

Forward detector system for the LUXE experiment

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05/12/19

Luxe Technical meeting

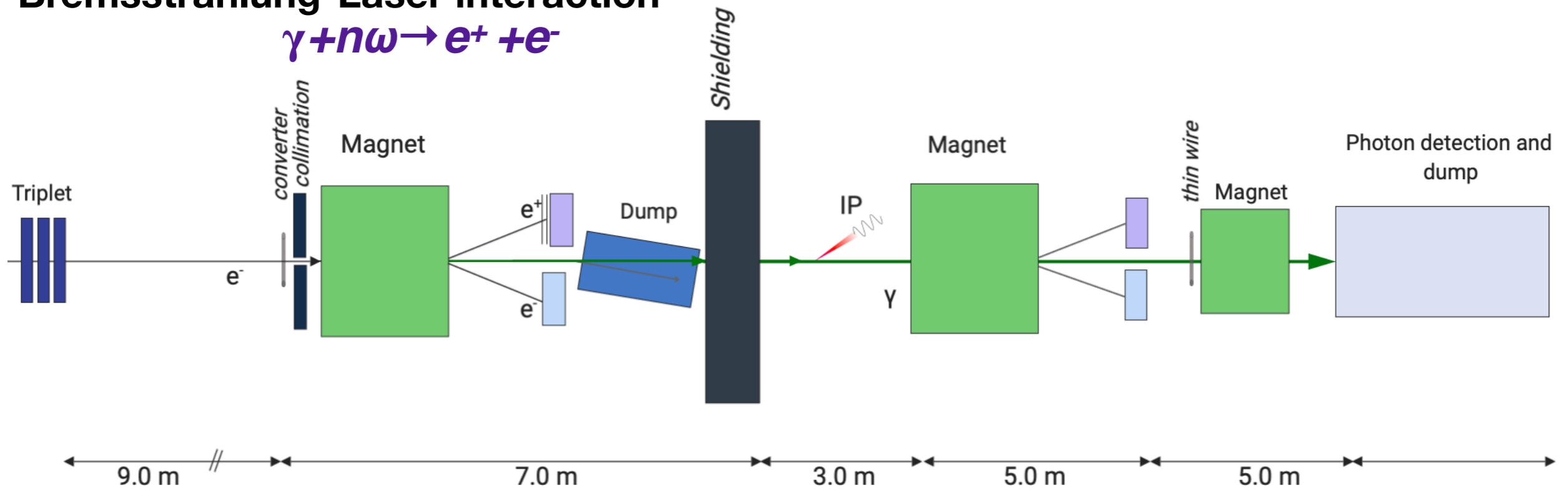
DESY Hamburg

The logo for the LUXE experiment, featuring the word "LUXE" in a bold, blue, sans-serif font. The letter "X" is stylized with a white starburst or spark effect at its center.

LUXE Set Up

Bremsstrahlung-Laser interaction

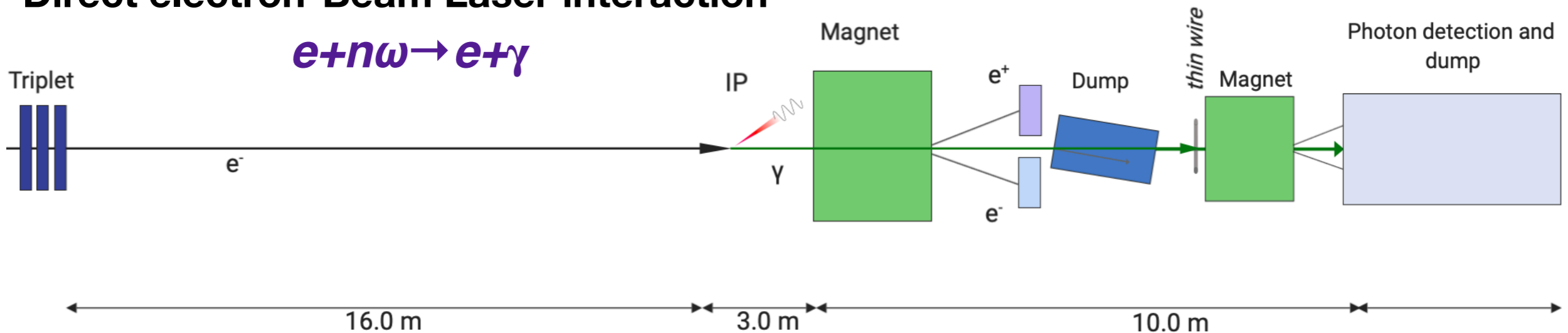
$$\gamma + n\omega \rightarrow e^+ + e^-$$



the non-linear pair-production process

Direct electron-Beam Laser interaction

$$e + n\omega \rightarrow e + \gamma$$



one-step and two-step trident
the non-linear Compton processes

Lead glass blocks from Hermes Experiment



Available: 6 calorimeter blocks w/ measures $9 \times 9 \text{ cm}^2$, length is 50 cm

TABLE 1. Chemical composition and calorimetric properties of F101 Lead Glass. Cerium is making the Lead Glass radiation hard, while also reducing its transparency.

Chemical Composition F101	weight %
PB_3O_4	51.23
SiO_2	41.53
K_2O	7.0
Ce	0.2
<hr/>	
Radiation Length	2.78 cm
Critical Energy	17.97 MeV
Refraction index	1.65
Molière Radius	3.28 cm

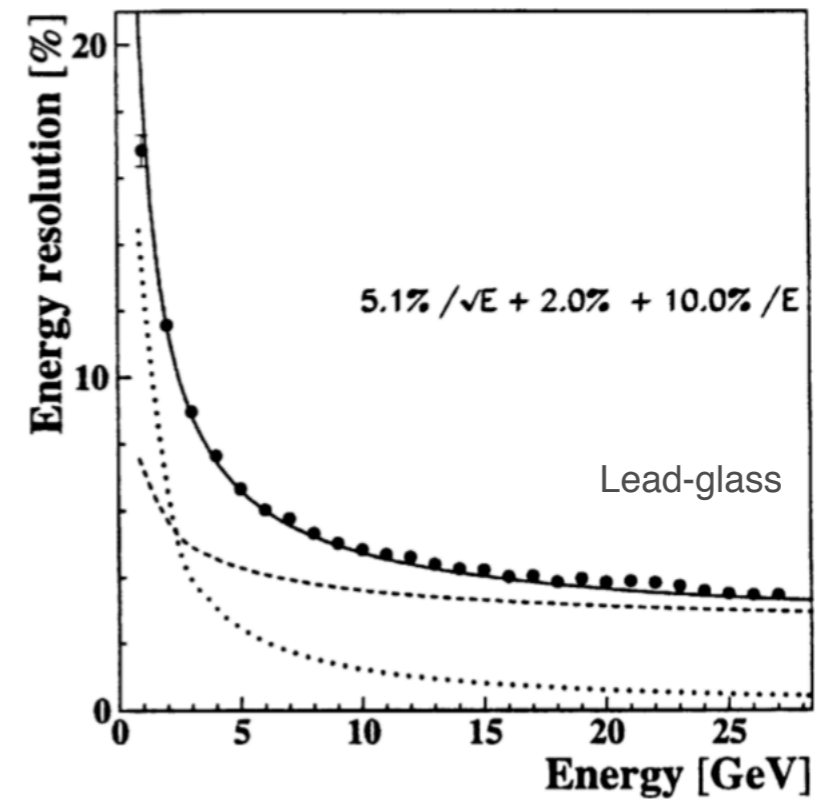
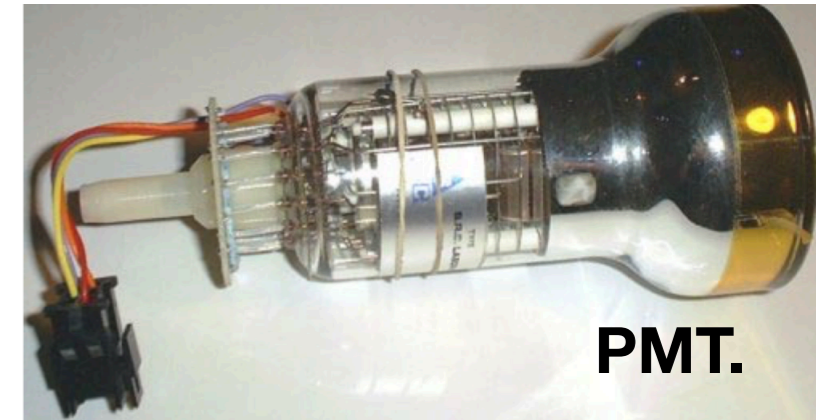
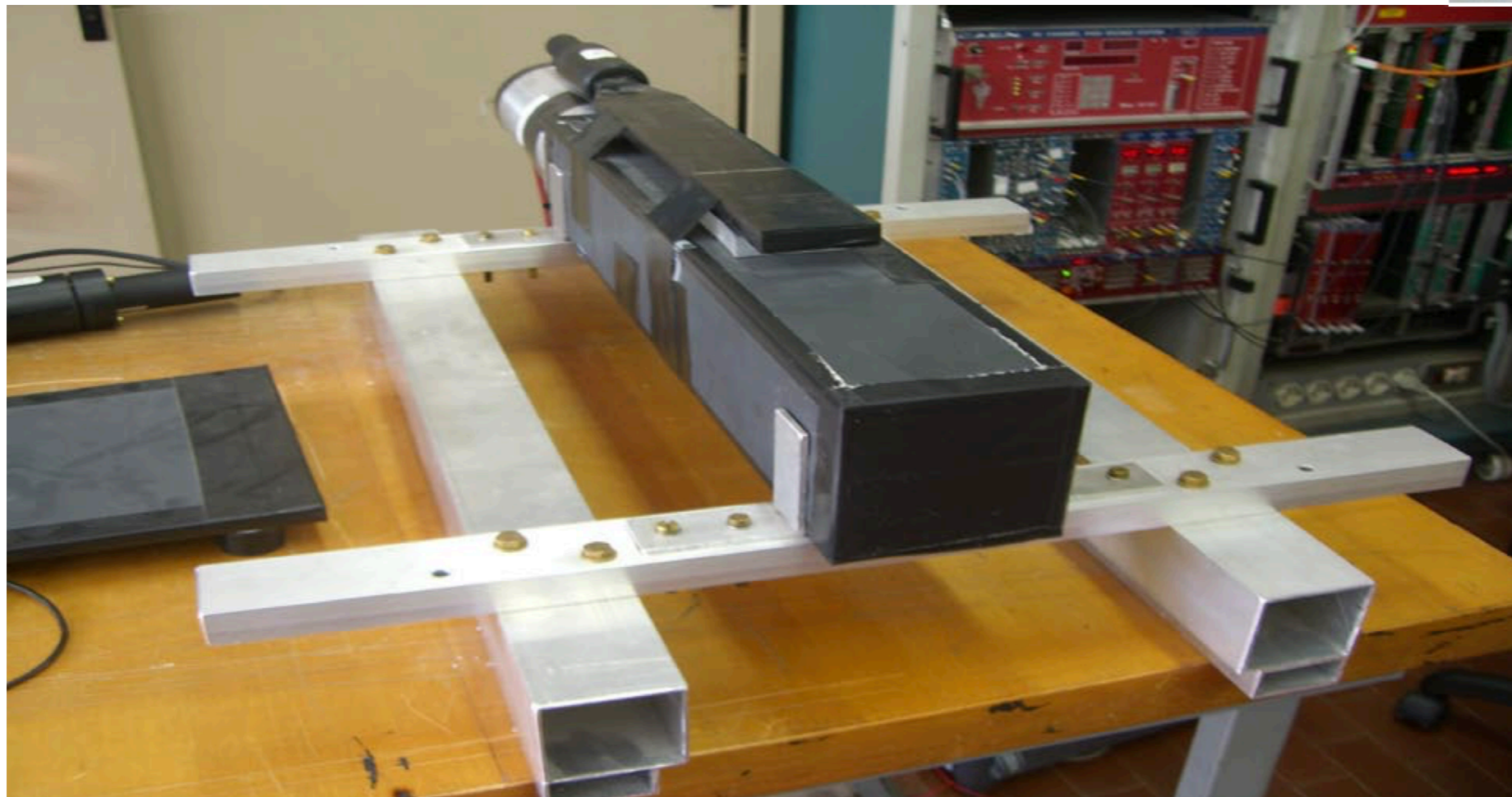


Fig. 5. Energy resolution of the calorimeter; the circles correspond to the 1996 data, the solid curve is the sum of the contributions from the lead-glass (dashed curve) and from the pre-shower (dotted curve) provided at test beam measurement [7].

Wrapped LG block

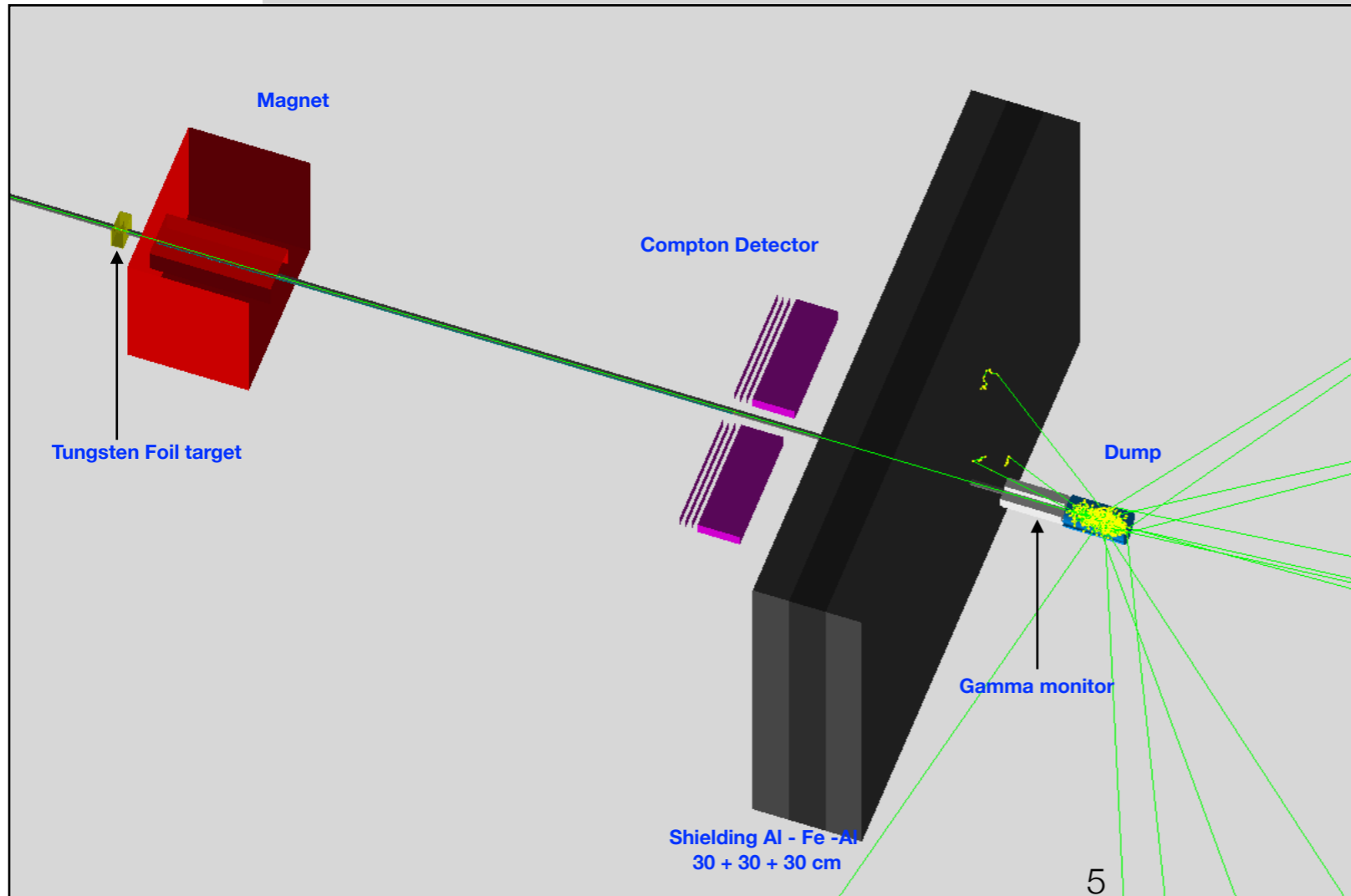
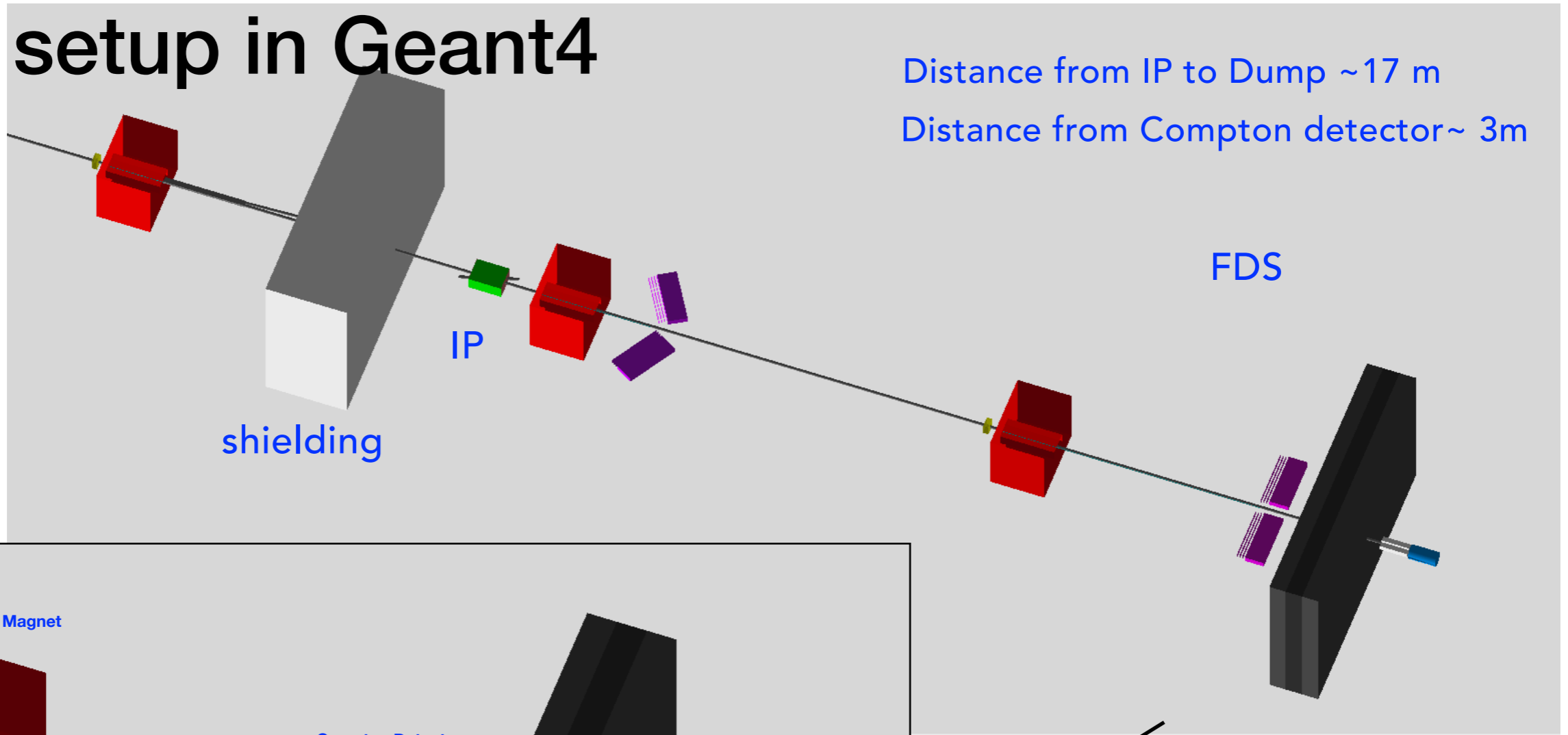


