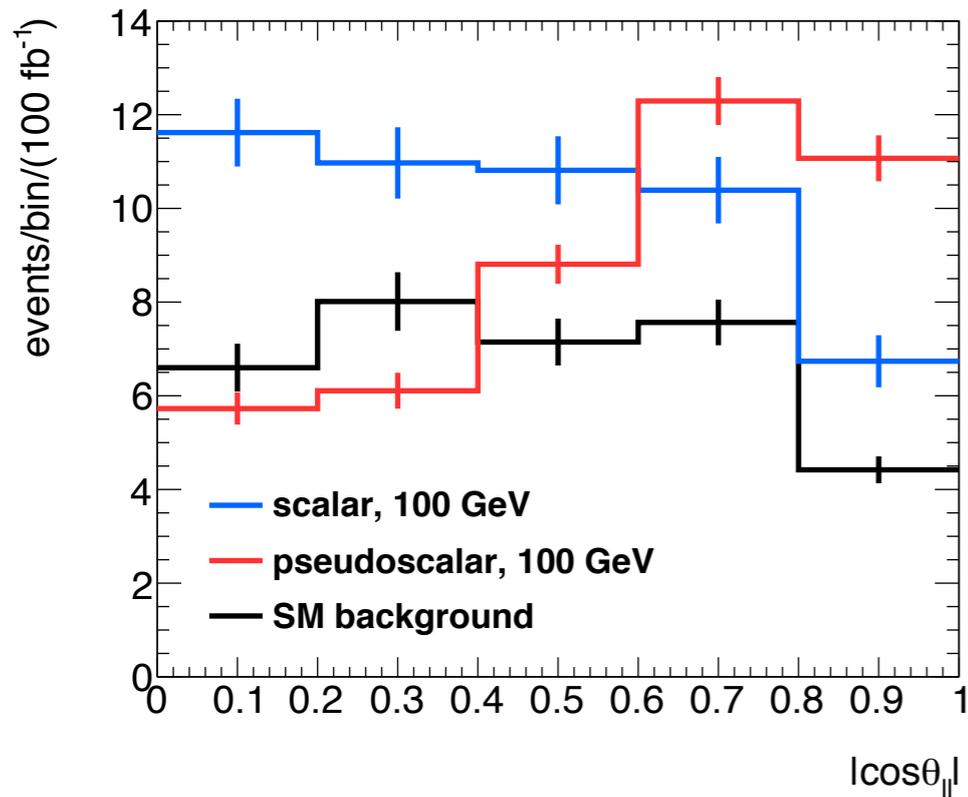


Multi-bin  
fit

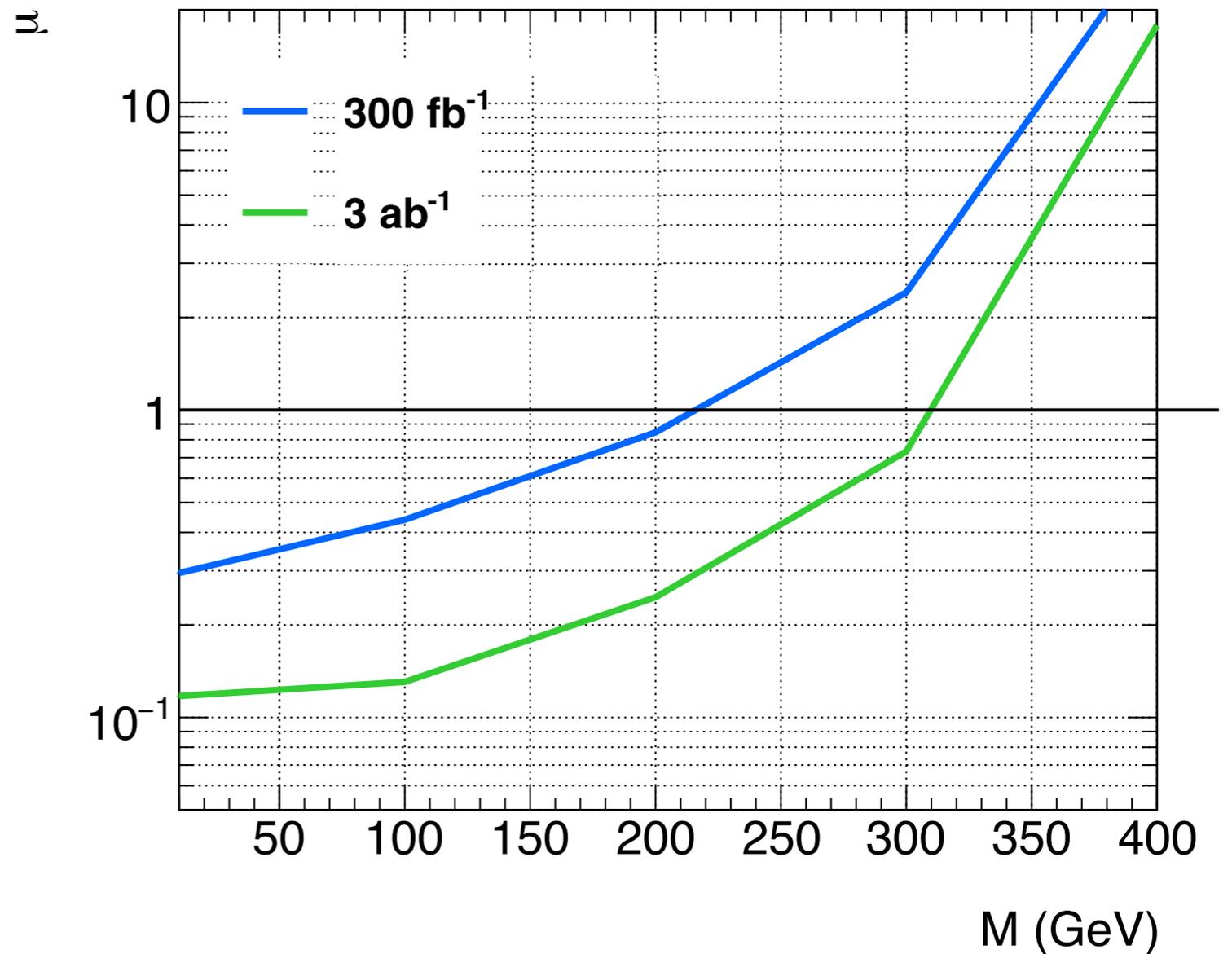


[Haisch,PP,Polesello 2017]

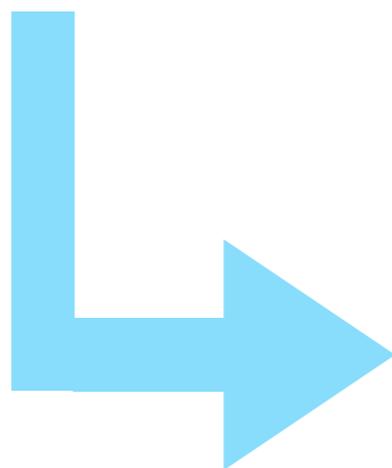
# Run 3 and HL-LHC outlook



Exclude scalar hypothesis in favour of the pseudo scalar one



Multi-bin  
fit



[Haisch,PP,Polesello 2017]