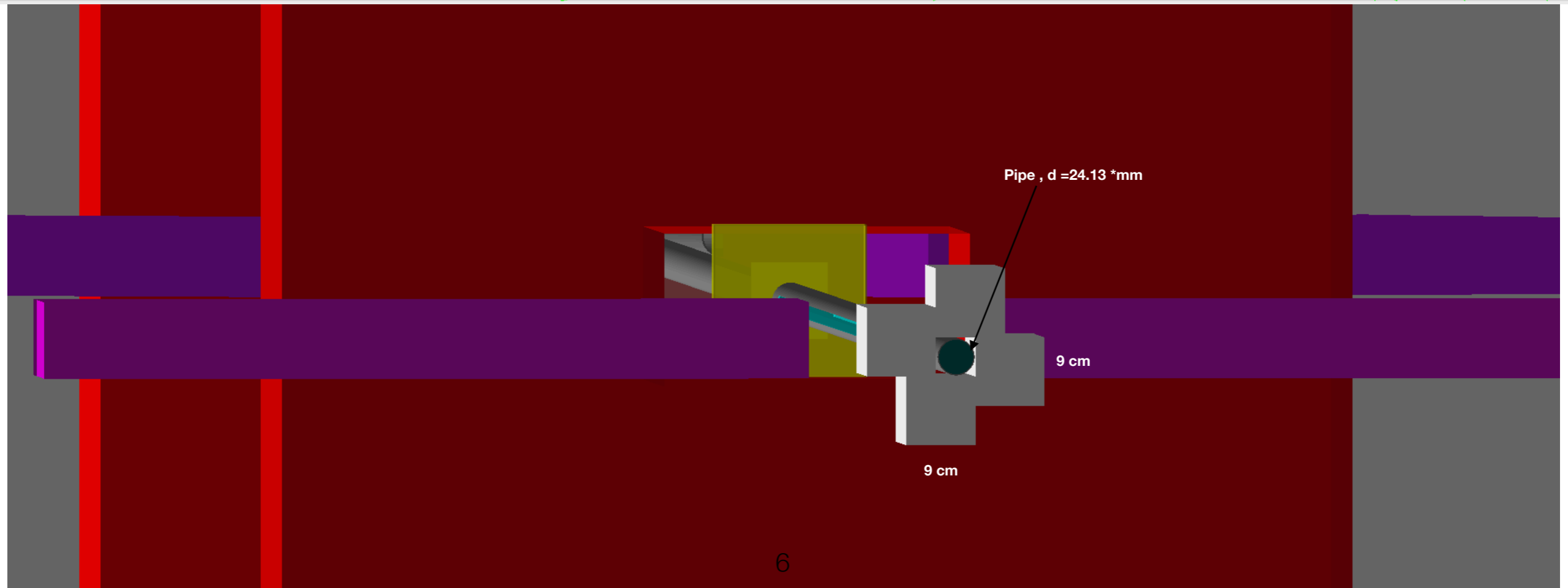
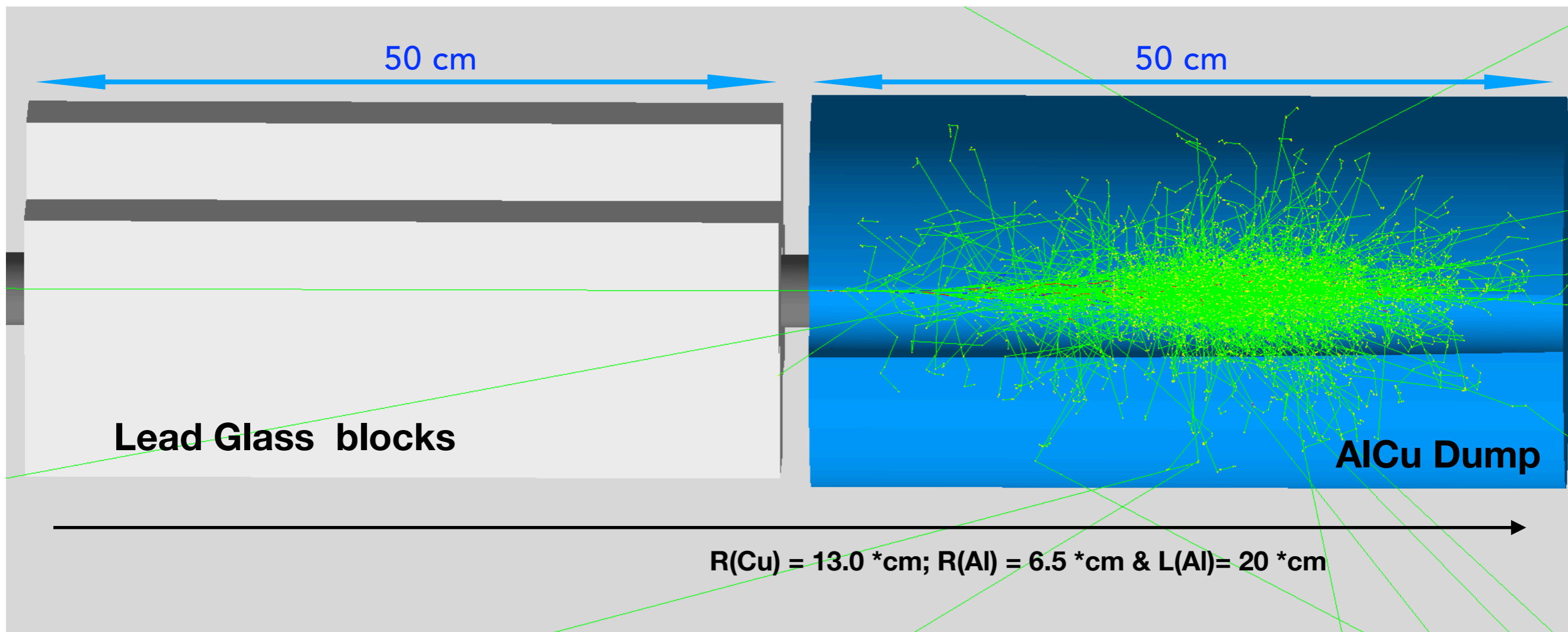
PMT.



blocks are wrapped with an aluminized mylar foil to reflect the Cherenkov photons, and a tedlar foil to make the blocks light tight.

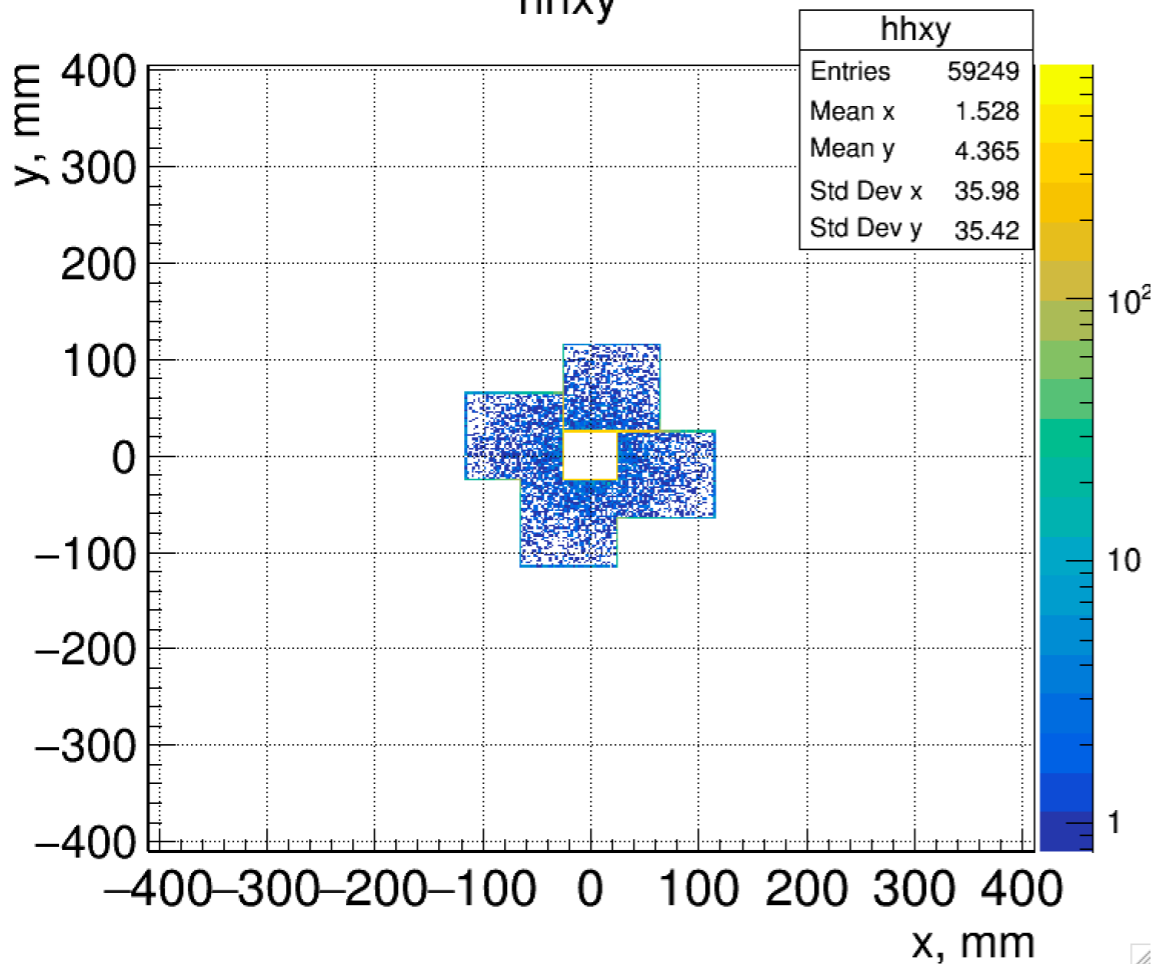
Gamma Monitor made of 4 Hermes GL blocks in Luxe setup in Geant4



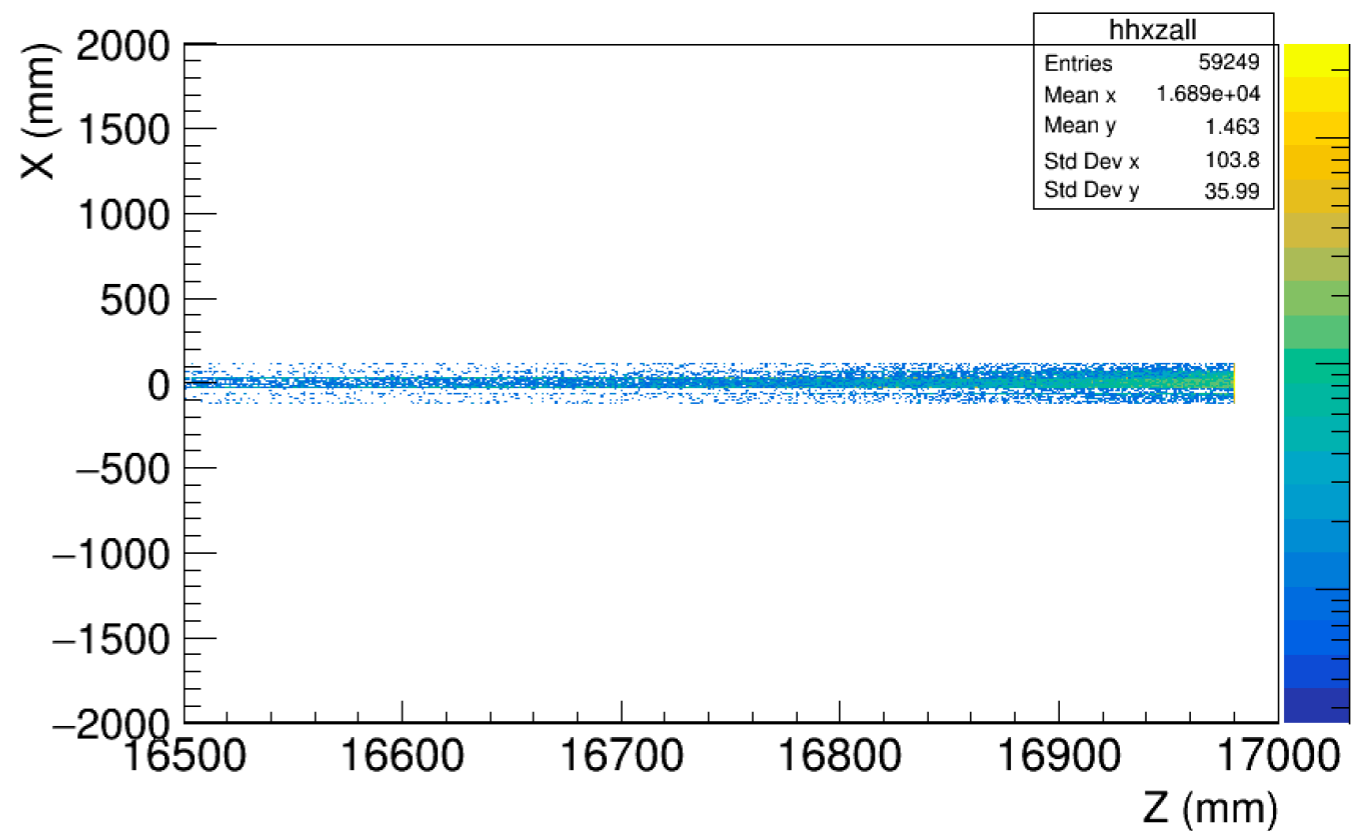


Control plots for LG monitor

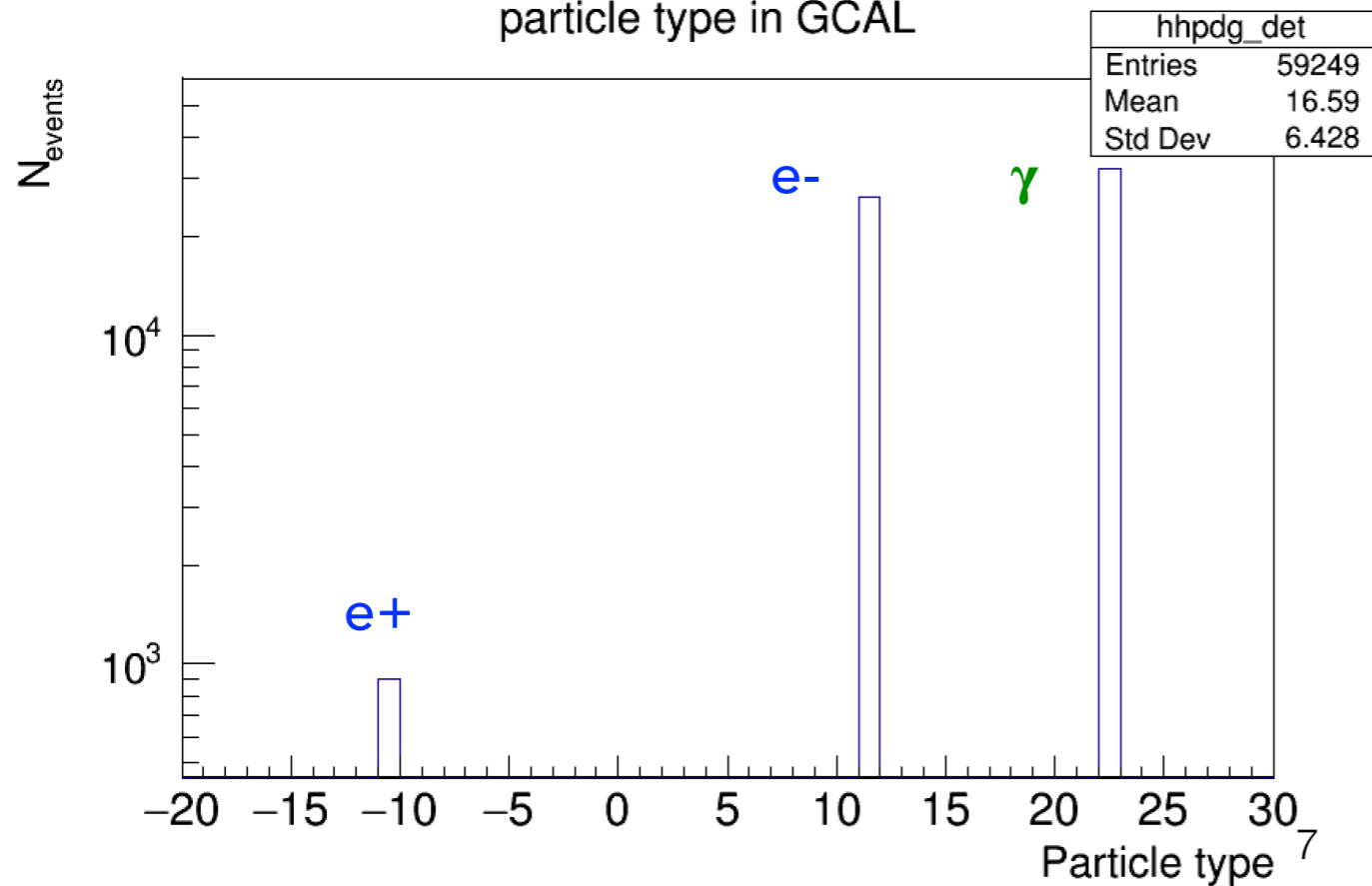
hhxy

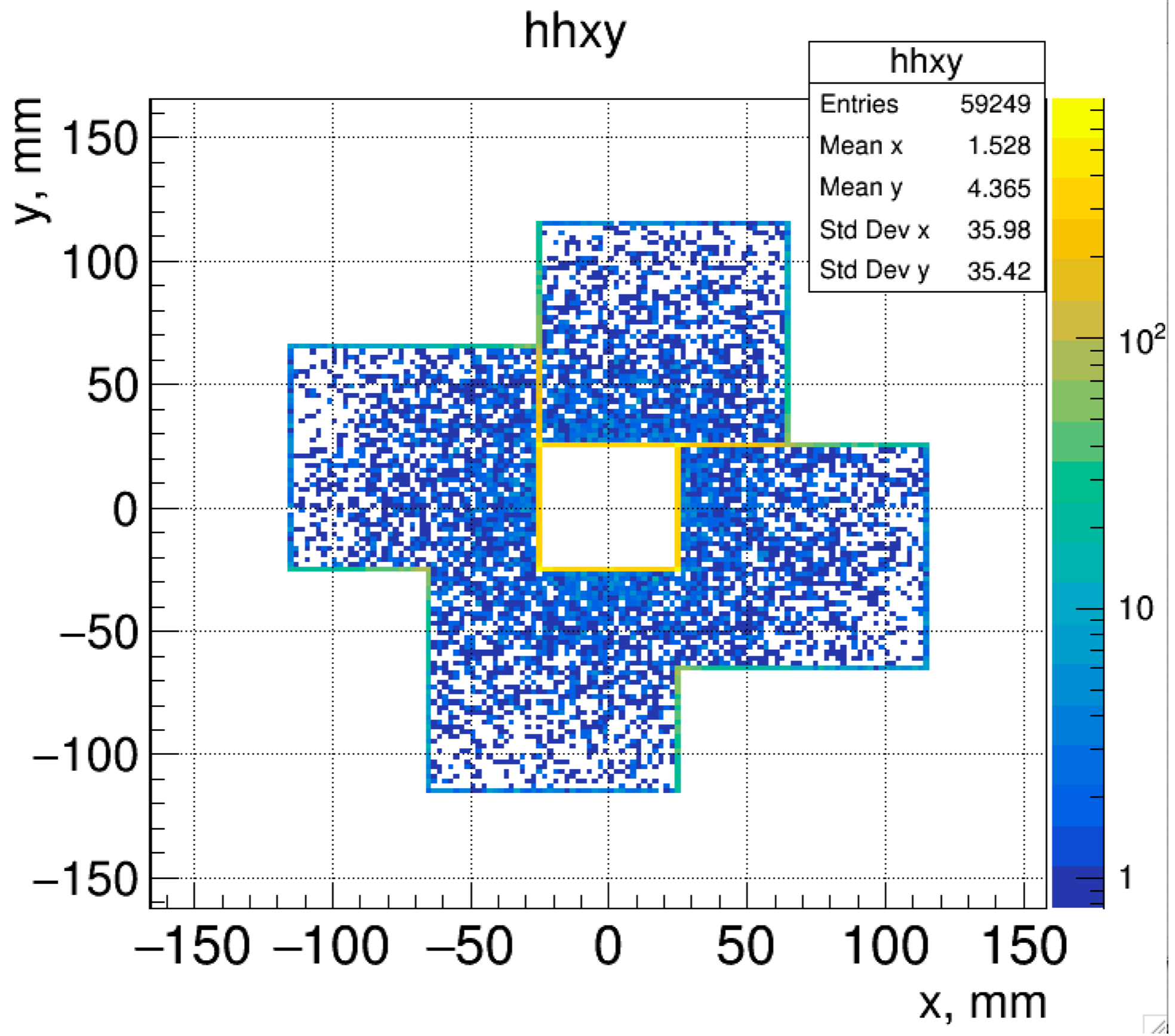


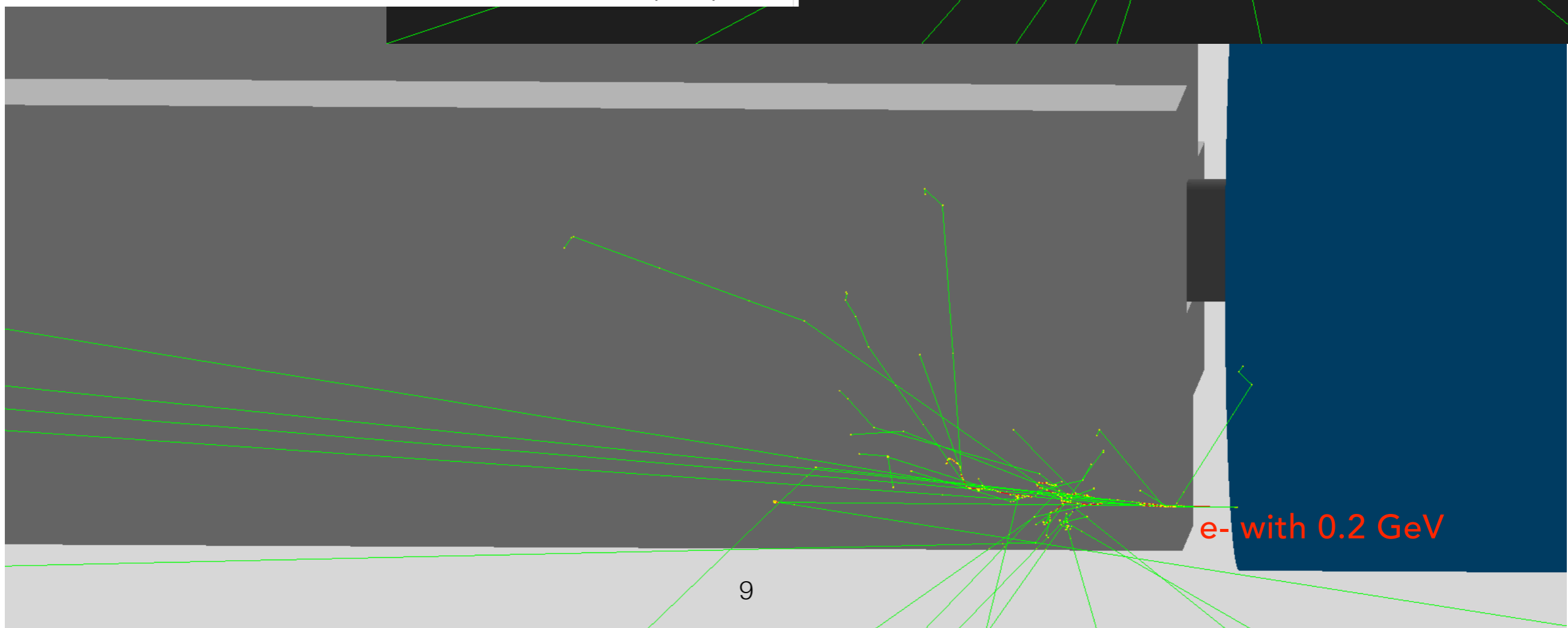
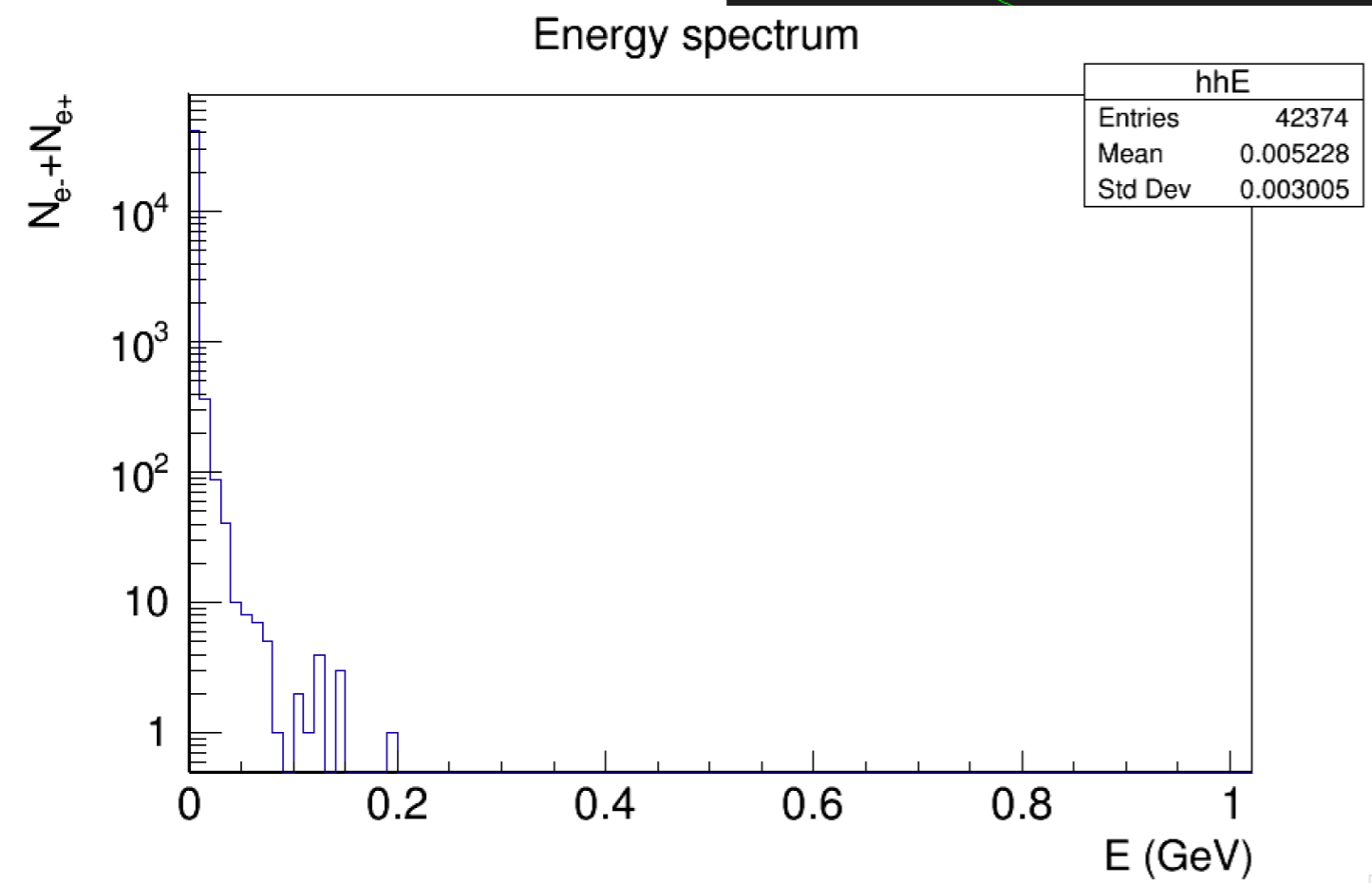
Hits xz distribution all



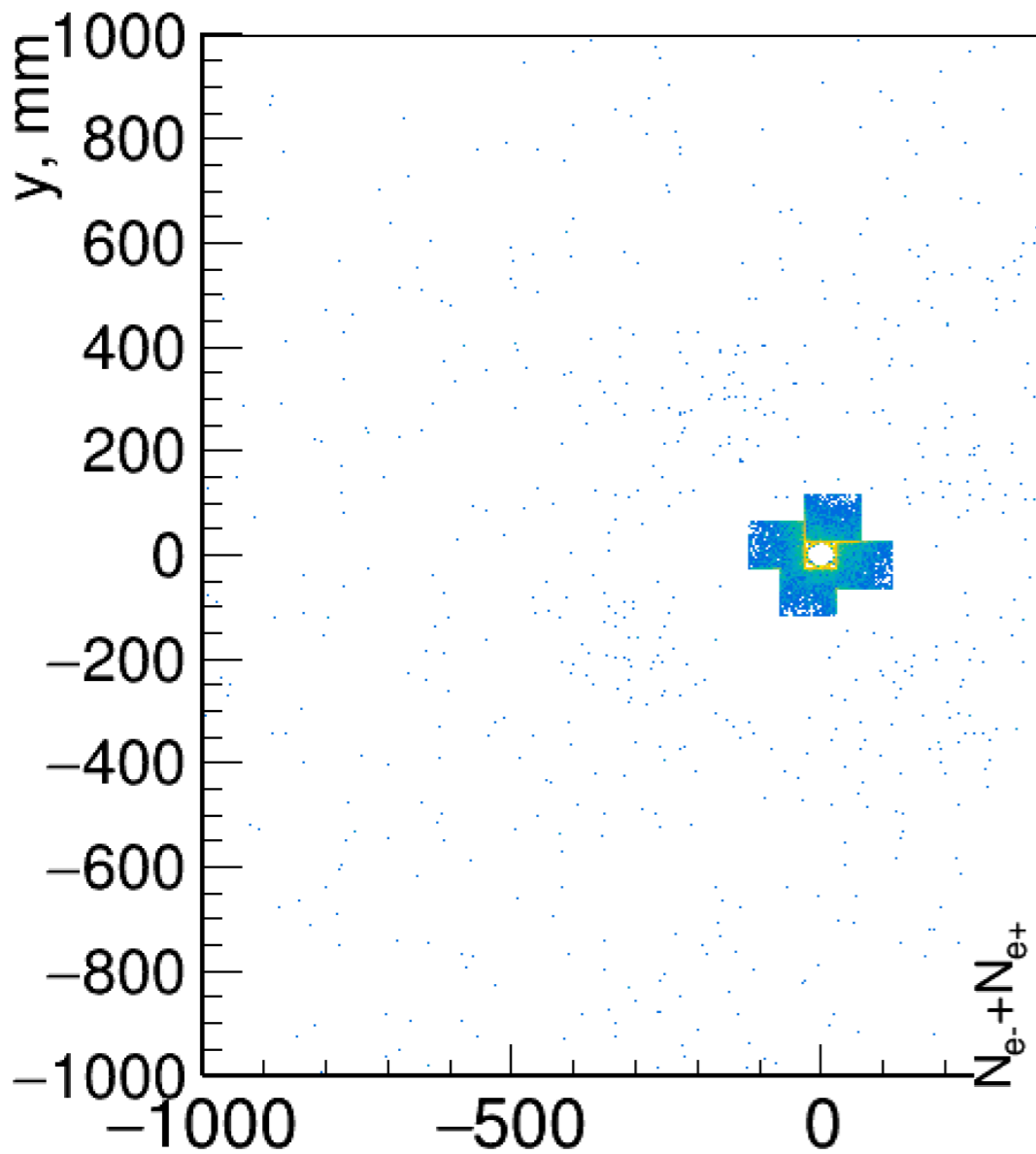
particle type in GCAL



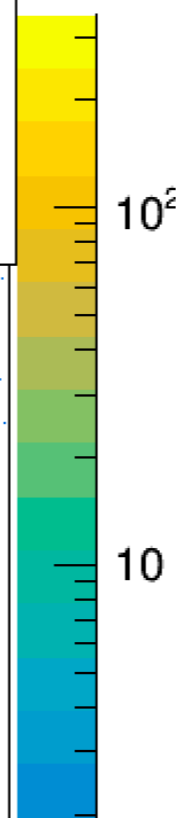




hhvtxy



hhvtxy	
Entries	59249
Mean x	1.69
Mean y	3.321
Std Dev x	69.01
Std Dev y	67.72

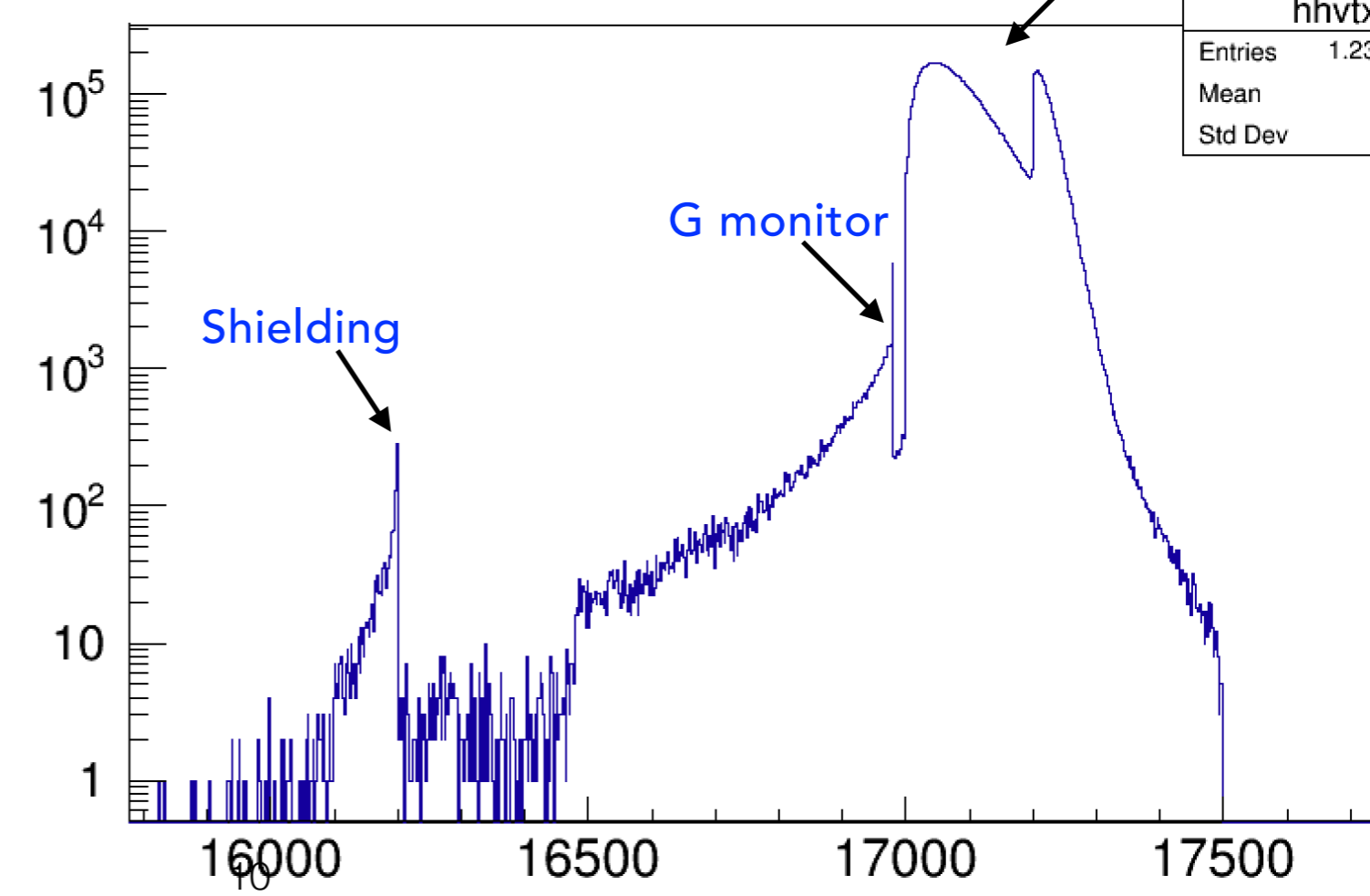


Vertexes of the particles

$N_{e^-} + N_{e^+}$

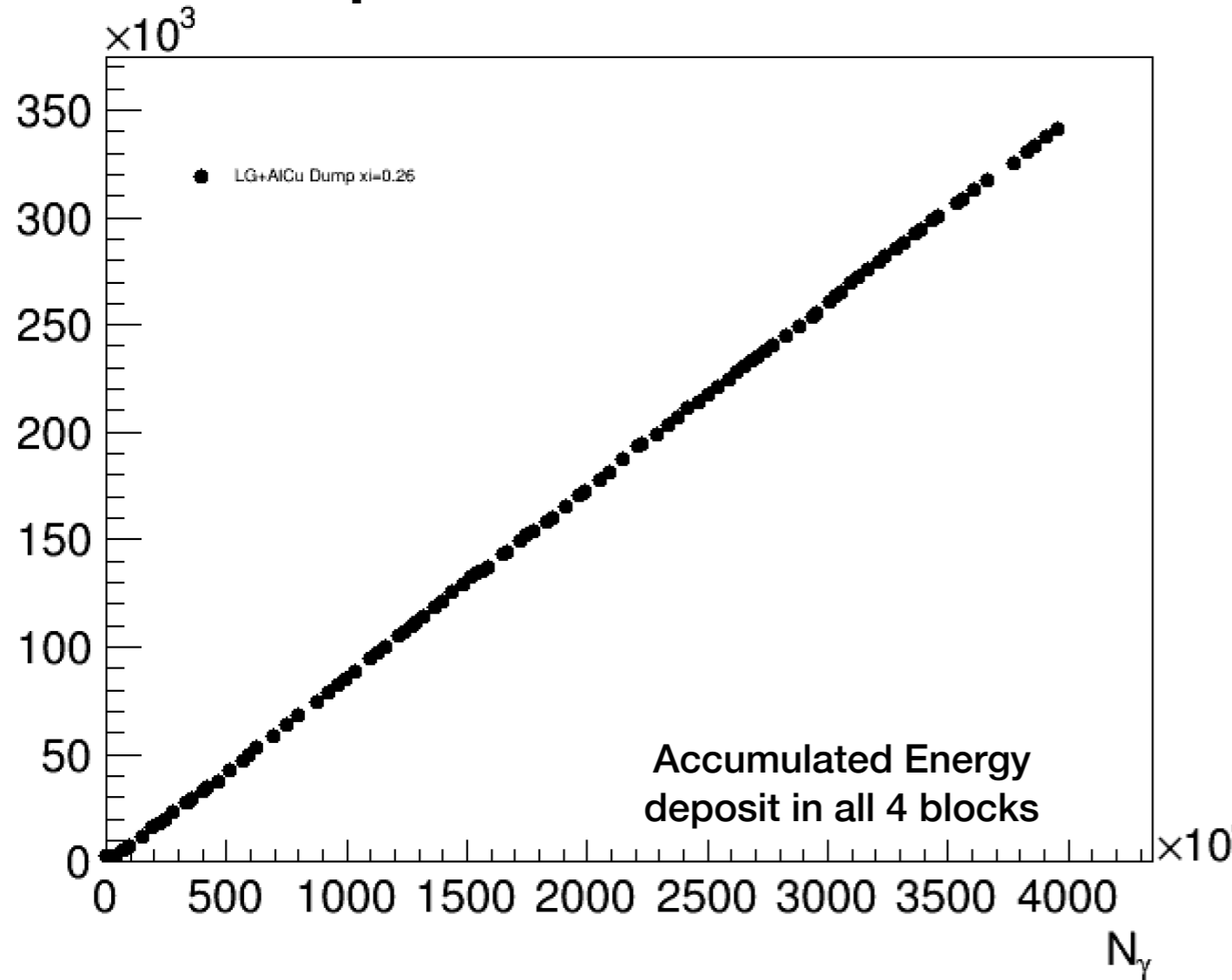
Vertex Z

ALCu Dump



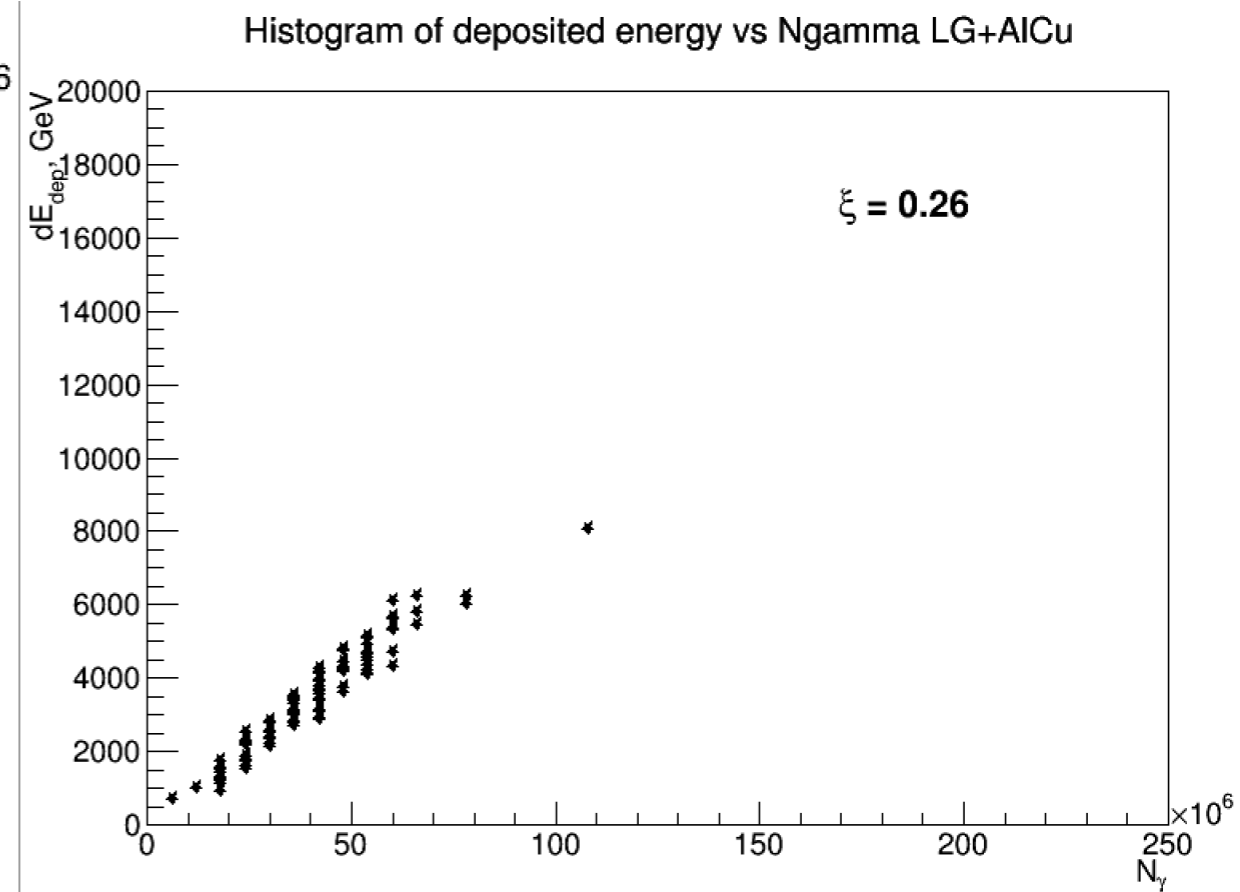
hhvtxz	
Entries	1.230929e+07
Mean	1.711e+04
Std Dev	75.69

The dependence of deposited energy on number of incoming photons for LG Gamma monitor and AlCu dump



The linear dependence of deposited energy on number of incoming photons in LG allows the usage of backscatters for estimating the photon flux

Energy deposit on N photons
Each point is one BX:



Outlook

- **Compton detector studies:**

- **Gamma monitor studies:**

- * The implementation in Luxe geometry of the LG Gamma Monitor made of Hermes LG blocks in front of Al-Cu Dump
- * Gamma Monitor was studied in GEANT4 w/ LG Monitor in front of Al-Cu Dump for the lowest intensity
- * The linear dependence of deposited energy on number of incoming photons allows the usage of backscatters for counting the photon flux for this configuration
- * The energy spectrum of backscatters is below 1 GeV and for the vast majority is below critical energy for the most detector materials

Further studies:

To run additional tests for LG Gamma Monitor to double check the performance

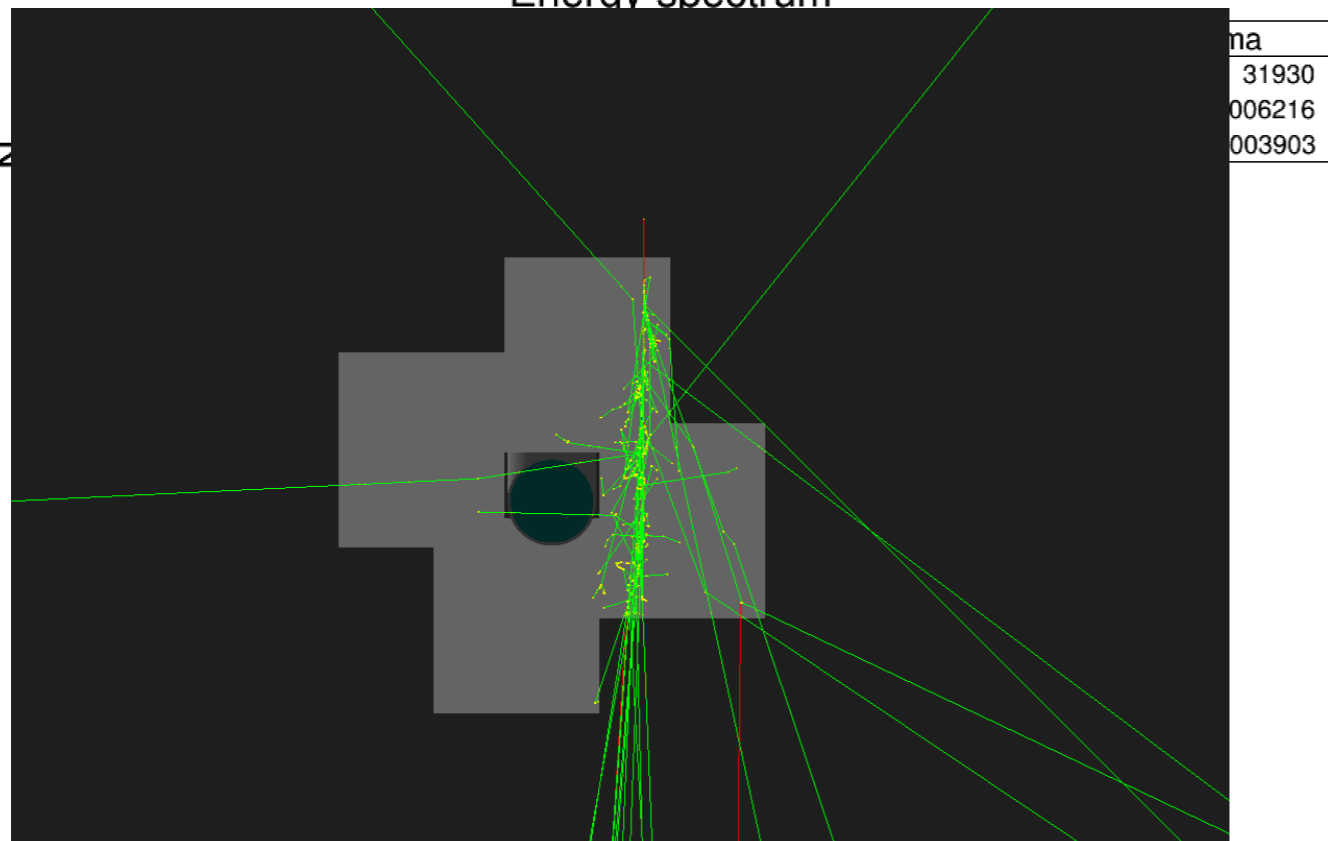
To implement the simulation of the electromagnetic shower and the resulting Cherenkov photon production

To implement mylar foil to reflect the Cherenkov photons, and a tedlar foil to block light

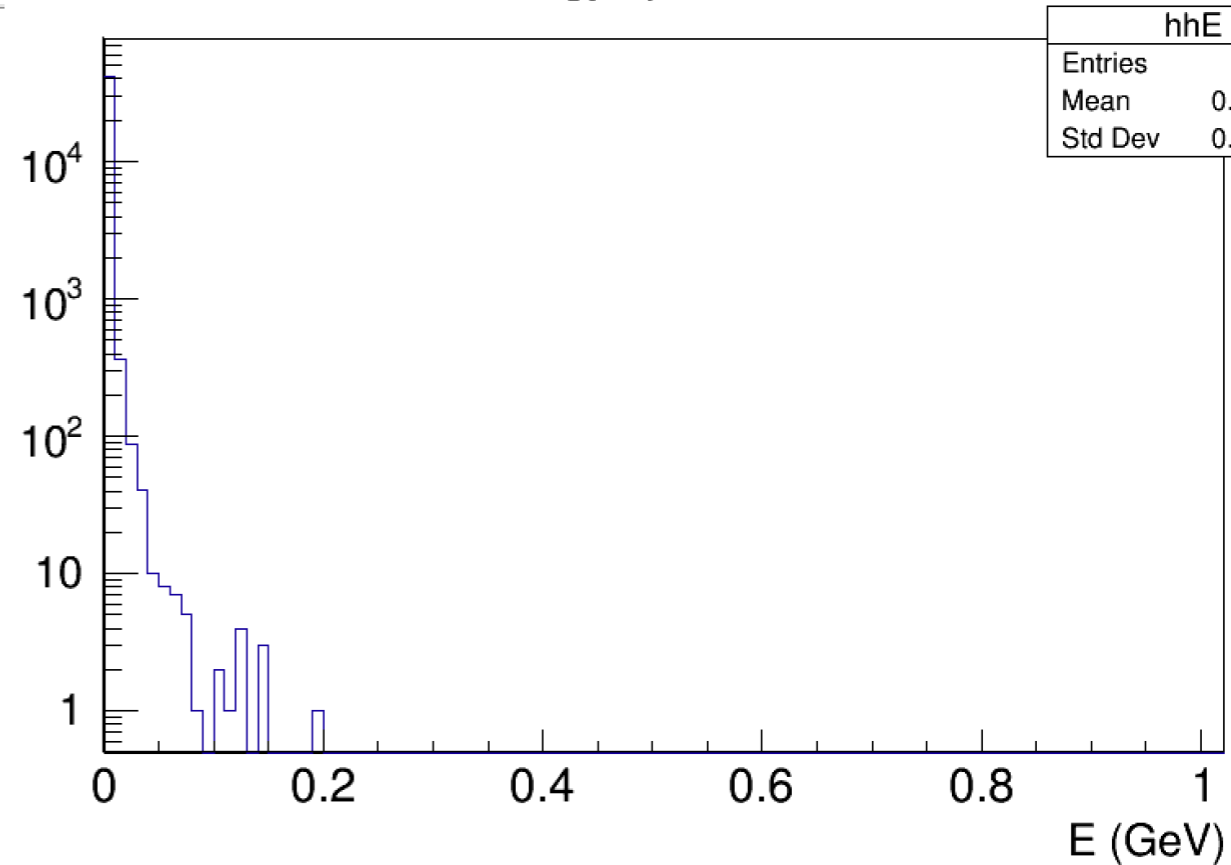
To study background

Back up

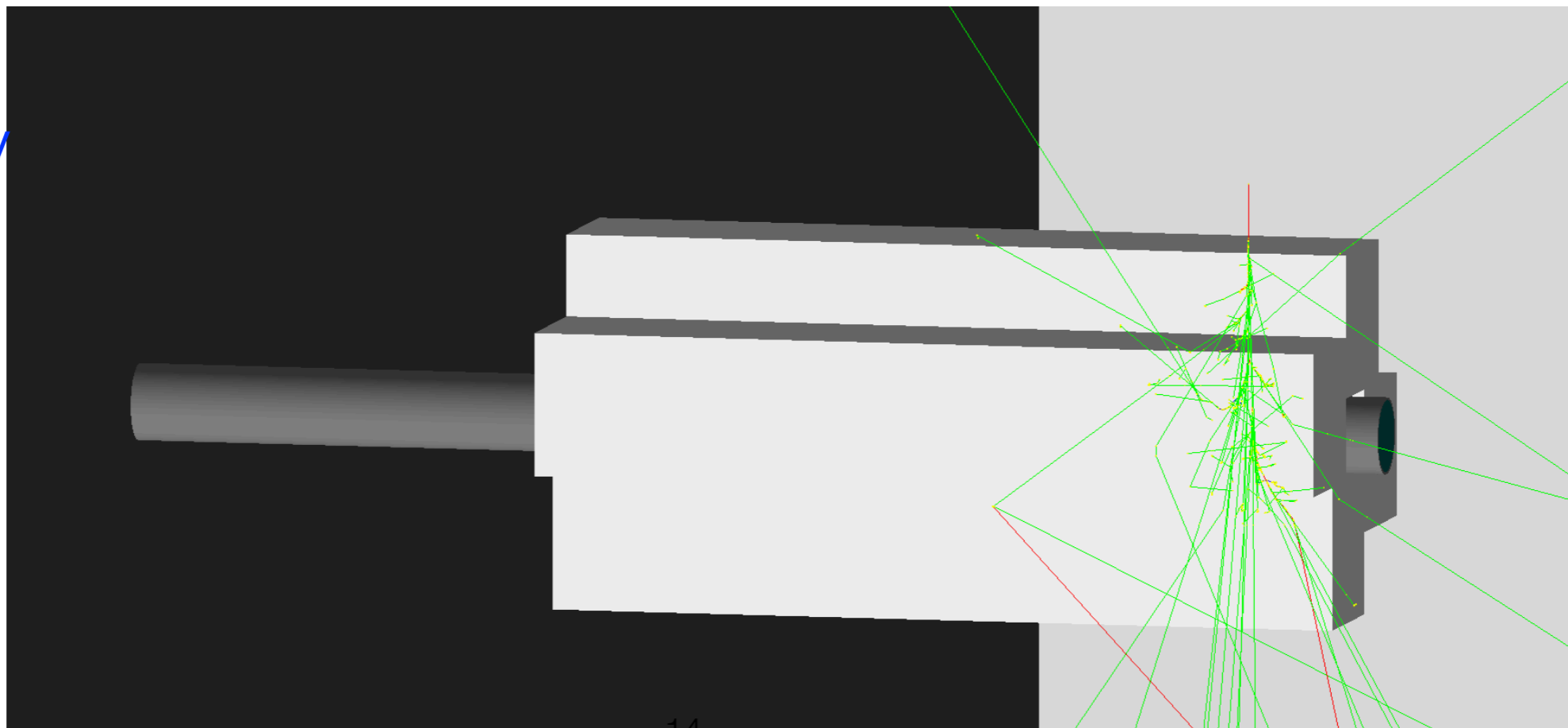
Energy spectrum



Energy spectrum

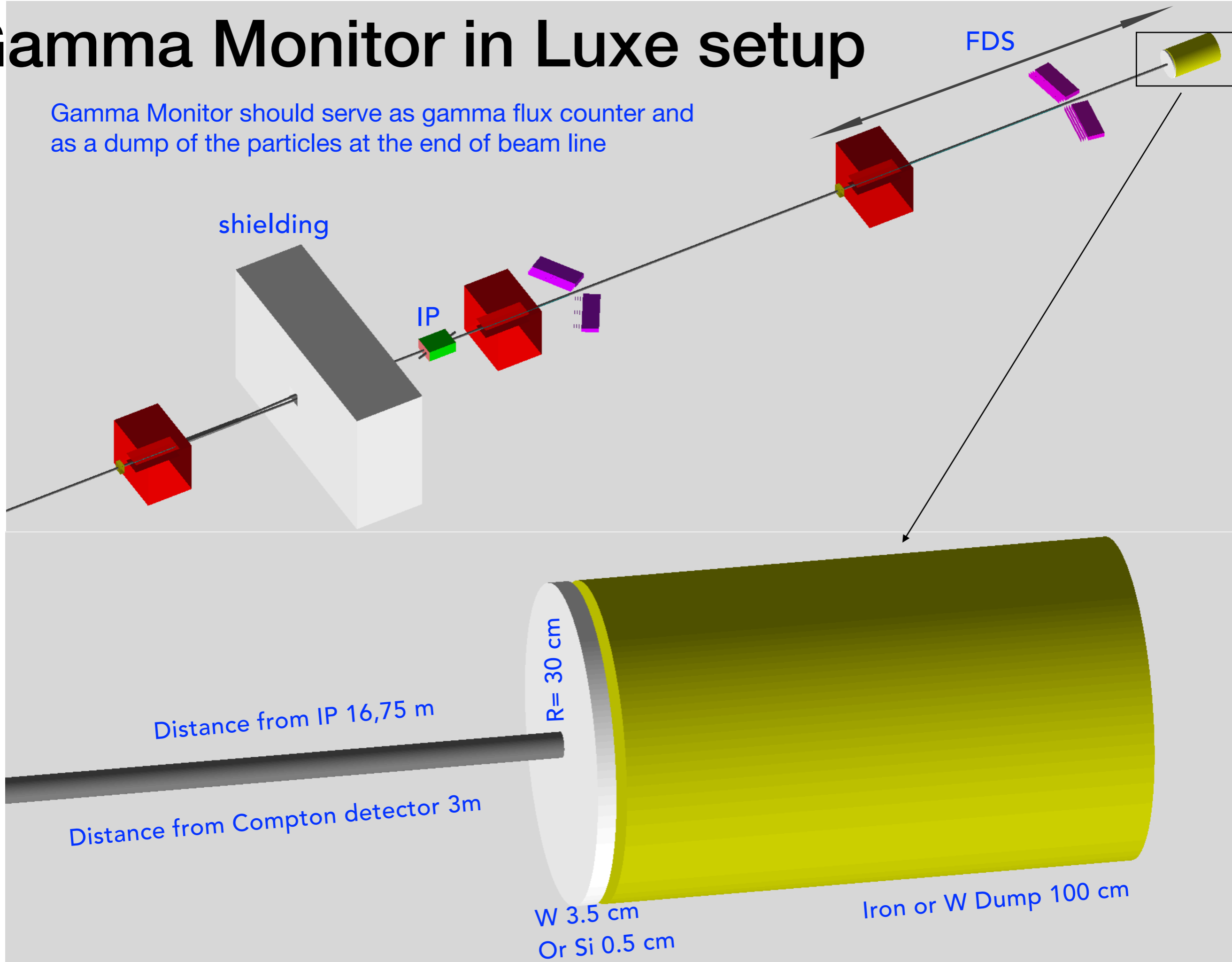


e- with 0.5 GeV

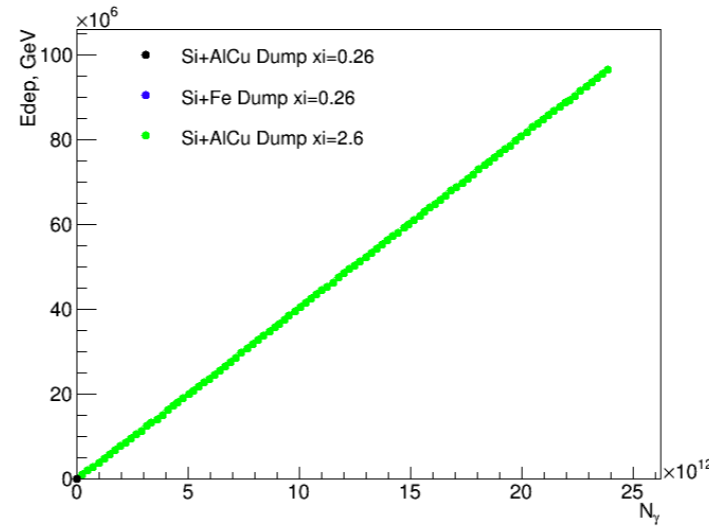
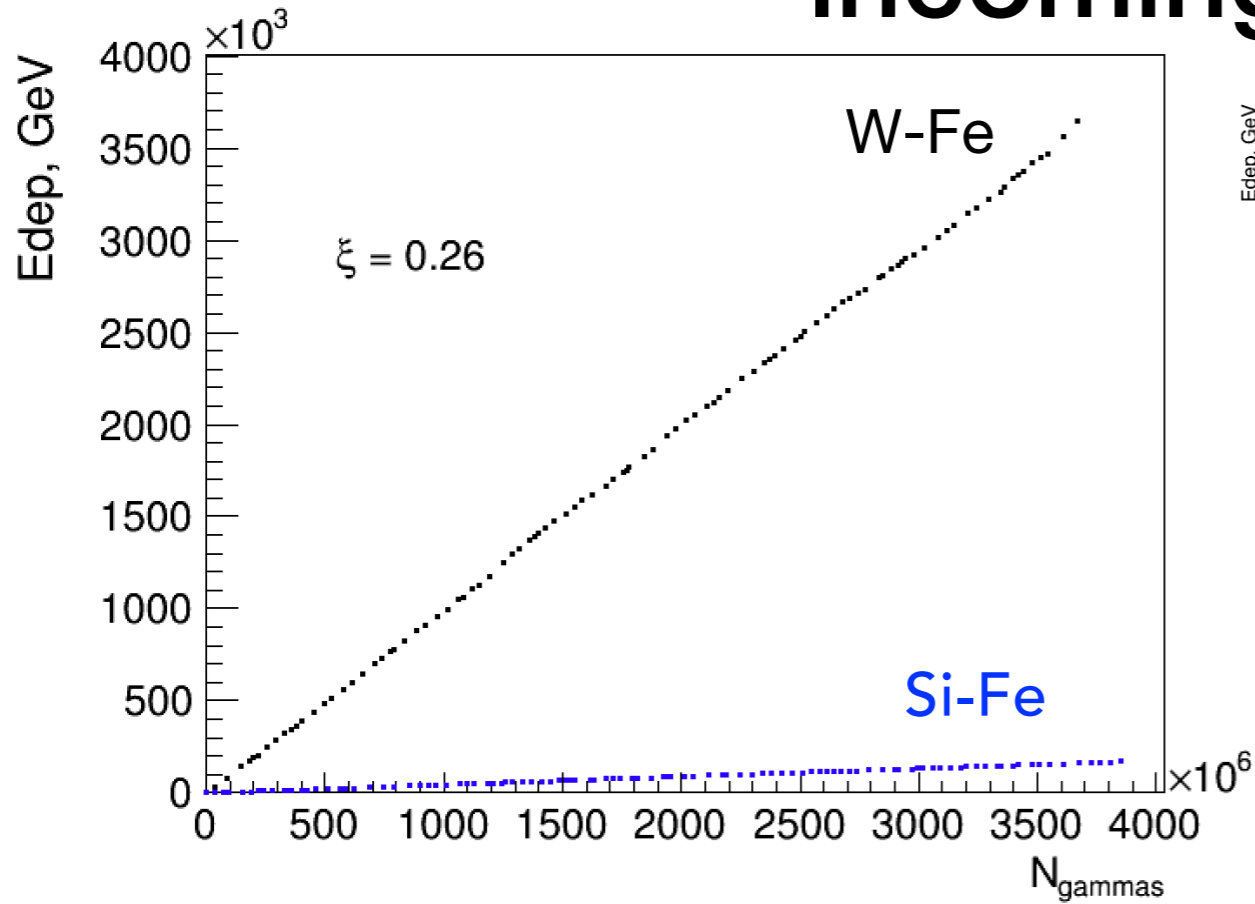


Gamma Monitor in Luxe setup

Gamma Monitor should serve as gamma flux counter and as a dump of the particles at the end of beam line

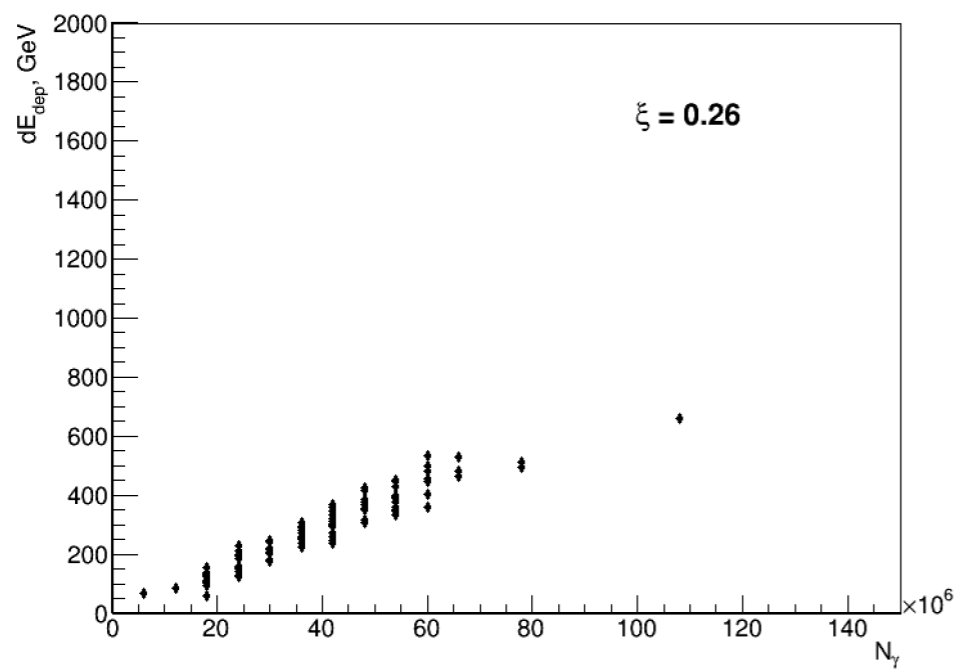


Energy dependence on number of incoming photons

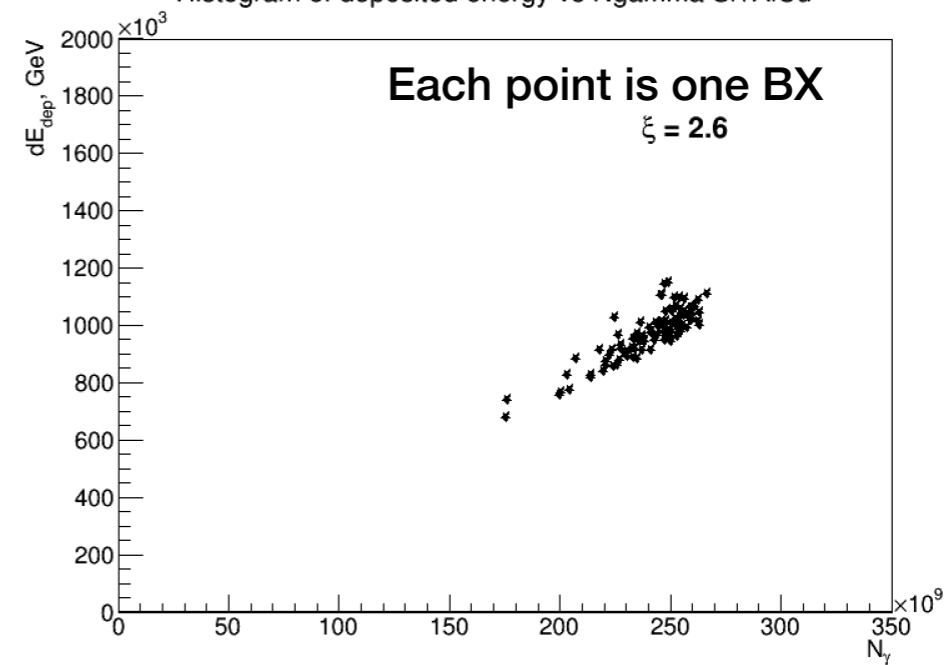


The linear dependence of deposited energy on number of incoming photons allows the usage of backscatters for estimating the photon flux

Histogram of deposited energy vs Ngamma Si+AlCu

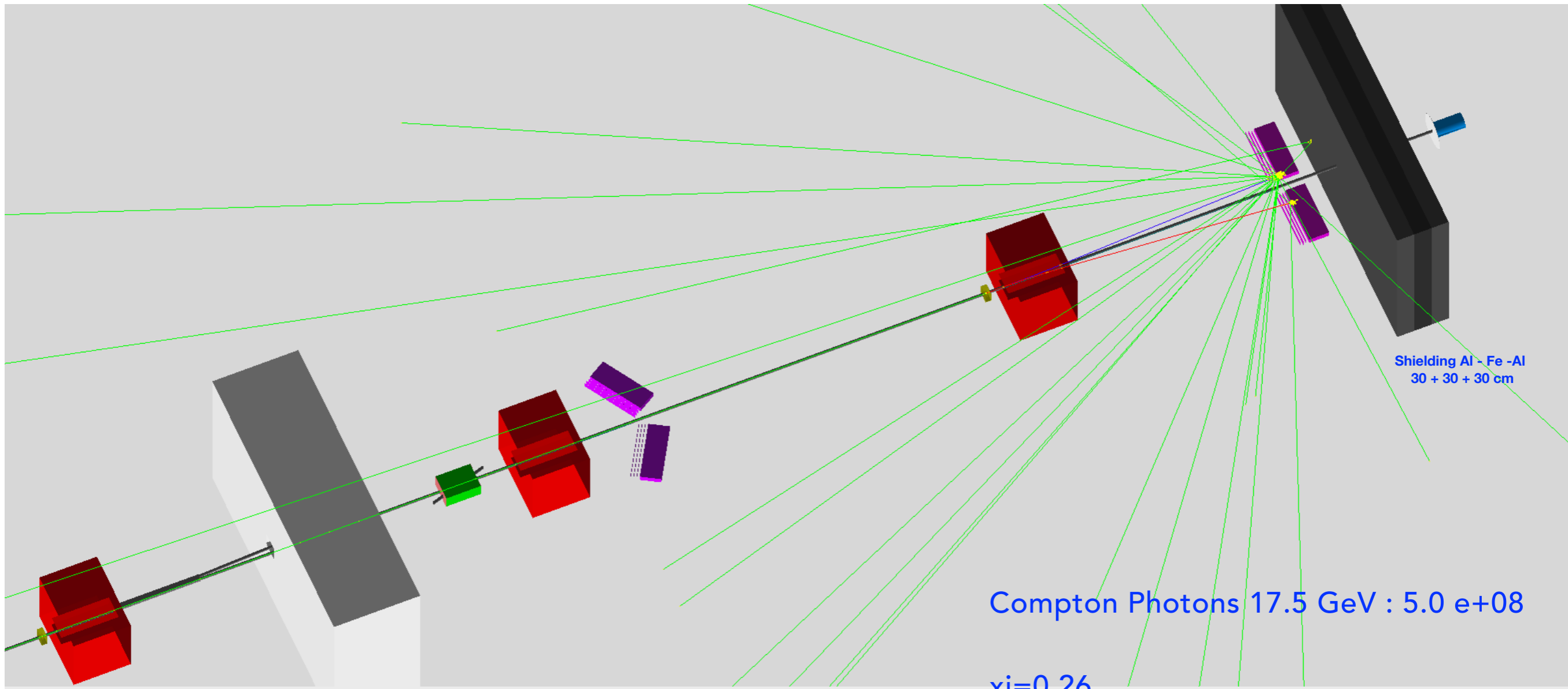


Histogram of deposited energy vs Ngamma Si+AlCu



In average one γ deposits ~ 1 keV; w/ the sigma 16.2 keV

Luxe setup with non-tilted Compton Detector



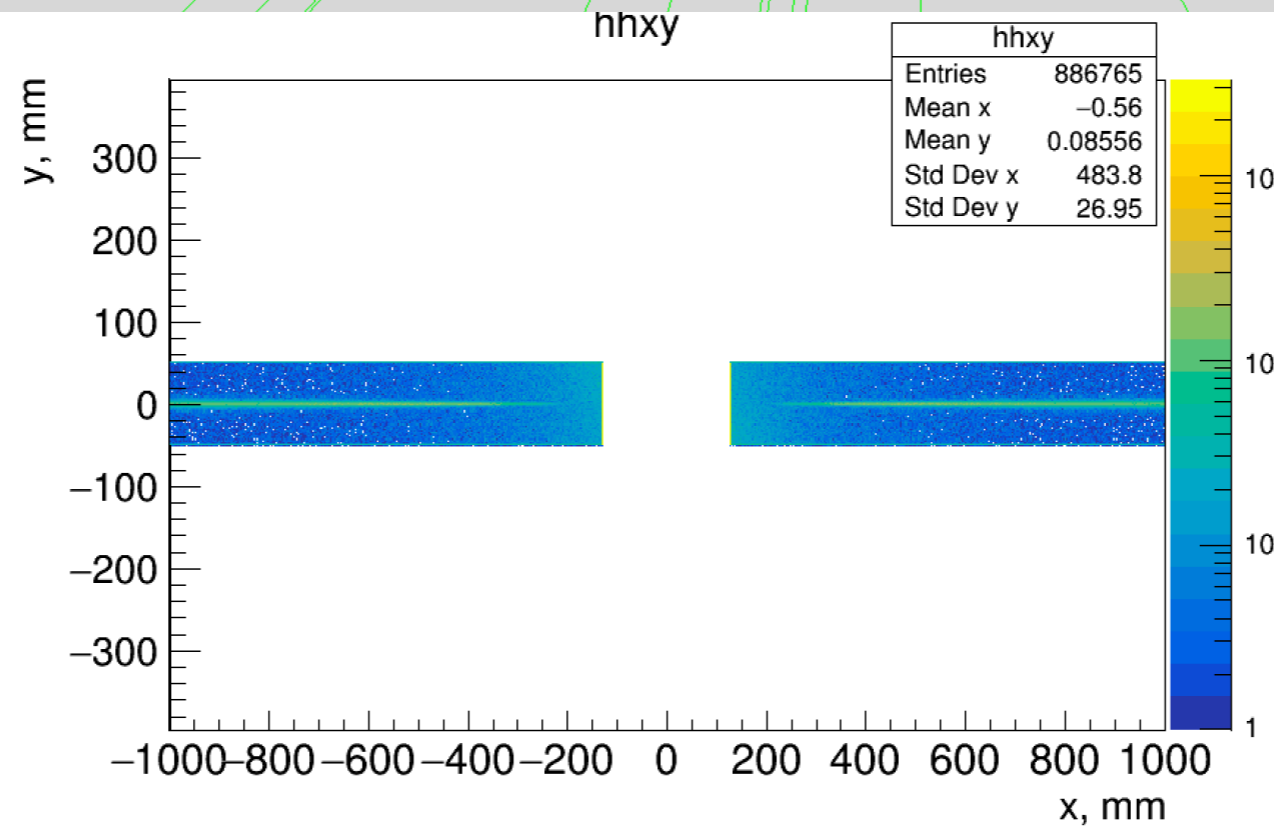
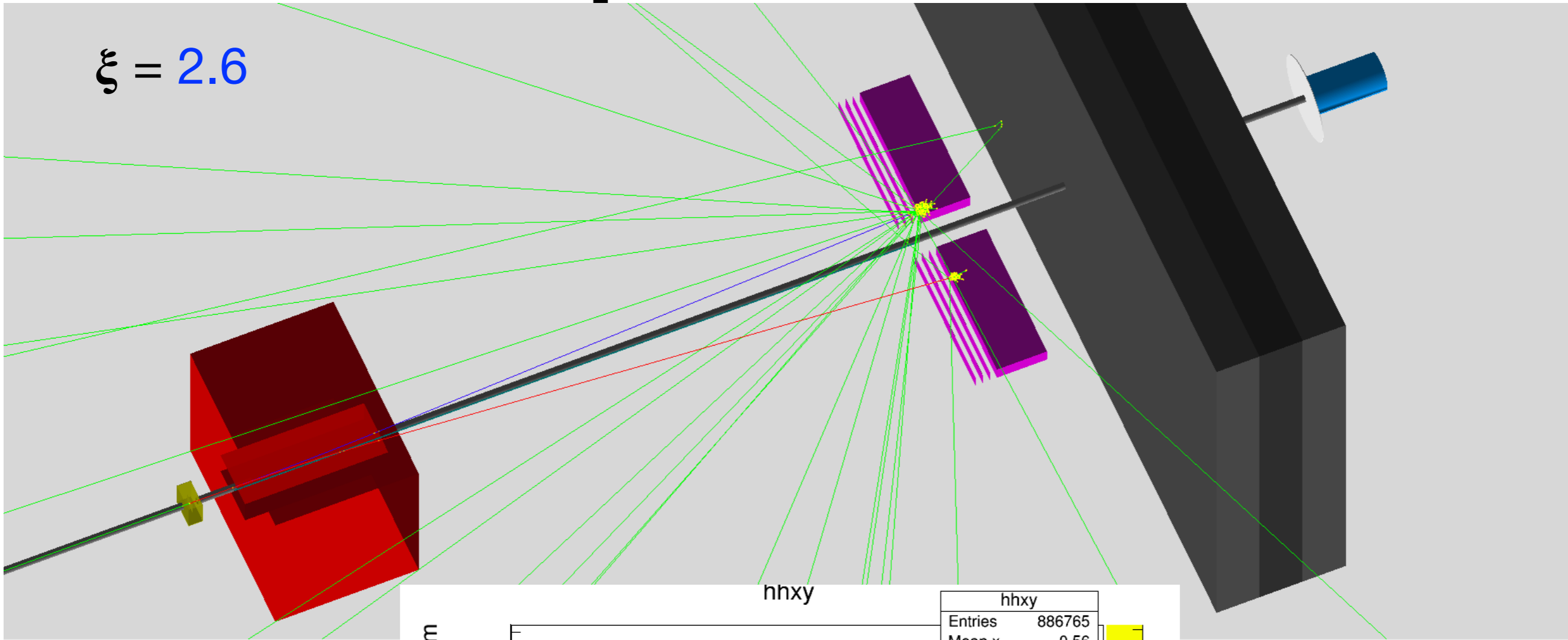
Compton Photons 17.5 GeV : 5.0 e+08

$\xi=0.26$

100 BX

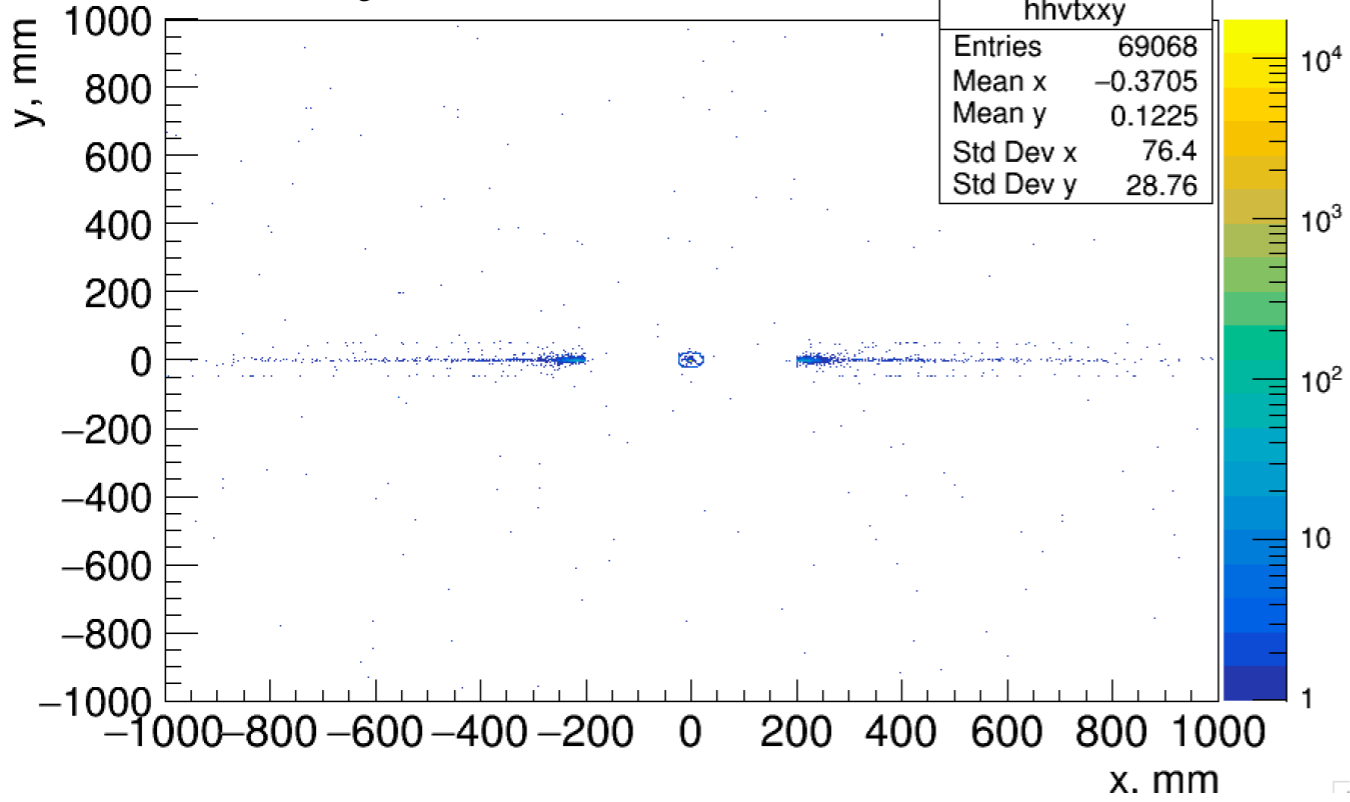
Target: W foil 10 um

Compton detector

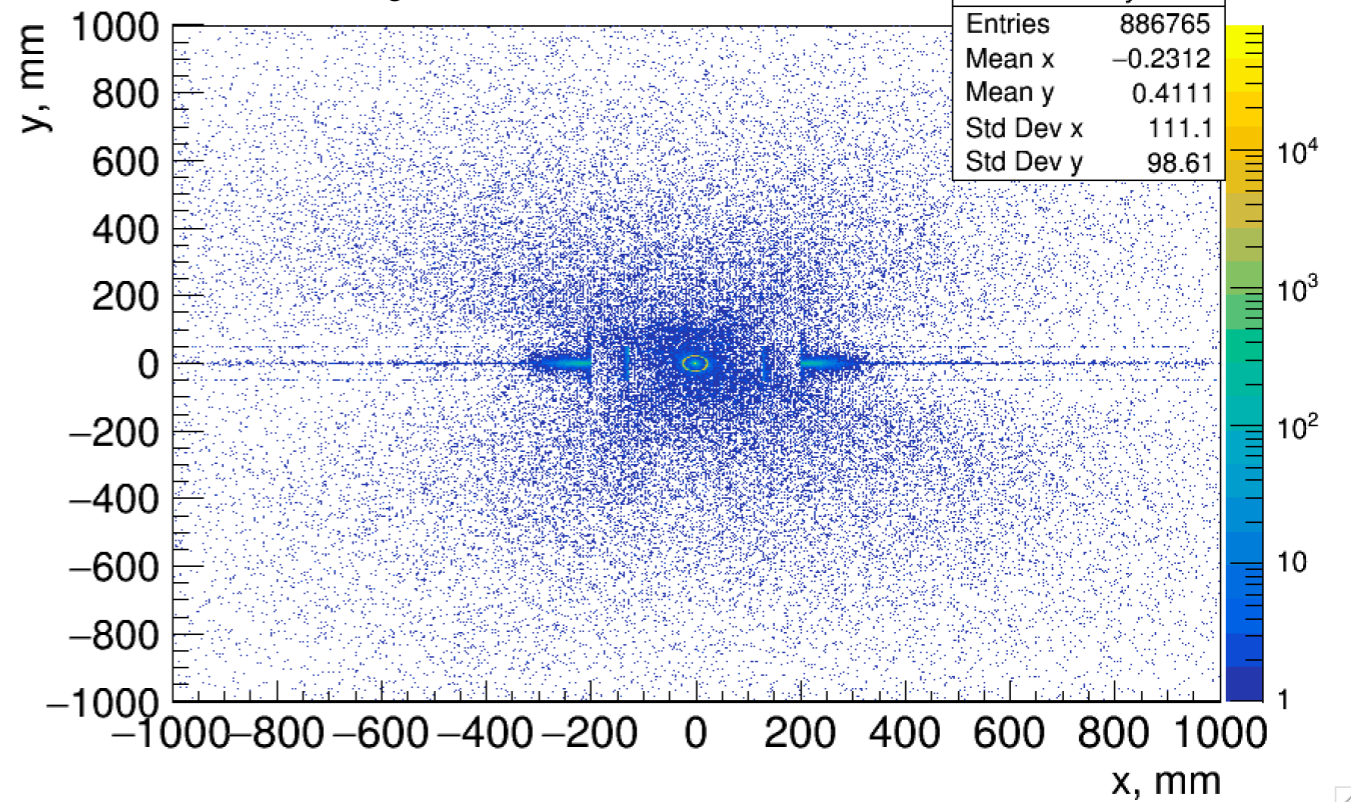


Vertexes in Compton detector

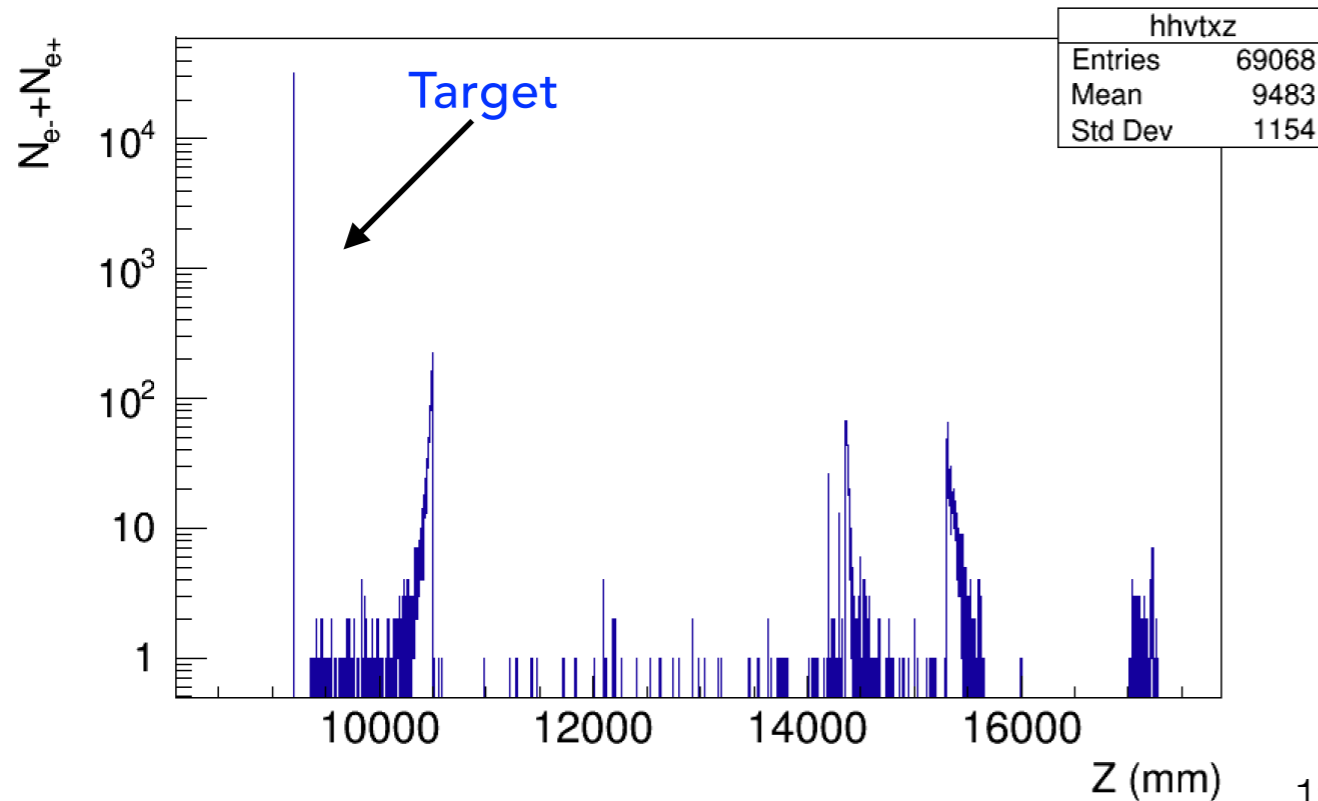
$\xi = 0.26$ nhvttxy



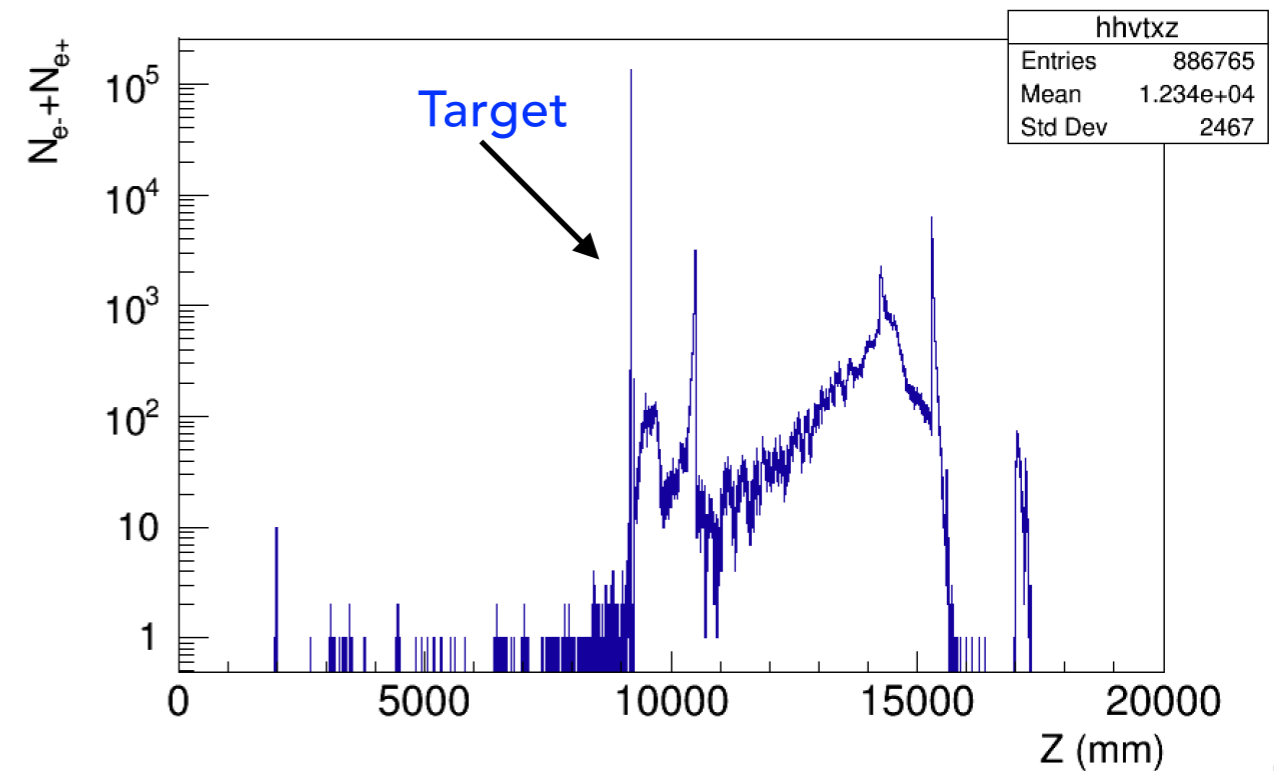
$\xi = 2.6$ hhvttxy



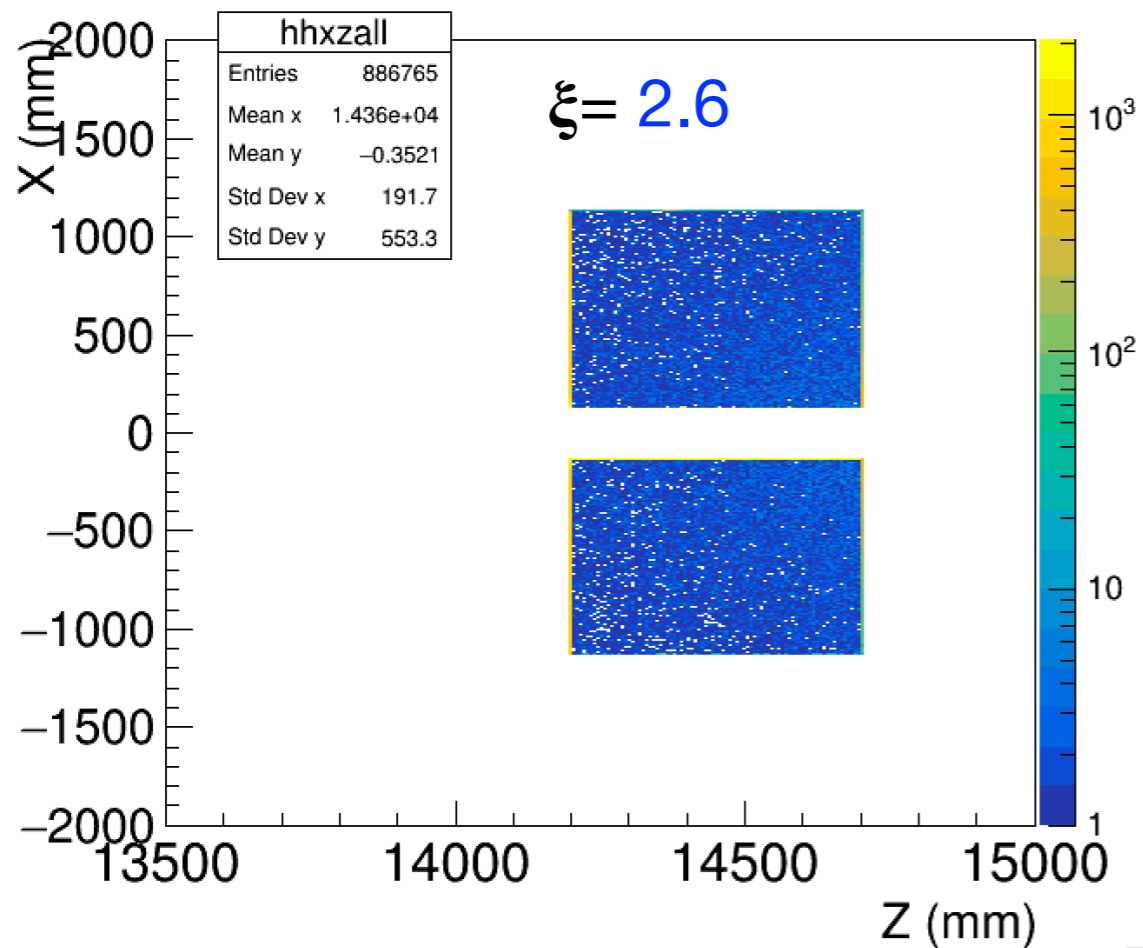
Vertex Z



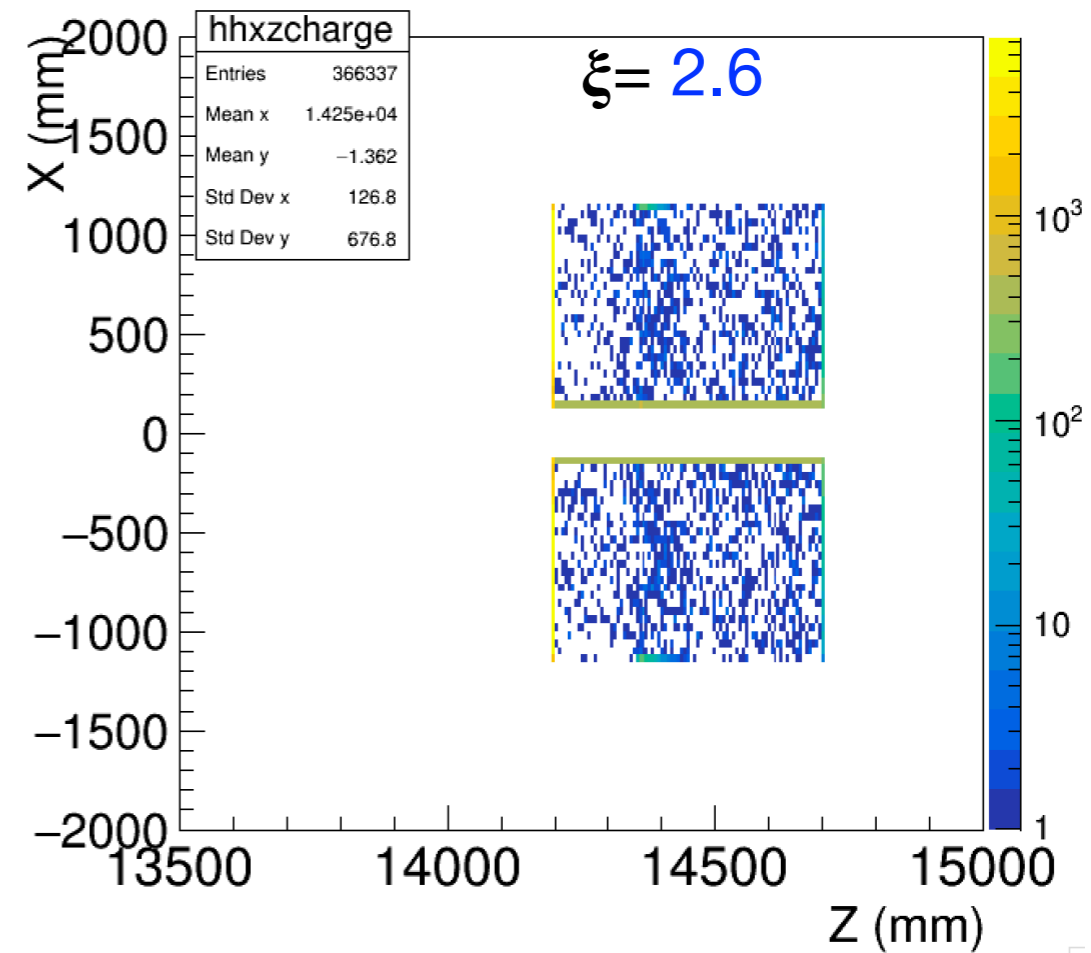
Vertex Z



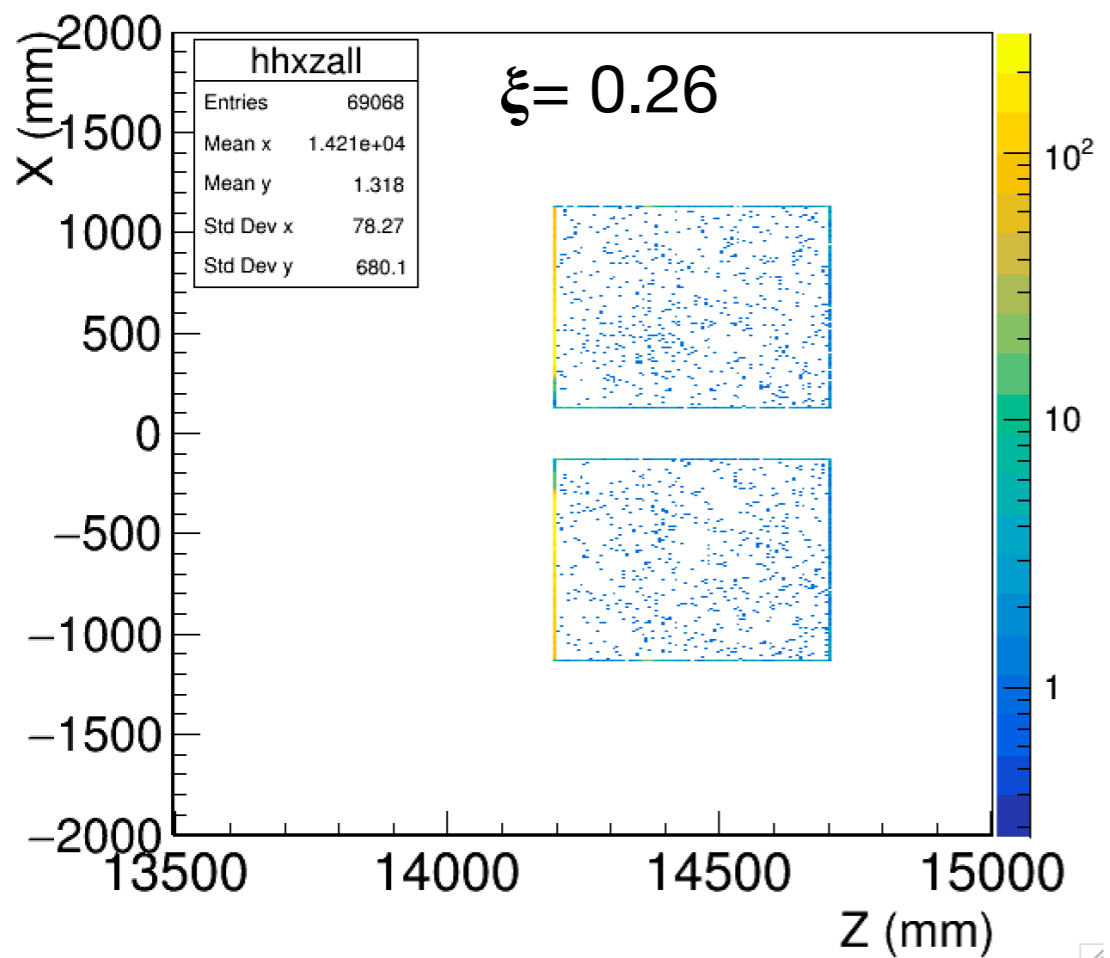
Hits xz distribution all



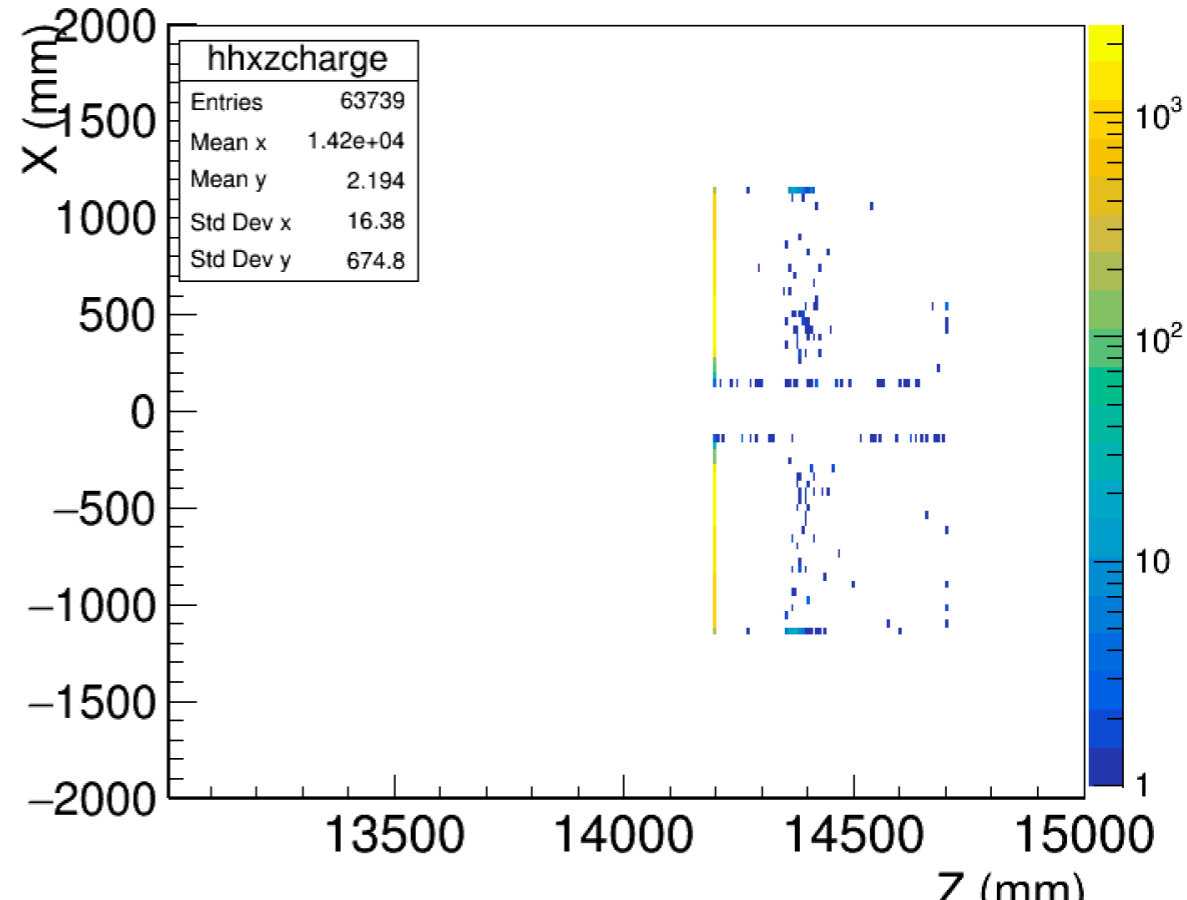
Hits xz distribution e-,e+



Hits xz distribution all

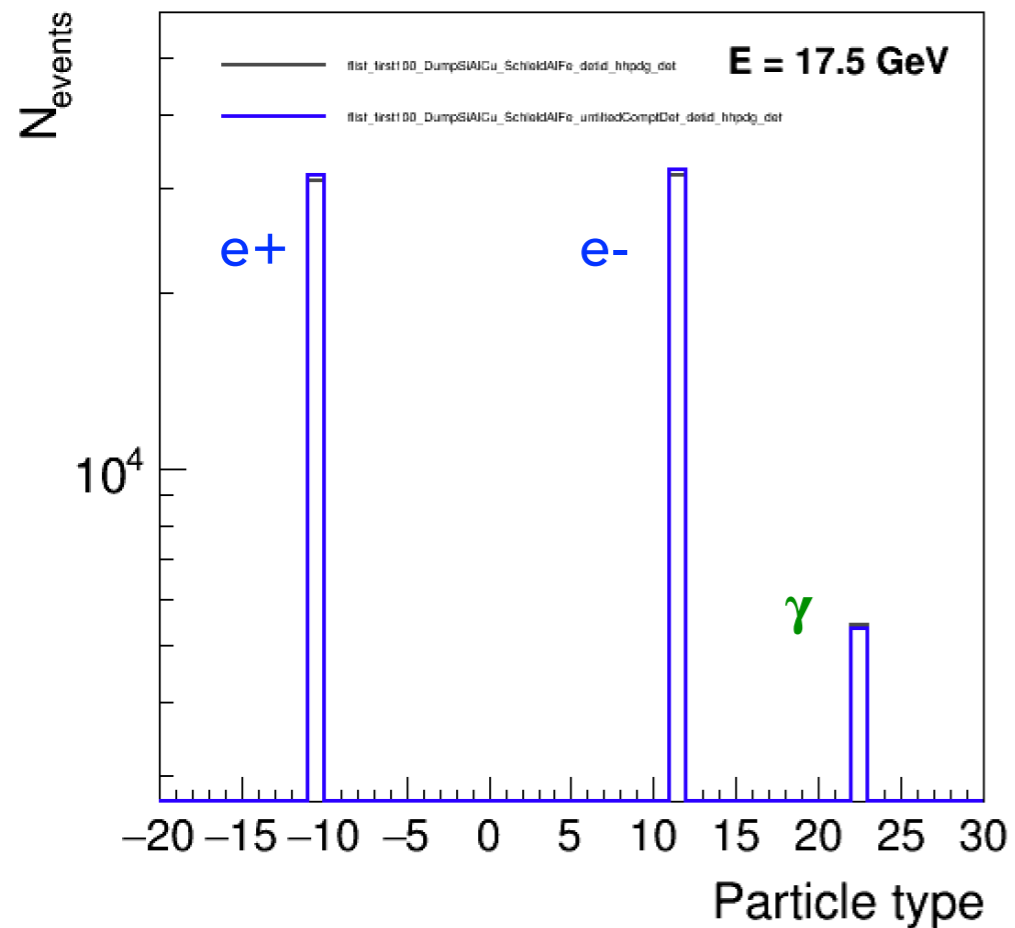
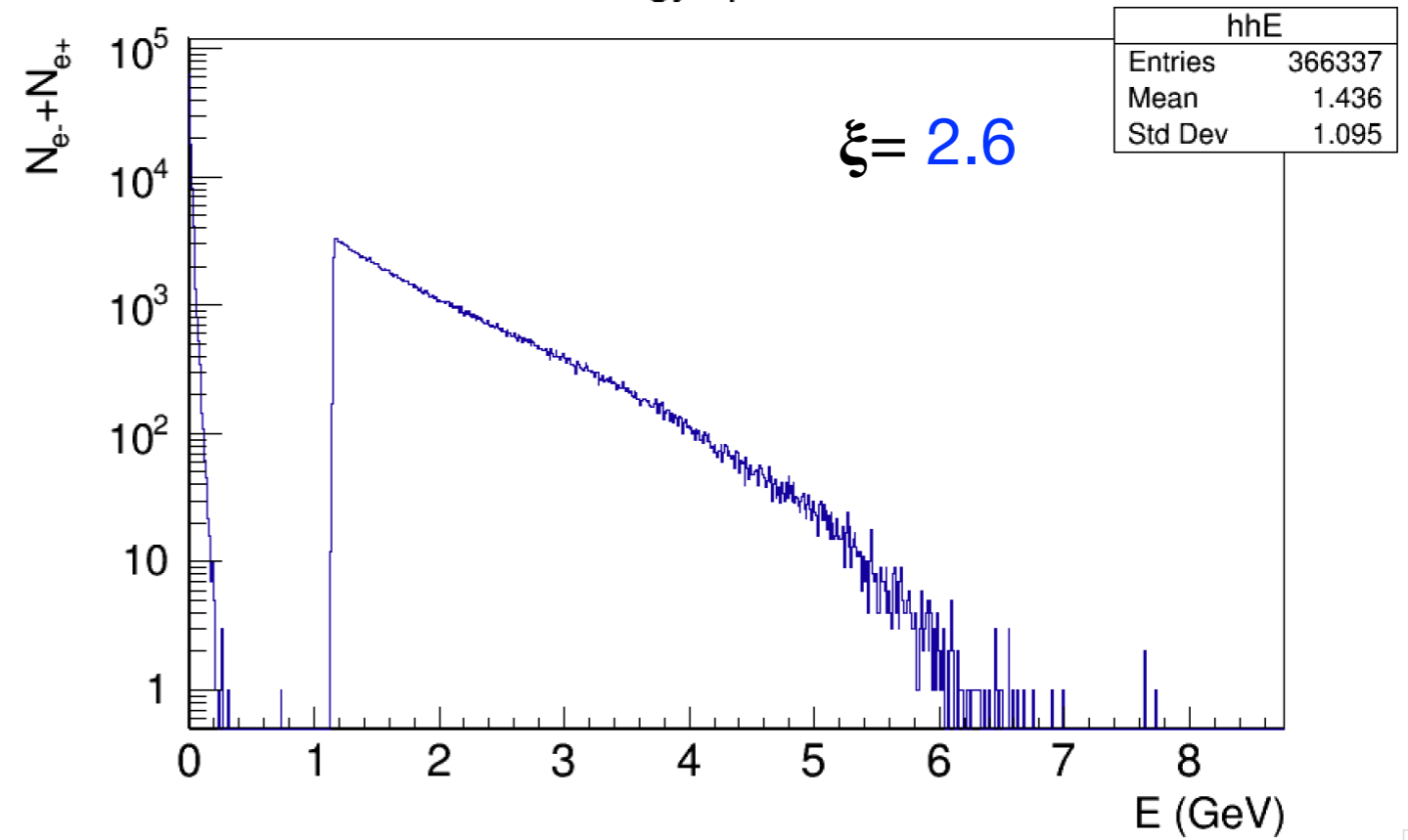
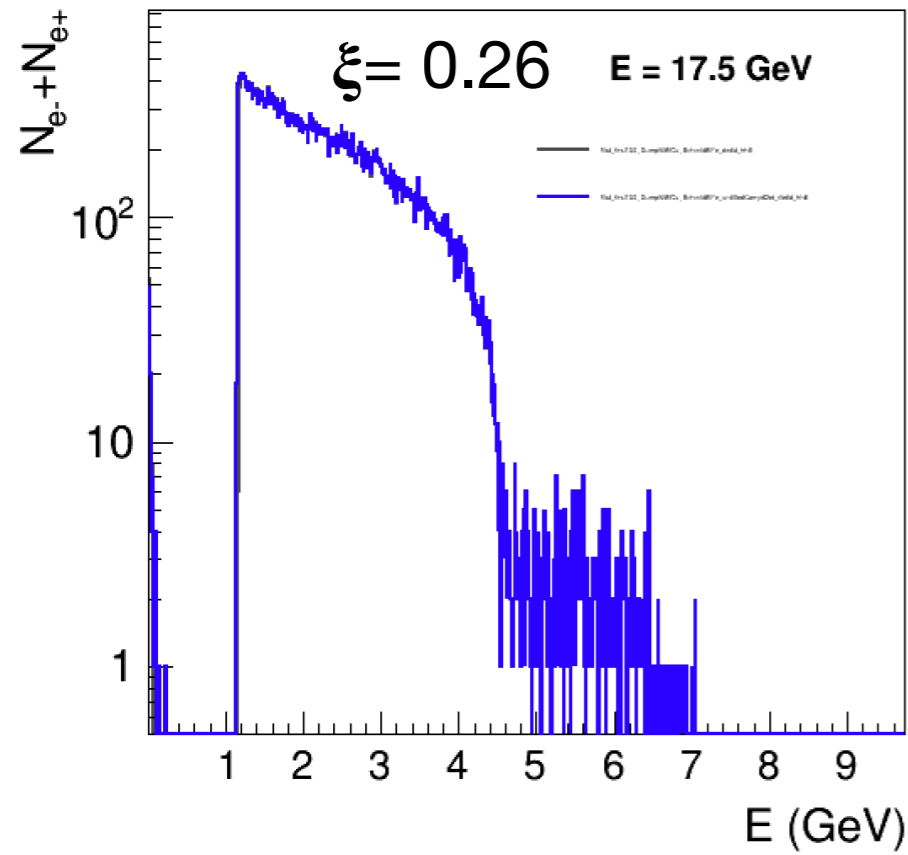


Hits xz distribution e-,e+

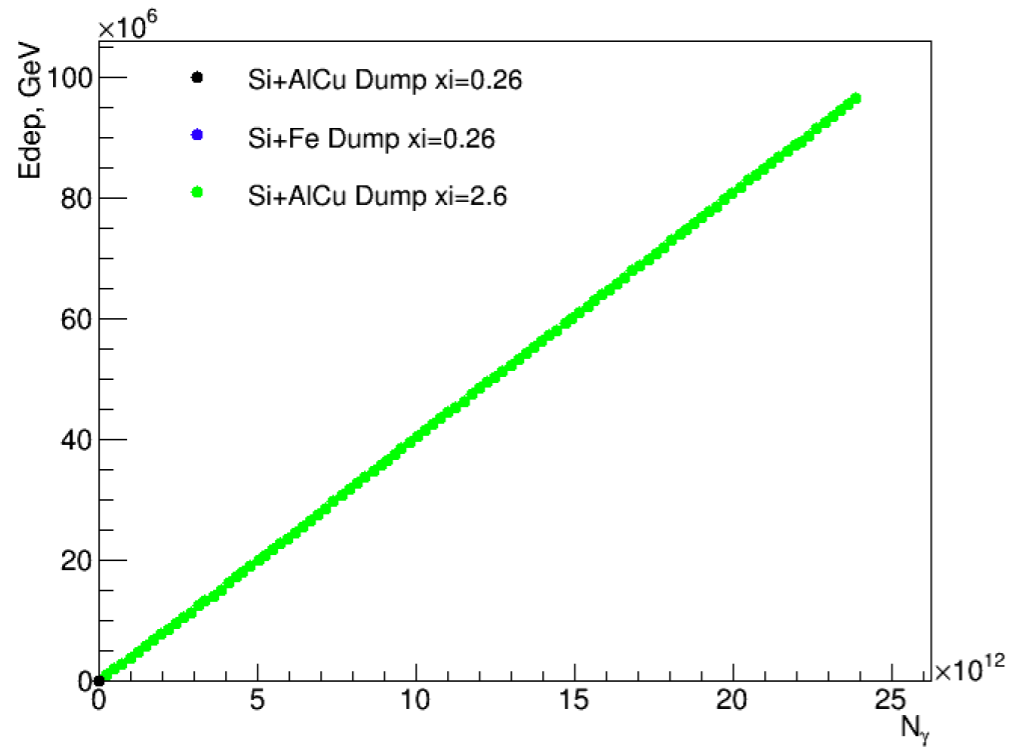


Compton detector: $\xi = 2.6$ vs 0.26

Energy spectrum

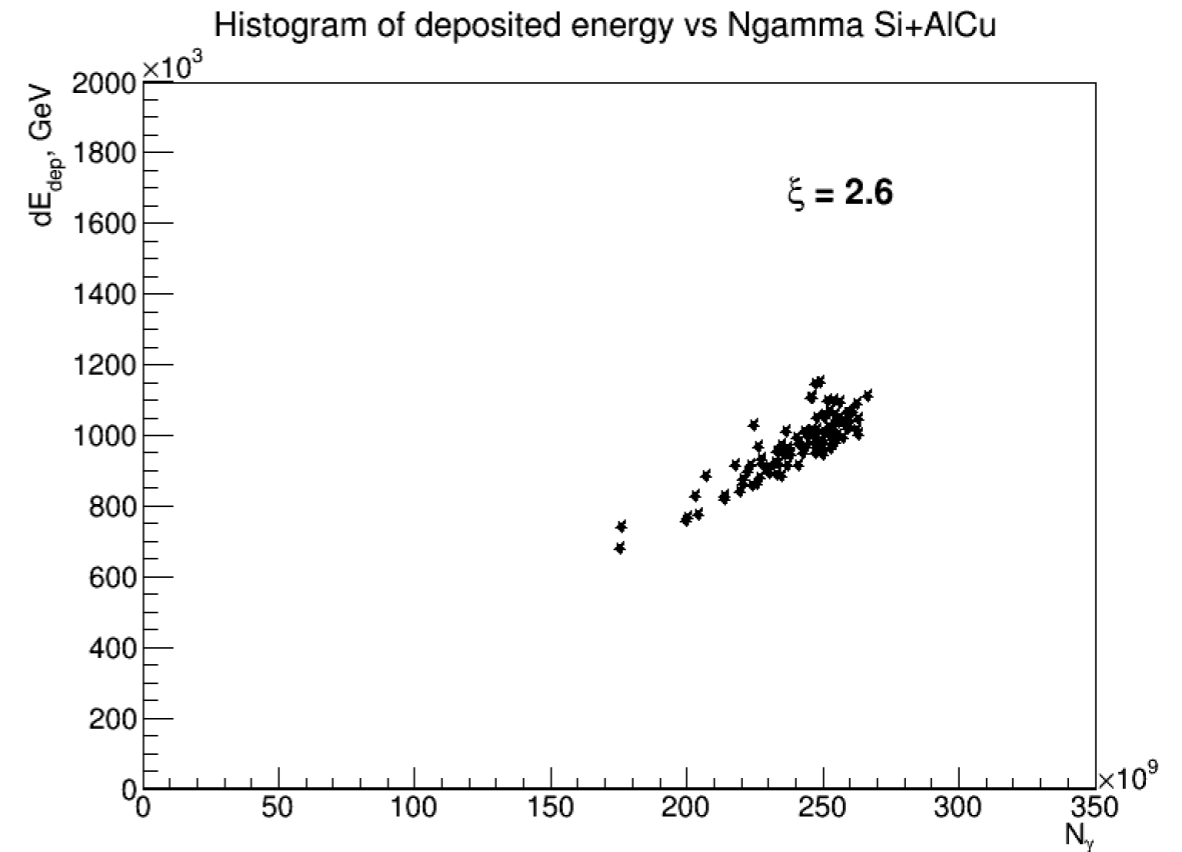
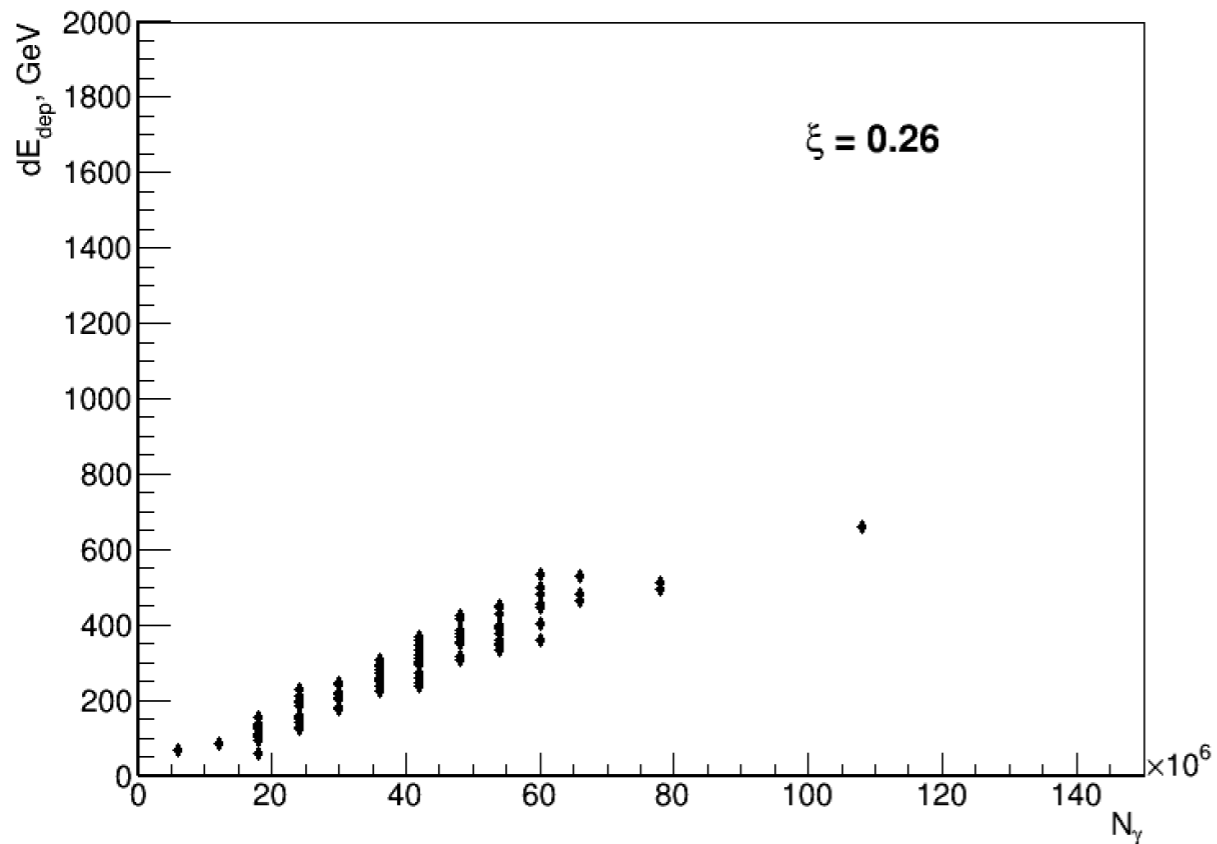


The dependence of deposited energy on number of incoming photons for Si Gamma monitor and AlCu dump for different laser intensities

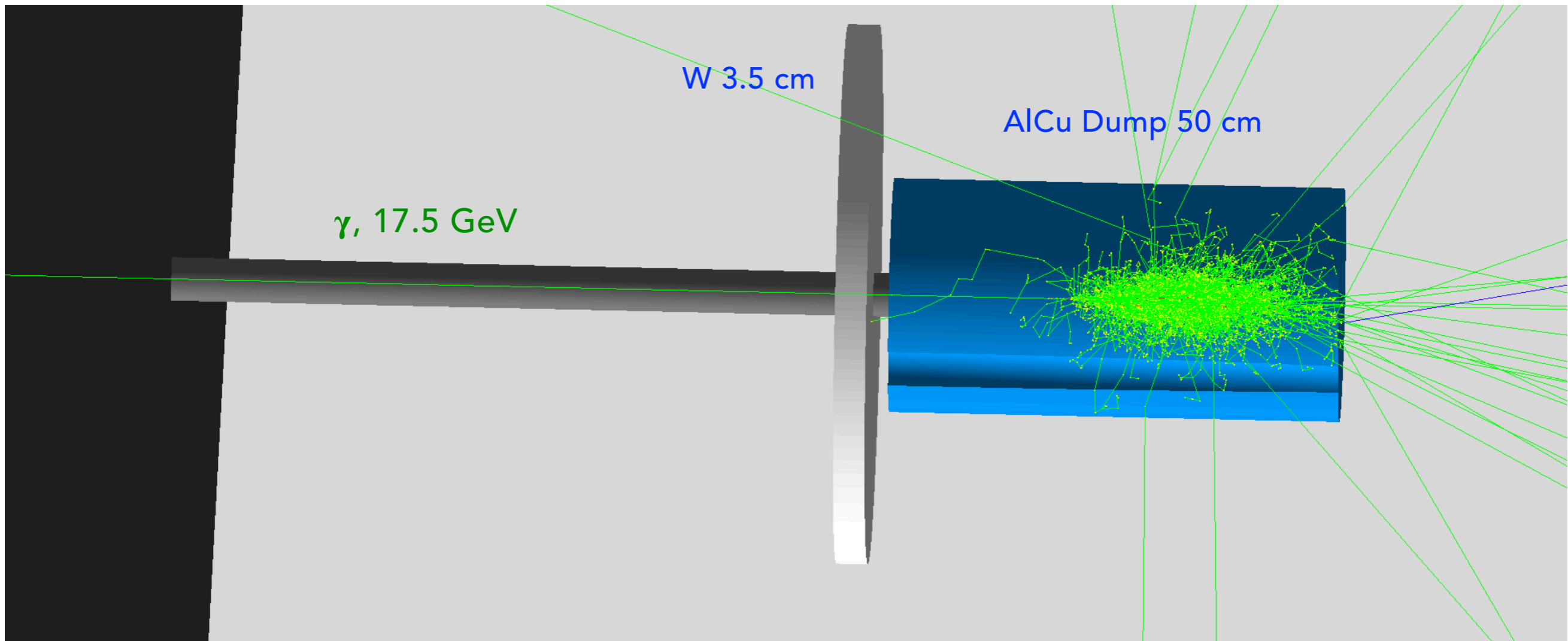


Histogram of deposited energy vs Ngamma Si+AlCu

Energy deposit on Nphotons
Each point is one BX



Tungsten Gamma Monitor in Luxe setup



Compton Photons, 100 BX

$\xi = 2.6$

Target: W foil 10 μm

The deposited energy on number of incoming photons for W Gamma monitor and AlCu dump

